

Light as a probe for ultrasmall effects of General Relativity

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Foreword: *General Relativity*

- The gravitational field is geometry
- Studying the gravitational interaction amounts to explore the geometry of space time
- Locally the most common (approximate) symmetry is chiral about a time-like axis: steadily rotating massive object

Rotating sources of gravity

Among the classical weak effects of GR there are the ones due to rotating masses, often called gravito-magnetic effects.

The direct measurements of these effects is important *per se* and could evidence deviations addressing one or another extended or alternative theory of gravity

Axially symmetric stationary space-time

Remote inertial observer; "polar" space coordinates;
(spherical symmetry in space)

$$ds^2 = g_{tt}dt^2 + g_{rr}dr^2 + g_{\theta\theta}d\theta^2 + g_{\varphi\varphi}d\varphi^2 + 2g_{t\varphi}dtd\varphi$$

Gravito-magnetic effects

$g_{\mu\nu}$'s do not depend on t and φ

$$\vec{h} \cdot d\vec{l}$$

$$2g_{ti}dx^i dt$$

Line element for light in approximated geographic terrestrial coordinates

$$g_{tt}dt^2 + g_{rr}dr^2 + g_{\theta\theta}d\theta^2 + g_{\varphi\varphi}d\varphi^2 + 2g_{t\varphi}dtd\varphi = 0$$



$$\left(1 - 2\frac{\mu}{r} - \frac{\Omega^2 r^2}{c^2} \sin^2 \theta\right) d\tau^2 - \left(1 + 2\frac{\mu}{r}\right) dr^2 - r^2 d\theta^2 - \left(1 + 2\frac{r^2 \Omega^2}{c^2} \sin^2 \theta\right) r^2 \sin^2 \theta d\varphi^2 + 2\left(2\frac{j}{r} - r^2 \frac{\Omega}{c} - 2\mu \frac{r\Omega}{c}\right) \sin^2 \theta d\tau d\varphi = 0$$

Laboratory coordinates

$$\begin{aligned}
 & \left(1 - \frac{\Omega^2 R^2}{c^2} \sin^2 \Theta - 2 \frac{\mu}{R} - 2 \frac{\Omega^2}{c^2} R \sin^2 \Theta (\zeta \sin \Theta + y \cos \Theta) + 2 \frac{\zeta \mu}{R^2} \right) d\tau^2 \\
 & - \left(1 + 2 \frac{\mu}{R} - 2 \frac{\zeta \mu}{R^2} \right) d\zeta^2 - \left(1 + \frac{\zeta}{R} \right)^2 dy^2 \\
 & - \left(1 + 2 \frac{R^2 \Omega^2}{c^2} \sin^2 \Theta + \frac{2}{R} \left(\zeta + y \frac{\cos \Theta}{\sin \Theta} \right) + 8 \frac{R \Omega^2}{c^2} (\zeta \sin \Theta + y \cos \Theta) \sin \Theta \right) dx^2 \\
 & + 2 \left(2 \frac{j}{R^2} \sin \Theta - \frac{R \Omega}{c} \sin \Theta - 2 \frac{\Omega}{c} (y \cos \Theta + (\zeta + \mu) \sin \Theta) \right) d\tau dx = 0
 \end{aligned}$$

General laboratory expression

$$\begin{aligned} & (1 + 2w(y, z; \Theta, \mu, \Omega, R))d\tau^2 \\ & - (1 + q(y, z; \Theta, \mu, \Omega, R))dx^2 \\ & - (1 + f(z; \Theta, \mu, \Omega, R))^2 dy^2 - dz^2 \\ & + 2h(y, z; \Theta, \mu, j, \Omega, R)d\tau dx = 0 \end{aligned}$$

