

Laboratori Nazionali del Gran Sasso

Princeton 10/15/2013
Stefano Ragazzi LNGS Director



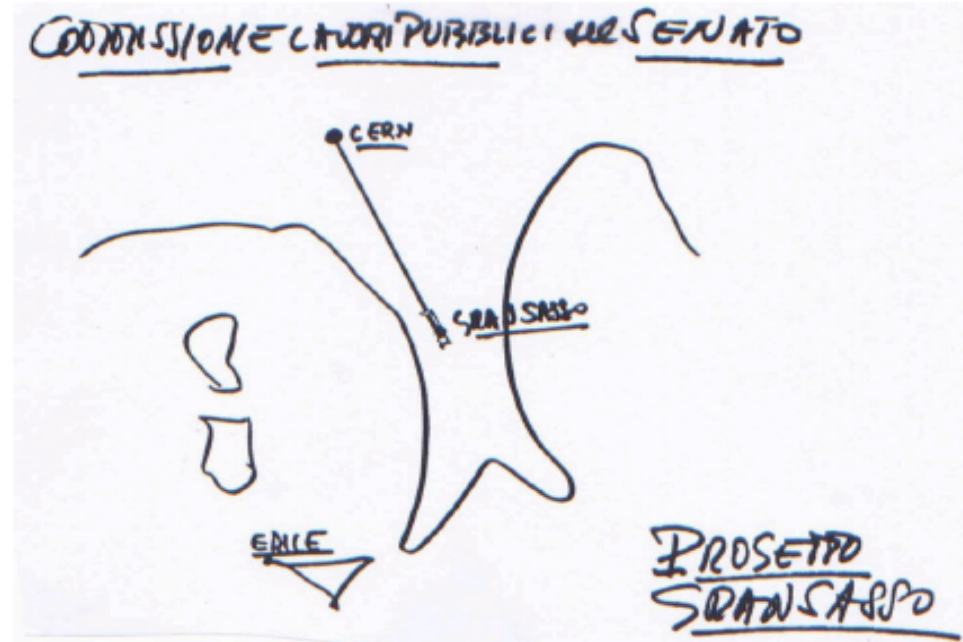
many thanks to

- Italian Consulate in Newark
- Princeton University and Physics Department
- All organizers
- All participants

**for organizing supporting and making this
meeting possible**

LNGS – early history

- 1979: proposal by A. Zichichi to Italian Parliament
- 1982: Approval of LNGS construction
- 1987: construction completed
- 1989: Start data taking of first large experiment (MACRO)



Note manoscritte di A. Zichichi presentate nella Seduta della Commissione Lavori Pubblici del Senato convocata con urgenza dal Presidente del Senato per discutere la proposta del Progetto Gran Sasso (1979).

To summarize, the scientific aims of the "Gran Sasso" laboratory are the study of:

- 1) nuclear stability;
- 2) neutrino astrophysics;
- 3) new cosmic phenomenology;
- 4) neutrino oscillations;
- 5) biologically active matter;
- 6) ground stability.

