CALOCUBE
SVILUPPO DI CALORIMETRIA OMOGENEA AD ALTA ACCETTANZA PER ESPERIMENTI DI RAGGI COSMICI NELLO SPAZIO

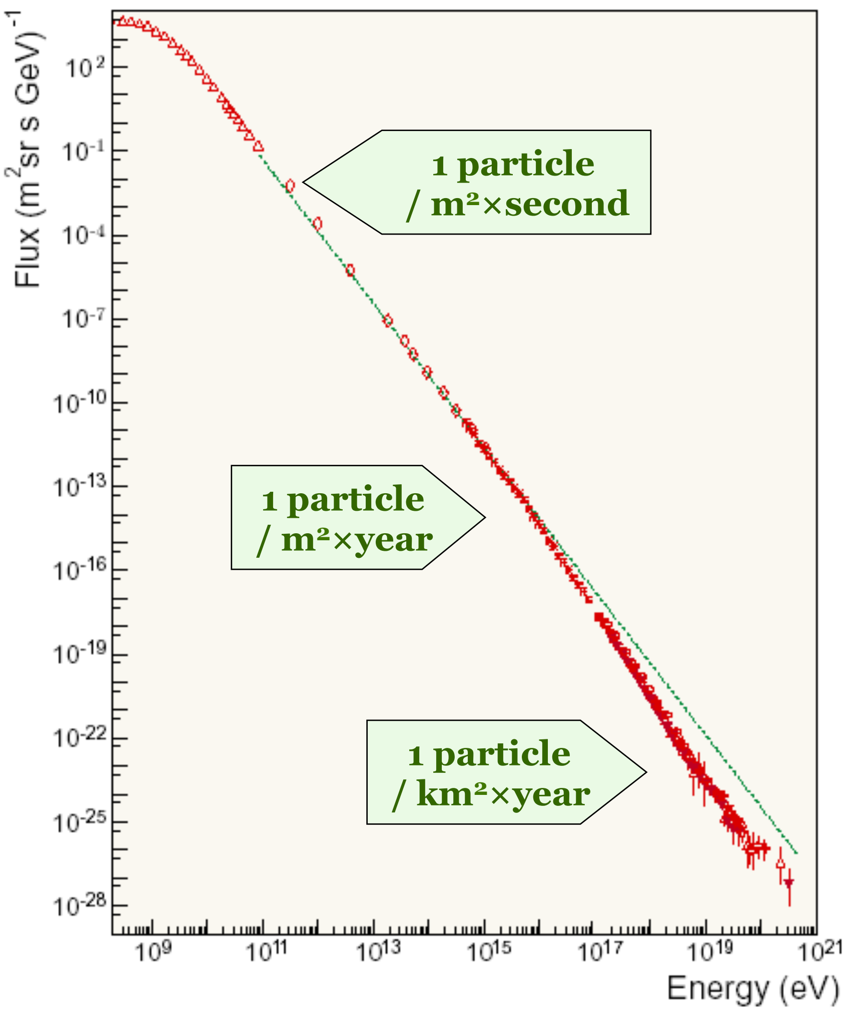
Call nell'ambito della CSN5 dell'INFN

Oscar Adriani, Responsabile Nazionale

PHYSICS MOTIVATION

And the starting point for the CaloCube proposal….

Some of the Cosmic-Ray 'mysteries'



High energy nuclei

- **"Knee" structure around ~ PeV**
  - Upper energy of galactic accelerators (?)
  - Energy-dependent composition
- **Structures in the GeV – TeV region recently discovered for p and He**
  - Composition at the knee may differ substantially from that at TeV
- **Spectral measurements in the knee region up to now are only indirect**
  - Ground-based atmospheric shower detectors
  - High uncertainties

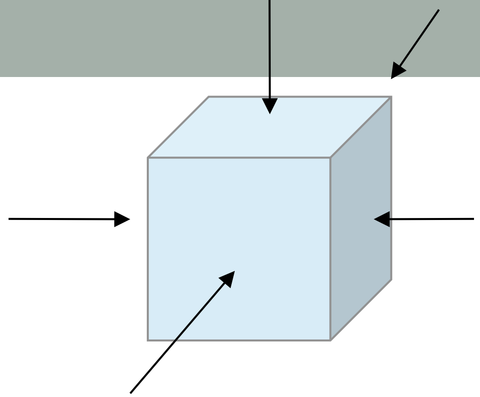
A direct spectral measurement in the PeV region requires great acceptance (few m^{2}sr) and good energy resolution for hadrons (at least 40%)

High energy Electrons+Positrons

- **Currently available measurements show some degree of disagreement in the 100 GeV – 1 TeV region**
- **Cutoff in the TeV region?**

Direct measurements require excellent energy resolution (~%), a high e/p rejection power (> 10^5) and large acceptance above 1 TeV

Our proposal for an 'optimal' CR detector



- **A 3-D, deep, homogeneous and isotropic calorimeter can achieve these design requirements:**
  - depth and homogeneity to achieve energy resolution
  - isotropy (3-D) to accept particles from all directions and increase GF
- **Proposal: a cubic calorimeter made of small cubic sensitive elements**
  - can accept events from 5 sides (mechanical support on bottom side) → GF * 5
  - segmentation in every direction gives e/p rejection power by means of topological shower analysis
  - cubic, small (~Moliére radius) scintillating crystals for homogeneity
  - gaps between crystals increase GF and can be used for signal readout
    - small degradation of energy resolution
  - must fulfill mass&power budget of a space experiment
    - modularity allows for easy resizing of the detector design depending on the available mass&power

Additional details….

- Exercise made on the assumption that the detector's only weight is ~ 1600 kg
  - Mechanical support is not included in the weight estimation
- The optimal material is CsI(Tl)

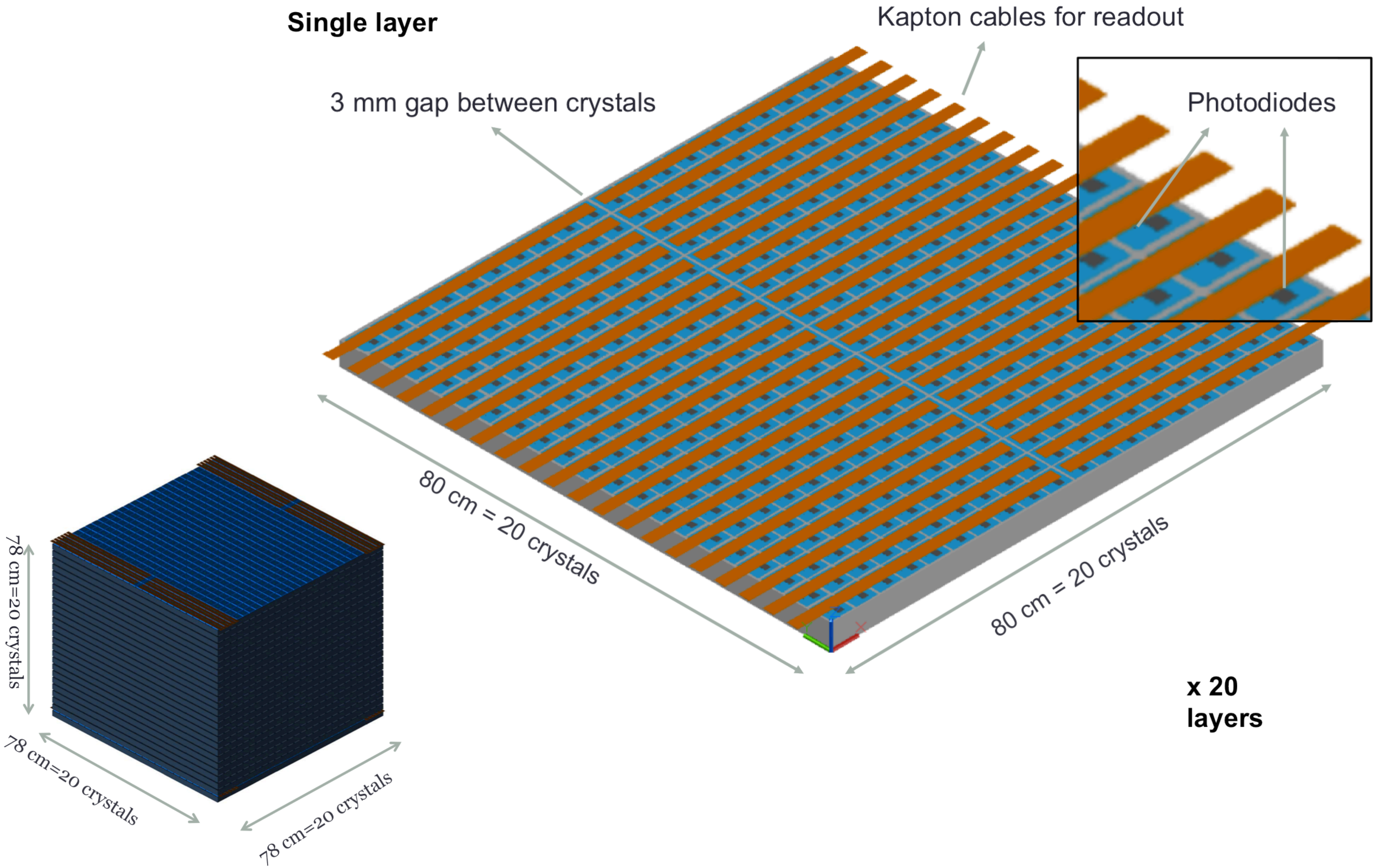
Density:	4.51 g/cm^3
X_0:	1.85 cm
Moliere radius:	3.5 cm
λ_I:	37 cm
Light yield:	54.000 ph/MeV
τ_{decay}:	1.3 μs
λ_{max}:	560 nm

- Simulation and prototype beam tests used to characterize the detector

N×N×N	**20×20×20**
L of small cube (cm)	3.6*
Crystal volume (cm^3)	46.7
Gap (cm)	**0.3**
Mass (Kg)	1683
N.Crystals	8000
Size (cm^3)	78.0×78.0×78.0
Depth (R.L.) " (I.L.)	39×39×39 1.8×1.8×1.8
Planar GF (m^{2}sr) **	1.91

(* one Moliere radius)
(** GF for only one face)

Mechanical idea



The readout sensors and the front-end chip

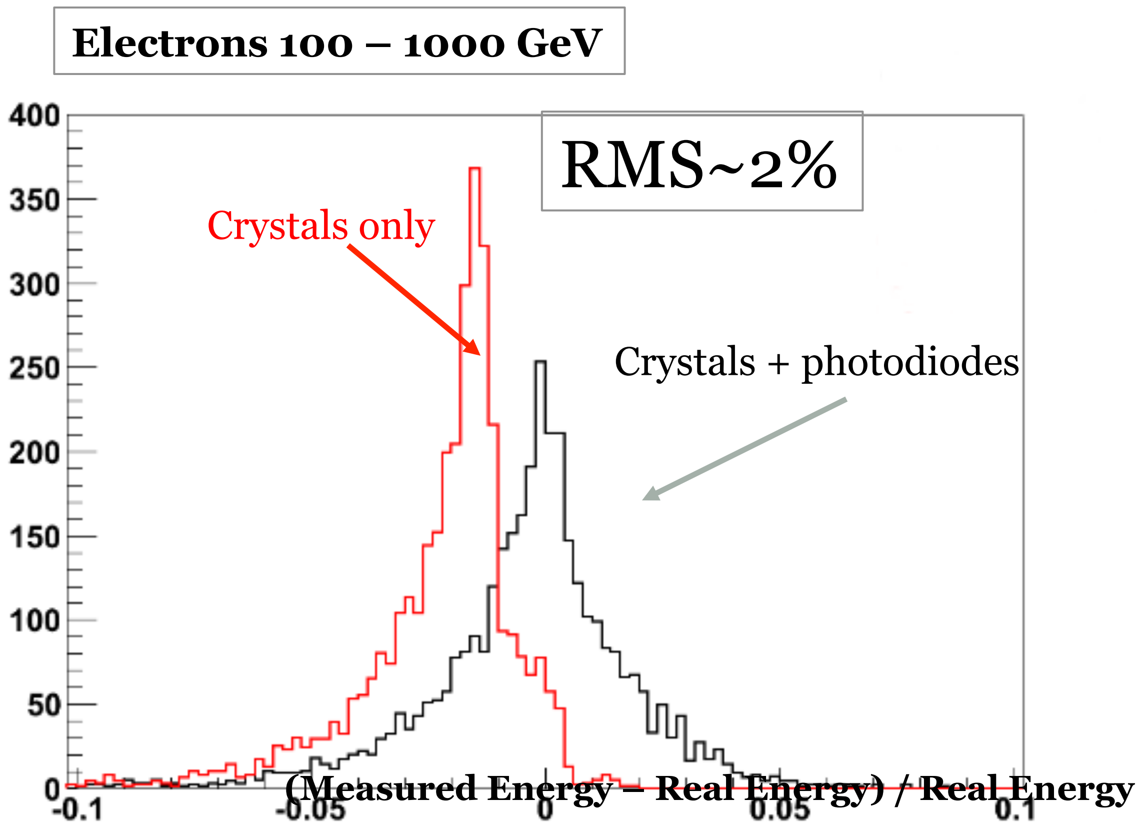
- Minimum 2 Photo Diodes are necessary on each crystal to cover the whole huge dynamic range 1 MIP→10^7 MIPS
  - Large Area Excelitas VTH2090 9.2 x 9.2 mm^2 for small signals
  - Small area 0.5 x 0.5 mm^2 for large signals

