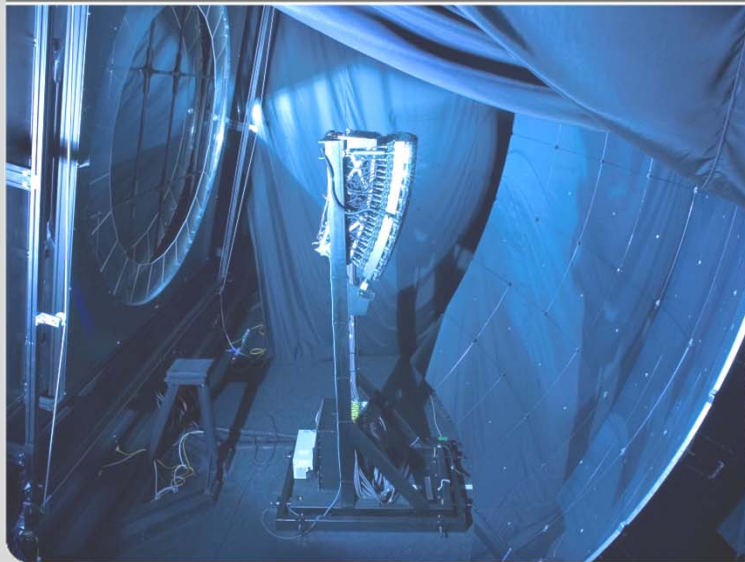


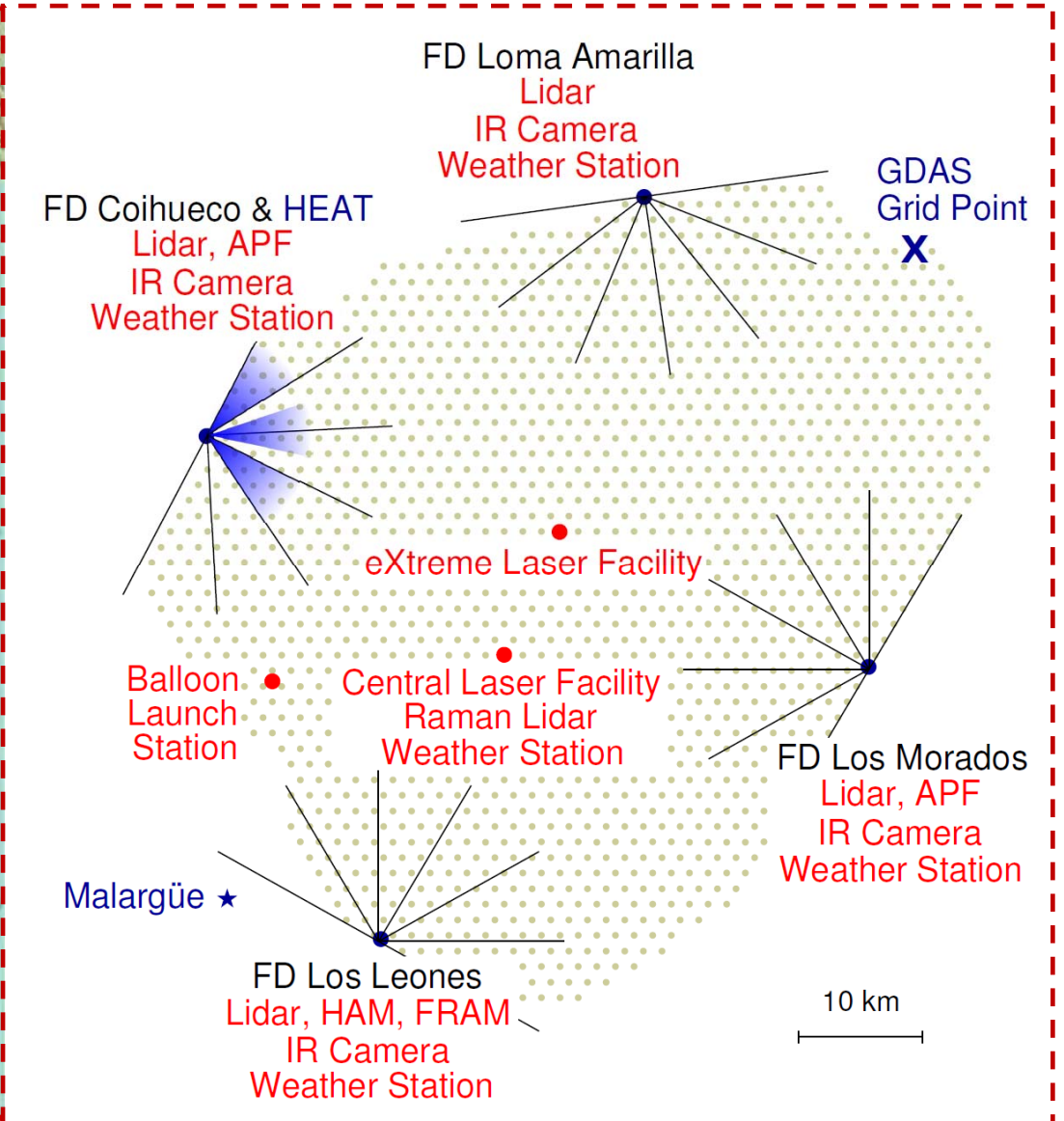
Atmospheric Monitoring at the Pierre Auger Observatory

Bianca Keilhauer for the Pierre Auger Collaboration

AtmoHEAD: Atmospheric Monitoring for High-Energy Astroparticle Detectors - Padova, 19-21 May 2014



The Pierre Auger Observatory



Main Detector Components

Surface Detectors

- ↑ 100 % duty cycle
- ↑ acceptance = geometric
- ↓ only last stage of shower development observed
- ↓ energy scale model dependent

Fluorescence Detectors

- ↓ ≈ 15 % duty cycle
- ↓ acceptance depends on distance and atmosphere
- ↑ observation of longitudinal shower development
- ↑ (almost) model independent

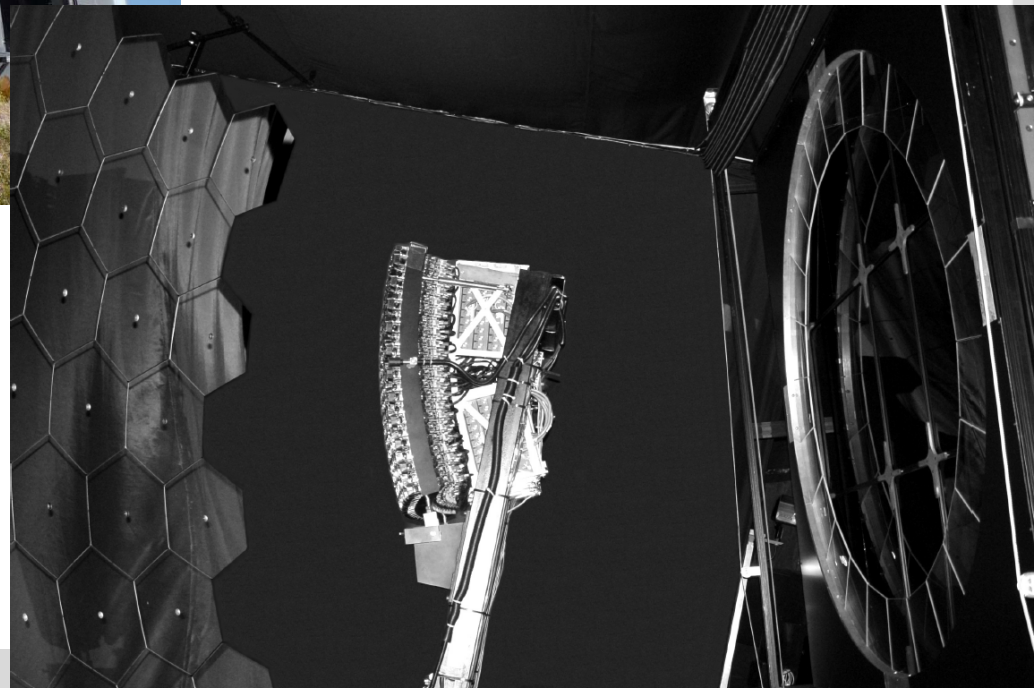


Main Detector Components



Fluorescence Detectors

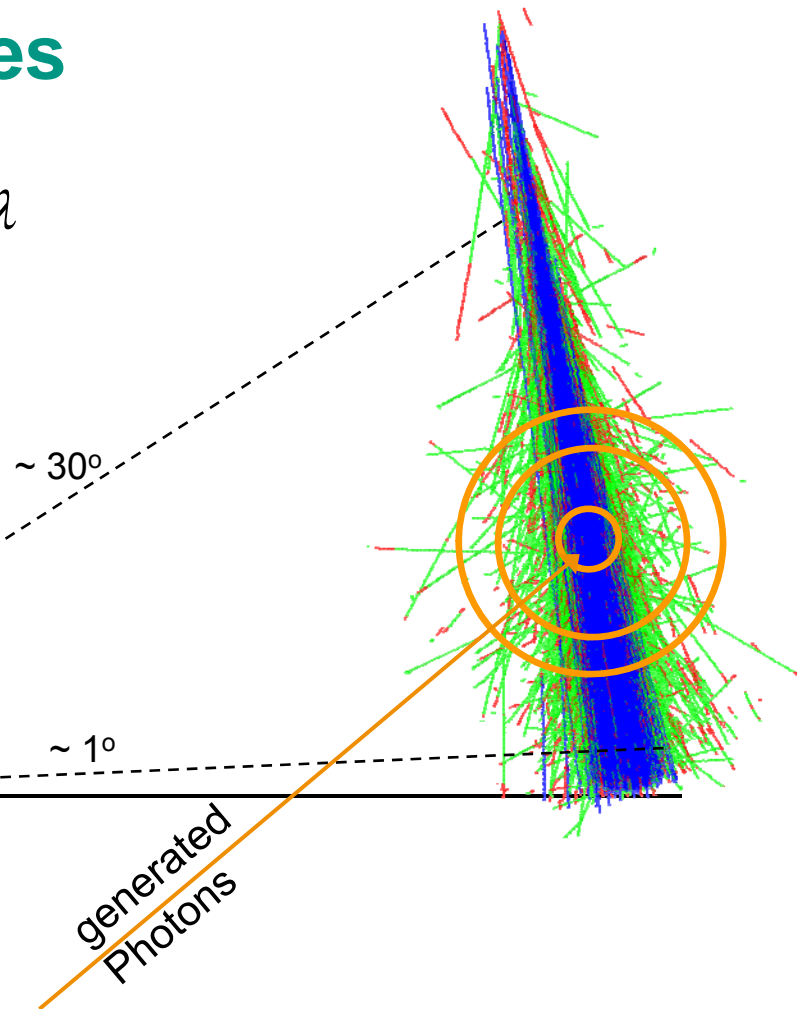
- ⬇️ $\approx 15\%$ duty cycle
- ⬇️ acceptance depends on distance and atmosphere
- ⬆️ observation of longitudinal shower development
- ⬆️ (almost) model independent



Measuring Principle of EAS with Fluorescence Telescopes

$$\frac{dN_\gamma}{dX} = \int \frac{d^2 N_\gamma^0}{dX d\lambda} \cdot \varepsilon_{FD}(\lambda) \cdot \tau_{atm}(\lambda, X) d\lambda$$

