



The EXO-200 Detector

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Frontier Detectors for Frontier Physics

11th Pisa meeting on advanced detectors

24-30 May 2009

Double Beta Decay ...

- Rare nuclear transition between same mass nuclei
 - Energetically allowed for even-even nuclei

$$bb\ 2n : \quad (Z, A) \rightarrow (Z + 2, A) + e_1^- + n_1 + e_2^- + n_2$$

Allowed in SM and
already observed!

$$\left[T_{1/2}^{2n}(0^+ \rightarrow 0^+) \right]^1 = G^{2n}(Q_{bb}, Z) M^{2n\ 2}$$

$$bb\ 0n : \quad (Z, A) \rightarrow (Z + 2, A) + e_1^- + e_2^-$$

Neutrinos are Majorana particles!

$$\Delta L = 2 \quad (Z, A) \rightarrow (Z + 2, A) + e_1^- + e_2^- + c$$

$$n \equiv n \quad m_n \neq 0$$

$$\left[T_{1/2}^{0n}(0^+ \rightarrow 0^+) \right]^1 = G^{0n}(Q_{bb}, Z) M^{0n\ 2} \langle m_{bb} \rangle^2 \quad \langle m_{bb} \rangle^2 = \sum_k m_k U_{ek}^2$$

... and its Role in Neutrino Physics

Double beta decay experiments are part of a massive effort to determine the nature and properties of neutrinos!

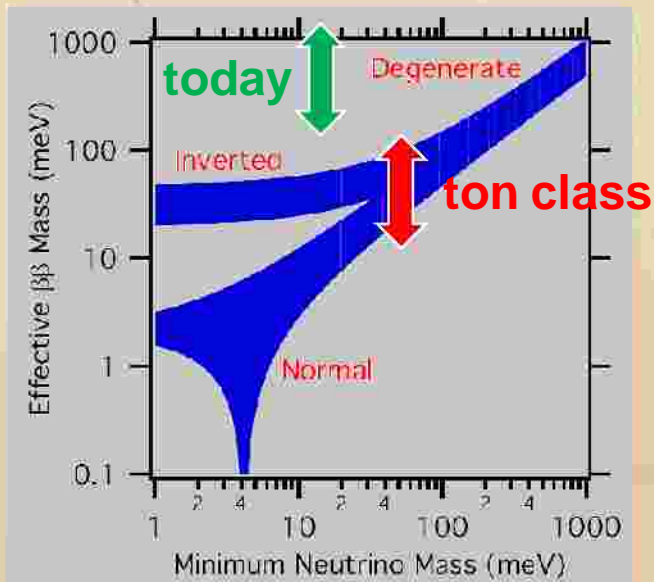
$\langle m_b \rangle < 2eV$ Beta Decay Endpoint $\Delta m_{23}^2 = (2.4^{+0.6}_{-0.5}) \times 10^{-3} eV$ Cosmological Constraints $\Sigma < 2eV$
 $\langle m_b \rangle^2 = \sum_k m_k^2 U_{ek}^2$ $q_{23} \approx 45^\circ$ Neutrino Oscillations Atmospheric and Reactor $\Sigma = \sum_k m_k = 92.5eV \times (\Omega_n h^2)$ Neutrino Oscillations Reactor and Beam CP violation

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{m1} & U_{m2} & U_{m3} \\ U_{t1} & U_{t2} & U_{t3} \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos q_{23} & \sin q_{23} \\ 0 & -\sin q_{23} & \cos q_{23} \end{pmatrix} \begin{pmatrix} \cos q_{13} & 0 & \sin q_{13} e^{-id} \\ 0 & 1 & 0 \\ -\sin q_{13} e^{-id} & 0 & \cos q_{13} \end{pmatrix}$$

$q_{12} \approx 34^\circ$ Neutrino Oscillations Solar and Reactor $q_{13} < 7^\circ$ Neutrinoless Double Beta Decay Only!
 $\Delta m_{12}^2 = (8.0^{+0.4}_{-0.3}) \times 10^{-5} eV$ $\langle m_{bb} \rangle < 0.7eV$

Experimental Requirements

Effective neutrino mass as a function of the smallest neutrino mass for various scenarios




