

Super-Kamiokande

Yoichiro Suzuki,

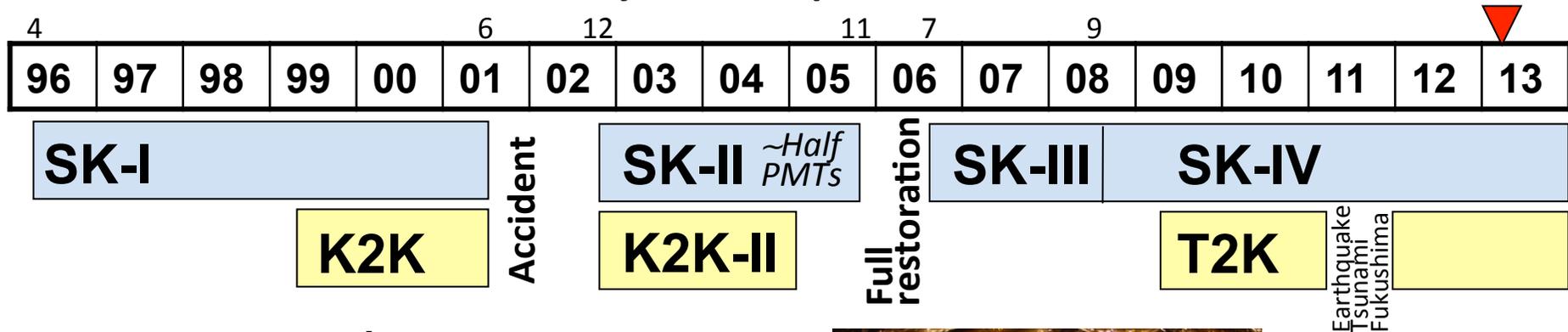
Kamioka Observatory, Institute for Cosmic Ray Research (ICRR),
the University of Tokyo

and

Kamioka Satellite, Institute for the Physics and Mathematics of the
Universe (IPMU), the University of Tokyo

Super-K will be 17 years old in this April

Brief history of Super-Kamiokande



- 1996: SK started
- 1998: Discovery of Atmospheric Neutrino Oscillation
- 2001: Discovery of Solar Neutrino Oscillation (w/SNO)
- 2004: Confirmation of the atmospheric neutrino oscillation by K2K.
- 2004: Discovery of the oscillatory behavior of the atmospheric neutrinos
- 2011: Indication of θ_{13} by T2K



Water Cherenkov
50,000 tons
11,129 PMTs
22.5 kt
fiducial mass



Protection case₂

SK Collaboration

	1998	2013
ICRR	27	23
IPMU	--	2
Fukuoka Tech	--	1
Gifu	1	1
KEK	9	11
Kobe	3	2
Kyoto	--	6
Miyagi	1	1
Nagoya	--	4
Niigata	8	--
Okayama	--	5
Osaka	5	1
Tohoku	13	--
SW Shizuoka	--	1
Tokyo	1	3
Tokyo Tech	5	--
Tokai	2	1

Japan	75	62
USA	55	31
Poland	1	1
Spain	--	1
Korea	--	5
China	--	4
Canada	--	8
Total	131	112

	1998	2013
Boston	14	4
BNL	1	1
Irvine	9	7
California State	2	3
Duke	--	5
George Mason	1	--
Hawaii	6	3
Los Alamos	1	--
Louisiana State	3	--
Maryland	4	--
Stony Brook	8	5
Washington	6	3
Warsaw(Poland)	1	1
UAM(Spain)	--	1
Chonnam(Korea)	--	3
Seoul National(Korea)	--	1
Sungkyunkwan(Korea)	--	1
Tsinghya(China)	--	4
Regina(Canada)	--	1
British C.(Canada)	--	3
Toronto(Canada)	--	2
TRIUMF(Canada)	--	2

- T2K is a separated collaboration
 - Not all the SK collaborators are involved in T2K
- New countries (Spain, Korea, China, Canada) joined

Status of Super-K

- People might say that SK has finished the role.
- But not true
 - ***Last year our funding agency asked us to provide a plan of SK for the next 10 years***
 - Many interesting subjects still to be solved
 - ***the number of collaborating institutions have increased for the last few years***
 - Good to learn Water Cherenkov technology for future

10 years plan

MEXT asked us to provide 10 years plan of Super-Kamiokande

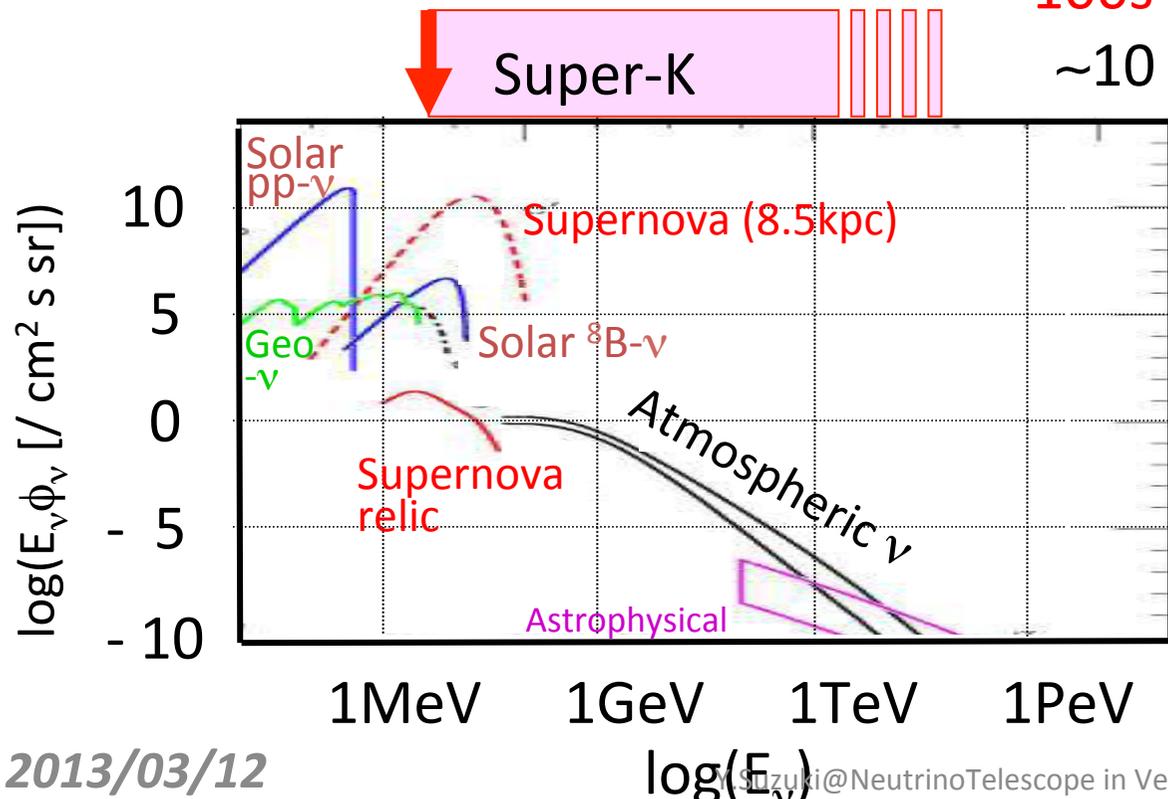
Subjects	2013	2014	2015	2016	2017	2018	2019	2020	2021	2022
Supernova Relic ν Entering a discovery region (<i>expected flux</i> : $10 \sim 22/\text{cm}^2/\text{s}$)	Improve water system									Discovery ?
	sensitivity			$24/\text{cm}^2/\text{s}$			$13/\text{cm}^2/\text{s}$			$9/\text{cm}^2/\text{s}$
Nearby SN neutrino burst World highest sensitivity		Improve DAQ system								
										Continuous observation
Mass hierarchy Precise measurement of atmospheric ν										Identification of MH ?
	sensitivity			2.0σ			2.3σ			2.5σ
Indication of CP Violation ? Look for ν_e appearance in SK/T2K(J-PARC)	ν_e appearance									
										Upgrade J-PARC (750kW)
										Combine results from reactor exp.
										Indication of CP ?
Sterile neutrinos ? Precise measurement of solar neutrinos	Reduce backgrounds									Indication of sterile neutrinos ?
	sensitivity			2.2σ			2.7σ			3.0σ
Nucleon decay search World highest sensitivity										
										Test of grand unified theories
										sensitivity 2.1×10^{34} yrs
Search for dark matter Neutrinos from the sun, the earth and so on										
										Sensitive to low mass WIMPs
										Improve sensitivity by factor 2

2013/03/12

Energy Range

- Trigger (Software):
 - 50% efficiency @ 3.3MeV
 - 100% eff. for $E_{kin} > 4.0$ MeV
- Trigger Rate
 - 3.5 kHz

- Solar neutrinos (< 15 MeV):
 - ~15 events /day
- Supernova ν 's (10~20 MeV):
 - ~8000 events @10 kpc
- Atmospheric Neutrinos (< a few 100s GeV):
 - ~10 events /day



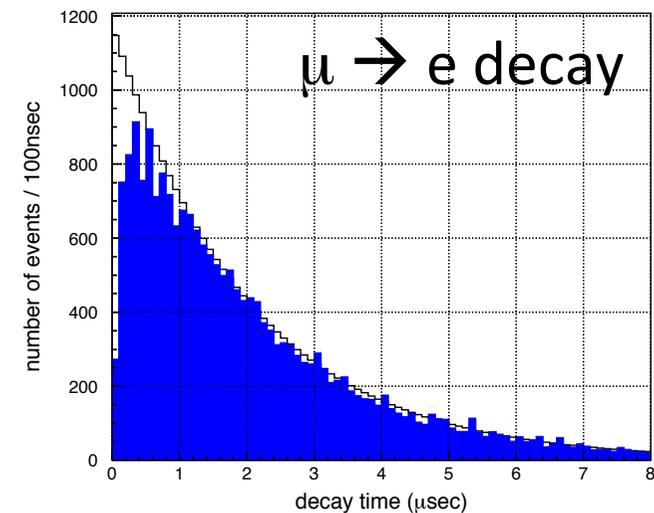
- 6 p.e. / MeV
- Resolution
 - (solar/supernova ν)
 - 14.2% @10MeV
 - (atmospheric ν)
 - $1.7+0.7/\sqrt{E(\text{GeV})}$ %
 - (single ring μ)

