



# V : WHERE WE STAND AND WHERE WE'RE HEADING TO


$\delta m_{12}^2$  

SOLARS+KAMLAND  
 $\delta m_{12}^2 = (7.9 \pm 0.7) 10^{-5} \text{ eV}^2$


$\theta_{12}$  

SOLARS+KAMLAND  
 $\sin^2(2\theta_{12}) = 0.82 \pm 0.055$

Addressed by accelerator neutrino experiments


$\delta m_{23}^2$  

ATMOSPHERICS  
 $\delta m_{23}^2 = (2.4 \pm 0.4) 10^3 \text{ eV}^2$

$\theta_{23}$  

ATMOSPHERICS  
 $\sin^2(2\theta_{23}) > 0.95$

$\theta_{13}$

  $\sin^2 2\theta_{13} = 0.1$

LSND/Steriles




$\delta_{CP}$



Mass hierarchy



$\Sigma m_\nu$  

BETA DECAY END POINT  
 $\Sigma m_\nu < 6.6 \text{ eV}$

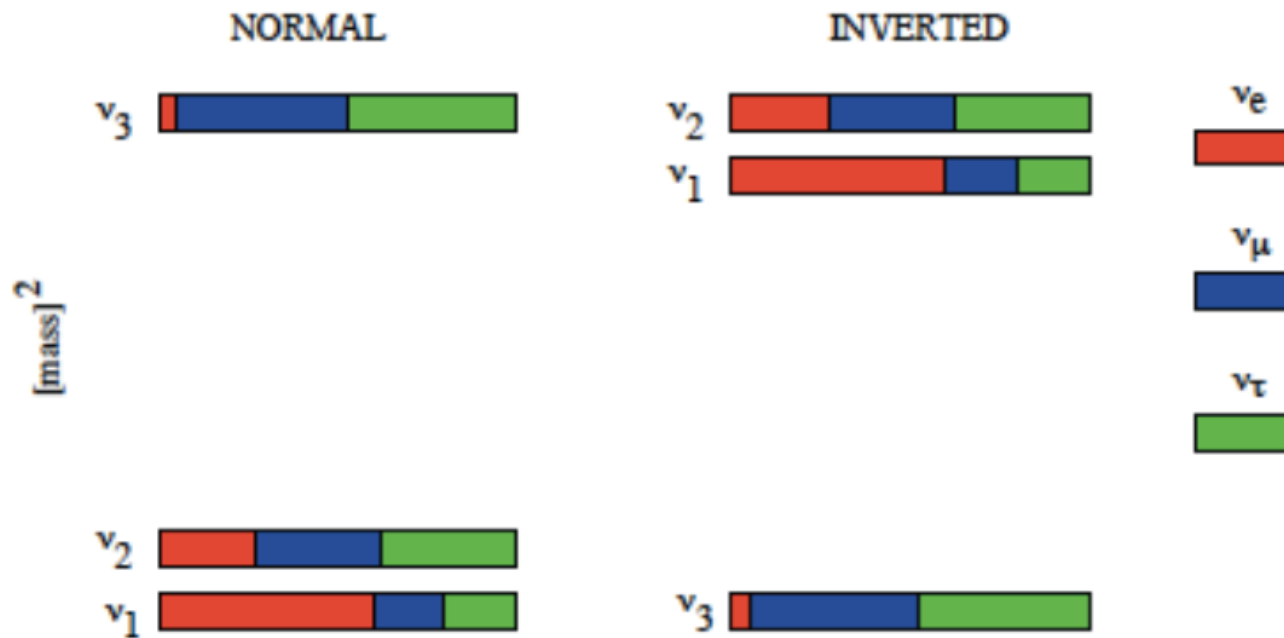


Dirac/Majorana



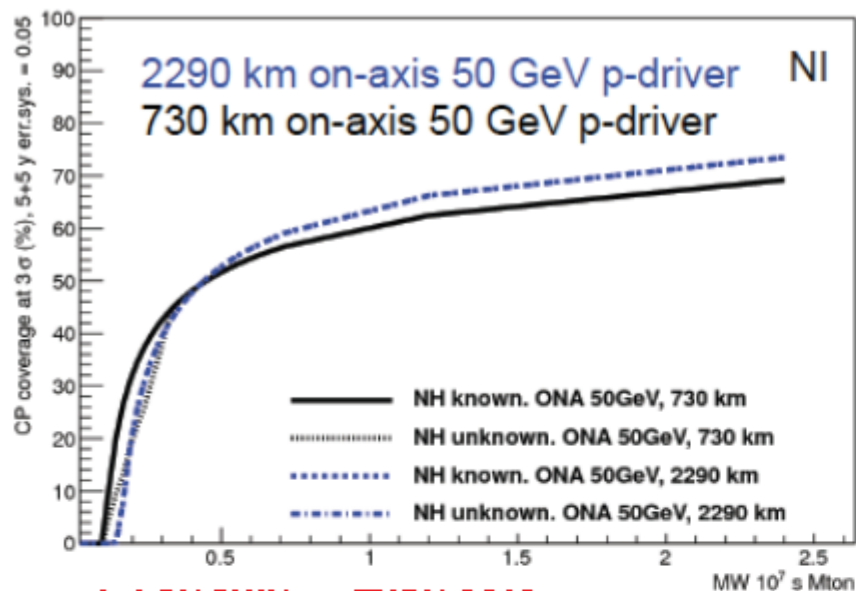
ACCORDING TO MY PERSONAL TASTE

