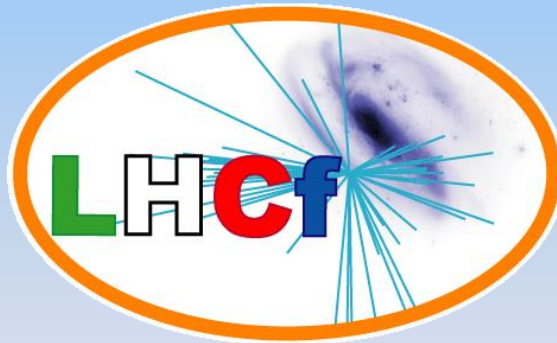


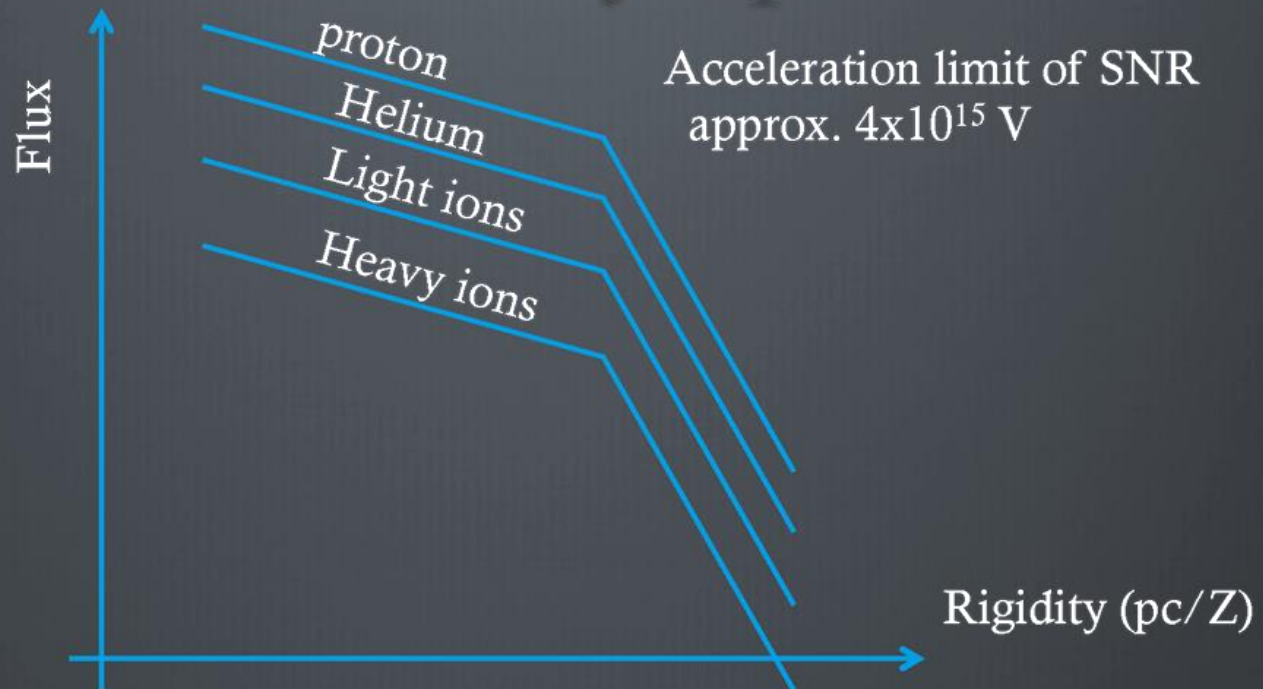
Cosmic-ray physics at LHC: the LHCf experiment



Massimo Bongi, for the LHCf collaboration

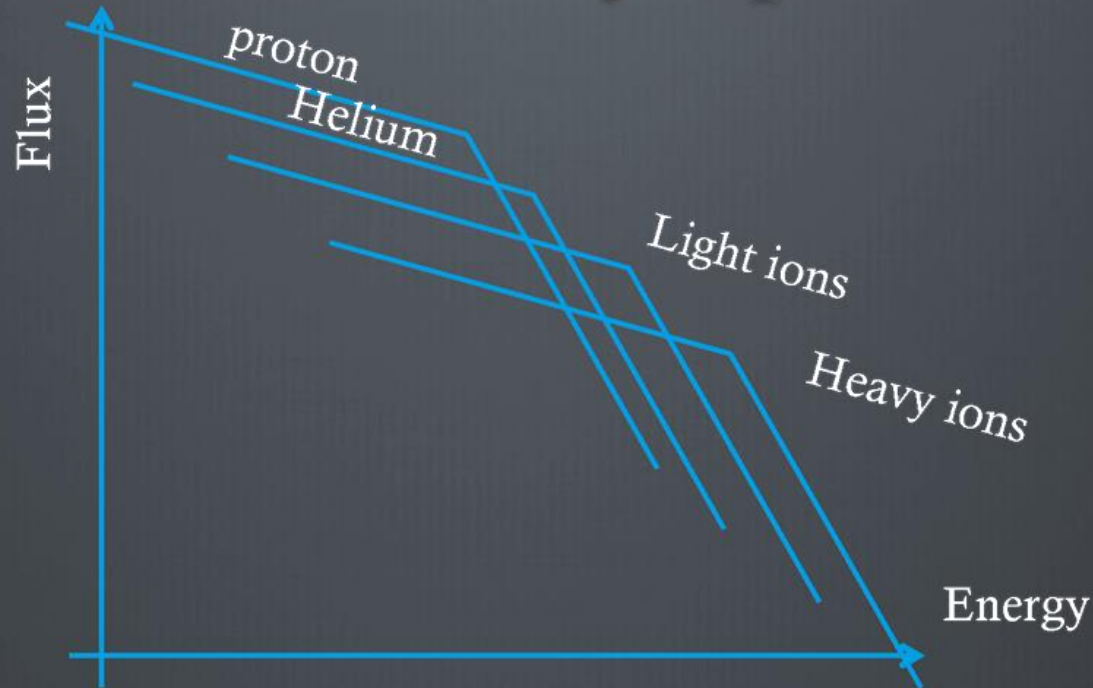
INFN Florence, 16th Dec 2013

Standard Scenario of the Cosmic-Ray Spectrum



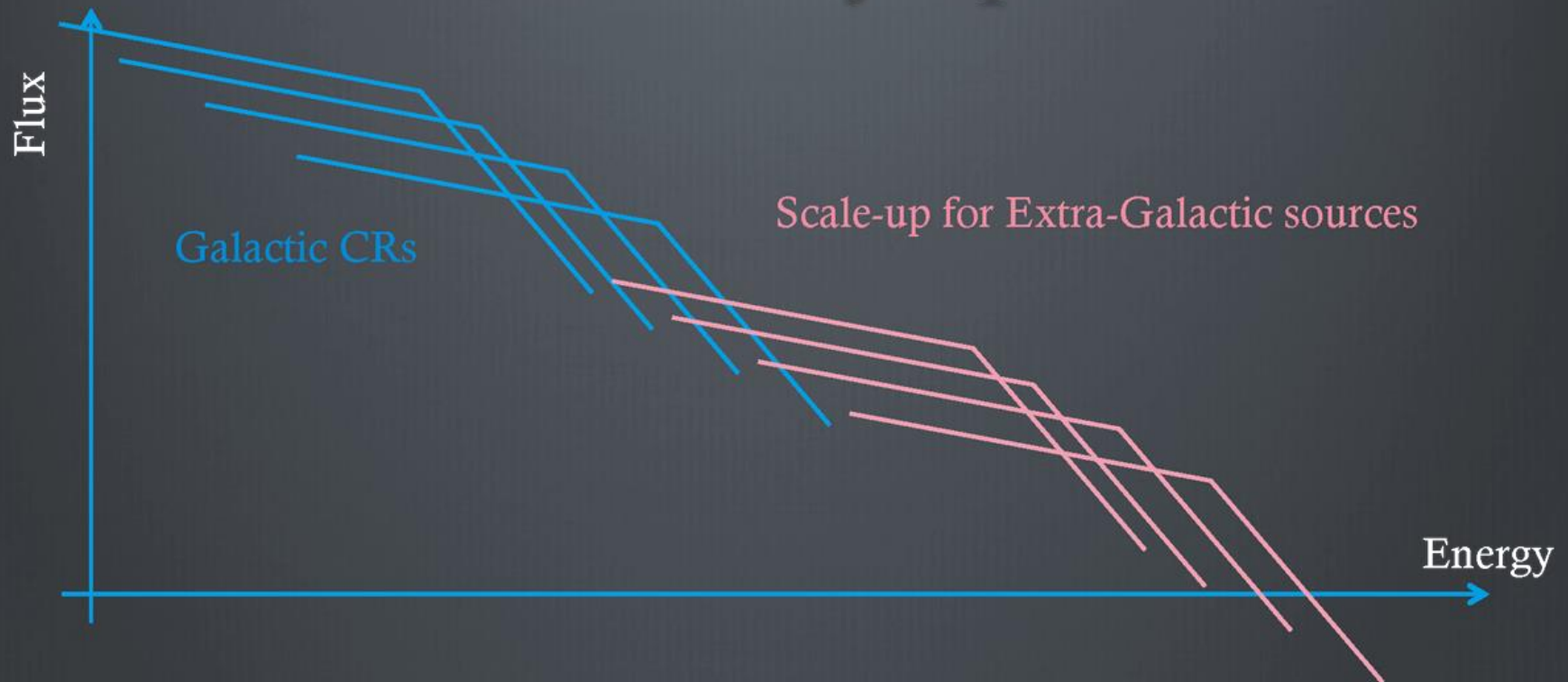
- ⊗ Electro-magnetic process \Rightarrow Same rigidity spectrum for different nuclei

Standard Scenario of the Cosmic-Ray Spectrum



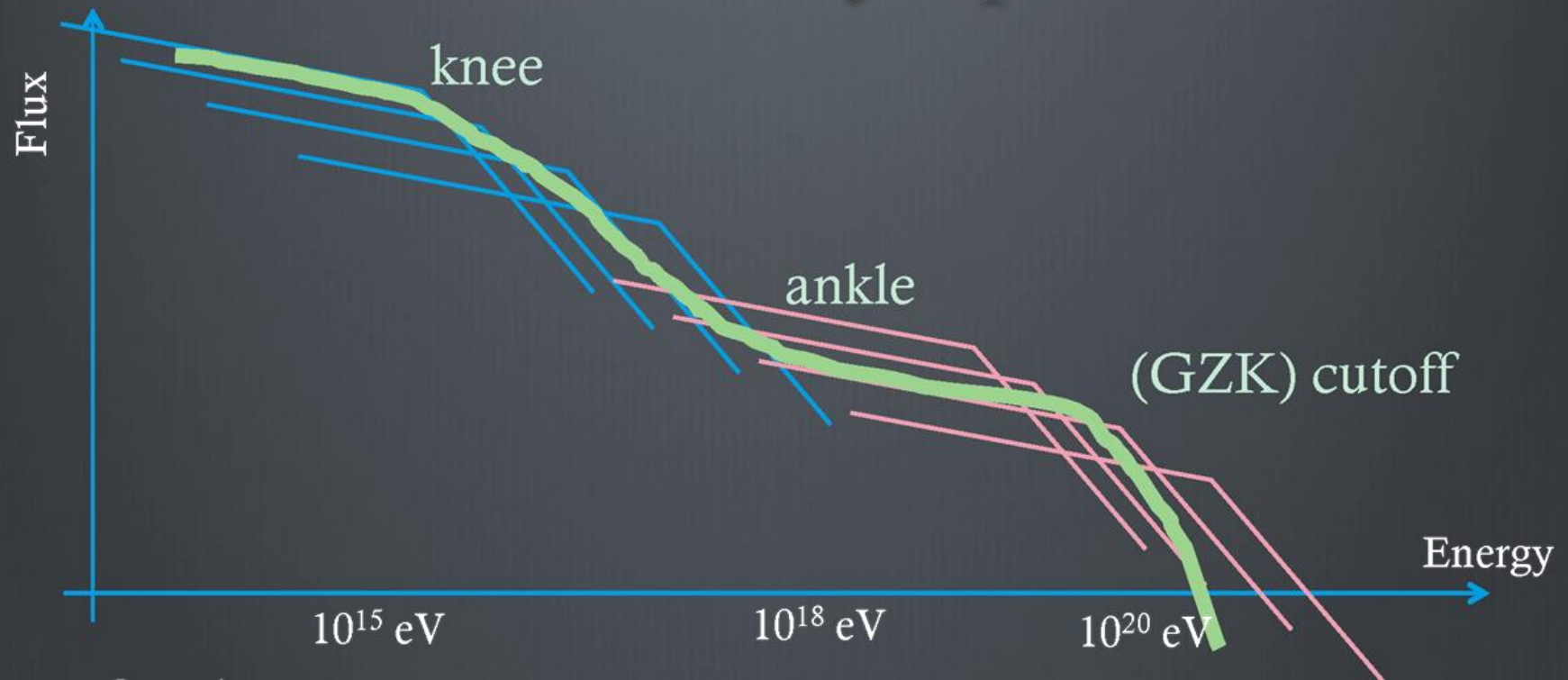
- ⊗ In term of 'Energy,' heavier particles have Z times higher energy than protons

