

*IX ATLAS Italia Physics and Upgrade Workshop,  
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Bologna - Dipartimento di Fisica e Astronomia -Italy*

*Physics prospects for  
Susy / BSM  
(Theory)*

*Bologna, 15 January 2014*



**Istituto Nazionale  
di Fisica Nucleare**

*Barbara Mele*

*Sezione di Roma*

# Outline

- **pp collisions: where we stand today**
- **guiding principles to go Beyond Standard Model**
- **SUSY : present status**
  - latest fits on CMSSM and NUHM**
- **extending mono-jet and VBF potential to Natural Compressed scenarios**
- **Outlook**

benefitted a lot from discussions  
with Giacomo Polesello and Tommaso Lari

# pp collisions: where we stand today

- LHC run at 7-8 TeV completed [  $\int L \sim 5 + 20 \text{ fb}^{-1} / \text{exp}$  ]  
( just initial LHC phase ! )  
amazing performance → results well above expectations...
- SM tested at high accuracy in a new  $\sqrt{s}$  range :  
QCD (many regimes), top physics, EW processes, flavor
- “direct” exploration of SM EWSB sector started up with  
observation of a (quite light) Higgs resonance !!!
- still a lot of room for a non-SM EWSB sector
- bounds on new heavy states predicted by many BSM  
models widely extended wrt pre-LHC era
- no real hint of BSM physics !
- SM Hierarchy-Problem solution getting harder...

# guiding principles to go BSM

- **SM** not enough !

beautifully successful at  $E < 1 \text{ TeV}$ ,

but has some “messy features” (flavour sector...),

and does not explain a number of things

(strong CP, neutrino sector, baryogenesis, Dark Matter...)

- crucial issue for Collider Physics (and LHC !) :

what is the expected  
**Energy Threshold ( $E_{\text{TH}}$ ) to go BSM ???**

# What is expected $E_{TH}$ to go BSM ?

- given by requirement of **Naturalness of EWSB scale** !  
quadratic divergences on mass parameters of fundamental scalars drive Higgs mass to the next energy threshold  $E_{TH}$   
(because of absence of symmetry protecting  $m_H$  in SM)  
→ to avoid Fine-Tuning of parameters, one expects roughly
$$E_{TH} \sim m_H / g_{\text{coupling}} \sim 1 \text{ TeV}$$
- this was (before LHC start-up), and **still is (!)**,  
a **ROBUST** statement !!!
- WARNING** : the way  $E_{TH}$  materializes (enters theory) depends on **actual SM extension** (nobody presently knows !)

how has LHC Run I affected this statement ?

# how **LHC Run I** has affected the statement $E_{\text{TH}} \sim m_{\text{H}} / g_{\text{coupling}} \sim 1 \text{ TeV}$

- in the last few decades a lot of theoretical speculations to build models satisfying this criterium ( **SuSy**, **Compositeness**, **Extra Dim.'s**, **Little Higgs**, .... )
- in general **good features** are introduced at the expenses of quite a number of **UNPREDICTED** new parameters
- in **PHENO** studies at colliders, most emphasis given to **SIMPLE** versions of models (preserving basic features) (→ less free parameters)
- after **LHC Run I**, **Simplest Versions of different models look quite Fine-Tuned !**

*theories must build up on experimental facts ....*

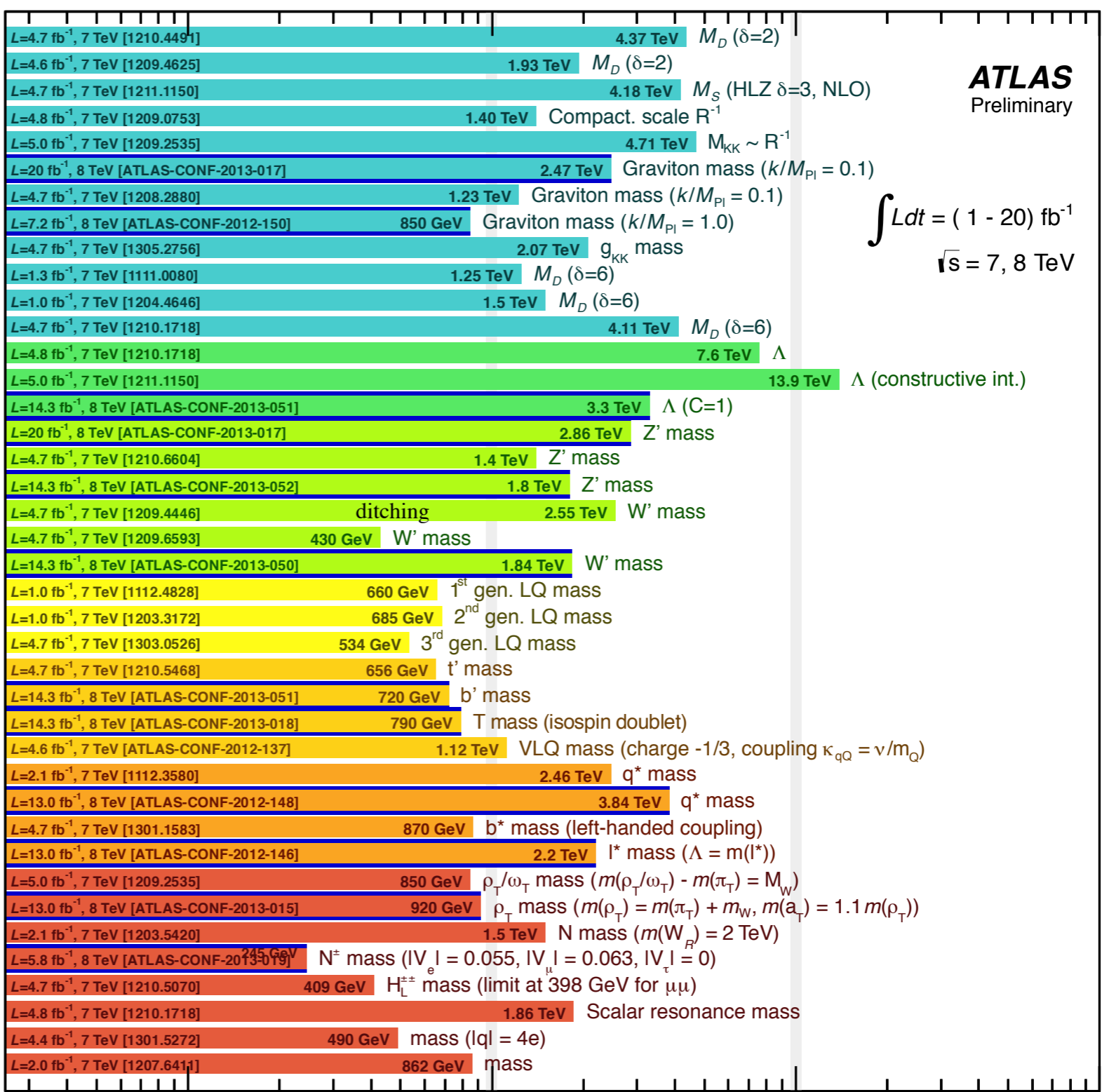
# BSM present status !

# ← how much we learned on EXP side →

how little we know on TH side

- Extra dimensions
  - Large ED (ADD) : mon
  - Large ED (ADD) : diphoton
  - UED : c
  - $S^1/Z_2$
  - R
  - RS1 : WW re
  - Bulk RS : ZZ
  - RS  $g_{KK} \rightarrow t\bar{t}$  (BR=0.925)
  - ADD BH ( $M_{TH}/M_D=3$ ) : SS
  - ADD BH ( $M_{TH}/M_D=3$ ) : le
  - Quantum black ho
- CI
  - qqqq contact in
  - qq
  - uutt CI : SS dilepto
- V
  - Z' (leptophobic topcolor
  - W' (
  - W' (
- LQ
  - Scalar LQ pair ( $\beta=1$ ) : kin.
  - Scalar LQ pair ( $\beta=1$ ) : kin.
  - Scalar LQ pair ( $\beta=1$ ) : kin.
- New quarks
  - 4<sup>th</sup> generati
  - 4th generation : b'b'  $\rightarrow$  SS dilepto
  - Vector-like qu
  - Vector-like
- Excit. ferm.
  - Excited quarks :  $\gamma$ -jet
  - Excited quarks : dijet
  - Excited b quark : W-
  - Excited leptons : l-
- Other
  - Techni-hadrons (LSTC)
  - Techni-hadrons (LSTC) : WZ resc
  - Major. neutr. (LRSM, no mix
  - Heavy lepton  $N^\pm$  (type III seesaw) : Z
  - $H_L^{\pm\pm}$  (DY prod., BR( $H_L^{\pm\pm} \rightarrow ll$ )=1)
  - Color octet scalar : dijet
  - Multi-charged particles (DY prod.) : high
  - Magnetic monopoles (DY prod.) : high

## ATLAS Exotics Searches\* - 95% CL Lower Limits (Status: May 2013)



ATLAS Preliminary

$$\int L dt = (1 - 20) \text{ fb}^{-1}$$

$$\sqrt{s} = 7, 8 \text{ TeV}$$

10<sup>-1</sup> 1 10 10<sup>2</sup> Mass scale [TeV]



today, **SuSy** is still the best candidate  
(among the ones we have thought of)  
to solve all problems  
connected to the TeV scale (and beyond !)

- predicts a light (fundamental) Higgs boson; radiat. EWSB
- stabilizes mass hierarchy;
- weakly coupled theory (coupl.s are known !) : allows accurate and consistent TH predictions even at scales  $\gg$ TeV
  - can in principle be extended up to  $M_{GUT}$ ,  $M_{Pl}$ , and support the desert hypothesis  $\rightarrow$  consistent with GUT
  - delicate impact on EWPT's and FCNC's (as needed by exp's)
- Dark Matter origin as a WIMP

# actually.... two weak points in SuSy

- one on the exp side :  
no susy partner observed in > 30 years of searches;  
present LHC mass bounds on  
squarks and gluinos  $\sim 1-1,7$  TeV ( in CMSSM ! )

- one on the theory side : →

makes implications  
of previous issue  
less dramatic !

vast arbitrariness in construction  
of theoretical models for SuSy breaking  
on which spectrum of SuSy partners  
is crucially based !

# a few robust constraints on mass spectrum :

- in order to stabilize SM mass hierarchy,  
(a few) SuSy-partner masses should be in the  $\mathcal{O}(\text{TeV})$  range
- (SuSy breaking) mass terms in SuSy Lagrangian should not spoil the good convergence properties of SuSy ( $\rightarrow$ soft)!  
 $\rightarrow$   $>100$  new parameters in MSSM (cf.  $\sim 10$  SM mass param's)
- FCNC's imply squarks and sleptons with same quantum #'s be either almost degenerate in mass or almost diagonal in Yukawa matrices !  
 $\rightarrow$  constrains # of free parameters
- many SuSy-breaking models proposed with reduced # of parameters (CMSSM, SuGra, NUHM, GMSB, pMSSM....)  
(none meets all challenges !)

# warning on “fashionable” SUSY models !

- simplicity ( $\rightarrow$  few parameters) not always a good guiding principle !  
[ex: simplicity in the SM would never lead to the observed fermion mass spectrum !]
- changing model (and # of parameter) in general affects pheno at LHC in a non-trivial way (different classes of signatures) !

