

Propagation of Cosmic Ray in the Heliosphere

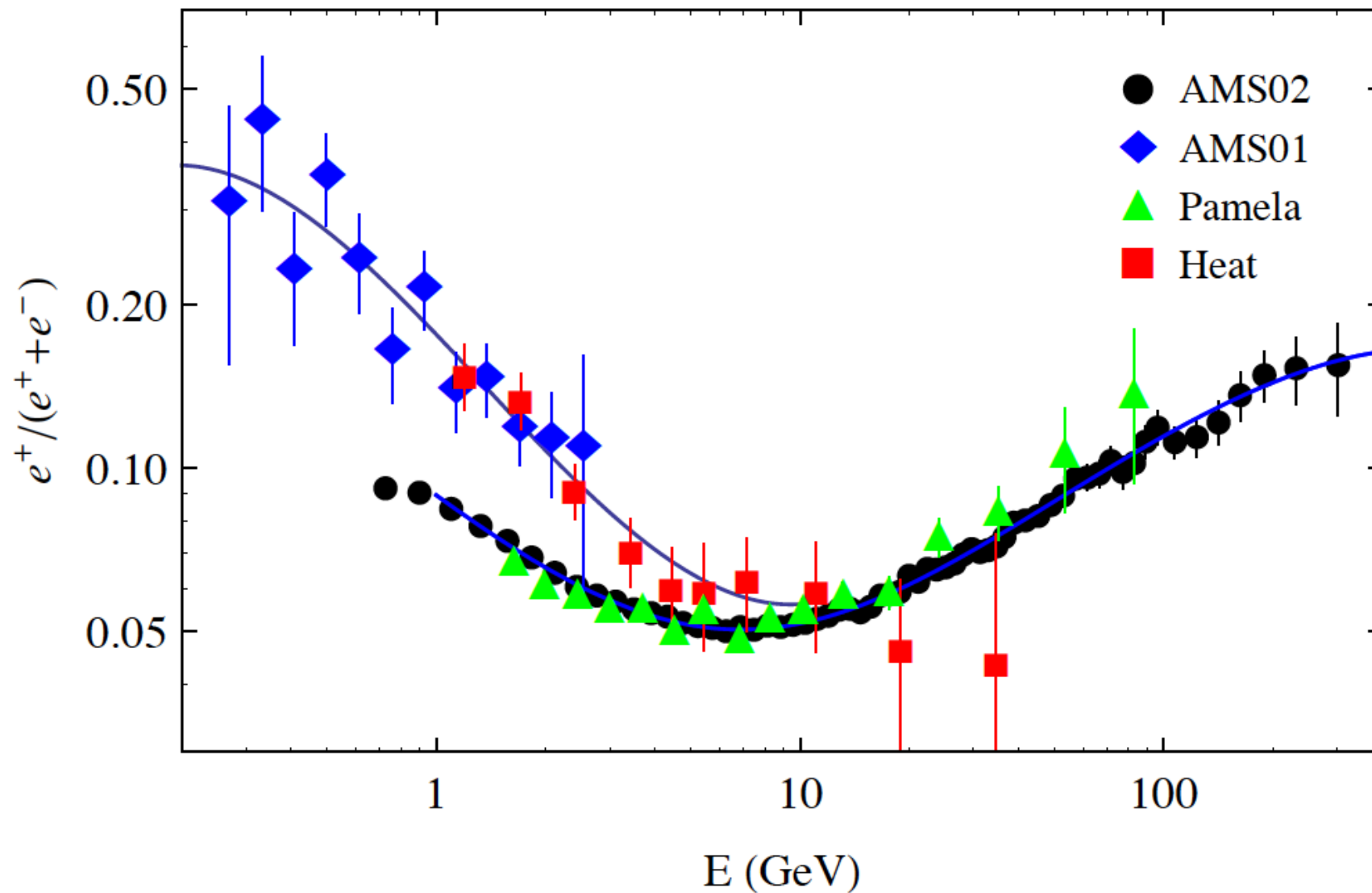
Paolo Lipari, INFN Roma

RICAP 2013

Roma, 23rd may 2013

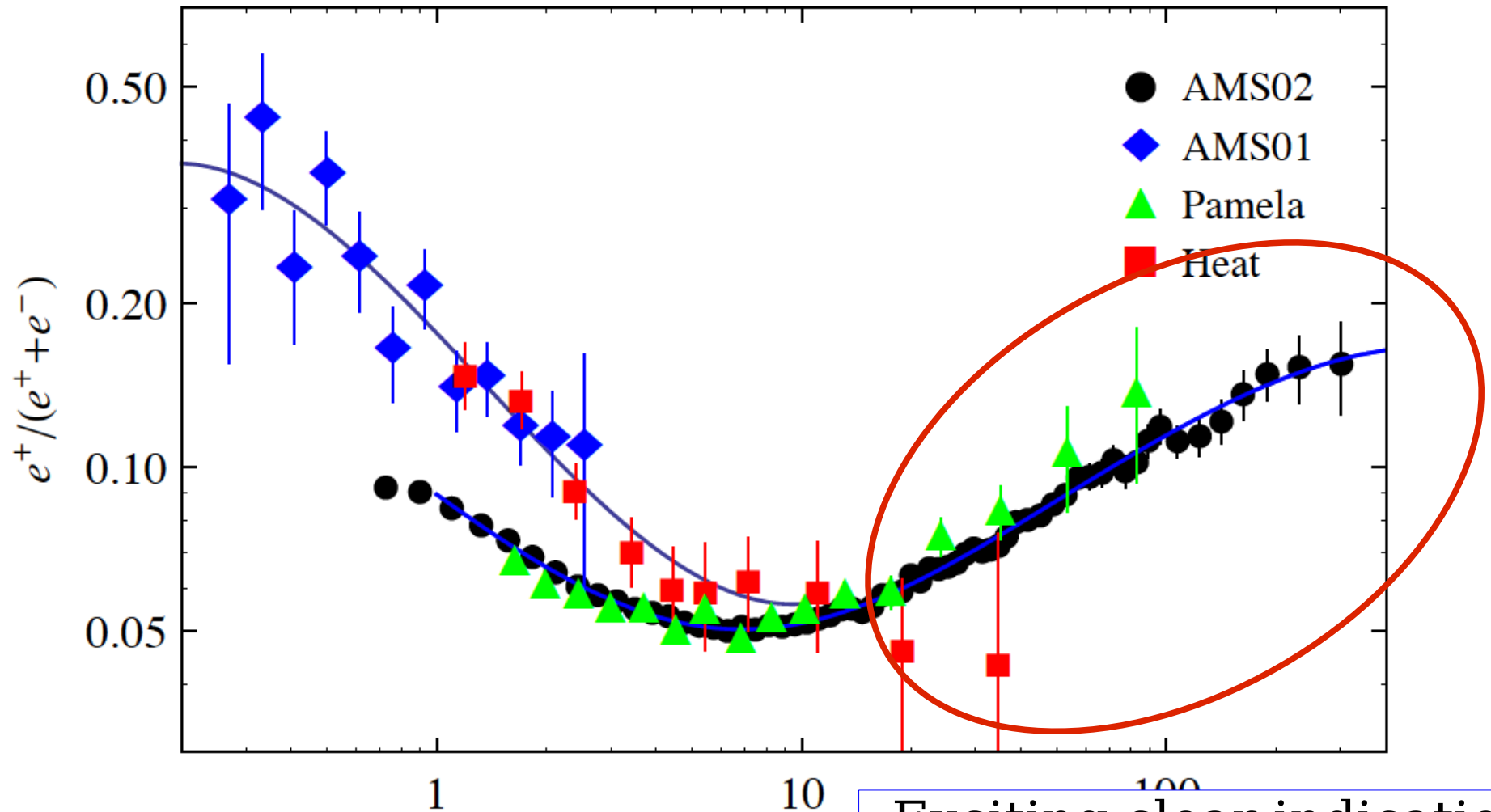
Recent Measurement of the of the ratio by AMS02.

$$\frac{e^+}{e^+ + e^-}$$



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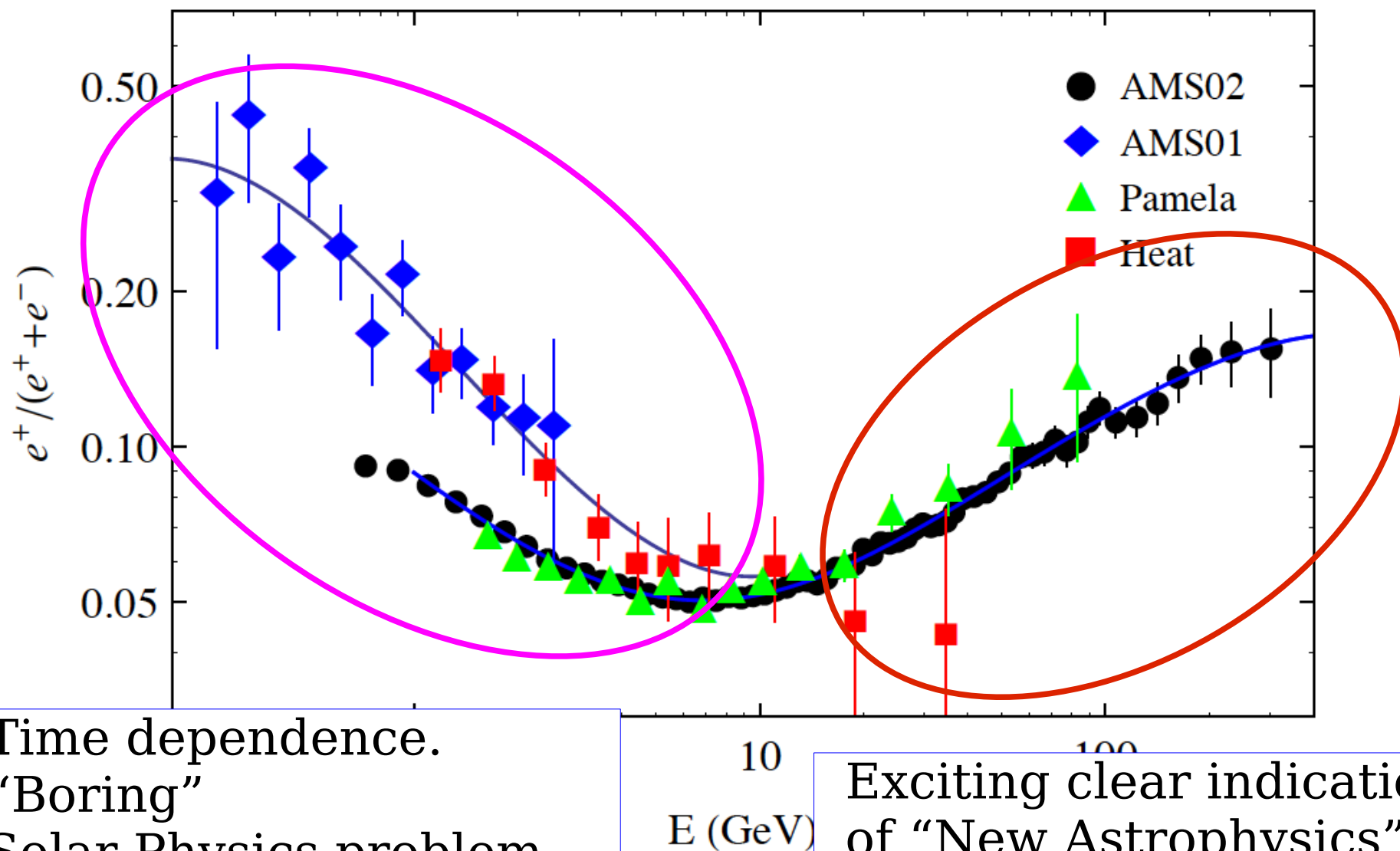
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Exciting clear indication of "New Astrophysics" or "New Physics" (DM)

Recent Measurement of the of the ratio by AMS02.

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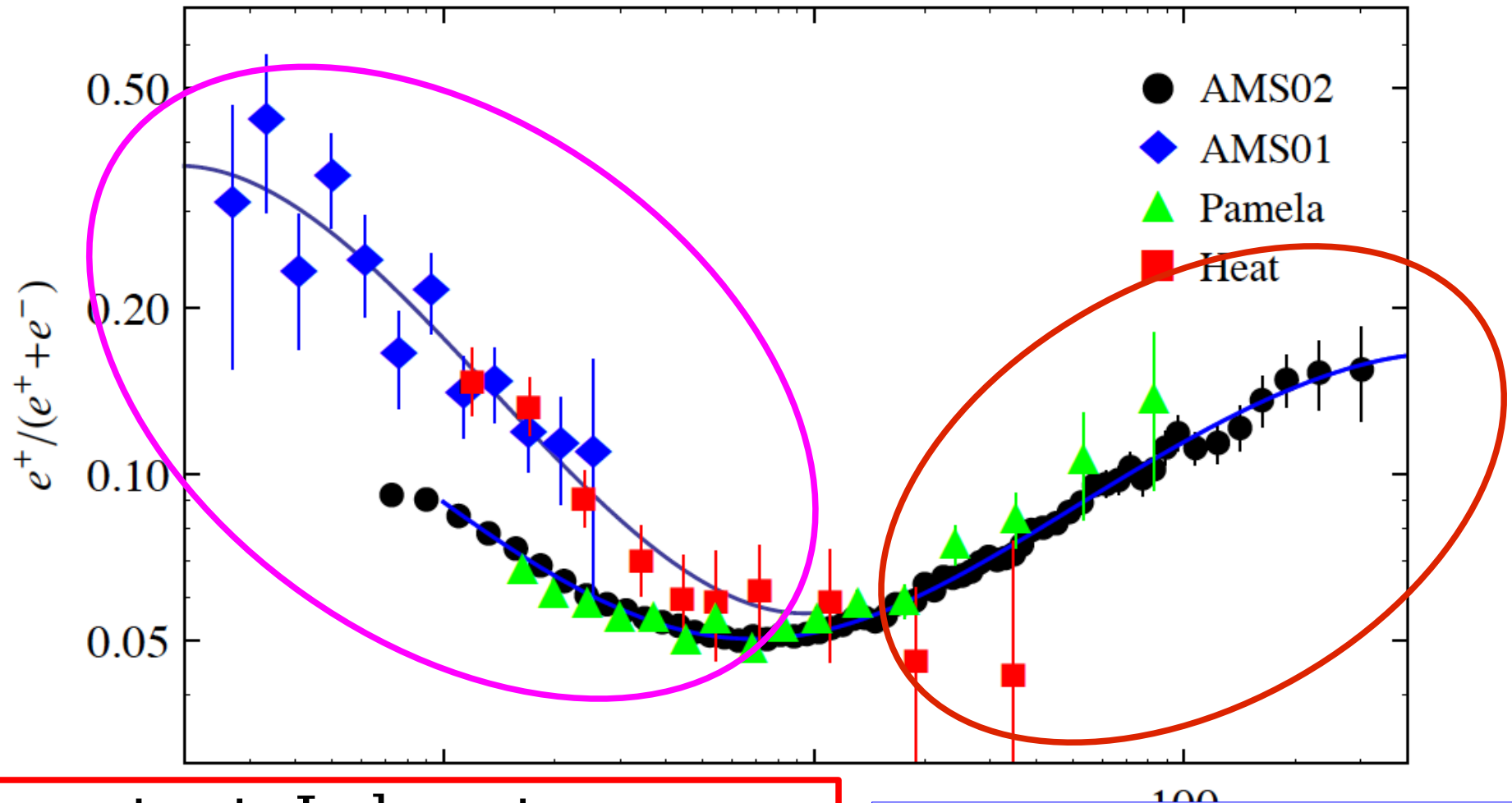


Time dependence.
“Boring”
Solar Physics problem.....

Exciting clear indication
of “New Astrophysics”
or “New Physics” (DM)

Recent Measurement of the of the ratio by AMS02.

$$\frac{e^+}{e^+ + e^-}$$



Important Laboratory for the study of Cosmic Ray Propagation!

Exciting clear indication of "New Astrophysics" or "New Physics" (DM)

Propagation in the Heliosphere

Propagation in the Milky Way

$$n(E) = \frac{1}{\beta c} \int d\Omega \phi(E, \Omega)$$

$$n(E) = q(E) \times T(E)$$

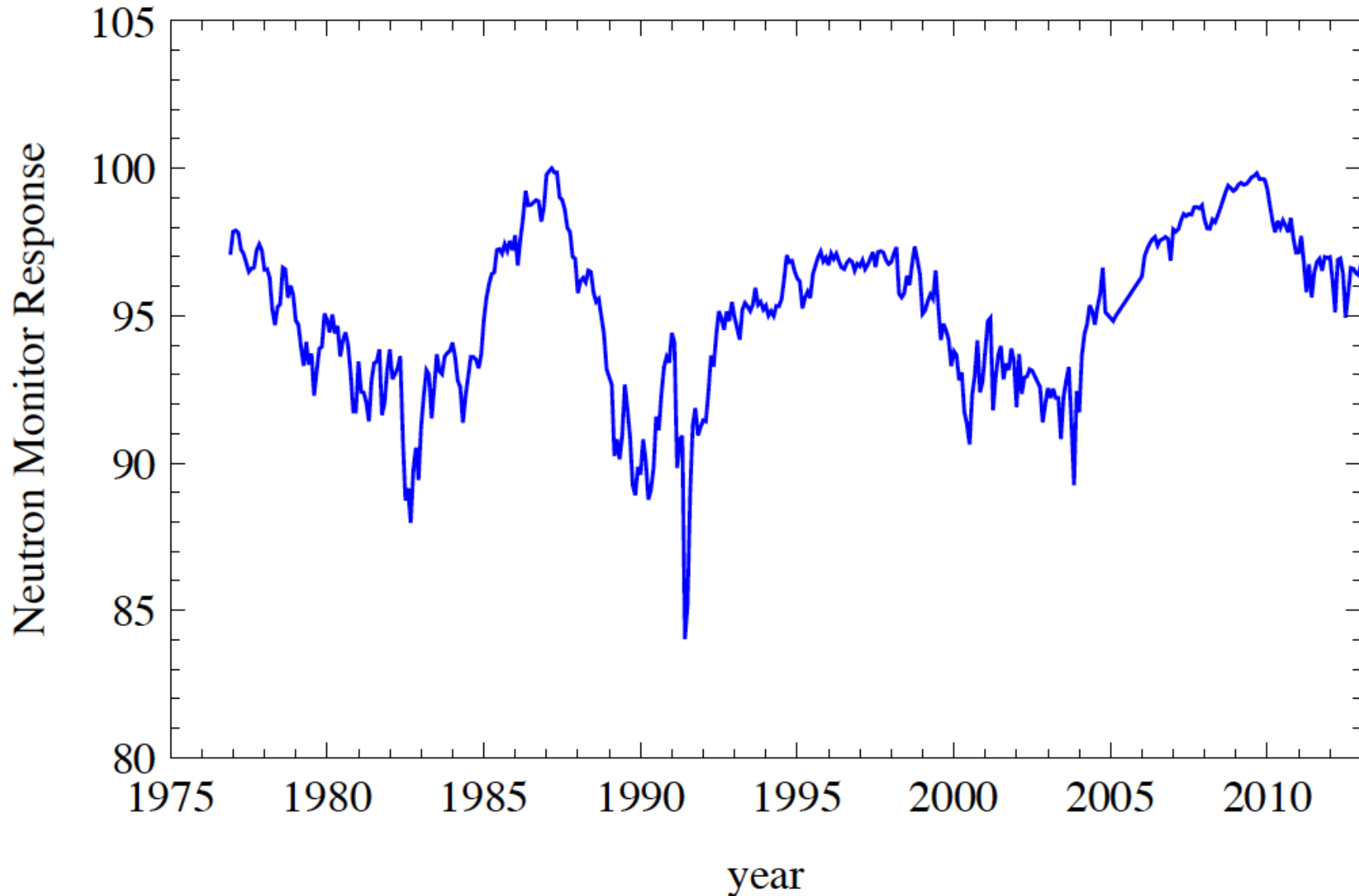
$$n(E) = q(E) \times \frac{L^2}{2 D(E)}$$