Standard Scenario of the Cosmic-Ray Spectrum



⊗ Over GCR max energy, Extra-galactic CRs appear

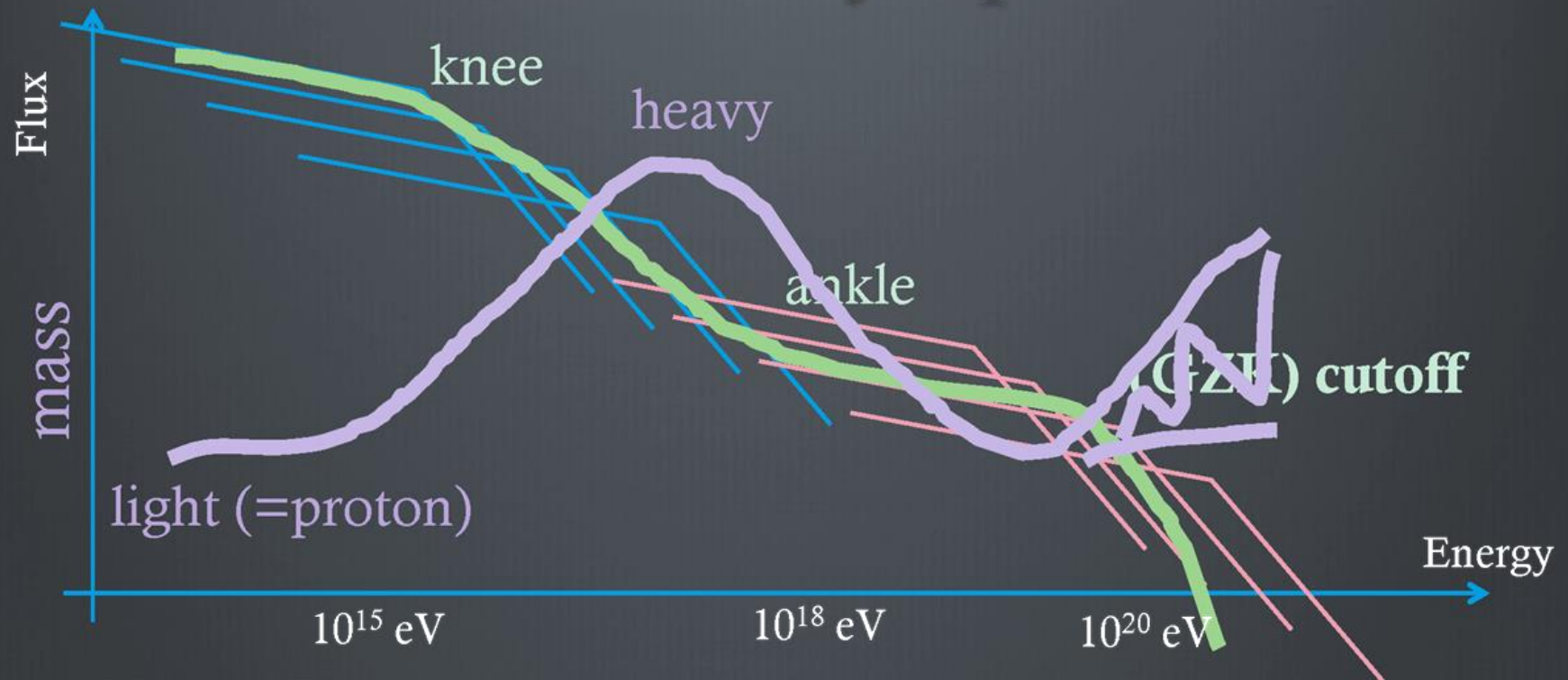
Standard Scenario of the Cosmic-Ray Spectrum



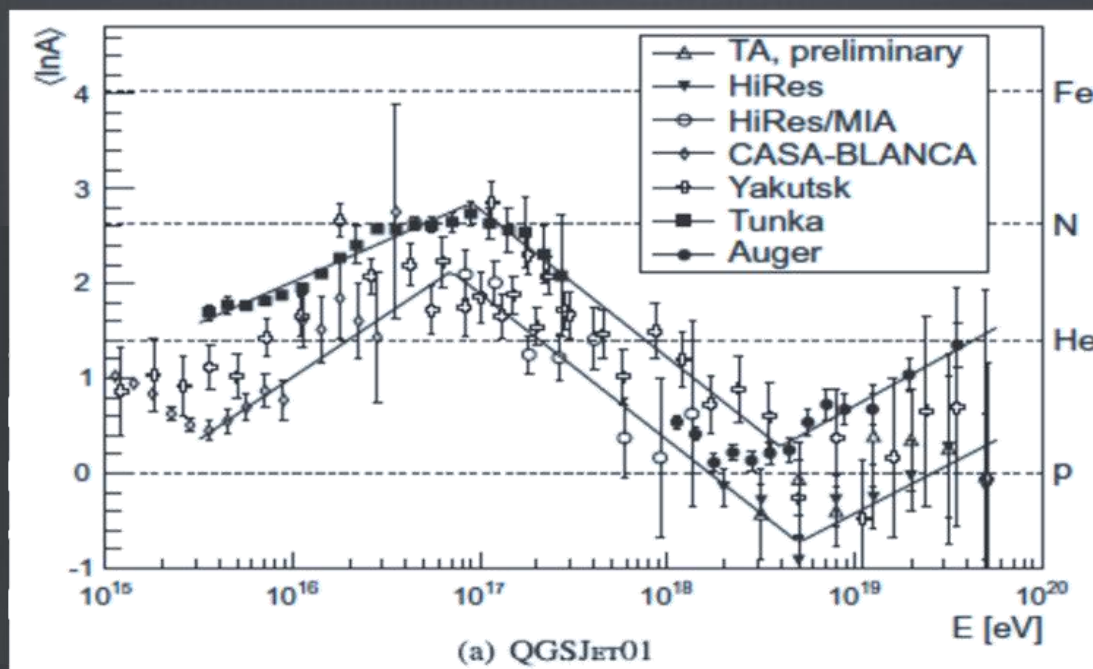
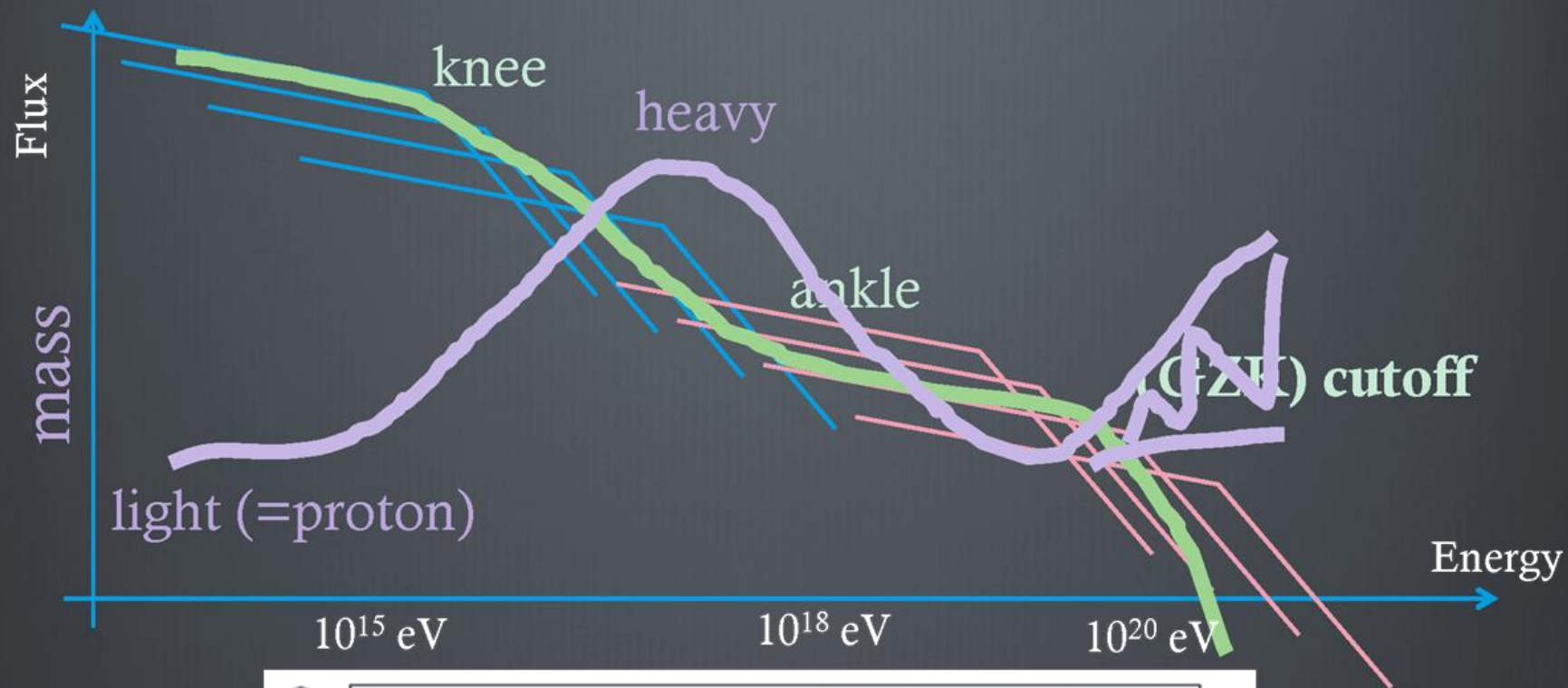
⊗ Questions

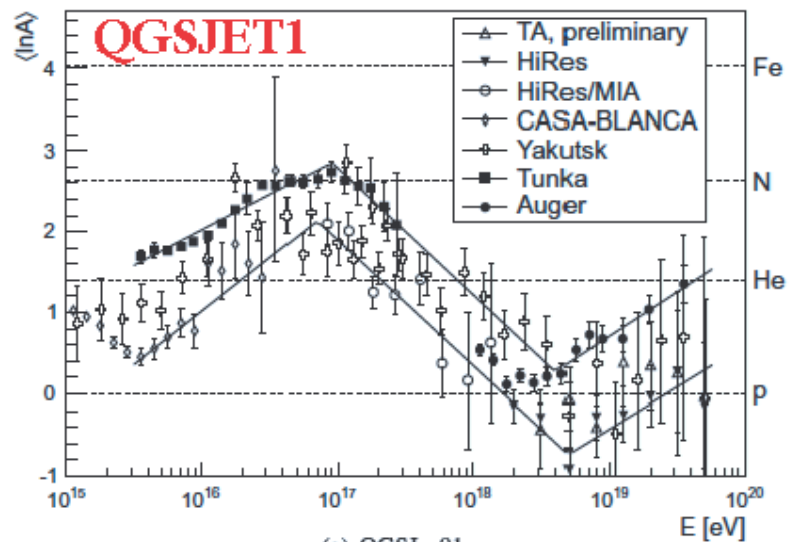
- ⊗ End of GCR
- ⊗ Turn over from GCR to EGCR
- ⊗ Cutoff (acc. Limit, proton GZK, ion GZK)

Standard Scenario of the Cosmic-Ray Spectrum

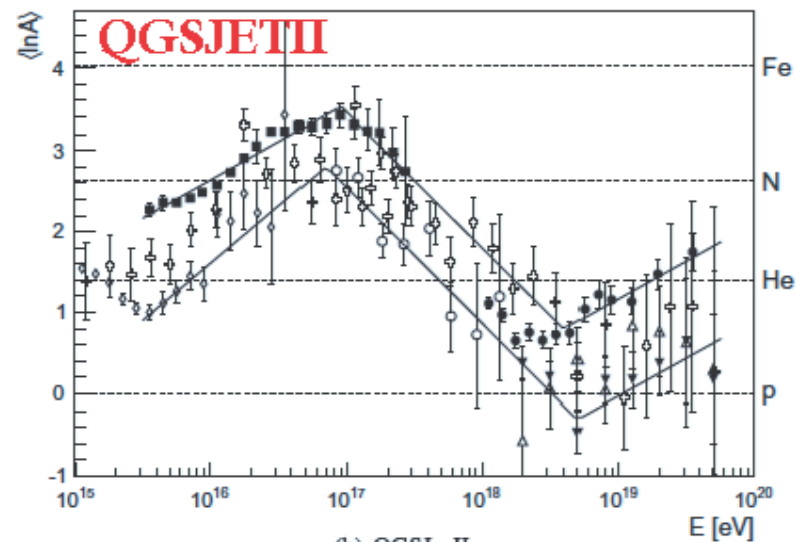


- ⊗ Mass vs. Energy
- ⊗ Light < knee
- ⊗ Light to heavy over knee
- ⊗ Heavy to light around ankle
- ⊗ Light or light to heavy around cutoff

