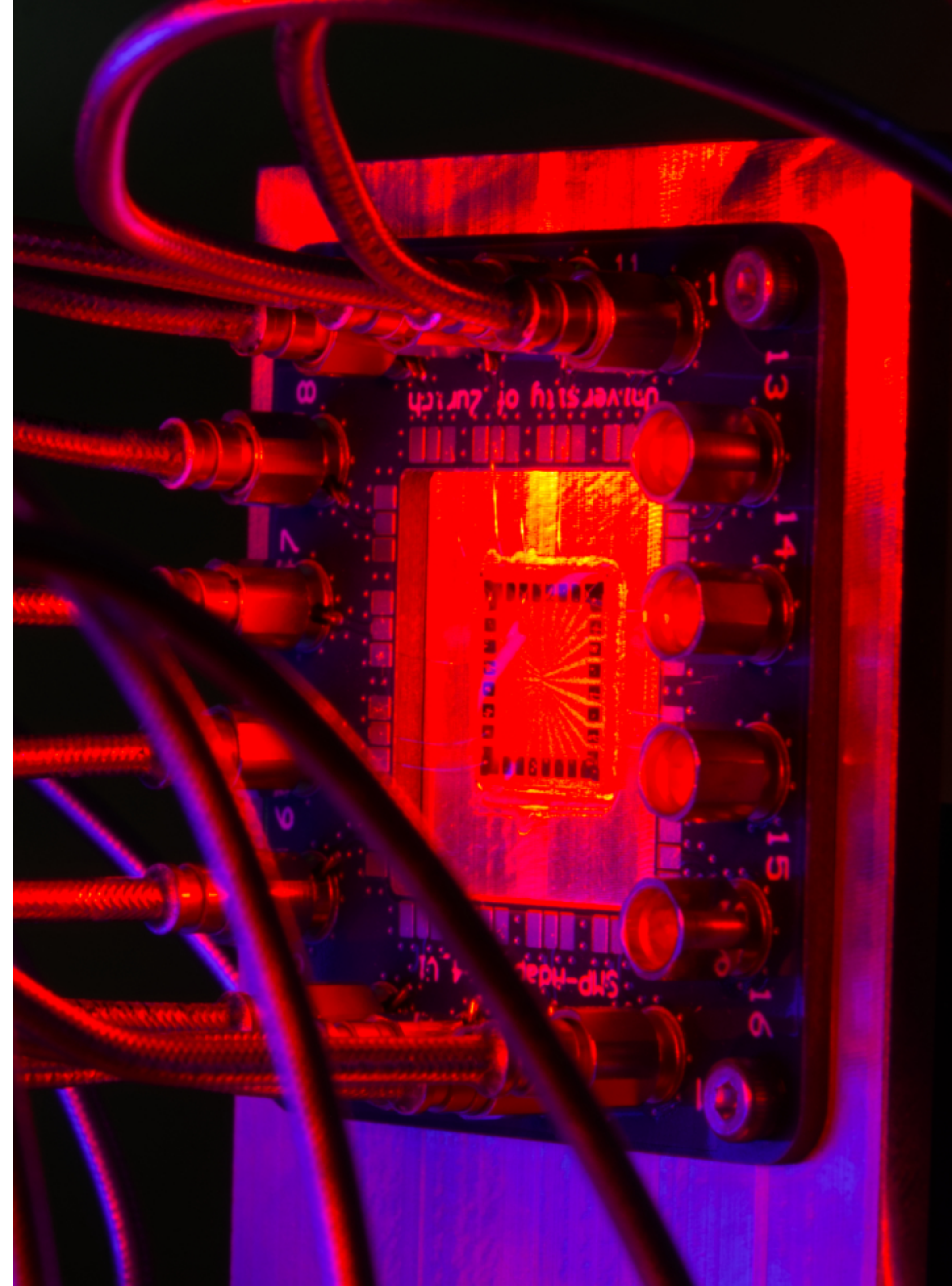


Search for DM with SNSPD: the QROCODILE experiment



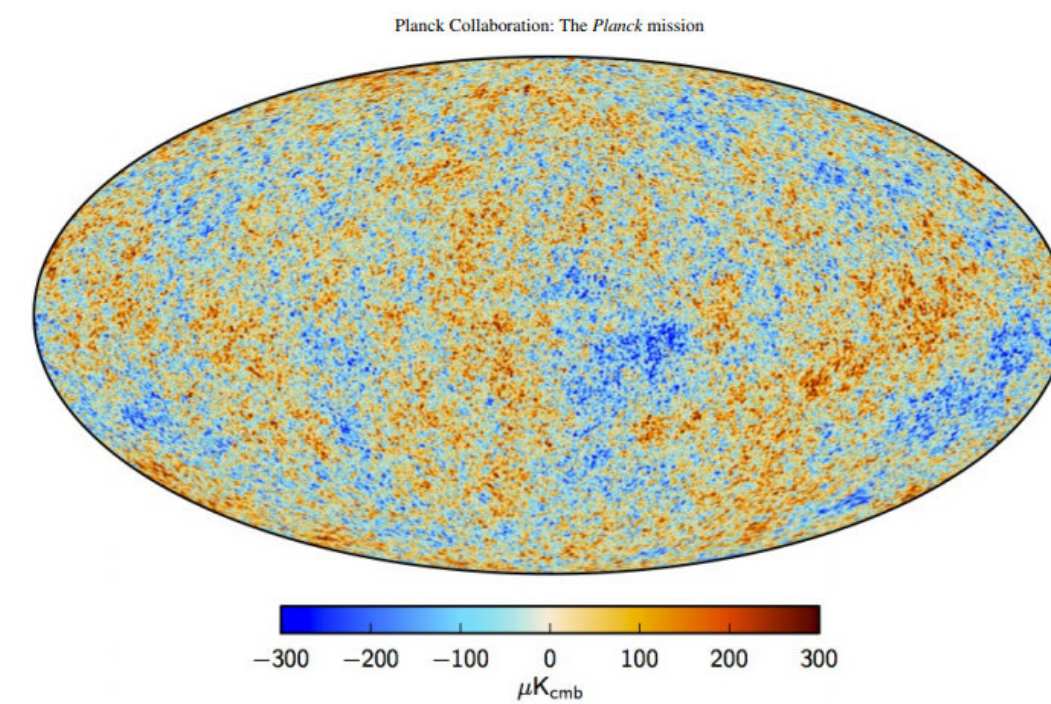
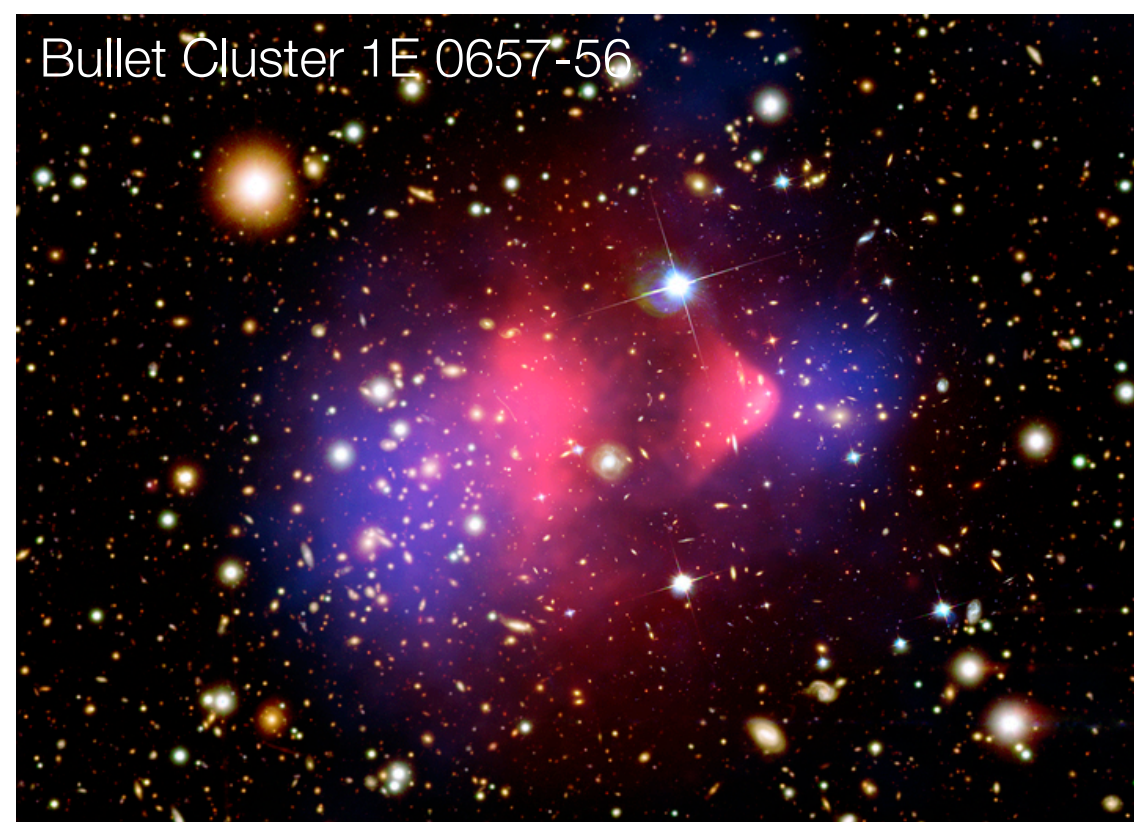
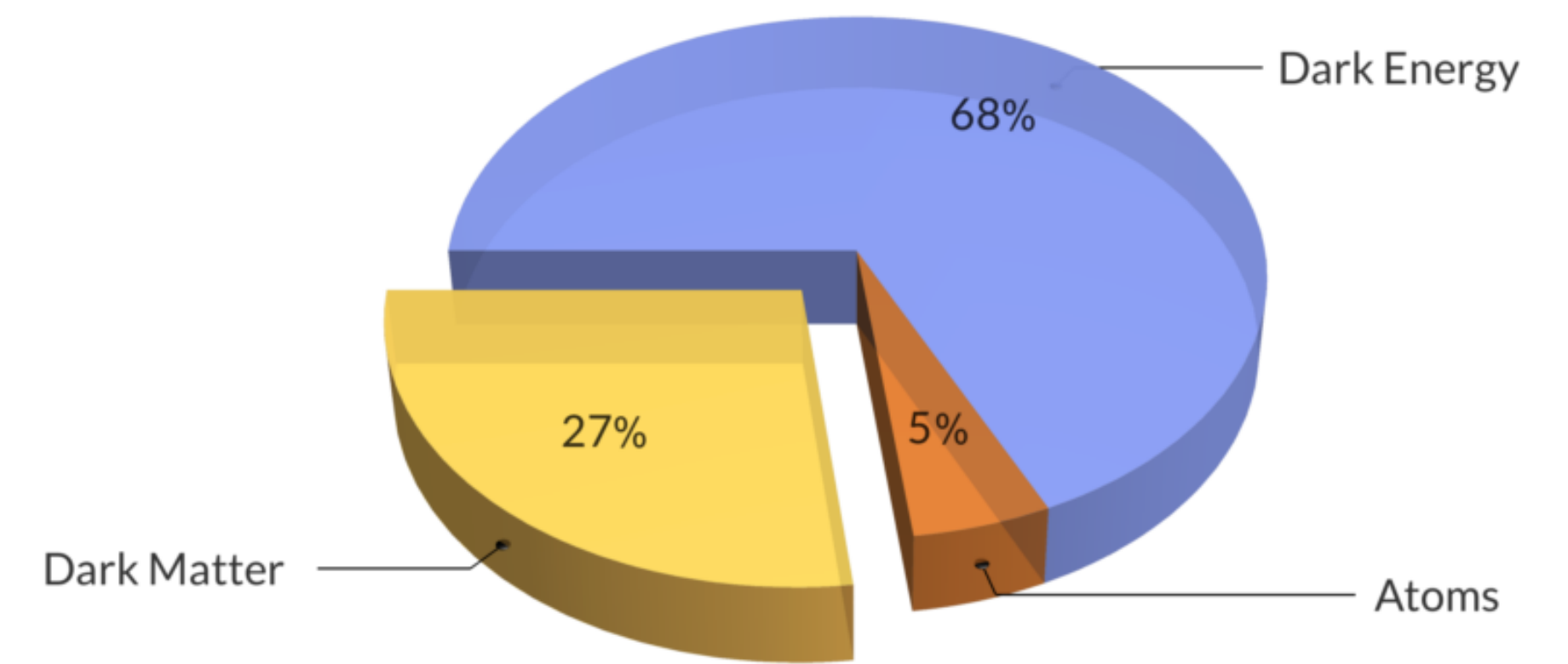
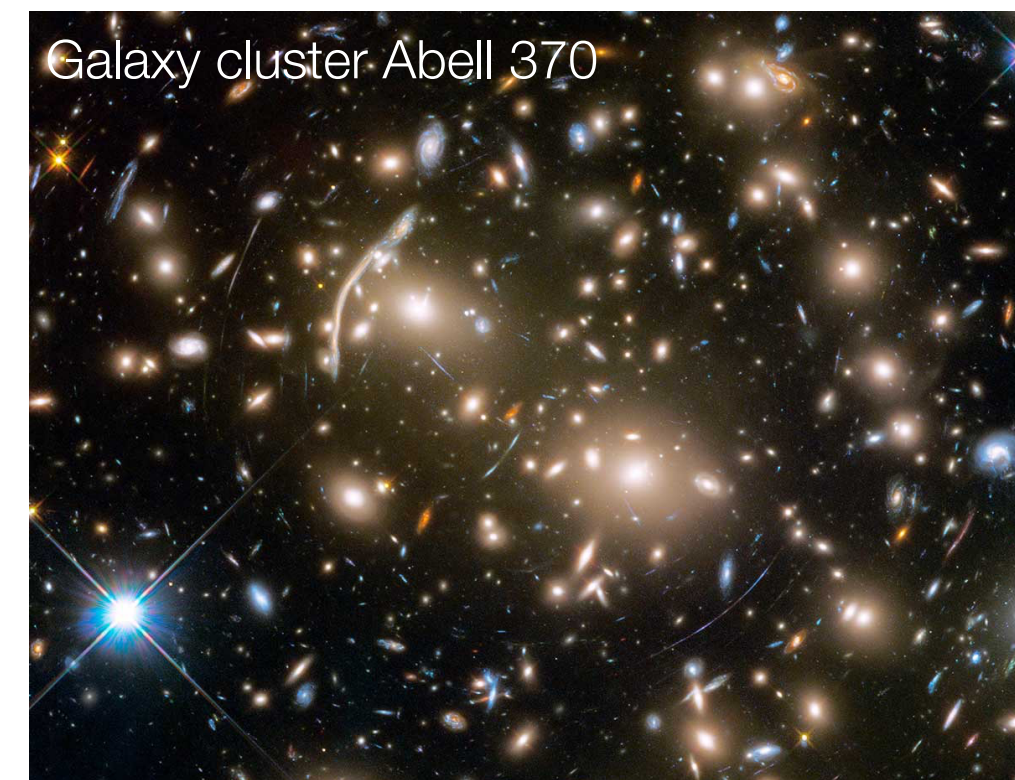
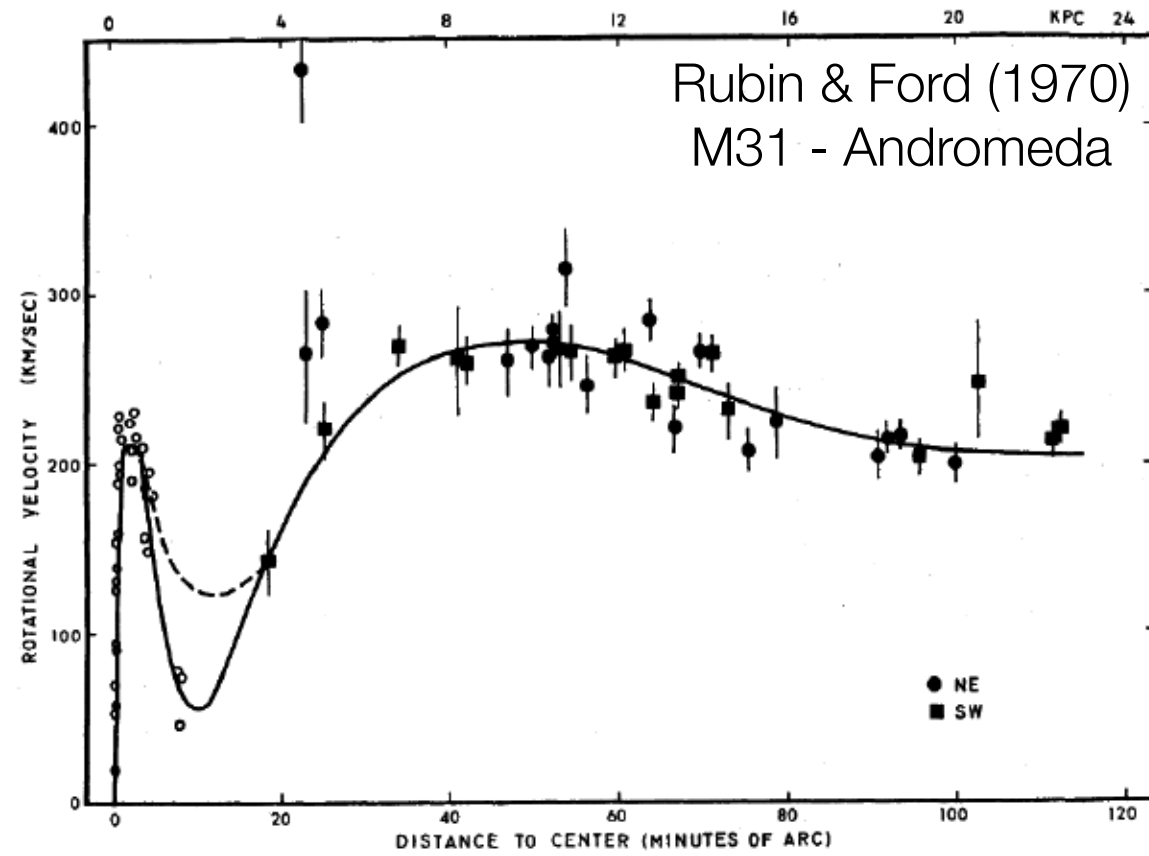
Chiara Capelli (University of Zurich)
On behalf of the QROCODILE collaboration

