

# Backgrounds: What Lies Beneath



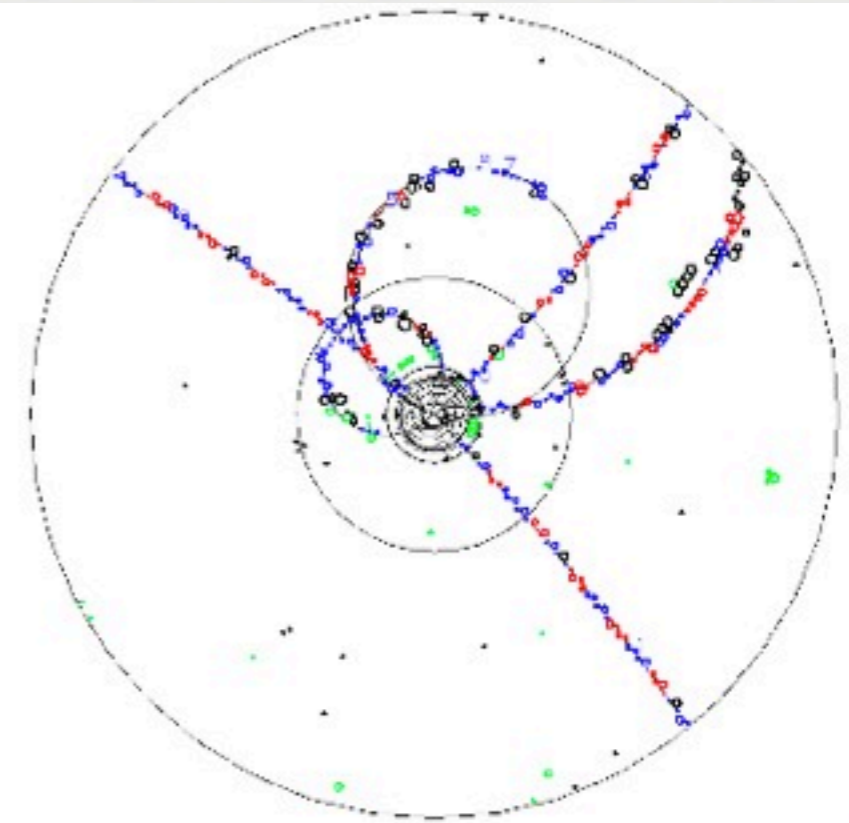
*E. Paoloni (INFN & Università di Pisa)  
for the MDI & background sim. Task Force*

# BACKGROUND & PERFORMANCES

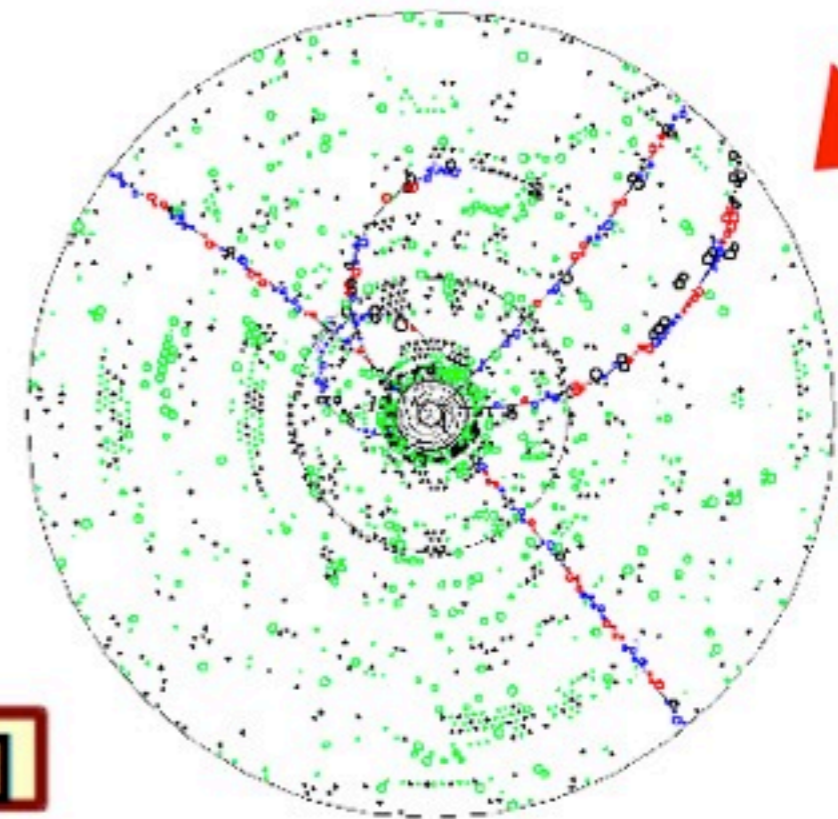
- Background effects on performances can be categorized in
  - Instantaneous ones (depending on the instantaneous rate):
    - Pattern recognition confusion
    - Resolution deterioration
  - Long term ones (depending on the integrated doses):
    - Detector aging => performances degradation
    - + backgrounds : (

# TRACKING

- Instantaneous ones:
  - Hit shadowing (SVT DCH)  
=> hit efficiency
  - Pattern recognition =>  
tracking efficiency
- background hit misassignment  
=> resolution degradation



