



# Smith-Purcell radiation as a longitudinal profile beam diagnostic

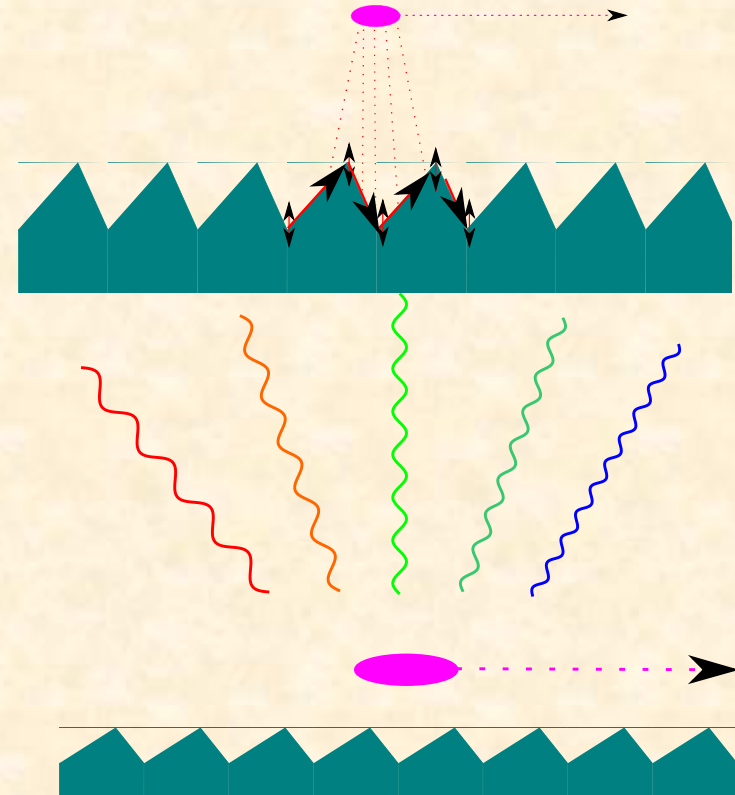
Nicolas Delerue  
with the E-203 & SPESO collaborations



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University of Oxford, John Fell Fund and by the ANR (France) under contract ANR-12-JS05-0003-01.*

# Coherent Smith-Purcell radiation

- An electron beam passing above a grating can induce a surface current on the grating.
- This will lead to the emission of radiation (called Smith-Purcell radiation).
- For bunches shorter than the emission wavelength this emission is coherent.



S.J. Smith and E.M. Purcell, Phys. Rev. **92**, pg. 1069, (1953)

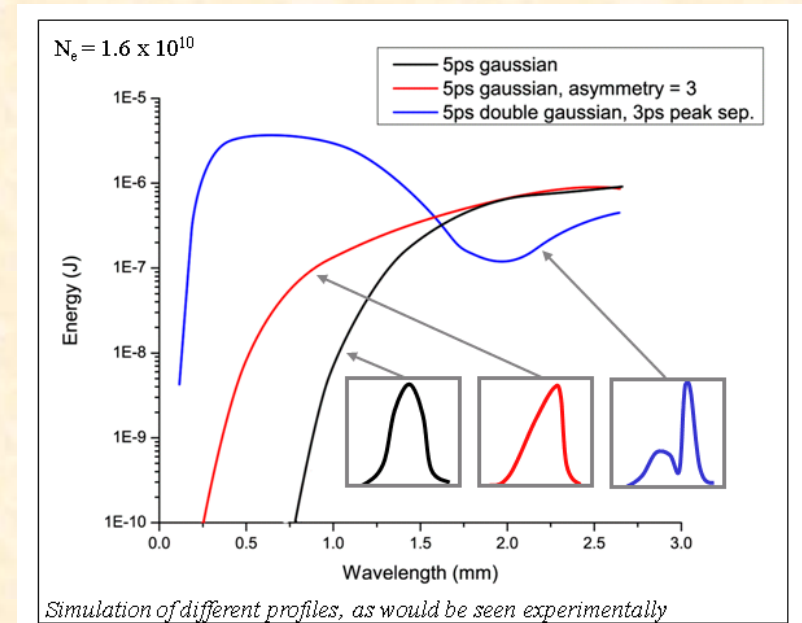
Ishiguro and Tako, Optica Acta (GB) 8 1961 25

$$\left( \frac{dI}{d\Omega d\omega} \right)_{N_e}(\Omega, \omega) \approx \left( \frac{dI}{d\Omega d\omega} \right)_{sp}(\Omega, \omega) \cdot [N_e + N_e(N_e - 1) |F(\omega)|^2]$$

# Coherent SPR as a longitudinal profile diagnostic

- Coherent SPR encodes the Fourier transform of the longitudinal profile.
- => It can be used as a single shot non destructive diagnostic.  
=> allows separate measurement of signal and background.
- Such diagnostic requires a measurement of the CSPR spectrum and then “FT inversion” to recover the spectrum (currently done using the Kramers-Kronig relation).

$$\left( \frac{dI}{d\Omega d\omega} \right)_{N_e} (\Omega, \omega) = \left( \frac{dI}{d\Omega d\omega} \right)_{sp} (\Omega, \omega) \cdot [N_e + N_e(N_e + 1) |F(\omega)|^2]$$



Phys. Rev. ST Accel. Beams 12, 032803 (2009)