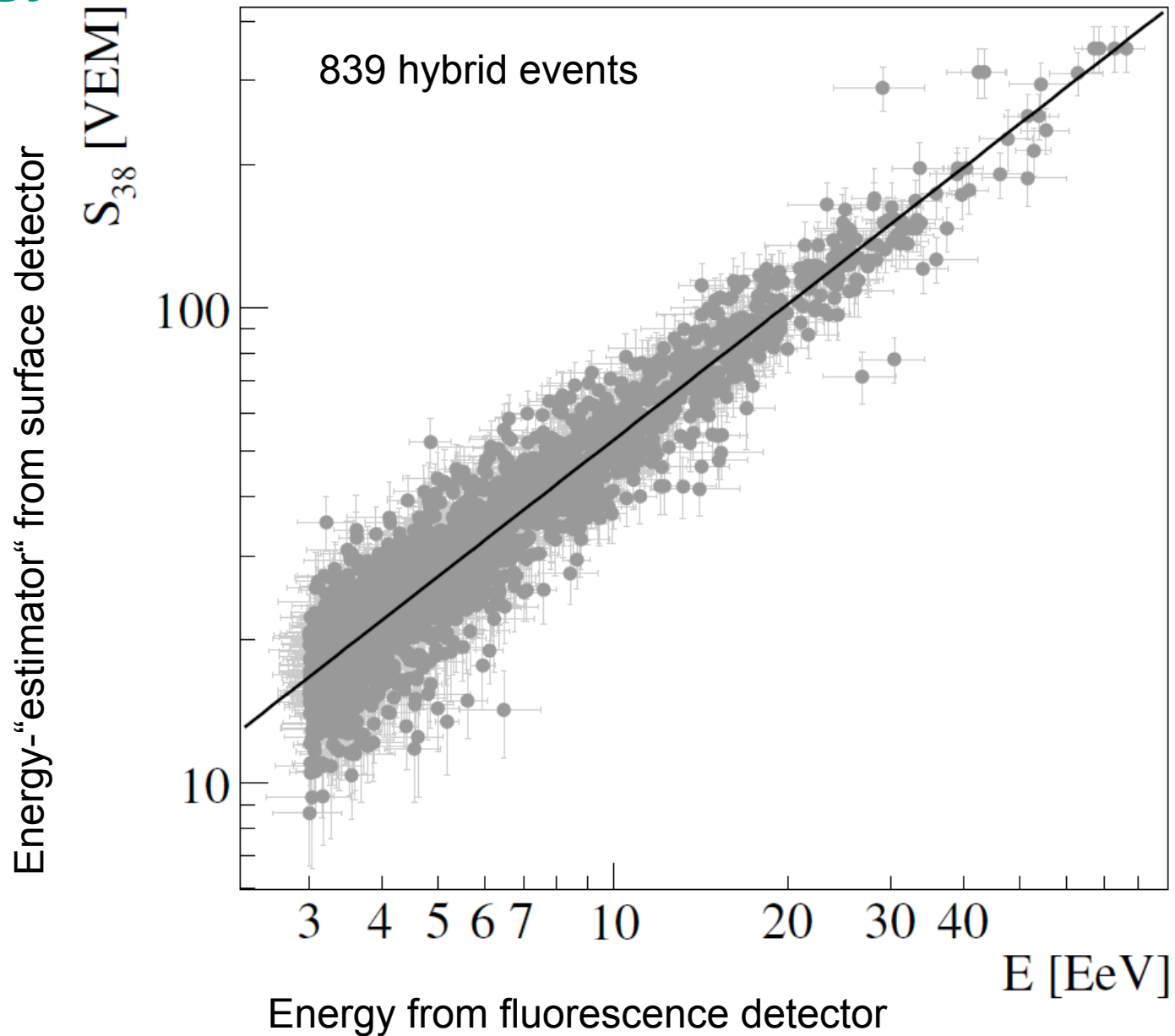
Photons
at Detector



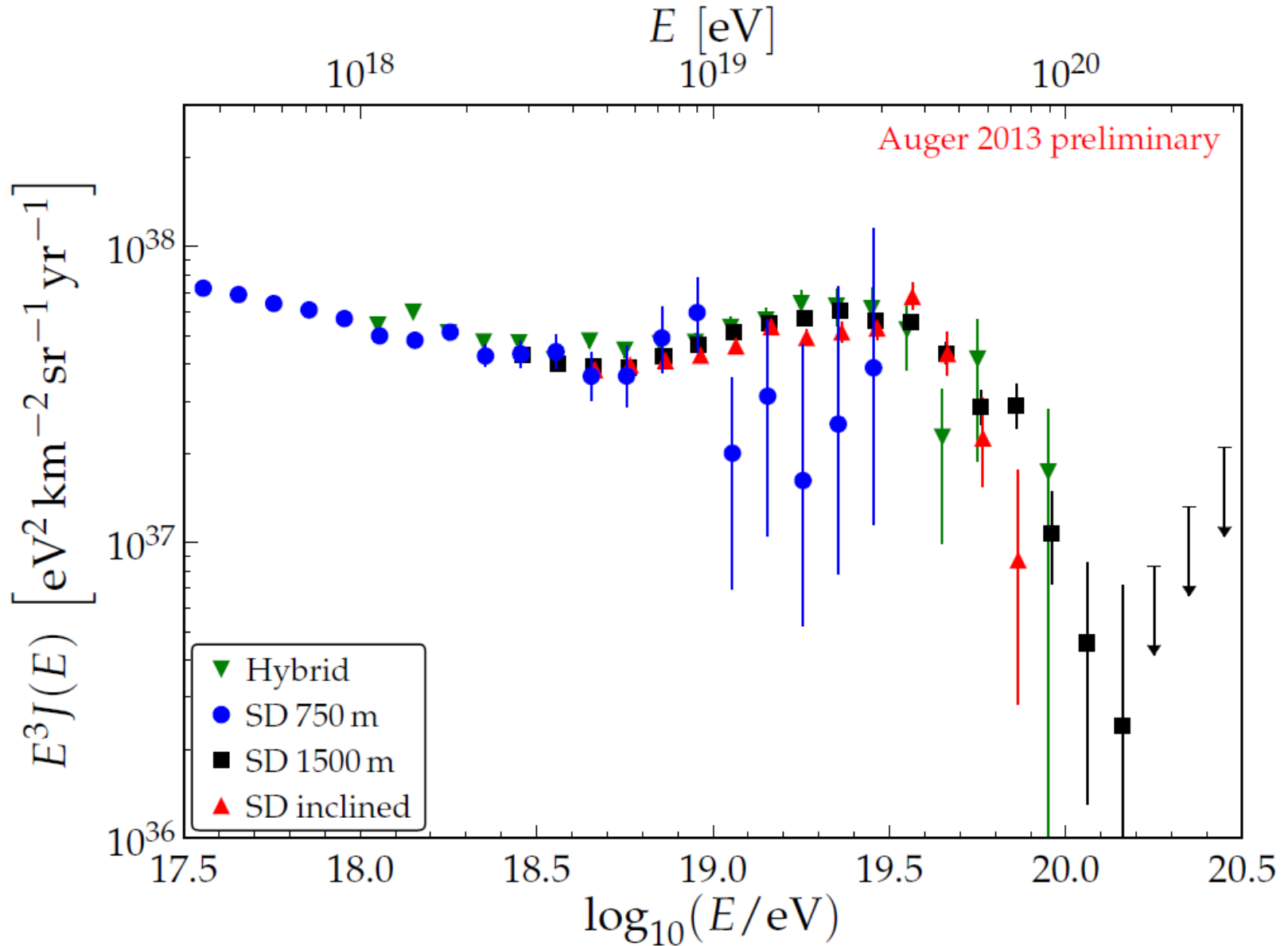
generated
Photons

$$\frac{d^2 N_\gamma^0}{dX d\lambda} = \int Y(\lambda, P, T, u, E) \cdot \frac{dN_e(X)}{dE} \cdot \frac{dE_{dep}}{dX} dE$$

