

# First results from the GERDA experiment

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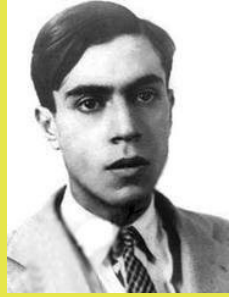


**LNGS seminar,**  
**21 February 2013**

# Outline

- **Double beta decay:  $2\nu\beta\beta$  and  $0\nu\beta\beta$**
- **The GERDA experiment**
- **First results: The  $2\nu\beta\beta$  half-life of  $^{76}\text{Ge}$**
- **A look into the future**

# Unveil the nature of the neutrino



**Majorana**  
 $\nu = \bar{\nu}$

or



**Dirac**  
 $\nu \neq \bar{\nu}$

**If  $0\nu\beta\beta$  is observed:**

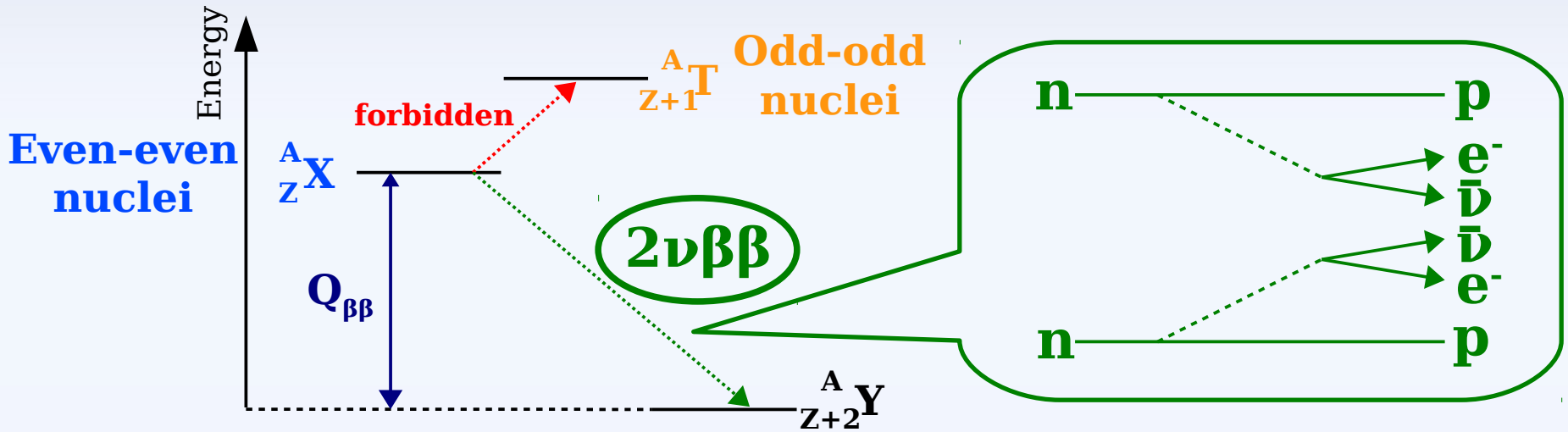
Lepton number violation  $\Delta L=2$

Neutrino has a Majorana mass term

Sheds light on absolute neutrino mass scale

Sheds light on neutrino mass hierarchy

# $2\nu\beta\beta$

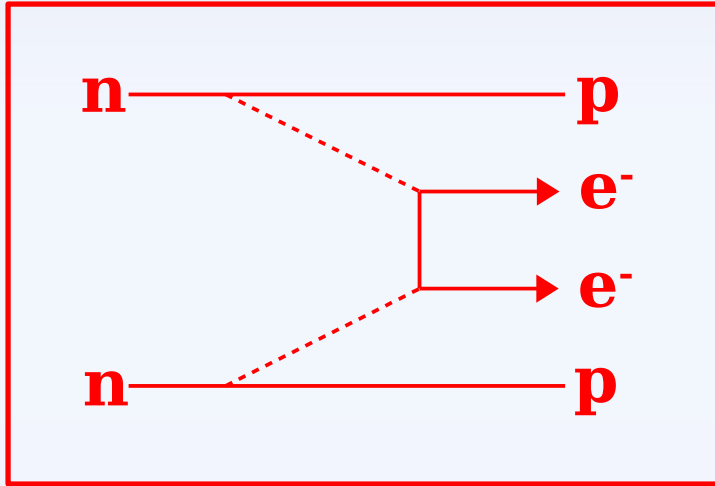


- allowed by SM
- $\Delta L=0$
- observed in many isotopes
- $T_{1/2}^{2\nu} \sim 10^{19} - 10^{21} \text{ yr}$
- $(T_{1/2}^{2\nu})^{-1} = G^{2\nu}(Q_{\beta\beta}, Z) \cdot |M^{2\nu}|^2$

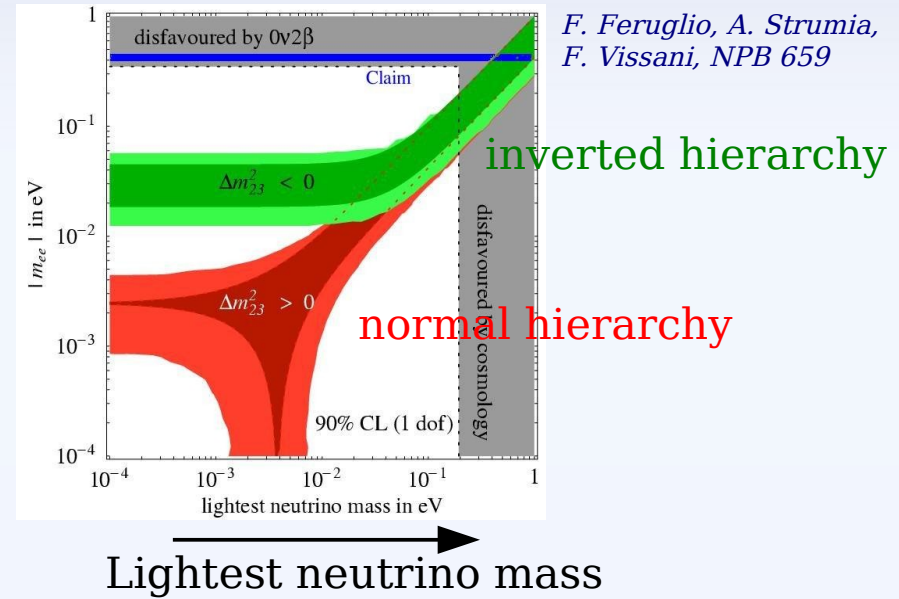
*Phase space*

*nuclear matrix element*

# $0\nu\beta\beta$



$|m_{\beta\beta}|$  ↑



● **Forbidden process in SM, needs Majorana neutrino**

●  $\Delta L=2$

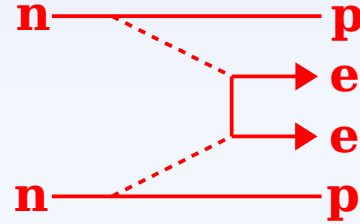
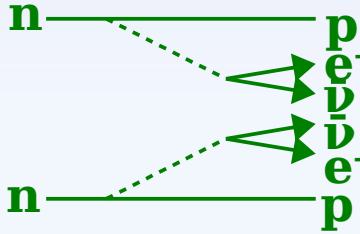
●  $(T_{1/2}^{0\nu})^{-1} = G^{0\nu}(Q_{\beta\beta}, Z) \cdot |M^{0\nu}|^2 \cdot \langle m_{\beta\beta} \rangle^2$

*Phase space* ( $\sim Q_{\beta\beta}^5$ )

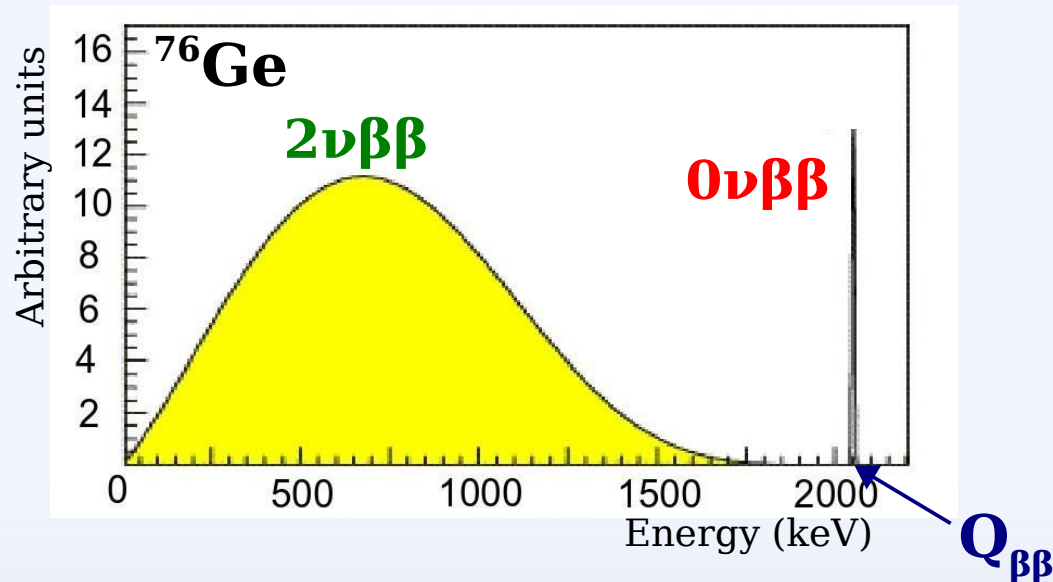
*nuclear matrix element*

$\langle m_{\beta\beta} \rangle^2 = |\sum_i |U_{ei}|^2 e^{i\alpha} m_i|^2$   
*Majorana neutrino mass*

# Experimental signature



➔ Measure **electron energy spectrum**:



# Searching in $^{76}\text{Ge}$

$$S \sim \epsilon \cdot f \cdot \sqrt{\frac{M \cdot t_{\text{run}}}{\text{BI} \cdot \Delta E}}$$

S: sensitivity

$\epsilon$ : efficiency

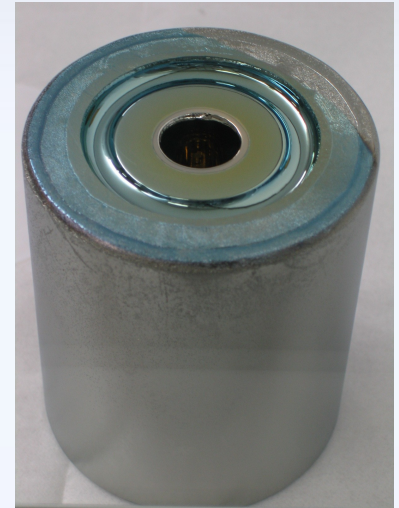
f: abundance of  $0\nu\beta\beta$  isotope

M: detector mass

$t_{\text{run}}$ : measurement time

BI: background index

$\Delta E$ : energy resolution at  $Q_{\beta\beta}$



*Germanium detector*

## Advantages of Germanium:

- **High  $\epsilon$** : Source = Detector
- **Small intrinsic BI**: High purity Ge
- **Excellent  $\Delta E$** : FWHM  $\sim$  (0.1-0.2)%
- Well-established technology

## Disadvantages of Germanium:

- **High external BI**:  $Q_{\beta\beta} = 2039\text{keV}$
- **Small f of  $^{76}\text{Ge}$** :  
7.8%  $\rightarrow$  Enrichment needed!
- Limited sources of crystal & detector manufacturers
- Small  $G^{0\nu}(Q_{\beta\beta}, Z)$

# Previous $^{76}\text{Ge}$ experiments

	<b>HdM</b>	<b>IGEX</b>
<b>Location</b>	LNGS	Homestake, Baksan, Canfranc
<b>Exposure [kg·yr]</b>	71.1	8.9
<b>Bg [cts/(keV·kg·yr)]</b>	$\geq 0.11$	0.17
<b><math>T_{1/2}</math> limit (90% CL) [yr]</b>	$1.9 \cdot 10^{25}$ [1]	$1.6 \cdot 10^{25}$ [2]

[1] *Eur. Phys. J. A12*, 147-154 (2001)

[2] *Phys. Rev. D 65*, 092007 (2002)

## **Claim of signal from part of HdM:**

$T_{1/2} (^{76}\text{Ge}) = (0.69 - 4.18) \cdot 10^{25}$  yr ( $3\sigma$ ) (Best fit:  $T_{1/2} (^{76}\text{Ge}) = 1.19 \cdot 10^{25}$  yr)