Superconducting Technologies for Dark Matter
Roma
May 11-13, 2026



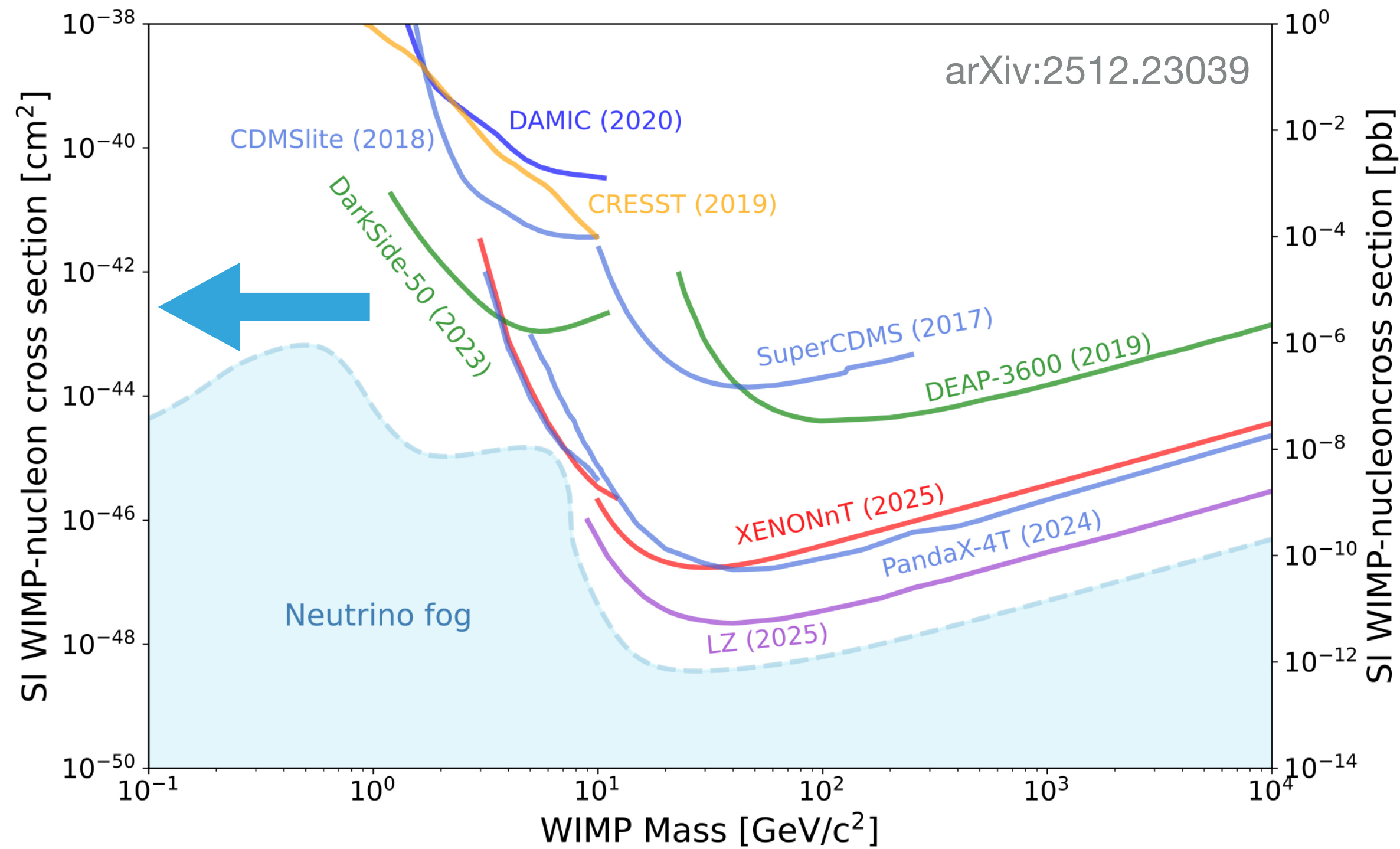
Dark matter

- Several cosmological observations from gravitational interaction (galaxies velocity dispersion, star rotation curves, gravitational lensing, galaxy clusters, CMB)



- DM constitutes ~84% of the matter in the Universe
- Non baryonic
- Non relativistic (cold)
- Electrically neutral
- Neither absorbing nor emitting light

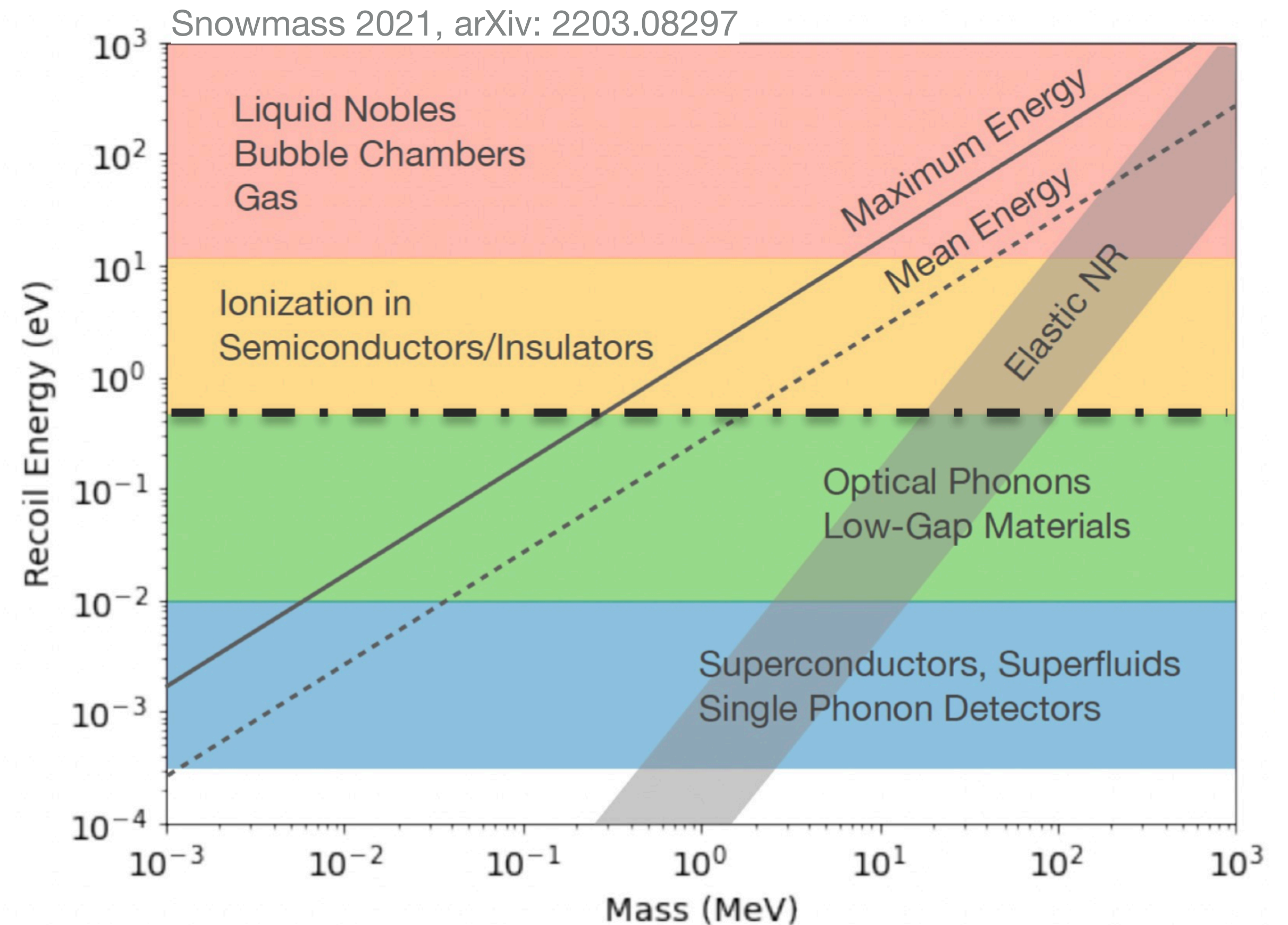
Low mass DM scenario



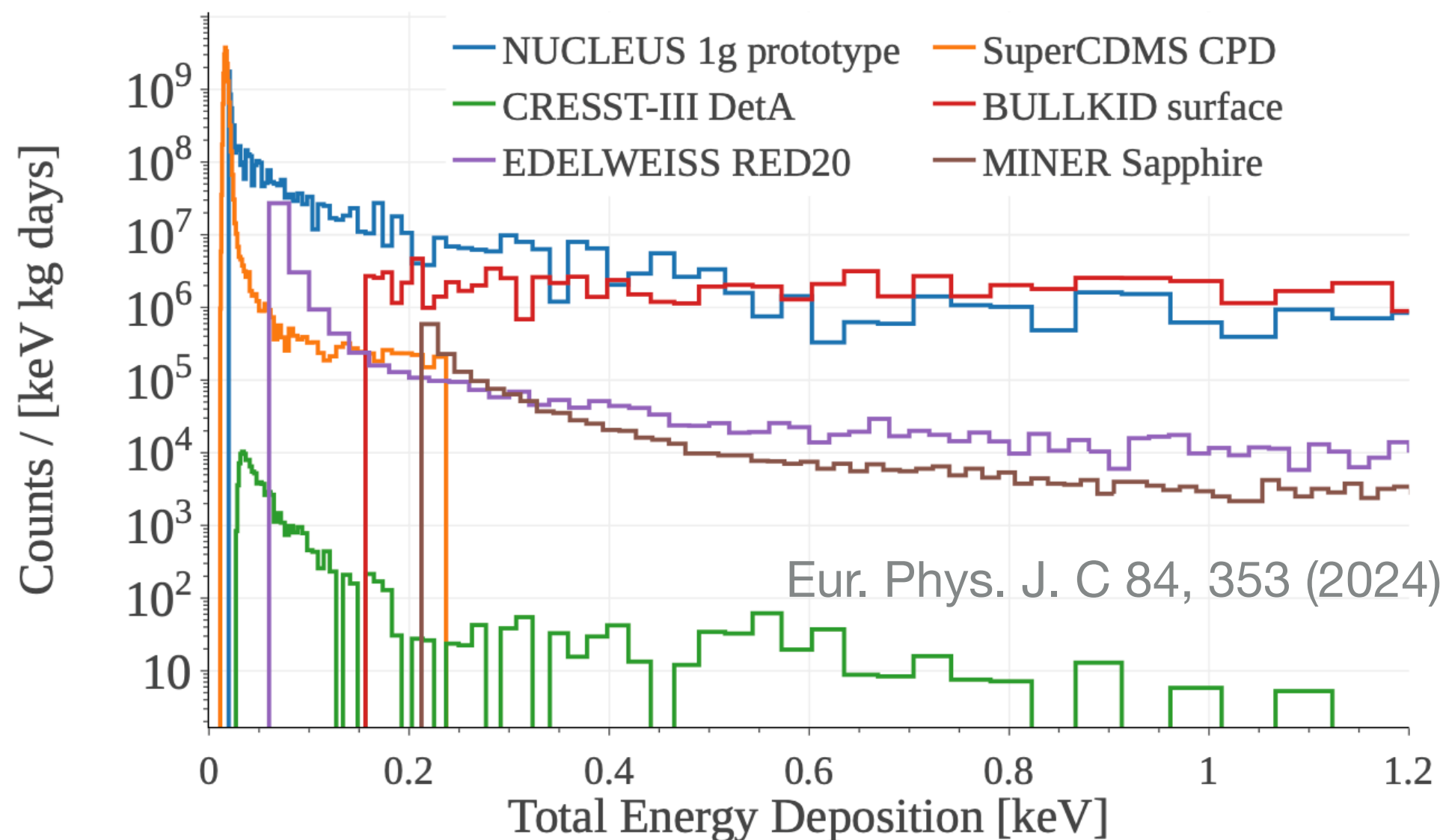
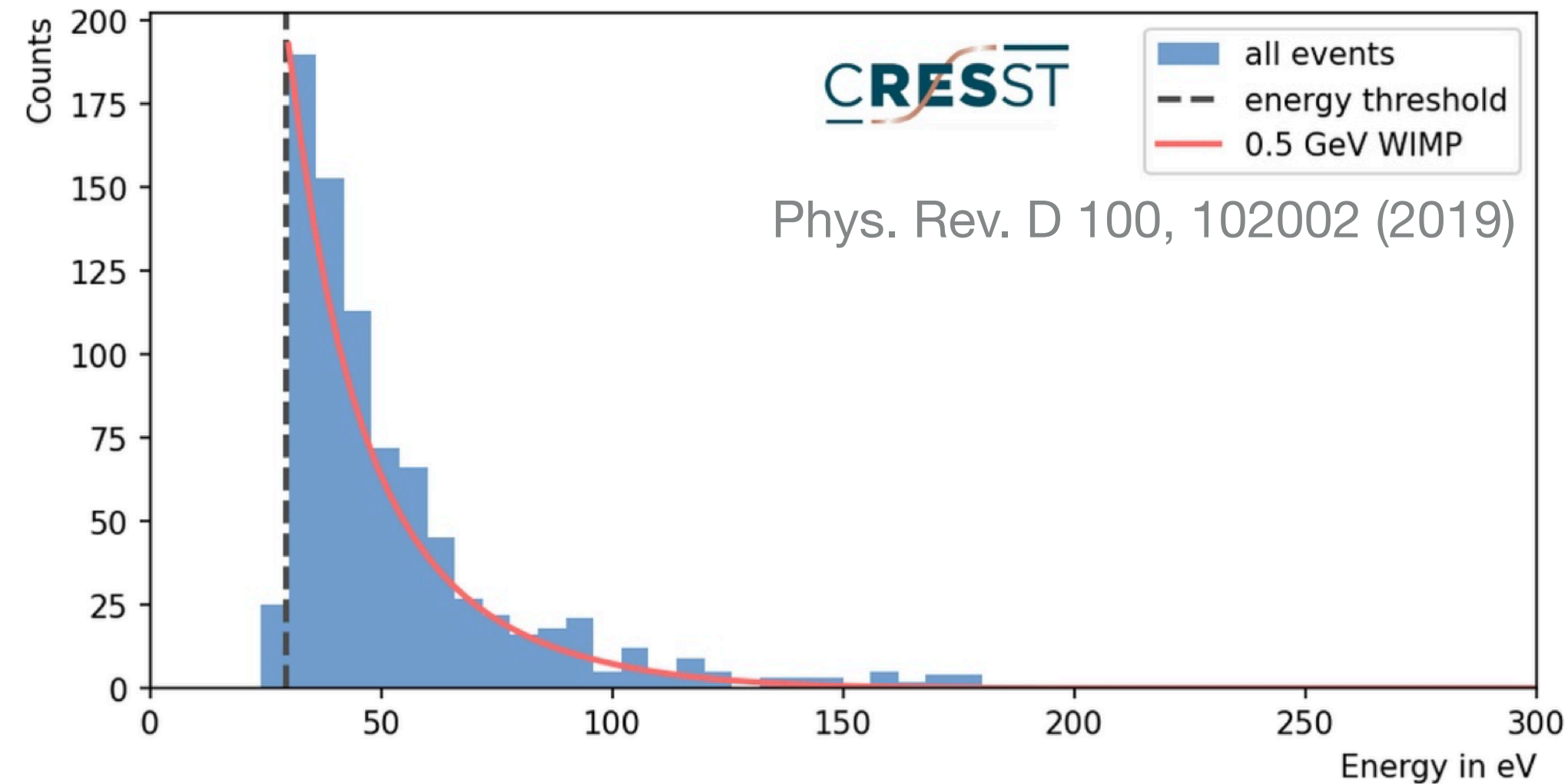
- Development of new technologies required to explore lower masses
- For sub-MeV masses needed sub-eV threshold