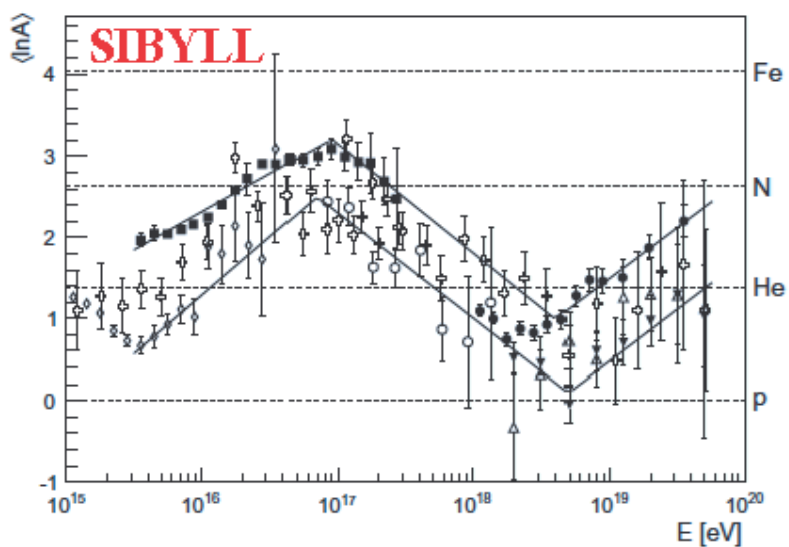




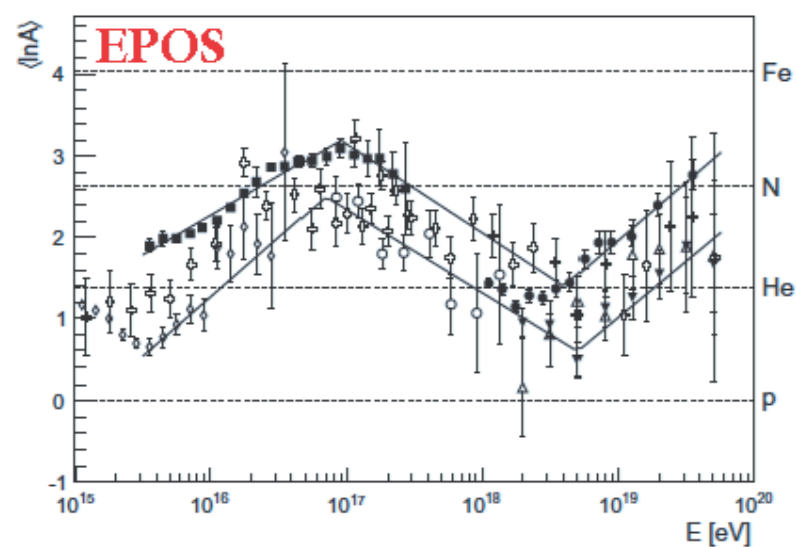
(a) QGSJET01



(b) QGSJETII



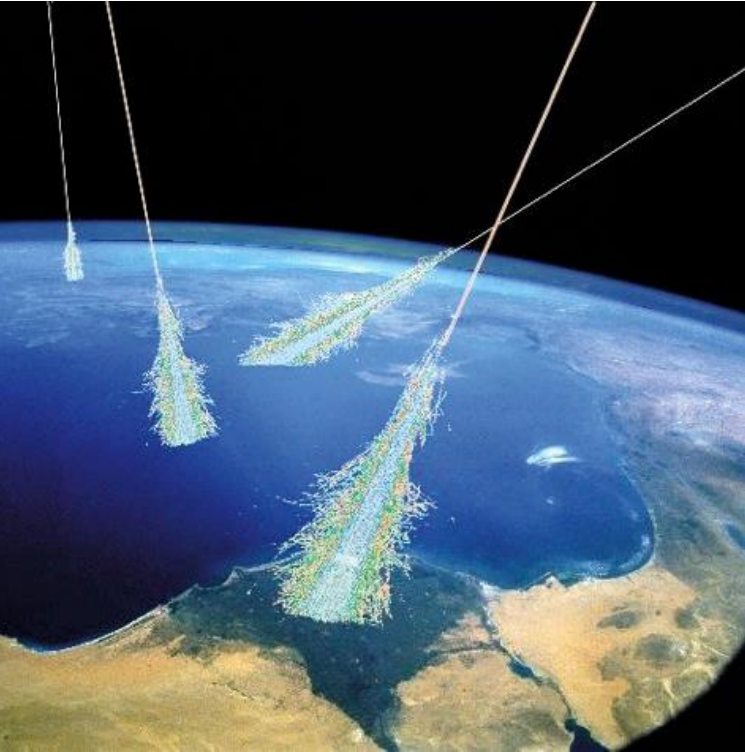
(c) SIBYLL2.1



(d) Epos1.99

(Kampert and Unger, Astropart. Phys., 2012)

High Energy Cosmic Rays



Extensive air shower (EAS) observation

- longitudinal distribution
- lateral distribution
- arrival direction



(air shower development)

Astrophysical parameters

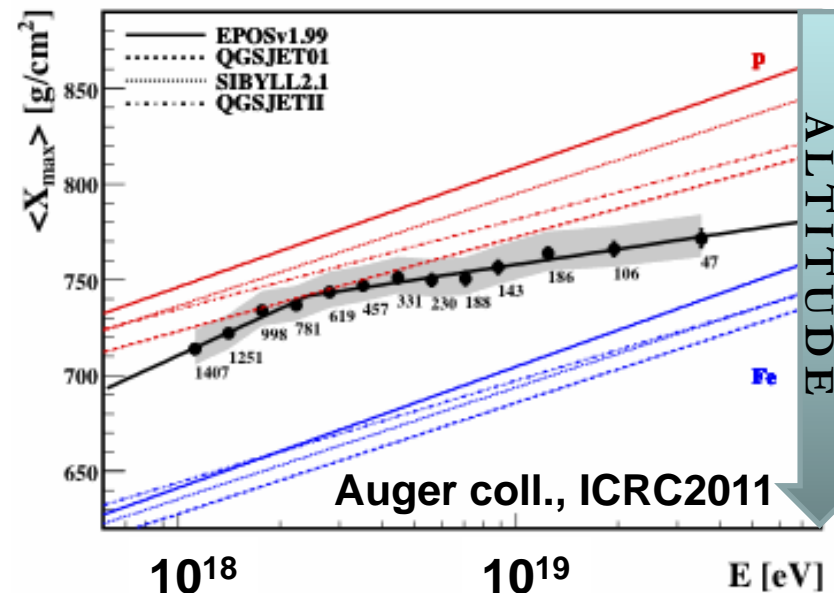
- spectrum
- composition
- source distribution

- X_{\max} : depth of shower maximum in the atmosphere
- $\langle X_{\max} \rangle$ gives information on the CR composition

Uncertainty of hadron interaction models

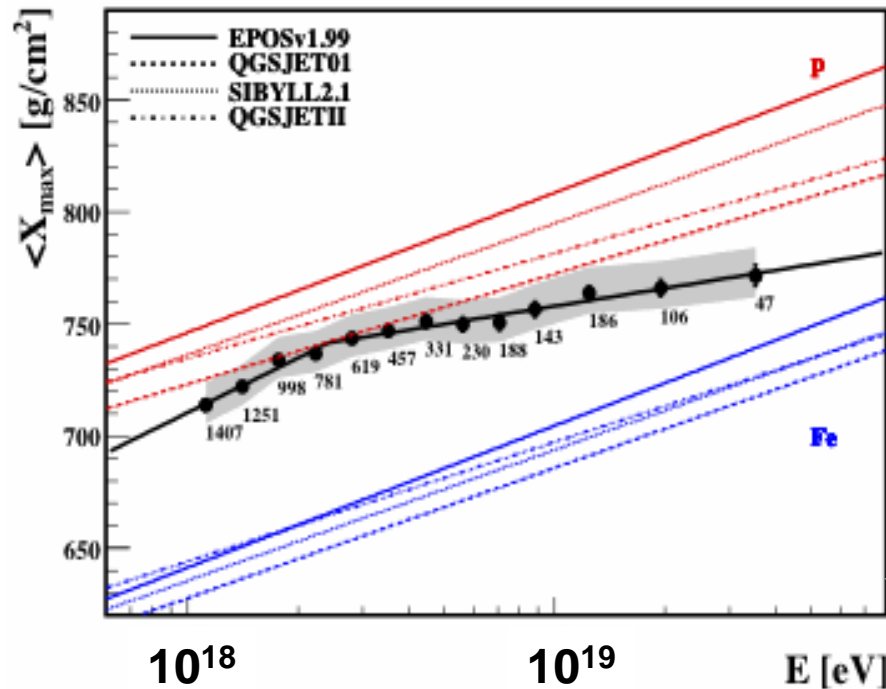


Uncertainty in the interpretation of $\langle X_{\max} \rangle$



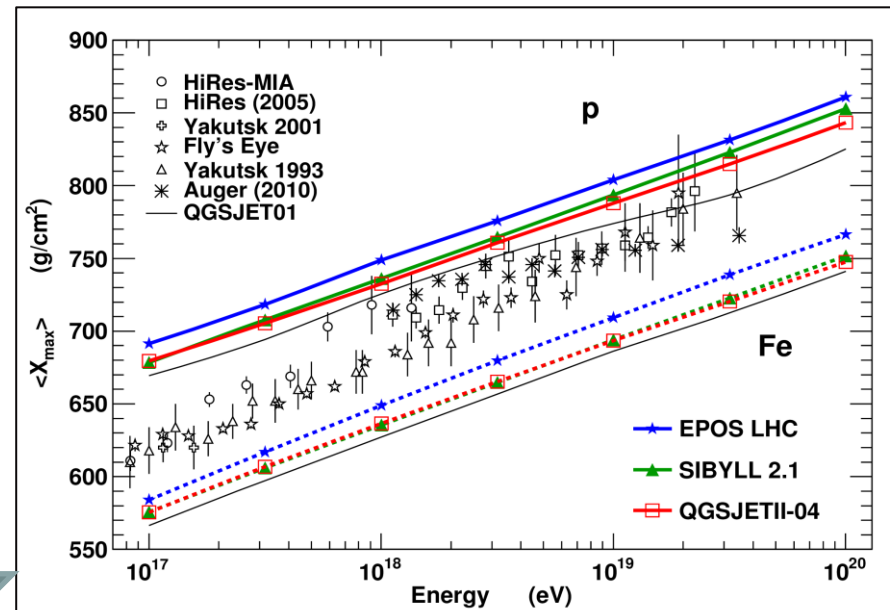
Tuning of hadron interaction models after the first LHC data

pre LHC



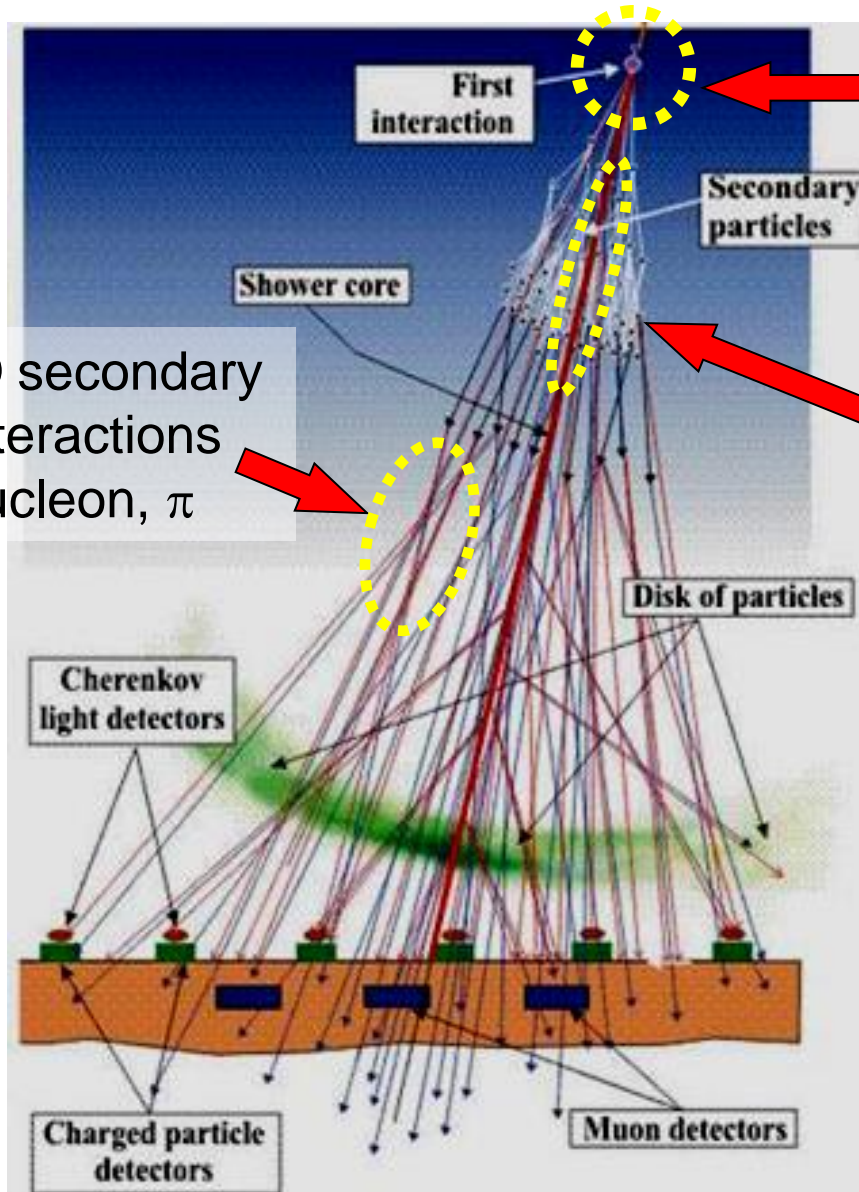
Auger coll., ICRC2011

post LHC



T. Pierog, Cosmic QCD 2013

How accelerator experiments can contribute?



① Inelastic cross section

If large σ : rapid development
If small σ : deep penetrating

② Forward energy spectrum

If softer rapid development
If harder deep penetrating

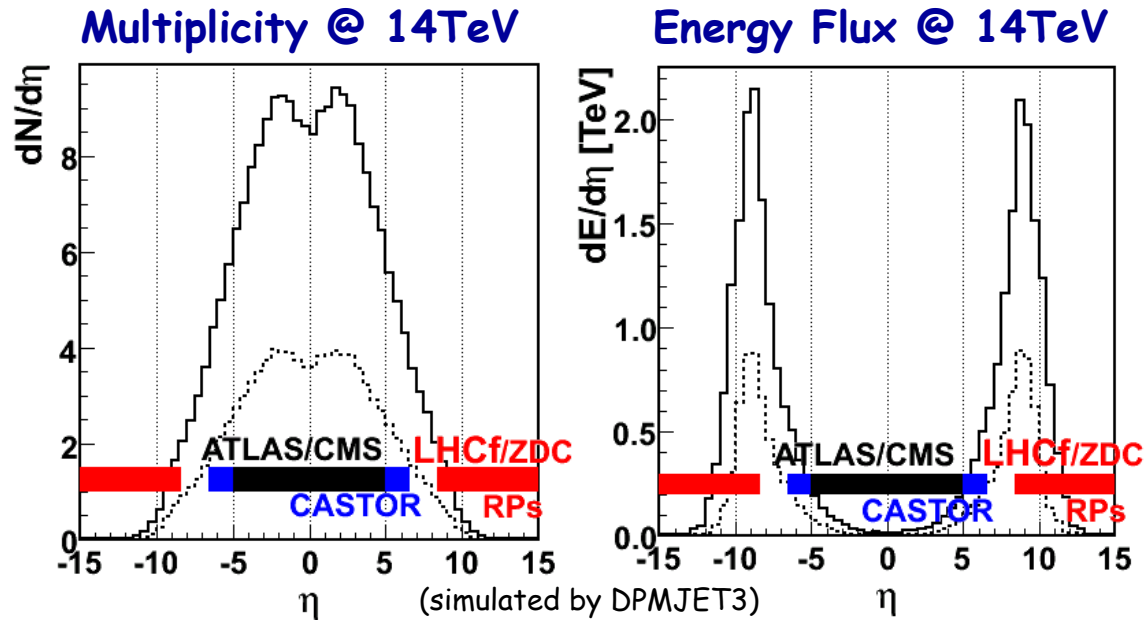
③ Inelasticity $k = 1 - \frac{E_{lead}}{E_{avail}}$

If large k (π^0 s carry more energy)
rapid development
If small k (baryons carry more energy)
deep penetrating