$$D(E) \simeq D_0 E^\delta$$

$$D_\perp(E)$$

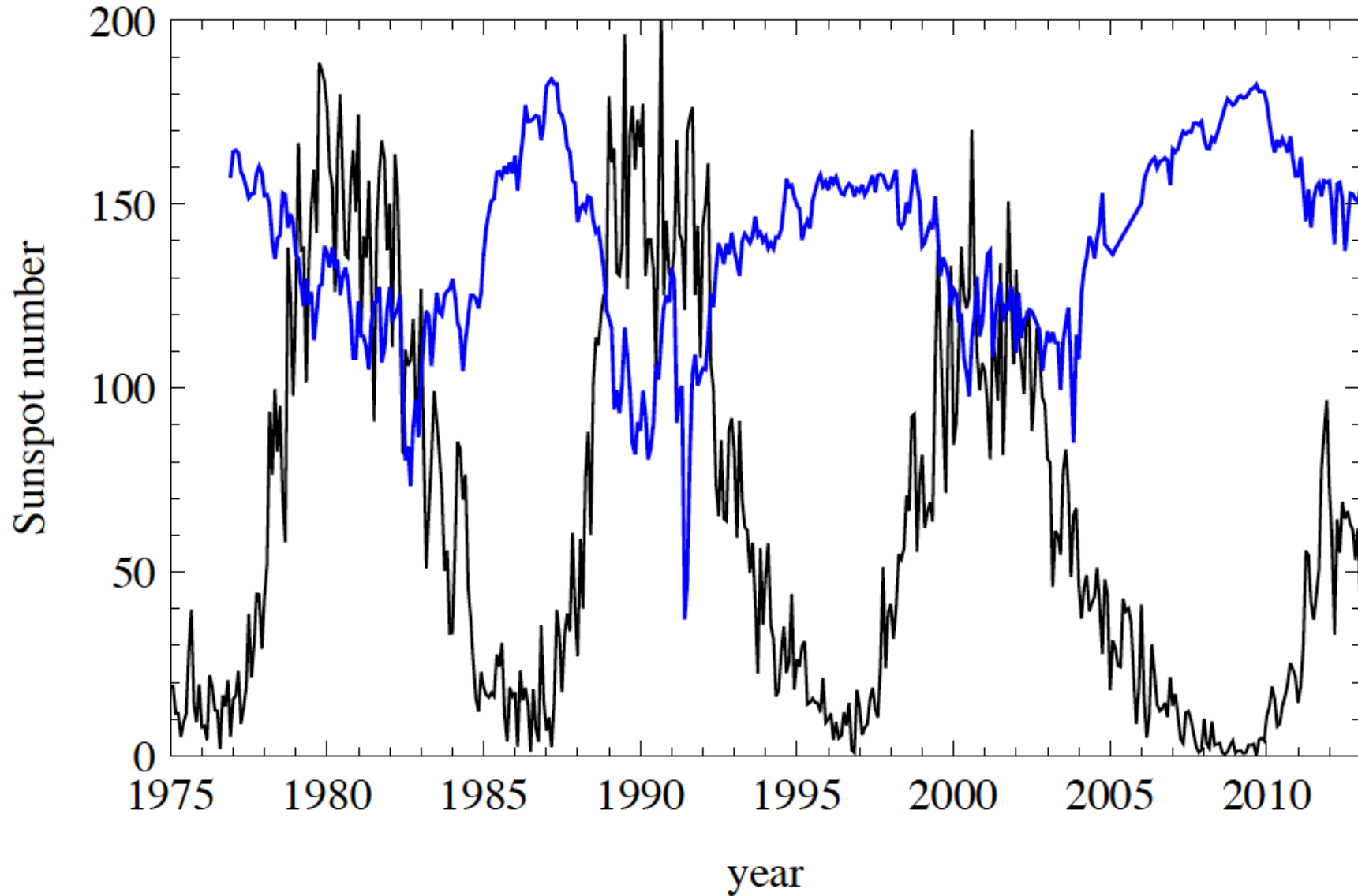
$$D_\parallel(E)$$

Time dependence of the Cosmic Ray Flux:

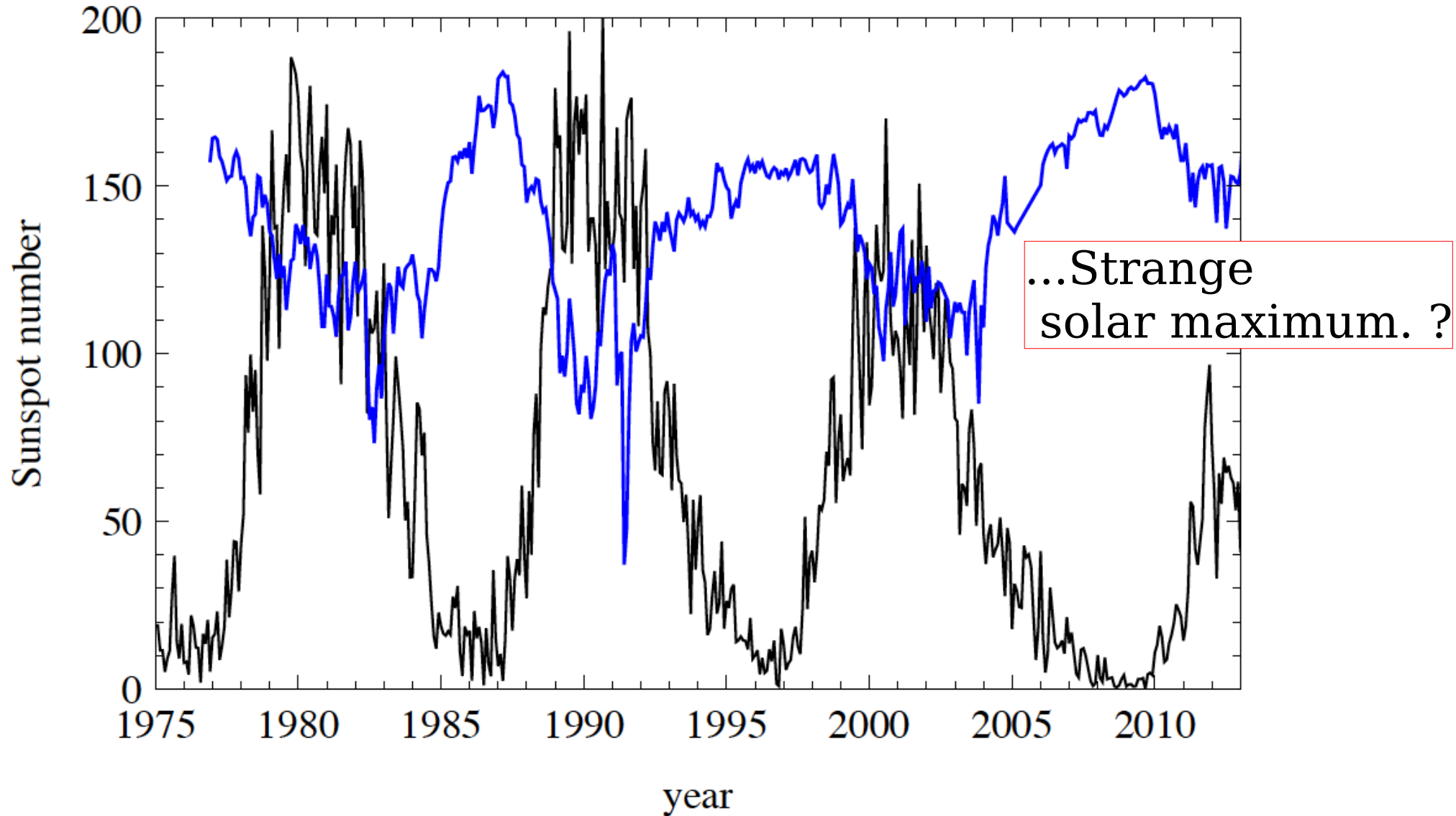


Hermanus (South Africa) Neutron Monitor

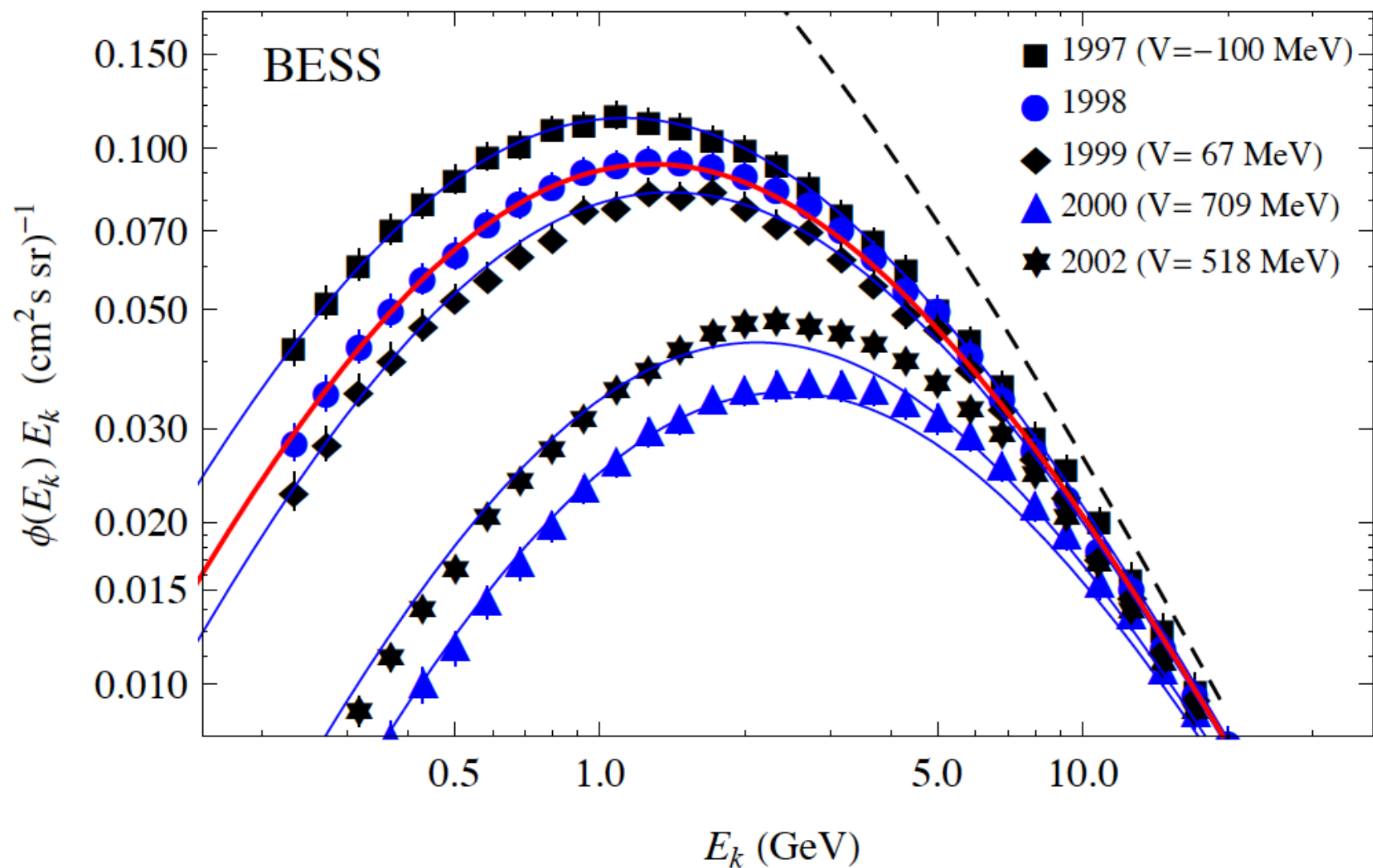
Correlation with Solar Activity (Sunspot Number)



Correlation with Solar Activity (Sunspot Number)



Precision Measurements of CR (proton spectra) at different times [BESS]



Flux in the “Local Interstellar Space” (LIS)
(outside the heliosphere)

$$\phi_{\text{LIS}}(E)$$

Flux at the Earth at time t

$$\phi(E, t)$$

Flux in the “Local Interstellar Space” (LIS)
(outside the heliosphere)

$$\phi_{\text{LIS}}(E)$$

Flux at the Earth at time t

$$\phi(E, t)$$

Phenomenological analysis.
Relation between 2 fluxes:

$$\phi(E, t_1)$$

$$\phi(E, t_2)$$

$$\phi(E, t_2) = \phi(E + V_{21}, t_1) \frac{E^2 - m^2}{(E + V_{21})^2 - m^2}$$

One parameter transformation.

Transformation: “Force - Field algorithm”

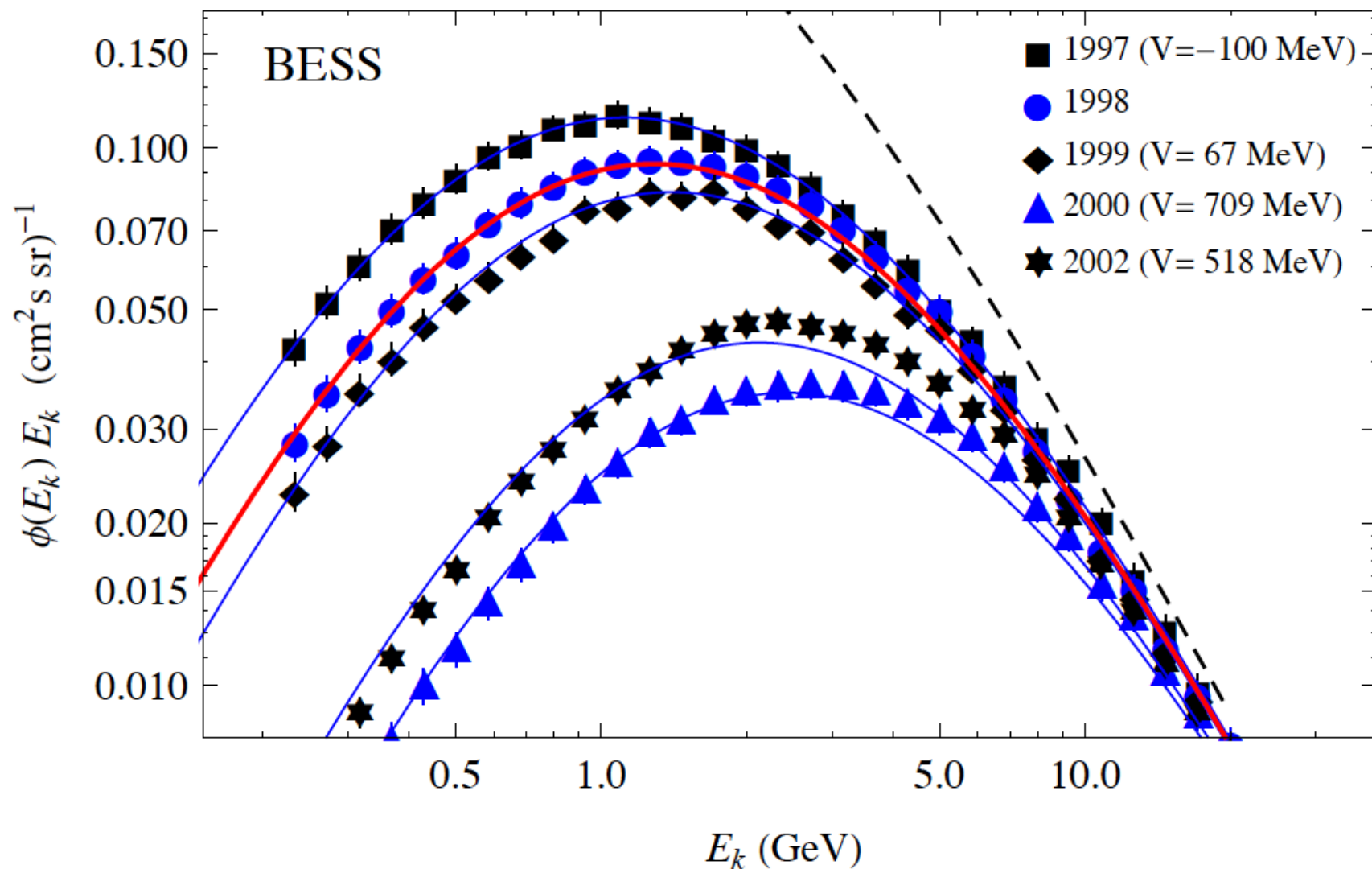
$$\phi(E, t_2) = \phi(E + V_{21}, t_1) \frac{E^2 - m^2}{(E + V_{21})^2 - m^2}$$

inversion: $V_{12} = -V_{21}$

$$\phi(E, t_1) = \phi(E + V_{12}, t_2) \frac{E^2 - m^2}{(E + V_{12})^2 - m^2}$$

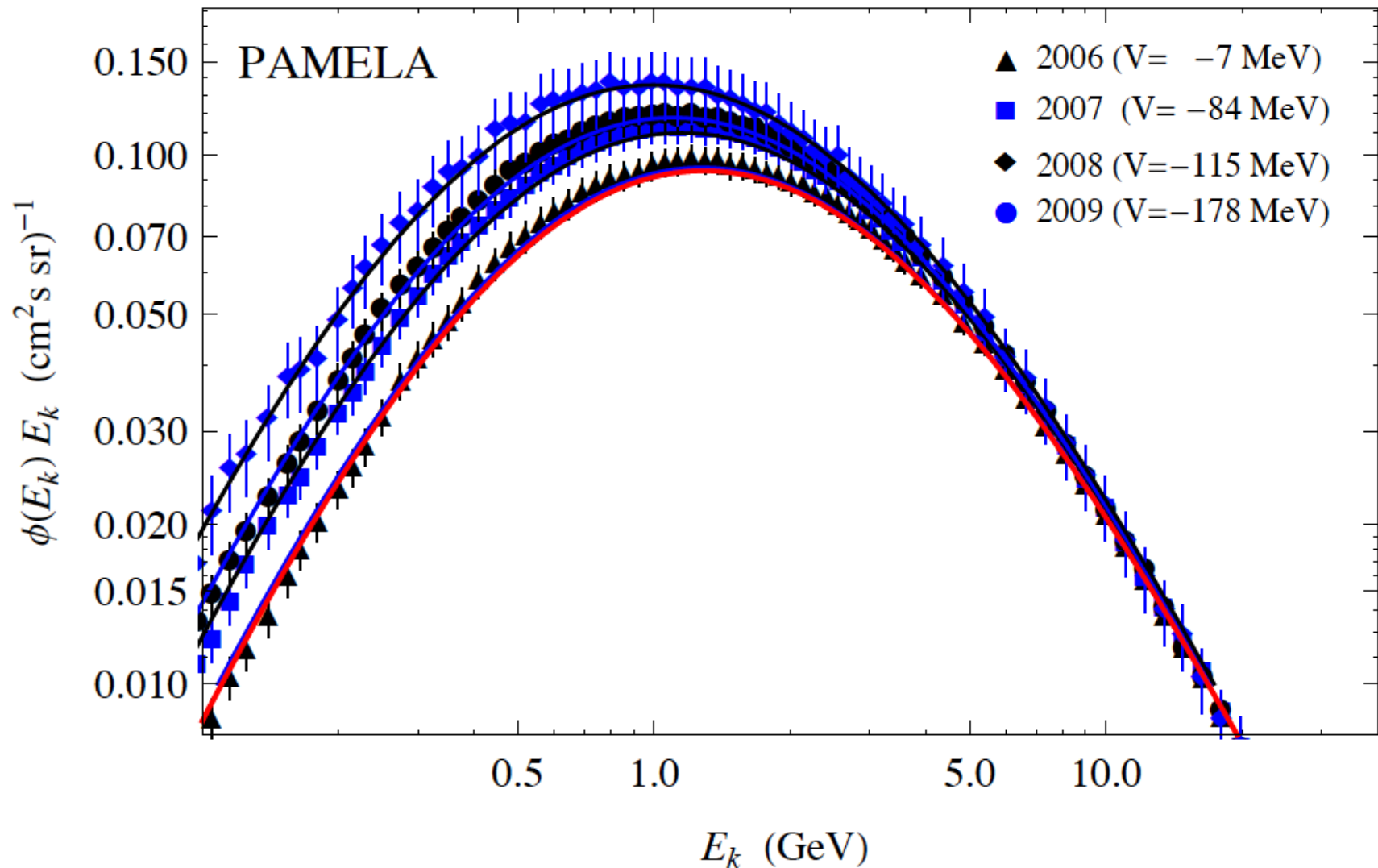
Algorithm is not perfect but
phenomenologically quite successful.

