

Solar Neutrinos with Borexino at LNGS

Current Results/Future Opportunities

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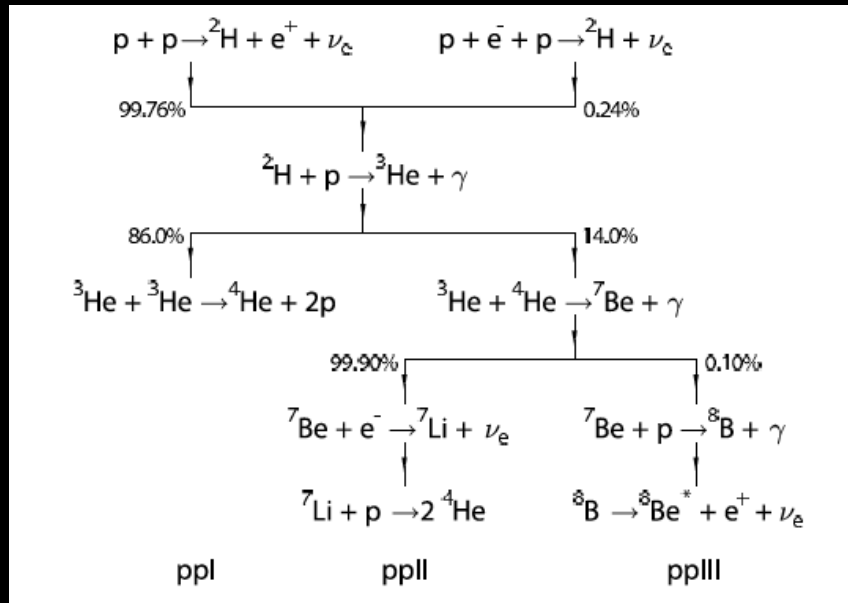
Borexino Collaboration



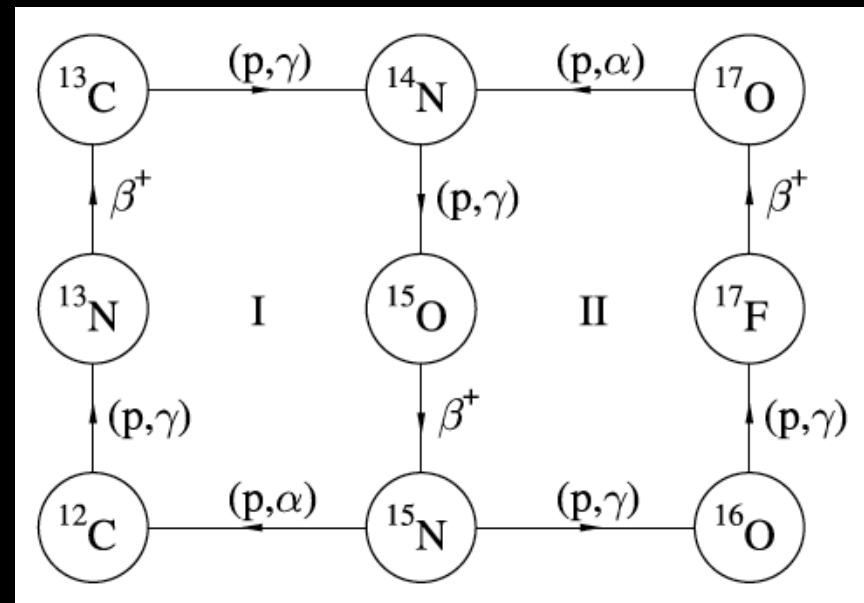
A 20+ year successful collaboration between U.S. and Italian + European Groups

