

Gamma-Spectroscopy Experiments with PRESPEC-AGATA at GSI

J. Gerl

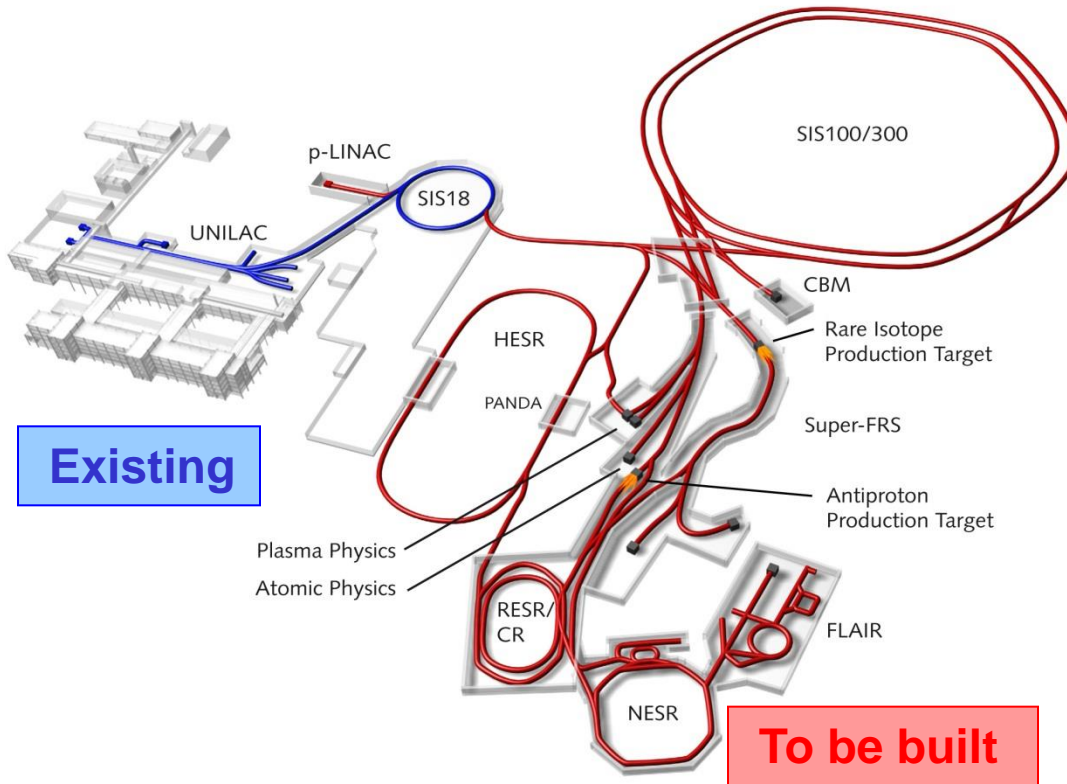
GSI Darmstadt, Germany

NSP13

Padova, Italia

10.6.2013

FAIR – The Science



Nuclear Structure Physics and Nuclear Astrophysics with RIBs

Structure of exotic nuclei far off stability;
 Nuclear synthesis in stars and star explosions;
 Fundamental interactions and symmetries

Hadron Physics with Antiproton Beams

Quark gluon structure and dynamics of “strong” interacting particles;
 Origin of the confinement and mass of hadrons
 Transversity measurement via polarized antiprotons and pol. protons

Physics of Nuclear Matter with Relativistic Nuclear Collisions

Studies of hadronic matter at high densities;
 Phase transitions in quark matter;
 Properties of neutron stars

Plasma Physics with highly Bunched Beams

Bulk matter at very high pressures, densities, and temperatures

Atomic Physics and Applied Science

Highly charged atoms; Low energy anti-protons
 Laser cooling

NUclear STructure Astrophysics and Reactions

What are the limits for existence of nuclei?

Where are the proton and neutron drip lines situated?

Where does the nuclear chart end?

How does the nuclear force depend on varying proton-to-neutron ratios?

What is the isospin dependence of the spin-orbit force?

How does shell structure change far away from stability?

How to explain collective phenomena from individual motion?

What are the phases, relevant degrees of freedom, and symmetries of the nuclear many-body system?

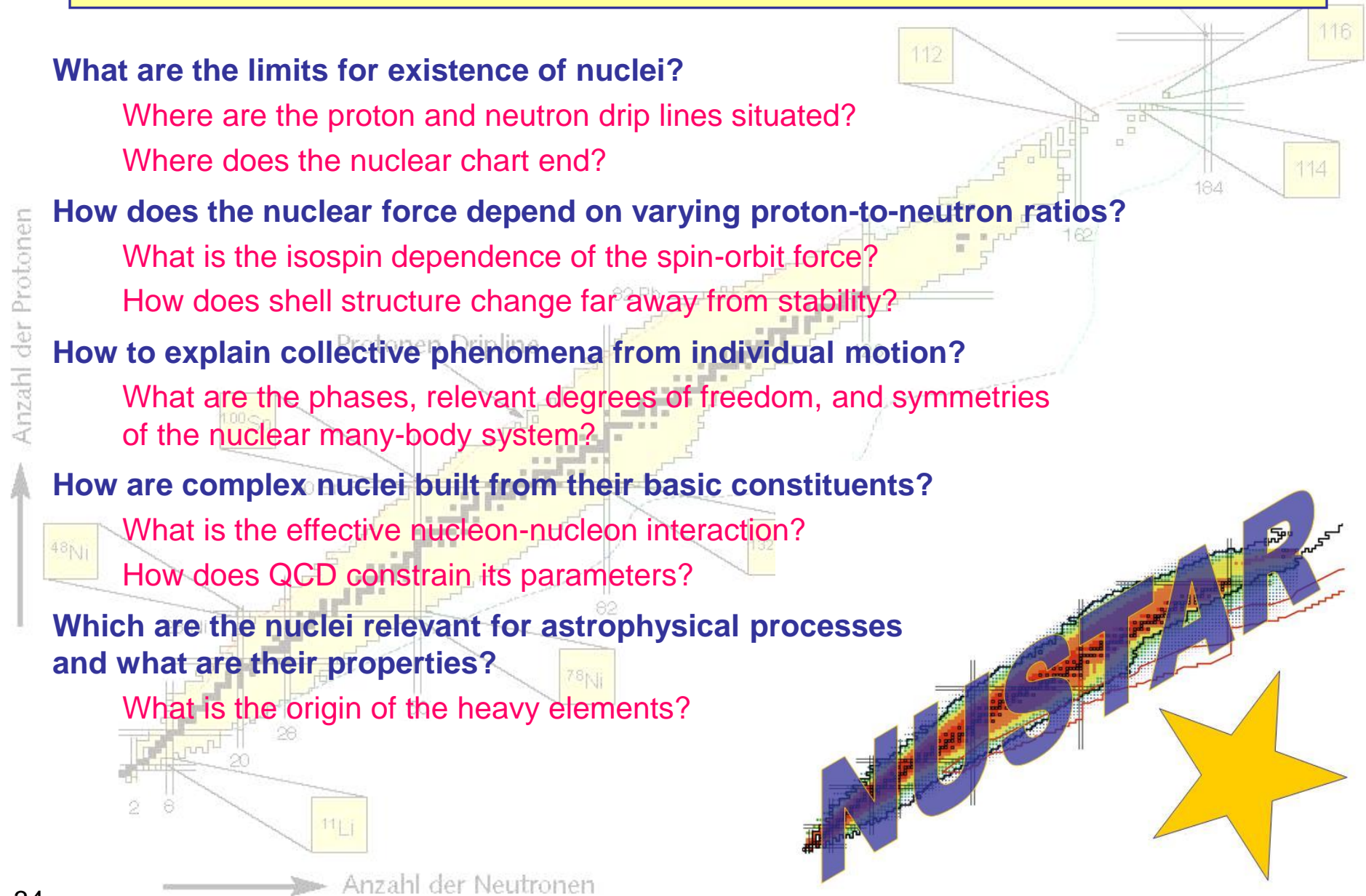
How are complex nuclei built from their basic constituents?

What is the effective nucleon-nucleon interaction?

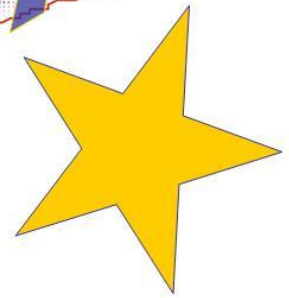
How does QCD constrain its parameters?

Which are the nuclei relevant for astrophysical processes and what are their properties?

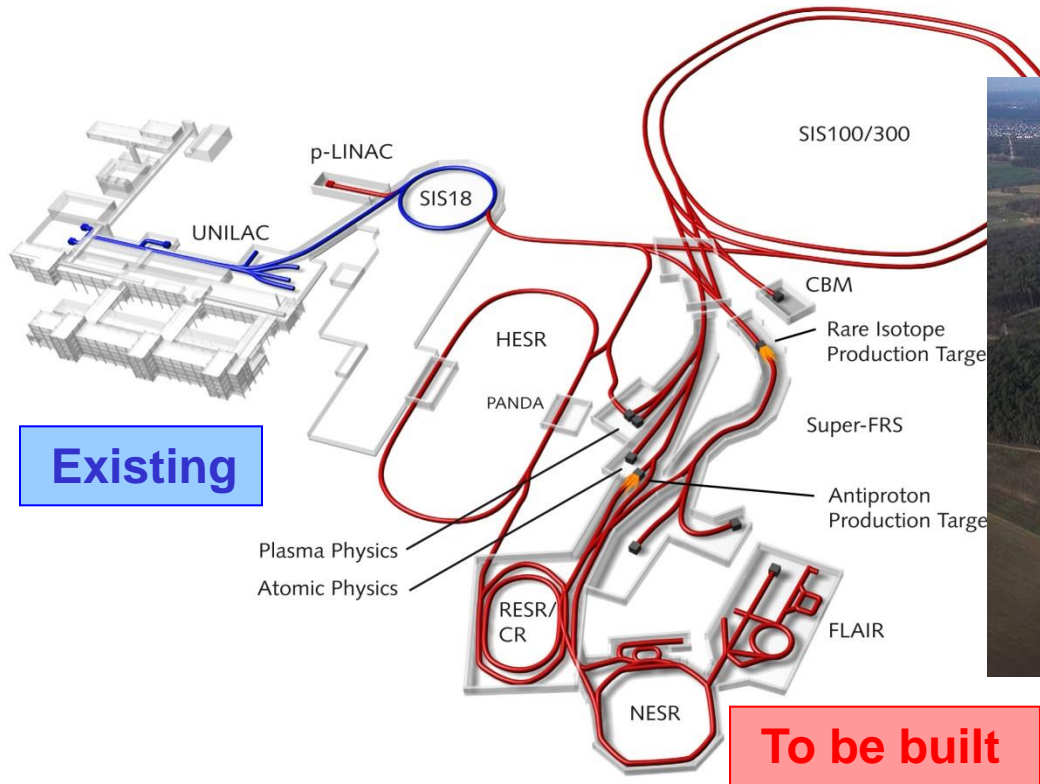
What is the origin of the heavy elements?



NUSSTAR

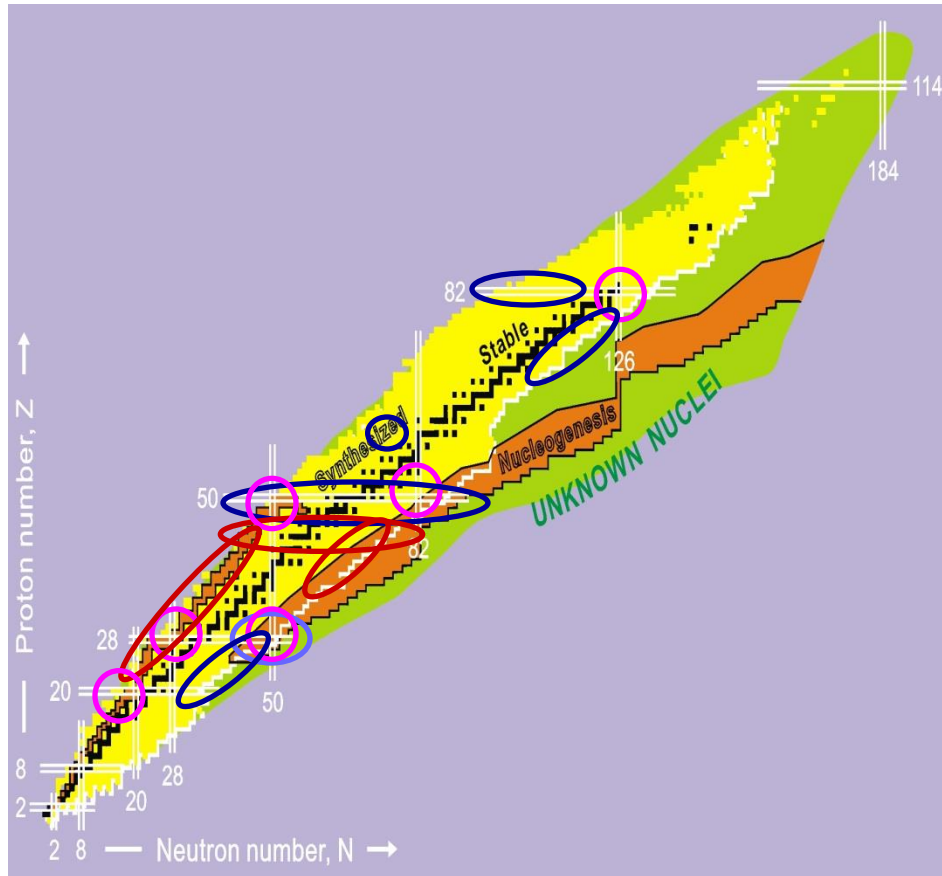


FAIR – Today's Reality



NUSTAR follows an evolutionary approach, constantly improving their instrumentation until FAIR becomes reality, using radioactive beams from the FRS at GSI for methodology development and to perform NUSTAR physics experiments