Precision Measurements of CR (proton spectra)
at different times. BESS instrument.
Fit 1998 + use Force-Field algorithm.

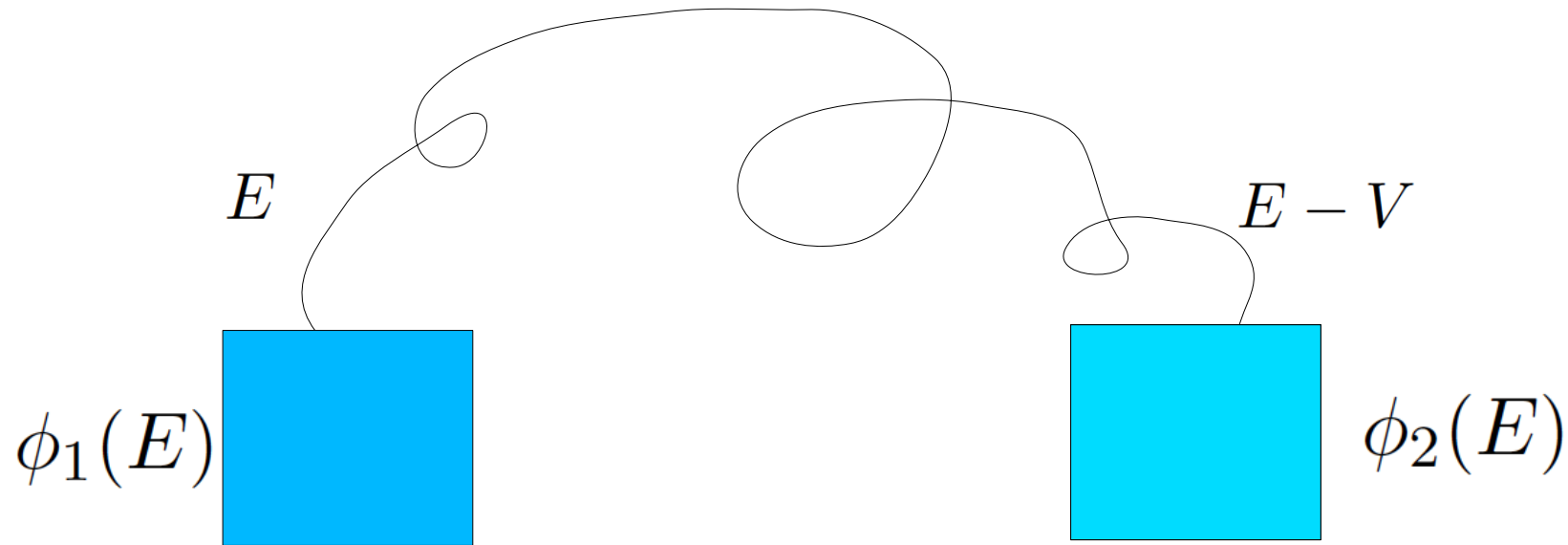


Pamela data: (proton spectrum)

[Compare with BESS-1998 fit (red line)
+ ForceField algorithm]



Physical Meaning of the success of the Force Field Algorithm:



$$\phi_1(E) = \phi_2(E + V) \frac{E^2 - m^2}{(E + V)^2 - m^2}$$

Relation between Flux and phase space density

$$\phi(E, \hat{p}, \vec{x}, t) = p^2 \rho(\vec{x}, \vec{p}, t)$$

$$\rho(\vec{x}, \vec{p}, t) = \frac{d^6 N}{d^3 x d^3 p}$$

Liouville Theorem

$$\rho[\vec{x}(t_0), \vec{p}(t_0), t_0] = \rho[\vec{x}(t), \vec{p}(t), t]$$

Physical Meaning of the phenomenological success of the Force Field Algorithm:

ALL particles of the same type

*(independently from
their energy and direction of arrival)*

$$V(t) = \Delta E(t)$$

lose approximately the same amount of energy
penetrating the heliosphere.

Physical Meaning of the phenomenological success of the Force Field Algorithm:

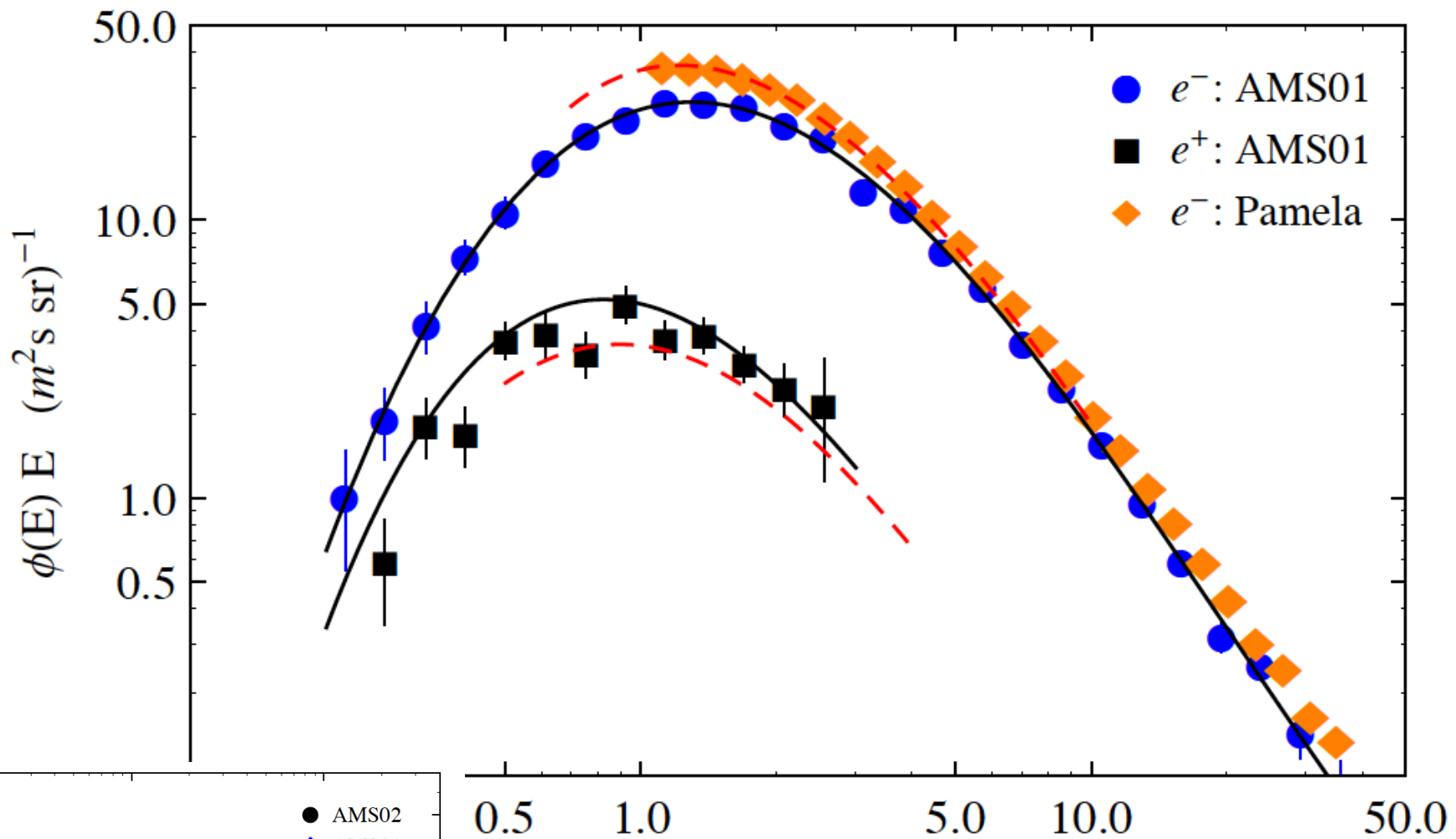
ALL particles of the same type

(independently from their energy and direction of arrival)

$$V(t) = \Delta E(t)$$

lose approximately the same amount of energy penetrating the heliosphere.

Particles of opposite electric charge lose a different amount of energy

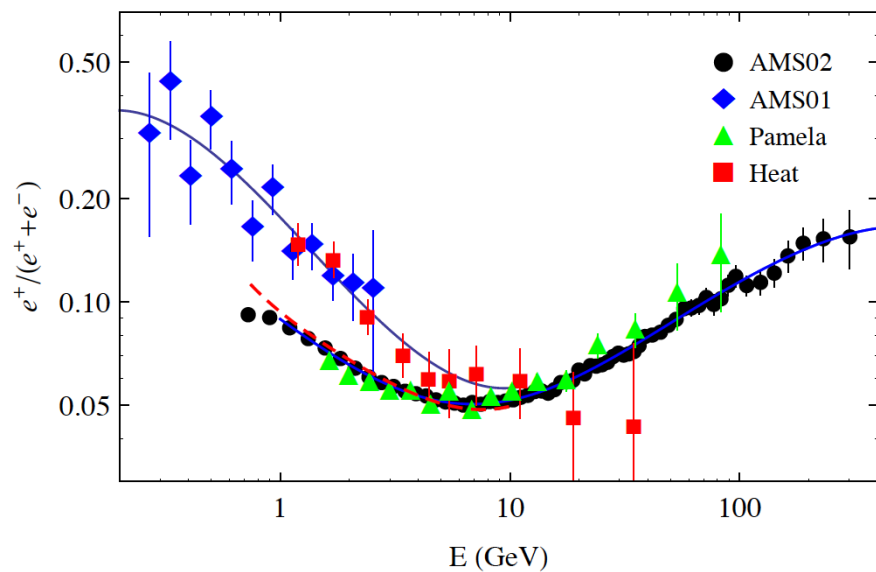


Fit AMS01 data, apply FF alg.

$$\Delta V_{e^-} = -127 \text{ MeV}$$

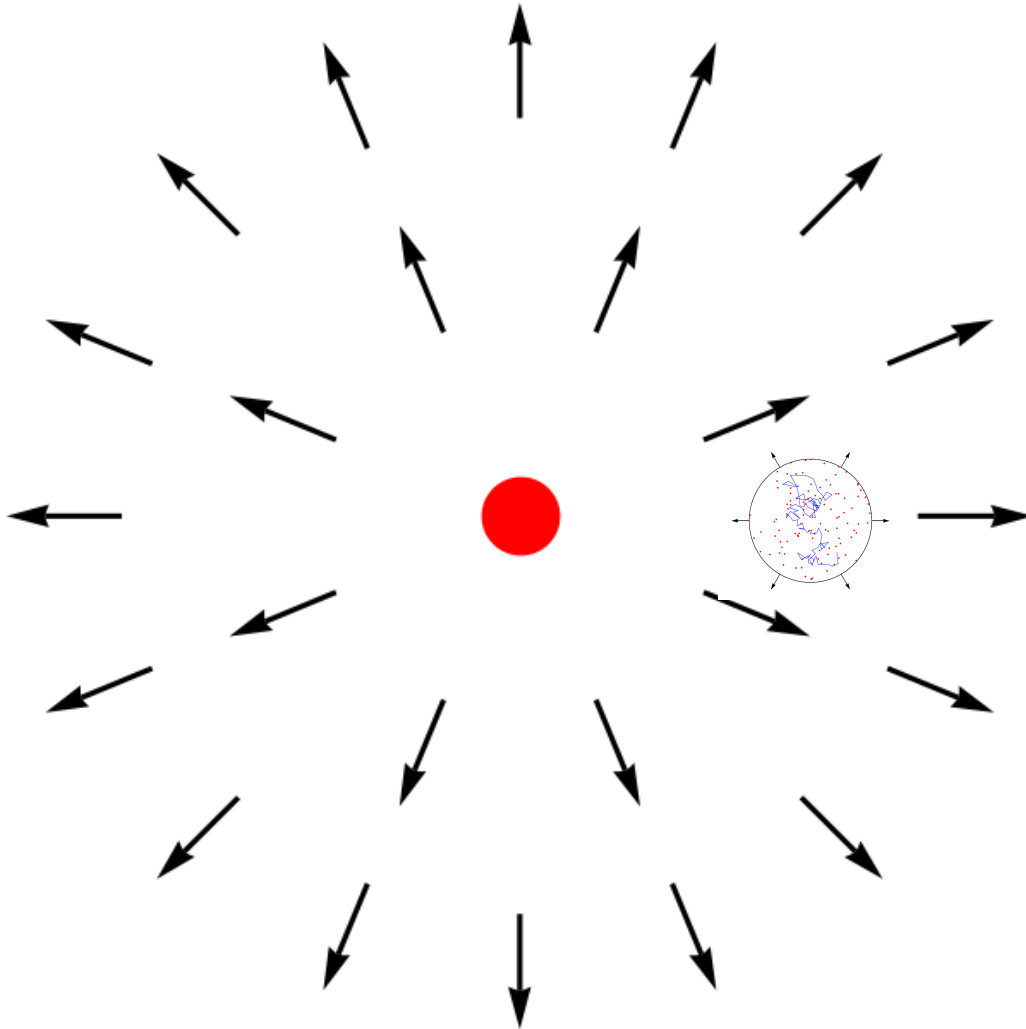
$$\Delta V_{e^+} = +127 \text{ MeV}$$

$$V_{e^-} \leftrightarrow V_{e^+}$$



Parker's (1965) original idea for Solar Modulations

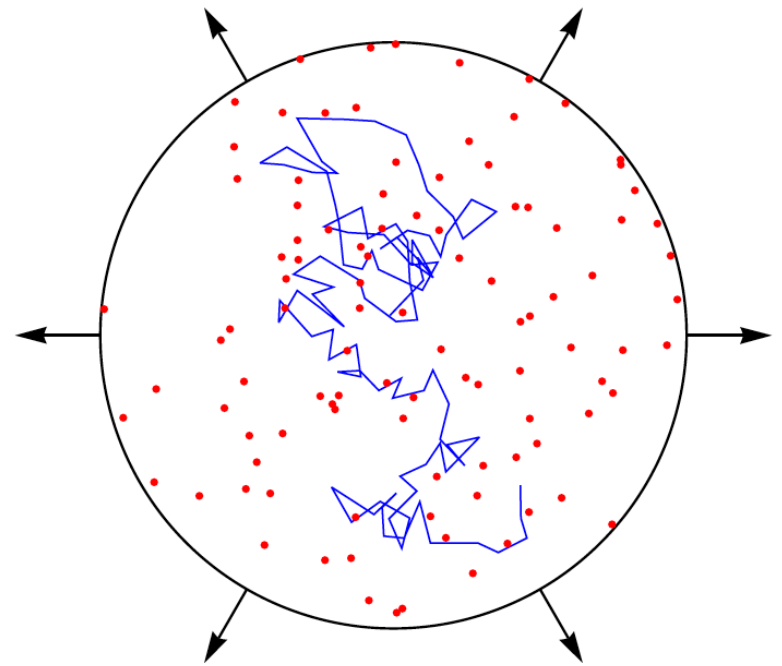
Radial Solar Wind (400 Km/s)



“adiabatic energy losses”
In expanding medium

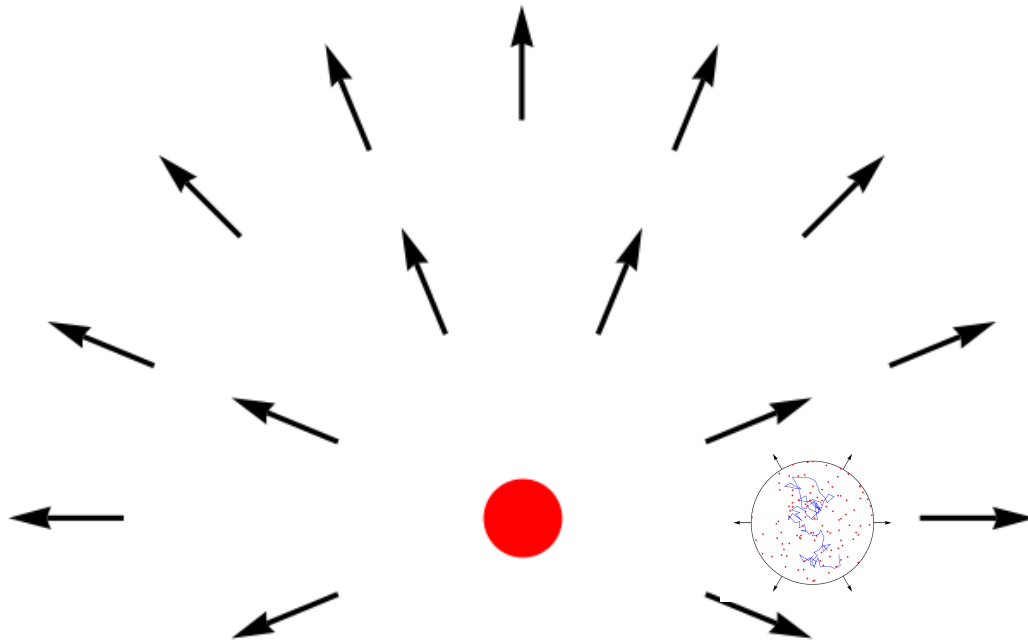
$$\left\langle \frac{dp}{dt} \right\rangle = - [\nabla \cdot \vec{v}_{\text{wind}}] \frac{p}{3}$$

$$\nabla \cdot \hat{r} = \frac{2}{r}$$



Parker's (1965) original idea for Solar Modulations

Radial Solar Wind (400 Km/s)



“adiabatic energy losses”
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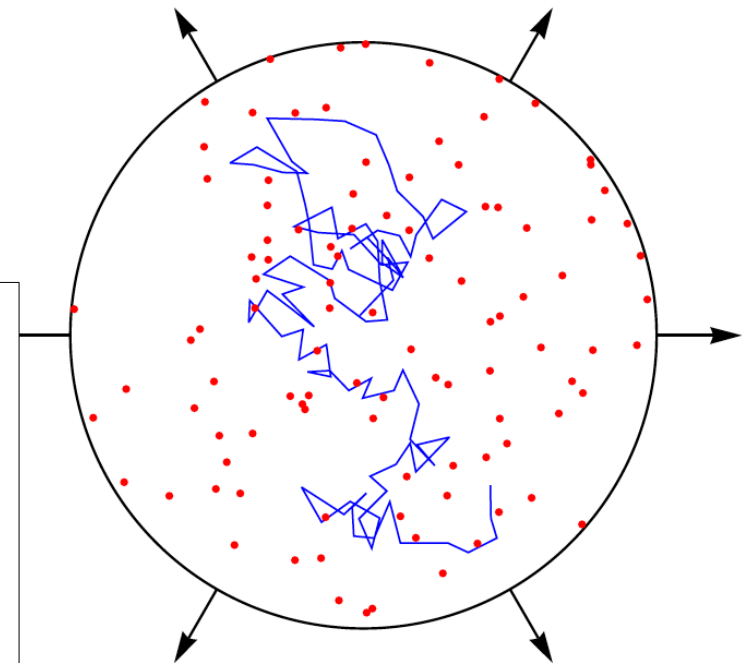
$$\left\langle \frac{dp}{dt} \right\rangle = - [\nabla \cdot \vec{v}_{\text{wind}}] \frac{p}{3}$$