Data Accumulated

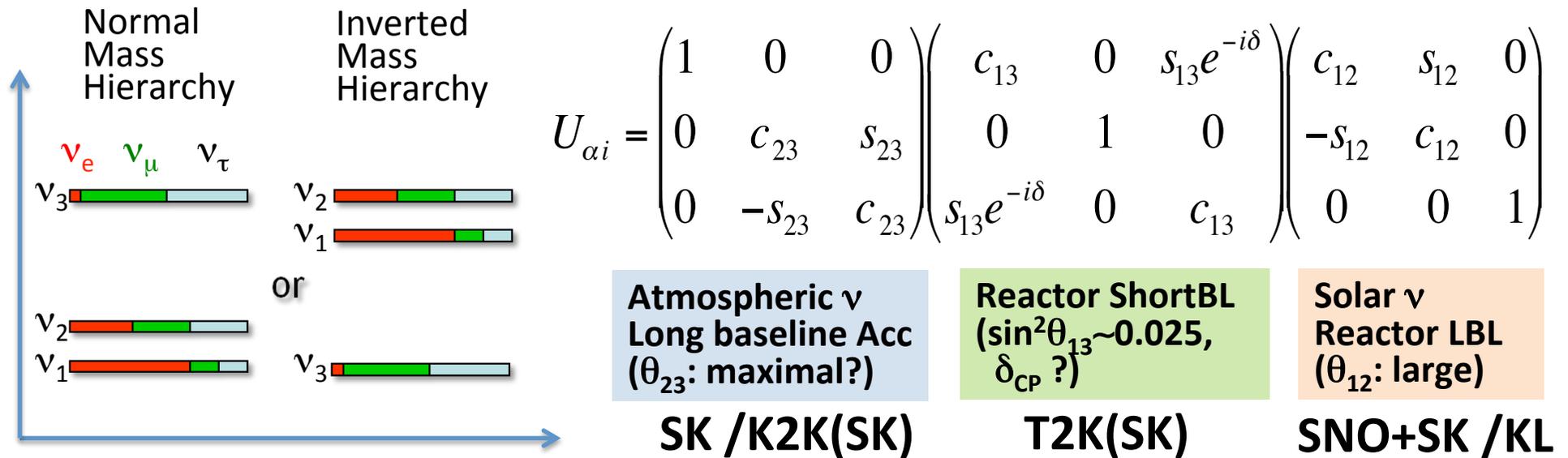
Phase	SK-I	SK-II	SK-III	SK-IV	Total
Periods	96-Apr ~01-Jun	02-Dec ~05-Nov	06-Jul ~08-Sep	08-Sep ~running	X
ID PMTs	11,146 (40%)	5,182 (19%)	11,129 (40%)	11,129 (40%)	
Electronics	ATM	ATM	ATM	QBEE	
Trigger	Hardware	Hardware	Hardware	Software	
Atm ν FC+PC (days)	1489 days	799 days	518 days	1097 days (< '12 Mar)	3903 days (< '12 Mar)
(# of ev.)	12,299+902	6,610+427	4,355+344	8,929+735	32,193+2,408
Atm ν up- μ (days)	1646 days	828 days	636 days	1097 days (< '12 Mar)	4207 days (< '12 Mar)
(# of ev.)	2,328	1,094	945	1,651	6,018
Solar ν (days)	1496 days	791 days	547.9 days	1069.3 days (< '12 Mar)	3904 days (< '12 Mar)
(# of ev.)	22,404 ev.	7,212.8 ev.	8,147.9 ev.	19,809.4 ev.	57,574.1 ev.
Proton decay	91.7 kt·yr	49.2 kt·yr	31.9 kt·yr	46.5 kt·yr (< '11 Mar)	220 kt·yr

Electronics update (SK-IV)

- QBEE(QTC-Based Electronics with Ethernet)
 - width of the output time pulse represent the integrated charge
 - Dynamic range: 0.2 – 2500 pC
 - 5 times better than the previous one
- Record every hit + software trigger
 - Higher efficiency for $\mu \rightarrow e$ decay
 - Detection of delayed 2.2 MeV γ -rays after neutron captures
- High event process:
 - Up to 6 Million events /sec without any event loss.
 - 100 times better than the previous one



Neutrino Oscillation



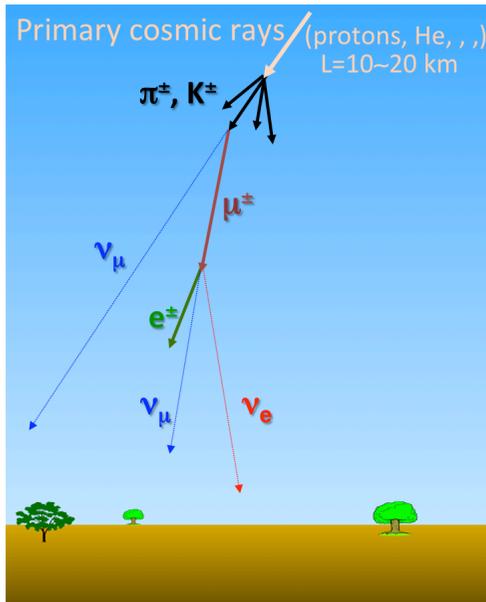
Super-K have not only discovered the neutrino oscillation, but also played the essential roles in discovering all the mixing angles.

Remaining Issues

- Octant of θ_{23} (if $\theta_{23} \neq \pi/4$); Mass hierarchy (sign of Δm_{13}^2), CPV

← SK atmospheric three flavor analysis may give some hints in near future

Atmospheric Neutrinos



- Wide range of baseline: 10 km ~ 13,000 km
- 5 orders of magnitude of energy range: 0.1 GeV ~ 10 TeV
- Variety of Matter Effect
- Four neutrino sources: $\nu_\mu, \bar{\nu}_\mu, \nu_e, \bar{\nu}_e$

Fully Contained (FC)

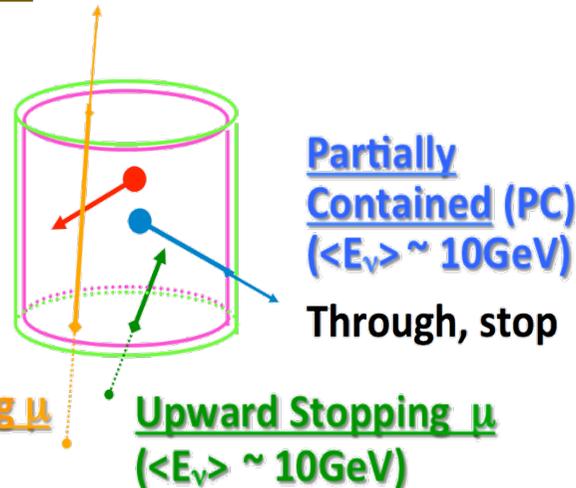
($\langle E_\nu \rangle \sim 1\text{GeV}$)

Sub-GeV: $E_{\text{vis}} < 1.33\text{GeV}$

Multi-GeV: $> 1.33\text{GeV}$

1R, MR, e-like, μ -like

ν_e -like, $\bar{\nu}_e$ -like, ...



Upward Through-going μ

($\langle E_\nu \rangle \sim 100\text{GeV}$)

Non-showering, showering

Many Event Categories

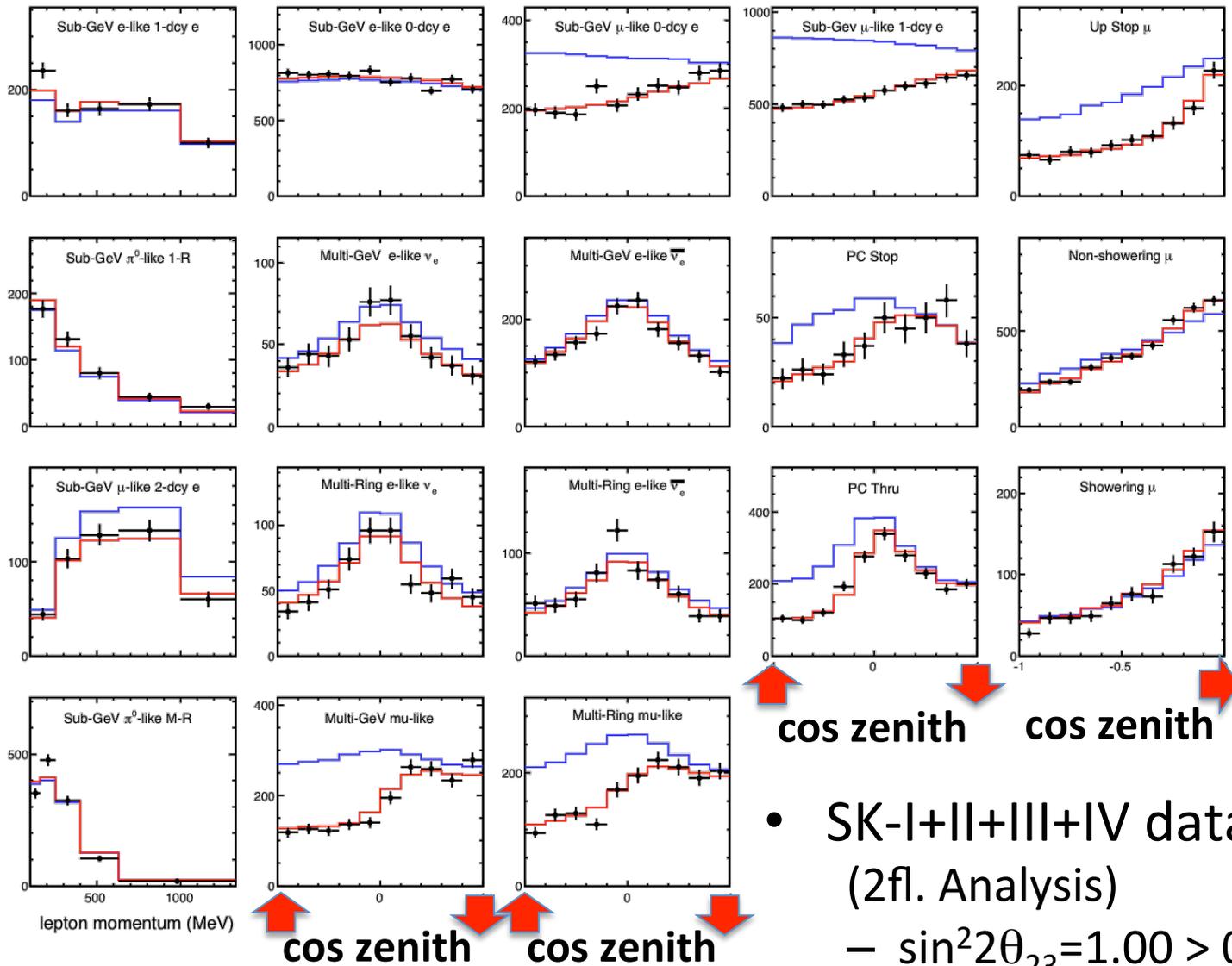
← Energy, topology, # of Rings, # of decay electrons, e-like, μ -like, ...