Calibration of hadron interaction models at LHC

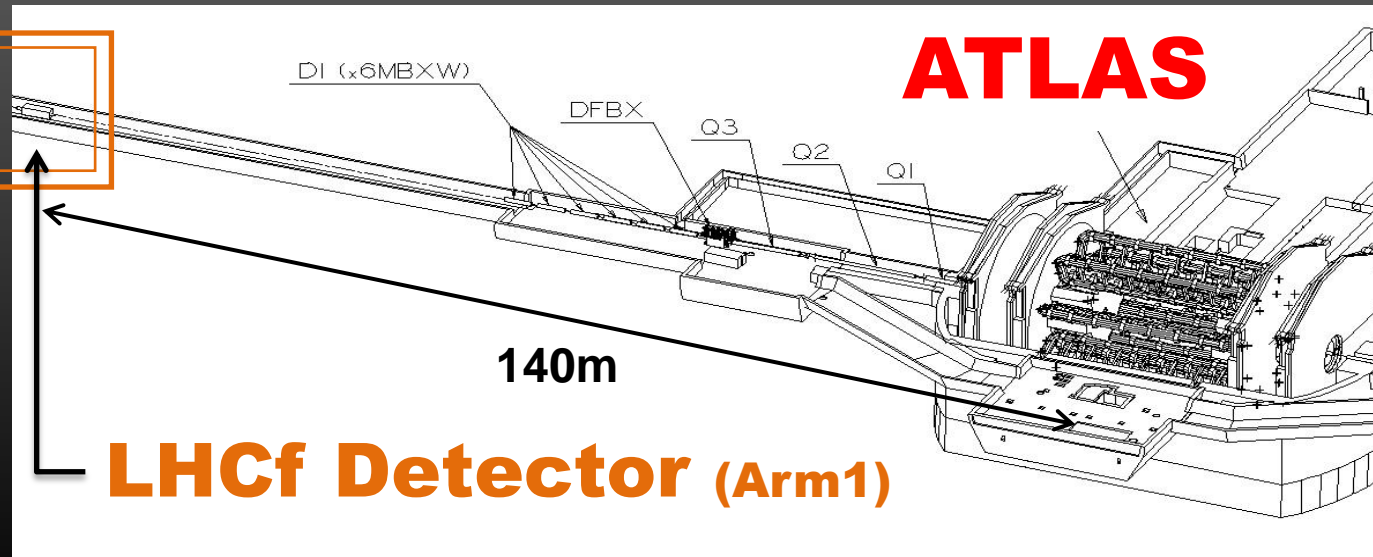
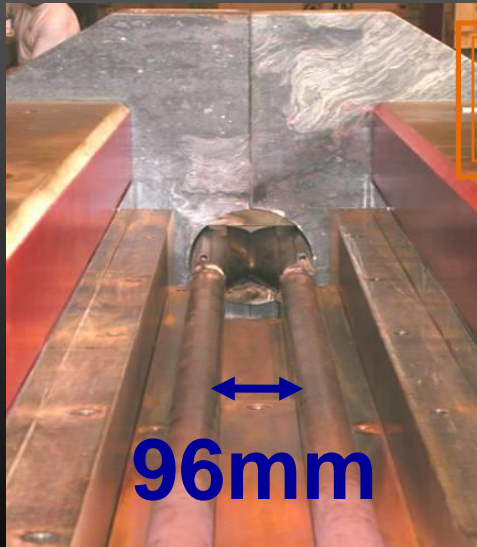
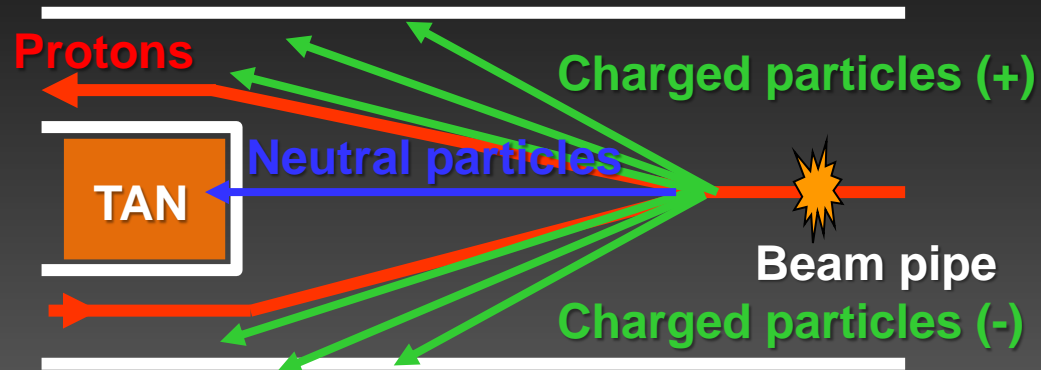
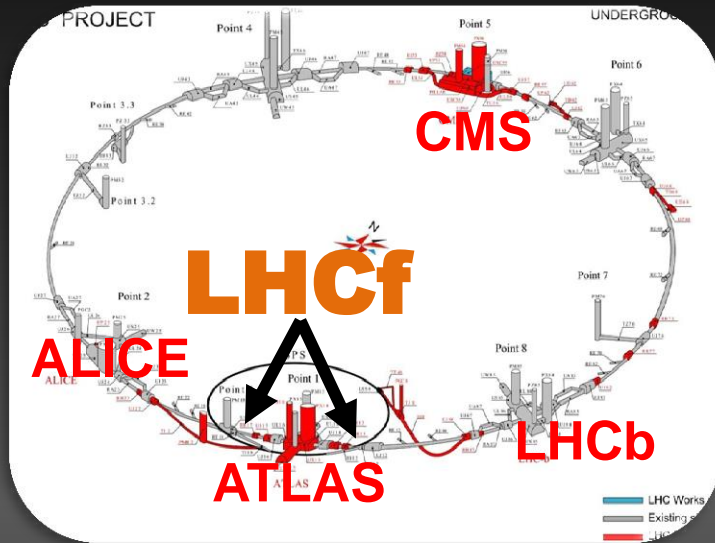
p-p 450 GeV + 450 GeV	→	$E_{\text{lab}} \sim 4 \cdot 10^{14} \text{ eV}$
p-p 3.5 TeV + 3.5 TeV	→	$E_{\text{lab}} \sim 3 \cdot 10^{16} \text{ eV}$
p-p 6.5 TeV + 6.5 TeV	→	$E_{\text{lab}} \sim 9 \cdot 10^{16} \text{ eV}$

- Total cross section ↔ TOTEM, ATLAS, CMS
- Multiplicity ↔ Central detectors
- Inelasticity/Secondary spectra ↔ Forward calorimeters (LHCf, ZDCs)



pseudo-rapidity $\eta = -\ln(\tan(\theta/2))$

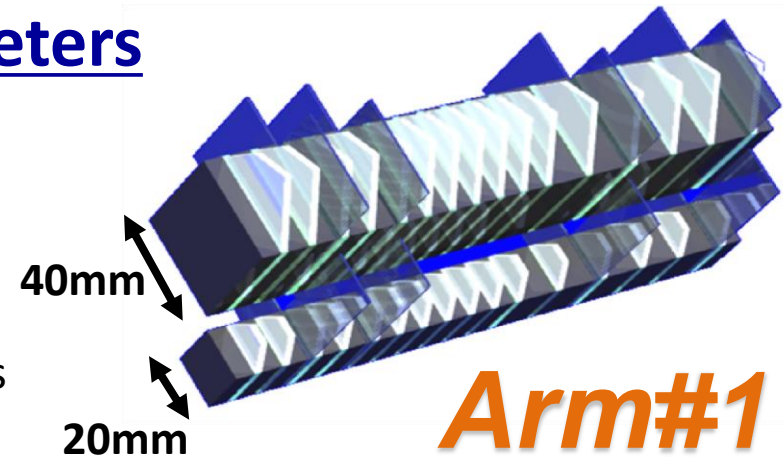
LHCf experimental set-up



LHCf detectors and performances

Sampling and imaging E.M. calorimeters

- **Absorber:** W layers (44 r.l , $1.55 \lambda_i$ in total)
- **Energy measurement:** plastic scintillator tiles
- **4 tracking layers** for imaging:
XY-SciFi (Arm#1) and XY-Silicon μ -strip (Arm#2)
- Each detector has two independent calorimeter towers
→ reconstruction of $\pi^0 \rightarrow \gamma\gamma$ events

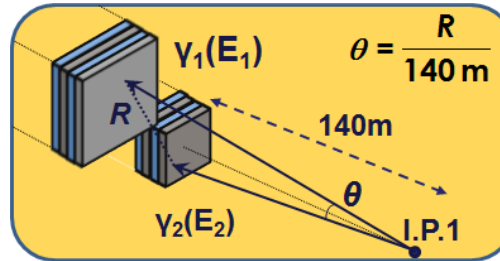


Performance

Energy resolution ($> 100\text{GeV}$)
< 5% for γ and $\sim 30\%$ for n

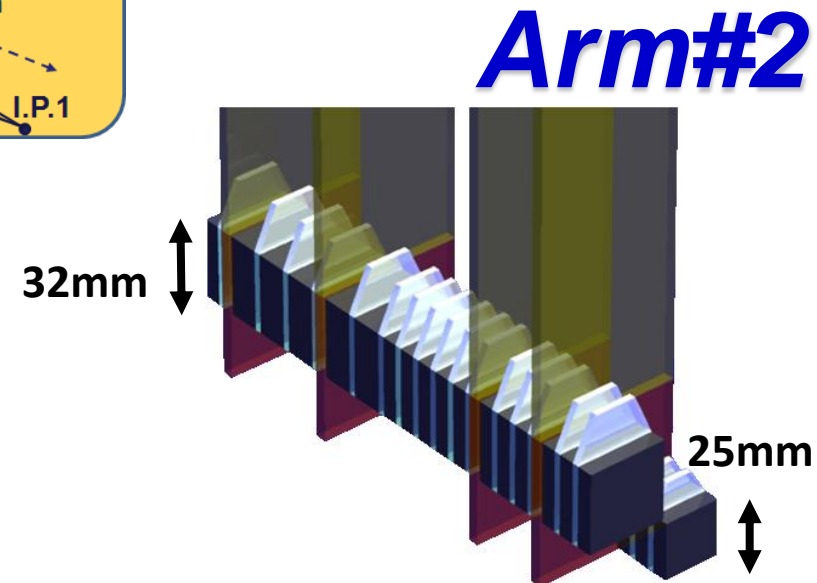
Position resolution

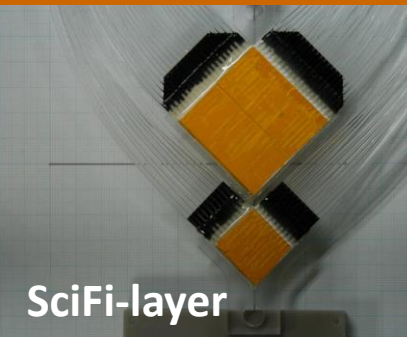
< $200\mu\text{m}$ (Arm#1) and $\sim 40\mu\text{m}$ (Arm#2)



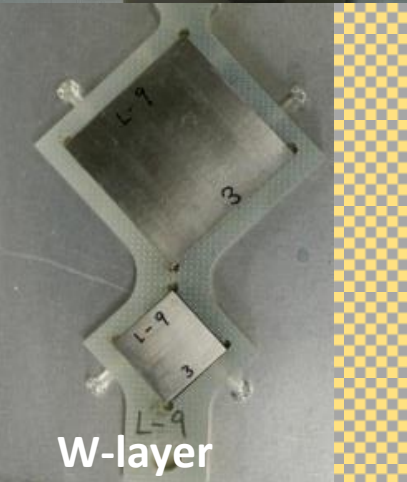
Front Counters

- thin scintillators $80 \times 80 \text{ mm}^2$
- monitoring of beam condition
- background rejection
- Van der Meer scan





SciFi-layer



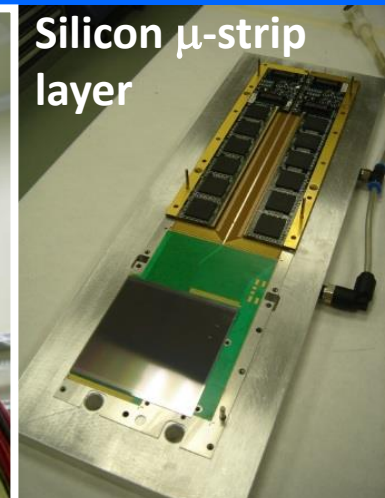
W-layer



Arm#1 Detector



Arm#2 Detector



Silicon μ -strip layer



Silicon read-out



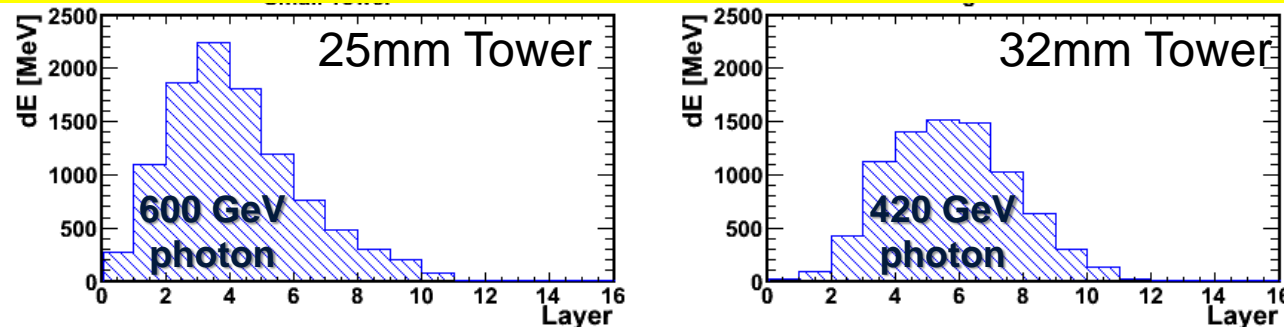
Arm#1 in the TAN



Arm#2 in the TAN

Detection of a π^0 in Arm#2

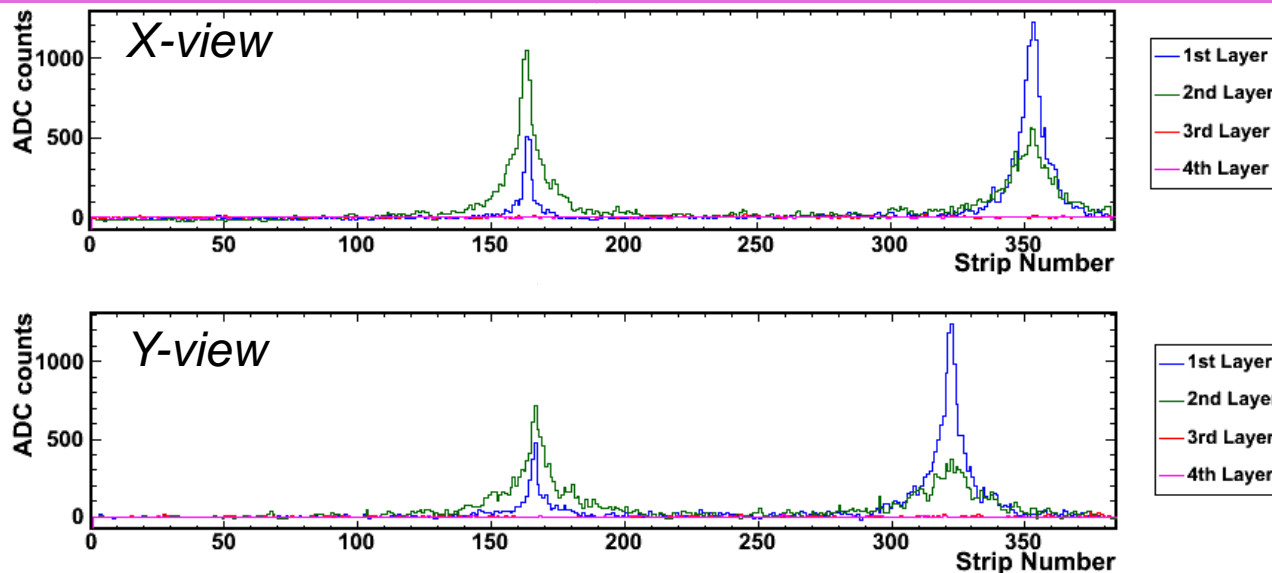
Longitudinal development measured by scintillator layers