Max. Energy for DM moving at escape velocity $\sim 544 \text{ km/s}$

Mean energy for DM moving at mean velocity $\sim 220 \text{ km/s}$



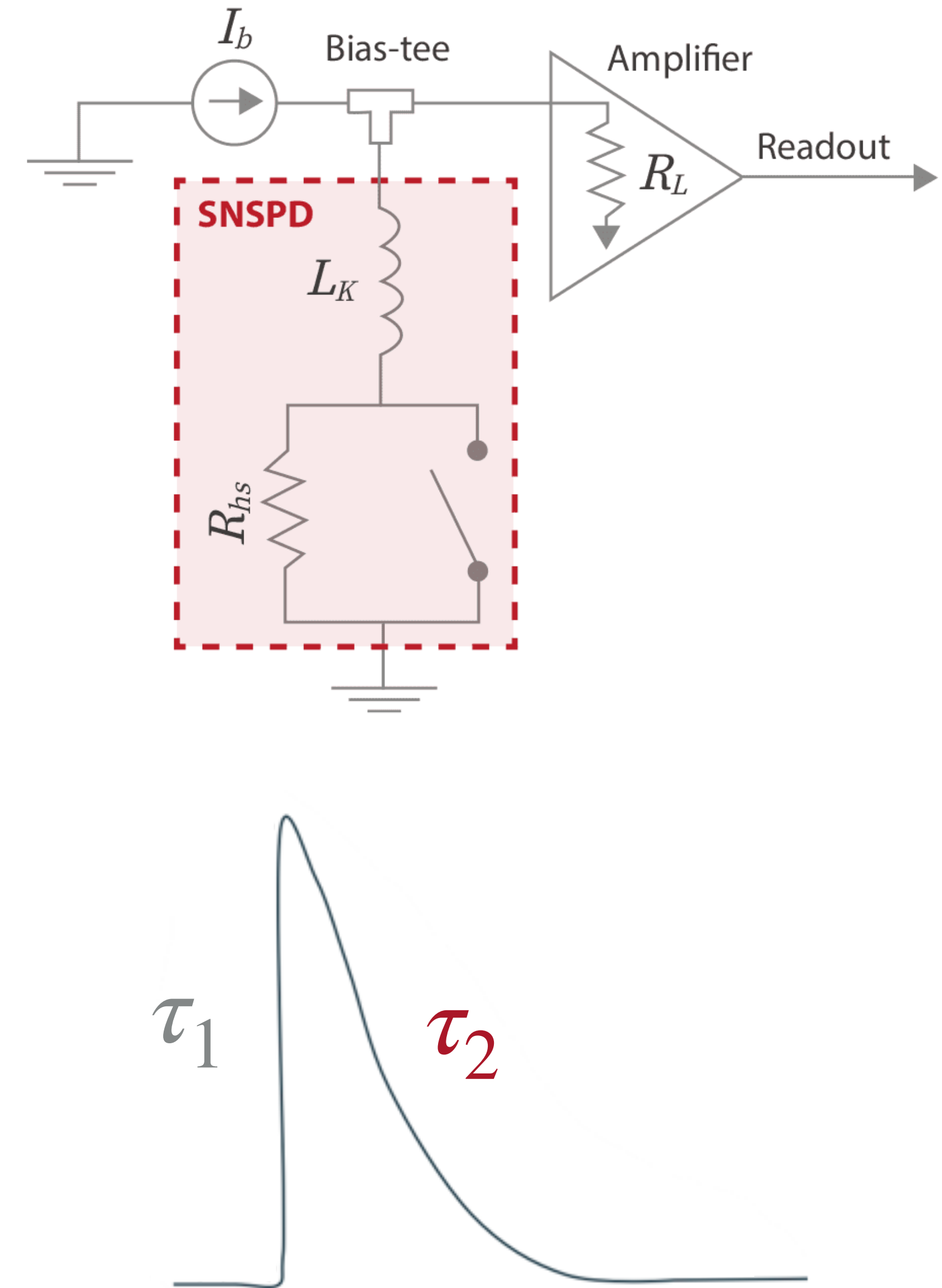
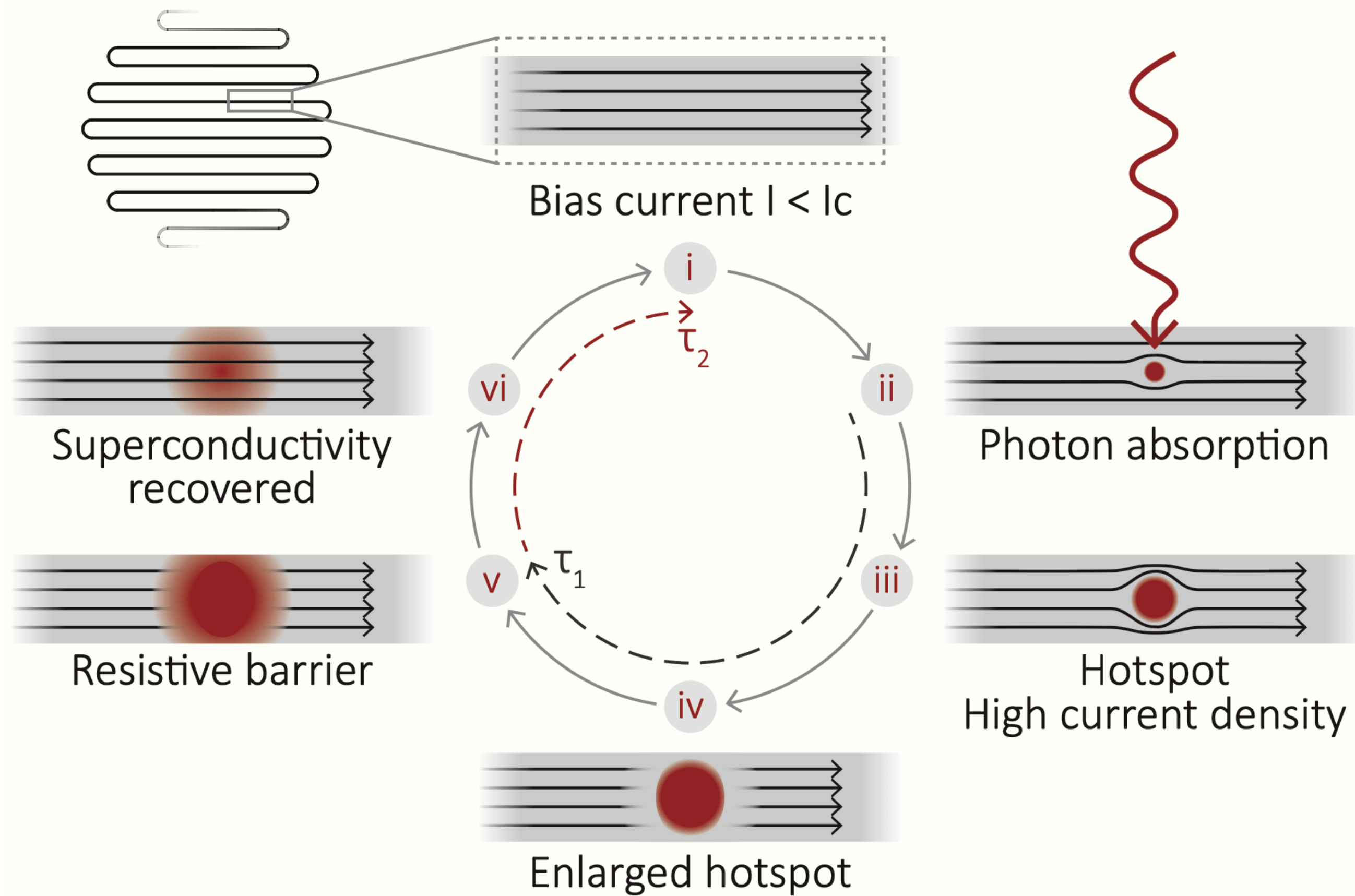
Low energy excess (LEE)



- Low energy excess not understood
 - Shape similar to light DM or CEvNS → limit sensitivity of current experiments
 - LEE not affected by charge amplification, decreases over time and it is reactivated with thermal cycles
- ➔ Most probably due to solid state effects:
- Stress induced by holding structures
 - Relaxation in crystal bulk
 - Events in sensor or sensor interface

SNSPD working principle

Superconducting Nanowire Single-Photon Detectors

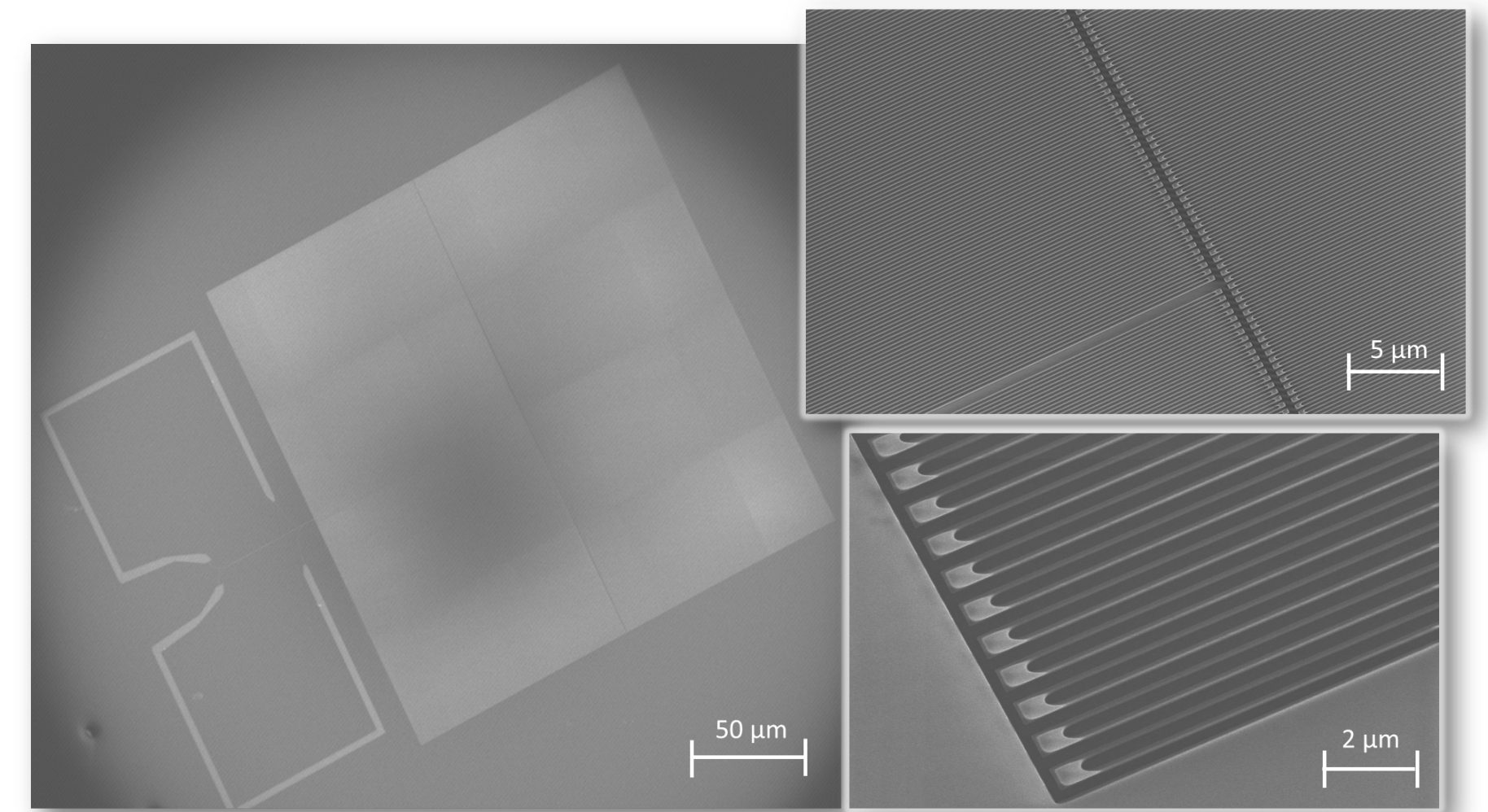


<https://www.singlequantum.com/technology/>

<https://www.idquantique.com/quantum-detection-systems/snsdp-technology/>

Superconducting Nanowire Single-Photon Detectors

- High detection efficiency (>90%)
- Low dark count rates
- Picosecond timing resolution
- WSi, MoSi: lower-gap materials → more quasiparticles per absorbed photon but with slower diffusion and recovery times - used for IR region
- NbN, NbTiN: higher-gap materials → respond faster but with less intrinsic energy spread



Credits to Boris Korzh

Timing jitter	Intrinsic photon number resolution	Efficiency	Array size	Maximum count rate	Dark count rate	Active area	Cut-off wavelength
18 ps	None	93%	64	1 Gcps	4 /s/mm ²	0.001 cm ²	5 μm
↓ 2.6 ps [1]	↓ 3-5 photons	↓ 98% [2]	↓ 4x10 ⁵ [3]	↓ 1.5 Gcps [4,5]	↓ 4x10 ⁻⁵ /s/mm ² [6]	↓ 0.1 cm ² [3]	↓ 29 μm [7]
1 ps	10	99 %	10 ⁷	10 Gcps	1x10 ⁻⁶ /s/mm ²	1 cm ²	100 μm

Records in 2016

Current records for isolated devices

Expected performance by 2030

[1] Korzh, Zhao et al, *Nature Photonics* 14, 250 (2020)

[2] Reddy et al, *Optica* 7, 1649 (2020)

[3] Oripov, Rampini, Allmaras, Shaw, Nam, Korzh, and McCaughan, *Nature* 622, 730 (2023)

[4] Craiciu, Korzh et al, *Optica* 10, 183 (2023)

[5] Resta et al, *Nano Letters* (2023)

[6] Chiles, *PRL* 128, 231802 (2022)