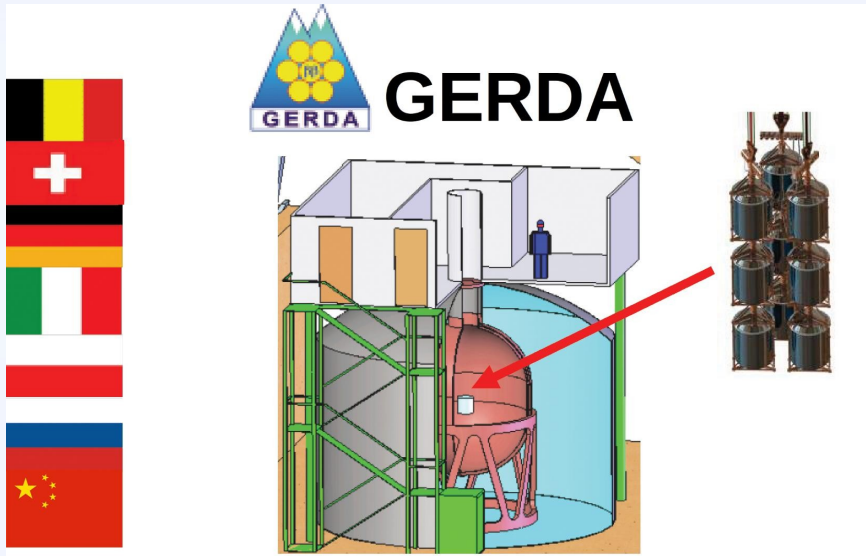
*Phys. Lett. B 586*, 198-212 (2004)



# GERmanium Detector Array (GERDA)

## The GERDA collaboration

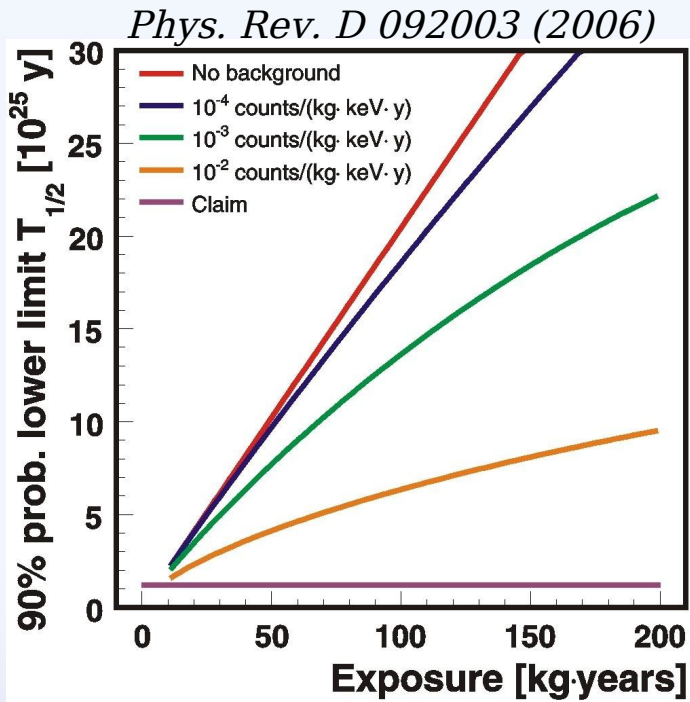
*112 members, 19 institutes, 7 countries*



- Bare  $^{enr}\text{Ge}$  array in liquid Argon
- Shield: high-purity liquid Argon /  $\text{H}_2\text{O}$
- **Phase I:** 18 kg enriched coaxial detectors ( $\sim 86\%$ )(HdM/IGEX)
- **Phase II:** add  $\sim 20$  kg new enriched BEGe detectors
- For future ton scale experiment: Merge with Majorana collaboration (already open exchange of knowledge and technologies)

# GERDA physics goal

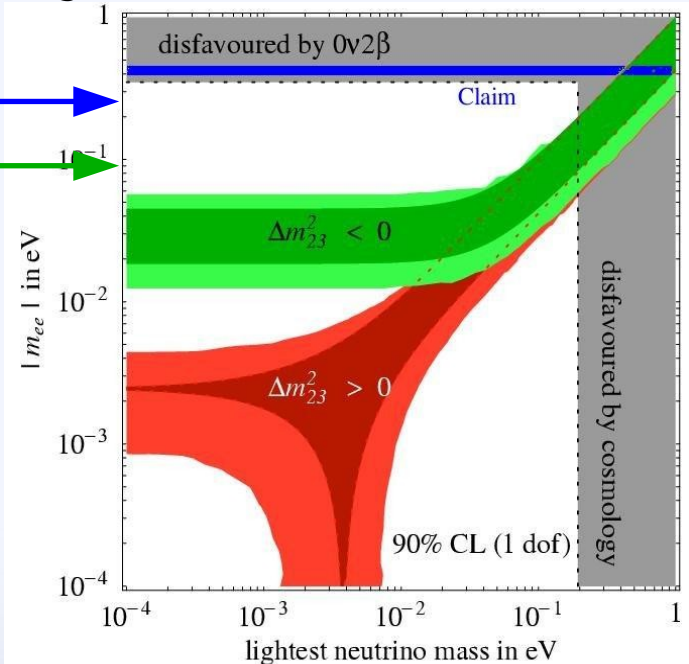
- Phase I:**
- reach sensitivity of  $T_{1/2} = 2 \cdot 10^{25}$  yr at 90% C.L.
  - $\langle m_{\beta\beta} \rangle \leq 0.23\text{-}0.39$  eV
  - $\rightarrow$  **check claim!**
- Phase II:**
- reach background of  $10^{-3}$  cts/(keV·kg·yr)
  - Exposure of 100 kg·yr  $\rightarrow T_{1/2} > 1.35 \cdot 10^{26}$  yr
  - $\langle m_{\beta\beta} \rangle \leq 0.09\text{-}0.15$  eV



*F. Feruglio, A. Strumia, F. Vissani, NPB 659*

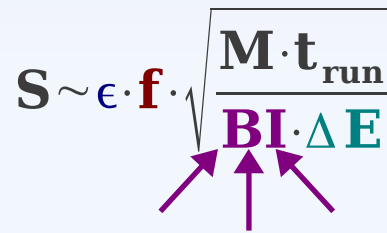
**Phase I**  $\rightarrow$

**Phase II**  $\rightarrow$



# Background reduction

**Key issue: Low background rate**

$$S \sim \epsilon \cdot f \cdot \sqrt{\frac{M \cdot t_{\text{run}}}{BI \cdot \Delta E}}$$


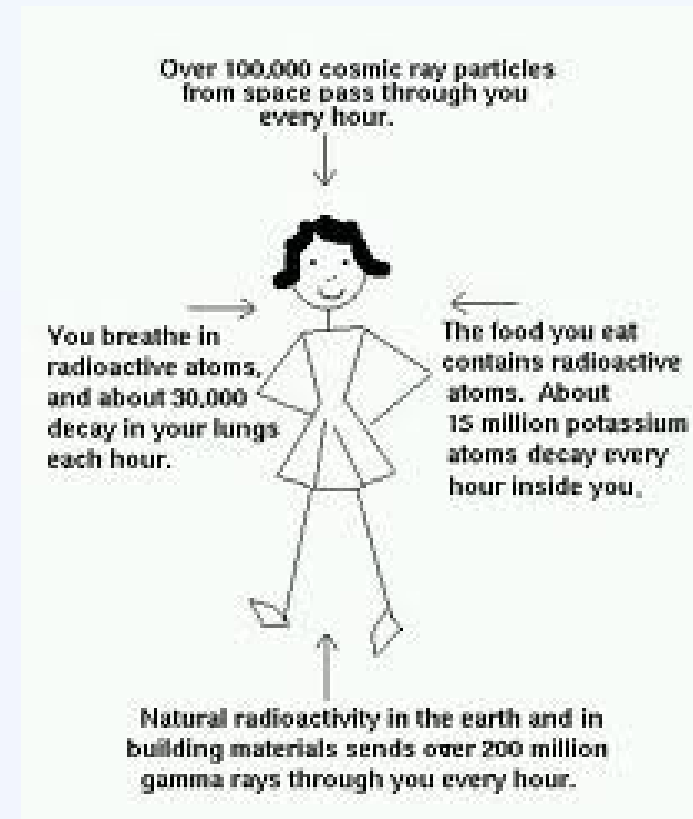
**Possible backgrounds:**

**External:**

- $\gamma$  from Th and U chain
- neutrons
- $\mu$  from cosmic rays (prompt and delayed)

**Internal:**

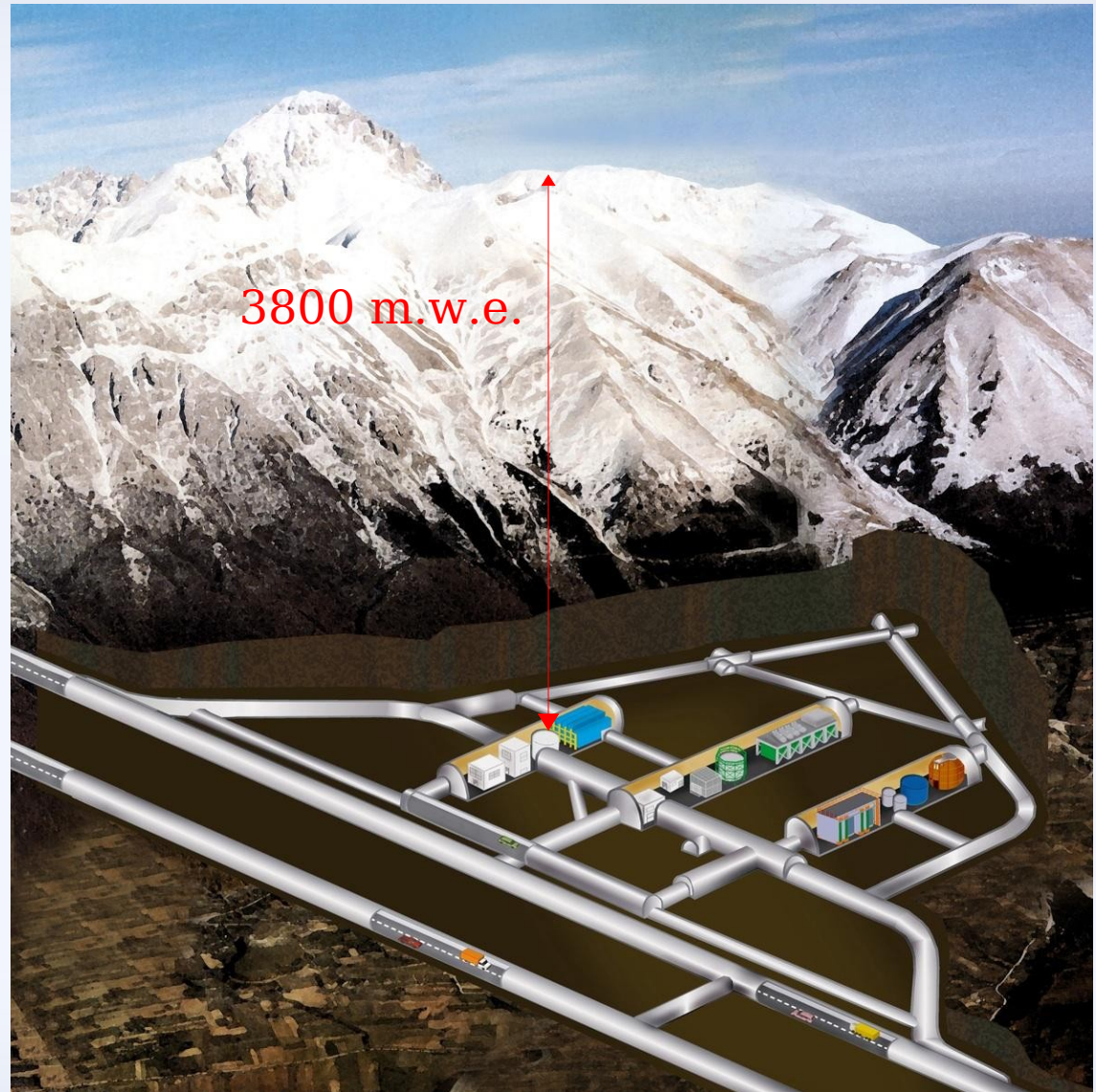
- cosmogenic  $^{60}\text{Co}$  ( $T_{1/2} = 5.3 \text{ yr}$ )
- cosmogenic  $^{68}\text{Ge}$  ( $T_{1/2} = 271 \text{ d}$ )
- Radioactive surface contamination