# Comparison between CSPR and Transition radiation

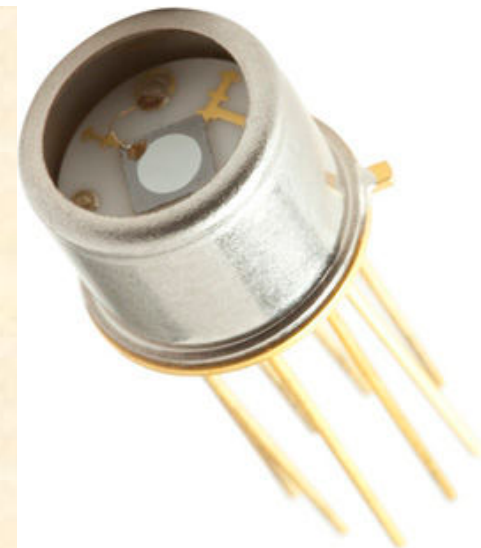
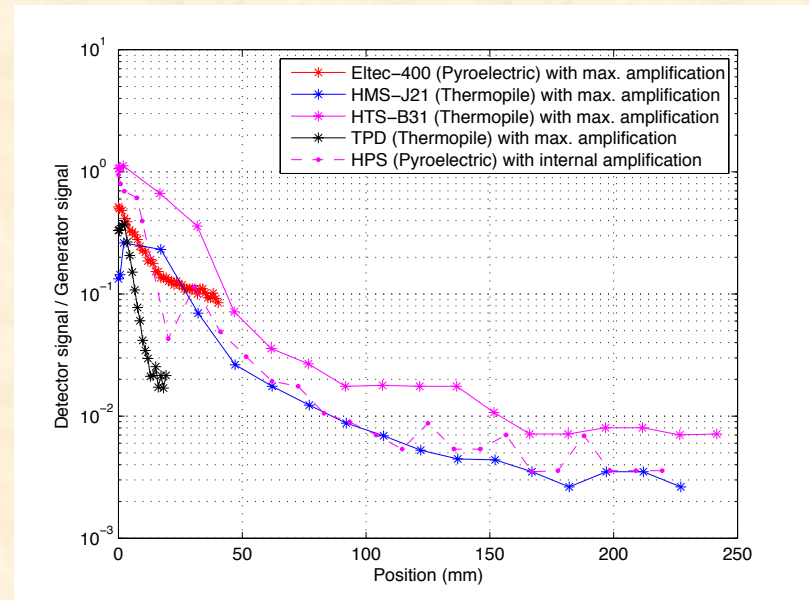
$$\left( \frac{dI}{d\Omega d\omega} \right)_{N_e}(\Omega, \omega) = \left( \frac{dI}{d\Omega d\omega} \right)_{sp}(\Omega, \omega) \cdot [N_e + N_e(N_e + 1) |F(\omega)|^2]$$

$$\left( \frac{dI}{d\Omega} \right)_{sp} = 2\pi q^2 \frac{Z}{\ell^2} \frac{n^2 \beta^3}{(1 - \beta \cos \theta)^3} R^2 \exp\left(-\frac{2x_0}{\lambda_e}\right)$$

- In both case there is a dependence on the square of the bunch charge.
- For CSPR, dependence on the bunch-grating separation (and grating pitch).
- In CSPR the radiation is dispersed spectrally  
=> it is possible to separate CSPR from other backgrounds (independent background subtraction).

# Wavelength and detectors

- This diagnostic is being developed for classical and novel accelerators.
- For a ps beam the CSPR is in the mm range.
- For 100fs beam the CSPR is in the far IR/THz range.
- We use pyroelectric detectors: good wavelength coverage but low responsivity (could use other detectors where available).