18 Categories

– 480 momentum/zenith angle bins (as of March, 2012) are used for the fits

→ Good to identify best suitable events for different kinds of oscillation studies

Atmospheric Neutrinos



- 18 Categories
- All consistent with the oscillation

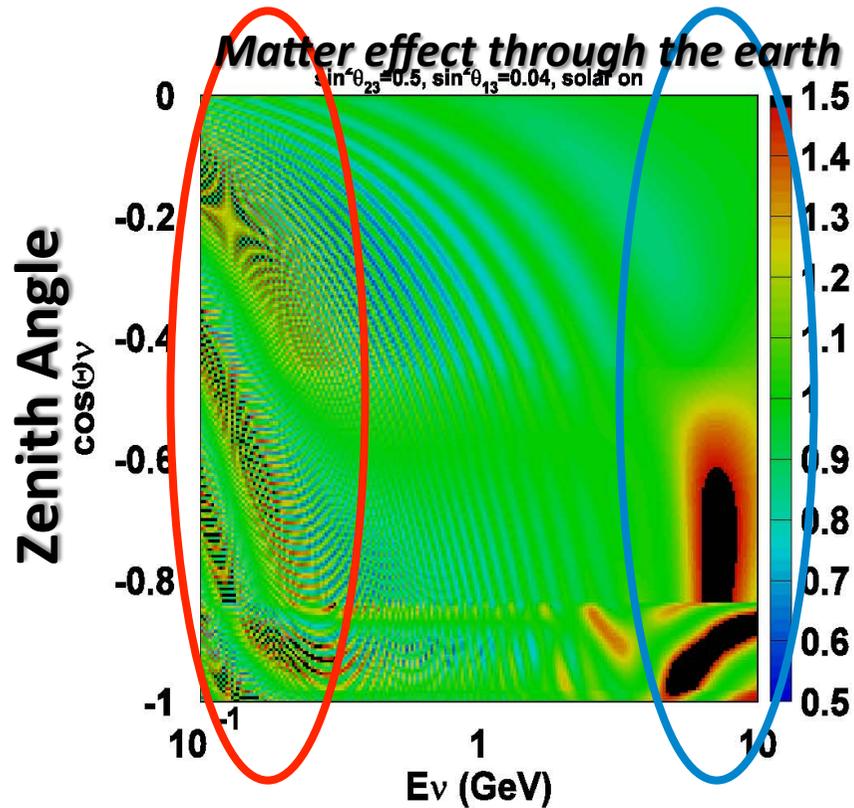
— w/o oscillation
— w/ oscillation

- SK-I+II+III+IV data (2fl. Analysis)
 - $\sin^2 2\theta_{23} = 1.00 > 0.96$ (90%, 1par)
 - $\Delta m_{23}^2 = 2.30^{+0.16}_{-0.22} \times 10^{-3} \text{ eV}^2$ (1σ)

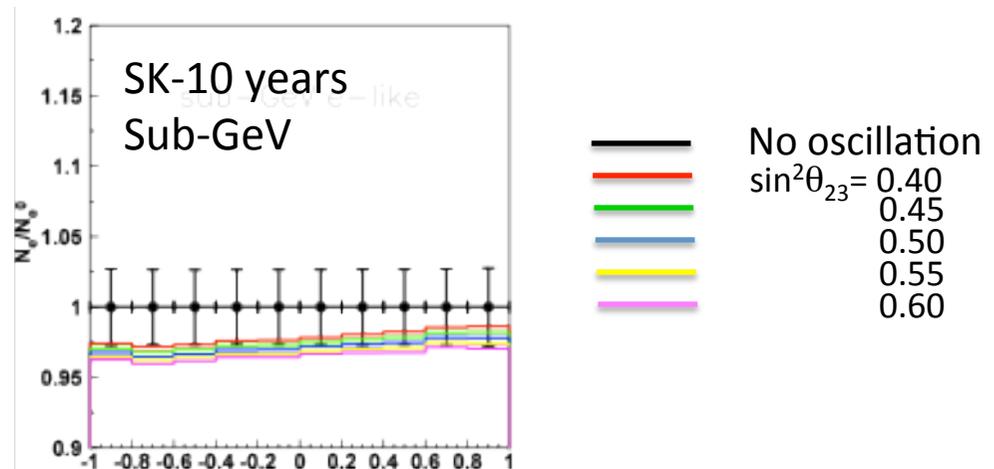
Three flavor analysis ν_e appearance in atmospheric- ν

$$\frac{\Psi(\nu_e)}{\Psi_0(\nu_e)} - 1 \cong P_2(r \cdot c_{23}^2 - 1) - r \cdot \tilde{s}_{13} \cdot \tilde{c}_{13}^2 \cdot \sin 2\vartheta_{23} (\cos \delta_{CP} \cdot R_2 - \sin \delta_{CP} \cdot I_2) + 2\tilde{s}_{13}^2 (r \cdot s_{23}^2 - 1)$$

$\tilde{}$: mixing angle in matter
 $P_2 = |A_{e\mu}|^2 : \nu_e \rightarrow \nu_{\mu\tau}$ in matter
 $R_2 = \text{Re}(A_{ee}^* A_{e\mu})$
 $I_2 = \text{Im}(A_{ee}^* A_{e\mu})$



1st term: solar term ($\theta_{12}, \Delta m_{12}$)
 mostly in low energy
 cancellation effect
 if $c_{23}^2=0.5, r=\nu_{\mu}/\nu_e=2@LE$
 a few %; octant may be seen



$s^2\theta_{12}=0.825, s^2\theta_{23}=0.4, s^2\theta_{13}=0.04$
 $\delta_{CP}=45^\circ, \Delta m^2_{12}=8.3 \times 10^{-5}, \Delta m^2_{23}=2.5 \times 10^{-3}$

Three flavor analysis

ν_e appearance in atmospheric- ν

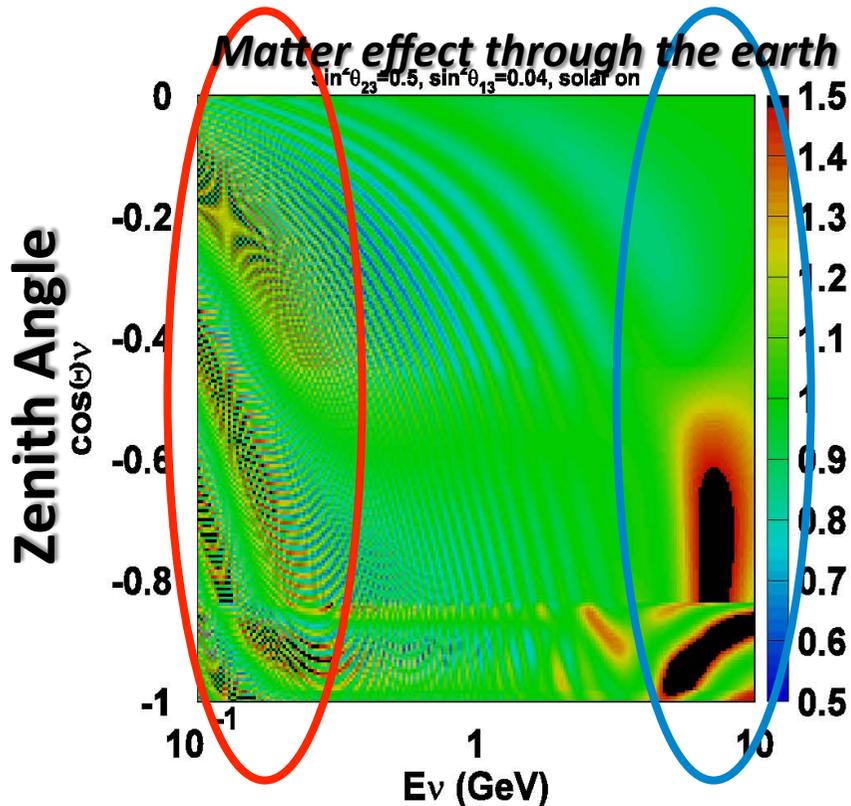
$$\frac{\Psi(\nu_e)}{\Psi_0(\nu_e)} - 1 \cong P_2(r \cdot c_{23}^2 - 1) - r \cdot \tilde{s}_{13} \cdot \tilde{c}_{13}^2 \cdot \sin 2\vartheta_{23} (\cos \delta_{CP} \cdot R_2 - \sin \delta_{CP} \cdot I_2) + 2\tilde{s}_{13}^2 (r \cdot s_{23}^2 - 1)$$

$\tilde{}$: mixing angle in matter

$P_2 = |A_{e\mu}|^2 : \nu_e \rightarrow \nu_{\mu\tau}$ in matter

$R_2 = \text{Re}(A_{ee}^* A_{e\mu})$

$I_2 = \text{Im}(A_{ee}^* A_{e\mu})$



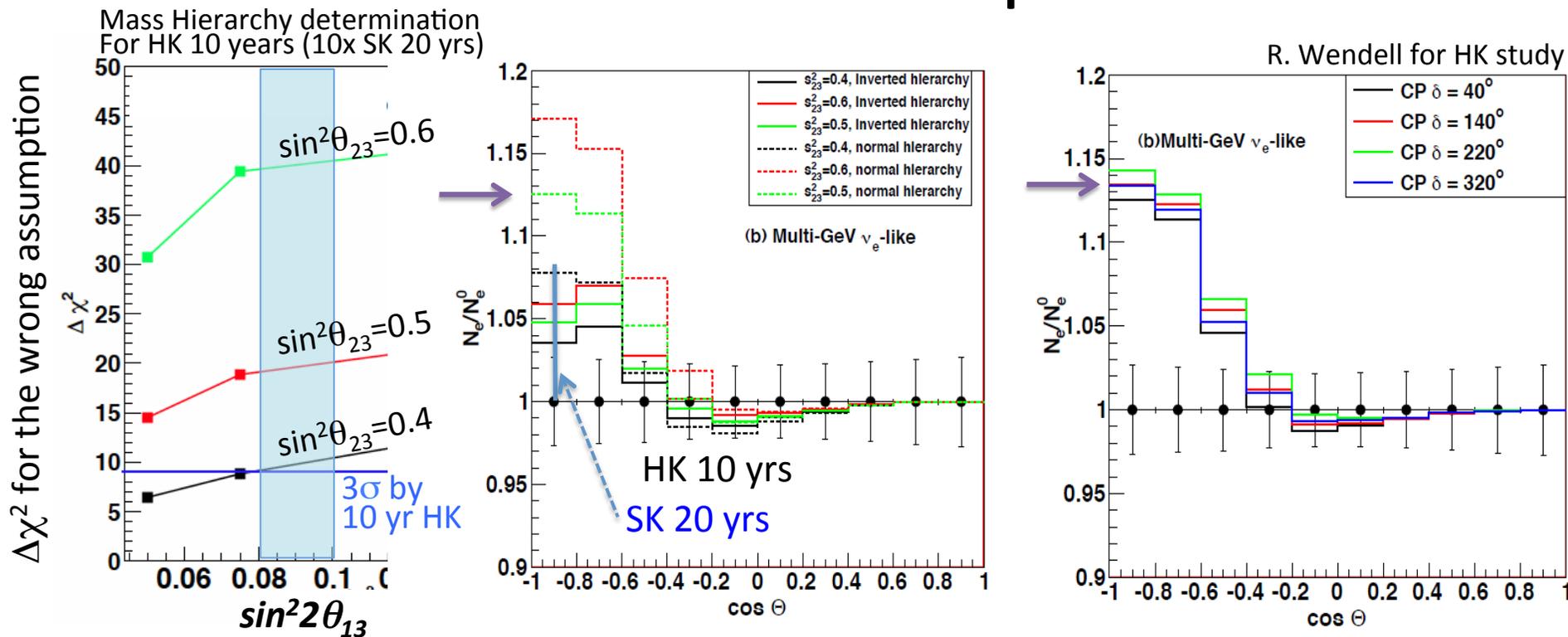
3rd term: θ_{13} term, Matter Effect Resonance for
 neutrinos + Normal MH
 antineutrinos + Inverted MH
 5~15% effect

2nd term: Interference: CP-Phase

→ Multi-GeV sample

$s^2\theta_{12}=0.825, s^2\theta_{23}=0.4, s^2\theta_{13}=0.04$
 $\delta_{cp}=45^\circ, \Delta m^2_{12}=8.3 \times 10^{-5}, \Delta m^2_{23}=2.5 \times 10^{-3}$

Multi-GeV sample



- Resonance Effects: 5~15%, Normal MH > Inverted MH
- Difference for MH is larger for $\cos^2\theta_{23} > 0.5$
- CP difference is ~1~2 % for SK → difficult to measure
- For MH → SK may indicate ~2 σ level effect in a few more years if 'the parameters are lucky for us'
- Since MH is a kind of on/off, therefore 2~3 experiments showing 2~3 σ effect may be enough to determine!