Belle - 2

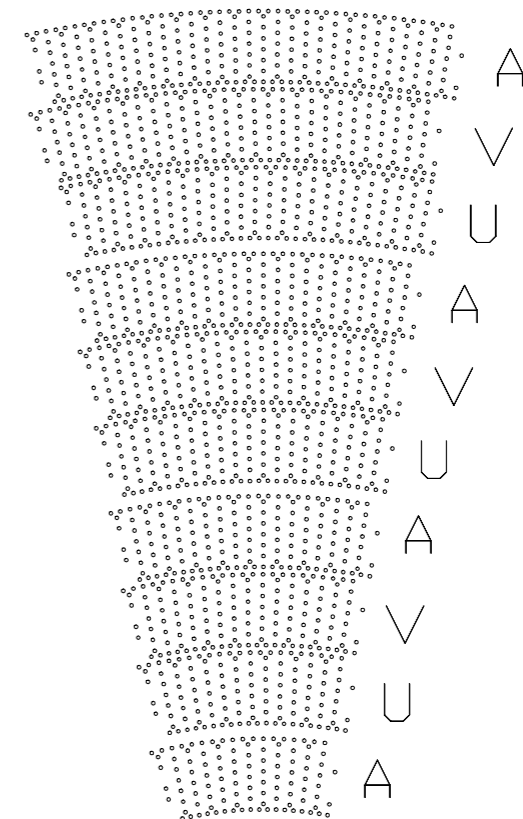
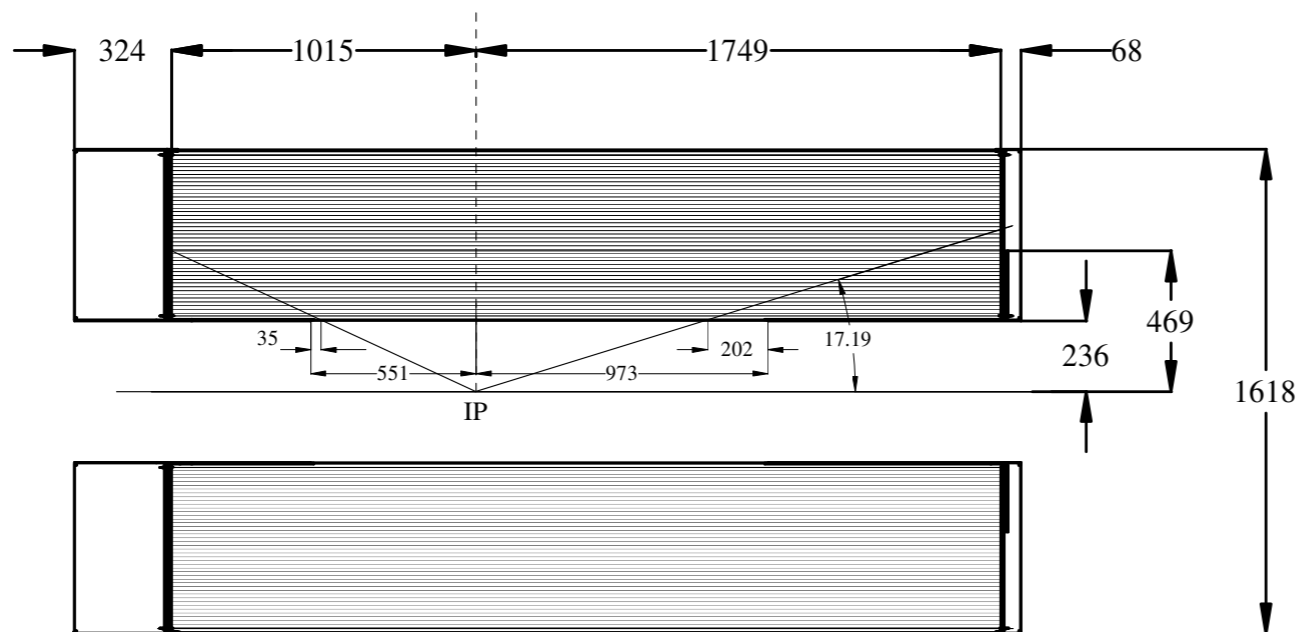


# BABAR EXPERIENCE (2005)

Performance Background Upgrade HiLumi Extra

## Drift Chamber design and operation

- Main tracking device in BaBar surrounding SVT
- 2.8m length, 1.6m diameter, 7104 sense wires
- 40 axial and stereo layers of 1-2cm hex cells
- Gas: 80:20 He:C<sub>4</sub>H<sub>10</sub>, 4000 ppm H<sub>2</sub>O vapor
- Operating voltage: 1930V



# BABAR BASELINE

Performance Background Upgrade HiLumi Extra

## Reconstruction

- Single hit resolution  $125 \mu\text{m}$
- Momentum resolution,  $p_T$  :  $0.45\% + 0.15\% p_T / (\text{GeV}/c)$
- Tracking efficiency :  $>95\%$  matching with SVT tracks
- $dE/dx$  resolution  $\approx 7.5\%$

# Tracking at high luminosity

- Mix Monte Carlo  $B \rightarrow D^{(*)}D^{(*)}$  with real random-trigger data
- Multiple triggers overlaid to match luminosity extrapolations
- Evaluate physics quality relative to current performance

Compared to design $3 \times 10^{33} \mathcal{L}$	$2 \times 10^{34}$ ( $3 \times \text{bkg}$ )	$4 \times 10^{34}$ ( $5 \times \text{bkg}$ )	$4 \times 10^{34}$ ( $10 \times \text{bkg}$ )	$100 \times 10^{34}$ ( $?? \times \text{bkg}$ )
Tracking eff. (%)	$98.6 \pm 0.1 \pm 0.7$	$97.4 \pm 0.1 \pm 1.0$		
Momentum $\sigma(p)/p = 4.7 \times 10^{-3}$	$+4.2 \times 10^{-5}$	$+5.5 \times 10^{-5}$		
$D^0 \rightarrow K^+ \pi^-$ (%)	$96.0 \pm 0.5$	$95.5 \pm 0.5$	$80 \pm 3$	
$D^0$ Mass $\sigma = 6.5 \pm 0.2 \text{ MeV}/c^2$	$6.5 \pm 0.2$	$6.4 \pm 0.2$	$7.0 \pm 0.3$	
$D^* \rightarrow D^0 \pi$ (%)	$84.4 \pm 1.1$	$75.0 \pm 1.3$	$25 \pm 2$	
$D^*$ Mass $\sigma = 0.80 \pm 0.03 \text{ MeV}/c^2$	$0.97 \pm 0.04$	$1.50 \pm 0.08$	$3.2 \pm 0.8$	

$$\text{Occupancy}(\%) = 0.61 + 0.17 I_{LER} + 3.97 I_{HER} + 0.42 \mathcal{L} + 0.21 I_{LER} \text{ (beam - beam)} + 0.03 \mathcal{L} \text{ (trickle inj.)}$$