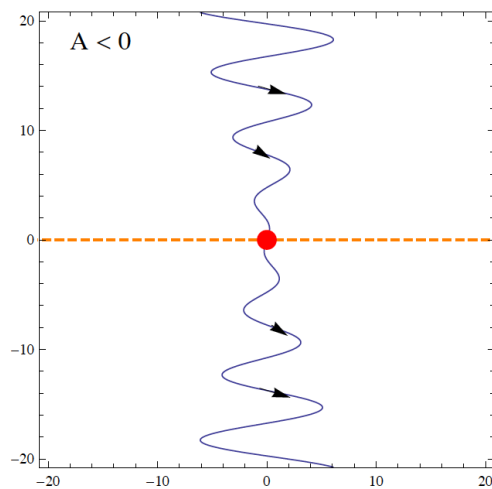
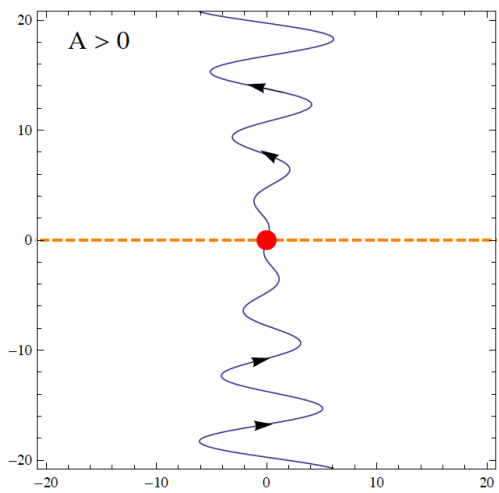
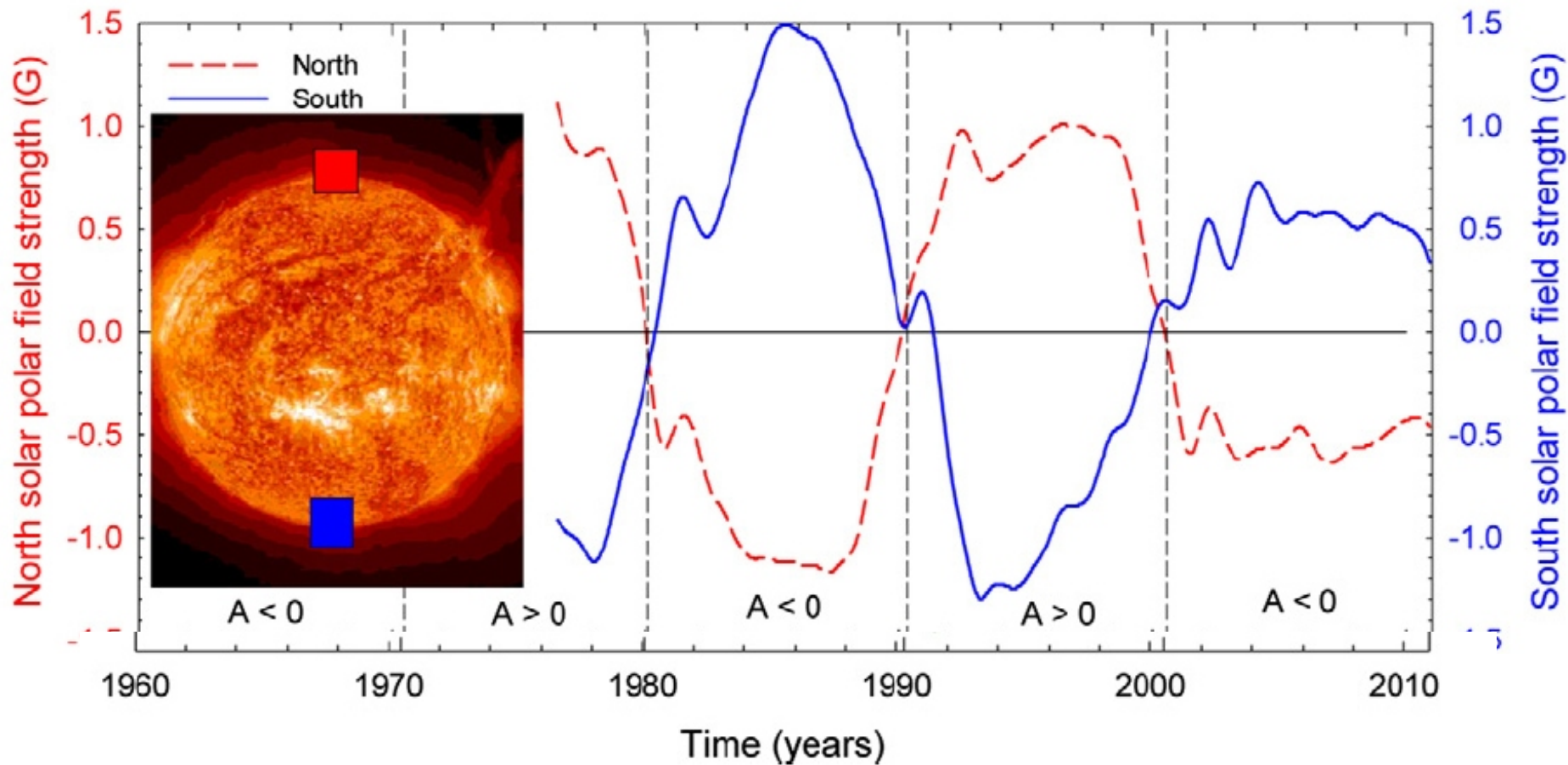
$$\nabla \cdot \hat{r} = \frac{2}{r}$$

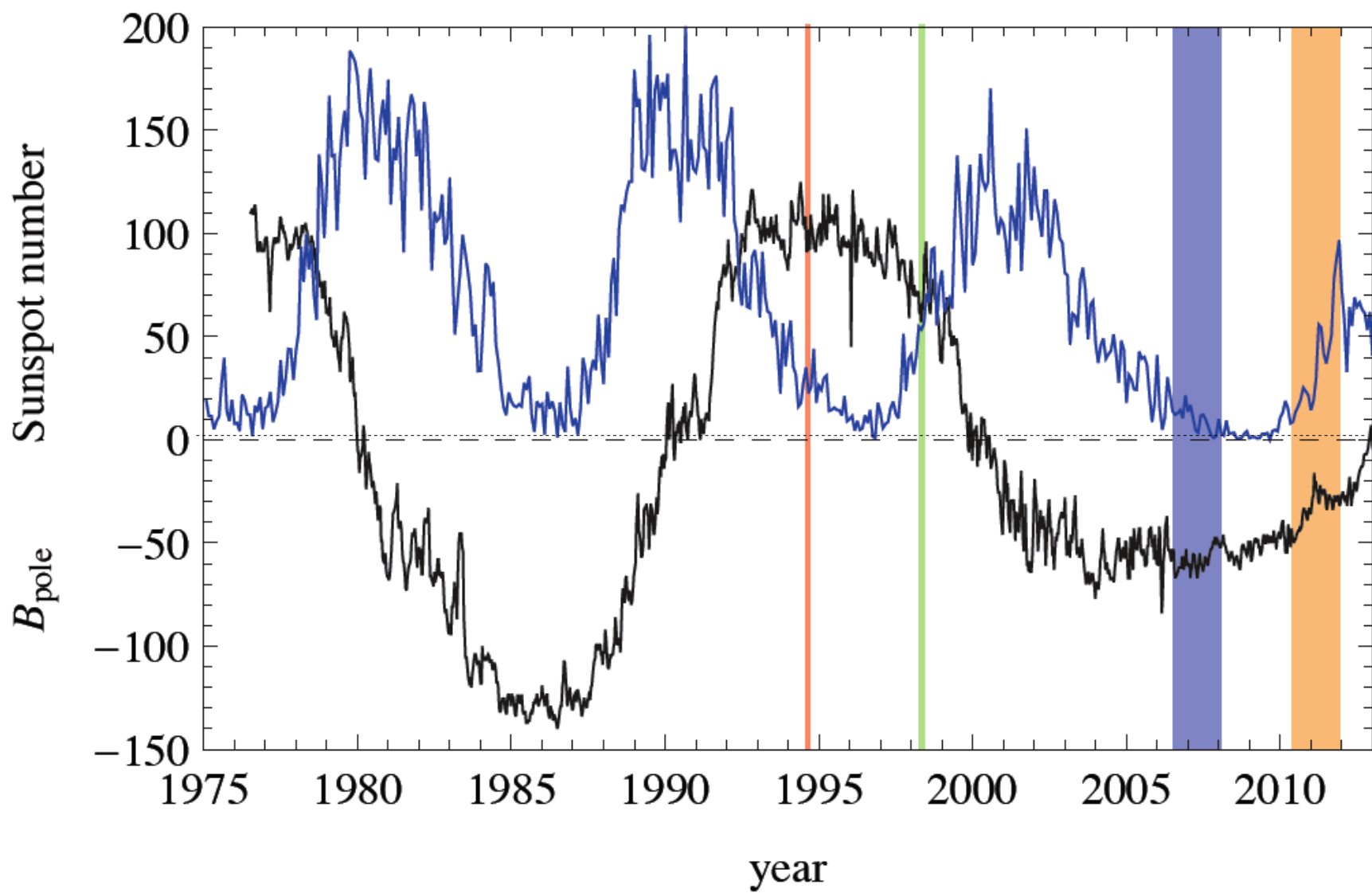
$$\Delta E = \text{const}$$

$$D(E) \propto E$$

$$T(E) \propto E^{-1}$$



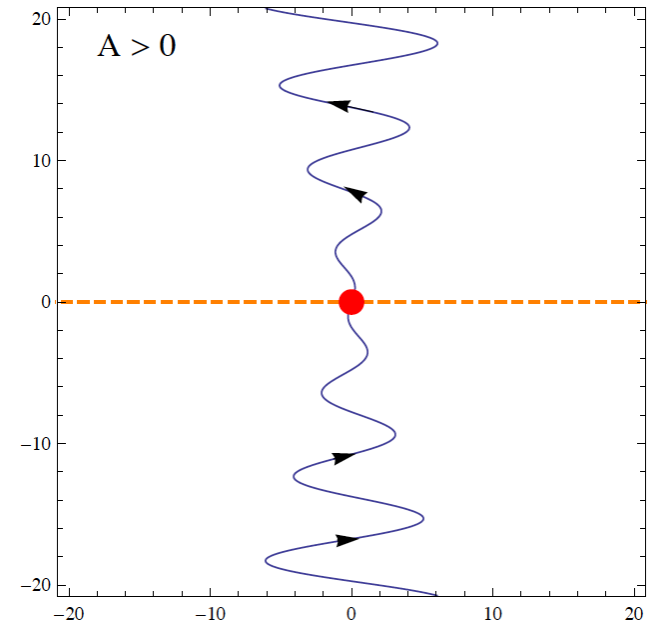
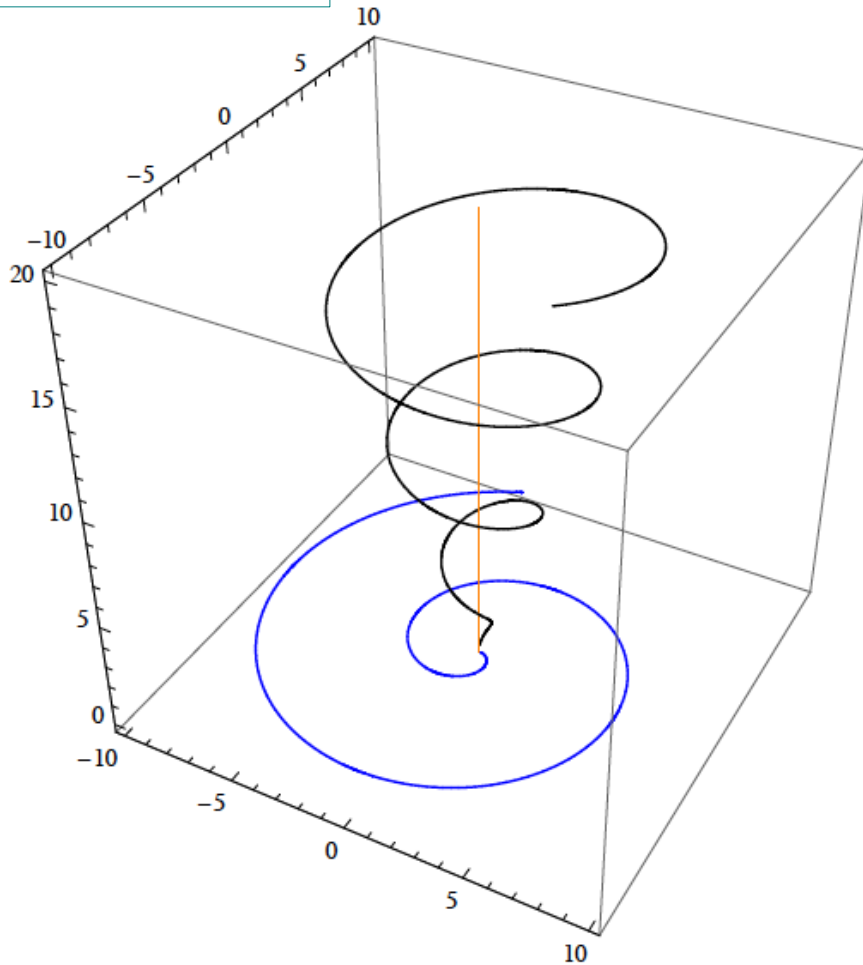




$$\vec{B} = A B_0 \left(\frac{r}{r_0} \right)^2 \left[\hat{r} - \frac{\Omega}{v_w} r \sin \theta \hat{\varphi} \right] S(\vec{r})$$

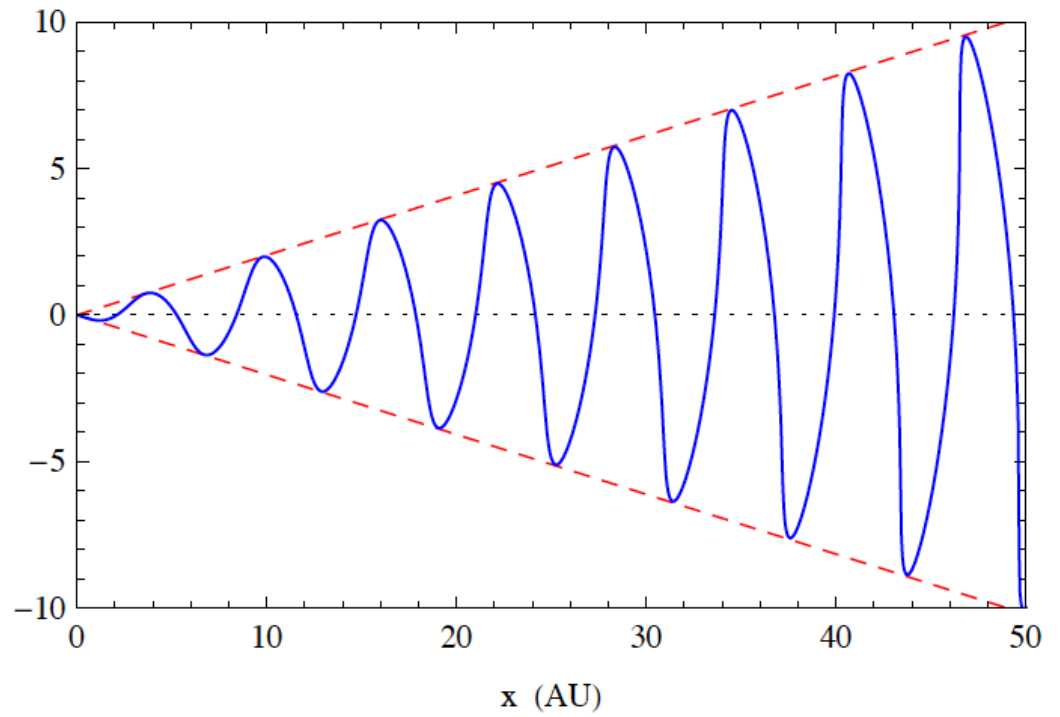
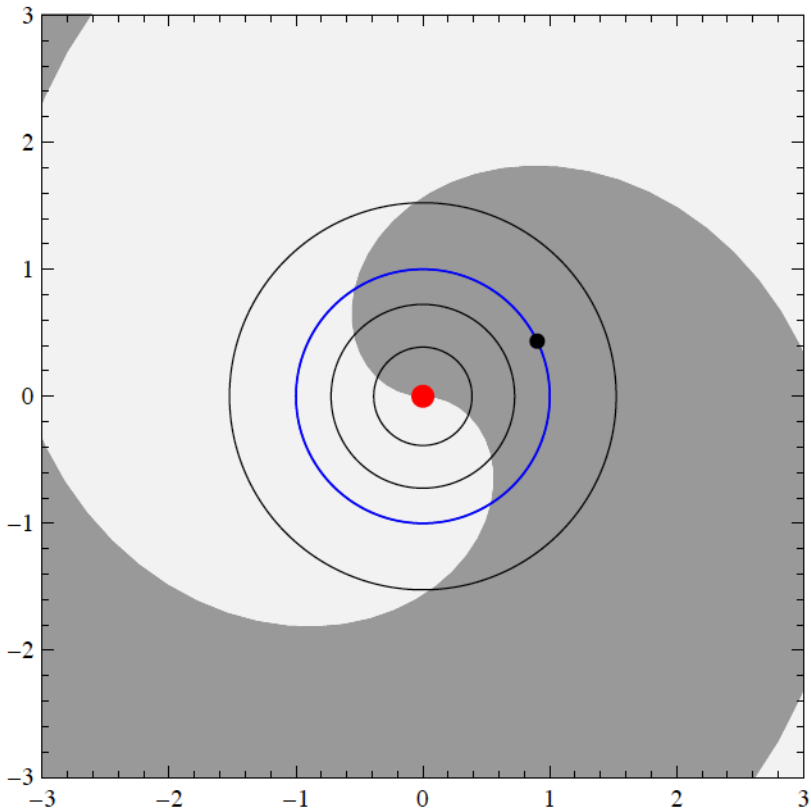
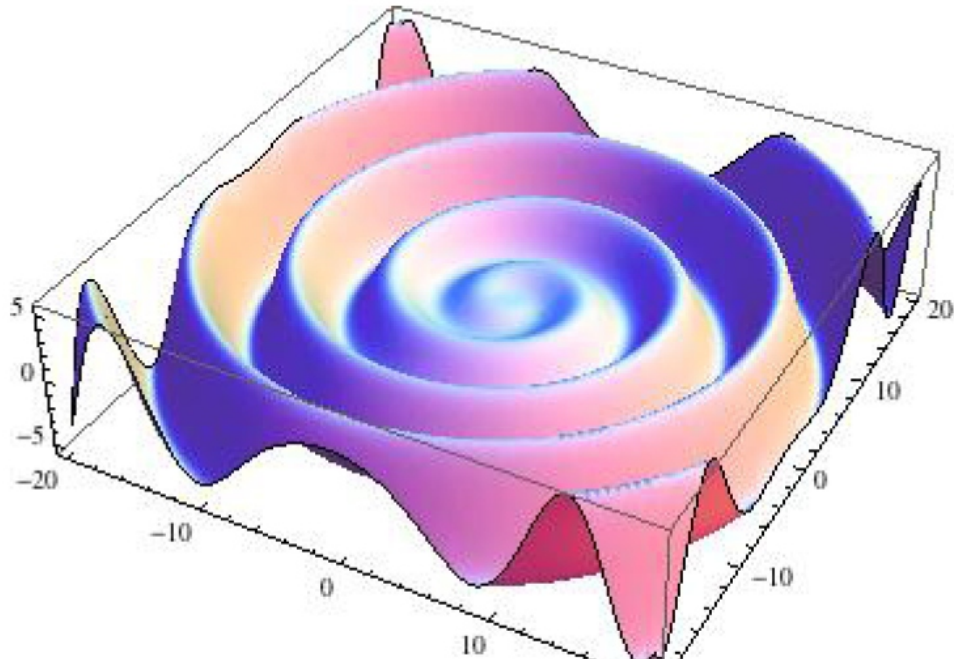
$$A = \pm 1$$

