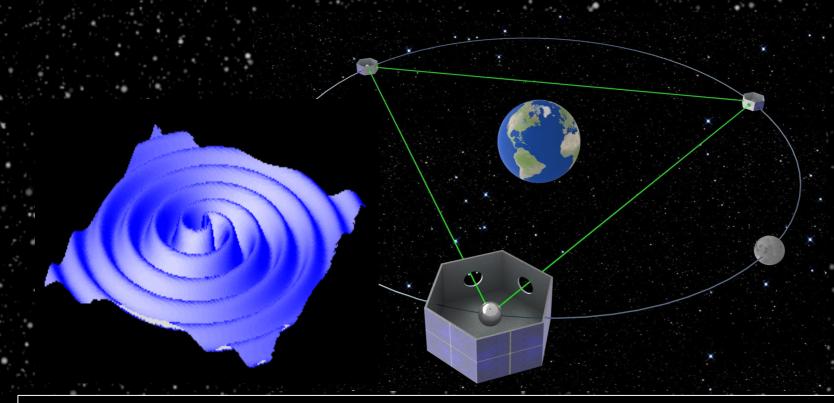
# Gravitational Wave Astrophysics The Next Frontier in Understanding the Universe



# LISA-2020 an Intermediate-Scale Space Gravitational Wave Observatory for This Decade

Gravitational Waves
Advanced Detectors Workshop
GWADW 20<sup>th</sup> -24<sup>th</sup> May 2013

Sasha Buchman
Stanford University
for the Space Sciences team

# LISA-2020 (LAGRANGE based) Collaboration

TELAND JUNIOR LINE STATE OF THE	NASA	LOCKHEED MARTIN	مدينة الملك عبد العزيز للعلوم و التقنية KACST	
Stanford	NASA Ames	Lockheed	KACST of	SRI
	Res. Center	Martin	Saudi Arabia	International
- Science	<ul> <li>Science orbit,</li> </ul>	■ Telescope,	<ul> <li>Science payload,</li> </ul>	- μN thrusters
<ul> <li>Payload lead</li> </ul>	Orb. injection,	■ Spacecraft	■ Tech. development	
GRS / IMS	Prop. mod.			

SSACHUSEI III	Haring Aller
MIT	University
	of Florida
- μN thrusters	• ATC
	■ DF control

#### **Outline**

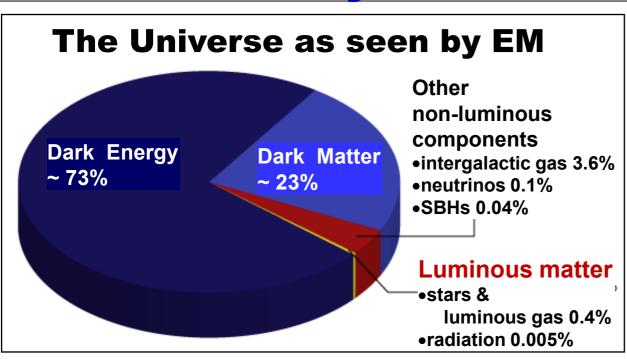
Why Gravitational Wave (GW) Astronomy?

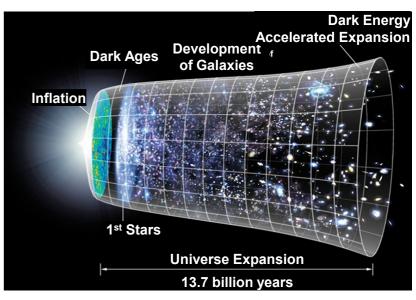
What Is the Status of GW Astronomy?

How Do We Go From Here to LISA 2020 ?

When Can We Fly 'THE NEXT' GWD in Space?

# Today's 'DARK' Universe





#### What do we really know?

- ➤ Universe known by EM; only ~0.5% of matter
- ➤ Continuous 'model improvements' last 100 years

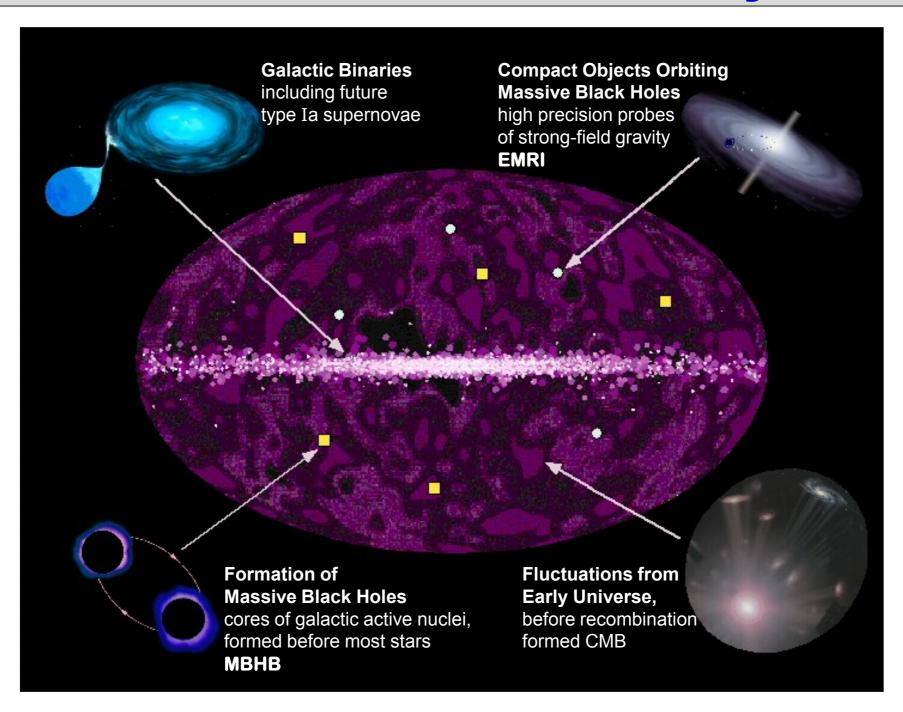
  BH Universes and paradoxes, CPT & LI violations, Unification?
- > GW 'see and interact' with 100% of matter
- > SM & GR used for converting EM data to Universe picture have 'issues'