# DCH NOTABLES FIGURES OF MERIT

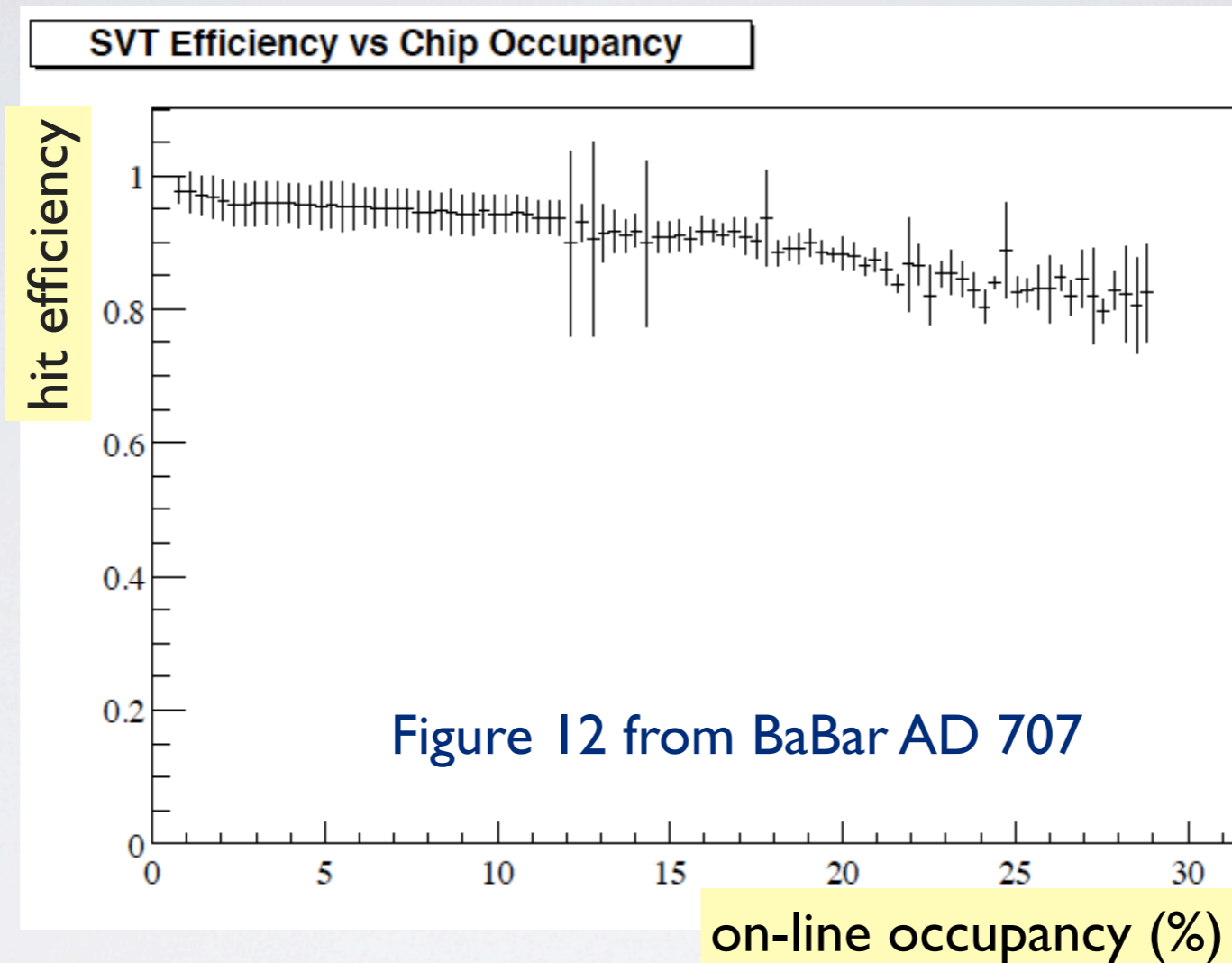
- Track reconstruction efficiency as a function of  $p_t$
- Track parameters resolution
- $dE/dx$  resolution
- How are they affected by backgrounds conditions?

# SVT STUDIES

- So far Isabelle Ripp-Baudot evaluated the effect of the background on a “SvtTrkHitAdd” like algorithm assuming DCH performances comparable to the BaBar one



# study of tracking performances with BaBar data

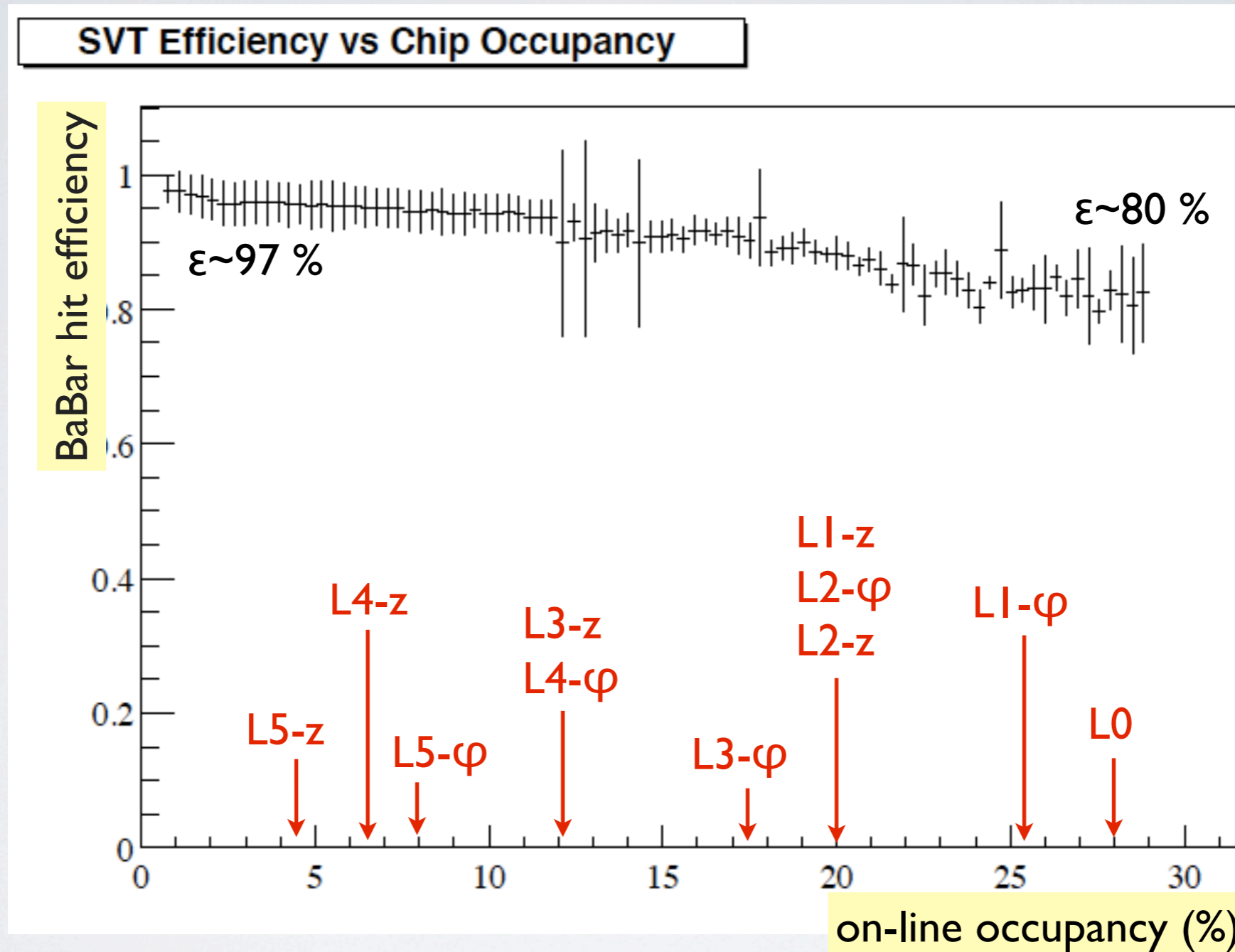


BaBar AD 707: Final Report of the SVT Long Term Task Force (2004):  
Study with **BaBar dimuon data** taken between Jan. and June 2003 (instantaneous luminosity increasing), of **hit efficiency** as a function of **chip on-line occupancy**.

→ how to translate this BaBar study to SuperB?

*Isabelle*

# on-line occupancy (3)



→ on-line occupancy in SuperB is 2 to 10x higher than in BaBar.