Not only
 $\tau_p \neq \infty$

LNGS

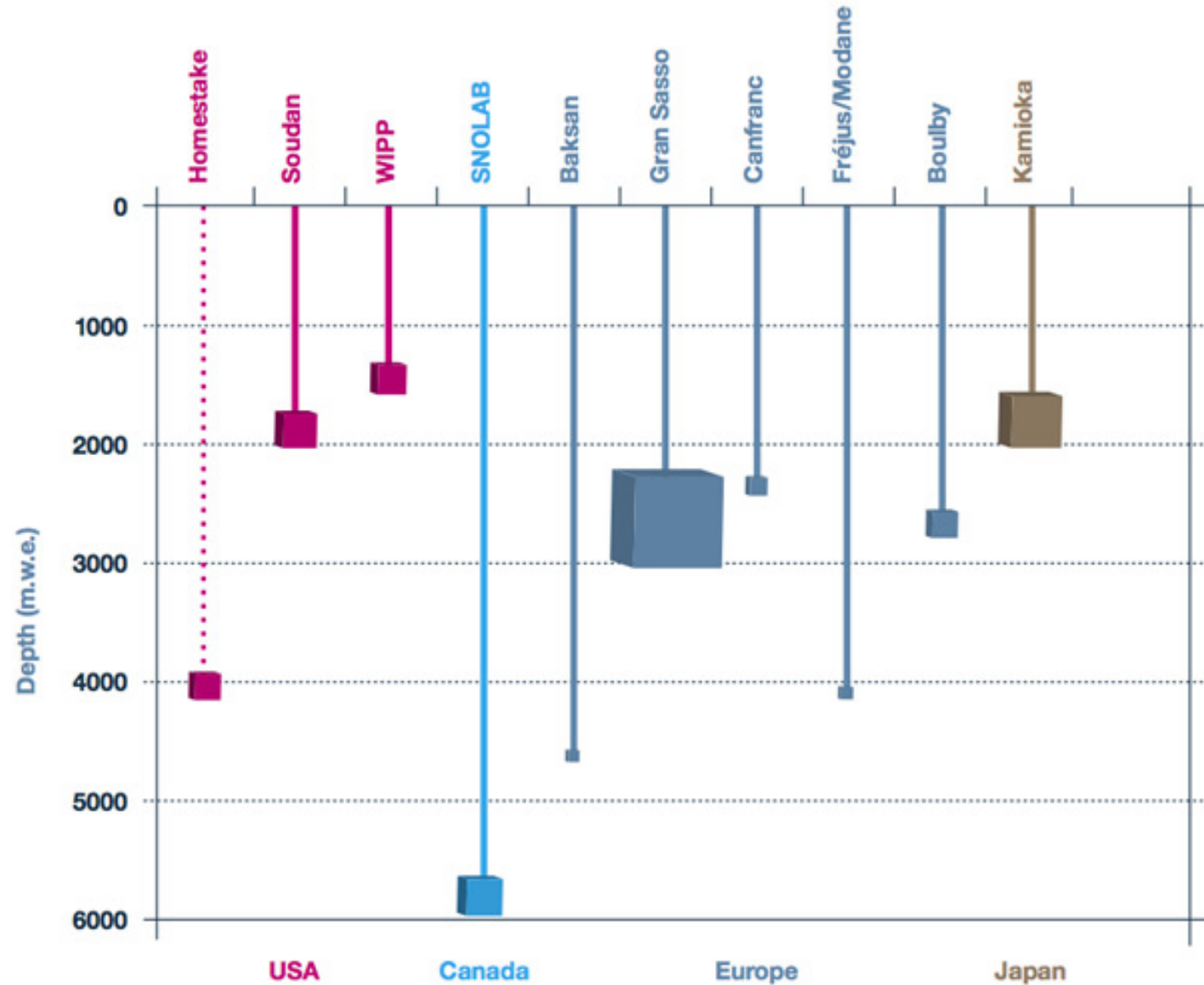


design and construction



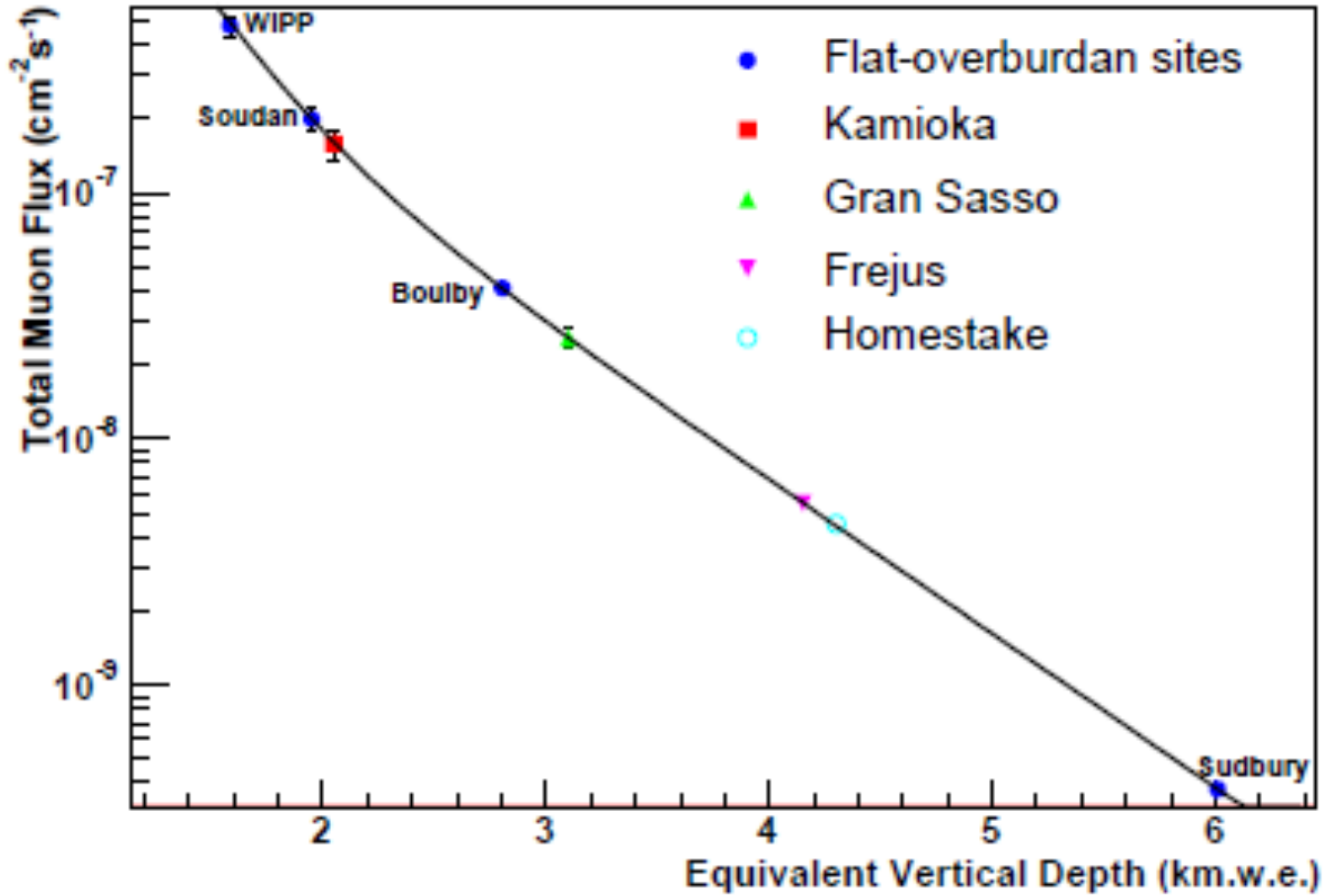
Underground Science Laboratories

- LNGS
 - largest
 - easiest to access



Plot adapted from <http://www.deepscience.org/contents/facilities.shtml>

Muon Flux versus depth

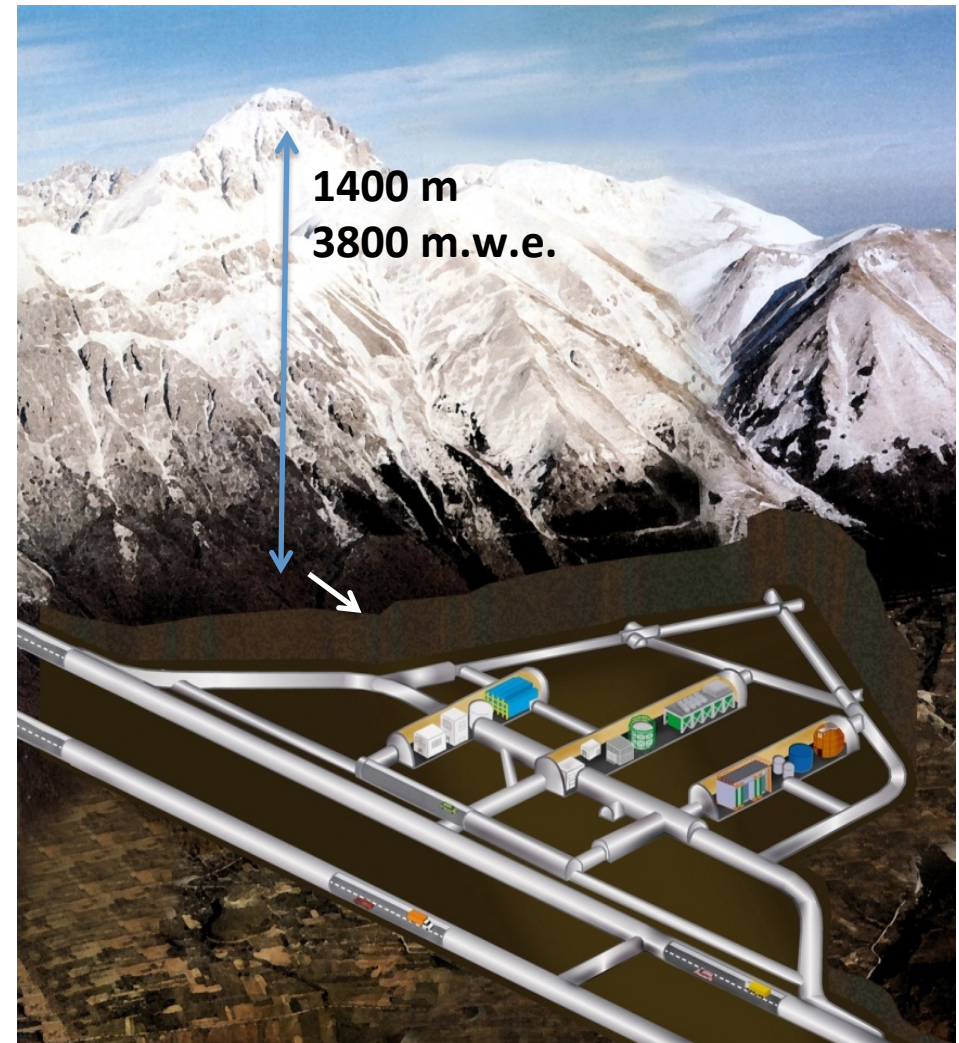


The LNGS Laboratory

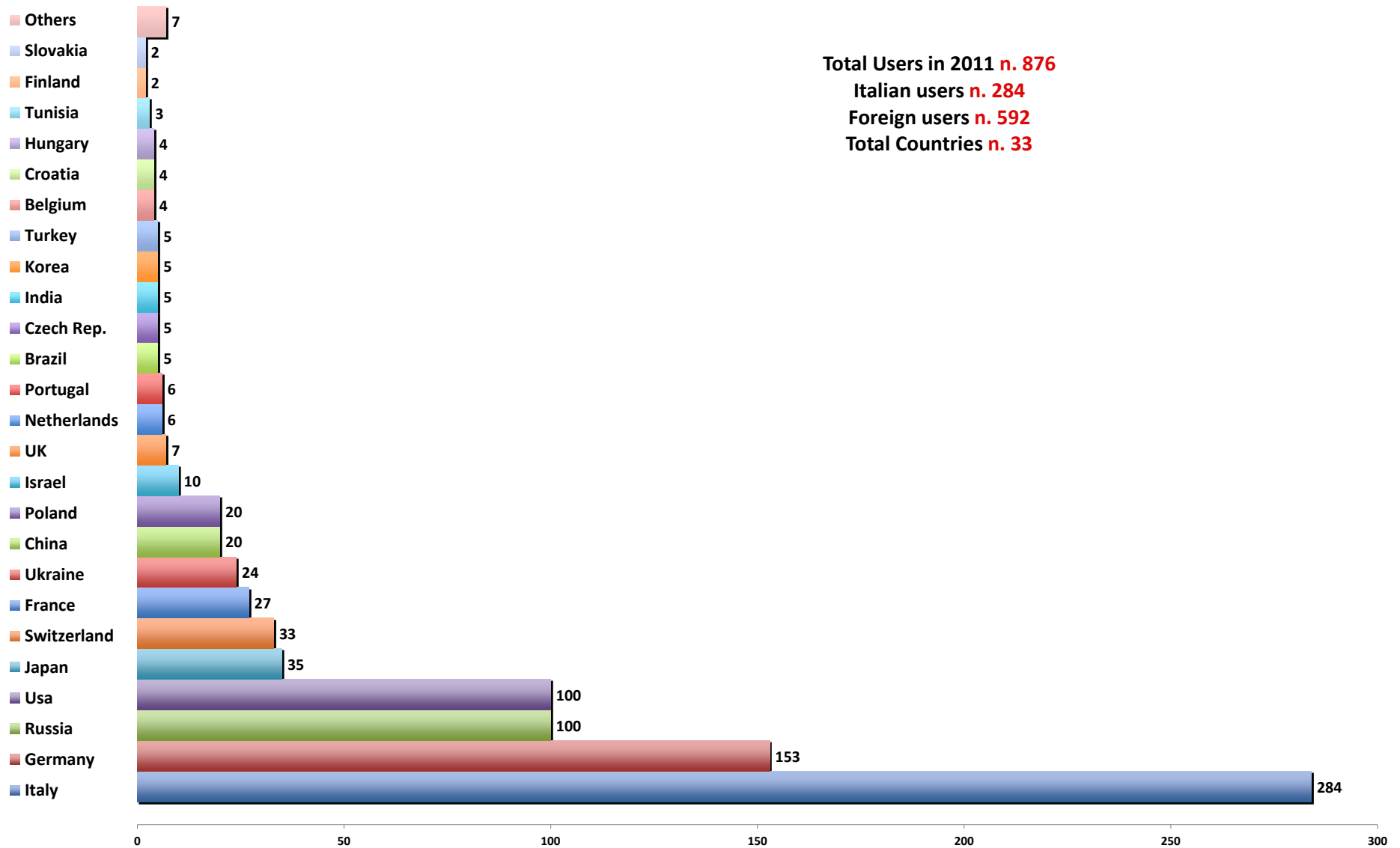
SHIELDING is the main facility

- Muon flux: $3.0 \cdot 10^{-4} \text{ m}^{-2}\text{s}^{-1}$
- Neutron flux:
 - $2.92 \cdot 10^{-6} \text{ cm}^{-2}\text{s}^{-1}$ (0-1 keV)
 - $0.86 \cdot 10^{-6} \text{ cm}^{-2}\text{s}^{-1}$ (> 1 keV)
- Rn in air: 20-80 Bq m⁻³
- Surface: 17 800 m²
- Volume: 180 000 m³
- Ventilation: 1 vol / 3 hours

Effective allocation of underground areas is essential



An International Laboratory



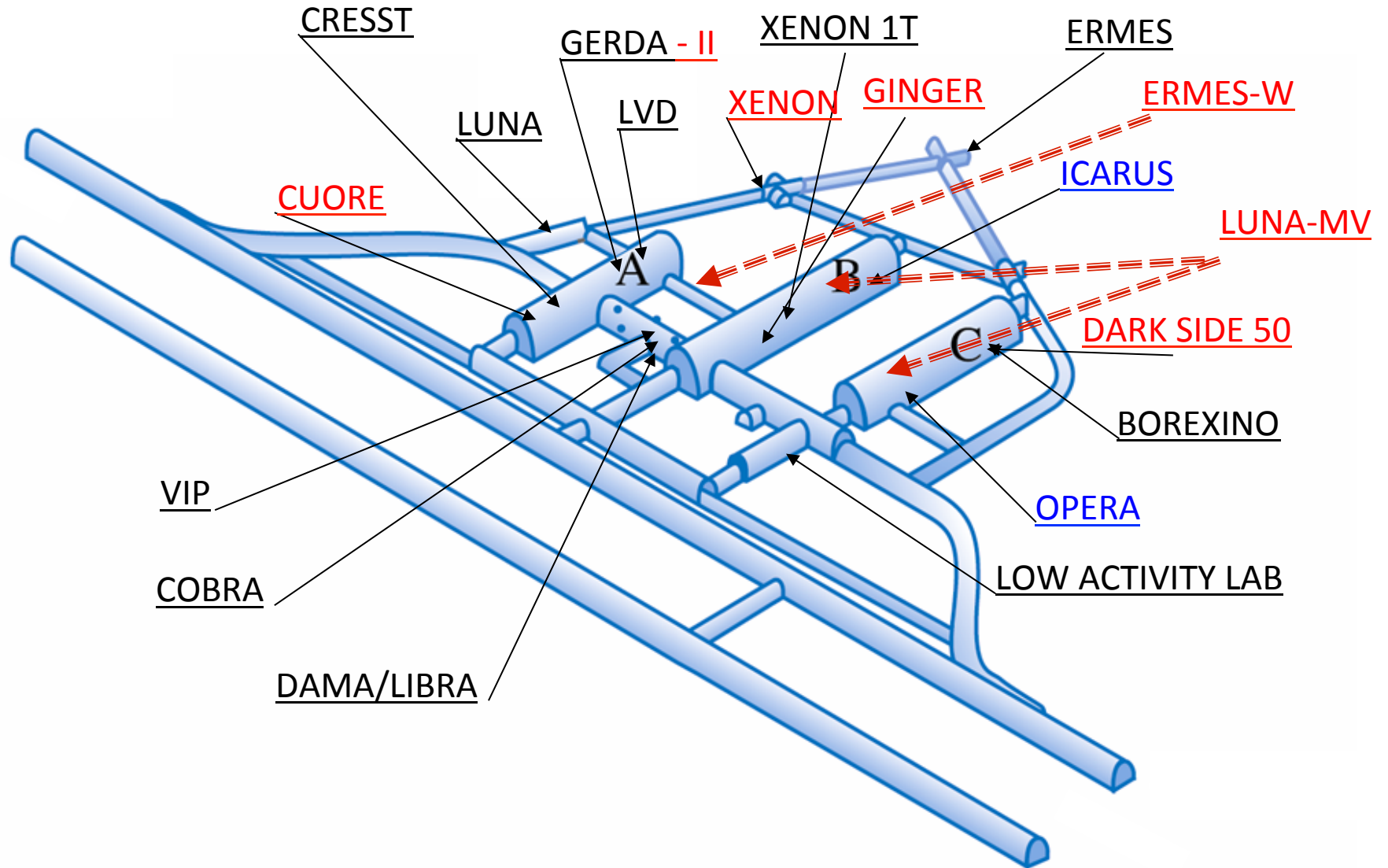
Personnel

- Research
 - 10 permanent staff (- 3 on leave)
 - 7 limited time staff