"Free" motion along a closed null path

Time of flight asymmetry:

$$\delta t = t_+ - t_- = -\frac{2}{c} \oint \frac{g_{0i}}{g_{00}} dx^i \neq 0$$

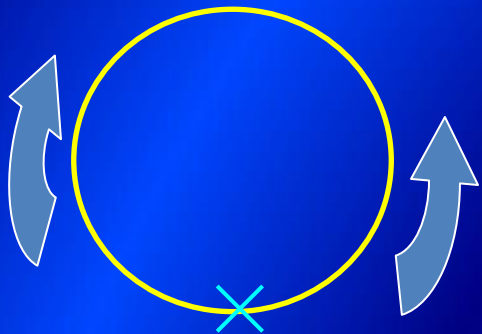
global coordinated time

$$\delta \tau = -2\sqrt{g_{00}} \oint \frac{g_{0i}}{g_{00}} dx^i$$

proper time of a fixed observer

Ring laser: light propagation

Closed loop (shape irrelevant)



Counterrotating light beams

Time of flight difference

$$\delta t = -2 \oint \frac{g_{0i}}{g_{00}} dx^i$$

$$\delta \tau = -2 \sqrt{g_{00}} \oint \frac{g_{0i}}{g_{00}} dx^i = -2 \sqrt{g_{00}} \int_S \vec{\nabla} \wedge \vec{h} \cdot \hat{n} dS$$

$$h_i = \frac{g_{0i}}{g_{00}}$$

Weak field approximation

Relevant parameters: $\mu = G \frac{M}{c^2};$ $a = \frac{J}{Mc}$

Earth: $\mu = 4.43 \times 10^{-3} \text{ m}$ $a = 3.95 \text{ m}$

$$ds^2 \cong \left(1 - 2\frac{\mu}{r}\right)c^2 dt^2 - \left(1 + 2\frac{\mu}{r}\right)dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\varphi^2 + 4c \frac{j}{r} \sin^2 \theta dt d\varphi$$

$$j = \mu a = G \frac{J}{c^3} \approx 1.75 \times 10^{-2} \text{ m}^2 \quad \vec{h} = 2 \frac{j}{r^2} \sin \theta \hat{\phi}$$

Earth-bound laboratory

$$ds^2 \cong \left(1 - 2\frac{\mu}{r} - \frac{r^2\Omega^2}{c^2}\sin^2\theta\right)c^2dt^2 - \left(1 + 2\frac{\mu}{r}\right)dr^2 - r^2d\theta^2 - r^2\sin^2\theta d\phi^2$$
$$+ 2\left(-\frac{\Omega r}{c} - 2\frac{\Omega\mu}{c} + 2\frac{j}{r^2}\right)cr\sin^2\theta dt d\phi$$

$$\frac{r\Omega}{c} \approx 2.47 \times 10^{-7}$$

Earth-centered co-rotating axes

Local Cartesian

$$\begin{aligned}
 & \left(1 - \frac{\Omega^2 R^2}{c^2} \sin^2 \Theta - 2 \frac{\mu}{R} - 2 \frac{\Omega^2}{c^2} R \sin^2 \Theta (\zeta \sin \Theta + y \cos \Theta) + 2 \frac{\zeta \mu}{R^2} \right) d\tau^2 \\
 & - \left(1 + 2 \frac{\mu}{R} - 2 \frac{\zeta \mu}{R^2} \right) d\zeta^2 - \left(1 + \frac{\zeta}{R} \right)^2 dy^2 \\
 & - \left(1 + 2 \frac{R^2 \Omega^2}{c^2} \sin^2 \Theta + \frac{2}{R} \left(\zeta + y \frac{\cos \Theta}{\sin \Theta} \right) + 8 \frac{R \Omega^2}{c^2} (\zeta \sin \Theta + y \cos \Theta) \sin \Theta \right) dx^2 \\
 & + 2 \left(2 \frac{j}{R^2} \sin \Theta - \frac{R \Omega}{c} \sin \Theta - 2 \frac{\Omega}{c} (y \cos \Theta + (\zeta + \mu) \sin \Theta) \right) d\tau dx = 0
 \end{aligned}$$