*Isabelle*

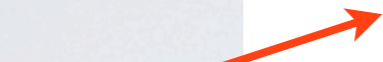
# estimation of BaBar hit detection efficiency (2)

measured y-axis  
of figure 12 from  
BAD 707



◆ Measured Babar hit efficiency = hit detection effi x hit to track effi

estimation of shadowing  
 $1 / (1 + R_{lost})$



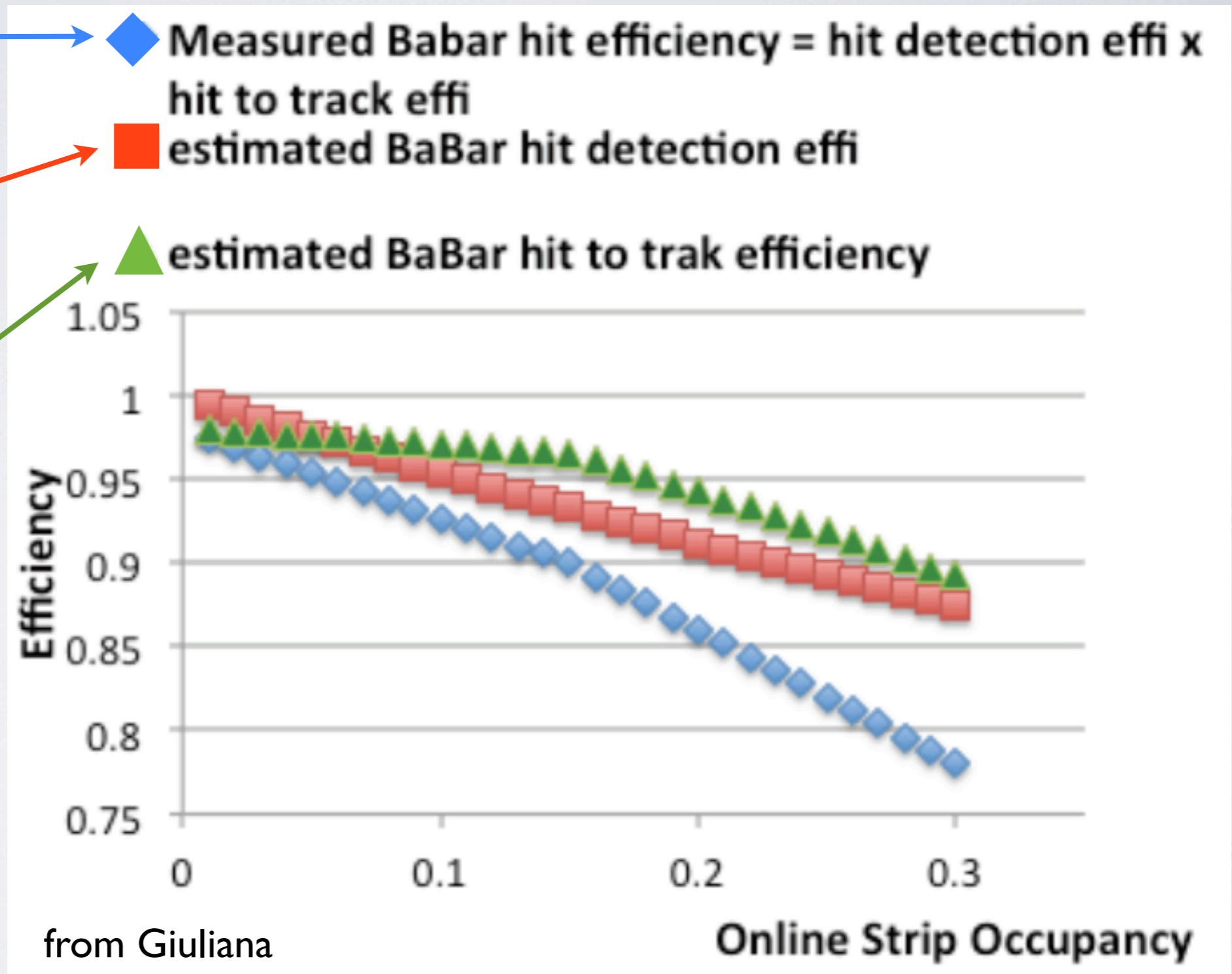
■ estimated BaBar hit detection effi



▲ estimated BaBar hit to trak efficiency

obtained through:

$$\blacktriangle = \blacklozenge / \blacksquare$$



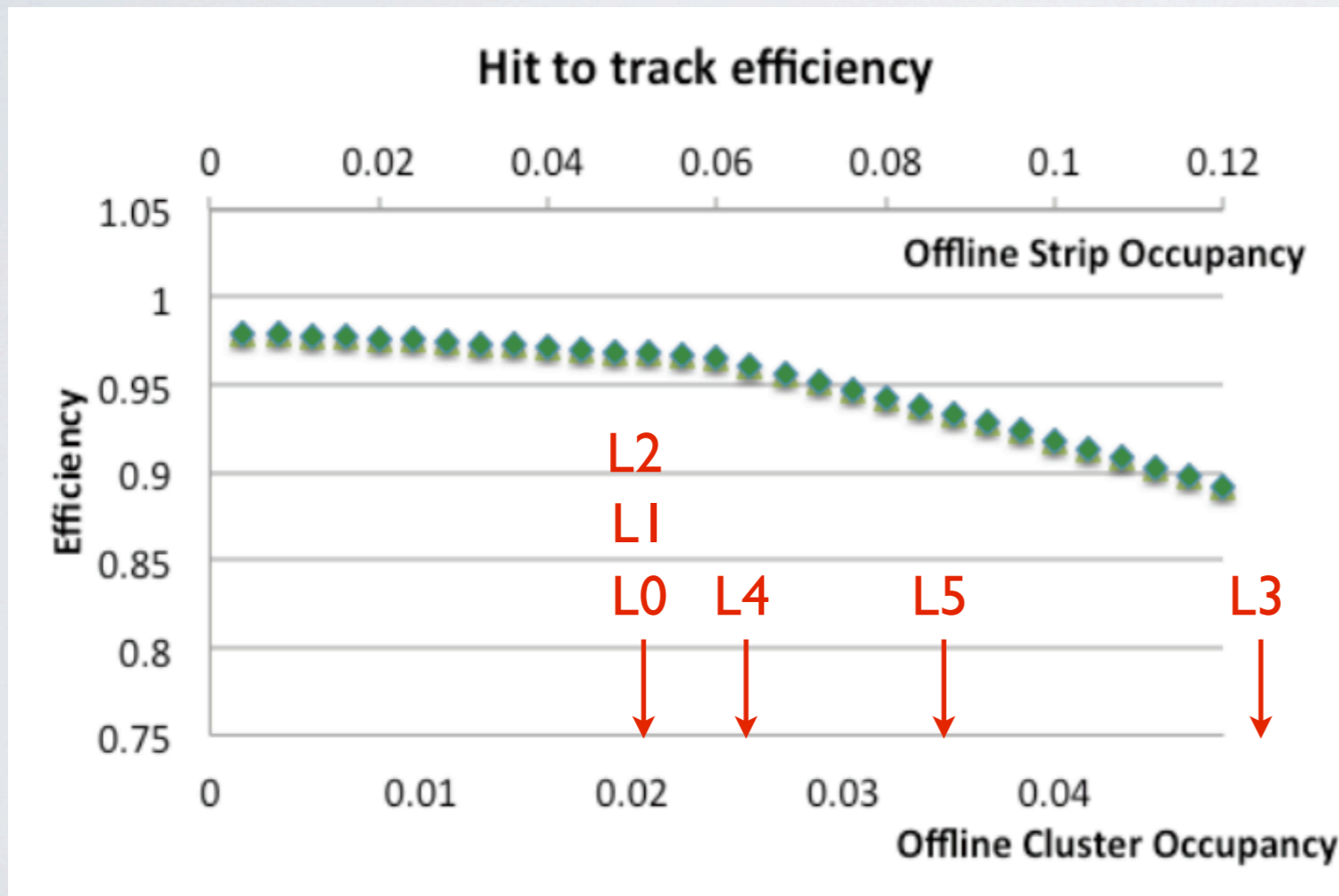
from Giuliana

Online Strip Occupancy

# hit-to-track matching efficiency as a function of off-line cluster occupancy (2)

And then see where SuperB Layers are on this curve:

Layer	View	Strip rate with renormalized area KHz estimates	time window used by neri) ns	offline cluster occupancy (x5 included)
0	1	9.32E+02	100	0.023
0	2	9.32E+02	100	
1	phi	8.479E+02	150	0.022
1	z	6.700E+02	150	
2	phi	6.649E+02	150	0.019
2	z	6.652E+02	150	
3	phi	5.770E+02	250	0.050
3	z	3.942E+02	250	
4	phi	1.241E+02	460	0.025
4	z	6.643E+01	460	
5	phi	8.034E+01	800	0.034
5	z	4.361E+01	800	



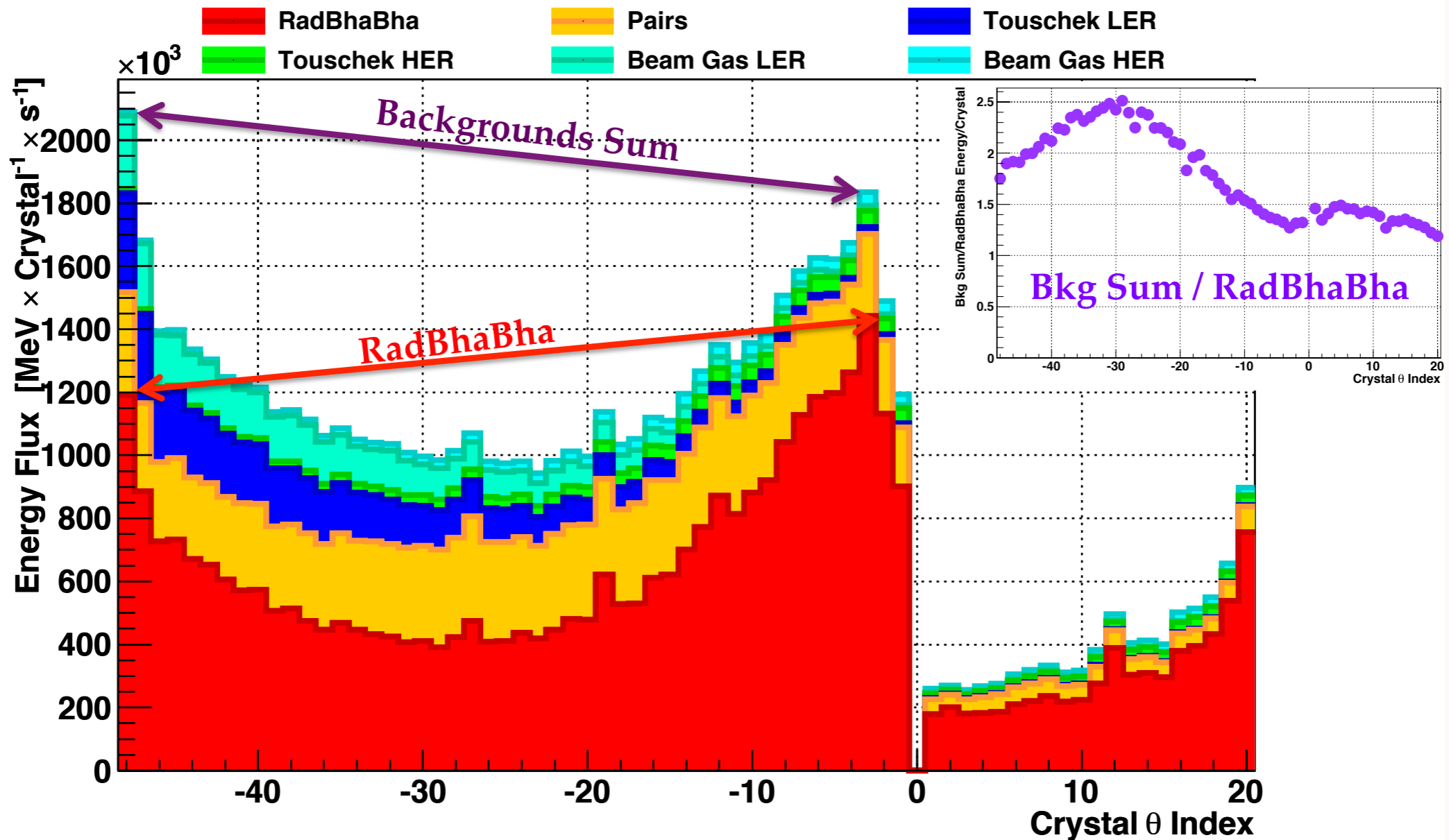
*Giuliana*

# estimation of SuperB hit efficiency

Layer	on-line strip occupancy (x5 included)	off-line cluster occupancy (x5 included)	hit detection efficiency (simulation) (x5 included)	hit-to-track matching efficiency (estimation from off-line cluster occ.)	total hit efficiency
0 $\varphi$	0.28	0.023	0.96	0.96	0.92
0 z	0.28		0.96		0.92
1 $\varphi$	0.25	0.022	0.88	0.96	0.84
1 z	0.20		0.89		0.85
2 $\varphi$	0.20	0.019	0.89	0.97	0.86
2 z	0.20		0.89		0.86
3 $\varphi$	0.20	0.050	0.77	0.88	0.68
3 z	0.17		0.86		0.76
4 $\varphi$	0.12	0.025	0.89	0.96	0.85
4 z	0.07		0.93		0.89
5 $\varphi$	0.08	0.034	0.86	0.93	0.80
5 z	0.04		0.91		0.85

