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Search for Sterile Neutrinos in the CC ν_μ mode at the new CERN-SBLNF

G. Marsella on behalf of Nessie coll.

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Outline

- Physics Potential and Motivations
- The new CERN Neutrino Facility (CENF)
- ICARUS LAr & NESSIE spectrometers
- Project Schedule

Neutrino Physics

Are neutrino a simple “carbon copy” repetition of quarks?

- Over the last several years, neutrinos have been the origin of an impressive number of “Surprises”:
 - Mass differences, once zero “by ignorance”, are important
 - Large mixing oscillations
 - Oscillations due to matter exist, due to neutral currents
- But this isn't all ! Important discoveries may be ahead:
 - CP violation in the lepton sector, following the $\theta_{13} \gg 0$ result
 - Majorana or Dirac ν 's ?; ν -less β -decay
 - Actual values of ν -masses
 - Right handed neutrinos and see-saw mechanisms
 - Sterile neutrino and other “surprises”
- Of course the astronomical importance of neutrinos in space is immense, so is their role in the cosmic evolution.
- Just a few eV neutrino may be the source of the dark mass.

Fitting all together?

there are three classes of data:

$\nu_e \rightarrow \nu_e$ disappearance	$\sin^2 2\theta_{ee}$
$\nu_\mu \rightarrow \nu_\mu$ disappearance	$\sin^2 2\theta_{\mu\mu}$
$\nu_\mu \rightarrow \nu_e$ appearance	$\sin^2 2\theta_{\mu e}$

- ▶ each combination of **two** sets is consistent
(they depend on different mixing parameters)
- ▶ **BUT:** strong tension if all three of them are combined

3+1 SBL oscillations

$$\sin^2 2\theta_{\mu e} \approx \frac{1}{4} \sin^2 2\theta_{ee} \sin^2 2\theta_{\mu\mu}$$

$\nu_\mu \rightarrow \nu_e$ app. signal **requires** also signal in both, ν_e and ν_μ disappearance
(appearance mixing angle quadratically suppressed)

How to disentangle all the anomalies

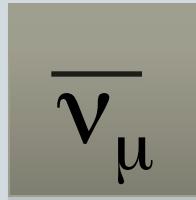
Short Base Line Neutrino Beam

Measure at L/E corresponding at $\Delta m^2 \sim 1 \text{ eV}^2$

with ≥ 2 sites



AND



in Apppearance (ν_e)

AND

in Disappearance (ν_μ)

L.Stanco[©]

- ✓ **Measure Electrons and Muons in a wide range (0.2 – 2 GeV)**
- ✓ **Measure Charge-Id**
- ✓ **Low systematics**
- ✓ **High rate capability (for both beams and detectors)**

Search for “anomalies” from neutrino and anti-neutrino oscillations at $\Delta m^2 \approx 1\text{eV}^2$ with muon spectrometers and large LAr-TPC imaging detectors.

Technical proposal.

(CERN-SPSC-2012-010 and SPSC-P-347)

ICARUS Collaboration

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(a) Contact Person

NESSiE Collaboration

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CERN Neutrino Facility

Letter of Intent for the New CERN Neutrino Facility

Draft 0 - v1.

January 11, 2013

Neutrinos CERN Community

ABSTRACT:

The nature and characteristics of neutrinos still represent one of the most intriguing questions of modern particle physics. This Letter of Intent presents a plan to build at CERN a Short Baseline Neutrino Facility capable to meet this challenge. The new beam line will make use of the SPS beam, it will be located in the CERN North Area of the Prevessin site and will comprise in a first step 2 experimental halls (far and near) which will host the detector facilities. The beam layout will be compatible with a long neutrino baseline which might be constructed in a second stage.

KEYWORDS: Neutrino, Short-baseline, Long-baseline, Sterile Neutrinos, ICARUS, NESSIE,

LAGUNA, LBNO, CERN.

Leader Project:
Marzio Nessi (Cern)

Deputy:
Rende Steerenberg (Cern)

- 200 pages for 5 Work-Packages,
- a TDR for Beam, Infrastructures and Experiments



The ICARUS-NESSiE P-347 proposal at the CERN-SPS

SPSC-P-347 (arXiv:1203.3432)

- L/E oscillation path lengths to ensure appropriate matching to the Δm^2 window for the expected anomalies
- NEAR and FAR sites
- “Imaging” LAr-TPC detector capable of identifying unambiguously all reaction channels
- Magnetic spectrometers to determine muon charge and momentum
- Interchangeable ν + anti- ν beams
- High rates due to detector large masses , in order to record relevant effects at the percent level ($>10^6 \nu_\mu$, $\sim 10^4 \nu_e$)
- Both initial ν_e and ν_μ components cleanly identified.

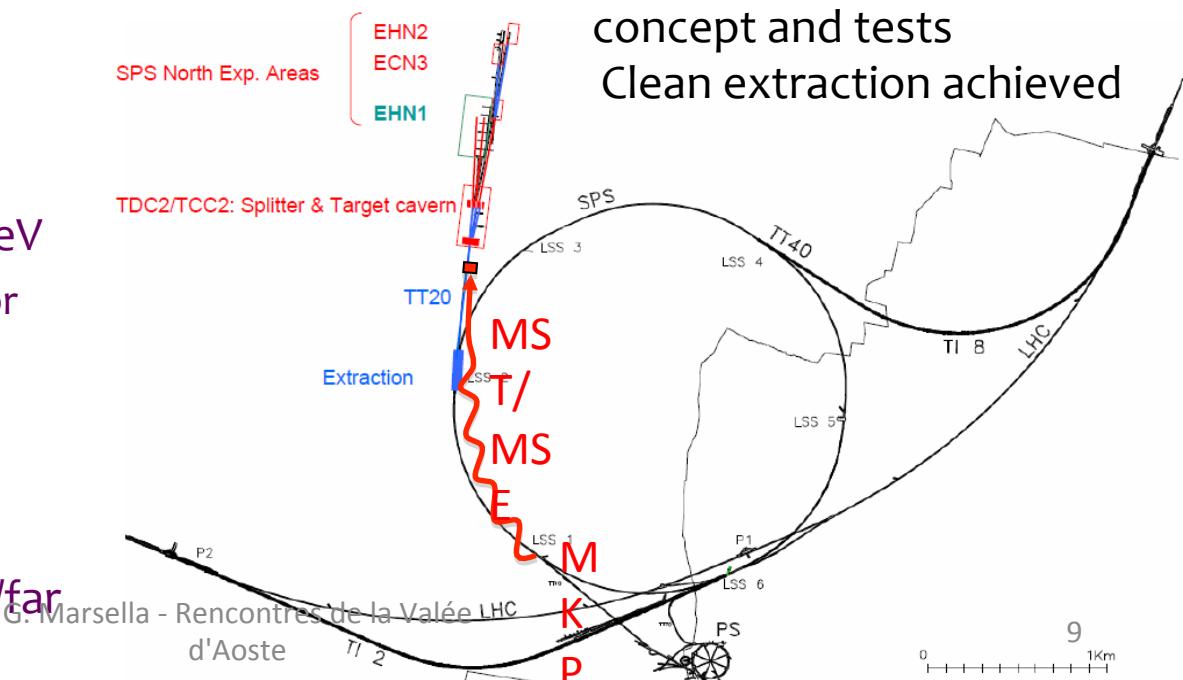
Preliminary SBL layout at the North Area

I. Efthymiopoulos - SBL2NA, September 26, 2012



Near site $L \sim 300$ m
(sufficient to shield from μ)
Far site $L \sim 1600$ m (still within CERN properties)
SPS p beam: 100 GeV ν beam ~ 2 GeV
 $4.5 \cdot 10^{19}$ pot/year (~stably achieved for CNGS)
target station ~ 11 m underground
 ν -beam pointing upwards (5 mrad slope, -3 m at far)
interactions/spill = 5 / 0.65 at near / far
($2 \cdot 10^{13}$ p/spill with "+" polarity)

SPS "non local extraction":
concept and tests
Clean extraction achieved



ICARUS-NESSiE

A coupled system of LAr detector
and muon Spectrometers

NEAR SITE @300 m :