 - 10 post-doc fellowships (8 INFN, 2 POR)
 - 8 graduate and undergraduate fellowships
- Technology and Engineering
 - 15 permanent staff (- 1 on leave)
 - 6 limited-time staff
 - 2 post-doc fellowships
 - 10 graduate and undergraduate fellowships
- Technicians
 - 24 permanent staff
 - 3 limited time staff
 - 2 fellowships
- Administration
 - 19 permanent staff
 - 3 limited time staff
 - 1 fellowship
- Other support staff
 - 7 permanent

LNGS – 25 years

- In 25 years LNGS evolved
 - from: UG space tailored around a few experiments
 - to: a multi purpose UG science facility

Busy laboratory



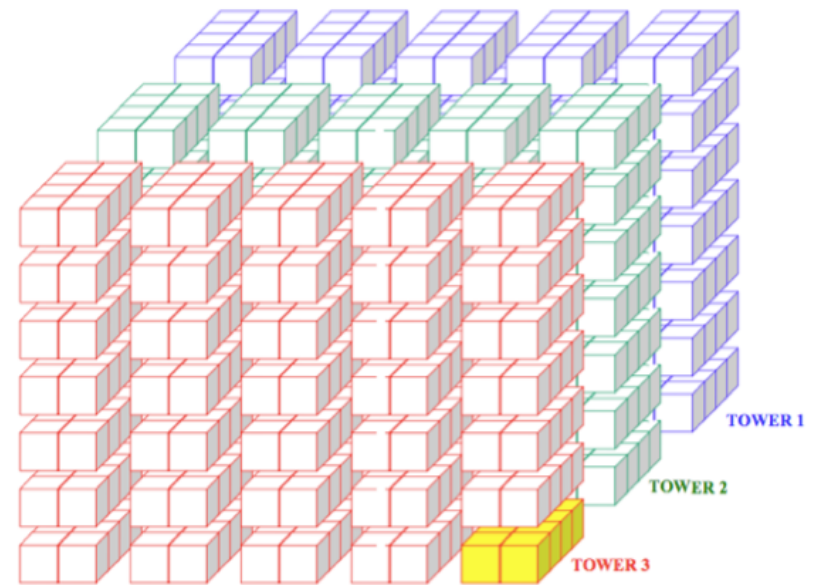
Neutrino physics

- neutrino oscillations
 - Opera, Icarus: the recent past
 - Borexino-SOX: the future
- solar physics
 - Borexino
- cosmogenic and geo neutrinos
 - LVD, Borexino
- double beta decay
 - GERDA: phase-I completed, phase-II in preparation (*Macolino*)
 - COBRA
 - CUORICINO, CUORE (2015)
 - LUCIFER R&D