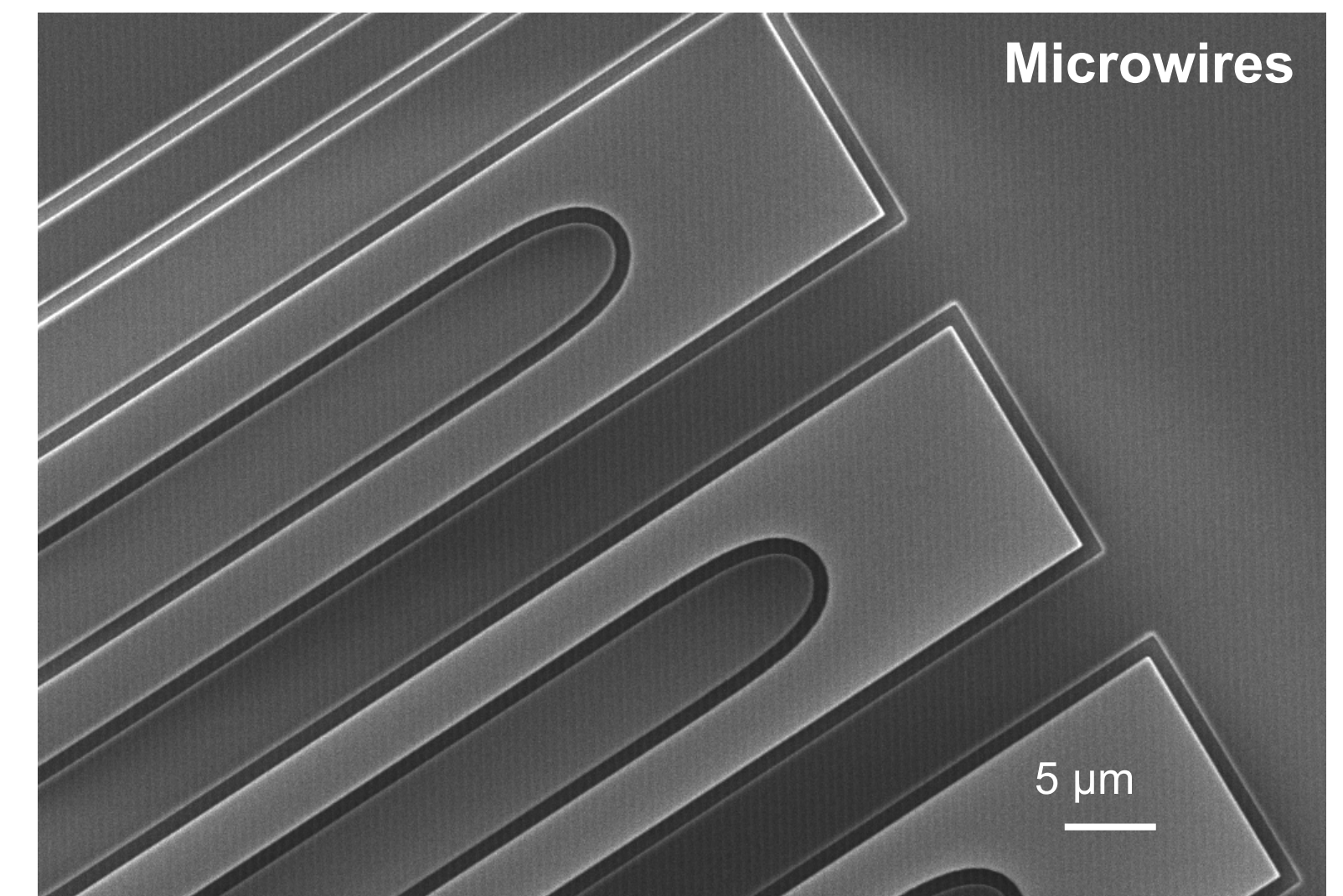
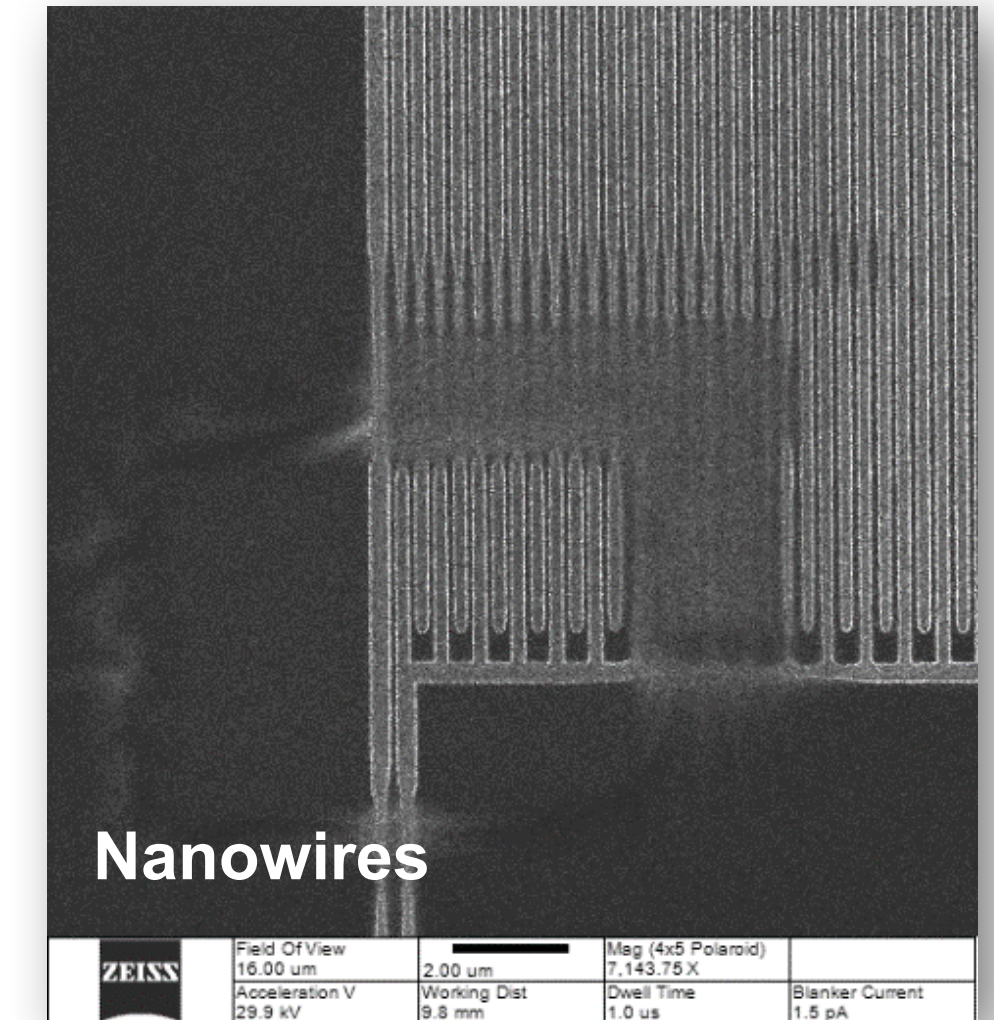
[7] Taylor, Walter, Korzh et al, *Optica*, (2023)

Requirements for dark matter search:

Yonit's talk yesterday

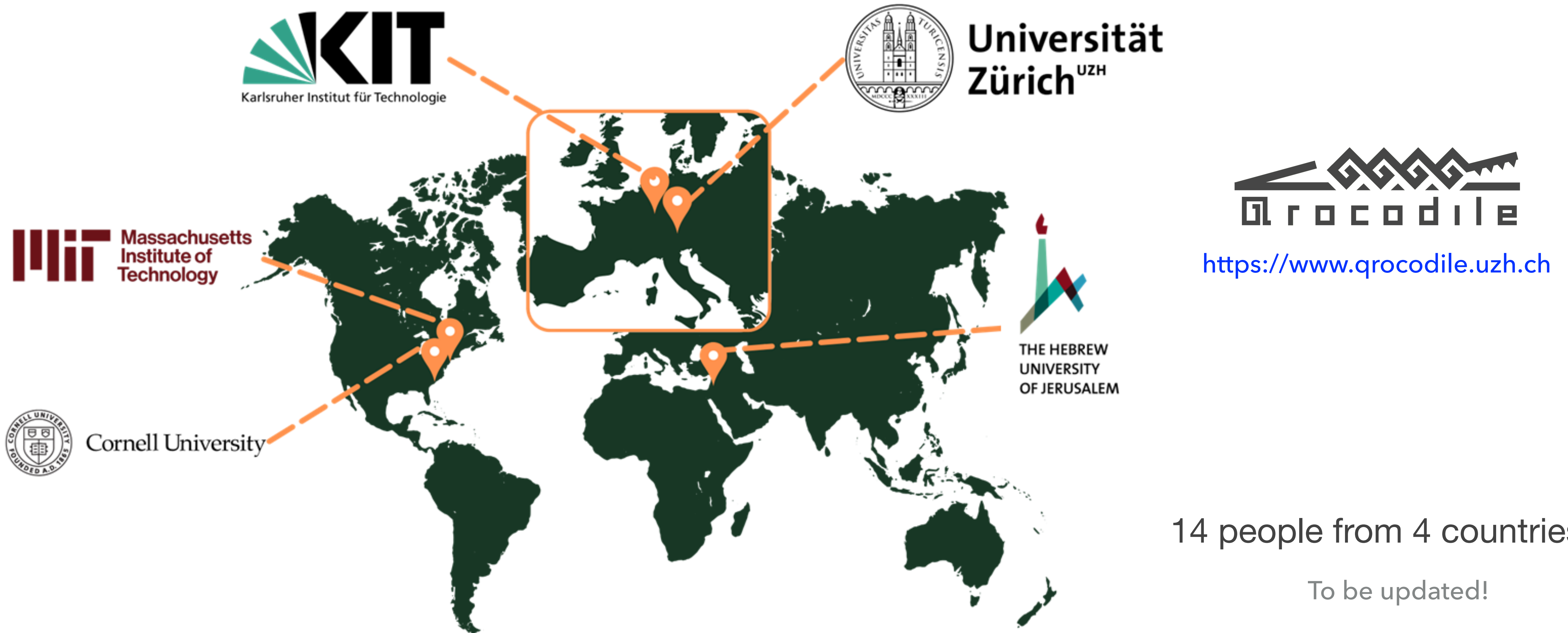
- Large-mass detector
- Mid-infrared detection
- Low dark count rate
- Mitigation of background

- small effective volume (active area $10 \times 10 \mu\text{m}^2$) → for DM microwires $1 \mu\text{m}$ wide on large area device $\gg 10 \times 10 \mu\text{m}^2$
- Complex fabrication
 - Non uniformity of wire production that affects critical current
 - Edge or internal structural defects



The QROCODILE collaboration

Quantum Resolution-Optimized Cryogenic Observatory for Dark matter Incident at Low Energy

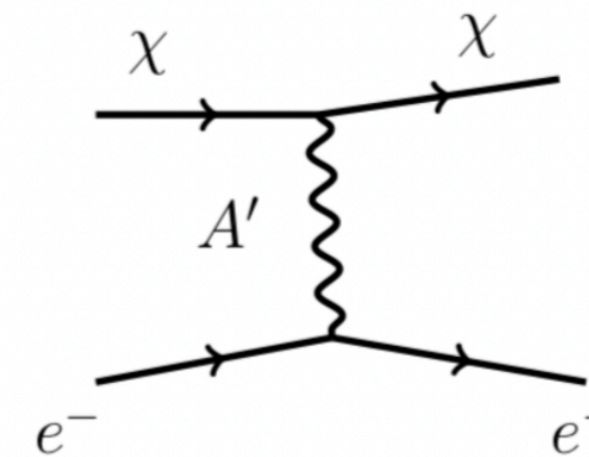
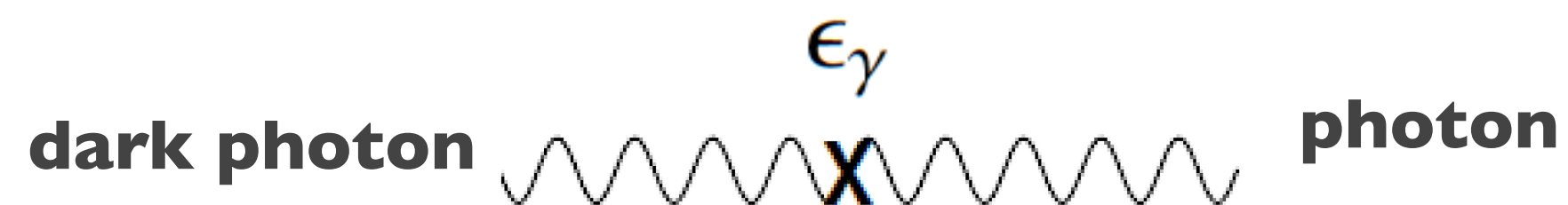


The QROCODILE experiment

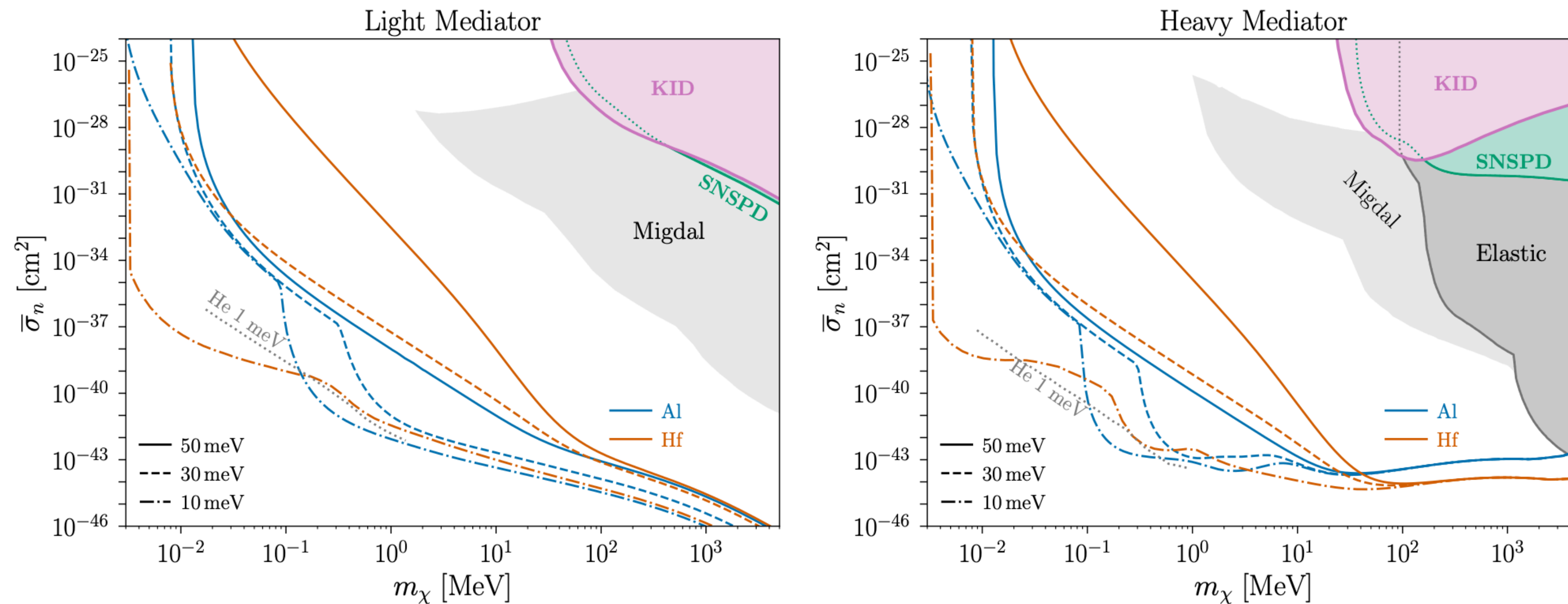
Phys.Rev.Lett. 135 (2025) 8, 081002

- DM particle with mass down to 30 keV
- SNSPD as target = sensor
- Sensitive to scatter and absorption of dark matter
- Coupling between phonons and quasiparticles in the detector → interaction with electrons and nucleons
- Directionality from anisotropy in the device geometry → reject background and select DM from Galactic halo

QROCODILE @ UZH cryostat

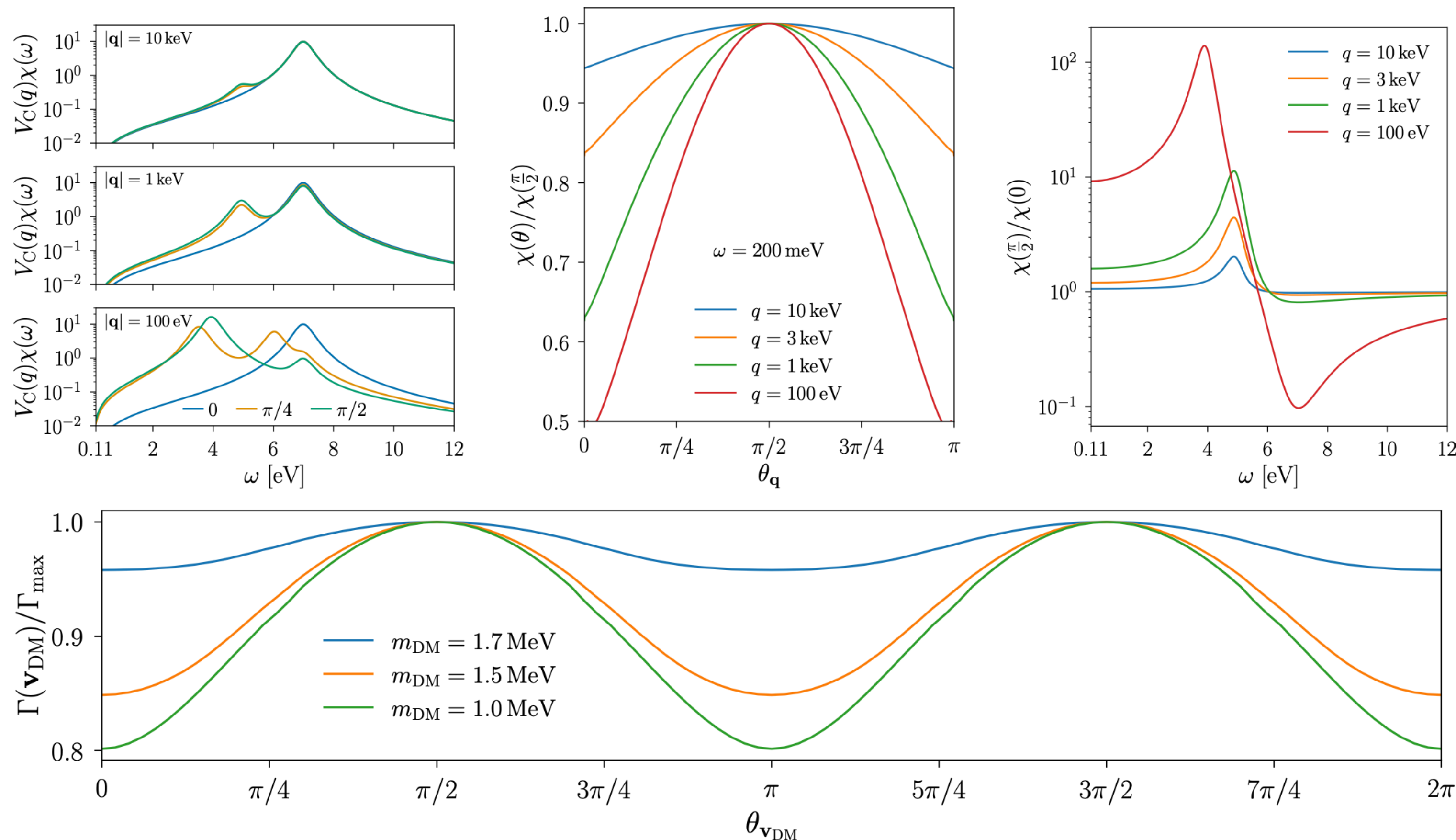


- At low masses interaction dominated by many-body processes rather than elastic scattering
- SNSPD sensitive to nuclear scattering via phonon production
- Phonons break Cooper pairs and produce free quasiparticles for $E > 2\Delta \rightarrow$ transition to the normal metal state
- Same detector sensitive to electronic and nuclear interaction