LAr mass = 119 t

Iron magnet mass = 840 t

5 interactions/spill

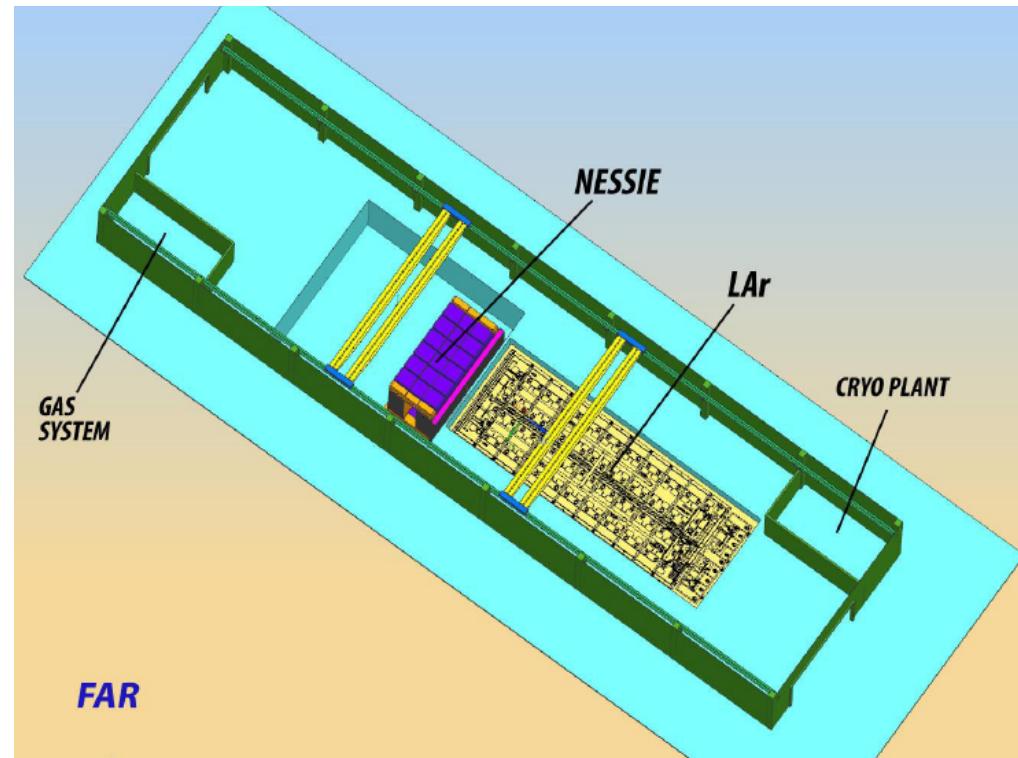
$2 \cdot 10^{13}$ p/spill (with “+” polarity)

FAR SITE @1600 m:

LAr mass = 476 t

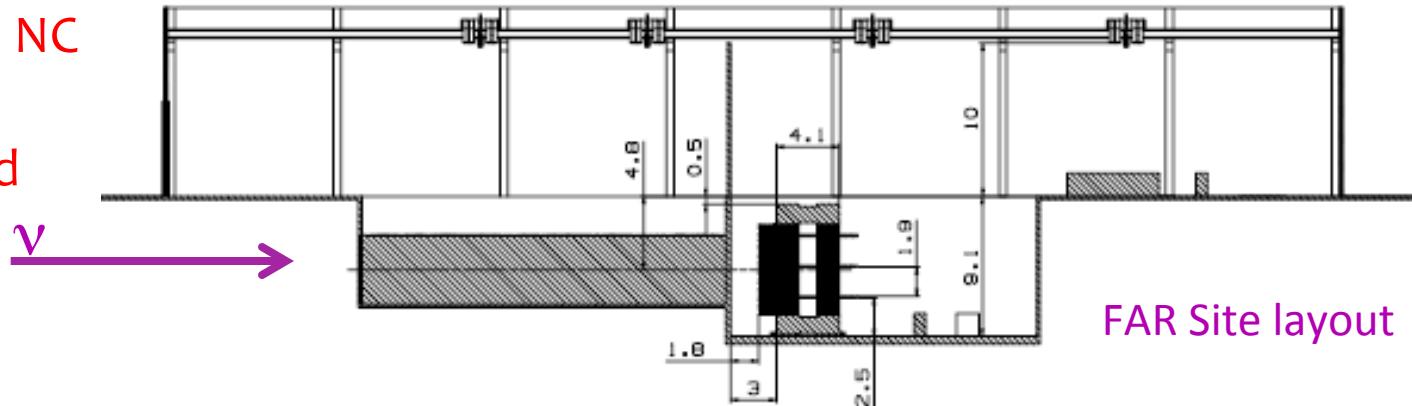
Iron magnet mass = 1515 t

0.65 interactions/spill

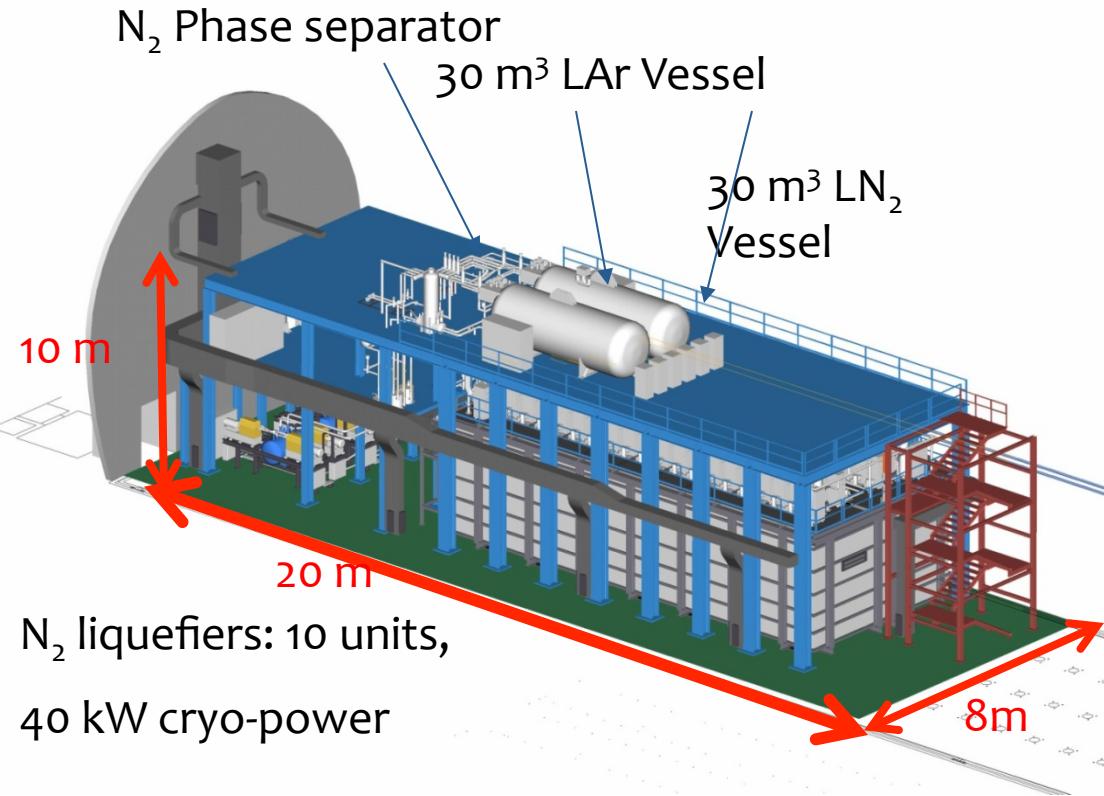


Observation ν_μ , ν_e CC, NC
channels

Charge separation and
muon momentum



ICARUS T600 at LNGS : LAr TPC status of the art



Two identical modules

- $3.6 \times 3.9 \times 19.6 \sim 275 \text{ m}^3$ each
- Liquid Ar active mass: $\sim 476 \text{ t}$
- Drift length = 1.5 m (1 ms)
- HV = -75 kV E = 0.5 kV/cm
- v-drift = 1.55 mm/ μ s

4 wire chambers:

- 2 chambers per module
- 3 R/O wire planes per chamber at 0, $\pm 60^\circ$
- 54000 wires, 3 mm pitch, 3 mm plane spacing
- 20+54 PMTs, 8" Ø, for scintillation light:
VUV sensitive (128nm) with wavelength shifter

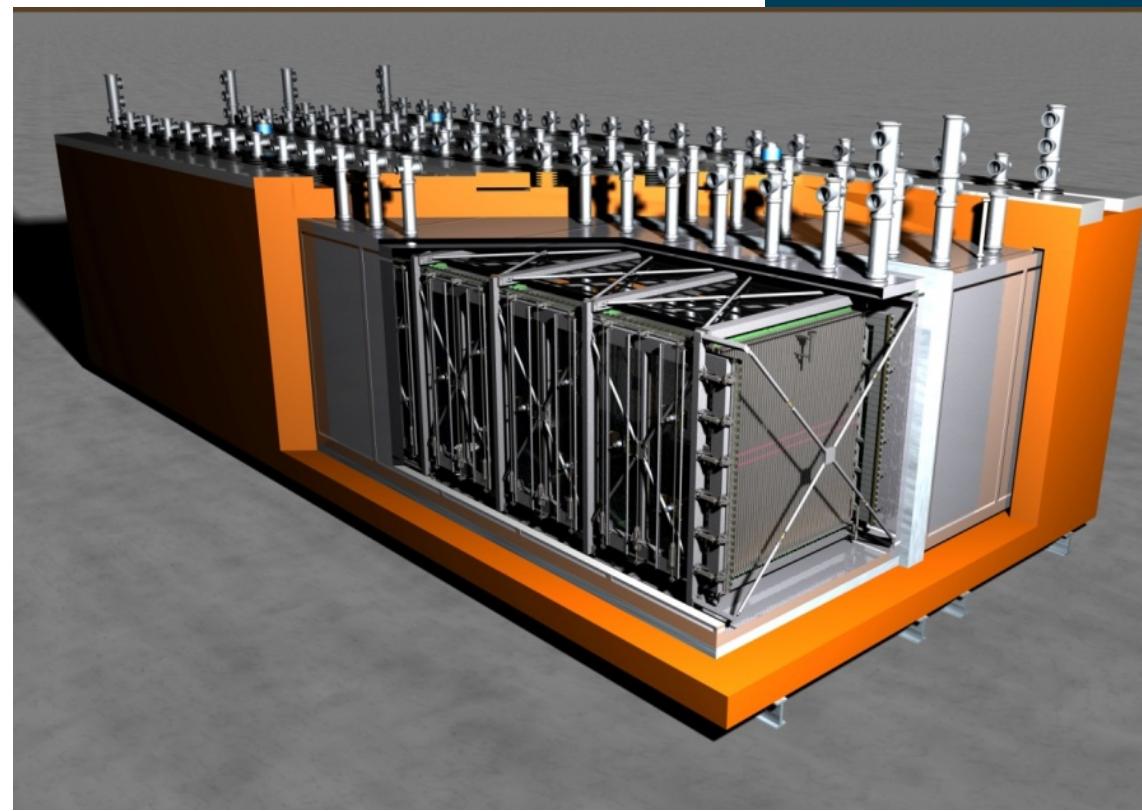
The ICARUS T600 as “Far” detector

- T600 will be transported to CERN in 2013, after decommissioning at LNGS, ensuring the new experiment operation again in 2016
- A large number of components will be disassembled and transported: inner detectors, electronics, ancillary systems, LN₂ liquefaction system

TPC's will be inserted in new vessels of extruded Al, vacuum-tight $< 10^{-5}$ mbar l s⁻¹ standing up to 1 bar internal overpressure

New external insulation based on industrial membrane tank concept: better performance and full spill containment. Expected heat loss: 6.6 kW (T600).