- Large Mass: at least 100 kg of source isotope
 - Scanning the quasi degenerate region
 - Ton scale for the inverted hierarchy region
 - Enrichment helps minimizing volume and improves source purity
- Very Low Background: 1 count per ton per year range
 - Survey, selection and purification of materials and components
 - Cleanroom assembly and detector operation
 - Deep underground installation and muon veto required
- Very Good Energy Resolution: in the 1% range
 - Limits the allowed double beta decay background
 - Increases signal to radioactive background ratio
- And ...
 - Large $Q_{\beta\beta}$ to have the signal out of the region densely populated by radioactive background for natural chains
 - Tagging the daughter isotope would eliminate most radioactive background or event topology and advanced kinematics details

EXO Project & EXO-200 Phase

- EXO project searches for double beta decay using ^{136}Xe
 - Ton scale implementation either as liquid or gas phase TPC
 - Relatively large Q value and straight forward enrichment technique
 - ^{136}Ba daughter tagging either in-situ or in external RF cage

$$\langle m_{bb} \rangle \propto \left(\frac{1}{Nt} \right)^{1/4} \img alt="sad alien" data-bbox="400 498 444 570"/>$$

No Background! $\langle m_{bb} \rangle \propto \sqrt{\frac{1}{Nt}}$ 

- EXO-200 is the first phase using 200 kg of 80% enriched Xe
 - Major R&D effort precursory to the ton-scale experiment
 - Exploration of the quasi-degenerate region with ^{136}Xe
 - Allowed double beta decay never observed in xenon!
 - No tagging but massive progress for radioactive background reduction and energy resolution improvement (easily scalable to future detectors)

EXO Collaboration

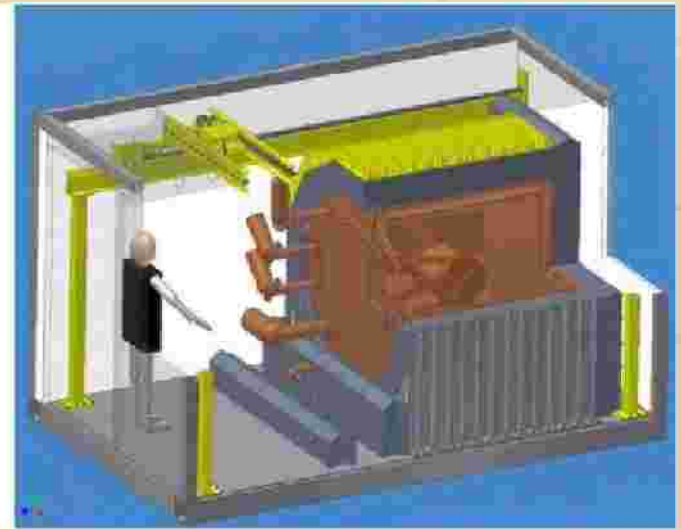
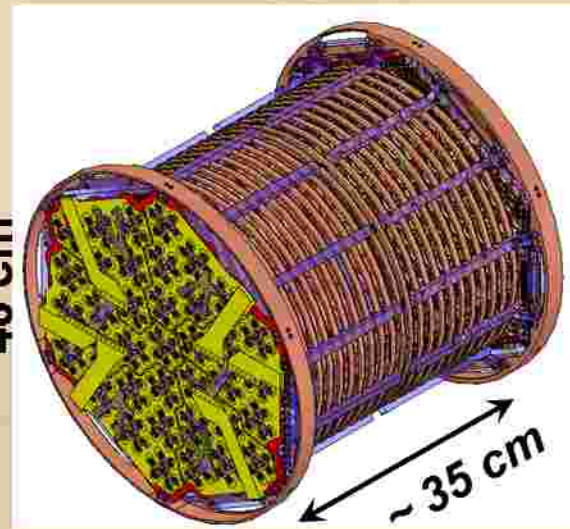
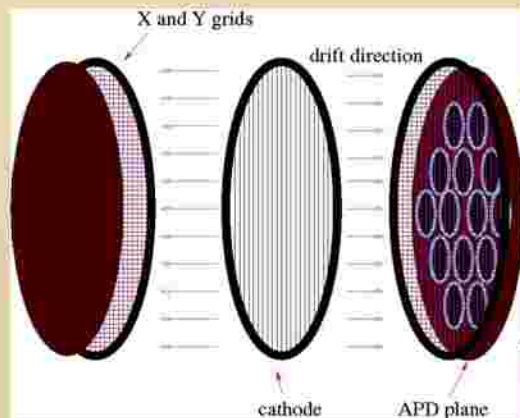
K.Barry, D.Leonard, E.Niner, A.Piepke from **Physics Dept, University of Alabama, Tuscaloosa AL, USA**
P.Vogel from **Physics Dept Caltech, Pasadena CA, USA**
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C. Benitez-Medina, S.Cook, W.Fairbank Jr., K.Hall, B.Mong from **Colorado State University, Fort Collins CO, USA**
M.Moe from **Physics Dept, UC Irvine, Irvine CA, USA**
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K.Kumar, A.Pocar, T. Daniels from **University of Massachusetts, Amherst, USA**
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N.Ackerman, M.Breidenbach, R.Conley, W.Craddock, J.Hodgson, D.McKay, A.Odian, C.Prescott, P.Rowson, K.Skarpaas, J.Wodin,
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J.Anthony, L.Bartoszek, R.DeVoe, P.Fierlinger, B.Flatt, G.Gratta, M.Green, S.Kolkowitz, F.LePort, M.Montero-Diez, R.Neilson, A. Rives,
K.O'Sullivan, K.Twilker from **Physics Dept, Stanford University, Stanford CA, USA**