- Front-End electronics: a big challenge!
- The CASIS chip, developed in Italy by INFN-Trieste, is very well suited for this purpose
  - IEEE TRANSACTIONS ON NUCLEAR SCIENCE, VOL. 57, NO. 5, OCTOBER 2010
- 16 channels CSA+CDS
- Automatic switching btw low and high gain mode
- 2.8 mW/channel
- 3.10^3 e^- noise for 100 pF input capacitance
- 53 pC maximum input charge

MC simulations

- **Fluka-based MC simulation**
  - Scintillating crystals
  - Photodiodes
    - Energy deposits in the photodiodes due to ionization are taken into account
  - Carbon fiber support structure (filling the 3mm gap)
- **Isotropic generation on the top surface**
  - Results are valid also for other sides
- **Simulated particles:**
  - Electrons: 100 GeV → 1 TeV
  - Protons: 100 GeV → 100 TeV
  - about $10^2 - 10^5$ events per energy value
- **Geometry factor, light collection and quantum efficiency of PD are taken into account**
- **Requirements on shower containment (fiducial volume, length of reconstructed track, minimum energy deposit)**
  - Nominal GF: $(0.78*0.78*\pi)*5*\varepsilon\ m^2sr = 9.55*\varepsilon\ m^2sr$

Electrons



Selection efficiency:
ε ~ 36%

GF_{eff} ~ 3.4 m²sr

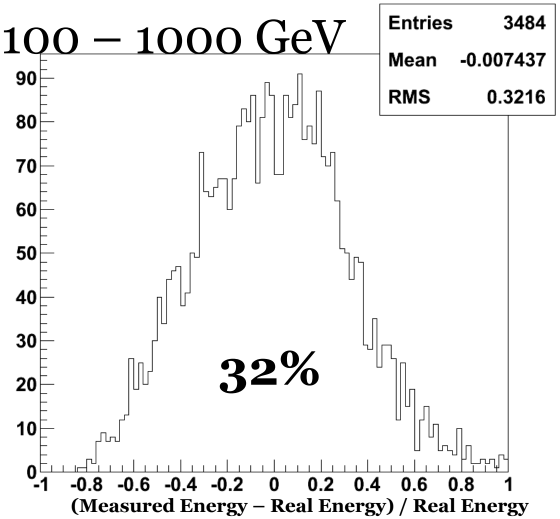
Non-gaussian tails due to leakages and to energy losses in carbon fiber material

Ionization effect on PD: 1.7%

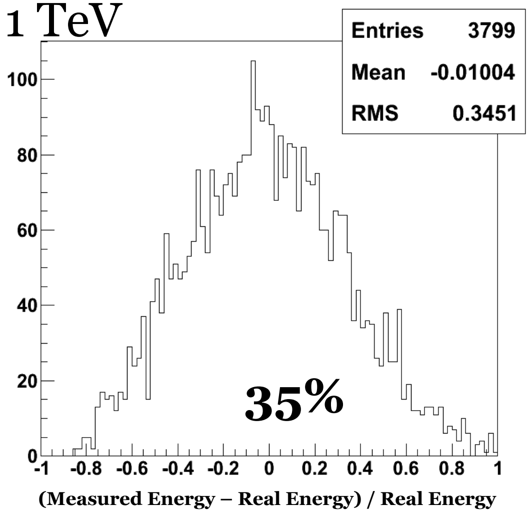
Protons

Energy resolution (correction for leakage by looking at the shower starting point)

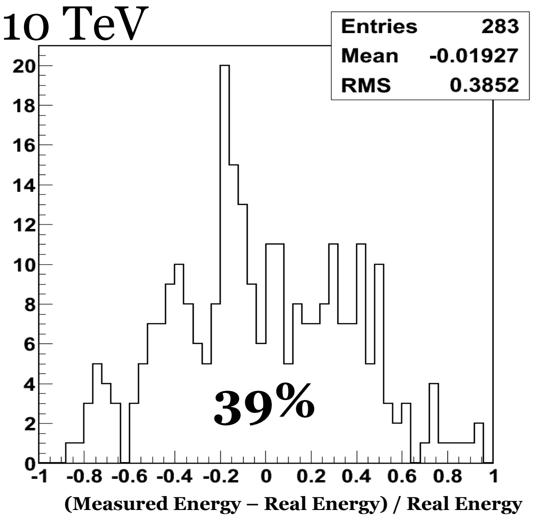
100 – 1000 GeV



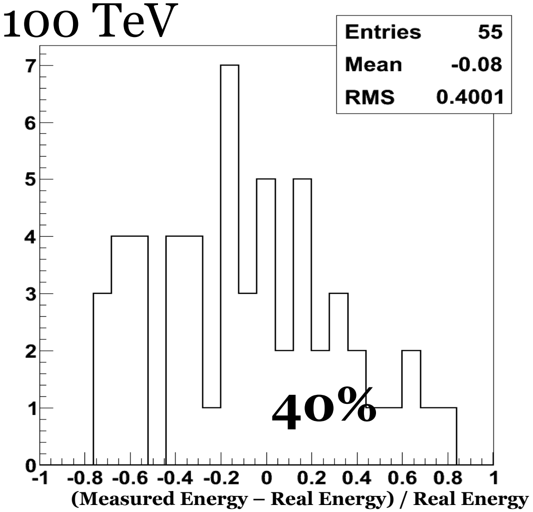
1 TeV



10 TeV



100 TeV



Selection efficiencies:

$\varepsilon^{0.1\text{-}1TeV}$ ~ 35%

ε^{1TeV} ~ 41%

ε^{10TeV} ~ 47%

$GF_{eff}^{0.1\text{-}1TeV}$ ~ 3.3 m²sr

Gf_{eff}^{1TeV} ~ 3.9 m²sr

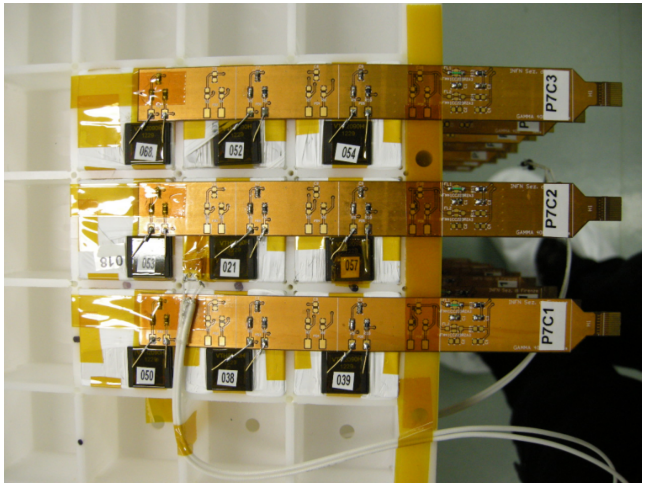
Gf_{eff}^{10TeV} ~ 4.5 m²sr

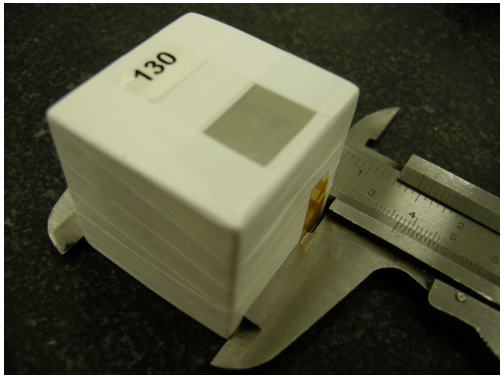
Proton rejection factor with simple topological cuts:

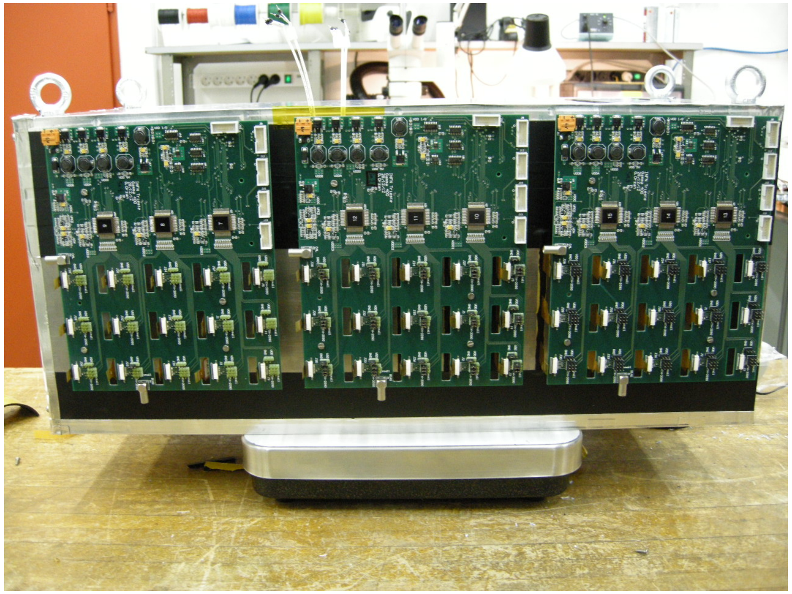
2.10^5-5.10^5 up to 10 TeV

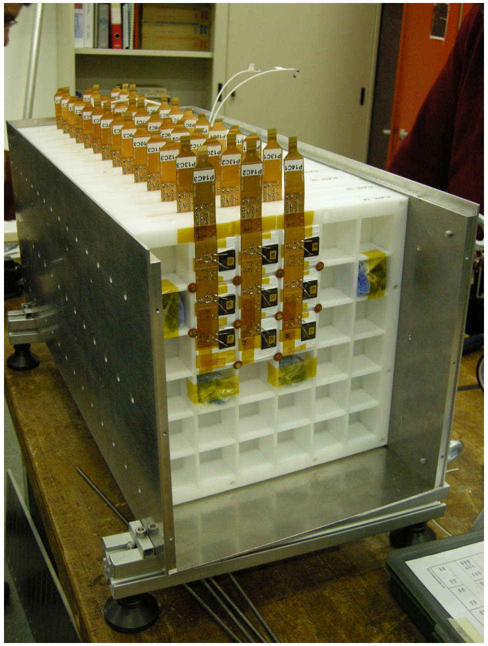
The prototype

14 Layers
9x9 crystals in each layer
126 Crystals in total
126 Photo Diodes
50.4 cm of CsI(Tl)
27 X_0
1.44 λ_I





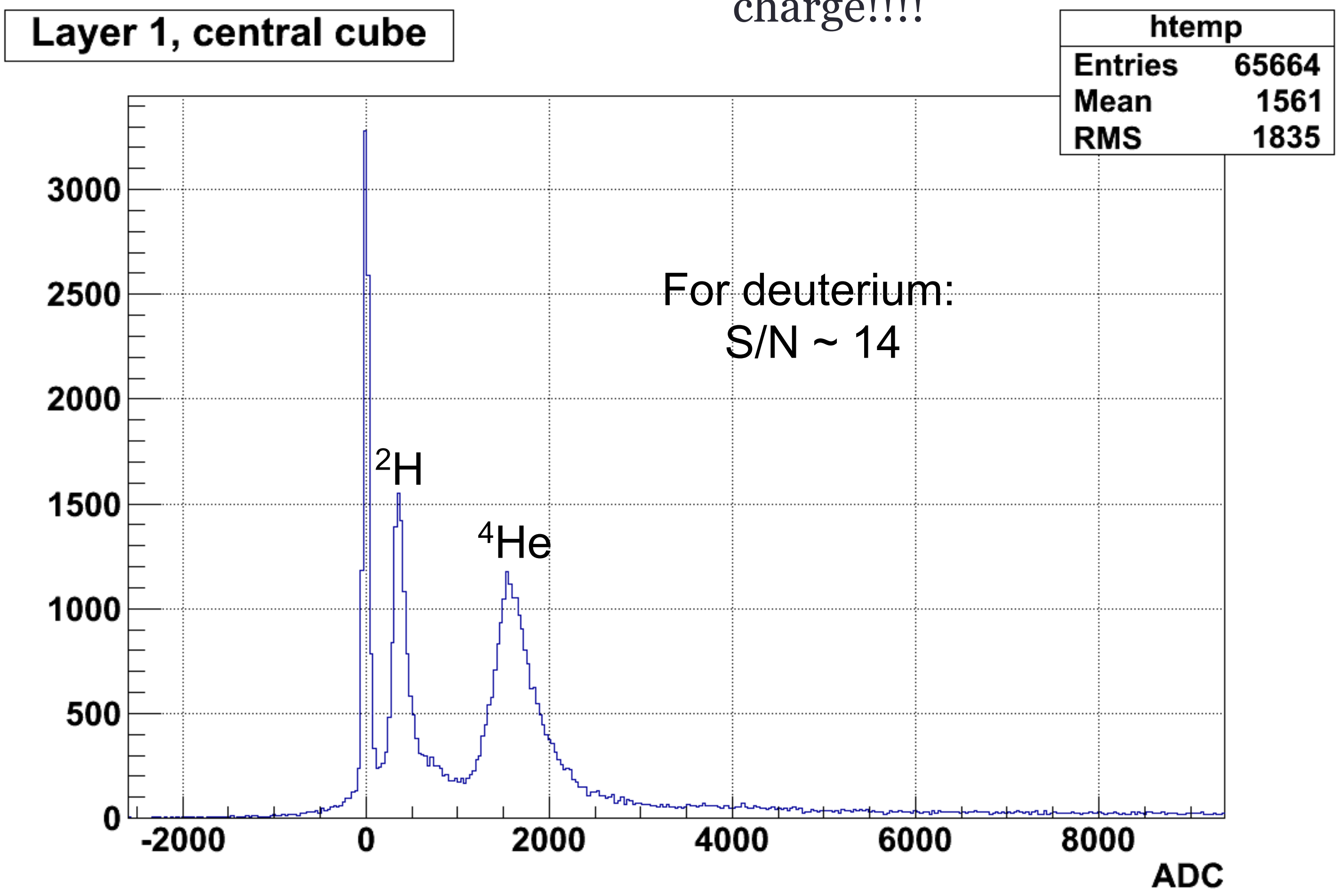




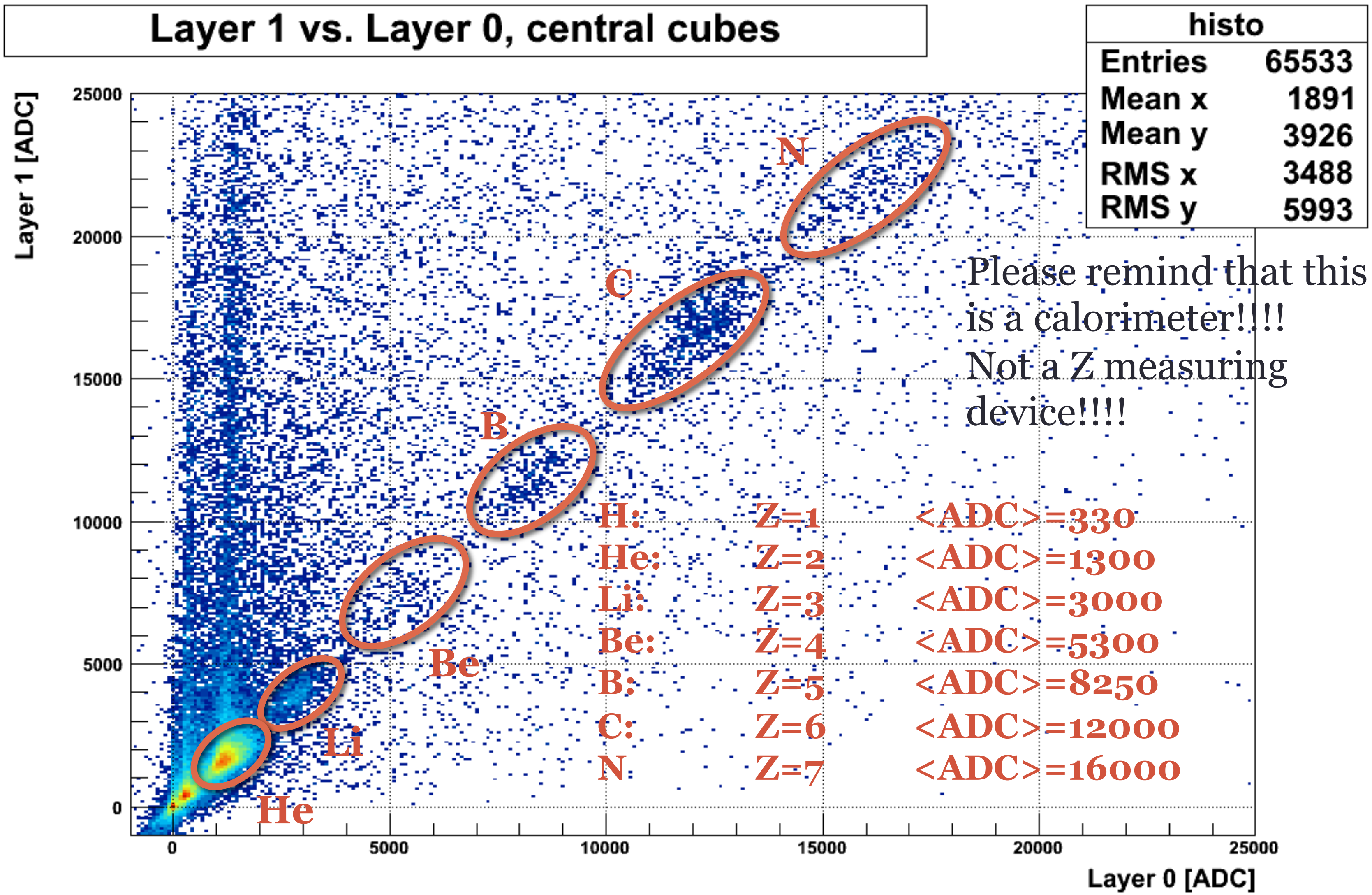
A glance at prototype's TB data

SPS H8 Ion Beam: Z/A = 1/2, 12.8 GV/c and 30 GV/c

Please note: we can use the data from a precise silicon Z measuring system located in front of the prototype to have an exact identification of the nucleus charge!!!!



A glance at prototype's TB data



HOW TO IMPROVE THE PERFORMANCES OF THE INSTRUMENT?

The CaloCube idea

Improve the existing Cubic Calorimeter concept to:

1. Optimize the hadronic energy resolution
   - Build up a really compensating calorimeter
     - Cherenkov light
     - Neutron induced signals
2. Optimize the charge measurement
   - Make use of the excellent results from the SPS test
   - Smaller size cubes on the lateral faces
   - Materials to reduce back scattering effect
3. Build up a prototype fully space qualified
   - Mechanics
   - Thermics

by developing highly innovative techniques that are one of the core interest of the INFN CSN5

The compensation technique

- Normally e/h>1
  - Some of the hadronic energy is lost
    - Neutrinos, muons, break up and excitation of nuclei, etc.
- Inside an hadronic shower there is a significant component of e.m. energy (π^0 decays!), different form event to event
- If e/h≠1, the hadronic energy measurement is worsen
- Idea:
  - Suppress the em signal or increase the hadronic energy signal
    - Almost impossible in homogeneous calorimeter
  - Find some estimator to measure the em fraction
    - Our idea: make use of the Cherenkov and neutron signals

How to detect Ch and neutrons?

- Insert in the calorimeter different types of cubes
  - CsI(T) for scintillation and Cherenkov light
  - Plastic scintillators for neutron thermalization
  - Aerogel radiators for Cherenkov light
- Use the timing information to discriminate btw scintillation (slow), Cherenkov (fast), and neutron (very slow) component
- Apply optical filters to discriminate btw Cherenkov and scintillation

THE ORGANIZATION

The groups involved

1. INFN
   1. Firenze
   2. Pisa
   3. Trieste/Udine
   4. Pavia
   5. Catania/Messina

2. External institutions
   1. FBK
   2. ST Microelectronics
   3. IMCB-CNR Napoli

Work Packages

- WP1: System design, software and simulation
  - Responsible: Oscar Adriani - Firenze
- WP2: Charge identifier system
  - Responsible: Paolo Maestro – Pisa
- WP3: Crystals and radiators
  - Responsible: xxx - Udine
- WP4: Optical treatments and light collection systems
  - Responsible: Sergio Ricciarini – Firenze
- WP5: Photodetectors
  - Responsible: Valter Bonvicini – Trieste
- WP6: Electronics
  - Responsible: Paolo Cattaneo – Pavia
- WP7: Beam tests
  - Responsible: Sebastiano Albergo – Catania
- WP8: Space qualification
  - Responsible: Guido Castellini - Firenze

The WP in the various groups

			FI	TS/UD	PI	PV	CT/ME	NA INFN/ NA CNR	FBK	ST
WP1 Adriani	**System design, software and simulation**	Definizione geometrie	x	x		x				
		Scelta cristalli	x							
		Tecniche analisi	x							
		Physics case (cenni)	x							
WP2 Maestro	**Charge identifier system**				x					
WP3 Udine	**Crystals and radiators**			x						
WP4 Ricciarini	**Optical treatments and light collection**	Deposizioni	x				x	x		
		Caratterizzazioni ottiche					x			
WP5 Bonvicini	**Photodetectors**	SiPM	X	x					x	x
		Fotodiodi	x	x					x	x
WP6 Cattaneo	**Electronics**	Front-end		x						
		Read-out	x			x				
WP7 Albergo	**Beam tests**		x				x			
WP8 Castellini	**Space qualification**	Meccanica			x					
		Termica	x							