Determination of **energy** from total energy release

PID from shape

Transverse profile measured by silicon μ -strip layers



Determination of the **impact point**

Measurement of the **opening angle** of gamma pairs

Identification of **multiple hit**

$$\text{Reconstruction of } \pi^0 \text{ mass } M_{\pi^0} \cong \sqrt{E_{\gamma 1} E_{\gamma 2}} \cdot \theta$$

LHCf Status

⊗ Done

- ⊗ 0.9, 2.76, 7 TeV pp collision, 5 TeV pPb collision data taking
- ⊗ Photon spectra at 0.9 and 7TeV published ▪ **PLB 715 (2012) 298**
- ⊗ π^0 spectra at 7 TeV published ▪ **PLB 703 (2011) 128**
- ⊗ Performance at 7TeV published ▪ **PRD 86 (2012) 092001**
▪ **IJMPA 28 (2013) 1330036**

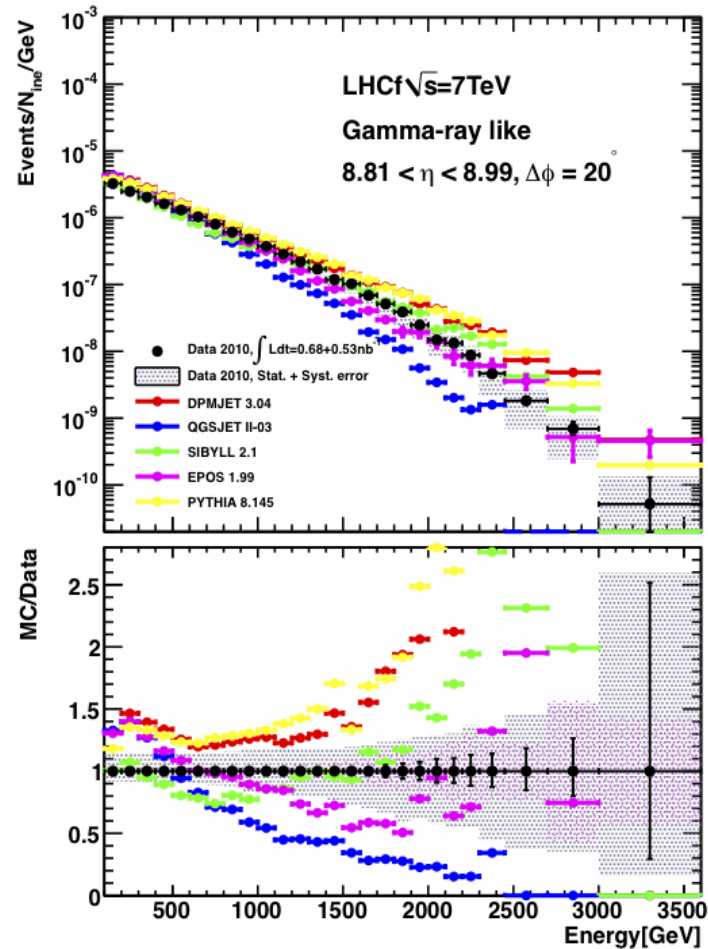
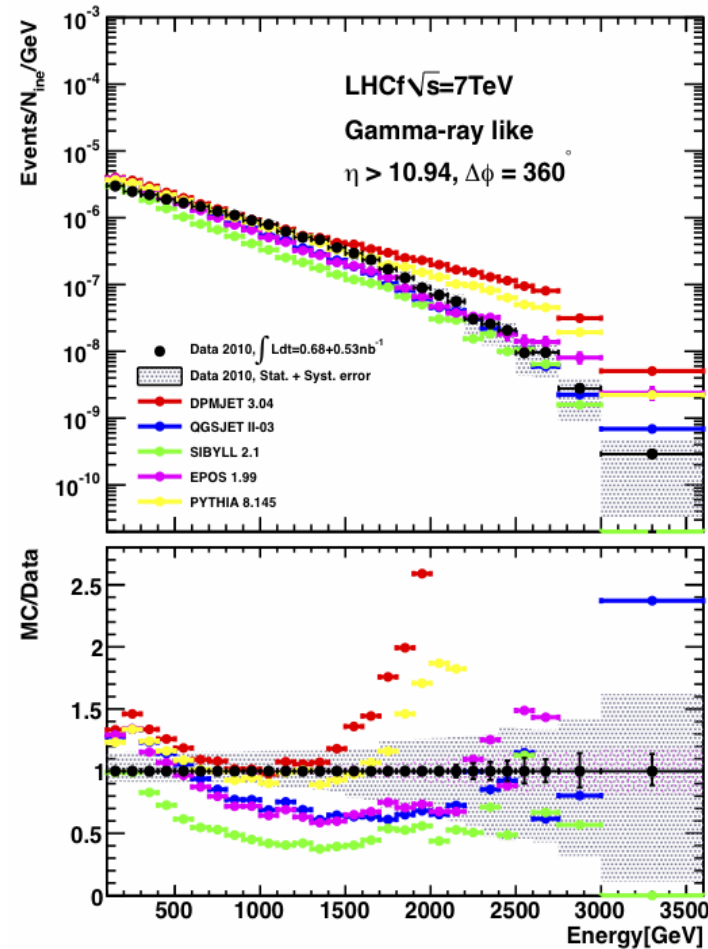
⊗ On going

- ⊗ Neutron spectra at 7TeV
- ⊗ π^0 and UPC spectra at 5TeV pPb
- ⊗ Detector upgrade for 13 TeV pp

⊗ Plan

- ⊗ 13TeV pp collision in 2015 (operation plan in discussion)
- ⊗ 0.5TeV pp at RHIC (LOI submitted)
- ⊗ Discussion for light ion collision at RHIC and LHC

Comparison of single γ data at $\sqrt{s} = 7$ TeV with hadronic interaction models (pre-LHC versions)



DATA

DPMJET 3.04

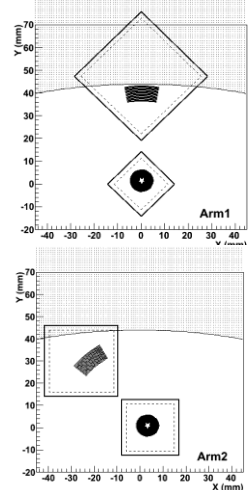
QGSJET II-03

SIBYLL 2.1

EPOS 1.99

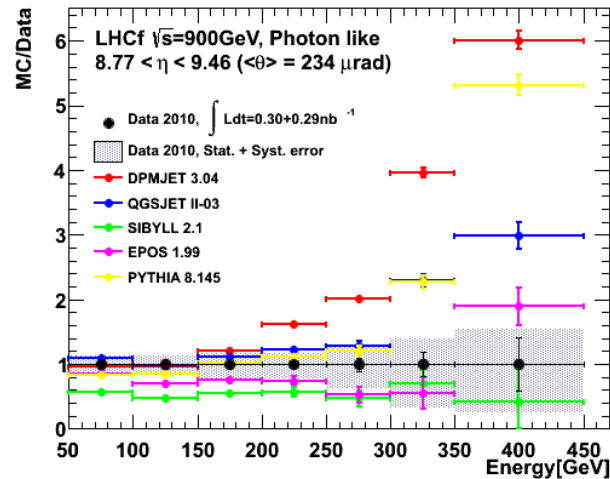
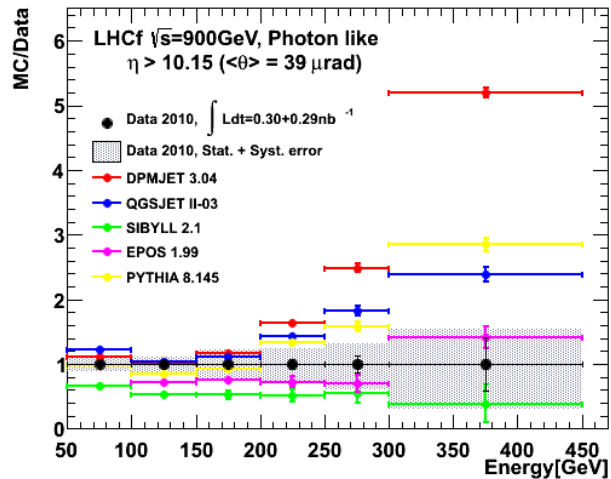
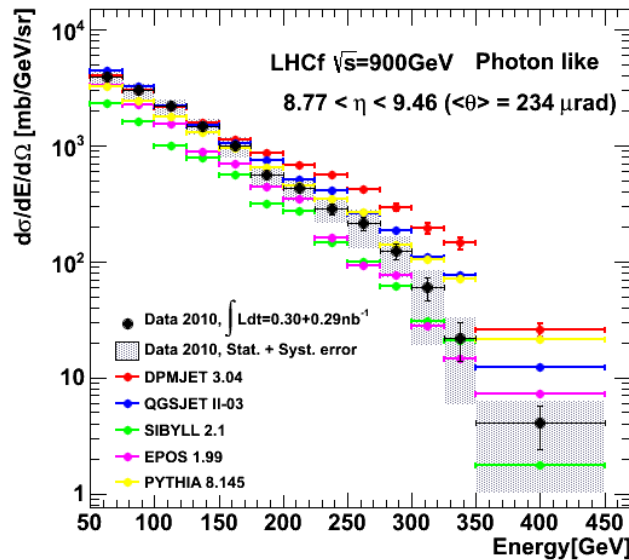
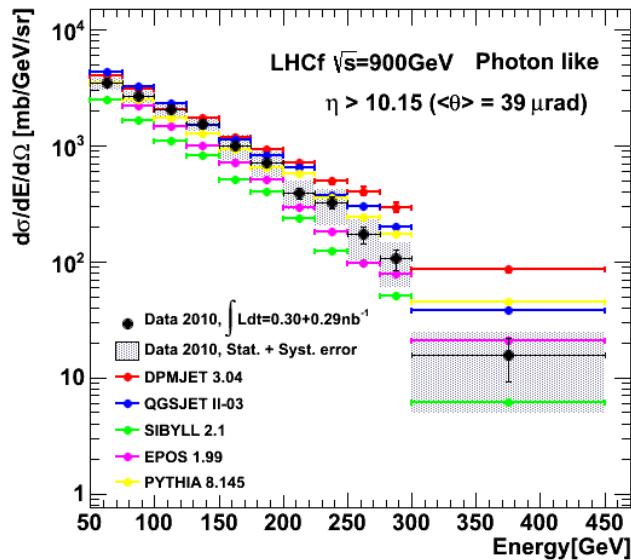
PYTHIA 8.145

Syst.+Stat.