Time of flight difference

$$\vec{\nabla} \wedge \vec{h} = 2 \left(\frac{\Omega}{c} r + \frac{\Omega \mu}{c} + 2 \frac{j}{r^2} \right) \frac{\cos \theta}{cr} \hat{r} - 2 \left(\frac{\Omega r}{c} - \frac{\Omega \mu}{c} - \frac{j}{r^2} \right) \frac{\sin \theta}{cr} \hat{\theta}$$

$$\delta\tau = -2\sqrt{g_{00}} \int_S \vec{\nabla} \wedge \vec{h} \cdot \hat{n} dA$$

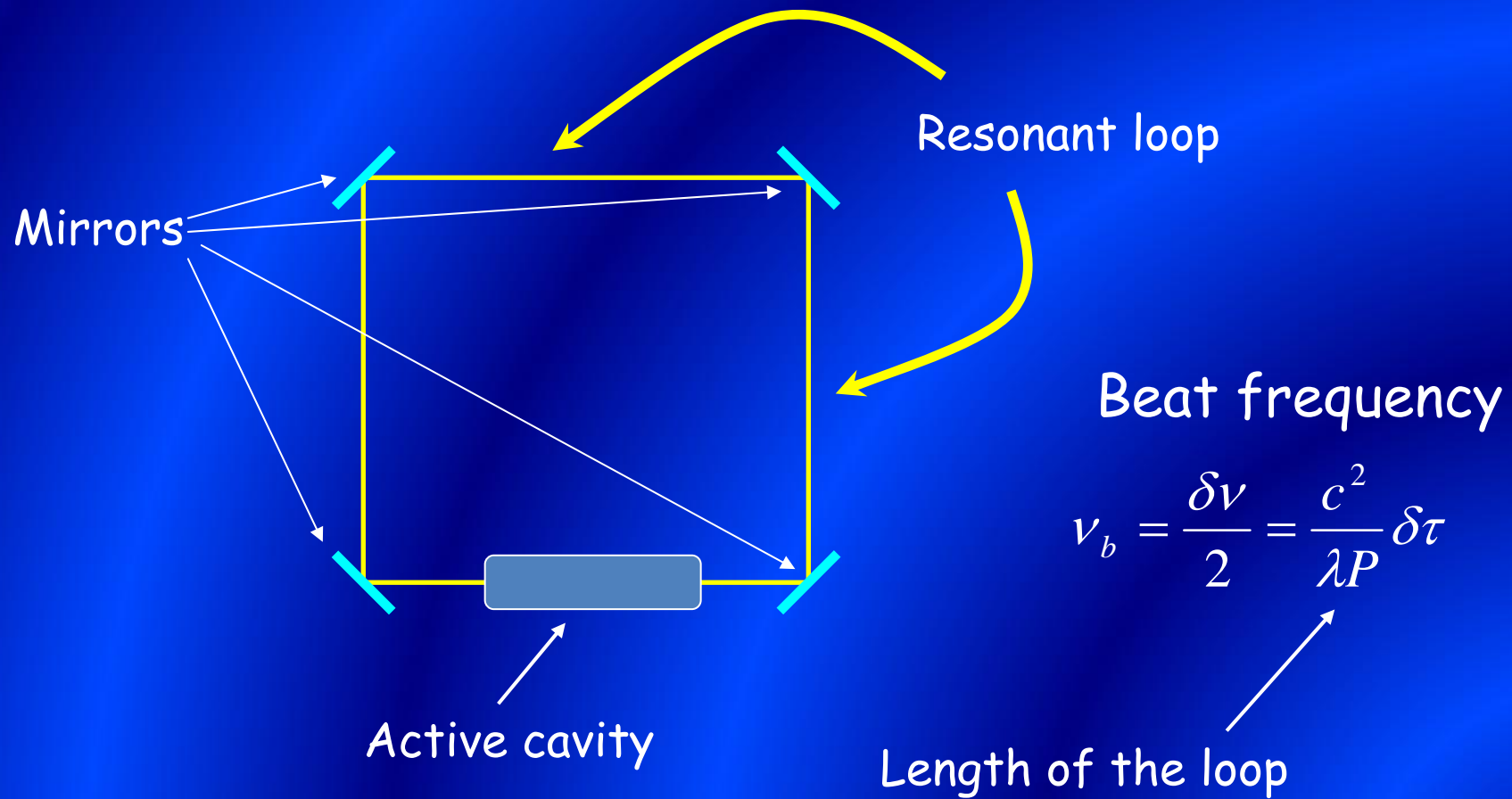
$$\delta\tau = -4 \left(\frac{\Omega R}{c} + 2 \frac{j}{R^2} \right) \frac{\cos \theta}{cR} A \hat{r} \cdot \hat{n} + 4 \left(\frac{\Omega R}{c} + 2 \frac{\Omega \mu}{c} - \frac{j}{R^2} \right) \frac{\sin \theta}{cR} A \hat{\theta} \cdot \hat{n}$$