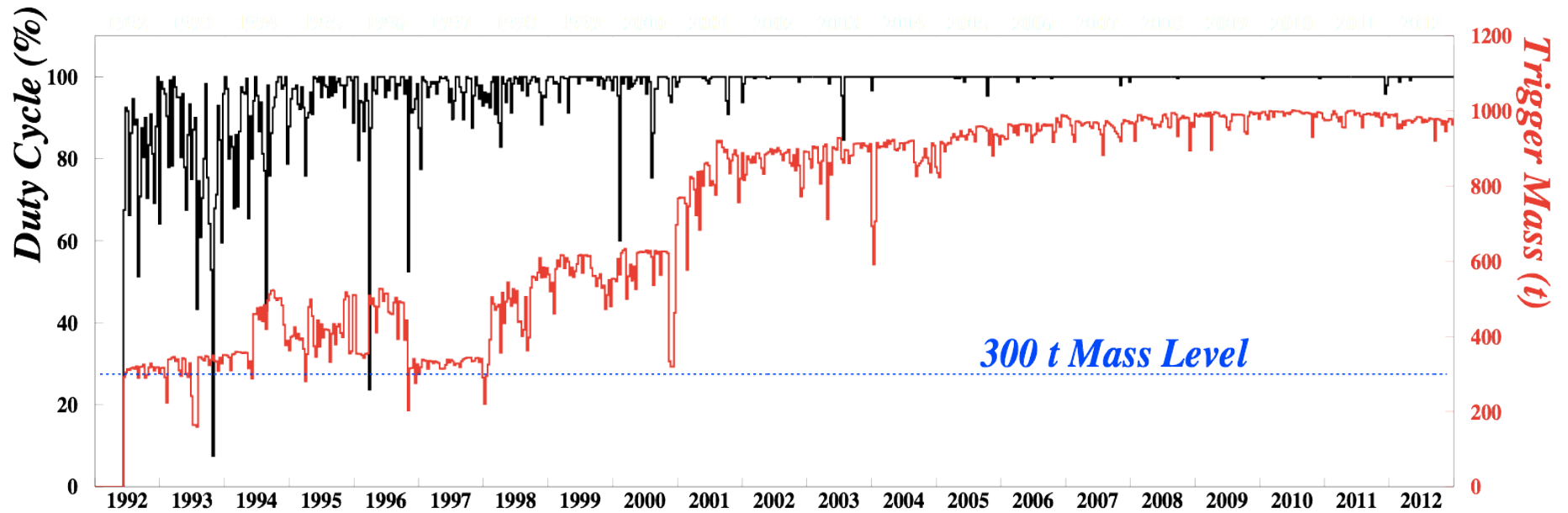
LVD



- A survivor of the first-generation LNGS experiments: waiting for a Supernova

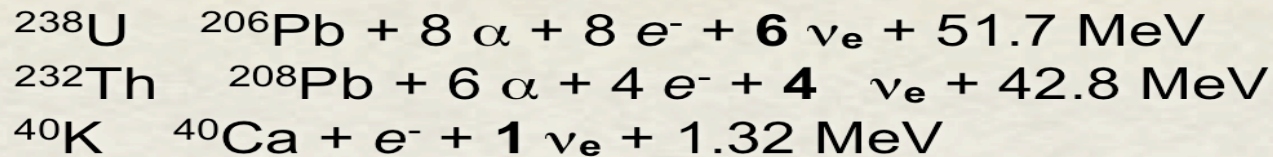
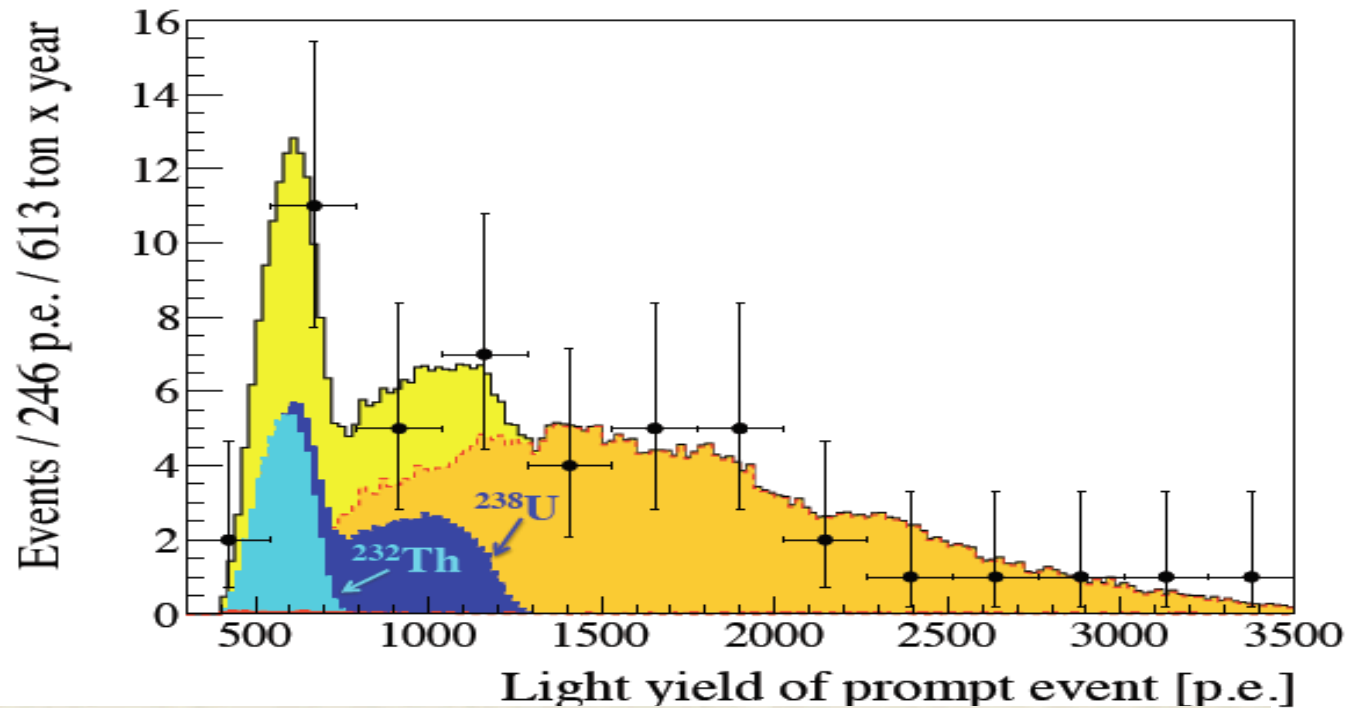