Solar Nuclear Fusion Cycles

The pp cycle

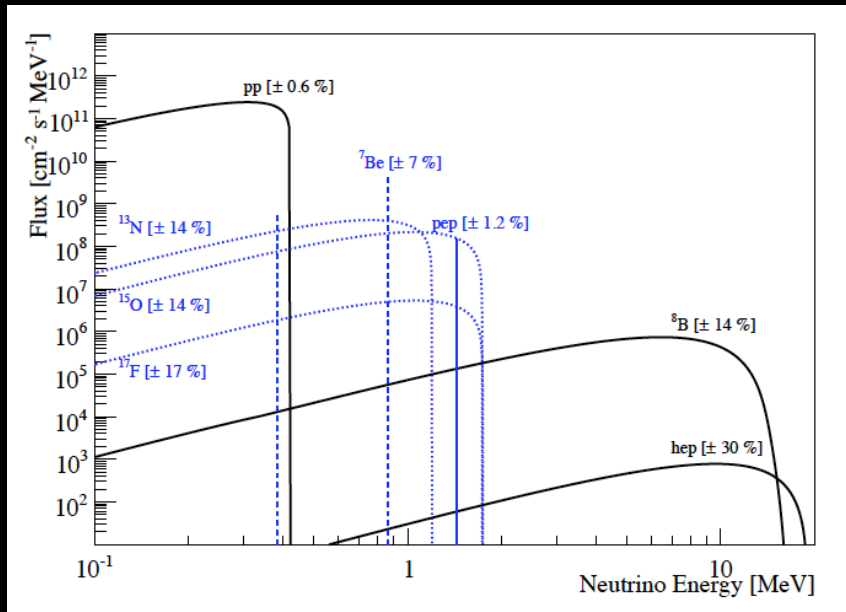


The CNO cycle

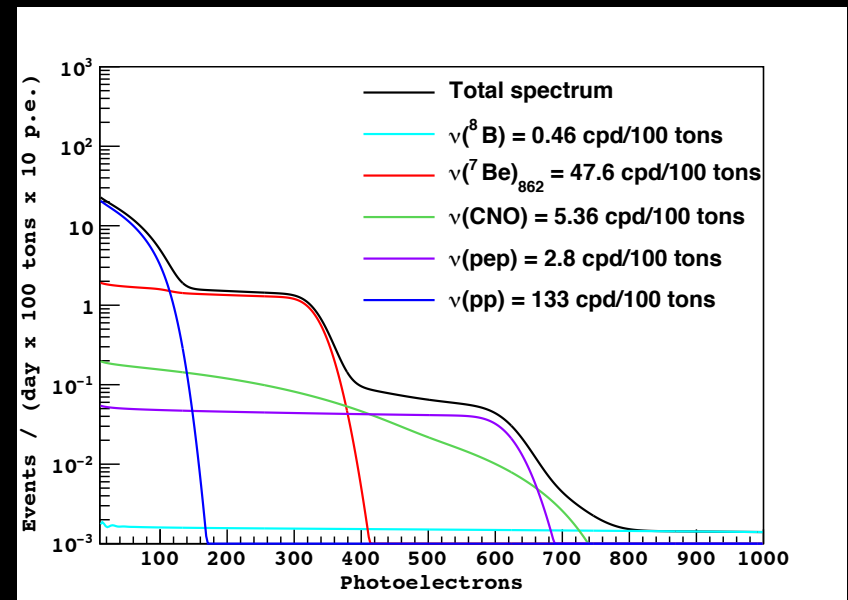


Solar Neutrino Spectra

Neutrino Energy Spectrum



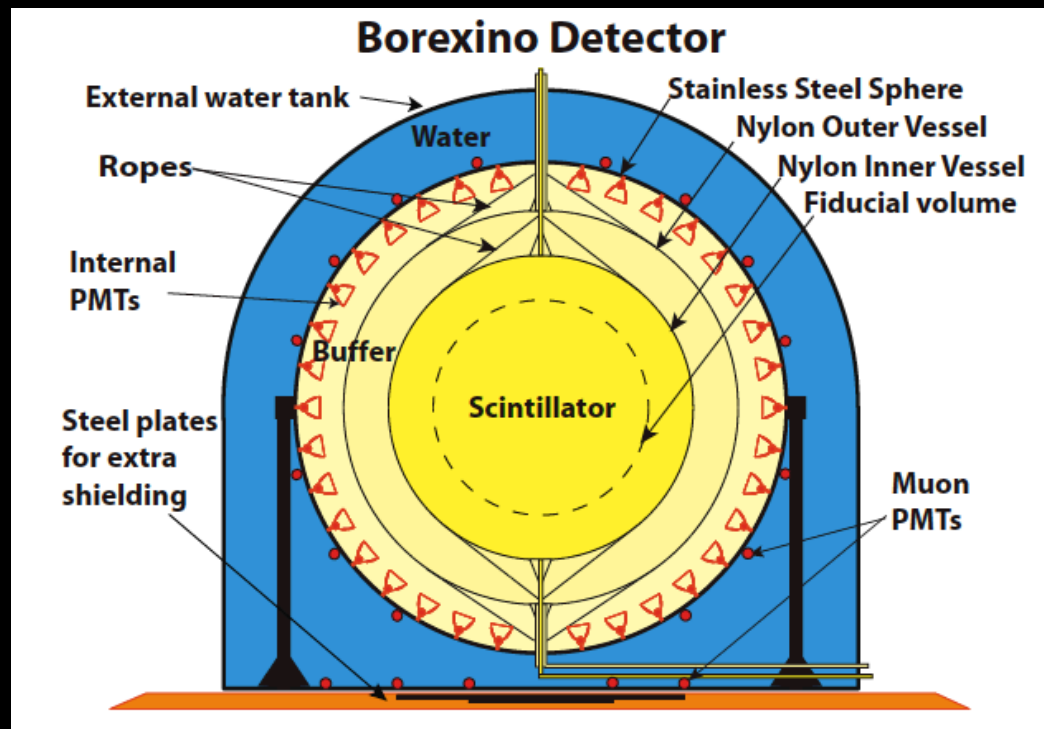
Neutrino-Electron Elastic Scattering Energy Spectrum



Overview of the Borexino Detector

(Mostly Active Shielding)

- Shielding Against Ext. Backgnd.
 - Water: 2.25m
 - Buffer zones: 2.5 m
 - Outer scintillator zone: 1.25 m
- Main backgrounds: in Liq. Scint.
 - $^{14}\text{C}/^{12}\text{C}$
 - 10^{-18} g/g. cf. 10^{-11} g/g in air CO_2
 - U, Th impurities
 - ^{222}Rn daught (^{210}Pb , ^{210}Bi , ^{210}Po)
 - ^{85}Kr
- Light yield (2200 PMT's)
 - Detected: 500 p_e/MeV (~4%)
- Pulse shape discrimination.
 - Alpha-beta separation



Borexino Energy Spectra

PRL 107 141302 (2011)

Data are based on 740.7 live days between May 16, 2007 and May 8, 2010.