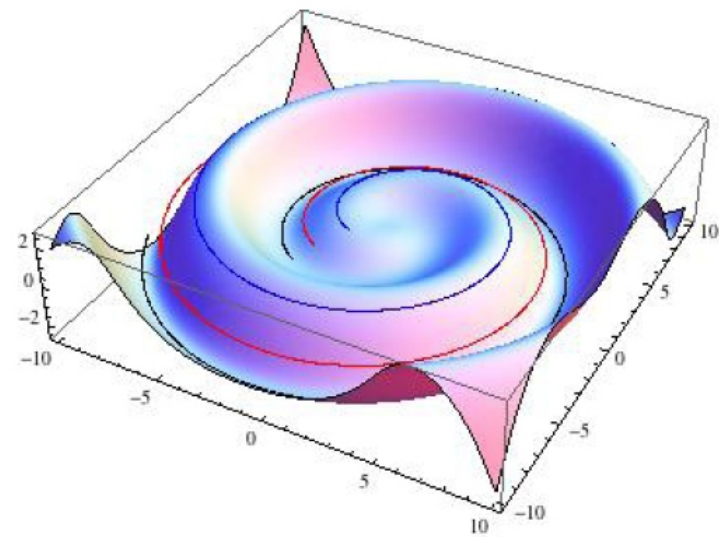
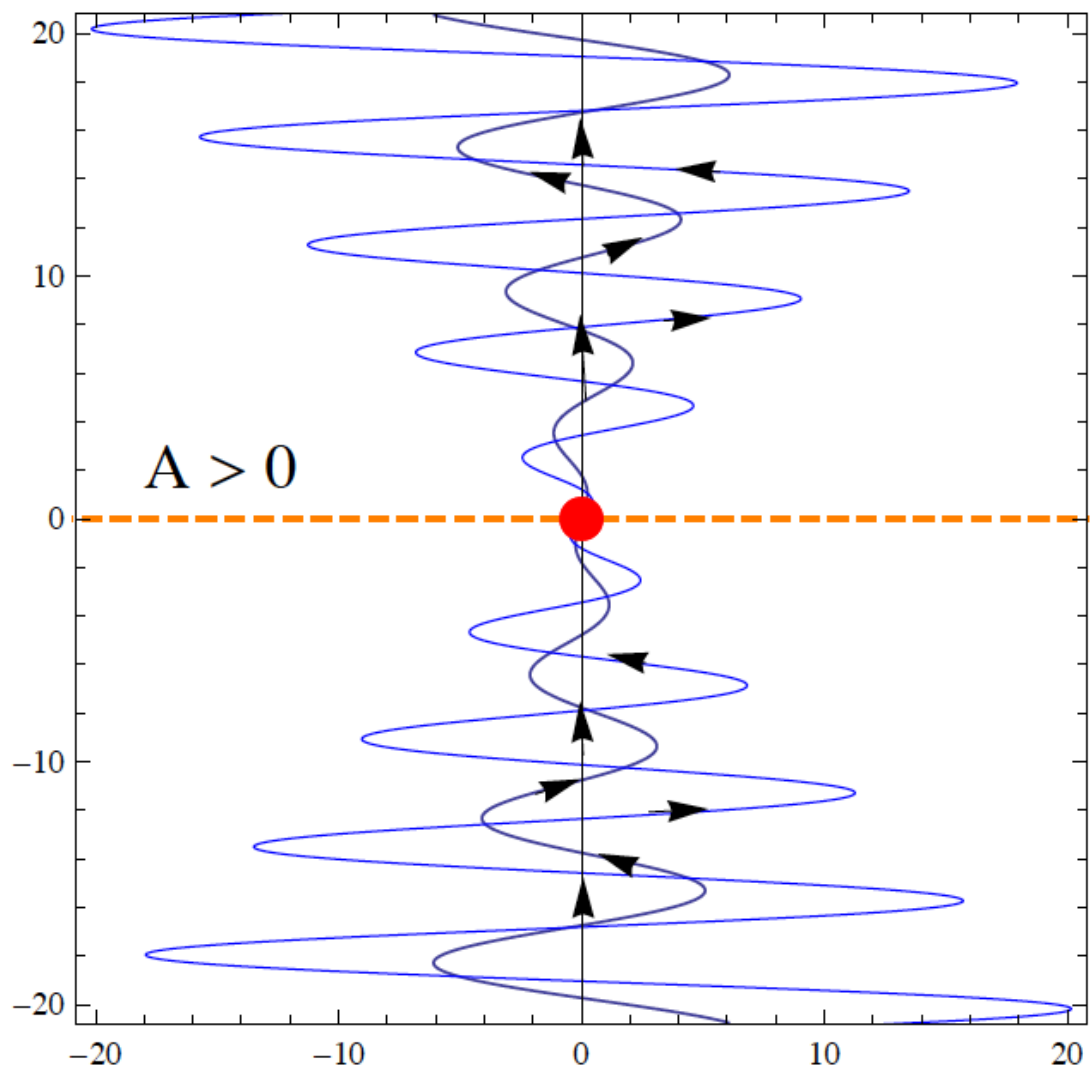
$$S(\vec{r}) = \pm 1$$



Heliospheric Current Sheet

(“Ballerina skirt”)





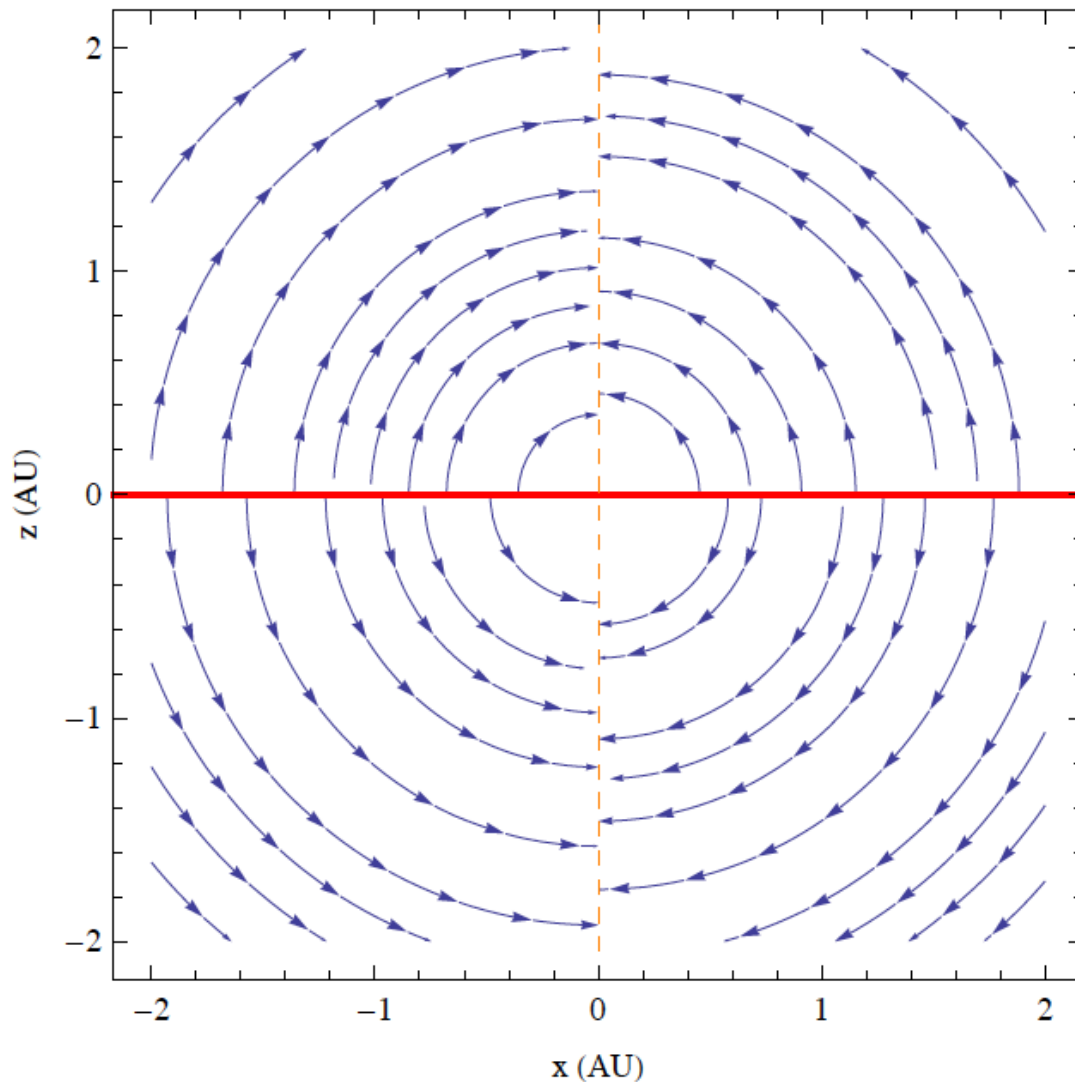
Electric Field Associated to the “regular” Magnetic Field

$$\vec{E}(\vec{x}) = -\frac{\vec{v}_w(\vec{x})}{c} \wedge \vec{B}(\vec{x})$$

Equivalent motivations for expression for electric field:

1. Net force on particle moving with the wind vanishes.
2. Field in wind frame is purely magnetic

$$\vec{E}(x, y, z) = \pm A B_0 \frac{\Omega r_0^2}{c r^3} \{x z, y z, -(x^2 + y^2)\}$$



Field with
Cylindrical symmetry

Stream-lines of the
“regular’ Electric Field

“Back-tracing” of particles observed at the Earth.

$$\frac{d\vec{x}}{dt} = \vec{v} = \frac{\vec{p}}{\sqrt{p^2 + m^2}}$$

$$\frac{d\vec{p}}{dt} = q \left(\vec{E} + \frac{\vec{v}}{c} \wedge \vec{B} \right)$$

$$\vec{E}(\vec{x}) = -\frac{\vec{v}_w(\vec{x})}{c} \wedge \vec{B}(\vec{x})$$

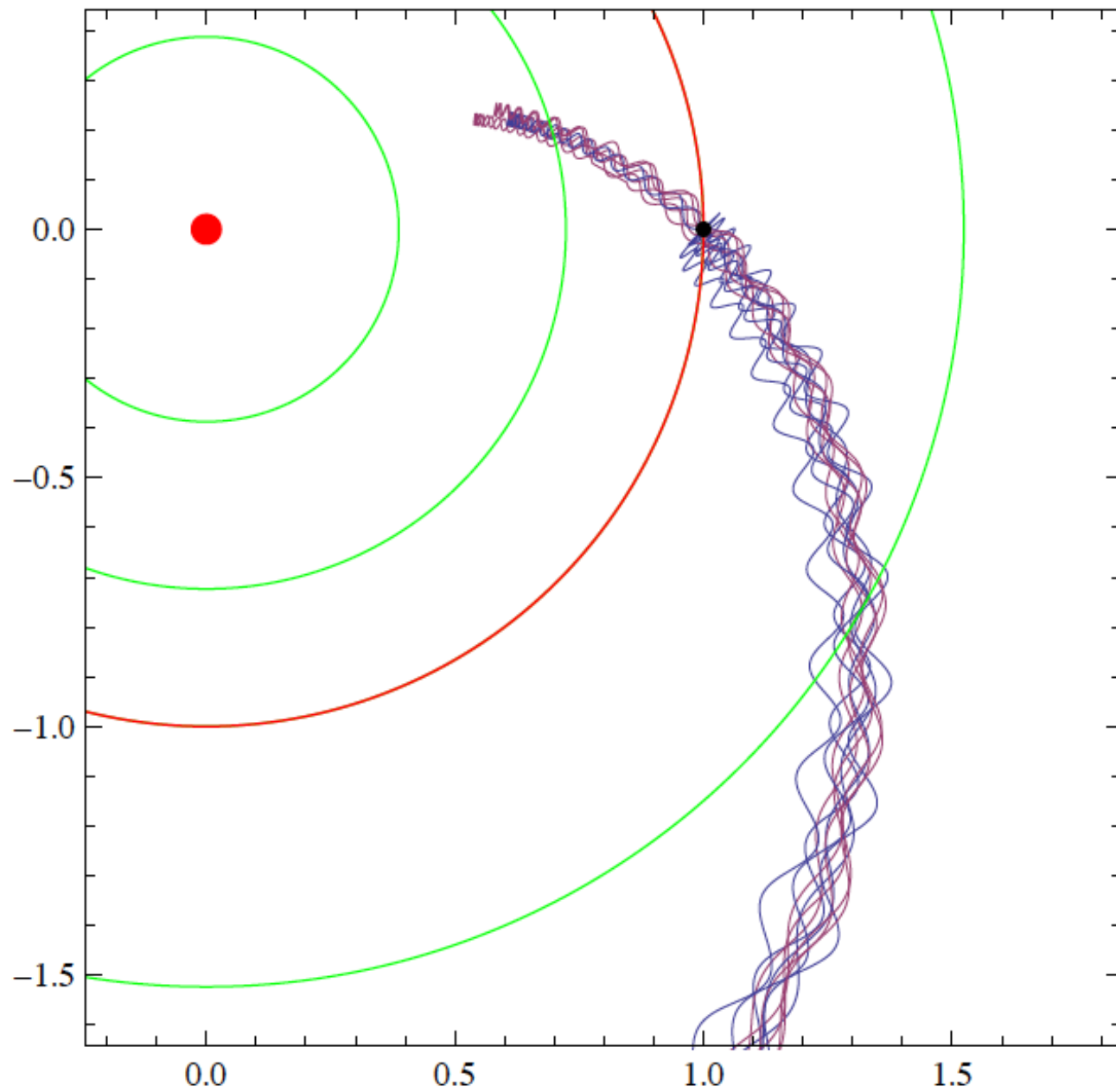
“Back-tracing” of particles observed at the Earth.
[Neglect random, turbulent field]

$$\frac{d\vec{x}}{dt} = \vec{v} = \frac{\vec{p}}{\sqrt{p^2 + m^2}}$$

$$\frac{d\vec{p}}{dt} = q \left(\vec{E} + \frac{\vec{v}}{c} \wedge \vec{B} \right)$$

Small perturbation

$$\vec{E}(\vec{x}) = -\frac{\vec{v}_w(\vec{x})}{c} \wedge \vec{B}(\vec{x})$$



“Backtracing”
of particles

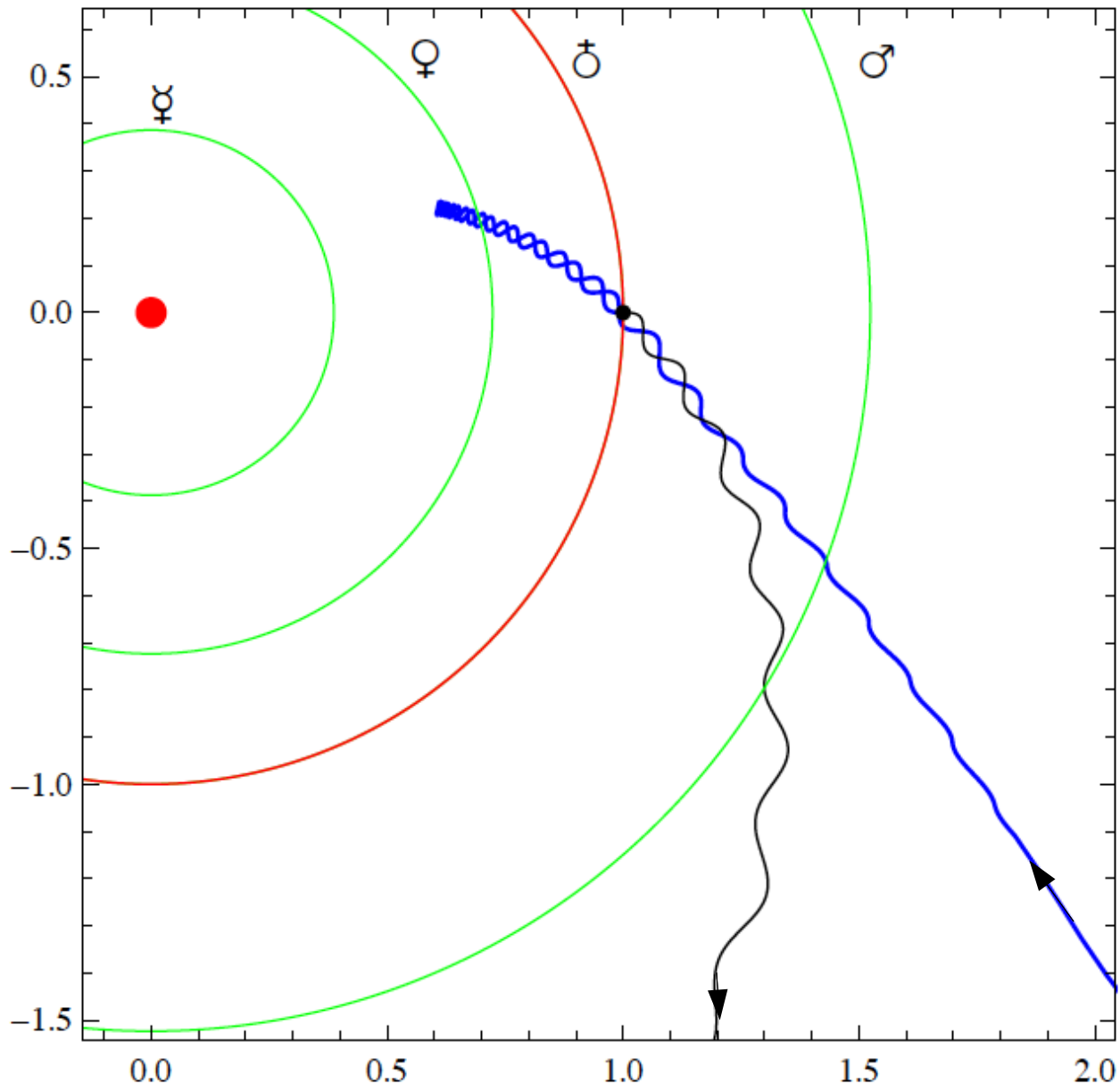
e^+ (12 GeV)