Seen in EM "understood" in GR

# Why GW Astronomy

- Gravitational Wave (GW) Astronomy Will Give the Answers About the Universe That EM Cannot Provide Not really supported by the space community
- ➤ The 10<sup>-4</sup> Hz to 1 Hz is the 'Richest' GW Range
  This Range Requires a Space GW Observatory
- > A Laser Interferometer Space Antenna (LISA) Is Necessary and Possible by 2020:
  - > Will Achieve the Most Important GW Science
  - > At "Affordable Cost" (\$500M)
- > Support of Science Community Critical for LISA Series

# The Gravitational-Wave Sky



#### **Conclusion #1**

Physics & Astrophysics are in a 'DARK' period; GW Astronomy is the most plausible SOLUTION

2

Status and prospects for GW Astronomy

#### Resolution and Sources of GW

> Earth 10 Hz to 1000 Hz, ~2016

Local (100 MPc range) Medium Resolution



> Astronomical Observations <10<sup>-7</sup> Hz, ~ **2017** 

**TBD Sources & Resolution** 



> Space Experiments 10<sup>-4</sup> Hz - 1 Hz, > **2030** 

Large # of Sources & Excellent Resolution



# GW Space Observatories Issues

#### With many caveats which are about 50% probable:

- > eLISA launch NOT BEFORE 2028 (means maybe after 2035)
- ➤ NASA LISA launch NOT BEFORE 2030 (means maybe after 2035)

  (Approval to Mission; 10 -20 years; Hubble, GP-B, LPF, WST ...)

#### Implications:

- > Delay in 'best' information required to understand the Universe
- Difficulty motivating students and scientists to join the field
- > Old technology and lack of program continuity
- > Loss of opportunity to perform in conjunction with LIGO/VIRGO/etc