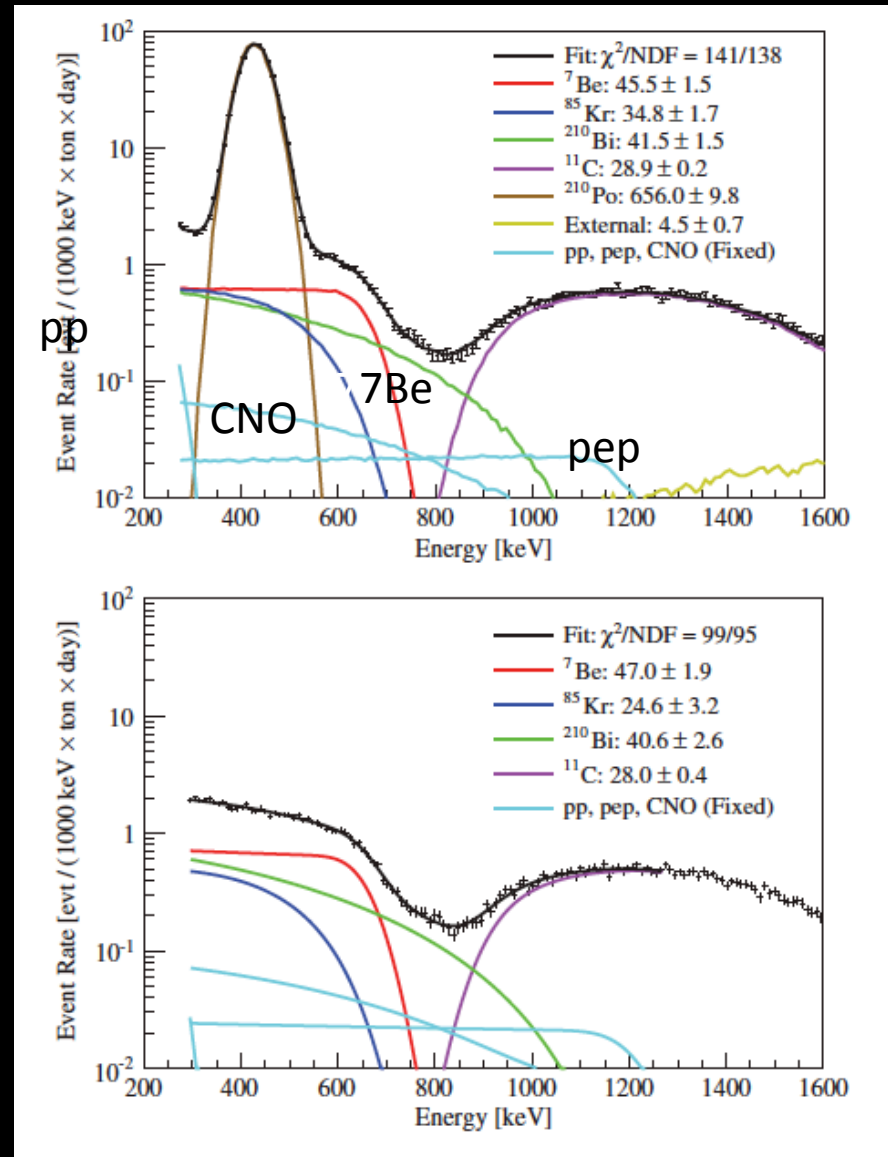
Prominent backgrounds are:

^{210}Po ^{210}Bi ^{85}Kr & ^{14}C (not shown)

CNO obscured mainly by ^{210}Bi due to similar shape.

The ^{210}Po alpha rate was very high but saved by alpha/beta pulse shape discrimination.

The pep was measured by applying cuts to reduce the ^{11}C background.



Borexino Neutrino Measurements

Solar Neutrino rates (cpd/t)

- ${}^7\text{Be}$: 0.460 ± 0.023 Phys. Rev. Lett. 107 141302 (2011)
- ${}^8\text{B}$: 0.0022 ± 0.0004 Phys. Rev. D 82, 033006 (2010)
- pep: 0.031 ± 0.005 Phys. Rev., Lett. 108, 051302 (2012)

Geo-neutrinos

- Total 14.3 ± 4.4 events Phys. Letts. B722, 295 (2013)

Electron Neutrino Survival Probabilities

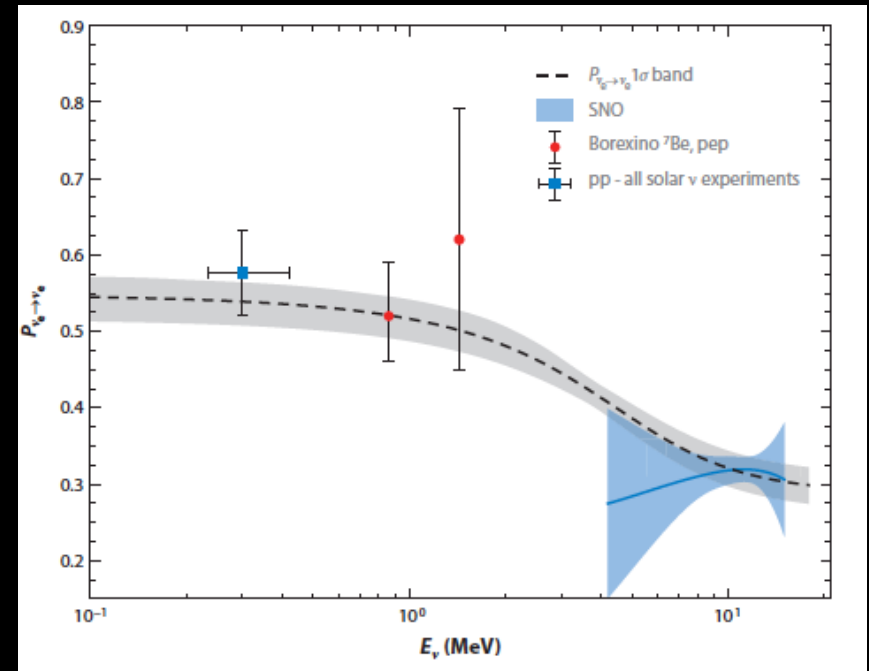
$P_{ee} = 1.0$ no neutrino oscillations

The MSW predicted transition from vacuum to matter-enhanced oscillations generally agrees with measurements of pp, ${}^7\text{Be}$, pep, and ${}^8\text{B}$ neutrinos.