- No model can reproduce the **LHCf data** perfectly
- **DPMJET**, **PYTHIA** are in good agreement at high- η for $0.5 < E_\gamma < 1.5\text{TeV}$, but harder for $E > 1.5\text{TeV}$
- **QGSJET**, **SIBYLL**, **EPOS** show reasonable agreement of shape for high η , but not for low η

Comparison of single γ data at $\sqrt{s} = 900$ GeV with hadronic interaction models (pre-LHC versions)



DATA

DPMJET 3.04

QGSJET II-03

SIBYLL 2.1

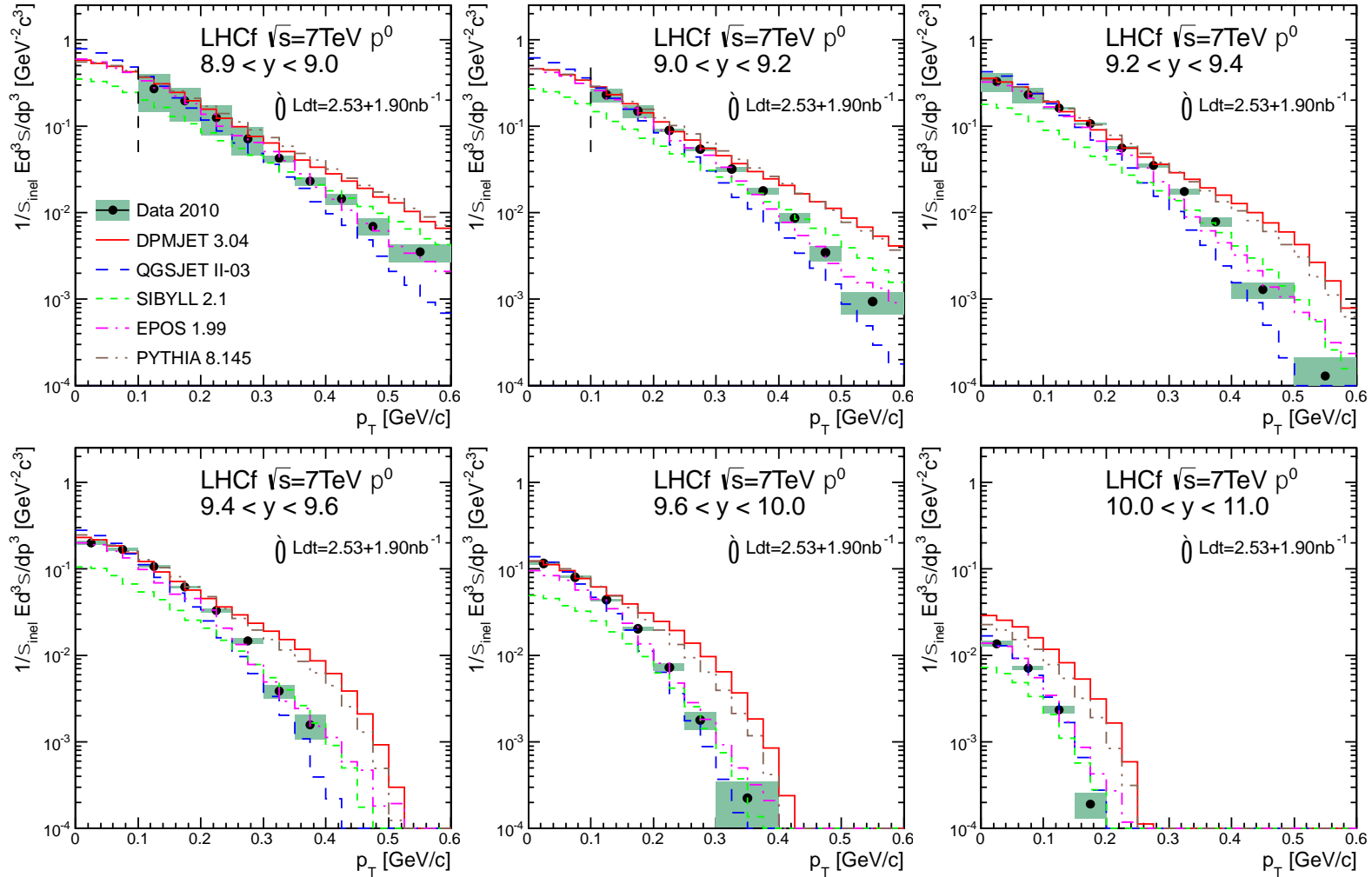
EPOS 1.99

PYTHIA 8.145

Syst.+Stat.

- No strong evidence of η -dependence
- SYBILL and EPOS show reasonable agreement of shape
- None of the models reproduces **LHCf data** within the error bars

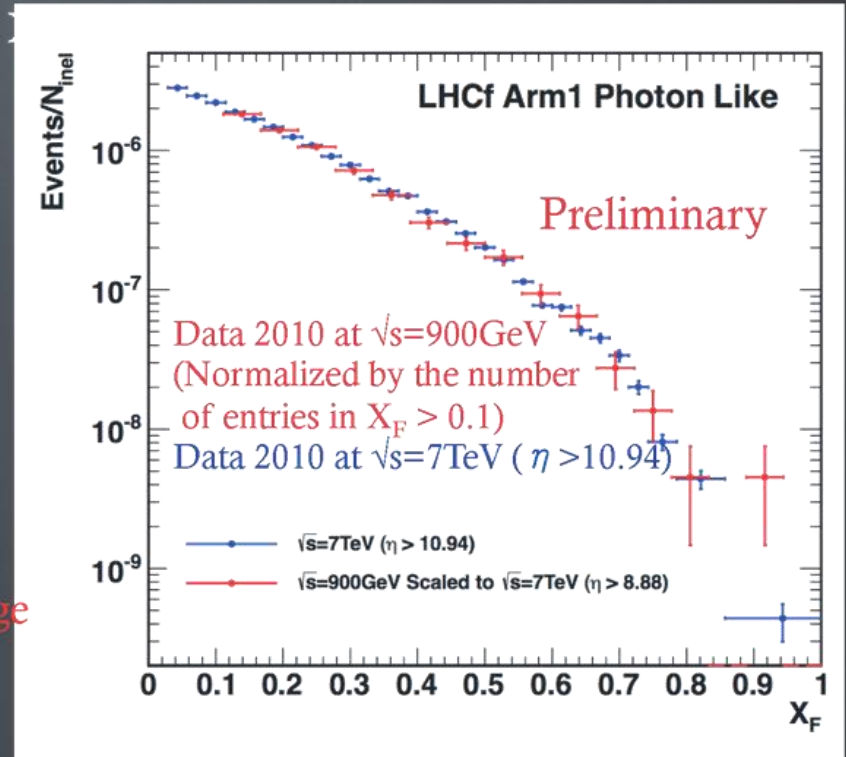
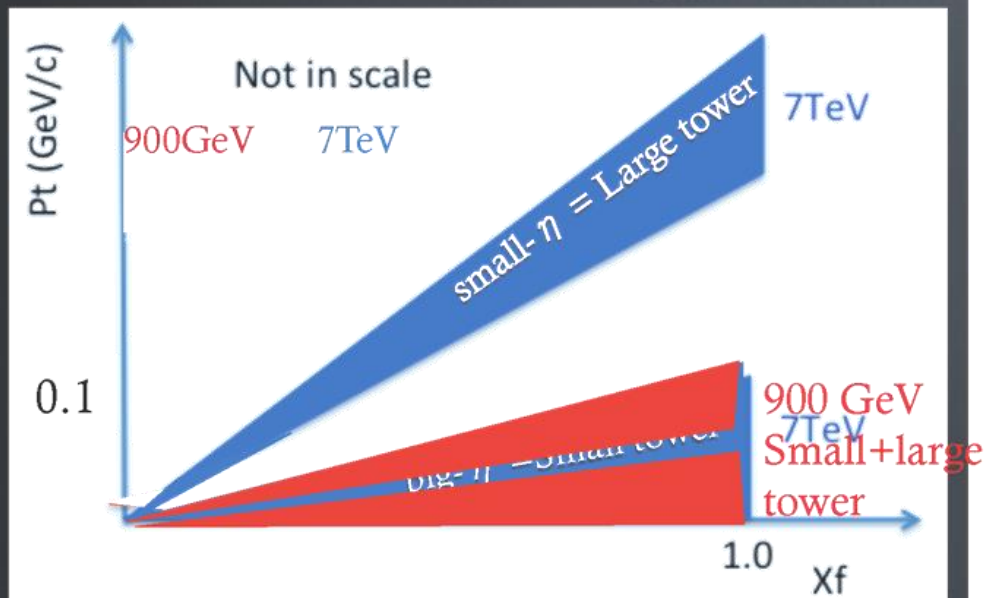
Comparison of π^0 data at $\sqrt{s} = 7\text{TeV}$ with hadronic interaction models (pre-LHC versions)



- **EPOS** shows the best agreement with data
- **DPMJET** and **PYTHIA** have harder spectra than data
- **QGSJET** has softer spectrum than data

900GeV vs. 7TeV

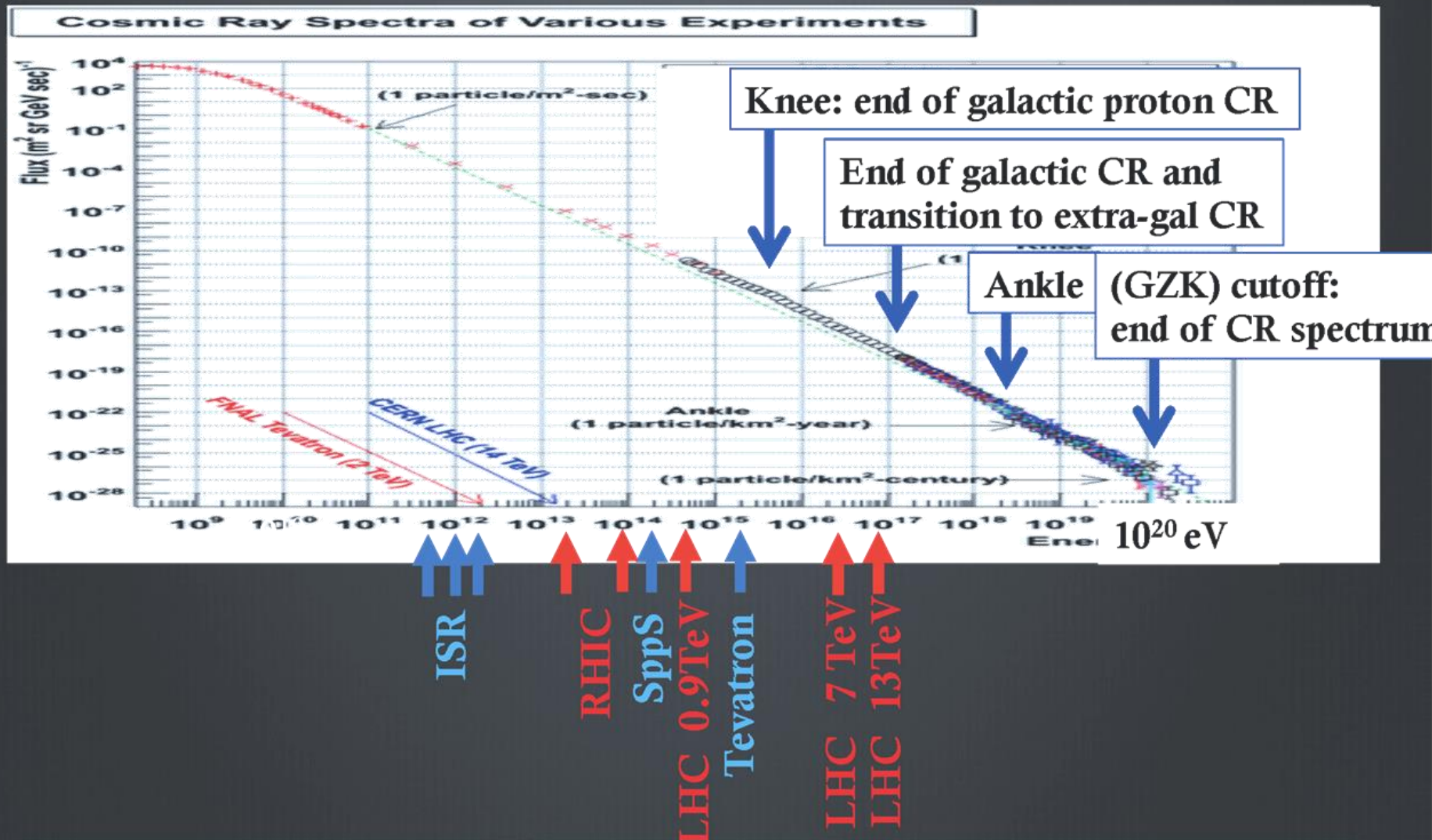
LHCf coverage in $X_F p_T$ plane
 $(X_F = E/E_{beam})$



- ✓ normalized by the number of entries in $X_F > 0.1$
- ✓ statistical errors only

Good agreement of X_F spectrum shape between 900 GeV and 7 TeV

Cosmic-ray spectrum & Colliders

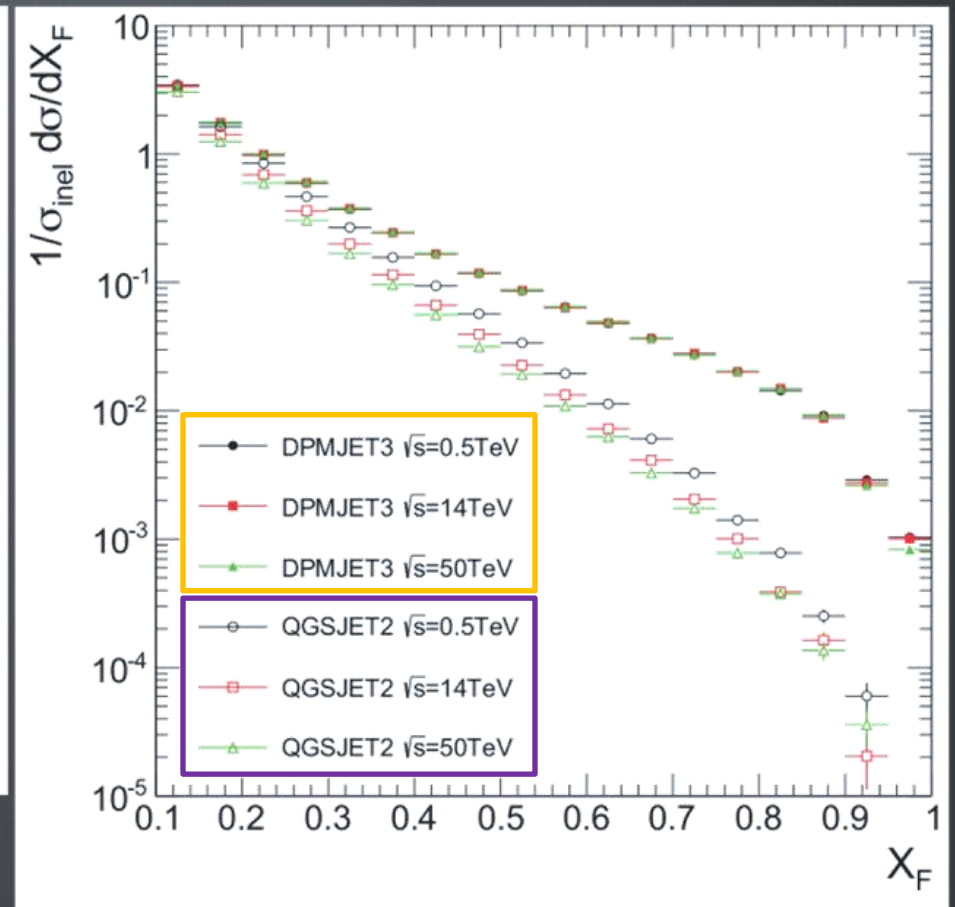
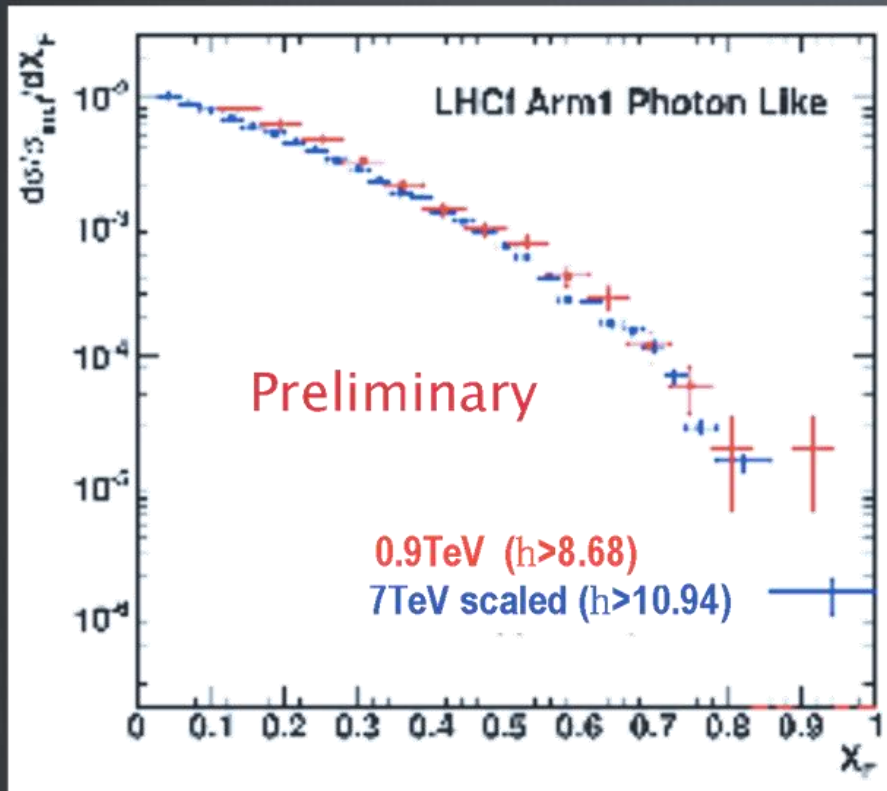