ICARUS Collab Courtesy



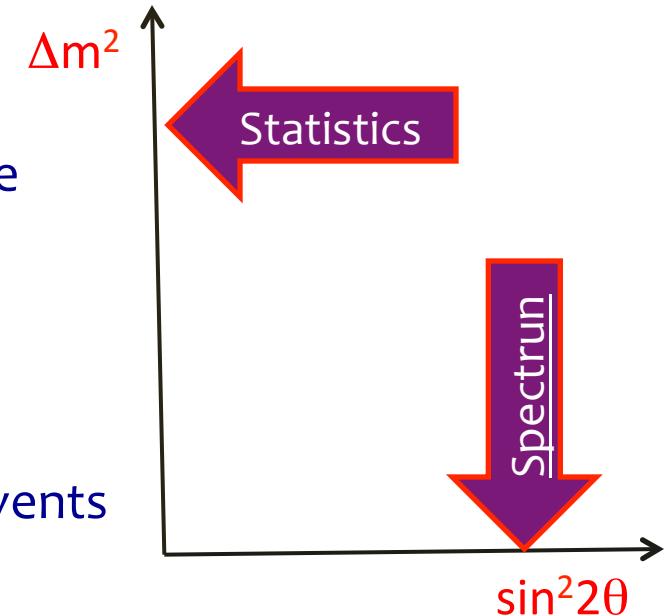


NESSiE

(Neutrino Experiment with SpectrometerS in Europe)

- Goals:

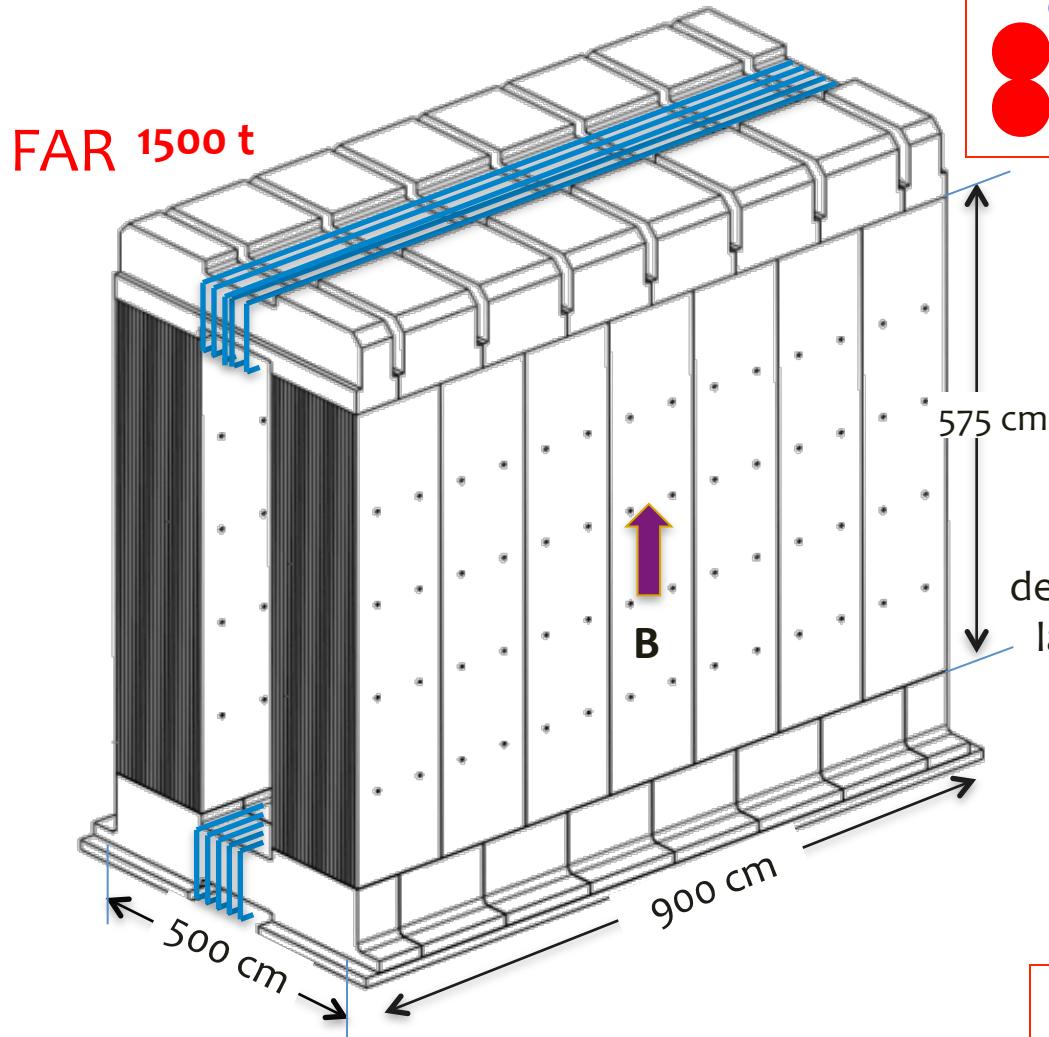
- Allow charge separation and momentum measurement of as many muons as possible escaping from LAr
(large statistics \leftrightarrow low $\sin^2 2\theta$)
- Go as low as possible in muon momentum
(low momenta \leftrightarrow low Δm^2)
- Possibility to also study (NESSiE) internal events
(coarser resolution w.r.t. LAr)



- Solution:

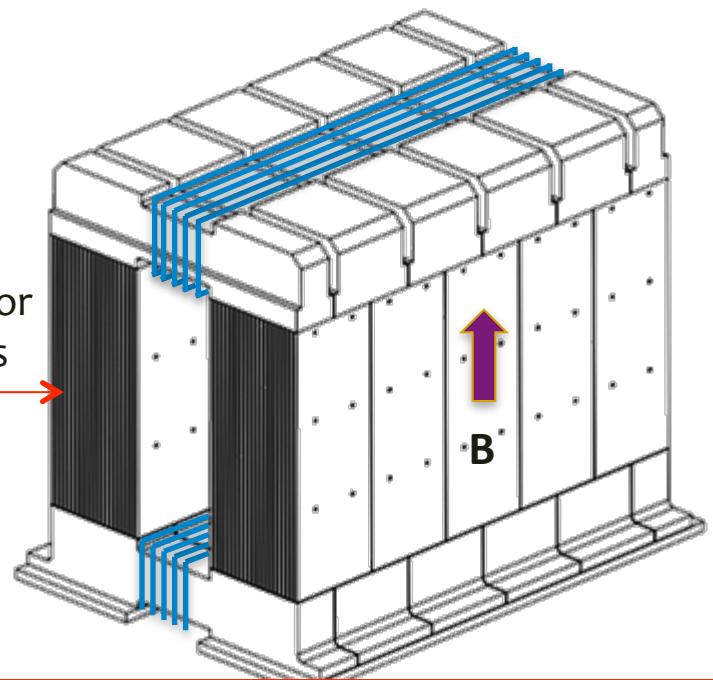
- Air-core magnets for low momentum muons escaping from LAr
($P_\mu < 0.5 \text{ GeV}/c$ in NESSiE $\leftrightarrow \langle E_\nu \rangle < 1 \text{ GeV}$ in LAr)
- Downstream massive iron dipolar magnets for higher momentum extension

Iron dipolar magnets (OPERA-like)



- design, construction/operation expertise acquired with OPERA
- digital RPC read-out
- Same design scheme for electronics