**LARGE  $\theta_{13}$**   $\rightarrow$  NOT ONLY ACCELERATORS, BUT ALSO REACTOR AND ATMOSPHERIC NEUTRINOS CAN PLAY A ROLE IN THE  **$\nu$  MASS HIERARCHY GAME**



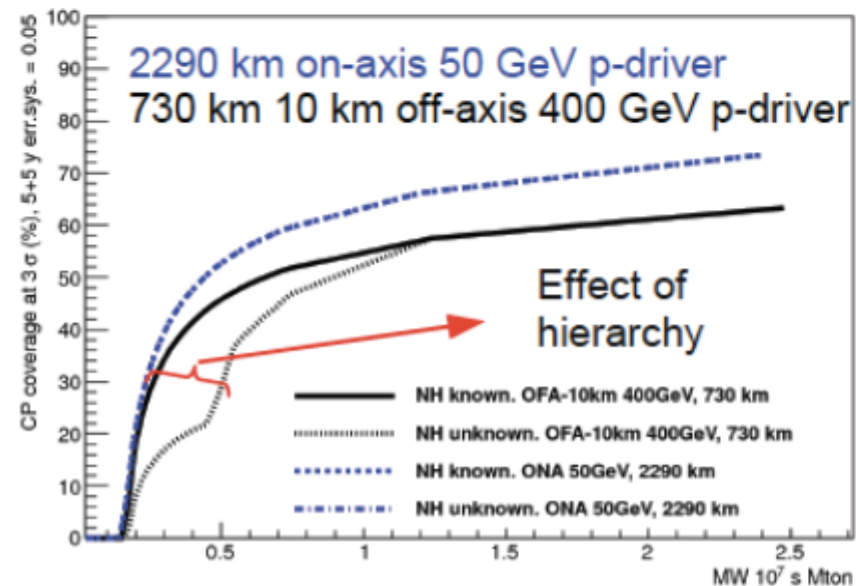
INO+NOVA+T2K, ICE-CUBE with PINGU, DAYABAY, future Long BaseLine

CP coverage at  $3\sigma$  (%), 5+5 y err.sys. = 0.05



A. LONGHIN, nuTURN 2012

CP coverage at  $3\sigma$  (%), 5+5 y err.sys. = 0.05



LAGUNA-LBNO observatory at Pyhasalmi (-1400m.)

2x50 kton LAr + 50 kton LSc  
879'000 m<sup>3</sup> excavation  
Design to be finalized within  
LAGUNA-LBNO by  $\approx$ 2014

Nominal beam power scenarios (700kW).

For  $\sin^2 2\theta_{13}=0.1$ , approximately  
(at 90%C.L.):

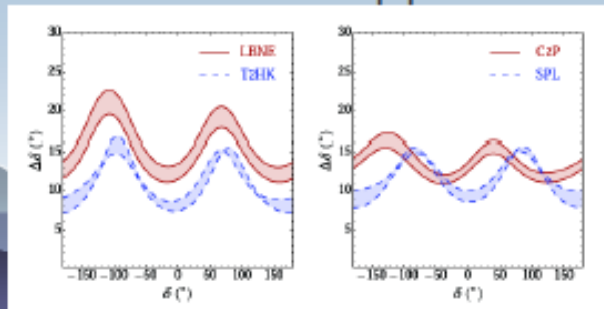
- MH: 100% coverage at  $>5\sigma$  in a few years of running
- CPV:  $\approx$ 60% coverage and evidence for maximal CP ( $\pi/2, 3\pi/2$ ) at  $2.9\sigma$  in 10 years

