

Direct reactions with SPES beams:
Nuclear magicity at $Z \sim 50$ and $N \sim 82$
n-capture cross section via surrogate method

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*SPES workshop, LNL - Italy
May 26÷28, 2014*



Outline

1 Around the magic ^{132}Sn

2 Nuclear astrophysics

- n-capture reaction



Collaboration

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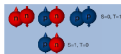


Nuclear Physics

high angular momentum

deformed nuclei

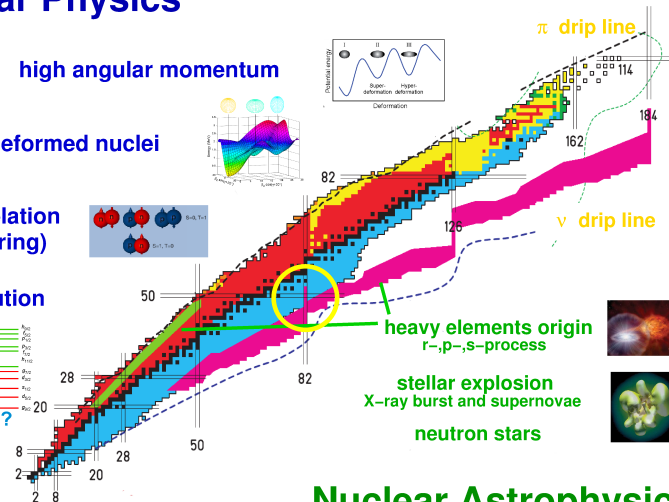
correlation
(pairing)



shell evolution



→ towards neutron-rich nuclei



heavy elements origin
r-, p-, s-process

stellar explosion
X-ray burst and supernovae
neutron stars



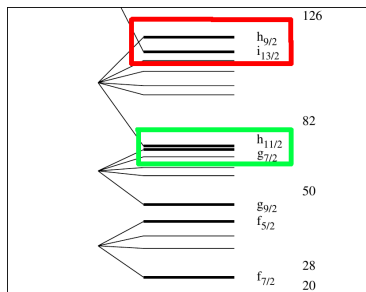
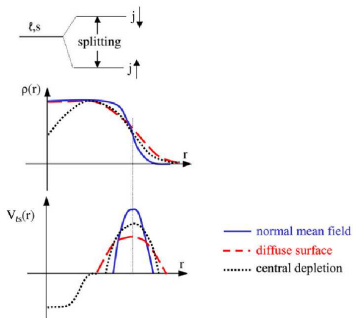
Nuclear Astrophysics



Around doubly-magic ^{132}Sn

SO term in exotic matter

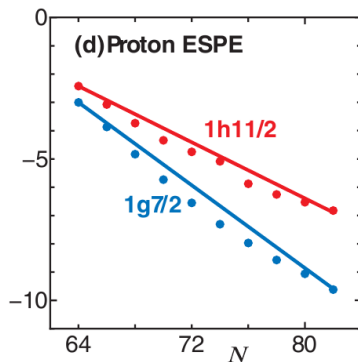
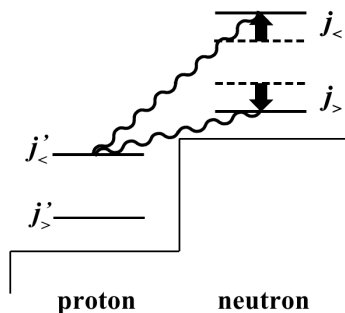
- SO-interaction scales with the derivative of the nucleon densities



Shift of the proton single-particle energies

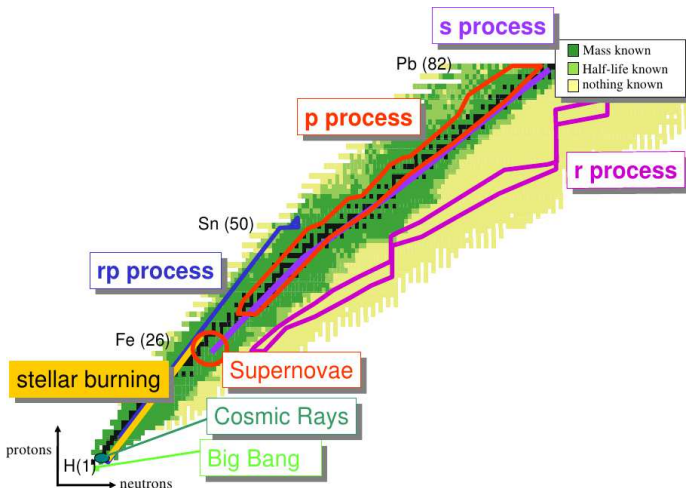
monopole part of the nuclear force

Attractive when spins of nucleons are antiparallel to their orbital angular momenta