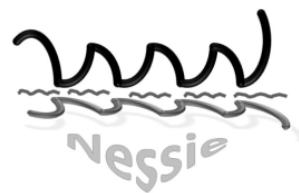
NEAR 840 t



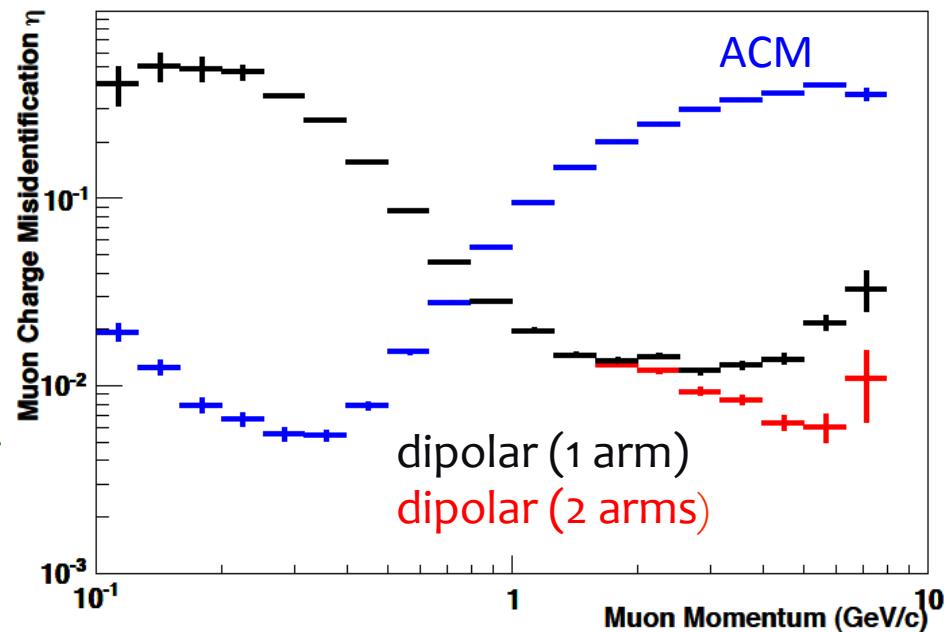
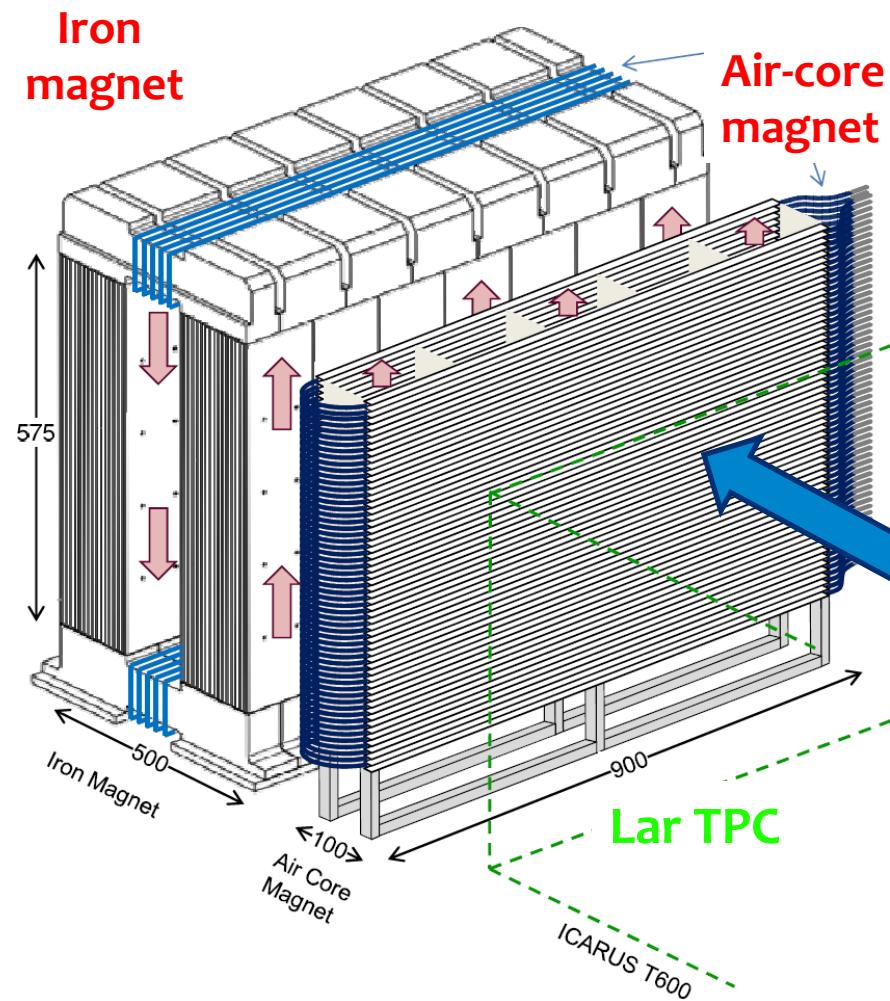
- 1800 + 700 m² of RPC (11 planes/arm)
- 20000 + 12000 digital channels
- Bending in 1.5 T uniform B-field and with range for stopping muons (~50%)

A massive instrumented target itself

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NESSiE Detector configuration (@ Near and Far locations)

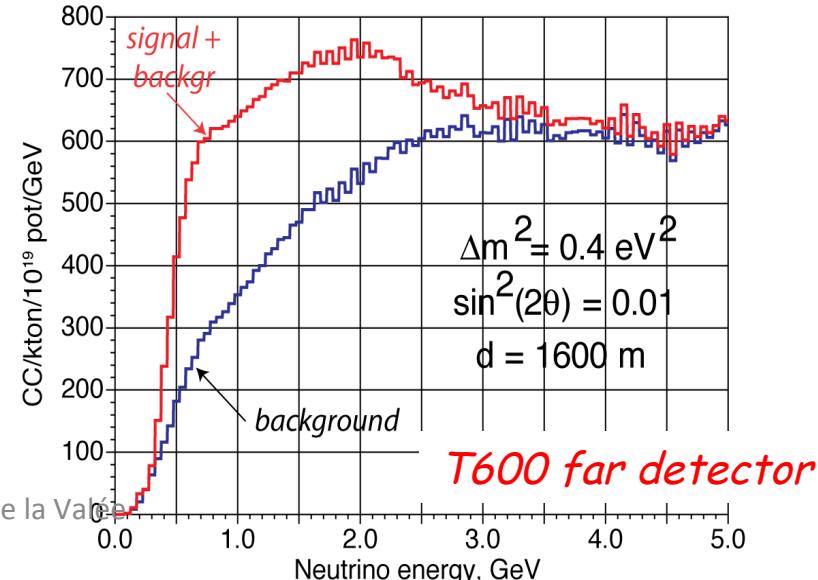
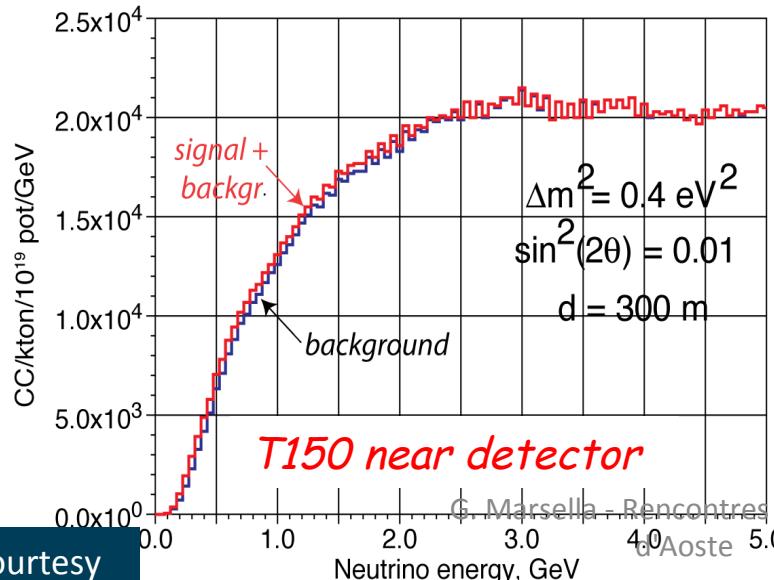


Charge mis-reconstruction at few % level in 0.5 -10 GeV range

(full simulation including selection, efficiency and reconstruction)

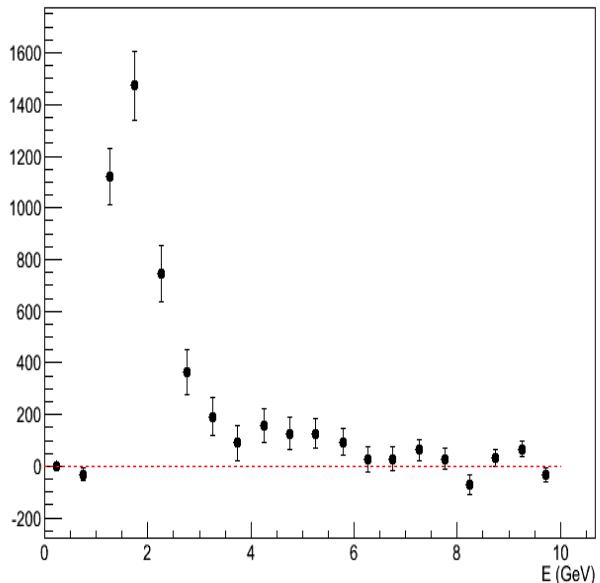
A possible expectation of LSND mass and mixing angle

- Our dual method, unlike LNSD and MiniBooNE, determines **both** the mass difference and the value of the mixing angle.
- Events rates are hereby, shown both for [background] and [oscill.+background] at d=300 m and d=1600 m and the optimal "predictions" from ICARUS et al.
 $\Delta m^2 = 0.4 \text{ eV}^2$, $\sin^2(2\theta) = 0.01$
- For d=1600 m, $E\nu < 5 \text{ GeV}$ and $4.5 \cdot 10^{19} \text{ pot (1 y)}$ a ν_e oscillation signal of ≈ 1200 events is expected above 5000 backgr. events