# Background reduction

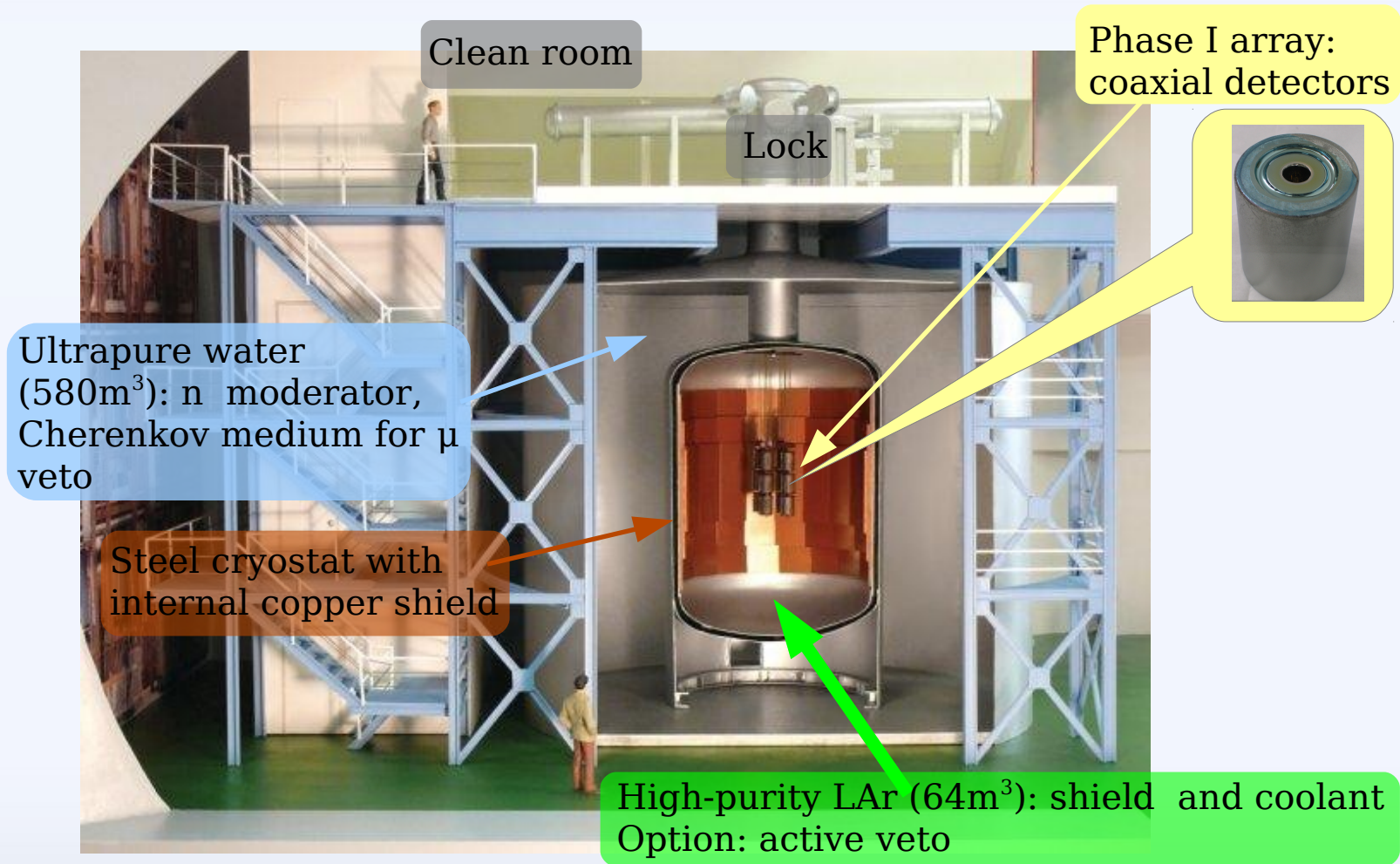
## GERDA situated in LNGS underground laboratories

- ➔ suppression of cosmic ray muons by factor  $10^6$  by overlaying rock
- ➔ hadronic components of cosmic rays removed



# Background reduction

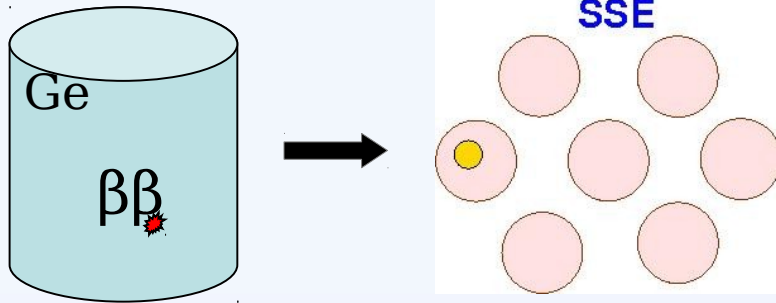
- Graded shielding against ambient radiation
- Rigorous material selection, avoid exposure above ground for detectors



arXiv: 1212.4067

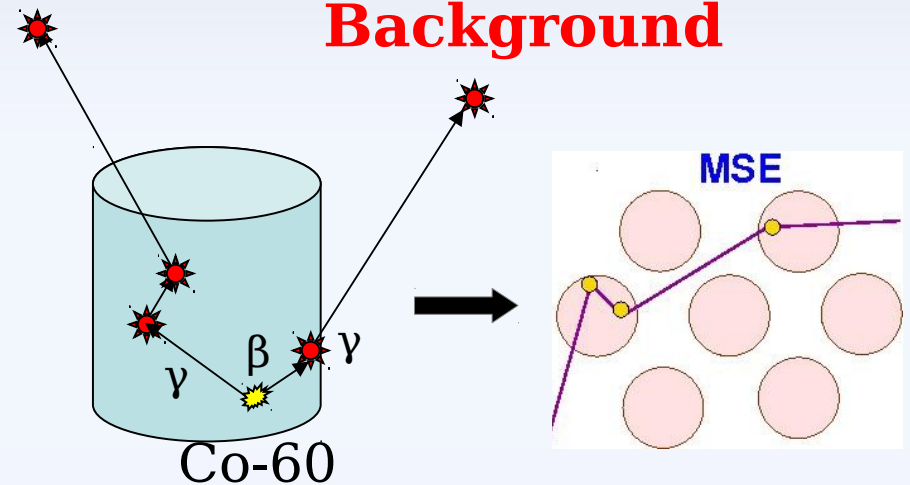
# Background reduction

## Signal



**Point-like (single-site)**  
energy deposition inside one  
HP-Ge diode (Range:  $\sim 1$  mm)

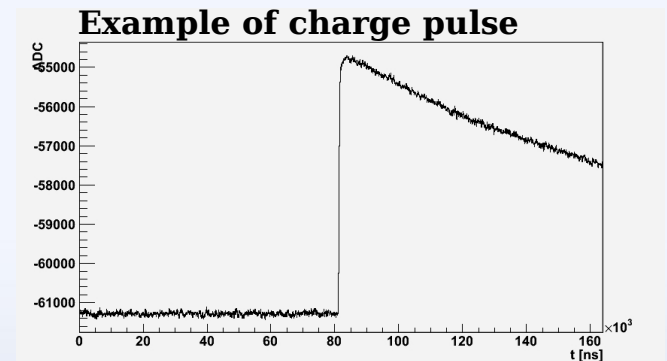
## Background



**Multi-site** energy deposition  
inside HP-Ge diode (Compton  
scattering)

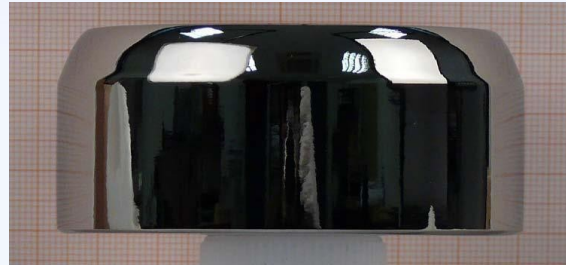
## Signal analysis:

- anti-coincidence between detectors
- pulse shape analysis (PSA)

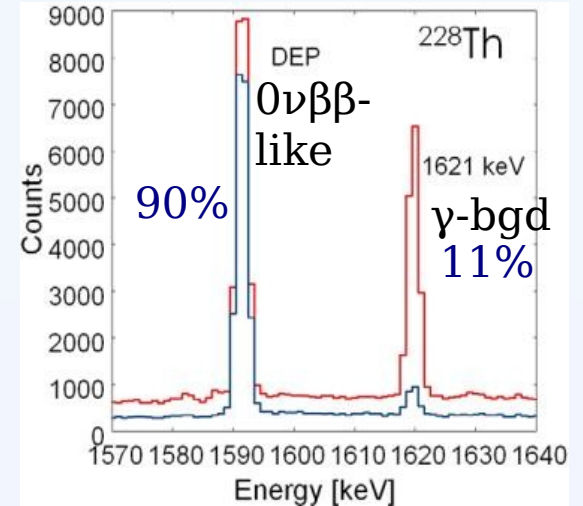
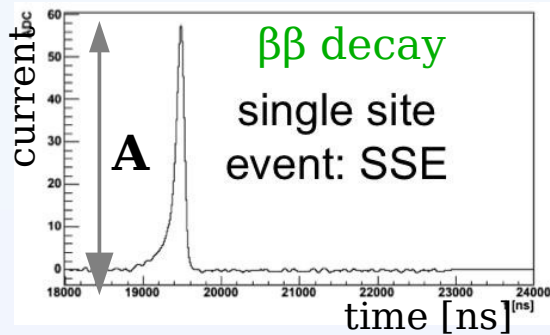


# Background reduction Phase II

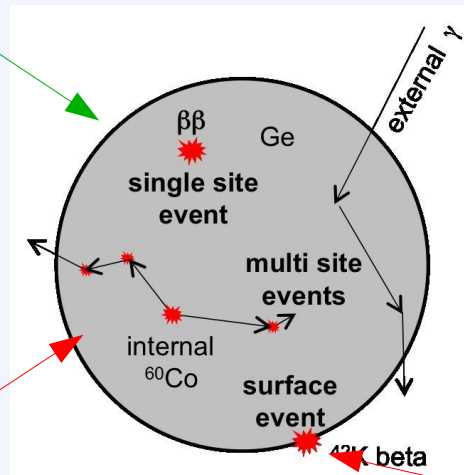
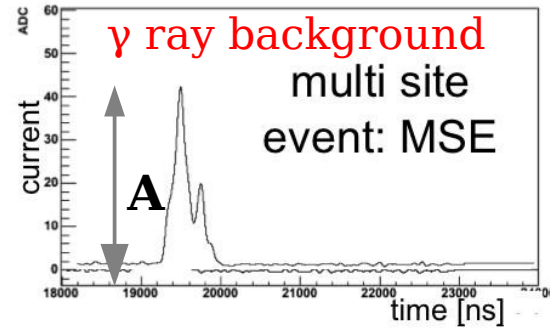
**BEGe detectors:** strongly non-linear field allows improved PSA



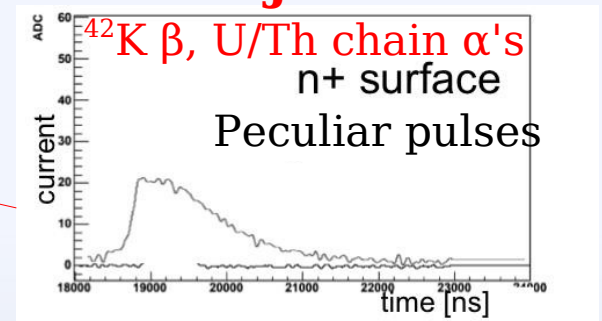
**ACCEPT**



**REJECT**



**REJECT**



# The GERDA experiment





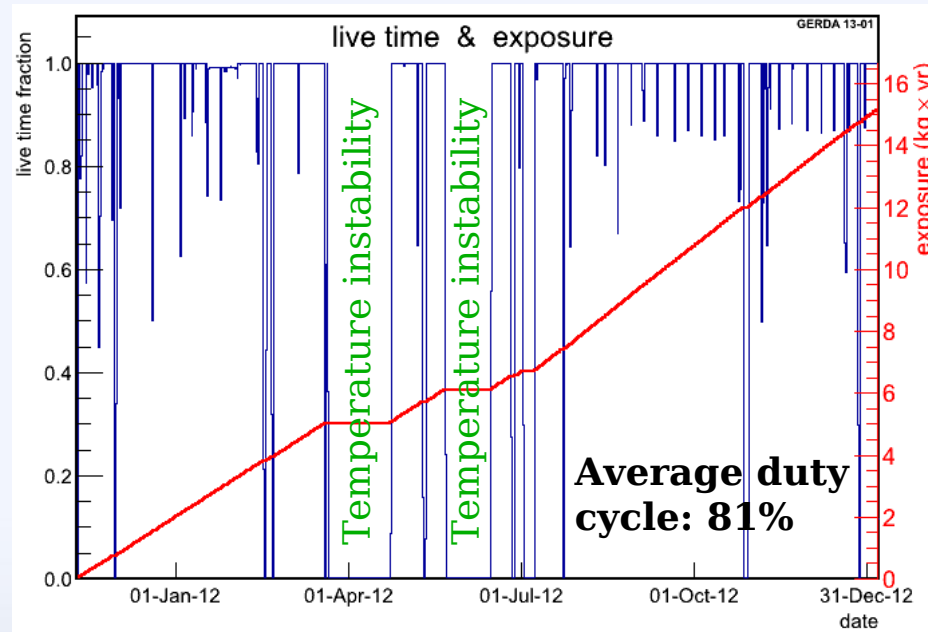
# Data taking

**9 November 2011:** Start of Phase I

All **8  $^{enr}\text{Ge}$  + 3  $^{nat}\text{Ge}$  coaxial detectors** deployed in GERDA  
(2  $^{enr}\text{Ge}$  detectors cannot be used for analysis due to high leakage current)