ToF difference in local Cartesian

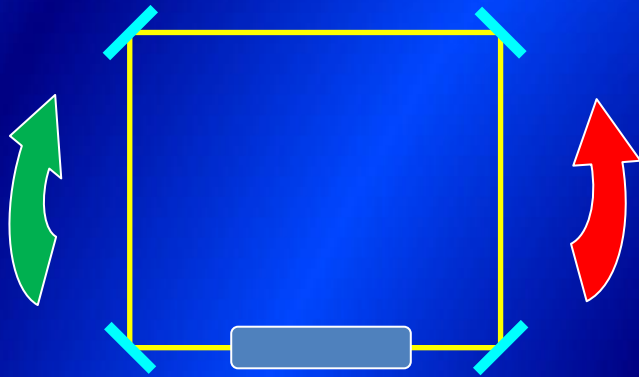
$$k_x = \frac{h}{1 + 2w} \quad \vec{\nabla} \wedge \vec{k} = \frac{\partial k_x}{\partial z} \hat{u}_y - \frac{\partial k_x}{\partial y} \hat{u}_z$$

$$\delta\tau = -2 \frac{\partial k_x}{\partial z} \hat{u}_y \cdot \hat{n} S + 2 \frac{\partial k_x}{\partial y} \hat{u}_z \cdot \hat{n} S$$

A ring laser



Stationarity condition



$$N = \frac{\tau_+}{T_+} = \tau_+ \nu_+$$

$$N = \frac{\tau_-}{T_-} = \tau_- \nu_-$$

Single mode



$$N = \frac{P}{\lambda}$$

Length of the path

$$\nu_b = \frac{\nu_+ - \nu_-}{2} \cong \frac{c^2}{2\lambda P} \delta\tau$$

Beat frequency

Expected signal (GR)

Normal in the meridian plane

$$v_b = 2 \frac{A}{\lambda P} \Omega_{\oplus} \left[\cos(\theta + \alpha) - 2 \frac{\mu}{R} \sin \theta \sin \alpha + \frac{GI_{\oplus}}{c^2 R^3} (2 \cos \theta \cos \alpha + \sin \theta \sin \alpha) \right]$$

Scale factor

$$v_b = 2 \frac{A}{\lambda P} \left[\vec{\Omega}_{\oplus} - 2 \frac{\mu}{R} \Omega_{\oplus} \sin \theta \hat{\theta} + \frac{c j_{\oplus}}{R^3} (2 \cos \theta \hat{r} + \sin \theta \hat{\theta}) \right] \cdot \hat{n}$$