	NEAR ($\bar{\nu}$ -bar)	NEAR(ν)	FAR($\bar{\nu}$ -bar)	FAR(ν)
$\nu_e + \nu_e\text{-bar (LAr)}$	35 K	54 K	4.2 K	6.4 K
$\nu_\mu + \nu_\mu\text{-bar (LAr)}$	2000 K	5250 K	270 K	670 K
Appear. test point	590	1900	360	910
ν_μ (LAr+NESSiE)	230 K	1200 K	21 K	110 K
ν_μ (NESSiE)	1150 K	3600 K	94 K	280 K
$\nu_\mu\text{-bar}$ (Lar+NESSiE)	370 K	56 K	33 K	6.9 K
$\nu_\mu\text{-bar}$ (NESSiE)	1100 K	300 K	89 K	22 K
Disappear. test point	1800	4700	1700	5000

NESSiE $\bar{\nu}_\mu^{\text{CC}}$ (non oscillated - oscillated)

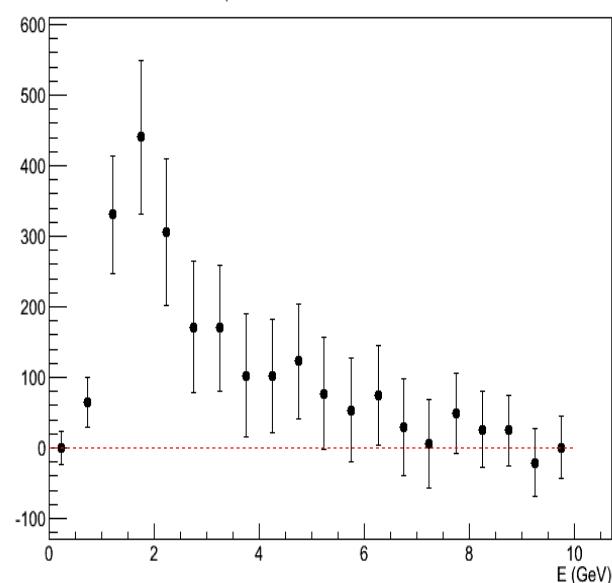


NESSiE ν_μ^{CC} (non oscillated - oscillated)

1 year of anti- ν run $4.5 \cdot 10^{11}$ pot (negative polarity)

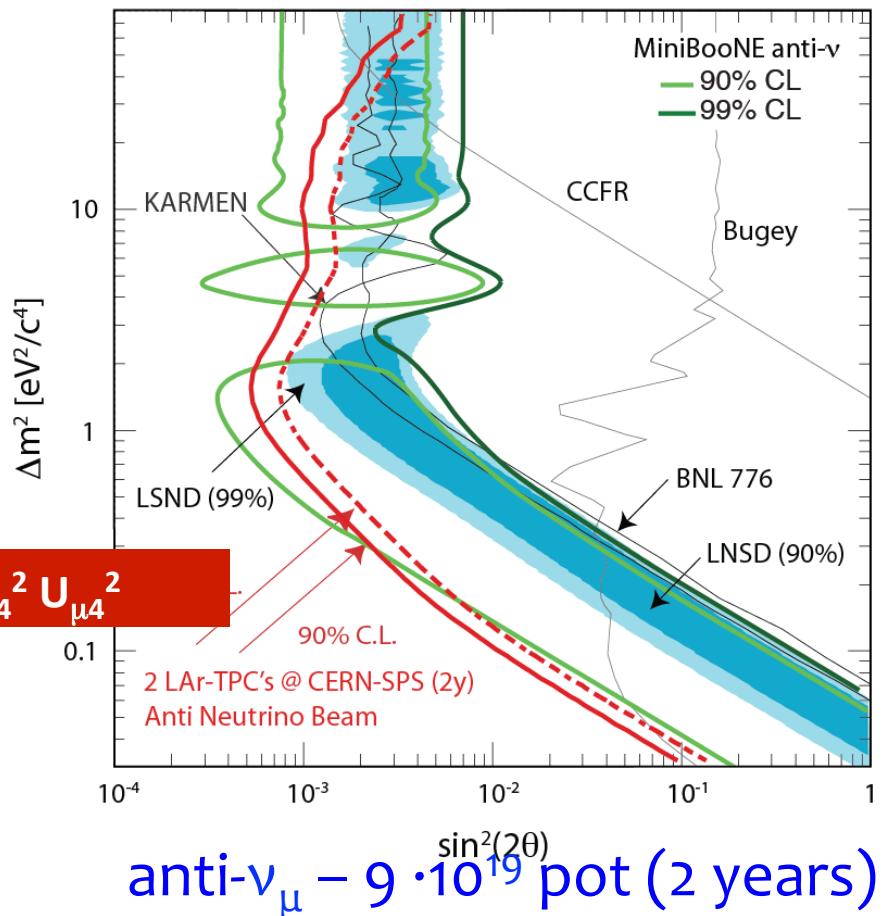
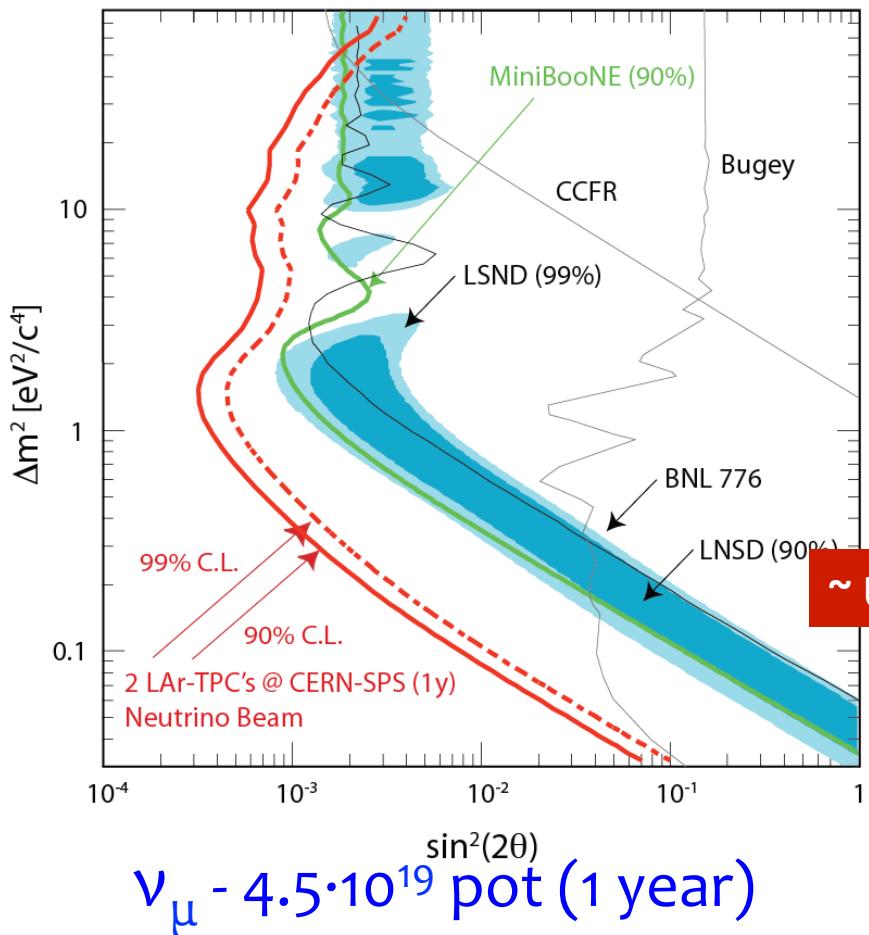
Disappearance signals

G. Marsella - Rencontres de la Vallée d'Aoste



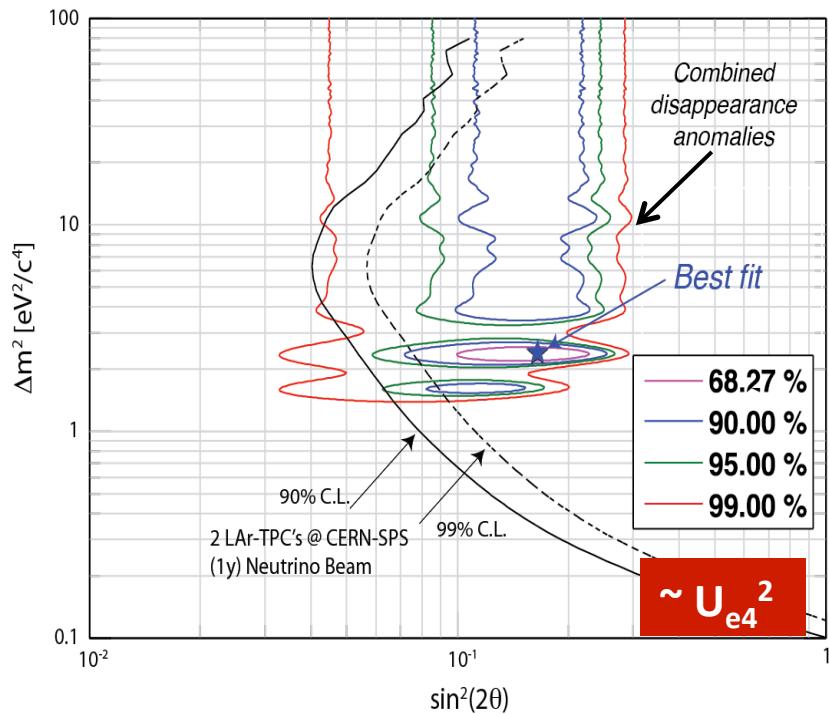
Physics reach: all channels explored

ν_e appearance



ν_e, ν_μ Disappearance

ν_e disappearance (1 year of ν_μ)