Nucleosynthesis processes

Implications on nuclear astrophysics

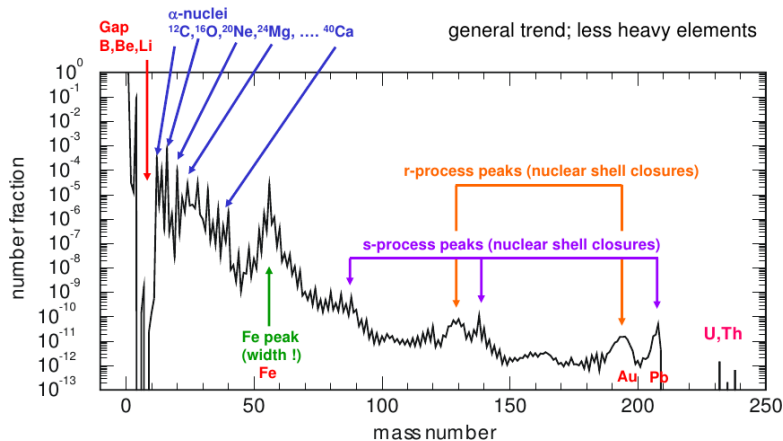


Burbidge, Burbidge, Fowler, Hoyle, Rev. Mod. Phys. 29 (1957) 547



Solar abundances

Implications on nuclear astrophysics



determined from solar and stellar spectra and from meteorites

■ N. Grevesse and A.J. Sauval, Space Science Reviews 85 (1998) 161



Nuclei around doubly closed shells: why?

In the independent particle SM the properties of a odd nucleus are determined by the odd unpaired nucleon.

This is especially true for magic numbers that corresponds to significant gap in SPE.

Test of the model and its prediction capacity:

- SPE
- TBME

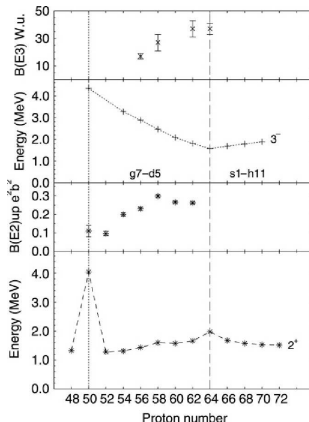
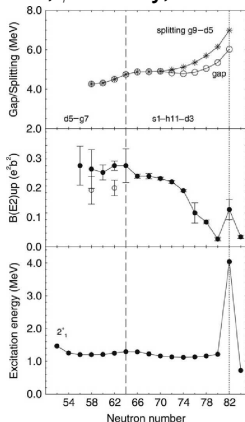
Around ^{132}Sn

- Doubly-magic nucleus (with extreme N/Z)
- Proximity of the r -process path
- Proximity of the continuum states
- Three-body forces, appearance of new (sub-)shells



Experimental background

■ SF, β decay, Coulomb excitation ... and transfer

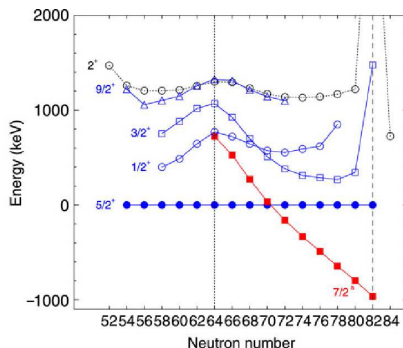
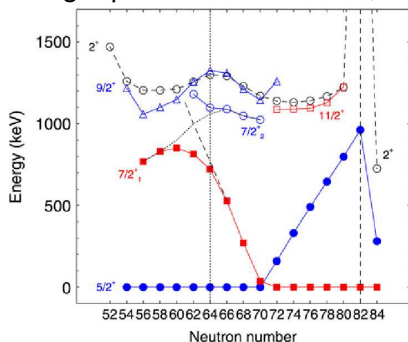