The centroid of SNO's ${}^8\text{B}$ low threshold data falls below the MSW curve with decreasing energy, but the significance is not high.

The transition region is sensitive to new physics. e.g non-standard interactions, sterile neutrinos.

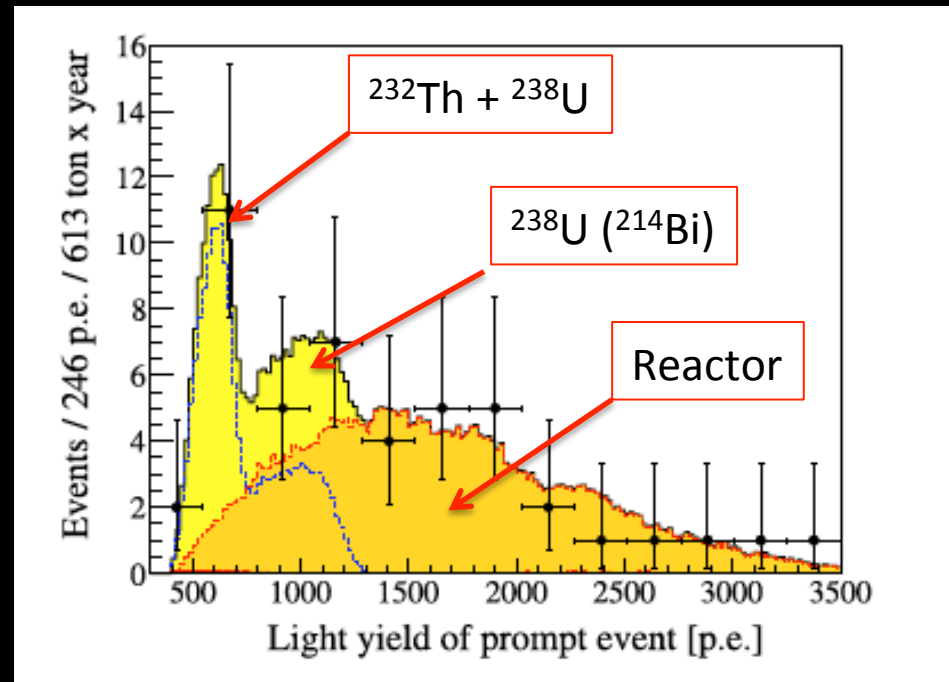
Improving pep, and low energy ${}^8\text{B}$ data motivated by sensitivity to new physics.



W.C. Haxton, R.G. Hamish Robertson,
and Aldo M. Serenelli, *Annu. Rev. Astron.*
Astrophys. 2013. 51:21–61

Geo-neutrinos

- Geo-neutrino data from 1353 days of data.
 - Physics Letters B 722 (2013) 295–300
- Geo-neutrino events in yellow; reactor events in orange.
- Antineutrinos detected by delayed coincidences from charged current reaction
 - $\bar{\nu} + p \rightarrow e^+ + n$
- Total 46 events detected
 - 14.3 ± 4.4 geo-neutrinos
 - 31.2 ± 6.5 reactor neutrinos



Success!

- Borexino delivered its original goal of measuring ${}^7\text{Be}$ neutrinos with 5% accuracy.
- Lower backgrounds than originally sought allowed measurements of ${}^8\text{B}$ and pep neutrinos and geo-neutrinos.
 - Original goal: $[\text{U}], [\text{Th}] \sim 5 \times 10^{-16} \text{ g/g}$
 - Achieved: $[\text{U}], [\text{Th}] \sim 2 \times 10^{-17}, < 7 \times 10^{-18} \text{ g/g}$
 - Background: 20 – 70 times lower than sought!

What's next?

- Lower background by scintillator re-purification.
- Solar program: detect all solar neutrinos!
 - pp neutrinos
 - CNO detection or upper limit
 - Improve accuracy of ${}^7\text{Be}$, pep
- SOX source experiment
 - Short baseline neutrino oscillations

Future Solar Neutrino Program

1. Detect/Measure CNO neutrinos

- Confirm proton burning process for stars more massive than Sun.
- CNO neutrino rate determines C, N abundances in core.
 - Clarify the “solar metallicity problem”
 - Haxton, Serenelli, and others suggest planetary accretion could explain difference in metal abundance in photospheric and helioseismology data.
 - High metallicity rate would support such possibilities.

2. Improve accuracy of ${}^7\text{Be}$ and pep measurements.

- Neutrino rates could provide accurate prediction of power produced by solar nuclear fusion reactions.
- Comparing neutrino-related power to total solar luminosity could reveal new physics? Sterile neutrino tension? (H. Robertson)

3. Non-standard interactions in vacuum- matter oscillations transition.

Going for lower background...