We have made ν_e ($\bar{\nu}_e$) enhanced samples

- ν_e → larger $\gamma=(E-E')/E$
- larger # of π
- diff. # of decay-e

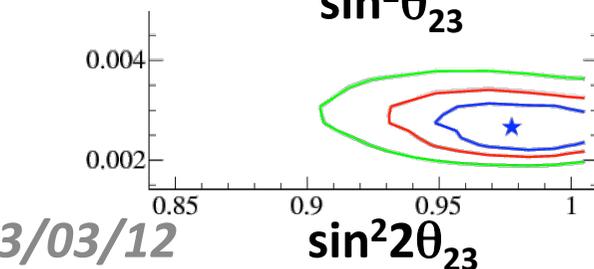
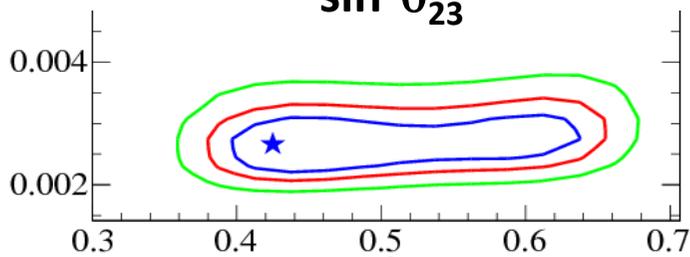
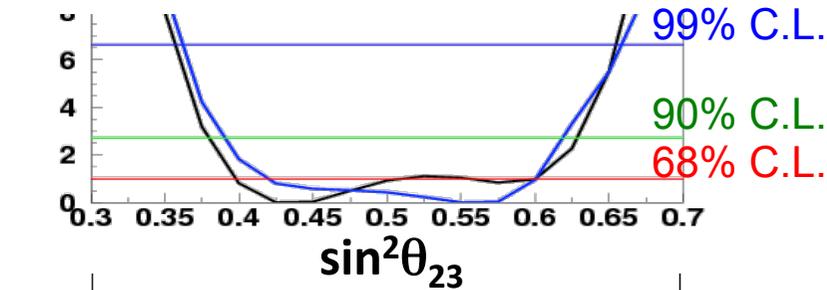
Current situation of atmospheric ν θ_{23} Octant

Super-Kamiokande atmospheric neutrino 3 flavor analysis

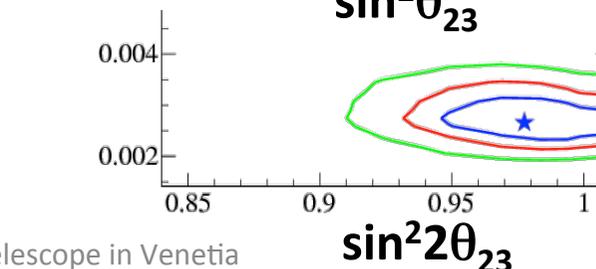
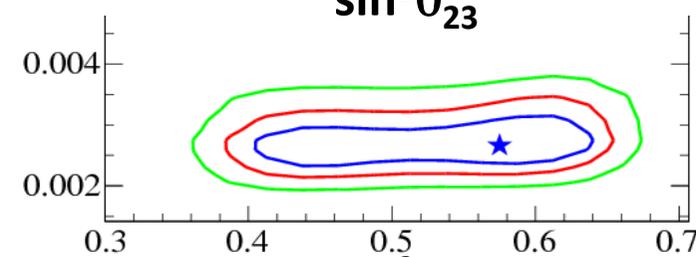
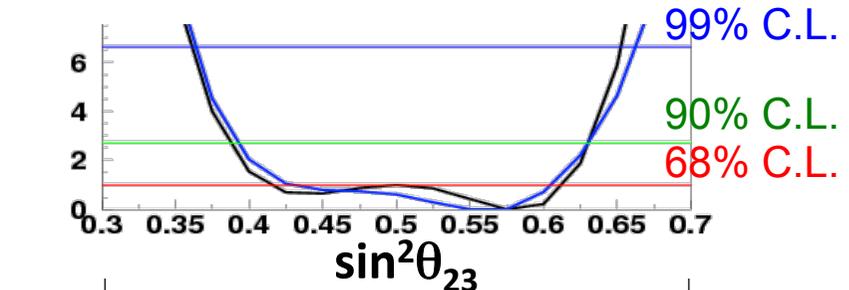
- θ_{13} free in the fitting
- θ_{13} fixed at the best value

*We may start to see
 1σ level effect??*

- Normal Mass Hierarchy
- Best fit value: 0.425



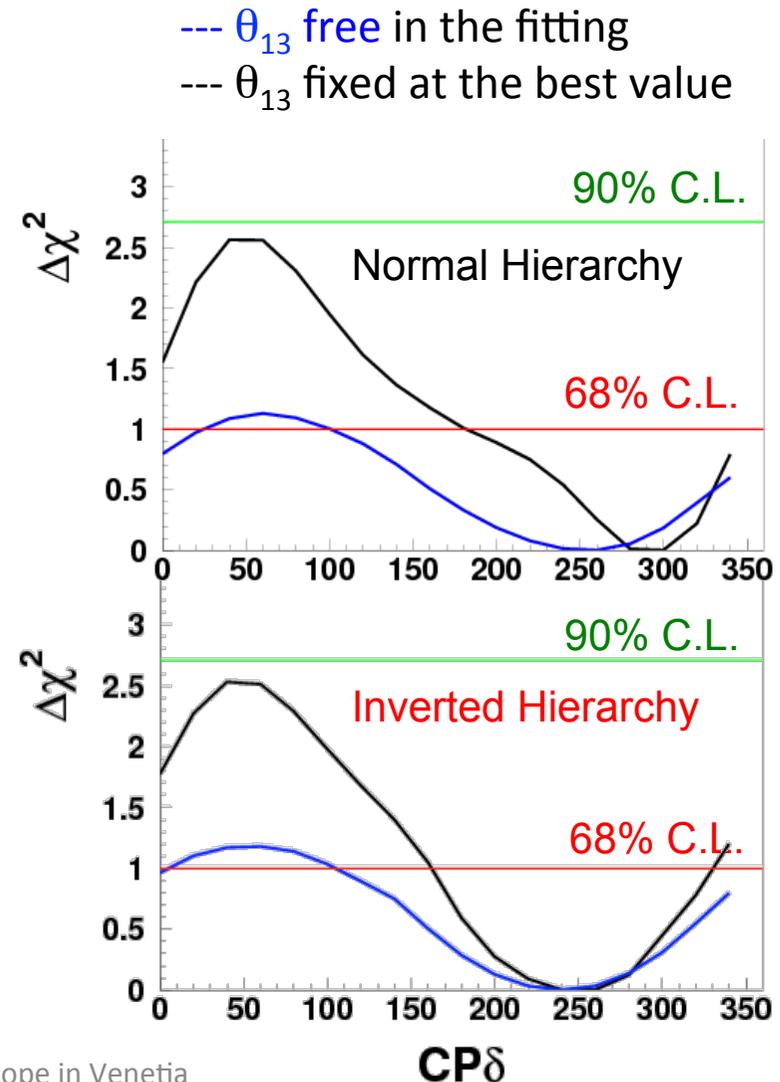
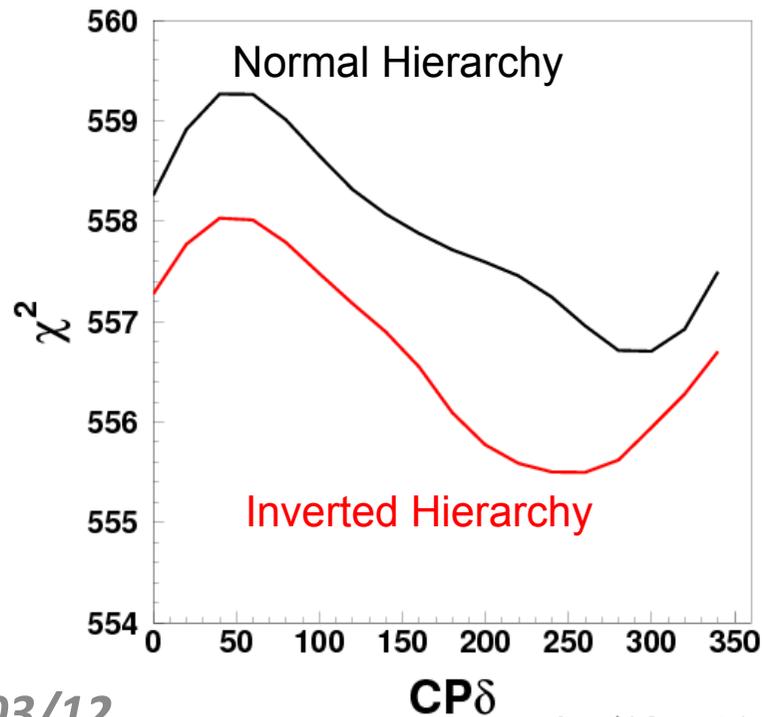
- Inverted Mass Hierarchy
- Best fit value: 0.575



Current situation of atmospheric ν Mass Hierarchy and CP phase

- There may be a hint in Atm ν (SK)
 - NH: $\chi^2_{\min} = 556.7 / 477$ dof
 - IH : $\chi^2_{\min} = 555.5 / 477$ dof

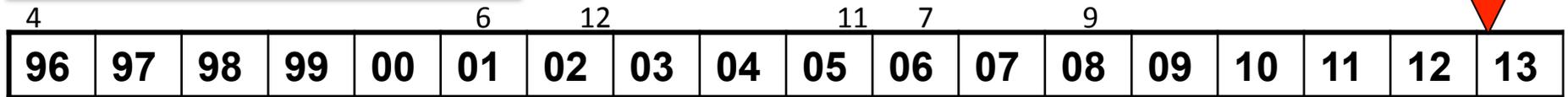
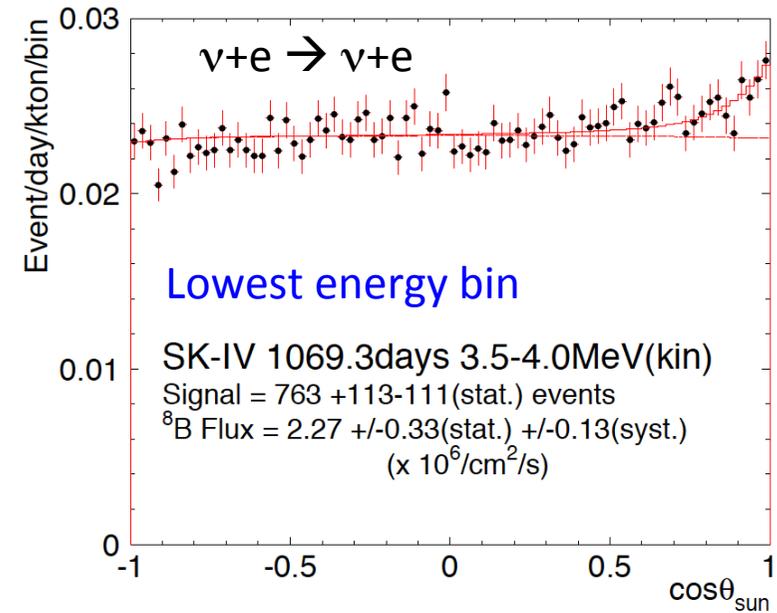
$$\chi^2_{\min}(\text{NH}) - \chi^2_{\min}(\text{IH}) = 1.2$$