# Long baseline projects

| Project       | Beam power MW | Fiducial Mass kt | Baseline km | MH        | CPV 90%CL, ( $3\sigma$ ) | Physics starts | Astrophysical program |
|---------------|---------------|------------------|-------------|-----------|--------------------------|----------------|-----------------------|
| LBNO          | 0.8           | 20-<br>>100      | 2300        | Excellent | 71 (44)                  | 2023           | Yes                   |
| T2HK          | 0.75          | 500              | 295         | Little    | 86 (74)*                 | 2023           | Yes                   |
| LBNE          | 0.7           | 10               | 1300        | OK        | 69 (43)                  | 2022           | No                    |
| Lund          | 5             | 440              | 365         | Some      | 86 (70)                  | >2019          | Yes                   |
| CERN-Canfranc | 0.8-4         | 440              | 650         | Some      | 80-88(80)                | >2020          | Yes                   |

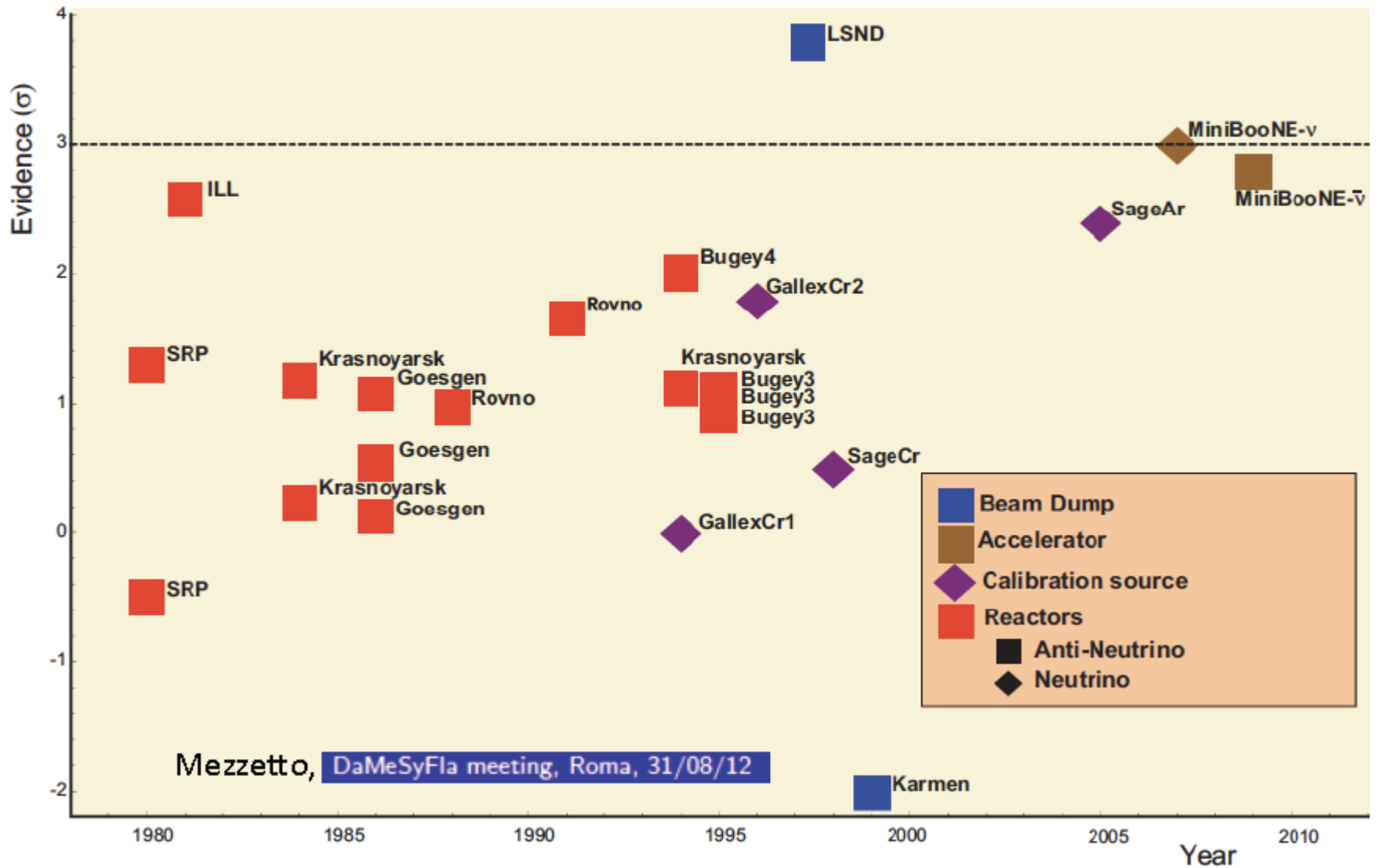
P. Coloma et al. hep-ph:1203.5651

\*: if mass hierarchy is known



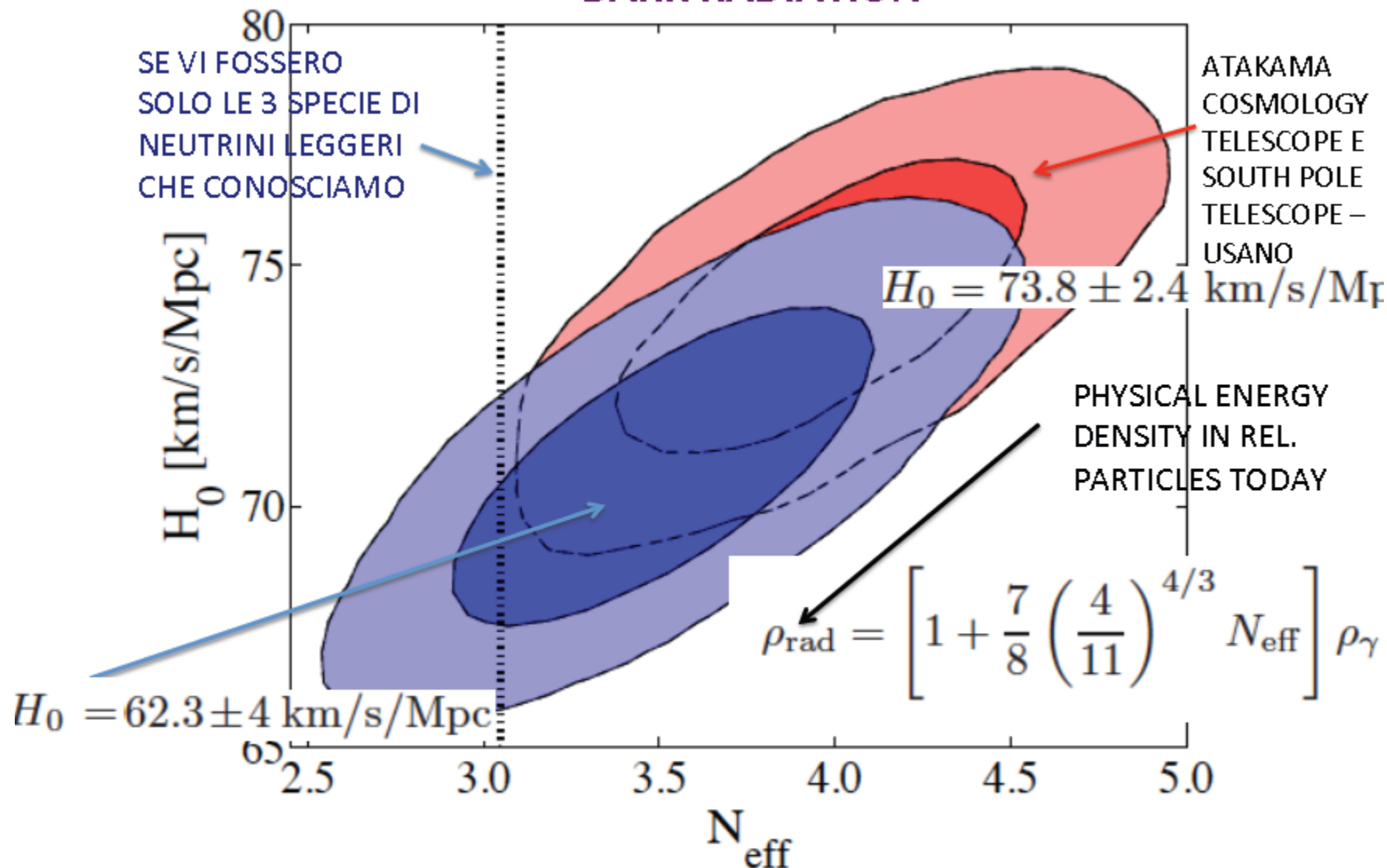
T2HK: 4MW, 500 kt  
 LBNE: 0.8 MW, 33 kt  
 C2P=LBNO : 0.8 MW, 100 kt