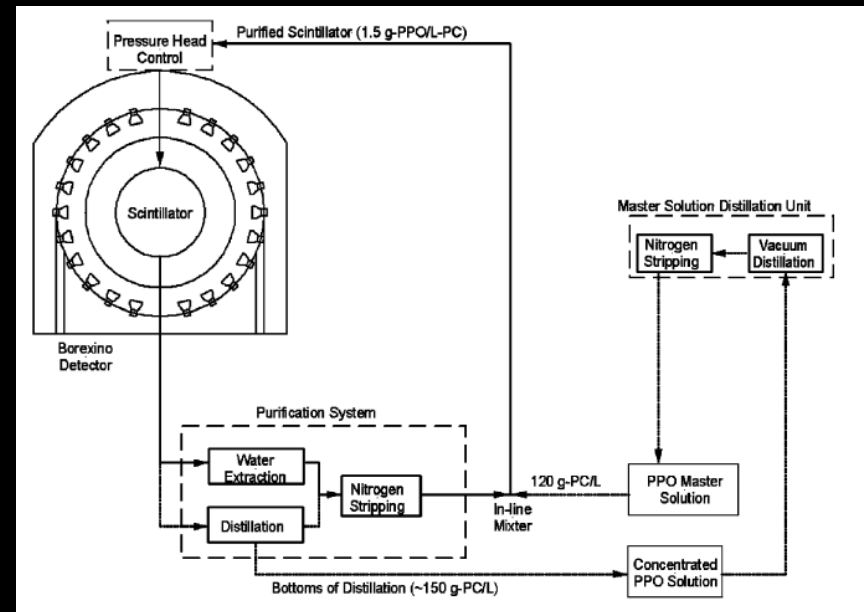
- Initial background was achieved by in-line purification as detector was being filled.
 - Distillation, nitrogen stripping and water extraction/ distillation of PPO.
- Same purification system was built for re-purification in loop mode.
- 2010: Re-purification started to achieve lower background.



The scintillator purification “skids”

Background Reduction with Loop Purification of Liquid Scintillator

- “Loop” purification is achieved by draining fluid from bottom of vessel, passing it through purification system, and returning to the top.
- Processes available are:
 - Water extraction or distillation
 - Nitrogen stripping (^{85}Kr)

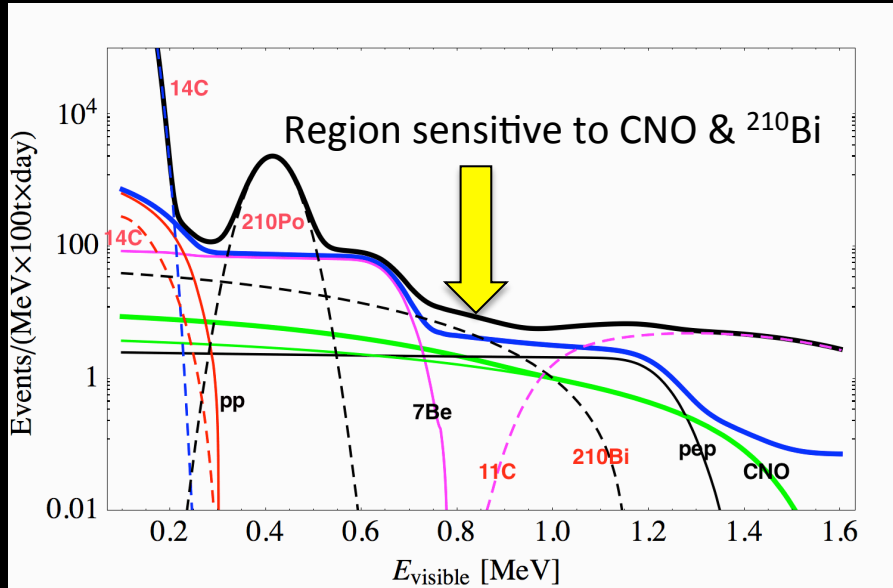


Results of 6 cycles of Re-purification

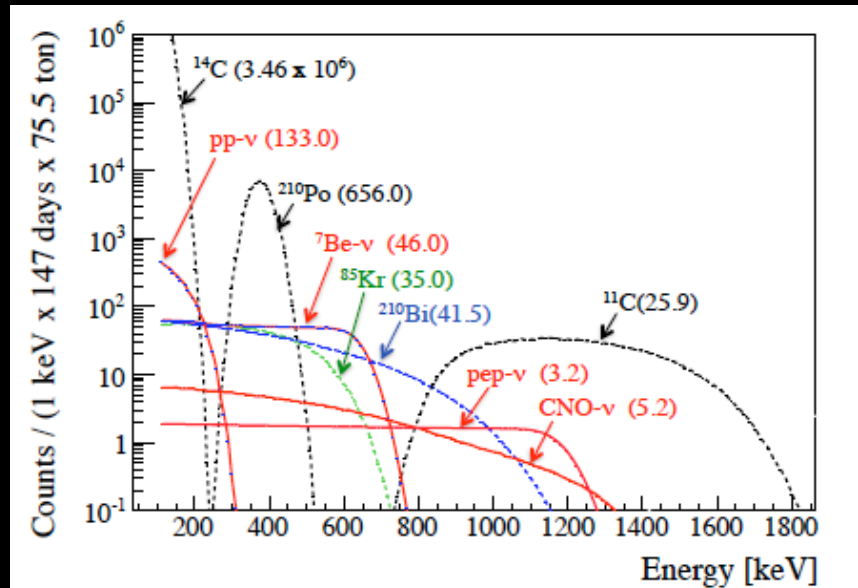
- ^{85}Kr : 30 cpd/100t \rightarrow < 5 cpd/100t
- ^{238}U (^{226}Ra): $[(53 \pm 5) \rightarrow (1.6 \pm 0.6)] \times 10^{-19}$ g/g
 ^{214}Bi - ^{214}Po Reduction factor = 33. 7 events/100t/yr.
- ^{232}Th : $[(3.8 \pm 0.8 \rightarrow < 1.2)] \times 10^{-18}$ g/g (95% CL)
 ^{212}Bi - ^{212}Po Reduction factor > 3 . 2 ev./100t in 600 d
- ^{210}Bi : 70 cpd/100t \rightarrow 20 ± 5 cpd/100t
- ^{210}Po : Essentially not reduced! **Why?**

Rates before purification are based on 153.6 ton-yr exposure taken in 740.7 d between May 16, 2007 and May 8, 2010. See Borexino Coll. arXiv 1308.0443v1.