- Details and other subjects of atmospheric neutrinos will be discussed by Chris in this afternoon

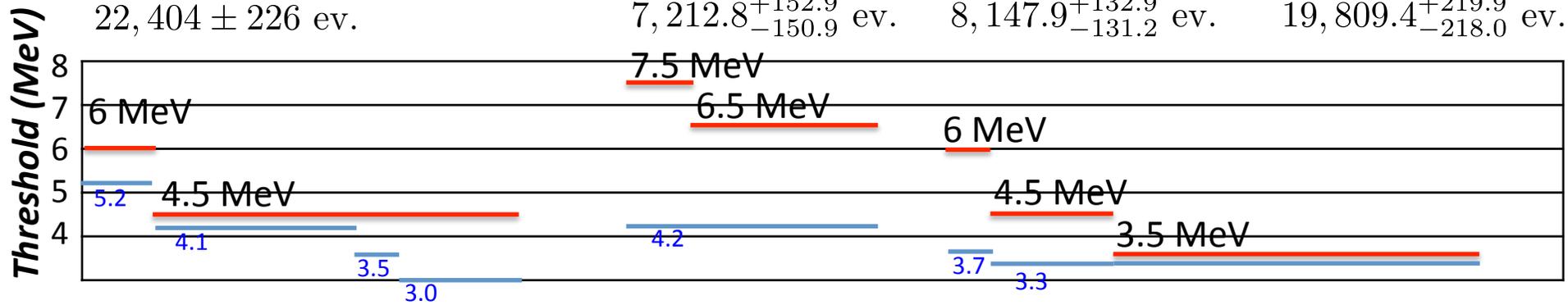
Solar neutrinos

- Total 3904 days
 - 57,574.1 solar neutrino events
- Analysis threshold: down to 3.5 MeV
 - Possible to 3.0 MeV in near future
- Fiducial volume
 - 22.5 kt (> 5.0 MeV)
 - 13.3 kt (4.5-5.0 MeV)
 - 8.8 kt (3.5-4.5 MeV)



SK-I	SK-II <i>~Half PMTs</i>	SK-III	SK-IV
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1496 days $22,404 \pm 226$ ev.	791 days $7,212.8^{+152.9}_{-150.9}$ ev.	547.9 days $8,147.9^{+132.9}_{-131.2}$ ev.	1069.3 days $19,809.4^{+219.9}_{-218.0}$ ev.
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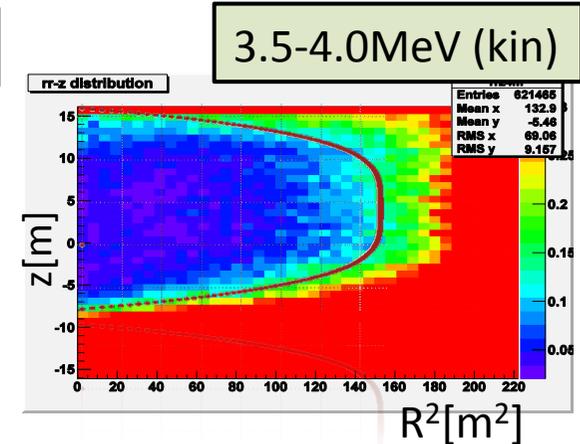
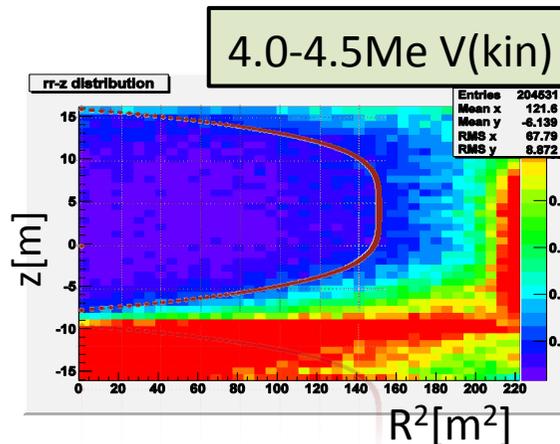
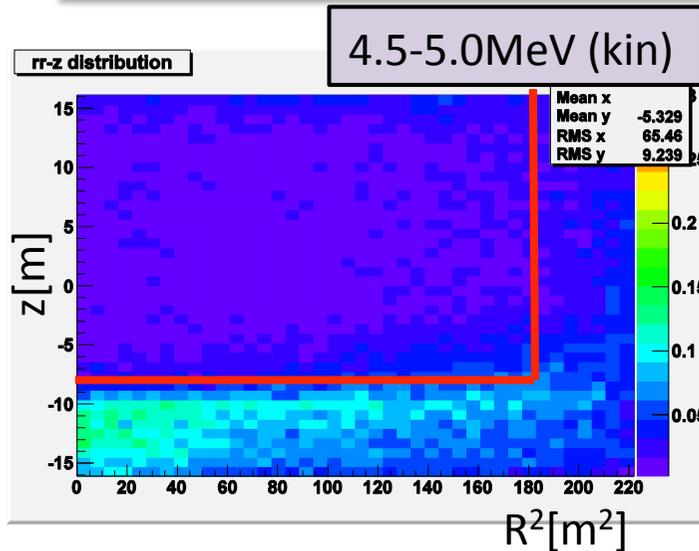
To reduce backgrounds

Tight fiducial volume cut

- Usual fiducial volume (> 5.0 MeV): 22.5 kt
- Need a limited fiducial volume for low energy to reduce backgrounds

- 4.5-5.0MeV(kin)
 - $z > -7.5$ [m], $R^2 < 180$ [m²]
 - ➔ 13.3 kt

- 3.5-4.5MeV(kin)
 - Use z dependent, but vertically symmetric fiducial cut
 - ➔ 8.8kt



- Keep R_n at the lower level of the water tank
 - Need to avoid convection of the inner water
 - ✓ make laminar flow ← temperature control of input water is important
- **Make Acrylic vessel to prevent R_n from sneaking into the fiducial volume**

Solar neutrinos

Flux measurements

- Many Improvement for the last few years (systematic errors)
 - Total 3.5% \rightarrow 1.7%
 - Fiducial volume: 1.3% \rightarrow 0.17%
 - Energy scale: 0.64% \rightarrow 0.54%
 - Others

Global Analysis (fixed: $\sin^2\theta_{13} = 0.025$)

Solar Global (SK Analysis)

$$\sin^2 \theta_{12} = 0.310^{+0.014}_{-0.015}$$

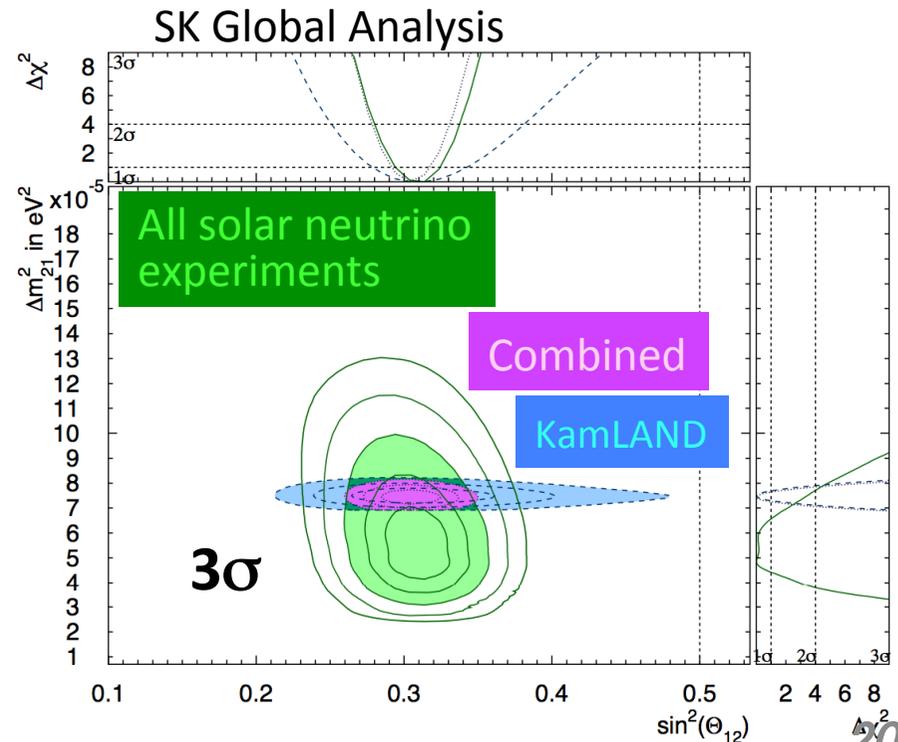
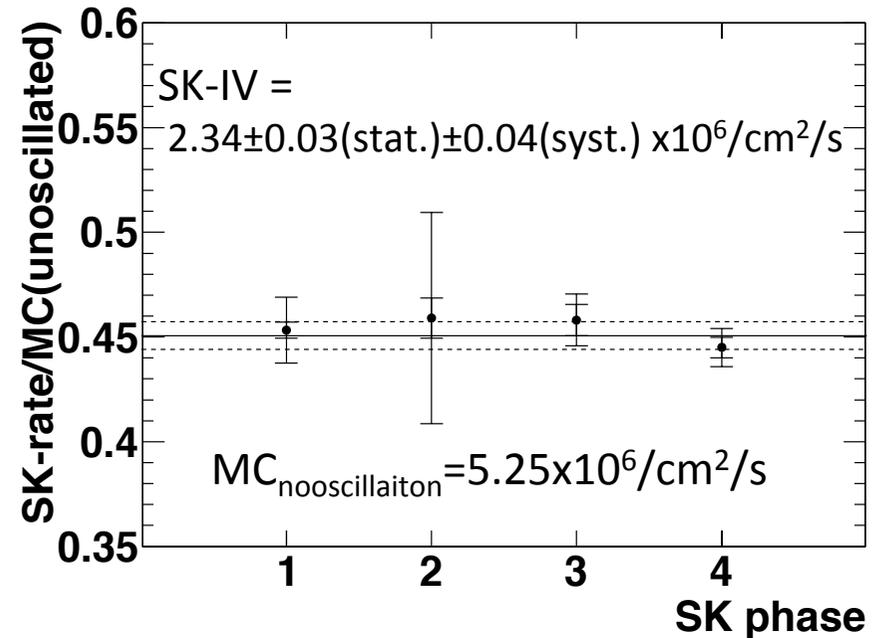
$$\Delta m_{12}^2 = 4.86^{+1.44}_{-0.52} \times 10^{-5} eV^2$$

