## The CRESST Dark Matter Search

Federica Petricca for the CRESST collaboration

XL meeting of the Gran Sasso Scientific Committee 29 – 30 October 2013, LNGS

#### Outline

- Status of the current run
- Outcome of the previous CRESST run
- Actions to achieve background reduction
- First glance through the collected data
- Schedule and Perspectives

#### Status

# Status of run 33: 18 detector modules mounted



Setting up detectors and calibration in June and July

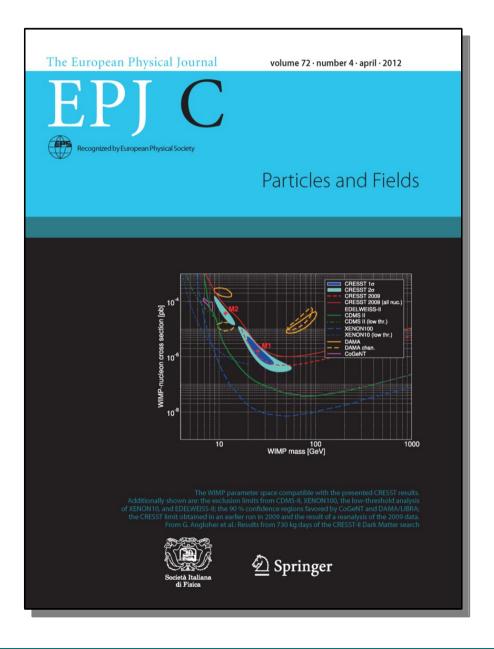
Heater of one Light Detector disturbs stability of carousel – used for temperature control

Data taking since July 30

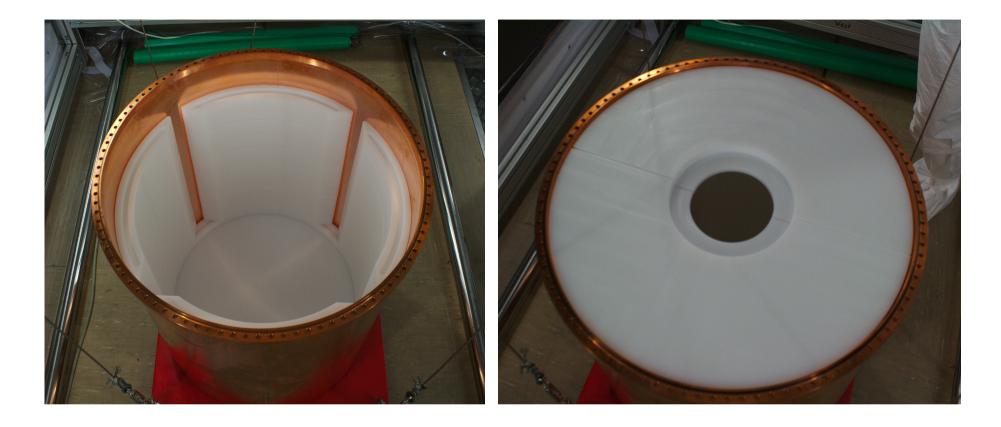
## The previous CRESST run

- Extended physics run from June 2009 to April 2011
- 8 CaWO4 modules used for Dark Matter analysis
- Net exposure after cuts: 730 kg days
- 67 events observed in WIMP search region
- Data analyzed with 2d likelihood fit of signal and background model

	M1	M2
$e/\gamma$ -events	$8.00\pm0.05$	$8.00\pm0.05$
$\alpha$ -events	$11.5^{+2.6}_{-2.3}$	$11.2^{+2.5}_{-2.3}$
neutron events	$7.5^{+6.3}_{-5.5}$	$9.7  {}^{+6.1}_{-5.1}$
Pb recoils	$15.0^{+5.2}_{-5.1}$	$18.7^{+4.9}_{-4.7}$
signal events	$29.4^{+8.6}_{-7.7}$	$24.2^{+8.1}_{-7.2}$
$m_{\chi} \; [\text{GeV}]$	25.3	11.6
$\sigma_{\rm WN} ~[{\rm pb}]$	$1.6 \cdot 10^{-6}$	$3.7 \cdot 10^{-5}$