EXO-200 Detector

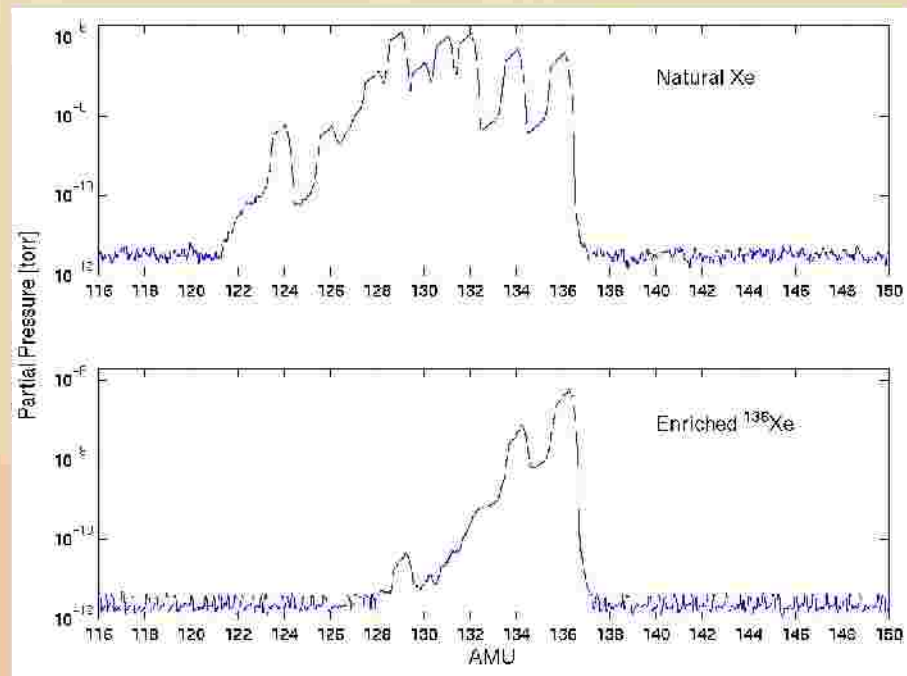
- Liquid xenon TPC with two cylindrical drift volumes
 - Charge collection using 114 by 114 wire planes (at 60° pitch)
 - Scintillation light readout using 37 groups of 7 bare (Large Area Avalanche Photodiodes) at both end caps
- High purity copper cryostat with external refrigeration-based cooling

~ 115 kg fiducial mass



Successful Enrichment Program

200 kg of 80% enriched Xe delivered in 2003!
Used mass-separating centrifuges in Russia
The other isotopes can be returned to provider!