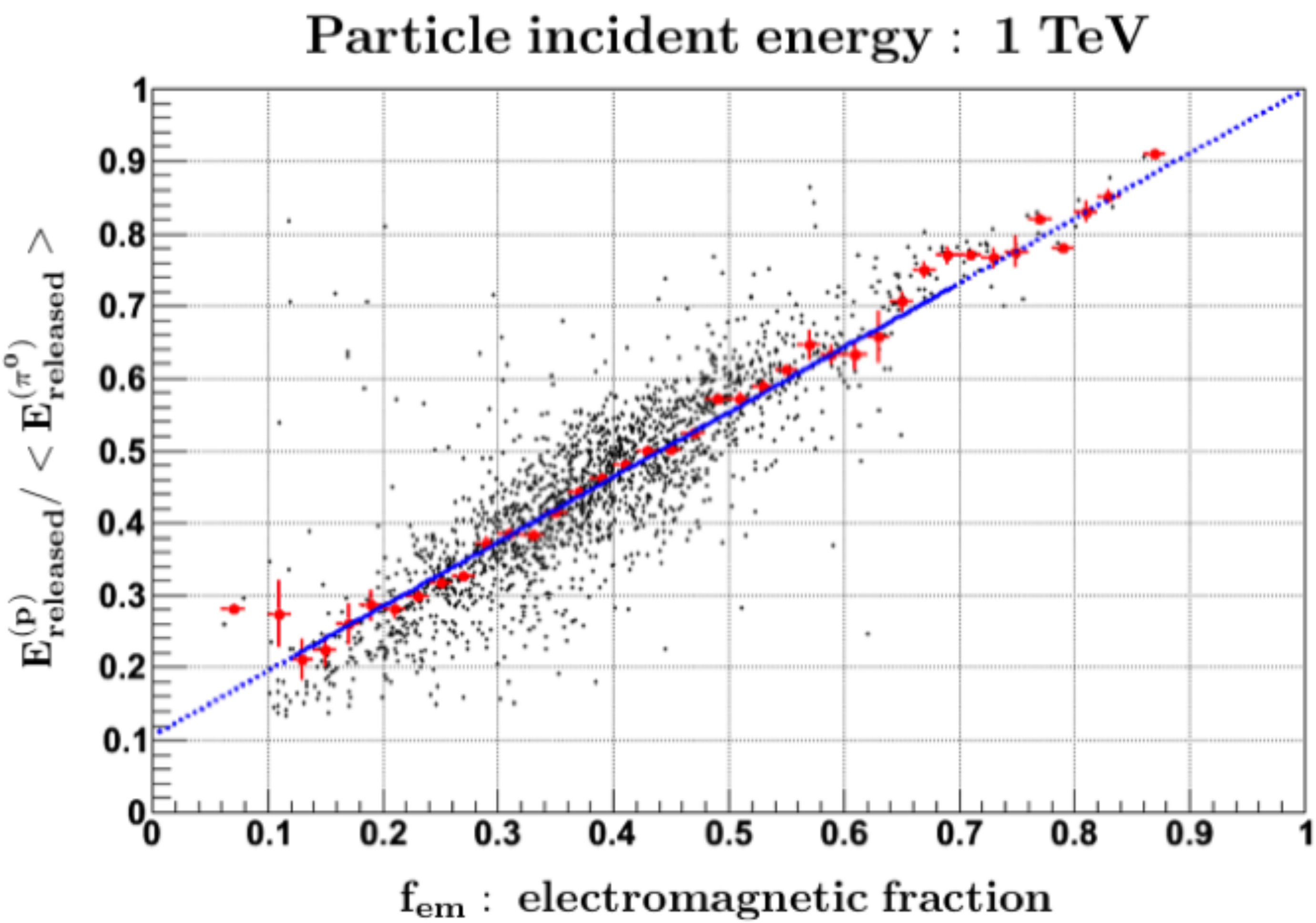
Anagrafica

Adriani	50
Bongi	70
Castellini	70 (0?)
Lenzi	30
Ricciarini	70
Starodubtsev	70
Detti	50

Richieste di servizi

- 2 m.u. servizio di elettronica
- 2 m.u. officina

BACKUP

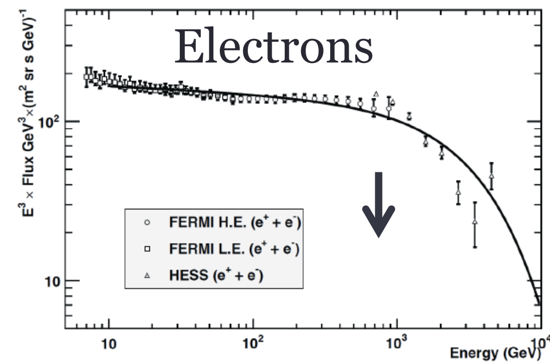


Particle incident energy : 1 TeV
$E^{(P)}_{released} / < E^{(\pi^0)}_{released} >$
f_{em} : electromagnetic fraction
0
0.1
0.2
0.3
0.4
0.5
0.6
0.7
0.8
0.9
1

What we can reach with this calorimeter?

Assumptions:

- 10 years exposure
- No direct closeby sources for electrons
- Polygonato model for protons/nuclei



Electrons

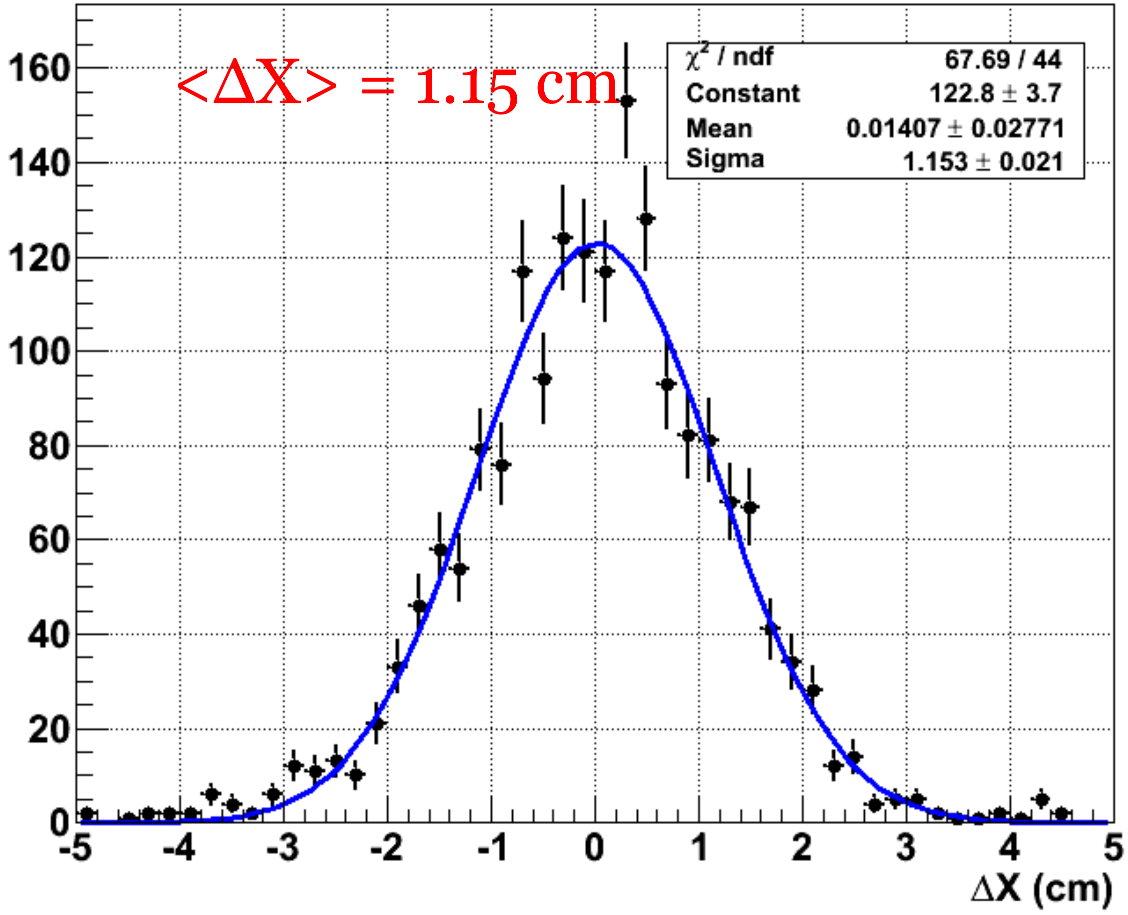
Gf_{eff} (m^2sr)	ΔE/E	Depth (X_0)	e/p rej. factor	E>0.5 TeV	E>1 TeV	E>2 TeV	E>4 TeV
3.4	2%	39	$>10^5$	~2.10^5	~4.10^4	~6.10^3	~7.10^2

~ knee

Protons and Helium

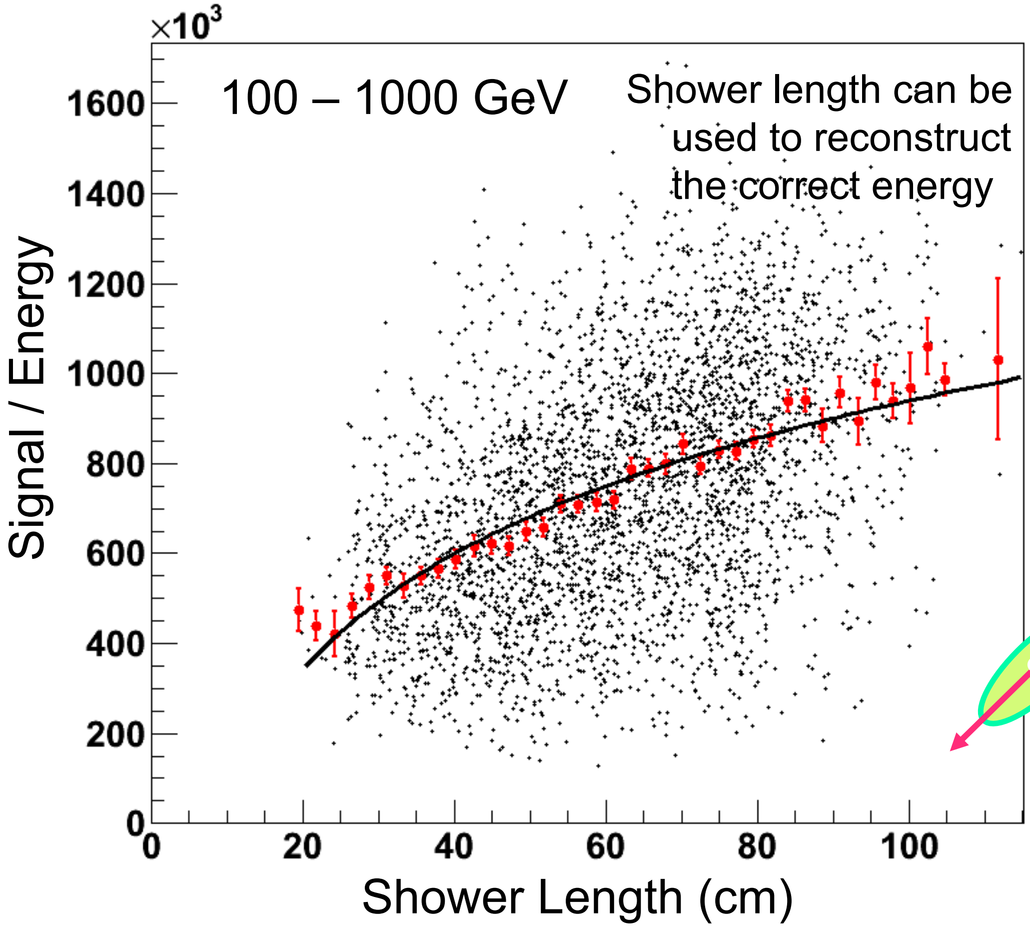
Gf_{eff} (m^2sr)	ΔE/E	Depth (λ_I)	E>100 TeV		E>500 TeV		E>1000 TeV		E>2000 TeV		E>4000 TeV	
			p	He	p	He	p	He	p	He	p	He
~4	40%	1.8	2.8x10^4	2.7x10^4	1.7x10^3	1.8x10^3	4.4x10^2	5.5x10^2	1.0x10^2	1.6x10^2	1.7x10^1	3.6x10^1