Few in this audience will have any chance to see LISA type science

### LISA 2020: GW Observatory for This Decade

Cost Reduce ~ 70%

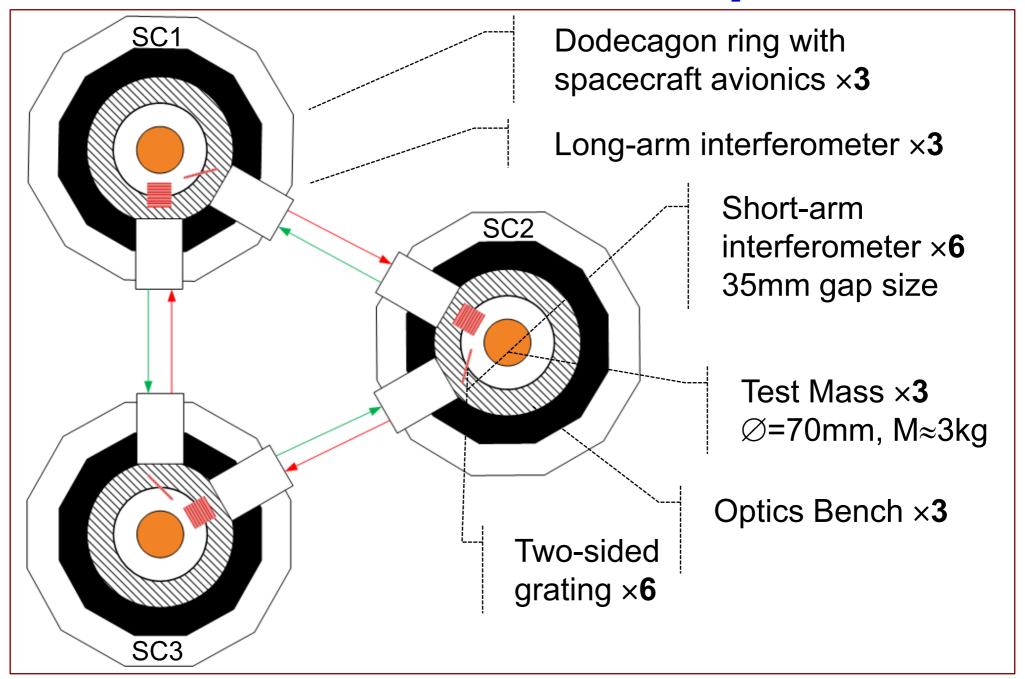
Complexity Reduce ~ 50%

Comm. Link Increase >100

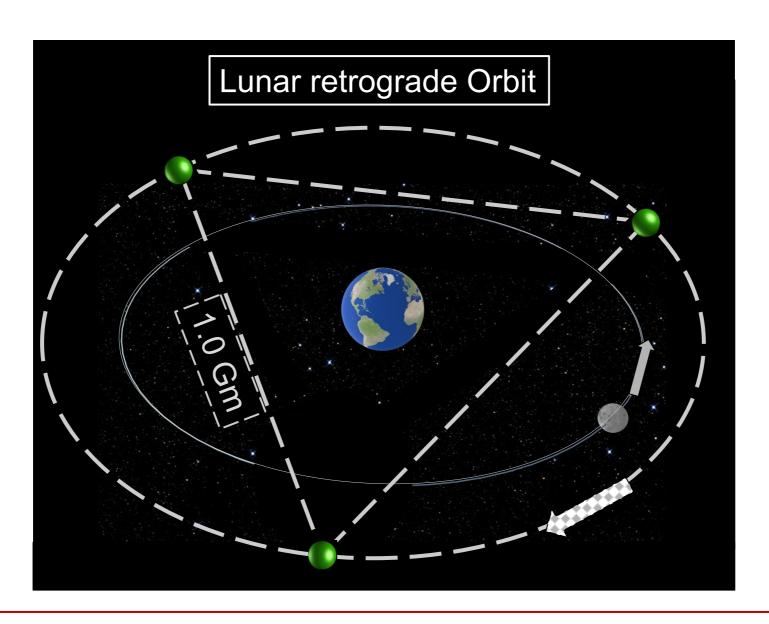
- ➢ Geocentric Orbit: ~ 50% Heliocentric cost
- > Reduced Requirements ~ ×30
- > Small sat approach to tech demonstrations
  - > 2013-2017 technology (LISA technology is older than 2000)
  - > Parallel, low cost, low risk, on small and cube satellites
  - > ~6 technologies at 1 M\$ 4 M\$ each
  - > Multiple institutions and international partners
- > Simplified Robust Inertial Sensor (LPF back-up)
  - Spherical, fully drag-free, optical sensing
- Metrology
  - > Reflective Optics with Gratings

#### LISA-2020 Schematics

# Based on LAGRANGE with 1/30 performance



#### **Overview of LISA-2020 Orbits**



Geocentric Orbits in Lunar Plane; Arm ~ 1 Gm

# Data Rate Estimate for Space Antennas

	GP-B	1 LISA or LISA-2020 SC	3 LISA SC vs GPB
Plan	0.35 GB/day (actual data rate)	0.011 GB/day (NASA) 0.004 GB/day (ESA)	0.033 GB/day (NASA) 0.013 GB/day (ESA)
System			
SC	SC (GPB 6 deg ctrl)	SC (LISA-2020 7 deg ctr) (LISA 7)	≈ (GPB)×3
Temperature	Cryogenics	μK control	≈ (GPB)×3
Propulsion	He thrusters	μN thrusters	≈ (GPB)×3
Pointing	1 telescope	2 telescopes	≈ (GPB)×3 <b>×2</b>
Test Masses	4 TM × 3 deg ctrl.	2 TM × 6 deg control (coupled)	≈ (GPB)×3 <b>×2</b>
Read-out	4 SQUID systems	4 pm interferometers	≈ (GPB)×3
BW	Meas BW 12.9 mHz	Meas. BW 0.1-100 mHz	≥ (GPB)×3
Formation	None	N/A	3 SC ???

GPB data rate ≤ 1 LISA/LISA-2020 SC data rate

LISA/LISA-2020 data rate ≥ 3 × GP-B data rate ≥ 1 GB/day

Estimated LISA/LISA-2020 data rate / Planned LISA data rate (ESA)\* ≥ 77

\*7 kbit/s for 8 hours every 2 days = 0.013 MB/day | ESA web site

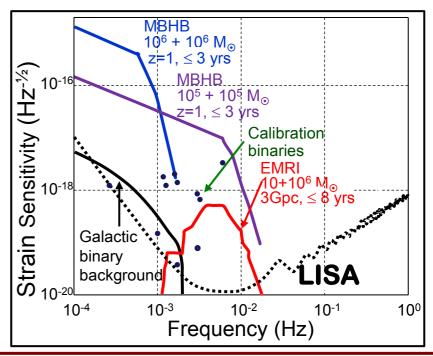
Comm Link
Increase > 100

#### **LISA & LISA 2020**

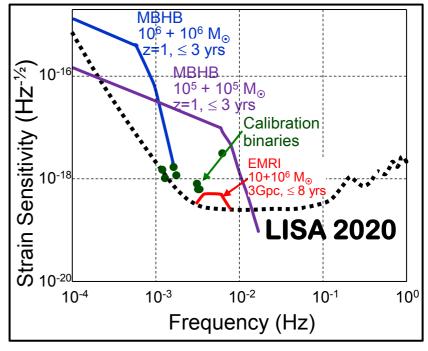
- ➤ LISA: 10<sup>-4</sup> 1 Hz GW in Space