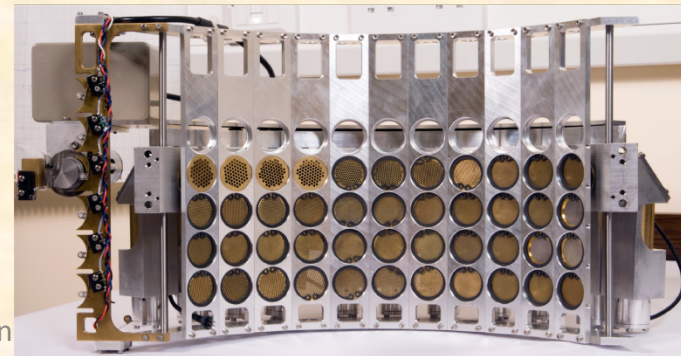
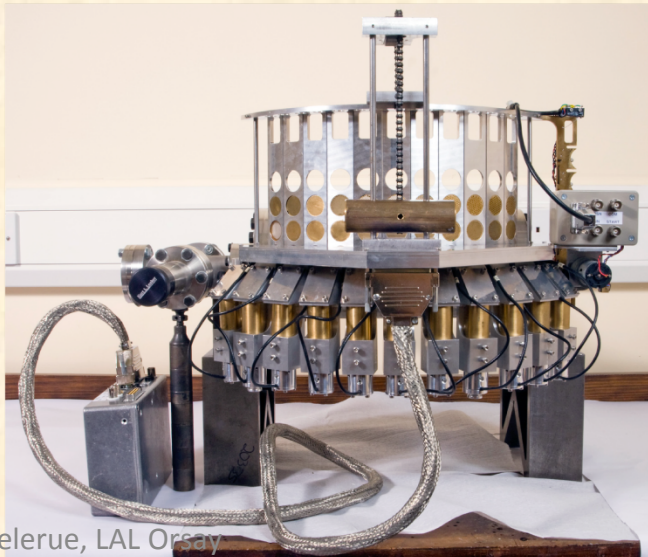
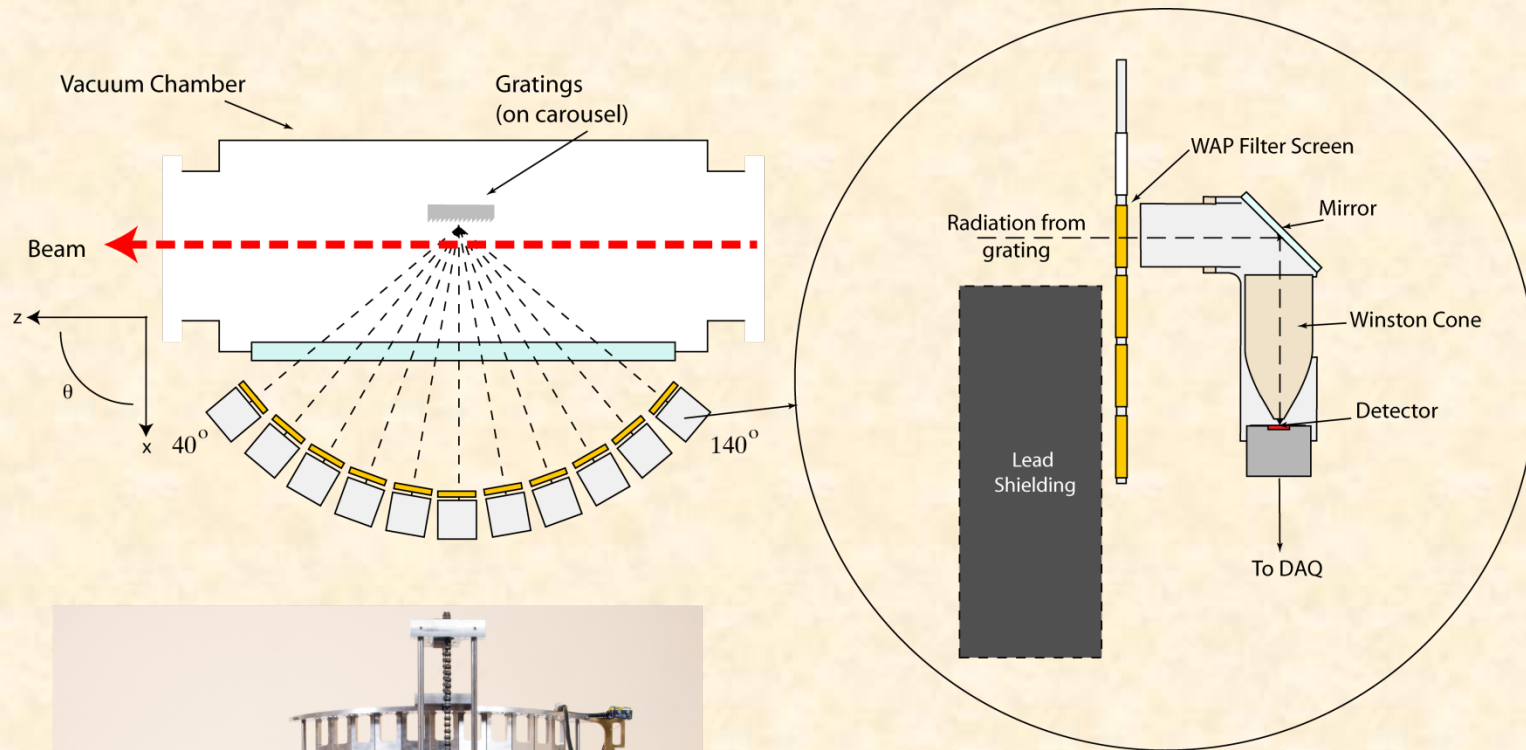


# Current R&D toward CSPR as a bunch length diagnostic

There are 2 collaborations working on using CSPR as a single shot bunch length diagnostics in the fs range:

- E-203 (Oxford, SLAC et al.): Experiment at FACET to demonstrate CSPR as a diagnostic in the fs-range using hardware previously validated during tests at End Station B.
- SPESO (LAL, SOLEIL et al.): Experiment at SOLEIL to map the distribution of CSPR and test various detection strategies.

