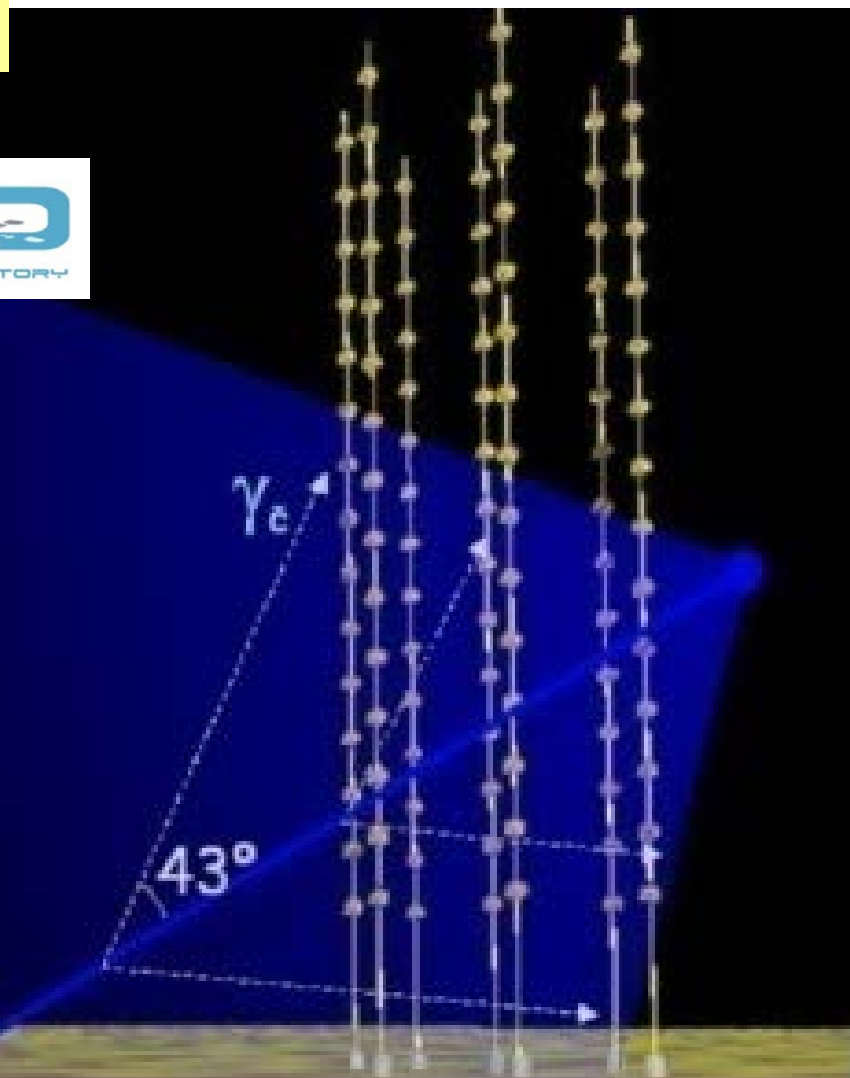




Primi risultati di
NEMO Phase-2 Tower

-> verso KM3NeT

Km3

“High Energy Neutrino Astronomy”

NEMO

KM3NeT-Italia

KM3NeT

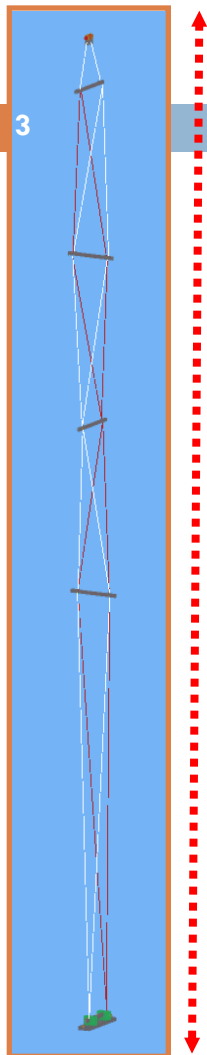
Breve resoconto delle attività nel 2013:

- NEMO: contributo alla messa in opera della Torre NEMO-Phase2 (INFN+MIUR)
- KM3NeT:
 - avviata la costruzione dell'apparato finanziato con fondi PON e progetto alla NEMO (prime 8 Torri, infrastruttura di fondo capace di ospitare fino a 32 Detection Units)
 - si contribuisce allo sviluppo del progetto KM3NeT “europeo” (moduli ottici “multi-PMT, stringhe)

Attività 2014: impegno del gruppo di Pisa, richieste finanziamento

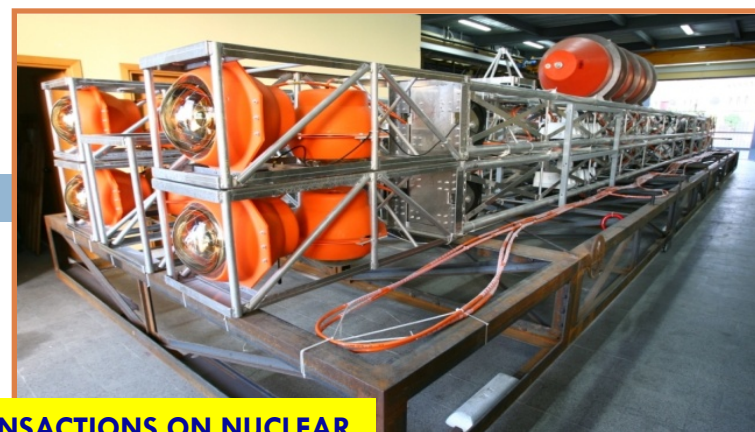
ANTARES

NEMO Towers, prototypes for KM3NeT

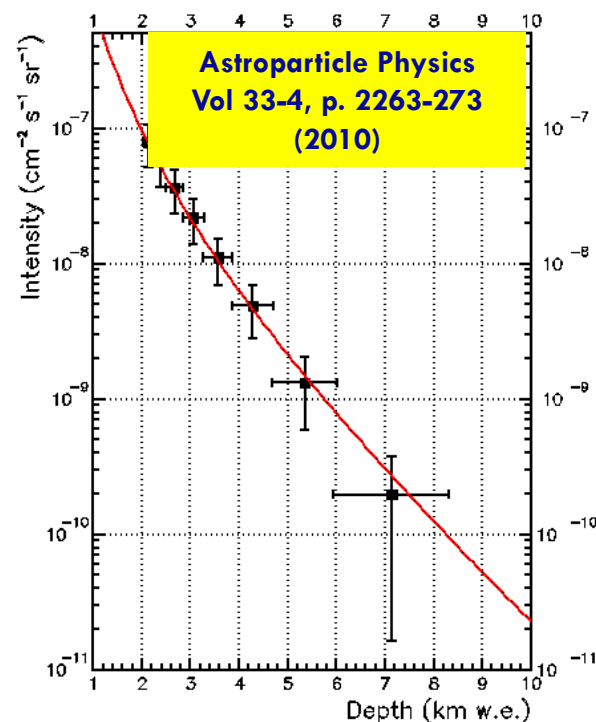


□ “Mini tower”, Dec. 06 at 2100m

- ▣ 4 floors @ 40 m
 - 4 PMs per floor
- ▣ 15 m long bars
- ▣ Deployment test (unfurling)
- ▣ NEMO electronics and readout (project and realization @ Roma)
 - Operated during 6 months
 - Data taking (muons)



IEEE TRANSACTIONS ON NUCLEAR SCIENCE, vol. 55; p. 233-240 (2008)



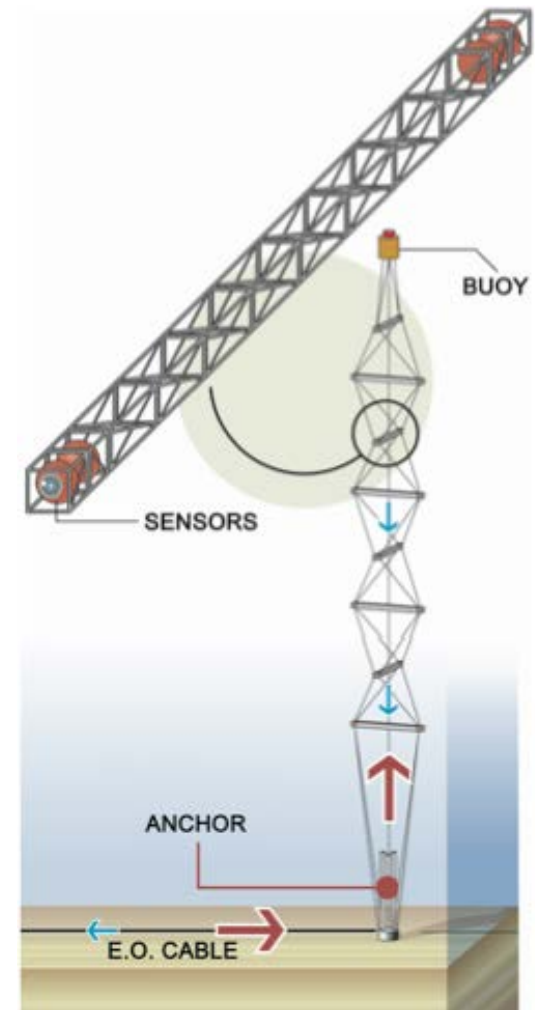
NEMO-Phase2 Tower deployment: short history

November 2012

-Teliri (MECMA) + INFN/INGV ROV: during a pre-inspection of the site, at 3500m depths, the ROV umbelical cable burned and the ROV was lost. Tower Deployment operation not even started.

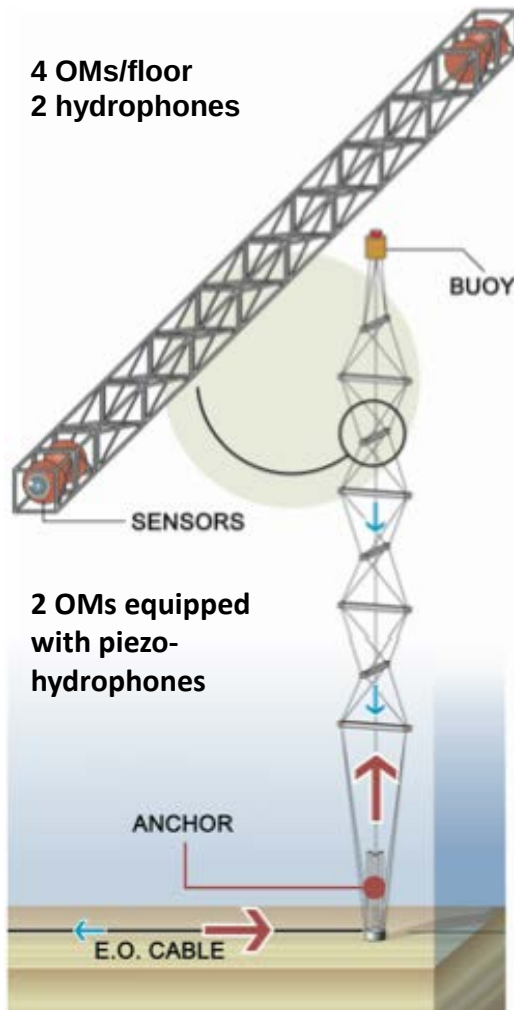
23 March 2013

-Campaign with FUGRO vessel (Nautical Tide) +ROV:
successful NEMO-Phase2 Tower deployment, connection and start of data taking



The NEMO Phase-2 tower: main components

5



- 8 floors
- 8 m bars, vertical dist. = 40 m, $H_{\text{tot}} = 450$ m
- 32 OM, 18 hydrophones
- oceanographic instrumentation

The OM: 10" Hamamatsu R7081, Front End Module, Time Calibration, LED beacons



Hydrophones: acoustic positioning and bioacoustics



Instrumentation on board the tower

6

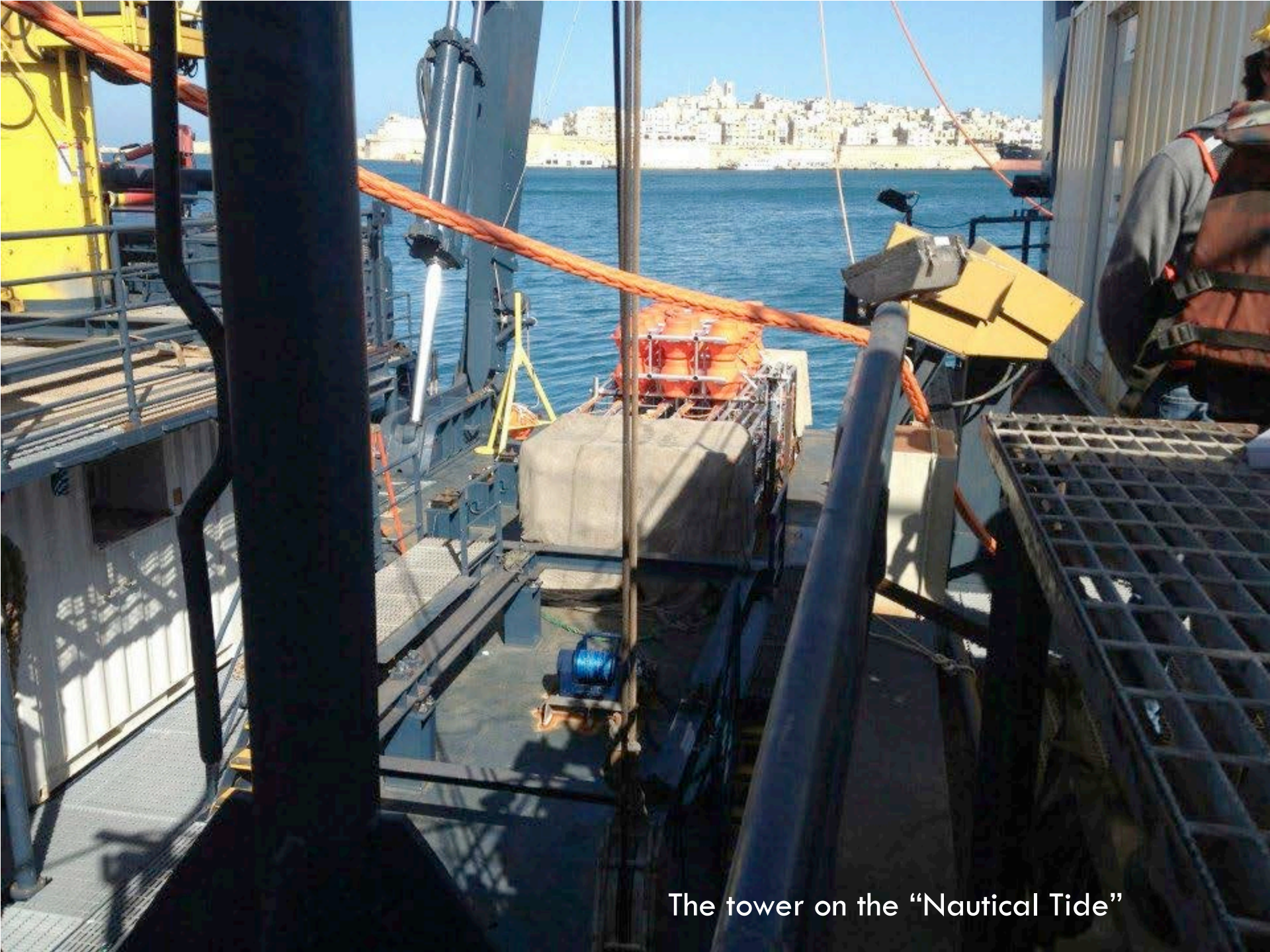
Floor	Acoustics	Calibration and environmental instrumentation
8	OAM (<i>ECAP, Erlangen</i>) [on OM 2 and 3 (downlooking)] OK	Porfido [on OM 1 and 4 (horizontal)]
7	FFR (<i>UPV, Valencia</i>) OK	Porfido [on OM 1 and 4 (horizontal)] Conductivity – Temperature – Depth (CTD) probe
6	OK	
5	Not integrated	Doppler Current Sensor (DCS)
4	OK	Cstar Nano-B 400 nm [on OM 2 and 3 (downlooking)]
3	OK	Nano-B 440 nm [on OM 2 and 3 (downlooking)]
2	OK	Nano-B 470 nm [on OM 2 and 3 (downlooking)]
1	OK	Conductivity – Temperature – Depth (CTD) probe Nano-B 470 nm [on OM 2 and 3 (downlooking)]
0	Monitoring station Not integrated	ACSA beacon Laser Beacon (<i>IFIC, Valencia</i>) Acoustic Beacon (<i>UPV, Valencia</i>)

Top ↑

↓ Bottom



NEMO-Phase2 under test in Malta before boarding on “Nautical Tide”



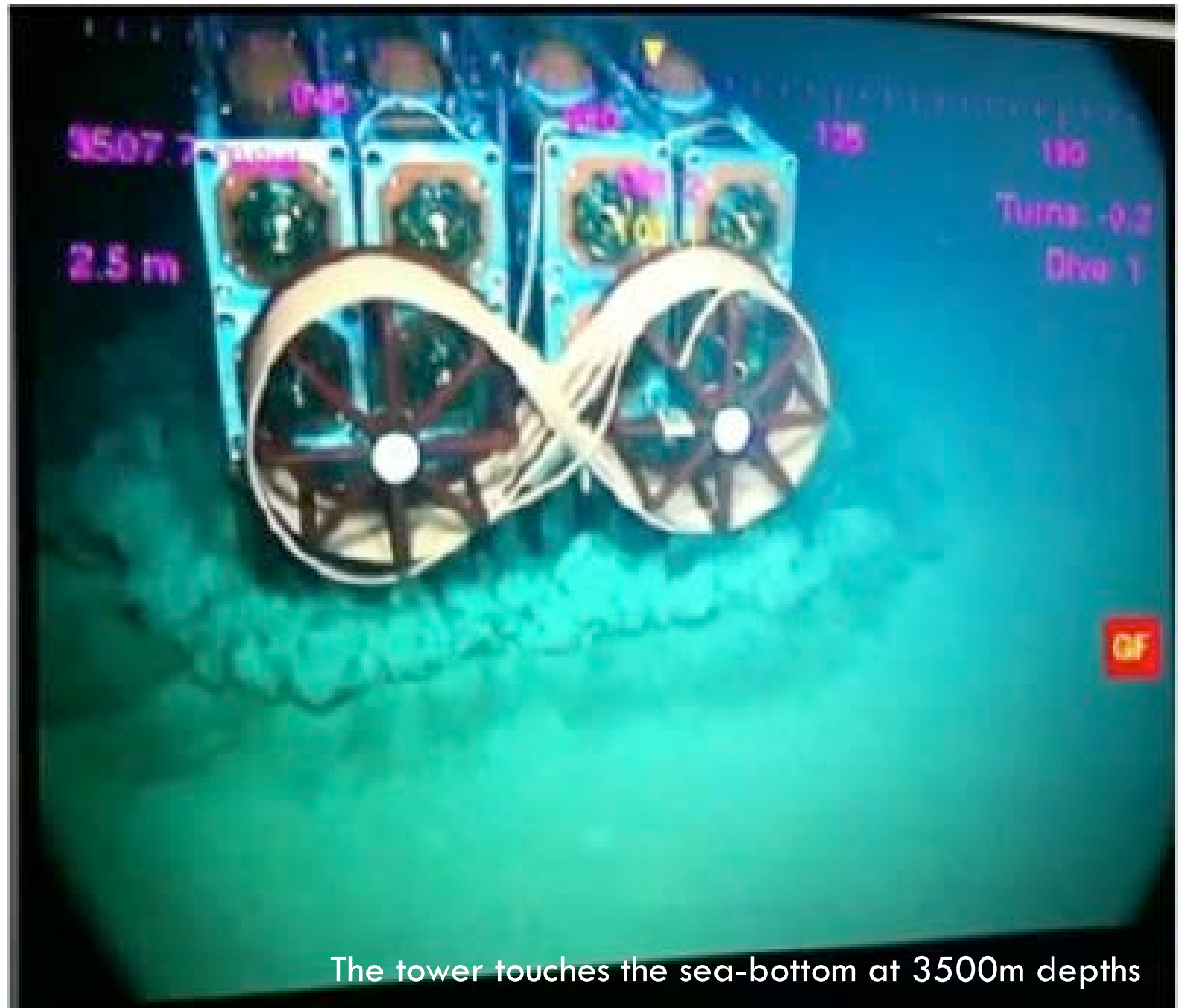
The tower on the “Nautical Tide”