  Laser Interferometer Space Antenna

  "Standard" since 1995
  - > Based on 20 yrs of studies by LISA team
  - Heliocentric Orbit with Three 5 Gm Arms
  - $\geq \delta h/h \approx 10^{-20}$
  - > Cost > 2 G€
  - > Launch AFTER 2030



- **► LISA 2020:** 10<sup>-4</sup> 1 Hz **GW in Space** 
  - Based on 10 yrs of studies by SU team
  - Geocentric Orbit with Three ~1 Gm Arms
  - $\geq \delta h/h \approx 3 \times 10^{-19}$
  - Cost ≈ 1/2 G\$
  - Launch Around 2020



# **LISA & LISA-2020**

	Orbit (Gm)	TM (ms <sup>-2</sup> Hz <sup>-1/2</sup> )	Metrology (pm Hz <sup>-1/2</sup> )
LISA 2020	0.7-1.0	10 <sup>-13</sup>	240
	Geocentric	Sphere × 1	Reflective
LAGRANGE	0.7-1.0 Geocentric	3×10 <sup>-15</sup> Sphere × 1	8 Reflective
LISA	5.0	3×10 <sup>-15</sup>	20
	Heliocentric	Cube × 2	Transmissive

Metric	LISA	LISA-2020
Total MBHB	110-220	20-40
MBHB z > 10	3-60	1-4
EMRIs	800	≤ 10
Total WDB	4 104	≤ 3×10³
WDB with 3D	8×10 <sup>3</sup>	≤ 10 <sup>2</sup>
Stochasic Background	1.0	≤ 0.2



# Principal Cost Savings Relative to LISA

#### 1. Orbit change: Geocentric (0.7 Gm – 1.0 Gm arm length)

- Requires 1 small propulsion module instead of 3
- Launch mass savings: ~ 3,000 kg
- Reduced operations & communications complexity
- Other orbits possible Earth-Sun L1?

#### 2. Reduced S/C mass from simplified payload components

- 1 GRS, 1 Laser, 1 optics bench, smaller (20 cm) telescopes
  - > 2 Lasers budgeted for redundancy (4 in LISA)
  - No credible TM failure mechanism
  - > TM sensing, charge control, spin-up, and drag-free have redundancy
- Launch mass savings: ~ 150 kg × 3 spacecraft
- Any 'available ' compact TM technology OK.

#### 3. LISA-2020 wet launch mass: ~2,000 kg (~5,000 kg for LISA)

- Historic trends show cost scales with mass
- Complex payloads are hard to cost

# Spacecraft & Mission Design by LM Off the shelve but too large

S/C based on existing LM S/C, TRL >6

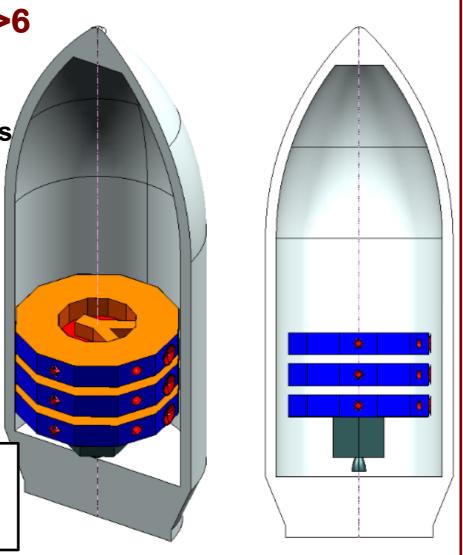
> ~3 m  $\times$  0.7 m, 300 kg, 500 W

Fixed 10 W antenna between telescopes

➤ Thermal design: GRS 10 µK at 1 mHz

- > ±50 K at exterior at 27.3 period
- > Thermal load radiated top/bottom
- Payload at center
- Launch mass: 2,070 kg
- > 4-7 month cruise
- 5 year lifetime

Concept of 3 SC & 1 Propulsion module In Launch Fairing



### LISA-2020 GWD Concept Study

#### LISA-2020 concept with heritage

Honeywell, DISCOS, LPF, ST-7, GP-B, STAR

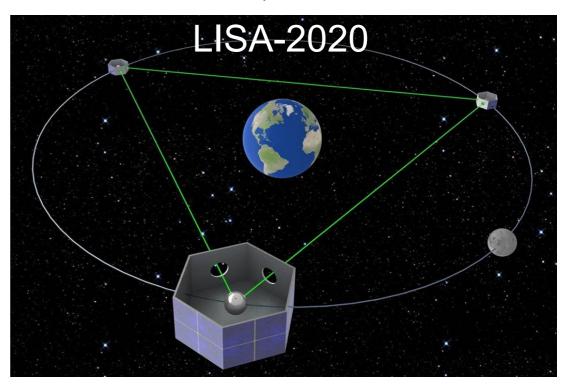
- > 3 drag-free spacecraft in geocentric orbit
- ➤ Minimized payload: 1 test-mass (sphere), 1 laser, 2 telescopes
- > Small sat approach to tech demonstrations

### 2t launch 3 00 kg 3 00 W

#### LISA-2020 maintains LISA science ~ 50%

- > 50% Complexity
- > 30% Cost
- > 10,000% Communications Band





# For and Against LISA 2020

- Advantages
  - ➤ GW Science ~2020
  - > Technology
  - > GW Community
- Obstacles
  - > Funding
  - Competition
    - > EM Astronomy has Data
    - > Planetary Science
  - ➤ Inadequate EPO



