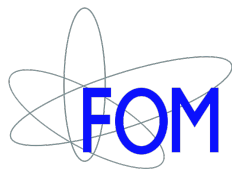


Test of Lorentz invariance in the weak decay of polarized atoms

Stefan E. Müller, E. Dijck, S. Hoekstra, J. Noordmans,
G. Onderwater, L. Willmann, H. Wilschut,
R. Timmermans, K. Yai*



KVI, University of Groningen/
* Osaka University

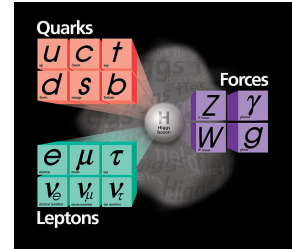


12th Meeting of the Working Group on Rad. Corrections and MC
Generators for Low Energies

Mainz - September 27-28, 2012

Lorentz symmetry is a fundamental basis of

- ▶ **the theory of Special Relativity**
- ▶ **the Standard Model of Particle Physics**



Connection to General Relativity and CPT symmetry

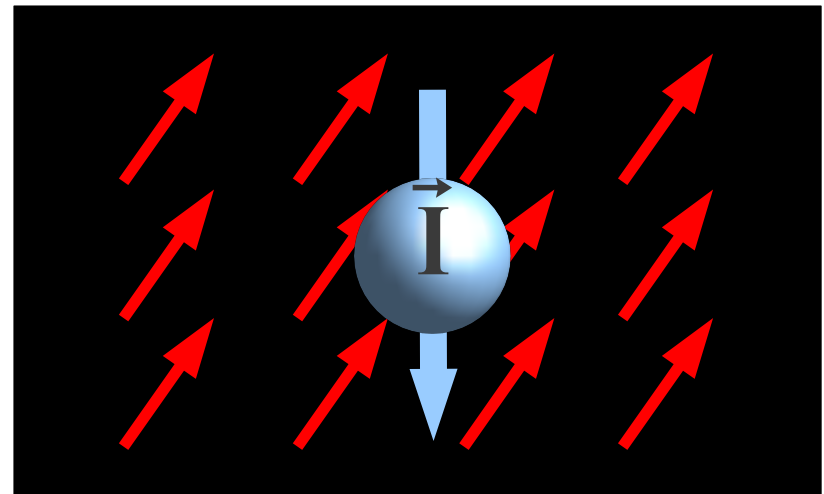
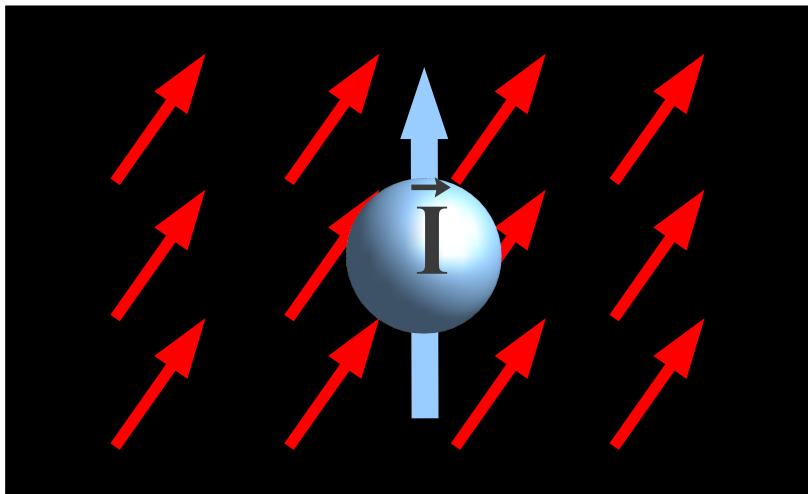
- ▶ **Lorentz symmetry breaking (LSB)**
 - Lorentz Symmetry spontaneously broken in Quantum Gravity models
 - “hidden” background fields → preferred direction
 - precision experiments can look for signatures of LSB
- ▶ **Many experimental tests, no evidence of LSB**
(mainly QED tests and gravity experiments)

**Weak decay sector essentially
unexplored!**

Lorentz Symmetry Breaking

2

- assume nuclei interact with Lorentz-violating background fields

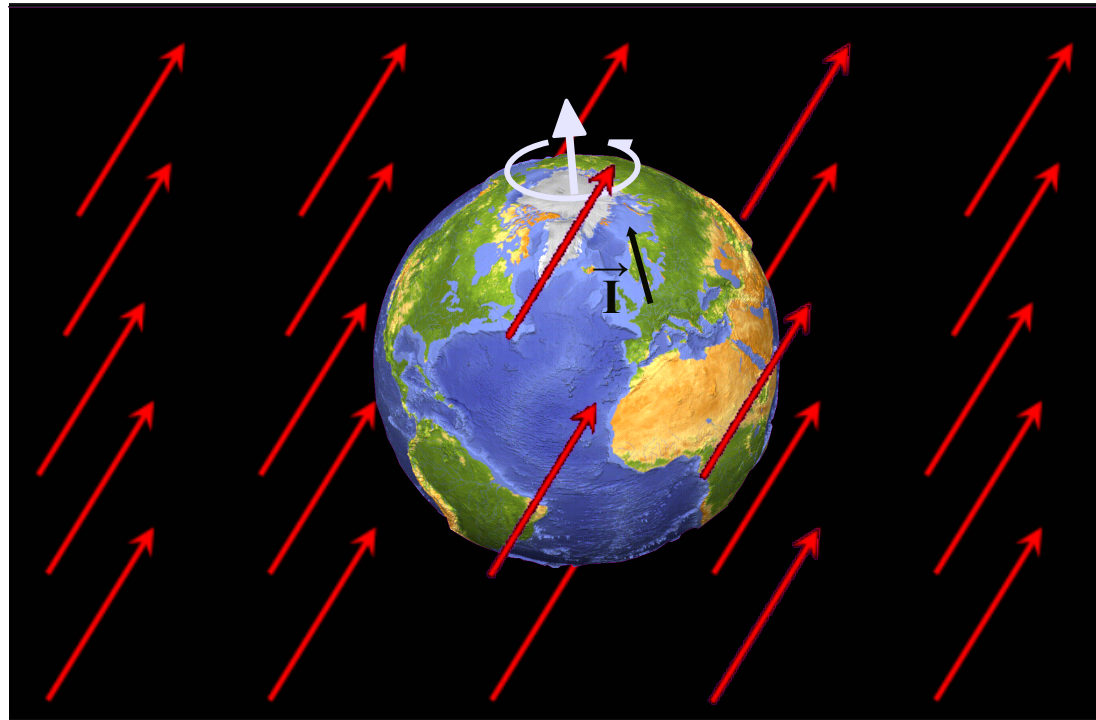


What is the change in the decay rate if the orientation of spin changes with respect to background fields?

Lorentz Symmetry Breaking

2

- assume nuclei interact with Lorentz-violating background fields

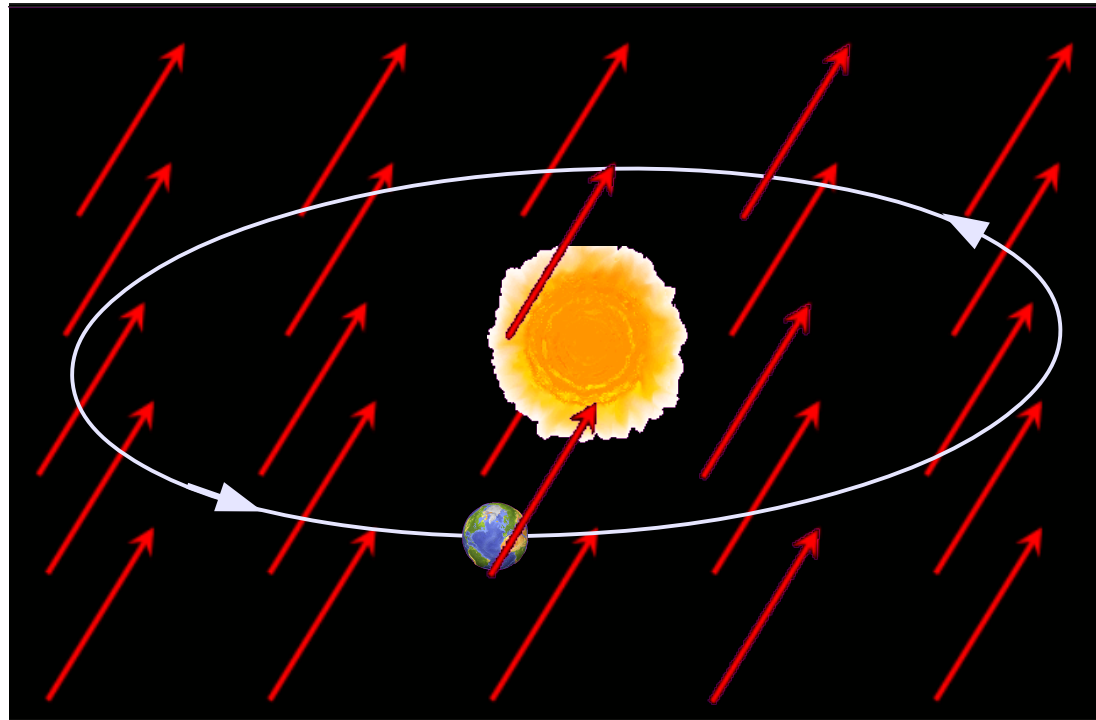


What is the change in the decay rate if the orientation of spin changes with respect to background fields?
- search for variations induced by **daily**, yearly or “deliberate” reorientation of spin

Lorentz Symmetry Breaking

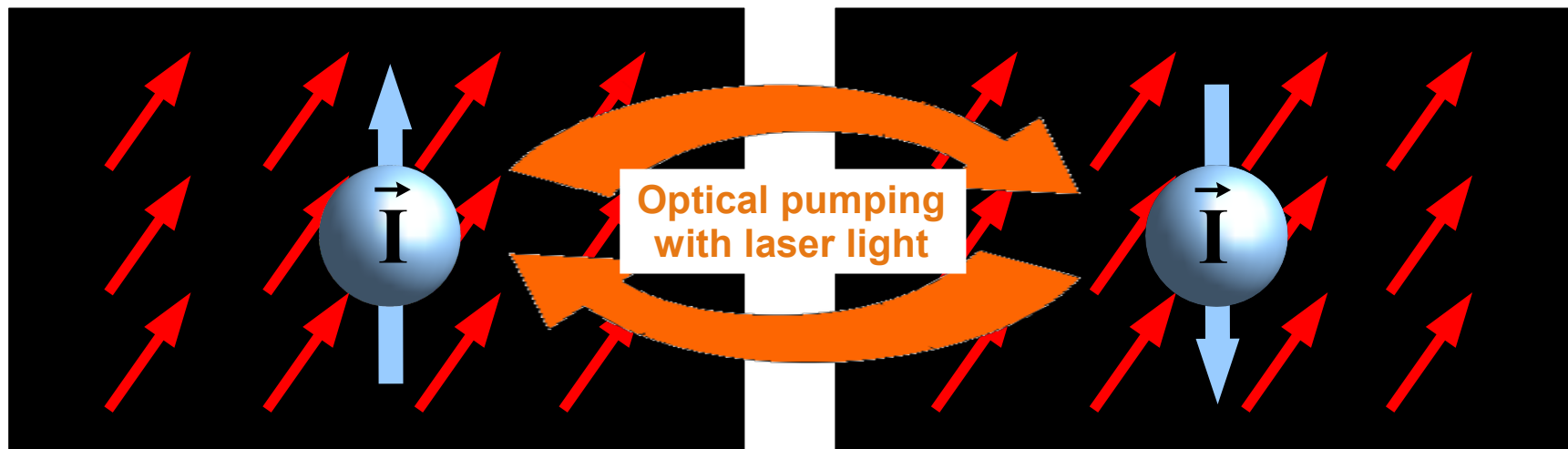
2

- assume nuclei interact with Lorentz-violating background fields



What is the change in the decay rate if the orientation of spin changes with respect to background fields?
- search for variations induced by daily, **yearly** or “deliberate” reorientation of spin

- assume nuclei interact with Lorentz-violating background fields

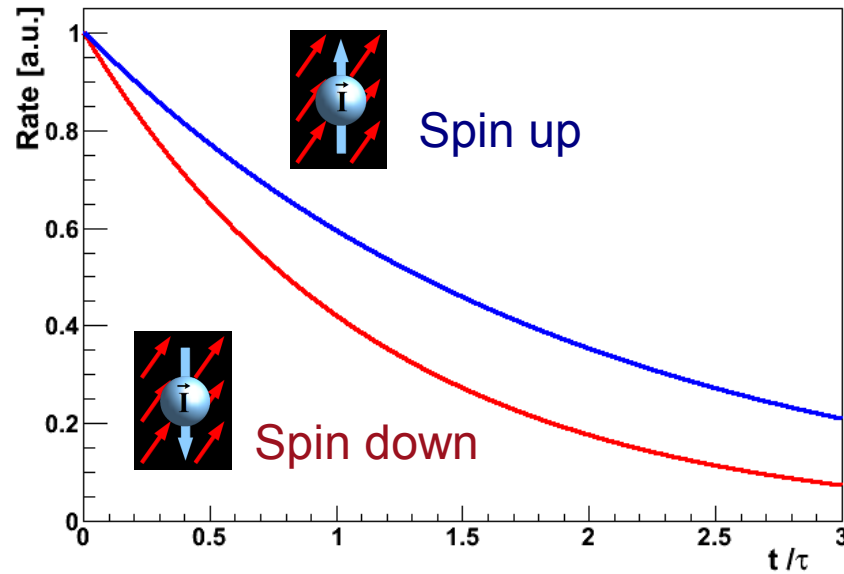


What is the change in the decay rate if the orientation of spin changes with respect to background fields?
- search for variations induced by daily, yearly or **“deliberate”** reorientation of spin