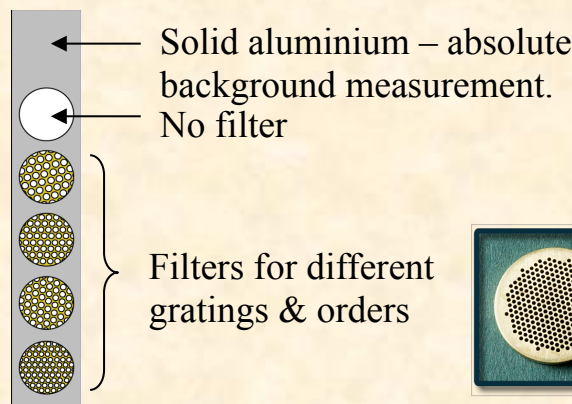
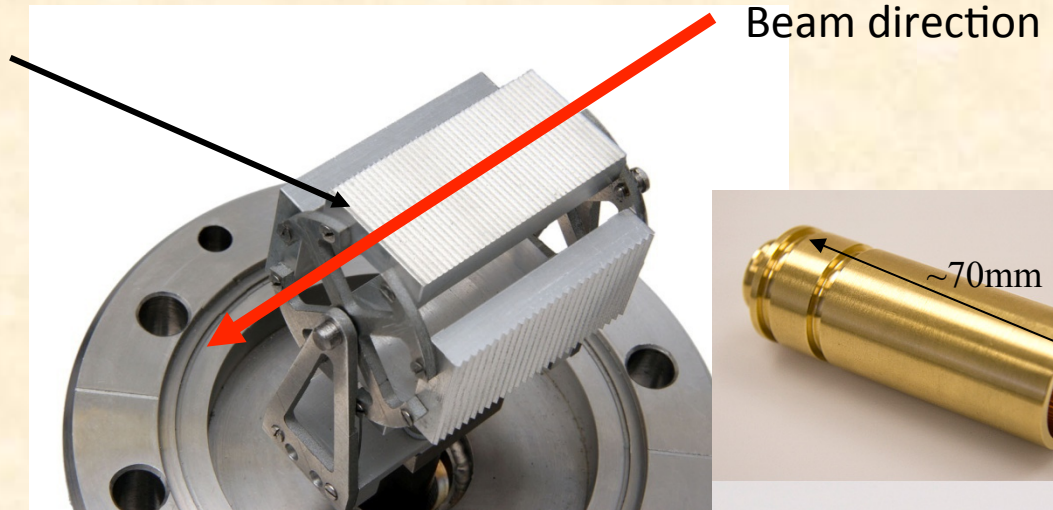
# E-203 Experimental apparatus (schematic)



# E-203 Experimental apparatus: gratings carousel

3 gratings  
1 blank piece of aluminium

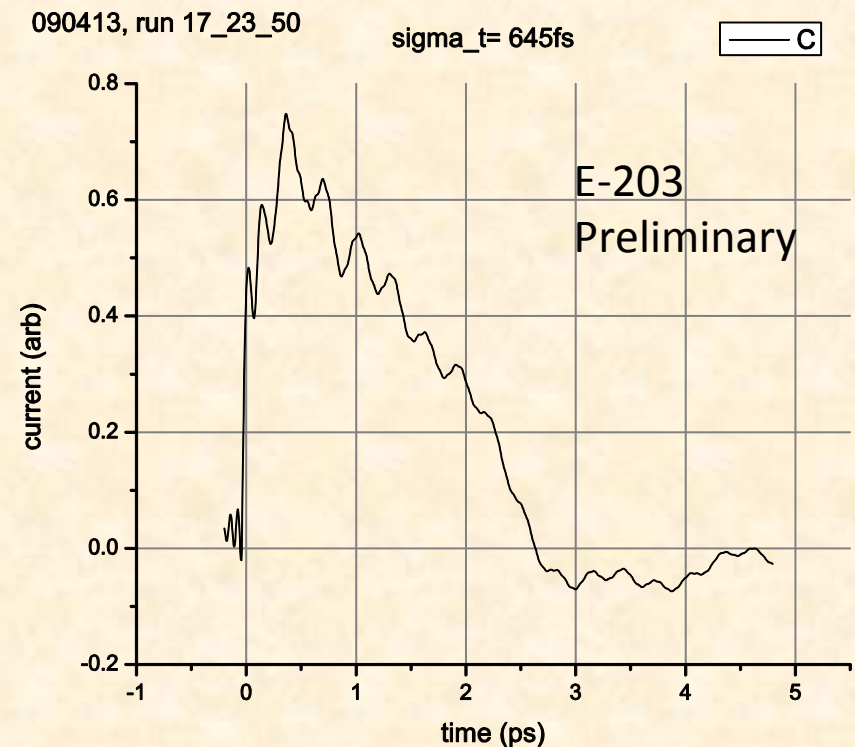
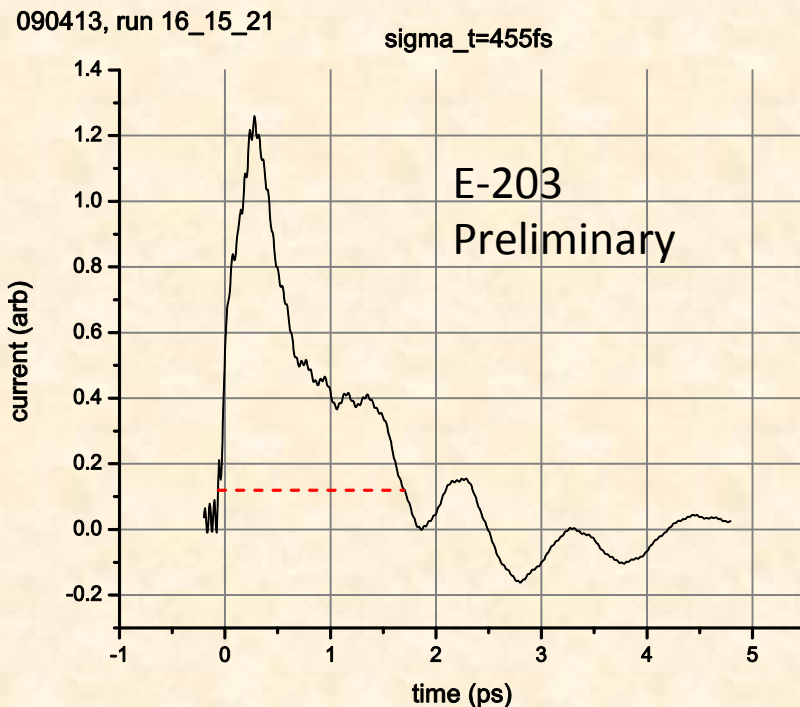
The carousel can rotate and offer three different gratings or one blank to the beam.