- ^{132}Sn behaves like a doubly-magic nucleus ($E2+$, $BE2$, $S2N$, gap, etc.)
- O.Sorlin et al., Prog.Part.Nucl.Phys. 61 (2008) 602 and reference therein
- RIKEN EURICA β -decay campaign



Z=51 proton orbits

- 1 core coupling ($2^+, 3^-$) : $9/2^+, 11/2^+, 3/2^+, 1/2^+$
- 2 mixing between different configurations: $7/2^+$
- 3 single-particle states: $7/2^+, 11/2^+$?

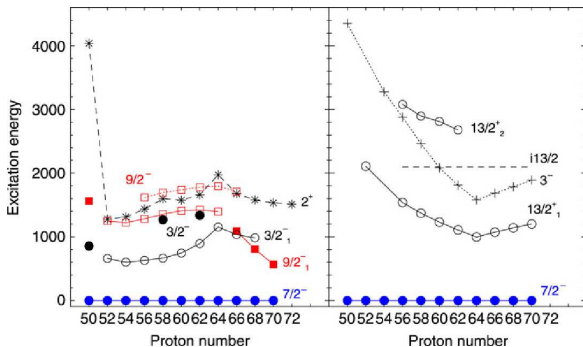


■ O.Sorlin et al., Prog.Part.Nucl.Phys. 61 (2008) 602 and reference therein



N=83 neutron orbits

- 1 core coupling ($2^+, 3^-$) with neutron $f_{7/2}$
- 2 neutron single particle states
- 3 mixing



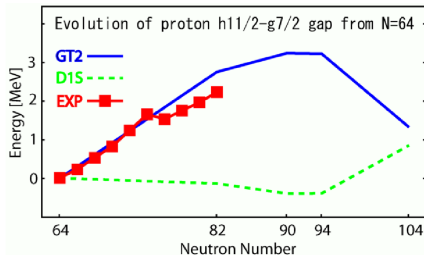
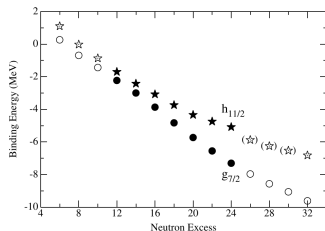
■ O.Sorlin et al., Prog.Part.Nucl.Phys. 61 (2008) 602 and reference therein



Transfer beyond Z=50



- Spectroscopic factors deduced for $l=4,5$ using DWBA model
- $7/2$ and $11/2$ states show a single-particle character
- Tensor part of interaction is fundamental to reproduce the exp trend



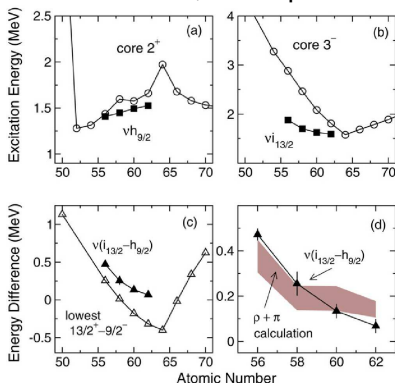
- J.P.Schiffer et al., Phys.Rev.Let.92, (2004) 162501
- T.Otsuka et al., Nucl.Phys. A 805 (2008) 127c



Transfer beyond N=82

$A X(^4\text{He}, ^3\text{He})^{A+1} X$ for $^{138}\text{Ba}, ^{140}\text{Ce}, ^{142}\text{Nd}, ^{144}\text{Sm}$

- Spectroscopic factors deduced for $l=5,6$ using DWBA model
- Energy difference fairly similar, no swap visible in the experiment
- Agreement LSSM calculations, discrepancies with MF.