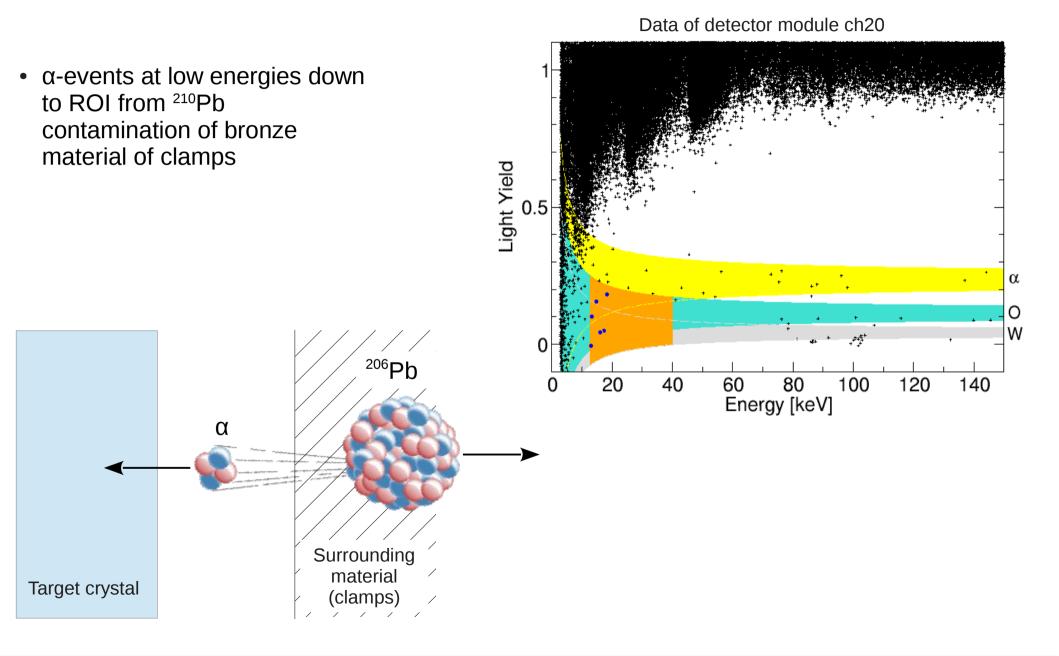
#### Passive background reduction – neutrons



Additional 5cm PE layer inside the Pb/Cu shield

• reduce background from neutrons originating in the Pb/Cu shield

#### The previous CRESST run – low energy $\alpha$

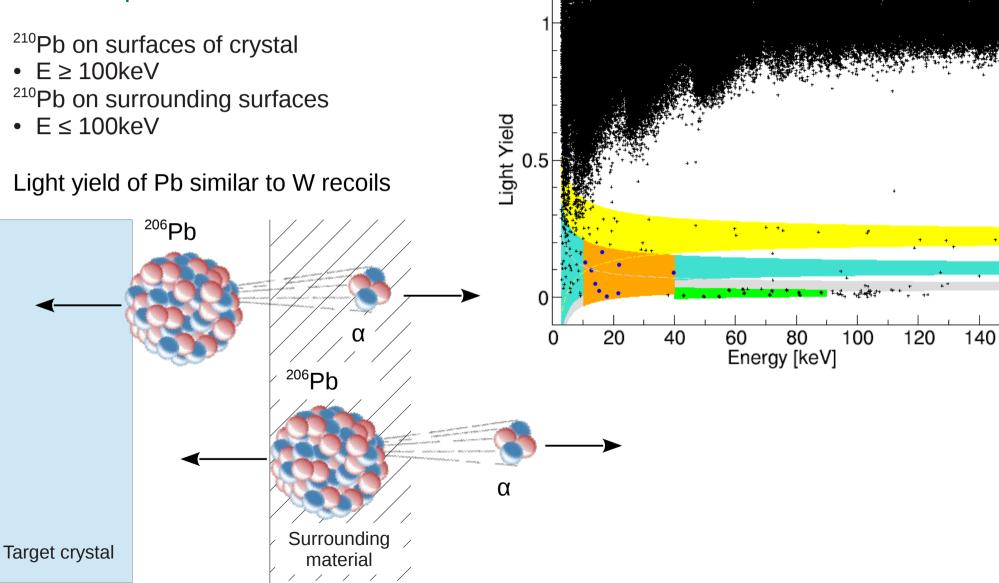


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# The previous CRESST run – <sup>206</sup>Pb recoils