LVD Duty factor



- Large mass
 - 1 kT
- Need high availability
 - ~100%

Geo Neutrinos with Borexino



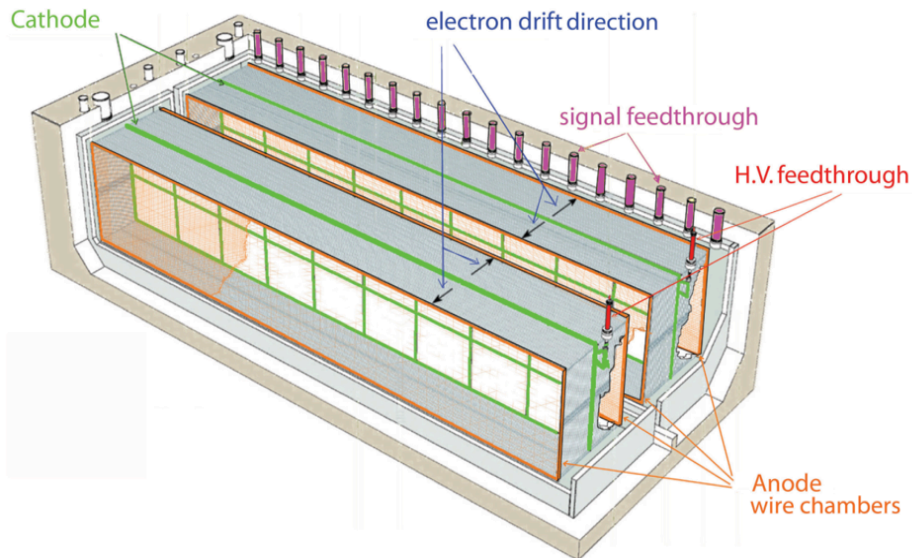
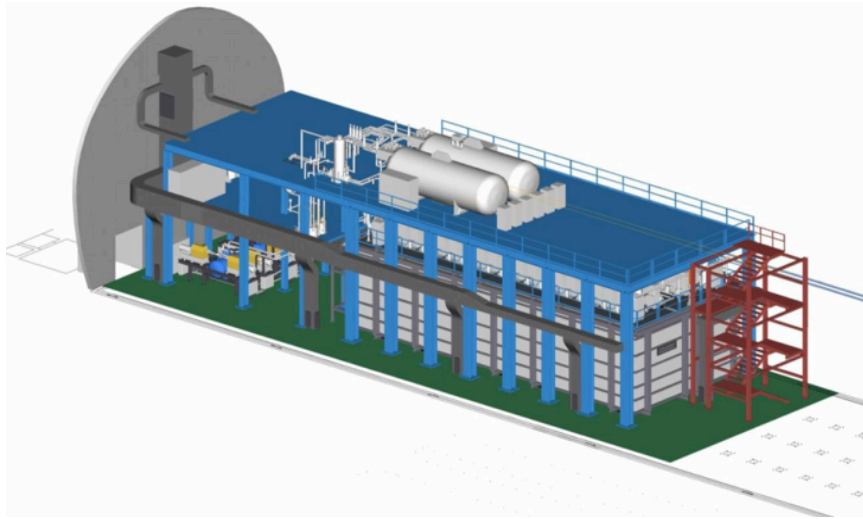
Purification level: U $\sim 1.6 \cdot 10^{-19}$ g/g

Th $< 1.2 \cdot 10^{-18}$ g/g

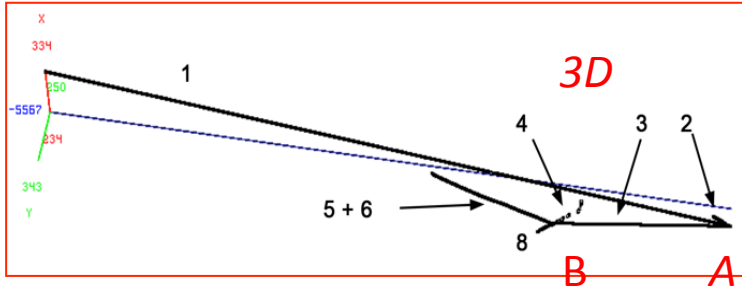
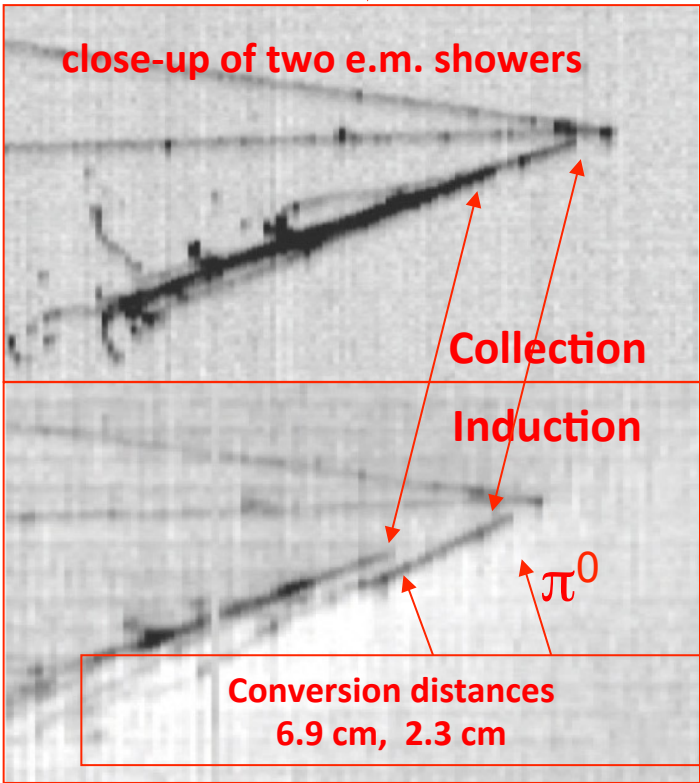
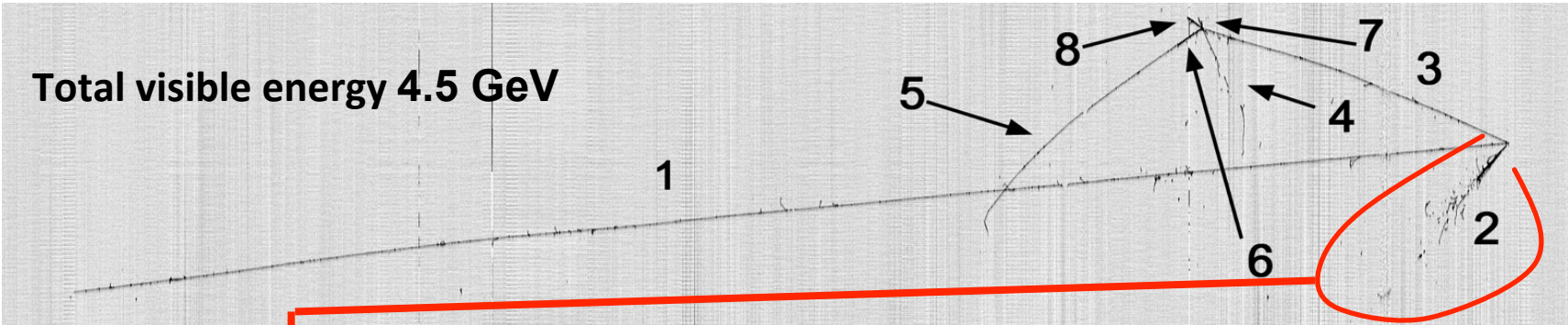
-> *applied* neutrino physics

ICARUS-T600 @ LNGS

0.77 kton LAr-TPC



The LAR-TPC: the electronic bubble chamber



Primary vertex (A):

very long μ (1), e.m.cascades(2), π (3)

Secondary vertex (B):