Nuclear Spectroscopy employing RIBs at GSI



Nuclear Shell structure

- $N \approx Z$
- $N \gg Z$

Nuclear shapes

- Quadrupole, Octupole, Triaxiality
- Shape transitions
- High K-isomers

Collective modes

- $N \gg Z$: GDR soft mode

Nuclear Symmetries

- mirror-isospin, pn-pair correlation

Nuclear astrophysics

- r, rp process

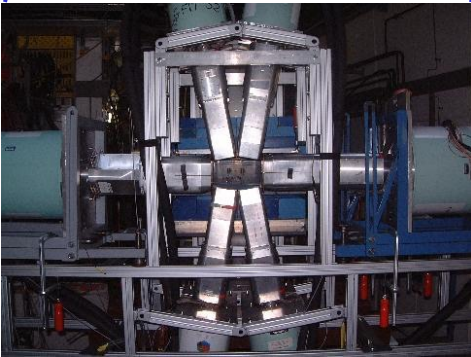
Coulomb excitation, Fragmentation and Decay studies using Rare Isotope Beams and high-resolution γ Spectroscopy

History...

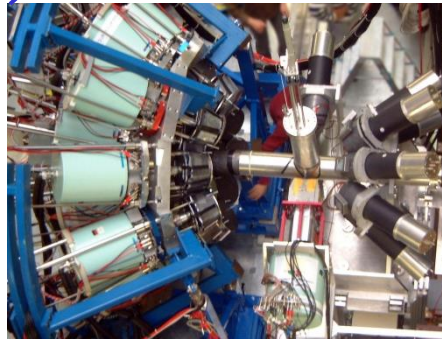


1998 2000 2002 2004 2006 2008

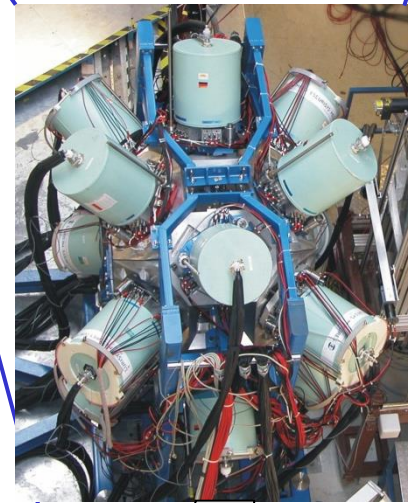
VEGA
Isomer campaign



RISING
Fast campaign



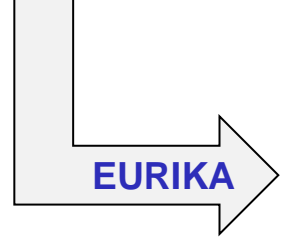
RISING
Stopped campaign



g-RISING



EURIKA



From RISING to HISPEC/DESPEC

RISING at GSI stopped in August 2009

Want to continue successful spectroscopy programme

HISPEC/DESPEC at FAIR starts in 2019

Need to commission and implement new instrumentation



Decay and In-beam spectroscopy programme at the FRS until HISPEC/DESPEC starts

Employing new instrumentation as it becomes available

Platform for coordinated test and commissioning of HISPEC/DESPEC components

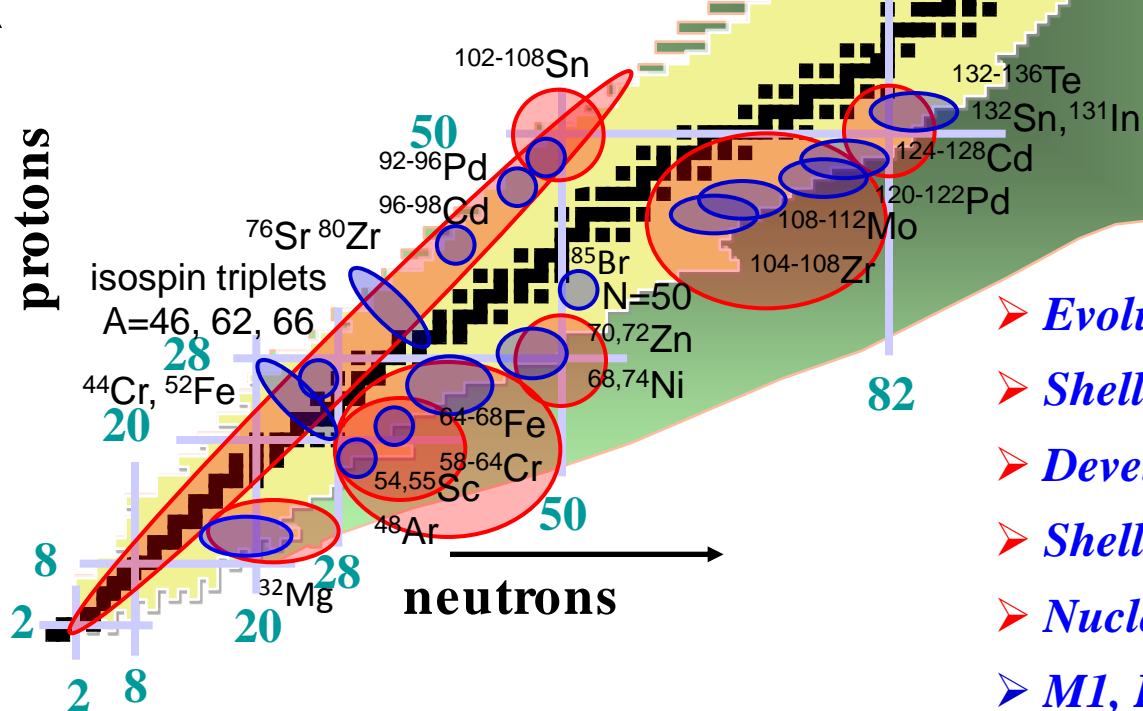
Organisational framework of the spectroscopy community at GSI/FAIR

PRESPEC-AGATA Physics Campaign 2012-2014

Physics workshop 4.-7.5.2010 in Istanbul
 → 34 LOIs, 12 proposals initially proposed,
 9 accepted by GSI PAC in 2011

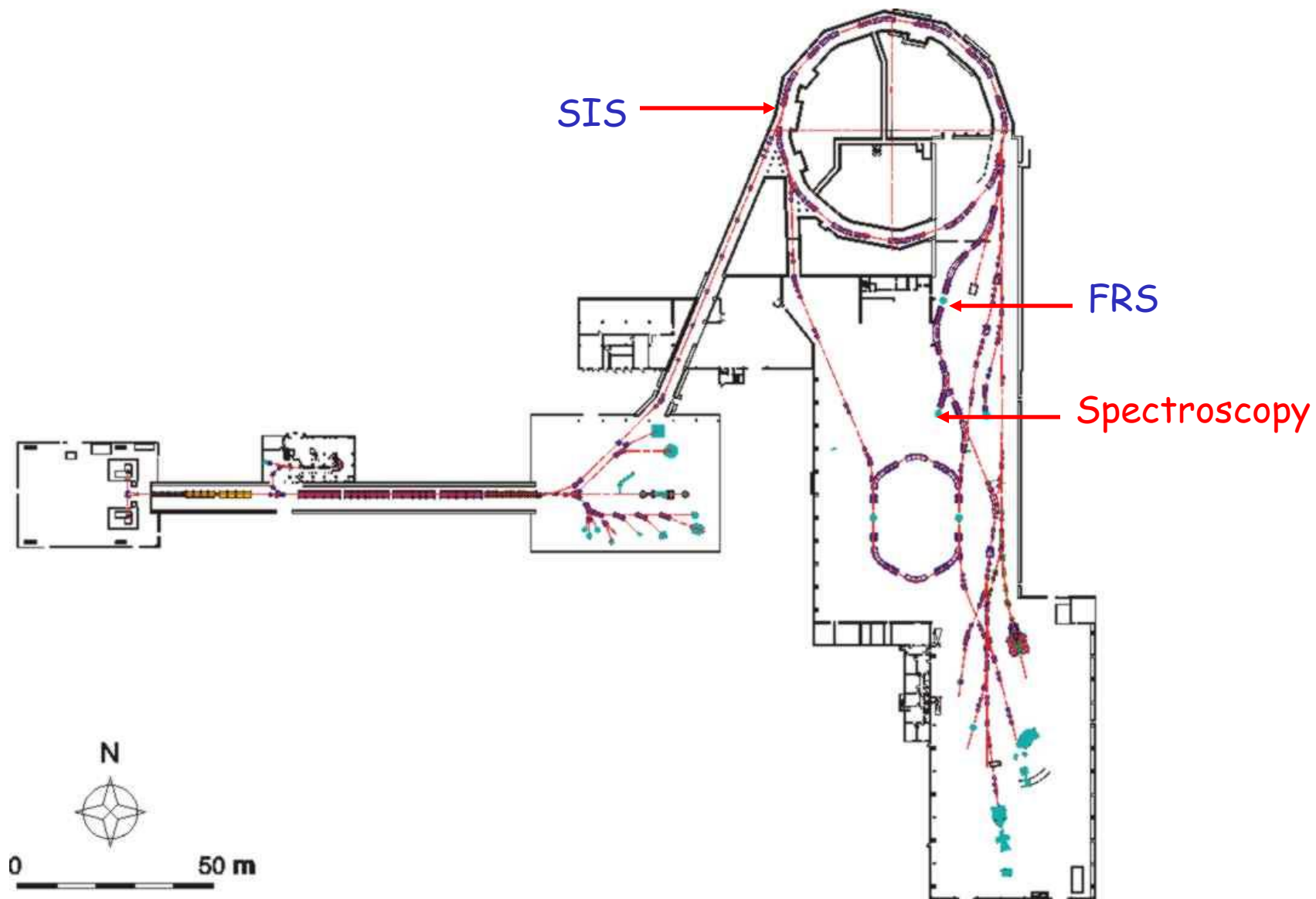
Experiments performed from 27.9. to 21.11.
 Coulomb excitation around ^{208}Pb
 Fine structure in Pygmy resonance
 Coulex on isomeric state in ^{52}Fe
 Lifetimes in heavy Zr-Mo isotopes

Physics workshop
 10.-11.12.2012
 TU-Darmstadt
 → 15 new LOIs



- *Evolution of collectivity near ^{208}Pb*
- *Shell structure near ^{78}Ni , ^{100}Sn , ^{132}Sn*
- *Development of nuclear collectivity*
- *Shell evolution in light nuclei*
- *Nuclear structure at the $N=Z$ line*
- *$M1$, $E1$, $E2$, $E3$ strength*