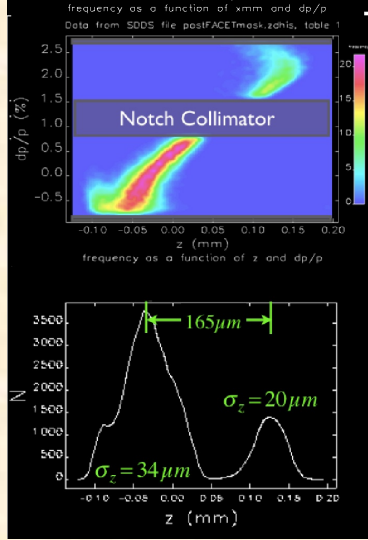




# Example of recent results from E-203

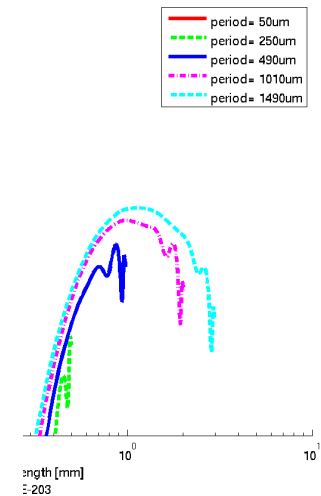
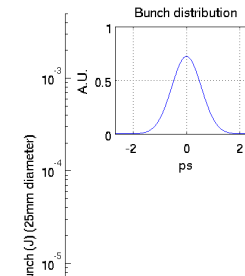
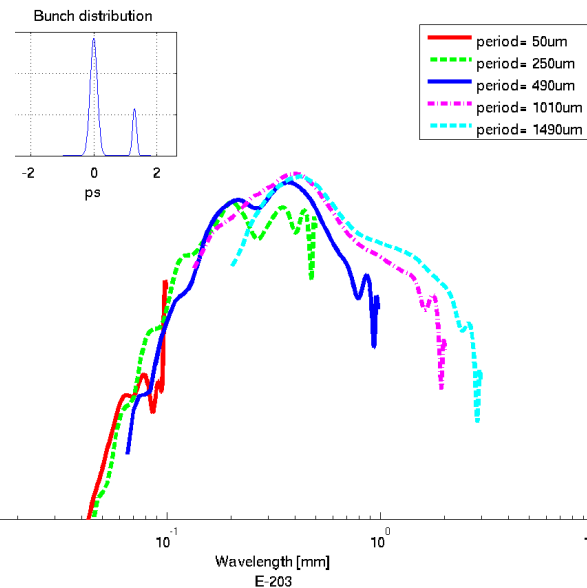
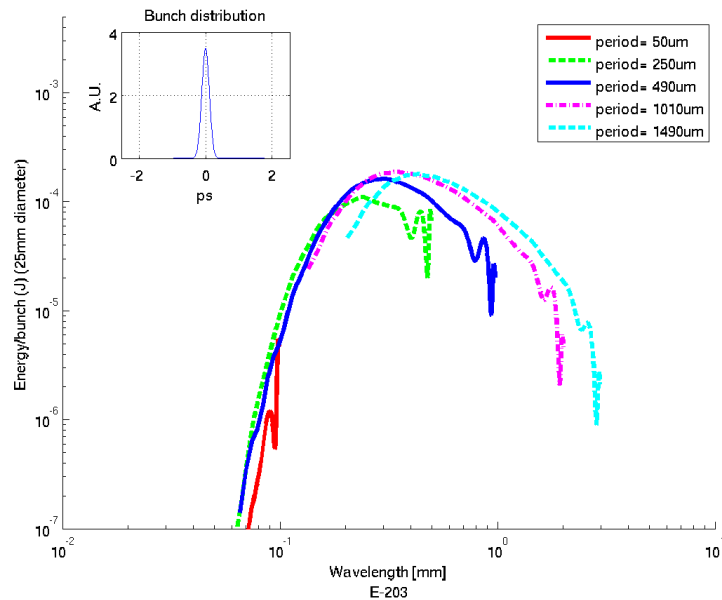
At FACET it is possible to change the settings of the compression chicane to compare high compression settings (left) and low compression settings (right) .





# Future measurement Notch collimator

- Simulations indicate that we should be sensitive to the effect of the notch. Measurements to come...

