## GERDA/LArGe

#### Detector calibration using <sup>228</sup>Th

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# GERDA designed for the detection of 0vββ



- Enriched <sup>76</sup>Ge crystal based detector array
- Contained inside a copper-lined steel cryostat
- Filled with LAr for cooling/shielding
- Surrounded by a copper shroud
- First phase data taking completed in 2013
- Preparing for Phase 2!



## LArGe

- Being used to prepare for GERDA phase 2 (designed for the detection of 0vββ)
- Similar to GERDA
  - 1 Natural BeGe detector
  - Steel cryostat w/1m<sup>3</sup> of LAr
  - Pb and Cu shielding instead of water
  - 9 PMT's in the LAr Cryostat to detect scintillation light - used for cooling the Ge and anticoincidence veto.



#### LArGe detector on Oct. 1, 2014



Primary Goal: To calibrate the LArGe detector's energy spectrum vs fADC channel response using <sup>228</sup>Th source

#### **Calibration objectives**

- To characterize the detector's resolution at the ROI for 0vββ (2039 keV)
- To examine the detector resolution at all energies
- To ease peak identification by lining up the fADC channel with the peak energy.



Source tube on the LArGe Detector









Page -1-

10/3/14

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#### Smoothed uncalibrated energy spectrum obtained from <sup>228</sup>Th source **10**<sup>5</sup> **Compton Edge** Double Escape 1592 keV 2614 keV (208TI) **10**<sup>4</sup> 510 keV (<sup>208</sup>TI) 583 keV (<sup>208</sup>TI) Single Escape 2103 keV 727 keV (212Bi) 10<sup>3</sup> ⊦ 860 keV (208TI) 1620 keV (212Bi) Mostly due to 511 keV 10<sup>2</sup> annihilation from pair production. SE and DE peaks are caused by 2614 keV 10 pair prodution/positron annihilation gammas escaping the detector without interacting. 1 6000 1000 2000 3000 4000 5000 0 **fADC** Channels



- Need to generate a function to shift the uncalibrated spectrum. Then, apply the function to all data points.
  - Perform a curve fit on gamma peaks using ROOT to find the mean channel value for the peak
  - Plot the mean channel values w.r.t. the nominal energy of the gamma peak given in literature
    - Perform a curve fit to generate a function we can use to translate the spectrum







- Most peaks fitted with a gaussian+pol1
- 2614 peak required const • +gauss+step function+tail
- Use peak fits to find: •
  - mean channel value •
  - FWHM (via  $\sigma$ ) •



Linear fit of mean channel value vs nominal energy of the peak (from literature)





#### Deviation of mean channel number curve vs fit





Uncalibrated Spectrum









Fit of FWHM points (using p0+p1\*sqrt(E))



 Using these parameters and the relation FWHM = 2.355σ, we can calculate the resolution we could expect at any energy level.





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Sanford Underground Research Facility



University of South Dakota





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