#### **Conclusion #2**

Physics & Astrophysics are in a 'DARK' period; GW Astronomy is a very plausible SOLUTION

2 A LISA-2020 Type Geocentric Medium GW Antenna Can Provide Excellent GW Data ~2020

3 Technology Development on Small Satellites

# Science & Technology Implementation on Small Satellites

#### Technology

- Gravitational Reference Sensors
- > Ultra-stable optics
  - > Precision navigation
  - Formation flying

#### **Education**

- Grad, Undergrad
- > 3-5 year projects
- > Student led tasks

- Science & Technology on Small Satellites
- > Education driven
- Internationalcollaborations

#### **Science**

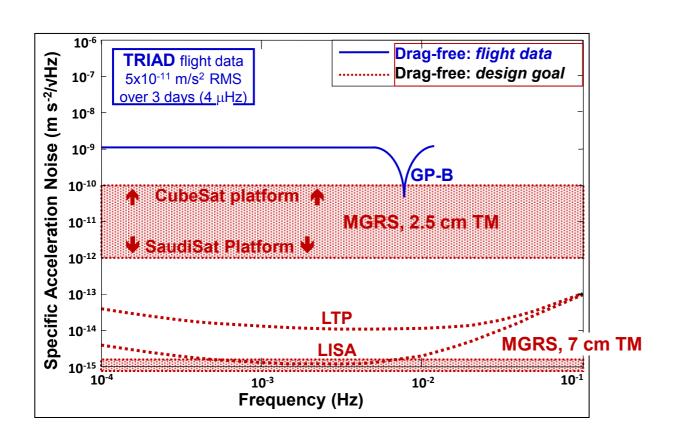
- Special/General Relativity
- Gravitational waves
- Earth Geodesy/Aeronomy

# Stanford Spherical 6 DOF TM Performance

#### 1. Control Spacecraft to follow TM

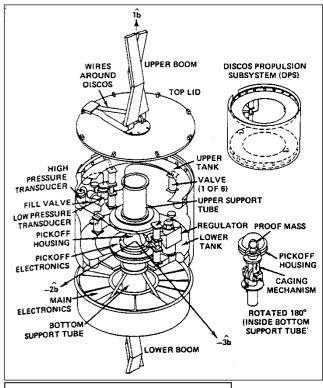
#### 2. Reduce External Disturbances

- Aerodynamic Drag
- Magnetic Torques
- > Radiation Pressure
- Gravitational Torques





GP-B Flight Gyroscope 2004



TRIAD Sensor 1972

# Applications of Drag-free Technology

	Category	Application	Drag-free Performance (m/sec <sup>2</sup> Hz <sup>1/2</sup> ), frequency (Hz)	Metrology (m)
Capability of Cubesat with 2.5 cm TM		Autonomous, fuel efficient orbit maintenance	$\leq 10^{-10}$ , near zero frequency <b>a,b</b>	≤ 10 absolute
	Navigation	Precision real-time on- board navigation	$\leq 10^{-10}$ , near zero frequency <sup>a</sup>	≤ 10 absolute <sup>a</sup>
		Formation flying	$\leq 10^{-10}$ , near zero frequency <sup>a</sup>	≤ 10 <sup>-9</sup> differential <sup>a</sup>
	Earth &	Aeronomy	$\leq 10^{-10}$ , $10^{-2}$ to 1 Hz <sup>a</sup>	1 absolute <sup>a</sup>
	Planetary	Geodesy, GRACE	$10^{-10}$ , $10^{-2}$ to 1 Hz <b>a</b> , <b>b</b> , <b>c</b>	10 <sup>-6</sup> differential <sup>a</sup>
	Science	Future Earth geodesy	$\leq 10^{-12}$ , $10^{-2}$ to 1 Hz <sup>a</sup>	≤ 10 <sup>-9</sup> differential <sup>a</sup>
	Fundamental Physics	Equivalence Principal tests	$\leq 10^{-10}$ , $10^{-2}$ to 1 Hz <sup>a</sup>	≤ 10 <sup>-10</sup> differential <sup>a</sup>
		Tests of general relativity	$\leq 10^{-10}$ , near zero frequency <sup>a</sup>	≤ 1 absolute <sup>a</sup>
7 cm TN	<b>Astrophysics</b>	Gravitational waves	3×10 <sup>-15</sup> , 10 <sup>-4</sup> to 1 Hz	≤ 10 <sup>-11</sup> differential

Notes: <sup>a</sup> Performance to be demonstrated by the drag-free CubeSat; <sup>b</sup> demonstrated; <sup>c</sup> non-drag-free Courtesy John Conklin

# Advantages of a Spherical TM

#### 1. No TM forcing or torquing

 Neither electrostatic support nor capacitive sensing required, reducing disturbances & complexity

#### 2. Large gap (35 mm)

> Disturbances reduced and/or spacecraft requirements relaxed

#### 3. Long flight heritage

 $\rightarrow$  Honeywell gyros, Triad I (5×10<sup>-11</sup> m/sec<sup>2</sup>), GP-B (4×10<sup>-11</sup> m/sec<sup>2</sup> Hz<sup>1/2</sup>)