# EMC STUDIES (STEFANO G.)

## Sum of Bkgs - 45 mm W

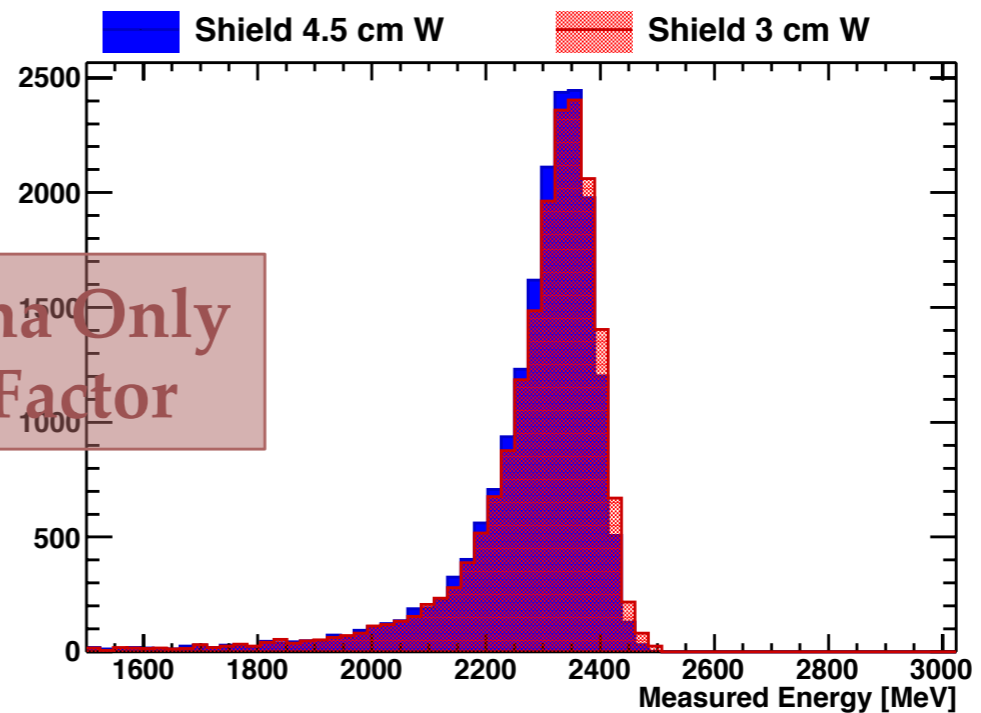