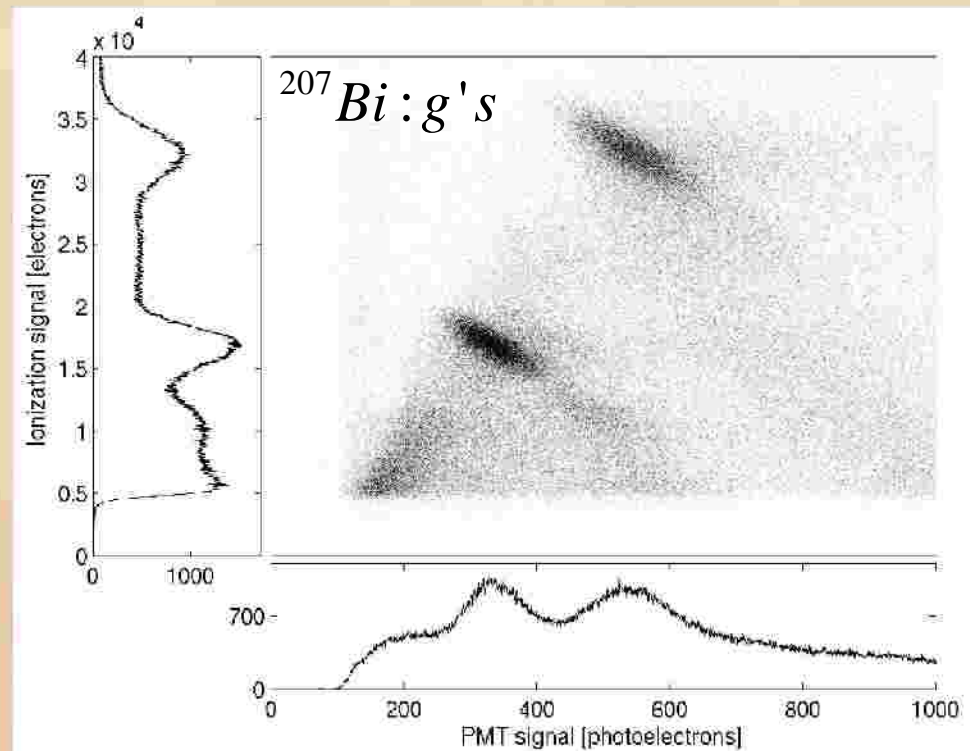
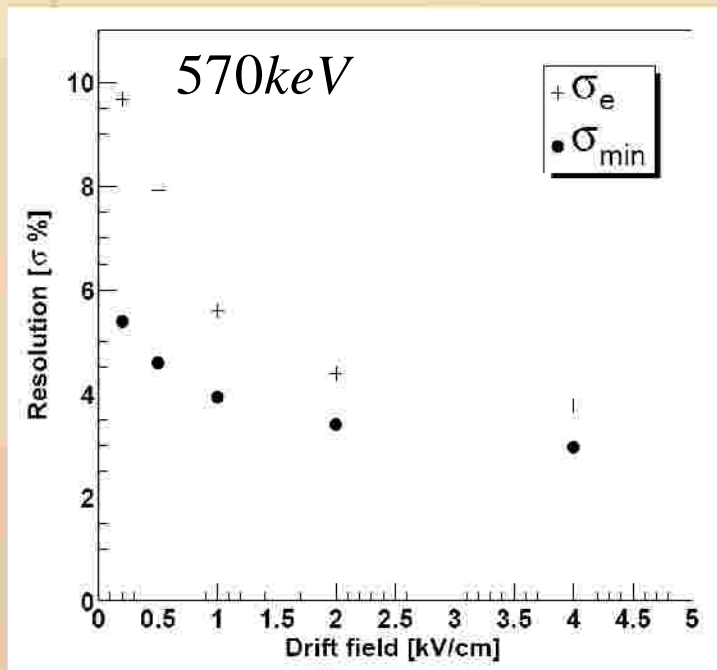
Radio-Purity Survey

- Large effort to determine the residual radioactive contamination of the materials employed for the construction of EXO-200 detector
 - Mass spectrometry (MS)
 - Neutron activation analysis (NAA)
 - Very sensitive but expensive, potential background from main elements
 - Alpha counting (evaluation of the ^{210}Pb concentration in the shield lead)
 - Glow discharge MS (GD-MS), inductively coupled plasma MS (ICP-MS)
 - ICP-MS has better sensitivity when pre-concentration procedures are employed but the samples have to be soluble in acids (preferably HNO_3)
 - Direct gamma counting
 - Large mass samples and long duration exposures are necessary
- Published database of characterized materials
 - ***Nucl. Instr. Meth. A 591, 490-509, 2008***
- Detailed Monte Carlo simulation of expected background