#### 4. Scalability

> Performance can be scaled up or down by adjusting TM and gap size

#### 5. Simplicity

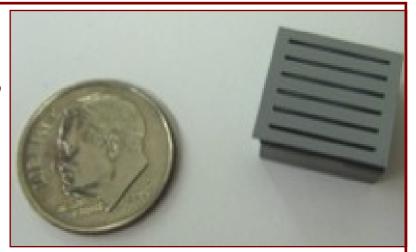
No cross coupling of degrees of freedom

#### 6. Simple flight-proven caging mechanism (DISCOS)

NEGATIVE: The Mirror Moves

# Micronewton Thrusters Design

- ➤ Drag-free & attitude via µN thrusters
- ➤ No existing thruster meets LISA noise, max thrust, and lifetime requirements
  - > LPF evaluating alternates to FEEPs
- ➤ MIT & SRI micro-fabricated ion thrusters as attractive alternative to Busek CMNT or Italian/Austrian FEEPs

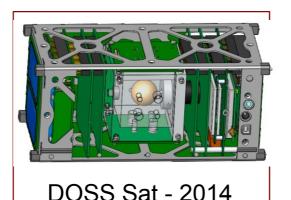


- > Micro-fabricated emission sites produce ions & electrons
- > "Digital propulsion": 100's 1,000's of independent emitters / cm<sup>2</sup>
  - > Single unit can produce forces + torques
- ➤ Huge dynamic range: ion production physics unchanged over 10<sup>-9</sup> to 1 N
- Up to 10,000 sec Isp
- > Prototype: 1 nN to 5 μN thruster ion source tested to 40 hr of operation
- Can be demonstrated on a 1U CubeSat
- > MIT uses capillarity; no moving parts

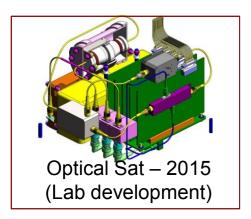
# Thrusters are a problem

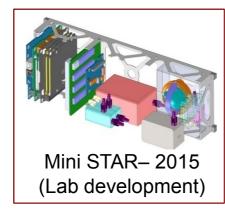
# Small Sats Technology Program

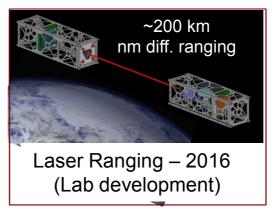










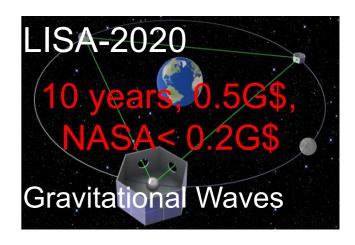


df/f ~ 10<sup>-12</sup>
1mm optical cavity
1 mm gas cell
25 cm<sup>3</sup>, 25 g, <100 mW

Mini clock Sat – 2016 (Lab development)



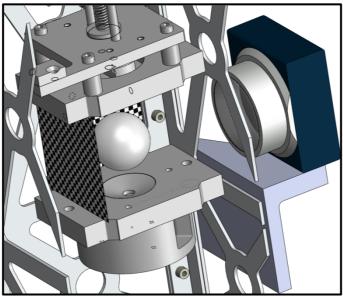


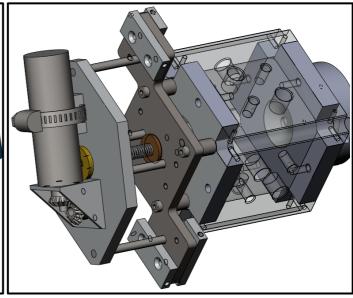


# Caging System - April 2013 Parabolic Flight

### MGRS, 2.5 cm TM, for Parabolic Flight Caging Test



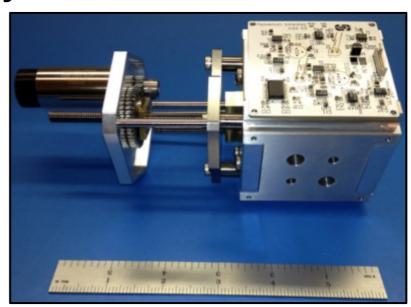




**Caging System Schematics** 



Housing

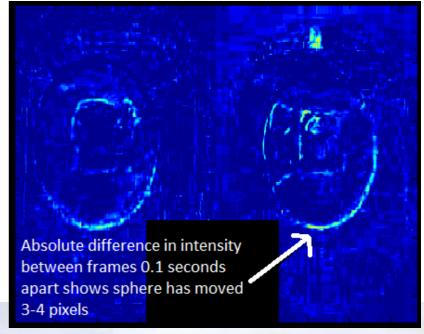


MGRS, Mechanical

**3 U Caging Fixture** 

# Caging System - April 2013 Parabolic Flight





# Flight Team (from left) April 22<sup>nd</sup> – 25<sup>th</sup>

- >Andreas Zoellner
- **≻Kirk Ingold**
- **≻Eric Hultgren**



Flight4-22a.m4v

#### **UV LED Small Satellite**

#### Technology Objectives

- Raise TRL levels  $(4/5 \rightarrow 8/9)$  for
  - Deep UV LEDs
  - ac charge control
- Beneficiaries:
  - LISA
  - GRACE follow-on
  - Drag-free CubeSat