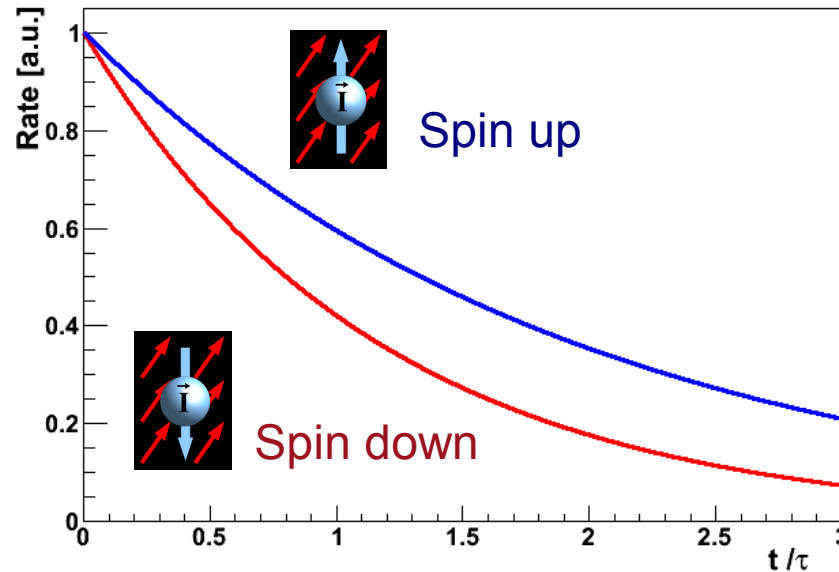
Experiment:

3

- **Change in decay rate** for different polarization orientations:



- **Change in decay rate** for different polarization orientations:



$$\frac{d\Gamma}{dE d\Omega} \sim \underbrace{\left(1 + A_0 \frac{\langle \vec{I} \rangle}{I} \cdot \frac{\vec{p}}{E} \right)}_{\text{SM}} + \xi_1 \left(1 + \xi_A \left(\hat{p} \cdot \frac{\langle \vec{I} \rangle}{I} \right) \right) \hat{p} \hat{n} +$$

$$+ \xi_2 \frac{\langle \vec{I} \rangle}{I} \hat{n} + \xi_3 \hat{p}_i \left(\frac{\langle \vec{I} \rangle}{I} \right)_j \rho^{ij}$$

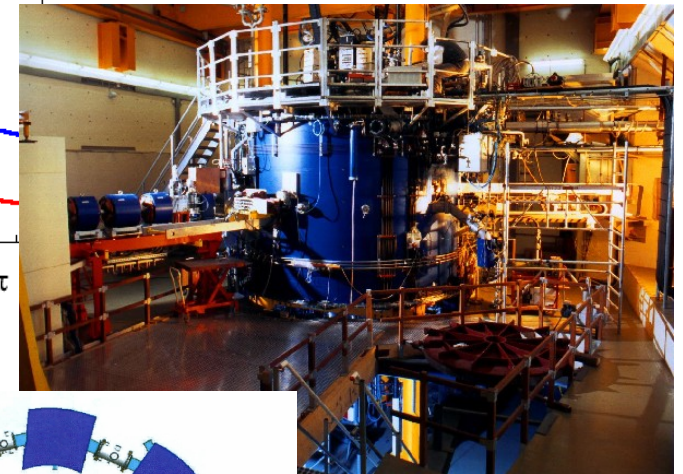
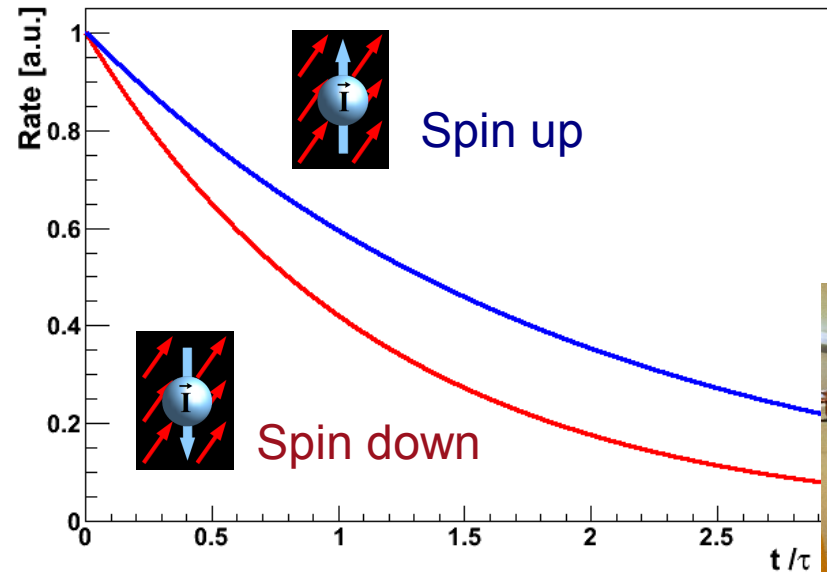
I = nuclear spin; p, E = electron momentum and energy

$\xi_{1,2,3,A}$ = coupling strength to LIV fields \hat{n}, ρ^{ij}

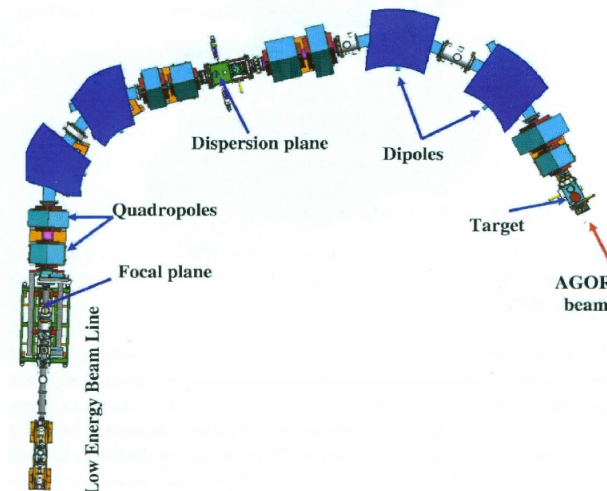
Experiment:

3

- **Change in decay rate for different polarization orientations:**



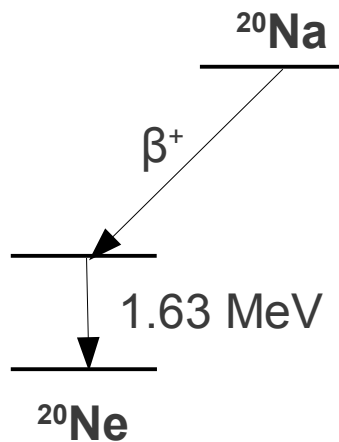
- **AGOR cyclotron at KVI**
Produce short-lived isotopes
- **TRIMUP isotope separator**
Clean isotope beam



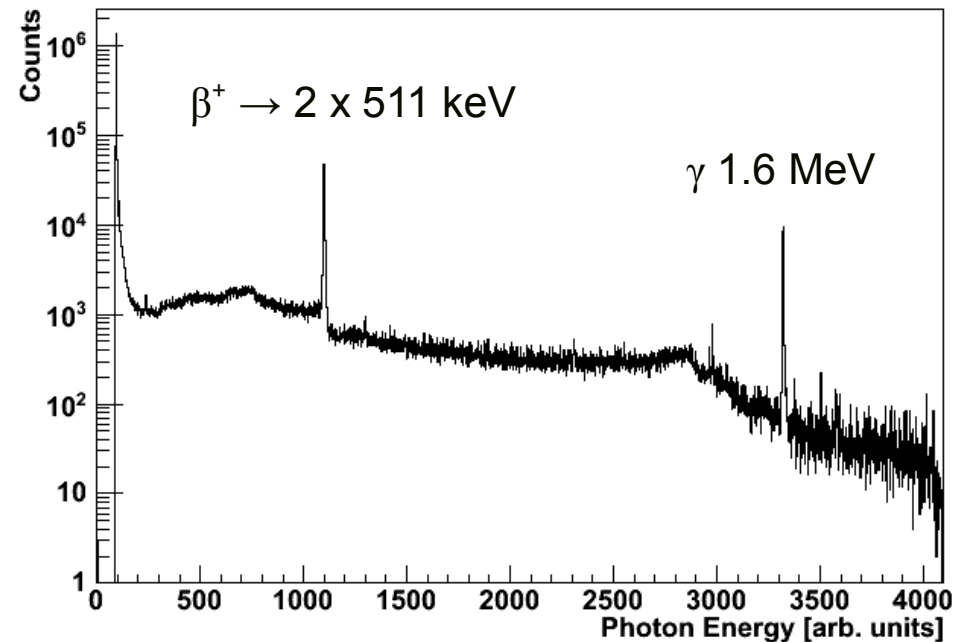
Choice of ^{20}Na :

- **Properties:** $2^+ \rightarrow 2^+$ (GT), β^+ , $\tau_{1/2} = 0.448\text{s}$, β -asymmetry parameter $A_0 = 1/3$
- **Produced** via $^{20}\text{Ne} + p \rightarrow ^{20}\text{Na} + n$ reaction: 10^6 decays/s
- **80% decay** to excited state of ^{20}Ne (1.63 MeV)

Level scheme



Gamma-energy spectrum

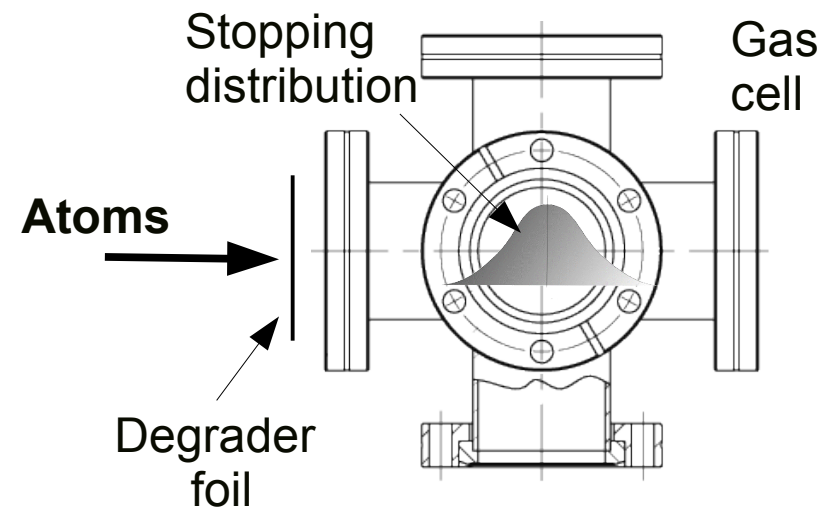


Experiment:

5

► Isotope beam stopped in buffer gas cell

- Aluminum foil degraders & buffer gas pressure (noble gas, 2atm)



Experiment:

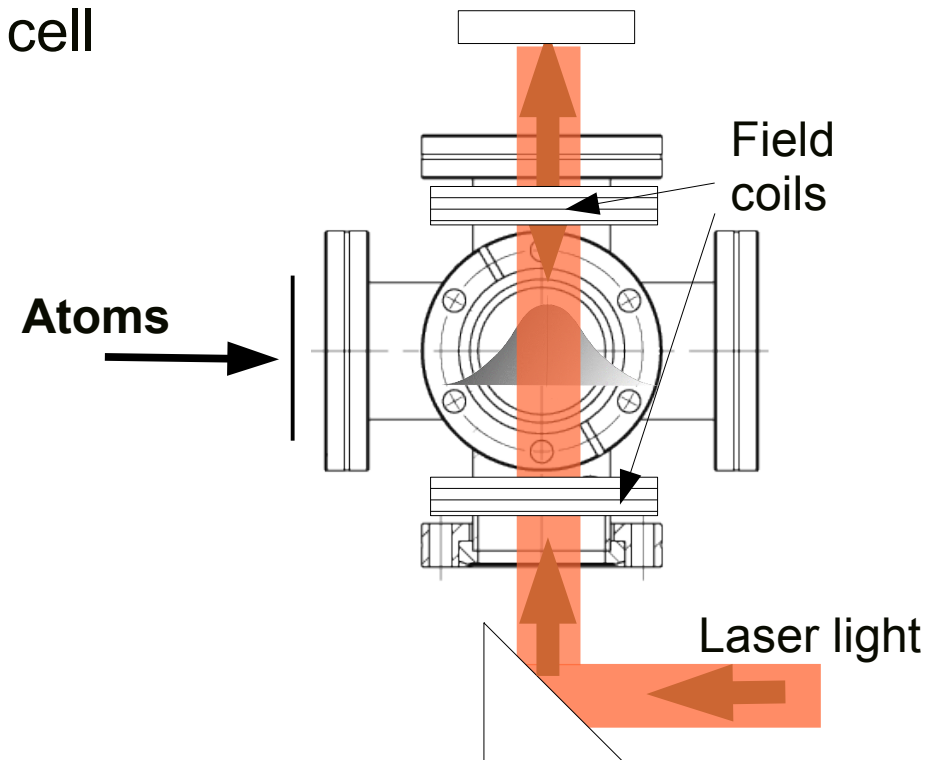
5

► **Isotope beam stopped** in buffer gas cell

- Aluminum foil degraders & buffer gas pressure (noble gas, 2atm)

► **Polarized nuclei** via optical pumping:

- magnetic holding field
- circularly polarized σ^\pm light



Experiment:

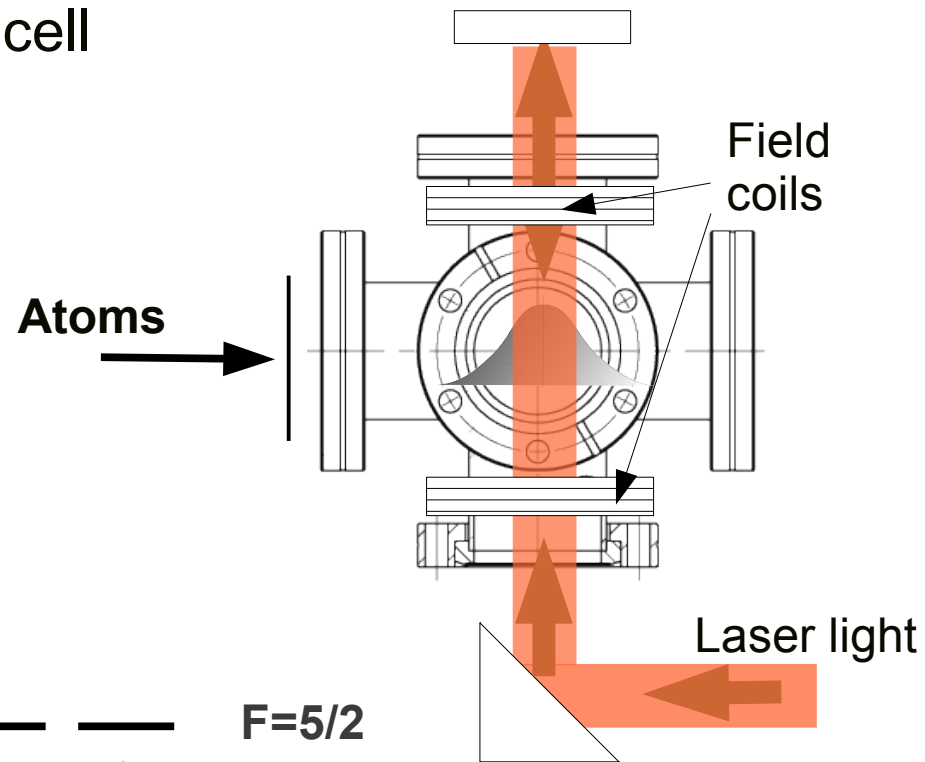
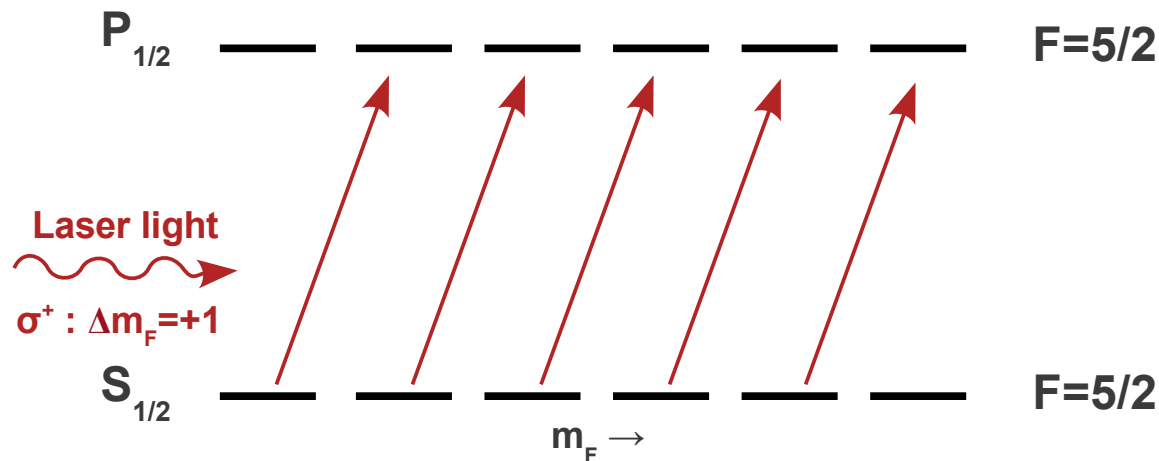
5

► **Isotope beam stopped** in buffer gas cell

- Aluminum foil degraders & buffer gas pressure (noble gas, 2atm)

► **Polarized nuclei** via optical pumping:

- magnetic holding field
- circularly polarized σ^\pm light



Experiment:

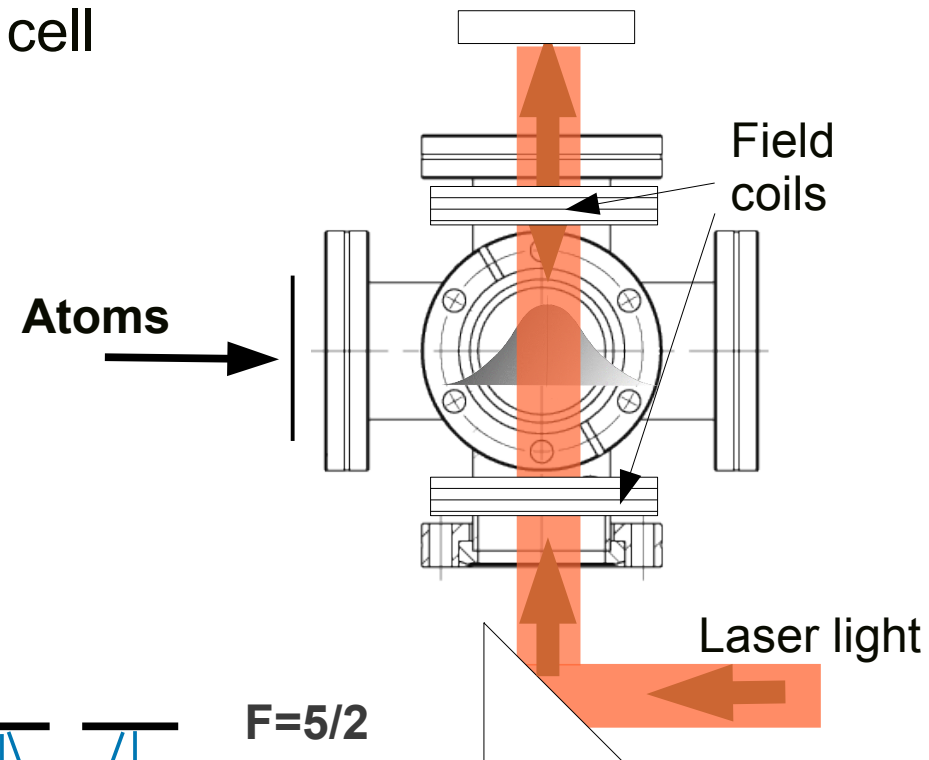
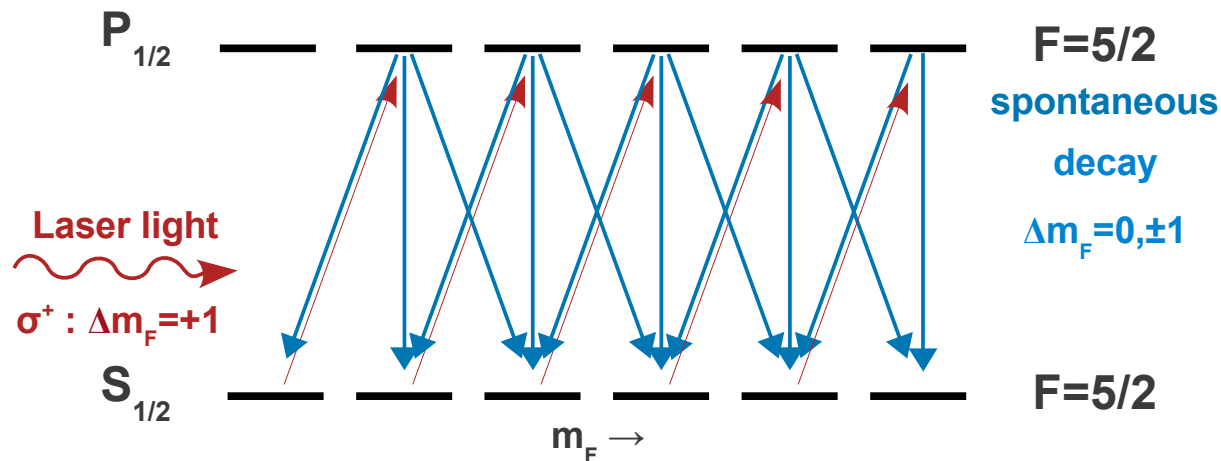
5

► **Isotope beam stopped** in buffer gas cell

- Aluminum foil degraders & buffer gas pressure (noble gas, 2atm)

► **Polarized nuclei** via optical pumping:

- magnetic holding field
- circularly polarized σ^\pm light



Experiment:

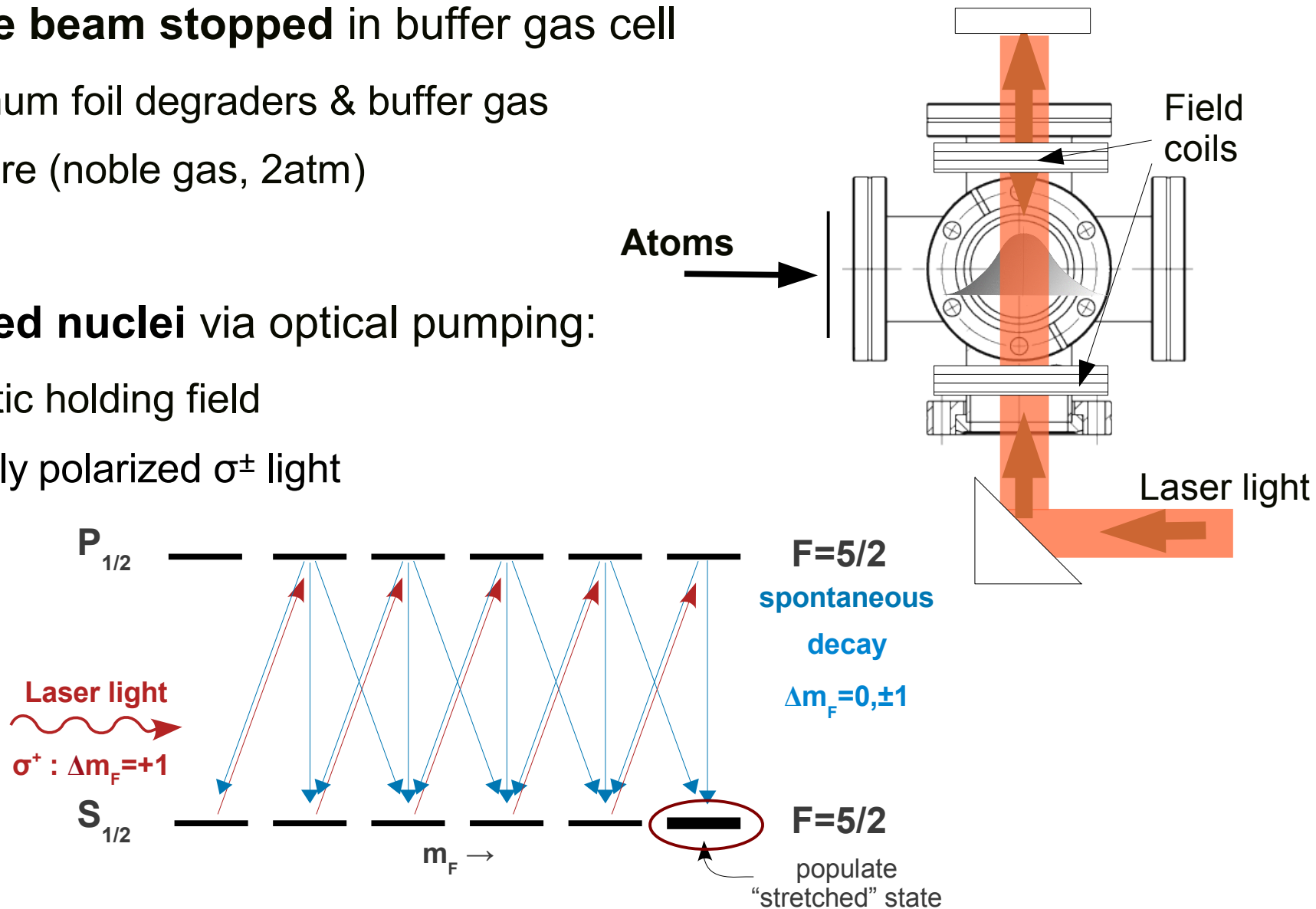
5

► **Isotope beam stopped** in buffer gas cell

- Aluminum foil degraders & buffer gas pressure (noble gas, 2atm)

► **Polarized nuclei** via optical pumping:

- magnetic holding field
- circularly polarized σ^\pm light



Experiment:

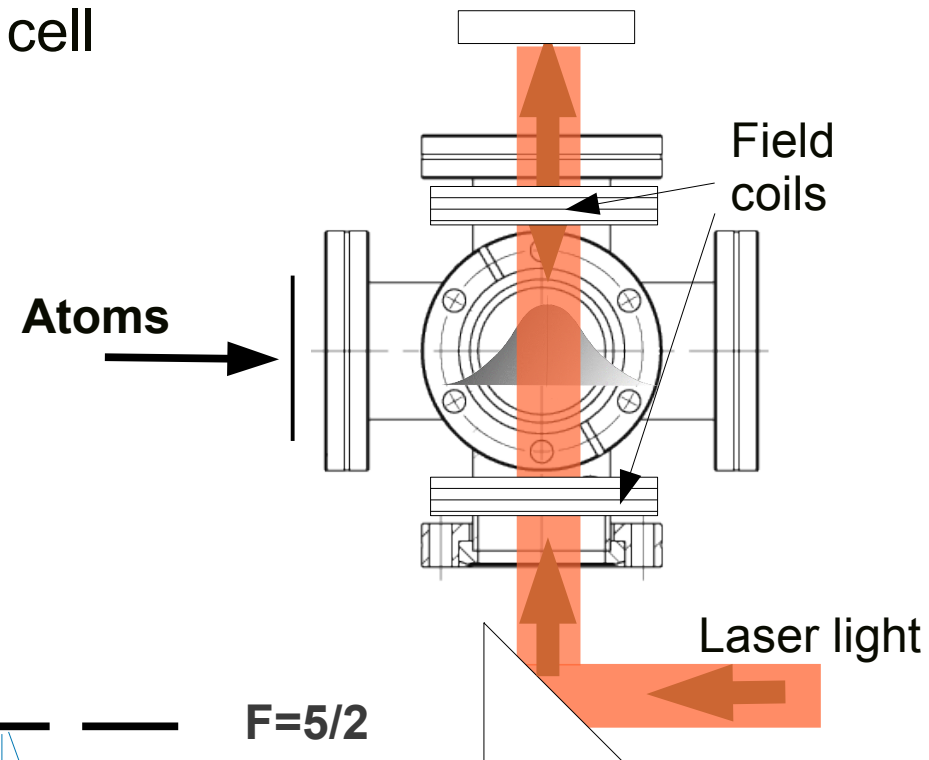
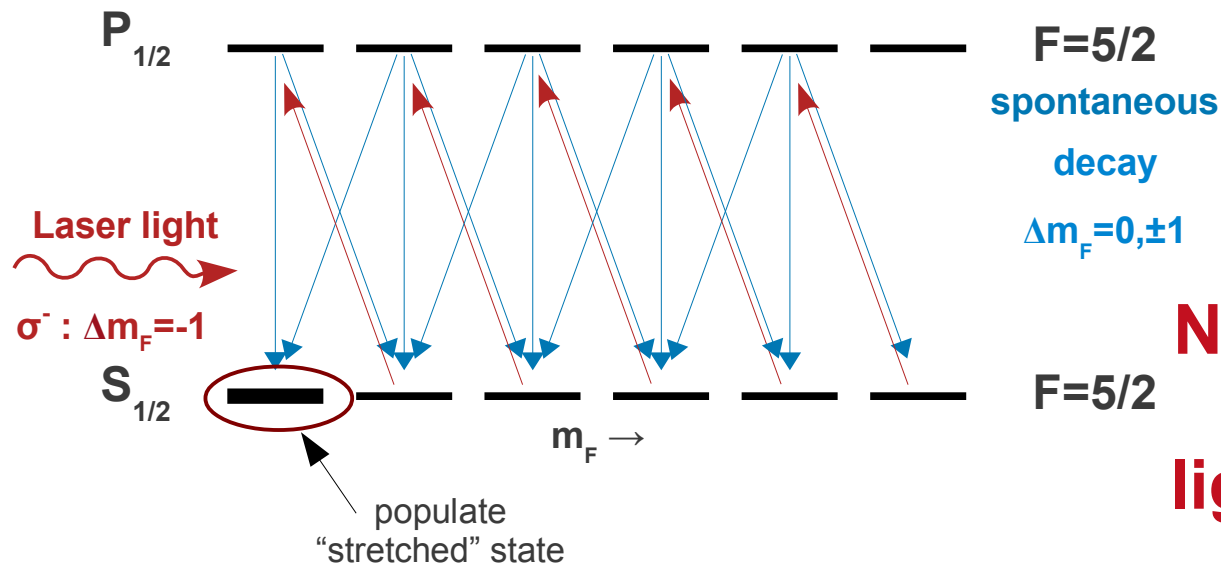
5

► **Isotope beam stopped** in buffer gas cell

- Aluminum foil degraders & buffer gas pressure (noble gas, 2atm)

► **Polarized nuclei** via optical pumping:

- magnetic holding field
- circularly polarized σ^\pm light



Nuclear spin follows light helicity

Experiment:

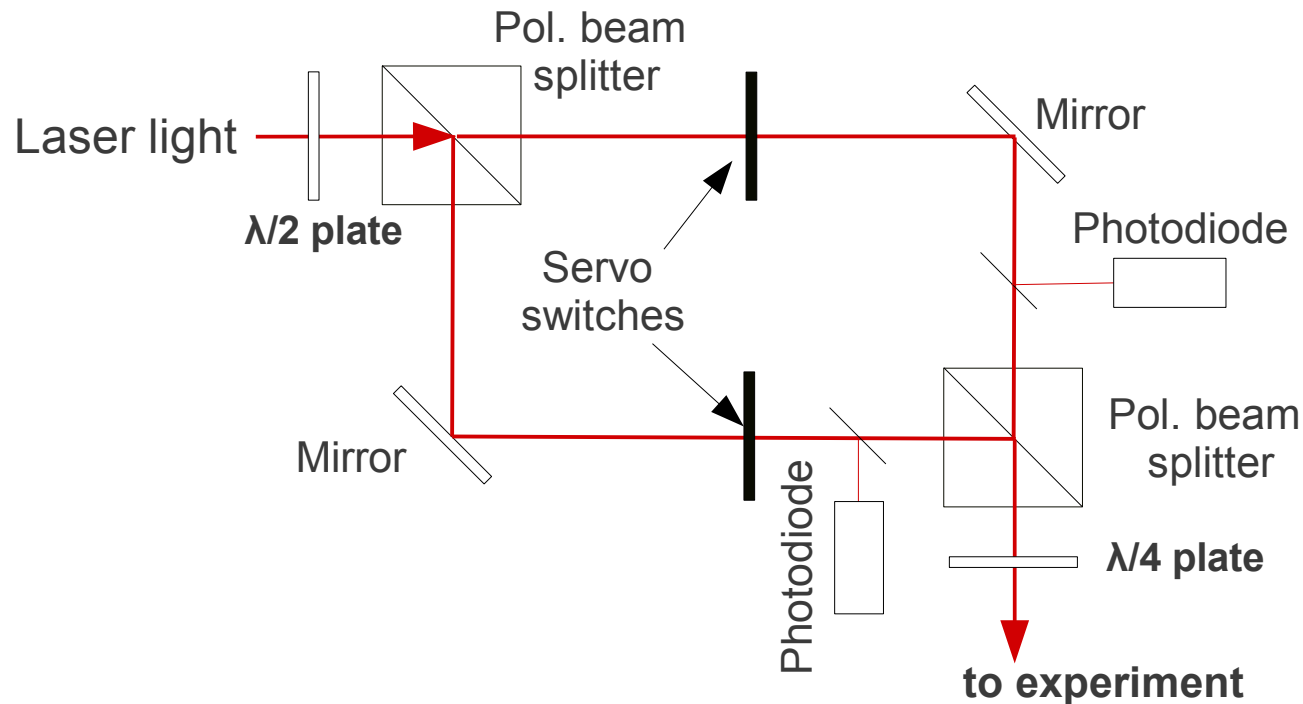
5

► **Isotope beam stopped** in buffer gas cell

- Aluminum foil degraders & buffer gas pressure (noble gas, 2atm)

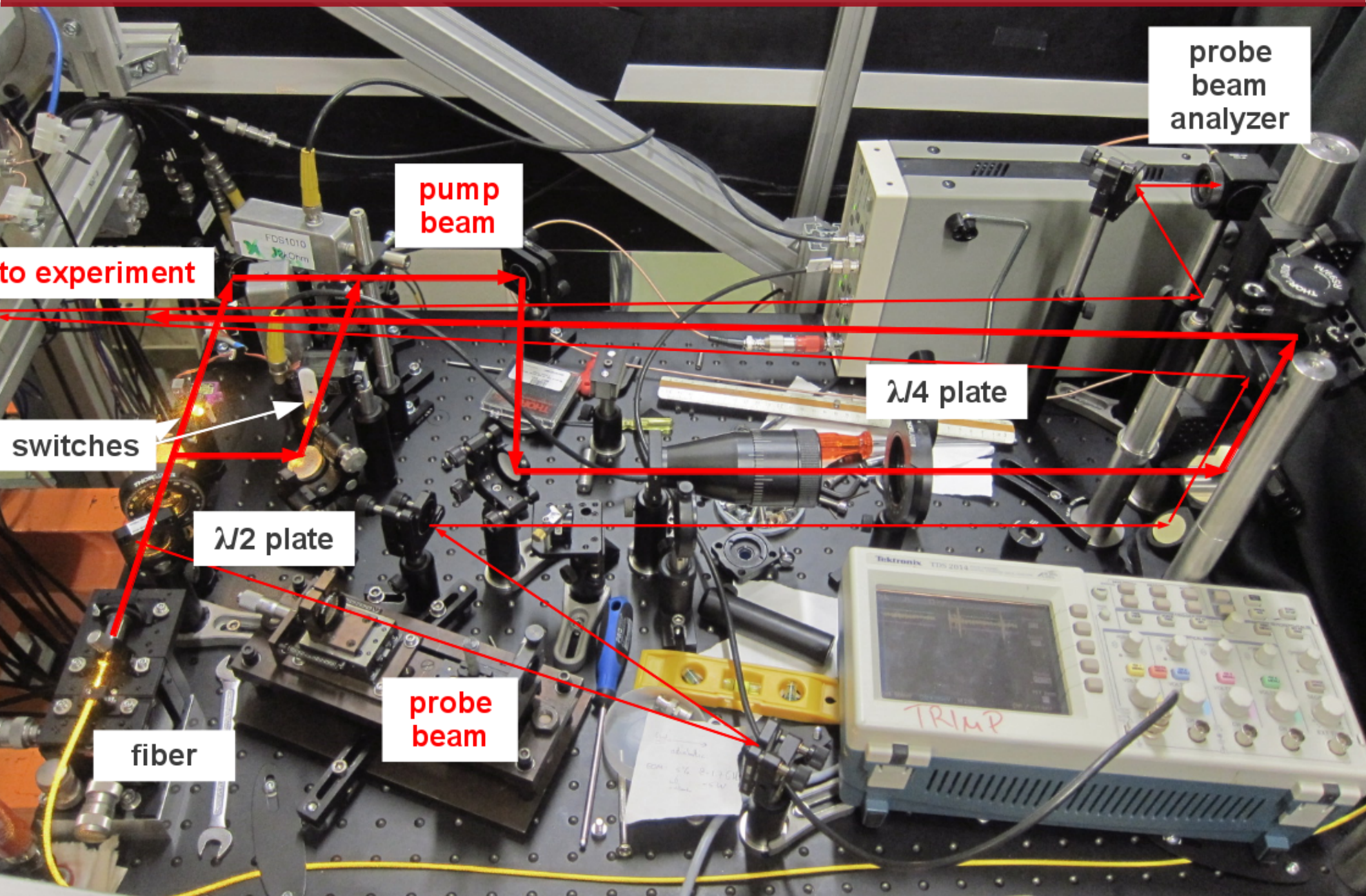
► **Polarized nuclei** via optical pumping:

- Switching polarization:



Experiment:

6

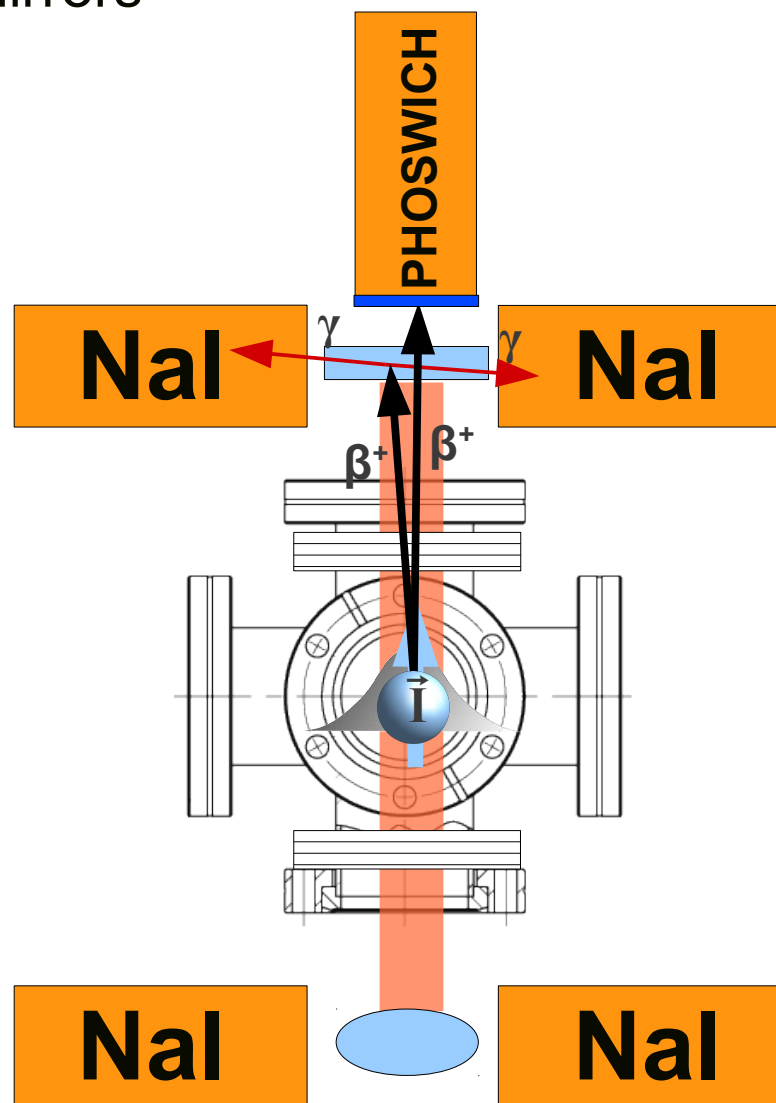


Measurement of polarization:

7

- **PHOSWICH detector** above target cell to detect β^+
- **Two pairs of NaI detectors** to measure 511 keV coincidences from β^+ particles stopped in mirrors above and below target cell

Use parity violating decay asymmetry of weak interaction to monitor nuclear polarization