( nicely illustrated by the  
ATLAS SUSY Summary plot  $\rightarrow$  )

# ATLAS SUSY Searches\* - 95% CL Lower Limits

Status: SUSY 2013

ATLAS Preliminary

$$\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1} \quad \sqrt{s} = 7, 8 \text{ TeV}$$

Model	$e, \mu, \tau, \gamma$	Jets	$E_T^{\text{miss}}$	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Mass limit	Reference	
Inclusive Searches	MSUGRA/CMSSM	0	2-6 jets	Yes	20.3	$\tilde{q}, \tilde{g}$ 1.7 TeV	$m(\tilde{q})=m(\tilde{g})$ ATLAS-CONF-2013-047
	MSUGRA/CMSSM	1 $e, \mu$	3-6 jets	Yes	20.3	$\tilde{g}$ 1.2 TeV	any $m(\tilde{q})$ ATLAS-CONF-2013-062
	MSUGRA/CMSSM	0	7-10 jets	Yes	20.3	$\tilde{g}$ 1.1 TeV	any $m(\tilde{q})$ 1308.1841
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	$\tilde{q}$ 740 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-047
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	$\tilde{g}$ 1.3 TeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-047
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^\pm \rightarrow qqW^\pm\tilde{\chi}_1^0$	1 $e, \mu$	3-6 jets	Yes	20.3	$\tilde{g}$ 1.18 TeV	$m(\tilde{\chi}_1^0)<200 \text{ GeV}, m(\tilde{\chi}^\pm)=0.5(m(\tilde{\chi}_1^0)+m(\tilde{g}))$ ATLAS-CONF-2013-062
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell/\ell\nu/\nu\nu)\tilde{\chi}_1^0$	2 $e, \mu$	0-3 jets	-	20.3	$\tilde{g}$ 1.12 TeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-089
	GMSB ( $\tilde{\ell}$ NLSP)	2 $e, \mu$	2-4 jets	Yes	4.7	$\tilde{g}$ 1.24 TeV	$\tan\beta<15$ 1208.4688
	GMSB ( $\tilde{\ell}$ NLSP)	1-2 $\tau$	0-2 jets	Yes	20.7	$\tilde{g}$ 1.4 TeV	$\tan\beta > 18$ ATLAS-CONF-2013-026
	GGM (bino NLSP)	2 $\gamma$	-	Yes	4.8	$\tilde{g}$ 1.07 TeV	$m(\tilde{\chi}_1^0)>50 \text{ GeV}$ 1209.0753
	GGM (wino NLSP)	1 $e, \mu + \gamma$	-	Yes	4.8	$\tilde{g}$ 619 GeV	$m(\tilde{\chi}_1^0)>50 \text{ GeV}$ ATLAS-CONF-2012-144
	GGM (higgsino-bino NLSP)	$\gamma$	1 $b$	Yes	4.8	$\tilde{g}$ 900 GeV	$m(\tilde{\chi}_1^0)>220 \text{ GeV}$ 1211.1167
	GGM (higgsino NLSP)	2 $e, \mu (Z)$	0-3 jets	Yes	5.8	$\tilde{g}$ 690 GeV	$m(\tilde{H})>200 \text{ GeV}$ ATLAS-CONF-2012-152
Gravitino LSP	0	mono-jet	Yes	10.5	$F^{1/2}$ scale 645 GeV	$m(\tilde{g})>10^{-4} \text{ eV}$ ATLAS-CONF-2012-147	
3 <sup>rd</sup> gen. $\tilde{g}$ med.	$\tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$	0	3 $b$	Yes	20.1	$\tilde{g}$ 1.2 TeV	$m(\tilde{\chi}_1^0)<600 \text{ GeV}$ ATLAS-CONF-2013-061
	$\tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0	7-10 jets	Yes	20.3	$\tilde{g}$ 1.1 TeV	$m(\tilde{\chi}_1^0) < 350 \text{ GeV}$ 1308.1841
	$\tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^\pm$	0-1 $e, \mu$	3 $b$	Yes	20.1	$\tilde{g}$ 1.34 TeV	$m(\tilde{\chi}_1^0)<400 \text{ GeV}$ ATLAS-CONF-2013-061
	$\tilde{g} \rightarrow b\tilde{t}\tilde{\chi}_1^\pm$	0-1 $e, \mu$	3 $b$	Yes	20.1	$\tilde{g}$ 1.3 TeV	$m(\tilde{\chi}_1^0)<300 \text{ GeV}$ ATLAS-CONF-2013-061
	3 <sup>rd</sup> gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^\pm$	0	2 $b$	Yes	20.1	$\tilde{b}_1$ 100-620 GeV
$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{\chi}_1^\pm$		2 $e, \mu$ (SS)	0-3 $b$	Yes	20.7	$\tilde{b}_1$ 275-430 GeV	$m(\tilde{\chi}_1^\pm)=2 m(\tilde{\chi}_1^0)$ ATLAS-CONF-2013-007
$\tilde{t}_1\tilde{t}_1$ (light), $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$		1-2 $e, \mu$	1-2 $b$	Yes	4.7	$\tilde{t}_1$ 110-167 GeV	$m(\tilde{\chi}_1^0)=55 \text{ GeV}$ 1208.4305, 1209.2102
$\tilde{t}_1\tilde{t}_1$ (light), $\tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$		2 $e, \mu$	0-2 jets	Yes	20.3	$\tilde{t}_1$ 130-220 GeV	$m(\tilde{\chi}_1^0)=m(\tilde{t}_1)-m(W)-50 \text{ GeV}, m(\tilde{t}_1)<m(\tilde{\chi}_1^\pm)$ ATLAS-CONF-2013-048
$\tilde{t}_1\tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^\pm$		2 $e, \mu$	2 jets	Yes	20.3	$\tilde{t}_1$ 225-525 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-065
$\tilde{t}_1\tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$		0	2 $b$	Yes	20.1	$\tilde{t}_1$ 150-580 GeV	$m(\tilde{\chi}_1^0)<200 \text{ GeV}, m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0)=5 \text{ GeV}$ 1308.2631
$\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$		1 $e, \mu$	1 $b$	Yes	20.7	$\tilde{t}_1$ 200-610 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-037
$\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^\pm$		0	2 $b$	Yes	20.5	$\tilde{t}_1$ 320-660 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-024
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$		0	mono-jet/c-tag	Yes	20.3	$\tilde{t}_1$ 90-200 GeV	$m(\tilde{t}_1)-m(\tilde{\chi}_1^0)<85 \text{ GeV}$ ATLAS-CONF-2013-068
$\tilde{t}_1\tilde{t}_1$ (natural GMSB)		2 $e, \mu (Z)$	1 $b$	Yes	20.7	$\tilde{t}_1$ 500 GeV	$m(\tilde{\chi}_1^0)>150 \text{ GeV}$ ATLAS-CONF-2013-025
$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$		3 $e, \mu (Z)$	1 $b$	Yes	20.7	$\tilde{t}_2$ 271-520 GeV	$m(\tilde{t}_1)=m(\tilde{\chi}_1^0)+180 \text{ GeV}$ ATLAS-CONF-2013-025
EW direct	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \tilde{\chi}_1^0$	2 $e, \mu$	0	Yes	20.3	$\tilde{\ell}$ 85-315 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-049
	$\tilde{\chi}_1^+\tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow \tilde{\ell}\nu(\ell\tilde{\nu})$	2 $e, \mu$	0	Yes	20.3	$\tilde{\chi}_1^\pm$ 125-450 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$ ATLAS-CONF-2013-049
	$\tilde{\chi}_1^+\tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow \tilde{\tau}\nu(\tau\tilde{\nu})$	2 $\tau$	-	Yes	20.7	$\tilde{\chi}_1^\pm$ 180-330 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}, m(\tilde{\tau}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$ ATLAS-CONF-2013-028
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow \tilde{\ell}_L\nu\tilde{\ell}_L(\tilde{\nu}\nu), \ell\tilde{\nu}\tilde{\ell}_L(\tilde{\nu}\nu)$	3 $e, \mu$	0	Yes	20.7	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$ 600 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$ ATLAS-CONF-2013-035
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 Z\tilde{\chi}_1^0$	3 $e, \mu$	0	Yes	20.7	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$ 315 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, \text{ sleptons decoupled}$ ATLAS-CONF-2013-035
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 h\tilde{\chi}_1^0$	1 $e, \mu$	2 $b$	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$ 285 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, \text{ sleptons decoupled}$ ATLAS-CONF-2013-093
	Long-lived particles	Direct $\tilde{\chi}_1^+\tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet	Yes	20.3	$\tilde{\chi}_1^\pm$ 270 GeV
Stable, stopped $\tilde{g}$ R-hadron		0	1-5 jets	Yes	22.9	$\tilde{g}$ 832 GeV	$m(\tilde{\chi}_1^0)=100 \text{ GeV}, 10 \mu\text{s}<\tau(\tilde{g})<1000 \text{ s}$ ATLAS-CONF-2013-057
GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu})+\tau(e, \mu)$		1-2 $\mu$	-	-	15.9	$\tilde{\chi}_1^0$ 475 GeV	$10<\tan\beta<50$ ATLAS-CONF-2013-058
GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$ , long-lived $\tilde{\chi}_1^0$		2 $\gamma$	-	Yes	4.7	$\tilde{\chi}_1^0$ 230 GeV	$0.4<\tau(\tilde{\chi}_1^0)<2 \text{ ns}$ 1304.6310
$\tilde{q}\tilde{q}, \tilde{\chi}_1^0 \rightarrow qq\mu$ (RPV)		1 $\mu$ , displ. vtx	-	-	20.3	$\tilde{q}$ 1.0 TeV	$1.5<c\tau<156 \text{ mm}, \text{BR}(\mu)=1, m(\tilde{\chi}_1^0)=108 \text{ GeV}$ ATLAS-CONF-2013-092
RPV	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e + \mu$	2 $e, \mu$	-	-	4.6	$\tilde{\nu}_\tau$ 1.61 TeV	$\lambda'_{311}=0.10, \lambda_{132}=0.05$ 1212.1272
	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e(\mu) + \tau$	1 $e, \mu + \tau$	-	-	4.6	$\tilde{\nu}_\tau$ 1.1 TeV	$\lambda'_{311}=0.10, \lambda_{1(2)33}=0.05$ 1212.1272
	Bilinear RPV CMSSM	1 $e, \mu$	7 jets	Yes	4.7	$\tilde{q}, \tilde{g}$ 1.2 TeV	$m(\tilde{q})=m(\tilde{g}), c\tau_{\text{LSP}}<1 \text{ mm}$ ATLAS-CONF-2012-140
	$\tilde{\chi}_1^+\tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow ee\tilde{\nu}_\mu, e\mu\tilde{\nu}_e$	4 $e, \mu$	-	Yes	20.7	$\tilde{\chi}_1^\pm$ 760 GeV	$m(\tilde{\chi}_1^0)>300 \text{ GeV}, \lambda_{121}>0$ ATLAS-CONF-2013-036
	$\tilde{\chi}_1^+\tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \tau\tau\tilde{\nu}_e, e\tau\tilde{\nu}_\tau$	3 $e, \mu + \tau$	-	Yes	20.7	$\tilde{\chi}_1^\pm$ 350 GeV	$m(\tilde{\chi}_1^0)>80 \text{ GeV}, \lambda_{133}>0$ ATLAS-CONF-2013-036
	$\tilde{g} \rightarrow qq\tilde{q}$	0	6-7 jets	-	20.3	$\tilde{g}$ 916 GeV	$\text{BR}(t)=\text{BR}(b)=\text{BR}(c)=0\%$ ATLAS-CONF-2013-091
	$\tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow bs$	2 $e, \mu$ (SS)	0-3 $b$	Yes	20.7	$\tilde{g}$ 880 GeV	ATLAS-CONF-2013-007
Other	Scalar gluon pair, $sgluon \rightarrow q\tilde{q}$	0	4 jets	-	4.6	$sgluon$ 100-287 GeV	incl. limit from 1110.2693 1210.4826
	Scalar gluon pair, $sgluon \rightarrow t\tilde{t}$	2 $e, \mu$ (SS)	1 $b$	Yes	14.3	$sgluon$ 800 GeV	ATLAS-CONF-2013-051
	WIMP interaction (D5, Dirac $\chi$ )	0	mono-jet	Yes	10.5	$M^*$ scale 704 GeV	$m(\chi)<80 \text{ GeV}, \text{limit of } <687 \text{ GeV for D8}$ ATLAS-CONF-2012-147

$\sqrt{s} = 7 \text{ TeV}$  full data  
 $\sqrt{s} = 8 \text{ TeV}$  partial data  
 $\sqrt{s} = 8 \text{ TeV}$  full data

10<sup>-1</sup> 1 Mass scale [TeV]

\*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 $\sigma$  theoretical signal cross section uncertainty.