Energy Calibration



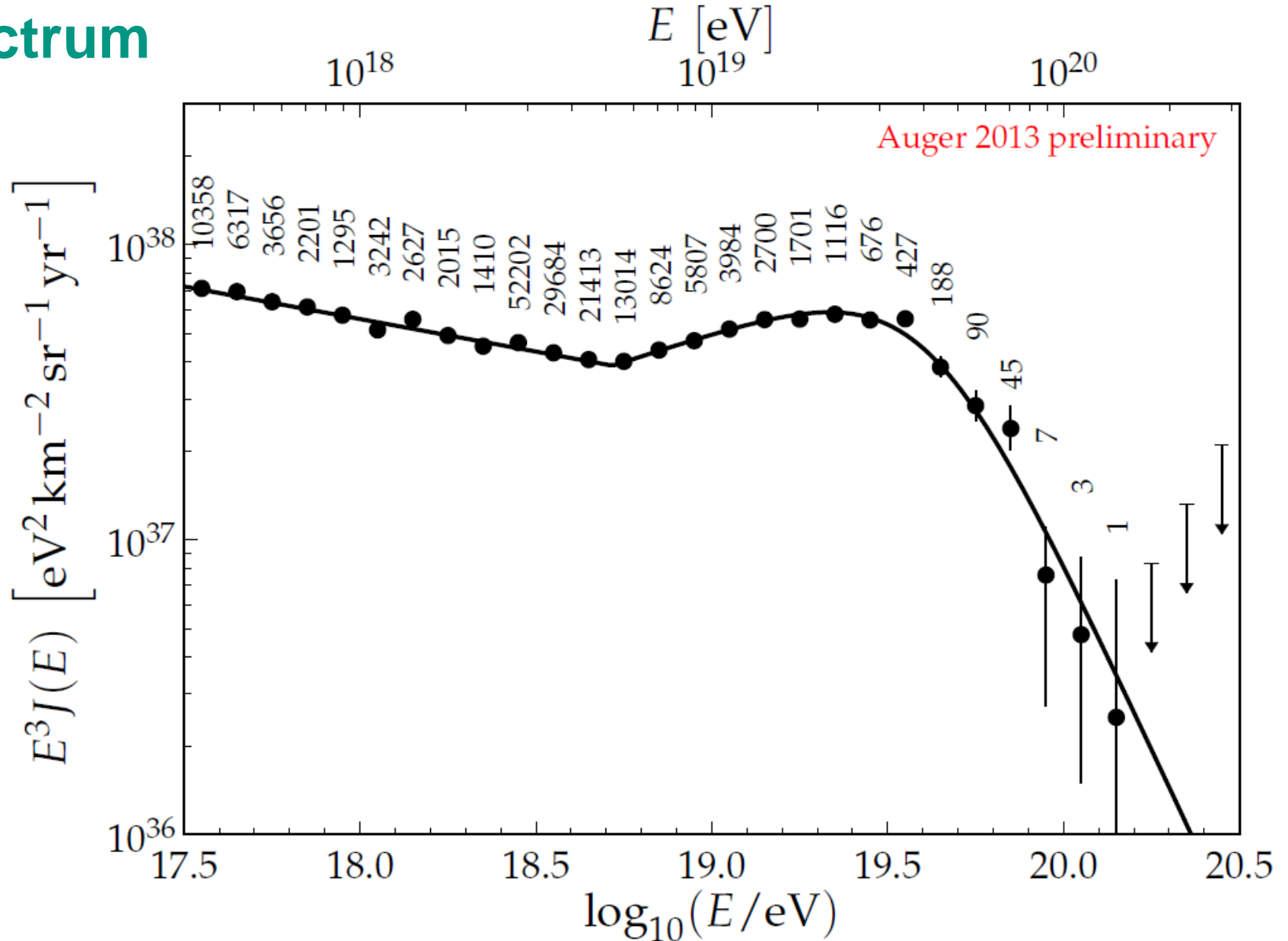
Flux spectra

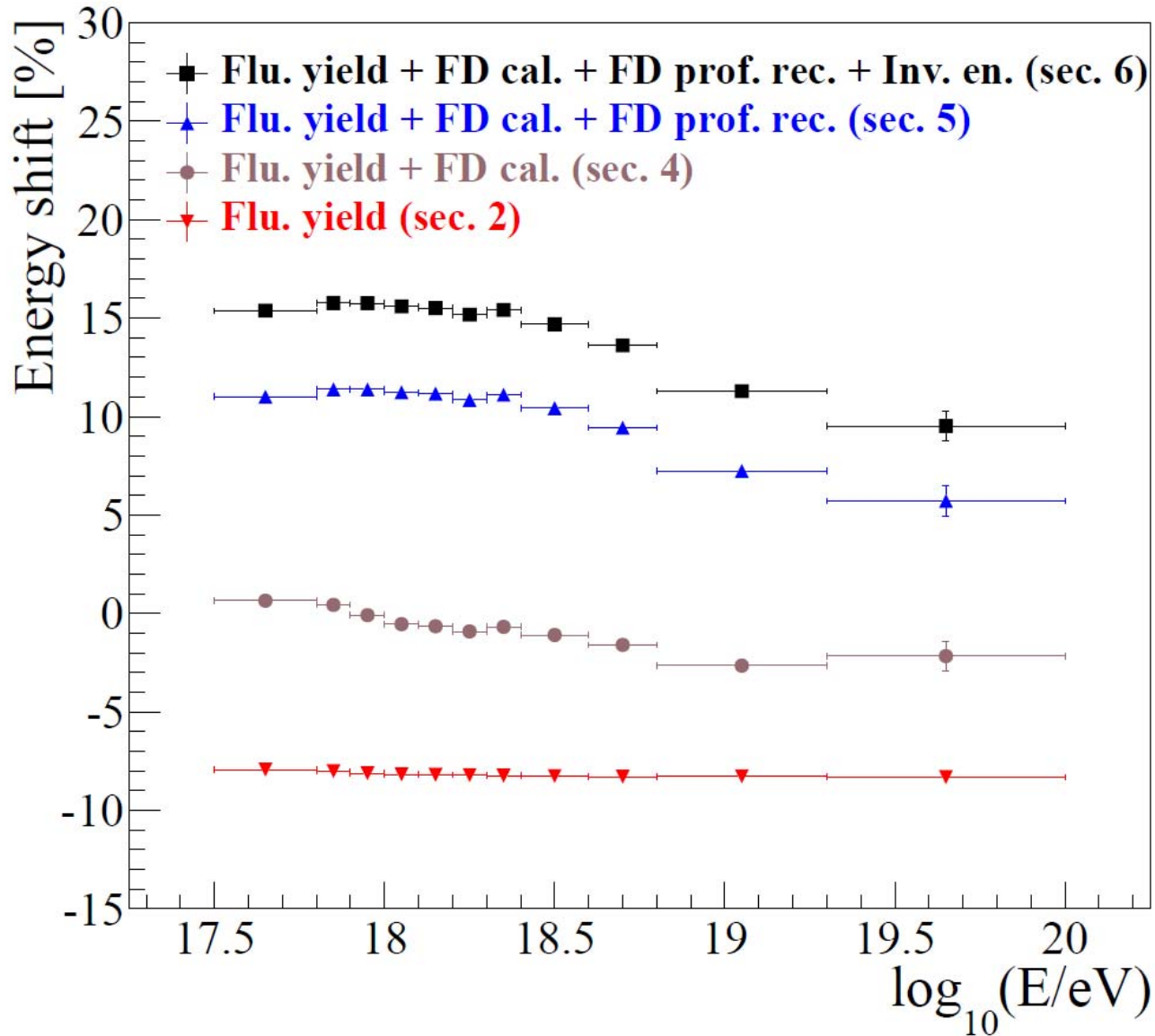


Combined Spectrum

A. Schulz *et al.* Energy spectrum measured at the Pierre Auger Observatory
33RD INTERNATIONAL COSMIC RAY CONFERENCE, RIO DE JANEIRO 2013

ICRC
2013





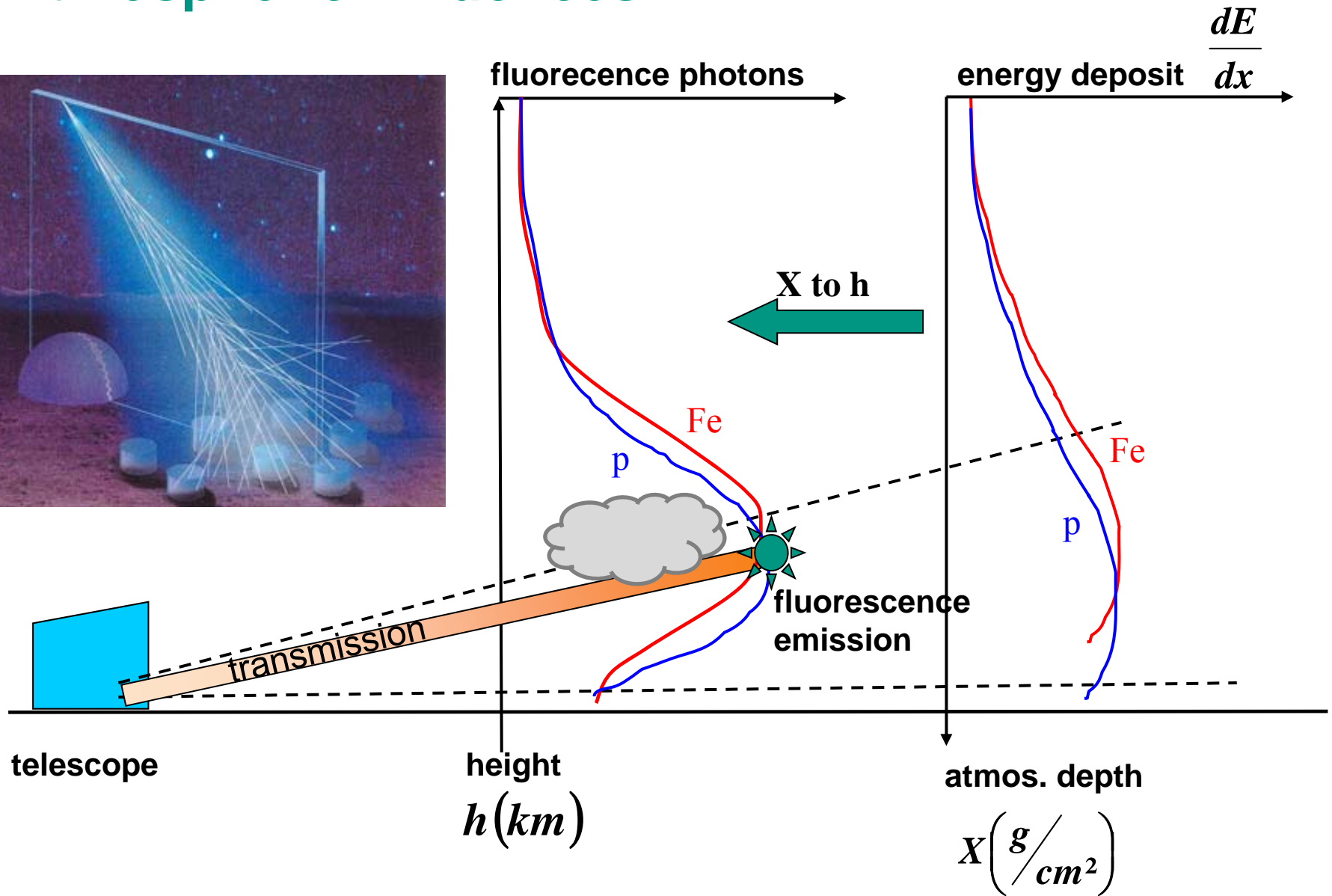
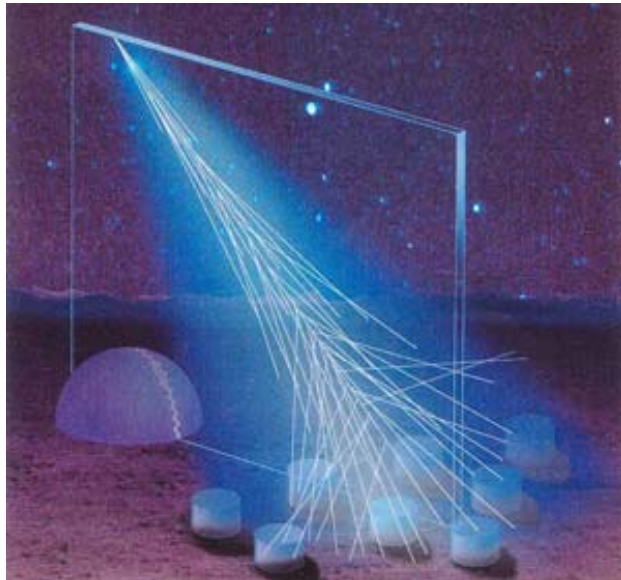
Energy Scale

shift and uncertainties

Changes in FD energies at 10^{18} eV	
Absolute fluorescence yield (sec. 2)	-8.2%
New optical efficiency	4.3%
Calibr. database update	3.5%
Sub total (FD calibration - sec. 4)	7.8%
Likelihood fit of the profile	2.2%
Folding with the point spread function	9.4%
Sub total (FD profile reconstruc. - sec. 5)	11.6%
New invisible energy (sec. 6)	4.4%
Total	15.6%

Systematic uncertainties on the energy scale	
Absolute fluorescence yield	3.4%
Fluor. spectrum and quenching param.	1.1%
Sub total (Fluorescence yield - sec. 2)	3.6%
Aerosol optical depth	3% ÷ 6%
Aerosol phase function	1%
Wavelength depend. of aerosol scatt.	0.5%
Atmospheric density profile	1%
Sub total (Atmosphere - sec. 3)	3.4% ÷ 6.2%
Absolute FD calibration	9%
Nightly relative calibration	2%
Optical efficiency	3.5%
Sub total (FD calibration - sec. 4)	9.9%
Folding with point spread function	5%
Multiple scattering model	1%
Simulation bias	2%
Constraints in the Gaisser-Hillas fit	3.5% ÷ 1%
Sub total (FD profile rec. - sec. 5)	6.5% ÷ 5.6%
Invisible energy (sec. 6)	3% ÷ 1.5%
Stat. error of the SD calib. fit (sec. 7)	0.7% ÷ 1.8%
Stability of the energy scale (sec. 7)	5%
Total	14%

Atmospheric Influences



Ground-based weather stations

- **measured:** mainly temperature T , pressure p , humidity u , wind
- every 5 min. a set of data
- transfer to database within a few days

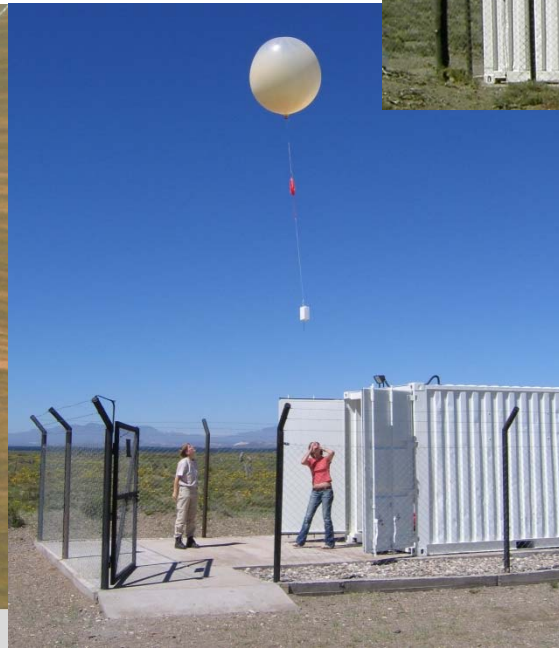
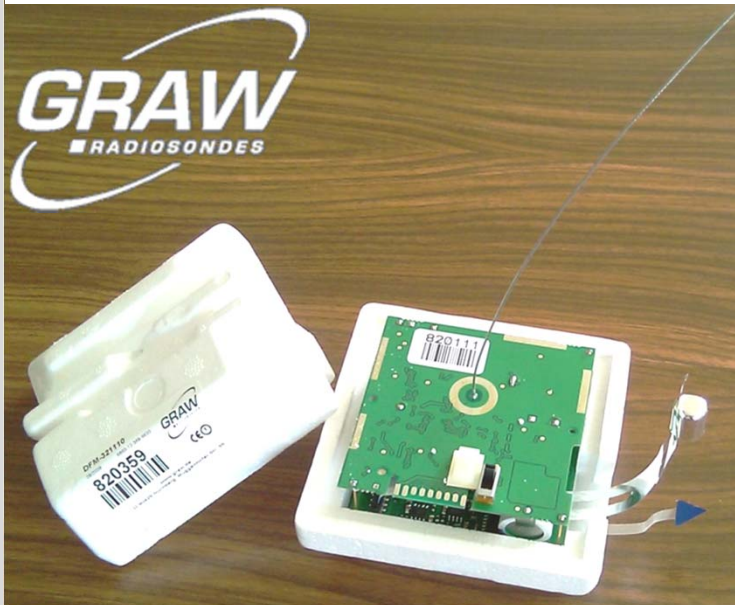