Backgrounds after & before Water Extraction + N₂ Stripping



After re-purification 2012-2013
(with ¹¹C cuts)



Before re-purification 2008-2010
Without ¹¹C cuts. See arXiv1308.0443v1.

Water Extraction Worked!!

- With current background we will make improvements in ${}^7\text{Be}$ and pep measurements, and improve the upper limit on CNO.
- However, lower background is needed for a strong thrust toward CNO and precise ${}^7\text{Be}$ and pep measurements.

“Purified” Water Surprises

- LNGS round water has relatively high levels of ^{222}Rn (10,000 Bq/m³) and its daughters ^{210}Pb - ^{210}Bi - ^{210}Po (1-10 Bq/m³)
- The inability to remove ^{210}Po and some problems with ^{210}Pb led to investigations of the water.
- Surprise #1
 - Water purification by de-ionization is not very effective in removing ^{210}Pb , and especially not effective for ^{210}Po .
- Surprise #2
 - ^{210}Po has a volatile compound that is not removed by simple distillation.

Surprise #1

- Reduction of ^{210}Pb (^{210}Bi) was inferred from ICPMS measurements of stable ^{208}Pb .
 - ^{208}Pb reduced by x 700 by de-ionization (modest)
 - Typical reduction is $\times 10^{5-6}$ for ionized salts.
- Reduction of ^{210}Po was measured by direct counting of alphas, using deposition on silver.
 - ^{210}Po reduced by X10 by de-ionization (very poor!)

Surprise #2

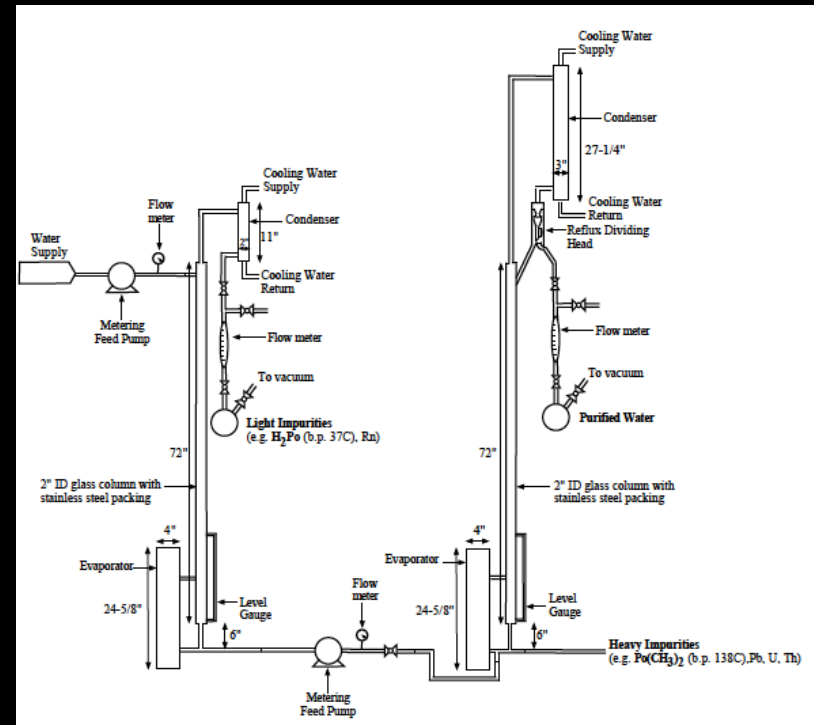
- Publications claim that some ^{210}Po in well water is a volatile compound produced by biological (bacteria) processes:
 - “Bio-Volatilization of Polonium: Results from Laboratory Analyses”, N. HUSSAIN, T. G. FERDELMAN, T. M. CHURCH, Aquatic Geochemistry 1: 175-188, 1995
 - “Impact of microorganism on polonium volatilization”, N. Momoshima, et. al. J. Radioanalytical and Nucl. Chem. 272, 413 (2007)
- Dimethyl polonium: Boiling point $\sim 138\text{ C}$ (est.)
- Poor removal of ^{210}Po by de-ionization and distillation due to volatile polonium compound could explain long-standing problems Borexino has had with this isotope.

Historical Problems with ^{210}Po

- Poor removal of ^{210}Po by de-ionization and distillation due to volatile polonium compound explains long-standing problems.
 1. Initial filling of Borexino (water was used on PPO)
 - ^{210}Po (6000 cpd/100t), ^{210}Pb (~15 cpd/100t)
 2. Water extraction:
 - Not effective to remove ^{210}Po
 - Reasonably effective to remove ^{210}Pb