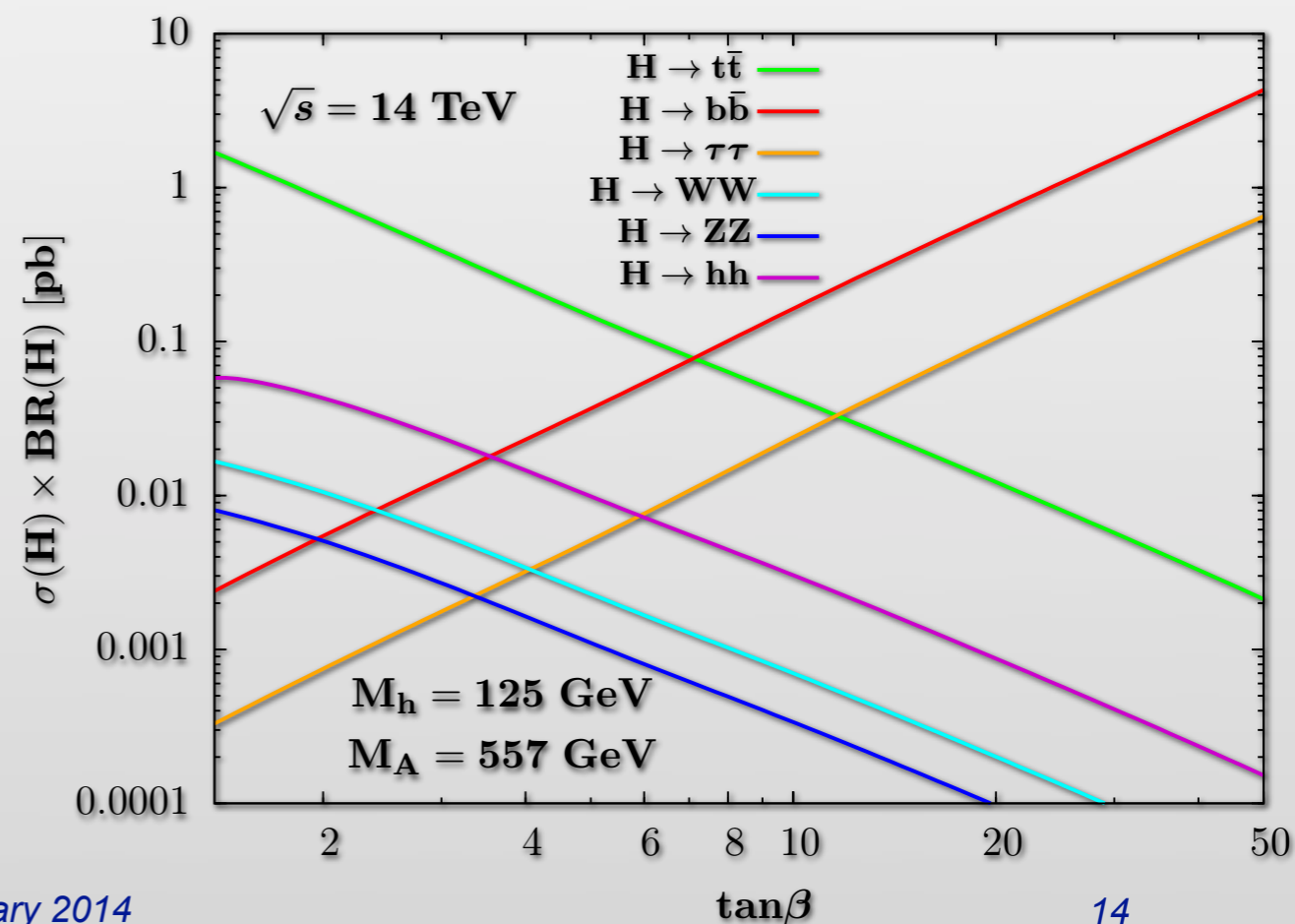
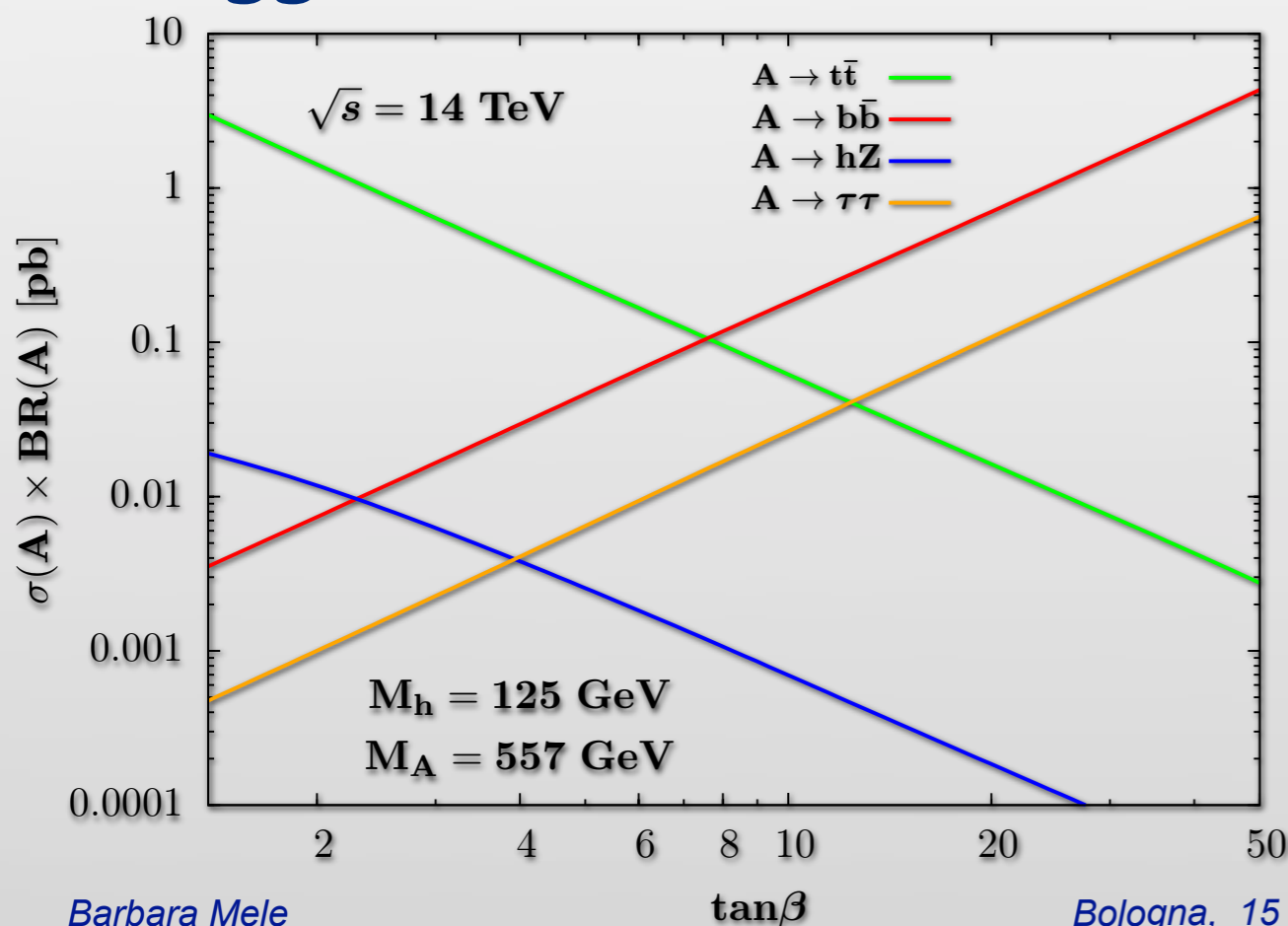
# Susy predicts extended Higgs sector !

2 Higgs doublets in MSSM  $\rightarrow$  5 physical fields  
 $h, H, A, H^+, H^-$

best fit for  $m_h \sim 125$  GeV and measured  $h$  signal strengths :  
 $M_H \sim 580$  GeV,  $M_A \sim M_{H^+} \sim 560$  GeV,  $\tan\beta \sim 1$

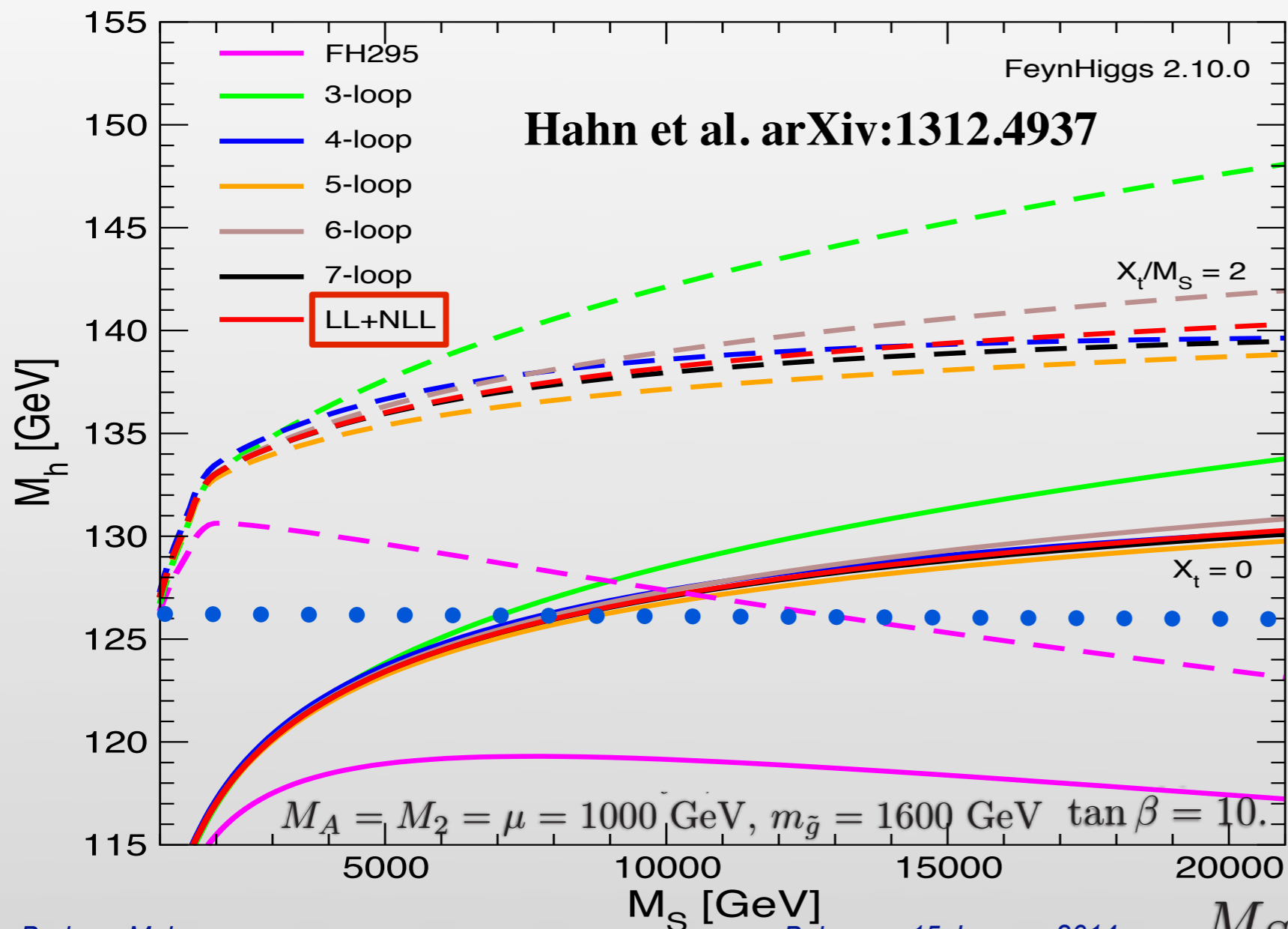
Djouadi et al, arxiv:1307.5205

$gg, bb \rightarrow A, H \rightarrow t\bar{t}$  could be accessible in next run !



# news on Susy Higgs-mass calculators

- new version of FeynHiggs (2.10.0) including resummation of LL and NLL in  $\log(m_{\text{stop}}/m_{\text{top}})$  through 2-loop RGEs
- achieves higher accuracy at large  $m_{\text{stop}}$



$\Delta M_h \sim + 2$  (5) GeV  
 at  $m_{\text{stop}} \sim 2$  (5) TeV !