The “Nautical Tide” ROV and
its launching system

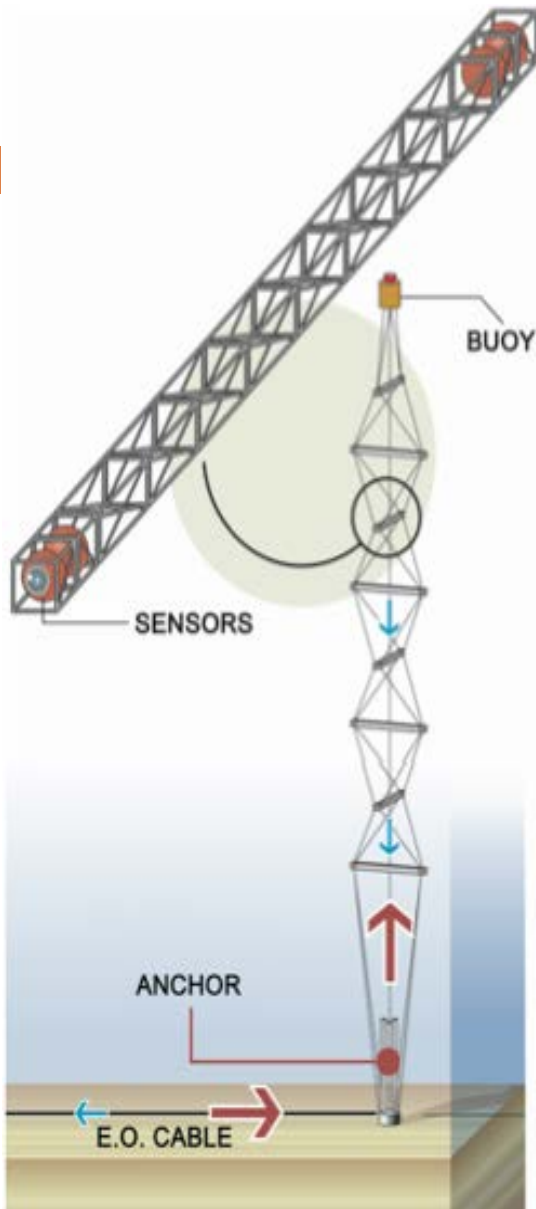
Deployment of the NEMO-Phase2 Tower from the “Nautical Tide”



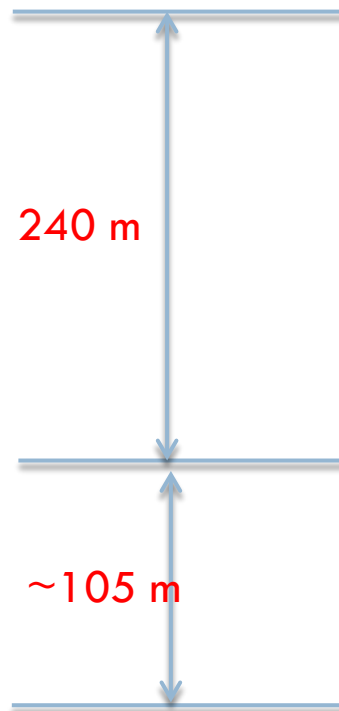


The tower touches the sea-bottom at 3500m depths





Mechanical
design

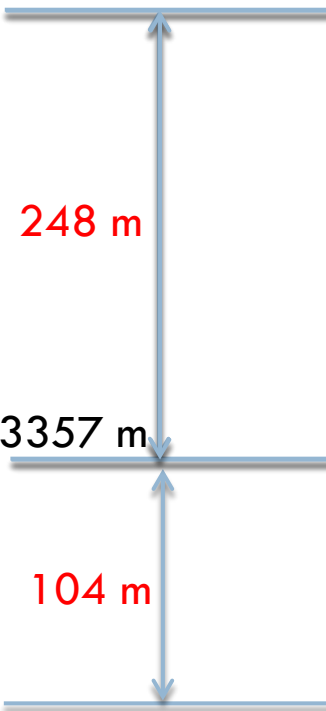


-3461.4 m
(FUGRO)

ROV movie

- 3109 m

- 3357 m



-3461.4 m
(FUGRO)

CTDs data

- 3107 m

- 3441 m

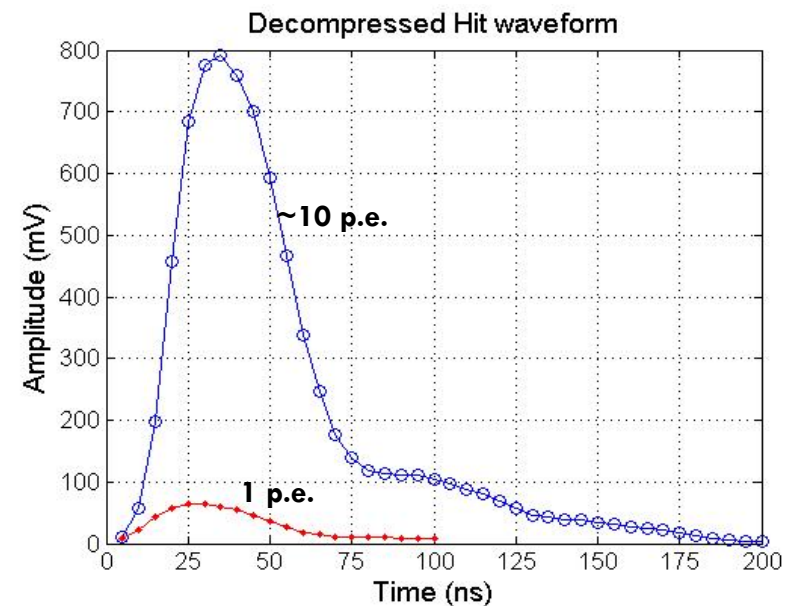
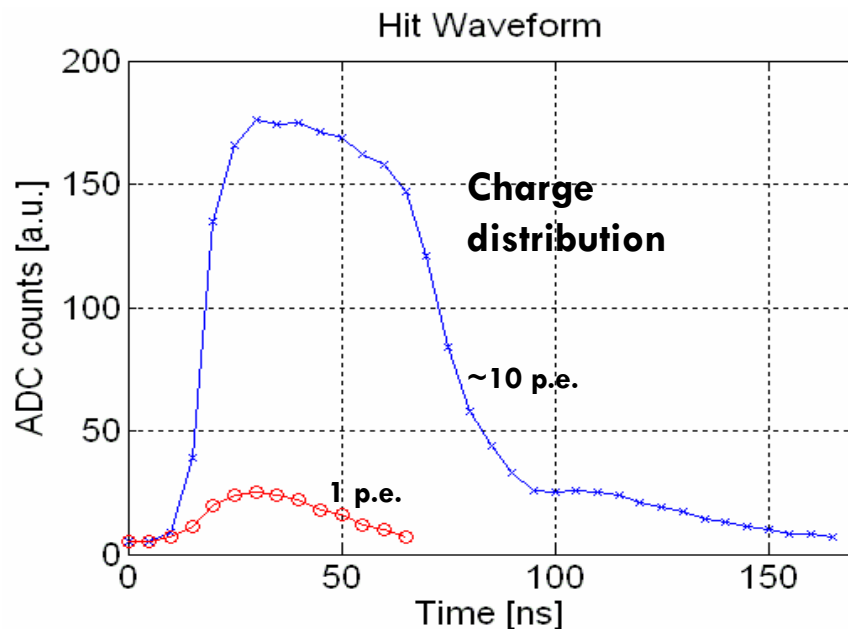


-3461.4 m
(FUGRO)

Signal digitization

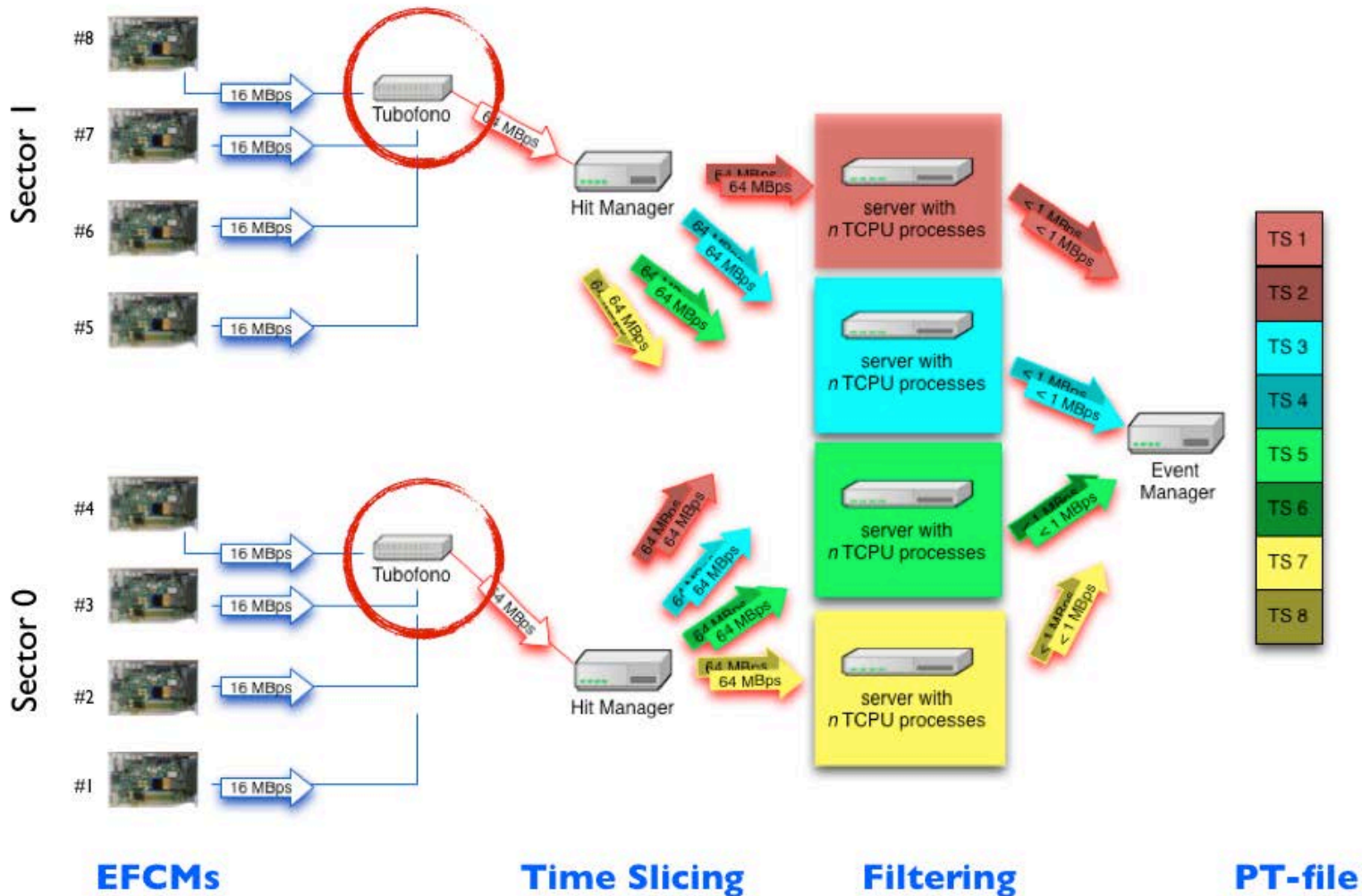
14

- PMT signal is compressed with a quasi-log law to increase the ADC input “dynamics”
 - ▣ Low voltage signals (up to few s.p.e.) sampled with high accuracy
 - ▣ High voltage signals sampled with lower accuracy