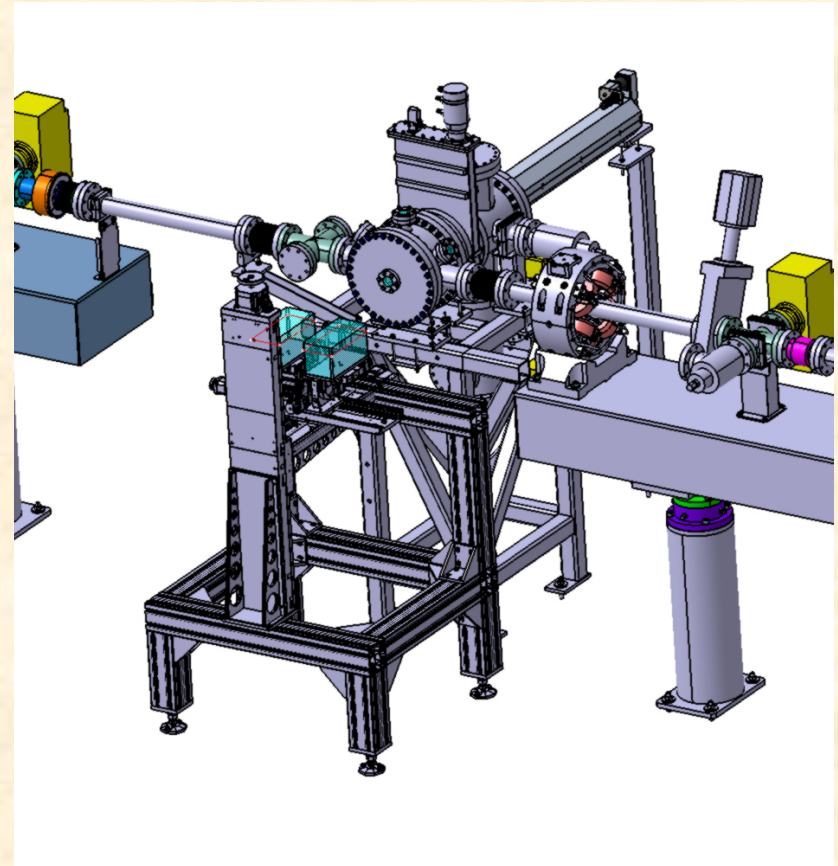


# Toward single shot

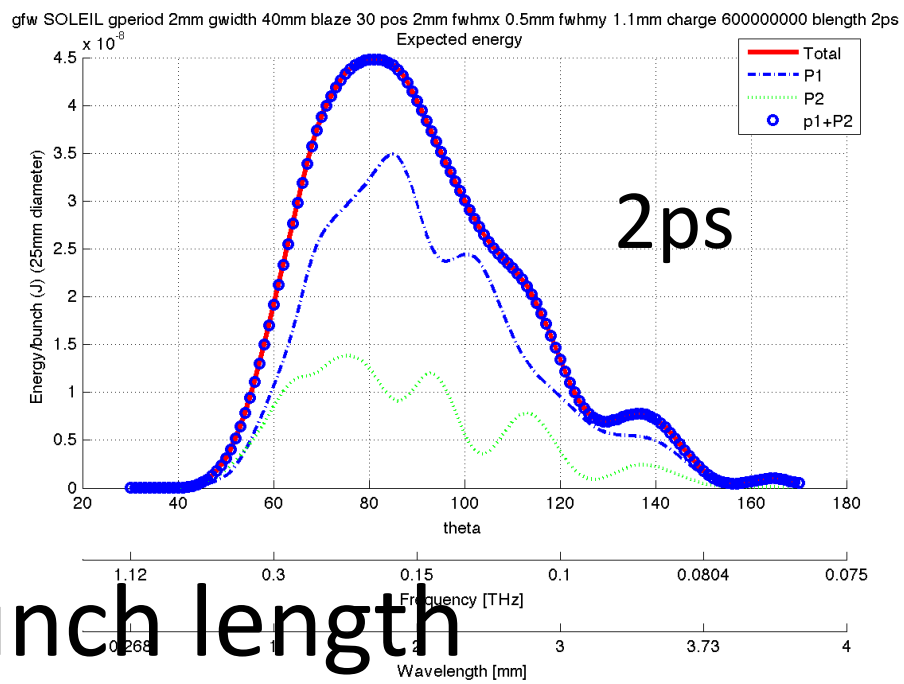
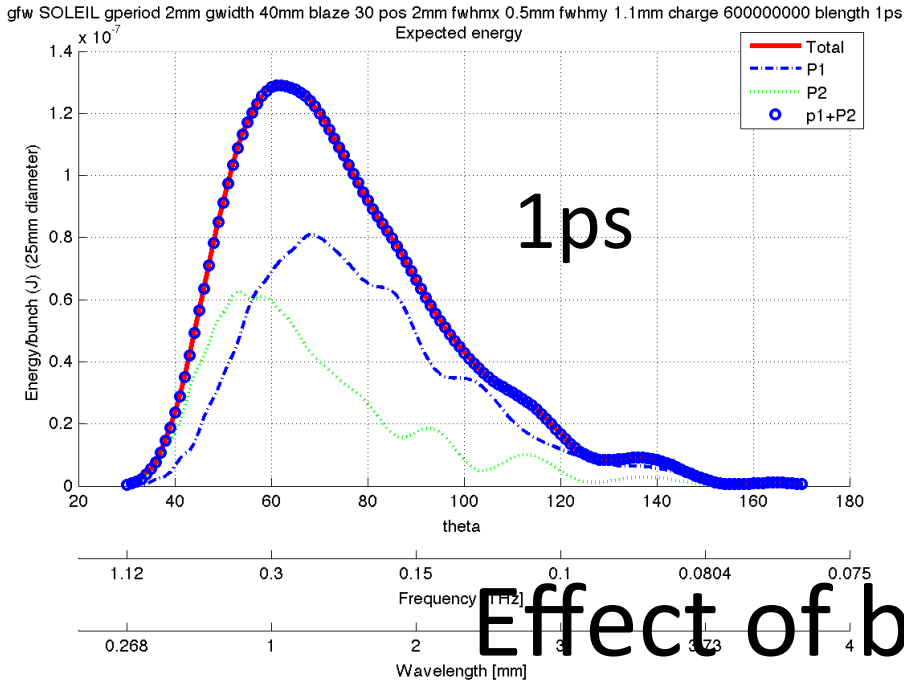
- At the moment our R&D has been mostly on classical accelerators.
- We do the background subtraction on a different shot than the signal measurement.  
=> we need a strategy to get signal and background in the same shot.
- To get a good profile reconstruction it is also useful to use several gratings (better dynamic range)  
=> need to develop a strategy to record signal from several gratings in a single shot.