Shower starting point resolution

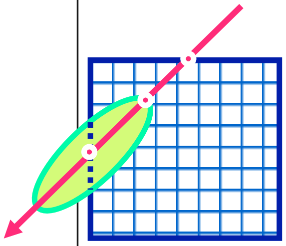


Protons

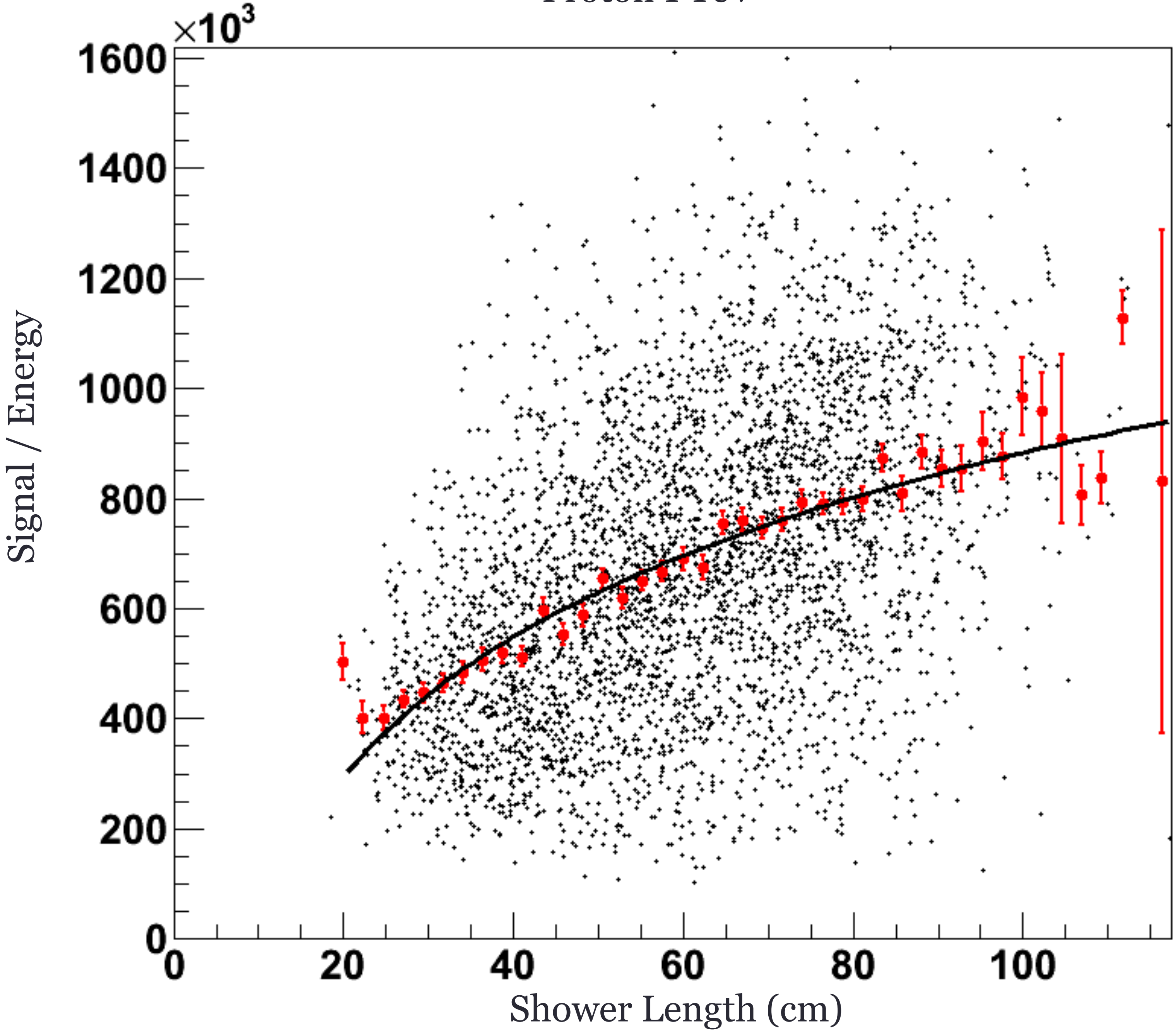
Energy estimation



Red points: profile histogram
Fitted with logarithmic function

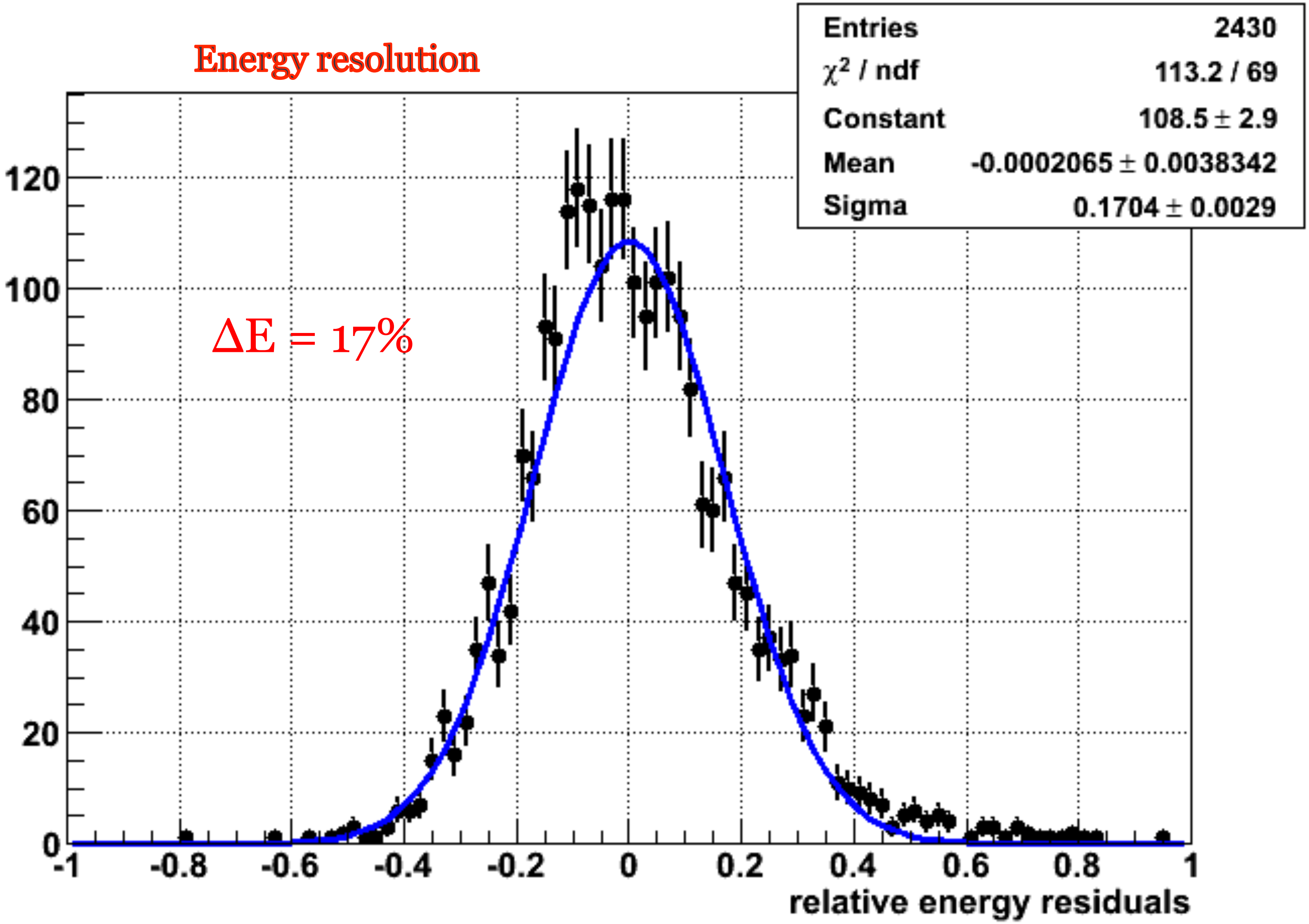


Proton 1 TeV

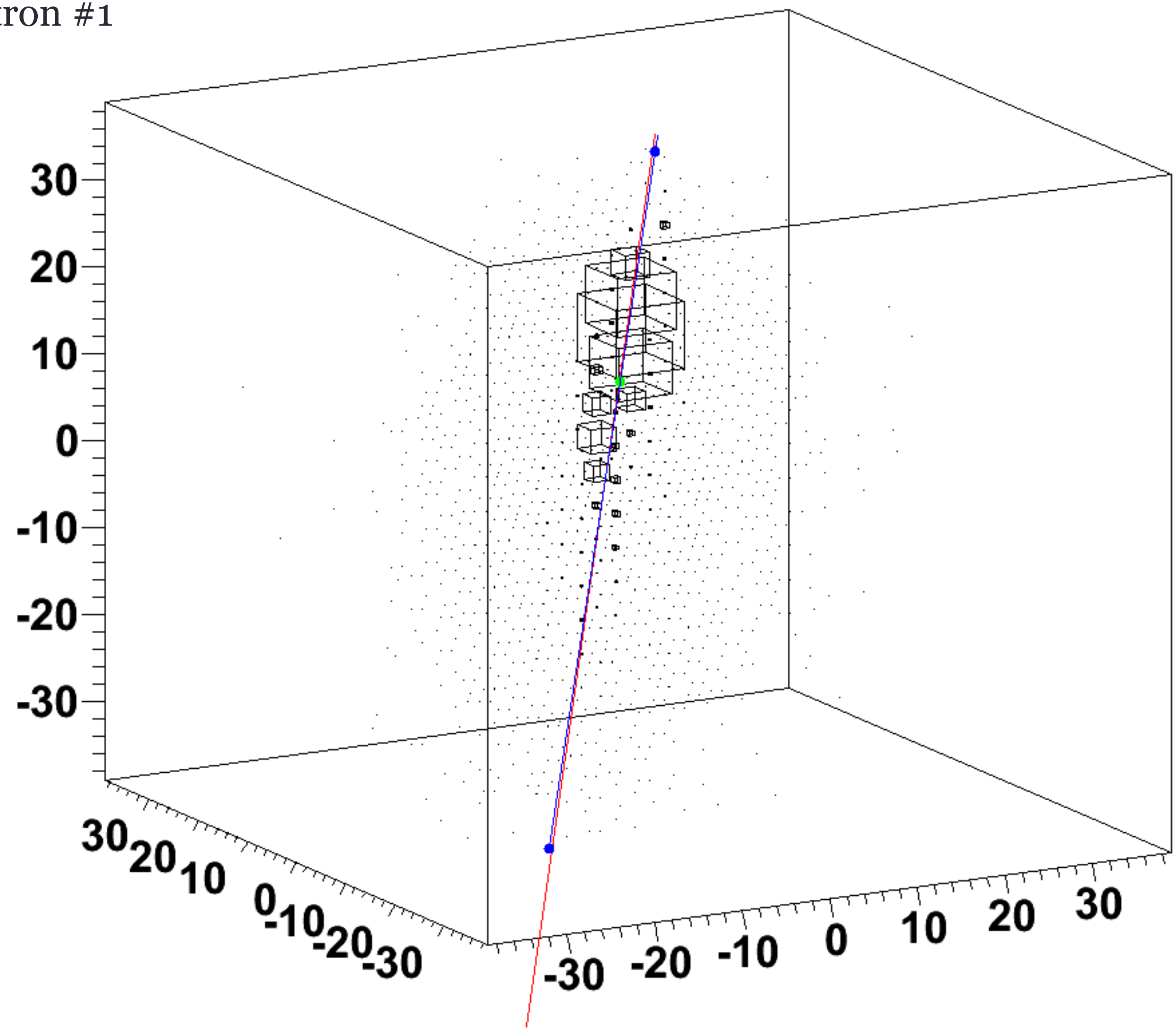


×10³
1600
1400
1200
1000
800
600
400
200
0
Signal / Energy
0
20
40
60
80
100
Shower Length (cm)