12 directions

$$B_0 = 10^{-4} \text{ G}$$

$$v_w = 400 \text{ km/s}$$

$$\sin \alpha = 0.05$$

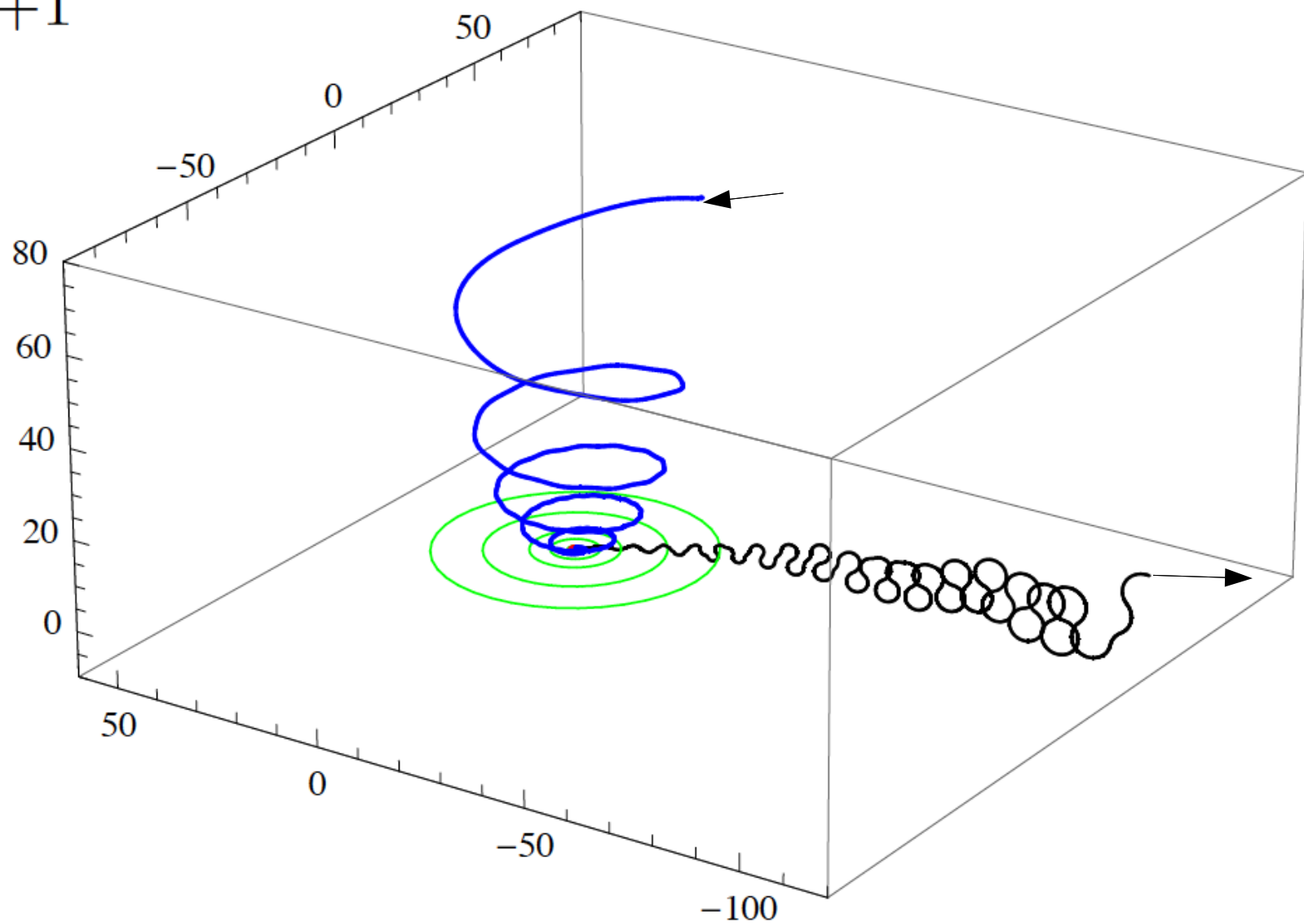


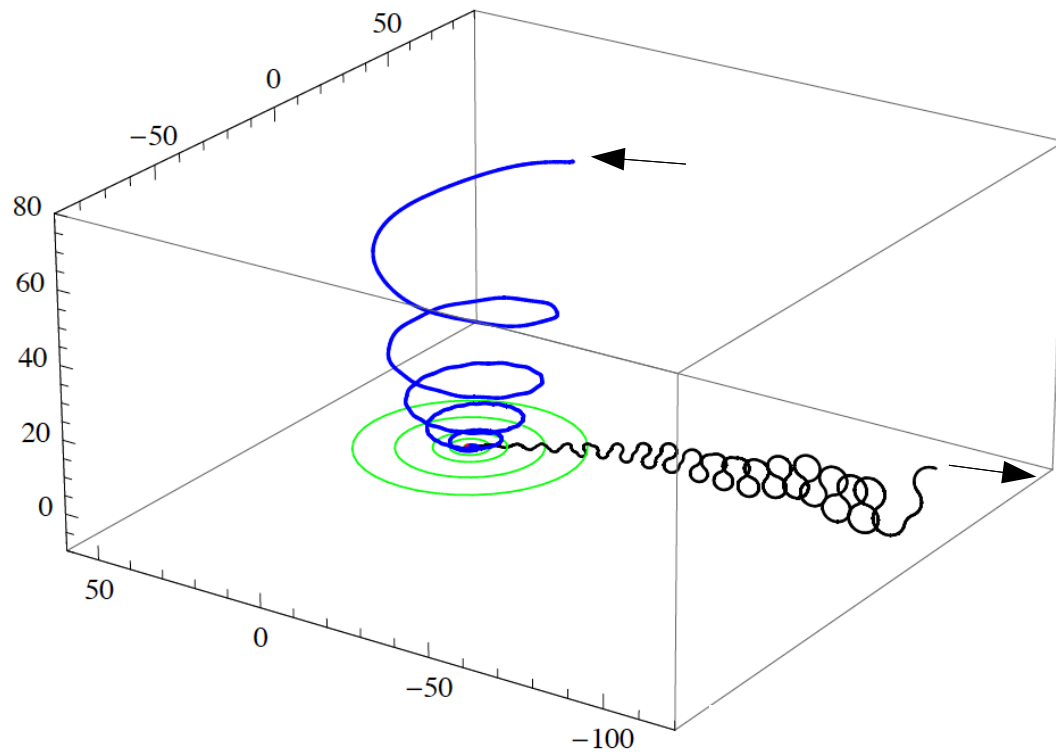
12 GeV
e-

$$A = +1$$

$$E_{\text{obs}} = 12 \text{ GeV}$$

$$q = +1$$



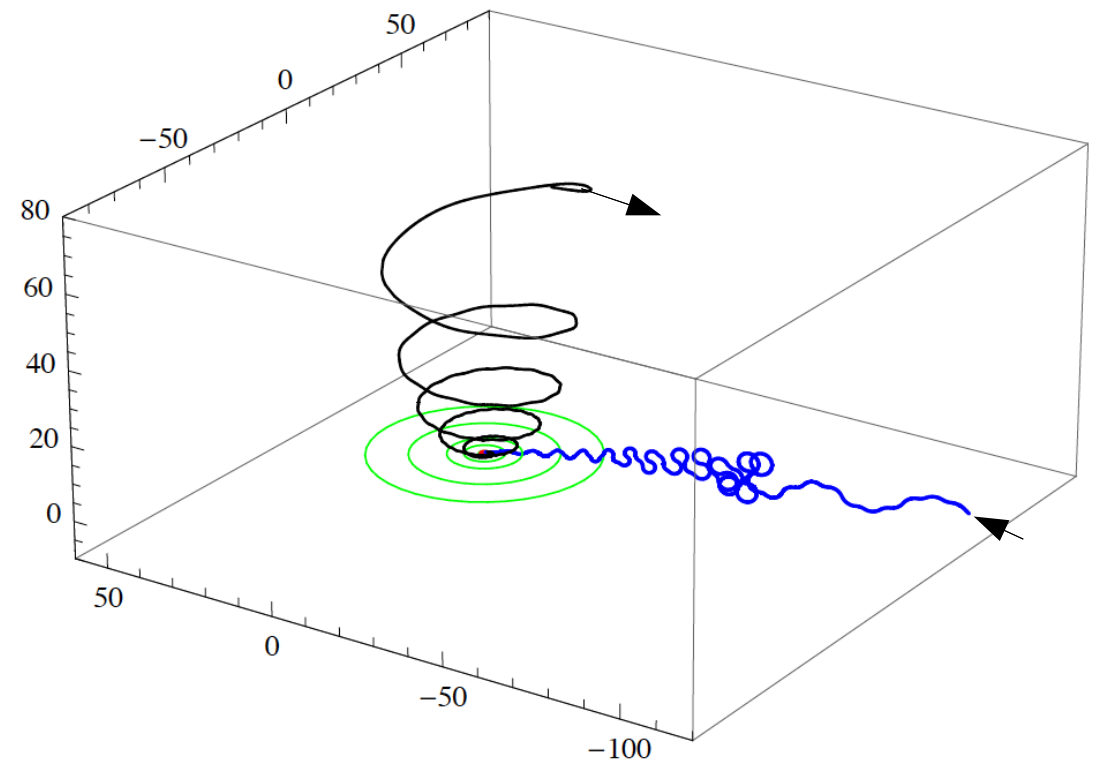


$$A = +1$$

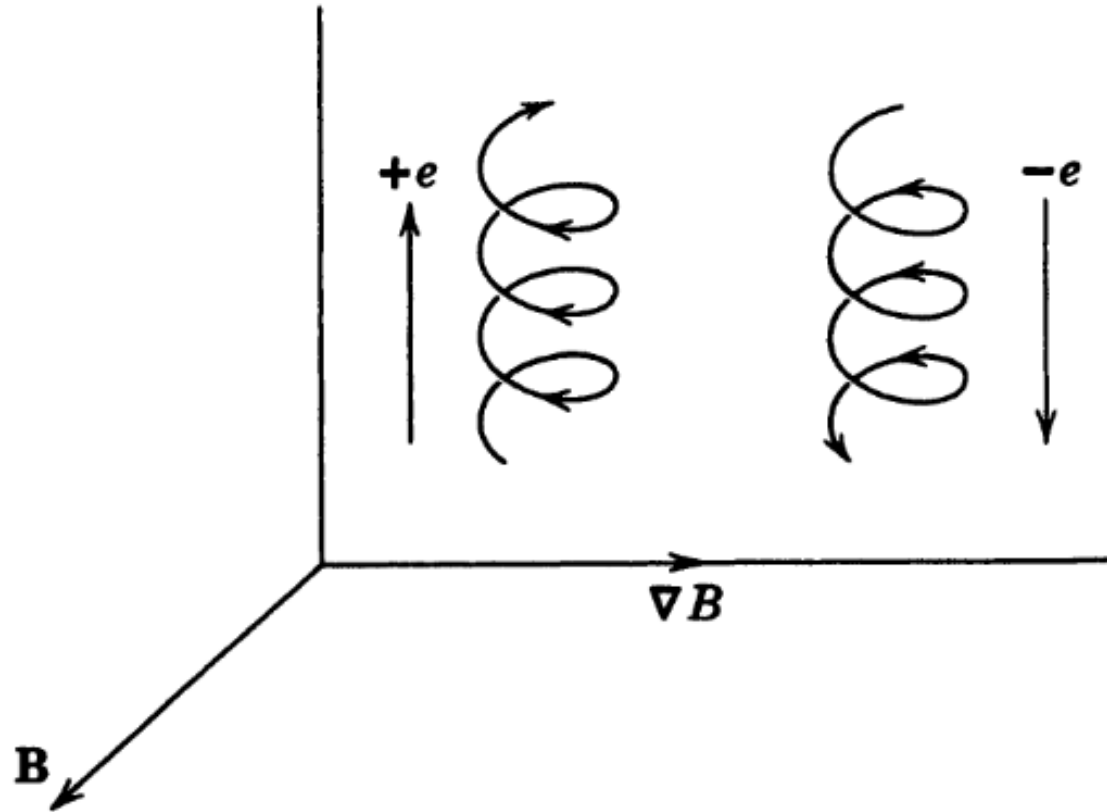
$$q = +1$$

$$A = +1$$

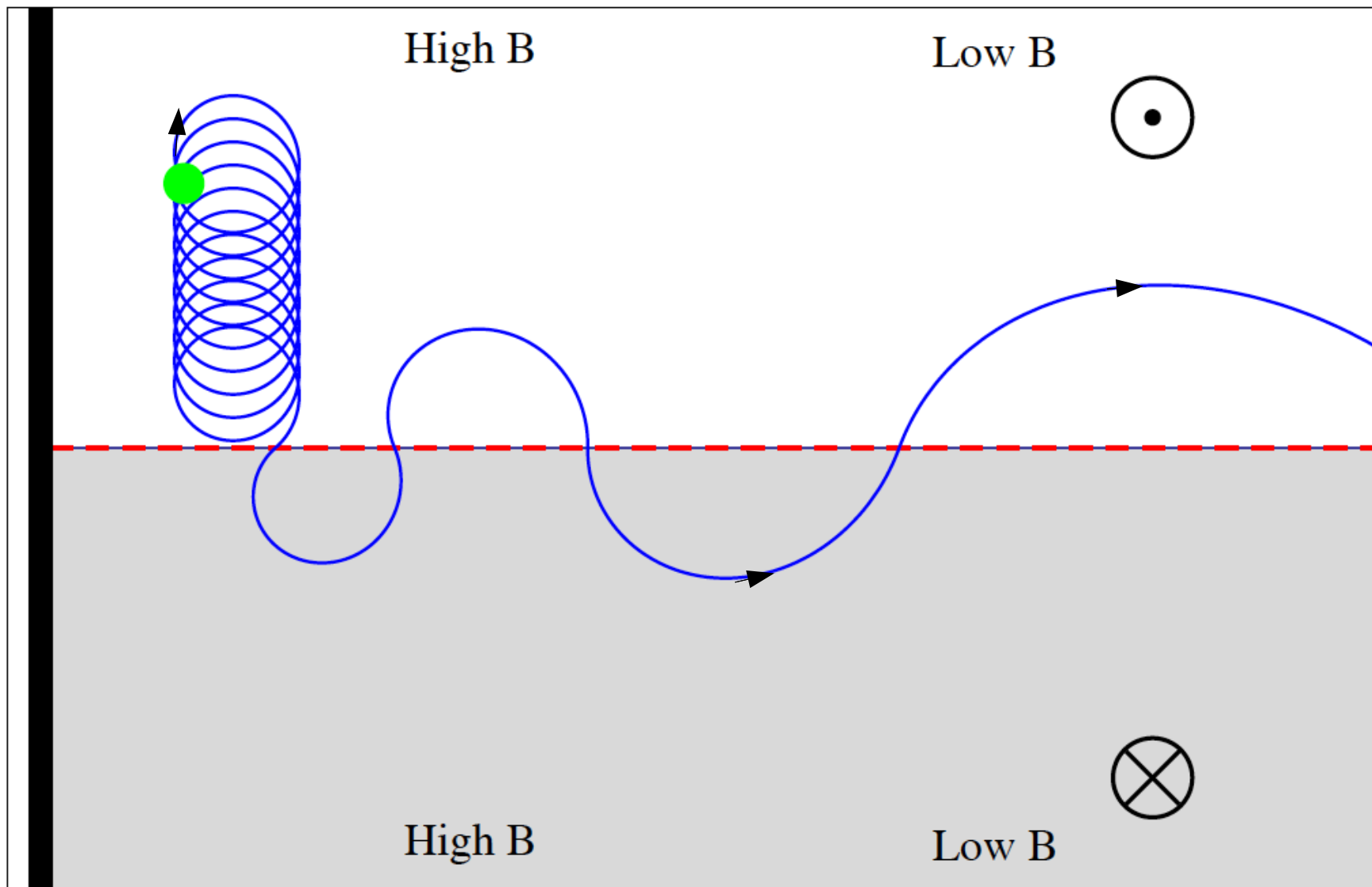
$$q = -1$$



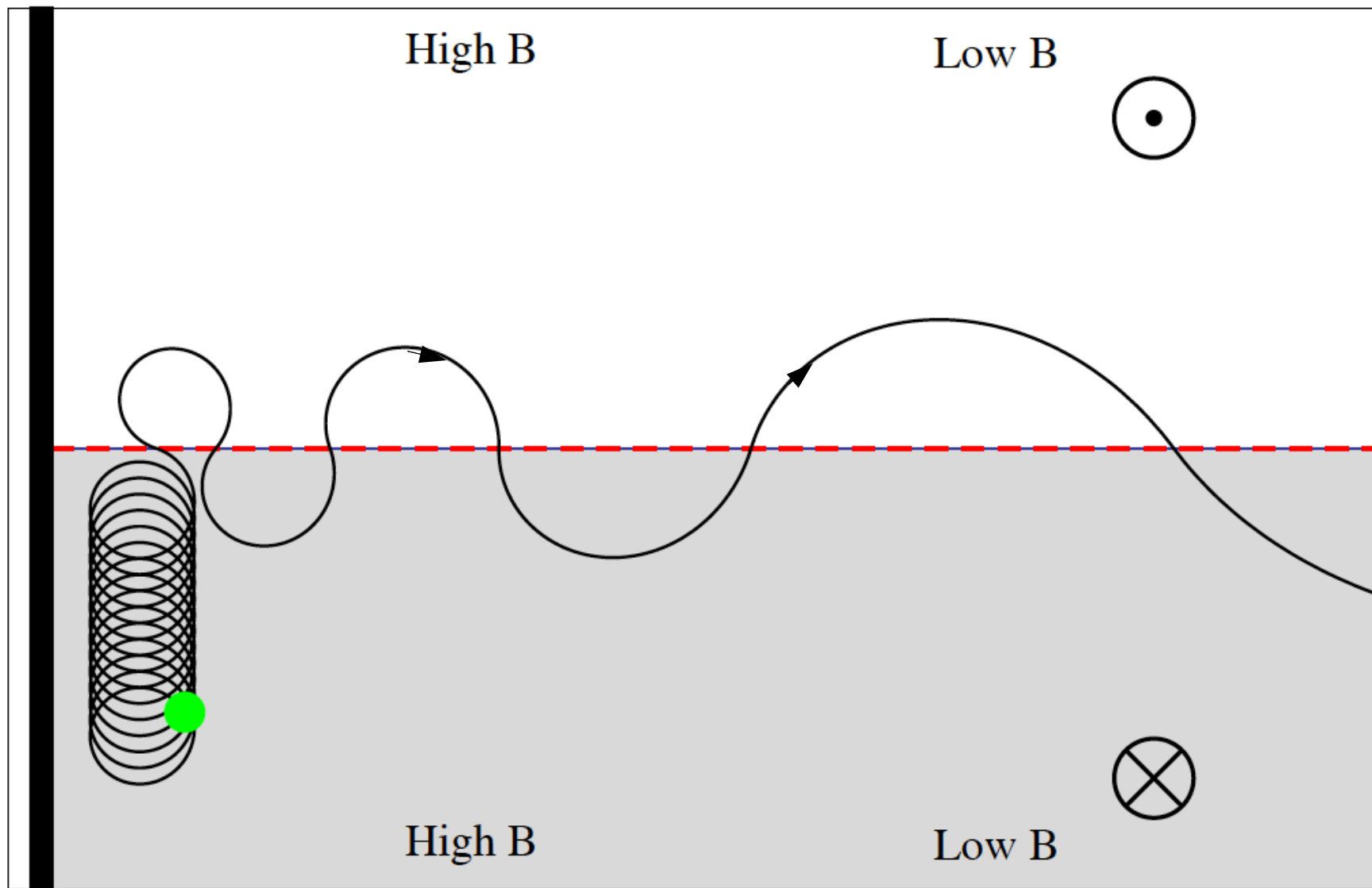
“Drift” of the guiding center of the particle gyration