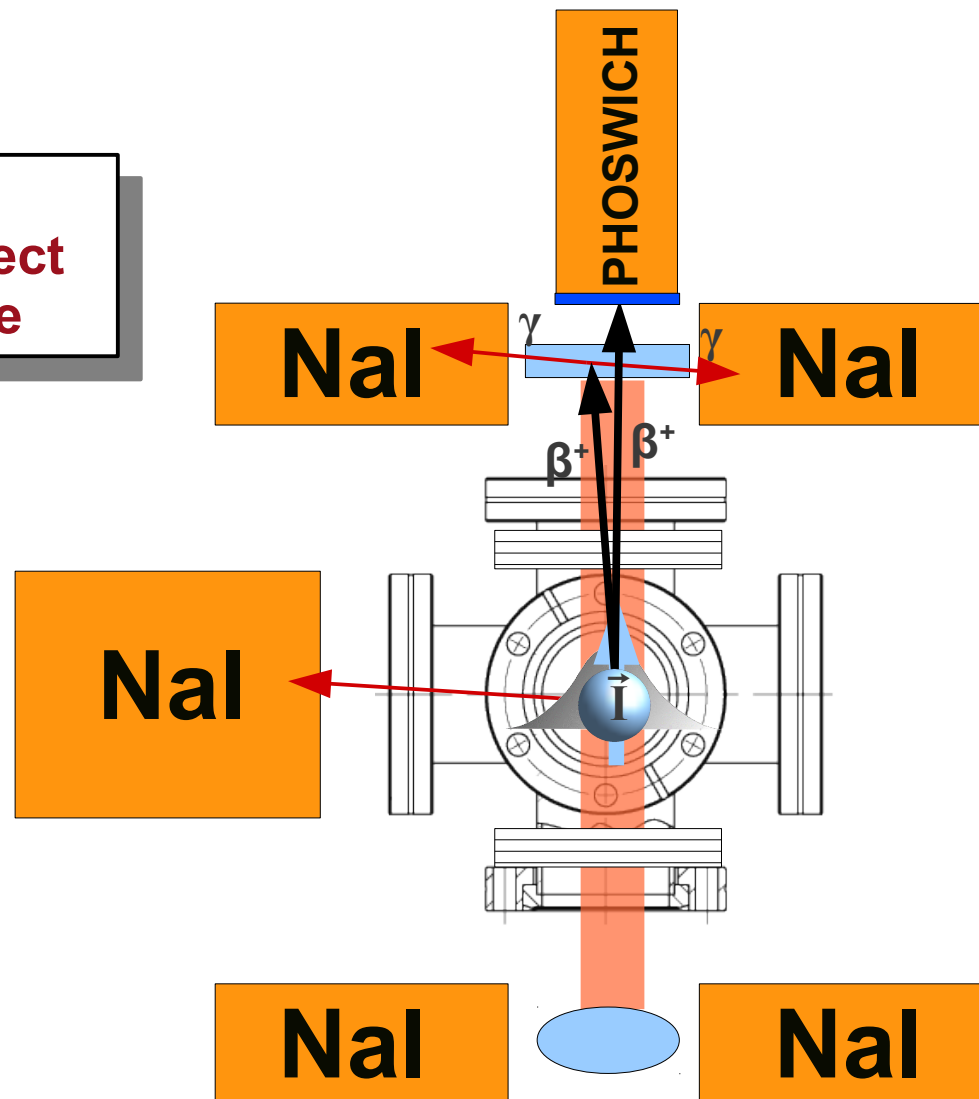


Measurement of lifetime:

8

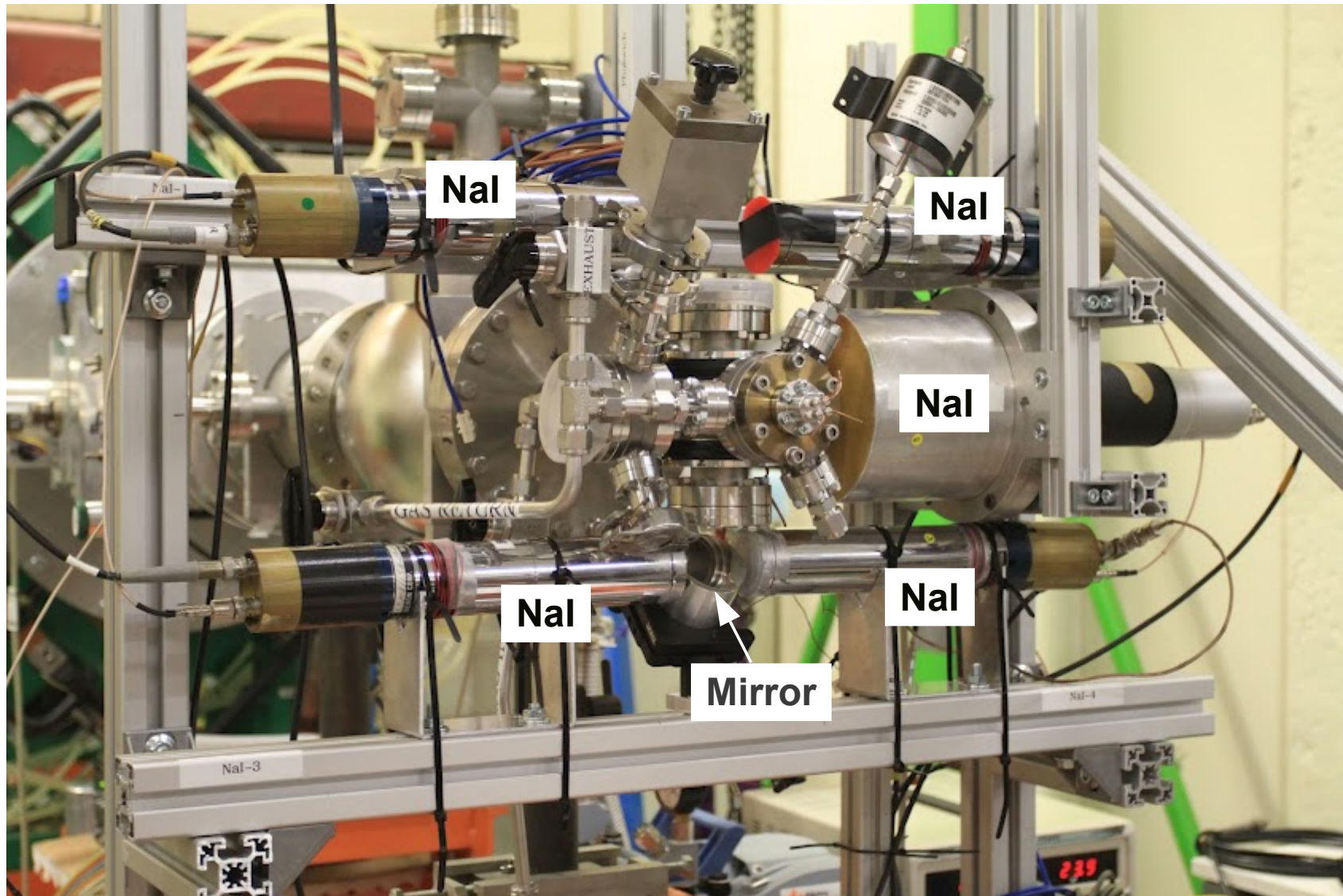
- **Additional NaI detector** for daughter particles decay photons
 $2^+ \rightarrow 0^+$ EM-decay of ^{20}Ne , parity conserving, Lorentz invariant

Use EM decay of ^{20}Ne
daughter nucleus to detect
changes in ^{20}Na lifetime



Experimental setup:

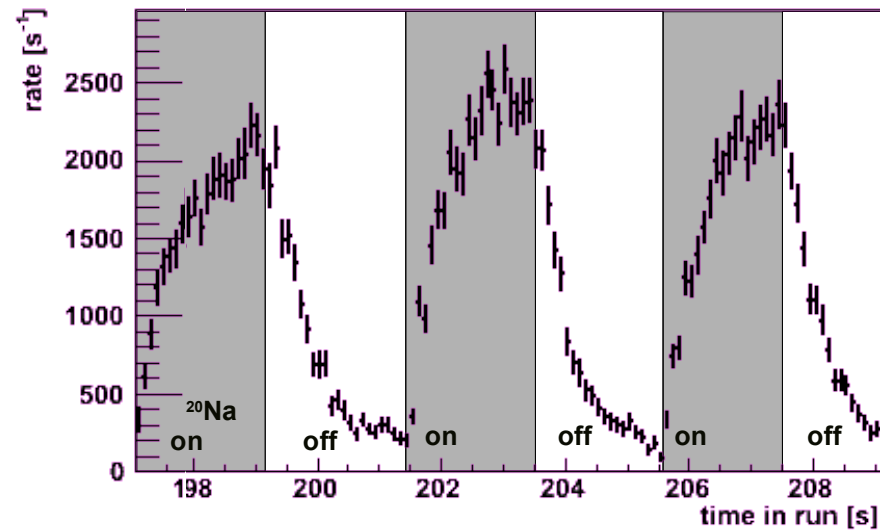
9



Polarization measurement:

10

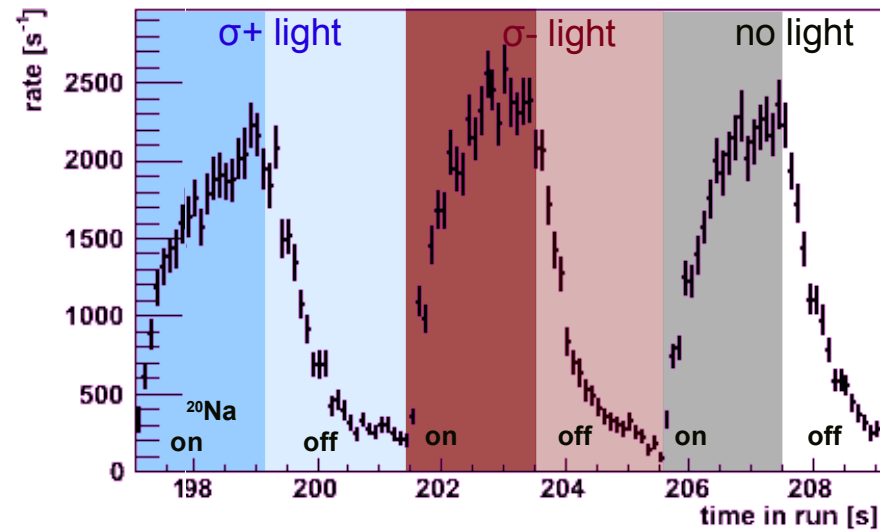
- **β^+ Rates** from PHOSWICH detector
 - 2s-on, 2s-off period of ^{20}Na beam:



Polarization measurement:

10

- **β^+ Rates** from PHOSWICH detector
 - 2s-on, 2s-off period of ^{20}Na beam:

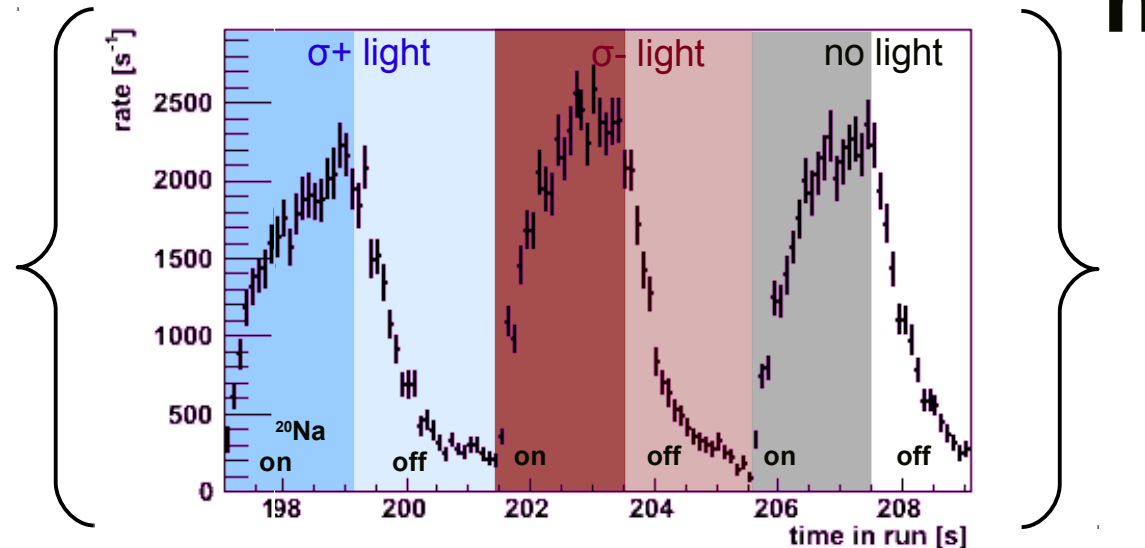


Polarization measurement:

10

► β^+ Rates from PHOSWICH detector

- 2s-on, 2s-off period of ^{20}Na beam:

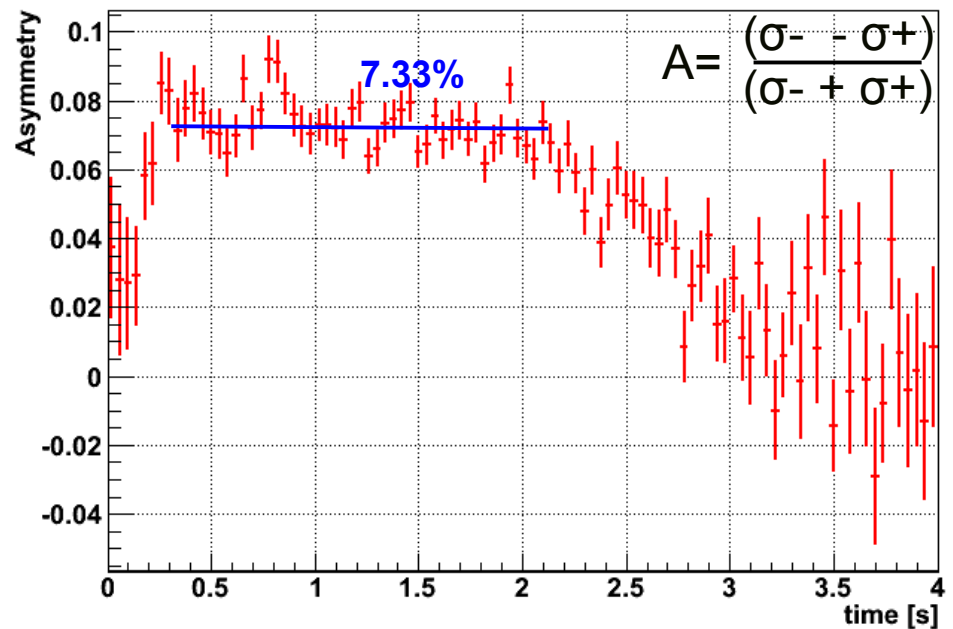
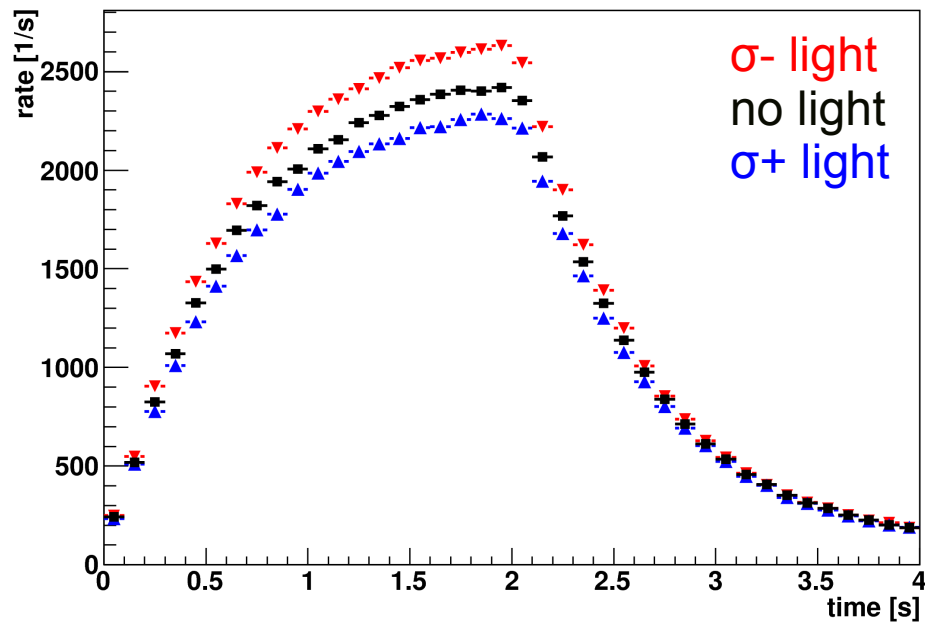
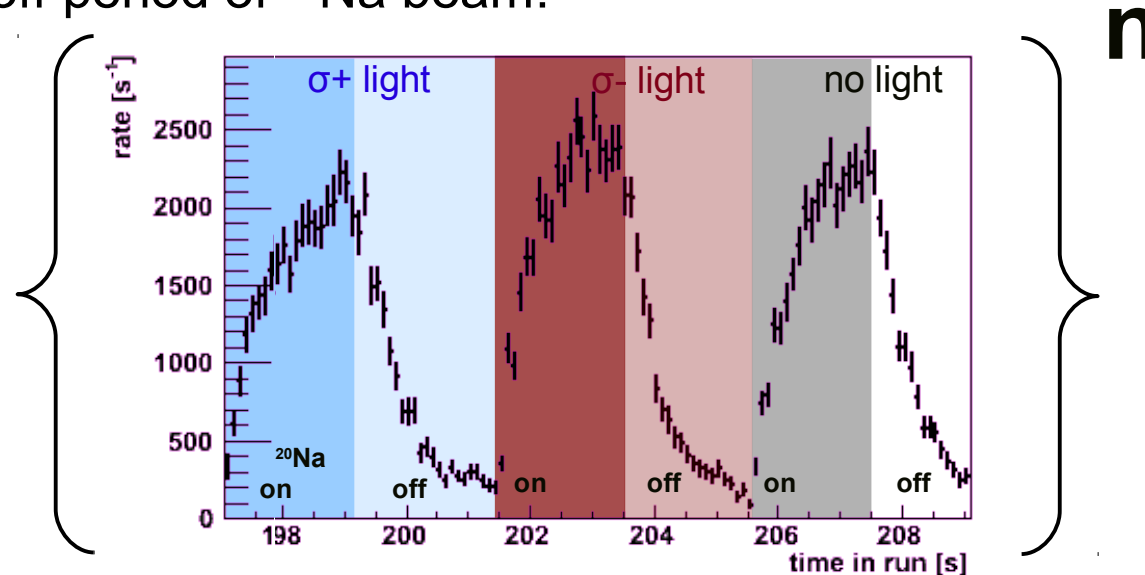


Polarization measurement:

10

► β^+ Rates from PHOSWICH detector

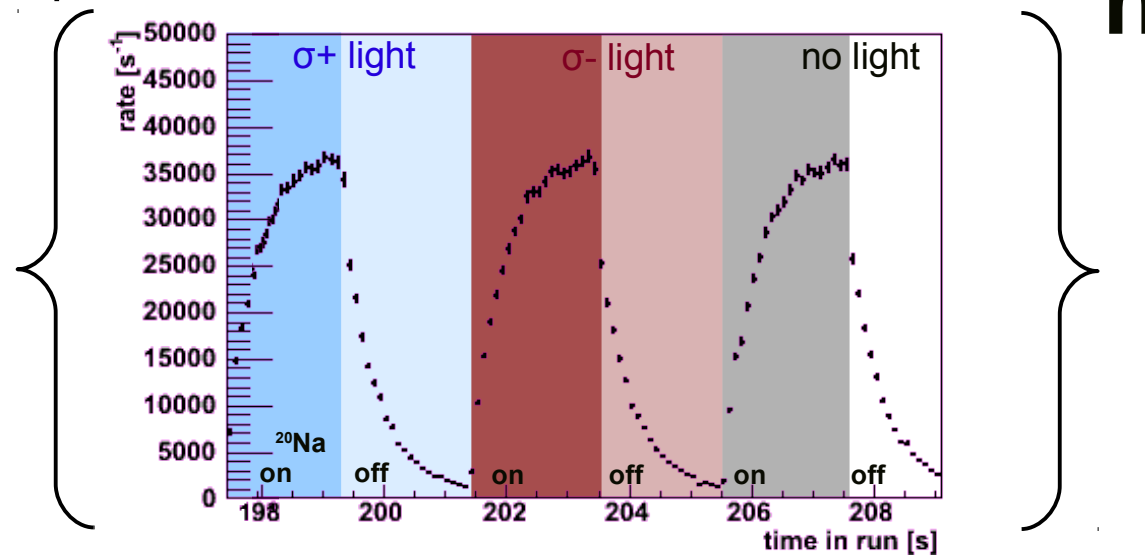
- 2s-on, 2s-off period of ^{20}Na beam:



Lifetime measurement:

11

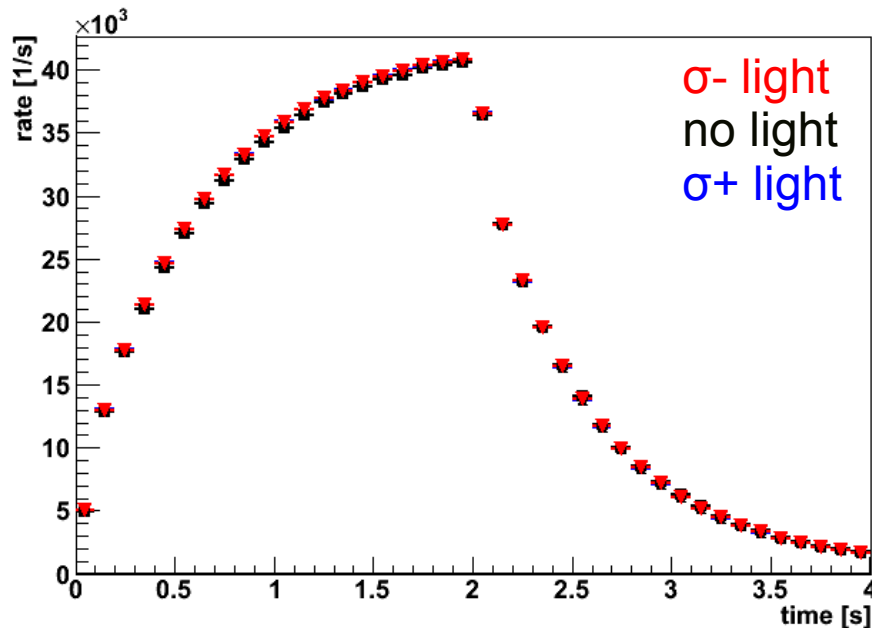
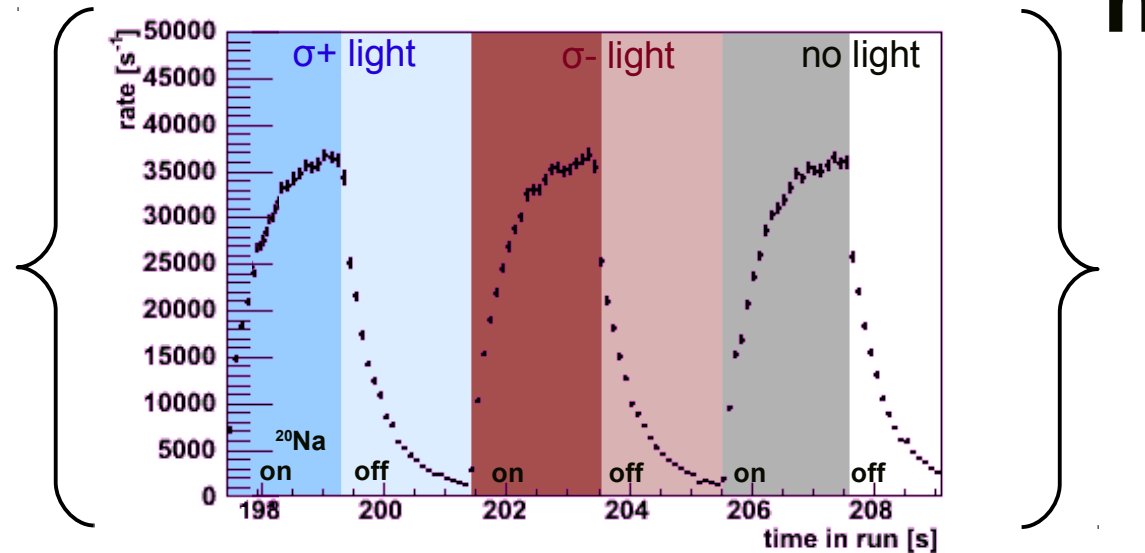
- γ Rates from NaI detector
 - 2s-on, 2s-off period of ^{20}Na beam



Lifetime measurement:

11

- γ Rates from NaI detector
 - 2s-on, 2s-off period of ^{20}Na beam

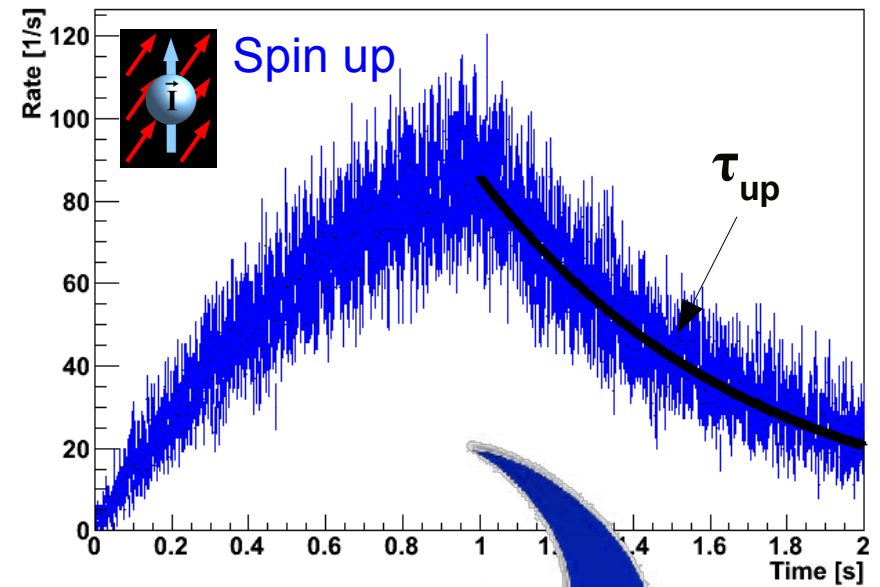
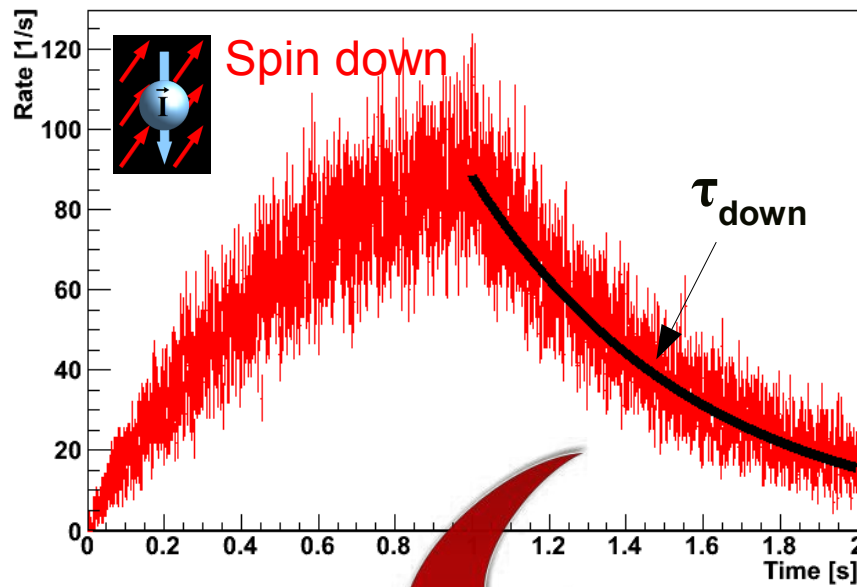


Lifetime-analysis:

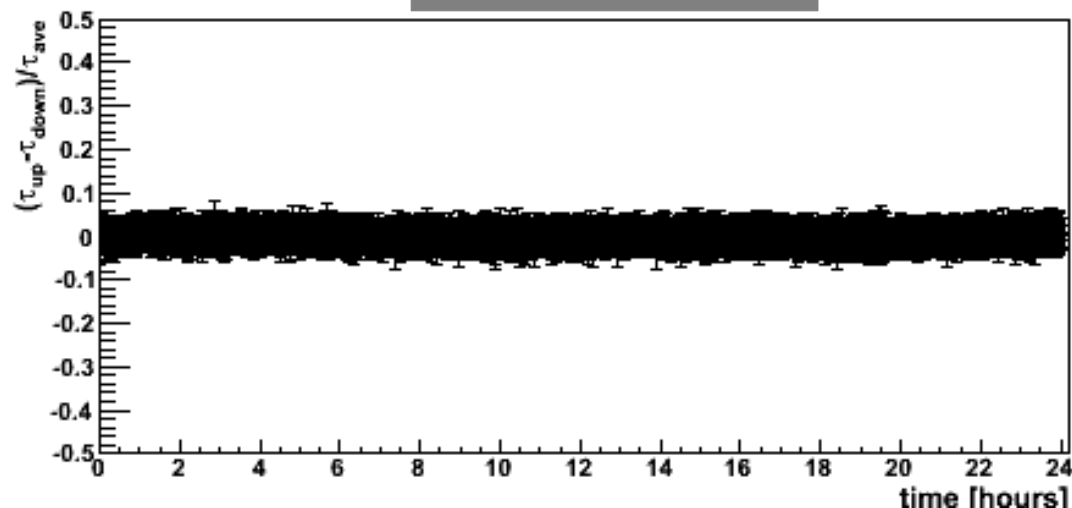
- compare lifetimes for σ^+ and σ^- case
- take into account time-dependence of polarization
- define and estimate systematic effects
- train algorithms on “no light” case

Data Analysis (simulation):

12

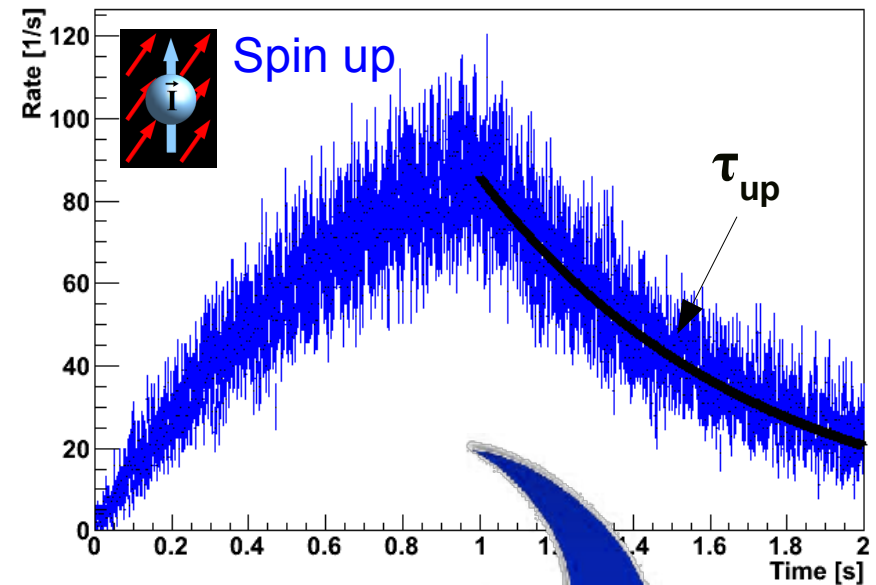
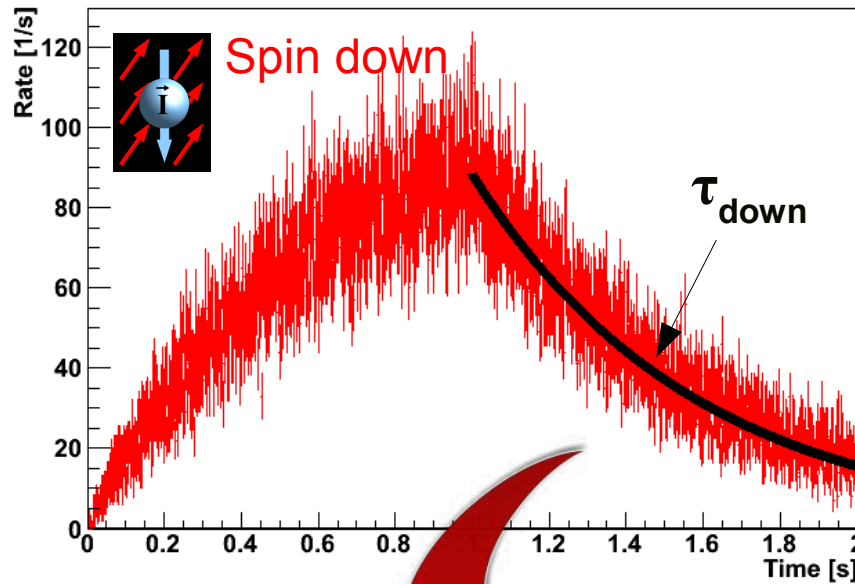


$$\Delta = \frac{\tau_{up} - \tau_{down}}{\tau_{ave}}$$



Data Analysis (simulation):

12

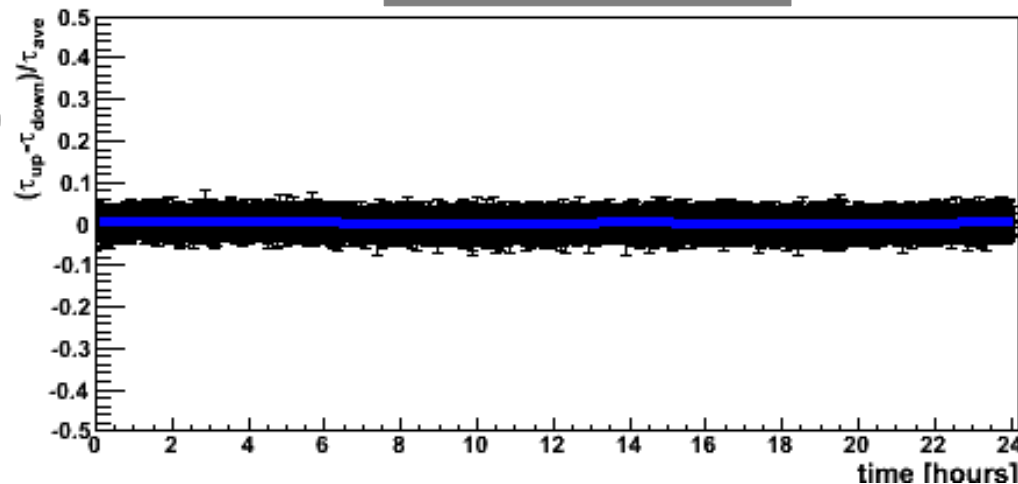


$$\Delta = \frac{\tau_{up} - \tau_{down}}{\tau_{ave}}$$

Fitfunction:

$$\Delta(t) = C + A_s \sin(\omega_{\oplus} t) + A_c \cos(\omega_{\oplus} t) + B_s \sin(2\omega_{\oplus} t) + B_c \cos(2\omega_{\oplus} t)$$

$$\omega_{\oplus} = 2\pi / T_{\text{sid. day}}$$

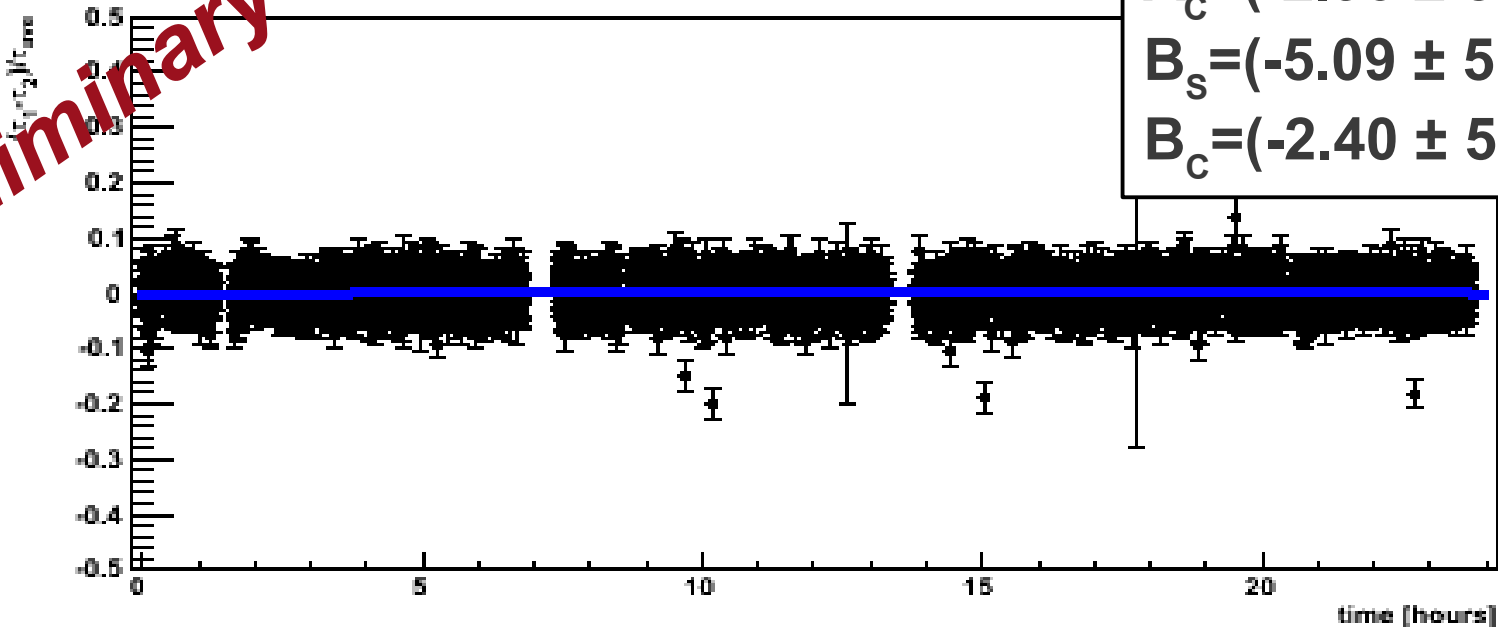


From fit:

$$\begin{aligned} C &= (0.84 \pm 1.31) \times 10^{-4} \\ A_s &= (14.9 \pm 1.85) \times 10^{-4} \\ A_c &= (13.6 \pm 1.85) \times 10^{-4} \\ B_s &= (20.3 \pm 1.85) \times 10^{-4} \\ B_c &= (0.69 \pm 1.85) \times 10^{-4} \end{aligned}$$

Analyzing 24h of non-polarized events:
- *fit lifetimes of 2 consecutive “no-light”-periods*

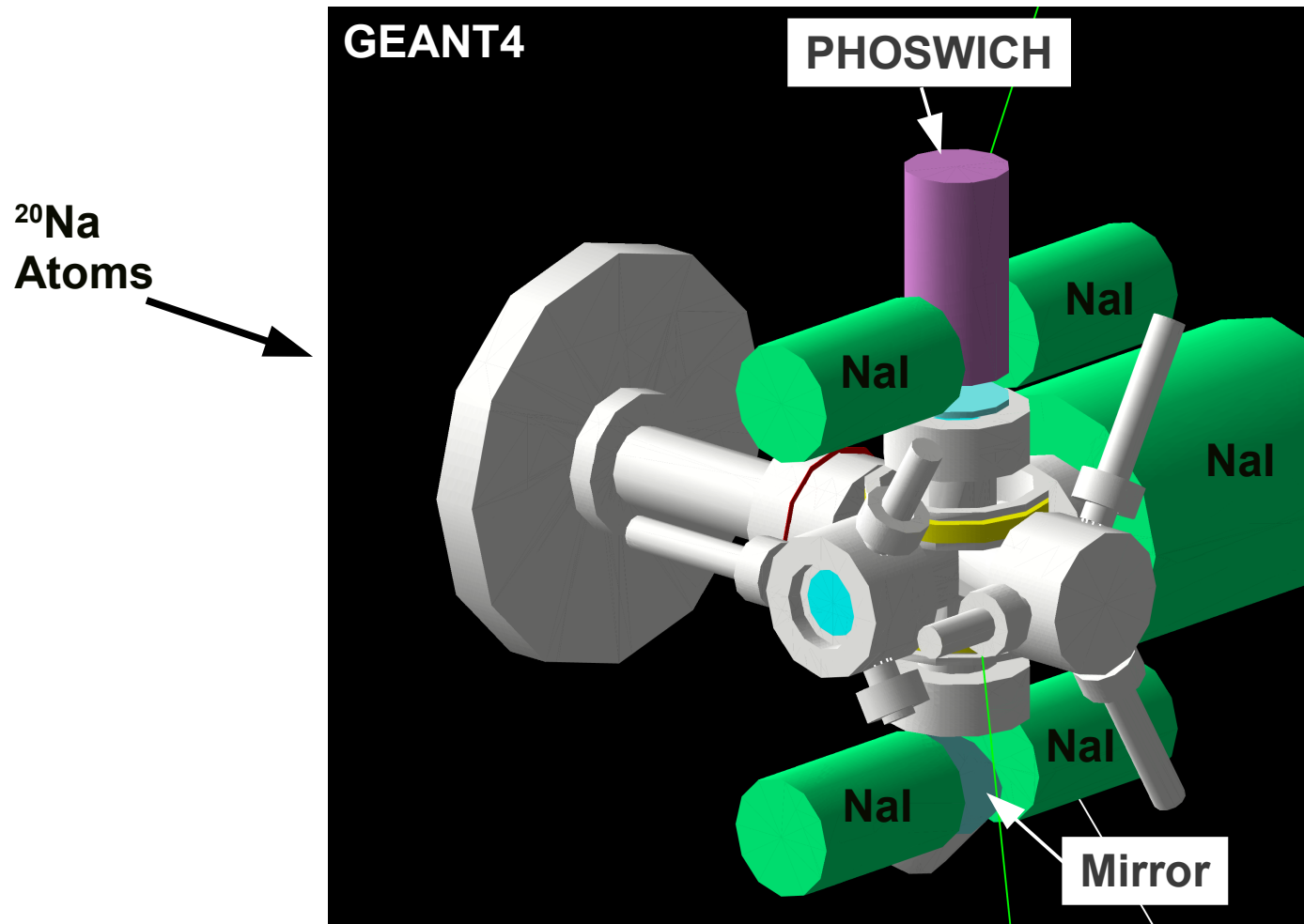
Preliminary



$$\begin{aligned} C &= (4.69 \pm 3.90) \times 10^{-4} \\ A_S &= (0.90 \pm 5.50) \times 10^{-4} \\ A_C &= (-2.90 \pm 5.55) \times 10^{-4} \\ B_S &= (-5.09 \pm 5.55) \times 10^{-4} \\ B_C &= (-2.40 \pm 5.56) \times 10^{-4} \end{aligned}$$