Area of the loop

Sagnac

$\vec{\Omega}_G$

$\vec{\Omega}_B$

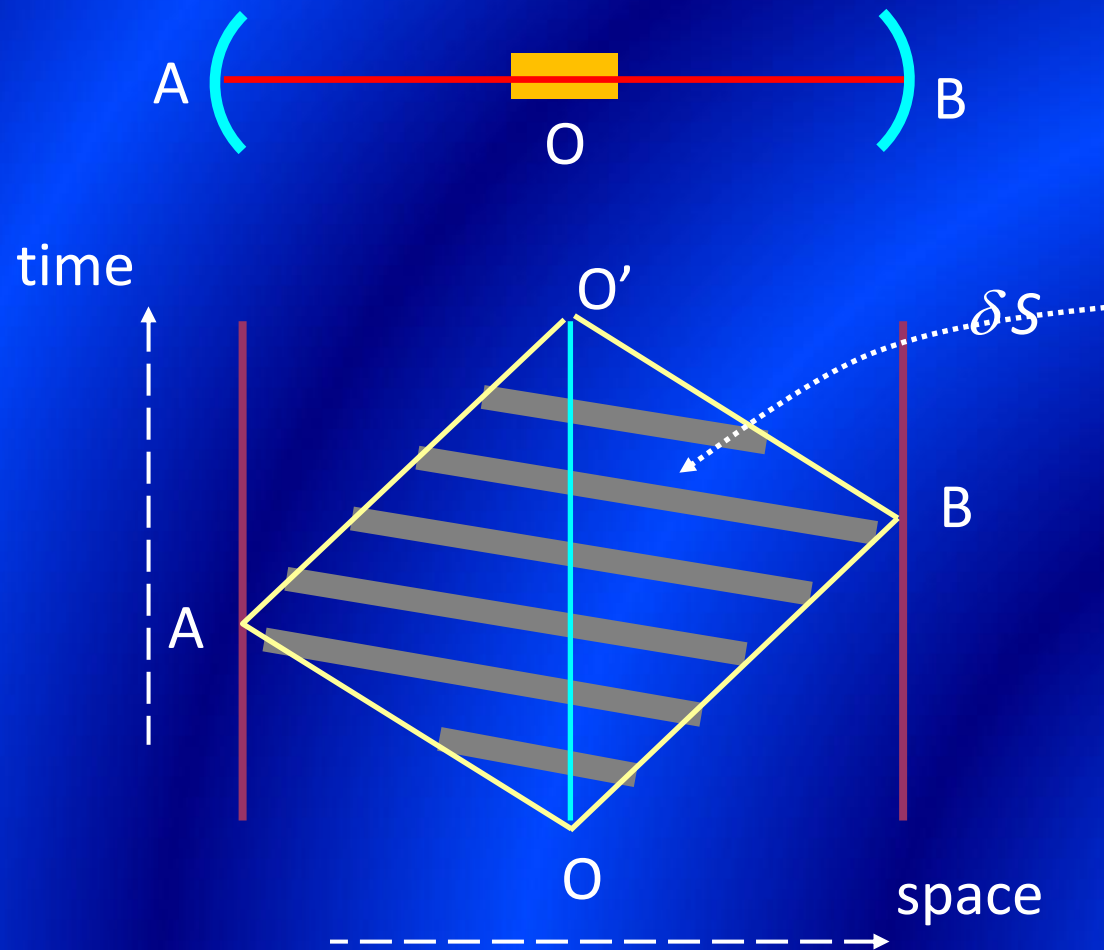
Orders of magnitude

$$\frac{\Omega_G}{\Omega} \approx \frac{\Omega_B}{\Omega} \approx 10^{-9}$$

$$\nu_b \sim 10^2 \text{ Hz}$$

Required accuracy for g-magnetic effects: a few prad/s

A linear cavity



The effect of curvature and anisotropy

$$\delta F^{\mu\nu} = \left(R^{\mu}_{\ \varepsilon 0 i} F^{\varepsilon\nu} + R^{\nu}_{\ \varepsilon 0 i} F^{\mu\varepsilon} \right) \delta S^{0i}$$

Riemann tensor

Electromagnetic tensor

Space-time area
spanned by the
cavity

Depends on the orientation of the cavity

Linearized Riemann tensor.