**FeynHiggs 2.10.0**

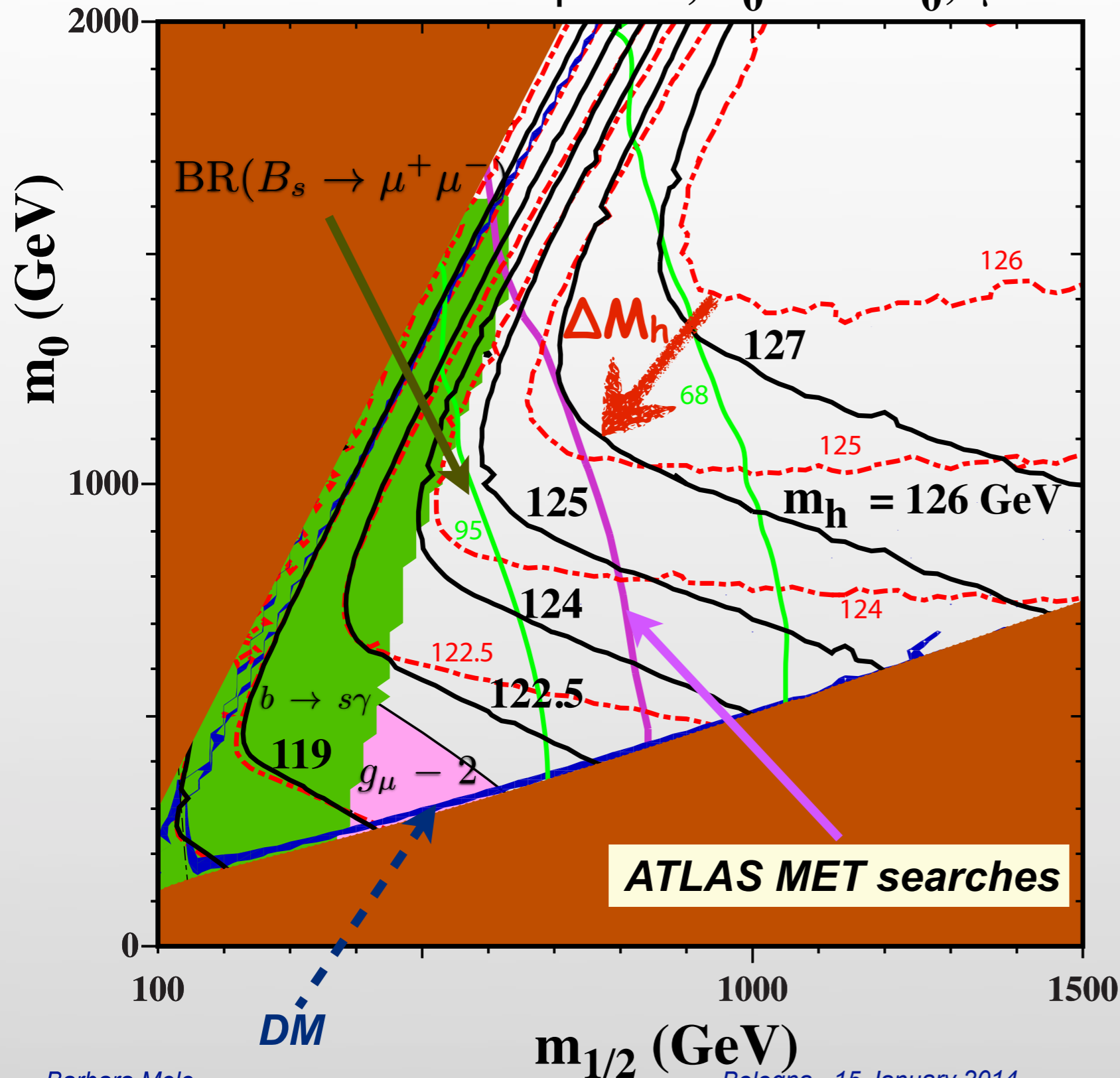
$M_h(\text{max}) \rightarrow m_{\text{stop}}(\text{max})$

**fixed order up to 2-loops**

# upward shift in $M_h$ relaxes constraints on CMSSM (and NUHM1)

Buchmueller et al. arXiv:1312.5233

$\tan \beta = 30, A_0 = 2.5m_0, \mu > 0$



improved consistency with all constraints !

large  $M_h$  contour differences at high  $m_{1/2}$

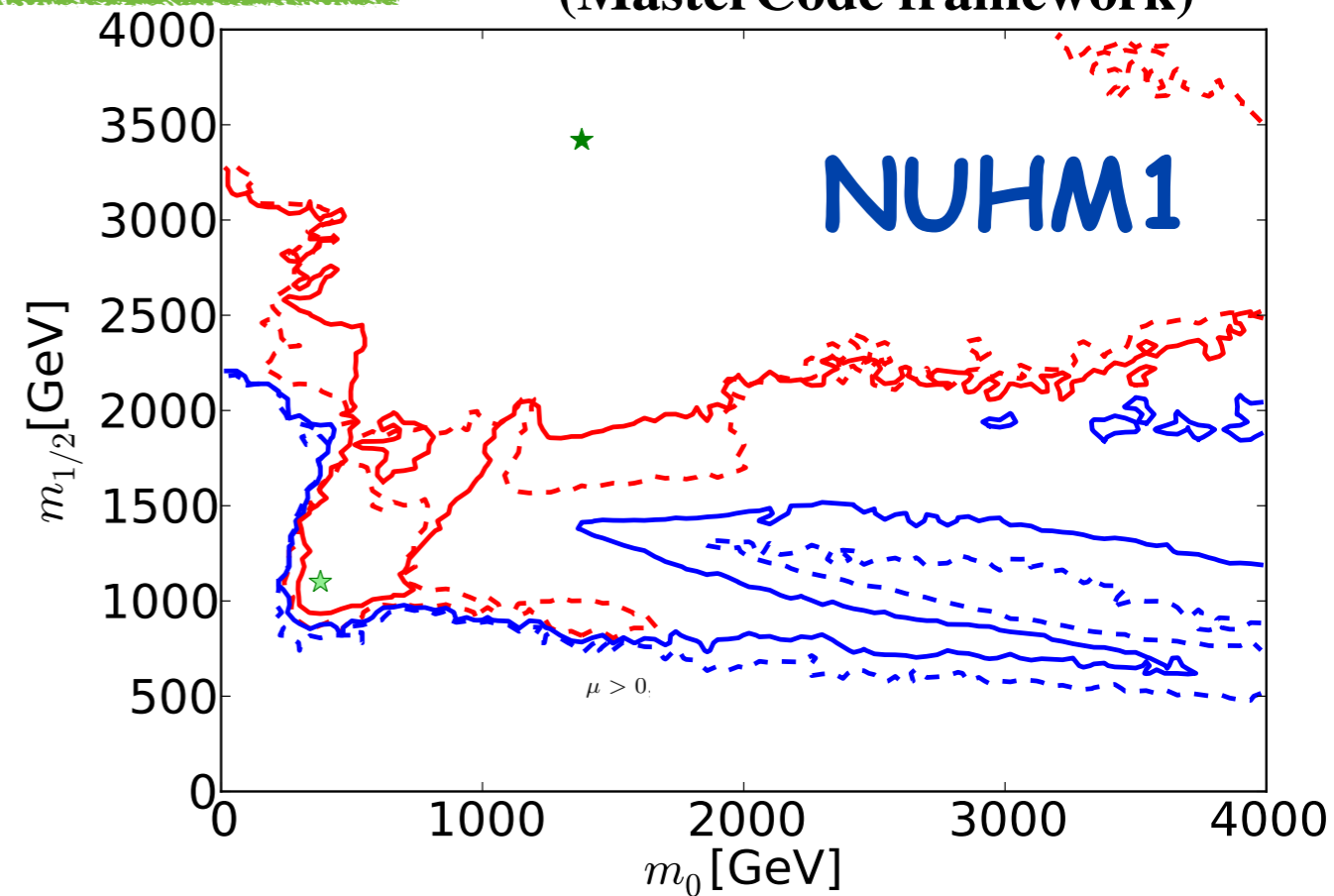
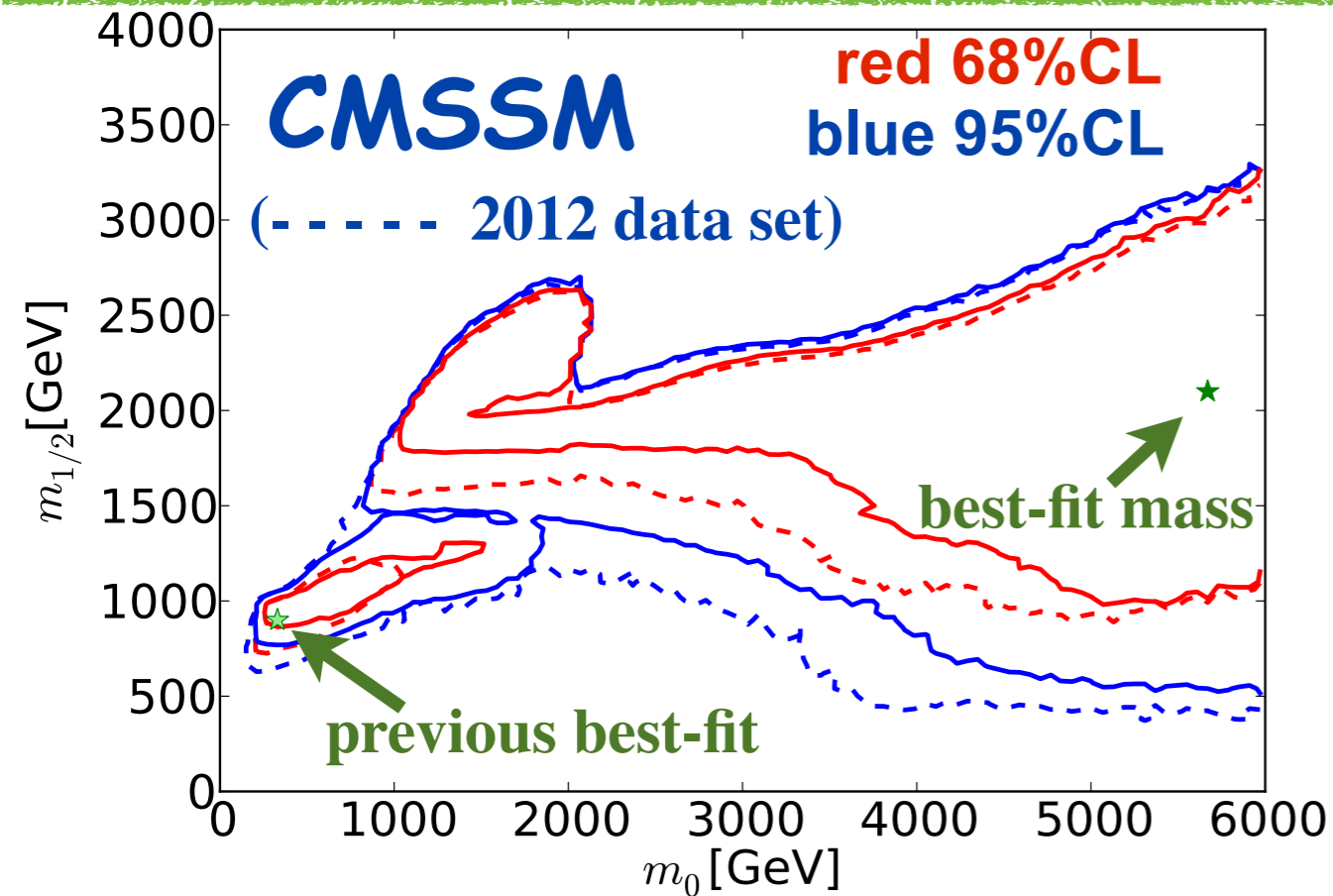
$\tan \beta < 10$  disfavored by  $M_h$  in CMSSM (not in NUHM) !