# A long standing set of anomalies



# INDICAZIONI DALLA COSMOLOGIA A FAVORE DI > 3 SPECIE DI NEUTRINI?

## “DARK RADIATION”



- The scientific impact of the search for sterile neutrinos has been highlighted: quoting Pilar Hernandez:

*If there are other light sterile states **we must know...***

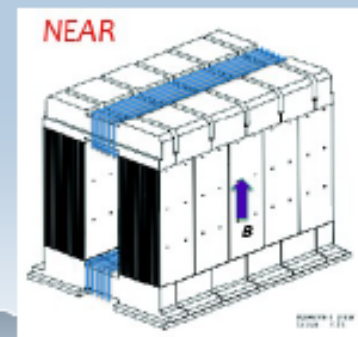
*Our predictions/constraints on*

- 1) matter-antimatter asymmetry*
- 2) large-scale structure, CMB*
- 3) nucleosynthesis*
- 4) supernova explosions*
- 5) the dark matter content of the Universe*
- 4) supernova explosions*
- 6) rate of neutrinoless double beta decay*

*depend on it !*

## *SPSC-P-347 (Icarus-Nessie)*

- ◆ Proposal (SPSC-P-347, 150 authors) of a comprehensive search for new neutrino states around  $\Delta m^2 \sim 1\text{eV}^2$  using a SPS 110 GeV proton beam in the NA
- ◆ with two LAr detectors, at 1600 m (ICARUS T600 now at Gran Sasso) and 300m (T150), supplemented by two spectrometers
- ◆ Method : two identical detectors, with imaging properties and complete final state reconstruction

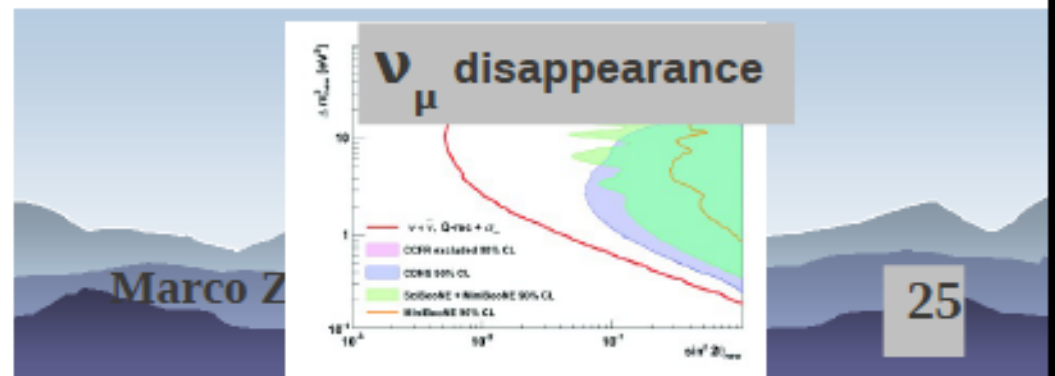
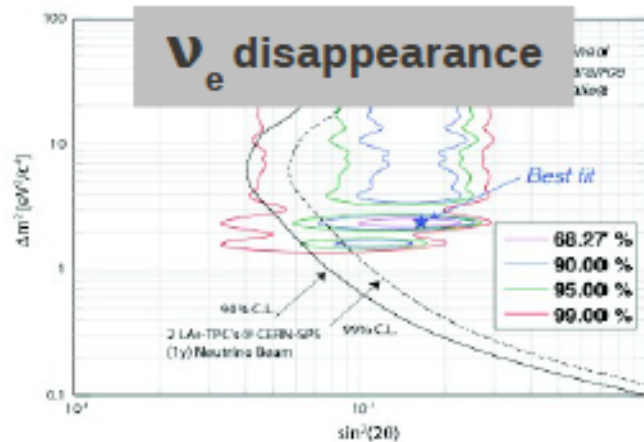
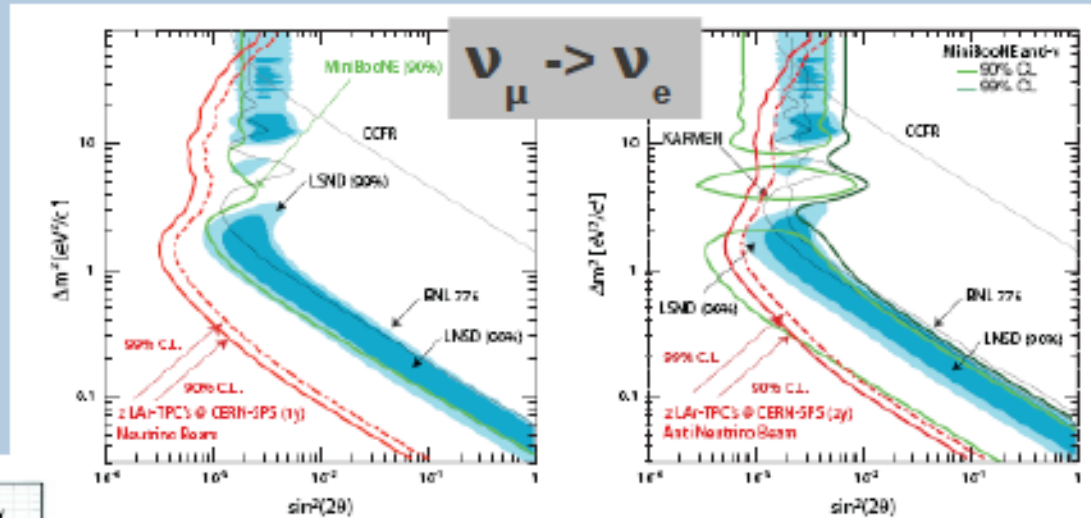