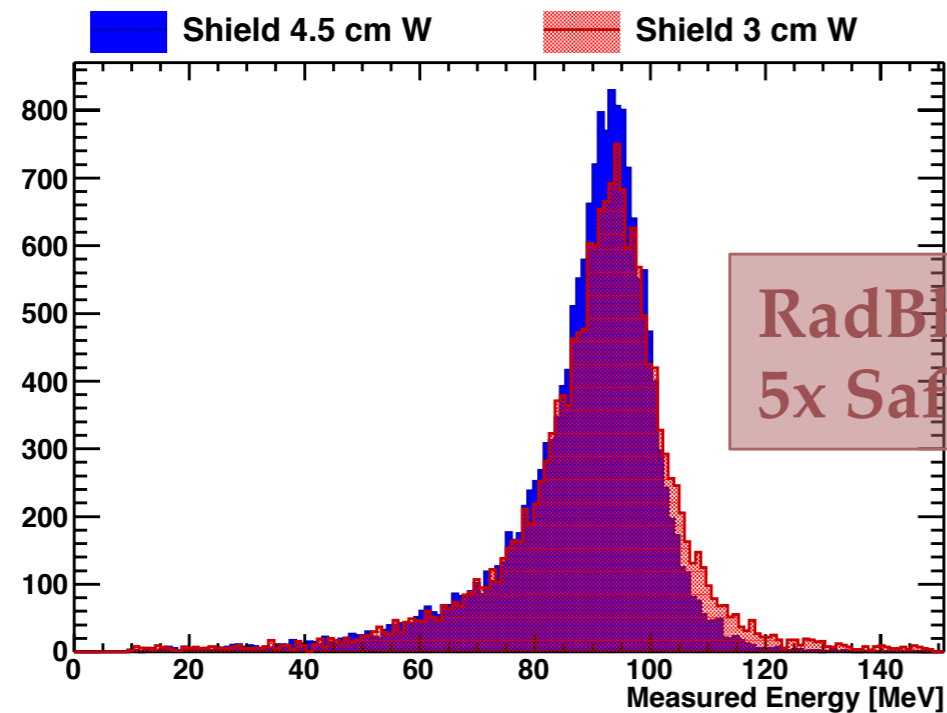
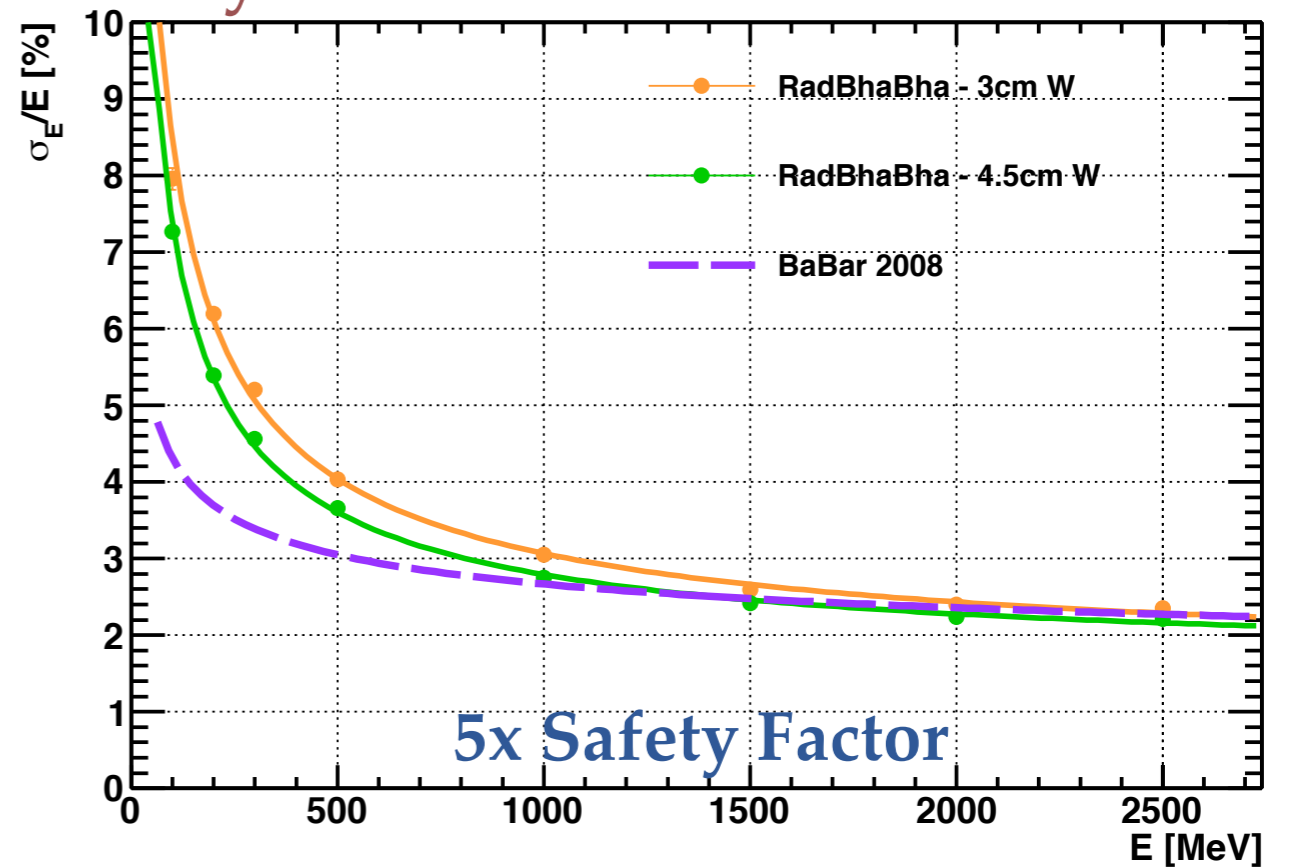
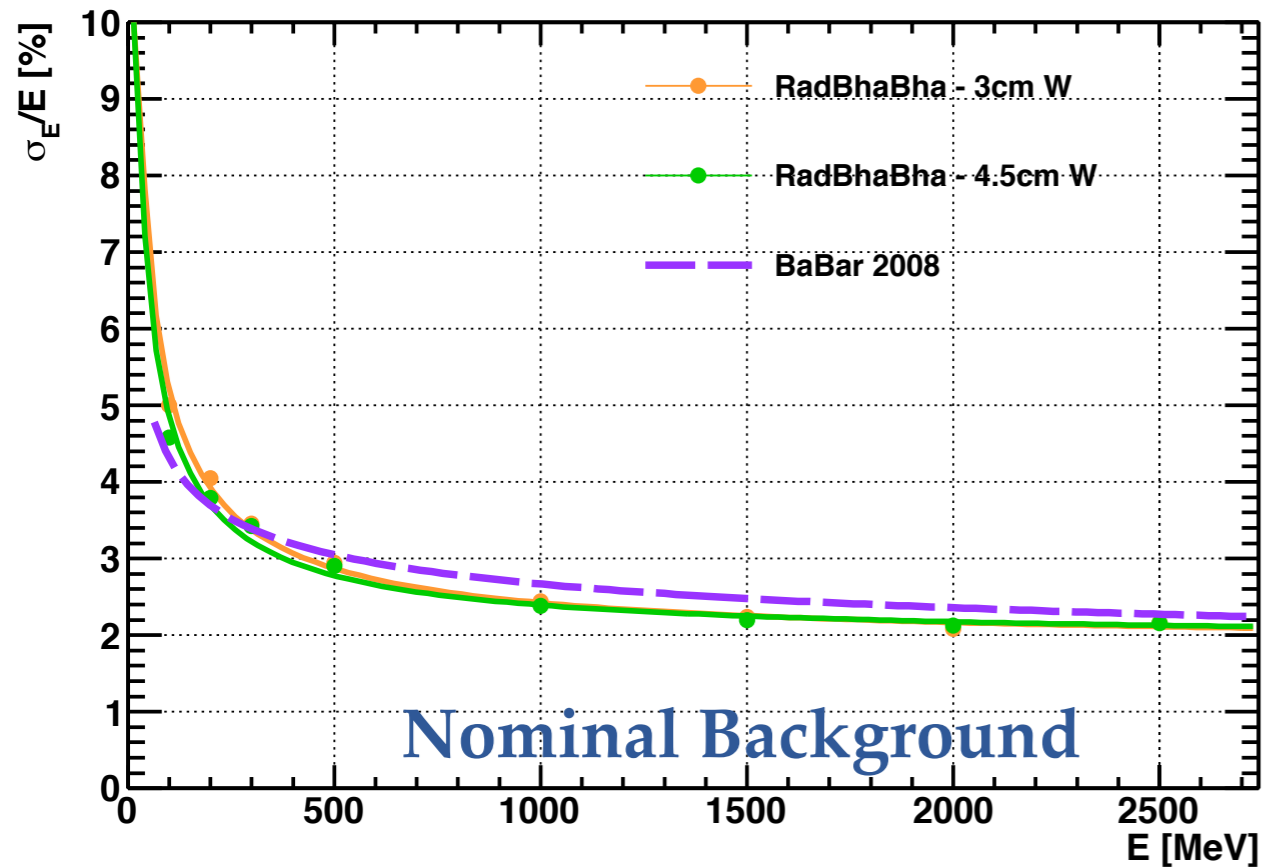