longest track (5) is a μ from stopping K (6)
 μ decay is observed

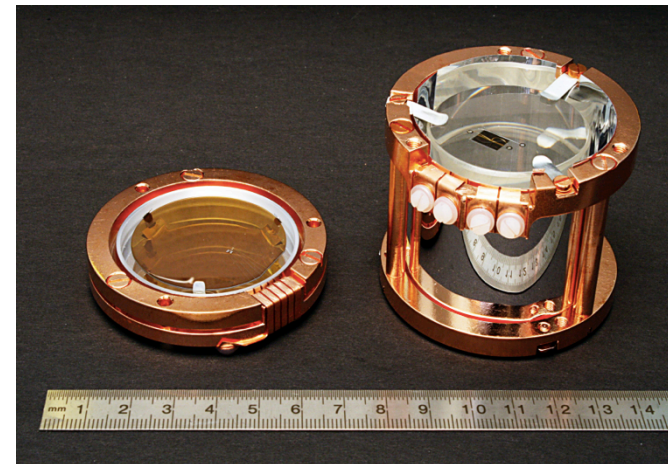
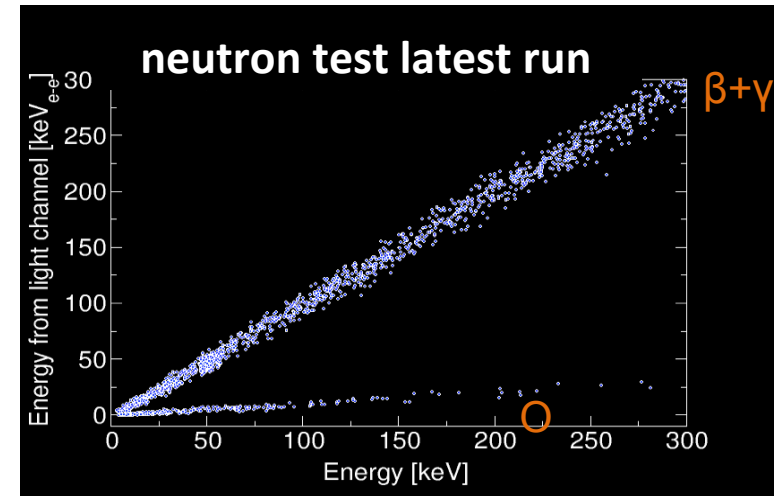
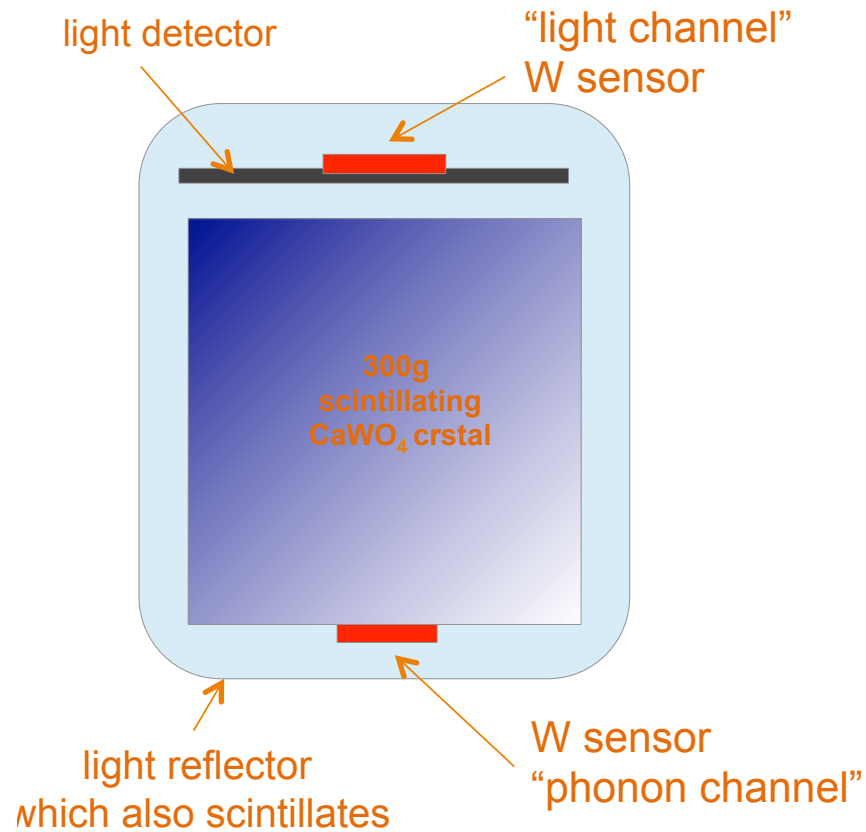
ICARUS T600 timeline

- **2010: Successful assembly and commissioning of ICARUS-T600:**
 - May 28th : first CNGS neutrino
 - October: start of physics runs
- **2011 – Dec 3rd, 2012:**
 - T600 data-taking with CNGS beam and cosmic rays
 - 2800 CC + 900 NC
 - Muons from upstream GS rock ≈ 12000 ev (≈ 8200 on TPC front face)
 - Intrinsic beam ν_e CC ≈ 26 ev
 - $\nu_\mu \rightarrow \nu_\tau$ with kinematic ν_τ -selection (~ 2 event $\tau \rightarrow e$) $\nu_\mu \rightarrow \nu_e$ (θ_{13}) from e-like CC events
 - excess at $E < 20$ GeV (~ 5 events CC)
 - Search for ν_s in LSND parameter space, studying e-like CC events at $E > 10$ GeV
 - Nov. 2011 & March 2012: bunched beam for ν velocity measurement
- **Dec.2012 – June 2013:**
 - Data-taking with cosmic rays
 - From June: decommissioning
- **2014 onward**
 - **ICARUS at CERN – R&D to prepare for a *future beam***

Dark Matter

- DAMA/LIBRA
- Cresst
- Xenon family
- DarkSide (*family*)

CRESST-II



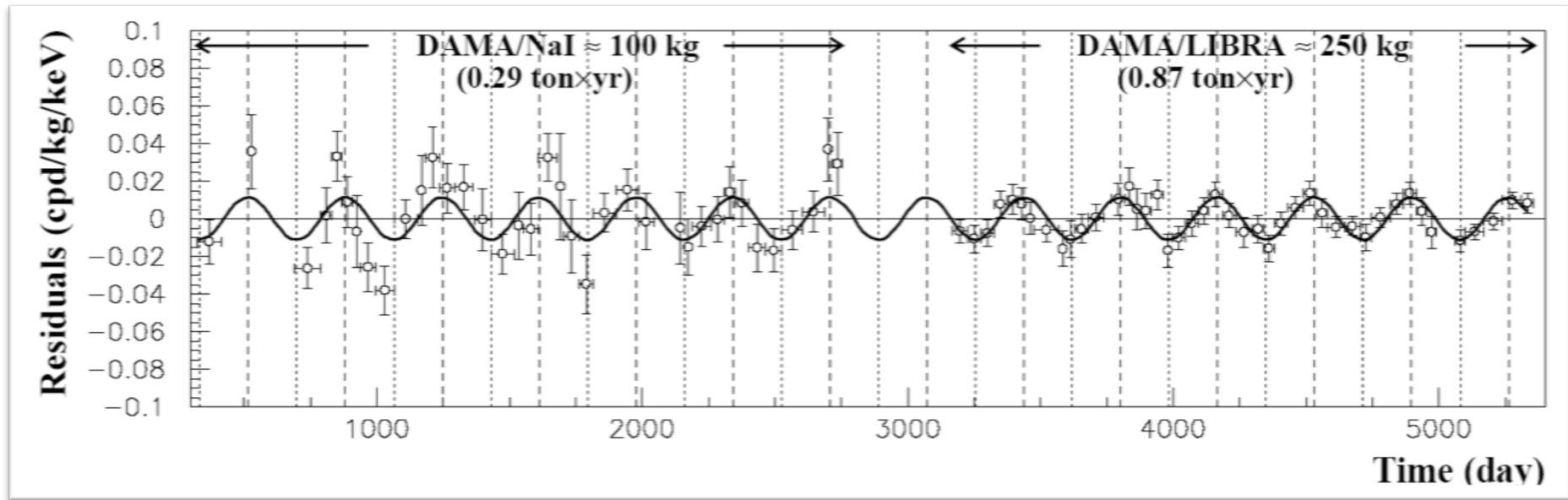
- phonon channel provides precise measurement of deposited energy
- Light channel distinguishes types of interaction
- Types of recoiling nuclei distinguished by different slopes in light energy plane

DAMA/LIBRA

- Ultrapure Na(Tl)
- Residual contamination
 - ^{232}Th , ^{238}U ~ 10 ppt
 - ^{40}K ~ 20 ppb