Energy resolution



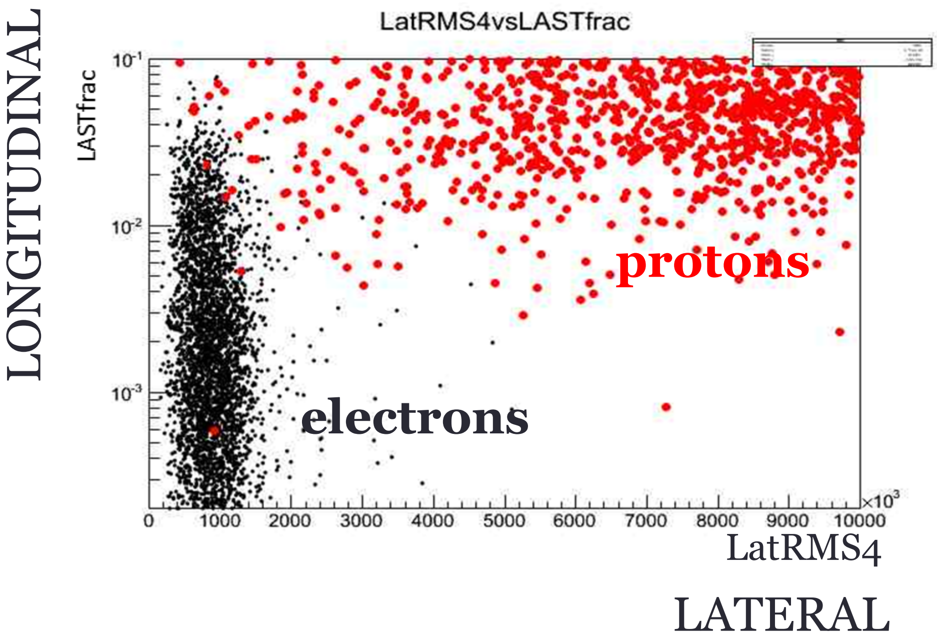
Electron #1



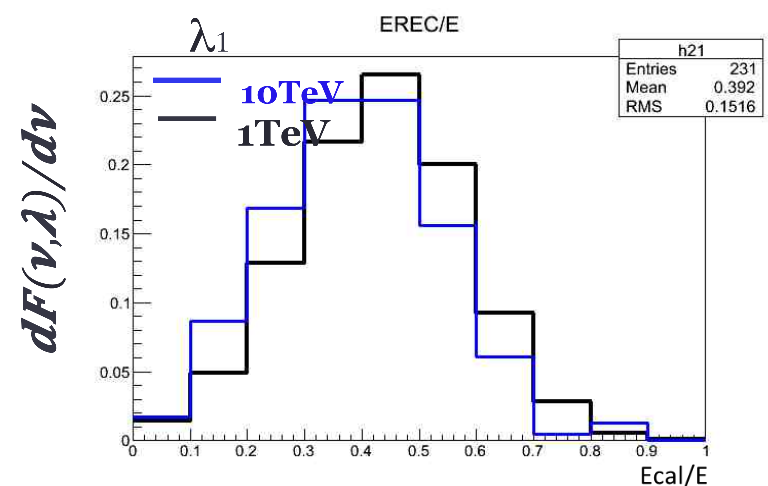
Proton rejection factor

Montecarlo study of proton contamination
using CALORIMETER INFORMATIONS ONLY

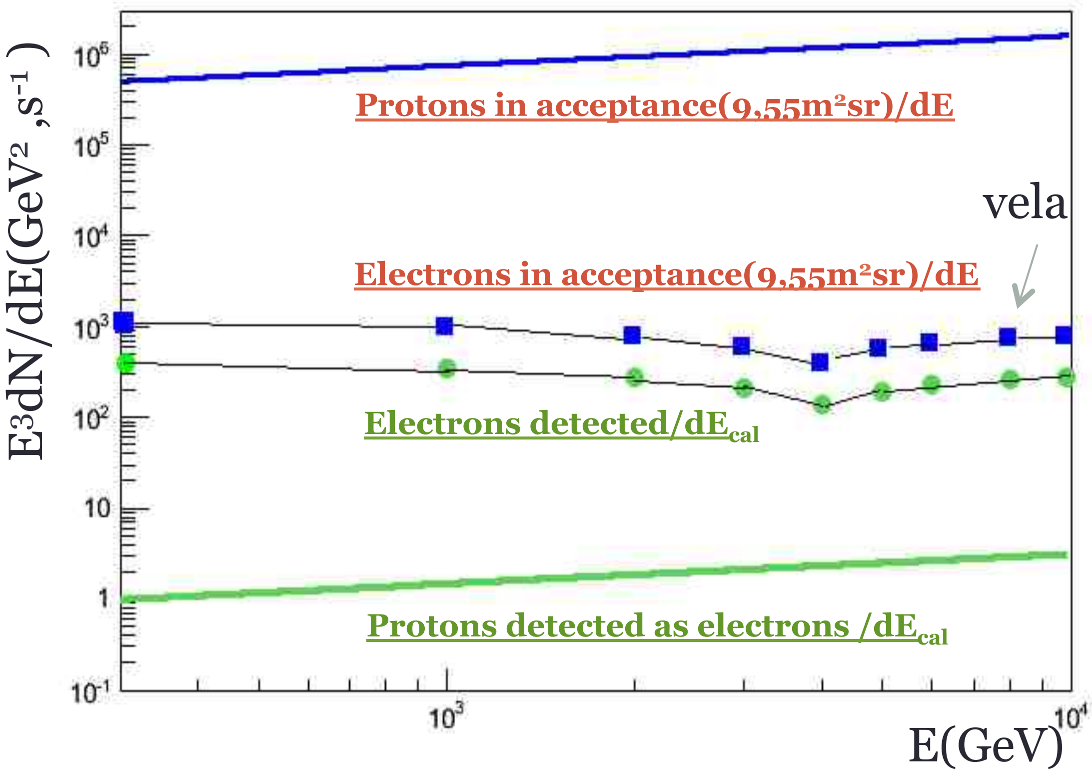
- **PARTICLES propagation & detector response simulated with FLUKA**
- **Geometrical cuts for shower containment**
- **Cuts based on longitudinal and lateral development**



- **155.000 protons simulated at 1 tev : only 1 survive the cuts**
- **The corresponding electron efficiency is 37% and almost constant with energy above 500gev**
- **Mc study of energy dependence of selection efficiency and calo energy distribution of misreconstructed events**



Proton rejection factor



Contamination :
0,5% at 1TeV
2% at 4 TeV

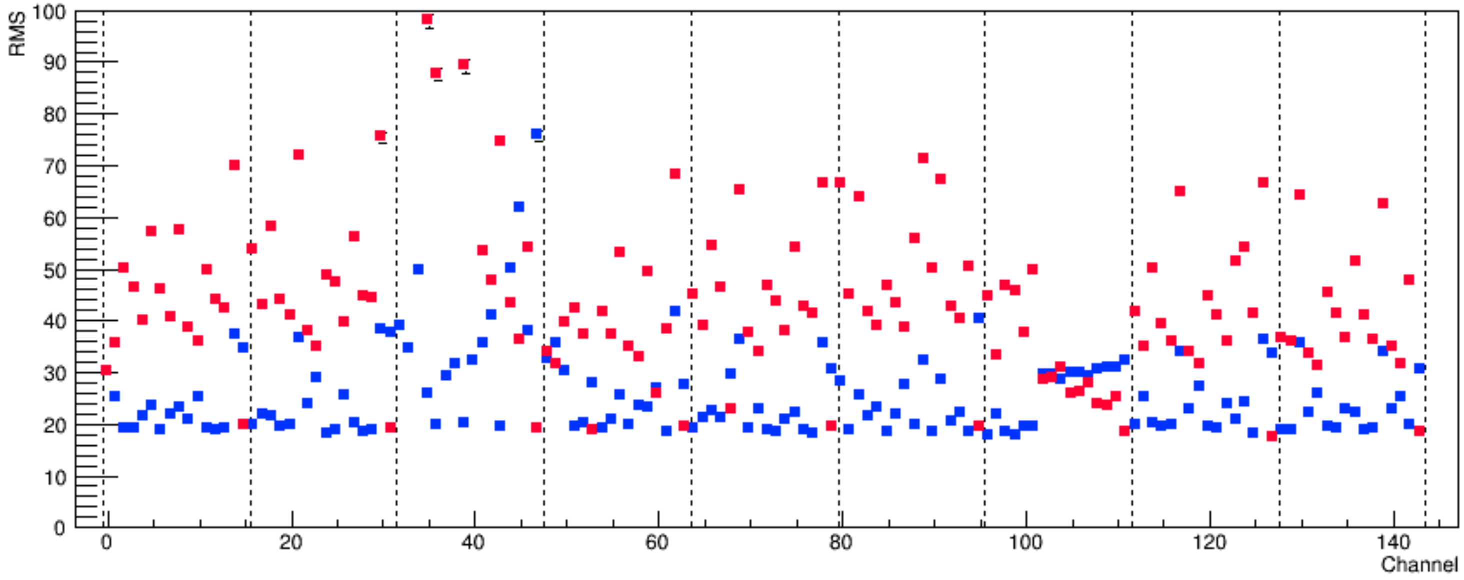
An upper limit 90% CL is obtained using a factor X 3,89