# new global fits in CMSSM and NUHM1

ATLAS  $E_{\text{miss}}^T$ , BR(B)'s, Mh, EWPT, DM density, LUX,  $g\mu-2$

Buchmueller et al. arXiv:1312.5250  
(MasterCode framework)



*best-fit points*

*overall likelihood function quite flat and precise location not very significant*

Model	Region	Minimum $\chi^2$	$m_0$ (GeV)	$m_{1/2}$ (GeV)	$\tan \beta$
$\mu > 0$	Low-mass	35.8	670	1040	21
	High-mass	35.1	5650	2100	51
NUHM1	Low-mass	33.3	470	1270	11
	High-mass	32.7	1380	3420	39

see also: Bechtle et al arXiv:1310.3045  
FITTINO framework

arXiv:1312.5426

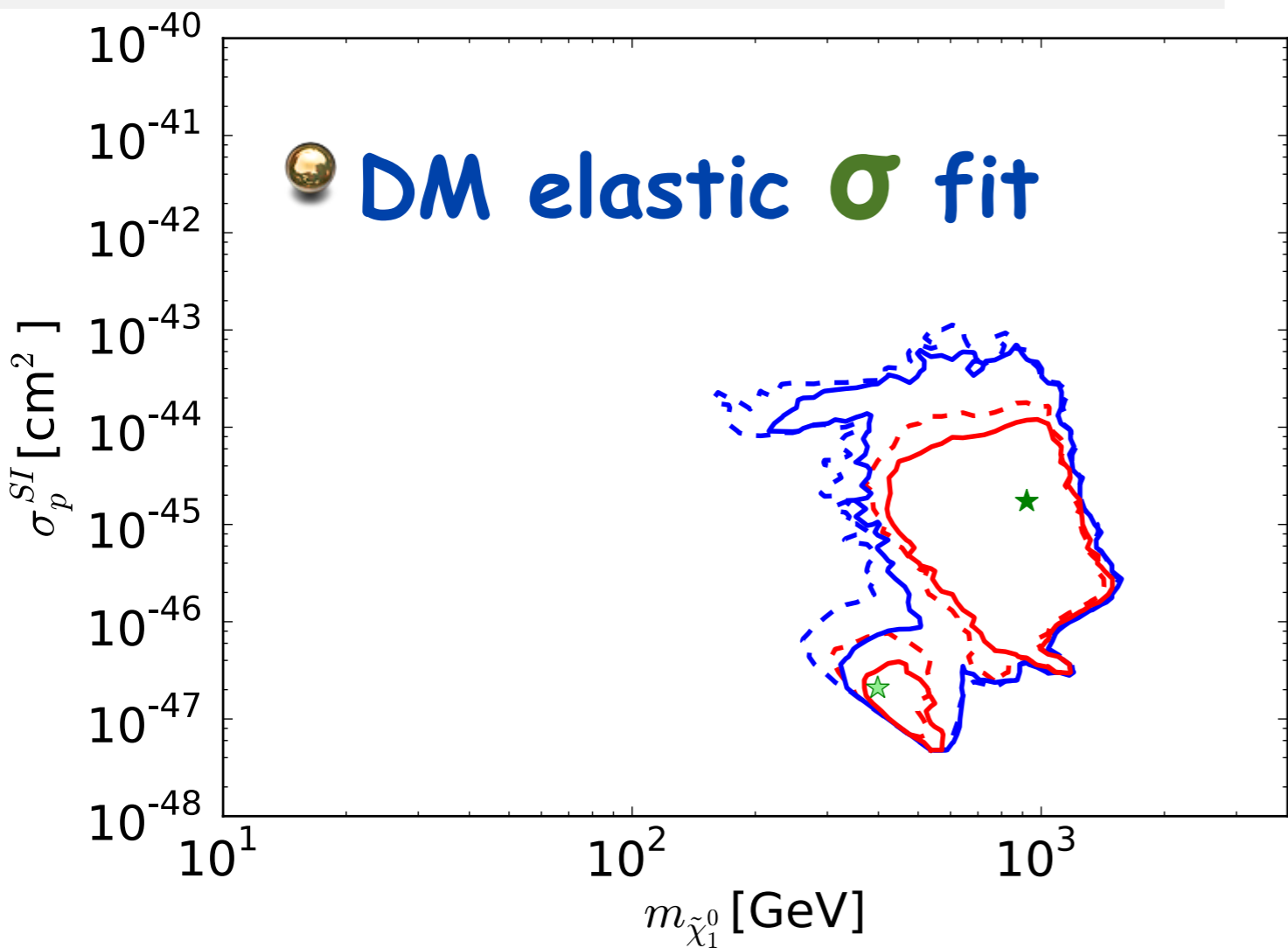
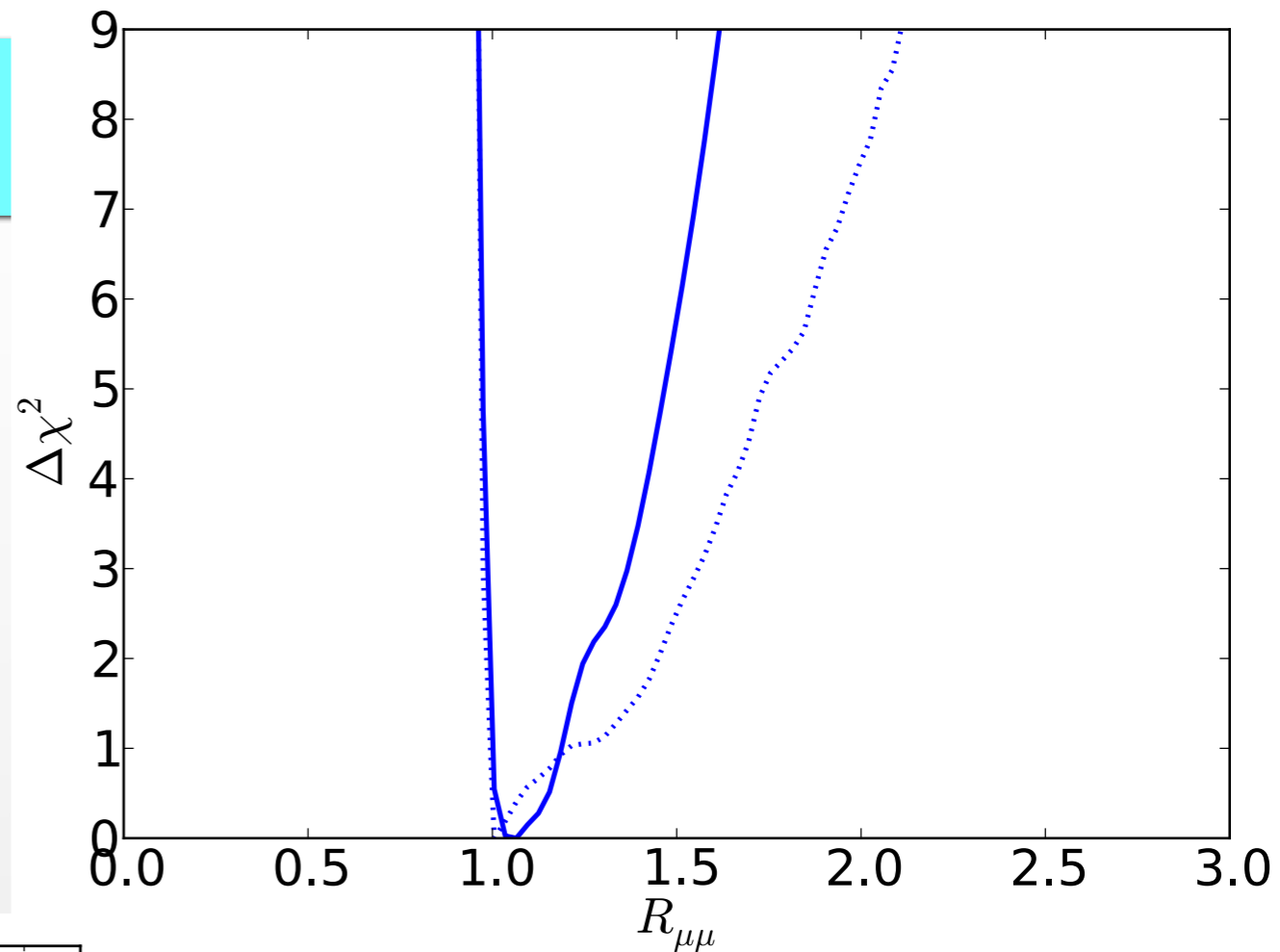
# contribution of most relevant observables to global $\chi^2$ at best-fit points at high and low masses

arXiv:1312.5250  
(MasterCode framework)

Observable	$\Delta\chi^2$ CMSSM $\mu > 0$ (high)	$\Delta\chi^2$ CMSSM $\mu > 0$ (low)	$\Delta\chi^2$ CMSSM $\mu < 0$ (high)	$\Delta\chi^2$ CMSSM $\mu < 0$ (low)	$\Delta\chi^2$ NUHM1 $\mu > 0$ (high)	$\Delta\chi^2$ NUHM1 $\mu > 0$ (low)	$\Delta\chi^2$ Standard Model
Global	35.1	35.8	36.6	38.9	32.7	33.3	36.5
$\text{BR}_{b \rightarrow s\gamma}^{\text{exp}/\text{SM}}$	0.52	1.58	0.37	0.00	0.54	0.02	0.57
$\text{BR}_{B \rightarrow \tau\nu}^{\text{exp}/\text{SM}}$	1.77	1.63	1.63	1.61	1.65	1.66	1.60
$\epsilon_K$	1.94	1.88	1.94	1.87	1.94	1.94	1.96
$a_\mu^{\text{exp}} - a_\mu^{\text{SM}}$	10.71	9.34	11.42	12.65	10.50	9.63	11.19
$M_W$	1.35	0.22	2.15	0.04	0.00	0.11	1.38
$M_h$	0.00	0.04	0.03	0.53	0.00	0.22	(1.5)
$R_\ell$	1.10	1.04	1.10	1.00	1.07	1.00	1.09
$A_{\text{fb}}(b)$	6.56	6.79	6.05	7.61	5.45	6.93	6.58
$A_\ell(\text{SLD})$	3.59	3.40	3.99	2.81	4.59	3.30	3.55
$\sigma_{\text{had}}^0$	2.52	2.55	2.56	2.51	2.59	2.56	2.54
LUX	0.03	0.07	0.66	0.07	0.00	0.07	-
ATLAS 20/fb	0.04	2.52	0.02	3.35	0.02	1.15	-
$B_{s,d} \rightarrow \mu^+ \mu^-$	0.51	0.46	0.13	0.11	0.22	0.35	0.15

global  $\chi^2$  functions vary slowly over most of param. space allowed by  $M_h$   
and  $E^{\text{T}}_{\text{miss}} + \text{jets}$  searches, with best-fit  $\chi^2/\text{dof}$  comparable to SM one !

# CMSSM

 $\mu > 0$ 

 **BR( $B_{s,d} \rightarrow \mu\mu$ )/SM fit**

**Buchmuller et al., arXiv:1312.5250**

**(MasterCode framework)**

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# 95% CL lower limits on masses

Buchmueller et al., arXiv:1312.5250

(MasterCode framework)

Sparticle	CMSSM $\mu > 0$	CMSSM $\mu < 0$	NUHM $\mu > 0$
$\tilde{g}$	1810	(2100) (3200) 3540	1920
$\tilde{q}_R$	1620	(1900) 6300	1710
$\tilde{t}_1$	750	(950) 4100	(650) 1120
$\tilde{\tau}_1$	340	(400) 4930	380
$M_A$	690	(1900) 3930	450

● large part of the preferred CMSSM and NUHM1 parameter regions accessible at future LHC runs !

# $E_{\text{miss}}$ searches strongly constrain **universal** soft Susy-breaking masses !

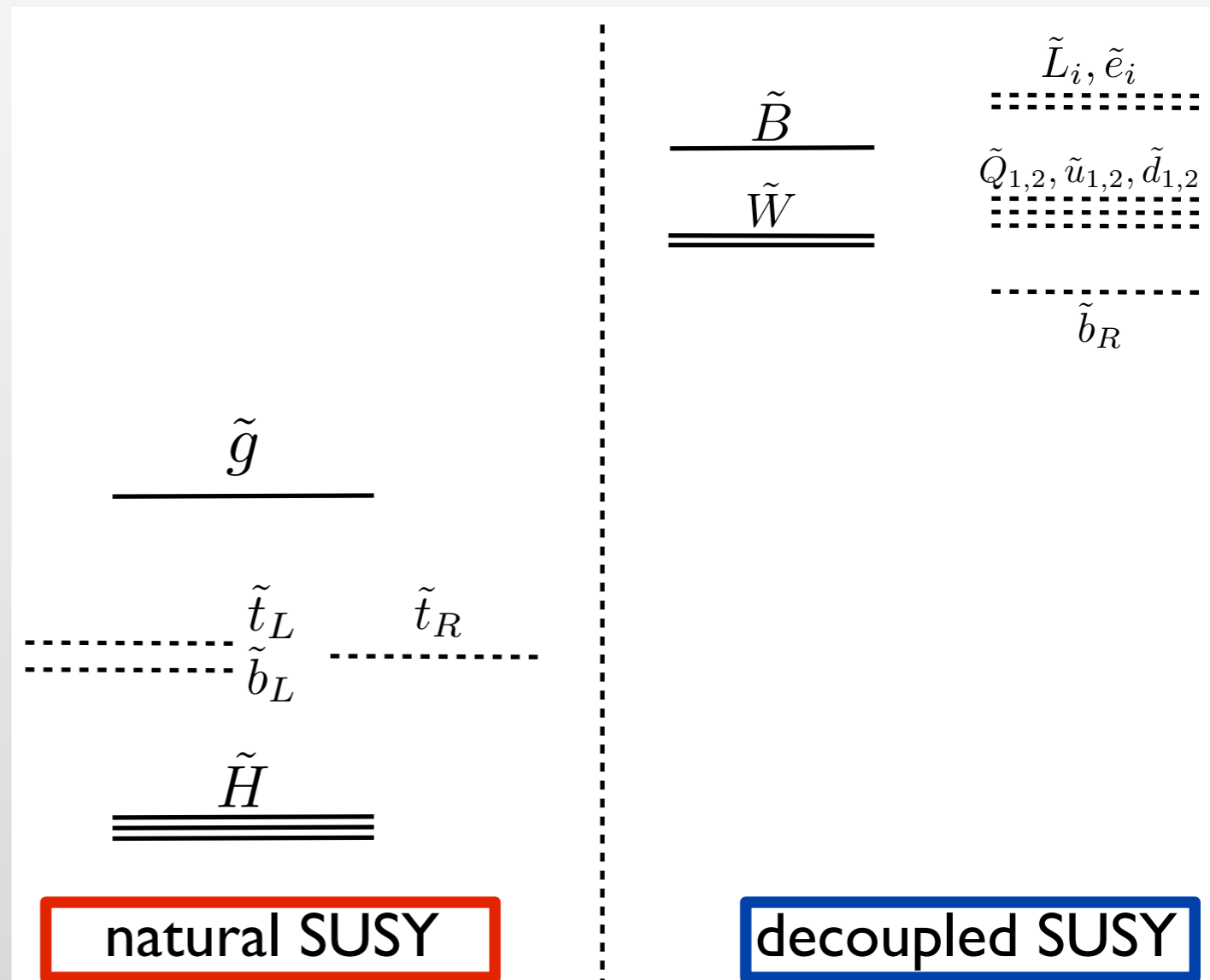
- high priority to Natural Scenarios' searches !
- light stop quarks (but also light higgsinos and gluinos) have a privileged role in stabilizing the Higgs mass
- bounds on  $m_{\text{stop}}$  much softer than ones on  $m_{\text{light-flavor}}$