Next steps:

- *determine polarization asymmetry*
- *analyze lifetimes for polarized nuclei*
- *evaluate and quantify systematic effects*



Simulations needed for:

- *detector acceptances*
- *study of systematic effects (stopping position of ^{20}Na atoms, detector alignment, etc.)*

$$\frac{d\Gamma}{dE d\Omega} \sim \left(1 + A_0 \frac{\langle \vec{I} \rangle}{I} \cdot \frac{\vec{p}}{E} \right) + \xi_1 \left(1 + \xi_A \left(\hat{p} \cdot \frac{\langle \vec{I} \rangle}{I} \right) \right) \hat{p} \hat{n} +$$

$$+ \xi_2 \frac{\langle \vec{I} \rangle}{I} \hat{n} + \xi_3 \hat{p}_i \left(\frac{\langle \vec{I} \rangle}{I} \right)_j \rho^{ij}$$

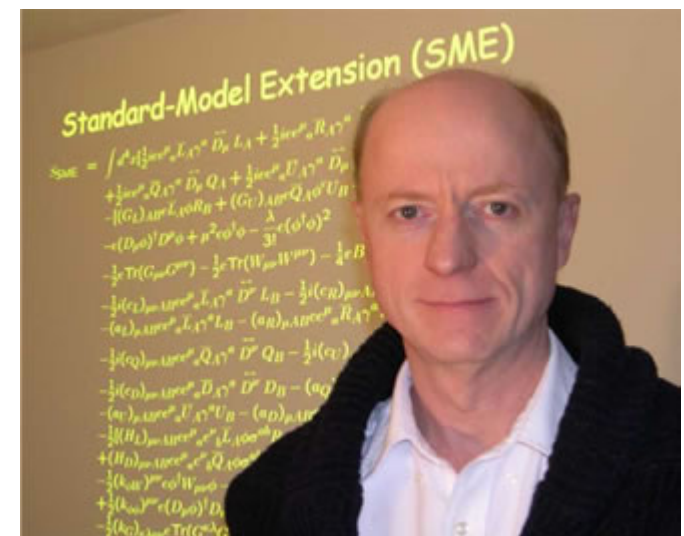
Experiment at KVI probes ξ_2

More general framework to compare with other experiments:

Standard Model Extension (SME)

D. Colladay, A. Kostelecký, PRD58 (1998) 116002)

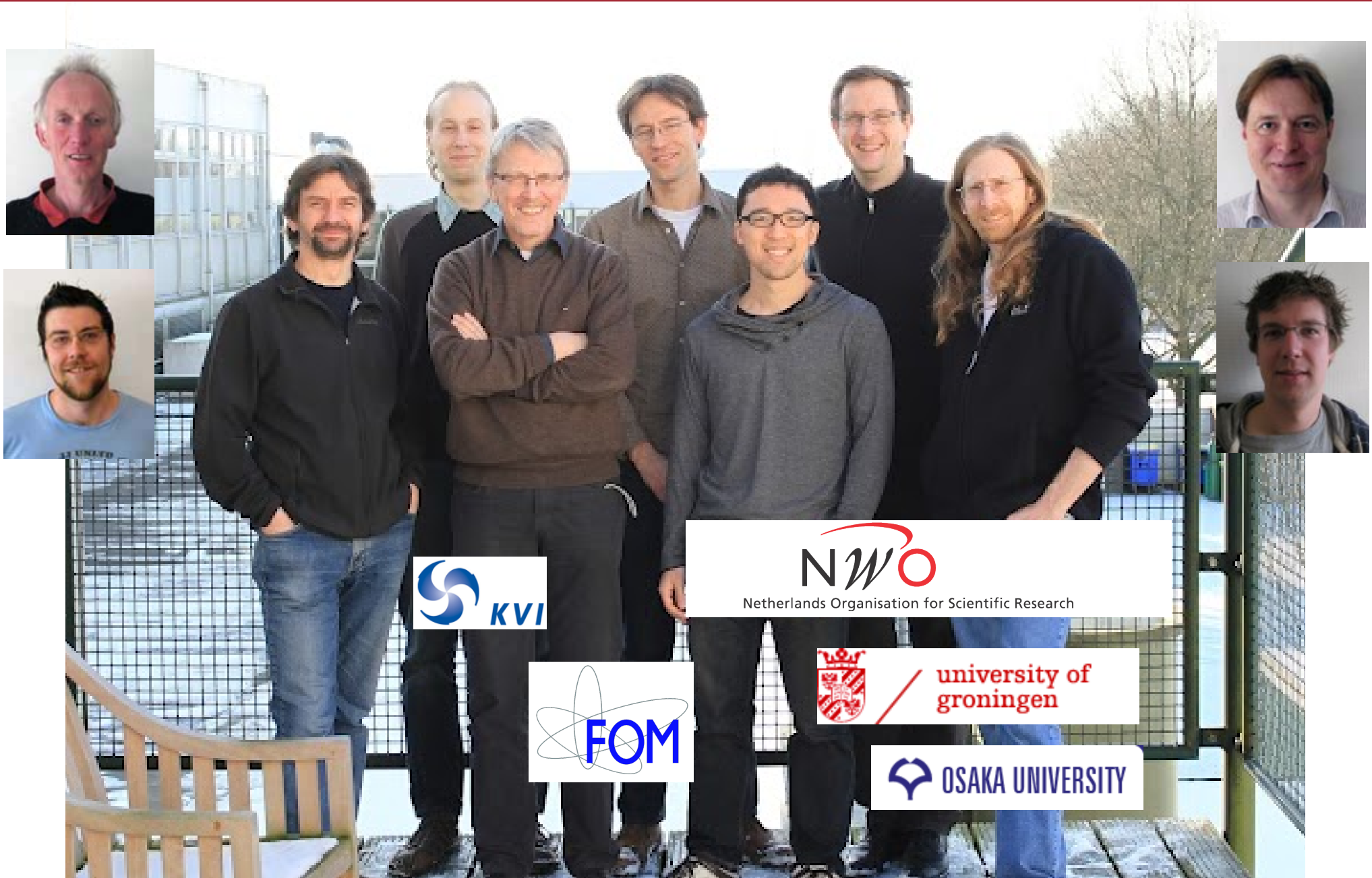
- relate ξ coefficients to SME parameters
- use galactical coordinates in sun-centered equatorial frame



Conclusions

- ▶ **Unique Test of LSB** using weak decay of polarized particles
Probe muon, neutron, radioactive isotopes,...
- ▶ **Combined effort** from theorists and experimentalists at KVI
Interpretation of observables in LSB framework (SME) underway
- ▶ **First dedicated experiment** studying LSB on polarized atoms
Polarization of nuclei achieved, several 24h-periods of data on disk
- ▶ **Outlook**
Lifetime analysis in progress, results expected soon

Thank you!



Greenberg's theorem:

O. W. Greenberg, PRL89 (2002) 231602

“If CPT invariance is violated in an interacting theory, then that theory also violates Lorentz invariance”

“Theories that violate CPT by having different particle and antiparticle masses must be nonlocal”

PDG2012 (“Tests of conservation laws”, L. Wolfenstein and C.-J. Lin):

“The best test comes from the limit on the mass difference between K^0 and \bar{K}^0 ”

Relating measurement to SME parameters:

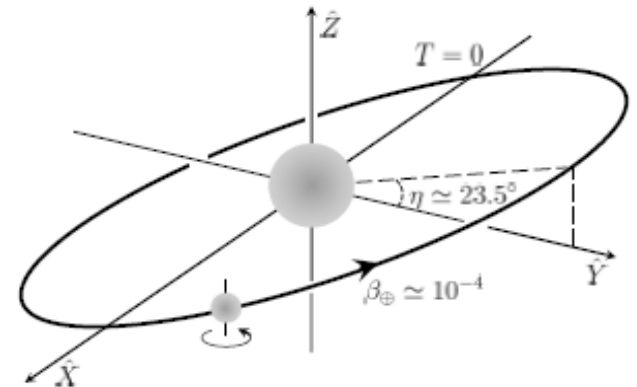
- ξ_2 measured in labframe for spin pointing in +z or -z direction - needs to be transformed into Standard Sun-Centered inertial reference frame

$$\xi_{2, LAB} \propto c_{LAB}^{z0} = c_{SCF}^{x0} \sin \chi \cos(\omega_{\oplus} t_{\oplus}) + c_{SCF}^{y0} \sin \chi \sin(\omega_{\oplus} t_{\oplus}) + c_{SCF}^{z0} \cos \chi$$

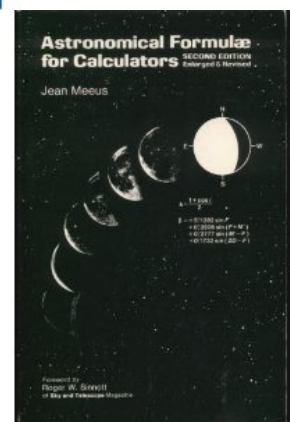
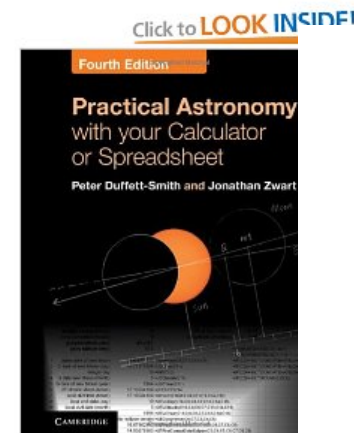
χ = KVI colatitude ($90^\circ - 53.25^\circ$)

$\omega_{\oplus} = 2\pi/T_{\text{sid. day}} = 2\pi/(23\text{h}56\text{m}04\text{s})$

t_{\oplus} = sidereal time

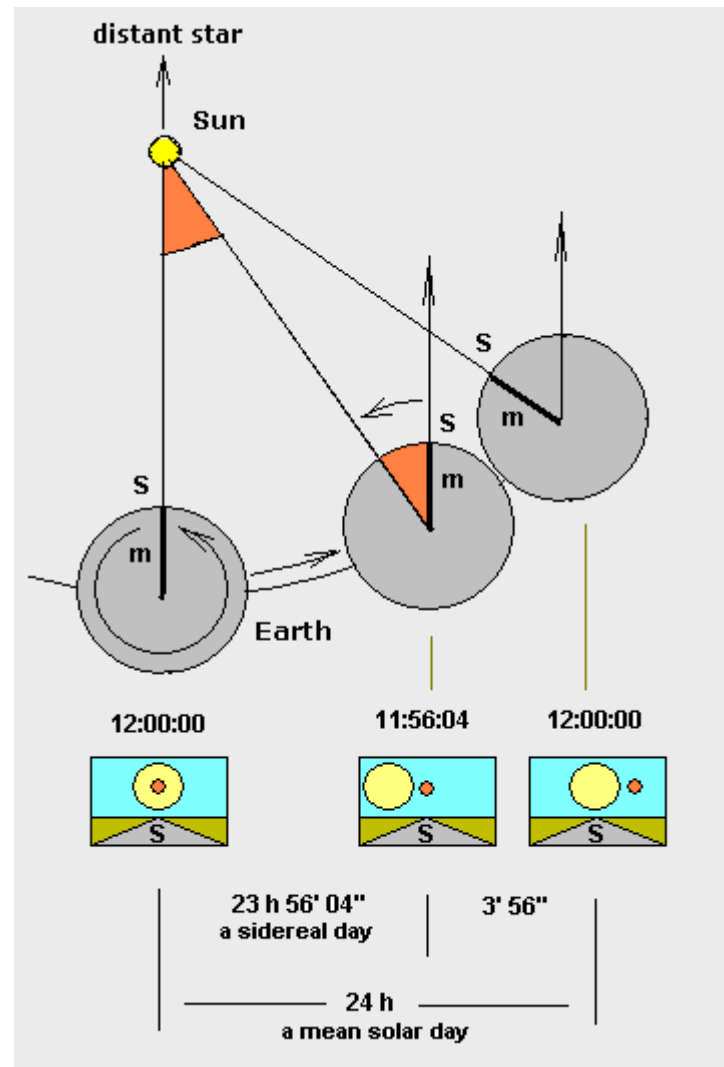


- need to express UNIX time (\sim UTC \sim UT) in sidereal time:
 - J. Meeus - Astronomical Formulae for calculators
 - P. Buffet-Smith, Practical Astronomy with your calculator
 - S. Aoki et al., Astron. Astrophys. 105 (1982) 359



Sidereal time:

- We consider the Lorentz-Violating fields constant and “fixed to a galactical reference frame”
earth moves while rotating → “Solar time” is not useful, need time independent of position of sun



from Wikipedia

Sidereal time:

- We consider the Lorentz-Violating fields constant and “fixed to a galactical reference frame”
earth moves while rotating → “Solar time” is not useful, need time independent of position of sun

Algorithm:

- get universal days elapsed since 01-01-2000 12h UT1
(Julian Day Number since JD 2451545)
- divide by 36525 to get fractional centuries:

$$T_U' = \frac{d_U'}{36525}$$

- Use this to obtain Greenwich Mean Sidereal Time (GMT)
at 0h UT1 in sidereal seconds:

$$\text{GMST1 of } 0^{\text{h}} \text{ UT1} = 24110^{\text{s}}54841 + 8640184^{\text{s}}812866 T_U' + 0^{\text{s}}093104 T_U'^2 - 6^{\text{s}}2 \cdot 10^{-6} T_U'^3 .$$

- divide by 3600 to get sidereal hours
- add longitudinal term $(\eta/360) \cdot 24$ with η KVI long. (6.53° E)
- convert fraction of hours since 0h to sidereal hours and add

$$1 \text{ sid. hour} = 1 \text{ h} \cdot (364.25/365.25)$$

- reduce result to 24 hours