Atmospheric State Variables

- **measured:** mainly temperature T , pressure p , and humidity u
- **derived variables:** density ρ , atmospheric depth X , and vapour pressure e

profiles from radio soundings

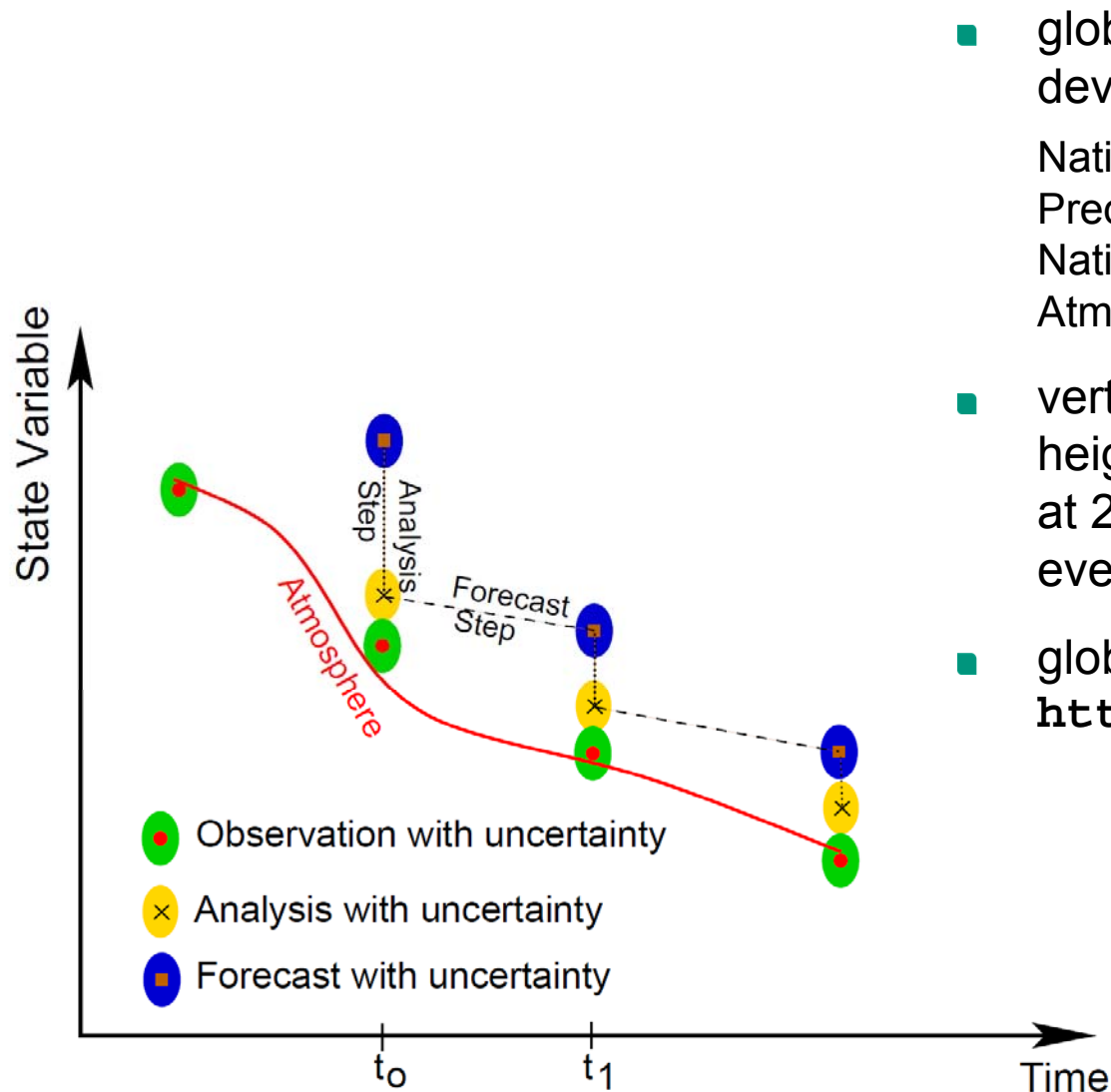
- intermittent launches
- 331 successful profiles until Dec. 2010
- 1st ascent in Aug. 2002
- now dedicated campaigns



Highschool and university project of Nebraska-Lincoln (UNL)
15 September 2012
maximal altitude about 95 000 feet (≈ 30 km a.s.l.)

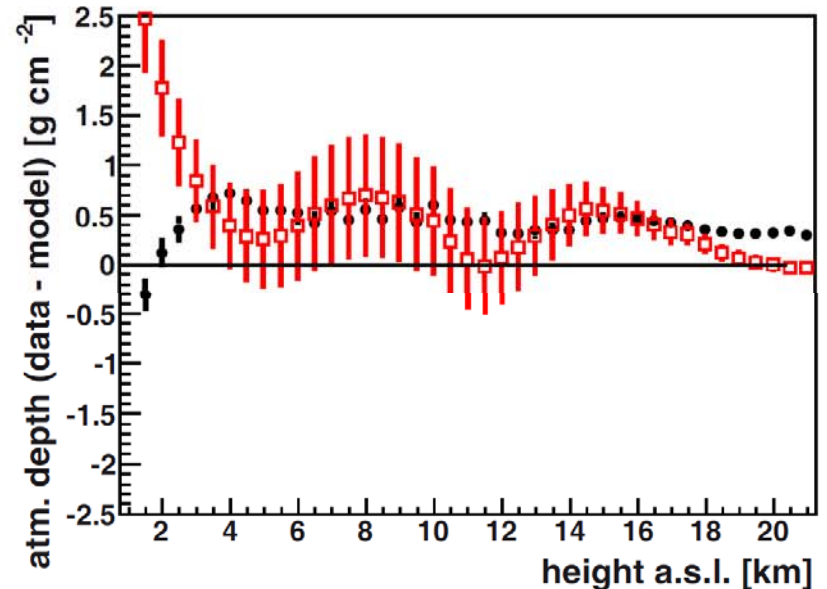
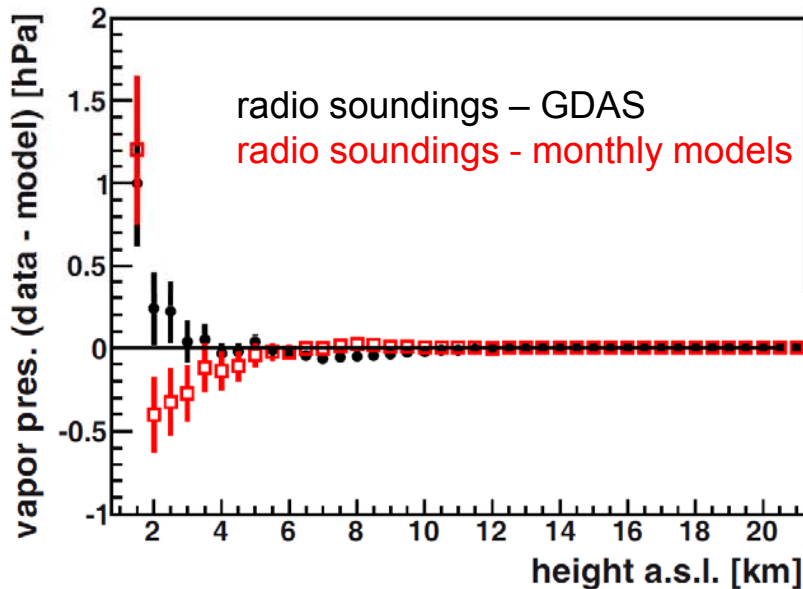
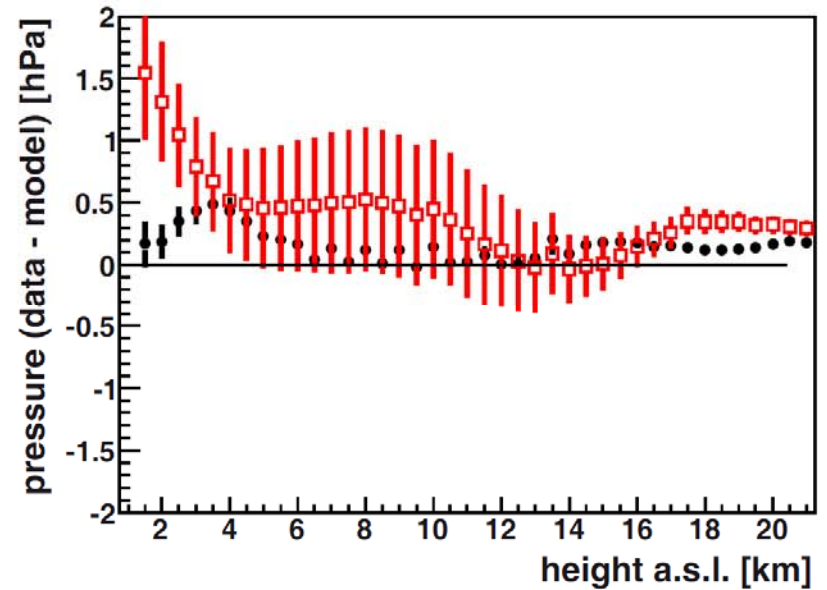
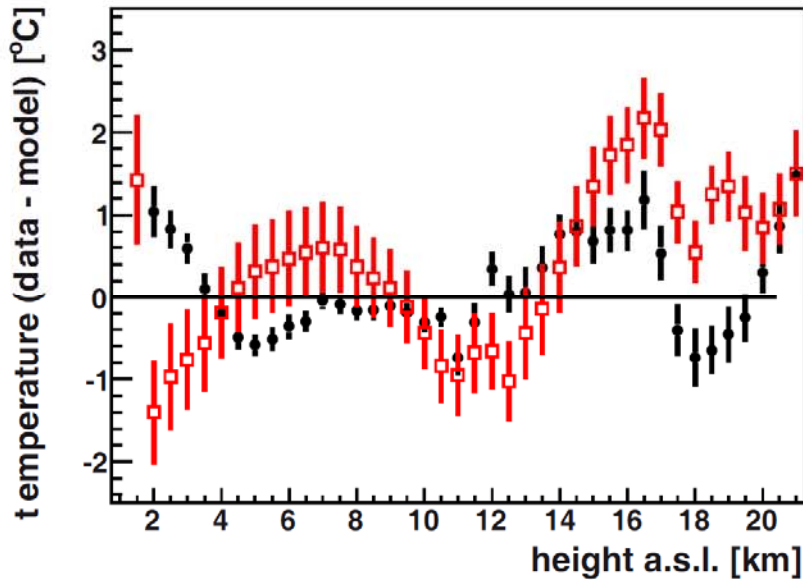