Layout of the GSI facility



In-beam Spectroscopy

production

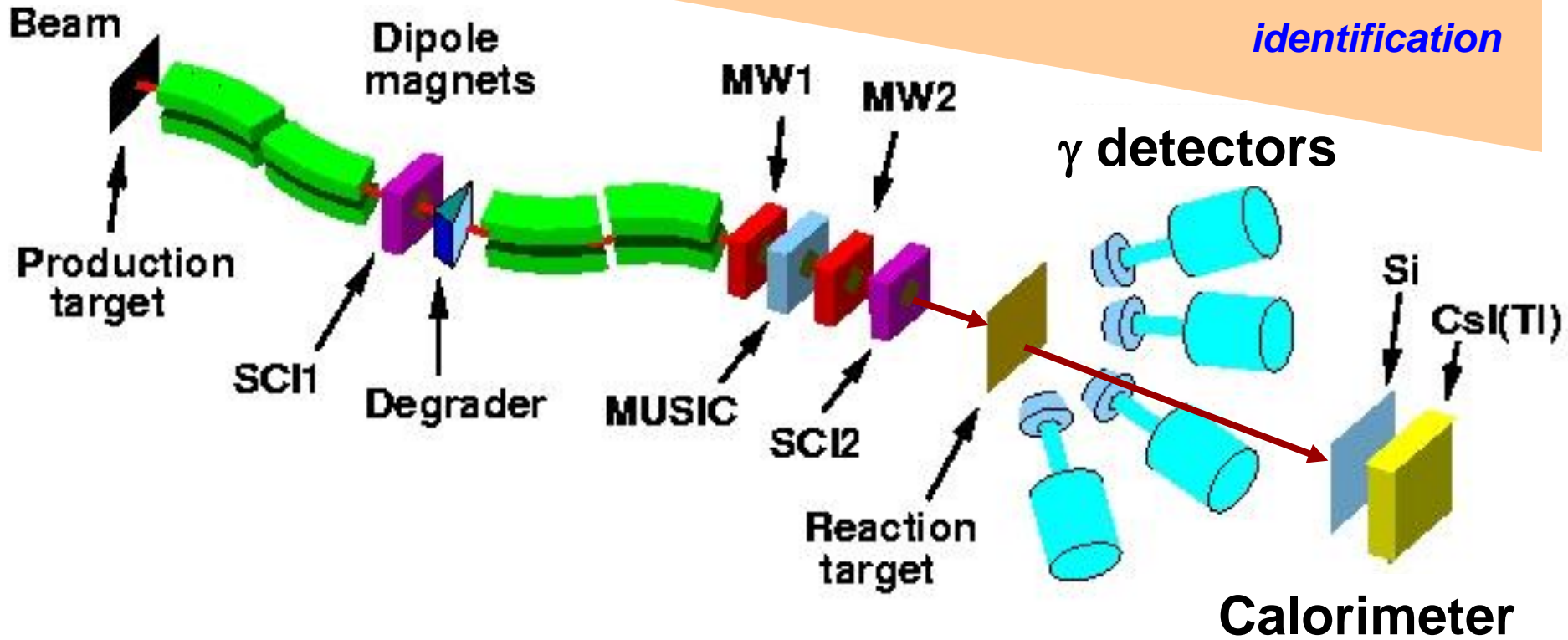
selection

identification

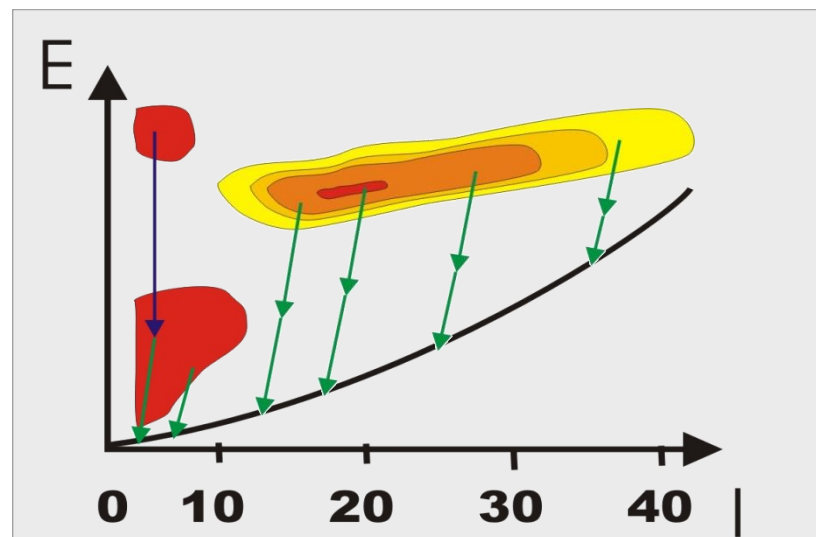
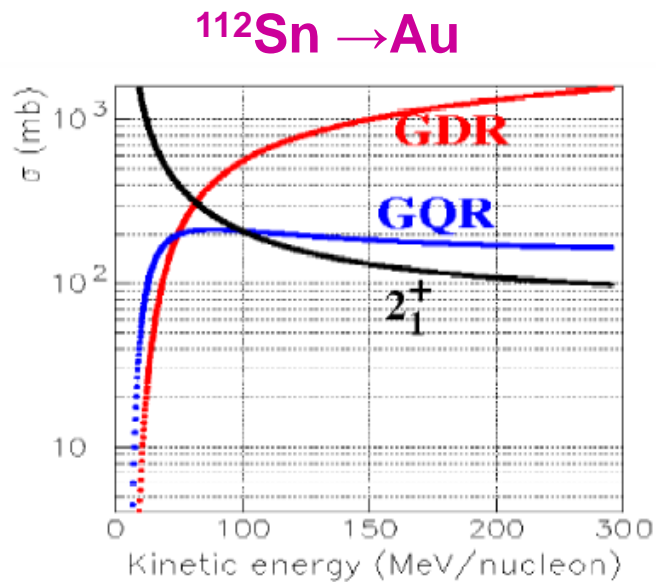
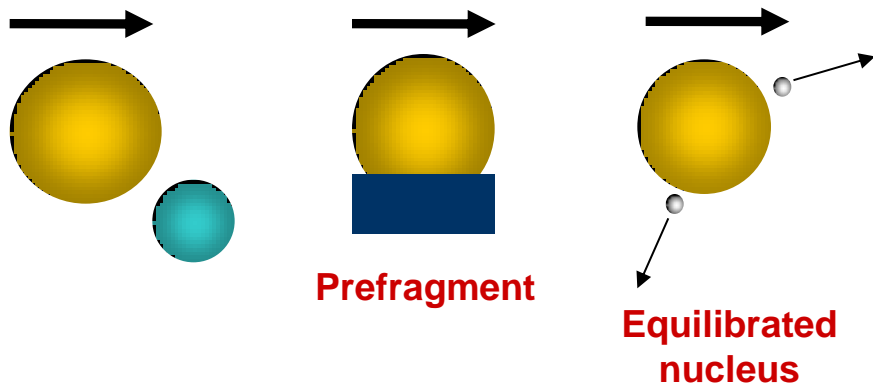
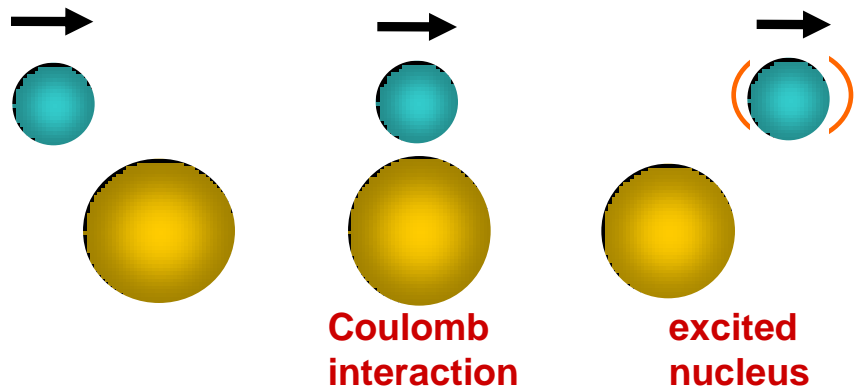
reaction

spectroscopy

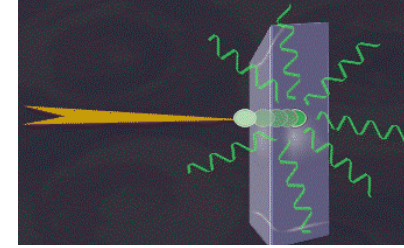
identification



Relativistic Coulomb excitation / fragmentation



Atomic Background Radiation Bremsstrahlung



➤ Radiative electron capture (REC)

capture of target electrons into bound states of the projectile:

$$\sigma \sim Z_p^2 \cdot Z_t$$

➤ Primary Bremsstrahlung (PB)

capture of target electrons into continuum states of the projectile:

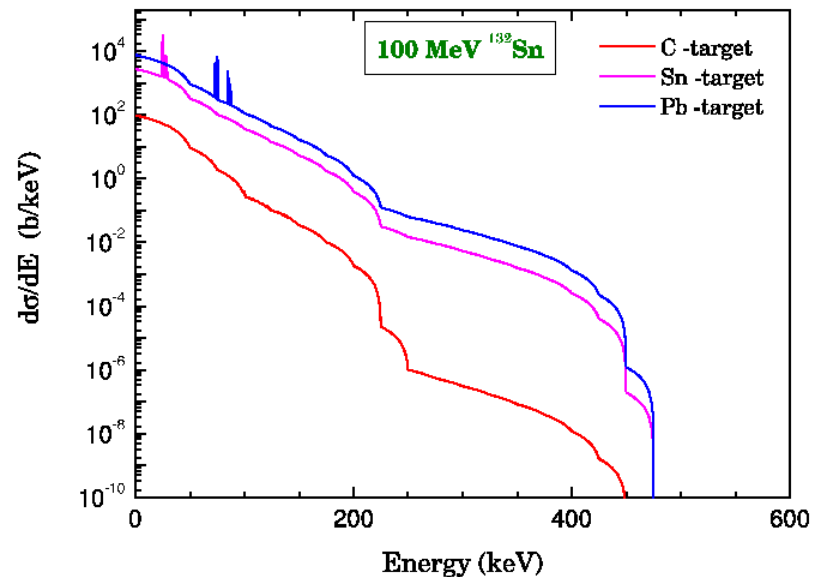
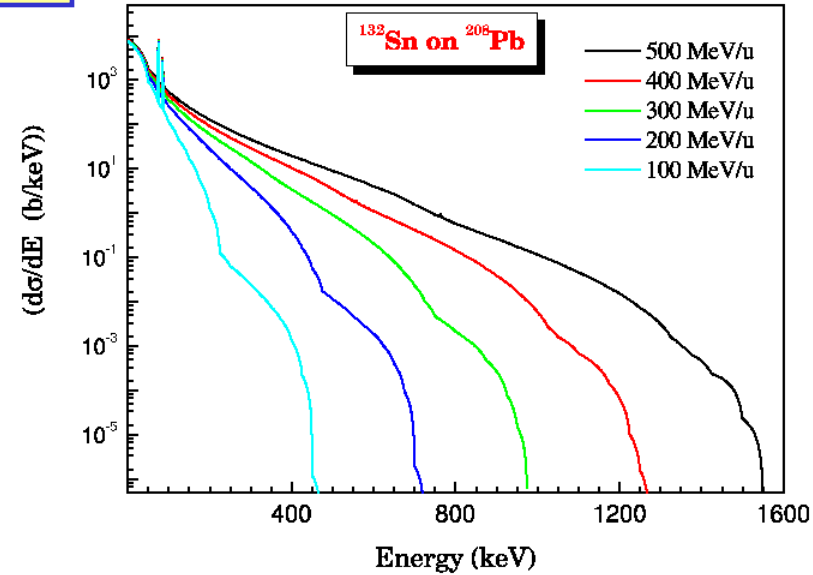
$$\sigma \sim Z_p^2 \cdot Z_t$$

➤ Secondary Bremsstrahlung (SB)