An example: cavity in the meridian plane

Length of the cavity

$$R^\phi{}_{r0\theta}$$

$$\delta F^{\vartheta\phi} \cong \left[\left(\frac{\mu\Omega}{cR^2} - 3\frac{j}{R^4} \right) \frac{\cos\vartheta}{\sin^2\vartheta} \right] \frac{l^2}{R} F^{\vartheta r}$$

Change in the radial component of the magnetic field

East-West component of the magnetic field

Rate of change of the EM field per bounce

Kinematic $\frac{\delta F}{F} \approx 10^{-7}$

Gravito-electric $\frac{\delta F}{F} \approx 10^{-8}$

Gravito-magnetic $\frac{\delta F}{F} \approx 10^{-14}$

Geometric effect in a cavity

k : null tangent vector

δk : change while moving back and forth

$$\delta k^\mu = R^\mu{}_{\alpha\nu\beta} k^\alpha S^{\nu\beta}$$



$$\delta k^\mu = R^\mu{}_{\alpha 0 i} k^\alpha S^{0i}$$

An example

East-West cavity

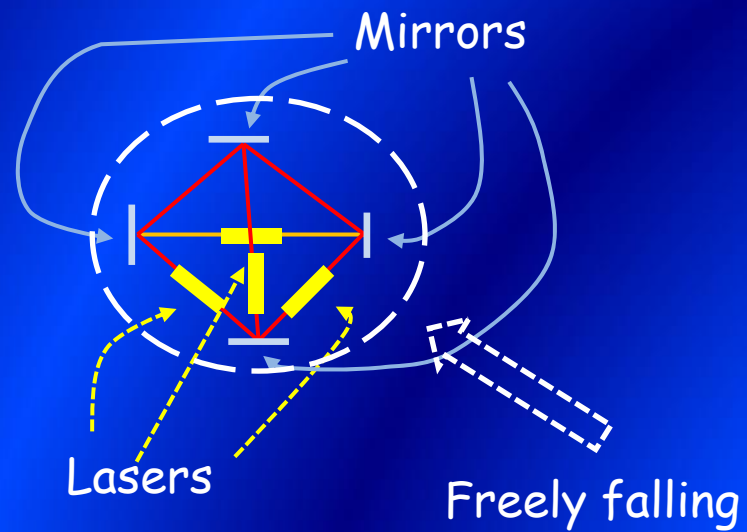
Frequency shift

$$\delta k^0 = R^0_{\phi 0 \phi} k^\phi S^{0\phi} \cong \left(\frac{\mu}{R} \sin^2 \Theta + \frac{\Omega^2 R^2}{c^2} \cos^2 \Theta \right) k^\phi S^{0\phi} = \left(\frac{\mu}{R} \sin^2 \Theta + \frac{\Omega^2 R^2}{c^2} \cos^2 \Theta \right) kl^2$$

$$\delta k^\phi = R^\phi_{\phi 0 \phi} k^\phi S^{0\phi} \cong \frac{\mu \Omega}{cR} \sin \Theta k^\phi S^{0\phi}; \quad \delta k^x \cong \frac{\mu \Omega}{cR^2} kl^2$$

$$\frac{\delta k^x}{k^x} \approx 10^{-29} \text{ rad}$$

Ring lasers in space



Tetrahedron with four
triangular ringlasers



Circular equatorial orbit (non-spinning device, spherical Earth)

$$\omega = \sqrt{G \frac{M}{R^3}}$$

$$ds^2 \cong \left(1 - 2\frac{\mu}{r}\right) c^2 dt^2 - \left(1 + 2\frac{\mu}{r}\right) dr^2 - r^2 d\phi^2 + 4cr \left(\frac{j}{r^2} - \left(\frac{\mu}{r}\right)^{\frac{3}{2}} \right) dt d\phi$$

$$v_b = \frac{cA}{\lambda PR} \left(\left(\frac{\mu}{R}\right)^{\frac{3}{2}} - 2\frac{j}{R^2} \right) \hat{\theta} \cdot \hat{n}$$