1 kg yr exposure
on existing devices

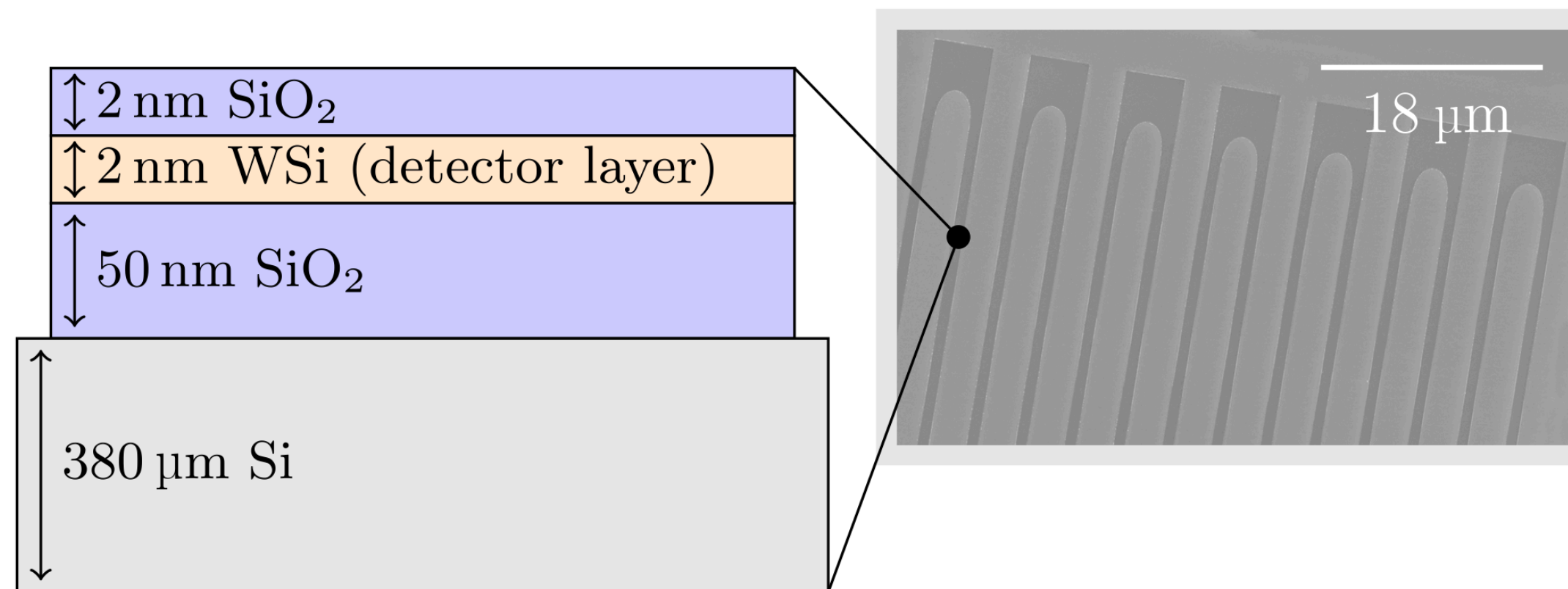
- Dynamic structure factor to model response of the target material
- Calculations assuming infinite bulk volume \rightarrow approximation fails for target dimensions $<$ inverse momentum transfer \rightarrow required geometric corrections



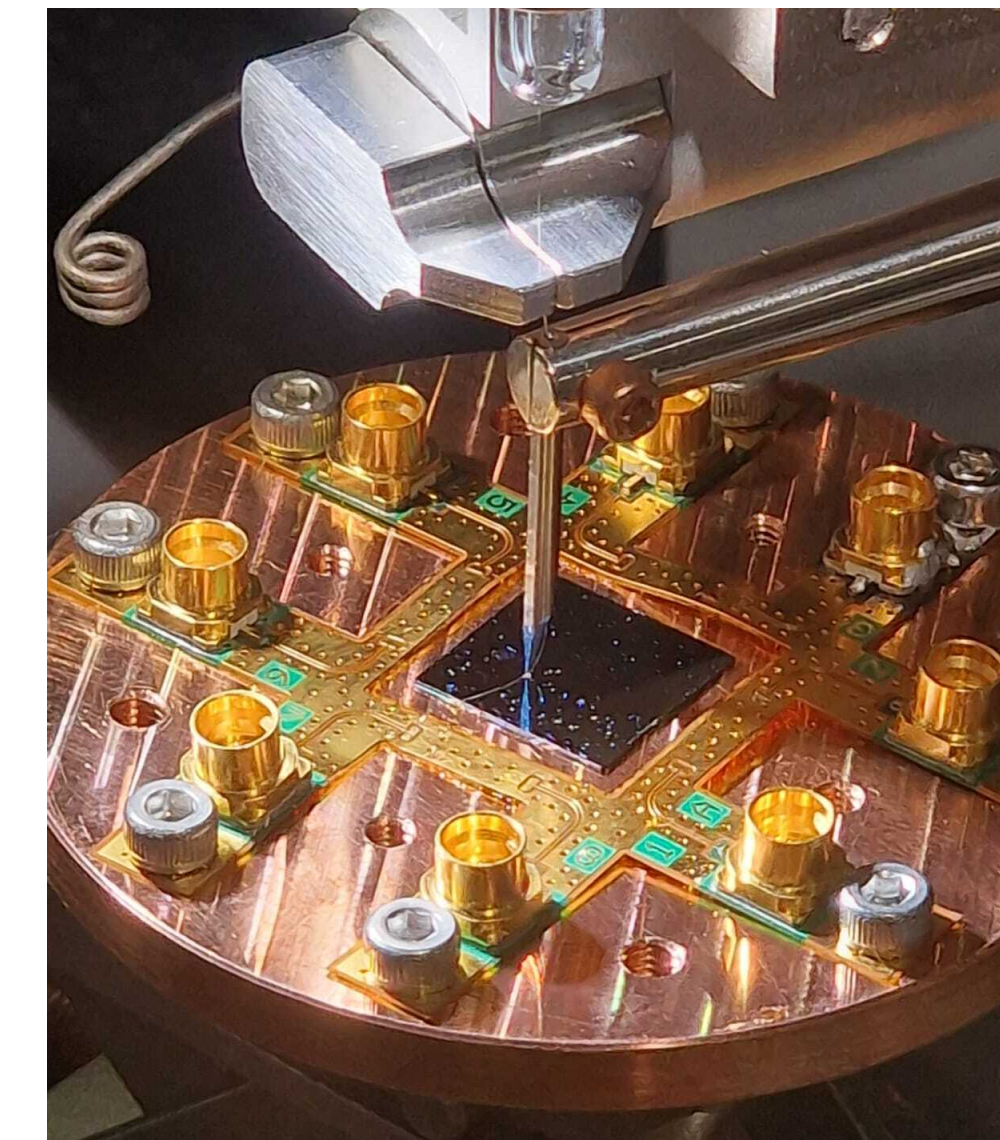
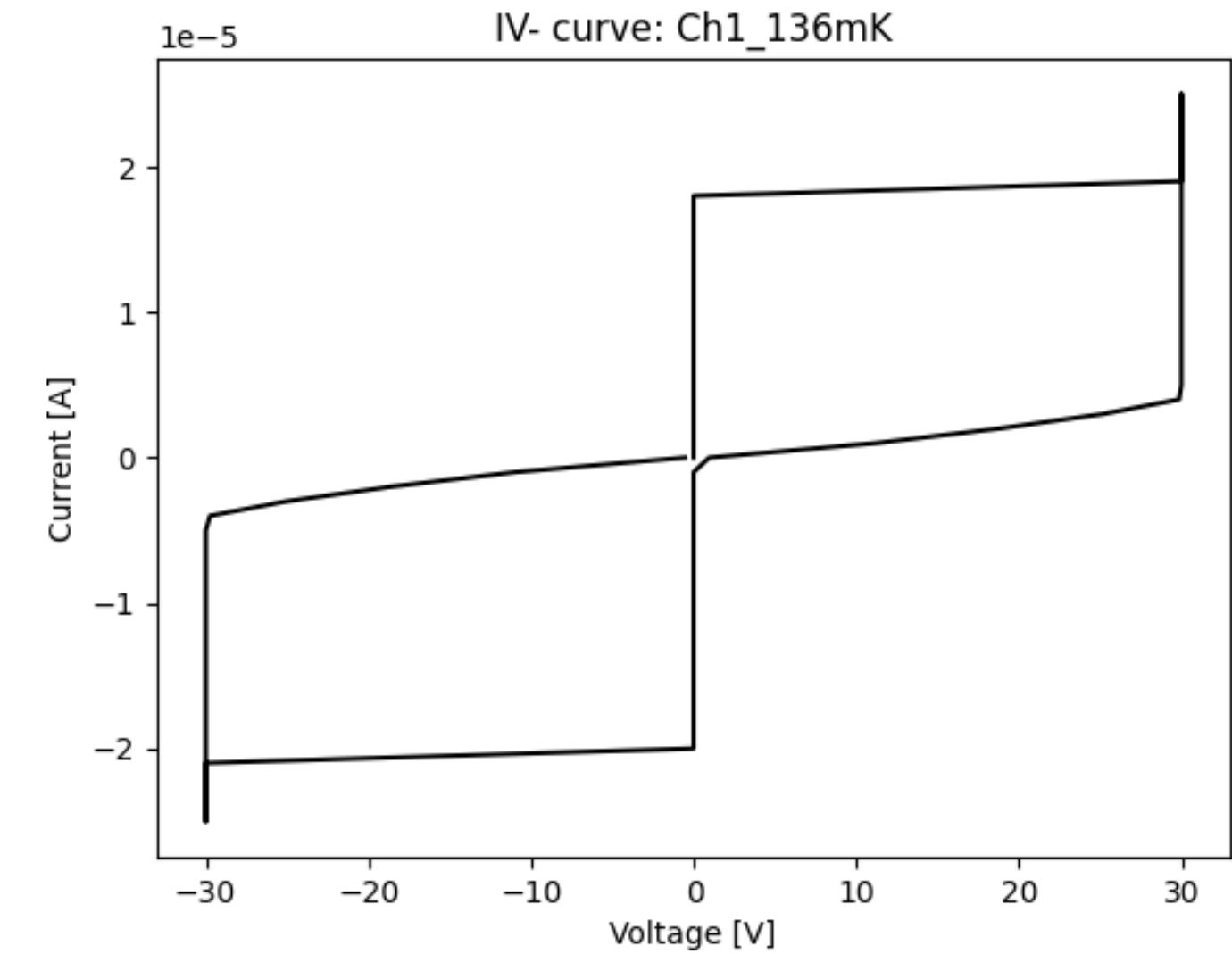
- Possible only for light masses:
 - 200 keV DM $\rightarrow 1/q \sim 1$ nm
 - 30 keV DM $\rightarrow 1/q \sim 10$ nm not negligible on 2 nm layer

The sensor

- 10x10 mm² chip size
- 600x600 μm² active area with 1 μm wire width
- 25% filling factor
- WSi 45:55 superconducting layer 2 nm thick
- Transition temperature at 1.73 K
- Si/SiO₂ substrate