KamLAND

$$\sin^2 \theta_{12} = 0.309^{+0.039}_{-0.029}$$

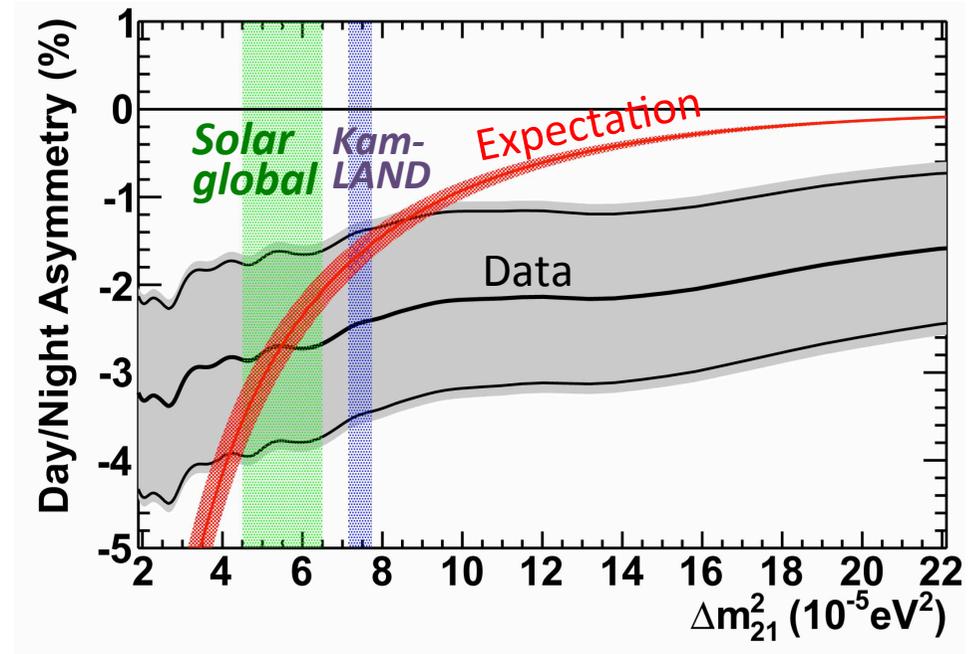
$$\Delta m_{12}^2 = 7.49^{+0.20}_{-0.19} \times 10^{-5} eV^2$$

- 1.5 σ difference between **KamLAND** and **solar neutrino experiments** in Δm_{12}^2



Day/Night Asymmetry

- Regeneration of ν_e through the earth:
- $A_{DN} = 2(D-N)/(D+N)$
~2~3% effect
- Night-time spectrum variation depends on $\Delta m_{12} \rightarrow$ Results (data) depend on Δm_{12}
- Δm_{12} from D/N analysis agree better with that from the solar global analysis

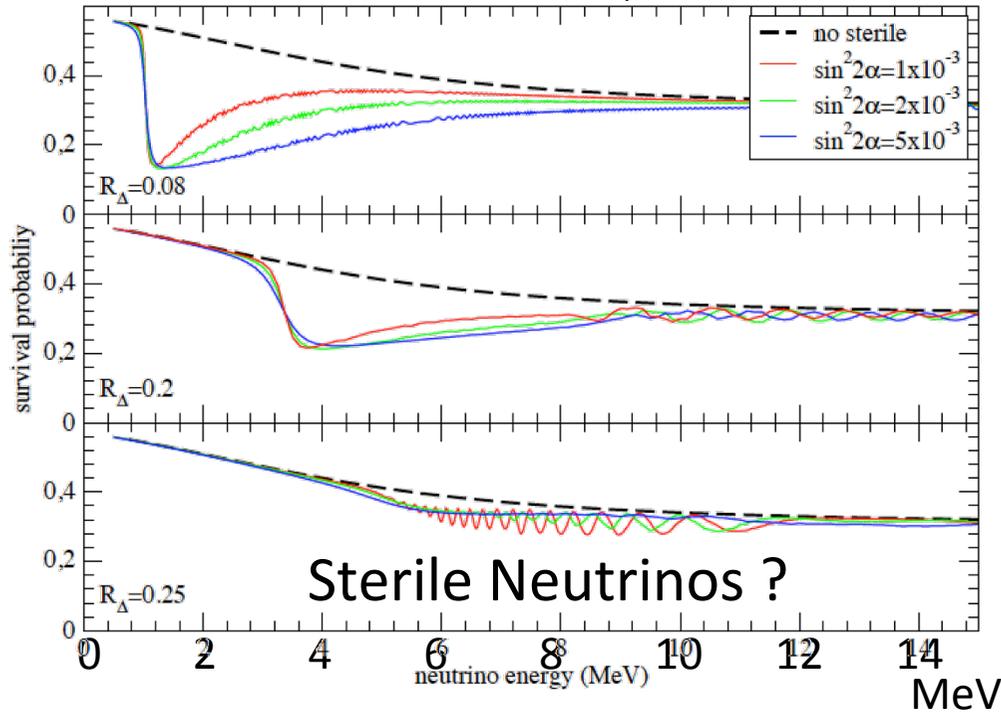


- SK-I, II, III, IV data
- $A_{ND} = -2.8 \pm 1.1 \pm 0.5\%$
 $\rightarrow 2.3 \sigma$

Solar neutrinos

- **No observation of up-turn yet**

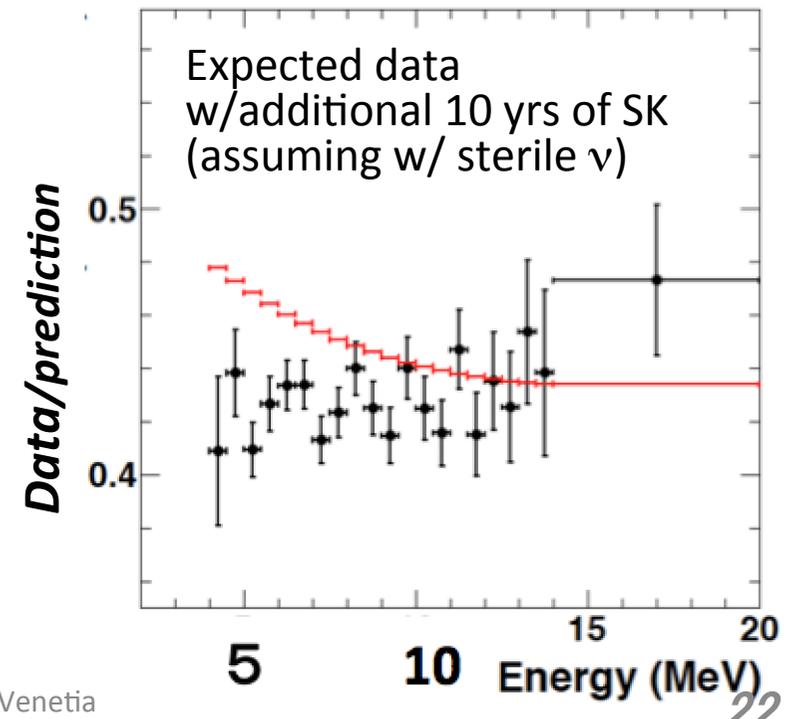
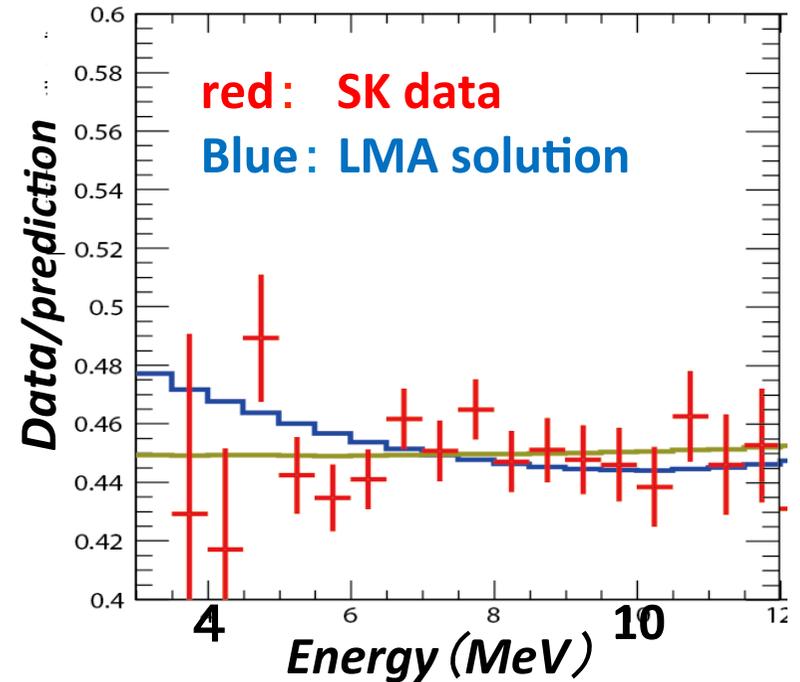
Holanda and Smirnov, arXiv: 1012.5627v2



- Upturn [remaining problem]
 - Lower energy threshold
 - Remove BG
- Borexino threshold: 3.0 MeV
- SK threshold: 3.5 keV

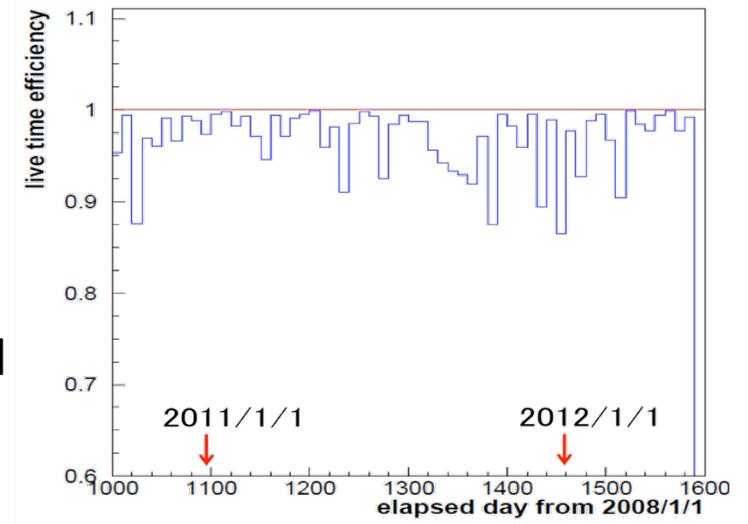
2013/03/12 aim to 3.0 MeV soon.