Stopping of high energy electrons in the target: $\sigma \sim Z_p^2 \cdot Z_t^2$

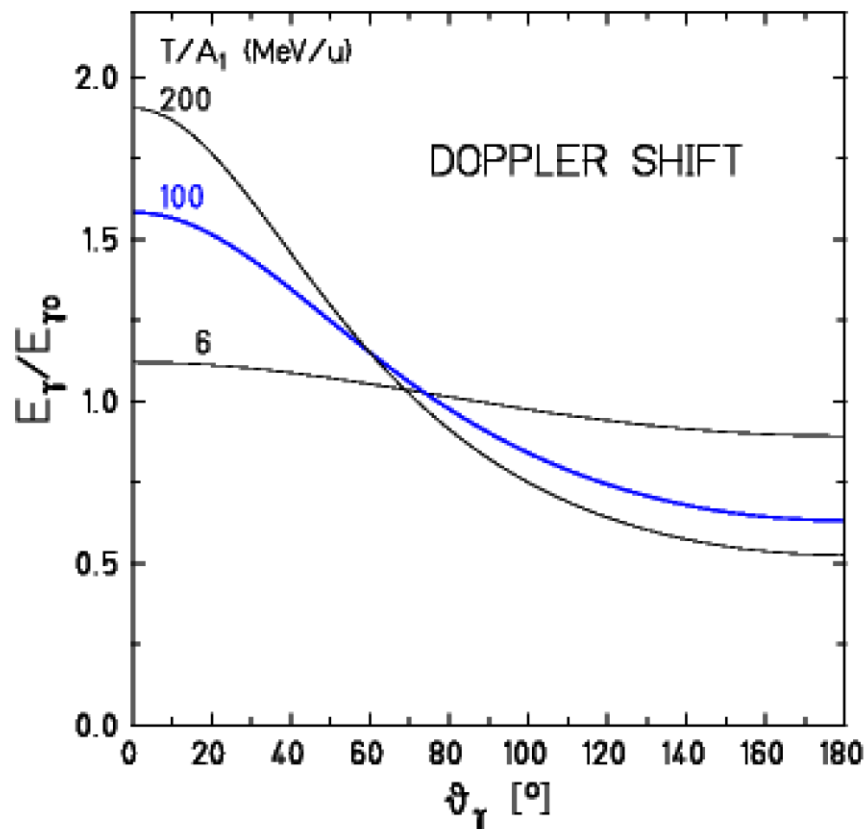


High granularity γ detector

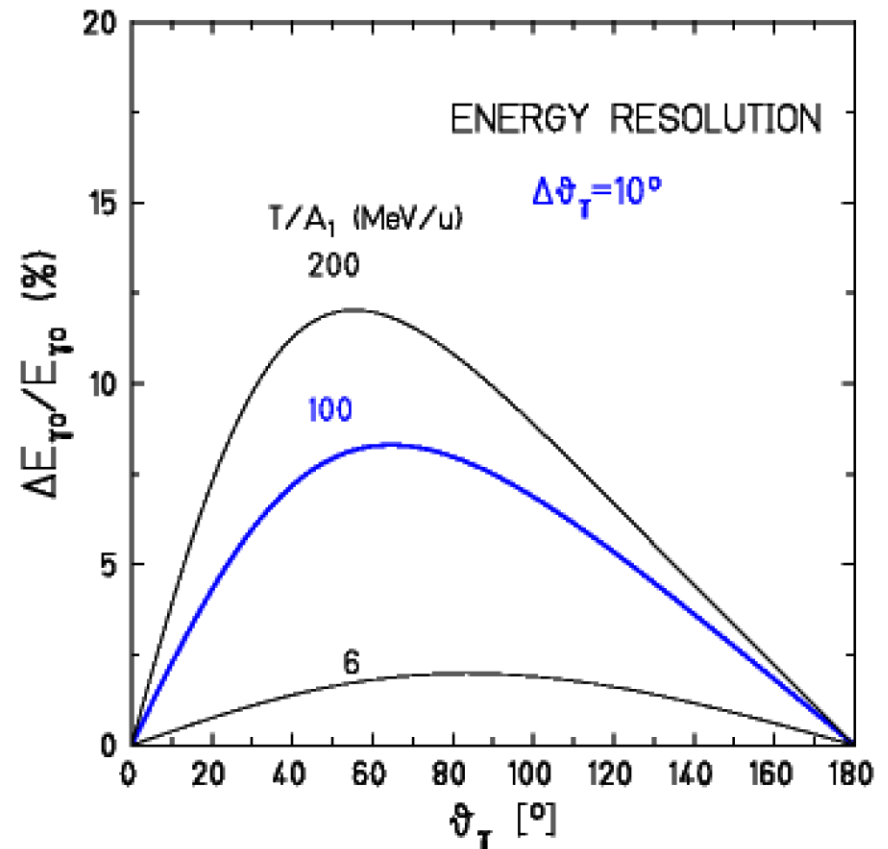


Doppler Effect

Doppler shift



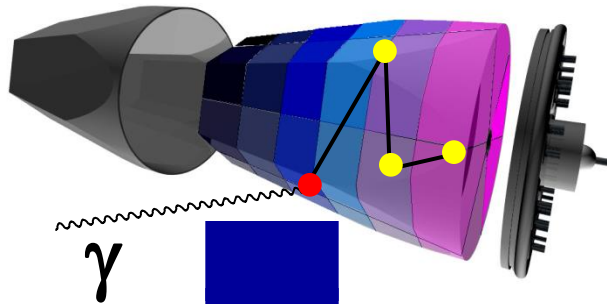
Doppler broadening



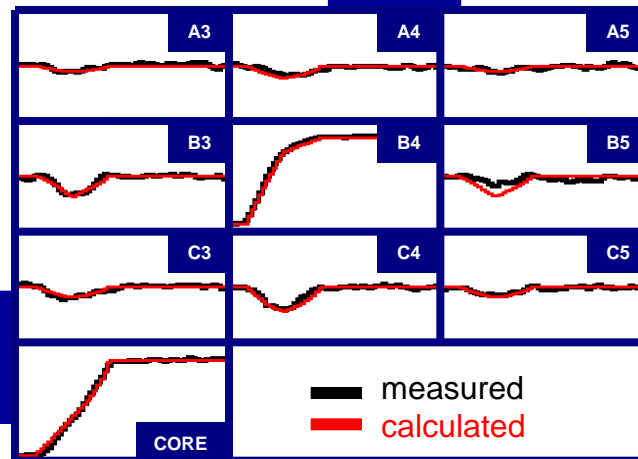
position sensitive γ detector

γ -ray Tracking Arrays

Electrically segmented HPGe detectors

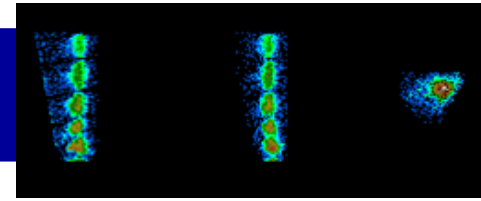


Digital electronics to record and process segment signals



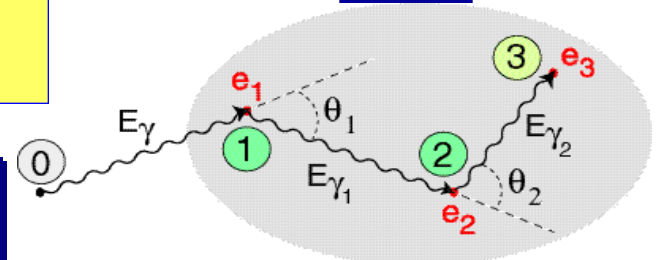
Identification of the individual interaction points

$$(x, y, z, E, t)_i$$



Deconvolution of the recorded waves through comparison with reference signals

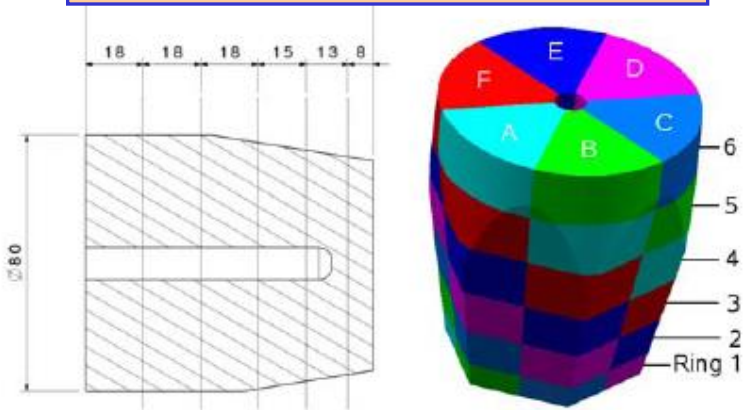
Reconstruction of tracks evaluating permutations of interaction points



Energy and direction of γ -rays

AGATA Detector unit

Large volume 36-fold segmented, encapsulated Ge detector

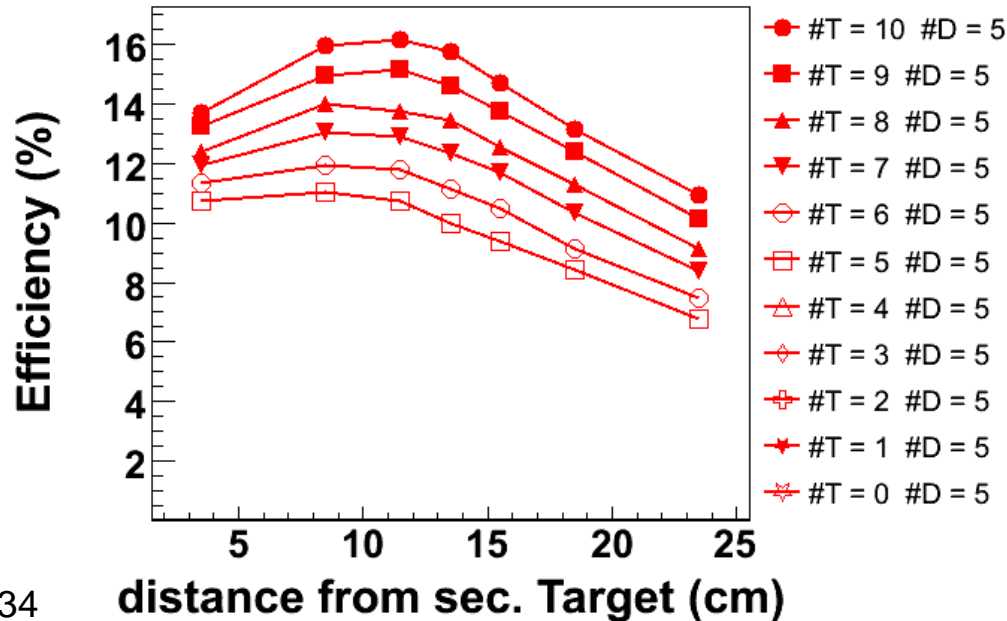
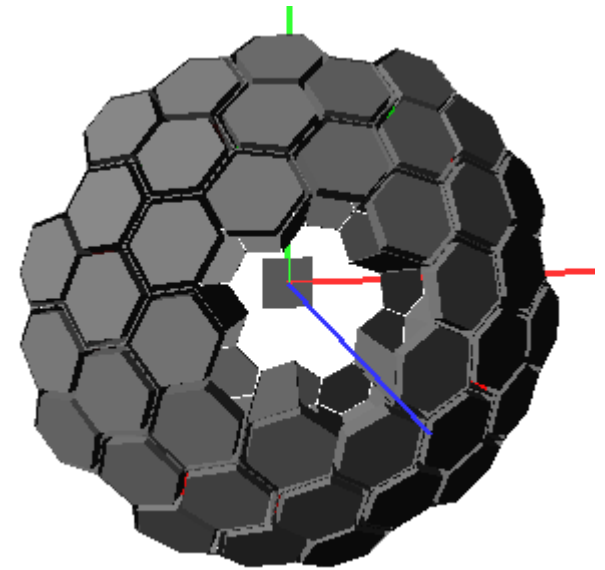
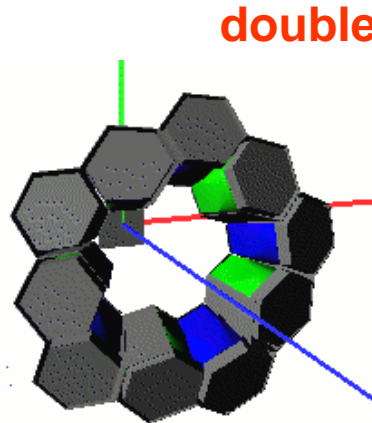
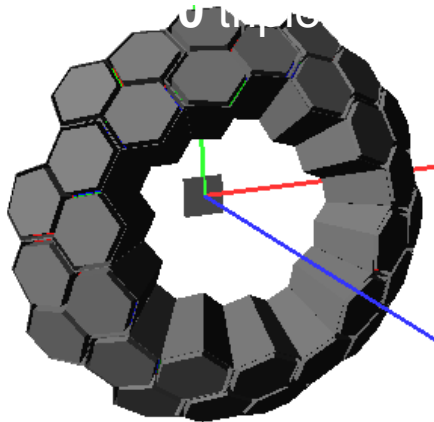


Triple Cluster unit



AGATA at GSI set-up

Challenge: FRS beam size!!!