Global Data Assimilation System - GDAS

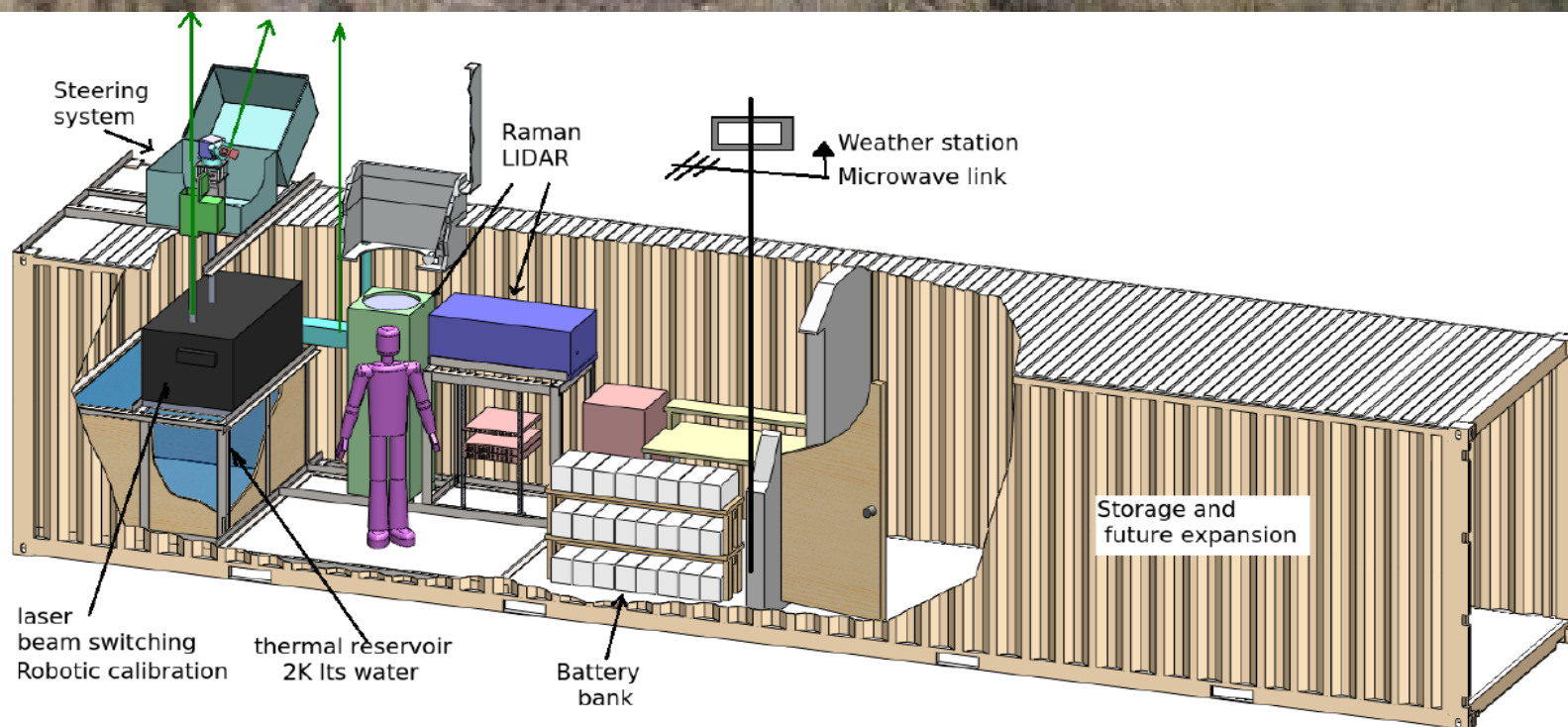


- global atmospheric model developed at NCEP
National Centers for Environmental Prediction (NCEP) at NOAA – National Oceanic and Atmospheric Administration
- vertical atmospheric profiles for height, temperature, humidity at 23 constant pressure levels every 3 hours since Dec. 2004
- global data publicly available at <http://ready.ar1.noaa.gov>

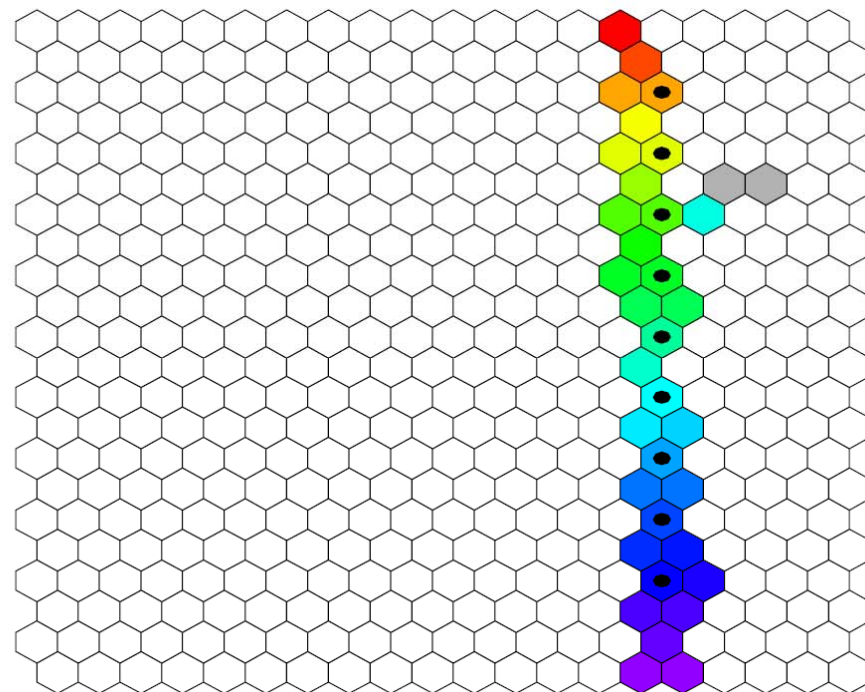
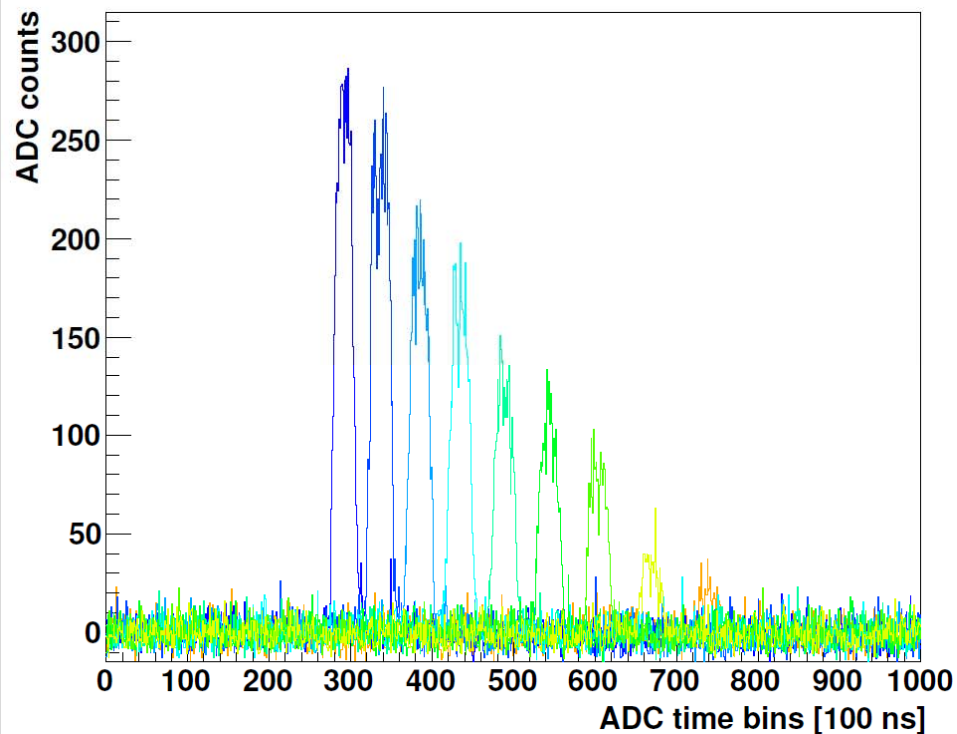
GDAS – comparison with local measurements



New Central Laser Facility with Raman lidar

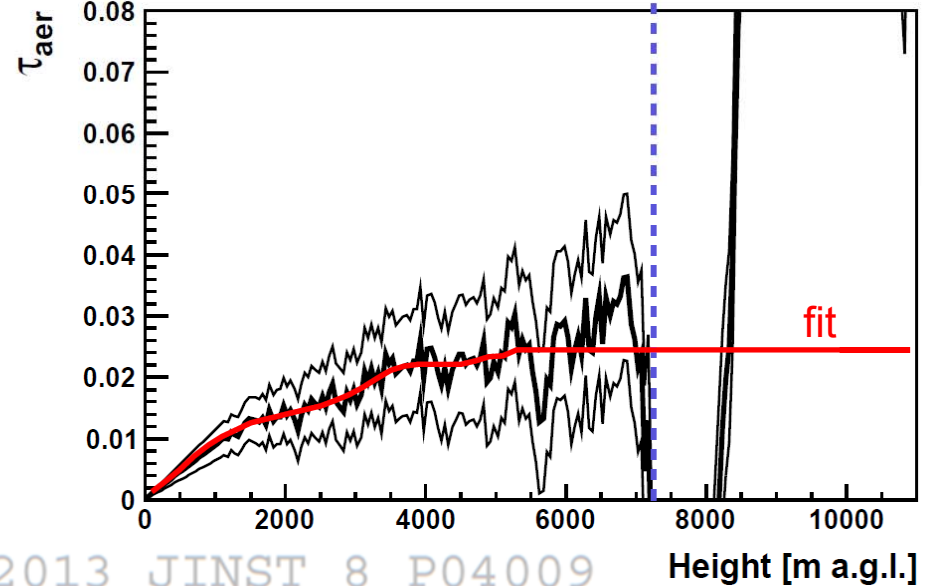
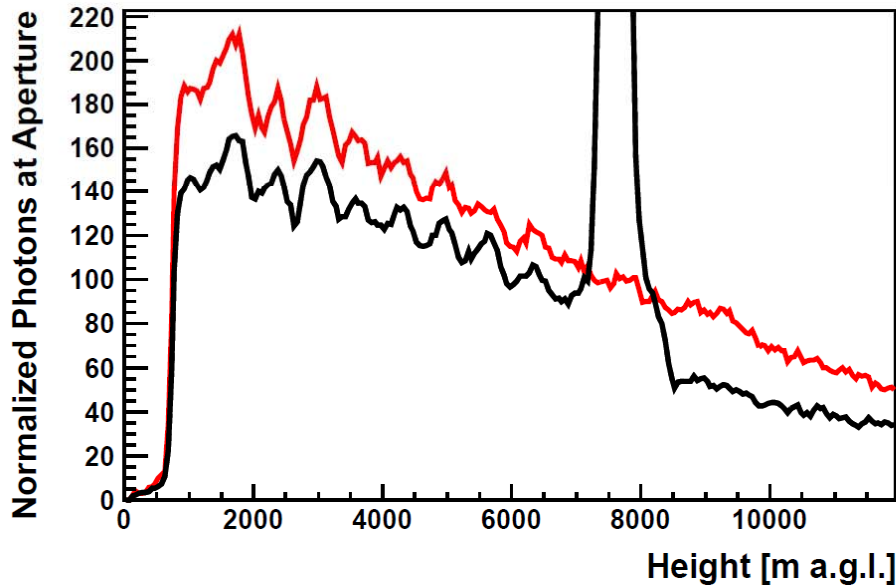
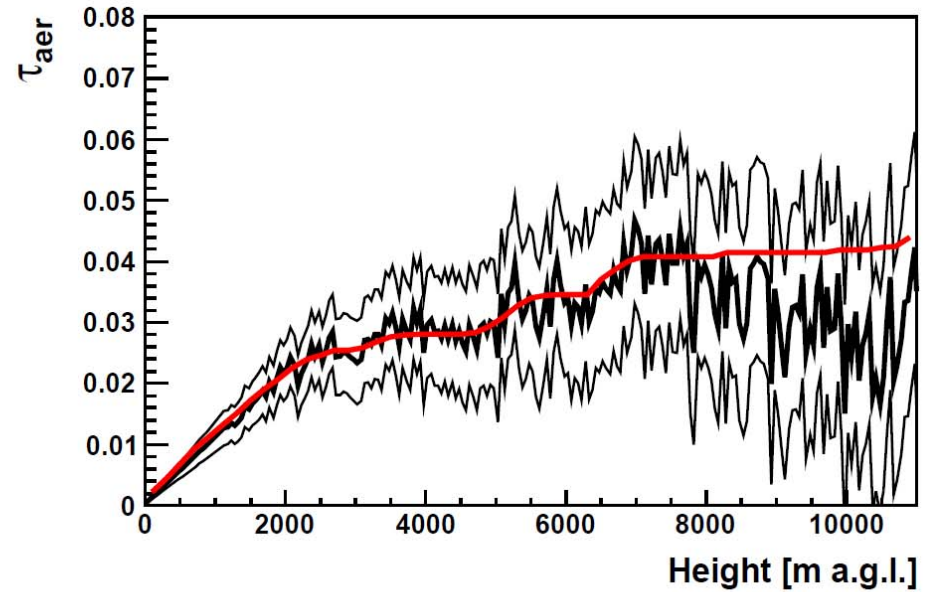
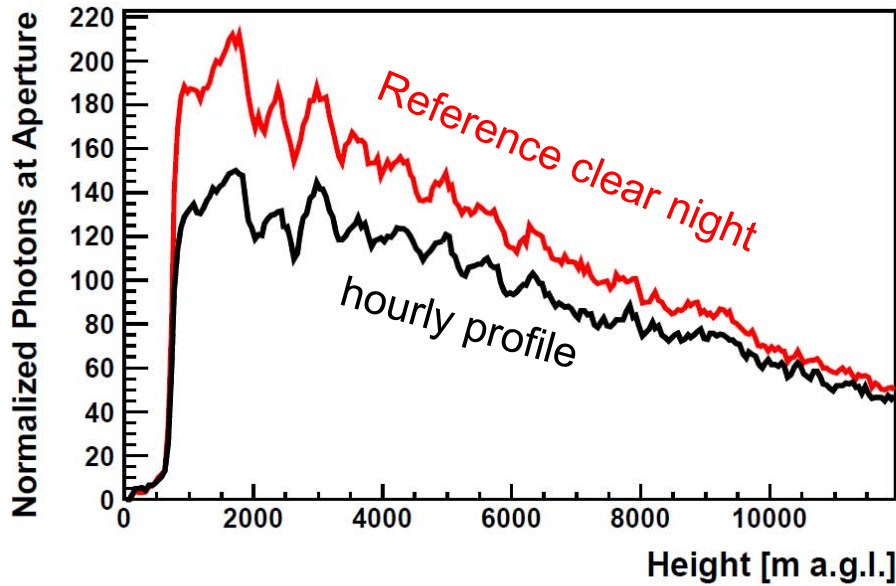