"Natural Scenario"

$m_{\text{stop}}, m_{\text{sbottom}} < 1 \text{ TeV}$

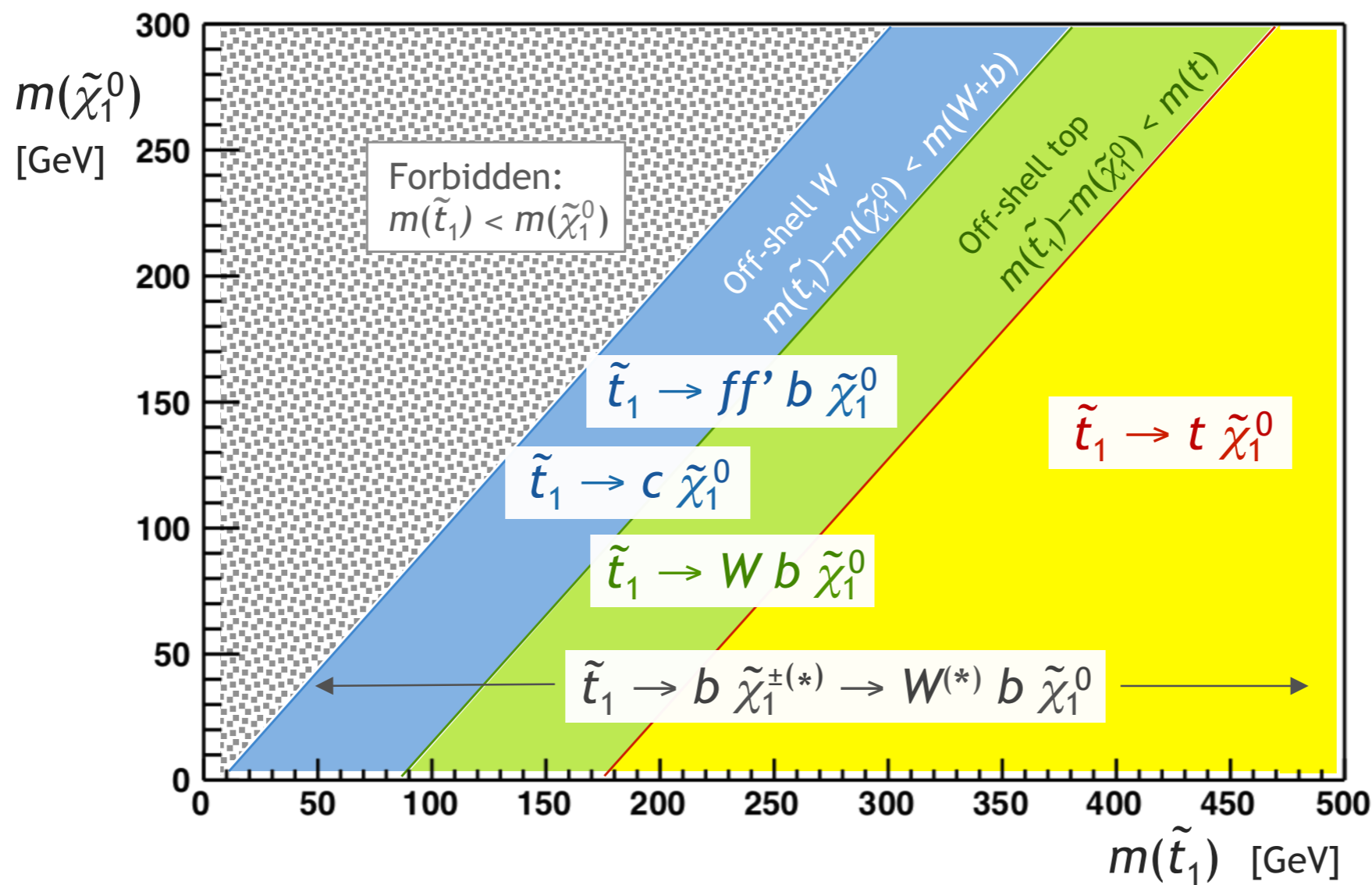
$m_{\text{gluino}} < 2 \text{ TeV}$

$m_{\text{higgsino}} \sim 100\text{-}300 \text{ GeV}$



# focus on stop pair production

- for **light stop**, **large cross sections**  
but  $E_{\text{visible}}$  and  $E_{\text{invisible}}$  released tends to be **smaller**  $\rightarrow$   
signal can be **missed** in searches tailored to more generic spectra
- huge background from **top pairs**
- different **mass hierarchies** require different search strategies
- small mass gaps** can give soft (untriggered)  $E_{(\text{in})\text{visible}}$  !!!



also :

$$\tilde{b}_1 \rightarrow b \tilde{\chi}_1^0$$

$$\tilde{b}_1 \rightarrow t \tilde{\chi}_1^{\pm}$$

$$\tilde{b}_1 \rightarrow b \tilde{\chi}_2^0 \rightarrow b h(Z) \tilde{\chi}_1^0$$

# bounds tend to evaporate for compressed spectra

$$m_{\tilde{t}} \cong m_t + m_{\tilde{\chi}_1^0}$$

$$m_{\tilde{t}} \sim 190 - 300 \text{ GeV}$$

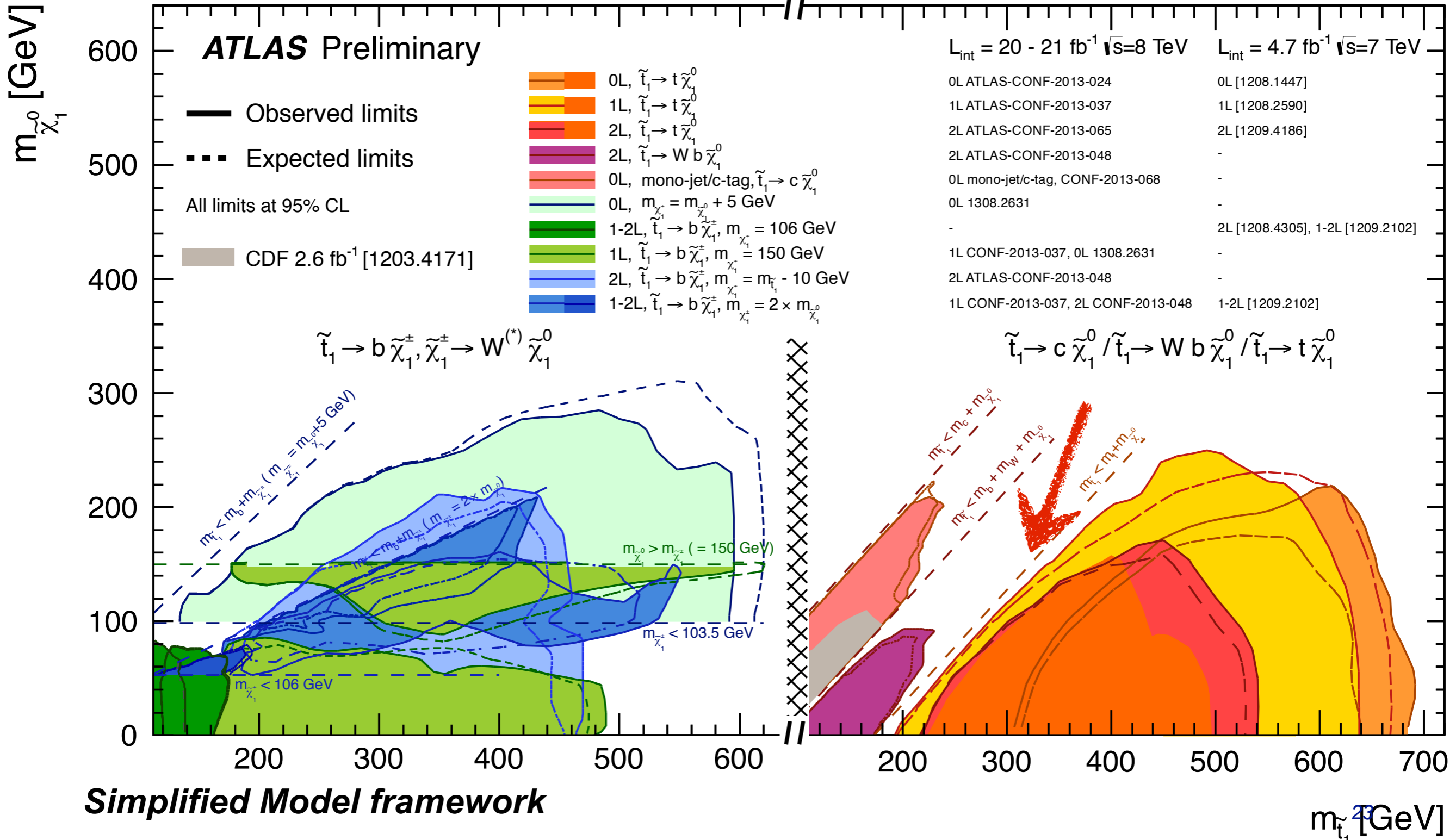
$$\Delta M = m_{\tilde{t}} - (m_t + m_{\tilde{\chi}_1^0}) \sim 5 - 10 \text{ GeV}$$

$$m_{\tilde{t}} \sim 300 - 450 \text{ GeV}$$

$$\Delta M \sim 15 - 20 \text{ GeV}$$

Status: SUSY 2013

$\tilde{t}_1, \tilde{t}_1$  production



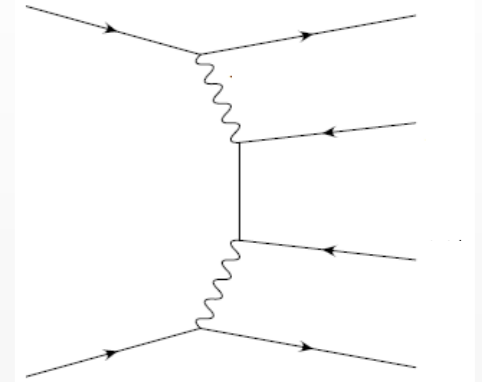
# compressed spectra : recent developments

- small mass gaps give softer  $E_T^{\text{miss}}$  and  $E_T^{\ell,j}$ ,  
reducing search sensitivity
  - boost typical  $E_T^i$  above the reconstruction threshold  
by considering same final state  
either in VBF topology or in mono-jet (mono- $\gamma$ ) events
    - trigger on extra jets (photon)
- two examples :
  - compressed-stop pairs with VBF tagging  
Dutta et al., arXiv:1312.1348 (1304.7779 , 1210.0964)
  - compressed charginos/neutralinos in mono-jet  
(mono- $\gamma$ ) events
    - Schwaller and Zurita, arXiv:1312.7350
    - Baer, Mustafayev, Tata, arXiv:1401.1162
    - Han et al., arXiv:1310.4274