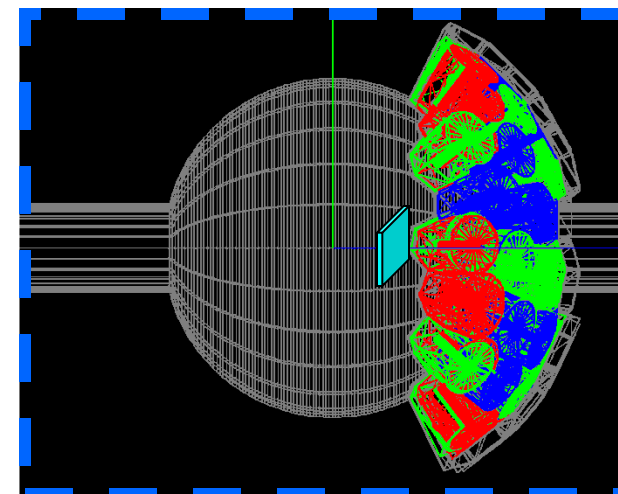
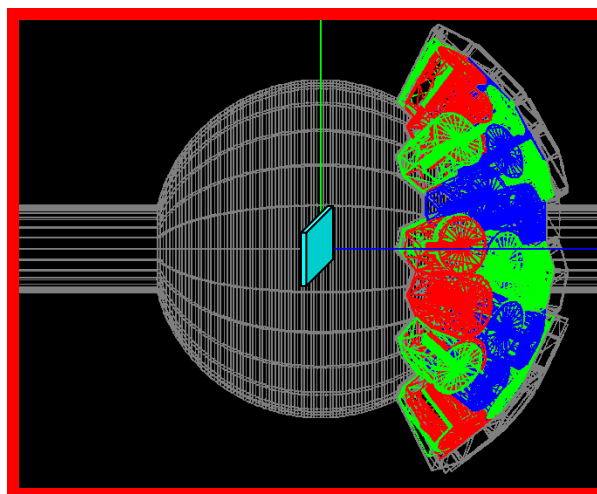
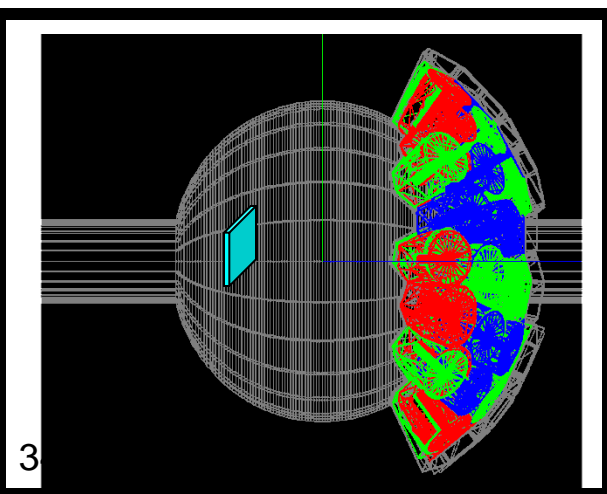
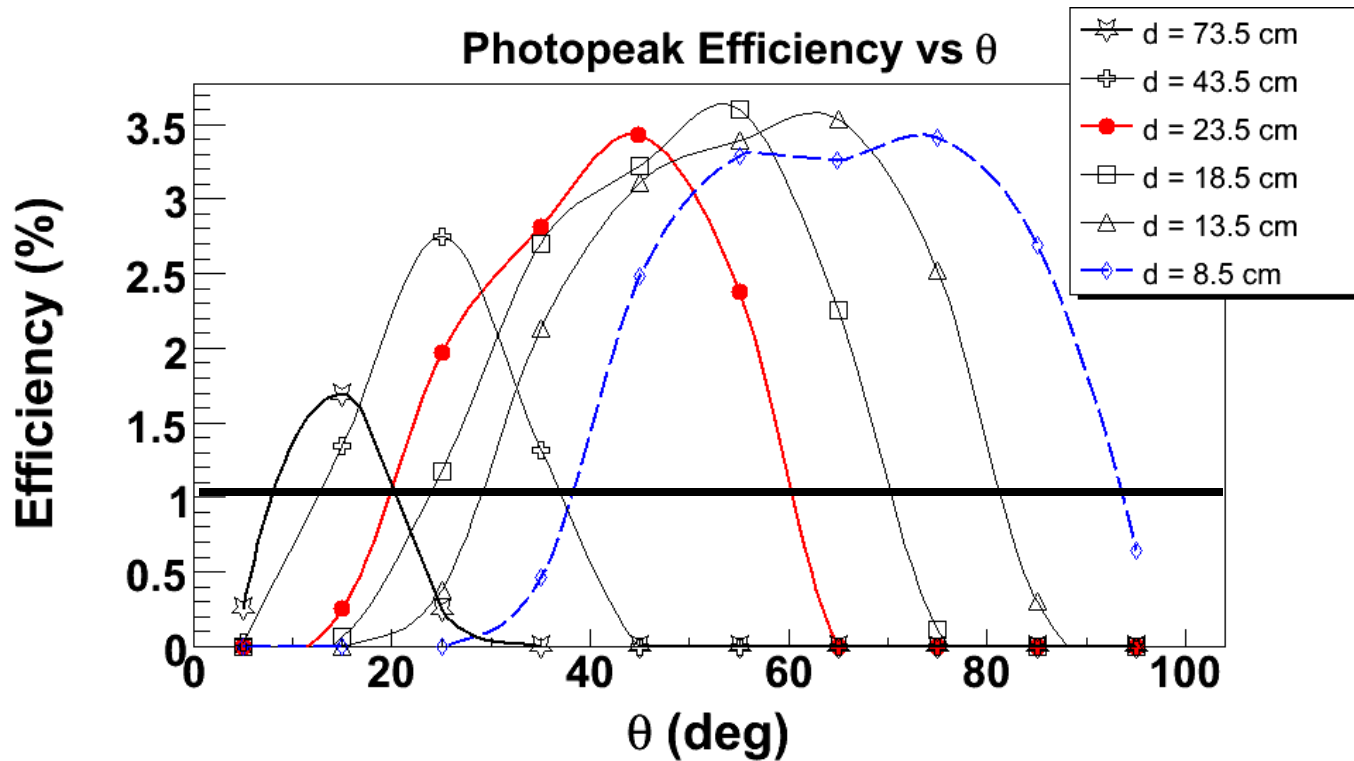


S2' Geometry:

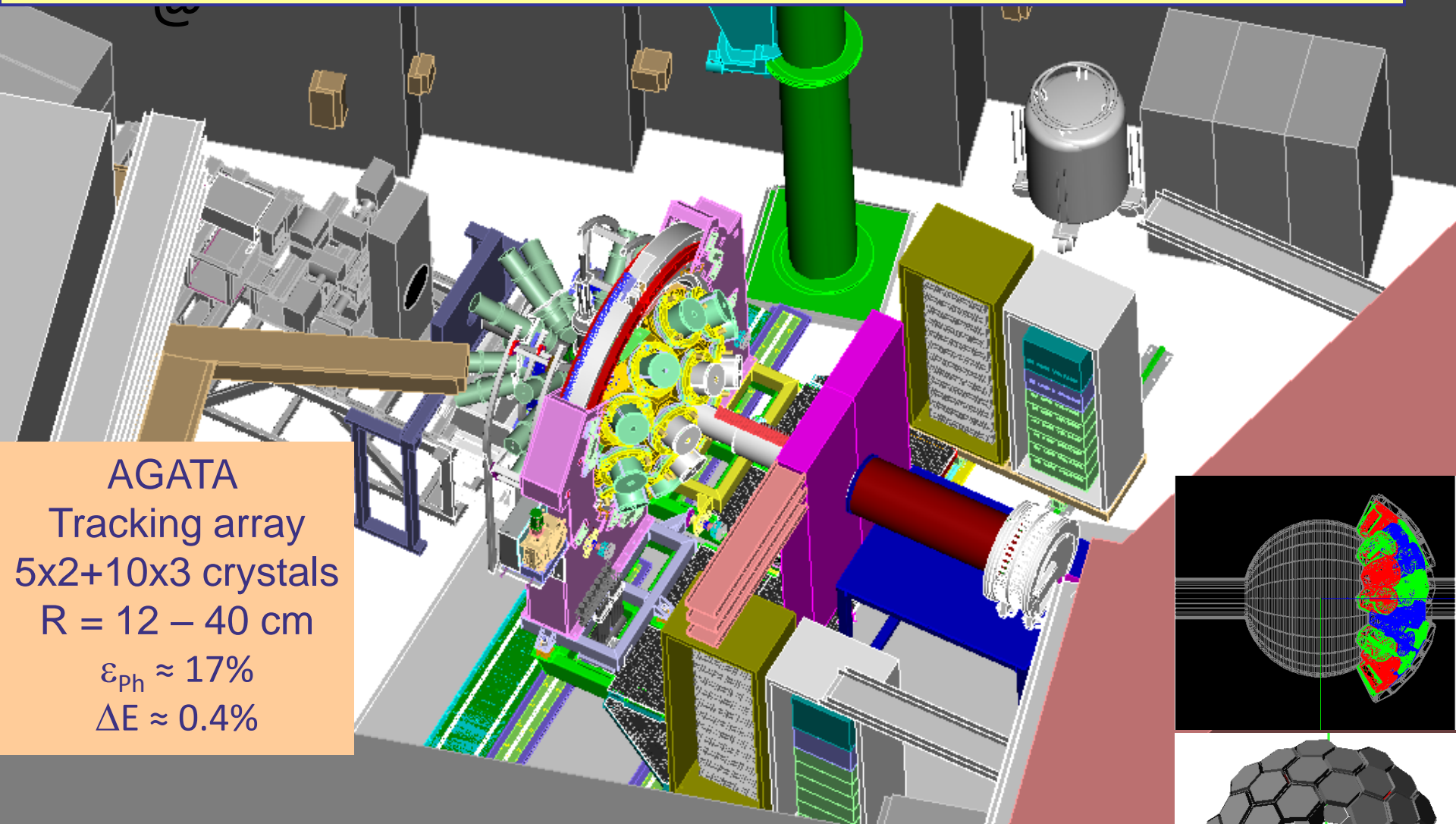
$P_{ph} \leq 17\%$; $\Delta E = 0.4\%$

(sensitivity gain 30x RISING)

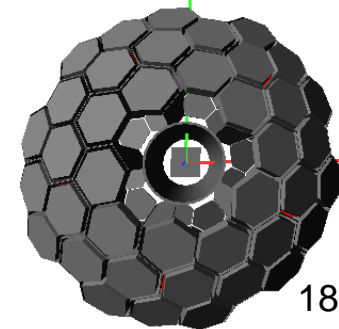
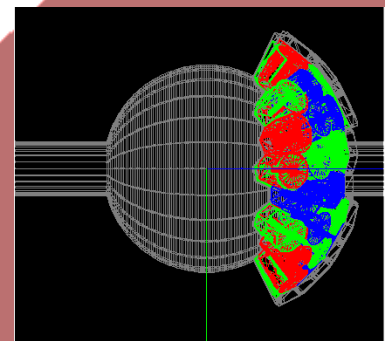
AGATA angular coverage



PRESPEC-AGATA Set-up = Early Implementation of HISPEC



AGATA
Tracking array
5x2+10x3 crystals
R = 12 – 40 cm
 $\epsilon_{ph} \approx 17\%$
 $\Delta E \approx 0.4\%$



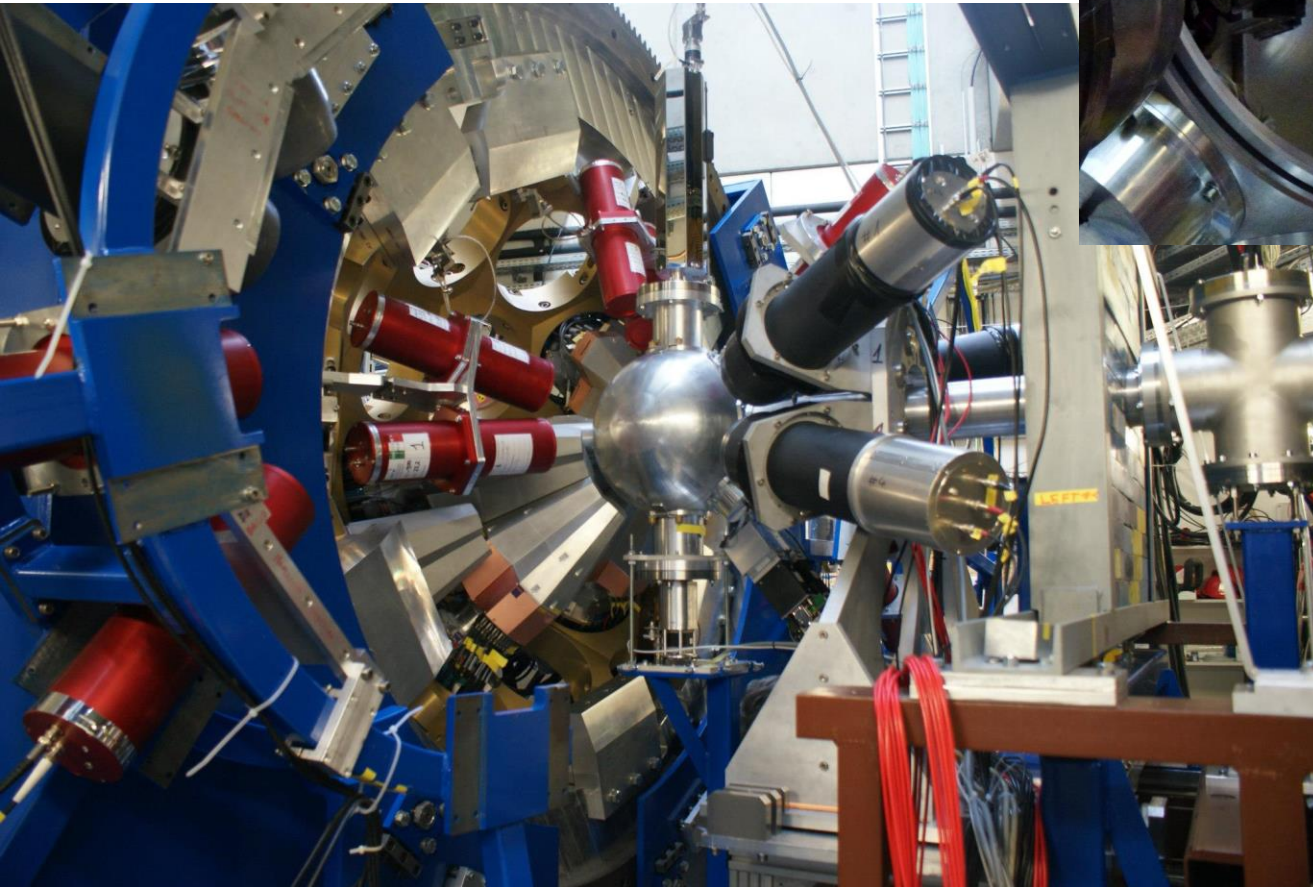
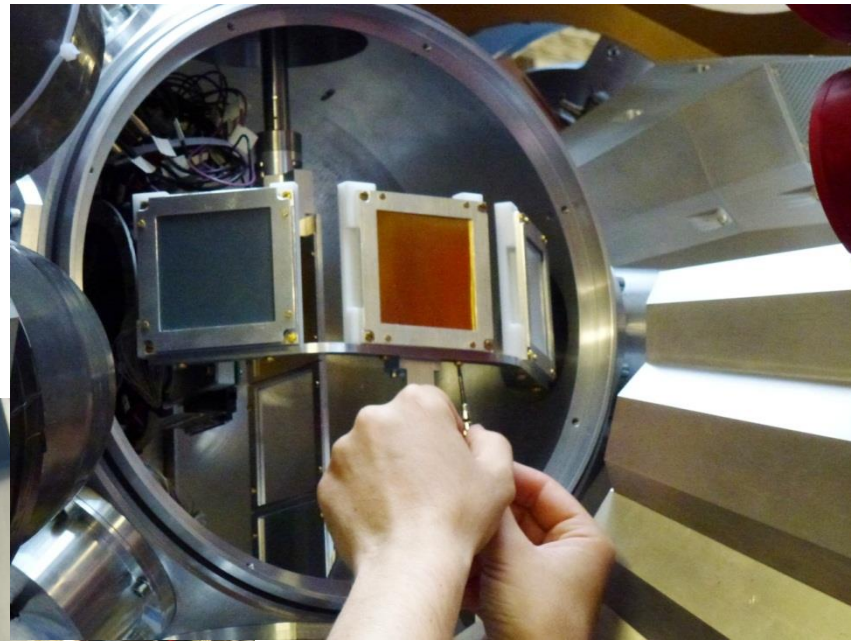
The Set-up in Reality