# Background of Pb recoils due to radon exposure

#### Module with highest <sup>206</sup> Pb background

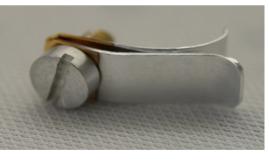


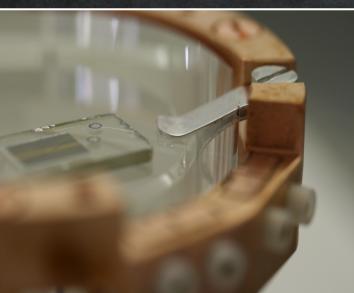
### Passive background reduction – low energy $\alpha$

New CuSn6:

- ultra pure (7n) Sn + low background Cu and careful monitoring of all production steps (\*)
- sputter coating with high purity AI







(\*) Thanks to the support of chemistry service and special techniques of LNGS.



### Passive background reduction – <sup>206</sup>Pb recoils

#### Background of Pb recoils due to radon exposure of clamps after production

- 1- Avoid any radon exposure of clamps and detector material after production
  - Detector modules assembled in clean room facility at Gran Sasso in radon free air (\*)
  - Detector modules mounted inside cryostat in radon free air
  - Reflecting foils and scintillator rings of 6 modules covered with Parylene (\*) to reset radon exposure history



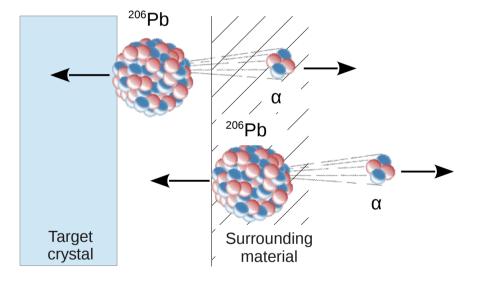
#### 12 conventional modules in present run entirely relying on radon prevention

(\*) Thanks to the CUORE collaboration for the radon free air plant and for the Parylene coating system The CRESST Dark Matter Search 29.10.2013

#### F. Petricca for the CRESST collaboration

### Active background reduction – <sup>206</sup>Pb recoils

- 2- Detect the emitted  $\alpha$  to veto the events
  - New detector layouts with active rejection of Pb recoils

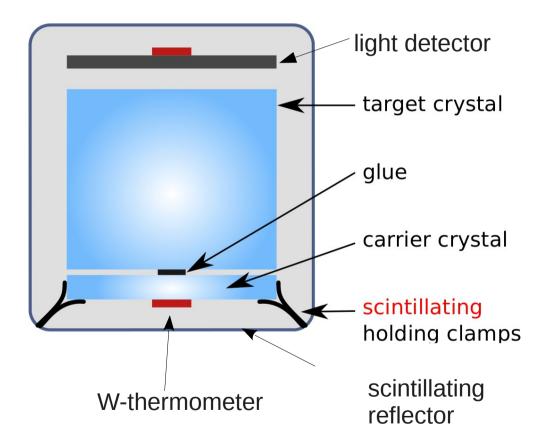


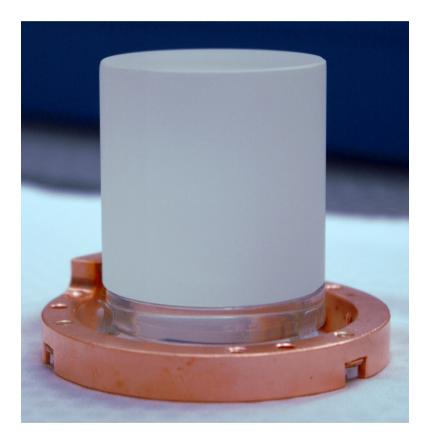
6 modules with 3 different designs in present run

### Active background reduction – Design I

Crystal clamped on carrier:

- fully scintillating holder design
- clamps covered with Parylene not in contact with the target crystal