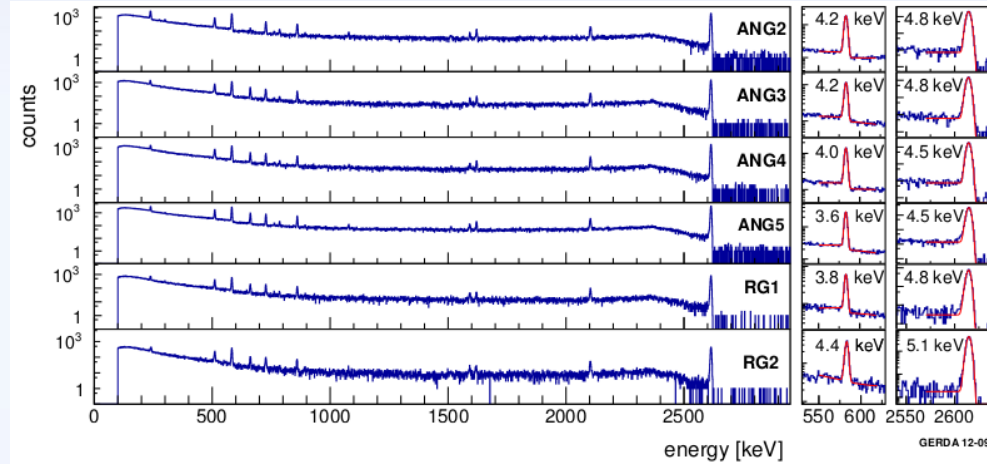
**7 July 2012:** Insert **5  $^{enr}\text{Ge}$  BEGe** detectors  
(Remove 2  $^{nat}\text{Ge}$  detectors)

**9 Nov 2011 - 5 Jan 2013:**  
**341 live days / 15.16 kg·yr enr exposure**



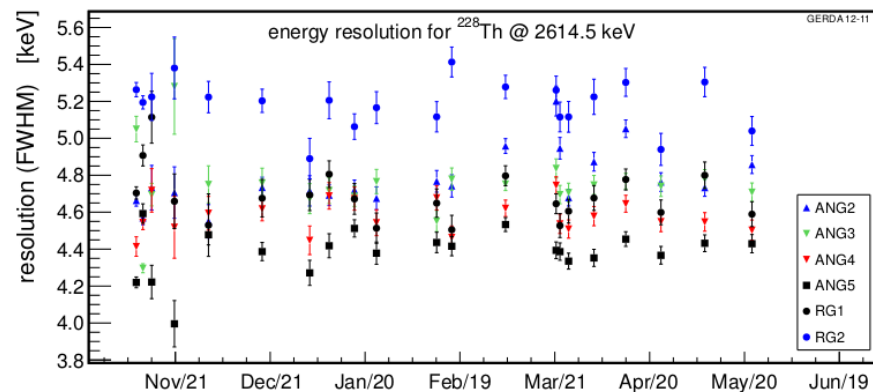
# Energy resolution

## Calibration spectra for $^{228}\text{Th}$ source

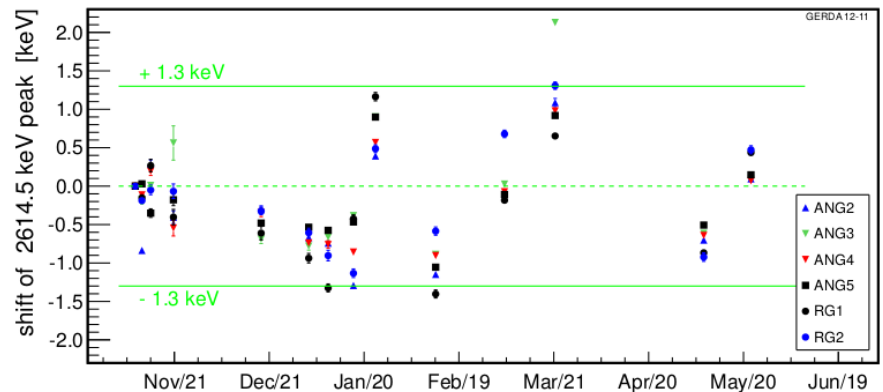


Mass weighted average for FWHM at  $Q_{\beta\beta} = 4.5$  keV

## Stability of the resolution

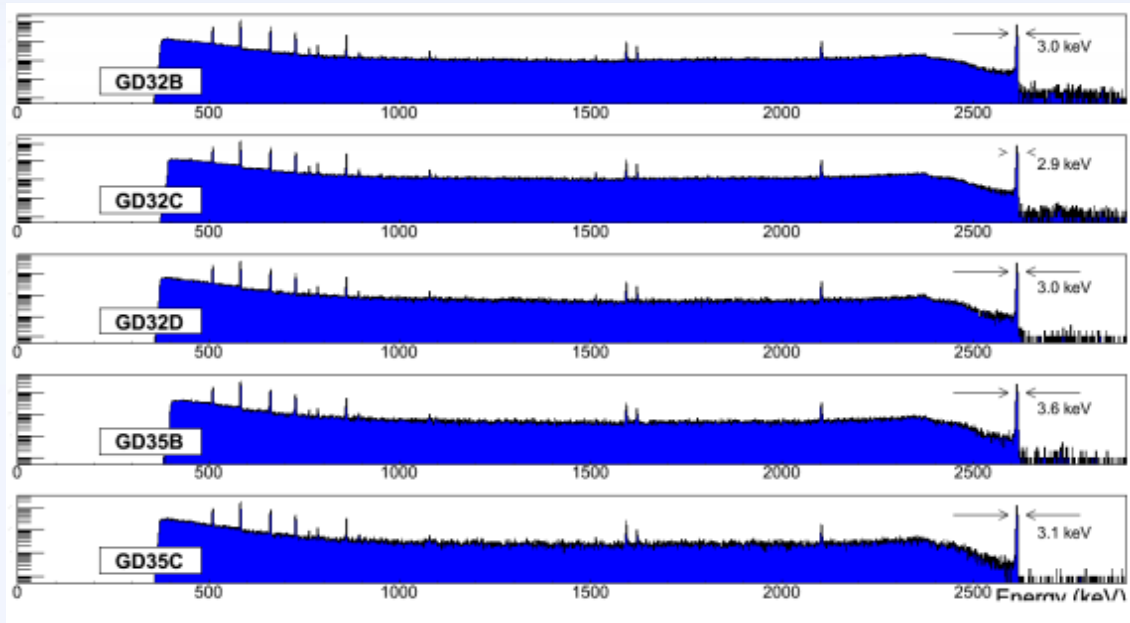


## Stability of the energy scale



# First BEGe's in GERDA

## Calibration spectra



Energy (keV)

3.0 keV

2.9 keV

3.0 keV

3.6 keV

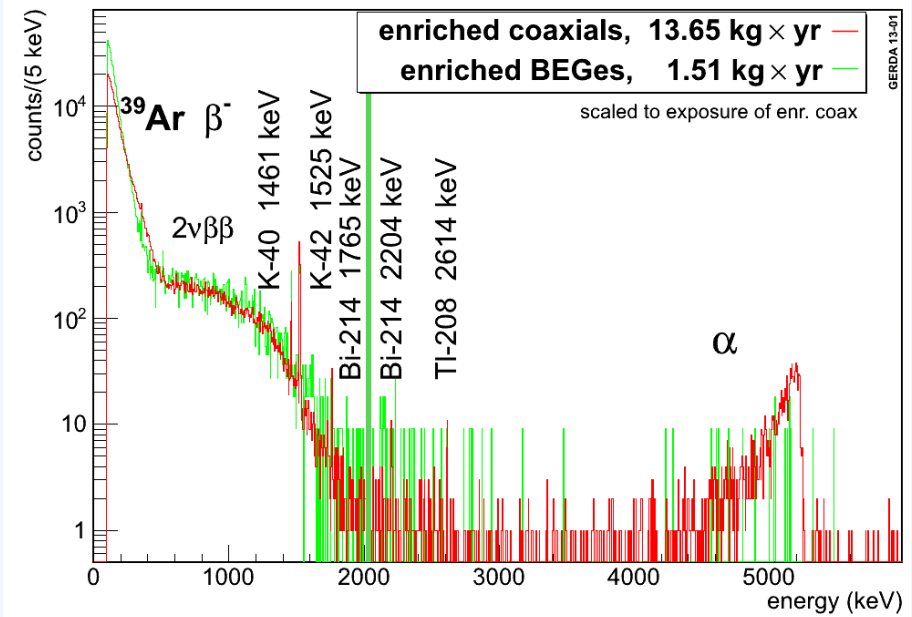
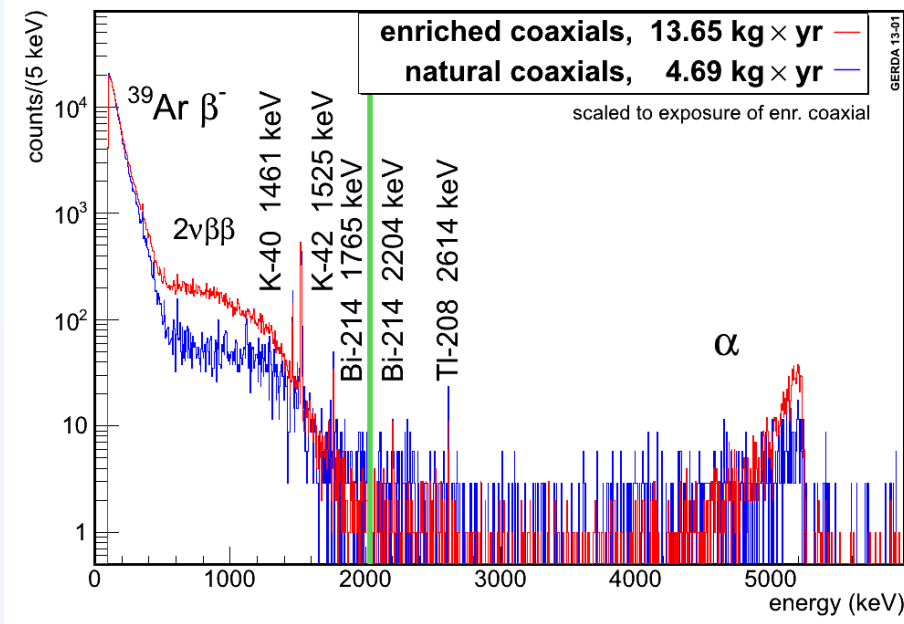
3.1 keV

## Energy resolution and PSA properties

Detector	E resolution [keV]	A/E res.	A/E res. HADES
Agamennone (GD32B)	$2.88 \pm 0.02$	1.5%	0.8%
Andromeda (GD32C)	$2.84 \pm 0.02$	1.7%	1.3%
Anubis (GD32D)	$2.96 \pm 0.04$	1.7%	1.6%
Achilles (GD35B)	$3.61 \pm 0.05$	1.9%	0.6%
Aristoteles (GD35C)	$3.09 \pm 0.06$	1.7%	1.7%



# Energy spectra



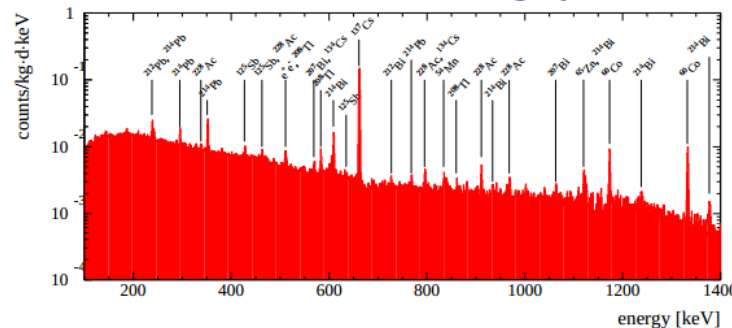
Data blinded between 2019 keV and 2059 keV