= = 0,5 x 10^6

X Electron Eff. ~ 2 x 10^5

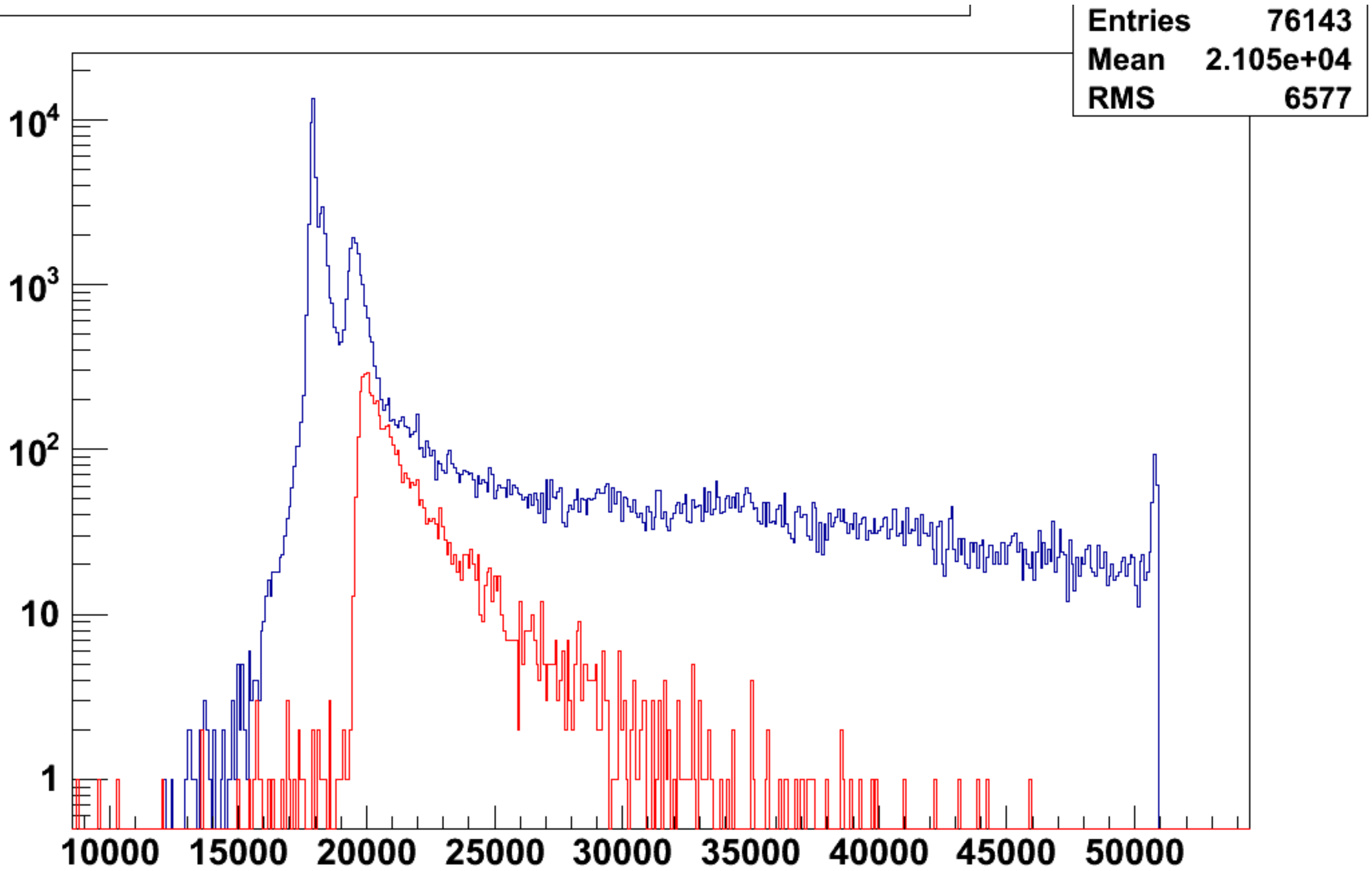
Noise

WITH and WITHOUT CN subtraction

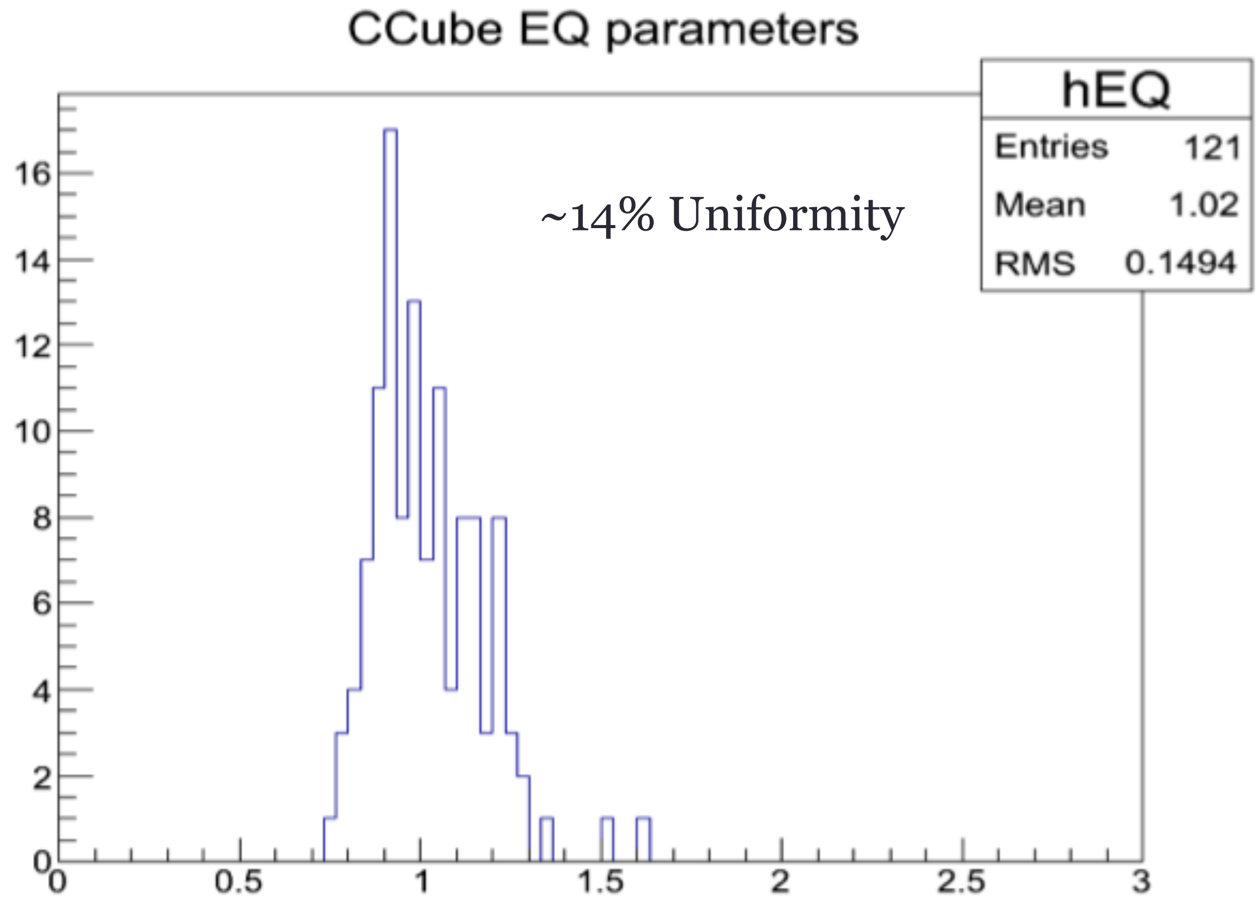


CN evaluated **without** disconnected channels

Signal in the central cube in High Gain (Blu) and Low Gain (Red)