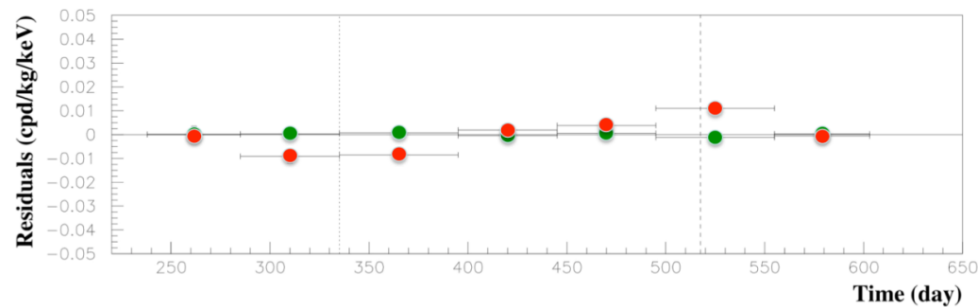


DAMA/LIBRA – annual modulation



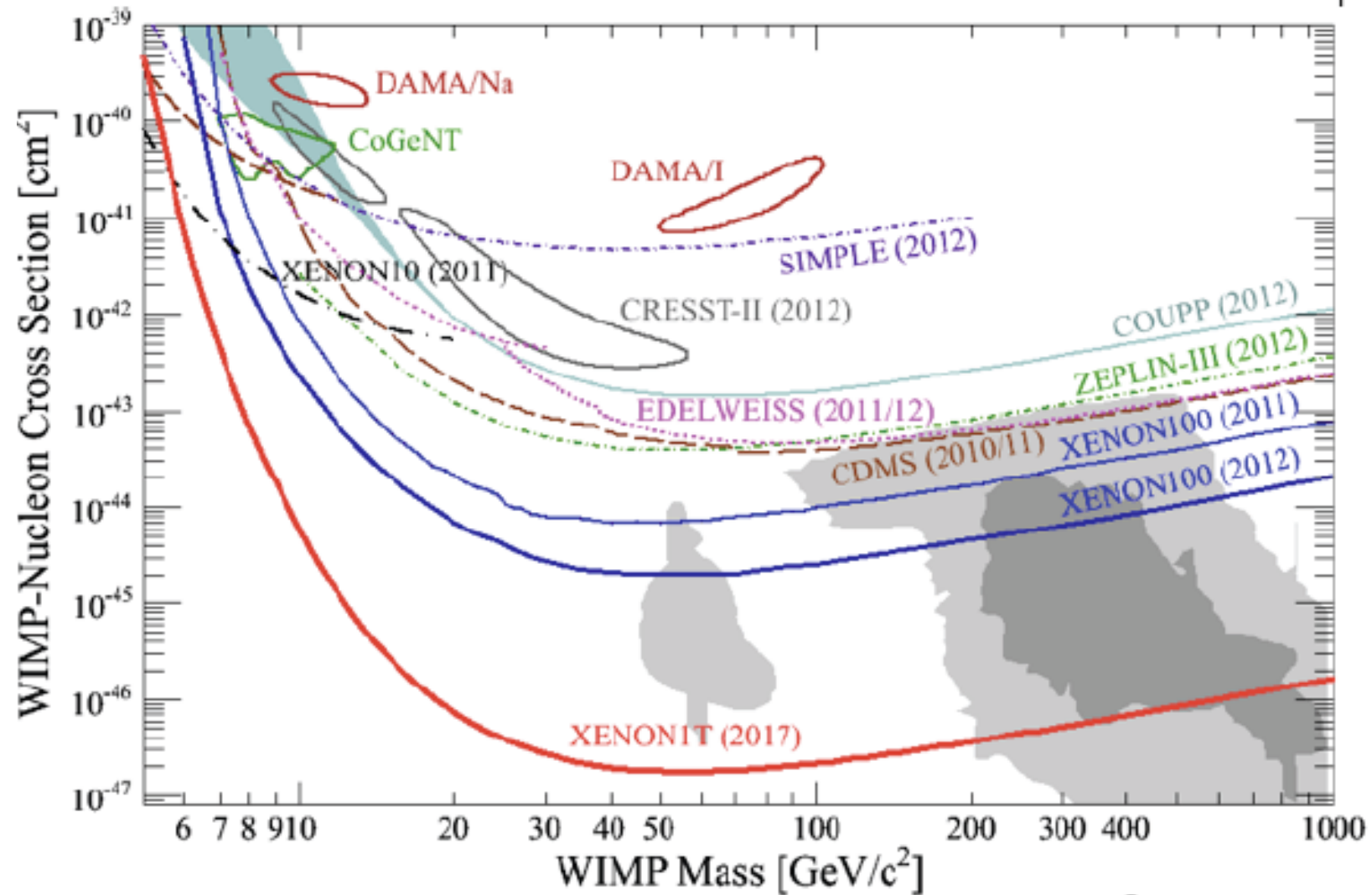
Comparison between **single hit residual rate (red points)** and **multiple hit residual rate (green points)**;
 $A = -(0.0006 \pm 0.0004)$ cpd/kg/keV

Multiple hits events = Dark Matter particle "switched off"
2-6 keV



DM searches

- no general consensus on DM detection by lab experiments
- **origin of Dama/Libra modulation MUST be clarified**



And more

- Test of fundamental principles
 - VIP: Pauli exclusion principle
 - Ginger (G-Gransasso) : Gyrolaser for General Relativity test: Earth gravitomagnetic effect (Lens-Thirring)
- Applications of trace elements measurements to environmental & earth sciences
 - ERMES-W
- Study of biological effects of low-doses on mammals
 - Cosmic Silence: testing the linear no-threshold hypothesis
- Outreach

OPEN DAY @ LNGS:1.500-2.000 visitors/year



European Researchers Night 2013 @ Galileium: 1.500 visitors



Educational Laboratories:
500-1000 students/years



Competition for schools: 1700 students/years



Visits at LNGS:
8000 visitors/year



Summer Schools for students
and teachers



Theatre performance in the underground facilities was followed by 2.5 Million people on a national TV channel.



Physics and Archaeology



Several tons (~6) of roman lead for shielding Cuore

2000 years old lead: free of ^{210}Pb (Romans effectively removed ^{238}U)

Many thanks to strong winds (or Alberto's ancestors?) which sunk the ship



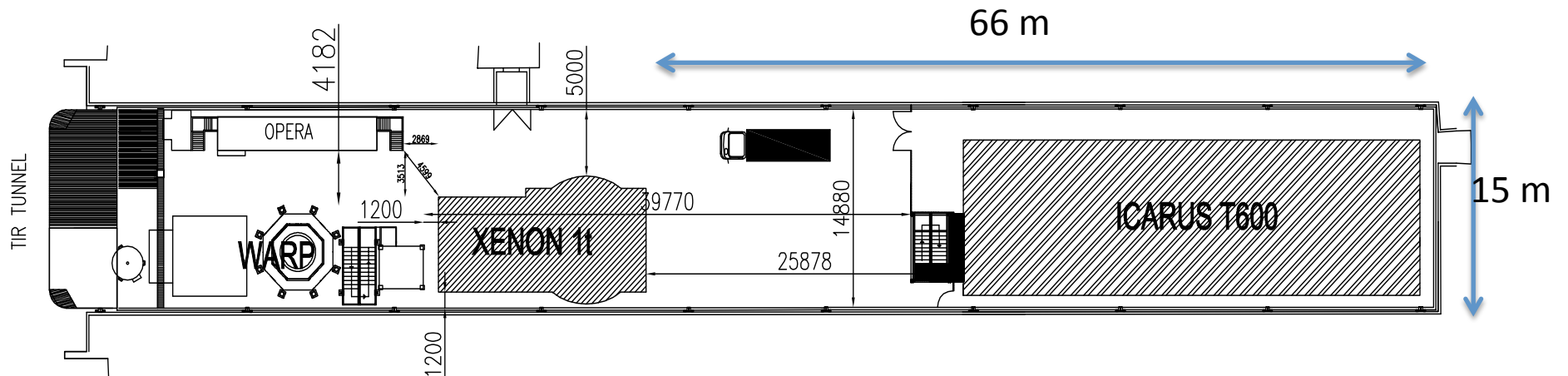
LNGS

- Room is being made free for new experiments

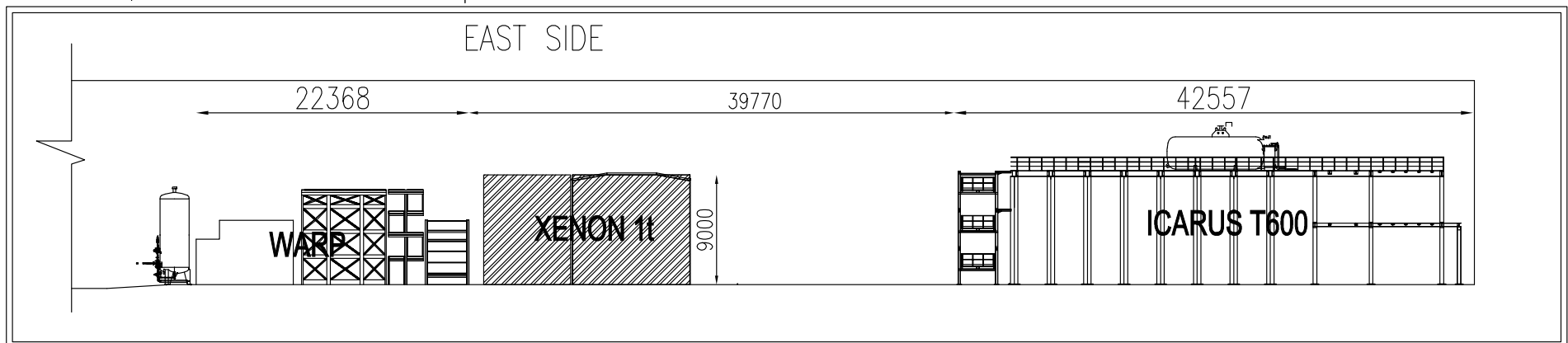
Hall b

- Warp area allocated to GINGER (G-Gransasso: Gyrolaser in General Relativity)
- Icarus decommissioning: June 2013 -> October 2014