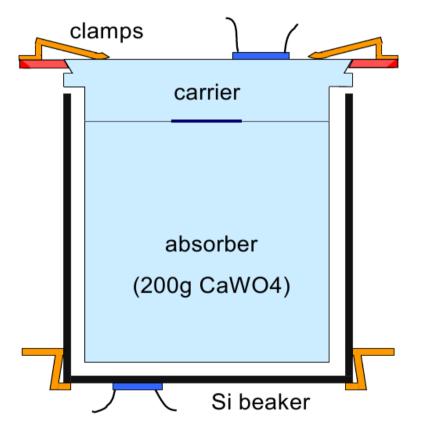


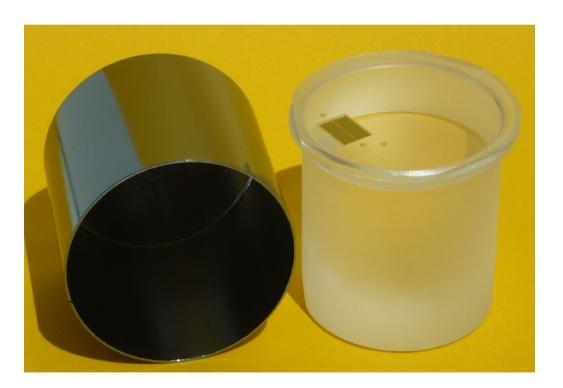
Additional light signal of the  $\alpha$  in the scintillating surrounding as veto

### Active background reduction – Design II

Silicon beaker light detector:

- light detector completely surrounding target crystal
- crystal held on carrier
- no line of sight between target crystal and clamps



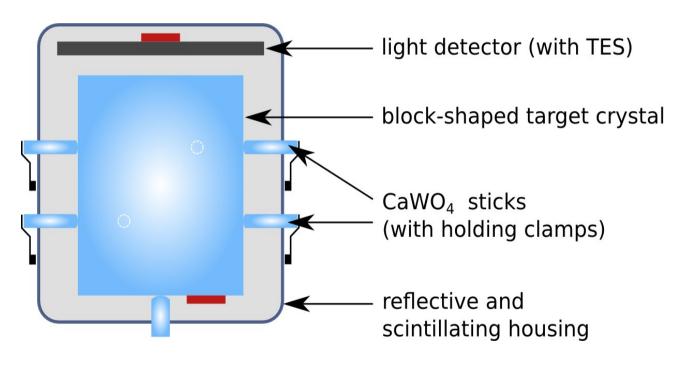


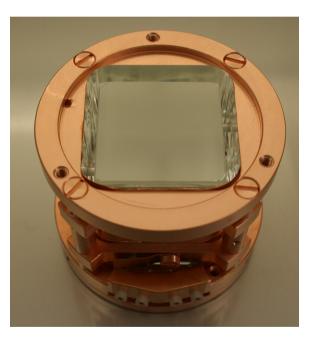
The  $\boldsymbol{\alpha}$  detected in the light detector

#### Active background reduction – Design III

Crystal held by sticks:

- fully scintillating holder design
- CaWO<sub>4</sub> sticks to hold the crystal
- block-shaped crystal





Additional light signal of the  $\boldsymbol{\alpha}$  in the scintillating surrounding as veto

Data until December used for defining cuts which then will be applied later in blind analysis

Data from July 30 to October 9

#### Very preliminary !!

General remarks:

- Exposure per detector between 10kg days and 16kg days
- Hardware trigger threshold ≤ 5keV Six with hardware trigger threshold ≤1keV

Data until December used for defining cuts which then will be applied later in blind analysis

Data from July 30 to October 9

Very preliminary !!

Ch37/38: TUM38/Petrus (15 kg-days)

Design I: Crystal clamped on carrier

Data until December used for defining cuts which then will be applied later in blind analysis

Data from July 30 to October 9

Very preliminary !!