CLF shot seen at FD Los Leones



2013 JINST 8 P04009

From light profiles to Vertical Aerosol Optical Depth



Further devices for aerosol properties

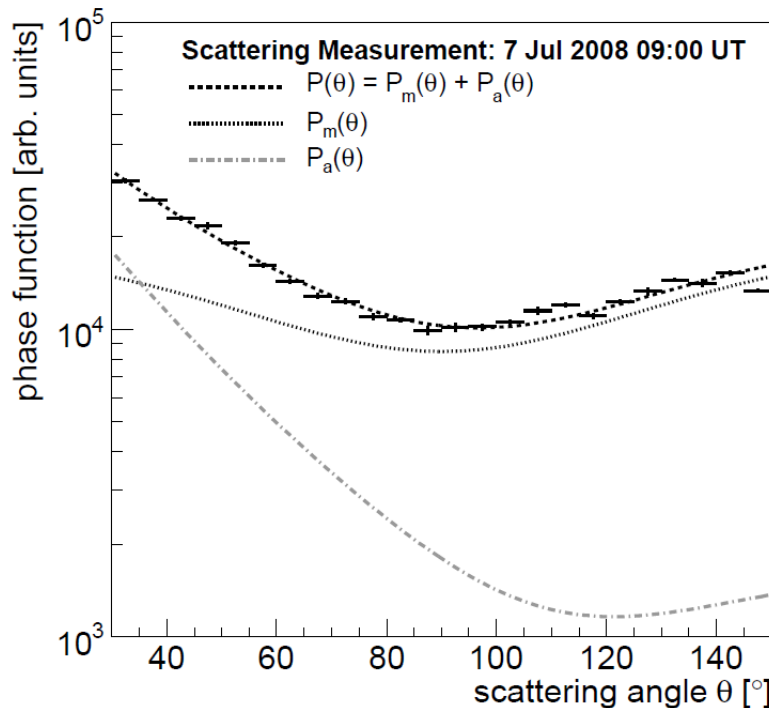
- HAM – Horizontal Attenuation Monitor

$$\tau_a(h, \lambda) = \tau(h, \lambda_0) \cdot \left(\frac{\lambda_0}{\lambda} \right)^\gamma$$

average at Auger side: $\gamma \approx 0.7 \pm 0.5$

- APF – Aerosol Phase Function monitor

$$P_a(\theta) = \frac{1 - g^2}{4\pi} \cdot \left(\frac{1}{(1 + g^2 - 2g \cos \theta)^{3/2}} + f \frac{3 \cos^2 \theta - 1}{2(1 + g^2)^{3/2}} \right)$$



average at Auger side: $g \approx 0.56 \pm 0.10$

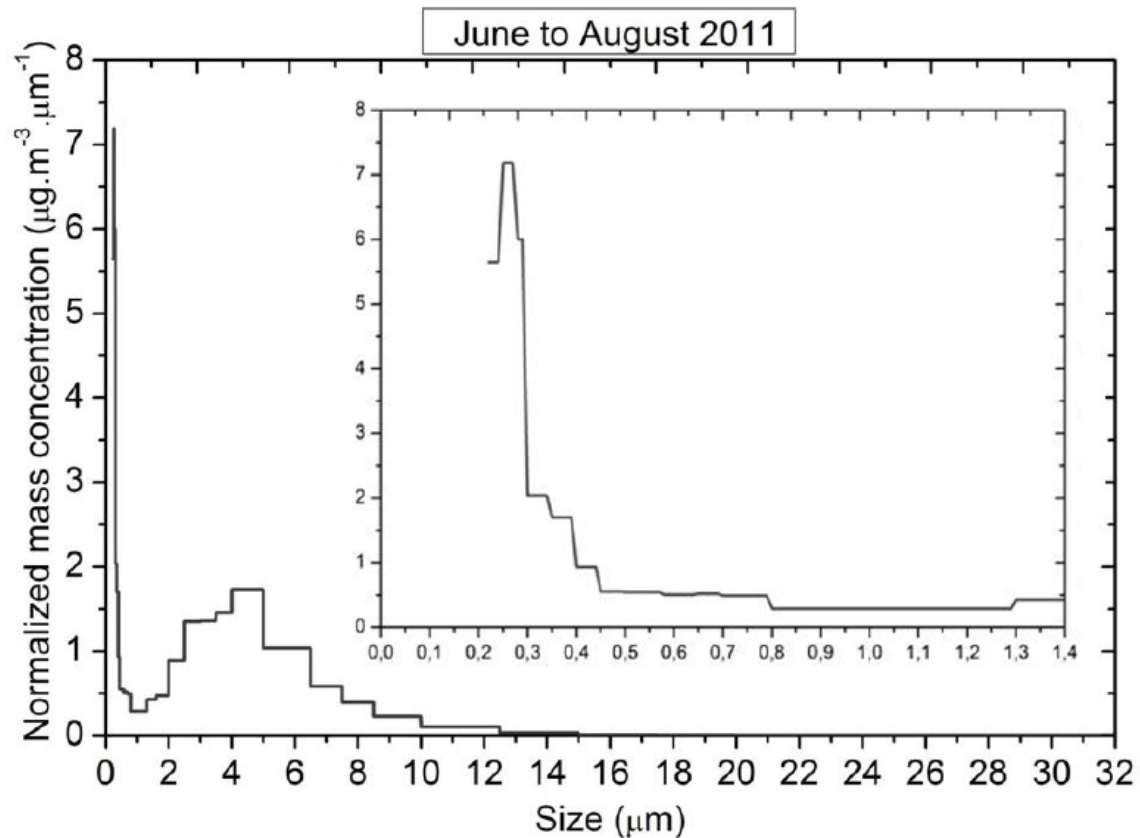
Further devices for aerosol properties

- Aerosol concentration, size, and elemental composition



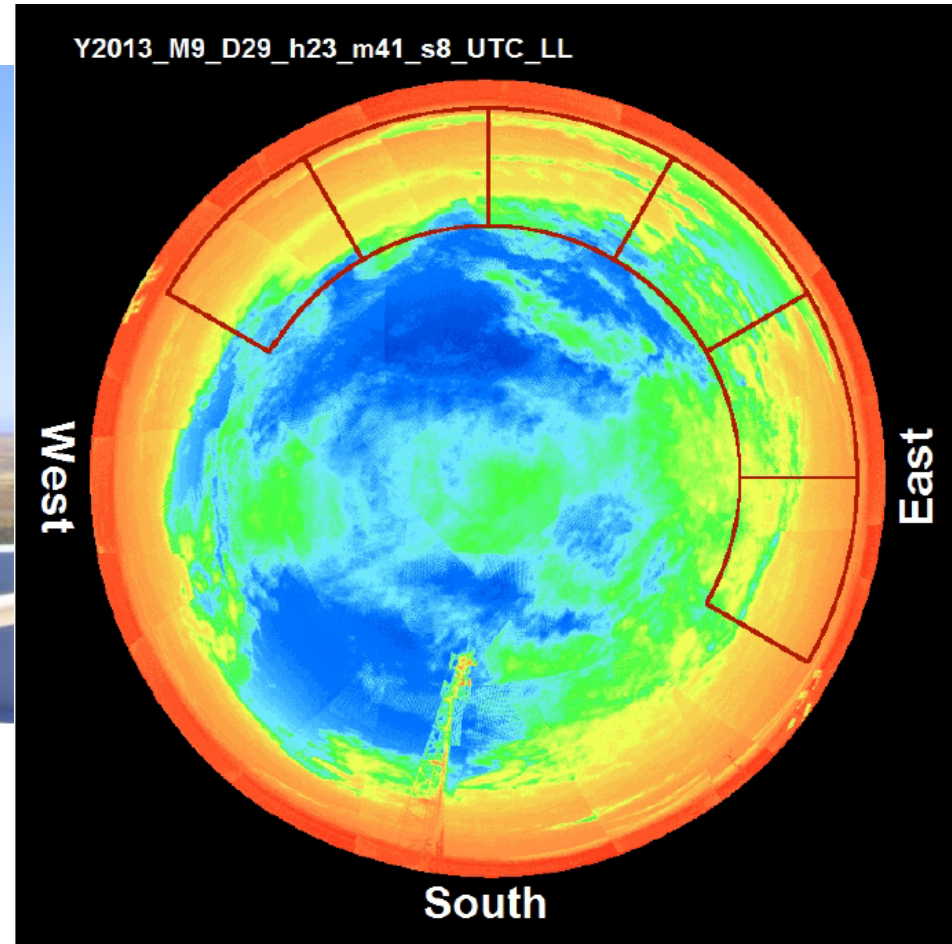
Andersen-Graseby 240 instrument @ Coihueco