Y.Suzuki@NeutrinoTelescope in Venetia



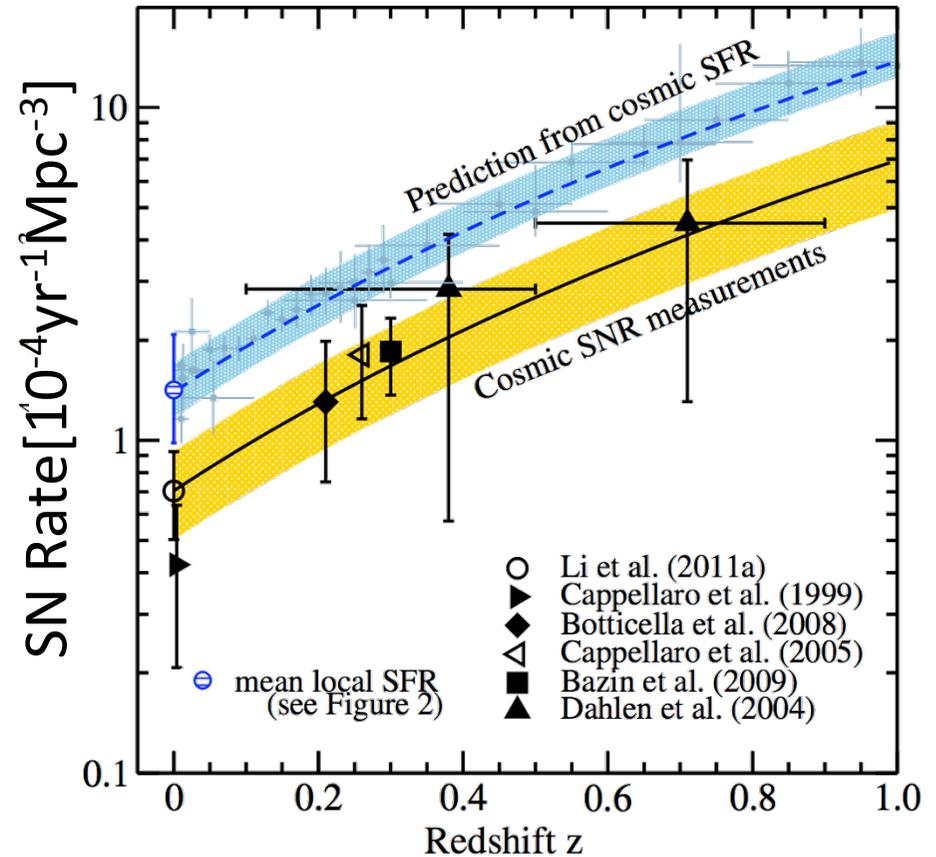
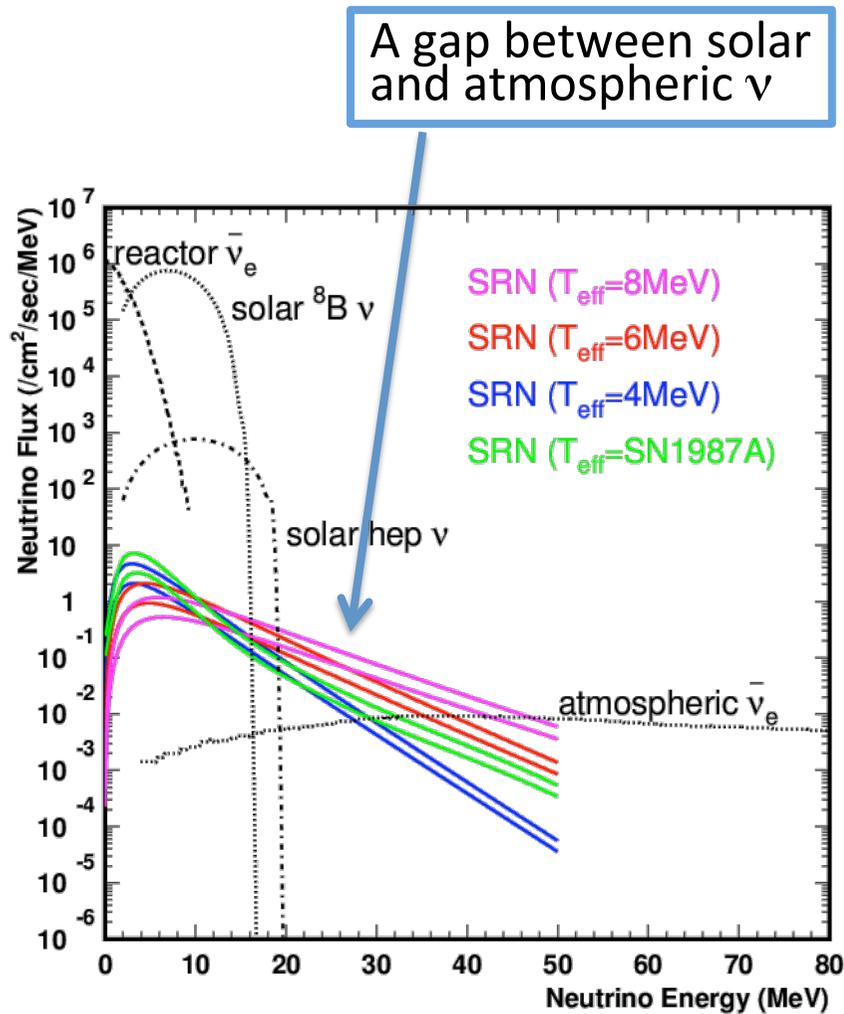
Supernova Neutrinos

- Burst search
 - Nearby SN → Detailed Explosion Mechanism
 - Additional information for Neutrino Oscillation
 - Mostly $\bar{\nu}_e + p \rightarrow e^+ + n$ and $\nu_x e \rightarrow \nu_x e$ interactions
 - Event by event reconstruction, time, energy and directionality(only for $\nu_x e$)
 - Expect 8000 neutrino events from the SN at 10 kpc
 - 2×10^7 neutrinos for Betelgeuse (640 light years)
 - Currently upgrading the electronics; 1) Sparse data taking, 2) energy flow
- SNWATCH: Continuous data taking
 - Minimizing dead time
 - Less down time for calibration
- SNEWS (Supernova Early Warning System)
 - Neutrino arrives 20-40 hours before the optical observation for Betelgeuse
 - Early warning to the observatories world wide
- We are preparing for the next Galactic Supernova



SRN(Supernova Relic Neutrinos)

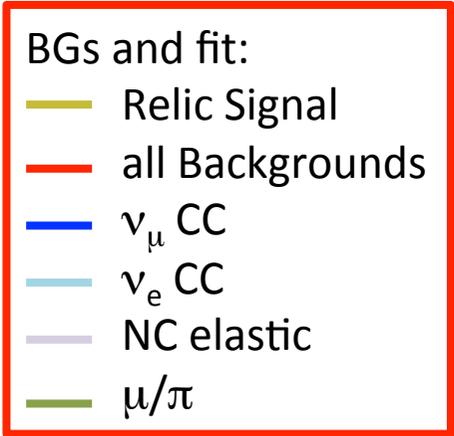
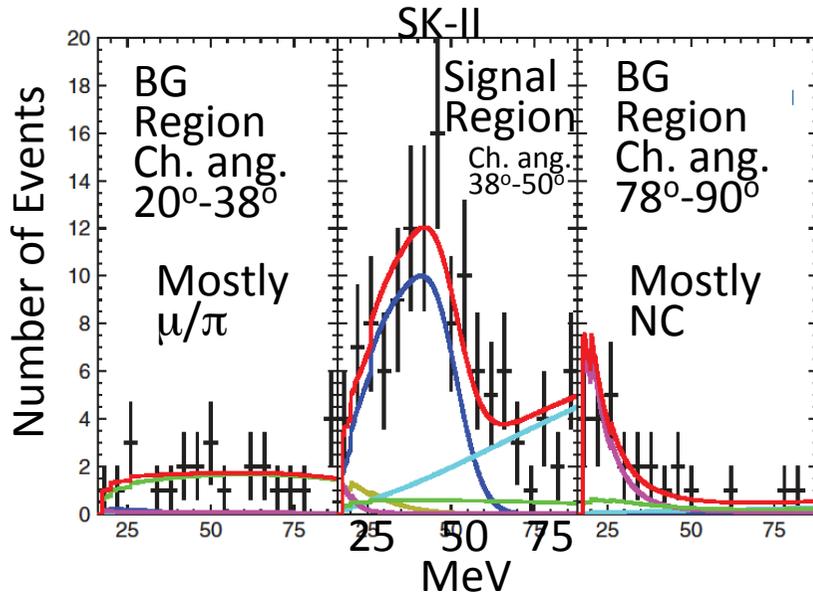
the diffuse supernova neutrino background from all the supernovae in the past



SRN carries information about Star Formation Rate (SFR), Initial Mass Functions (IMF) and so on.

SNRate: factor 2 mismatch
pred. SFR \leftrightarrow meas. SNRate

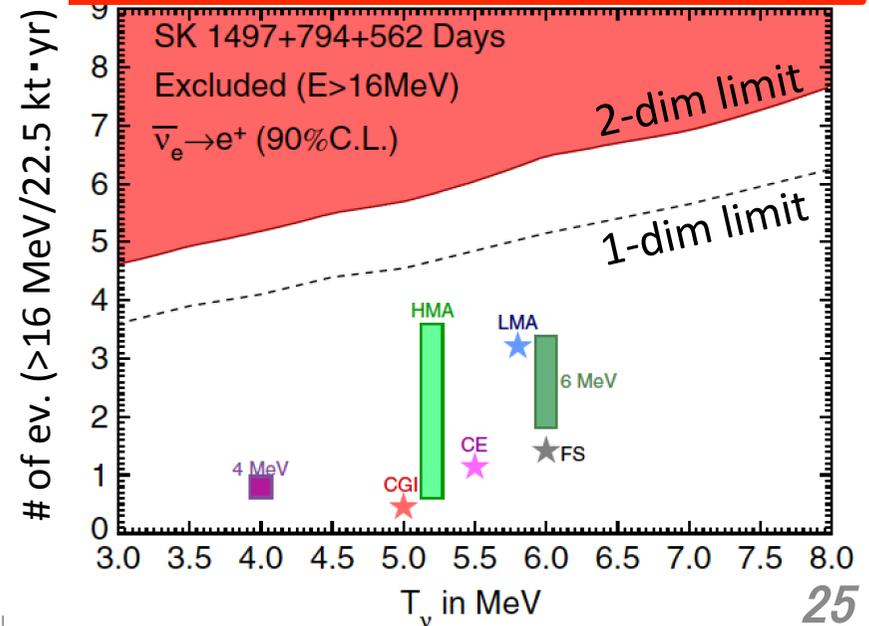
SK result on SRN

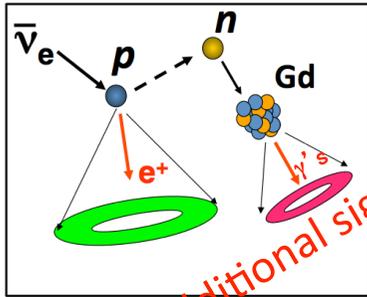


- μ/π : heavier and low energy
→ small Cherenkov angle events.
- NC elastic: multiple γ 's (Isotropic events).
→ 'Large Ch.' angle

- Improvement
 - Reduce spallation BG
 - Improved the muon/spallation correlation
 - Select 16-30MeV Interval
 - Lower energy threshold by 2 MeV
 - Increase efficiency by 20%
- Two new BG channels are taken
 - Multiple region of Cherenkov angle distributions for the evaluation

Flux limit: 2.8-3.1 $\nu_e/\text{cm}^2/\text{s}^2$ (>16 MeV)





Future Project GadZOOKs!

- To detect SRN
 - Need to reduce BG
- Gadzooks: to identify neutrons in

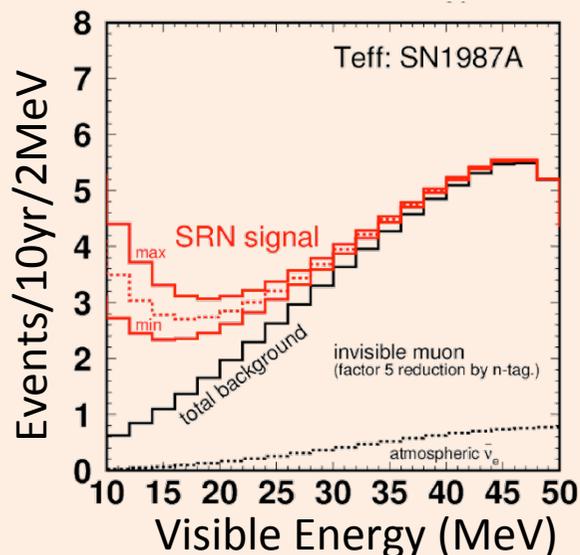
$$\bar{\nu}_e + p \rightarrow e^+ + n$$
 - Coincidence of e^+ and n
- 0.2% Gd sulfate
 - 49,000 barns (5 order larger than p) for thermal cap.
 - γ cascade of 8 MeV
 - $\Delta t \sim 20 \mu s$

- In summer, the 240 PMTs will be mounted
- Full test w/ the PMT will be done in Fall, 2013.