Ch55/56: VK28/Zam (9.7 kg-days)

Design II: Silicon beaker light detector

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Data until December used for defining cuts which then will be applied later in blind analysis

Data from July 30 to October 9

**Very preliminary !!** 

1.5 Light Yield 50 0 50 100 150 200 250 300 0 Energy [keV]

Ch7/8: TUM40/Michael (12.6 kg-days)

Design III:Crystal held by sticks

Data until December used for defining cuts which then will be applied later in blind analysis

Data from July 30 to October 9

#### Very preliminary !!

Three modules with active background rejection

- 37.3 kg days
- reference regions defined as in run 32

Neither  $\alpha$  nor <sup>206</sup>Pb recoils backgrounds in the reference regions

With rate of previous run:

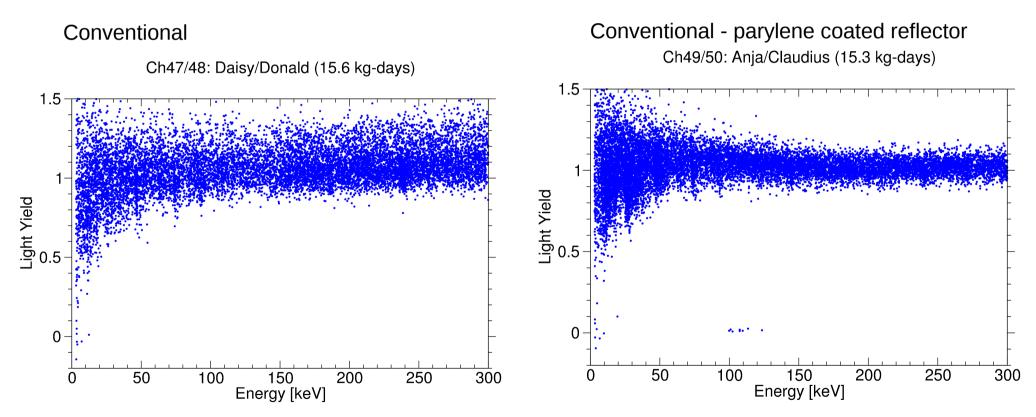
- ~5  $\alpha$ 's expected
  - 5 <sup>206</sup>Pb recoils expected

### First glance at data - Conventional modules

Data until December used for defining cuts which then will be applied later in blind analysis

Data from July 30 to October 9

#### Very preliminary !!



- <sup>206</sup>Pb recoils backgrounds at  $\geq$  100keV expected from crystal surface
- · background in conventional modules lower than in previous run

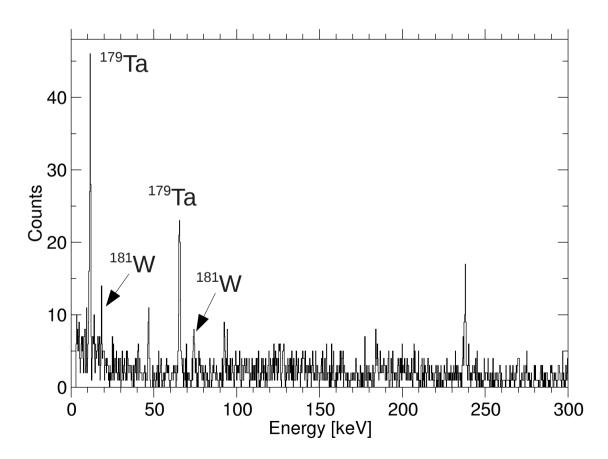
### First glance at data – Self grown crystals

Data until December used for defining cuts which then will be applied later in blind analysis

#### Data from July 30 to October 9

#### Very preliminary !!

- low intrinsic background level
  Possible reduction of discrimination threshold in Dark Matter analysis
- lines from cosmogenic activation
- excellent energy resolution
- precise energy calibration at low energies



- Data taking since July 30th 2013
  Stable running conditions during night time/weekend Stability of running affected during day time
- Data until December used for defining cuts (\*)
- Expect ~2000 kg-days of data within 2 years
- Either confirm or reject low mass WIMP scenario with high confidence
- Competitive limit if no low mass WIMP is found

In parallel:

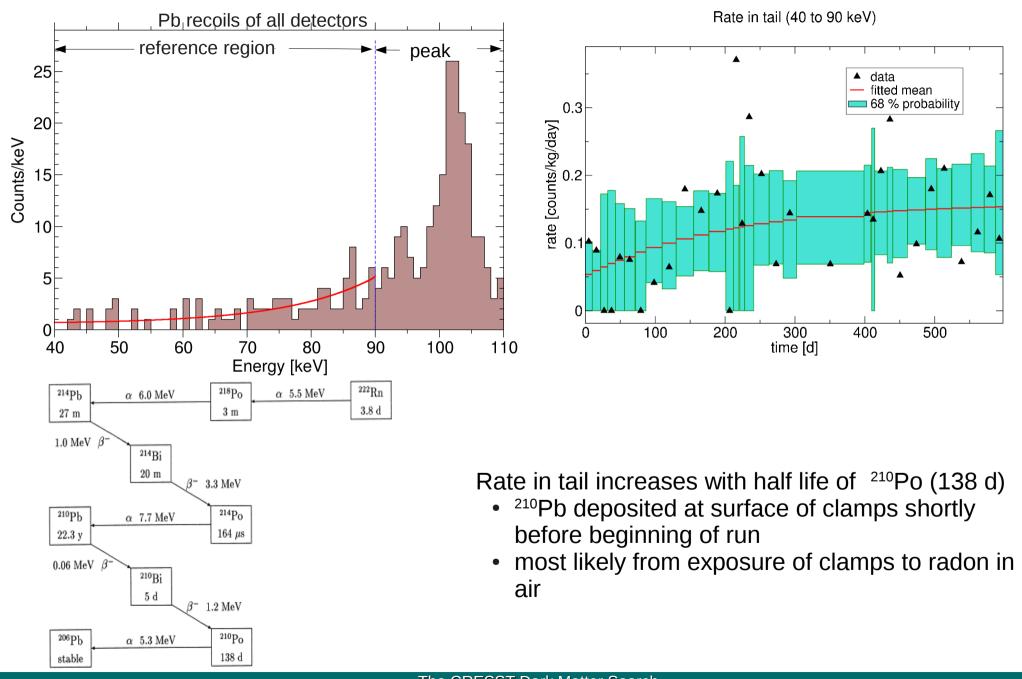
Detectors' development program to go beyond the achievements of run33

(\*)In case of overwhelming superiority of the new designs the possibility of replacing the conventional ones will be considered

#### **Pictures**



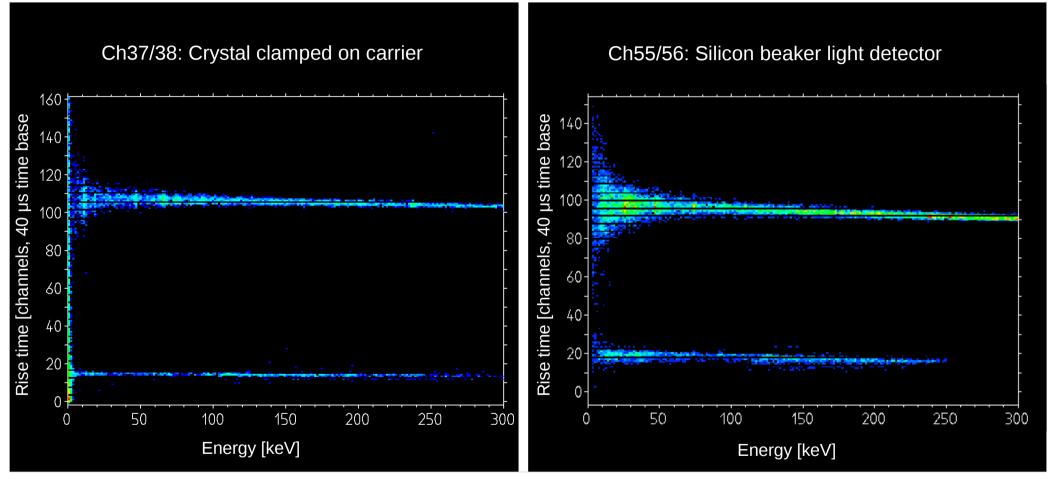
# Passive background reduction – <sup>206</sup>Pb recoils



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# **Carrier events discrimination**

Simple rise time to pulse maximum efficiently distinguishes events in thermometer carrier



Identify possible stress-relaxation events from contact between Parylene coated clamps and carrier Identify possible recoiling nuclei in the carrier crystal originating from the clamps