# SPSC-P-347

With two years negative and one year positive focusing, the following channels can be studied in neutrino and antineutrino modes. Very complete coverage of the region of the anomalies

- $\nu_{\mu} \rightarrow \nu_e$
- $\nu_e$  disappearance
- $\nu_{\mu}$  disappearance



# Summary of the $\nu$ session

- $\nu$  mass and mixings confirmed by many experiments and remain, with dark matter, the only present evidence of beyond the Standard Model physics.
- As the highest priority we should determine the unknown oscillation parameters and look for surprises. CP violation and the  $\nu$  mass hierarchy could be keys to the matter/antimatter asymmetry of the Universe.
- A large and effective European community exists in this area.
- Long baselines are optimal for determining the mass hierarchy, real advantage of the CERN  $\rightarrow$  Pyhäsalmi baseline and, to a lesser extent, LBNE.
- The CERN  $\rightarrow$  Pyhäsalmi baseline is also near optimal for a Neutrino Factory.
- Shorter ( $\sim$ hundreds of kilometres) baselines with huge detectors would allow very high statistics measurements more helpful for CP violation, particularly if hierarchy is known. This is the case of T2HK (also European alternatives such as CERN  $\rightarrow$  Frejus, CERN  $\rightarrow$  Canfranc, or ESS-based  $\nu$  beam)
- For best performance and synergy an experiment of each category is needed  $\rightarrow$  Coherence with efforts in other regions. Coordination and cooperation with our international colleagues mandatory.
- Anomalies in a range of phenomena at lower energies perhaps point to sterile neutrinos, and a proposed experiment at CERN would be highly competitive.
- More sophisticated future projects, which EUROnu has concluded should be a Neutrino Factory, necessary to achieve the desired sensitivity to the CP phase and probe new physics.
- R&D including projects such as MICE and nuStorm (which may also offer a definitive test for sterile neutrinos) should be supported.
- Experiments in absolute neutrino mass, especially in neutrinoless double-beta decay, are also a top priority.
- Hadron production, neutrino cross-section, and other support measurements will be essential to reach the neutrino oscillation sensitivity goals.

## My general comments for discussion II

- **Small scale experiments** should still be possible for some precision measurements. But, may be not for long.....
- **$e^+e^-$  colliders** can now make a concrete physics programme on precision physics with “Higgs”-like particle and top (and some more)
- Concrete physics cases for **both long and short baseline neutrino programme** can be made. In principle, many of the technology seems to exist for those (i.e. **more D than R** needed). Long baseline facilities might be a concrete example of coordination between PP and APP.
- Projects to deepen our knowledge in Standard Model are being proposed
- Lets start the discussion, if anytime left....