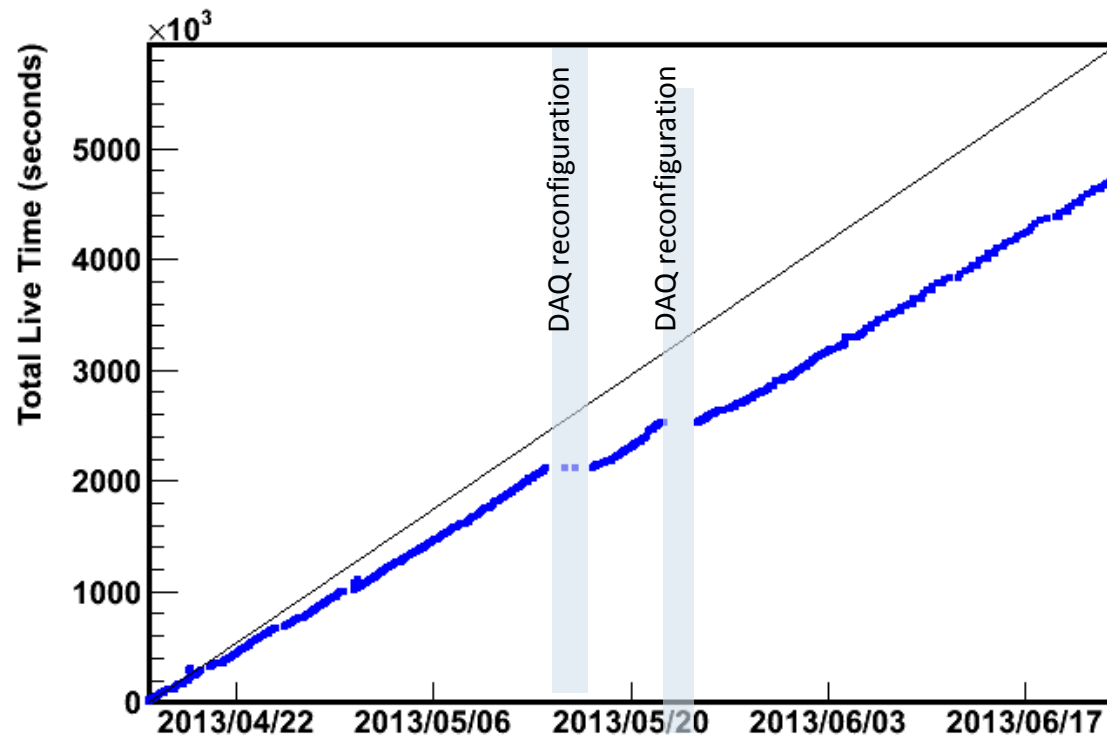
Trigger and DAQ system

15



Live time

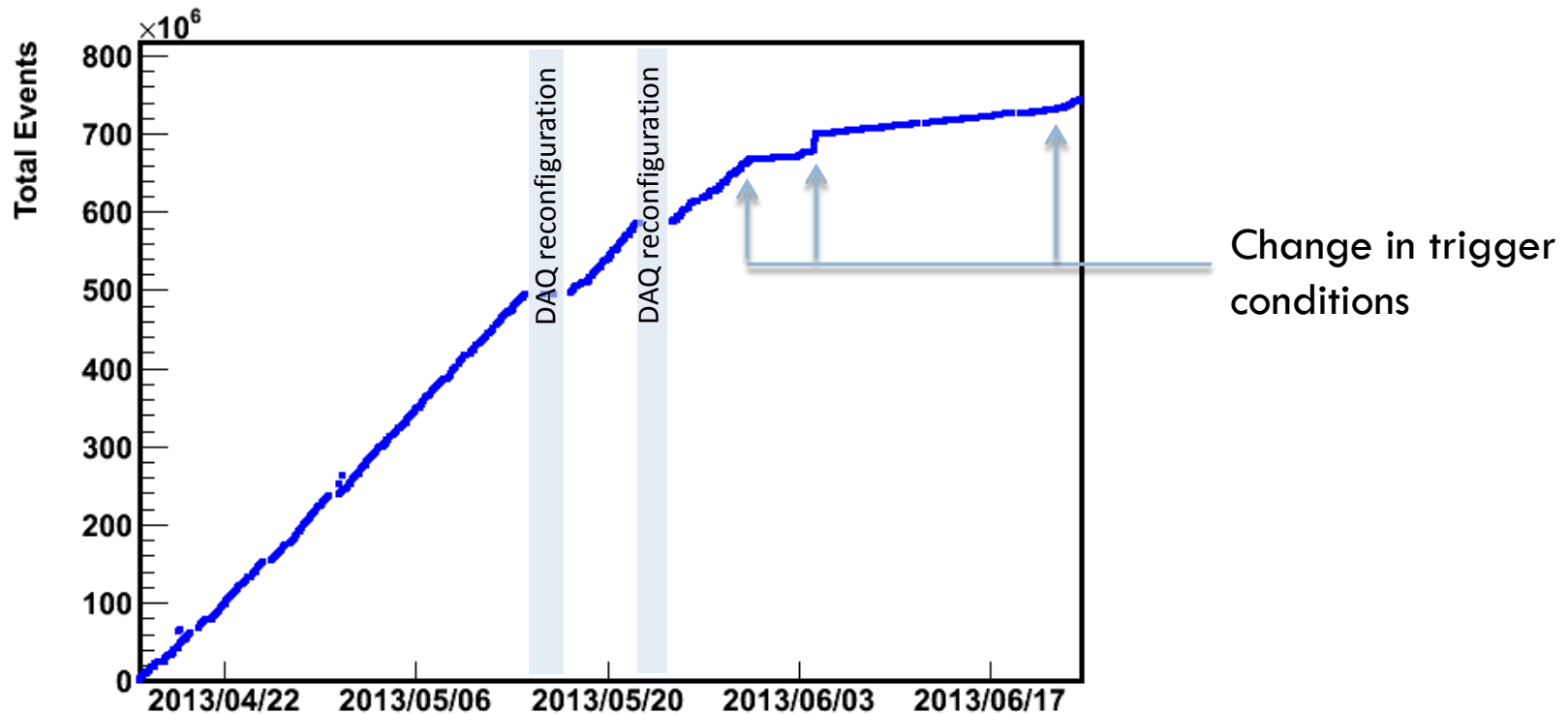
16



Live time since april 16 2013

Accumulated events

17

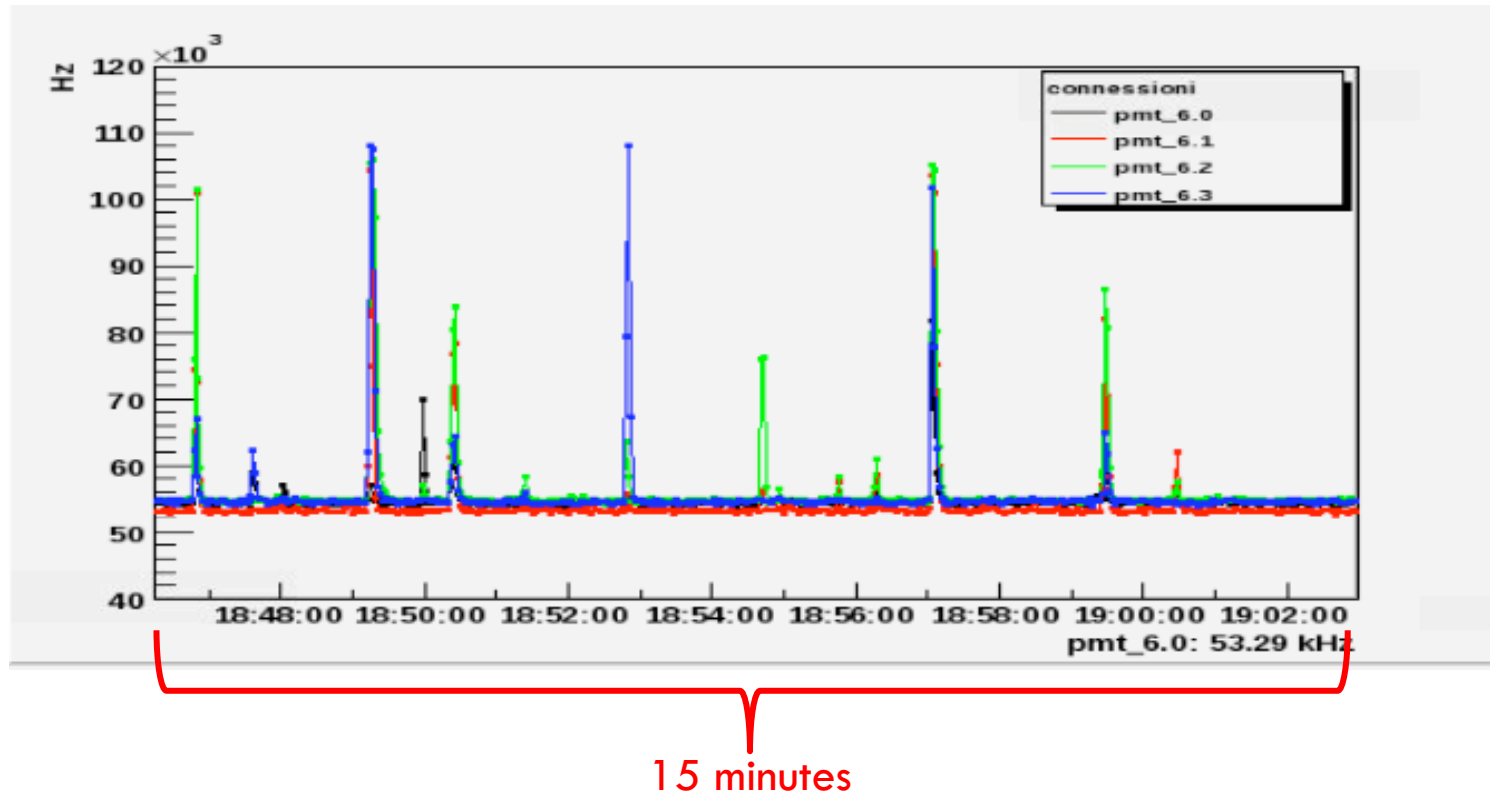


Total number of events accumulated since april 16 2013

Background rates

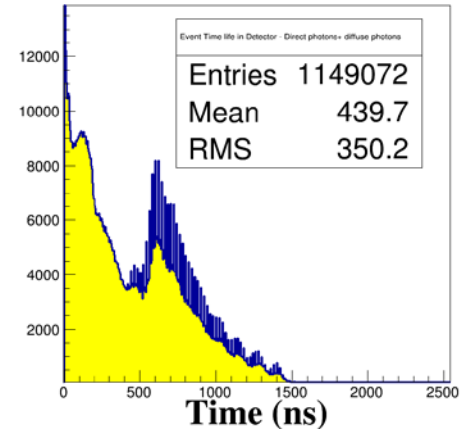
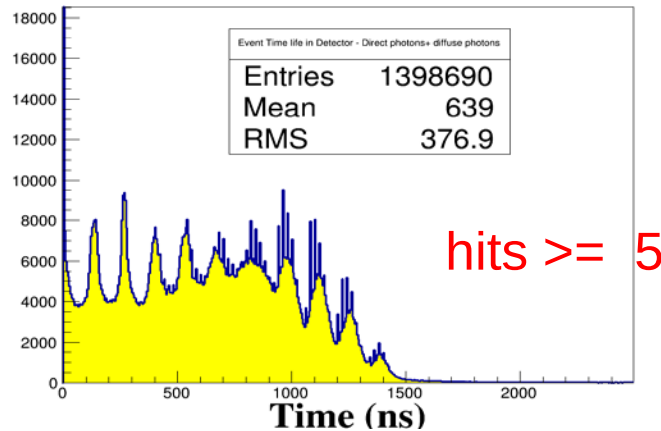
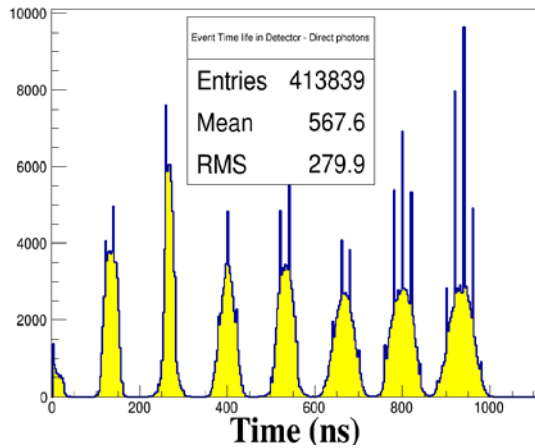
18

A typical snapshot of background rates from DAQ

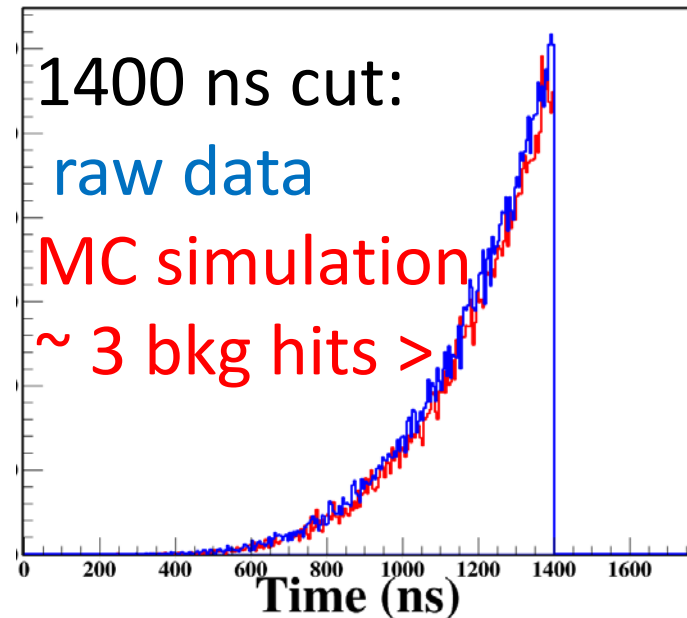
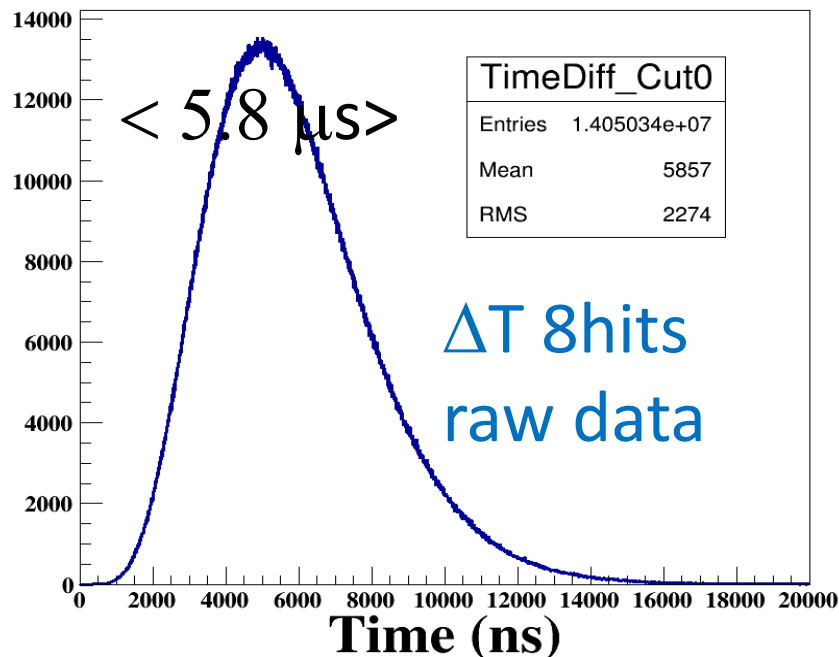


Dead time = 0 for rates $< 100\text{kHz}$ -- saturation effects above 120 kHz

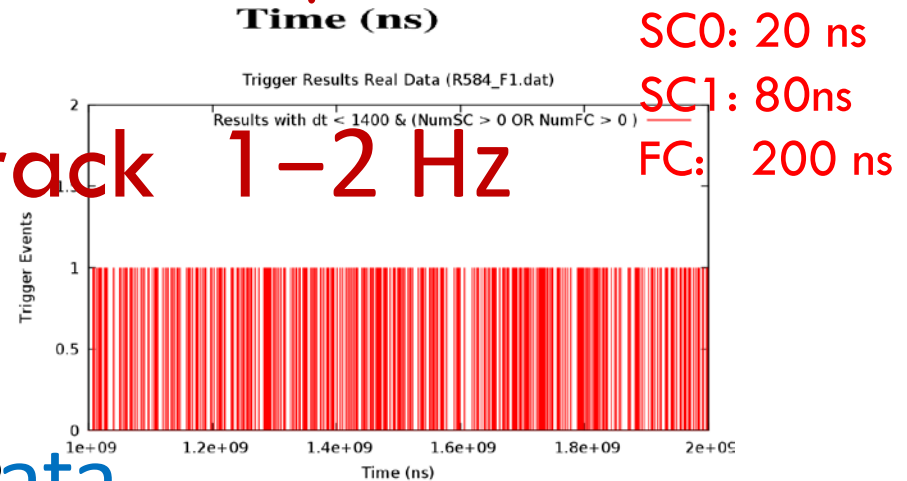
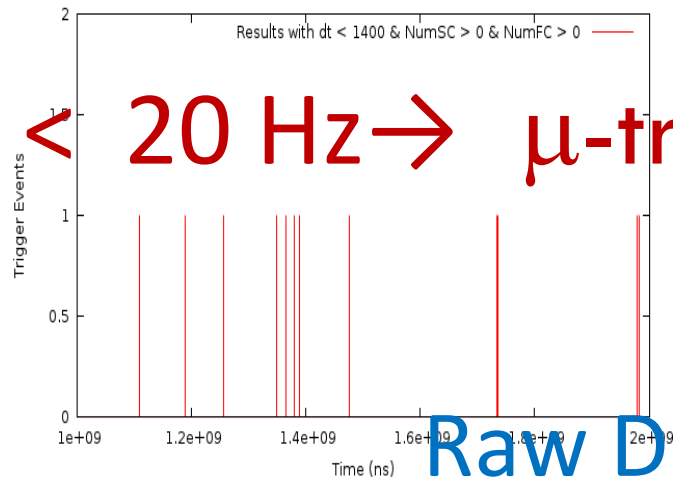
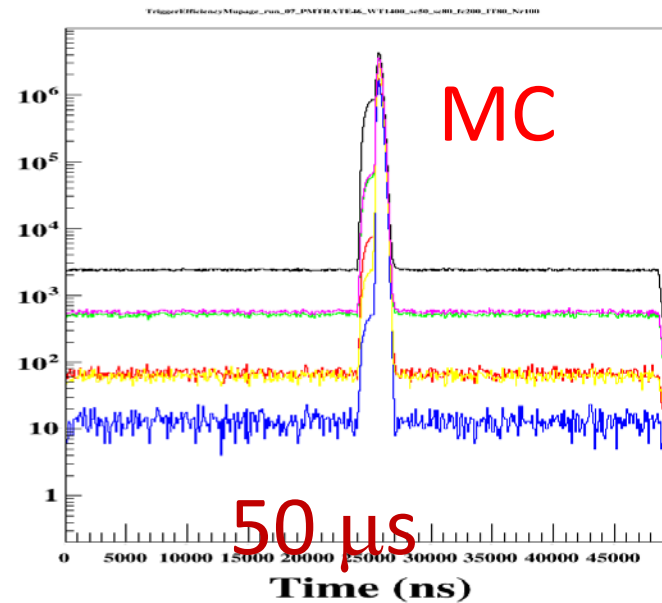
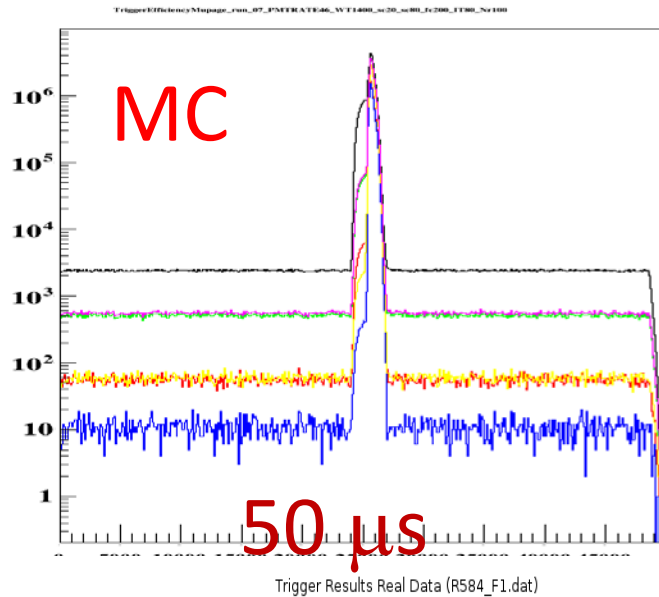
Studio di trigger -> GPU



8 floor: max $\Delta h = 280$ m.



Studio di trigger -> GPU



Raw Data

Studio di trigger -> GPU

<v> kHz	Time Win ns	<T> ns	<Nhit> ns	<ΔT> ns	Win kHz	Win+Sc kHz	Win+FC kHz	Win+SC+F C KHz
45.8	1400	>5450	2.9	1173	2.8	0.5	2.2	0.4
50.3	1400	>5040	3.1	1275	4.6	0.7	3.7	0.6
54.8	1400	>4670	3.2	1280	7.2	1.0	5.7	1.0
59.3	1400	>4340	3.4	1340	11	2.0	8.7	1.5
63.8	1400	>4040	3.6	1310	15.8	2.9	12.5	2.1

SC0: 50 ns

SC1: 80 ns

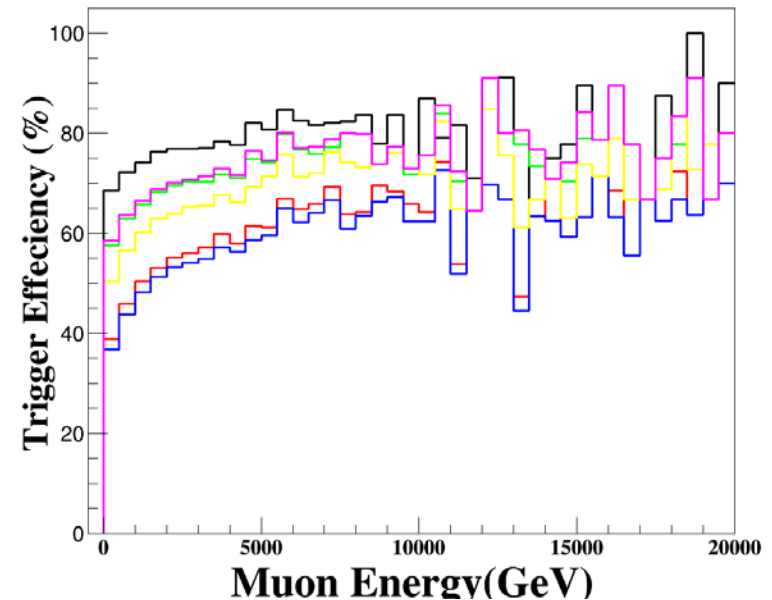
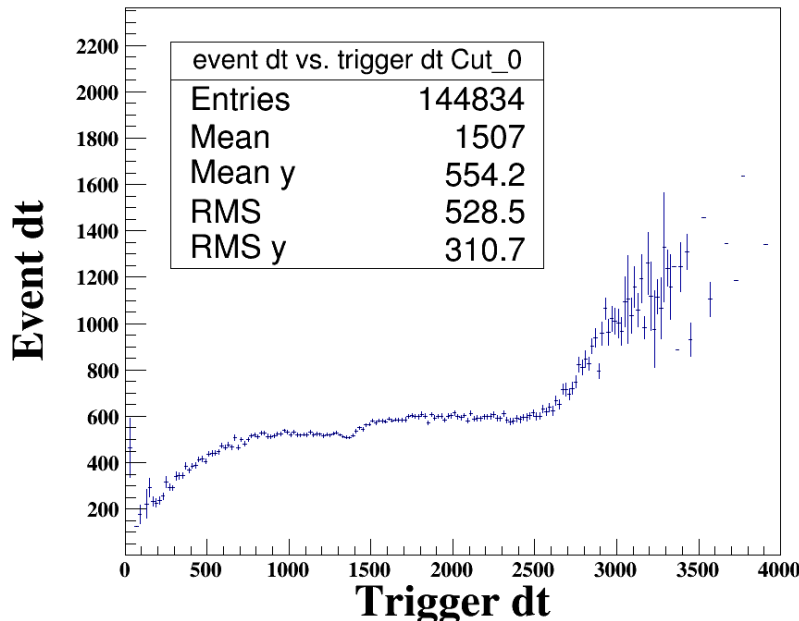
FC: 200 ns

45.8	1350		2.8	1213	2.2	0.4	1.8	0.3
50.3	1350		3.0	1225	3.8	0.7	3.1	0.5
54.8	1350		3.1	1231	6.0	1.2	4.8	0.9
59.3	1350		3.2	1245	7.4	1.5	6.1	1.1
63.8	1350		3.5	1257	13	2.6	11	1.9

Lo studio è completo e
proporremo di adottare
questo trigger.

Ultimo passo: tracking dei μ .
Implementazione su GPU .