LYCCA

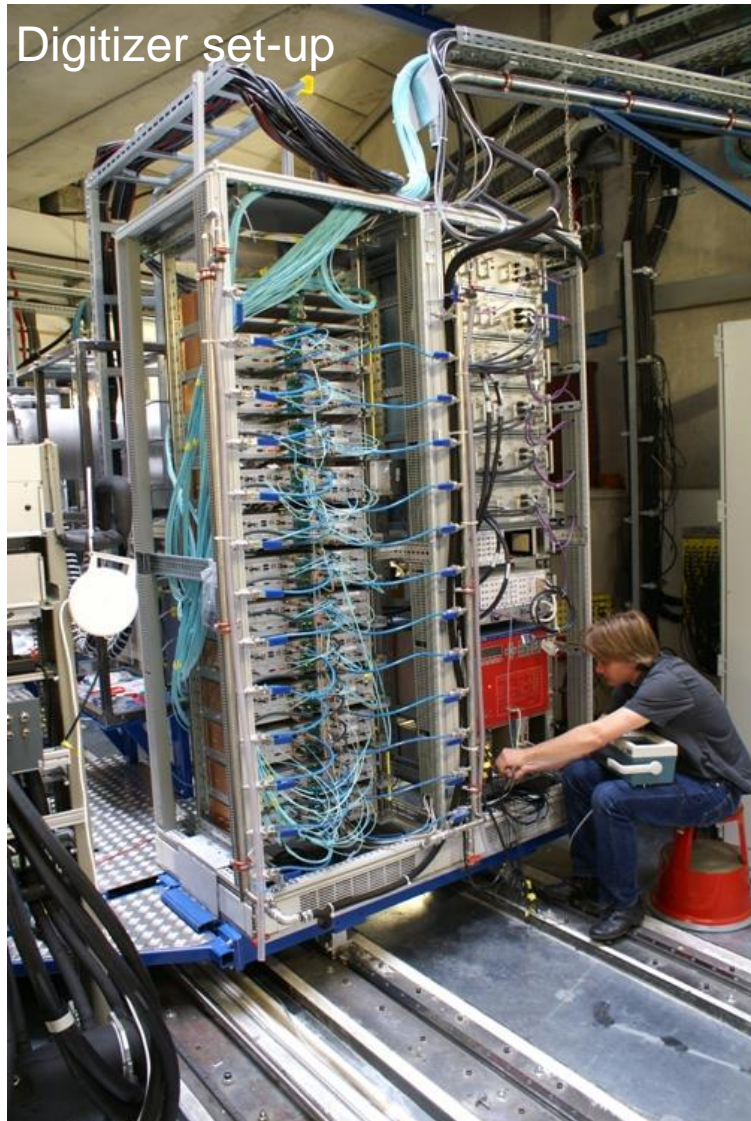
AGATA

Hector

Target chamber



EDAQ



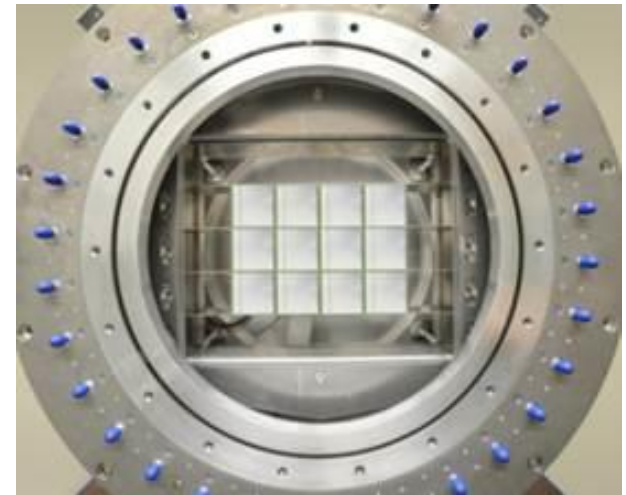
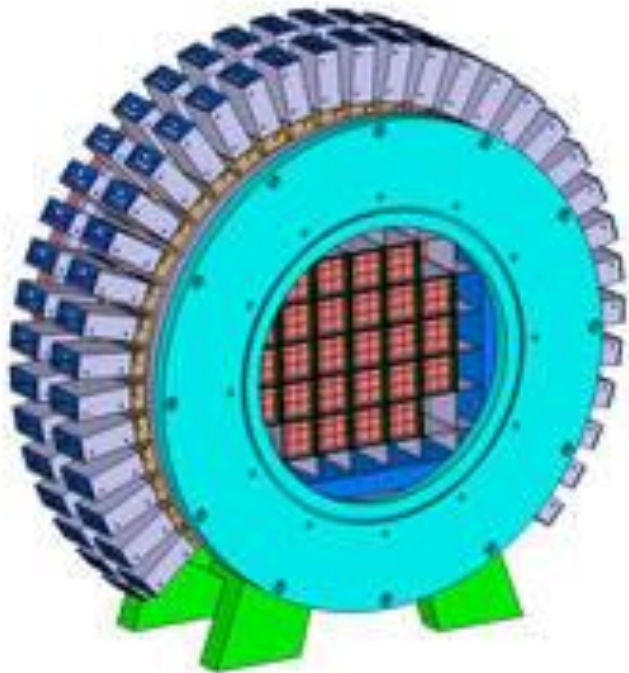
>2000 Channels (mainly high-resolution)
 ≈ 300 Gbyte/s (front-end)
 ≈ 1 Tbyte/d (after trigger and pre-processing)