# compressed stop in VBF topology

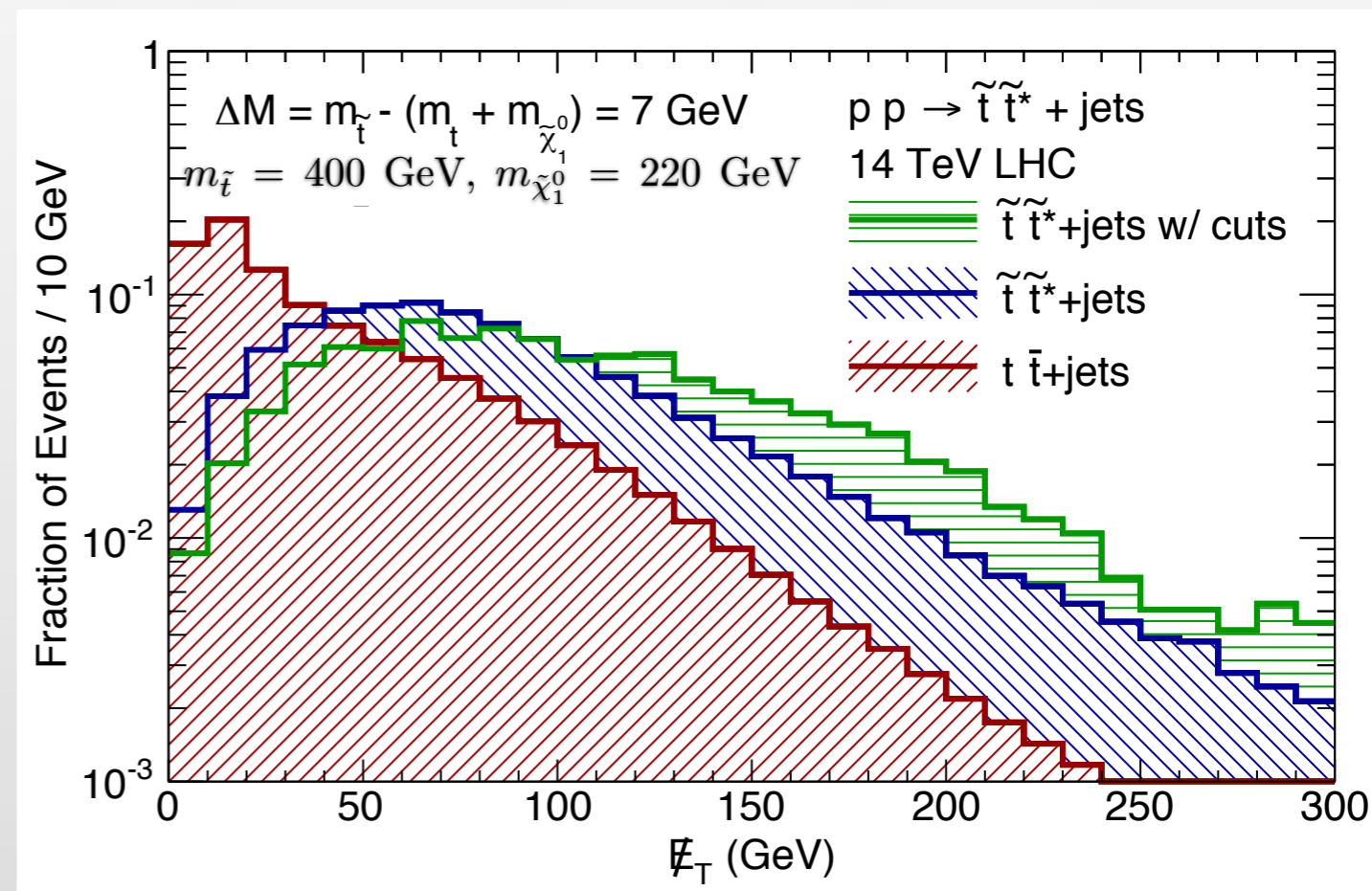
- in inclusive  $\tilde{t}\tilde{t}^* + \text{jets}$  production, ask for 2 forward jets with large  $M_{\text{inv}}(\text{JJ})$



$(m_{\tilde{t}}, m_{\tilde{\chi}_1^0})$	Selection	Signal	$t\bar{t} + \text{jets}$
		fb	fb
	VBF	95.7	16774
(300, 120)	1 lepton	22.1	3587
	2 $b$ -jets	9.70	1612
	$\cancel{E}_T > 50$	8.00	924
	VBF	25.2	16774
(400, 220)	1 lepton	5.93	3587
	2 $b$ -jets	2.84	1612
	$\cancel{E}_T > 100$	1.48	337
	VBF	7.50	16774
(500, 320)	1 lepton	1.69	3587
	2 $b$ -jets	0.74	1612
	$\cancel{E}_T > 150$	0.27	123

$$m_{\tilde{t}} \Rightarrow m_t + m_{\tilde{\chi}_1^0}$$

Dutta et al., arXiv:1312.1348

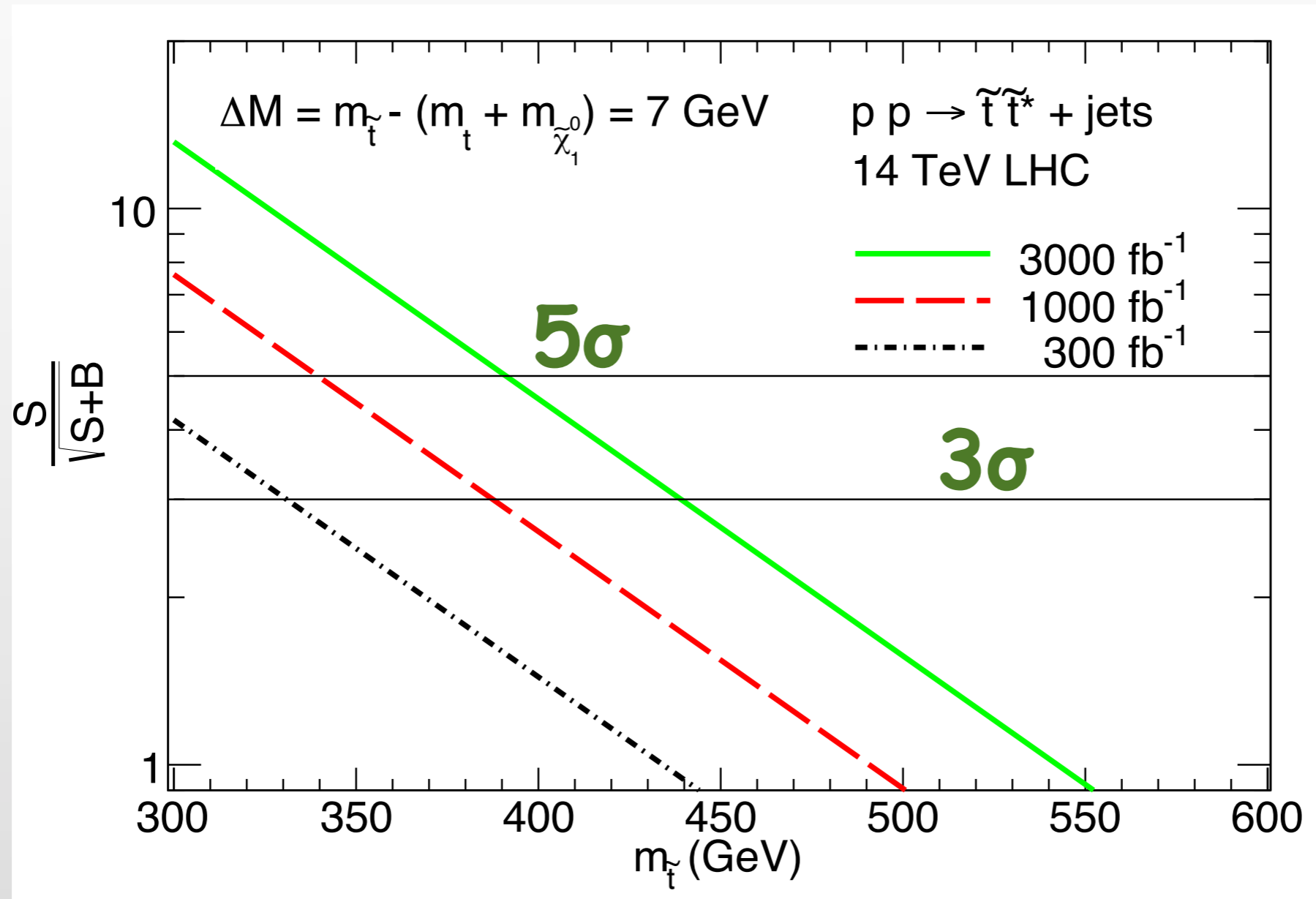


**$E_T^{\text{miss}}$  cut very effective in enhancing S/B**

for  $\Delta M \sim 7 \text{ GeV}$

LHC discovery potential up to

$m_{\text{stop}} \sim 340 \text{ (390) GeV}$  with  $1000 \text{ (3000) fb}^{-1}$



for stealthy stop ( $m_{\tilde{\chi}^0} \sim 0$ ),  $E_T^{\text{miss}}$  cut loses performance

# compressed $X^+/X^0$ in **mono-jet**

Schwaller and Zurita, arXiv:1312.7350

- extend strategies for DM searches to degenerate  $X^+/X^0$  whose  $\chi' \rightarrow \ell\ell'\chi_1^0$  products are too soft to pass trigger requirements

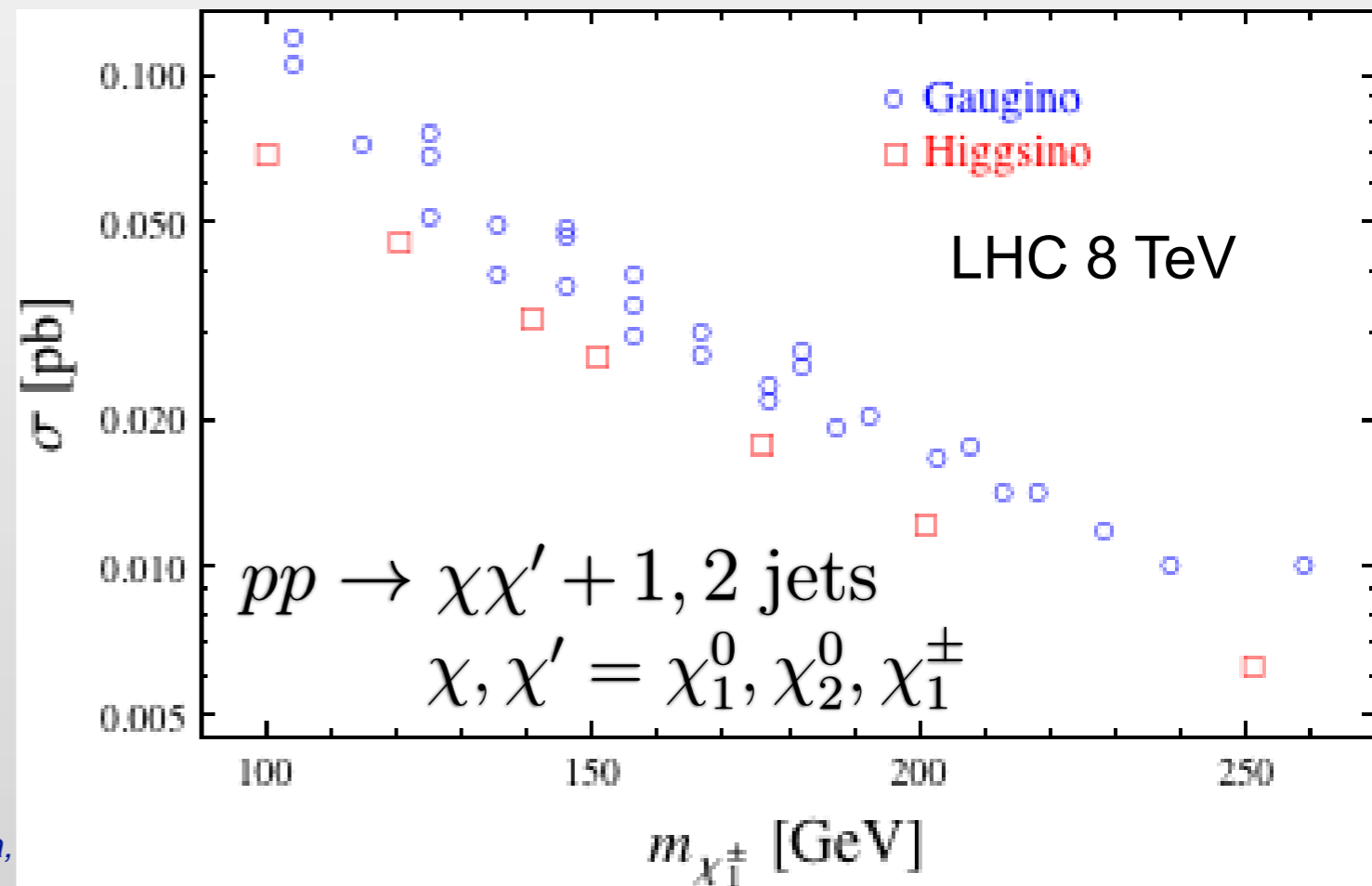
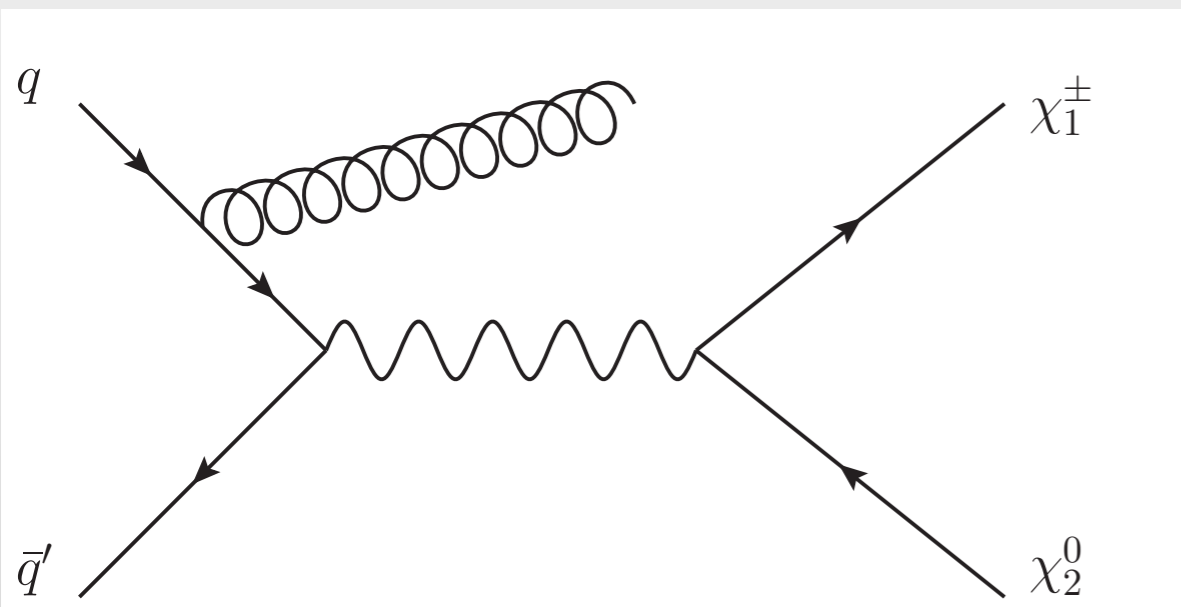
- two possibilities studied:

$$\mu \gg M_1, M_2 \sim M_Z$$

$$M_1, M_2 \gg \mu \sim M_Z$$

degenerate gauginos

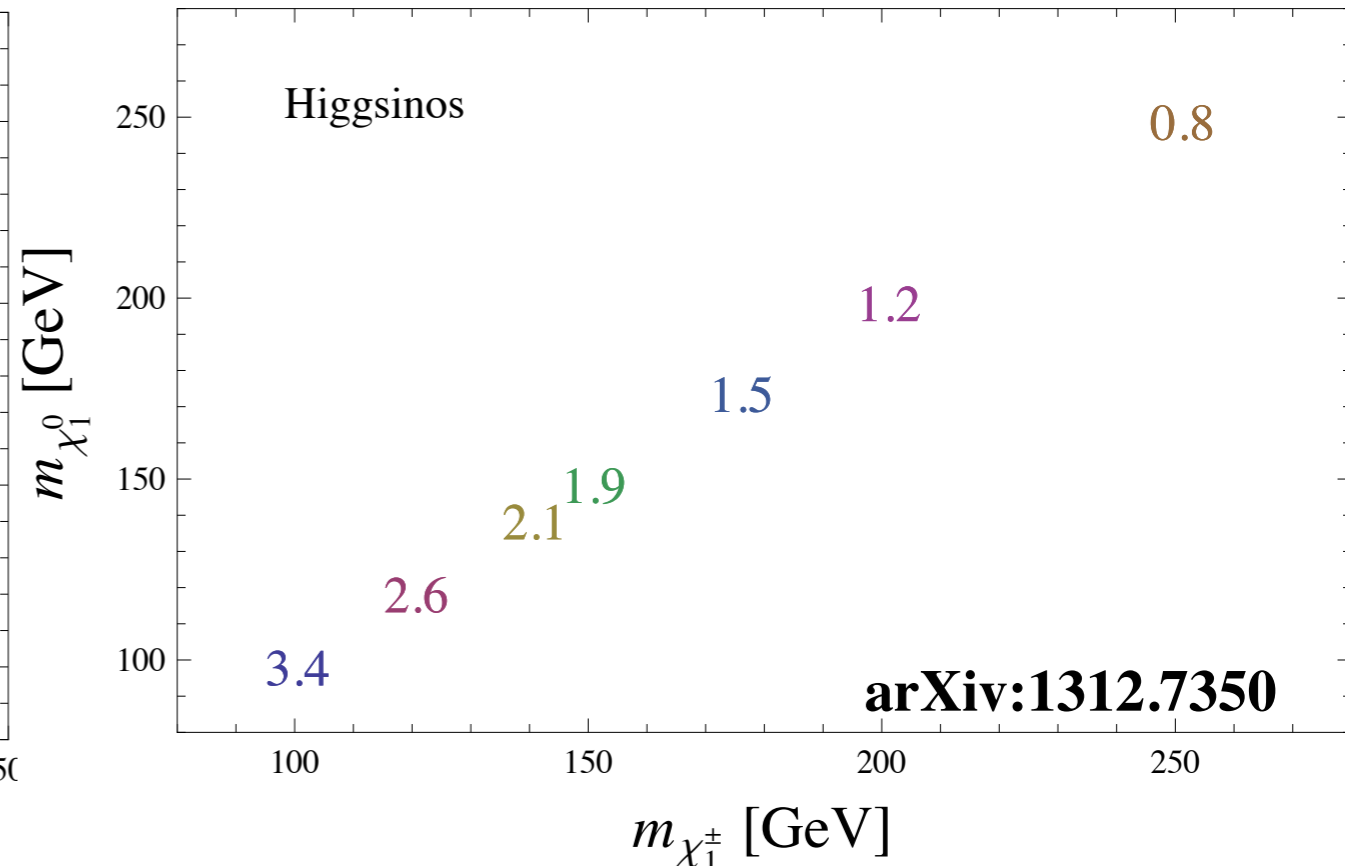
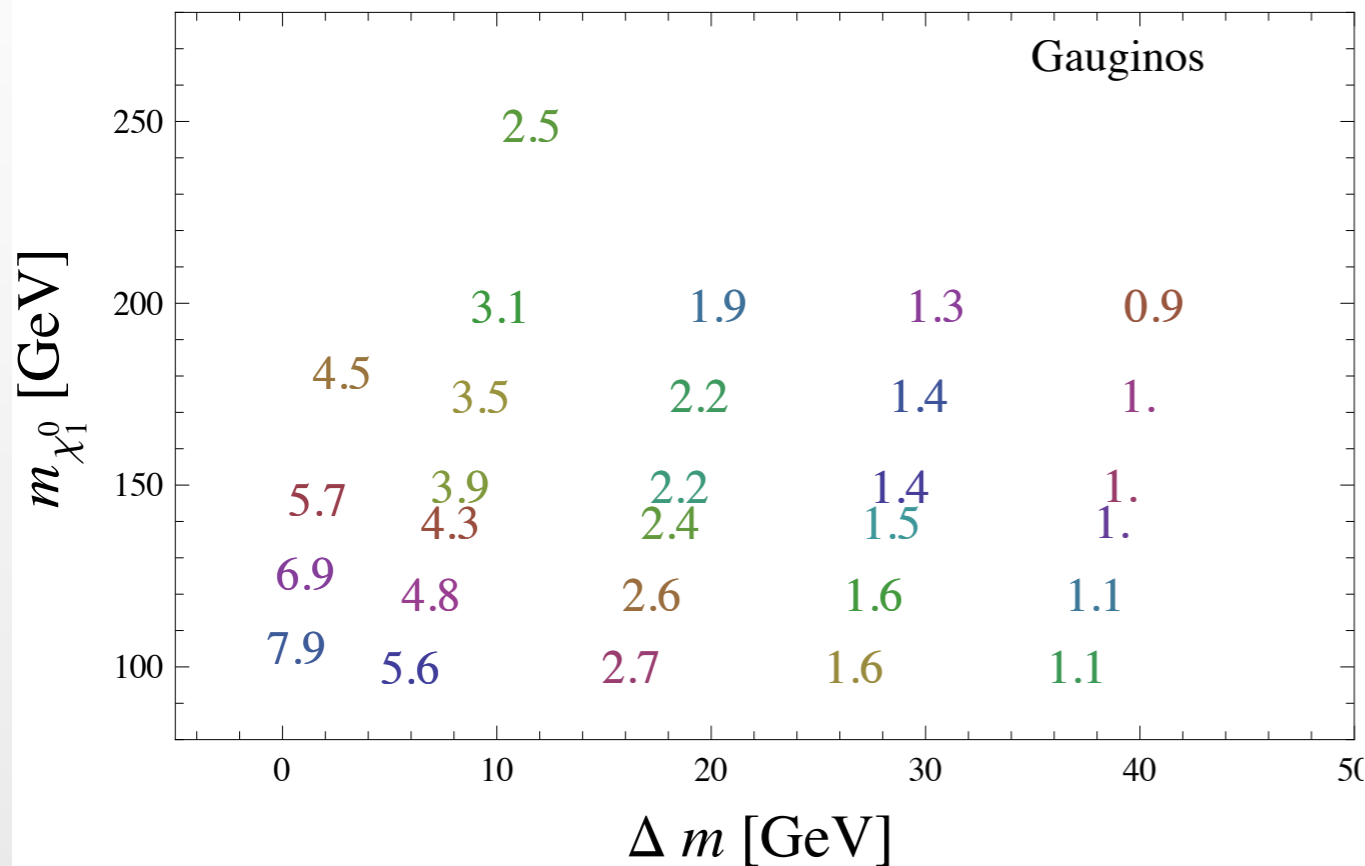
“ higgsinos (Natural Scen.)



● addition of soft leptons crucial to reach sensitivity at 14 TeV (for  $\Delta M \sim 10-30$  GeV)

Significance ( $300 \text{ fb}^{-1}$ ), syst = 1%

Significance ( $300 \text{ fb}^{-1}$ ), syst = 1%



● at 14 TeV for 300 (3000)  $\text{fb}^{-1}$   
 can exclude up to 250 GeV gauginos,  
 for  $\Delta M < 10-(40)$  GeV, and  
 up to 150 ( 200 ) GeV higgsinos

# what if a BSM signal comes out at LHC ?

- a **Great Breakthrough** in particle physics, although not a confirmation of any single theory model ! ! !
- just the start-up of exploration of the “next layer of theory”
- considerable degeneracy in the expected phenomenology for quite a number of BSM models (eg. missing  $P_T$  from many models with a WIMP candidate)
- for any single TH model to be credited, it will have to overcome the “anomaly-fitting phase” and enter the “prediction phase” !
- it will take time and a lot of effort to advance in theory !

# Outlook (1)

- SM built up on decades of unexplained experimental facts...  
By early 70's it was a complete framework whose many crucial predictions had to be checked out in following 40 years, guaranteeing a tremendous outcome for collider physics (and proving the model an extraordinary success !)
- today no BSM framework is as mature as SM in 70's !  
just a number of theoretical suggestions ...  
(with no guarantee of success for anyone !)
- nevertheless, the statement that "new phenomena at the TeV scale are needed to make the EWSB scale Natural" is still robust after LHC Run 1
- in which format ?

# Outlook (2)

- many alternatives open: new resonances, but also deviations in inclusive cross sections / decay rates / Kinem. distributions / correlations
- use theory suggestions as “benchmark frameworks” for search strategies !
- crucial also to extend “signature-driven” studies, as unbiased as possible by theoretical prejudice !
- today exp data definitely needed to advance in theory
- we are now in a thrilling phase: LHC is an extraordinary (and highly complex) tool that has still to develop its full potential ... (lot of effort and inventiveness will be needed)

**big advances in the exploration of TeV scale  
in next few years “guaranteed” !**