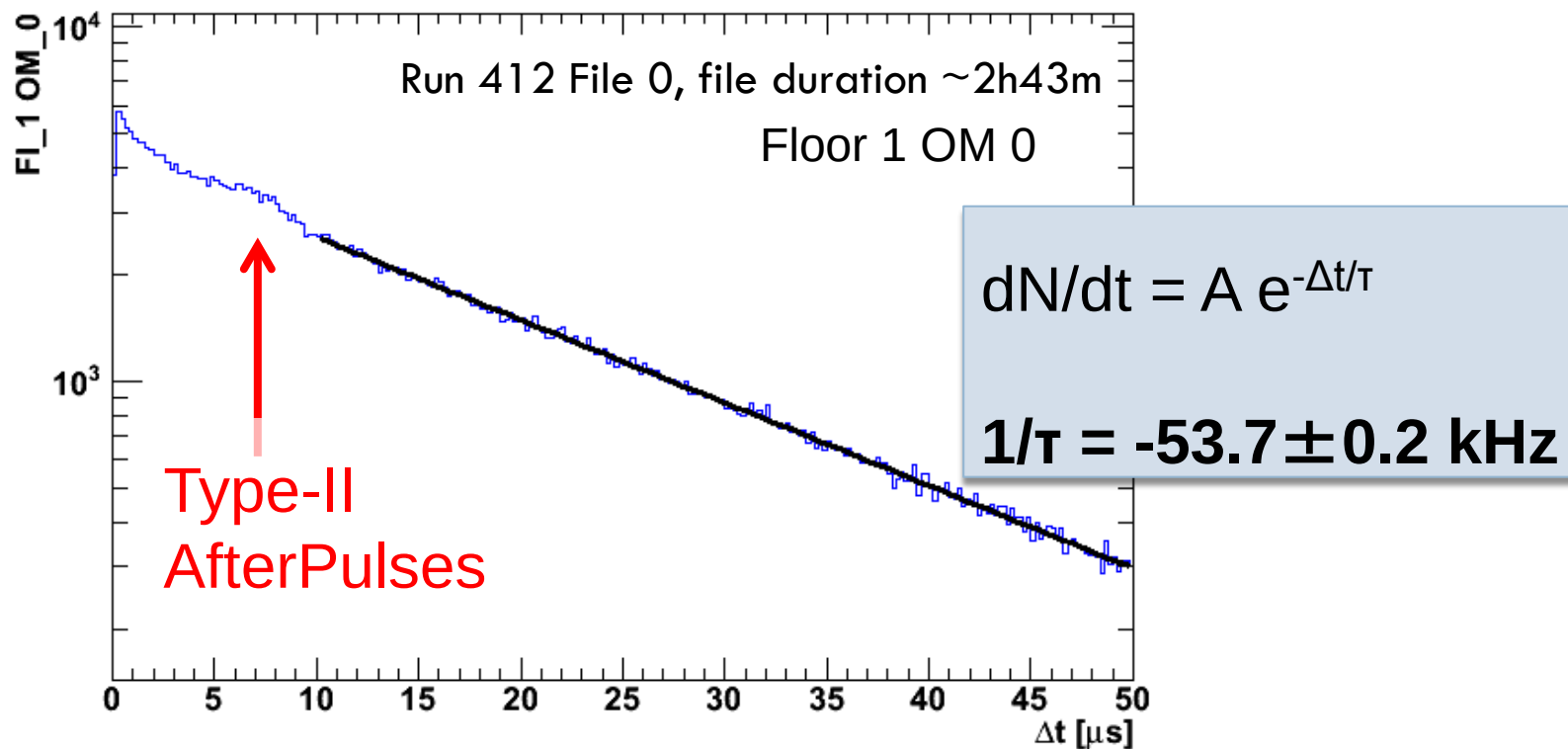
event dt vs. trigger dt



Singles rates

22

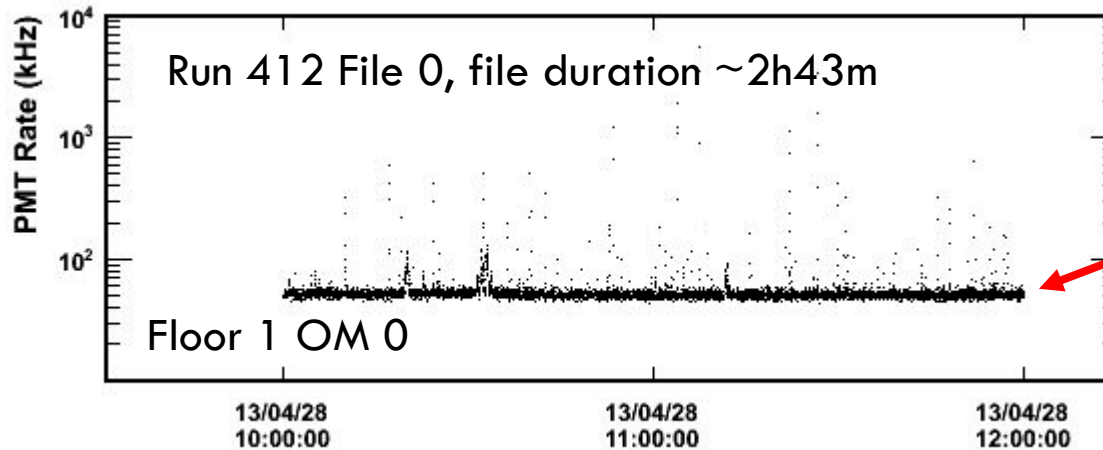
Distribution of time differences between consecutive hits



Similar behaviour and values in all OMs

Singles rates

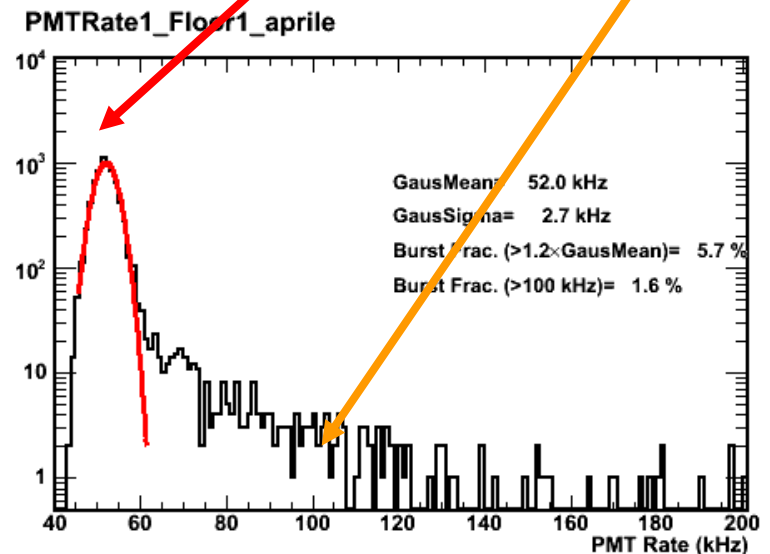
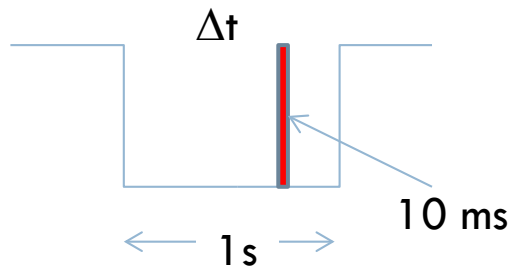
23



Baseline value = 52 kHz

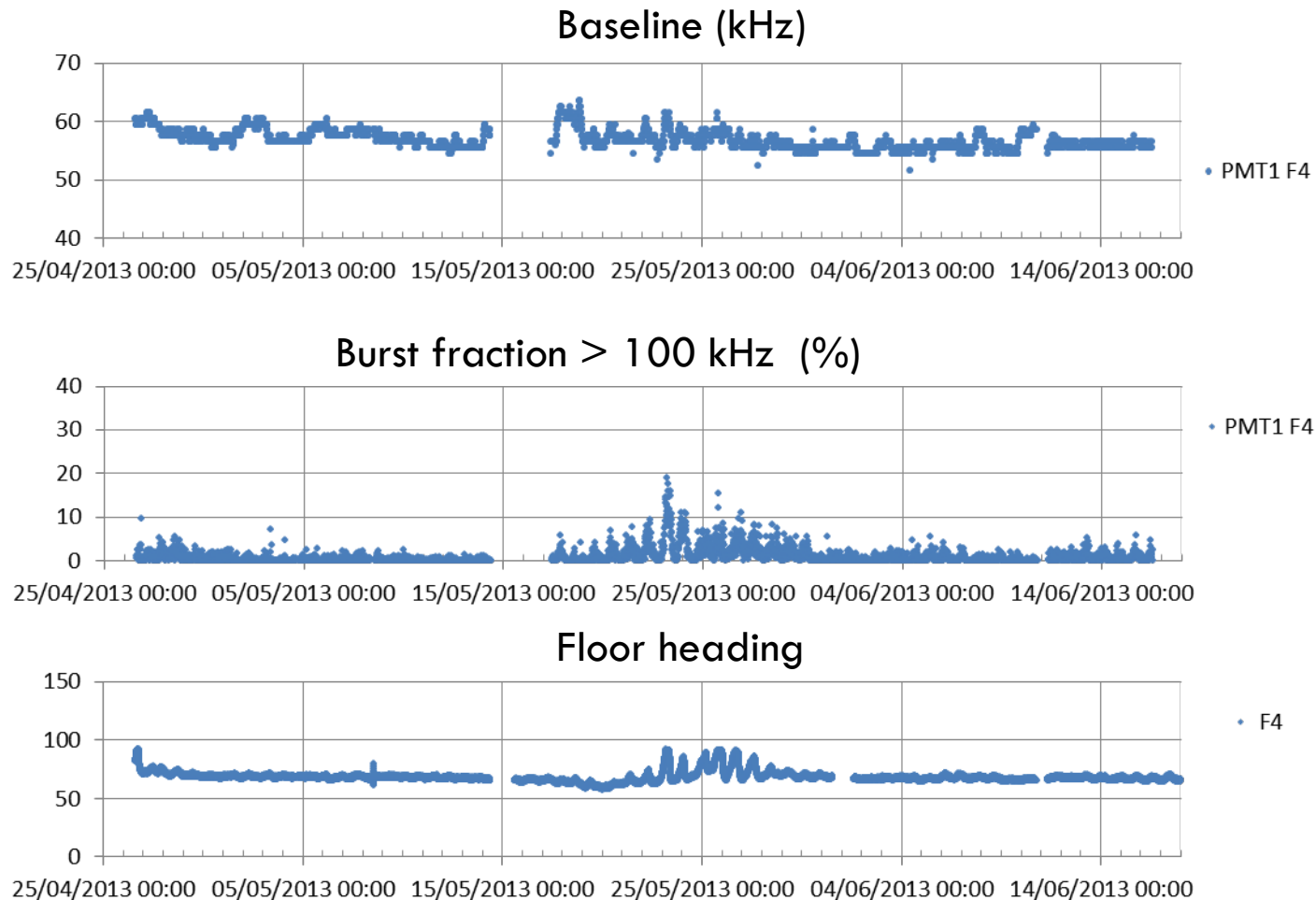
Burst >100 kHz = 1.6%

Data rates are sampled once per second
 by the front-end electronics
 Rate measured in a time window $\Delta t = 10$ ms
 No "dead time" effects



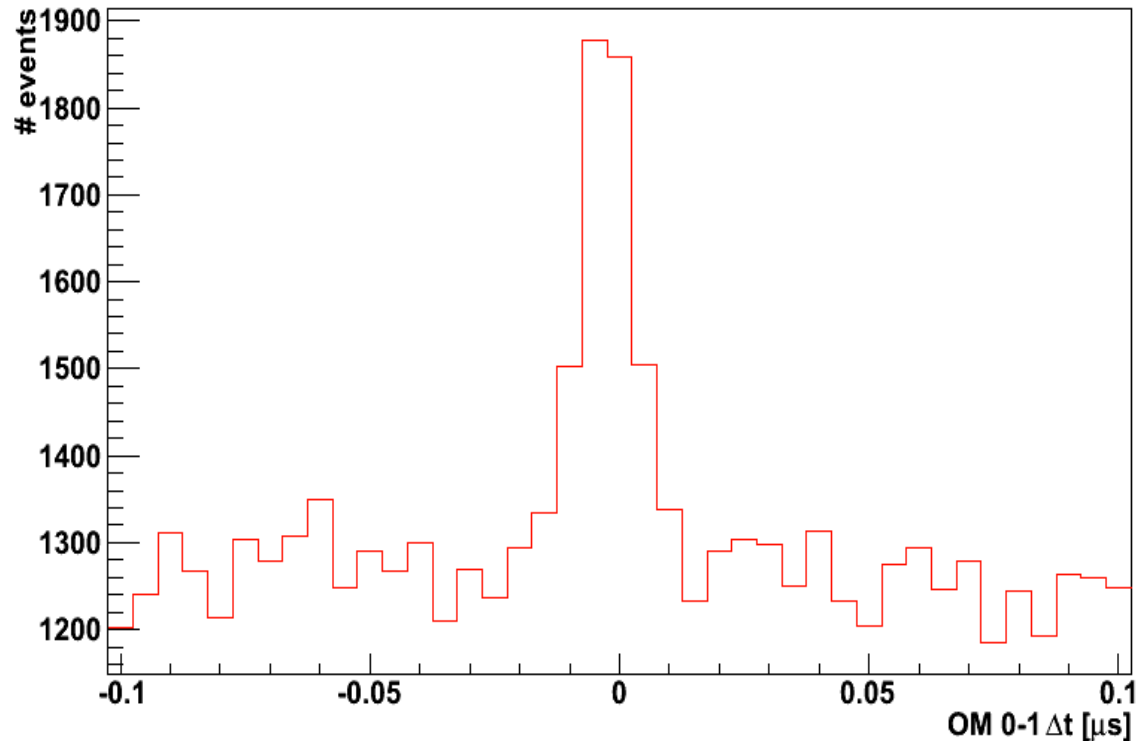
Long term behaviour of the background rates

24



Coincidences between close OMs

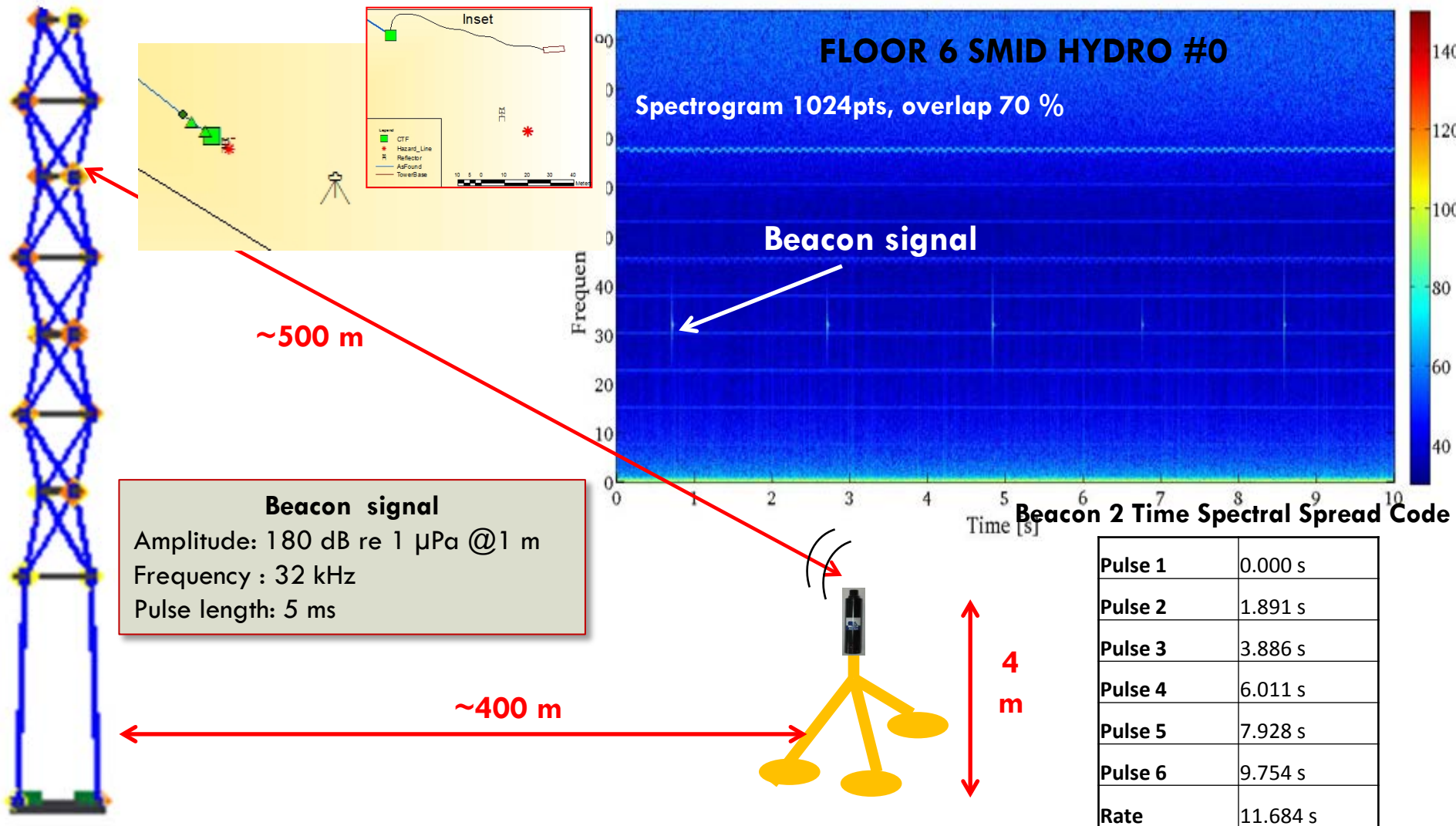
25



Run 489 (files 0-4)+
run 491 (files 0-3) + run
499 (files 0-1) +
= total livetime ~ 28 h.