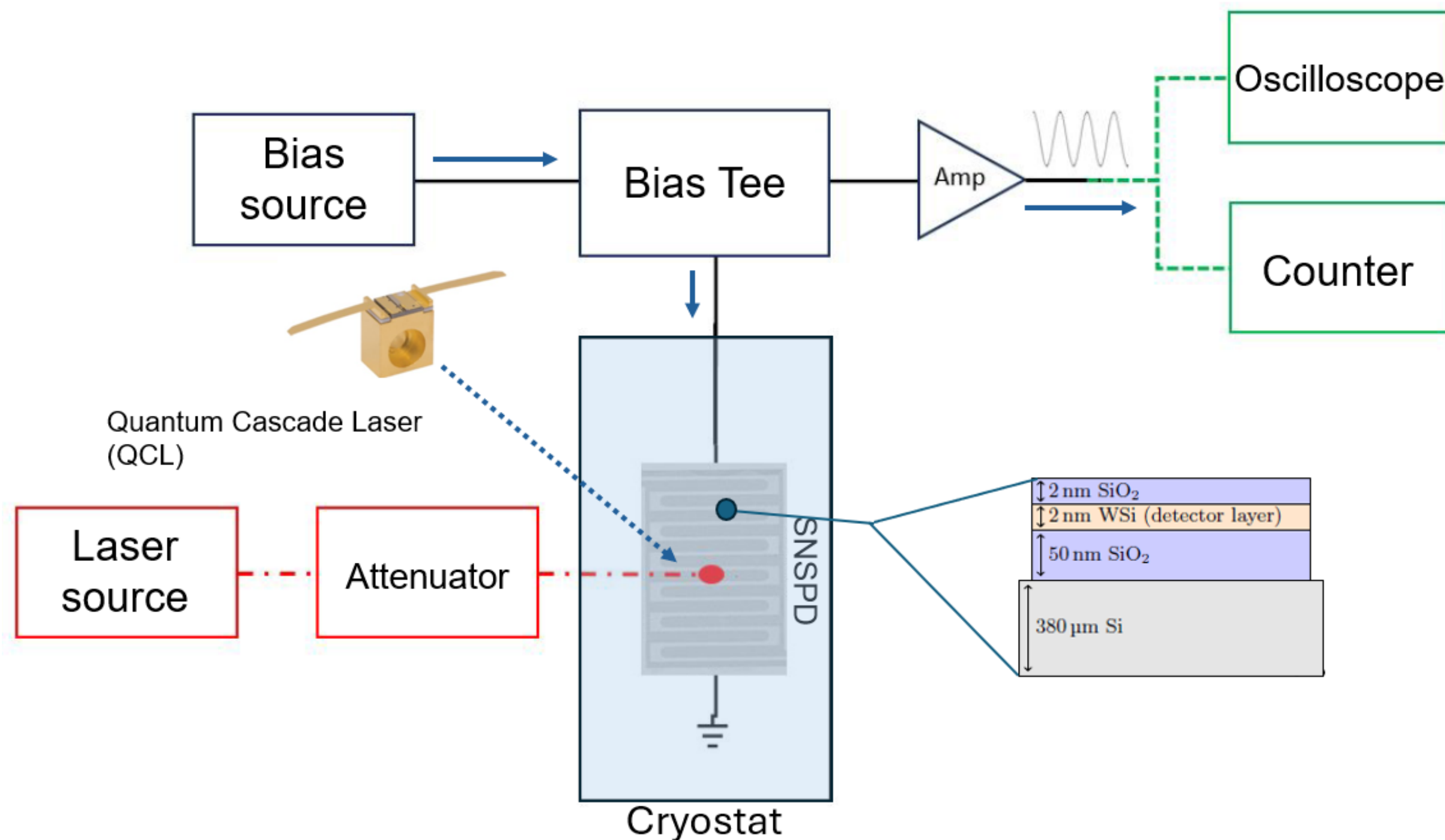


Phys.Rev.Lett. 135 (2025) 8, 081002

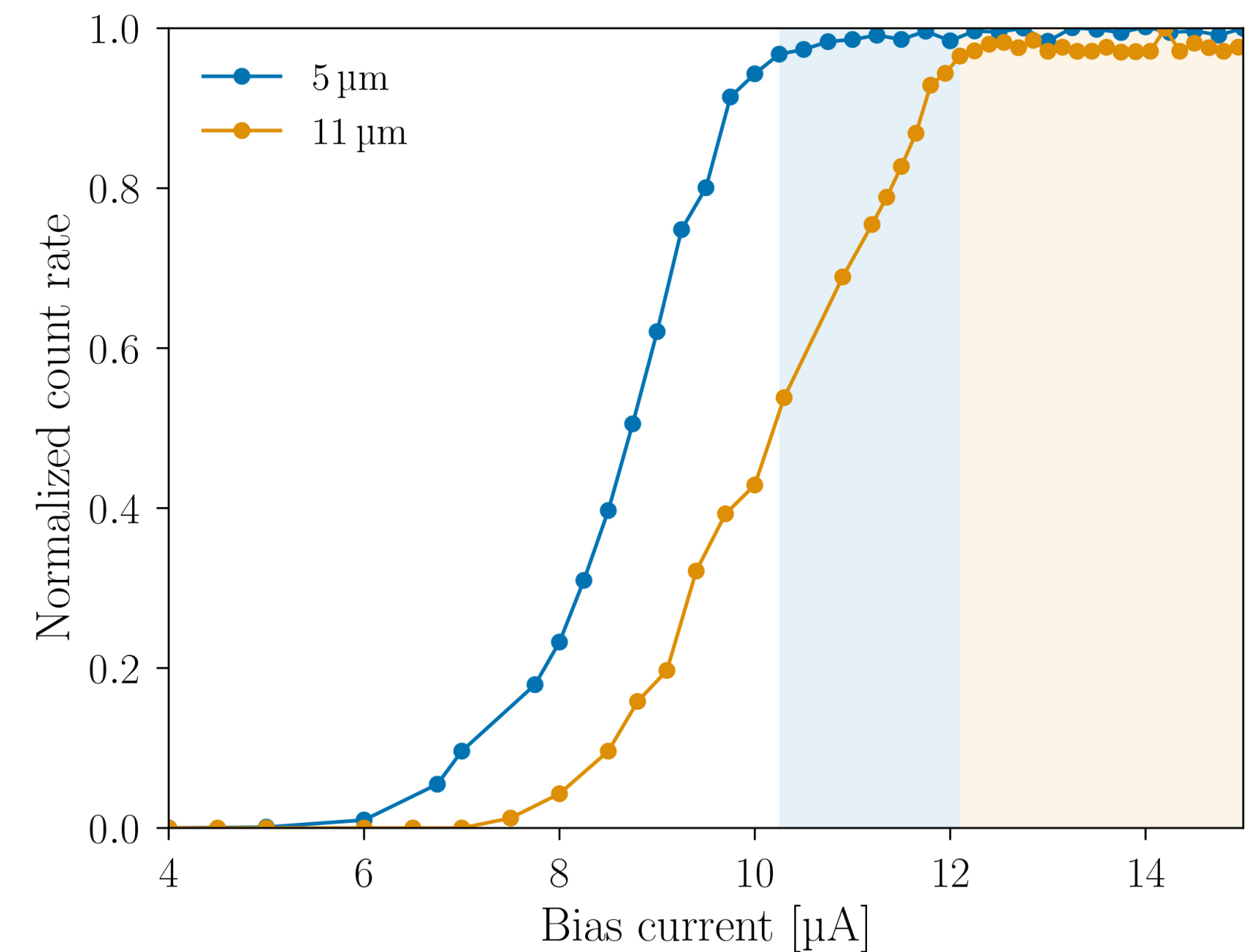
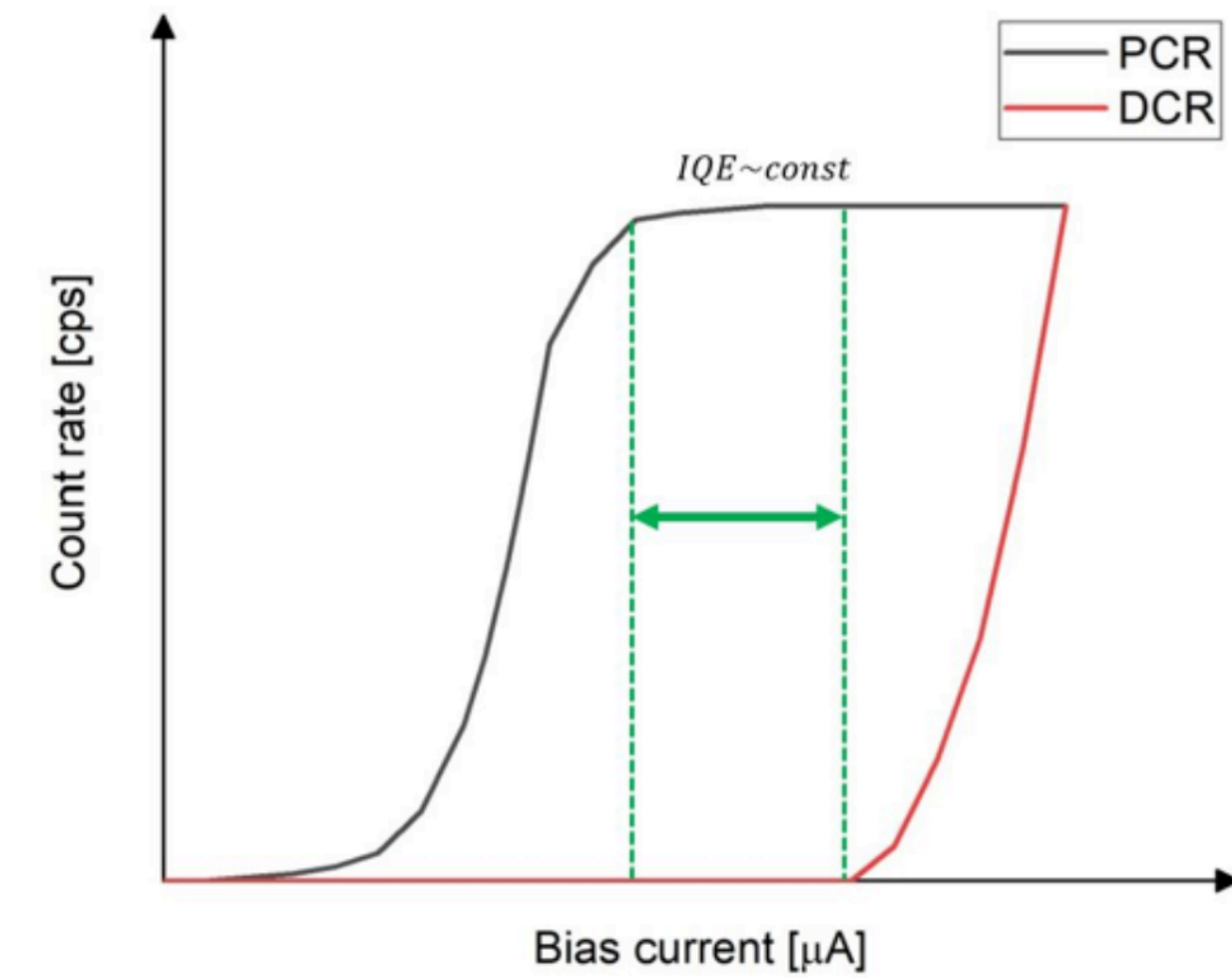


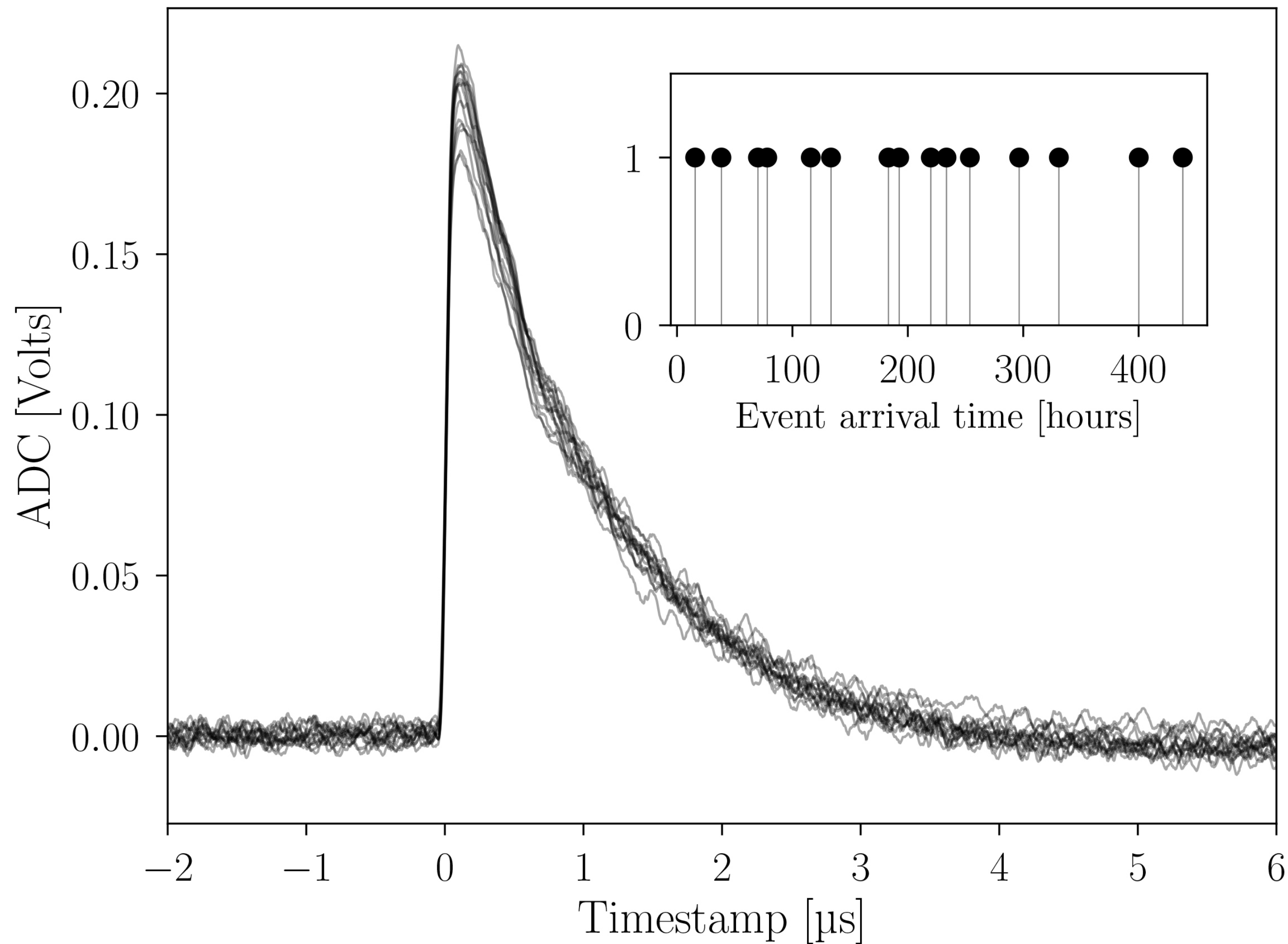
Calibration

- Calibration in mid-infrared region with 5 μm and 11 μm lasers \rightarrow corresponding to 0.25 eV and 0.11 eV
- Detection efficiency up to 98%
- Set working point at 12.2 μA and 100 mK

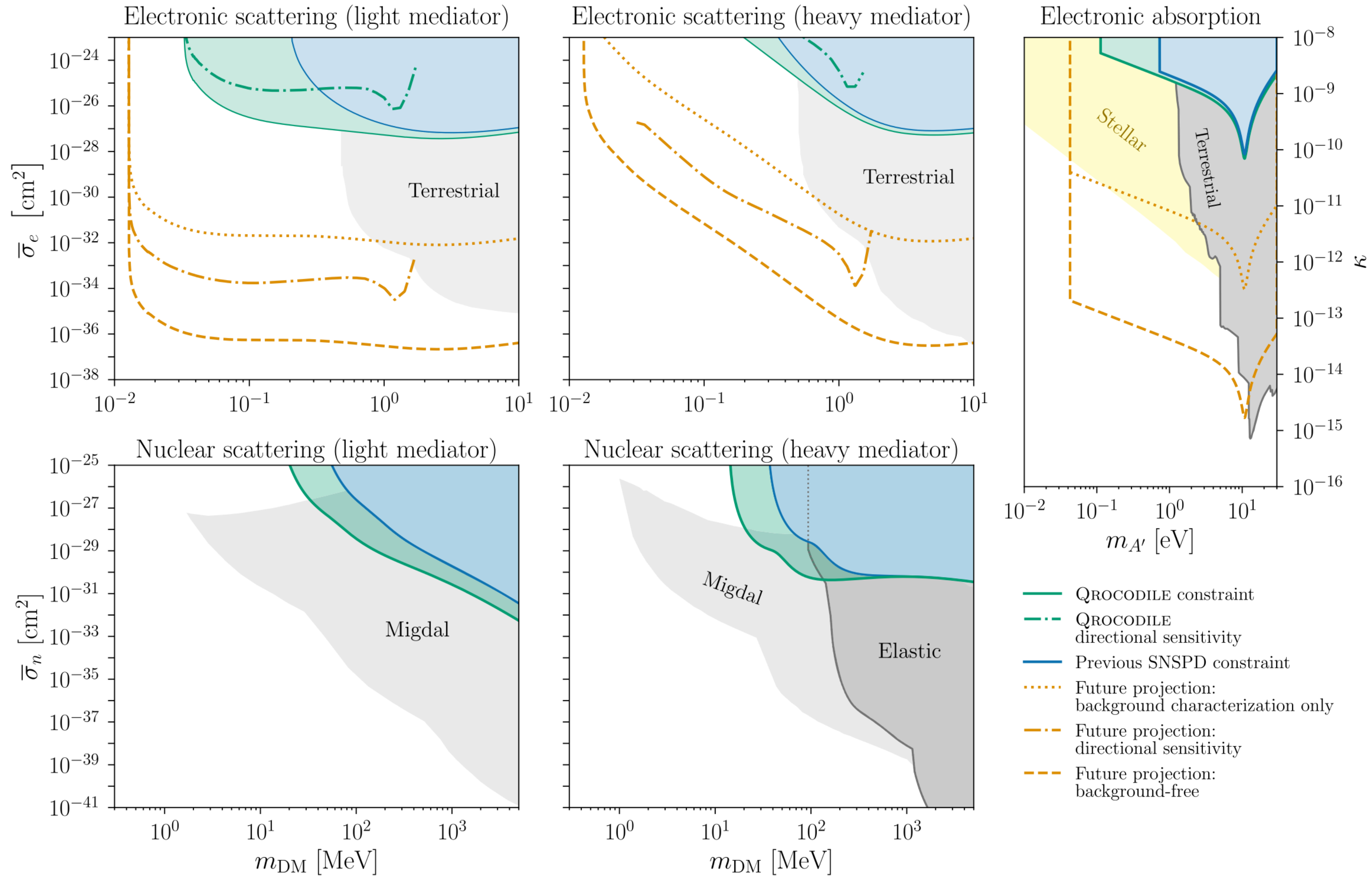


Phys.Rev.Lett. 135 (2025) 8, 081002

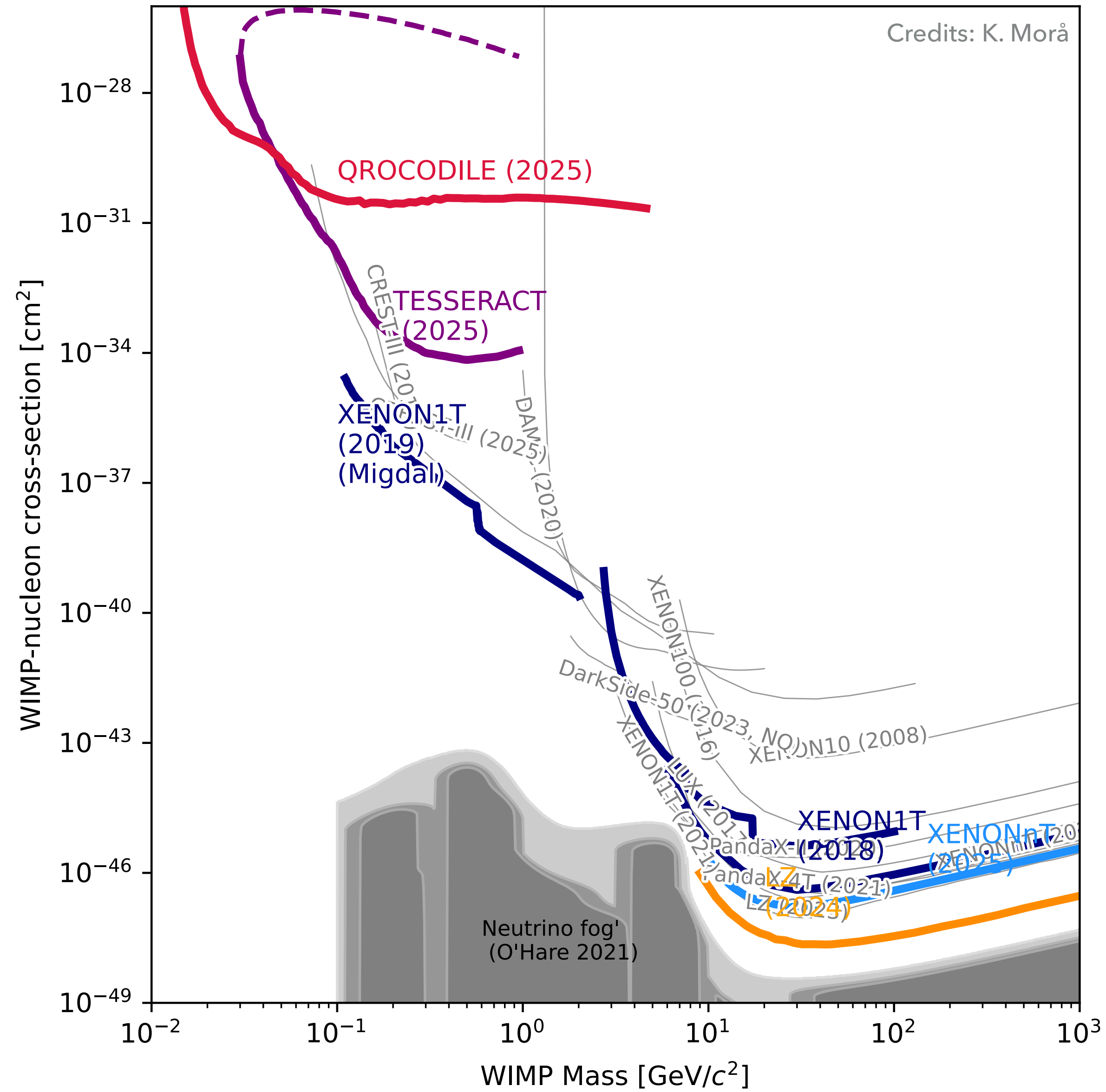




- Livetime: 415.15 hours
- 15 individual non periodic pulses = 10^{-5} Hz
- Pulse shape not energy correlated
- Not enough statistics to detect anisotropy in rate
- Constraints from total count rate
- 95% C.L. from Feldman-Cousins
- Projection for 10^7 pixels and threshold of 0.04 eV (29 μm)