Improving the Energy Resolution

Strong anti-correlation between ionization and scintillation signals in liquid xenon has been observed!

$$\frac{\Delta E}{E} = 1.4\% @ Q_{bb} = 2479keV$$

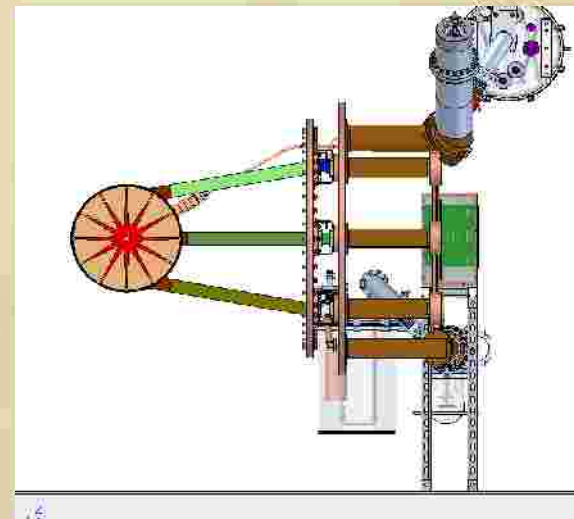
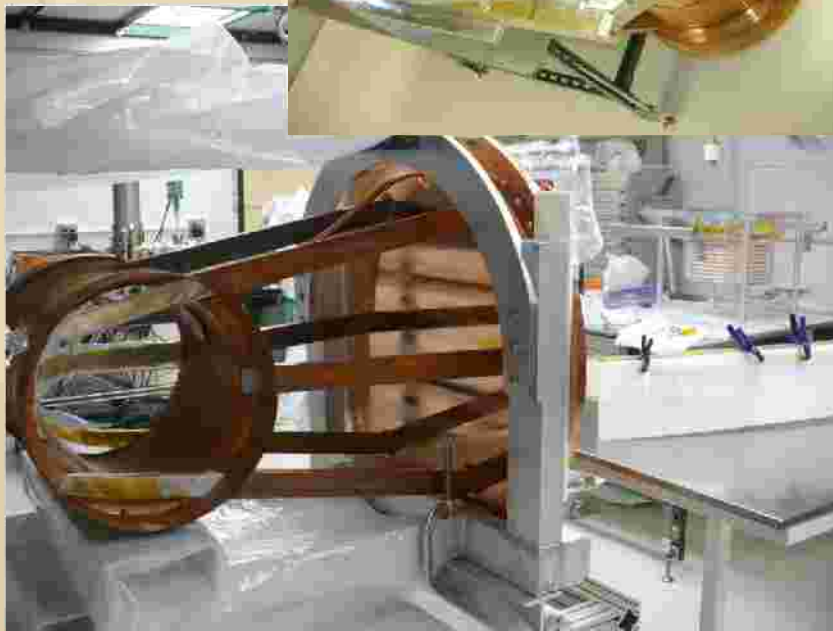




EXO-200 Detector

EXO-200 Chamber

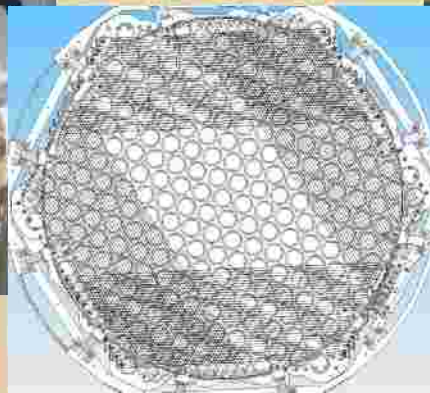
Ultra low radioactivity copper!
Shielded surface transport and storage
Only 1.5 mm thickness to reduce mass
PLC-based real-time pressure control
e-beam welded components
TIG welding for the final assembly