Acoustic detection: status

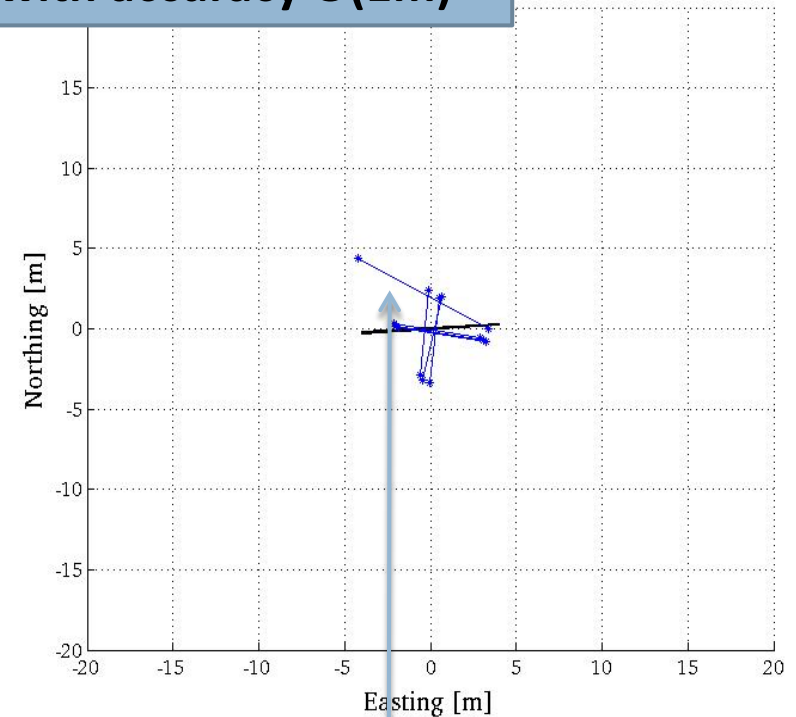
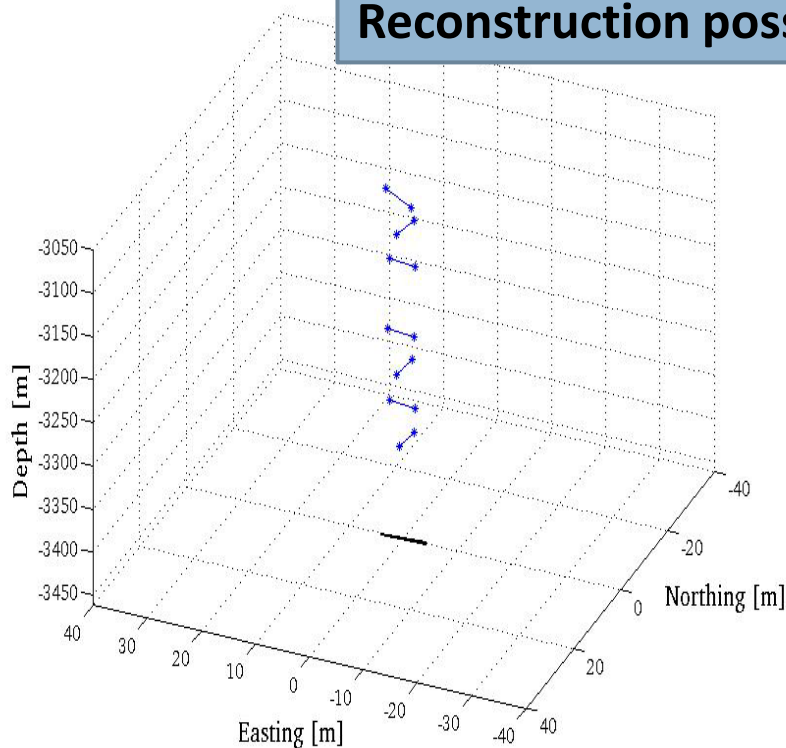
26



Reconstruction with acoustics pos.

27

Reconstruction possible with accuracy $O(1m)$



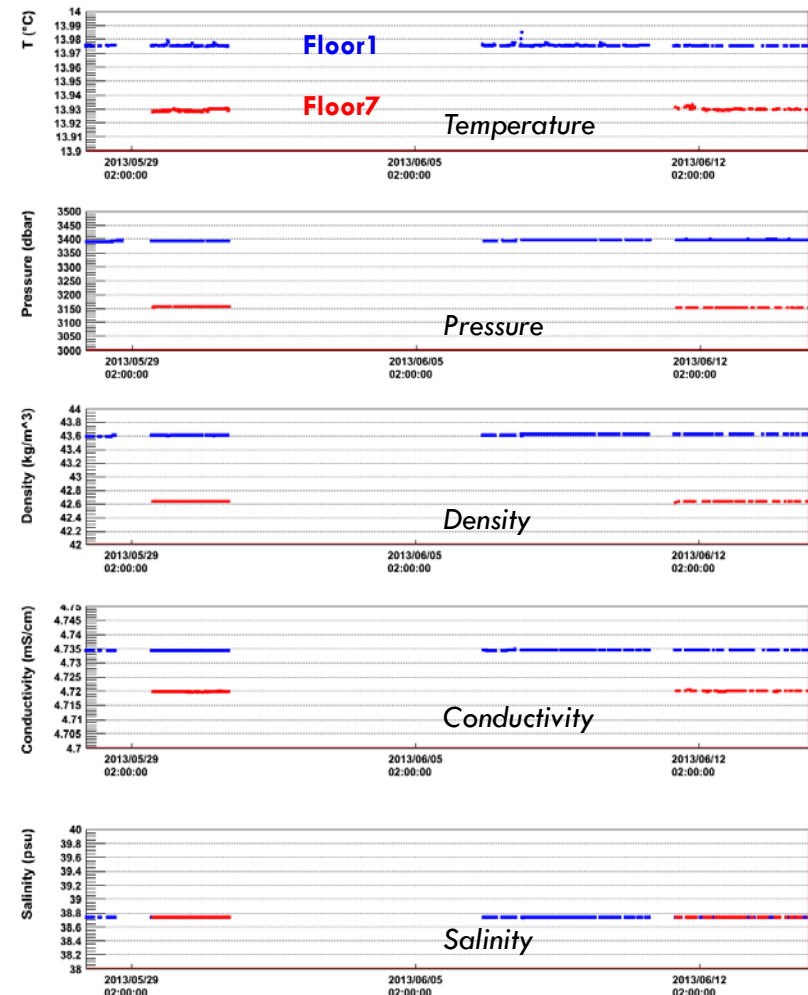
Not longer bar! Just different
hydropone position (piezo).

07/05/2013
00:00 UTC

CTD probes

28

- Two CTD (Conductivity, Temperature, Depth) probes (Floor 1 and Floor 7)
- Sampling every 10 m
- Both working, but some inconsistencies on the measured pressure value of the CTD on floor 1 (calibration problem?)



Summary

29

- 8 storey tower deployed on March 23 2013
- Operational since then
- Still in commissioning phase
- First results
 - ▣ Very low background rates (compatible with 40K background with few bioluminescence bursts)
 - ▣ Acoustic positioning limited by the incompleteness of the LBL, but accuracy $O(1\text{m})$ reachable
- Next steps
 - ▣ Fully understand some inconsistency in the instrumentation (CTD and currentmetre) data
 - ▣ Muon track reconstruction

KM3NeT in Italia

30

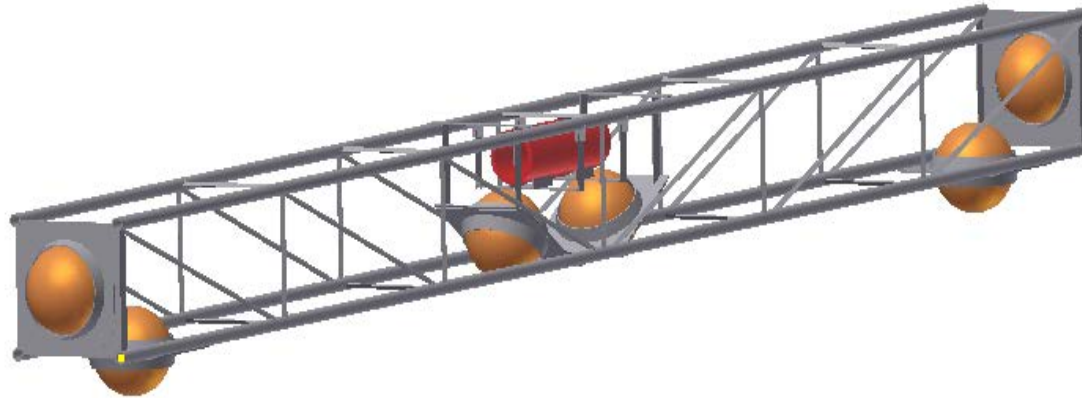
- Progetto approvato con 110 punti su 120 (quarto in graduatoria)
- Finanziato per € 20.800.000 (fondi da utilizzare entro Dicembre 2014)
- Progetto rimodulato in funzione del finanziamento ricevuto
 - Costruzione di un primo modulo di torri (28) → 8
 - Engineering Array / primo modulo del nodo italiano di KM3NeT
 - Da potenziare successivamente (fino ad almeno 100 DU)

Km3 – KM3NeT-Italia nel 2013-2014

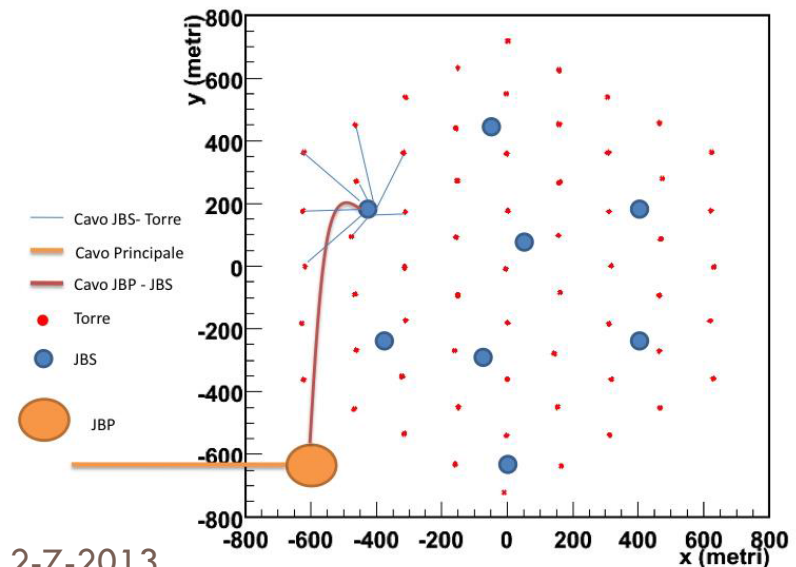
- Nell'ambito del consorzio europeo KM3NeT partecipiamo alla validazione di DU basate su OM multi-PMT e stringhe.
- **Realizzazione di 8 Torri con tecnologia basata su “NEMO-Fase2”**
- **Realizzazione di una infrastruttura sottomarina capace di ospitare Torri e/o Stringhe (fino a 32 DUs)**
- **Con l'obiettivo di ospitare a Capo Passero il massimo numero di DUs siamo pronti a completare il detector con “stringhe” se:**
 - **stringa+OM multi-PMT sarà stata validata entro una data ragionevole per poter passare alla nuova produzione industriale**
 - **I partner europei (NIKHEF, ...) contribuiranno con un numero di stringhe almeno uguale a quello realizzato da km3 IT.**

NEMO-Fase2 → Km3 - meccanica

- Progetto rivisitato: 6 PMT/piano (8 torri di 14 piani)



- Distanza fra piani: 20m
- N. di idrofoni/piano: 2
- Layout: schematicamente indicato nei vari documenti di KM3NeT e nel proposal PON
- Distanza fra le Torri: la minima compatibile con l'esigenza di sicurezza nel deployment (100m ?)



KM3NeT-Italia:

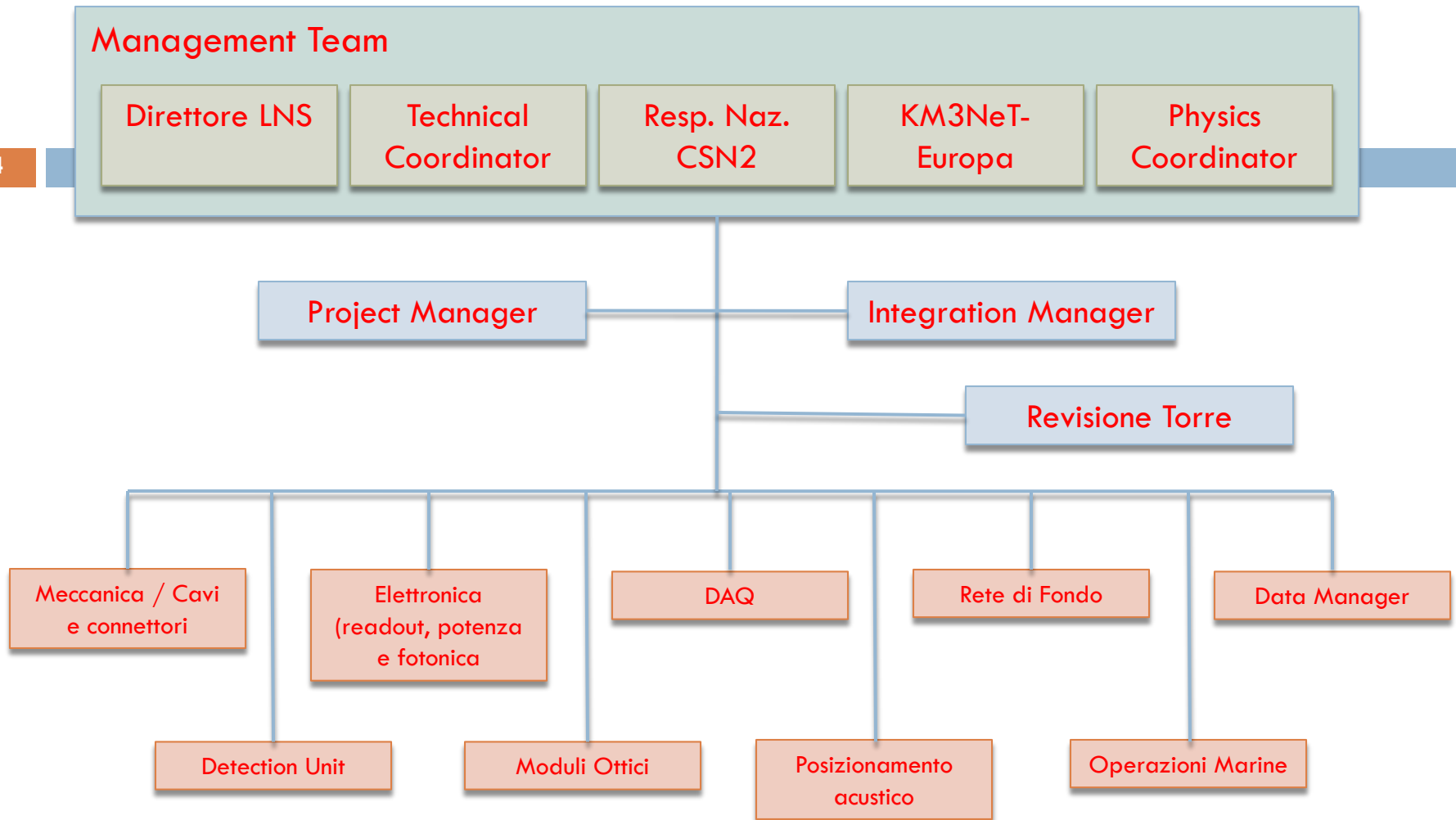
Activity in 2013

All elements of NEMO-Phase2 Towers reviewed and scrutinized in order to prepare the industrialization of the detector construction and assembly:

NEMO-Phase2 → KM3NeT-Italia

Preparing the construction tools and the test-benches for the production control.

Launched/ing tenders for the construction of detector elements and for the Towers integration.



Sezioni Coinvolte: LNS, Roma1, Na, Ct, Bo, Ge, Pi, Ba, LNF

Percentuali 2012:

Nome	Cognome	posizione	Km3 14
Nicolò	Beverini	PA	50
Bachir	Bouhadeb	Ass. INFN	100
Massimo	Calamai	Assegno Ricerca	40
Enzo	Flaminio	PO in pensione	?
Mauro	Morganti	PS Accad.	100
Enrico	Maccioni	Tecnico Univ.	50
Francesco	Francesconi	Tecnico Univ.	40
Fabrizio	Raffaelli	INFN	
Giuseppe	Terreni	INFN	
			380

NEMO Pisa: preventivi 2014									
	M	I	M	E	Con s	Ap p.	Inv	T.	Tot.
Meeting di collaborazione: 2CM*3persone	7,2								
Istitution Board + missioni istituzionali 5x1	4,0								
Partecipazione a raccolta dati C.P. 2X2=4 settimane	6,8								
Montaggio struttura di oscuramento 1x3	4,0				2,0				
Partecipazione a gruppi di lavoro per analisi dati	4,0								
Partecipazione a integrazione delle Torri 5x1	5,0				4,0				
Partecipazione a gruppi di lavoro Km3-Europa: 3x1				4,5					
Meeting di collaborazione Km3-Europa: 3x1				4,5					
Altro*					0,0				
Trasporti e Assicurazioni (struttura Oscuramento ...)								1,5	
	31,								
Totale	0	9,0	6,0	0,0	0,0	0,0	1,5	47,5	

* da concordare con la collaborazione prevista la prossima settimana

La costruzione delle Torri è tutta finanziata dal budget PON.
Non sono finanziate le missioni nè le attrezzature necessarie per la realizzazione di test-benches (consumo + impianti) e trasporti/assicurazioni:

Risorse di Sezione 2012: Nemo + Antares

- Progettazione mecc. ~ 2mu

Contributo alla progettazione all'ancora della linea PPM-DOM

- Officina mecc. ~ 2 mu

Montaggio struttura di oscuramento

Supporto officina per montaggio torri

- Elettronica ~ 0

- Calcolo G.Terreni che svolto un lavoro fondamentale seguendo tutte le fasi del progetto del DAQ a terra: da eFCM a Data Storage .

Se la Sezione mette a disposizione le risorse
possibilità di avere una linea di test.

Attività 2014

NEMO

- Contributo al commissioning del TRIDAS
- Sviluppo del trigger per la selezione degli eventi
- Test del trigger con GPU
- Raccolta dati a CP
- Analisi dei dati

KM3 Italia

- Contributo alla fase di Test, integrazione e qualificata degli 8 piani di torre
- Installazione copertura oscuramento

KM3 Europa

- Meccanica dell'ancora e test del prototipo + Trigger + altro ?????