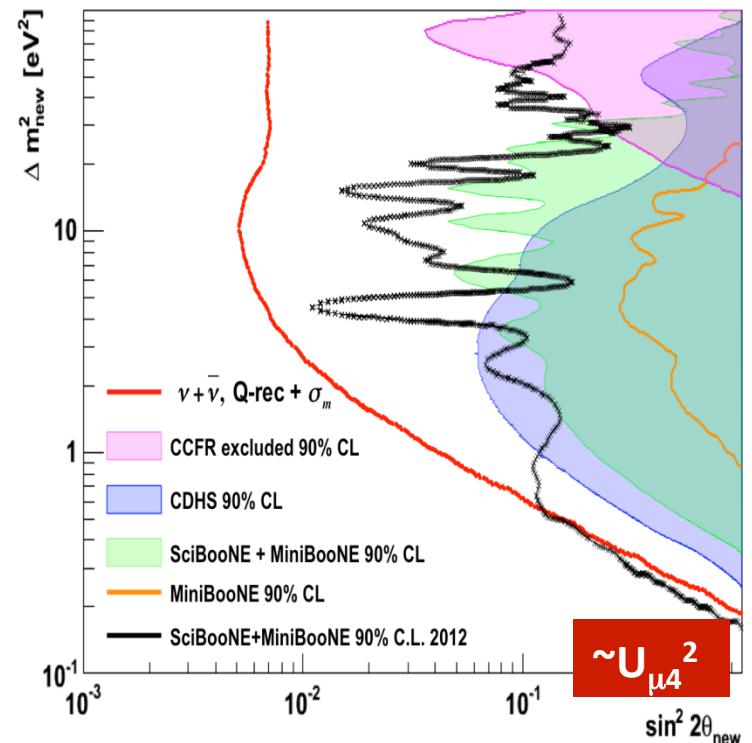
The reactor/Gallium anomalies fully addressed

(*) $4.5 \cdot 10^{19}$ pot (1 year)

3/2/13

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ν_μ disappearance ($2 + 1$)



2 year run (*) with $\bar{\nu}$ would address CPV in one shot with NESSiE+LAr

Status of approval and foreseen

At CERN

- Group established by CERN in order to realize SPS based new short-baseline ν beam in the North Area (project leader M. Nessi)
- Scientific Approval (SPSC) middle of January 2013
- Feasibility document submitted to CERN Directorate on February 7th 2013
- Resource Approval (Research Board) on March 4th 2013
- Final Approval (CERN Council) in June 2013

INFN: Currently Major contributor to the experiments (not beam)

- Scientific approval
- Next step: evaluation by the Technical Scientific Committee (CTS) as for costs, manpower
- In-Kind contribution of Opera Spectrometers

Time schedule: neutrino beam by June 2016

activity/year	2013	2014	2015	2016	2017	2018
Far Detector	civil engineering + Infrastructure		Detectors installation + calibration		Commissioning with beam	
Near Detector	civil eng. + Infrastruct.		Detectors installation + calibration			
Primary beam line	CE permits	civil engineering + components preparation	INFRA	Installation		
Target/Dump facility	CE permits	civil engineering + components preparation	INFRA	Installation		

ICARUS & NESSIE Installation

SBL-v Facility construction

physics : data taking

3.5 years from now

Conclusions

Possibility of exciting discoveries of BSM Physics with vast consequences or, at least, a complete clarification of present anomalies.

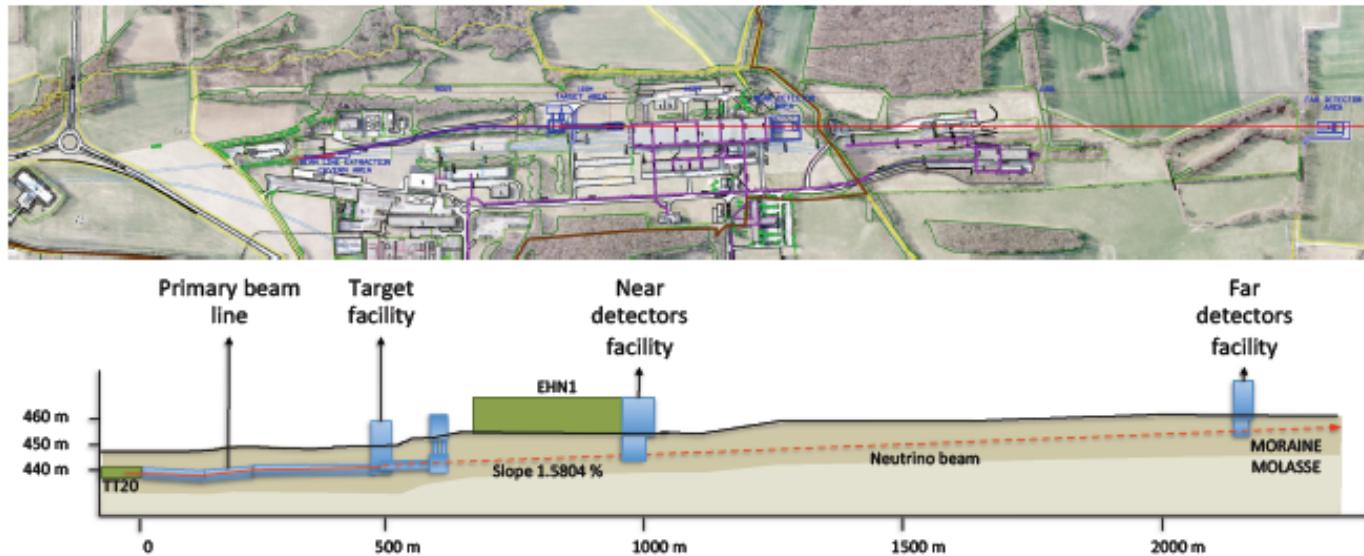
Favorable time scale thanks to the use of existing/running detectors (or reasonable extensions).

Opportunity for a revival of neutrino activity in Europe.

Large room for contributions !

BACKUP SLIDES

CERN Neutrino Facility (CENF)



Layout of the facility at CERN. Bottom: Vertical cut showing the TT20, the EHN1 and the main components of the facility

Effective SBL Oscillation Probabilities in 3+1 Schemes

$$P_{\nu_\alpha \rightarrow \nu_\beta} = \sin^2 2\vartheta_{\alpha\beta} \sin^2 \left(\frac{\Delta m_{41}^2 L}{4E} \right)$$

(no CP violation)

$$P_{\nu_\alpha \rightarrow \nu_\alpha} = 1 - \sin^2 2\vartheta_{\alpha\alpha} \sin^2 \left(\frac{\Delta m_{41}^2 L}{4E} \right)$$

$$\sin^2 2\vartheta_{\alpha\alpha} = 4|U_{\alpha 4}|^2 (1 - |U_{\alpha 4}|^2)$$

Perturbation of 3ν Mixing

$$|U_{e4}|^2 \ll 1, \quad |U_{\mu 4}|^2 \ll 1, \quad |U_{\tau 4}|^2 \ll 1, \quad |U_{s4}|^2 \simeq 1$$

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix}$$

↑
SBL

$$\sin^2 2\vartheta_{\alpha\alpha} \ll 1$$



$$|U_{\alpha 4}|^2 \simeq \frac{\sin^2 2\vartheta_{\alpha\alpha}}{4}$$

Effective SBL Oscillation Probabilities in 3+2 Schemes

$$\phi_{kj} = \Delta m_{kj}^2 L / 4E$$

$$\eta = \arg[U_{e4}^* U_{\mu 4} U_{e5} U_{\mu 5}^*]$$

$$P_{\substack{(-) \\ \nu_\mu \rightarrow \nu_e}}^{(-)} = 4|U_{e4}|^2|U_{\mu 4}|^2 \sin^2 \phi_{41} + 4|U_{e5}|^2|U_{\mu 5}|^2 \sin^2 \phi_{51} \\ + 8|U_{\mu 4} U_{e4} U_{\mu 5} U_{e5}| \sin \phi_{41} \sin \phi_{51} \cos(\phi_{54} - \eta)$$

$$P_{\substack{(-) \\ \nu_\alpha \rightarrow \nu_\alpha}}^{(-)} = 1 - 4(1 - |U_{\alpha 4}|^2 - |U_{\alpha 5}|^2)(|U_{\alpha 4}|^2 \sin^2 \phi_{41} + |U_{\alpha 5}|^2 \sin^2 \phi_{51}) \\ - 4|U_{\alpha 4}|^2|U_{\alpha 5}|^2 \sin^2 \phi_{54}$$

- More parameters: 7 *many coupled measurements !!*
- CP violation

[Sorel, Conrad, Shaevitz, PRD 70 (2004) 073004; Maltoni, Schwetz, PRD 76 (2007) 093005; Karagiorgi et al, PRD 80 (2009) 073001; Kopp, Maltoni, Schwetz, PRL 107 (2011) 091801; Giunti, Laveder, PRD 84 (2011) 073008; Donini et al, arXiv:1205.5230]

M. Laveder®

3+1 SBL oscillations

appearance

$$P_{\mu e} = \sin^2 2\theta_{\mu e} \sin^2 \frac{\Delta m_{41}^2 L}{4E} \quad \sin^2 2\theta_{\mu e} = 4|U_{e4}|^2 |U_{\mu 4}|^2$$

disappearance ($\alpha = e, \mu$)