# Background lines

isotope	energy [keV]	natGe (3.17 kg·yr)		enrGe (6.10 kg·yr) *		HdM (71.7 kg·yr) rate [cts/(kg·yr)]	Rate HdM, enr coaxial
		tot/bck [cts]	rate [cts/(kg·yr)]	tot/bck [cts]	rate [cts/(kg·yr)]		
<sup>40</sup> K	1460.8	85 / 15	21.7 <sup>+3.4</sup> <sub>-3.0</sub>	125 / 42	13.5 <sup>+2.2</sup> <sub>-2.1</sub>	181 ± 2	13
<sup>60</sup> Co	1173.2	43 / 38	< 5.8	182 / 152	4.8 <sup>+2.8</sup> <sub>-2.8</sub>	55 ± 1	11
	1332.3	31 / 33	< 3.8	93 / 101	< 3.1	51 ± 1	> 48
<sup>137</sup> Cs	661.6	46 / 62	< 3.2	335 / 348	< 5.9	282 ± 2	
<sup>228</sup> Ac	910.8	54 / 38	5.1 <sup>+2.8</sup> <sub>-2.9</sub>	294 / 303	< 5.8	29.8 ± 1.6	
<sup>208</sup> Tl	968.9	64 / 42	6.9 <sup>+3.2</sup> <sub>-3.2</sub>	247 / 230	2.7 <sup>+2.8</sup> <sub>-2.5</sub>	17.6 ± 1.1	11
	583.2	56 / 51	< 6.5	333 / 327	< 7.6	36 ± 3	
<sup>214</sup> Pb	2614.5	9 / 2	2.1 <sup>+1.1</sup> <sub>-1.1</sub>	10 / 0	1.5 <sup>+0.6</sup> <sub>-0.5</sub>	16.5 ± 0.5	11
	352	740 / 630	34.1 <sup>+12.4</sup> <sub>-11.0</sub>	1770 / 1688	12.5 <sup>+9.5</sup> <sub>-7.7</sub>	138.7 ± 4.8	11
<sup>214</sup> Bi	609.3	99 / 51	15.1 <sup>+3.9</sup> <sub>-3.9</sub>	351 / 311	6.8 <sup>+3.7</sup> <sub>-4.1</sub>	105 ± 1	~10
	1120.3	71 / 44	8.4 <sup>+3.5</sup> <sub>-3.3</sub>	194 / 186	< 6.1	26.9 ± 1.2	
	1764.5	23 / 5	5.4 <sup>+1.9</sup> <sub>-1.5</sub>	24 / 1	3.6 <sup>+0.9</sup> <sub>-0.8</sub>	30.7 ± 0.7	
	2204.2	5 / 2	0.8 <sup>+0.8</sup> <sub>-0.7</sub>	6 / 3	0.4 <sup>+0.4</sup> <sub>-0.4</sub>	8.1 ± 0.5	

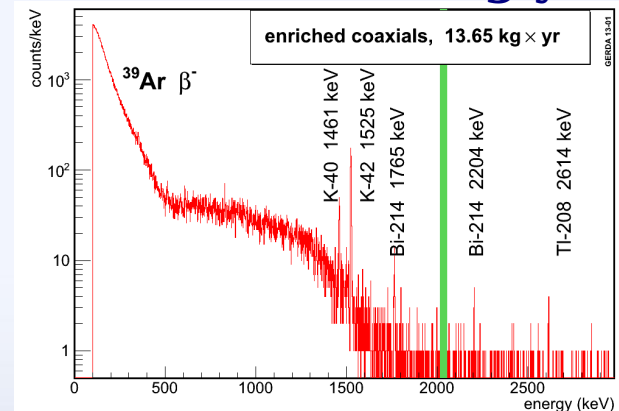
\* only coaxial detectors

## HdM 47.4 kg·yr



taken from *Eur. Phys. J. A12*, 147-154 (2001)

## GERDA 13.65 kg·yr



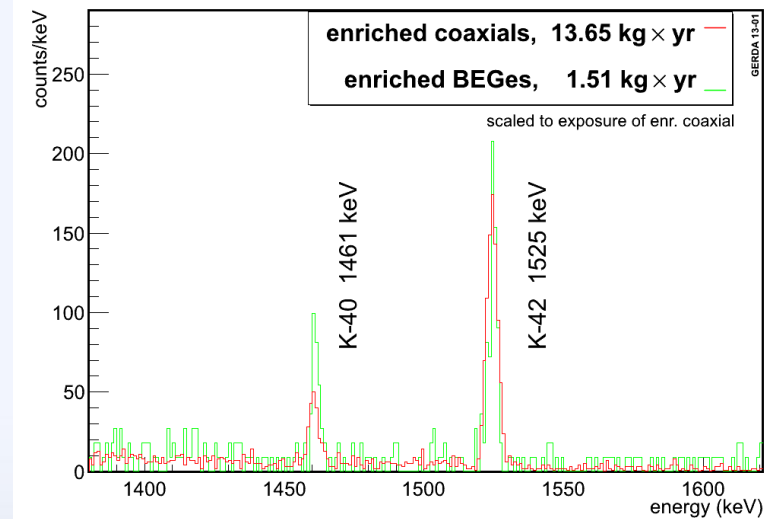
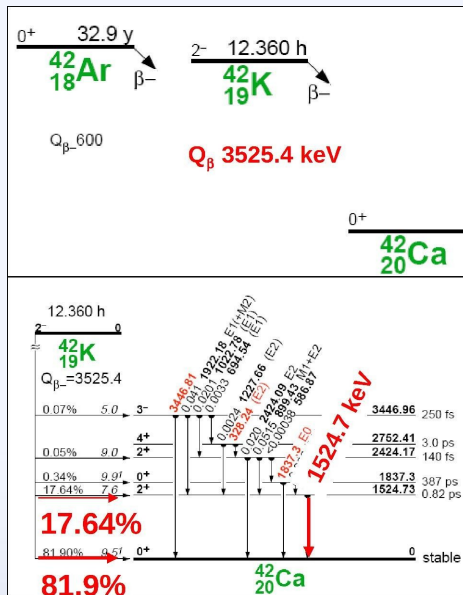
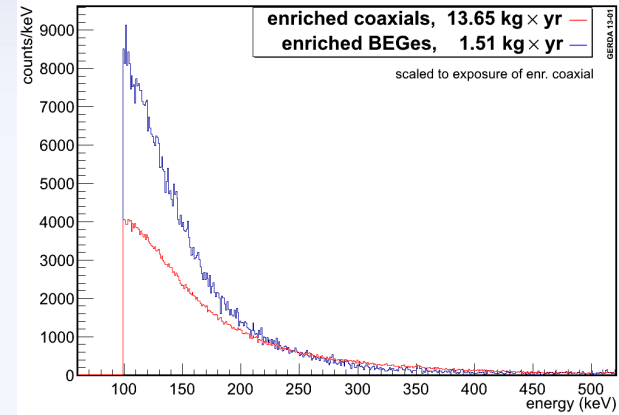
# Background from Argon

## <sup>39</sup>Ar

- 1.01 Bq/kg,  $T_{1/2} = 269$  yr
- pure  $\beta$  emitter, Q-value=565 keV  
→ **below region of interest**

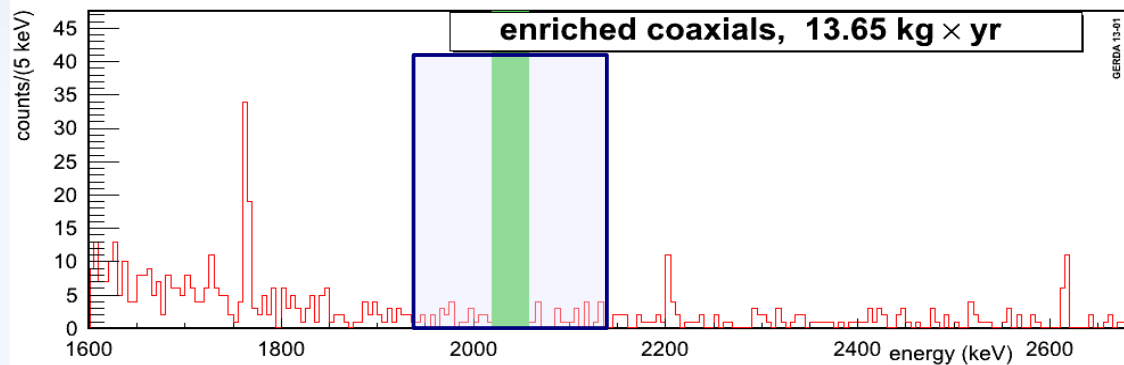
## <sup>42</sup>Ar

**GERDA proposal:**  $^{42}\text{Ar}/^{\text{nat}}\text{Ar} < 3 \times 10^{-21}$  (*Barabash et al. 2002*)  
**GERDA measurement:** Count rate at 1525 keV  $\sim 2$  times expectation



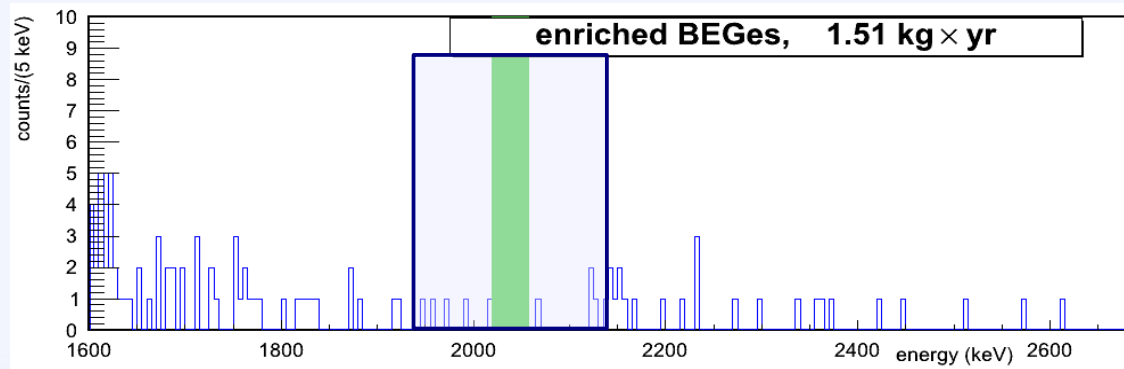
# Region of Interest

Background rate in ROI ( $Q_{\beta\beta} \pm 100$  keV, **blinded window excluded**)



**Enriched coaxials:  $0.022 \pm 0.003$  cts/(keV·kg·yr)**  
( $0.017 \pm 0.003$  cts/(keV·kg·yr) excluding 1.30 kg·yr period with higher background following detector substitutions in July)

➔ **factor ~6-8 lower than previous experiments (HdM, IGEX)**

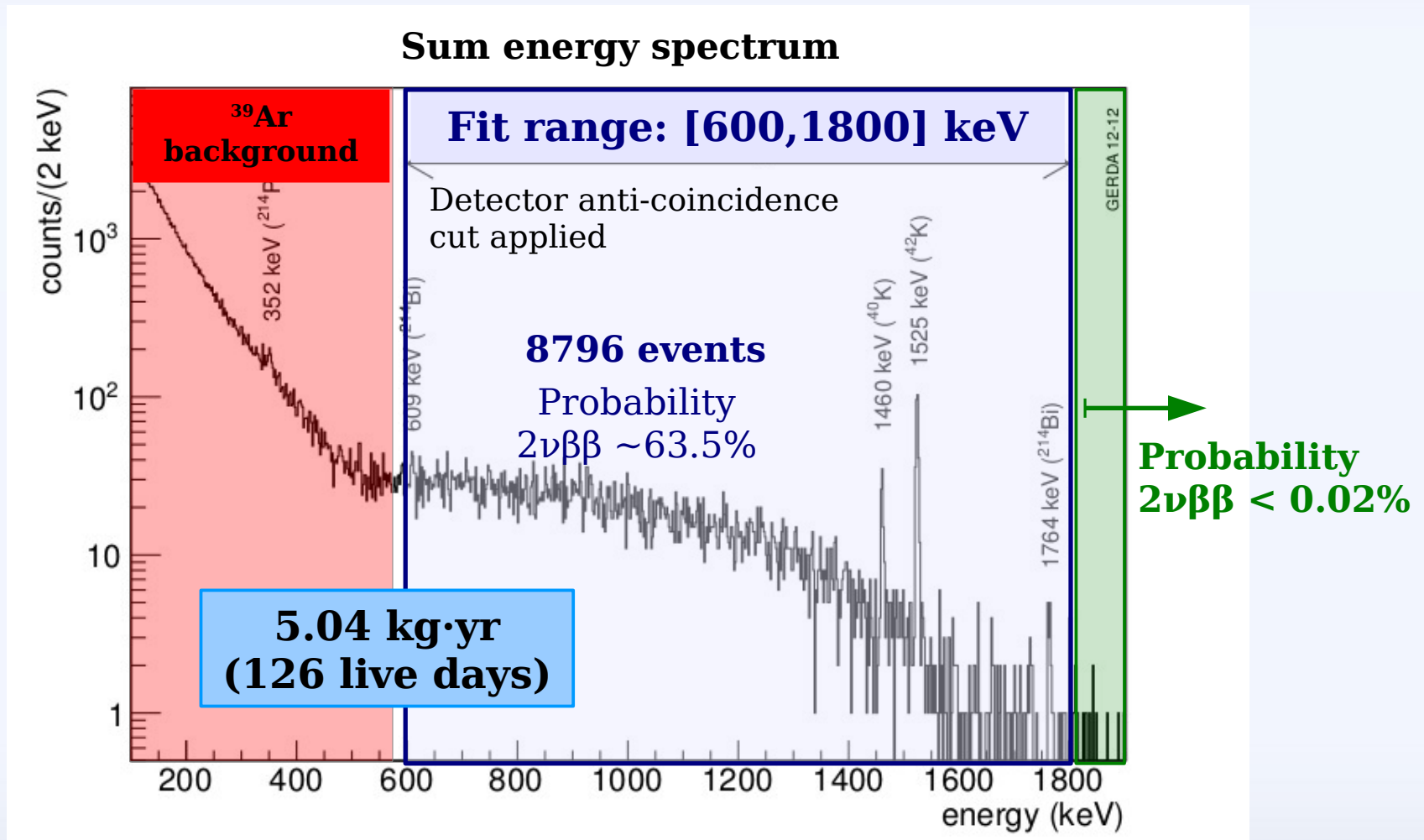


**Enriched BEGes:  $0.041^{+0.015}_{-0.012}$  cts/(keV·kg·yr)**

(No pulse shape analysis applied yet)