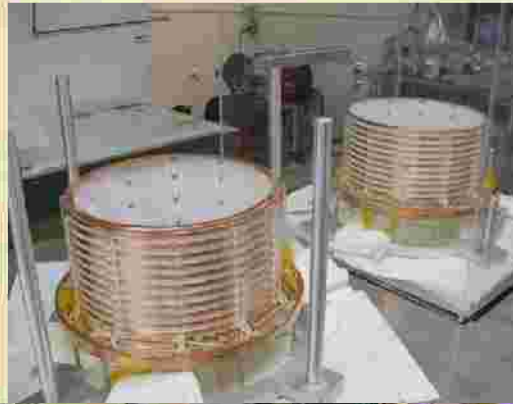
Charge and Light Readout



Photo-etched phosphor-bronze cathode
Induction & charge collection wire grids
259 LAAPD (37 groups of 7) per plane
- 1.6 cm active diameter
- very clean and low mass
- QE > 1 @ 174 nm
- gain 100× to 150× @ ~ 1500V
Radial Teflon UV light reflectors

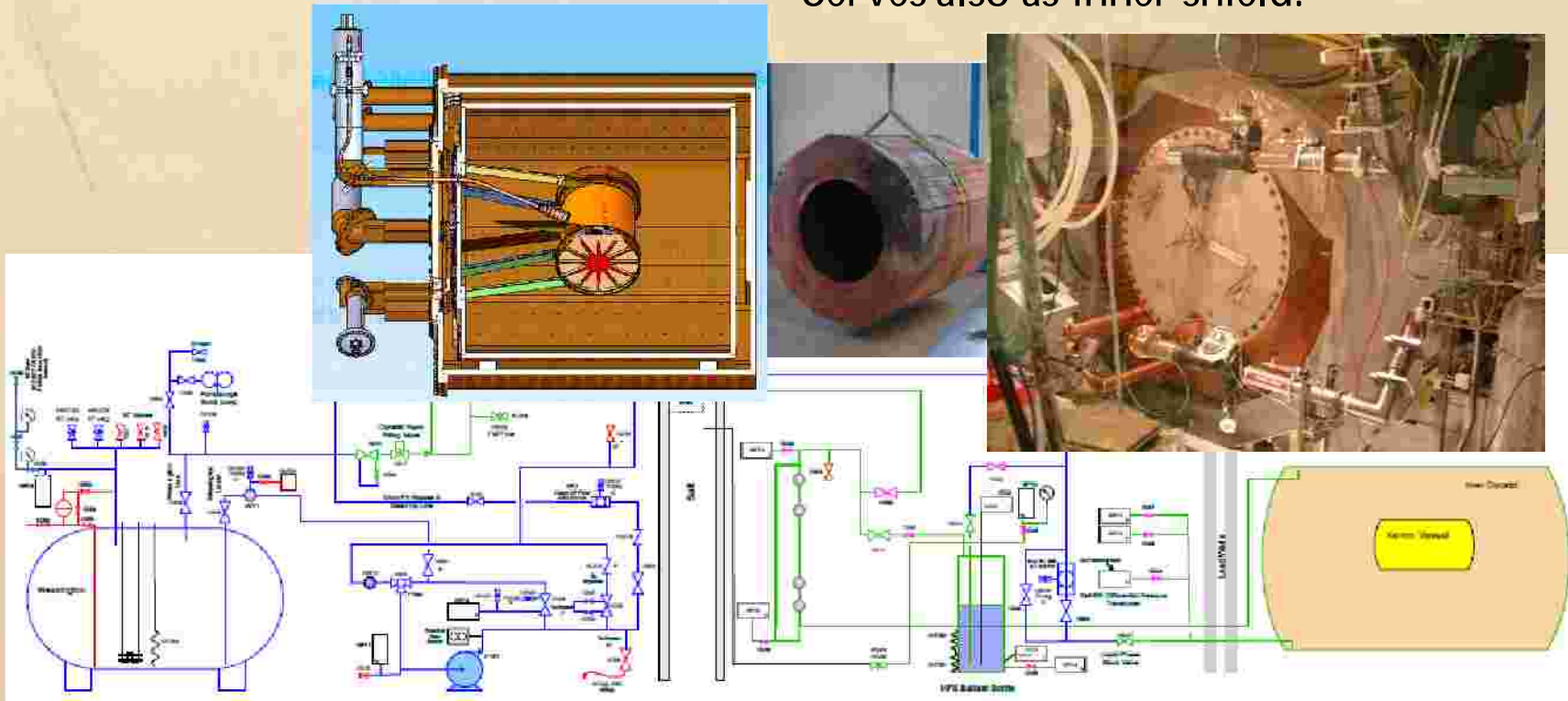


Chamber Assembly



Cryostat and Cooling System

Refrigeration based cooling (3×1500 W PolyCold units)
4.2 tons of high purity heat transfer fluid (3M HFE-7000)
Serves also as inner shield!