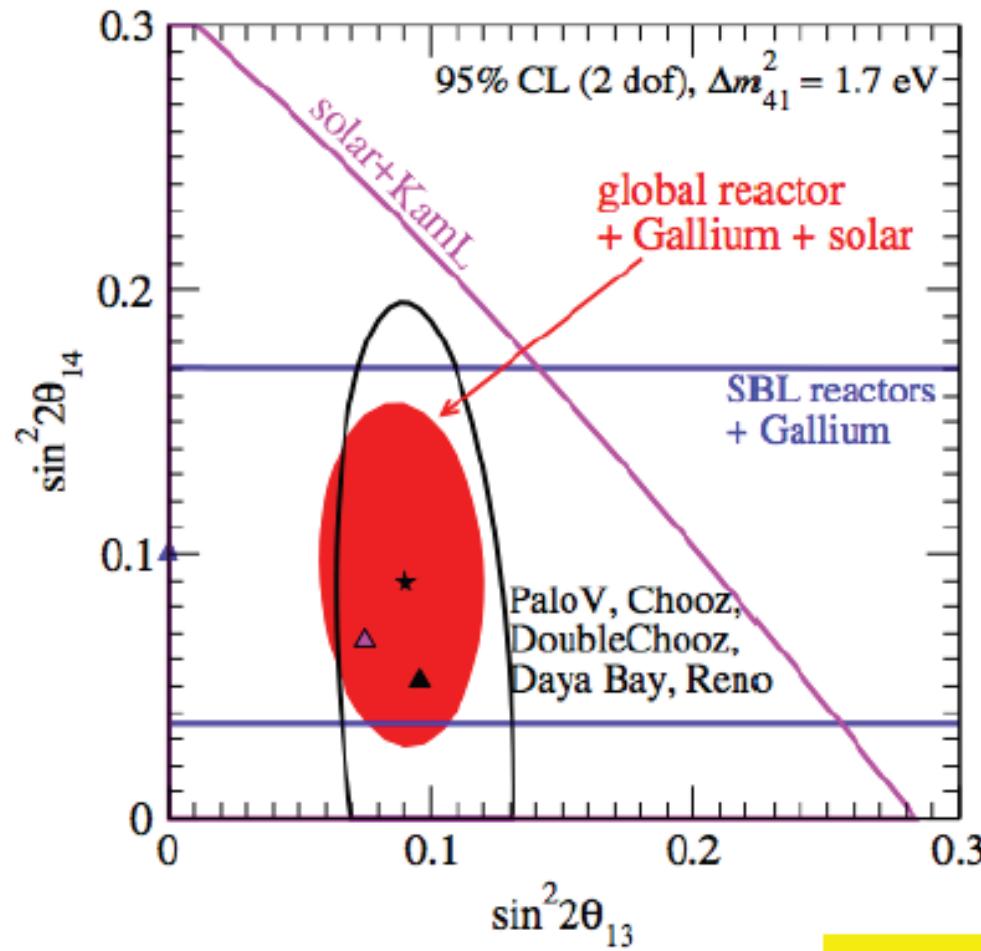
$$P_{\alpha\alpha} = 1 - \sin^2 2\theta_{\alpha\alpha} \sin^2 \frac{\Delta m_{41}^2 L}{4E} \quad \sin^2 2\theta_{\alpha\alpha} = 4|U_{\alpha 4}|^2 (1 - |U_{\alpha 4}|^2)$$

$$\sin^2 2\theta_{\mu e} \approx \frac{1}{4} \sin^2 2\theta_{ee} \sin^2 2\theta_{\mu\mu}$$

$\nu_\mu \rightarrow \nu_e$ app. signal requires also signal in both, ν_e and ν_μ disappearance
(appearance mixing angle quadratically suppressed)

T. Schwetz[©]

*TENSIONS also for the 3+2 scheme, due to disappearance bounds, particularly for ν_μ
On top of that, coupling to standard 3-flavour scenario is weak :*

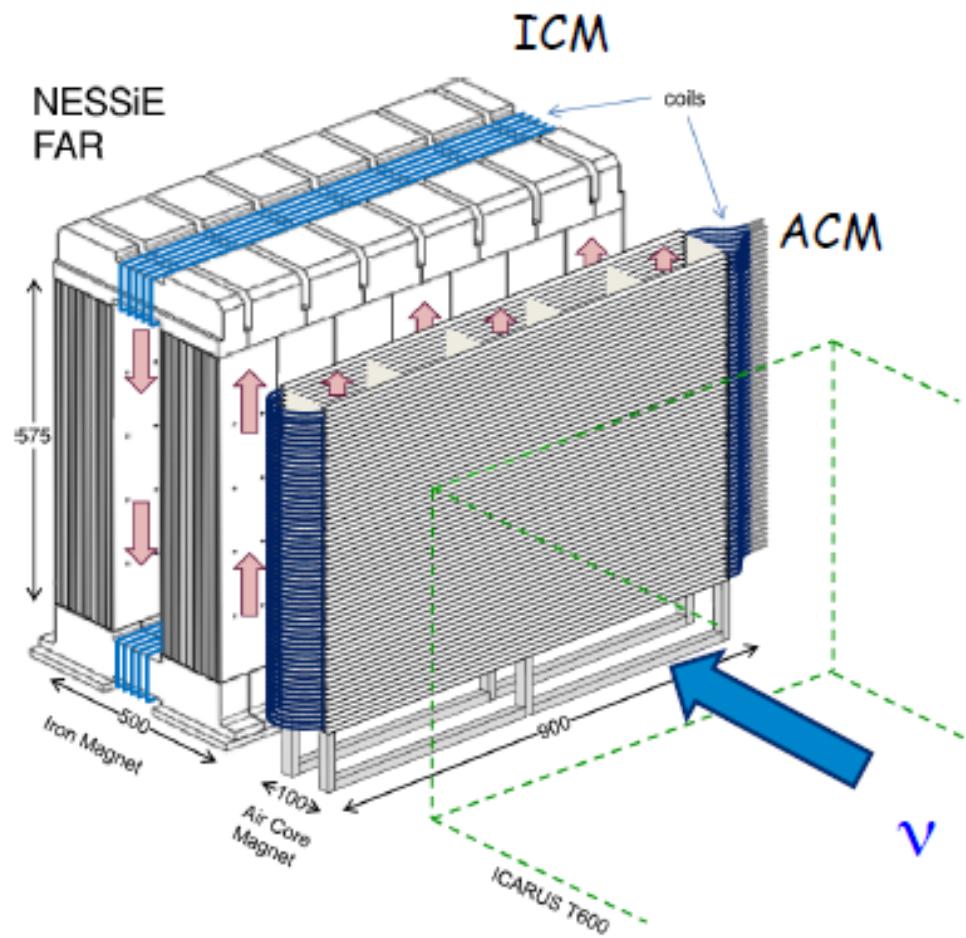


TENSIONS...

*it is a game to play
VERY carefully !*

Nessie main facts

- Two Iron spectrometers (**ICM**),
1500 + 800 t, composed by:
 - 48 yoke blocks, , $4.5 \times 0.6 \times 1$ m, 25t
 - 480 slabs, 2 – 3 t, 1.25m x 3.5–6 m
 - 1800 + 700 m² of RPC
- «sandwich style» assembly to be made in situ, one piece per time
- 20,000+12,000 digital channels
- Two **ACM** preassembled and installed in one shot
- Precision Trackers preassembled and installed in one shot
- Near Nessie movable aside on air-pad
- 1 + 0.5 MW , 10 kA, power
(summed up for ACM and ICM)



Optimized ! Reduced by almost a factor of 2 !!!

Useful ICARUS-NESSiE Refs

• C. Rubbia et al., ICARUS/CNGS2 Coll., Physics Programme for ICARUS after 2012 SPSC-M-773
09/03/2011