KM3NeT in Europa

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Consorzio internazionale che coinvolge più di 300 ricercatori tra una decina di stati europei

- Obiettivo: costruire una grande infrastruttura sottomarina il cui principale scopo è la costruzione del più sensibile telescopio per neutrini di alta energia, complementare a IceCube

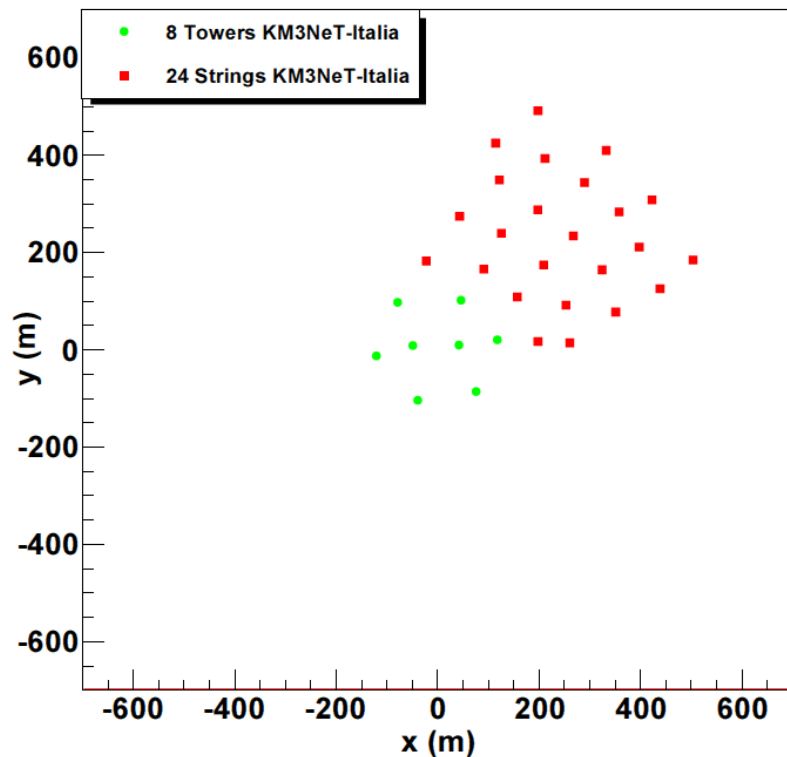
- 'Research Infrastructure' inserita nella roadmap ESFRI EU



Layout

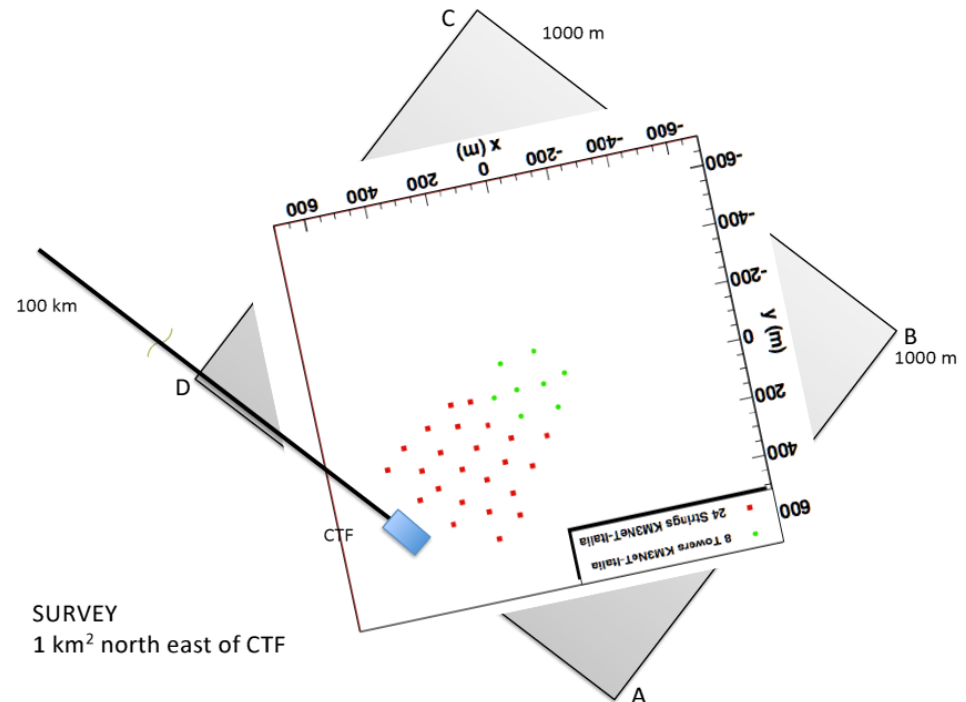
40

footprint KM3NeT CapoPassero 2



R. Coniglione mail
Approved by Km3Net

CAPO PASSERO SITE SKETCH

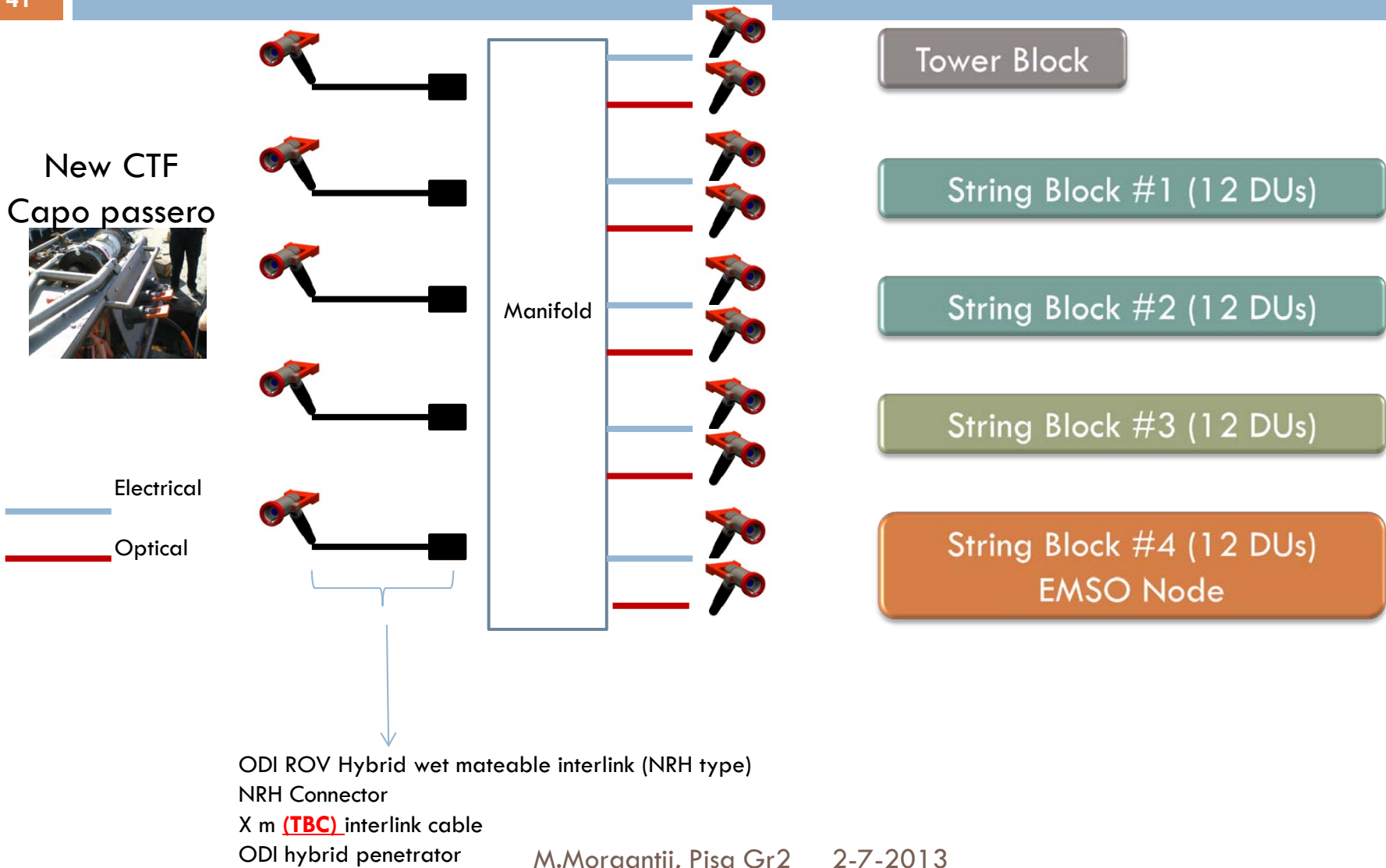


K. Leismuller (Sea Operation Manager Km3Net-IT) mail
Draft 01

CTF position according to: maintenance requirement of the MEOC, central position of the towers block

General Scheme

41



Seafloor Network

String and Tower configuration

42

- Constrain
 - ▣ $MVC = 8.5 \text{ kW}$
 - ▣ 8 Tower means
 - 4 fibres
 - $P_{\text{block}} = 200W_{\text{@tower}} \times 8 + P_{\text{Jb}} = 1.6 \text{ kW} + .2\text{kW} = 1.8\text{kW}$
 - ▣ String
 - 200W
 - #2 fibres shore → offshore
- Input available for string
 - ▣ 6.7 kW
 - ▣ 16 fibres

Seafloor Network

String capability

43

□ Optical

▣ 16 fibres

■ 2 fibres shore → offshore

■ 14 fibres offshore → shore

■ 4 string (18 DOM)@ fibres = 56 strings

□ Electrical

▣ $P_{jb} = 200W \rightarrow 4 \text{ JB (3 JB string + 1 JB EMSO)} = .8 \text{ kW}$

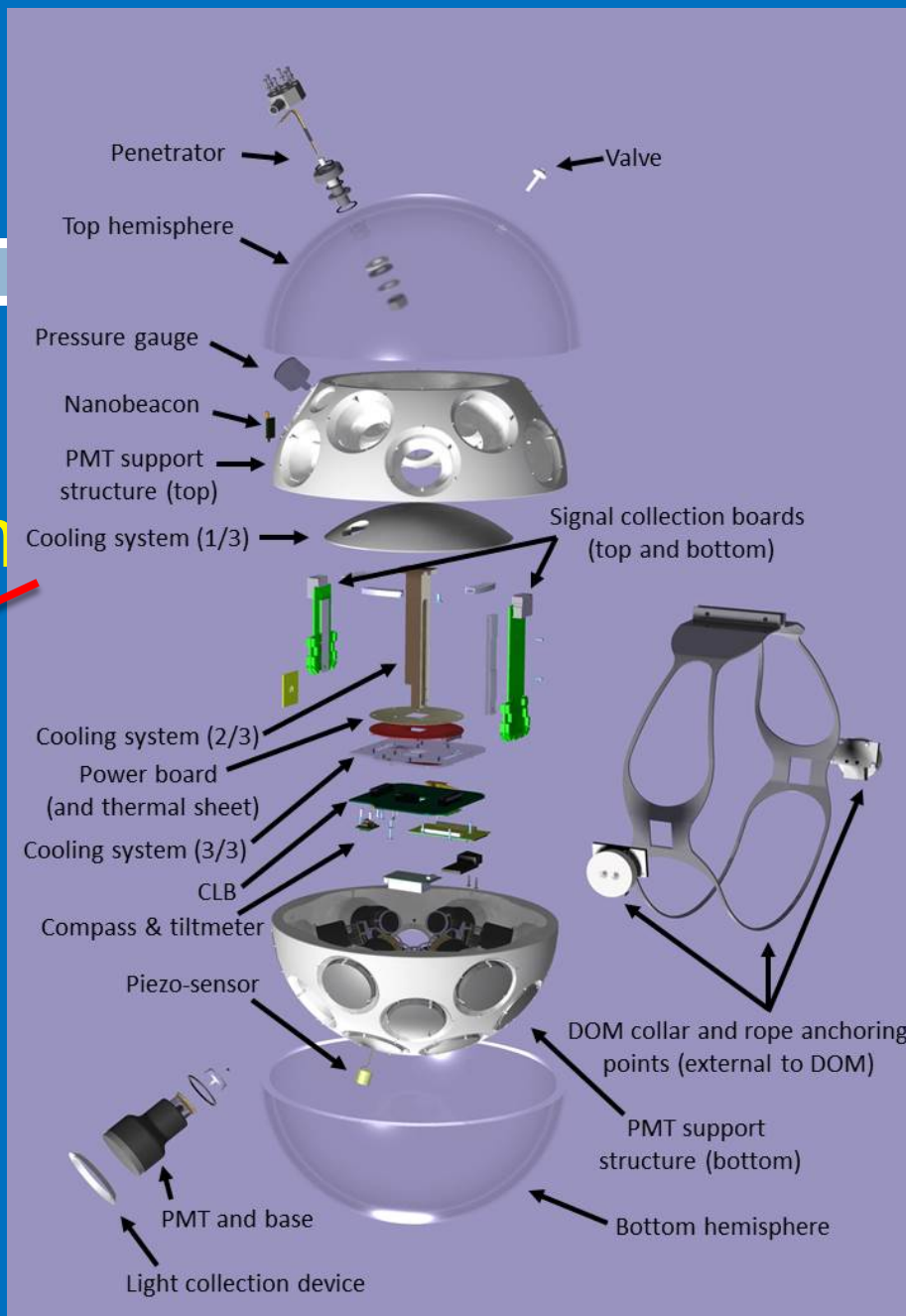
▣ EMSO Node = .8 kW

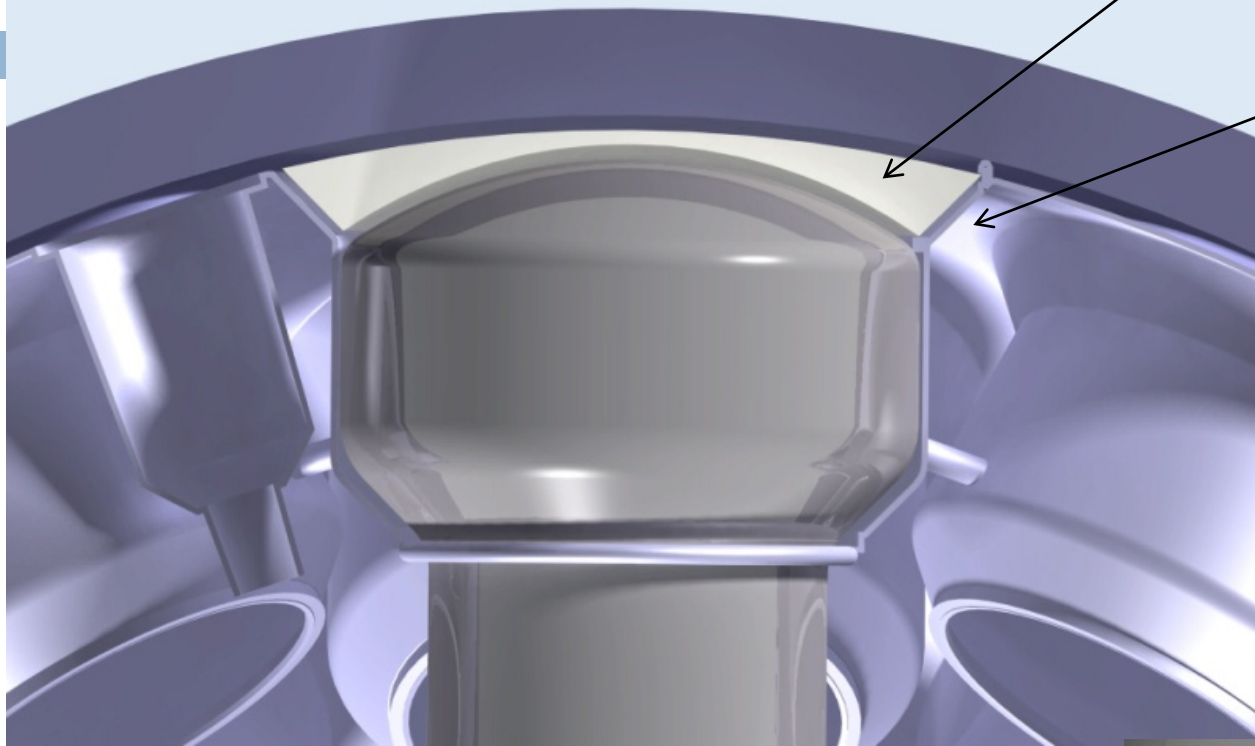
▣ P available = 5.1 kW → **26 strings**

IT'S OUR LIMIT!!!

The DOM breakdown structure

The DU structure





PMMA-lens

aluminium coating
deposited by vacuum
evaporation.

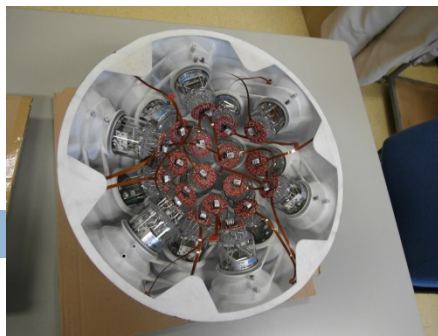
Cross-section with PMMA
“lens”, possibly exchangeable
with a pressed aluminum or
PMMA reflector ring.





PPM-Dom's ready for
electrical test.

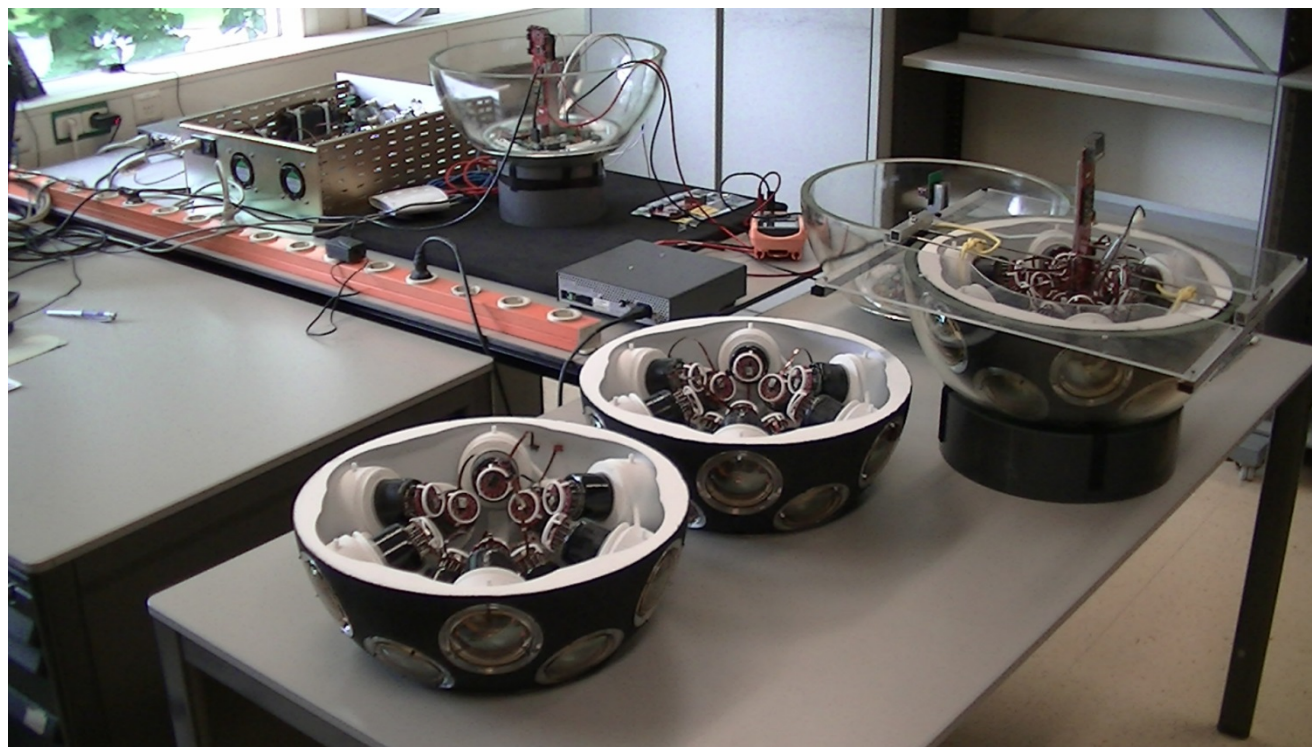
46



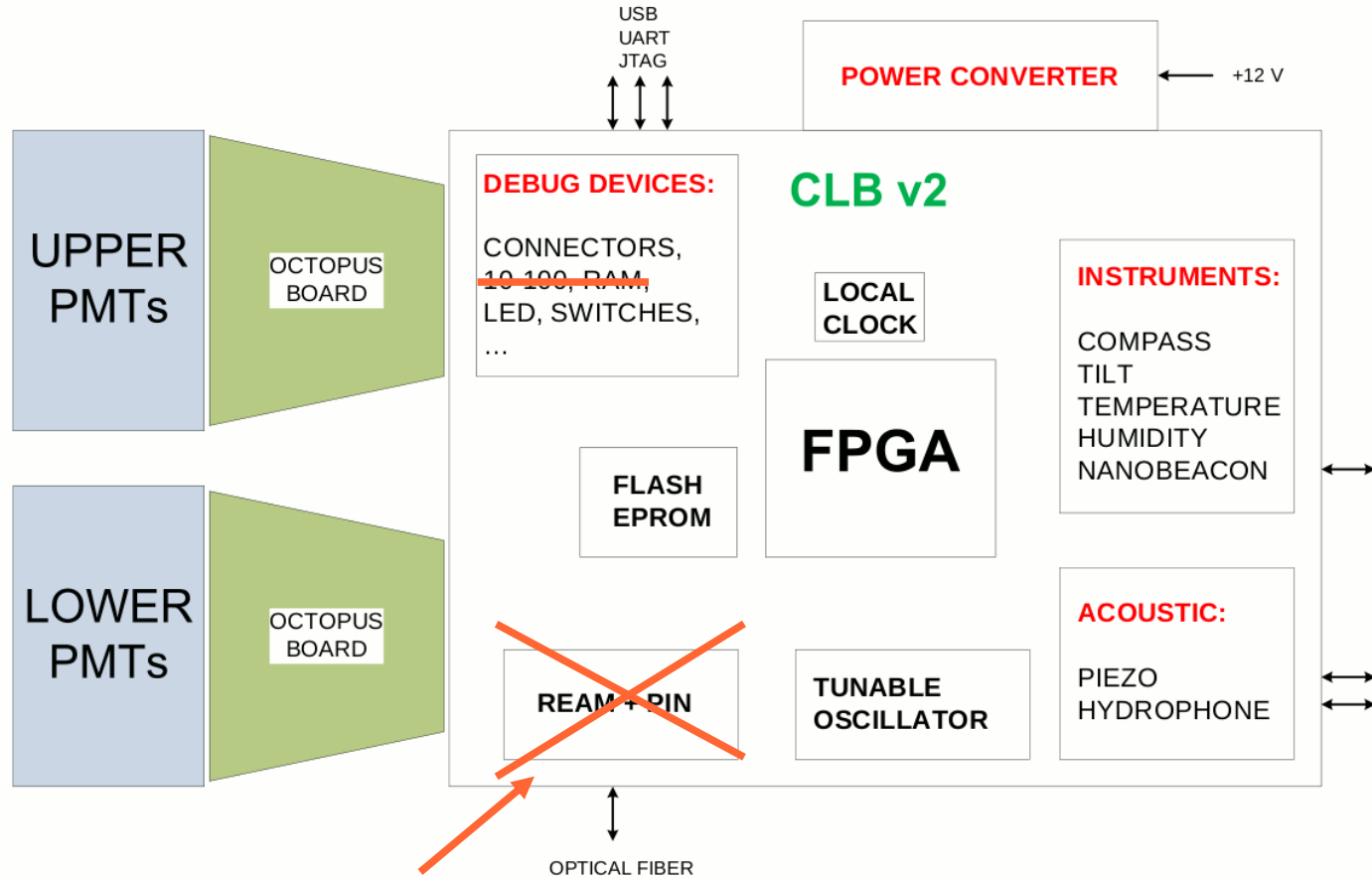
With Hamamatsu-
pmt's.