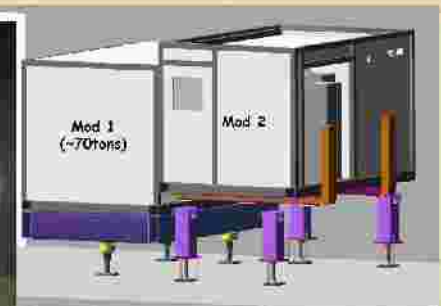


WIPP Installation

Experimental Area

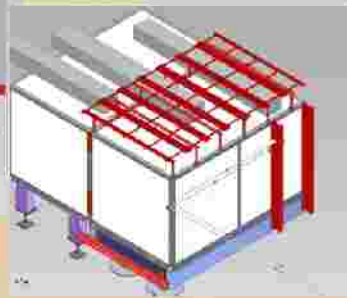
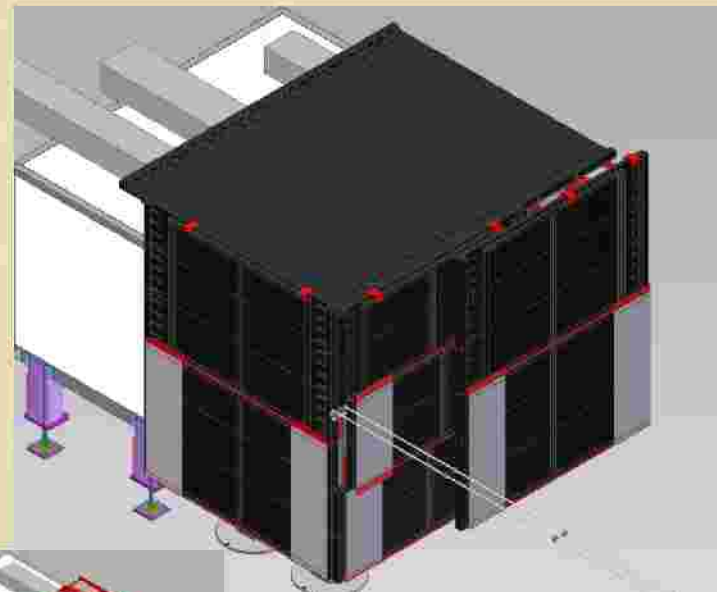
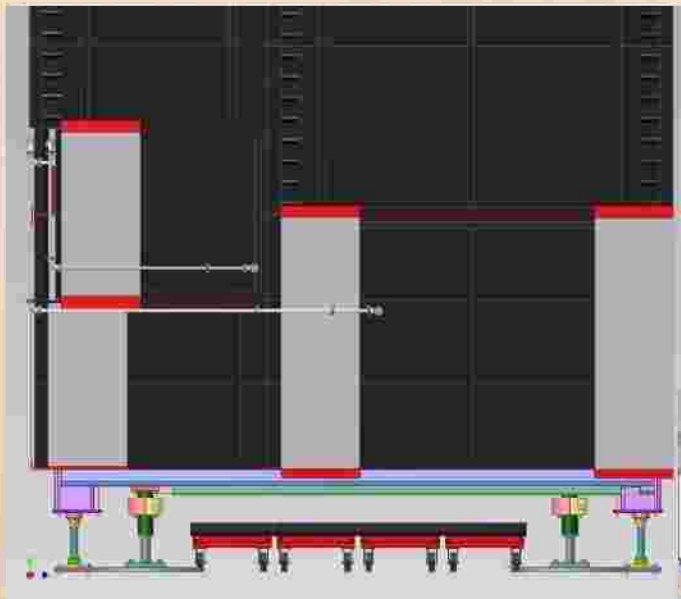
Waste Isolation Pilot Plant, Carlsbad, New Mexico
~ 1600 m.w.e. (muon flux reduction by ~ 10 \times)

Large and wide remote experimental area available!