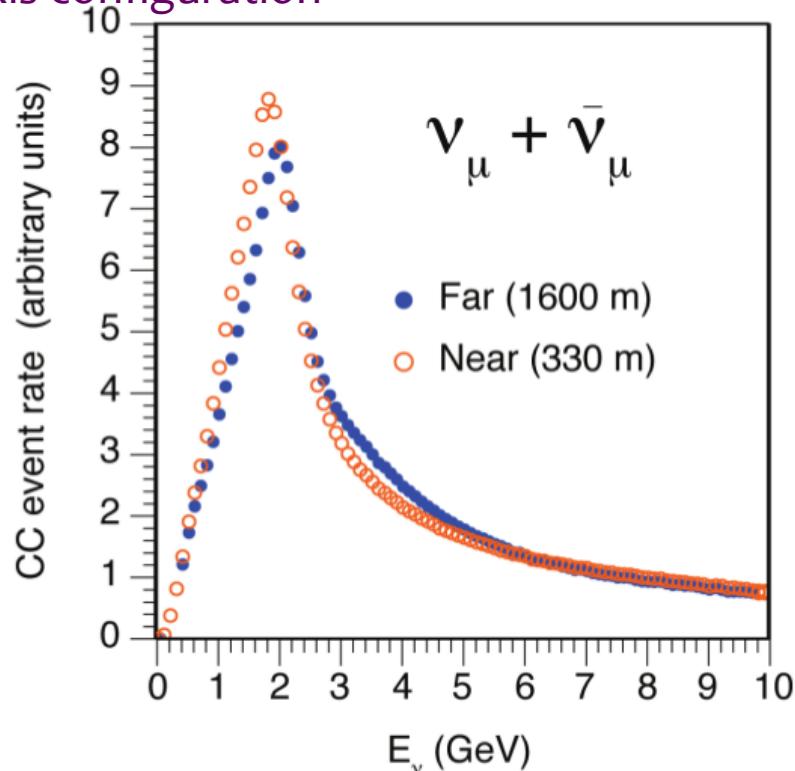
- C. Rubbia et al., ICARUS/CNGS2 Coll., A comprehensive search for “anomalies” from neutrino and anti-neutrino oscillations at large mass differences ($\Delta m^2 \approx 1\text{eV}^2$) with two LAr-TPC imaging detectors at different distances from the CERN-PS SPSC-P-345 14/10/2011
- P. Bernardini et al., NESSiE Coll. Prospect for Charge Current Neutrino Interactions Measurements at the CERN-PS SPSC-P-343 11/10/2011 arXiv:1111.2242v1
- M. Antonello et al. ICARUS-NESSiE Coll. Search for “anomalies” from neutrino and anti-neutrino oscillations at $\Delta m^2 \approx 1\text{eV}^2$ with muon spectrometers and large LAr-TPC imaging detectors. Technical proposal. SPSC-P-347 arXiv:1203.3432
- M. Antonello et al. ICARUS-NESSiE Coll. Search for anomalies in the neutrino sector with muon spectrometers and large LArTPC imaging detectors at CERN. Contribution to the European Strategy for Particle Physics. Krakow 10-12 September 2012. 28/09/2012 arXiv: 1208.0862v2



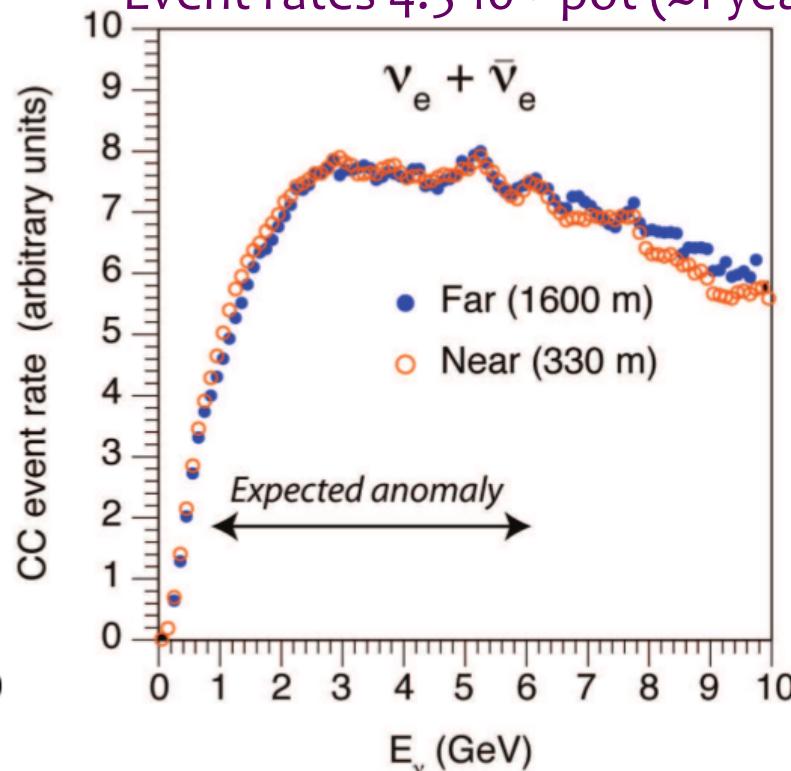
I. Efthymiopoulos - SBL2NA, September 26, 2012

Tentative beam-line and ν CC spectra

On-axis configuration



Event rates $4.5 \cdot 10^{19}$ pot (≈ 1 year)



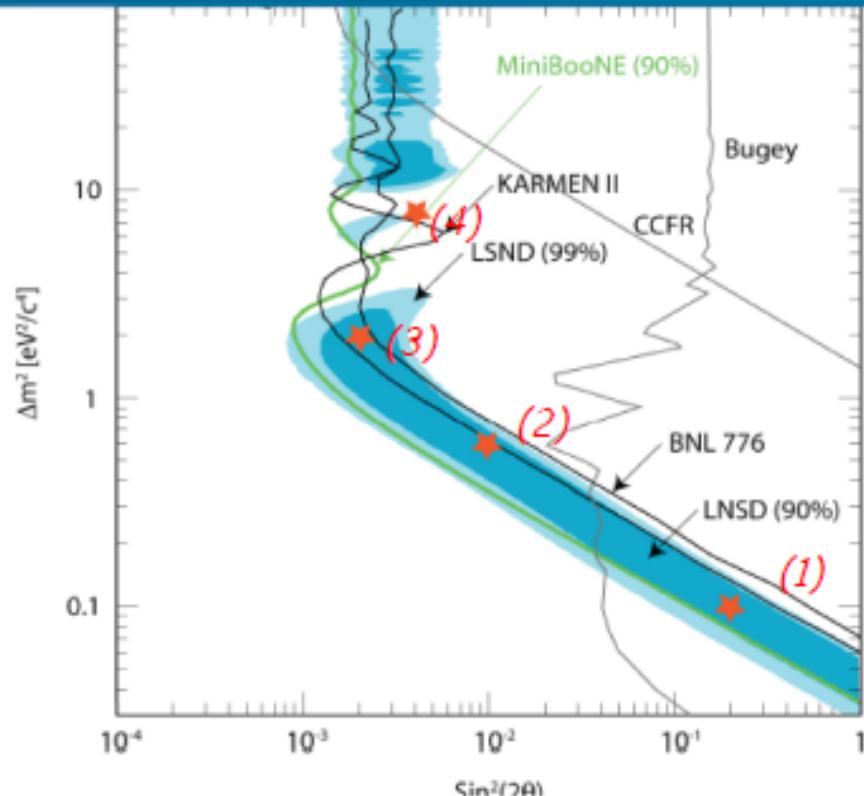
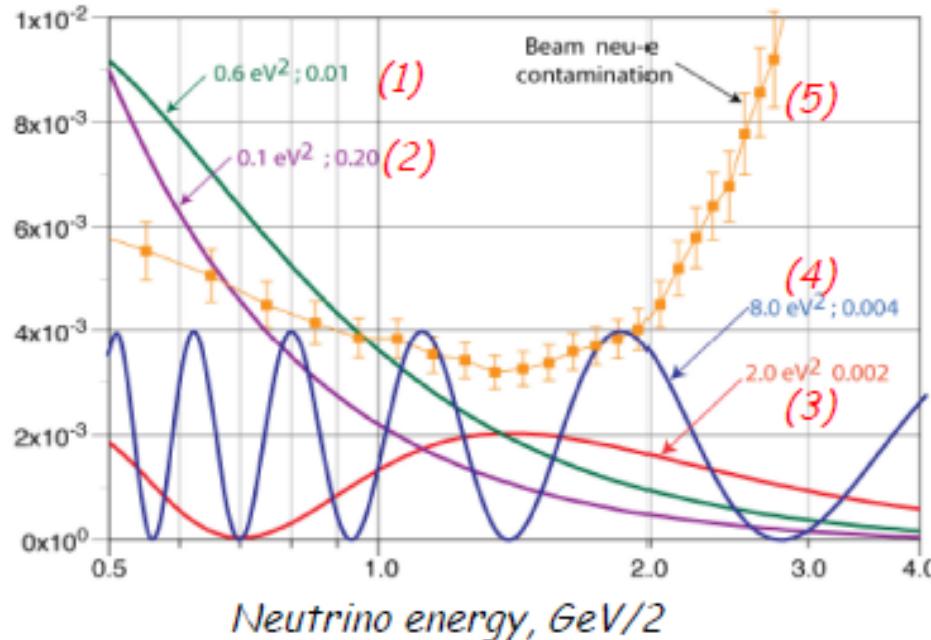
Unoscillated ν_e fluxes are \sim identical \rightarrow N/F deviations = oscillations
The oscillated signals are clustered below 6 GeV of visible energy

LAr-TPC and Spectrometer Trigger

- Expected ν event rate in each LAr-TPC in FAR position : 0.18 (0.10) per 10.5 μs spill ($2 \cdot 10^{13}$ pot) for positive (negative) focusing.
- Additional beam related events: 0.15 (0.08)/spill/TPC. In spill Cosmic ray: ~0.11 μ/spill/TPC easily recognized by the different event topology.
- Coincidence of PMT signals within SPS p-spill gate will be used to trigger event acquisition separately on each T600-TPC.
- In T150 p-spill signal will be used to trigger data collection, due to multiple neutrino interactions (5 (2.5) events/spill/TPC).
- Moreover PMT signals & charge deposition in TPC wires (SuperDaedalus boards) will also be used for event localization/ data reduction.
- Spectrometer: few (several tens) events/spill in FAR (NEAR), triggerless mode within spill gate, events are time-stamped, DAQ transmission in inter-spill time.

LSND direct determination of mass and mixing angle

Two neutrino oscillation pattern at 1.6 km



- The present method, unlike LNSD and MiniBooNE, determines **both** the mass difference and the value of the mixing angle.
- Very different and clearly distinguishable patterns (1-4) are possible, depending on the values in the $(\Delta m^2 - \sin^2 2\theta)$ plane.
- The intrinsic ν -e background (5) is also shown.