#### Mission Design

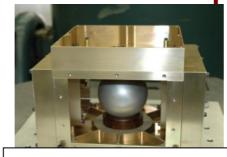
Spacecraft: Saudi Sat

55 kg 50 W Saudi Sat 3

- Russian launch Nov 2013
- 2 month mission
- Fully funded (\$1.5M)

#### **Payload**

- Isolated "test mass"
- 16 UV LEDs & photodiodes
- Charge amp
- Voltage bias plates
- ac charge control electronics



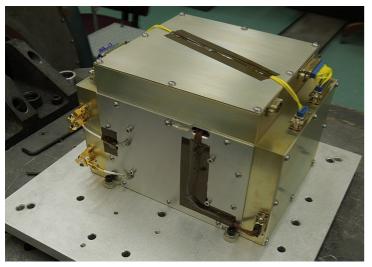
222×277×180 mm; 6.5 kg

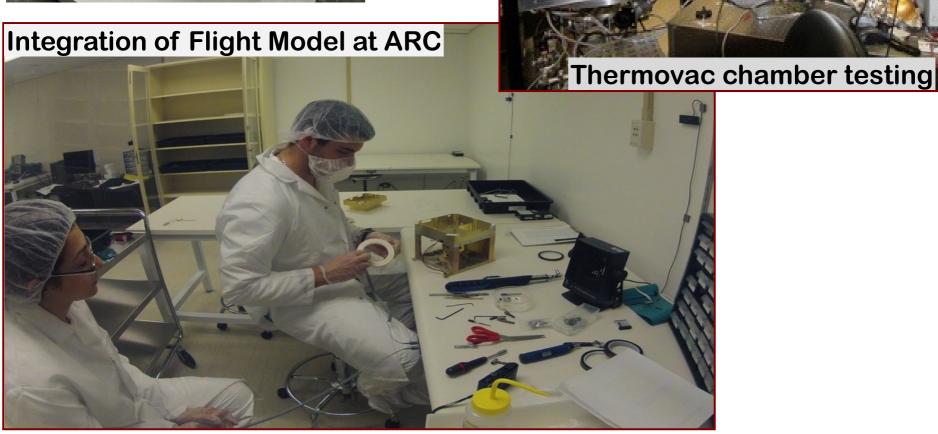
#### Management

- NASA Ames: Flight payload, PM, SE, SMA
- Stanford: Payload design, SOC
- KACST: Spacecraft, Launch, MOC

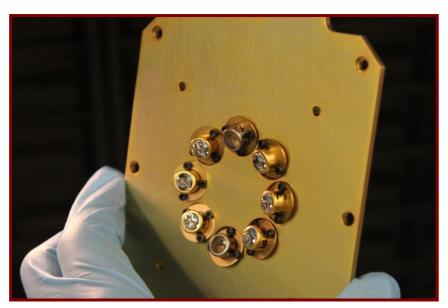
Demonstrates unconventional international collaboration

# **UV LED Instrument Integration and Test**





# **UV LED Instrument Components; 2013 Launch**



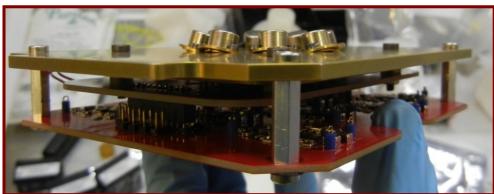
■ Payload completion: May 2013

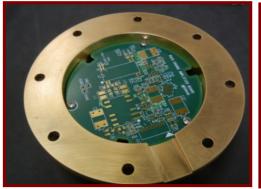
■ PL-SC Com Interface: May 2013

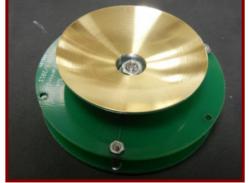
■ Spacecraft CDR: May 2013

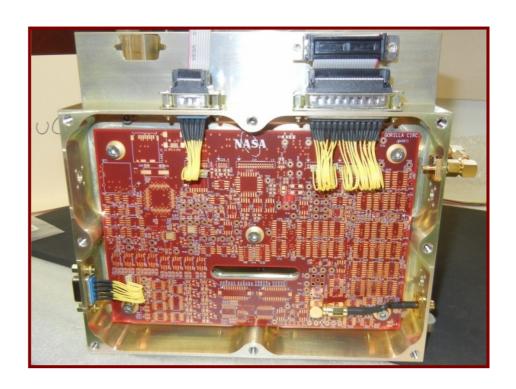
■ Payload Integration: Jun 2013

■ Russian launch: Nov 2013





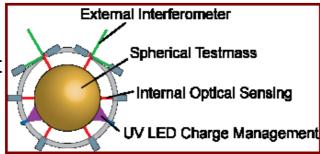




# Differential Optical Shadow Sensor (DOSS)

#### Technology Objectives

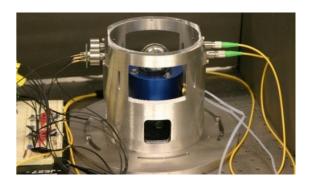
- Raise TRL level for miniature highsensitivity displacement sensor
  - ■nm/Hz<sup>1/2</sup> sensitivity
  - No forcing
  - Non-contact