Current scenario

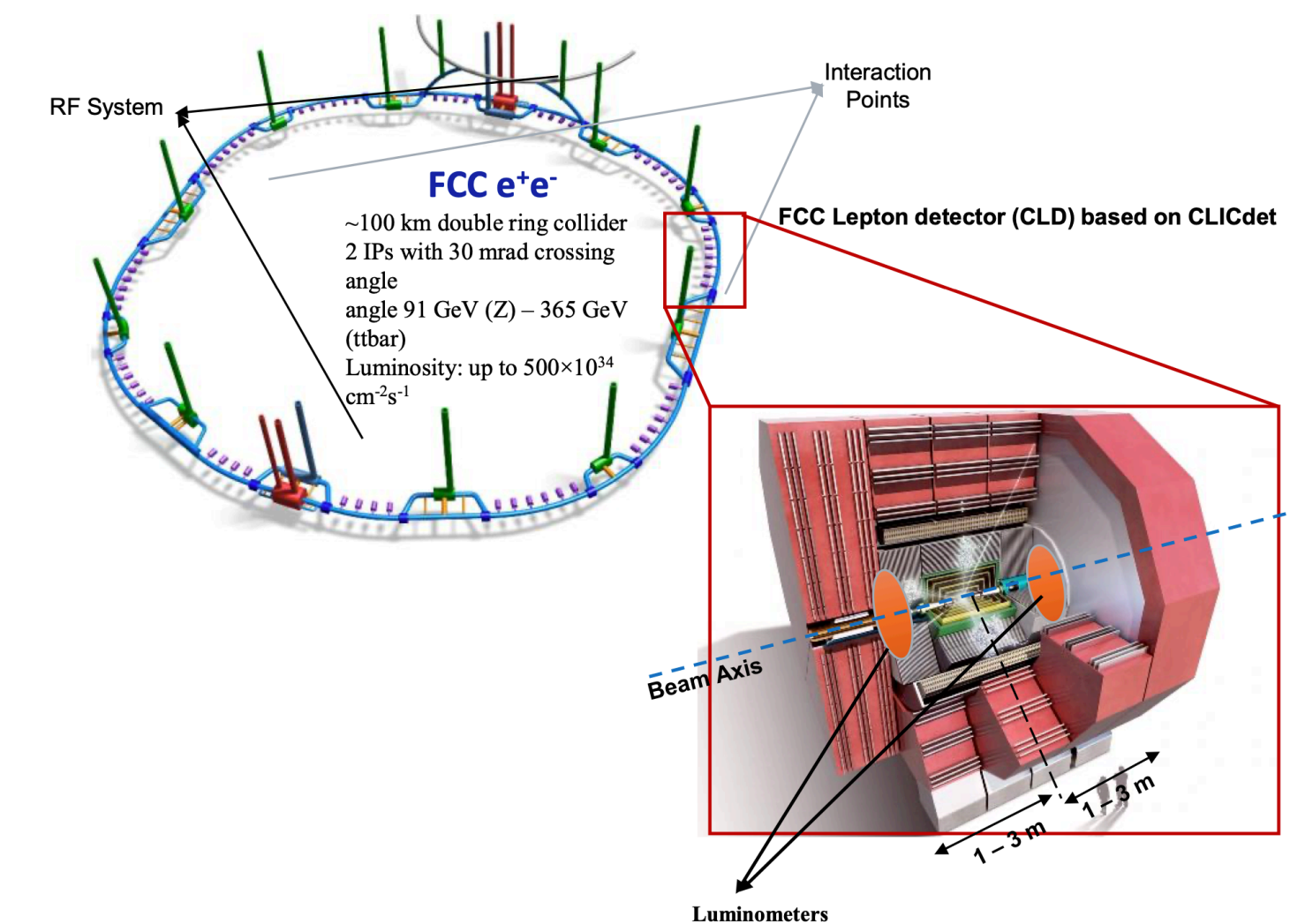


Next plans

Temporary downtime during migration to alternative sensor suppliers

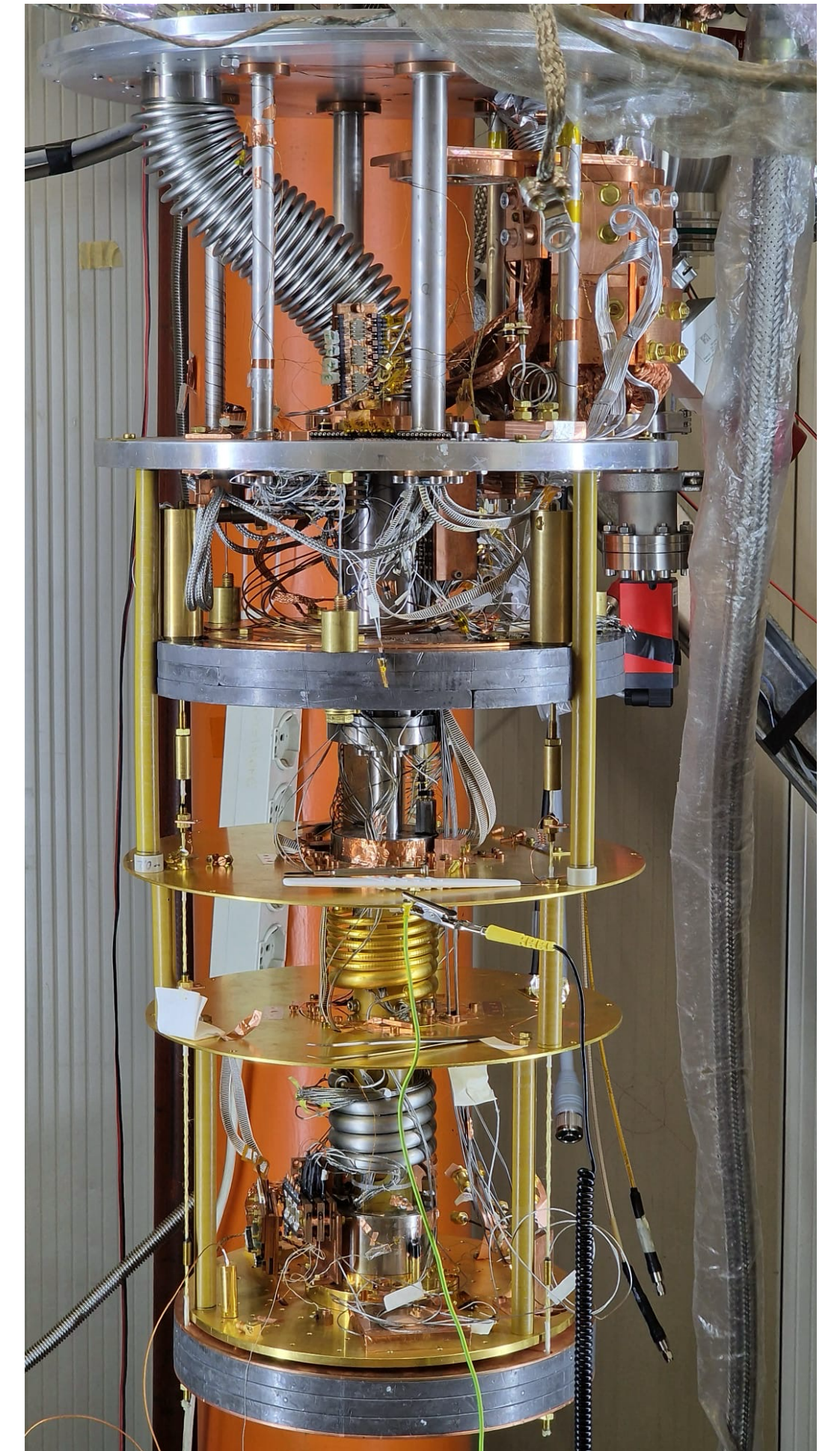
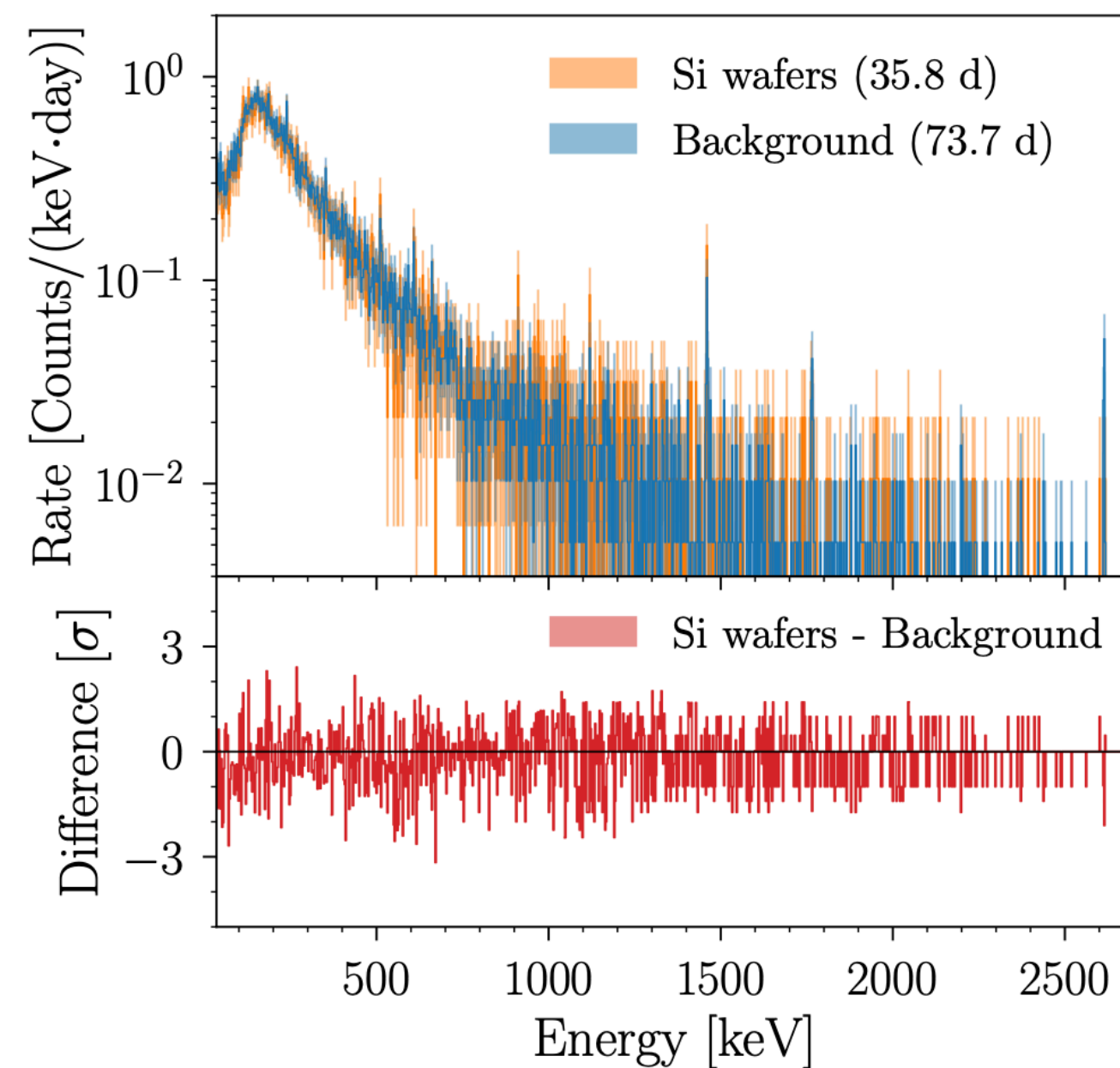
New sensors production:

- Moving UZH production to EPFL facility
- WSi chip from JPL - 1x1 mm with 8 pixels
 - Already fabricated, in Geneva for encasing
- Uni Geneva joining - production of MoN films, 800x800 um with 4 pixels
 - Validation of the production line
- Collaboration with DRD5 and the UZH CMS group (development for FCC)

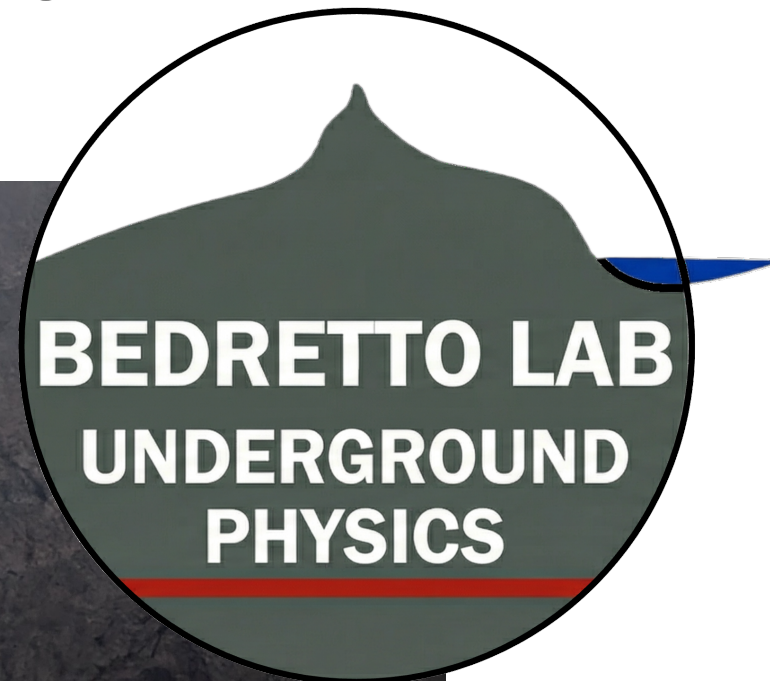
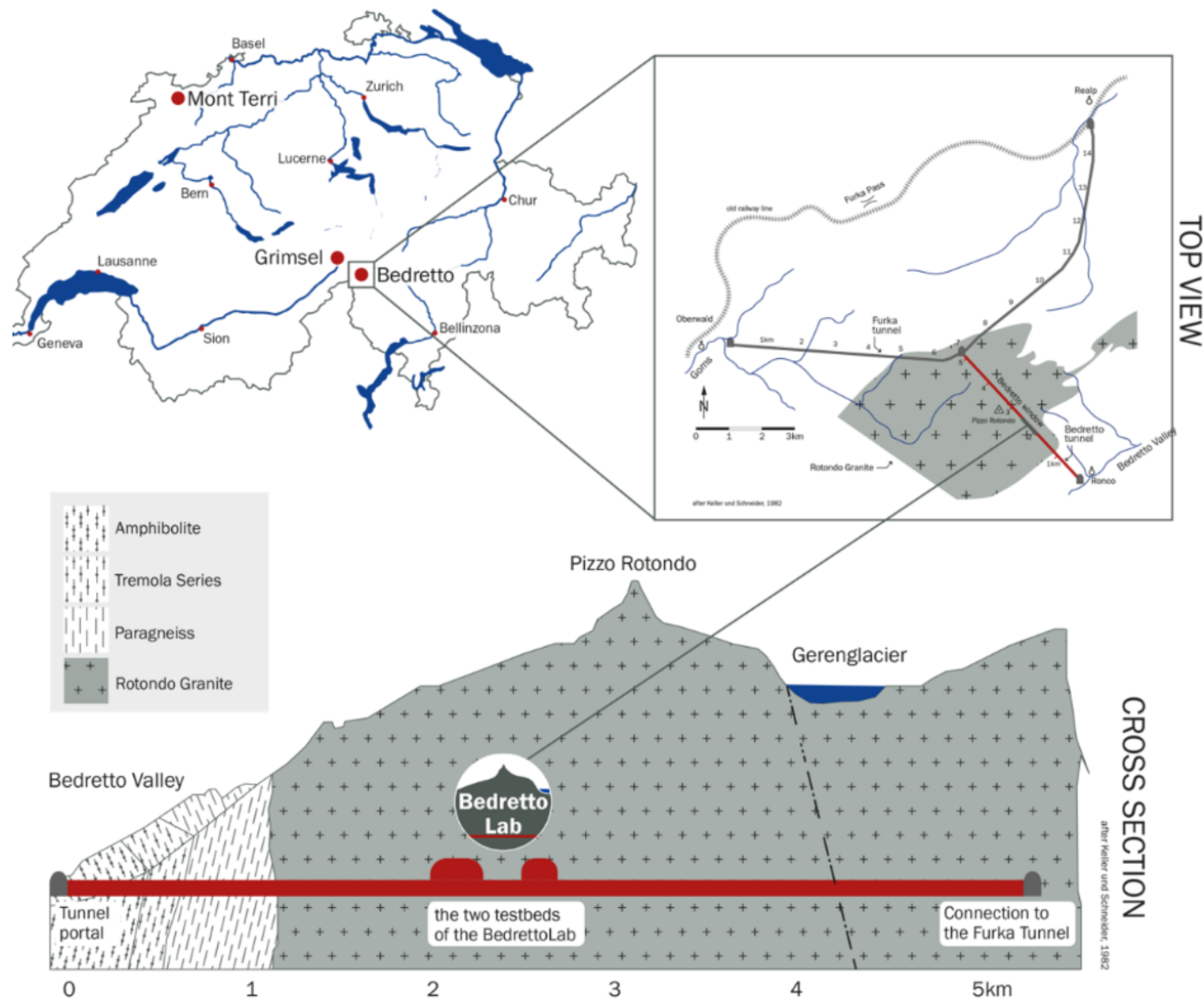


Origin of the events?