- R&D: EGADS (2009 ~): Test Facility
 - 200 ton main tank
 - Selective filtration system
 - Cleaning unwanted impurities
 - Keep Gd in the water
 - Transparency measurement
- Currently Gd doped water is circulated by increasing the Gd concentration gradually

Merits of GadZOOKs



Supernova Relic Neutrinos

- Assume 1/5 reduction of the backgrounds (efficiency for neutron capture: 90% and cut efficiency 74%)
- ➔ We expect 20~40 SRN events between 10 and 30 MeV for $T_{\text{eff}} = \text{SN1987A}$
- ➔ More events for larger T_{eff}

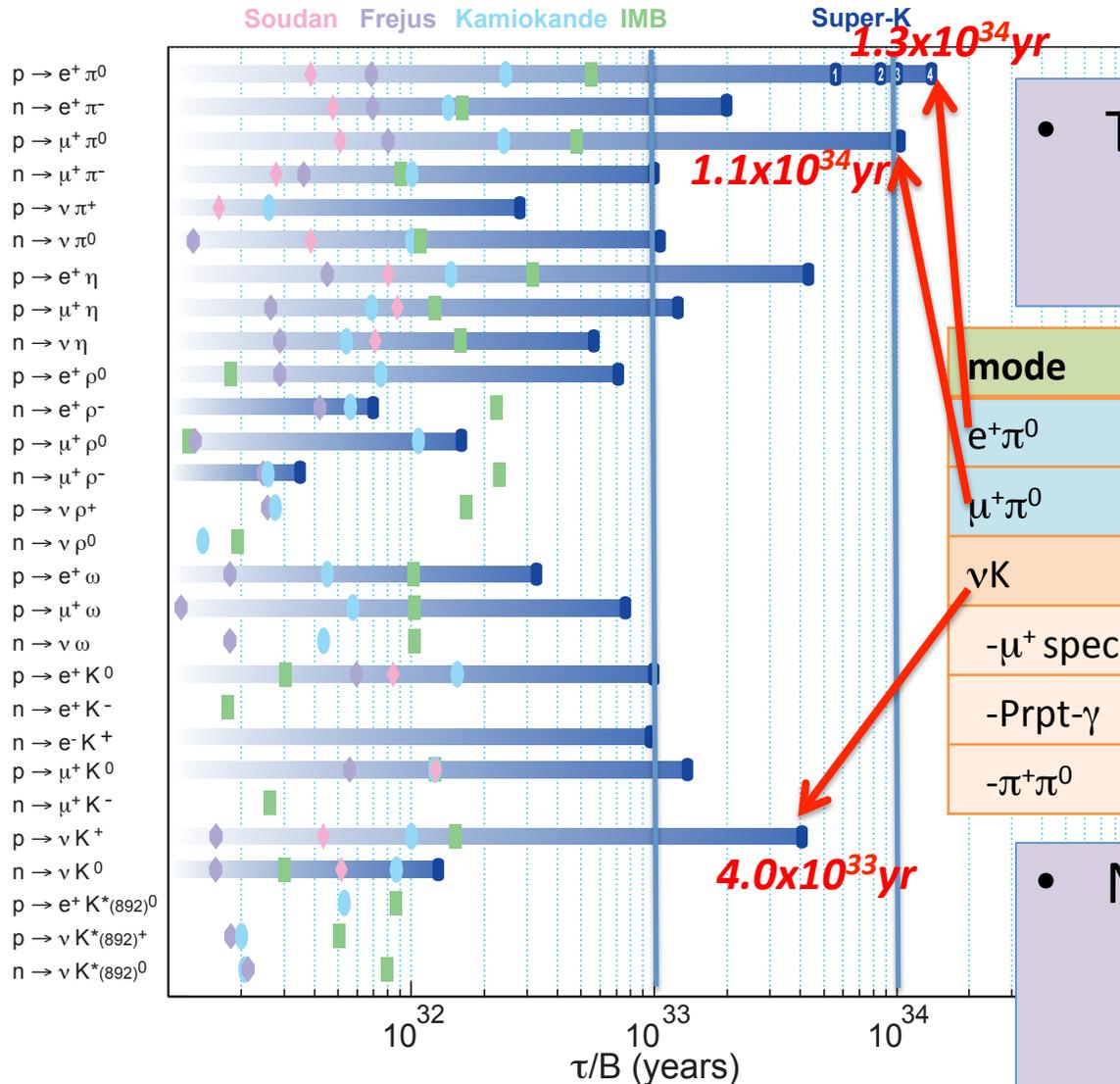
A. Odrzywolek et al., astro-ph/0311012

*Si-burning
for $20x M_{\odot}$*

Burnig Phase	T_c [MeV]	L_{ν} [erg/s]	Duration	Total energy [erg]	Average ν energy [MeV]	Ev. Rate for SK day-1
C	0.07	7.4×10^{39}	300 yrs	7×10^{49}	0.71	
Ne	0.146	1.2×10^{43}	140 days	1.4×10^{50}	0.99	
O	0.181	7.4×10^{43}	180 days	1.2×10^{51}	1.13	
Si	0.319	3.1×10^{45}	2 days	5.4×10^{50}	1.85	41 at 1 kpc

- For Betelgeuse (0.2 kpc), expect 1000 events / day
 - difficult to detect e^+ from $\bar{\nu}_e + p \rightarrow e^+ + n$, but
 - easy to observe 8 MeV γ 's from neutron capture

Nucleon Decay

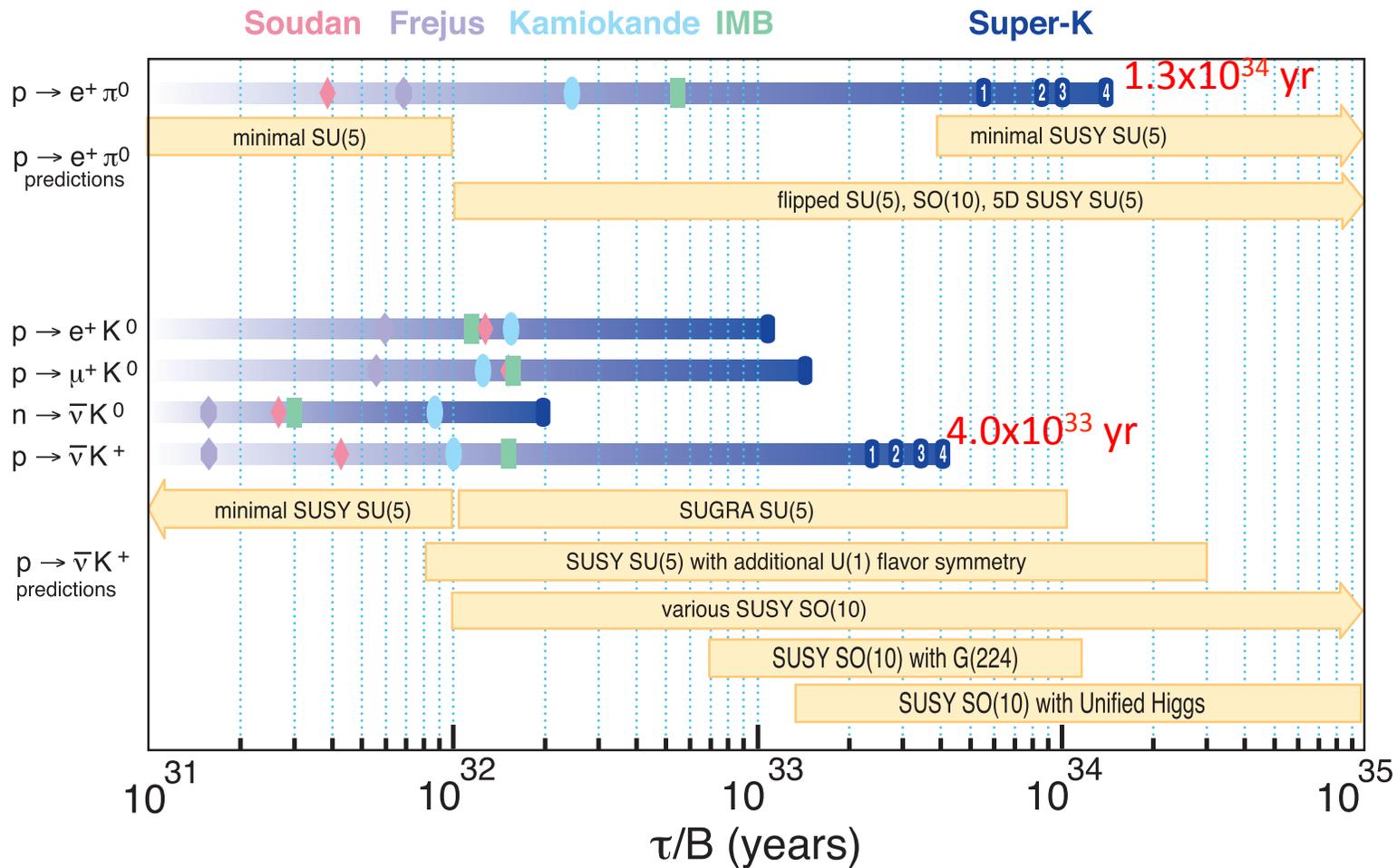


- The largest detector
 → 220kt·yr maximum exposure
 → No nucleon decay indications !

mode	selection	εBr (%)	obs	BG	physics
$e^+\pi^0$	P_{tot}, m_p	44-45	0	0.5	$\tau \sim M_x^4$
$\mu^+\pi^0$	P_{tot}, m_p	35-44	0	0.6	$\tau \sim M_x^4$
νK					SUSY-G
$-\mu^+$ spec	p_μ	36-44	*	*	
$-Prpt-\gamma$	$p_\mu, E\gamma, \Delta t$	6-8	0	0.4	
$-\pi^+\pi^0$	E_{π^0}, π^+ id.	5-8	0	1.2	

- New electronics (SK-IV)
 - Increase the efficiency for Michel electrons
 - Improved efficiency for $\mu^+\pi^0, \nu K^+$

Nucleon Decay



- **Background level < 1 in many decay modes**
 - Sensitivity: proportional to the exposure in future
- **We continue to search for proton decay**

Summary

- After the discovery of the neutrino oscillations we have further developed the oscillation study.
 - Precisely determined oscillation parameters (atmospheric and solar neutrinos)
 - Obtained evidence for tau appearance and tested CPT in atmospheric ν
 - Obtained positive indication of Day/Night effect (solar ν)
 - Made many efforts to reduce low energy backgrounds
 - Improved the machinery for detecting nearby SN burst
 - Mostly finished the the feasibility study of GadZOOKs

17 year old Super-K is still very active

- We will continue data taking at least for the next ten years
 - to study MH, CPV, upturn
 - to observe SN Relic neutrinos
 - to look for neutrino bursts from SNe and protons to decay
- We hope that we will hand them to Hyper-Kamiokande in very near future