# Measurement of $T_{1/2}^{2\nu}$ : Data set

First 126 live days of the 6  $^{enr}\text{Ge}$  detectors:





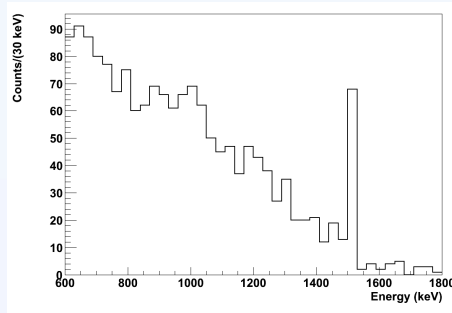
# Measurement of $T^{2\nu}_{1/2}$ : Fit

## Binned maximum likelihood approach

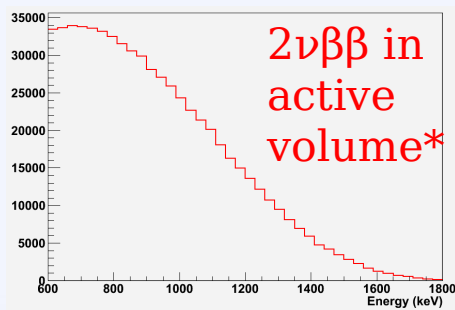
### Ingredients:

- 6 energy spectra from  $^{\text{enr}}\text{Ge}$  detectors (30 keV bins)

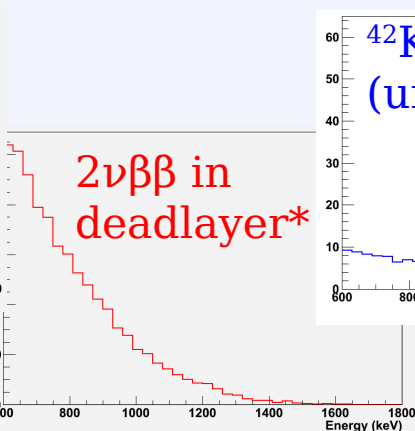
Single detector energy spectrum:



- Model: Simulated spectra of  $2\nu\beta\beta$ ,  $^{42}\text{K}$ ,  $^{40}\text{K}$ ,  $^{214}\text{Bi}$  for each detector

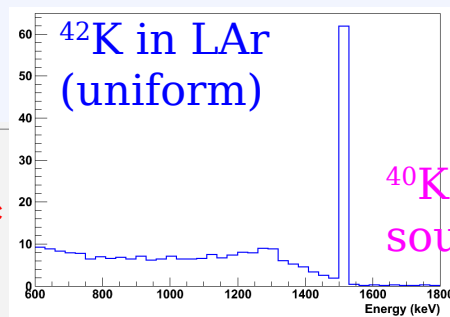


$2\nu\beta\beta$  in active volume\*

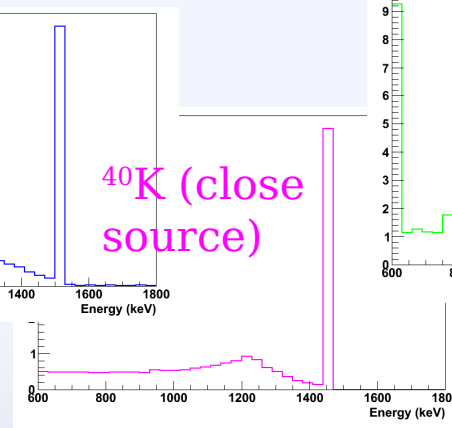


$2\nu\beta\beta$  in deadlayer\*

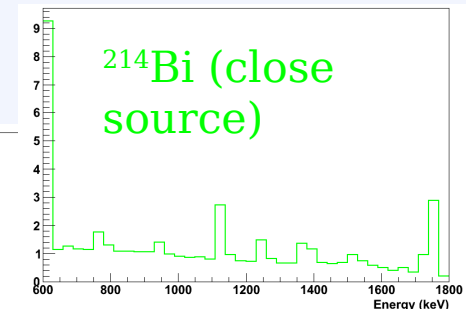
\*Distribution from DECAY0 (V. I. Tretyak)



$^{42}\text{K}$  in LAr (uniform)



$^{40}\text{K}$  (close source)



$^{214}\text{Bi}$  (close source)

# Measurement of $T_{1/2}^{2\nu}$ : Fit

## Binned maximum likelihood approach

### Ingredients:

- Information on active masses and enrichment fractions:

detector	total mass (g)	active mass (g)	$^{76}\text{Ge}$ isotopic abundance (%)
ANG2	2833	$2468 \pm 121 \pm 89$	$86.6 \pm 2.5$
ANG3	2391	$2070 \pm 118 \pm 77$	$88.3 \pm 2.6$
ANG4	2372	$2136 \pm 116 \pm 79$	$86.3 \pm 1.3$
ANG5	2746	$2281 \pm 109 \pm 82$	$85.6 \pm 1.3$
RG1	2110	$1908 \pm 109 \pm 72$	$85.5 \pm 2.0$
RG2	2166	$1800 \pm 99 \pm 65$	$85.5 \pm 2.0$

uncorrelated      correlated

**Average active mass fraction:  $(86.7 \pm 4.6(\text{uncorr.}) \pm 3.2(\text{corr.}))\%$**

**Average enrichment fraction:  $(86.3 \pm 2)\%$**

# Measurement of $T_{1/2}^{2\nu}$ : Fit

## Binned maximum likelihood approach

### Tool:

- Bayesian Analysis Toolkit BAT  
*Caldwell A., Kollar D., and Kröninger K. 2009 Comput. Phys. Comm. 180 2197*

### Directions:

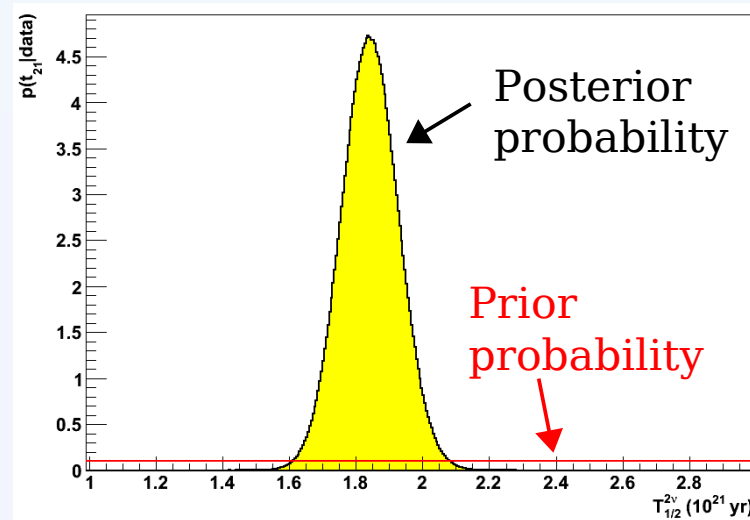
- Define the parameters:
  - Active detector masses (6+1)
  - Fraction of enrichment in  $^{76}\text{Ge}$  (6)
  - Background contributions (3x6)
  - $T_{1/2}^{2\nu}$  common to all detectors (1)
- Run the fit
- Integrate over all nuisance parameters to derive posterior for  $T_{1/2}^{2\nu}$

} nuisance parameters

# Measurement of $T_{1/2}^{2\nu}$ : Fit result

$$T_{1/2}^{2\nu} = (1.84^{+0.09}_{-0.08}) \cdot 10^{21} \text{ yr (smallest interval 68\%)}$$

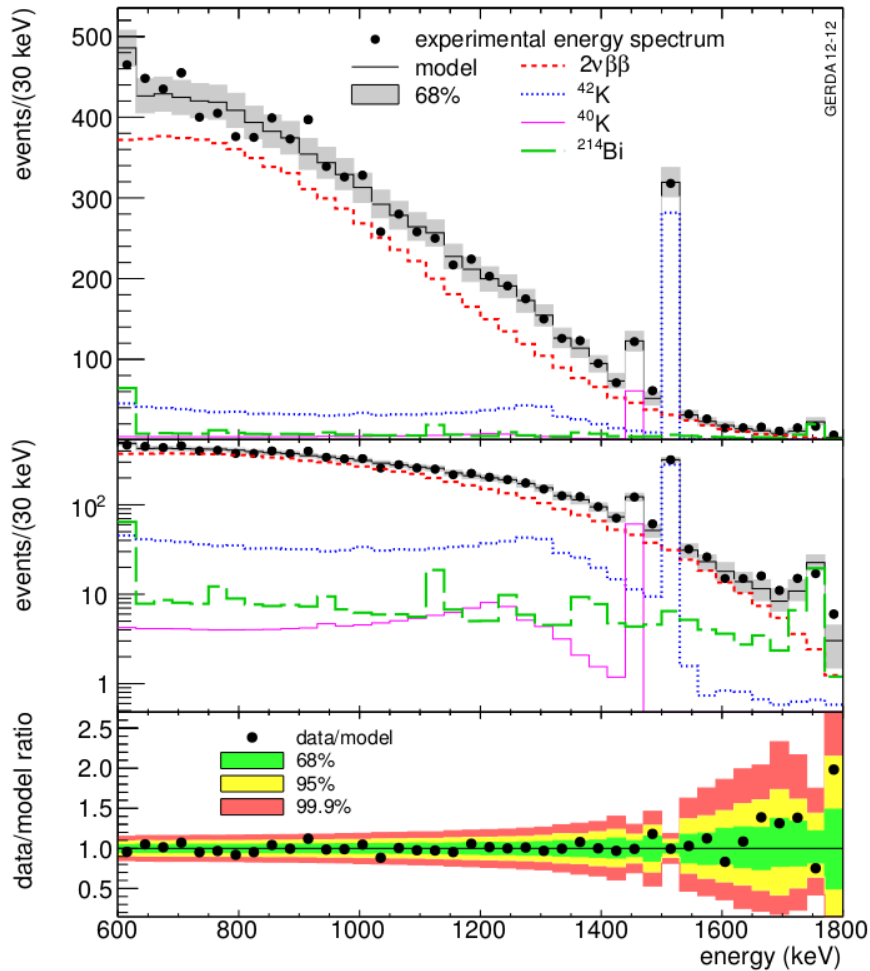
Uncertainty includes uncertainties on nuisance parameters, especially on active masses and enrichment fractions



## Crosscheck:

Fit each detector separately  $\rightarrow$  results mutually consistent ( $\chi^2/\nu=3.02/5$ )

# Measurement of $T_{1/2}^{2\nu}$ : Fit result



**Data: 8796 events**

**Best fit model: 8797.0 events**

**$2\nu\beta\beta$ : 80%**

**$^{42}\text{K}$ : 14%**

**$^{214}\text{Bi}$ : 4%**

**$^{40}\text{K}$ : 2%**

**Signal to background:  
4:1**

**p-value: 0.77**

# Measurement of $T_{1/2}^{2\nu}$ : Systematics

$^{60}\text{Co}$ ,  $^{228}\text{Ac}$ ,  $^{208}\text{Tl}$  ???

Source positions

Decay distribution model

Dimensions, materials, ...