Perfect (or best at least) understanding up to 10^{17} eV helps CR physics

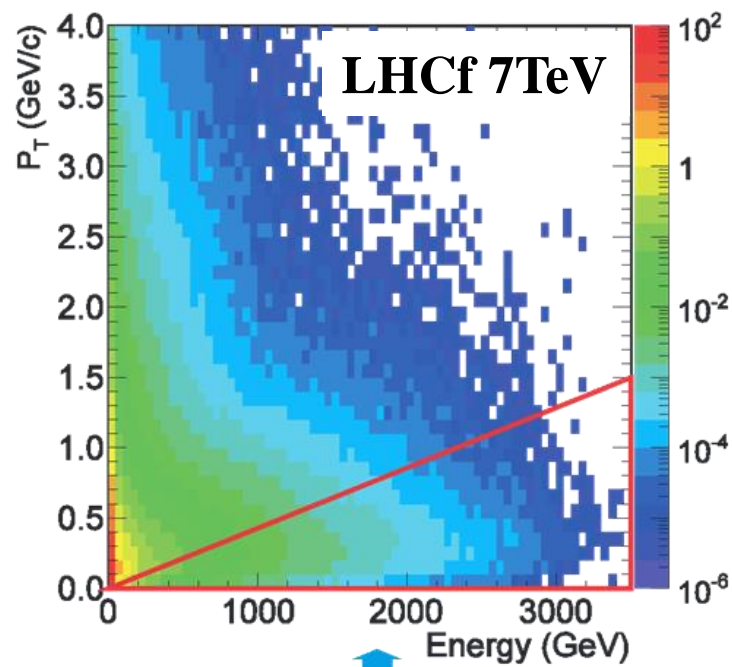
x_F scaling : a key for extrapolation

LHC single gamma data
(900GeV pp / 7TeV pp)

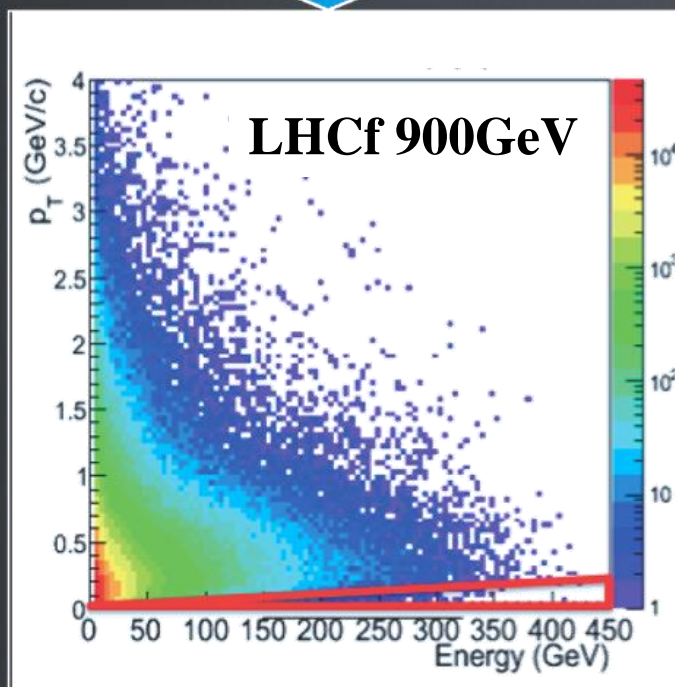
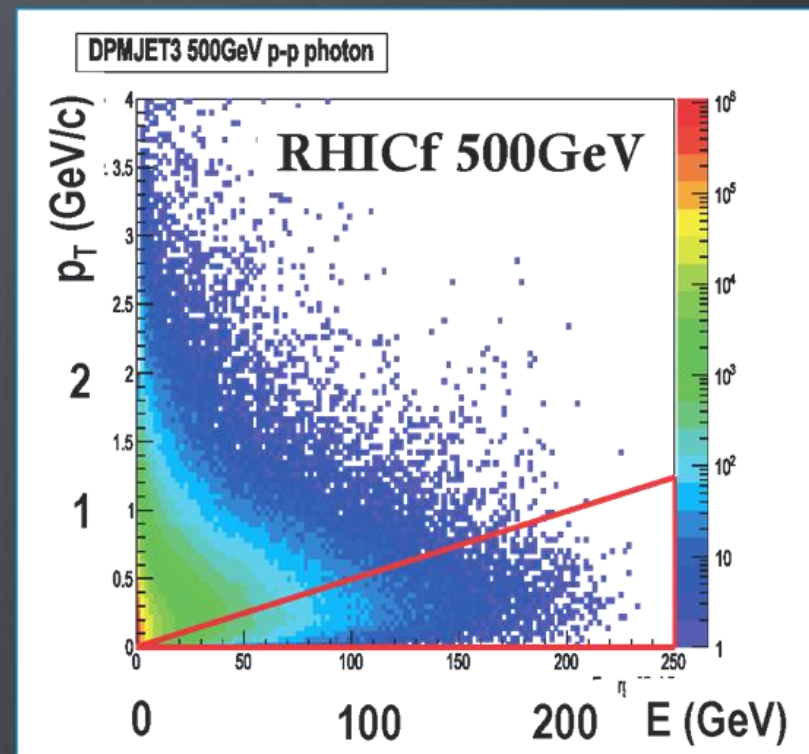
Expected from models
(5TeV, 14TeV and 50TeV)



But this comparison done in very limited phase space..



RHICf 500GeV
 Similar phase space to
 LHCf 7TeV



LHCf: future plan

$p\text{-}p \sqrt{s} = 13\text{TeV}$ at LHC (2015)

Main target: measurement at the LHC design energy.

Study of energy scaling by comparison with $\sqrt{s} = 900\text{ GeV}$ and 7 TeV data.

Upgrade of the detectors for radiation hardness.

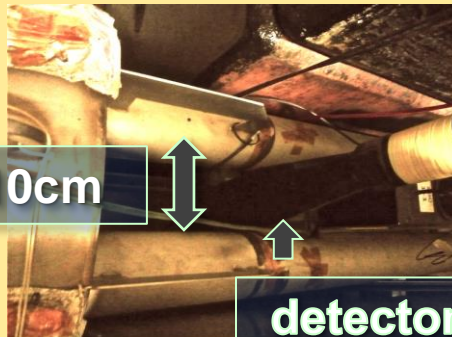
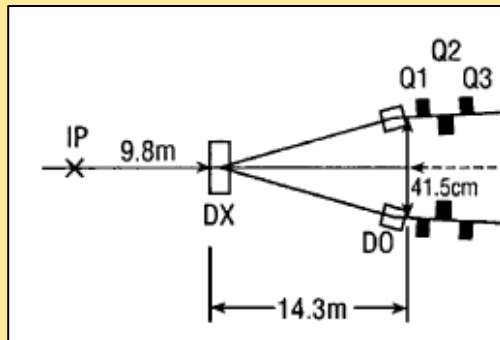
$p\text{-light ions}$ (O, N) at the LHC (2019?)

It allows studying HECR collisions with atmospheric nuclei.

RHICf experiment at RHIC (*Relativistic Heavy Ion Collider @ Brookhaven*)

Lower collision energy, ion collisions.

LOI to the RHIC committee submitted.



$p\text{-}p$ collisions:

- Max. $\sqrt{s} = 500\text{ GeV}$
- Polarized beams

Ion collisions:

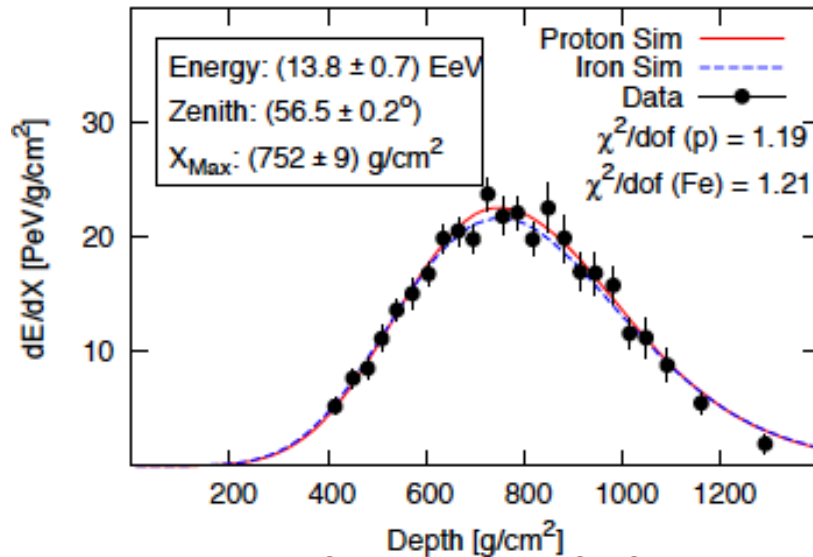
- Au-Au, d-Au
- Max. $\sqrt{s} = 200\text{ GeV}$
- Possible, d-O,N ($p\text{-O,N}$)
 - ➔ Cosmic ray – Air @ knee energy

Conclusions

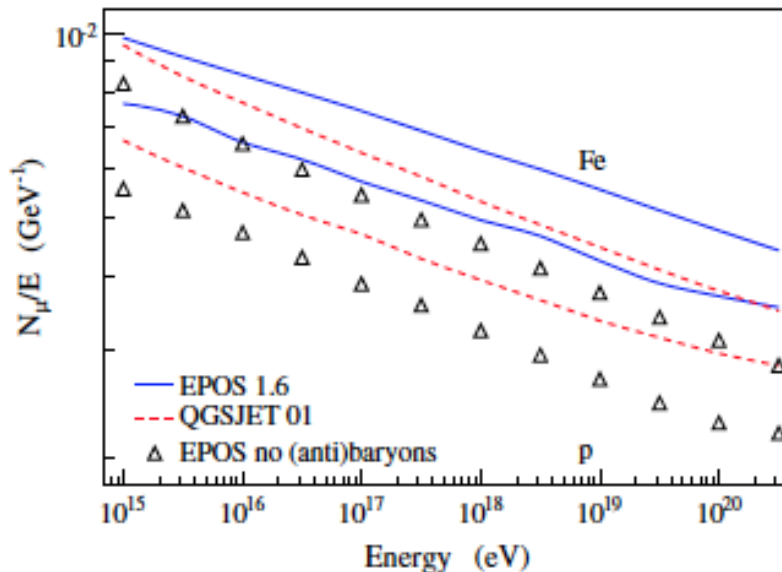
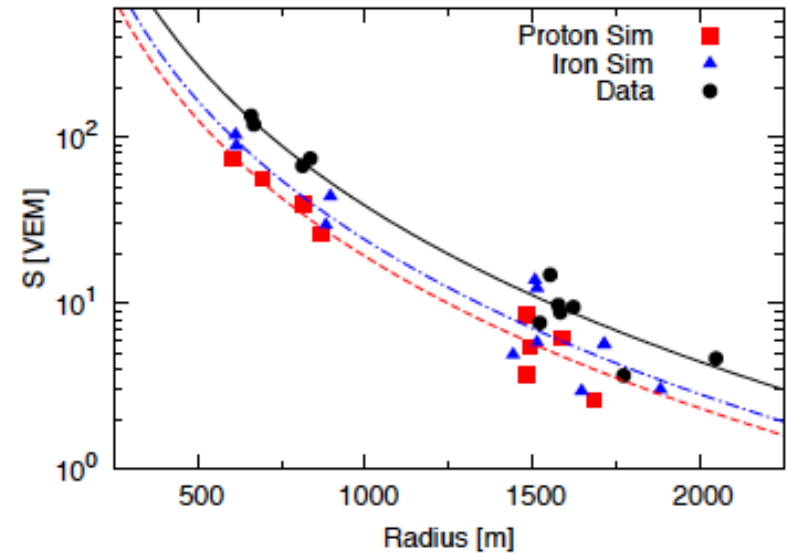
- LHCf is a small experiment at LHC dedicated to **forward physics**
 - Important for High Energy Cosmic-Ray Physics
- We have published **spectra of photons and neutral pions** for p-p interactions at $\sqrt{s} = 7$ TeV and $\sqrt{s} = 900$ GeV
 - None of the hadron interaction models that we have considered can reproduce the data within the errors, but data lie anyway between the models
 - On-going data analysis for the hadron component (neutrons)
- p-Pb run at the beginning of 2013
 - Successful data taking in p-remnant and Pb-remnant side
 - Common operations with ATLAS
 - On-going data analysis
- Future plan
 - Complete the **upgrade of the detectors** for radiation hardness
 - Data taking for **p-p collisions at $\sqrt{s} = 13$ TeV** (2015)
 - Run **p-light ions** at LHC (2019?)
 - Operations at **RHIC** (p-O or p-N at lower energies)

BACKUP SLIDES

Muon excess at Pierre Auger Obs.