$$q = +1$$

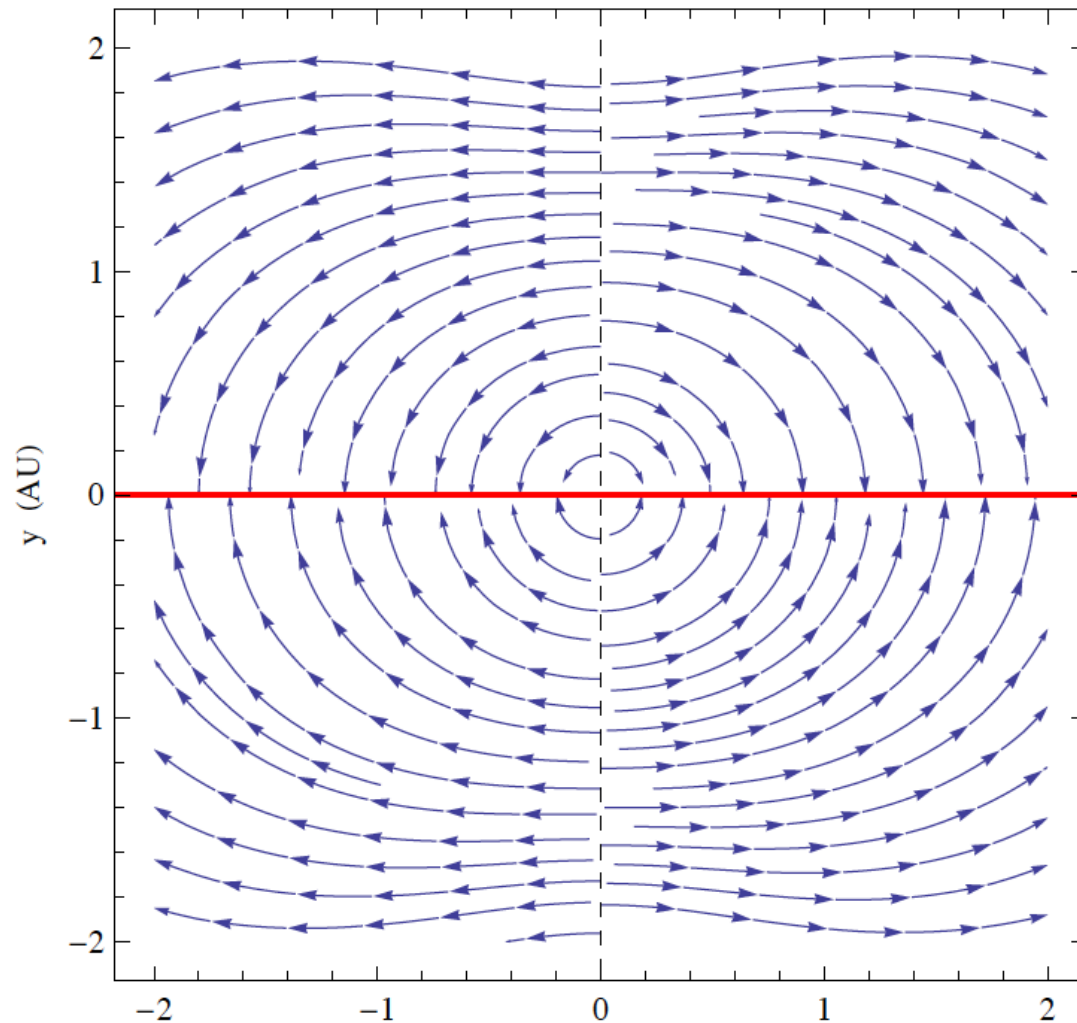


$$q = +1$$

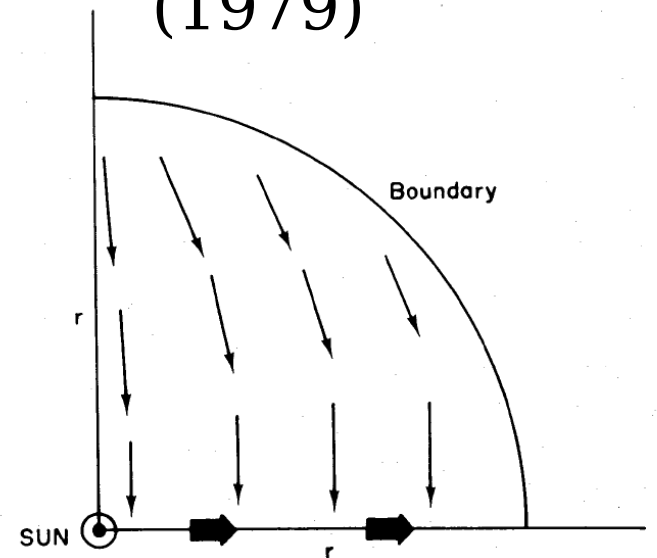


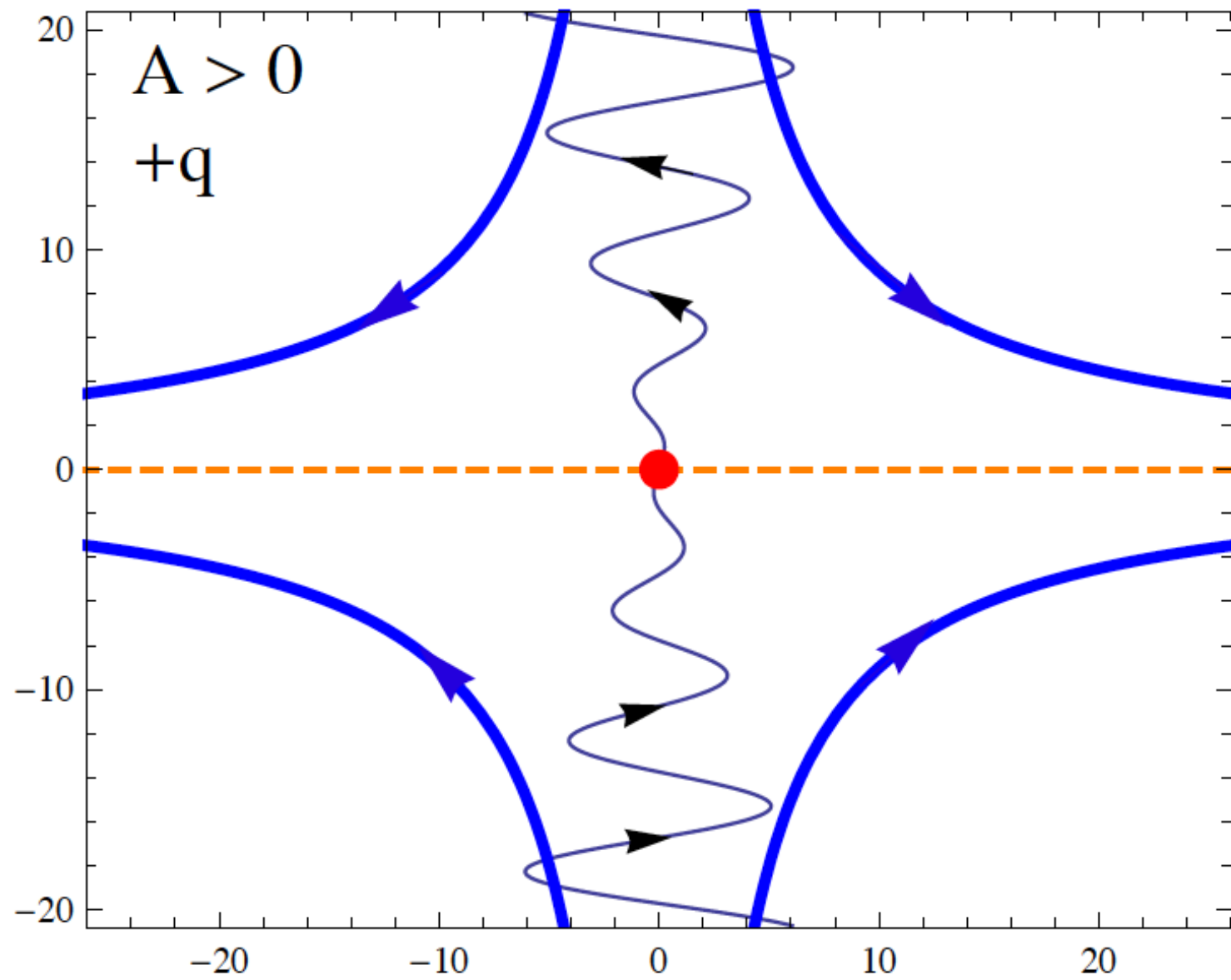
Stream lines of the drift velocity

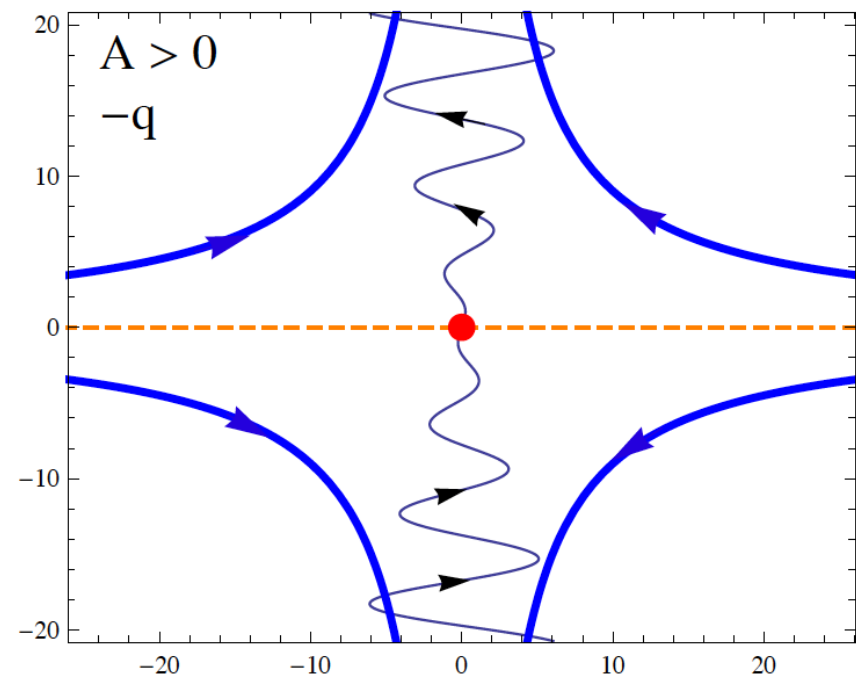
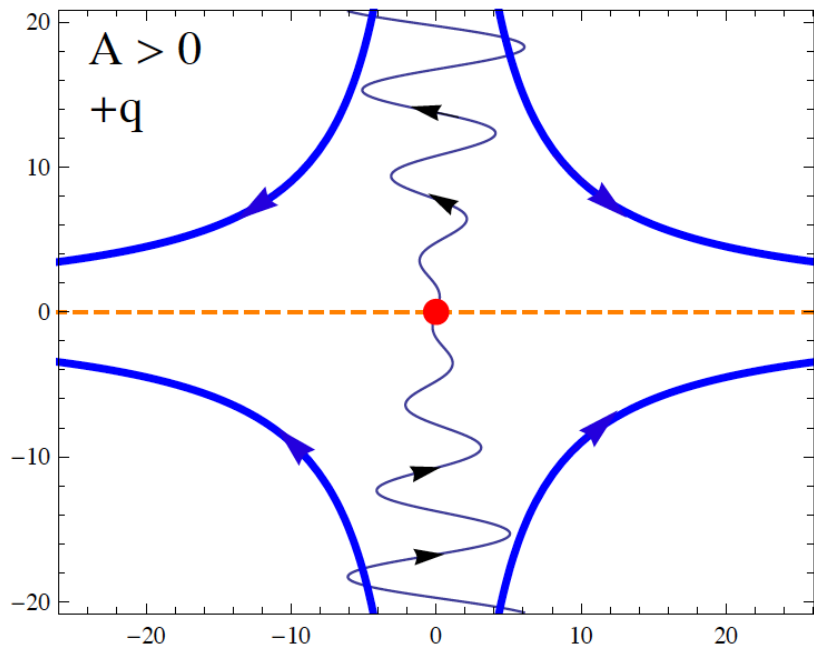
$$q A > 0$$



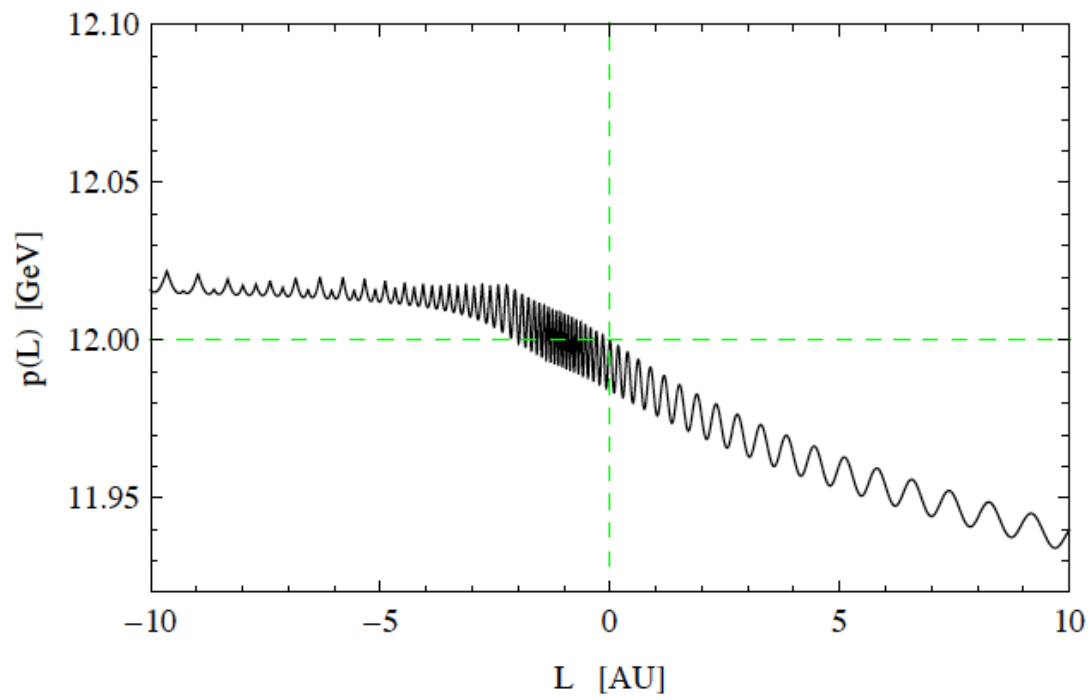
Jokipii + Kopriva
(1979)