$\sim 10^{-14}$

$\sim 10^{-16}$

Orbital gyro



Arrival times difference

Equatorial orbit

$$\delta\tau \cong -2 \frac{A}{cR} \left(\left(\frac{\mu}{R} \right)^{\frac{3}{2}} - 2 \frac{j}{R^2} \right) \approx 10^{-2} T$$

Pulses

Gaia and GAME

- *GAI*A: catalog of ~ billion stars, μ s accuracy. Dynamic model of the gravitational field in the solar system needed. Areas close to the major objects (Sun, Jupiter, ...) excluded.
- *GAME*: comparison of the apparent positions of stars on opposite sides of the sun or Jupiter. Look for asymmetries.

Experimental tests of General Relativity (historical)

- Precession of the perihelion of Mercury
 - 1915
 - Known value: $43''/\text{Julian century}$
 - Value computed by Einstein: $43''/\text{Julian century}$
- Lensing by the Sun
 - Theoretical value according to GR (1915)
 - Observed by Eddington (1919)
- Gravitational redshift
 - Recognized in the spectral lines of Sirius B (W. S. Adams, 1925)
 - Measured on Earth: Pound and Rebka experiment (1959)

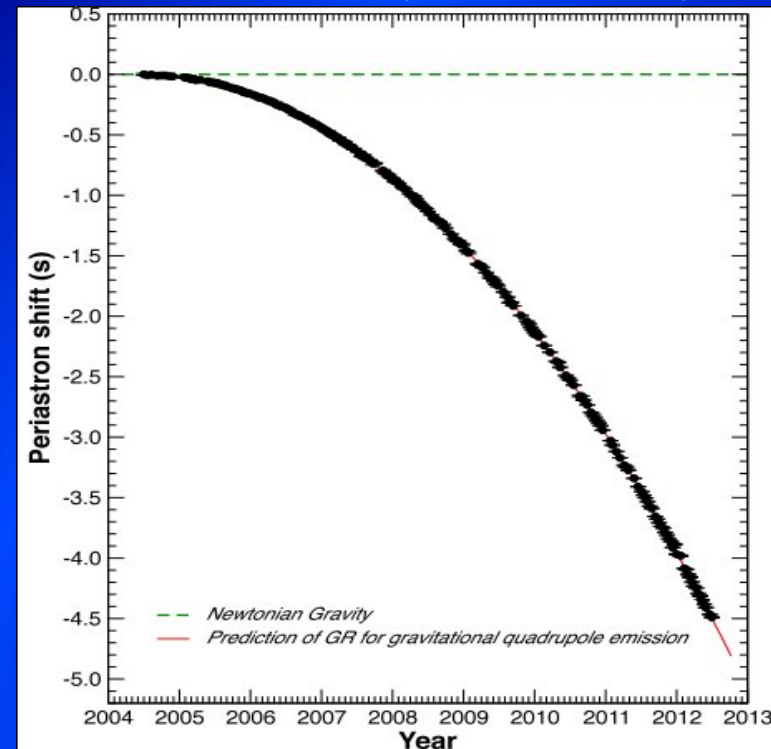
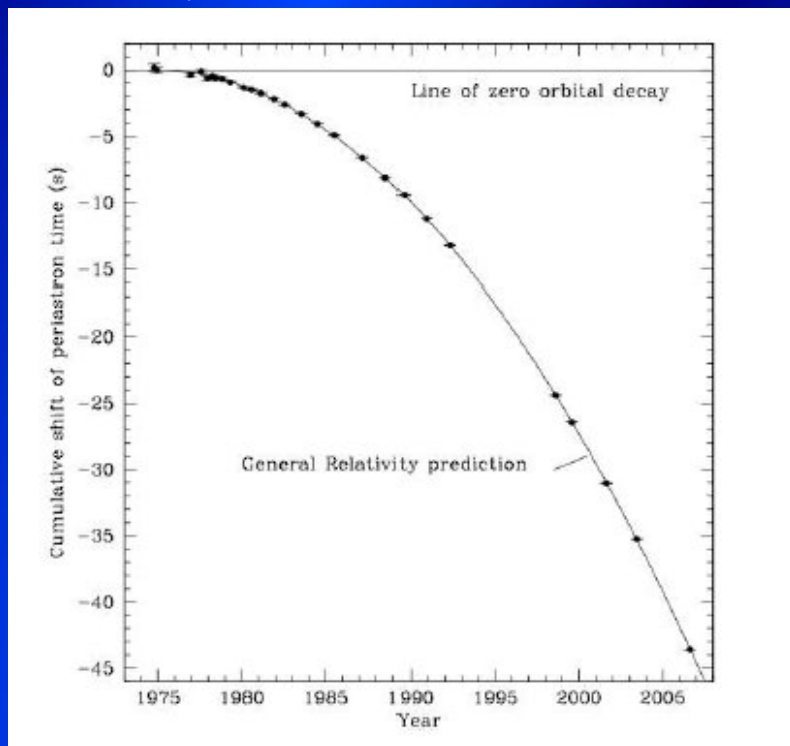
Observation of binaries