With ETEL-pmt's.



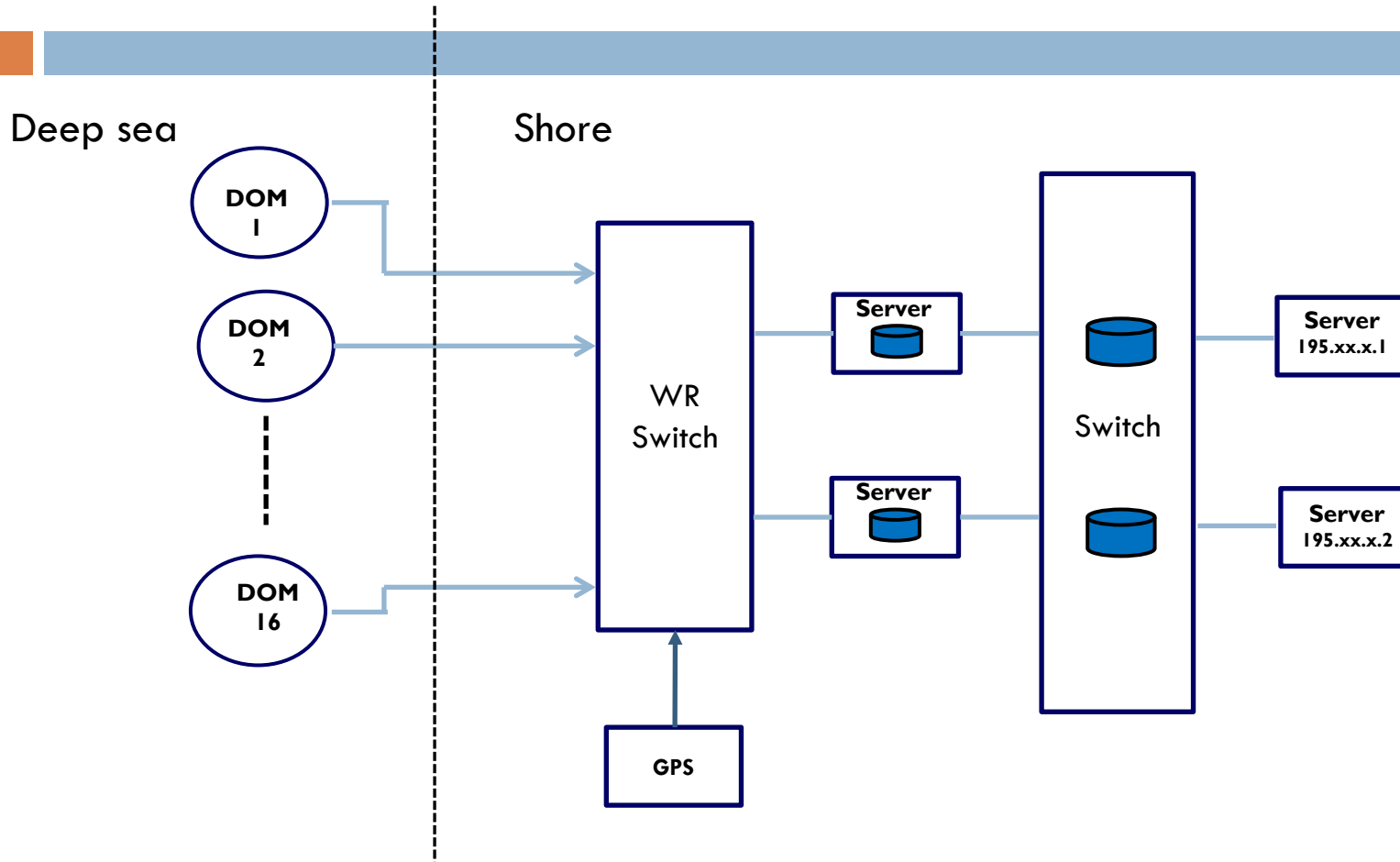
Schematic



Changed with
LASER

Shore Station diagram

48



1 WR switch accommodates up to 16 DOMs

Easily expandable structure to facilitate sets of 16-tuples of DOMs

Qualifica del progetto “multi-PMT (DOM) su stringhe”

- **Development Plan e Risk Analysis** avviati – da completare in ~2 mesi
- **PPM-DOM (Pre-production Model del Digital Optical Module):** montato (da tempo) sull'instrumentation line di ANTARES di prossima installazione



PPM-DOM



**Instrumentation line di
ANTARES imbarcata per il
deployment**



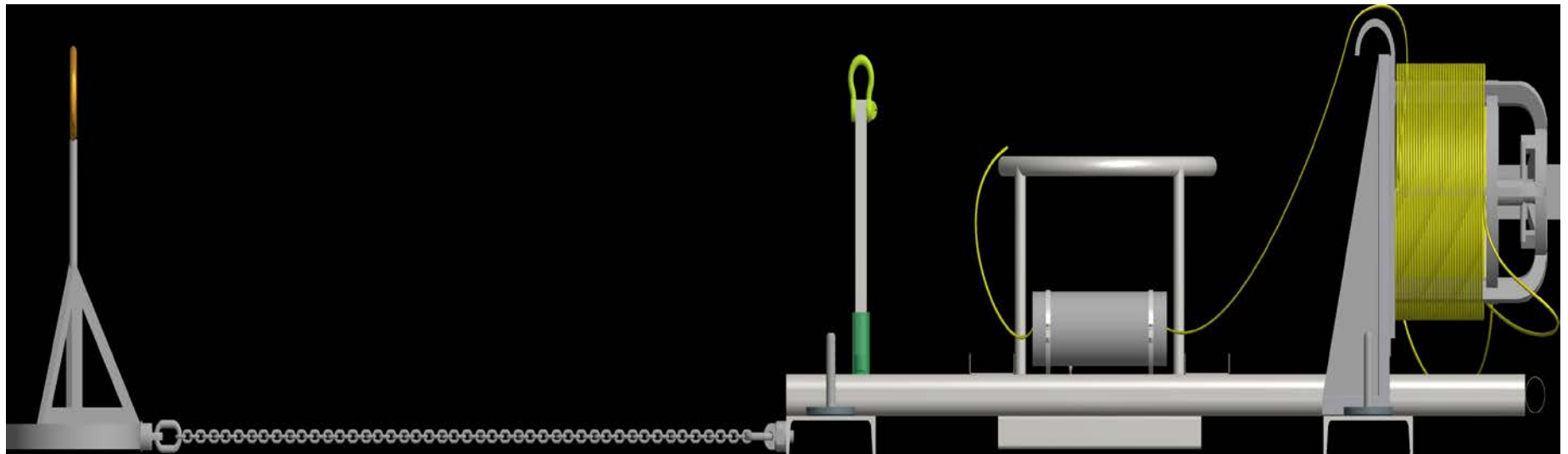
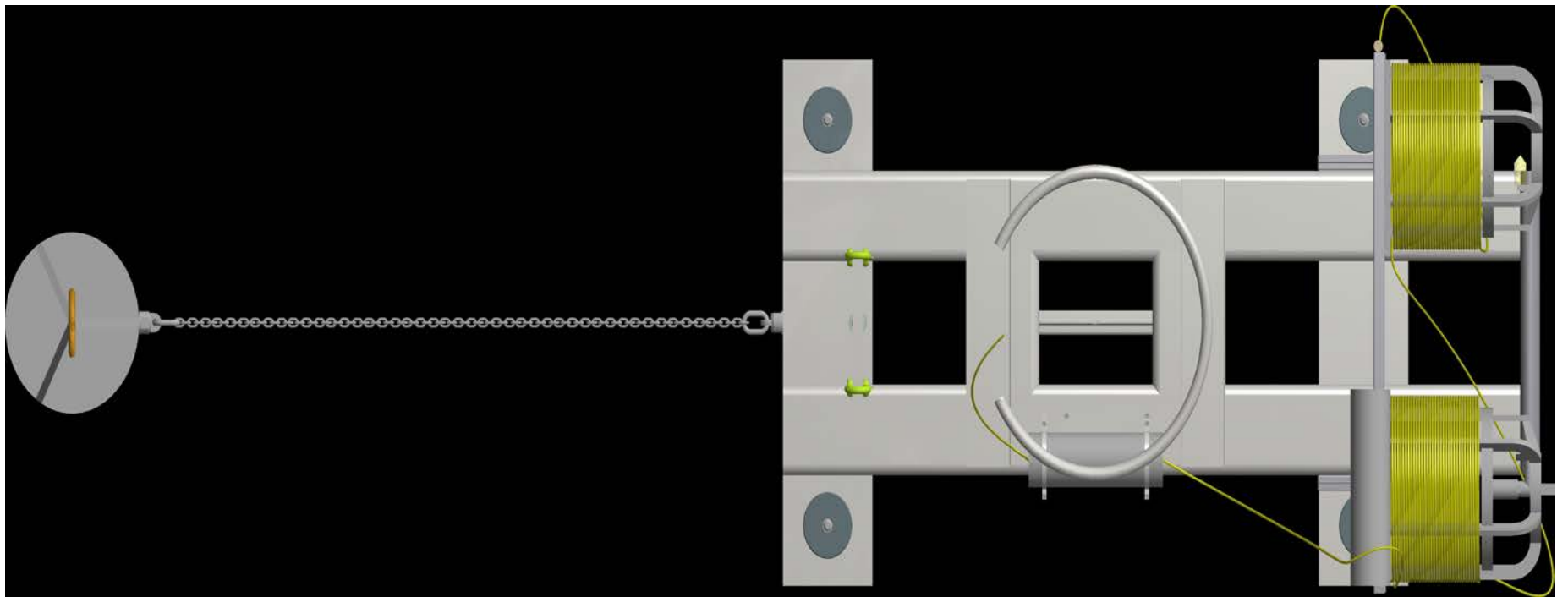
**Lanciatore con stringa
KM3NeT in fase di deployment**

- **Test di installazione delle stringhe:** ongoing in Spagna - Tra i goals della campagna c'è anche il test del cavo di dorsale e la tenuta delle sfere dei moduli ottici

- **PPM-DU (Pre-production Model della Detection Unit):** mini-stringa dotata di 3 DOM completamente strumentati, da installare in autunno



**Recupero del lanciatore dopo
l'installazione della stringa**



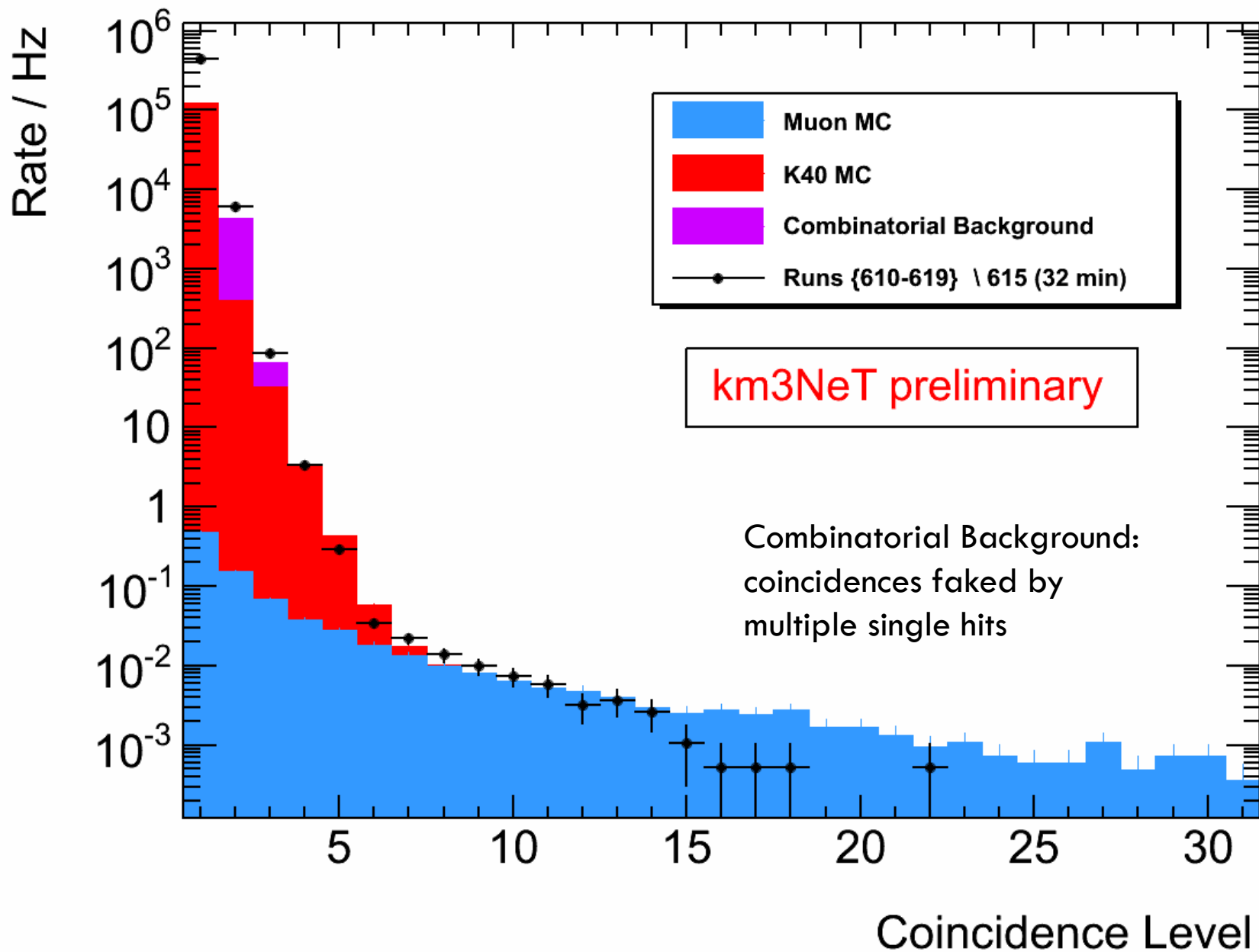
From prototypes to mass production

51



- **PPM-DOM installed in ANTARES and operated smoothly since mid-April**
- **3 new DOMs to be built for the PPM-DU**
- **Prototype experience precious:**
 - proof of concept
 - bench-mark of integration
 - experience in calibration, operation, data taking and analysis
- **Big step needed to go to mass production**
 - new design (electronics, optics)
 - qualification required for some items (penetrator, internal structure)
 - scale factor needed for mass production







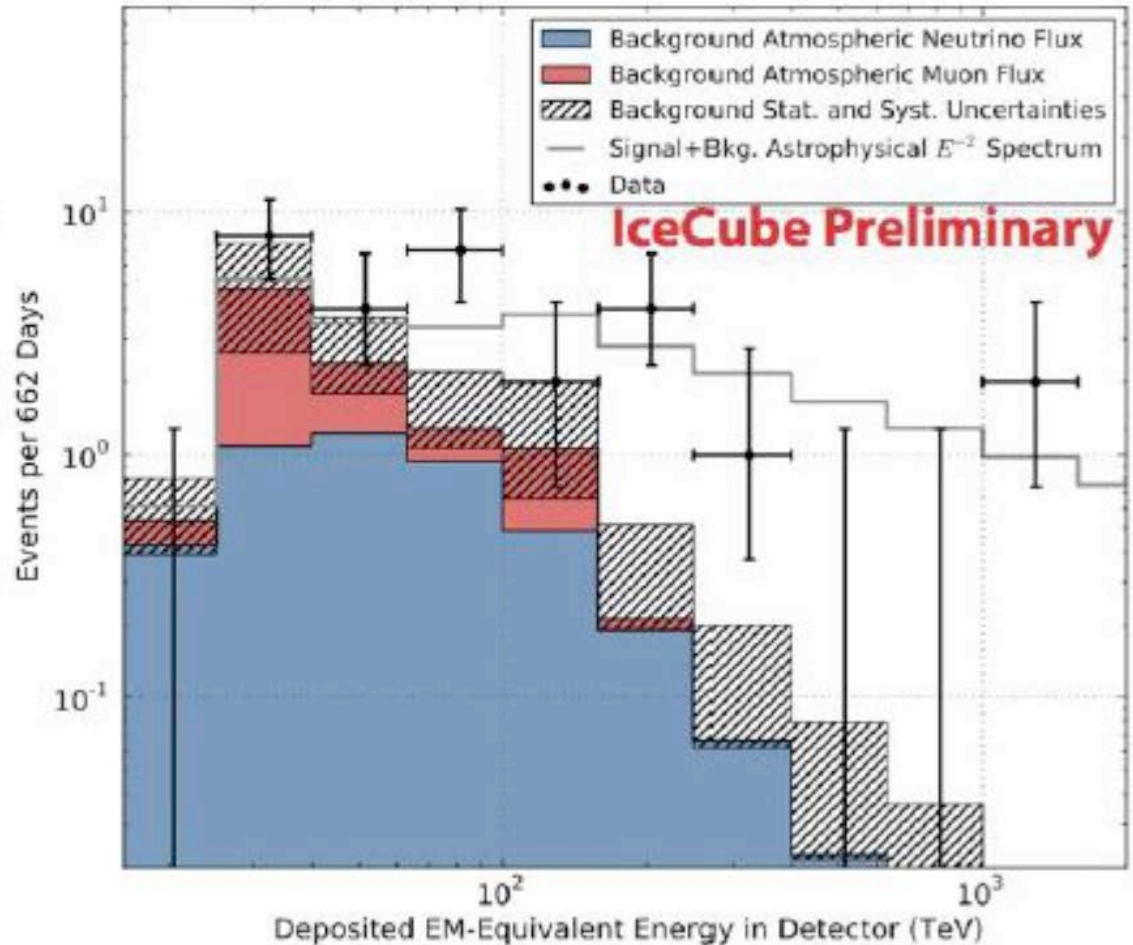
Schedule: assumptions

- **300 DOMs (Italy) + line 1 (France) are needed by end of 2014**
- Typical 5-month tendering time for Italian tenders
- Productions for line 1 can be faster
- **Production and integration times based on best estimates available today**
- Two DOM integration sites are active (working at a speed of 5 DOMs/week/site)