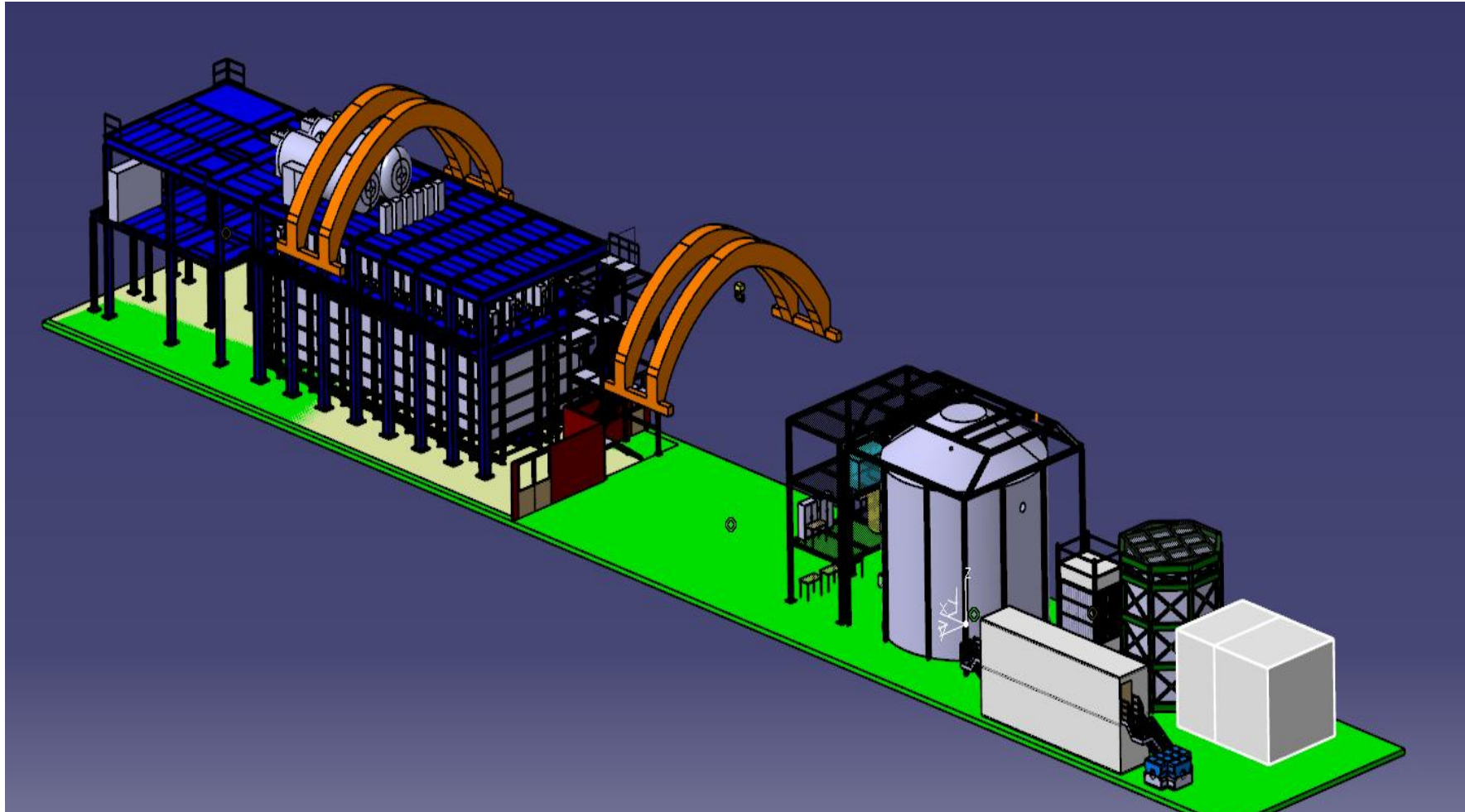
PLANT VIEW



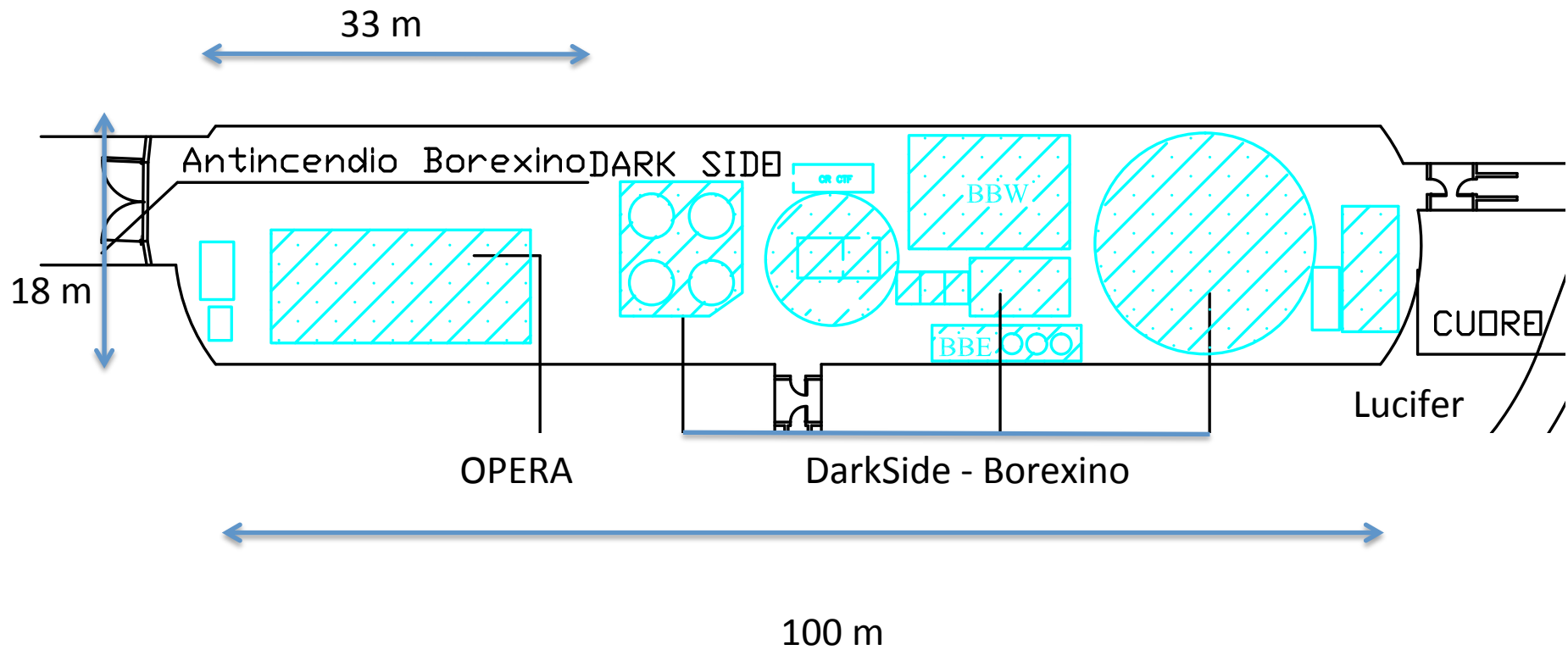
EAST SIDE



Hall B



Hall C – present status



Which future?

- 3rd generation $\beta\beta$ (IH region) and DM
 - Technique? new xtals? Xenon?
 - Size?
 - Bkg reduction?
 - Can the existing active shieldings be considered as facilities for future experiments?
 - Can/should LNGS invest on a general-purpose active shielding?
- What else?
 - Ginger takes advantage of tremendous reduction of seismic noise underground: more experiments/measurements?
 - Great expertise in ultra purification and measurement of trace elements -> look for applications

Prepare the future

- Stimulate international initiatives and attract international investment
 - e.g. Gerda, Xenon, DarkSide ...
- LNGS is a key research facility in the world
 - open to international access with transparent and unbiased procedures
- Coordination and synergies with other UG labs
 - EUROPE: EULAB-ERIC an important (fundamental) step

ERIC

- Ongoing discussions to constitute a distributed European Research Infrastructure Consortium including (most of) European UG labs
- Consistent savings on taxes
- European procurement/tendering rules simpler than Italian PA rules
- **Most important: hire personnel**

Conclusions

- A national lab and international asset
- Room is being made available for the equivalent of > 3 Xenon-1T experiments
- **Science Proposals renewing and strengthening a long standing and successful INFN-US collaboration are welcome**