Muon Veto

20× muon induced background reduction (99,7% detection efficiency)



Extensive Monte Carlo simulations to optimize the veto configuration!

Expected Performance

- Very low radioactive background
 - Careful selection of materials, optimized custom design
 - Manufacturing, handling and installation in cleanrooms

- Very good energy resolution

$$\frac{S}{B} = \frac{m_e}{7Q_{bb}} \left(\frac{E}{\Delta E} \right)^6 \frac{T_{1/2}^{2n}}{T_{1/2}^{0n}}$$

Chamber underground installation in august 2009
 Physics runs starting in 2010, 2 years run time!

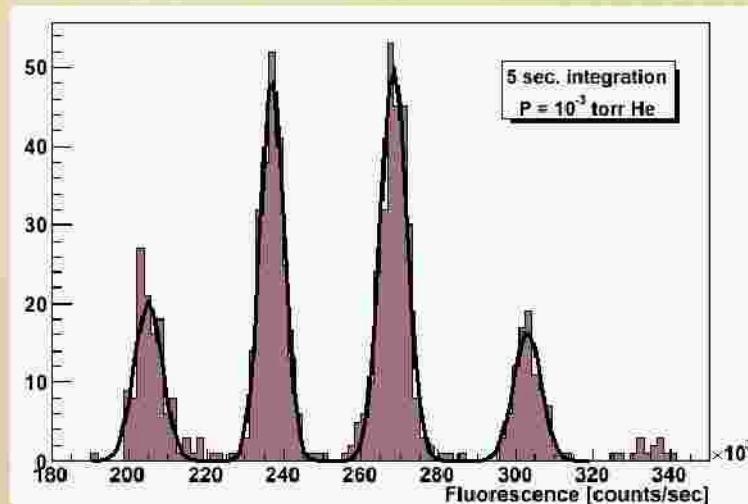
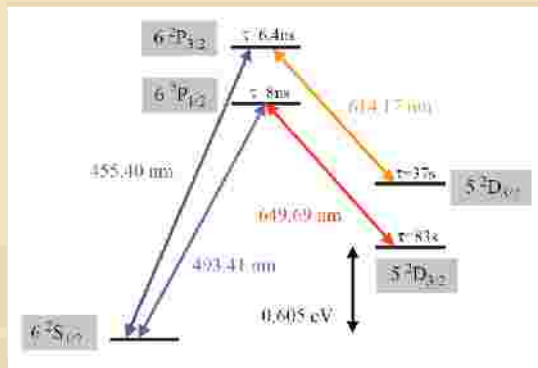
$$T_{1/2}^{2n} > 1.2 \times 10^{24} \text{ y @ 90\% C.L.}$$

| Case | Mass (ton) | Eff. (%) | Run Time (yr) | s_E/E @ 2.5 MeV (%) | Radioactive Background (events) | $T_{1/2}^{0?}$ (yr, 90% CL) | Majorana mass (meV) | |
|---------|---------------|-------------|---------------------|-----------------------------|---------------------------------------|--------------------------------|------------------------|------------------|
| | | | | | | | QRPA ¹ | NSM ² |
| EXO-200 | 0.2 | 70 | 2 | 1.6 | 40 | $6.4 * 10^{25}$ | 133 | 186 |

1) Rodin et. al., Nucl. Phys. A 793 (2007) 213
 2) Caurier et. al., arXiv:0709.2137v1



Ba⁺ Tagging



RF cage with low pressure buffer gas

- Ba⁺⁺ ? Ba⁺ conversion expected
 - Ionization potentials:
 - Xe⁺ = 12.13 eV vs. Ba⁺ = 5.21 eV
 - Xe⁺⁺ = 21.21 eV vs. Ba⁺⁺ = 10.00 eV
 - Solid Xe band gap (*Phys. Rev. B10 4464 1974*)
 - $E_G = 9.22 \pm 0.01 \text{ eV}$
 - "Liquid Xe ionization potential" close to E_G (*J. Phys. C: Solid State Phys. Vol. 7 1974*)
 - 9.28 to 9.49 eV range
 - Use of additives for gas based detectors

Conclusion

- EXO-200 detector soon operational!
- The largest neutrino-less double beta decay detector!
- Successful large scale xenon enrichment proven!