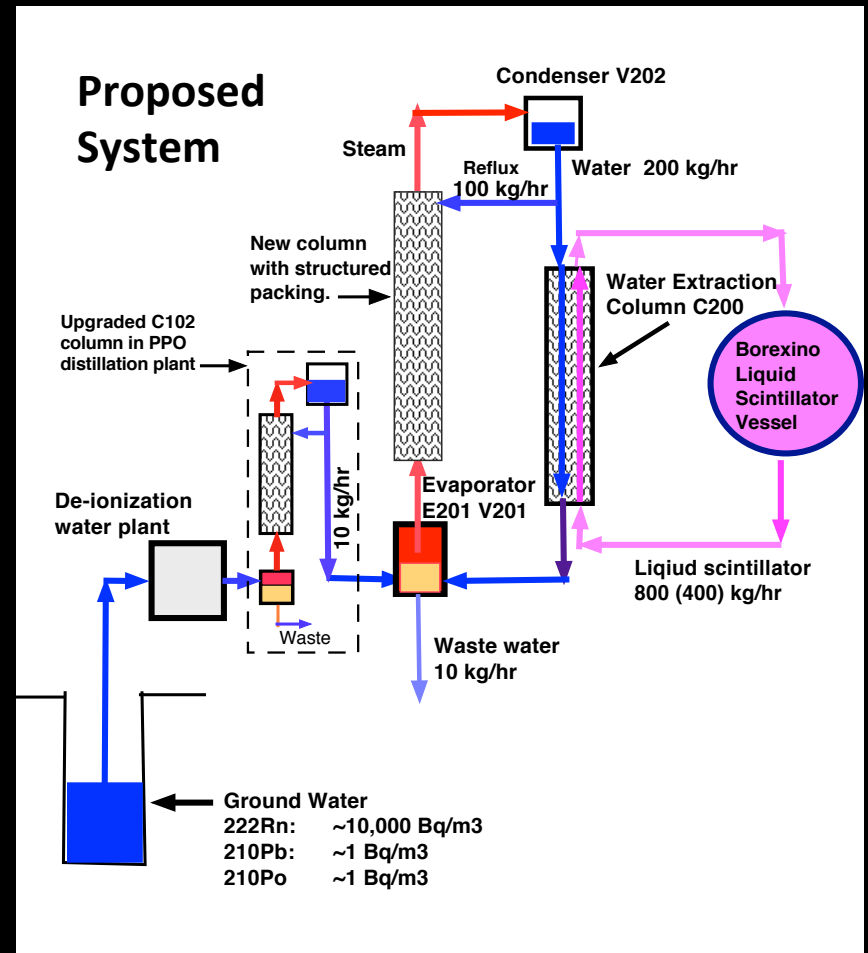
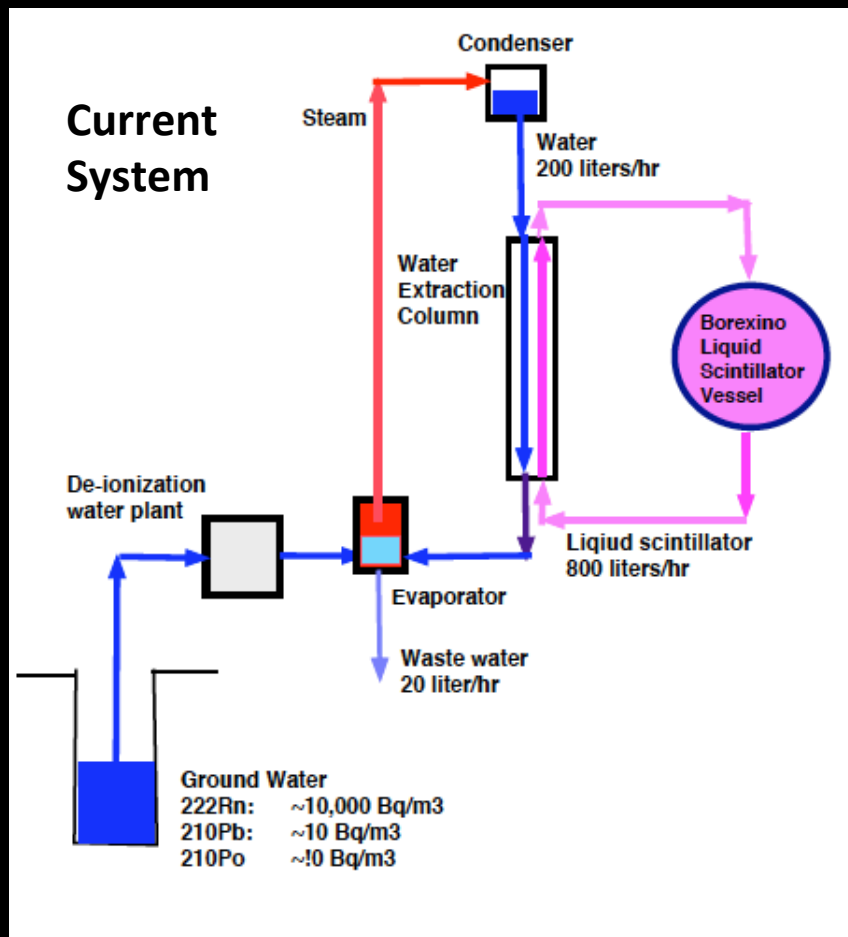
New Distillation System Tested to Remove ^{210}Po and ^{210}Pb from Water.

- A small-scale distillation system was designed and tested at Princeton to remove dimethyl polonium from ground water.
 - A tall column with structured packing and high reflux was designed for x1000 reduction.
- Princeton well water, noted for its high radon levels, made it easy to measure the polonium before and after distillation.
- Results:
 - When system was operated as simple evaporator, the ^{210}Po was not separated. Explains many problems.
 - When operated with full column and high reflux, the output water showed no measurable ^{210}Po .
 - Reduction factor > 300 .