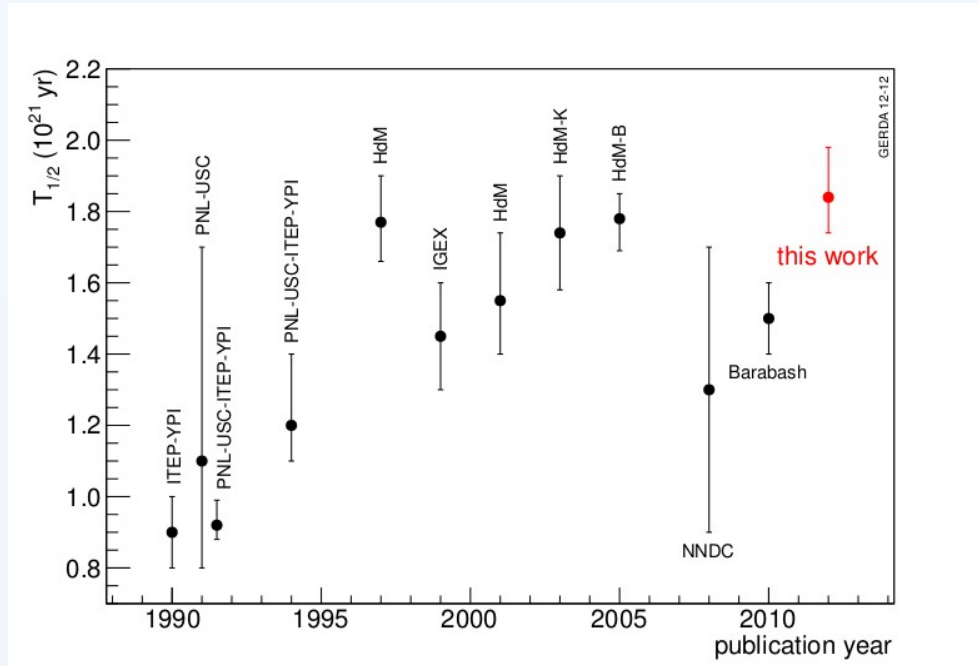
Validation of GEANT4 processes

Source of uncertainty	Uncertainty on $T_{1/2}^{2\nu}$ [%]
Non-identified background components	+5.3
Energy spectra from $^{42}\text{K}$ , $^{40}\text{K}$ , $^{214}\text{Bi}$	$\pm 2.1$
Shape of the $2\nu\beta\beta$ decay spectrum	$\pm 1$
Precision of the Monte Carlo geometry model	$\pm 1$
Accuracy of the Monte Carlo tracking	$\pm 2$
Data acquisition and selection	$\pm 0.5$
<b>Total</b>	<b>+6.2</b>
	<b>-3.3</b>

# Measurement of $T_{1/2}^{2\nu}$ : Result

Final result:

$$(1.84^{+0.09}_{-0.08 \text{ fit}} \text{ } ^{+0.11}_{-0.06 \text{ syst}}) \cdot 10^{21} \text{ yr} = (1.84^{+0.14}_{-0.10}) \cdot 10^{21} \text{ yr}$$



- Superior signal-to-background ratio  
→ uncertainty comparable to previous measurements despite much smaller exposure
- Good agreement with re-analysis of HdM data  
*HdM-K: Nucl. Instrum. Methods A 513, 596 (2003)*  
*HdM-B: Phys. Part. Nucl. Lett. 2, 77 / Pisma Fiz. Elem. Chast. Atom. Yadra 2, 21 (2005)*

# Measurement of $T_{1/2}^{2\nu}$ : Matrix element

Calculate  $M^{2\nu}$

(with phase space factor from [1]):

$$0.133^{+0.004}_{-0.005} \text{ MeV}^{-1}$$

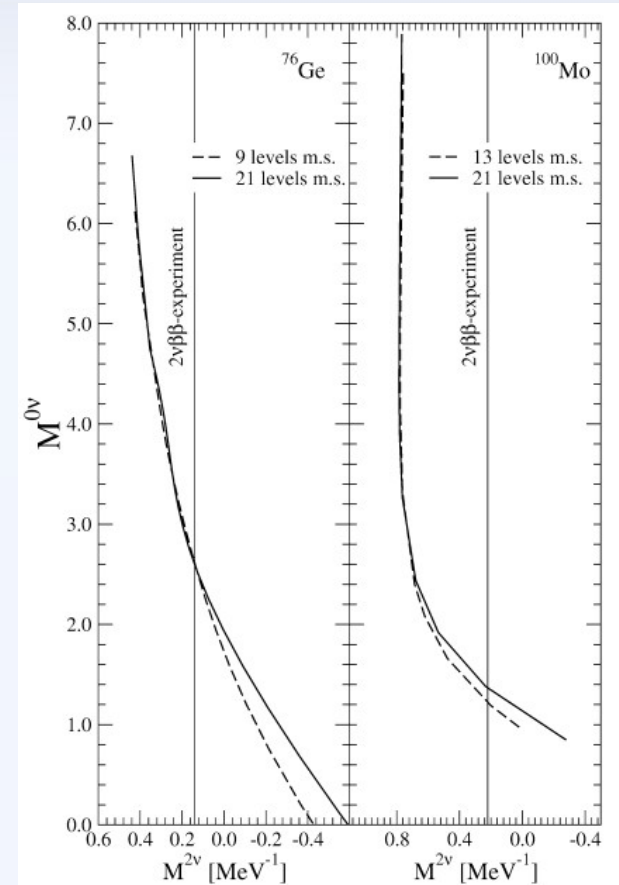
→ decrease by 11% compared to [2]

→ well consistent with  $M^{2\nu}$  derived from (d, $^2\text{He}$ ) and ( $^3\text{He}$ ,t) charge exchange reactions

**Relation between  $M^{2\nu}$  and  $M^{0\nu}$  [2]:**

→ **Decreasing  $M^{2\nu}$  → decreasing  $M^{0\nu}$**

→ **Increase of predicted  $T_{1/2}^{0\nu}$  by 15%**



Relation between  $M^{2\nu}$  and  $M^{0\nu}$   
(taken from [2])

[1] Kotila J., and Iachello F. 2012 Phys. Rev. C 85 034316

[2] Rodin V. A., Fässler A., Šimkovic F. and Vogel P. 2006 Nucl. Phys. A 766 107;  
erratum: 793 (2007) 213

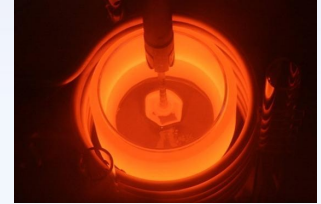


# Phase II detector production

Production of 30 new  $^{enr}\text{Ge}$  BEGe detectors (~20 kg):



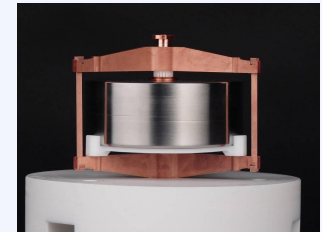
Crystal pulling at Canberra: Oak Ridge, TN, USA



Transport to Mol, Belgium



BEGe detector diode production: Olen, Belgium



BEGe acceptance tests: Hades, Belgium  
(Tests of detector performance and properties)



Electrical contacting : Olen, Belgium



**Phase II detectors almost ready!**

# Liquid Argon readout

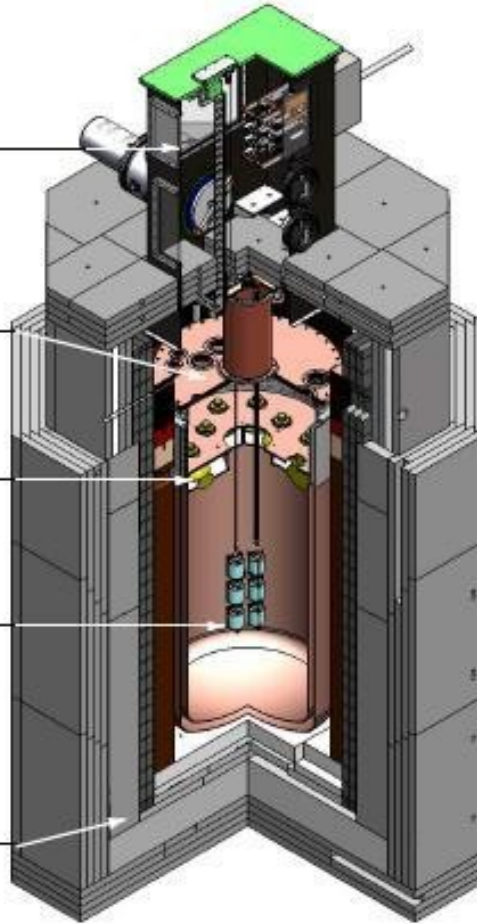
**lock**  
for Ge-detector deployment

**copper cryostat**  
inner  $\varnothing = 90$  cm, height = 205 cm  
LAr volume = 1 m<sup>3</sup> (1.4 t)  
coated with WLS mirror foil

**PMTs**  
9  $\times$  8" ETL 9357  
coated with WLS

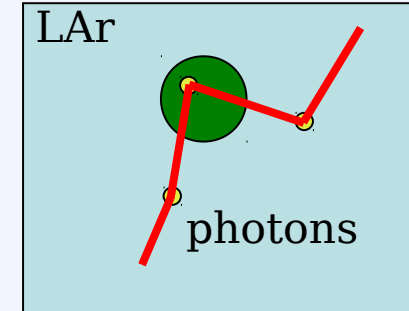
**detector strings**  
up to 3 strings  
(9 Ge-detectors)

**graded shield**  
15 cm copper  
10 cm lead  
23 cm steel  
20 cm polyethylene

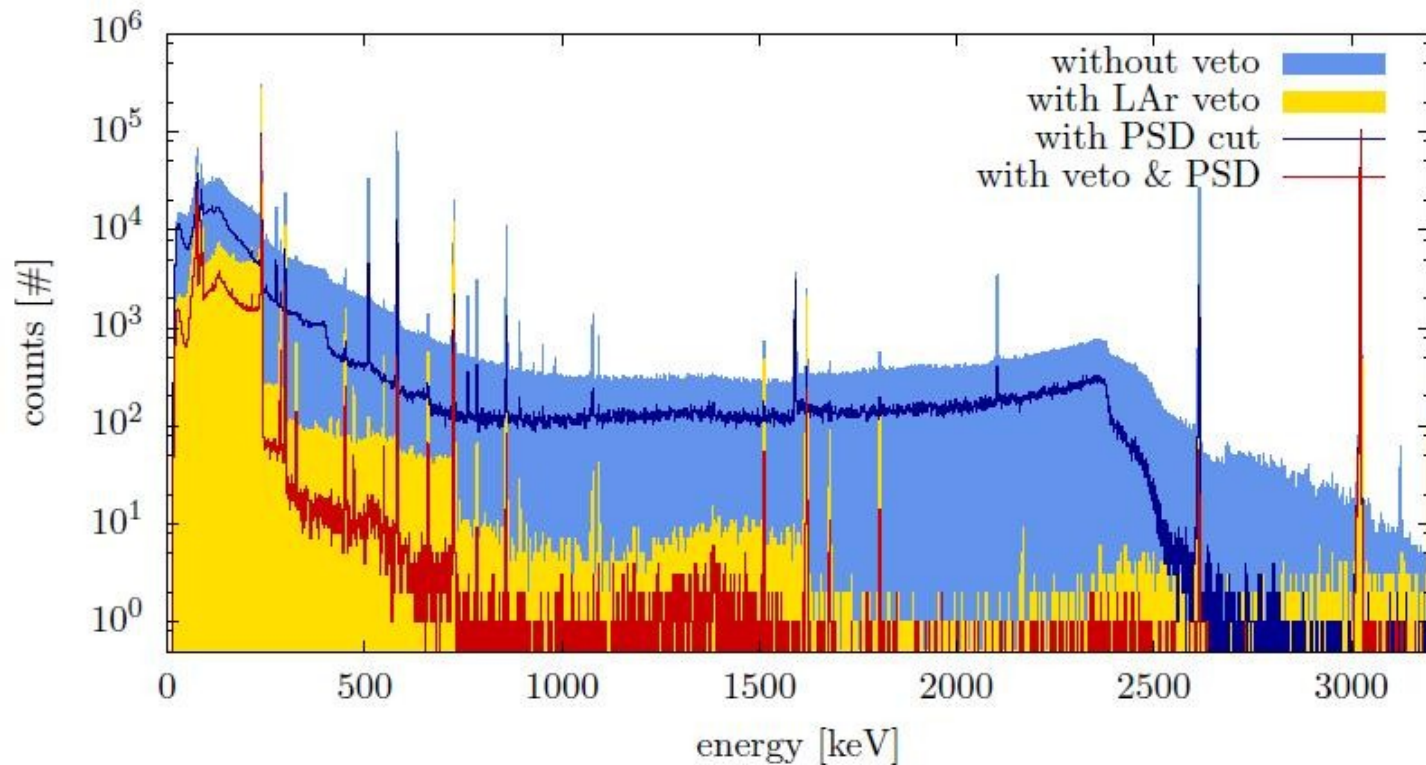


## LArGe: Low-background test facility at LNGS

Detection of coincident LAr scintillation light to discriminate background



# Phase II background reduction



## Operation of Phase II proto-type detector in LArGe:

Combining PSD of BEGe detector and LAr veto:

Measured suppression factor for a  $^{228}\text{Th}$  source at  $Q_{\beta\beta}$ :  $\sim 5 \cdot 10^3$

Also other designs using SiPM's investigated

→ **avoid additional background!**

# Conclusions

- GERDA Phase I collected about 15 kg·yr of enr data so far
- Background in ROI about factor 1/10 compared to previous experiments
- First result using 5 kg·yr of data:  $T_{1/2}^{2\nu} = (1.84^{+0.14}_{-0.10}) \cdot 10^{21} \text{ yr}$   
*The GERDA Collaboration, J. Phys. G: Nucl. Part. Phys. 40 (2013) 035110*
- Phase II detectors produced and tested

**2013 will be very exciting year:**

- **Unblinding**
- **$T_{1/2}^{0\nu}$  results**
- **Transition to Phase II**

# **Additional material**

# HdM claim

Data acquisition and analysis of the  $^{76}\text{Ge}$  double beta experiment in Gran Sasso 1990-2003

H.V. Klapdor-Kleingrothaus<sup>\*1</sup>, A. Dietz, I.V. Krivosheina<sup>2</sup>, O. Chkvorots

*Phys. Lett. B 586, 198-212 (2004)*

BI 0.11 cts / (keV·kg·yr)

$28.75 \pm 6.87$  events (bgd: ~60)

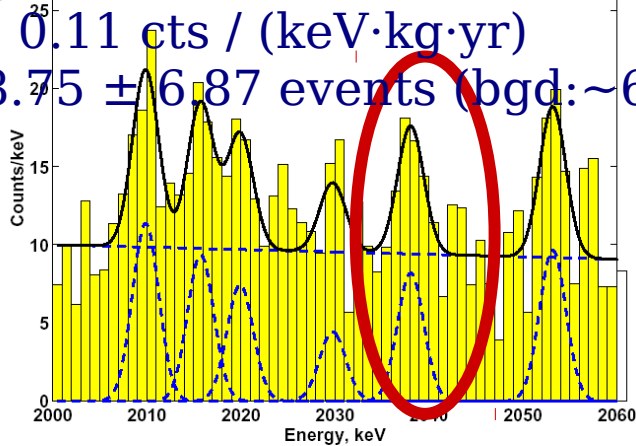


Fig. 17. The total sum spectrum of all five detectors (in total 10.96 kg enriched in  $^{76}\text{Ge}$ ), for the period November 1990–May 2003 (71.7 kg year) in the range 2000–2060 keV and its fit (see Section 3.2).