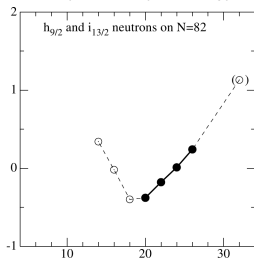
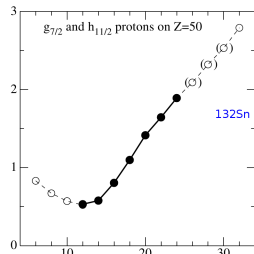


- B.P.Kay et al., Phys.Lett.B658(2008) 216
- T.Otsuka et al., Phys.Rev.Lett.95 (2005) 232502



SO evolution with neutron excess

Interpretation



- density (rms and diffuseness) are most similar
- proton and neutron Fermi surface are equal

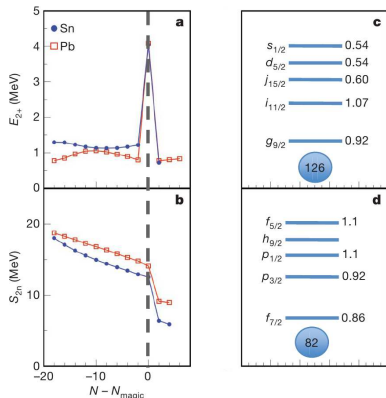
■ J.P.Schiffer et al., Phys.Rev.Let.92, (2004) 162501



LETTERS

The magic nature of ^{132}Sn explored through the single-particle states of ^{133}Sn

K. L. Jones^{1,2}, A. S. Adekola¹, D. W. Bardayan¹, J. C. Blackmon⁴, K. Y. Chae¹, K. A. Chipps³, J. A. Cizewski², L. Erikson⁵, C. Harlin⁶, R. Hatarik², R. Kapler¹, R. L. Kozub⁷, J. F. Liang⁴, R. Livesay³, Z. Ma¹, B. H. Moazen¹, C. D. Nesaraja¹, F. M. Nunes⁸, S. D. Pain², N. P. Patterson⁶, D. Shapira¹, J. F. Shriner Jr², M. S. Smith⁴, T. P. Swan^{2,6} & J. S. Thomas⁹



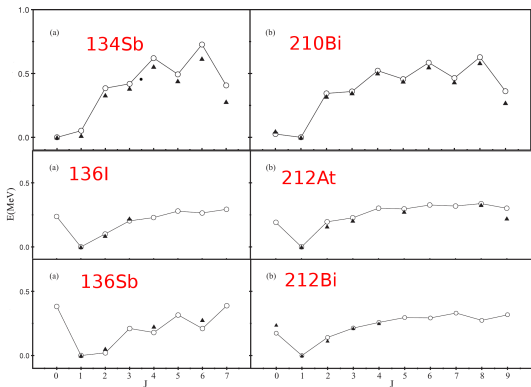
- single particle strength (S) concentrated in one state
- large discontinuities in numerous observables



LSSM calculations

Theoretical predictions outside $N=82$, compared with $N=126$

- Few valence particle nuclei above the doubly-closed magic 132Sn core
- realistic interaction
- good results for isotopes of Sn, Sb, Te, I, Xe, Cs with different interactions for $134 \leq A \leq 138$ and $50 \leq Z \leq 56$



- L. Coraggio et al., Phys.Rev.C(R) 80 (2009) 021305

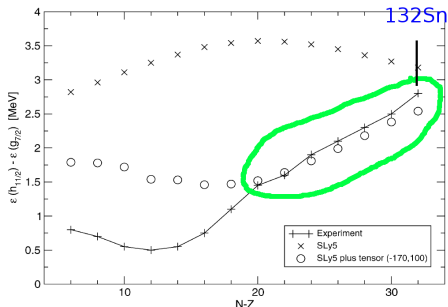


HFB calculations

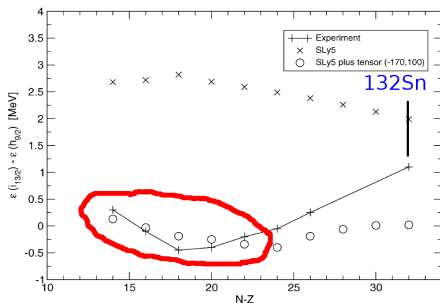
Theoretical predictions outside Z=50 and N=82

■ tensor term in the effective Skyrme interaction

Protons on Z=50 core



Neutrons on N=82 core



■ G.Colò et al., Phys.Lett.B646 (2007) 227



What we can probe? and how?