# SPESO: Smith Purcell Experiment at SOLEIL

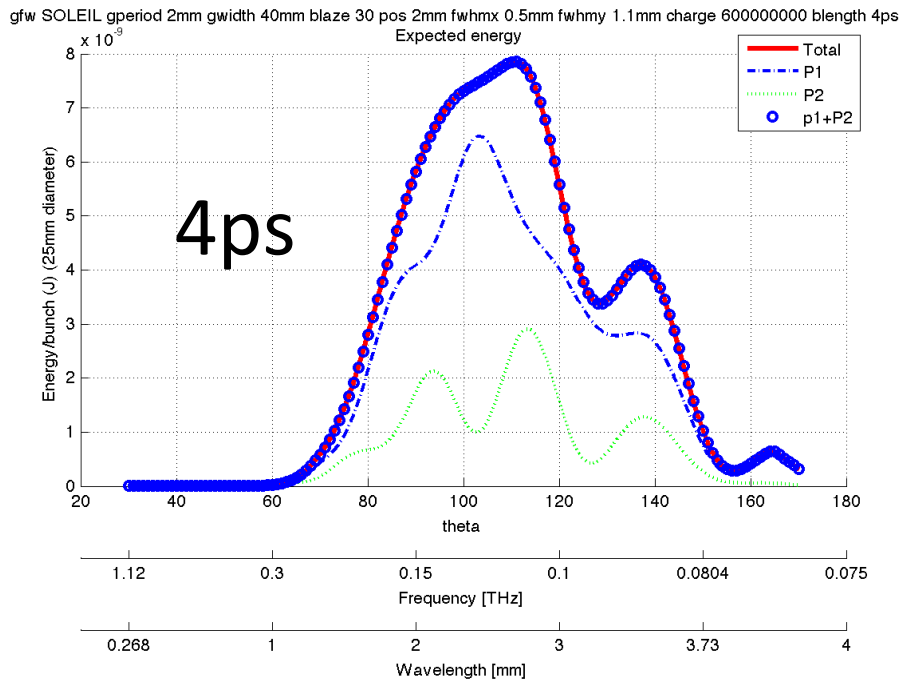
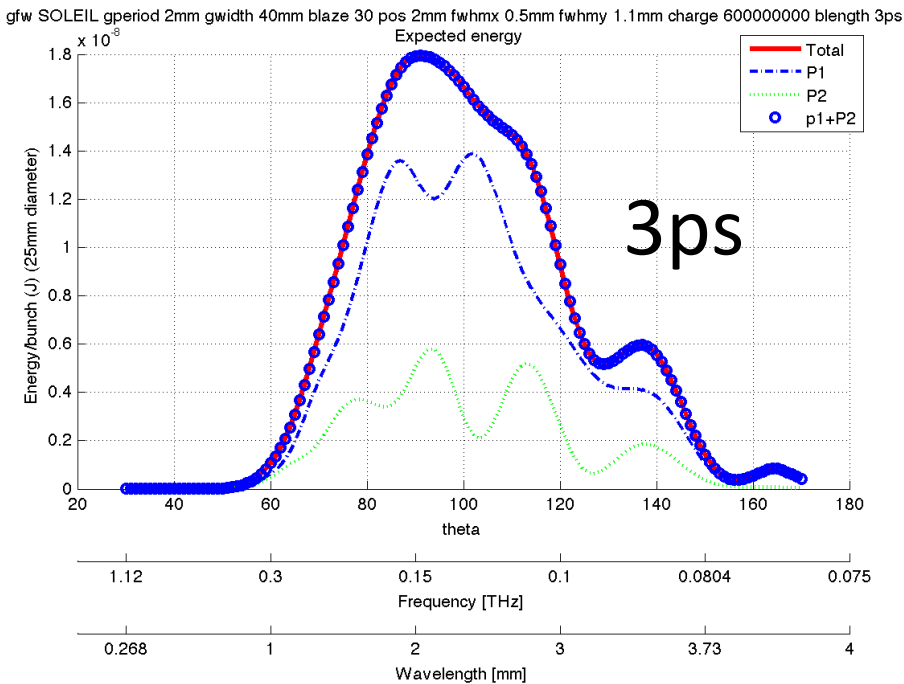
- To make a single shot measurement we need to be able to measure background and signal at the same time.
- => need to understand the 3D distribution of CSPR to find where to measure the background radiation. Polarisation measurements are also important.
- SPESO: 5D robot to scan the distribution of CSPR.
- Aim: build a single shot monitor, including background subtraction, based on CSPR in 2014.