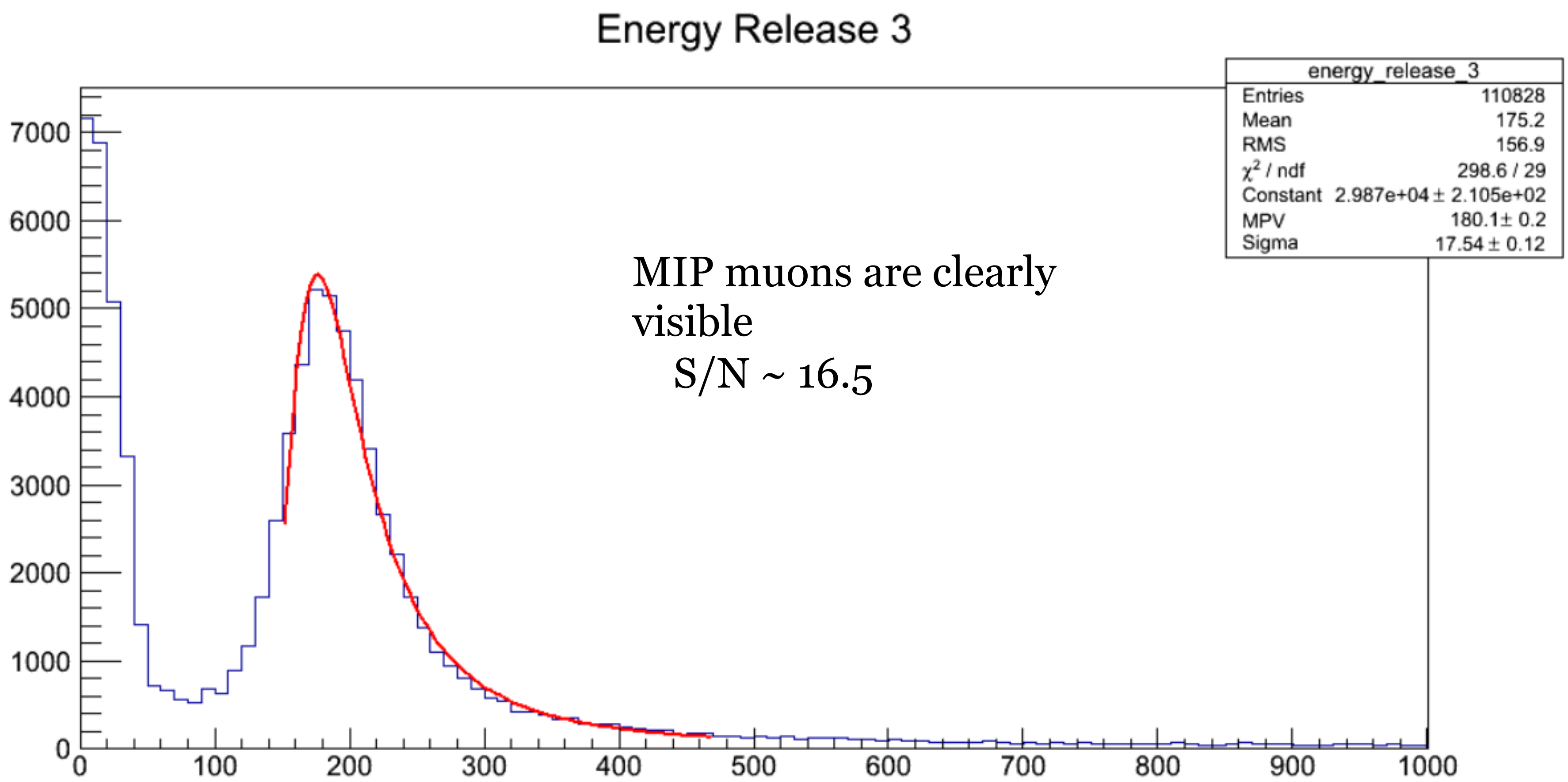


Response uniformity of the crystals



Pre-prototype test

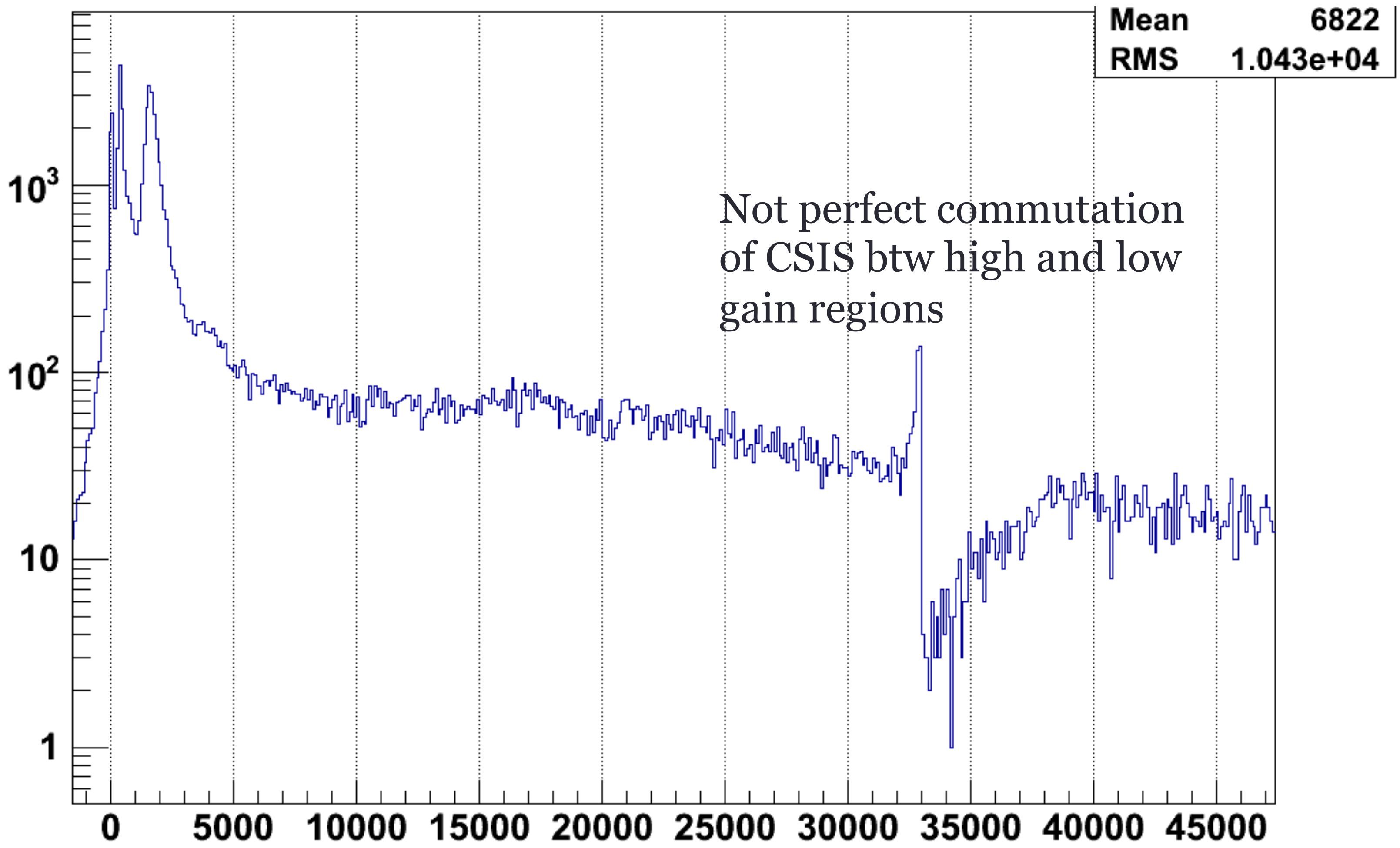
Muon beam



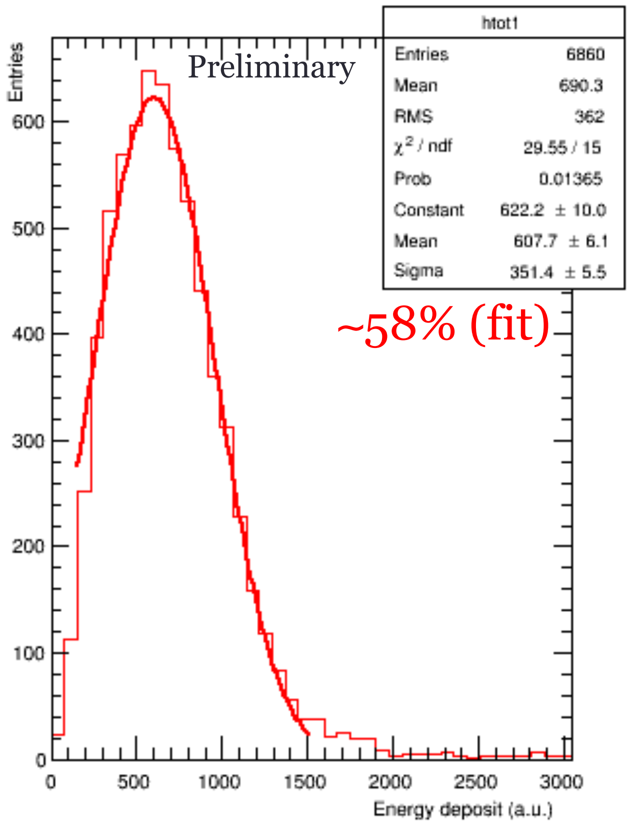
MIP muons are clearly visible

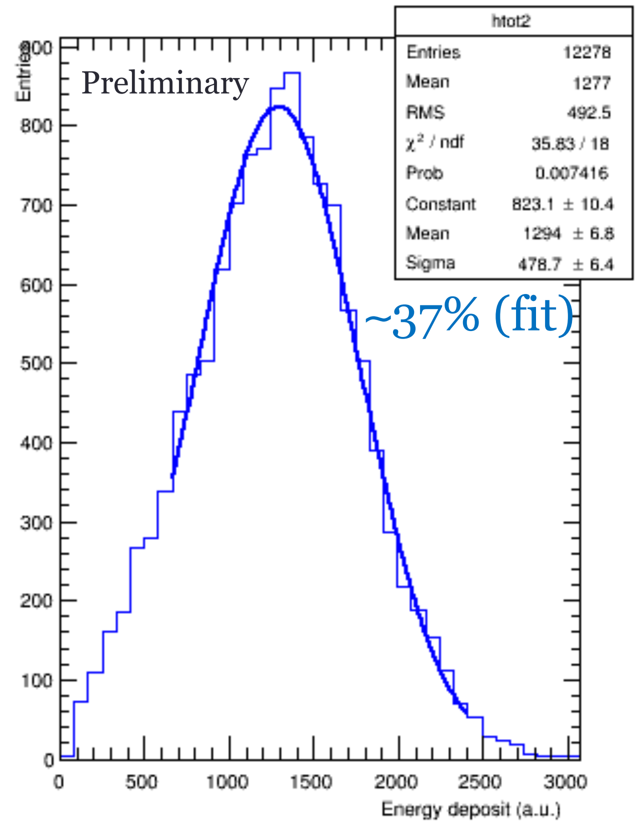
S/N ~ 16.5

Matching region btw high and low gain



Energy resolution (very rough)



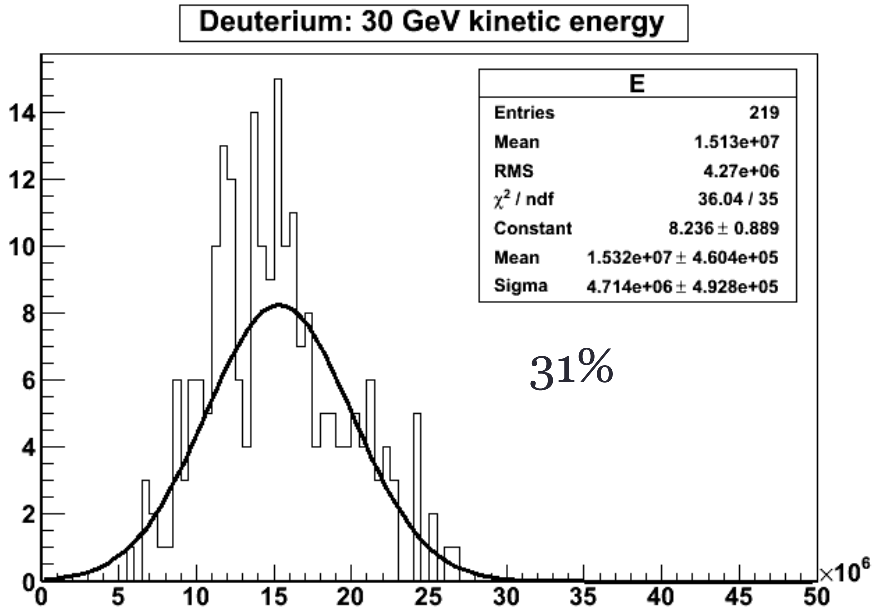


Z=2
Z=1

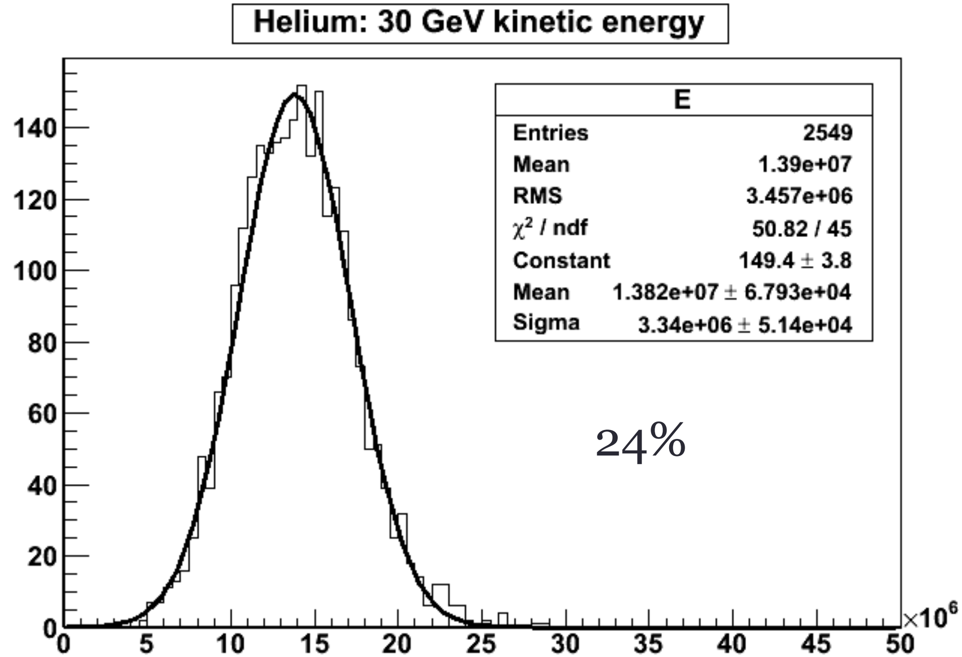
30 GV/c
Starting-layer ==2

No cuts and no corrections on the incident position

Expected resolution from simulation



Particle hitting the center of the crystals



Dual-readout calorimetry with a full-size BGO electromagnetic section

N. Akchurin [a], F. Bedeschi [b], A. Cardini [c], R. Carosi [b], G. Ciapetti [d], R. Ferrari [e], S. Franchino [f], M. Fraternali [f], G. Gaudio [e], J. Hauptman [g], M. Incagli [b], F. Lacava [d], L. La Rotonda [h], T. Libeiro [a], M. Livan [f], E. Meoni [h], D. Pinci [d], A. Policicchio [h,1], S. Popescu [a], F. Scuri [b], A. Sill [a], W. Vandelli [i], T. Venturelli [h], C. Voena [d], I. Volobouev [a], R. Wigmans [a,*]

Dual readout –> BGO: scintillation + Cherenkov

Filter: 250 ÷ 400 nm for Cherenkow light
>450 nm for Scintillator light

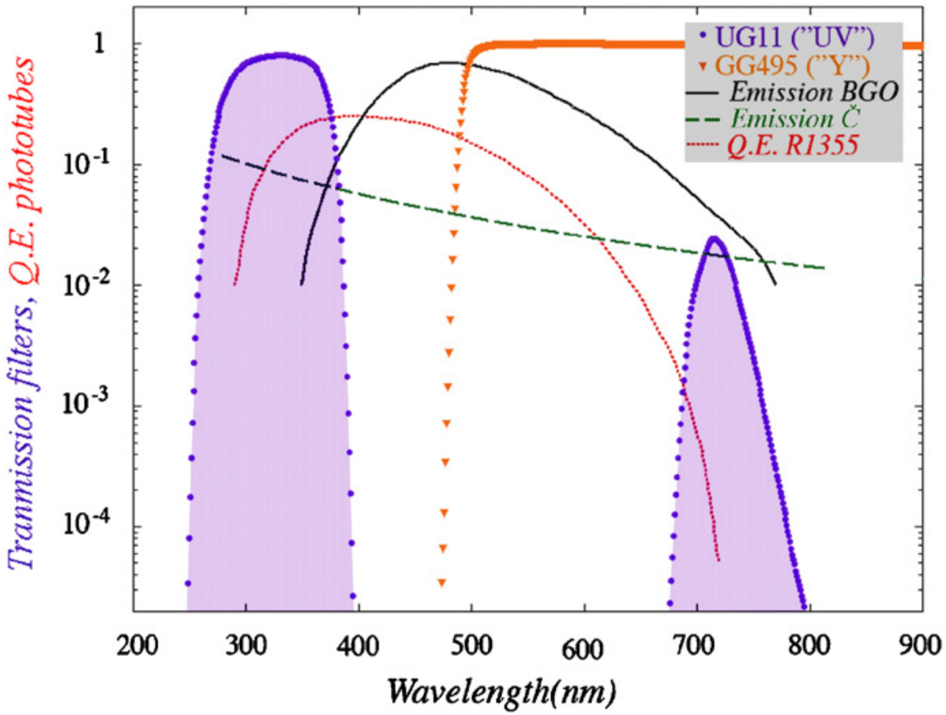
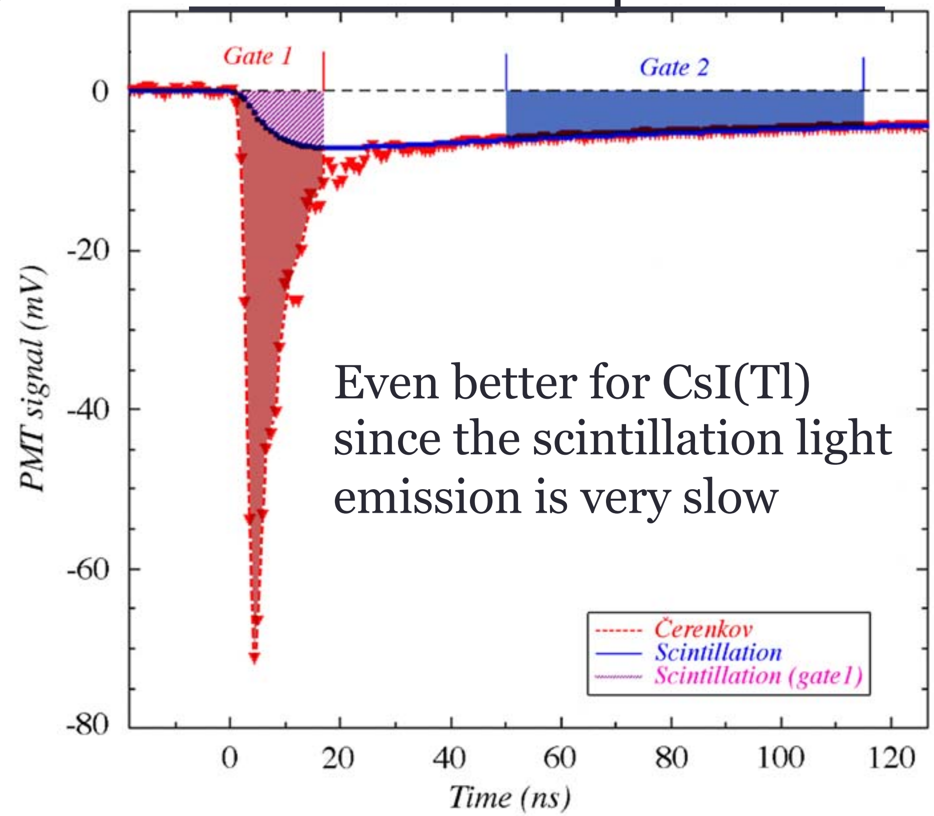


Fig. 14. Light transmission as a function of wavelength for the two filters used to read out the BGO crystal. The light emission spectrum of the crystal, the spectrum of the Cherenkov light generated in it and the quantum efficiency of the PMTs used to detect this light are shown as well. The vertical scale is absolute for the transmission coefficients and the quantum efficiency, and constitutes arbitrary units for the light spectra.

Hardware compensation



Even better for CsI(Tl) since the scintillation light emission is very slow

Fig. 5. The time structure of a typical shower signal measured in the BGO em calorimeter equipped with a UV filter. These signals were measured with a sampling oscilloscope, which took a sample every 0.8 ns. The UV BGO signals were used to measure the relative contributions of scintillation light (gate 2) and Cherenkov light (gate 1).