Laser Aerosol spectrometer and Dust Monitor Grimm 1.109

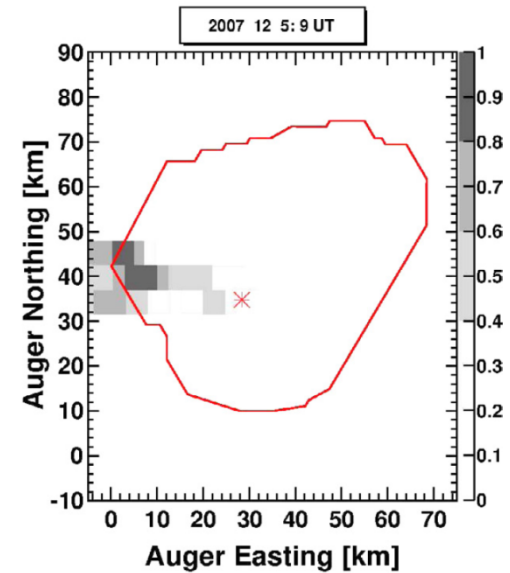
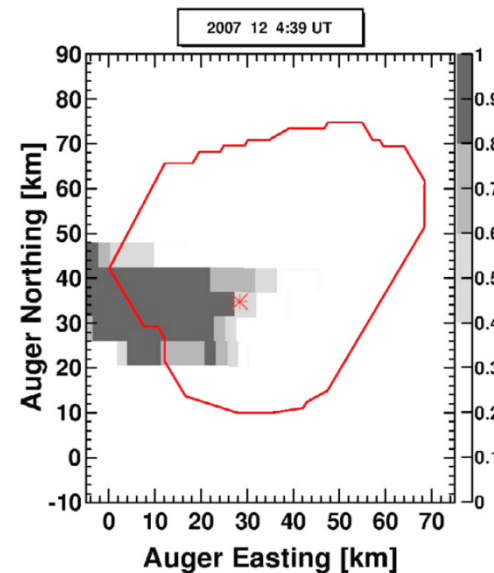
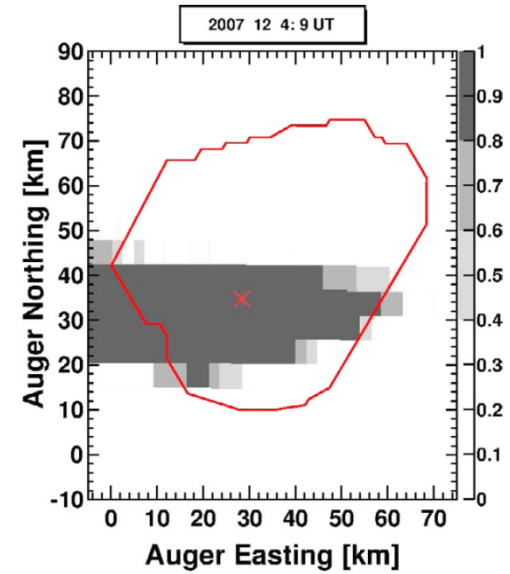
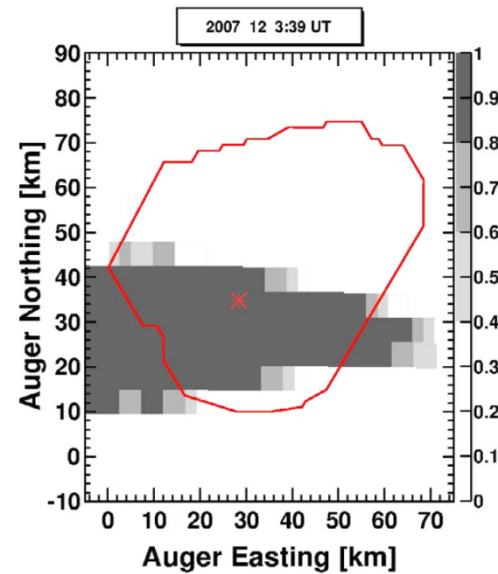
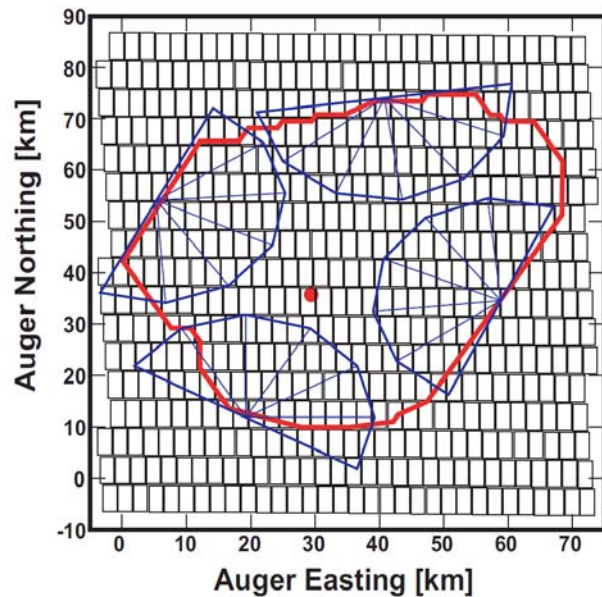


Infrared Cloud Cameras

Gobi-384 radiometric microbolometer array infrared(IR) camera

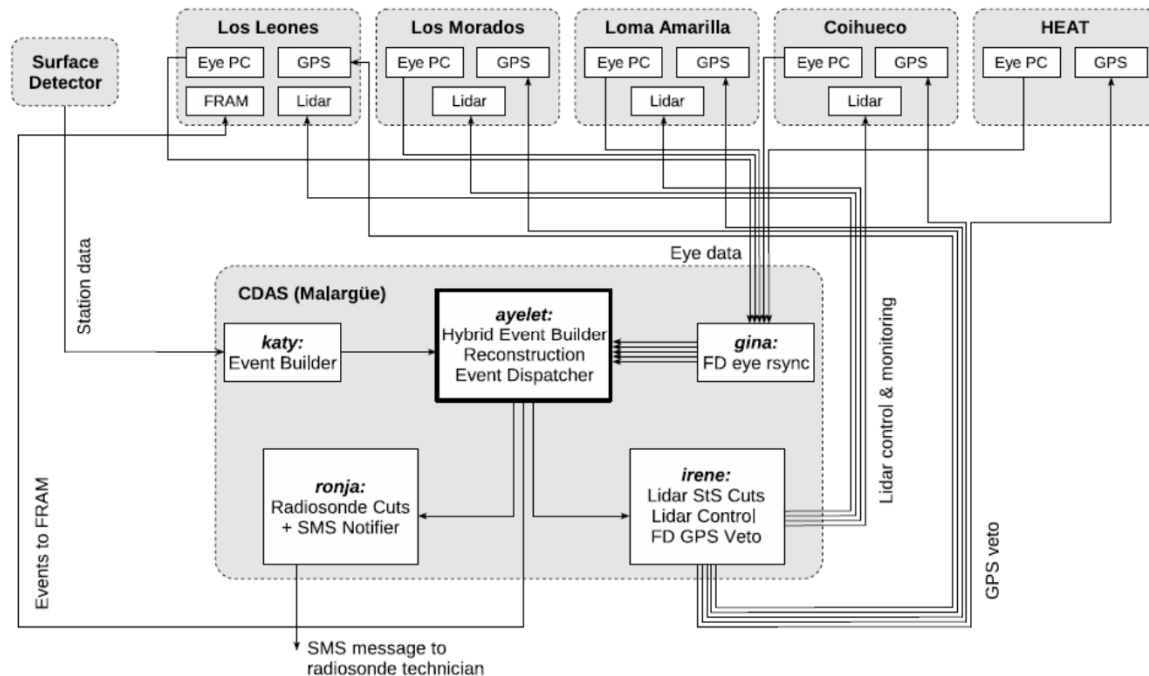


Cloud Identification with GOES-12

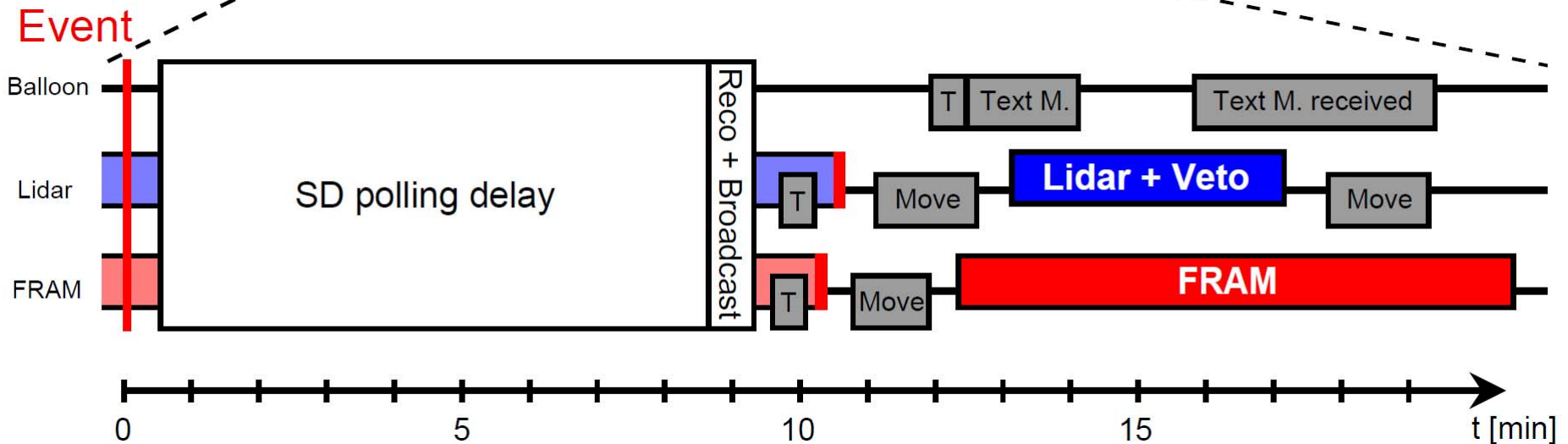
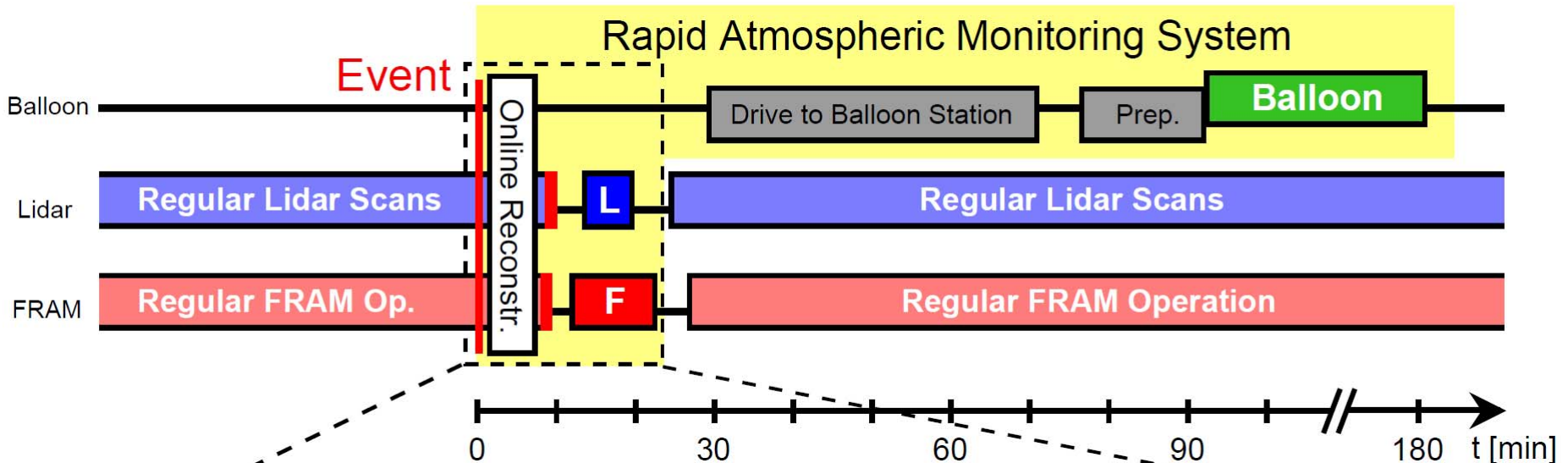