Pierre Auger Collaboration, ICRC
2011 (arXiv:1107.4804)



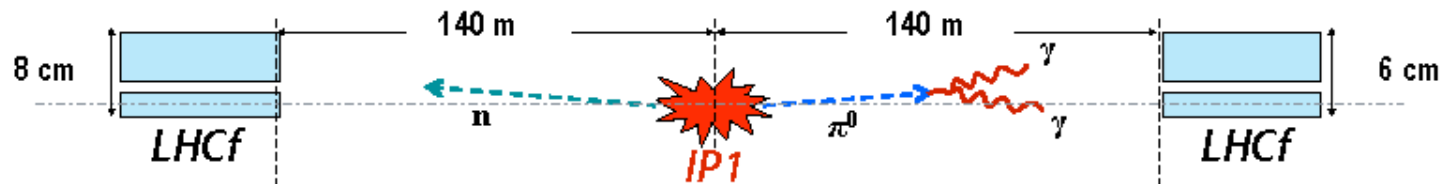
Auger hybrid analysis

- event-by-event MC selection to fit FD data (top-left)
- comparison with SD data vs MC (top-right)
- **muon excess in data even for Fe primary MC**

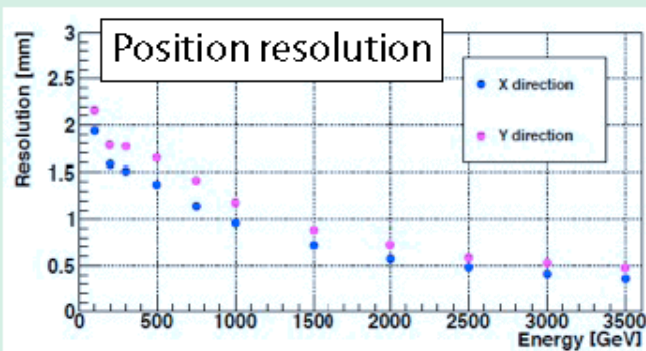
EPOS predicts more muon due to larger baryon production
=> importance of baryon measurement

Pierog and Werner, PRL 101 (2008) 171101

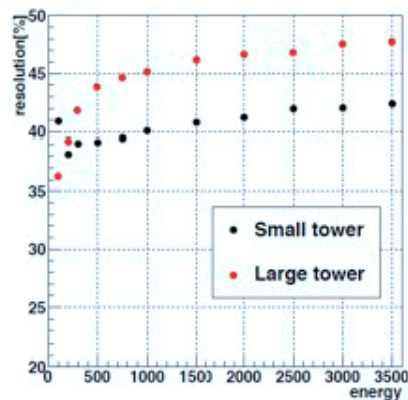
Detector performances



Hadronic shower (MC)

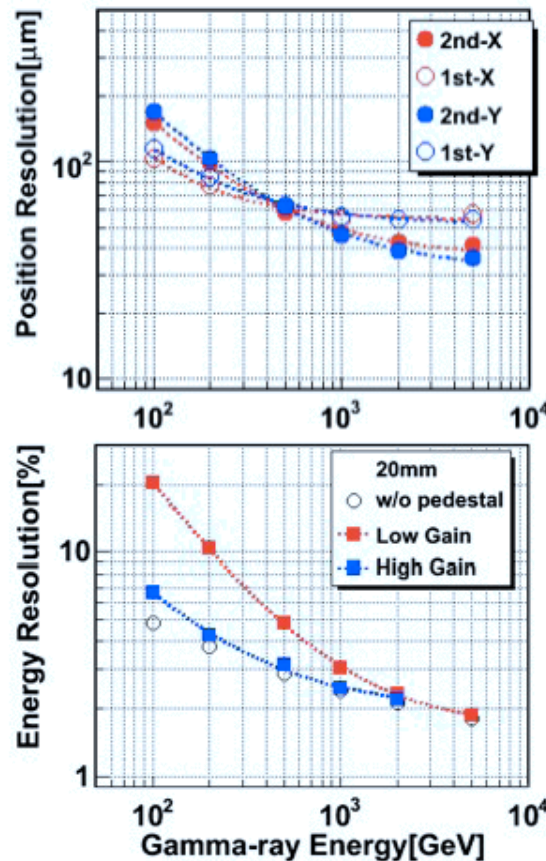


Energy resolution



$\sigma_E/E > 40\%$ because of 1.6λ

EM shower (MC)

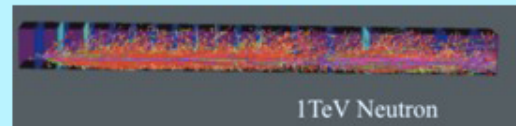


PID technique

400GeV photon

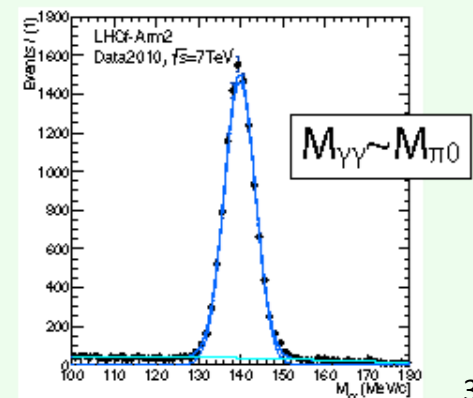


1TeV neutron



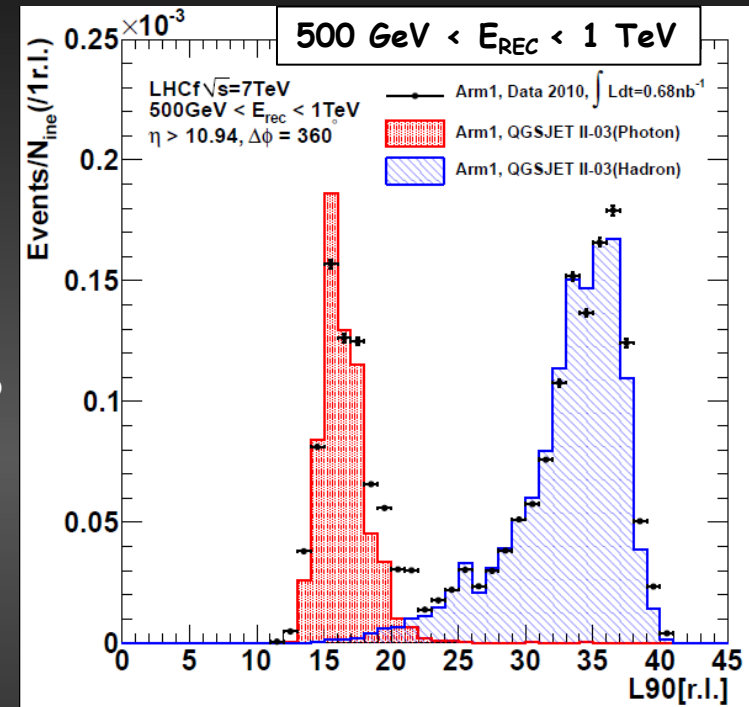
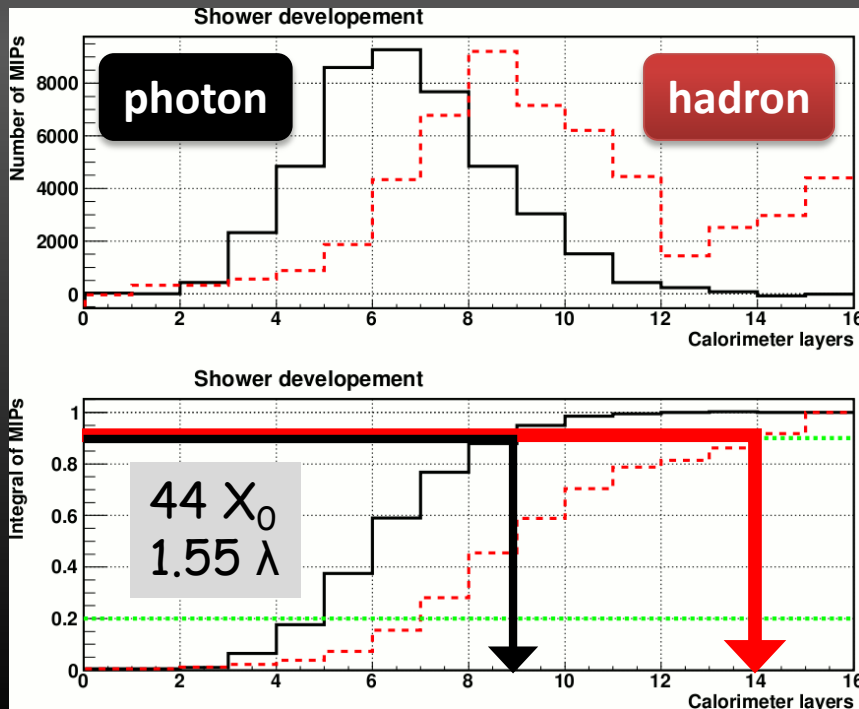
Identification of incoming particle by shower shape

π^0 reconstruction



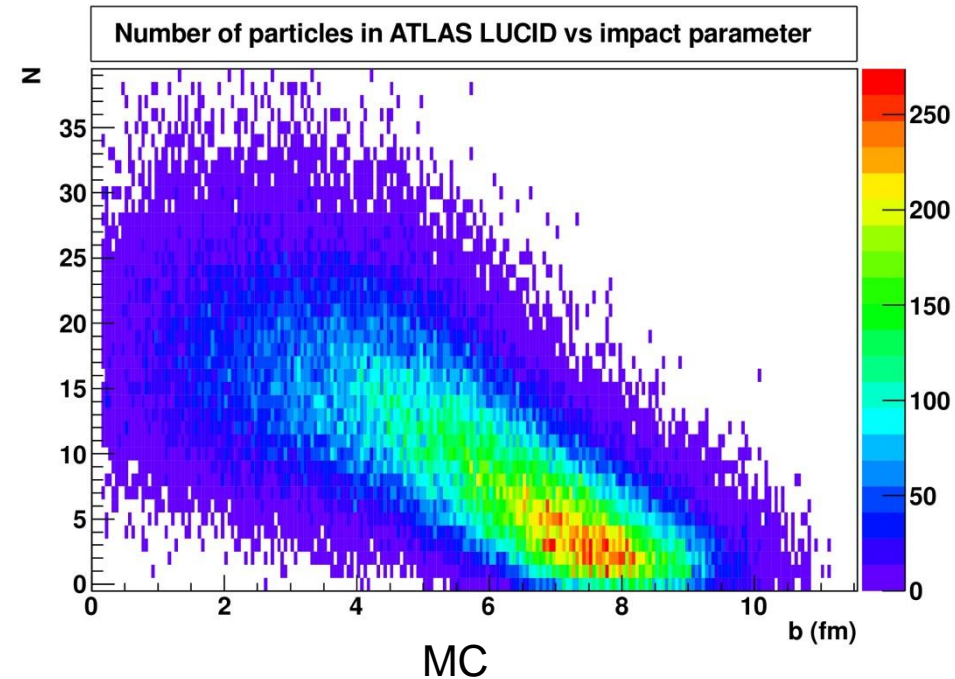
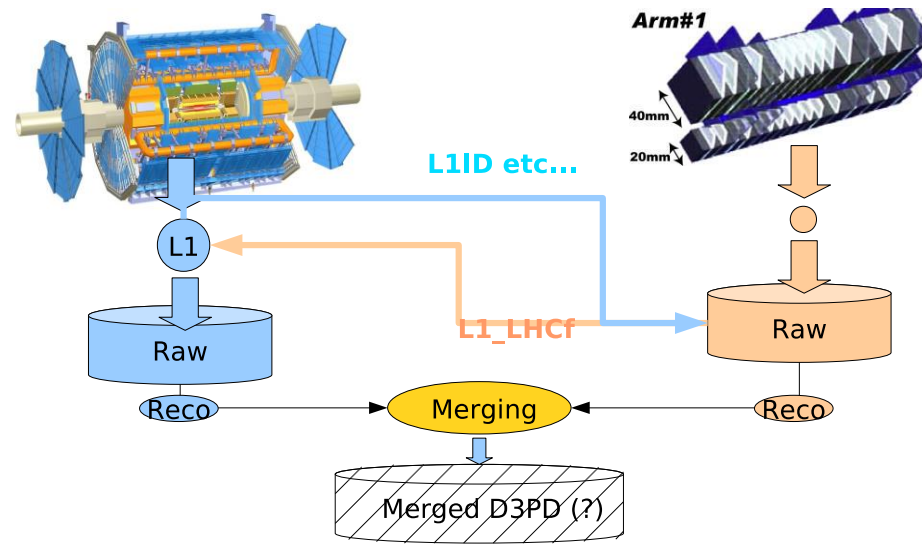
Particle identification

- $L_{90\%}$: longitudinal position containing 90% of the shower energy
- Photon selection based on $L_{90\%}$ cut
- Energy dependent threshold in order to keep constant efficiency $\varepsilon_{PID} = 90\%$
- Purity $P = N_{phot}/(N_{phot}+N_{had})$ estimated by comparison with MC
- Event number in each bin corrected by P/ε_{PID}



- MC photon and hadron events are independently normalized to data
- Comparison done in each energy bin
- LPM effect is switched on

Common trigger with ATLAS



impact parameter vs. # of particles in ATLAS LUCID

- LHCf signal has been used to trigger ATLAS
- Impact parameter may be determined by ATLAS
- Identification of forward-only events