#### Mission Design

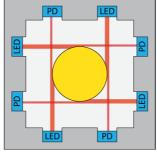
- 3U CubeSat
- Any orbit
- Launch ~ 2014
- 1 month ops
- Payload funded





#### **Payload**

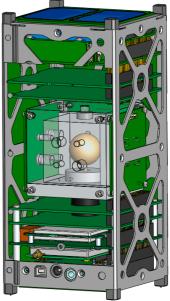
- Light source:
  - SLED, 1545 nm
- InGaAs quad-photodiode
- Ultra-low current Difet amp





#### Management

- Stanford & KACST: Payload, CubeSat structure
- I&T & Launch: pending



# DOSS & ADCS on 3U Cubesat; 2014 Launch

Secondary Payload

**ADCS Evaluation** 

**Electrical Power System** 

Volume Margin

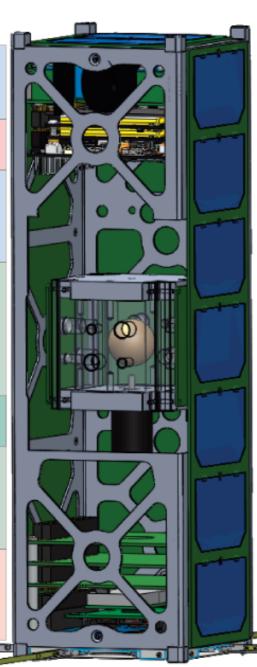
Differential Optical Shadow Sensor

**Motor Drive** 

**Payload Processor** 

Motherboard, CPU, Radio, Antenna

Main Payload

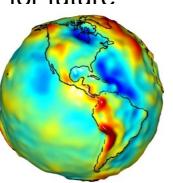


Bus

# The Drag-free CubeSat

#### Science

- Aeronomy, space weather
- Demo < 10<sup>-10</sup> m/sec<sup>2</sup> for future
  - Planetary Geodesy
  - Earth observation
  - Gravity science
  - Gravity-waves

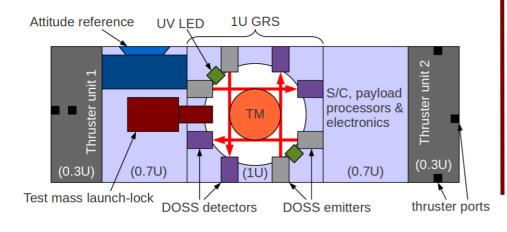


#### Mission Design

- 3U CubeSat
- Secondary launch via P-POD
- Launch ready ~ 2015
- 1-2 month drag-free ops in low g environment < 10<sup>-8</sup> m/s<sup>2</sup>

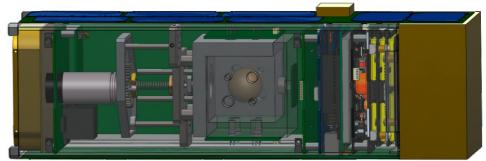
#### Payload (back-up version)

Drag-free sensor + micro-thrusters



#### Management

- NASA ARC: PM, SE, SMA, MO
- Stanford: Payload design, drag-free control, data analysis





# The Drag-free CubeSat

4 kg 6 W 3U Cube

Caging System

Rate Gyro and GPS

Thruster

Motherboard, CPU and Radio

Shadow Sensor UV LED

Payload with Test Mass **Electrical Power System** 

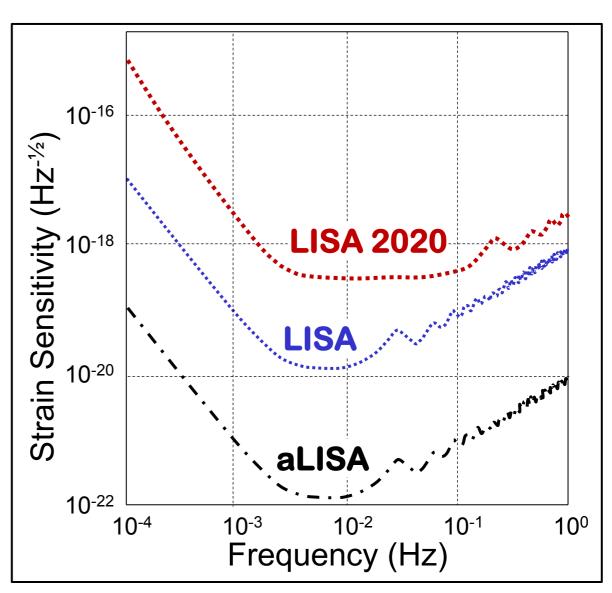
**ADACS** 

# **GWD** in Space

> LISA 2020: ~2020

>LISA: ~2030

> aLISA: ~2045



#### **Conclusions**

Physics & Astrophysics are in a 'DARK' period; GW Astronomy is a very plausible SOLUTION

A LISA-2020 Type Geocentric Medium GW Antenna Can Provide Excellent GW Data ~2020

Technology Development on Small Sats Provides the Road to LISA-2020 & Significant Science

# Thank you for your attention