11 VME Crates + AGATA DAQ

LYCCA

Position sensitive
 ΔE -E calorimeter with
ToF capability

Detects projectile-like reaction product
after the secondary target

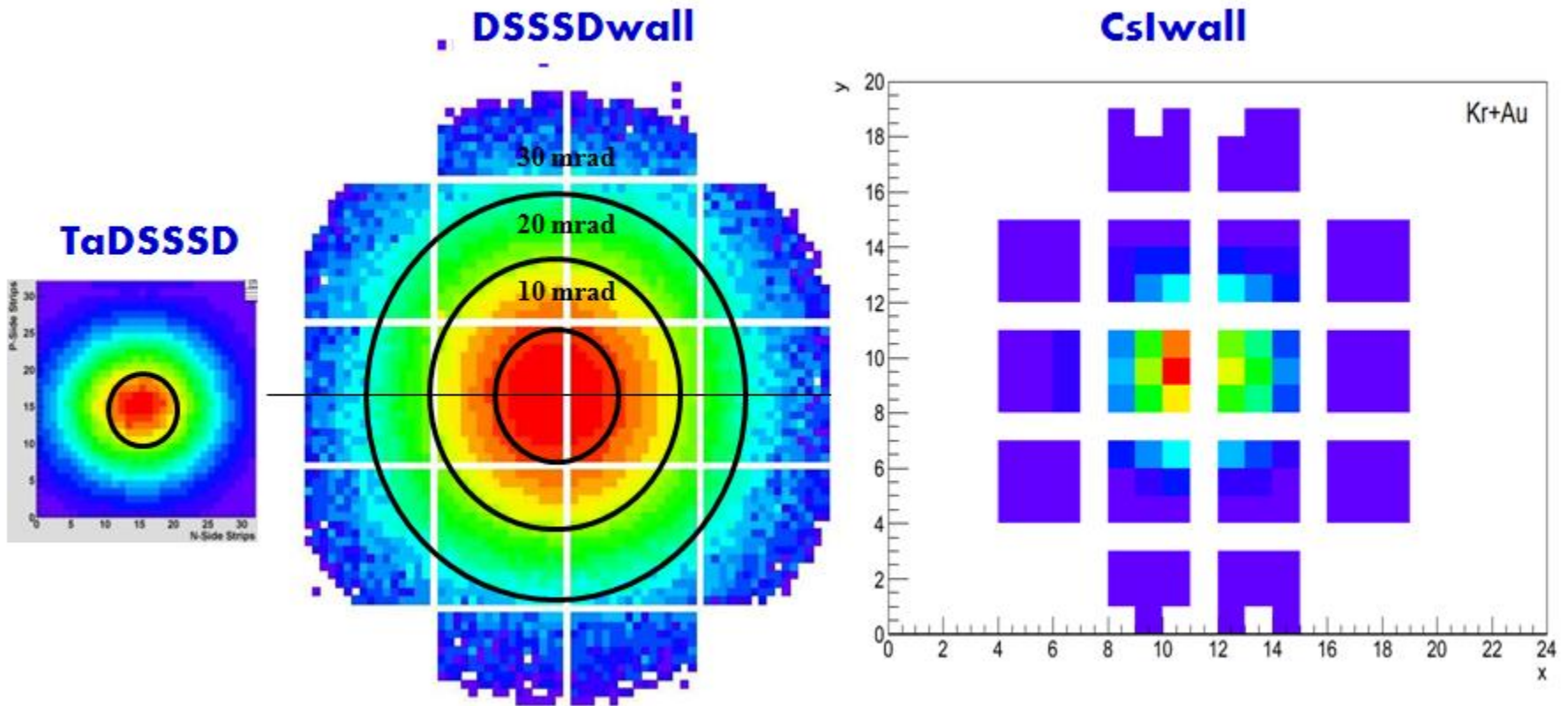


12 DSSDs + 12x9 CsI(Tl)

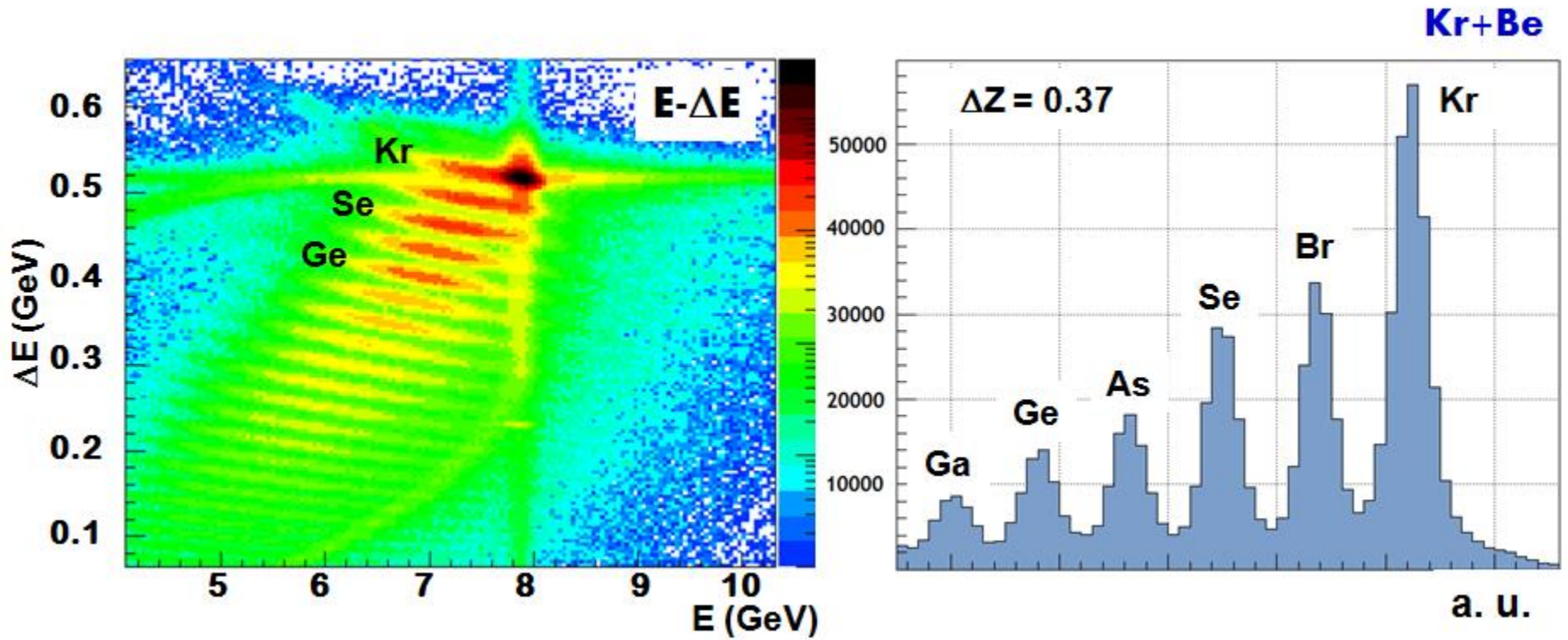


Tof Plastic membrane

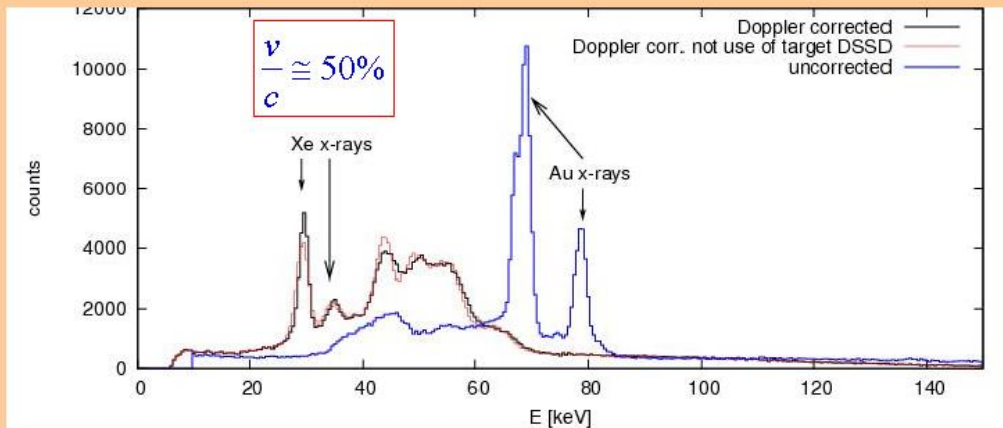
Position measurement



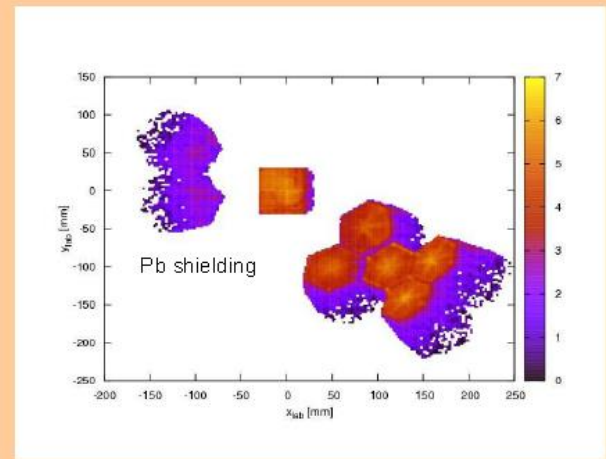
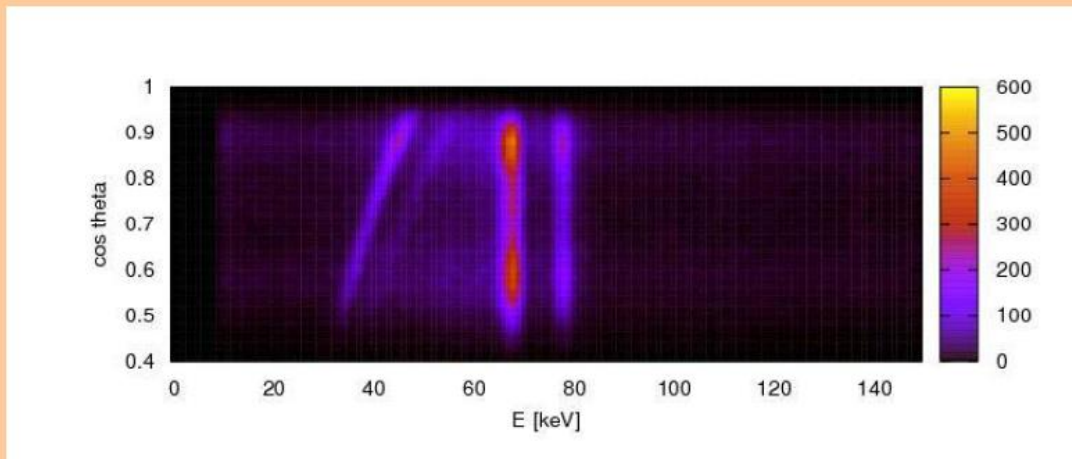
Z- determination with Kr beam



First Doppler correction with AGATA



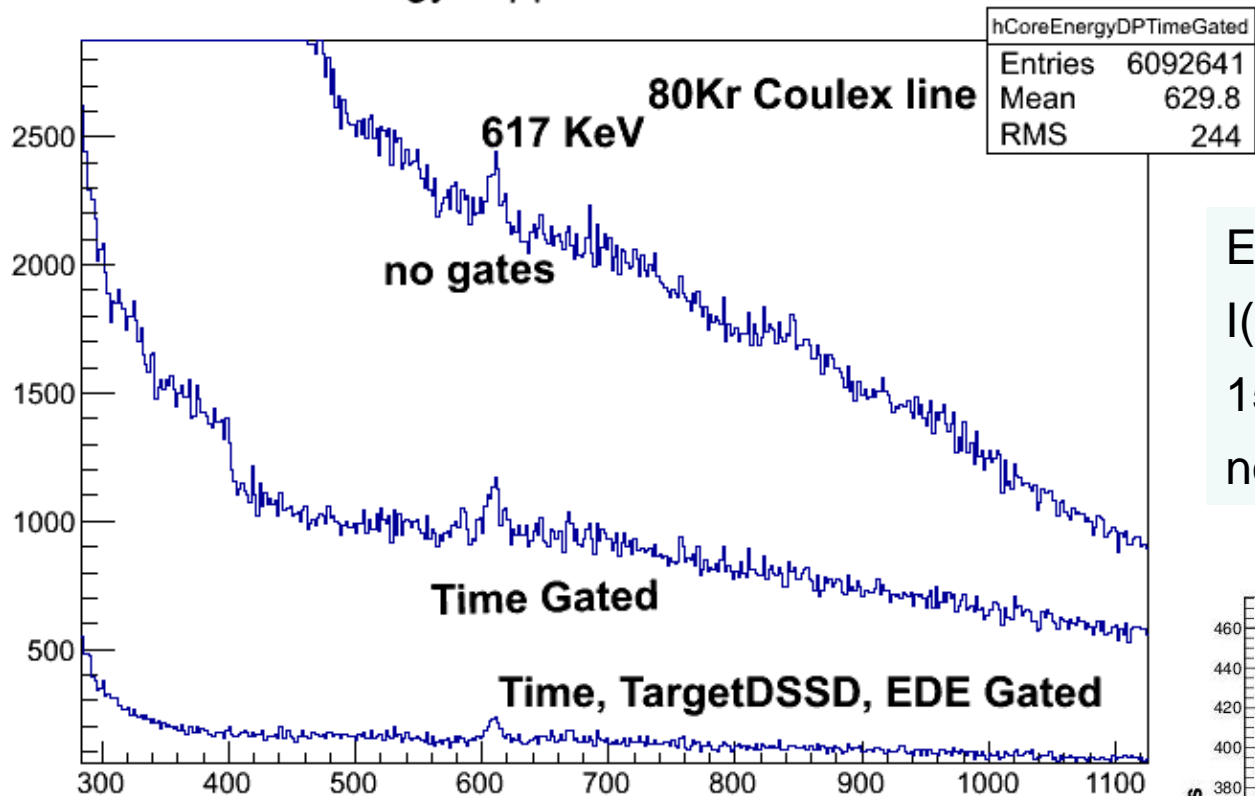
**atomic
excitations**



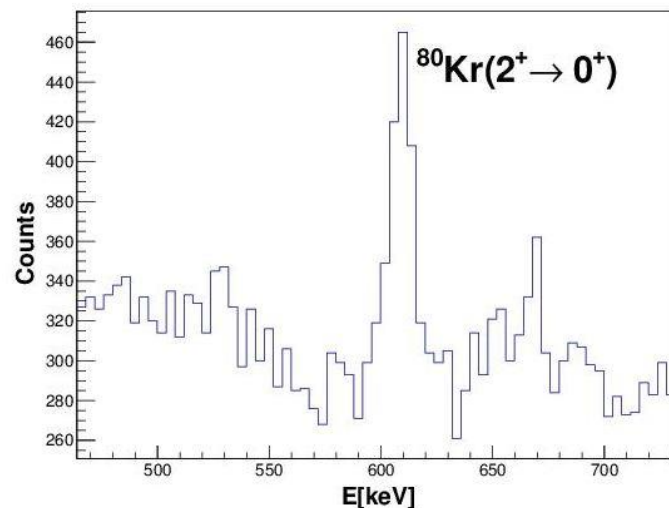
hit pattern of AGATA detectors

^{80}Kr Coulomb Excitation

CoreEnergyDopplerCorrectedTimeGated

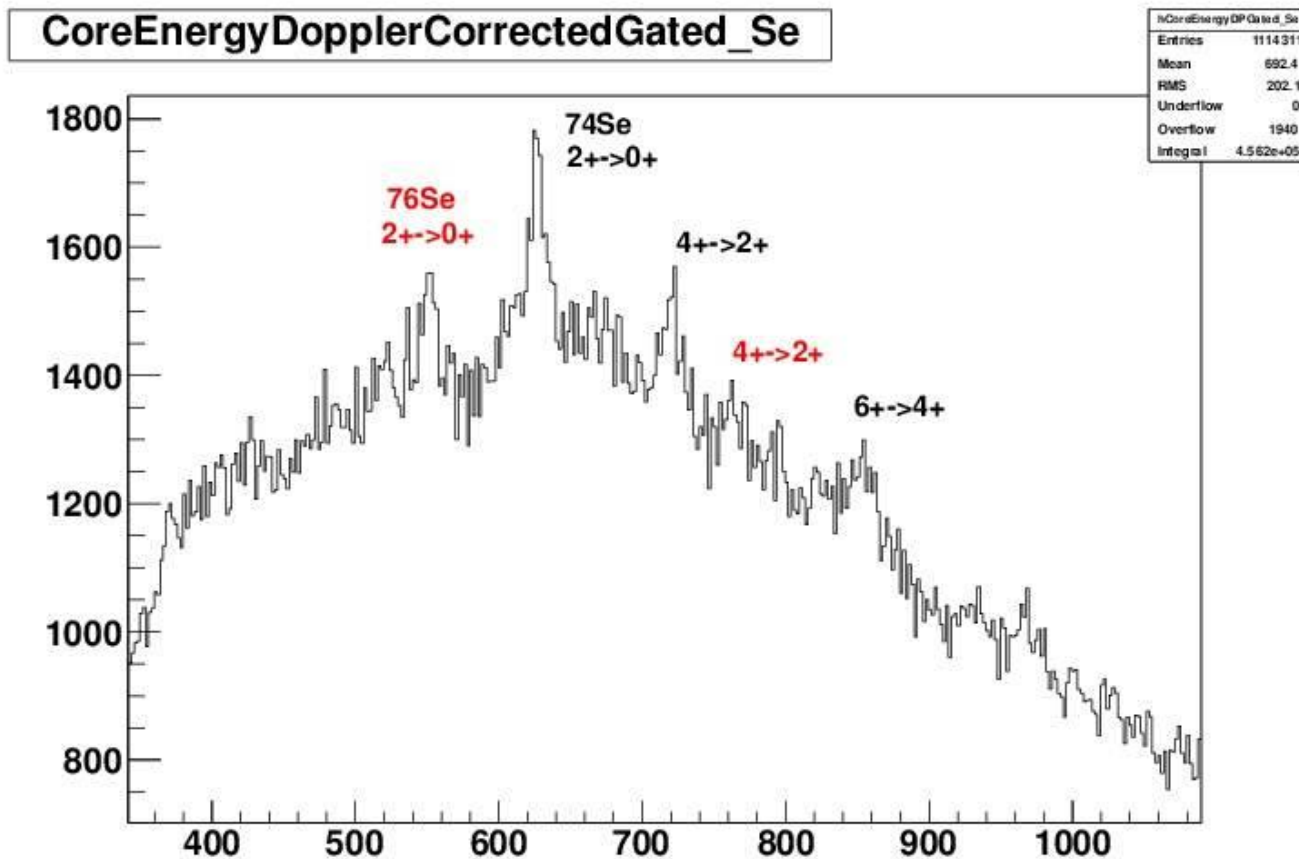


$E(\text{Kr}) = 150 \text{ AMeV}$
 $I(\text{max}) = 30 \text{ kHz}$
15 AGATA crystals
nominal target position