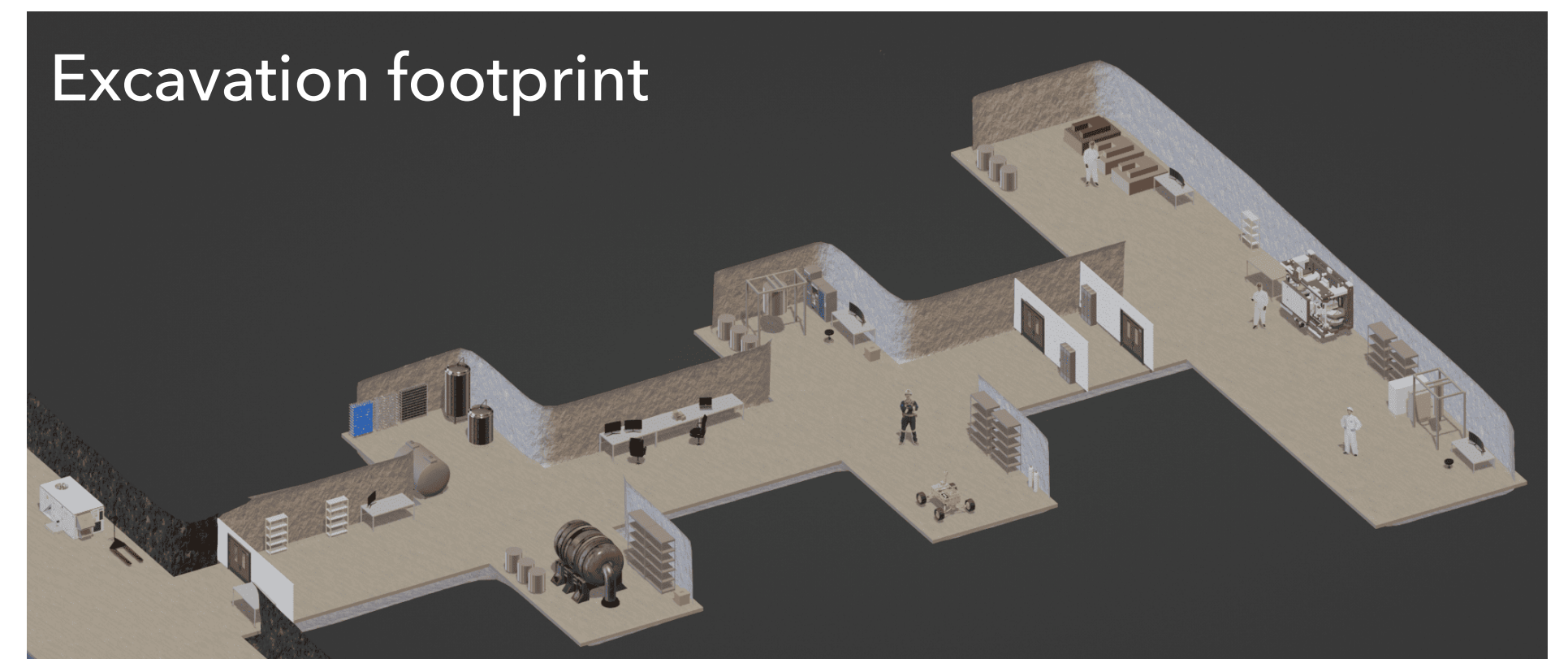
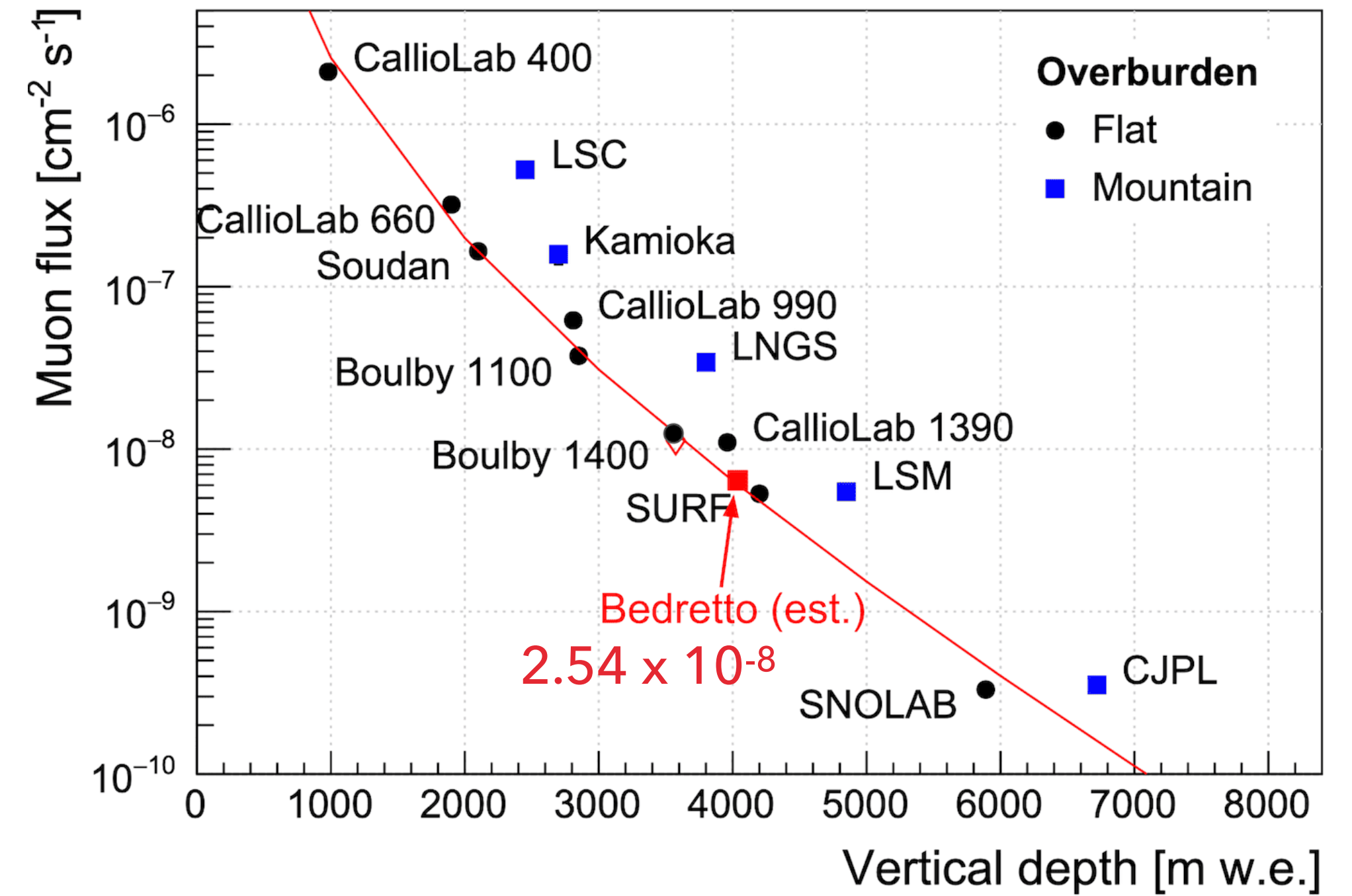
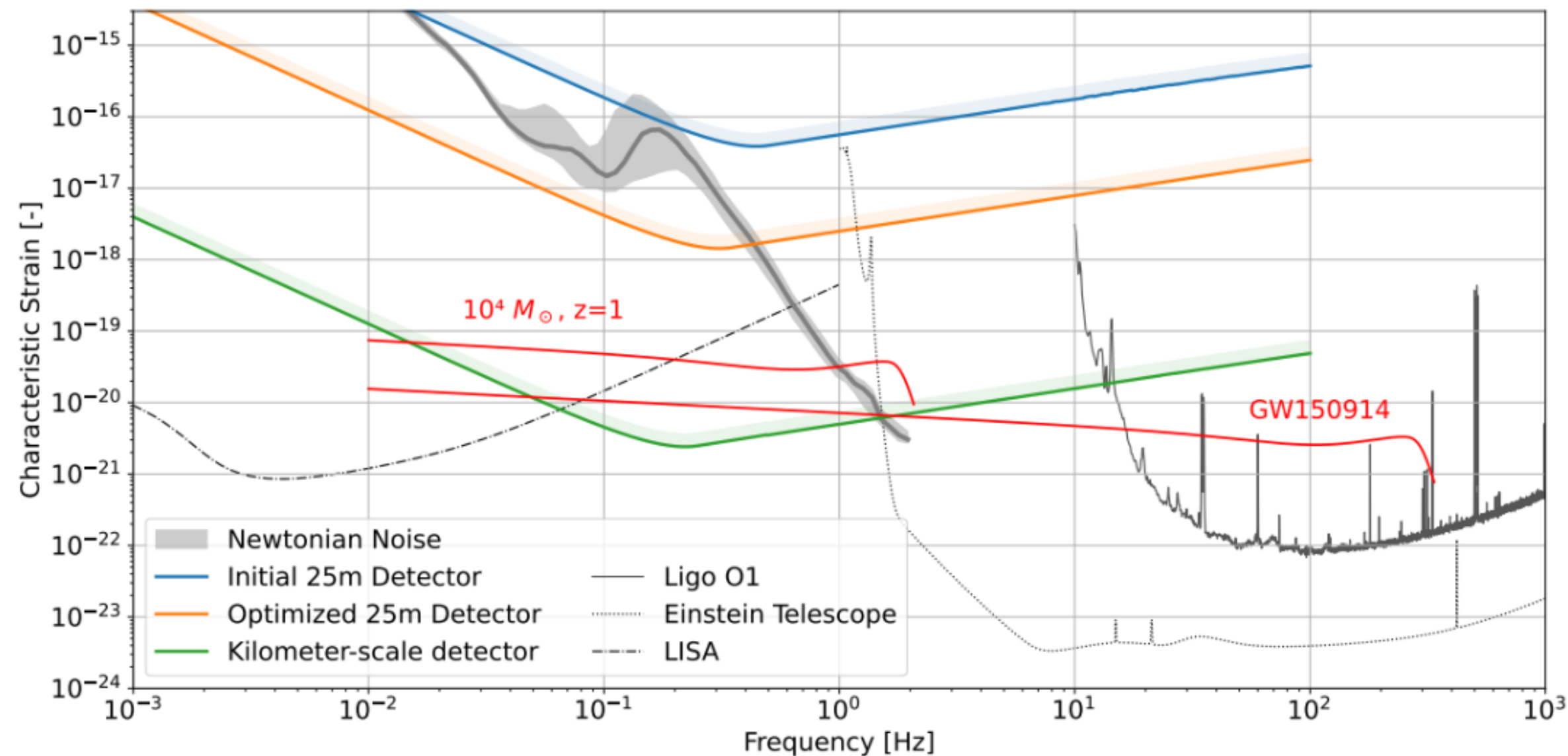
1. Measurement in underground lab to disentangle background contribution from cosmic rays
 - IETI facility at LNGS in Italy
 - Test runs to setup readout and electronics
2. Material radioessay for contaminant tracing
3. MonteCarlo simulations



- Underground tunnel in Switzerland - 5.2 km in side access of Furka train tunnel
- ~1000-1500 m deep, three caverns 6 m wide and 50-100 m long
- Since 2019 used by geophysics groups at ETH - since 2025 particle physics involved

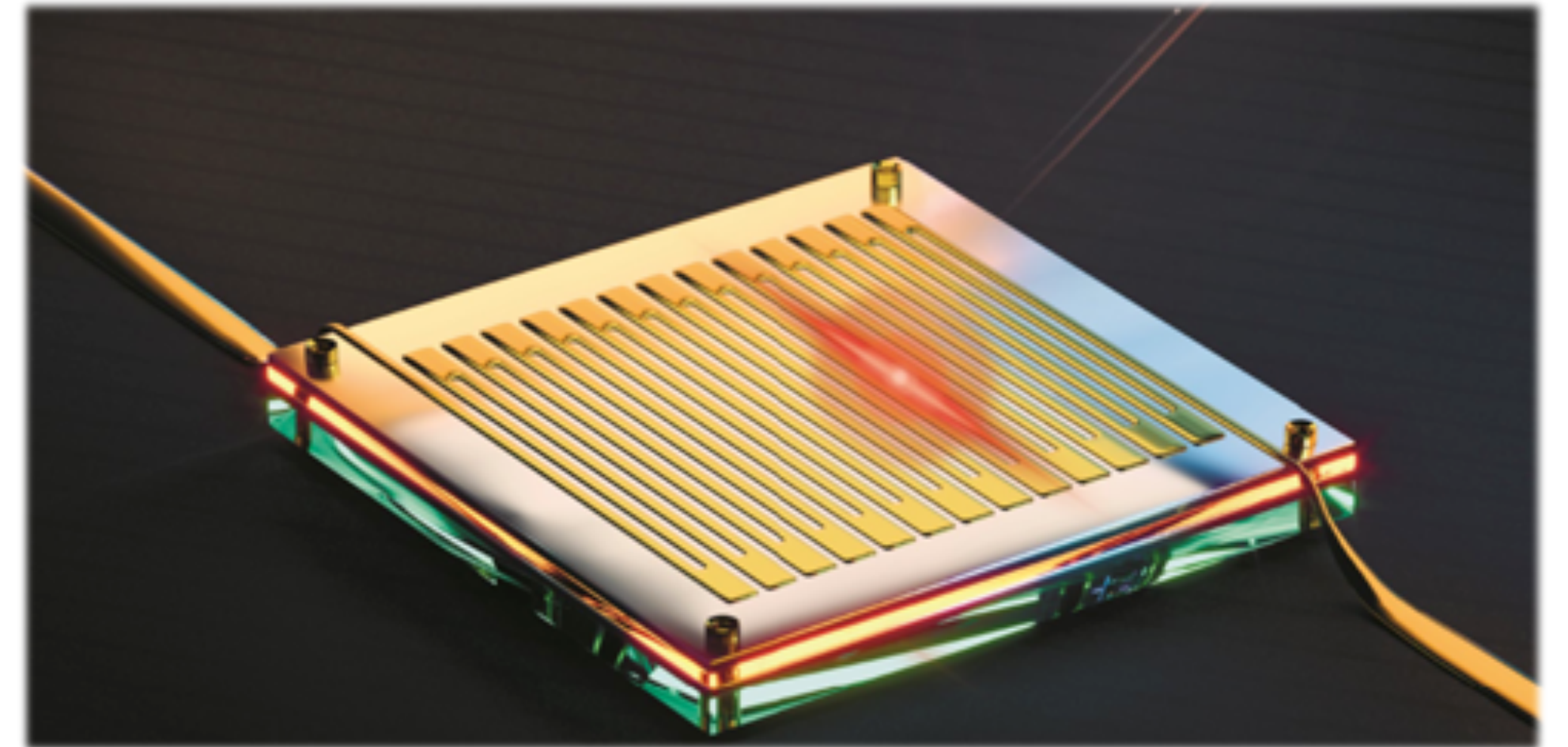


- Radioactivity background characterisation campaign
- Mid-size astroparticle experiments in ~2 years
 - Dilution cryostat for SNSPDs and quantum sensors
 - Gravitational waves interferometer
 - Small scale TPC for R&D, HPGe low background counter,...



Conclusion and outlook

- Large area SNSPDs promising technology for low mass dark matter search
 - Simple readout
 - Easily scalable to larger masses
- First large area chip in aboveground measurement probed new parameter space
- Planned underground measurement
- Producing new sensors to optimise design and threshold



Phys.Rev.Lett. 135 (2025) 8, 081002

Thank you!