Rapid Atmospheric Monitoring

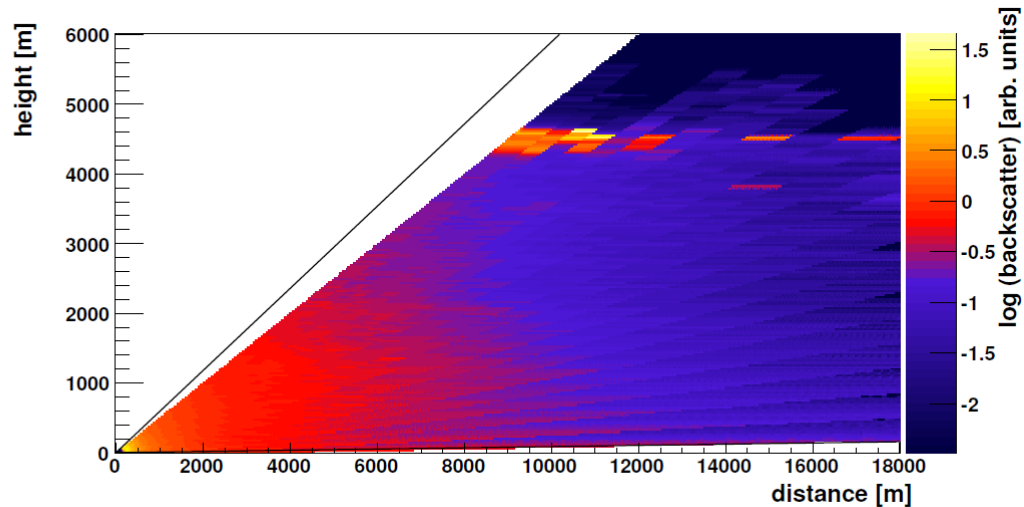
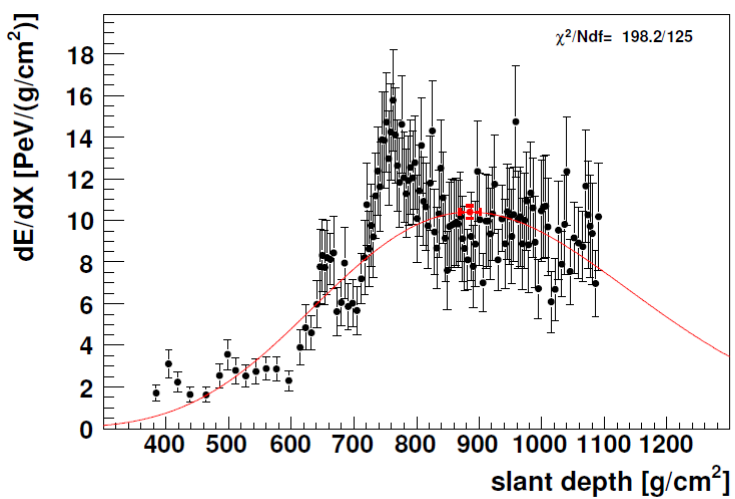
- Atmospheric observations shortly after detection of EAS of interest
- Fast online-reconstruction
- Participating instruments: lidar, radio soundings, FRAM
- programme started in spring 2009



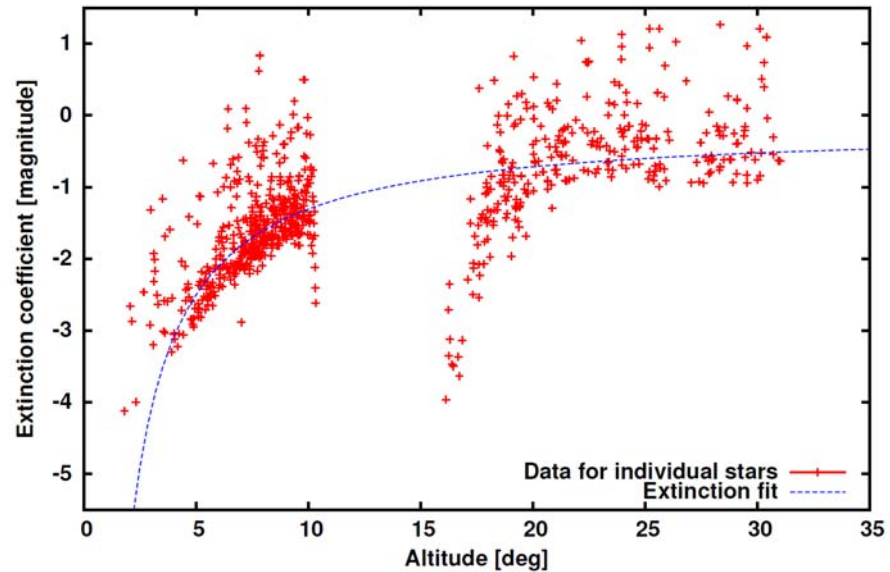
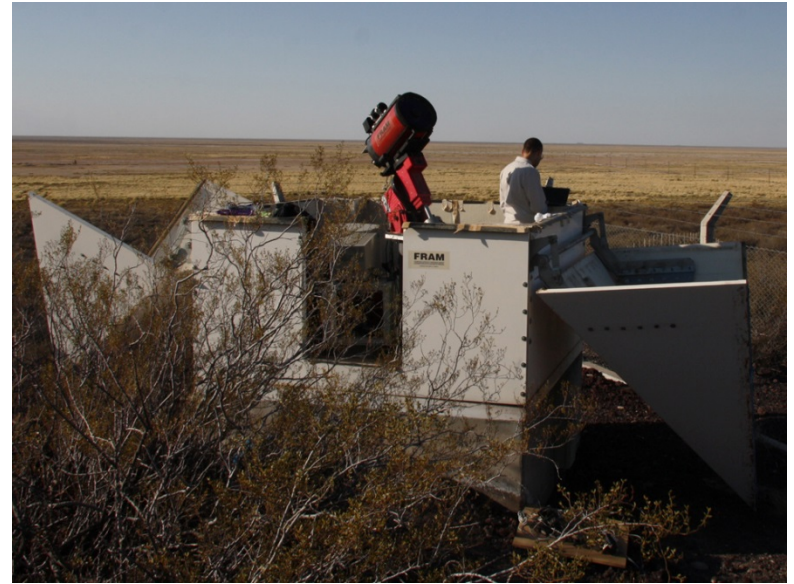
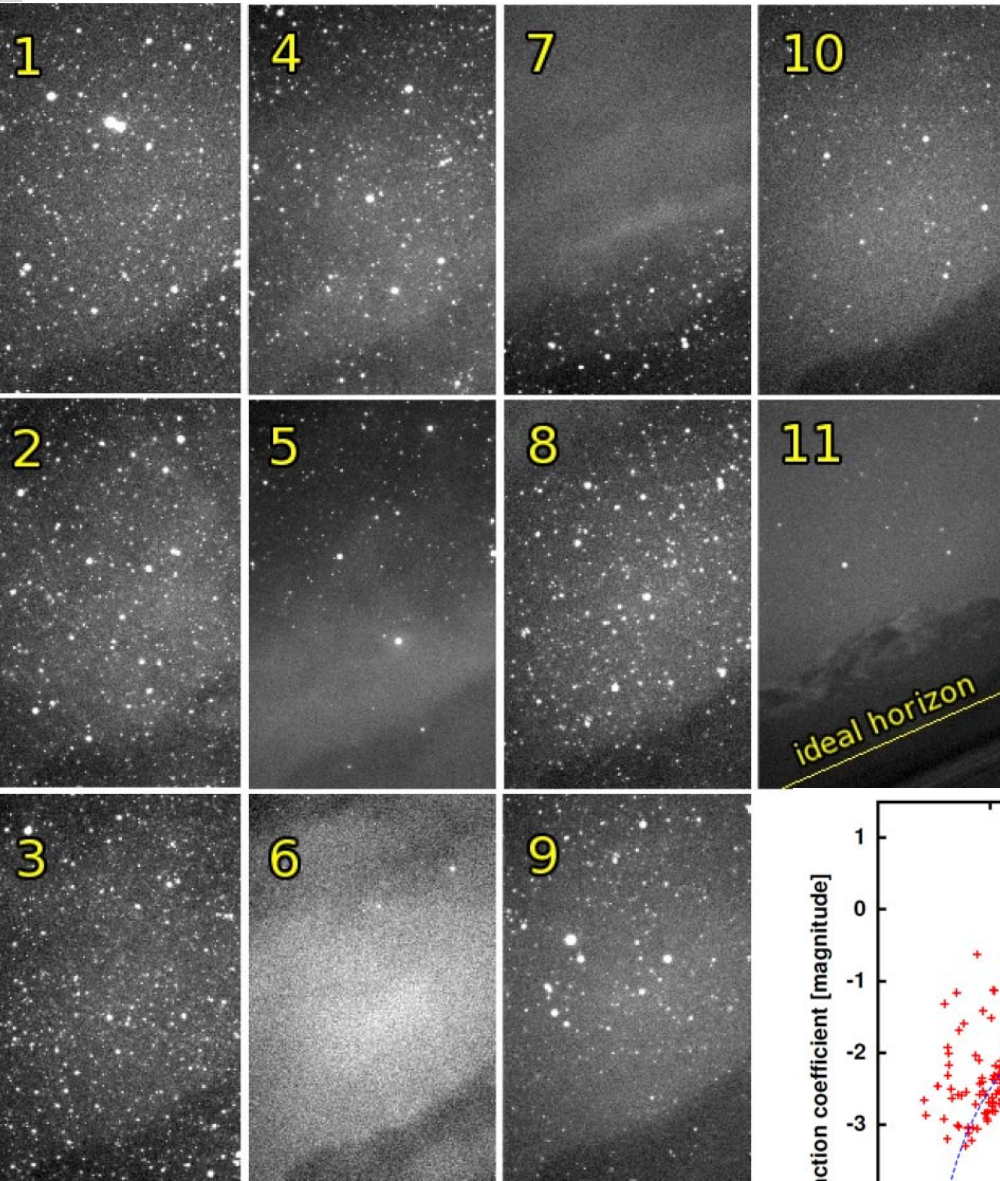
Rapid Atmospheric Monitoring

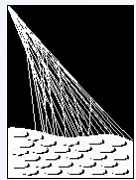


Lidars



FRAM – (F/Ph)otometric Robotic Atmospheric Monitor



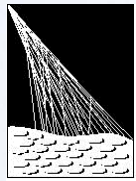


PIERRE
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Instrument Summary



Category	Variable	Frequency	Instrument(s)
State	At ground: Pressure, Temp., Wind, Humidity	5 min	Weather Stations
	Profile: Pressure, Temp., Humidity	3 hours	GDAS ^a
Aerosols	Vert. Optical Depth(z)	hourly	CLF, XLF + FD
	Phase Function	hourly	2 APF units
	Ångström Coefficient	hourly	FRAM (HAM)
Clouds	Presence in FD pixels	15 min	4 Cloud Cameras
	Behind FD sites	15 min	4 lidar stations
	Along select tracks	avg. 1/night	FRAM, lidar
	Above CLF/XLF	hourly	CLF, XLF + FD



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Summary

- Actual **atmospheric profiles** must be known for a reliable EAS reconstruction
- **Many devices** are installed at the Pierre Auger Observatory
- Many data of **atmospheric conditions** are available for area Malargüe
- **EAS reconstruction** reaches a higher level of precision

