## Measurement of astrophysically important excitation energies of <sup>58</sup>Zn with GRETINA

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Type I X-ray bursts are thermonuclear explosions taking place on an accreting neutron star in a lowmass X-ray binary system. The accreted material, consisting of H/He, heats up under high pressure and at a certain temperature a thermonuclear runaway is ignited. The main energy generation is driven by the rapid proton capture process (rp-process), which synthesizes elements up to  $A \approx 100$  via fast  $(p,\gamma),(\alpha,\gamma)$  and  $(\alpha,p)$  reactions, and slower  $\beta^+$  and electron-capture decays. Along the reaction path of the rp-process, several waiting points affecting element and energy production have been identified. Of special interest is doubly-magic <sup>56</sup>Ni. It has been shown that the reaction <sup>57</sup>Cu(p, $\gamma$ )<sup>58</sup>Zn determines the effective lifetime of <sup>56</sup>Ni since the electron-capture lifetime of <sup>56</sup>Ni is larger than 1000 s and <sup>56</sup>Ni is in  $(p,\gamma) - (\gamma,p)$  equilibrium with <sup>57</sup>Cu at rp-process temperatures. Proton capture on <sup>57</sup>Cu is the only open break-out reaction channel within typical burst timescales. So far, the rate was only calculated theoretically with large uncertainties due to the unknown level structure of <sup>58</sup>Zn.

At the National Superconducting Cyclotron Laboratory we studied the astrophysically important excitation energies of <sup>58</sup>Zn, which determine the <sup>57</sup>Cu(p, $\gamma$ )<sup>58</sup>Zn rate. The secondary <sup>57</sup>Cu beam was produced by many-nucleon transfer reactions of a primary <sup>58</sup>Ni beam impinging on a Be target, and further purification and separation of the beam was performed using the A1900 fragment separator. The secondary beam was transported to the S800 large-acceptance spectrometer, where a CD<sub>2</sub> target with a thickness of 200 mg/cm<sup>2</sup> was installed and surrounded by the next-generation gamma-ray detector GRETINA. Excited levels in <sup>58</sup>Zn were populated using a (d,n) reaction in the secondary target, and the de-excitation  $\gamma$ -rays were measured.

The presentation will focus on the details of the study and present the measured excitation energies of astrophysically important  $2^+$  states in  ${}^{58}$ Zn and their astrophysical implications.