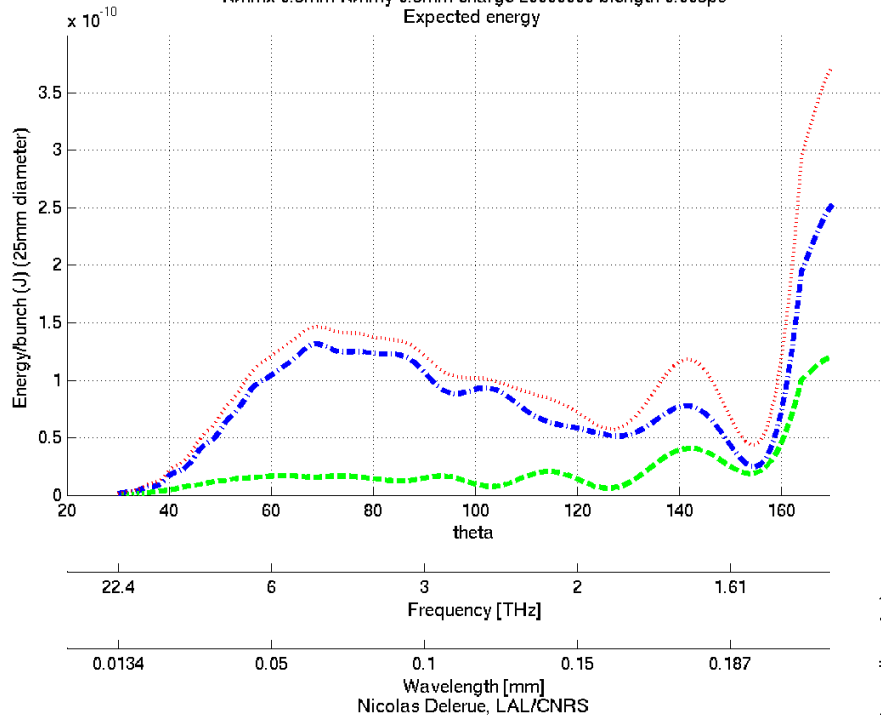


# Effect of bunch length



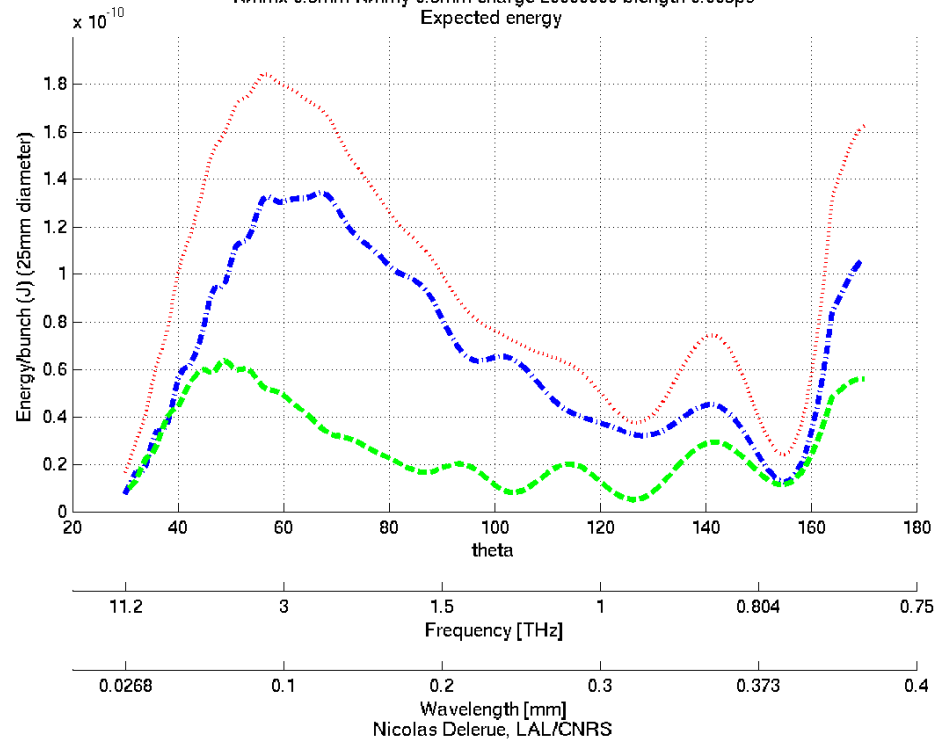
# 5fs bunch / 1pC

gfw DACTOMUS gperiod 0.1mm gwidth 20mm blaze 30 pos 1mm  
fwhmx 0.5mm fwhmy 0.5mm charge 20000000 blength 0.005ps  
Expected energy



- Very low signal level  
=> will require good amplification and background control.

gfw DACTOMUS gperiod 0.2mm gwidth 20mm blaze 30 pos 1mm  
fwhmx 0.5mm fwhmy 0.5mm charge 20000000 blength 0.005ps  
Expected energy



# Outlook

- CSPR can be used as a non interceptive longitudinal bunch profile diagnostic suited for the ps/fs range.
- Currently being tested on classical accelerators (FACET, SOLEIL,...).
- To make it a single shot diagnostic we need to find how to measure the background independently of the signal.
- Measurements to address this issue to start soon at SOLEIL.