- Nov 1990 - May 2003
- 71.7 kg·yr
- $4.2\sigma/6\sigma$  evidence for  $0\nu\beta\beta$

- $(0.69 - 4.18) \cdot 10^{25}$  yr ( $3\sigma$ )

Best fit:  $1.19 \cdot 10^{25}$  yr

*Phys. Lett. B 586, 198-212 (2004)*

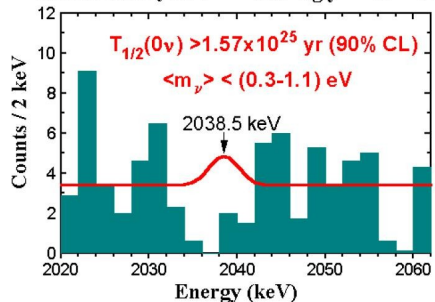
$2.23^{+0.44}_{-0.31} \cdot 10^{25}$  yr

*Mod. Phys. Lett. A 21, 1547-1566 (2006)*

Criticism in *arXiv: 1210.7432*

## Comparison: IGEX

116.75 mole.years - 8.87 kg.y in  $^{76}\text{Ge}$



- $m_{\beta\beta} = (0.24-0.58)$  eV  
(best fit 0.44 eV) /  
 $0.32 \pm 0.03$  eV

**Note: statistical significance depends on background model!**

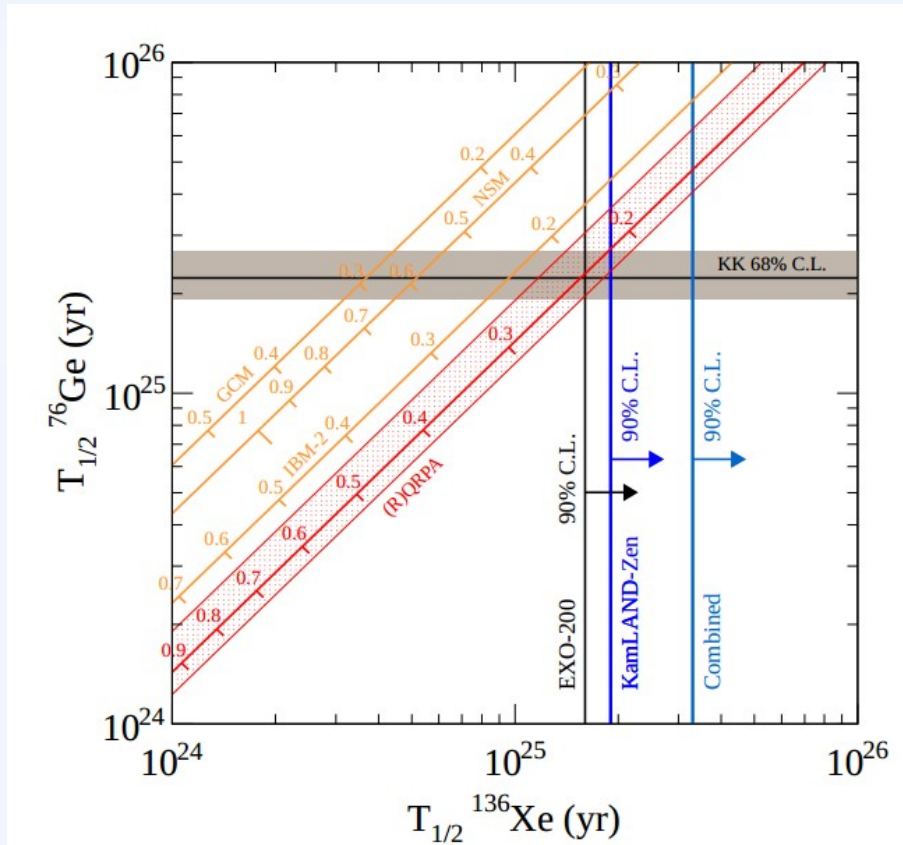
# Detector parameters

detector	total mass (g)	active mass (g)	<sup>76</sup> Ge isotopic abundance (%)	$T_{1/2}^{2\nu}$ ( $10^{21}$ yr)
ANG2	2833	2468±121±89	86.6±2.5	1.99 <sup>+0.14</sup> <sub>-0.15</sub>
ANG3	2391	2070±118±77	88.3±2.6	1.69 <sup>+0.15</sup> <sub>-0.14</sub>
ANG4	2372	2136±116±79	86.3±1.3	1.94 <sup>+0.14</sup> <sub>-0.15</sub>
ANG5	2746	2281±109±82	85.6±1.3	1.79 <sup>+0.12</sup> <sub>-0.14</sub>
RG1	2110	1908±109±72	85.5±2.0	1.94 <sup>+0.18</sup> <sub>-0.14</sub>
RG2	2166	1800±99±65	85.5±2.0	1.93 <sup>+0.16</sup> <sub>-0.16</sub>

$\chi^2/\nu=3.02/5$

# KamLAND-Zen and EXO-200

From arXiv:1211.3863v1 [hep-ex] (2012):



**“Combined result for  $^{136}\text{Xe}$  refutes the  $0\nu\beta\beta$  detection claim in  $^{76}\text{Ge}$  at  $> 97.5\%$ ”**

Comparison depends on matrix calculations  
→ GERDA will check claim independently from calculations



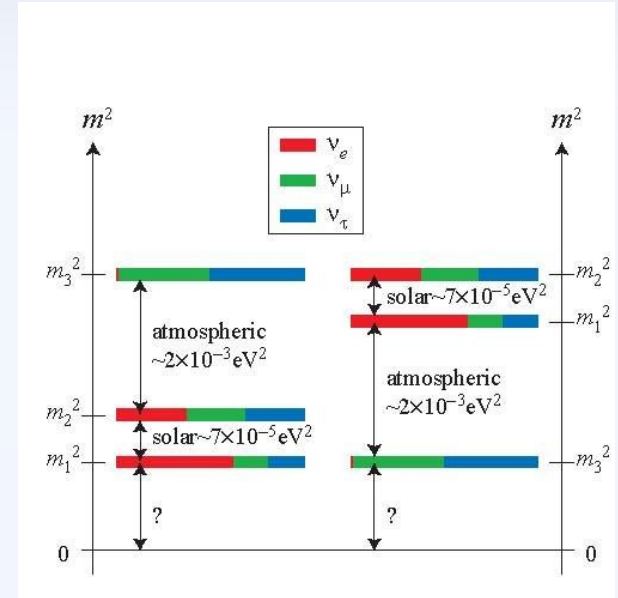
# Neutrino properties

## Neutrino Mixings

Weakly interacting and mass eigenstates are independent basis

$$\begin{bmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \\ |\nu_\tau\rangle \end{bmatrix} = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{bmatrix} \begin{bmatrix} |m_1\rangle \\ |m_2\rangle \\ |m_3\rangle \end{bmatrix}$$

$$U_{\nu i} = \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta_{13}} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta_{13}} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta_{13}} & s_{23}c_{13} \\ s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta_{13}} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta_{13}} & c_{23}c_{13} \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 \\ 0 & e^{i\frac{\alpha_{21}}{2}} & 0 \\ 0 & 0 & e^{i\frac{\alpha_{31}}{2}} \end{bmatrix}$$



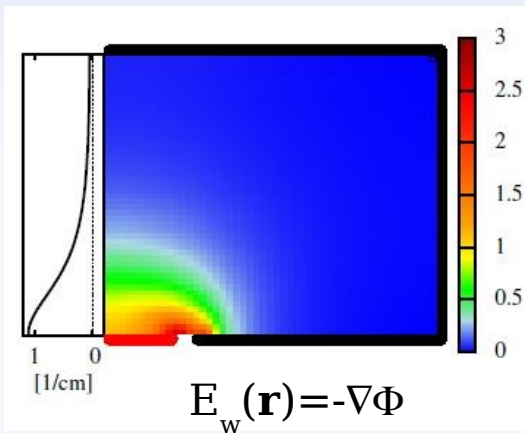
## What we know:

- $m_2^2 - m_1^2 = \Delta m_{\text{sun}}^2$
- $m_2^2 - m_1^2 = \Delta m_{\text{atm}}^2$
- $\theta_{12} = \theta_{\text{sun}}$
- $\theta_{23} = \theta_{\text{atm}}$
- $\theta_{13}$

## What we do not know:

- Absolute mass scale
- Mass hierarchy
- Phases ( $\delta_{13}$ ,  $\alpha_{21}$ ,  $\alpha_{31}$ )
- Nature of the neutrino mass (Dirac or Majorana)

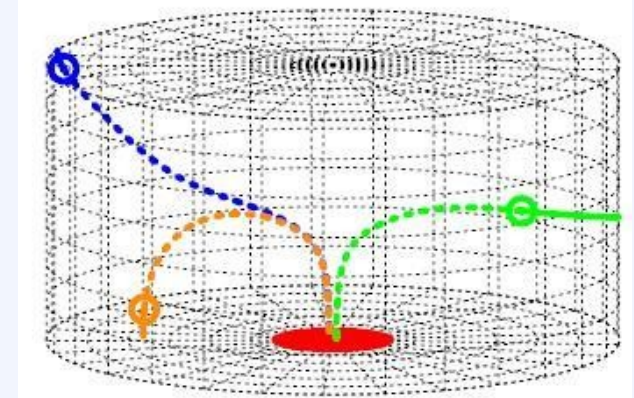
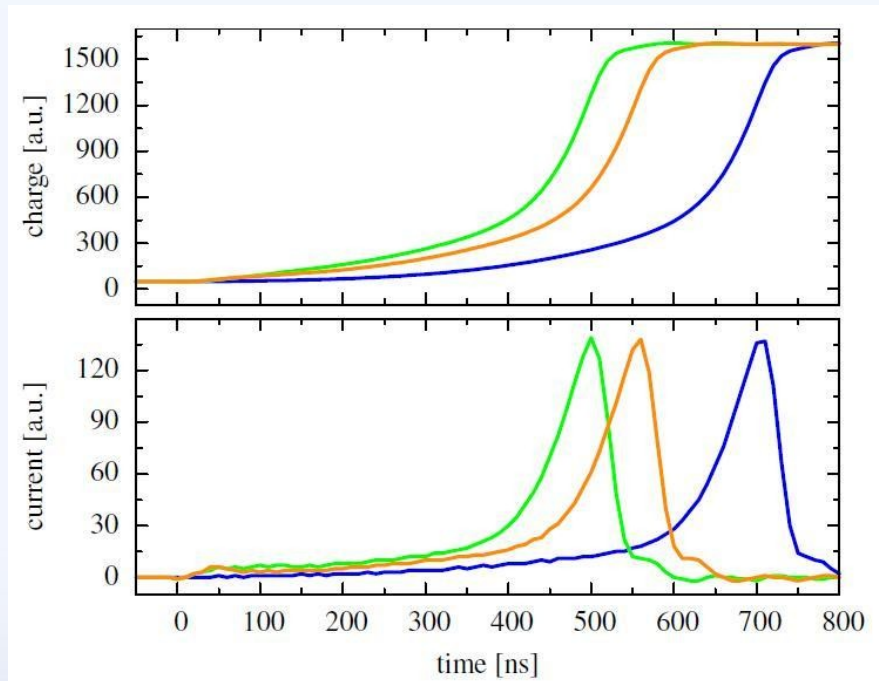
# Phase II detectors



## Shockley-Ramo Theorem:

$$Q(t) = -q \cdot \Phi_w(\mathbf{r}(t))$$

- ..... anode
- cathode
- electrons
- - - holes
- ⊙ interaction point



# Background from Argon

## Add a mini-shroud (MS):