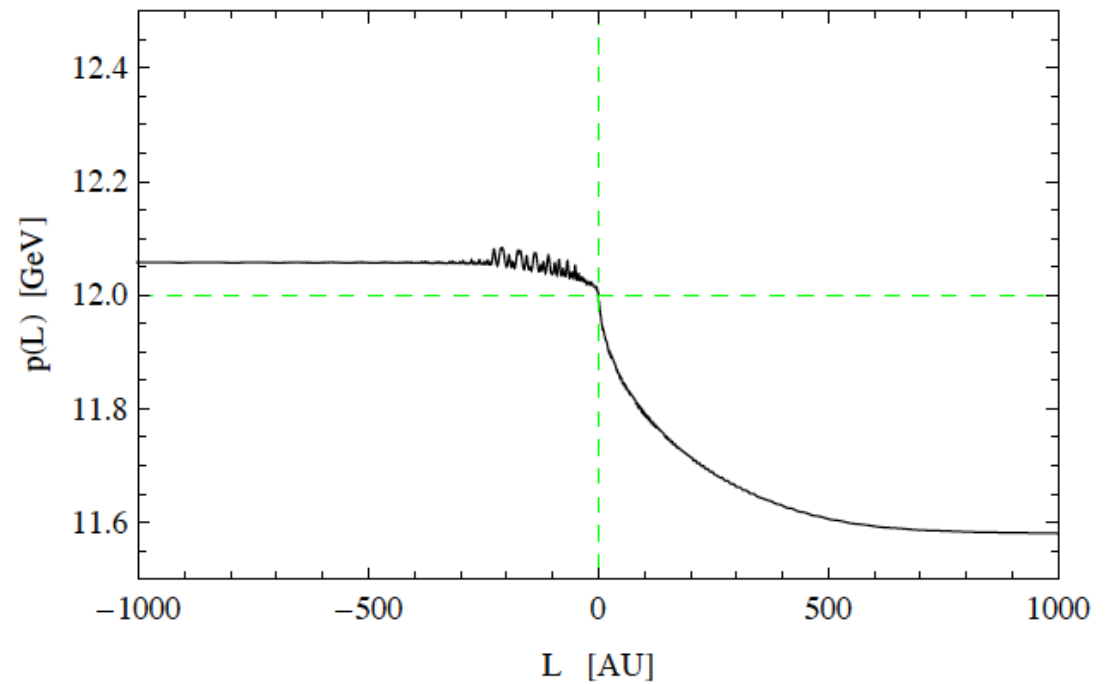




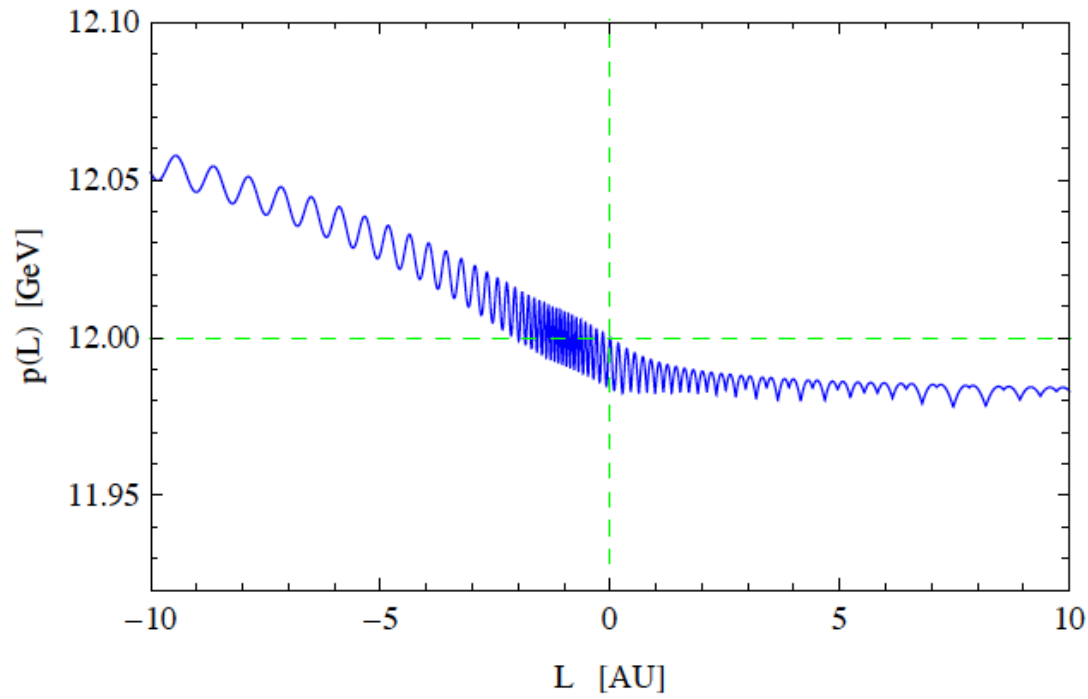
Energy Evolution



$$A = +1$$
$$q = +1$$

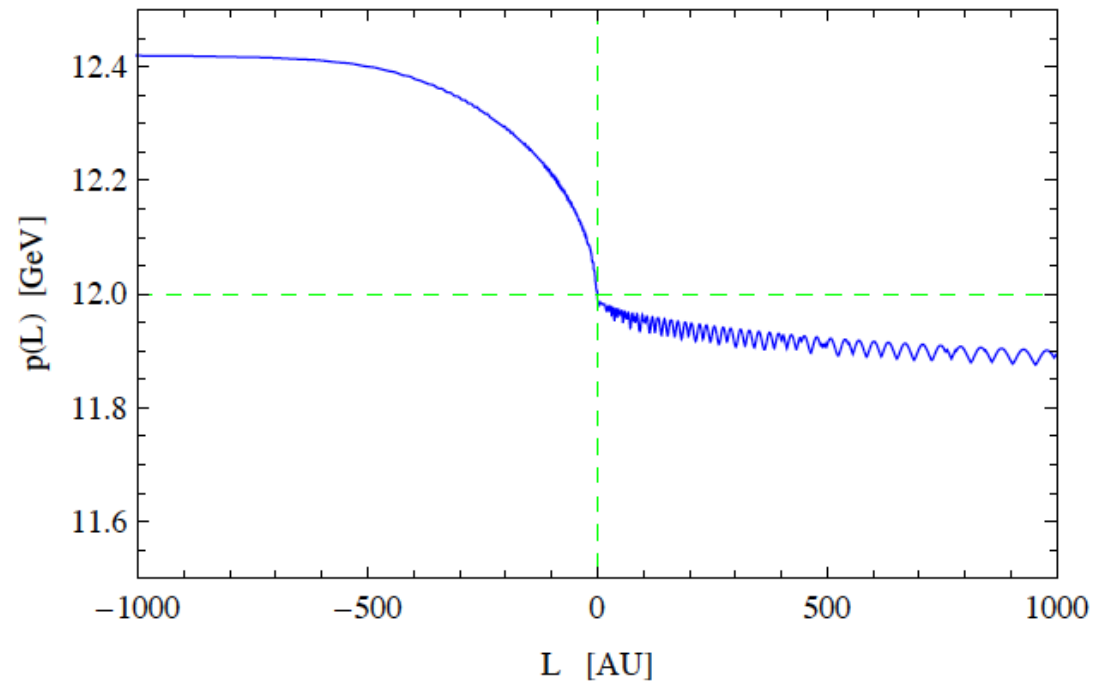


Energy Evolution



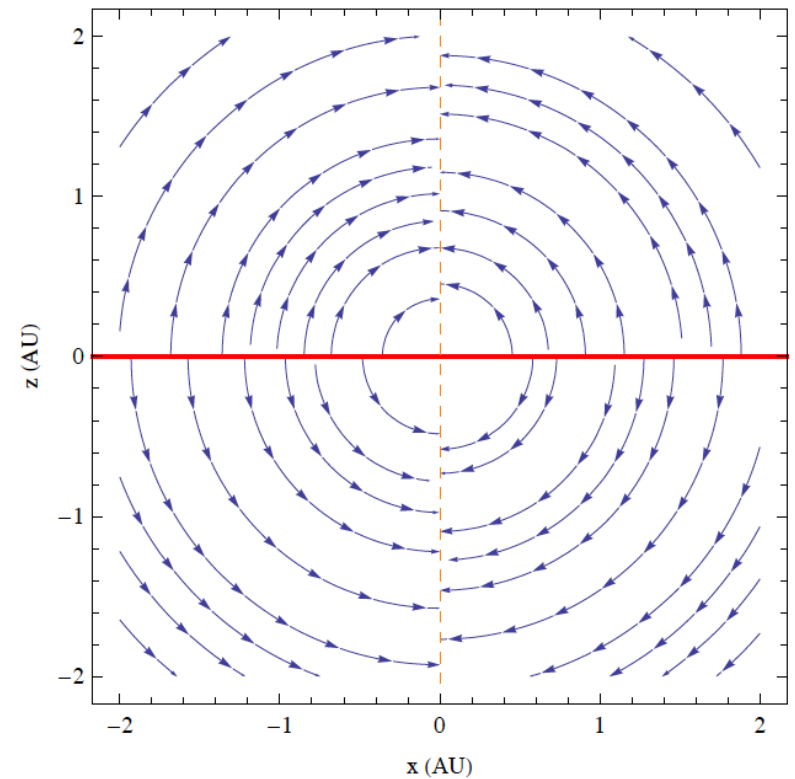
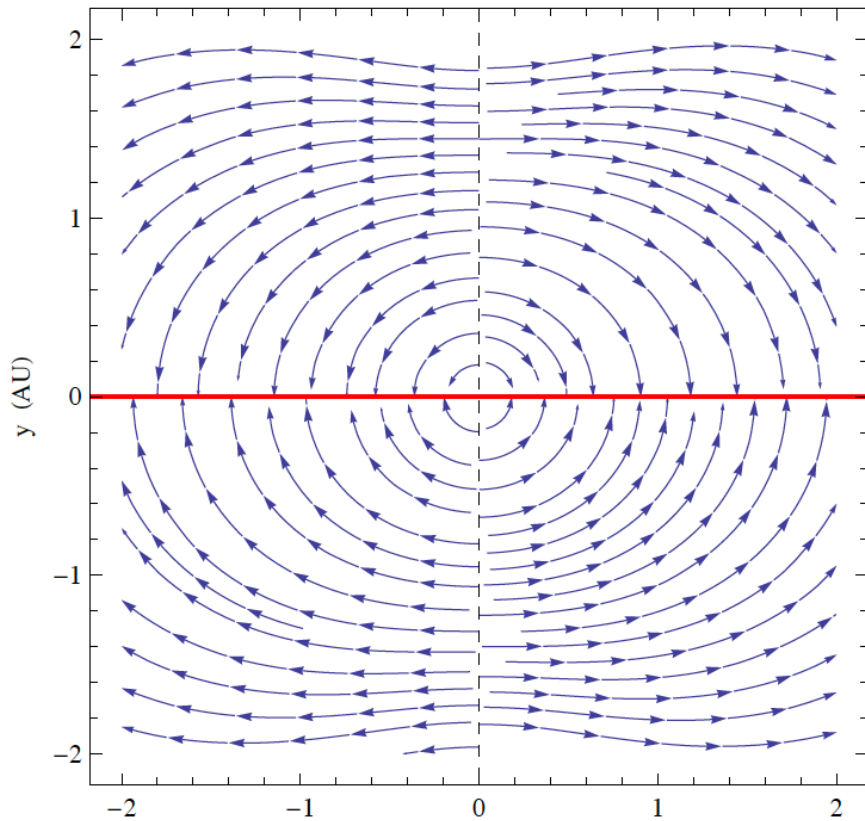
$$A = +1$$

$$q = -1$$



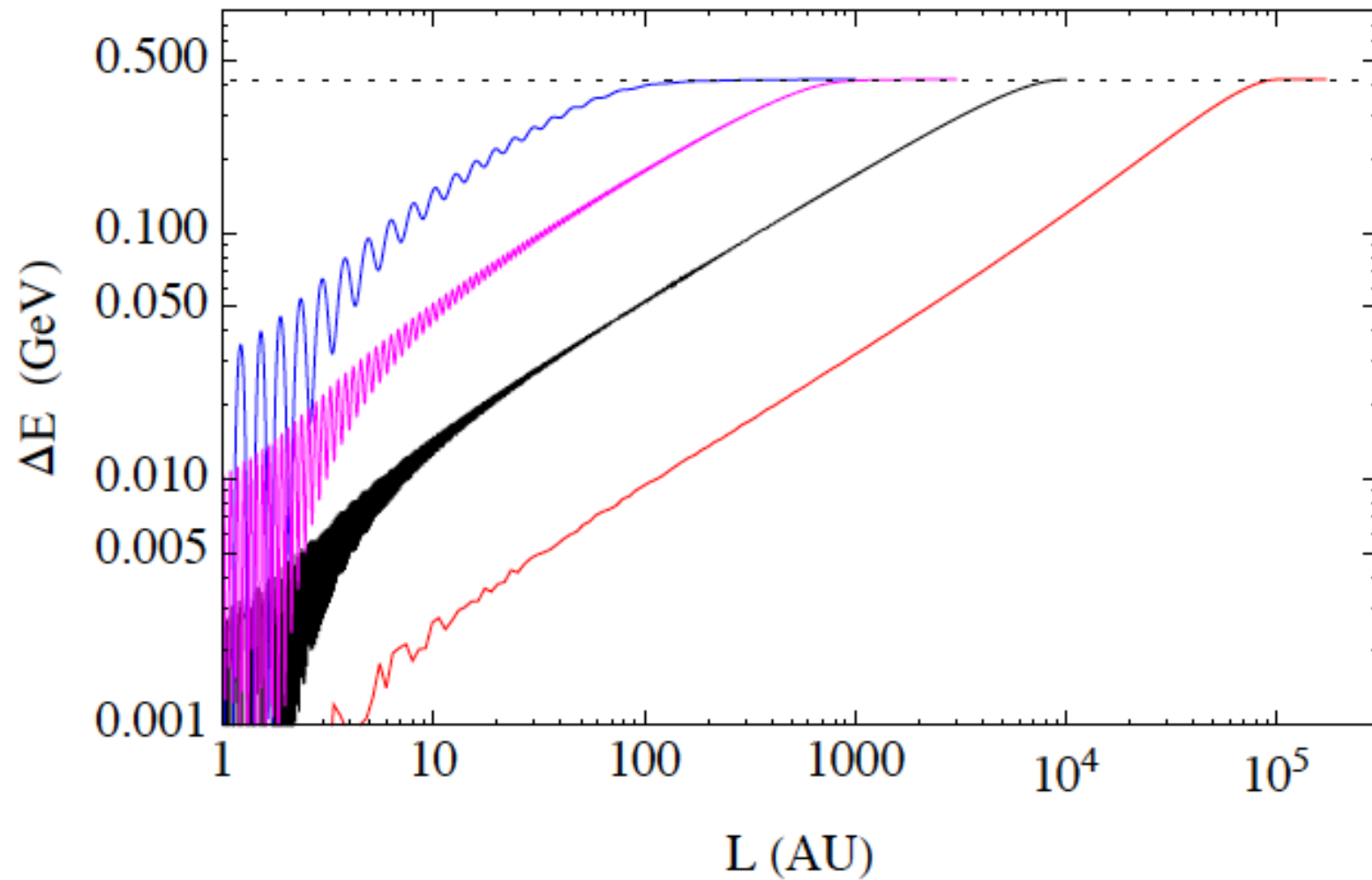
Stream lines of the drift velocity

$$q A > 0$$



Stream-lines of the
“regular” Electric Field

Back-Tracing e^+ ($A=+1$) 1,3,10,20 GeV



$$\vec{E}(x, y, z) = \pm A B_0 \frac{\Omega r_0^2}{c r^3} \{x z, y z, -(x^2 + y^2)\}$$

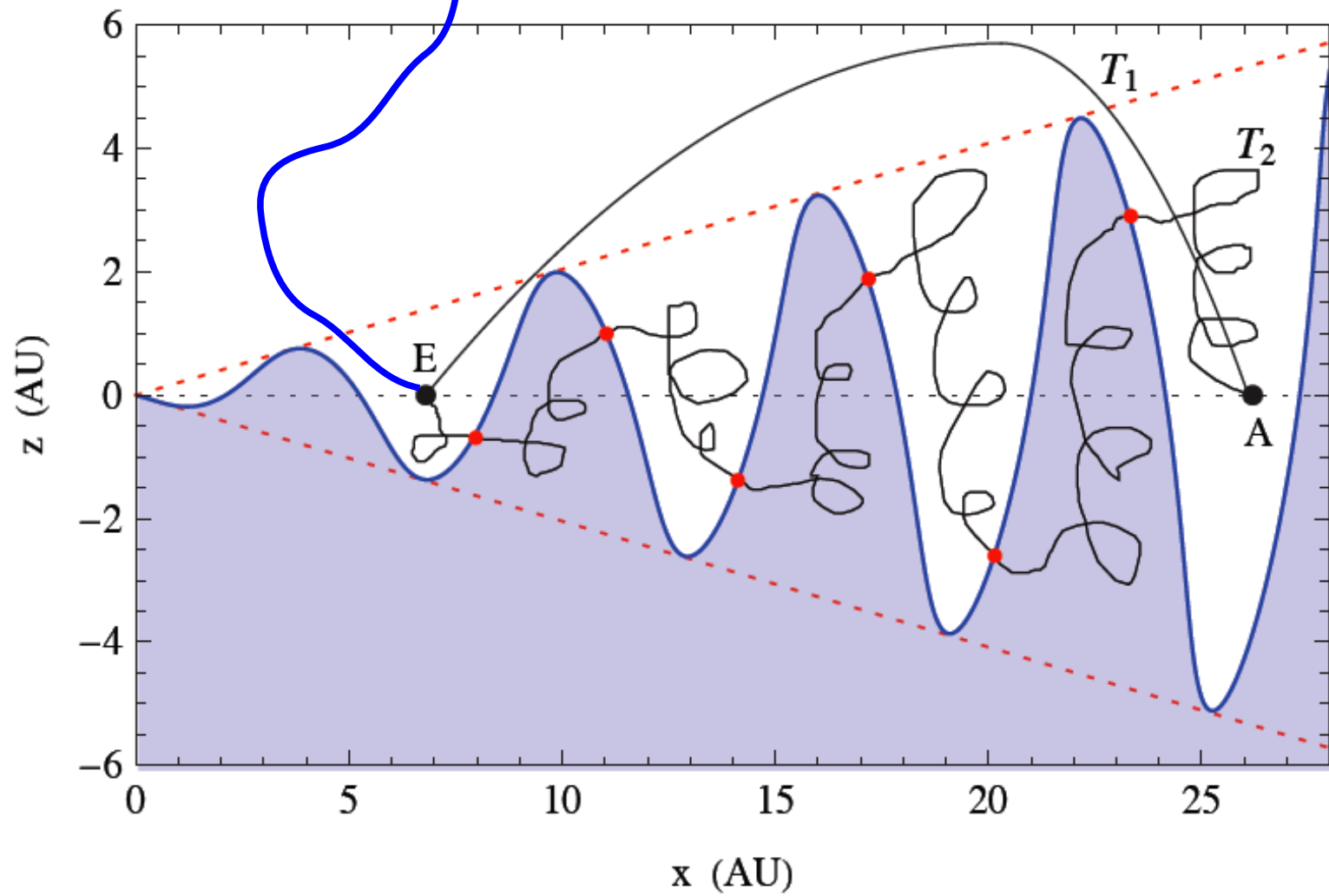
$$\vec{E}(x, y, z) = -\nabla V(x, y, z)$$

$$V(x, y, z) = \mp A B_0 \frac{\Omega r_0^2}{c} \frac{z}{r}$$

Trajectory that does not cross
the Heliospheric Current Sheet

$$\int_T d\vec{s} \cdot \vec{E}(s) = V(\vec{x}_f) - V(\vec{x}_i)$$

$$\lim_{z \pm \infty} V(x, y, z) = \mp A B_0 \frac{\Omega r_0^2}{c}$$

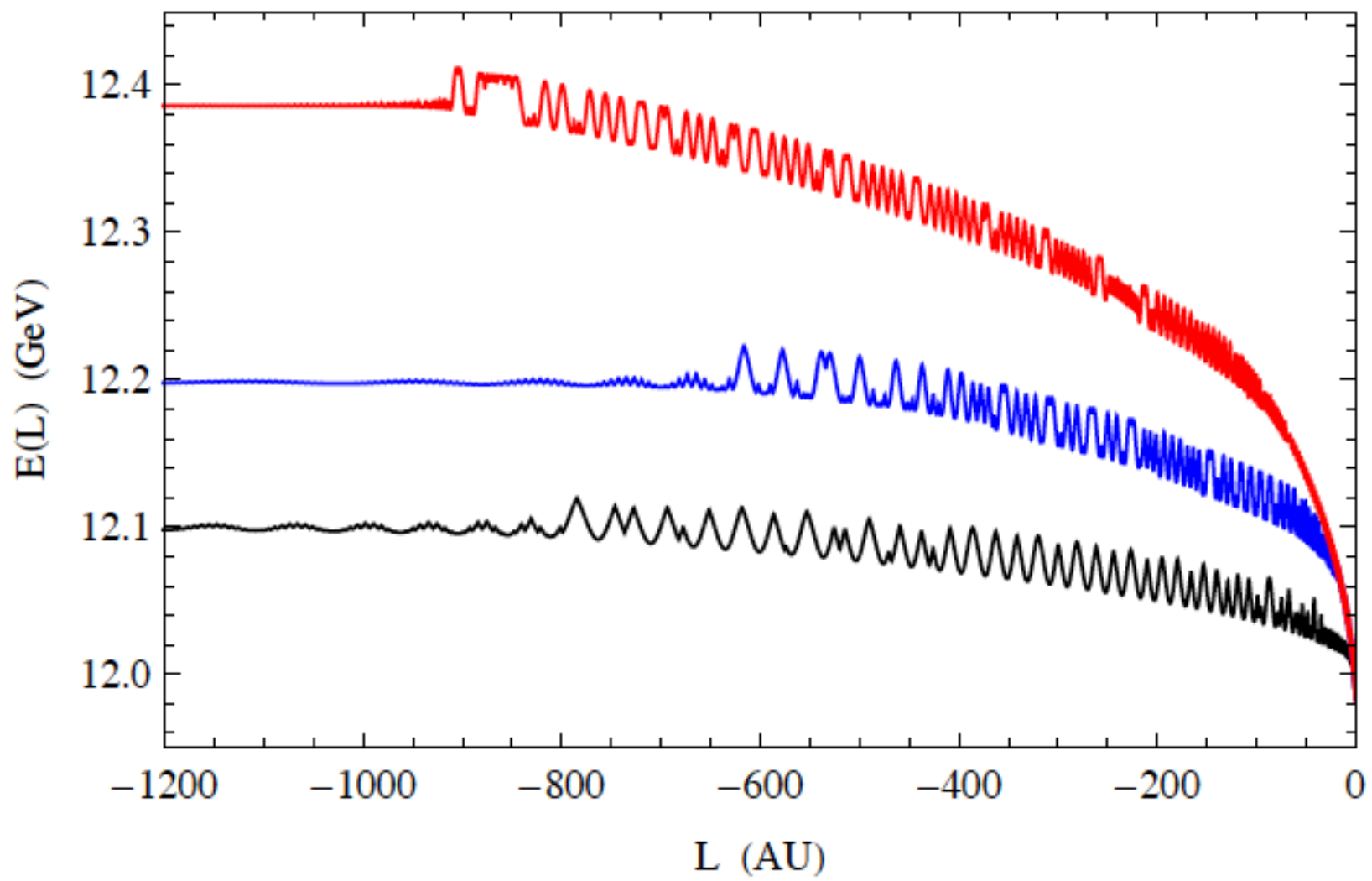


$$V(x, y, z = 0) = 0$$

$$\Delta E|_{qA>0} = |q_e| B_0 \frac{\Omega r_0^2}{c} \simeq 0.422 \left(\frac{B_0}{10^{-4} \text{ Gauss}} \right) \text{ GeV}$$

$$\Delta E|_{qA<0} \simeq 2 \sin \alpha \Delta E|_{qA>0}$$

Energy Losses attributable to
the regular heliospheric electric field



Concluding Remarks:

The Heliosphere is a fundamental “Laboratory” to study the propagation of relativistic particles. Many problems relevant for Milky Way propagation can be studied in detail.

The AMS02 data with their high statistics (together with various measurement of the (time varying) properties of the heliospheric environment) can provide very important information.

The “regular” Heliospheric Magnetic Field has associated a regular electric field that is very important in the energy evolution of the particles that penetrates the heliosphere.