n states (d,p) for neutron pickup, (d,t) for neutron stripping

p states ($^3\text{He},d$) for proton pickup, (t, α) or (d, ^3He) for proton pickup

pairing (t,p) or (p,t) for n-n pairing

pairing (p, ^3He) or (d, α) for n-p pairing

target

- CH_2/CD_2 loaded carbon
- cryogenic gaseous
- jet



Proposed reactions at SPES

Expected SPES-beam intensity: $10^{5\div 8}$ pps (10^5 required on target for transfer reactions at ISOLDE)

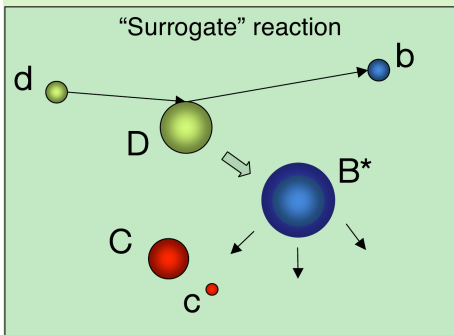
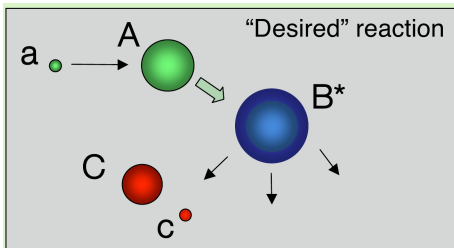
Systematic measurements in the region

- **(d,p)** : ^{133}Sn , ^{134}Sn , ^{133}Sb , ^{131}In
- **(d,t)** : ^{131}Sn , ^{134}Sn , ^{131}In
- **(d, ^3He)** : ^{131}Sn , ^{133}Sn , ^{131}In

complemented by Coulomb excitation measurements?



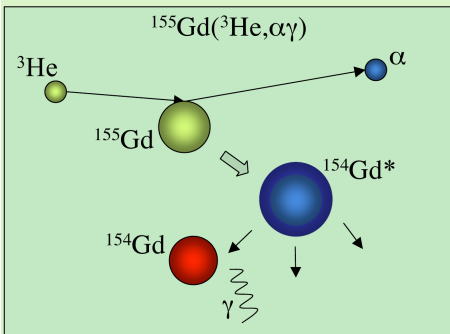
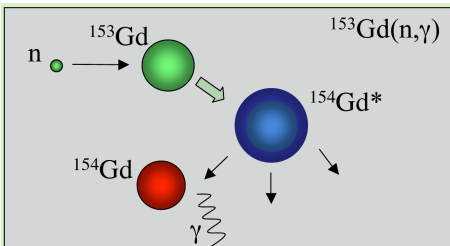
Indirect Determination of Cross Sections



The Surrogate Nuclear Reactions approach is an indirect method for determining XS of CN reactions difficult to measure directly.



(n,γ) cross section

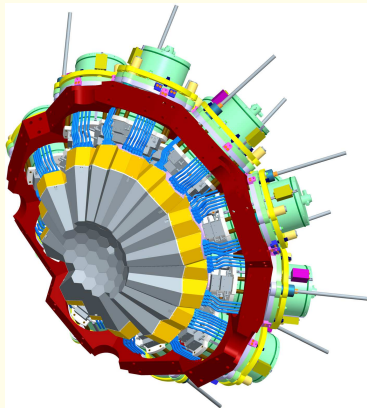


Various direct-reaction mechanisms can be employed to create the compound nucleus of interest.

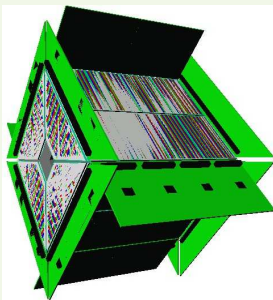


Detector Setup

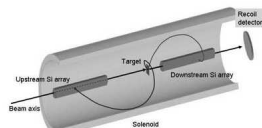
AGATA, GALILEO



TRACE, SPIDER



Solenoid



Summary and conclusions

- SPES RIB:
 10^{5-8} pps
- $N \sim 82$, $Z \sim 51$:
Spectroscopic factor in Sb, Sn, In
- Detection Setup:
AGATA, TRACE, SOLE, (SPIDER, DANTE)