- Periapsis precession
 - PSR B1913+16 (Hulse and Taylor pulsar):
4.2°/year (compatible with GR)
 - PSR J0737-3029 (double pulsar): 16.9°/year
- Geodetic (de Sitter) precession
 - PSR J0737-3029B (double pulsar):
 - GR $(5.0734 \pm 0.0007)^\circ/\text{year}$
 - measured $(4.77^{+0.66}_{-0.65})^\circ/\text{year}$

A. Possenti and M. Burgay, private communication (2011)

Period decay (GW emission)

- 0.2% agreement with the quadrupole GW emission rate. PSR B1913+16 (Hulse and Taylor pulsar), after 33 years of data taking. Right: double pulsar, 9 years



Past and present tests

- Weak effects:
 - Geodetic precession
 - Inertial frame dragging
- Direct measurement in space
 - Gravity Probe B: 19%
 - LAGEOS I and II: 10%
 - LARES, under way → 1%
- Lunar laser ranging, pulsars in binary systems, double pulsar,

Extended or alternative theories: the PPN parameters

$$ds^2 = \left(1 - 2\frac{GM}{c^2 r}\right) d\tau^2 - \left(1 + 2\gamma\frac{GM}{c^2 r}\right) dr^2 - r^2 d\vartheta^2 \\ - r^2 \sin^2 \vartheta d\phi^2 + (4 + 4\gamma + \alpha_1) \frac{GJ}{2c^3 r} \sin^2 \vartheta d\tau d\phi$$

$$2M \rightarrow (1 + \gamma)M \qquad 2J \rightarrow \left(1 + \gamma + \frac{\alpha_1}{4}\right)J$$

Comparison with other techniques

- The experiments using mechanical gyroscopes in space sense a time and position depending gravitational field (along the orbit)
- With respect to matter waves, light has a local propagation velocity in vacuo which is a universal constant.
- Laser sources are far more stable and long lasting than the sources of matter waves

Light is a good probe of the structure of space-time

- In the case of GINGER the ring-lasers are fixed to the earth, so the gravitational field (both gravitoelectric and gravitomagnetic component) is constant.
- The side of the loops is far smaller than the local curvature radius of space-time: in practice the measurement is local.

Conclusion

- Closed paths of light in space (and space-time) evidence general relativistic effects either by
 - Interferometric techniques
or
 - Beat frequency measurements
- The GINGER experiment is likely to verify the inertial frame dragging on Earth within a few years and with the 1% accuracy or better
- It would be useful and important to integrate different approaches and techniques in order to reveal possible deviations from GR.