Small prototype distillation system with two 6-foot column 2 l/hr capacity.
Students : B.Russell, C. Aurup, W. Taylor

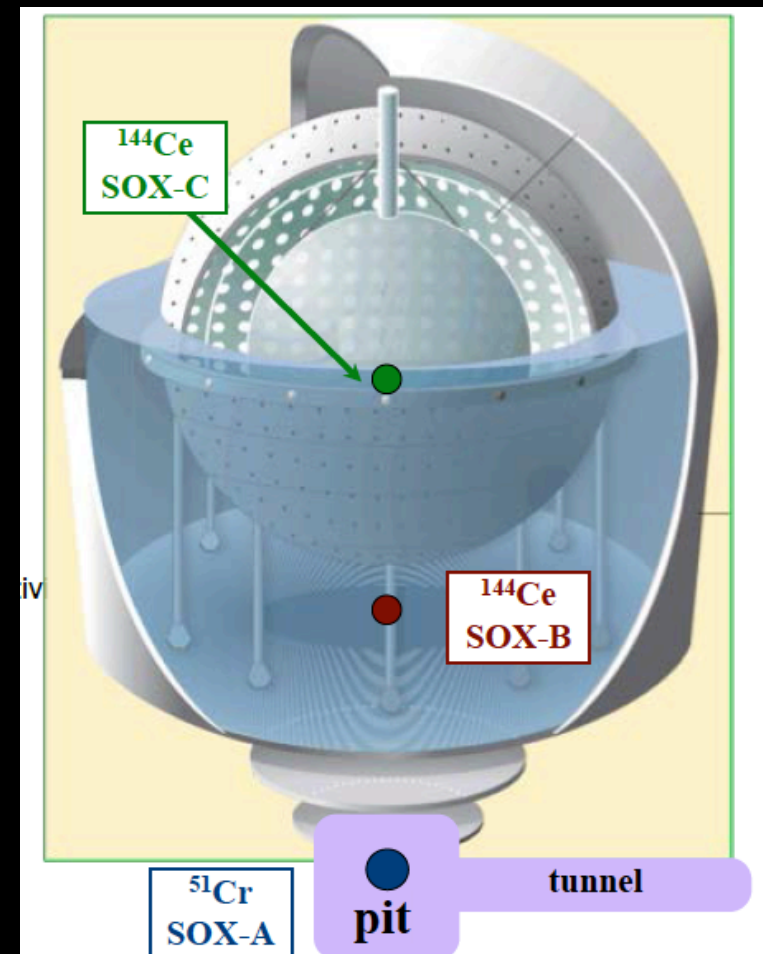
Current & Proposed Upgraded Water Extraction Systems



The Borexino SOX Experiment

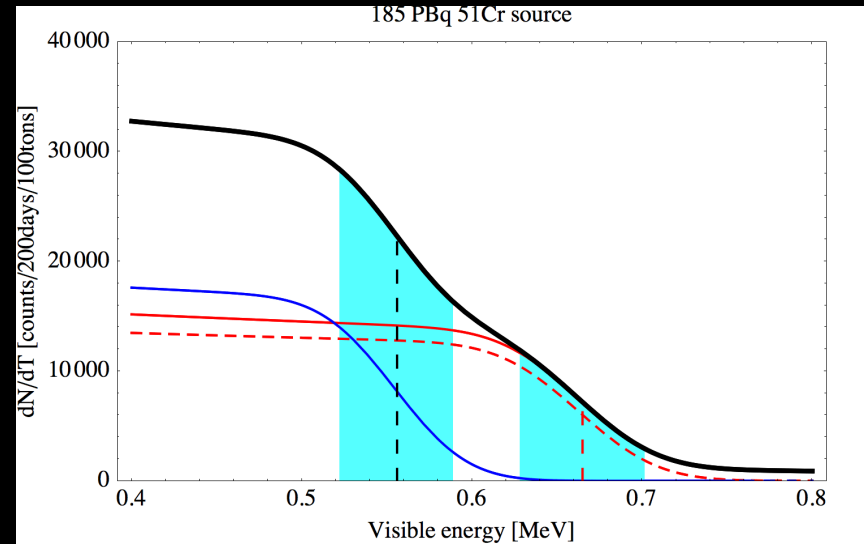
Short distance ν_e Oscillations with BoreXino

- The SOX-A experiment will deploy a 10 MCi ^{51}Cr source below Borexino to search for short baseline neutrino oscillations, due to sterile neutrinos.
- The lower background needed for solar neutrinos will benefit SOX with higher signal/background ratio.
- The mono-energetic ^{51}Cr neutrino will benefit the solar program by providing an excellent energy calibration
- SOX-A is funded by the European Research Council.



SOX ^{51}Cr Source and Solar Neutrinos

- The SOX ^{51}Cr source will provide a precise energy calibration for the ^7Be measurement.
- The 2.7% uncertainty due to the energy scale will be reduced significantly (<1%).
- With lower background, and more accurate energy calibration, a 3% measurement of the ^7Be rate could be possible.
- Uncertainty in pep will also be reduced significantly.



Main systematic errors in ^7Be measurement:

Fiducial volume:	+0.5/-1.3 %
Energy response:	2.7%
Fit methods:	2.0%
Statistical uncertainty:	3.3%

Status of Solar/Neutrino Science

- Bethe's pp fusion cycle
 - Observed 3 of the 4 neutrinos.
 - pp still not detected.
 - Precise comparison of neutrino & EM luminosity needed.
- Neutrino oscillations
 - Solar neutrinos consistent with MSW theory
 - Precise data on Vacuum-matter effect transition needed.
- CNO physics
 - Solar metallicity problem still unexplained
- A precision phase of Borexino is coming.
 - Stay tuned- it could address outstanding issues.

The End