**Produzione e integrazione da definire:
possibile partecipazione di Pisa**

The 2013 IceCube result

- ▶ Harder than expected from atmospheric background
- ▶ Merges well into expected backgrounds at low energies
- ▶ Potential cutoff at $1.6^{+1.5}_{-0.4}$ PeV





ANTARES Results

**multi-messenger
astronomy**
(gamma astronomy,
optical astronomy,
gravitational
waves...)

- point sources
- diffuse flux

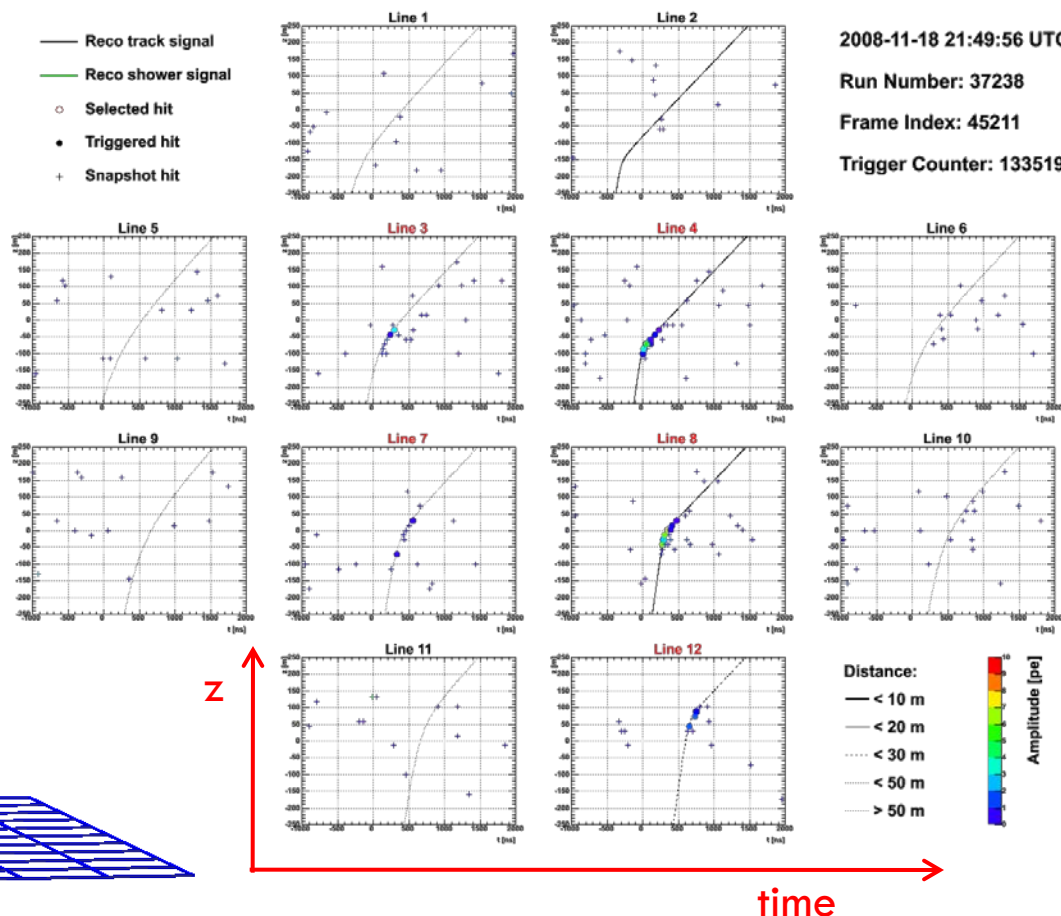
- **monopoles**
Astropart. Phys. 35 (2012)
634-640
- nuclearites
- dark matter



ANTARES Event Display a Neutrino Candidate

56

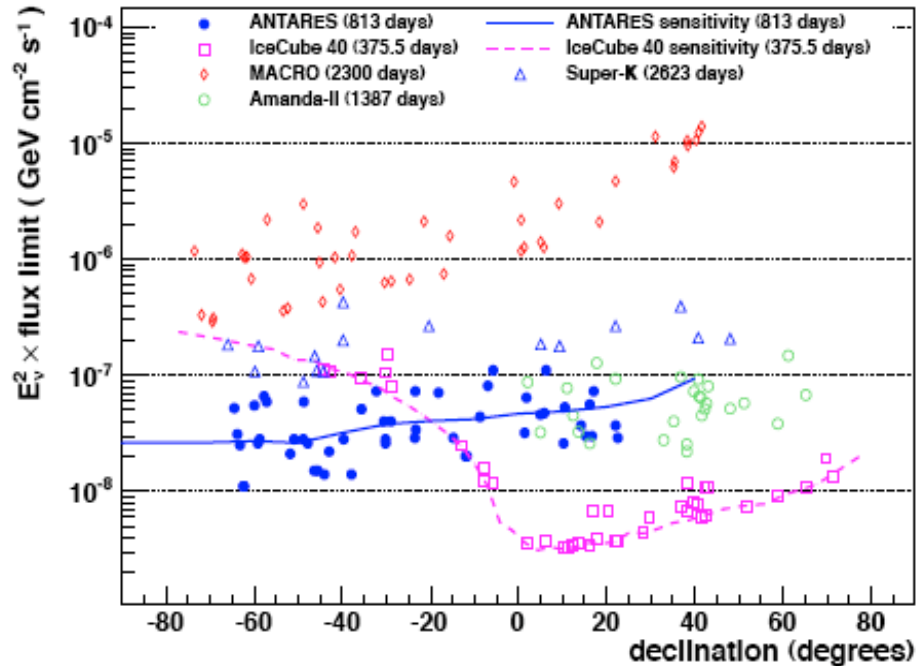
reco. μ energy: 5-10 TeV
decl.: $-6^{\circ}23'20.98''$
RA: 12h57 23.18



Point-like Sources

Astroph. J. Lett. 743 (2011) L14-L19

+ new analysis submitted to *ApJ*



No significant cosmic neutrino sources have been observed

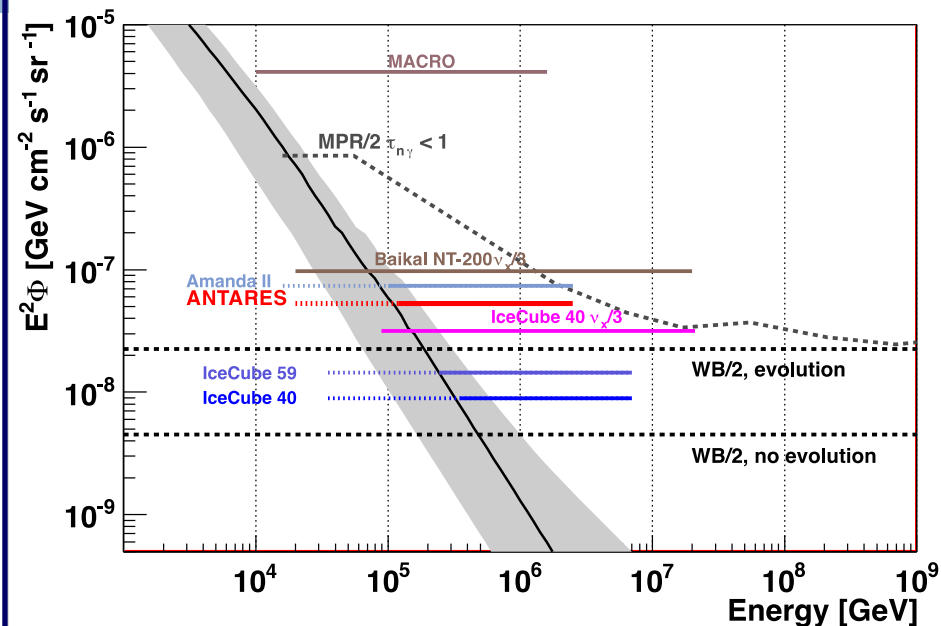
The BEST limit in the Southern Sky!

$$1-10 \times 10^{-8} (E/\text{GeV})^{-2} \text{ GeV}^{-1} \text{ s}^{-1} \text{ cm}^{-2}$$

[$\delta < 48^\circ$ (the part of the sky always visible)]

Diffuse Cosmic ν Flux

Phys. Letter B 696, 16-22 (2011)



No excess of high energy events found over expectation from atmospheric ν

Flux Upper Limit (90% CL)

$$E^2 \Phi(E) < 5.3 \times 10^{-8} \text{ GeV cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$$

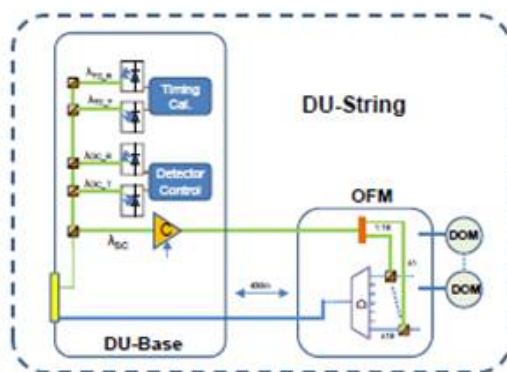
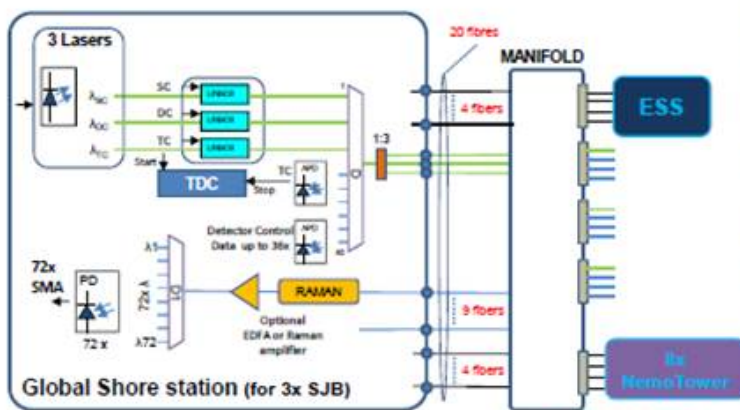
[for $20 \text{ TeV} < E < 2.5 \text{ PeV}$]

ANTARES: Papers

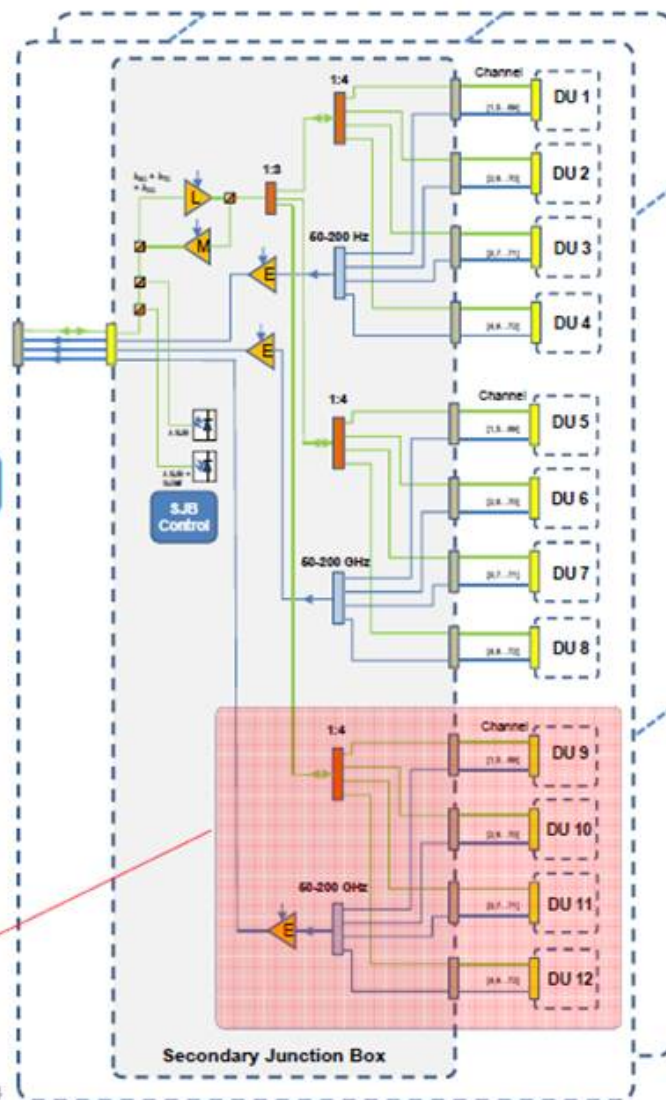
18	A FIRST SEARCH FOR COINCIDENT GRAVITATIONAL WAVES AND HIGH ENERGY NEUTRINOS USING LIGO, VIRGO AND ANTARES DATA FROM 2007 (Antares+ Ligo+ Virgo)	JCAP_039P_0413, published as: JCAP06(2013)008	Astrophysics
1	AUGER AND HE NEUTRINO COINCIDENCES (Antares+Auger)	Submitted to APP	Astrophysics
2	First search for neutrinos in correlation with gamma-ray bursts with the ANTARES neutrino telescope, S. Adrian-Martinez et al., JCAP 03(2013)006	Astrophysics	
3	HIGH-FREQUENCY INTERNAL WAVE MOTIONS AT THE ANTARES SITE IN THE DEEP WESTERN MEDITERRANEAN	In preparation	
4	A TIME DEPENDENT SEARCH FOR NEUTRINO EMISSION FROM MICROQUASARS WITH THE ANTARES	In prepararion, S. Galatà, C. Distefano	
5	SEARCH FOR NUCLEARITES WITH THE ANTARES DETECTOR	In preparation	
6	2008-2011 GRB analysis	Julia Schmid, Colas Rivière	
7	The Fermi Bubbles analysis	V. Kulikovskiy et al.	
12	Search for dark matter from the Sun	Submitted to APP	



Preliminary Design Phase-1 Capo Passero : SFP Lasers, 72 Channels, Separate fiber for SC, DC, TC



3x Secondary Junction Box



4 extra DUs are only possible when power system is updated!

G. Kieft
Nikhef
Amsterdam
Electronics
Technology



KM3NeT Phase 1 - Optical Network for Capo Passero - Designed by Nikhef Optical Team - 2013-08-24

25 Jun 2013 Collaboration Meeting Catania

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From Gerard Kieft's presentation in optical system session

M.Morgantii, Pisa Gr2 2-7-2013



Multi-Messenger Searches

60

Potential astrophysical sources are predicted to emit very faint neutrino signal. The Multi-Messenger Approach increases the **discovery potential**, by observing with different probes; the **significance**, by coincident detection; the **efficiency**, by relaxed cuts.



AGN Flares

(ν emission from γ -flaring blazars)

[ANTARES ↔ FERMI]



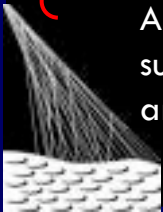
arXiv:1111.3473 [astro-ph.HE]
sub. to *Astrop.Phys*

[ANTARES ↔ VIRGO
LIGO]

common working group (GWHEN)
arXiv:1205.3018 [astro-ph.HE]
sub. to *Astrop.Phys*

[ANTARES ↔ UGER]

Adrian-Martinez et al.,
subm to *Astropart. Phys.*
arXiv:1202.6661



TAToO

(Telescopes – ANTARES
Target of Opportunity)

**Optical follow-up of
neutrino alerts for
transient source search**
(GRBs, SNe).

Analysis in progress!

[ANTARES



↔ Optical Telescopes
TAROT & ROSTE + more]



Ageron et al., *Astrop.Phys* 35 (2012) 530-536



GCN

(GRB Coordinat. Network)

[ANTARES

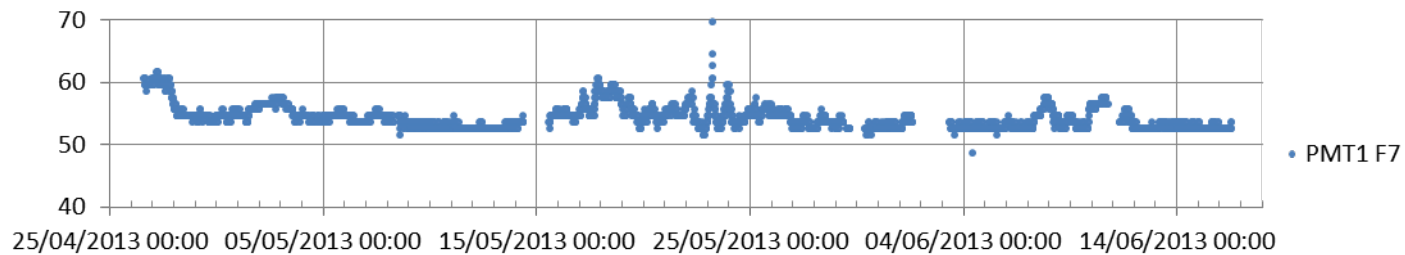


↔ CN]

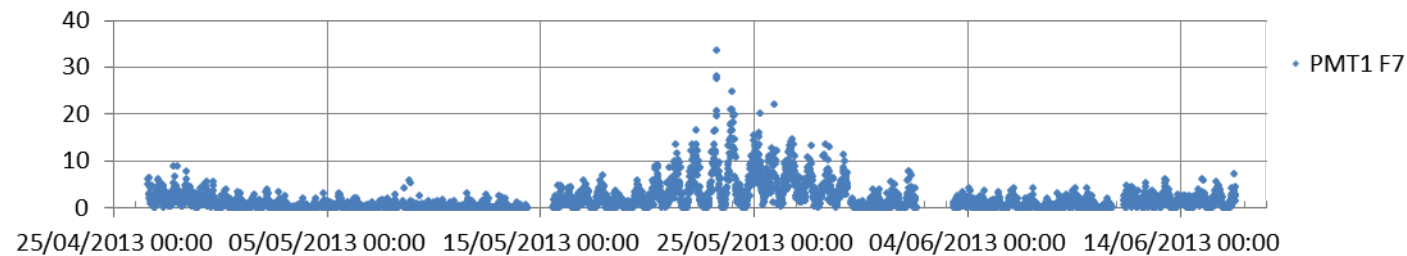
Long term behaviour of the background rates

61

Baseline (kHz)



Burst fraction > 100 kHz (%)



Floor heading