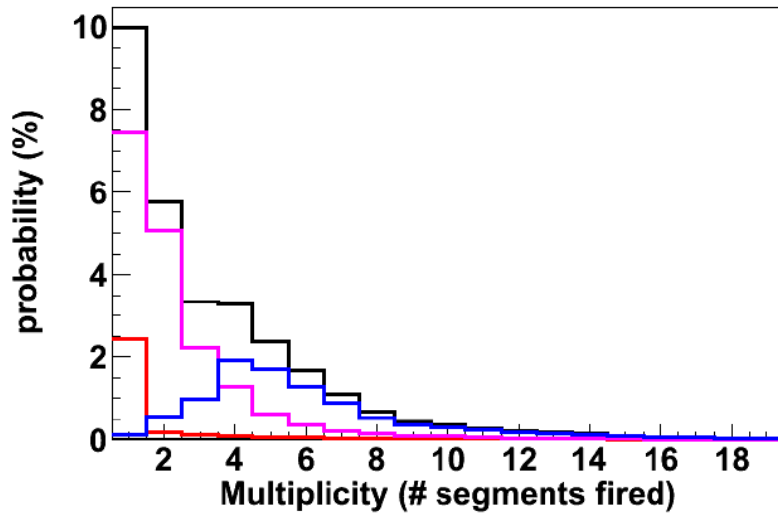
Preliminary analysis without full
AGATA tracking

^{80}Kr secondary fragmentation

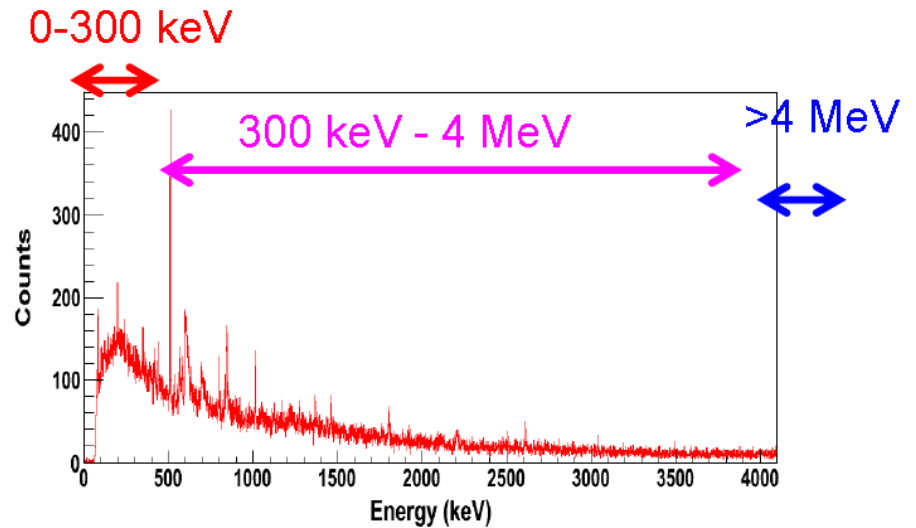
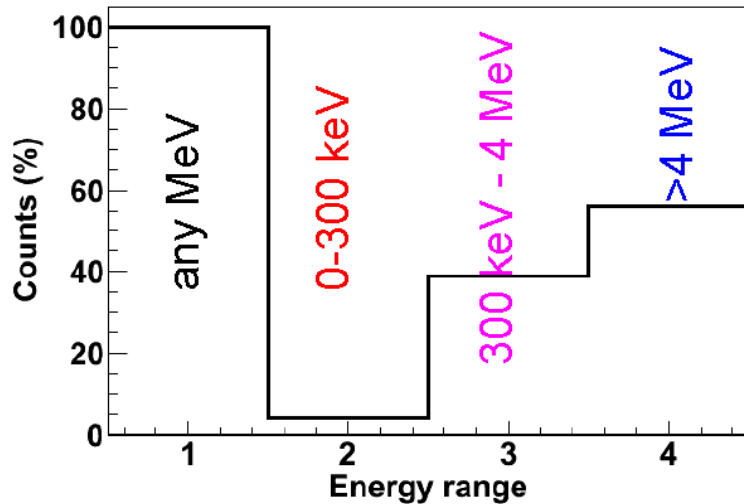
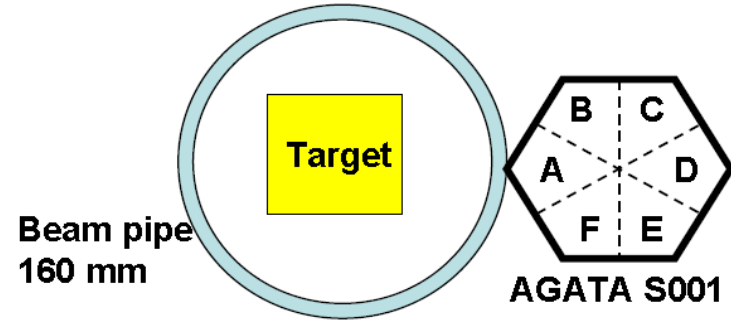


Preliminary analysis without
LYCCA mass gate

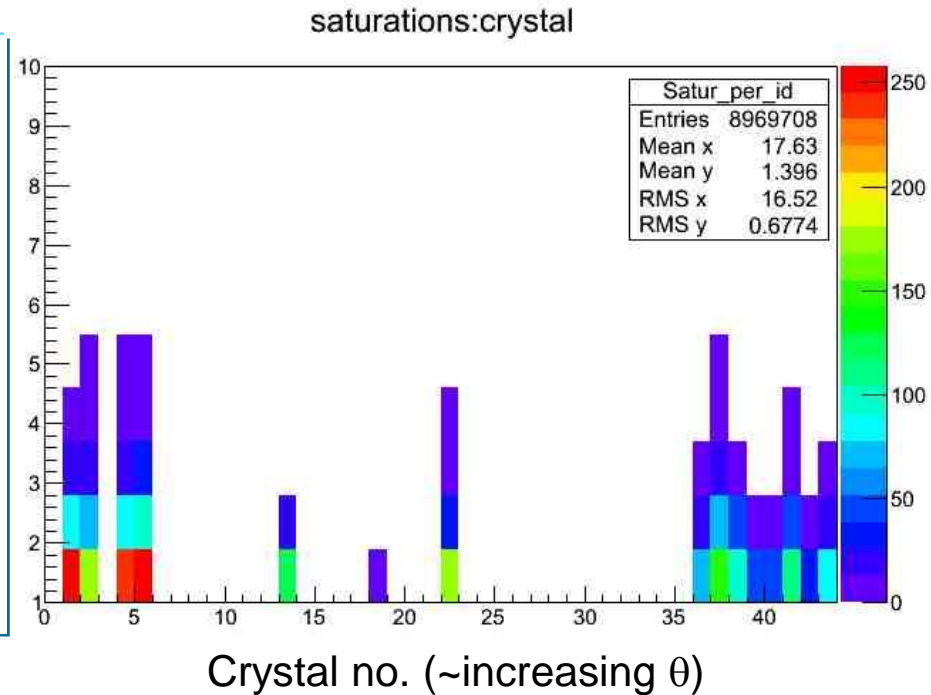
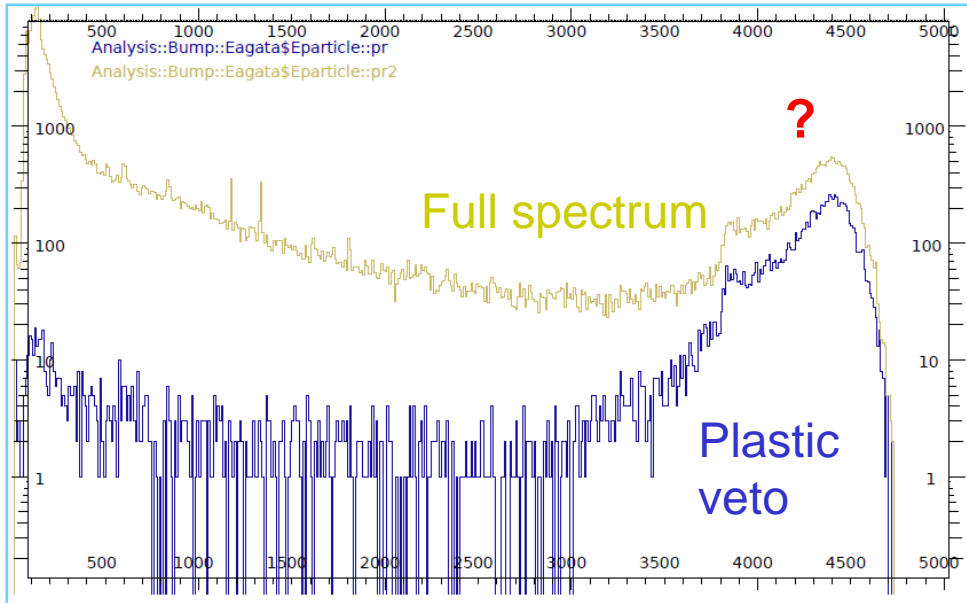
AGATA multiplicity distribution



- Cross-section view of the setup:



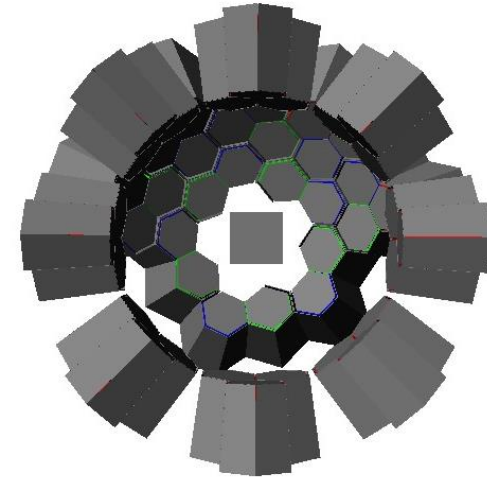
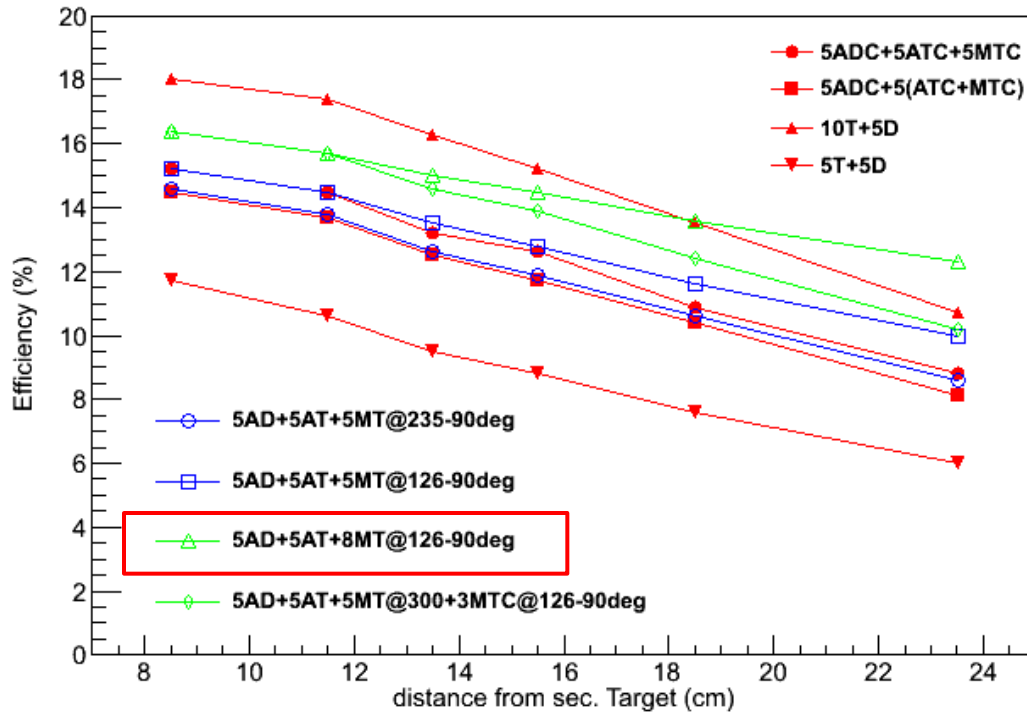
AGATA High-energy bump



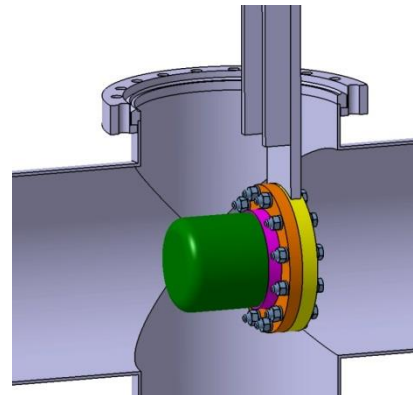
→ Protonen, $E_p \leq 120$ MeV

Campaign 2014

1. AGATA + MINIBALL



2. Hydrogen target



Conclusions

- NUSTAR follows an evolutionary approach and first experiments have already started
- NUSTAR instrumentation is fairly advanced and ready for FAIR
- PRESPEC-AGATA = HISPEC is the most complex nuclear spectroscopy experiment in the world
- The commissioning was successfully performed in 2012
- Gamma-tracking detectors boost the sensitivity for subtle nuclear structure effects by at one order of magnitude
- AGATA allows to detect and discriminate all kinds of background events
- First PRESPEC-AGATA experiments were performed in Fall 2012, more to come in 2014

A Great Collaboration