• 1/6/2012

EMC FullSim Studies - Stefano Germani • 3

# Barrel Resolution vs Shield

## RadBhaBha Only



RadBhaBha Only  
5x Safety Factor

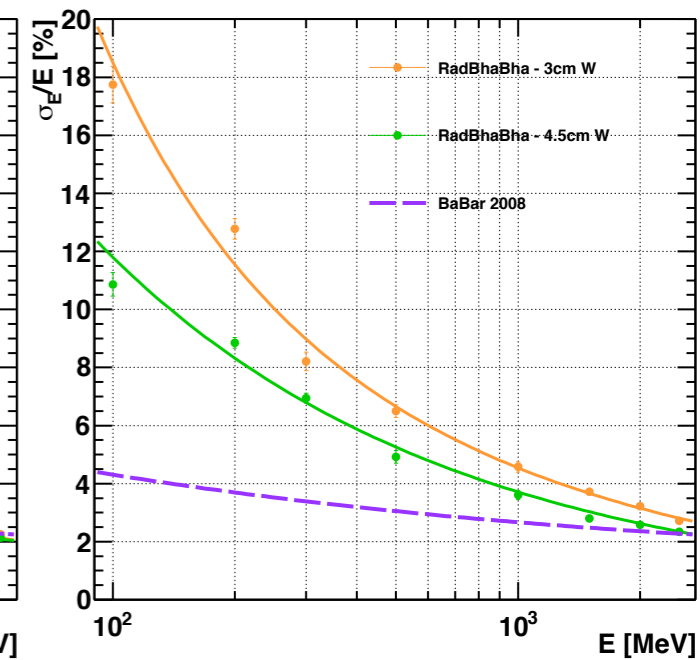
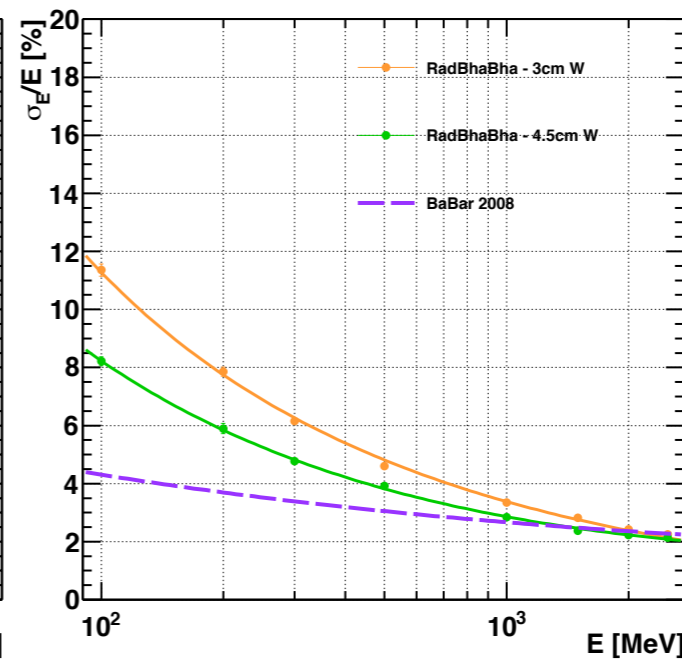
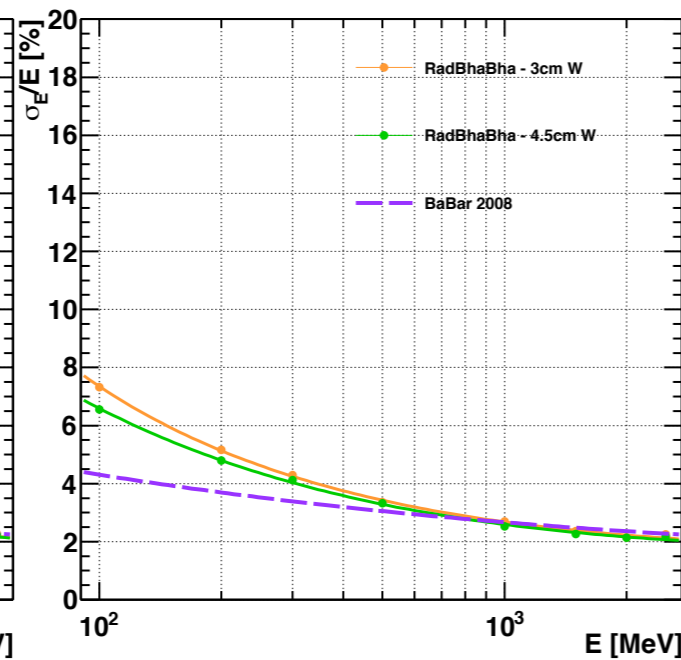
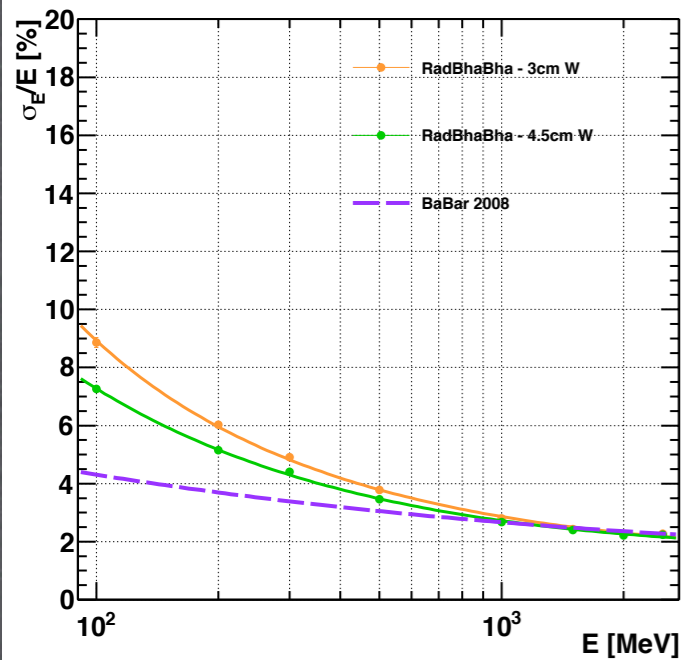
# Shield Effect vs Theta (x5)

Theta 100-140

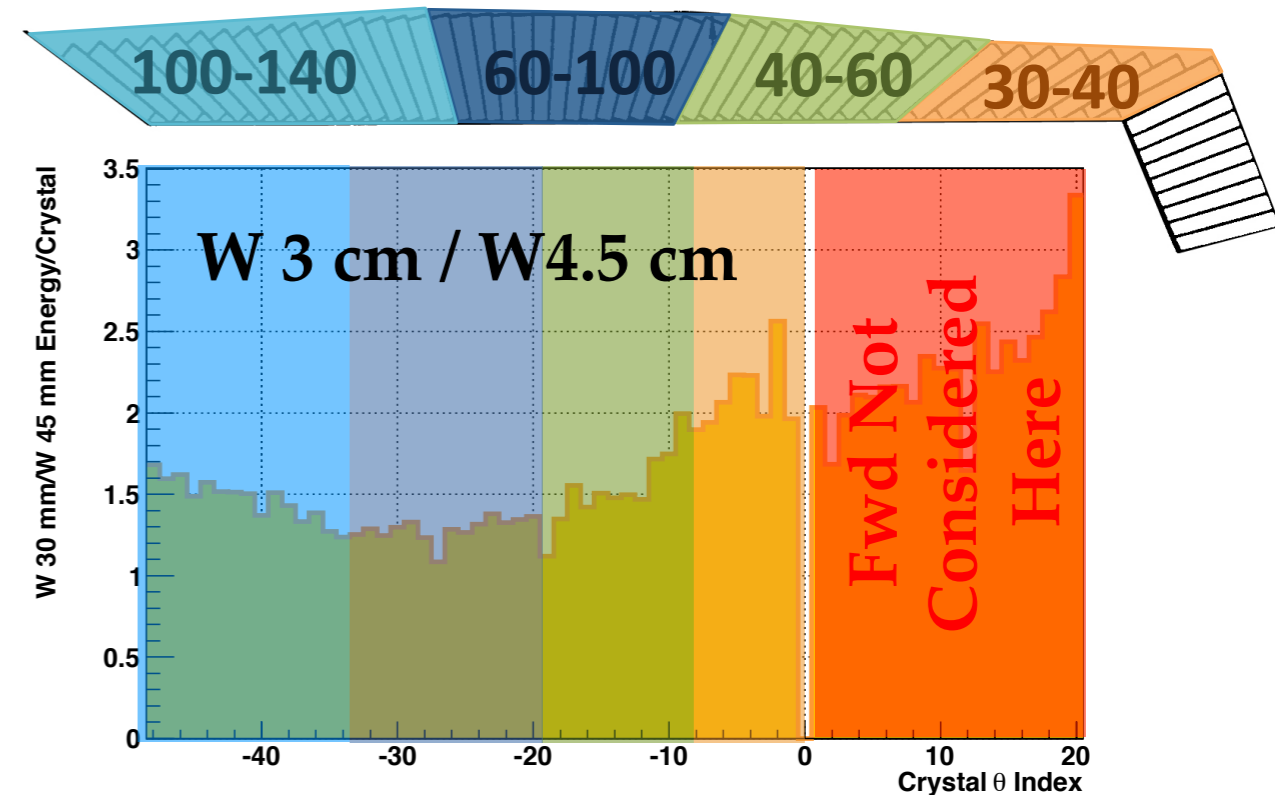
Theta 60-100

Theta 40-60

Theta 30-40

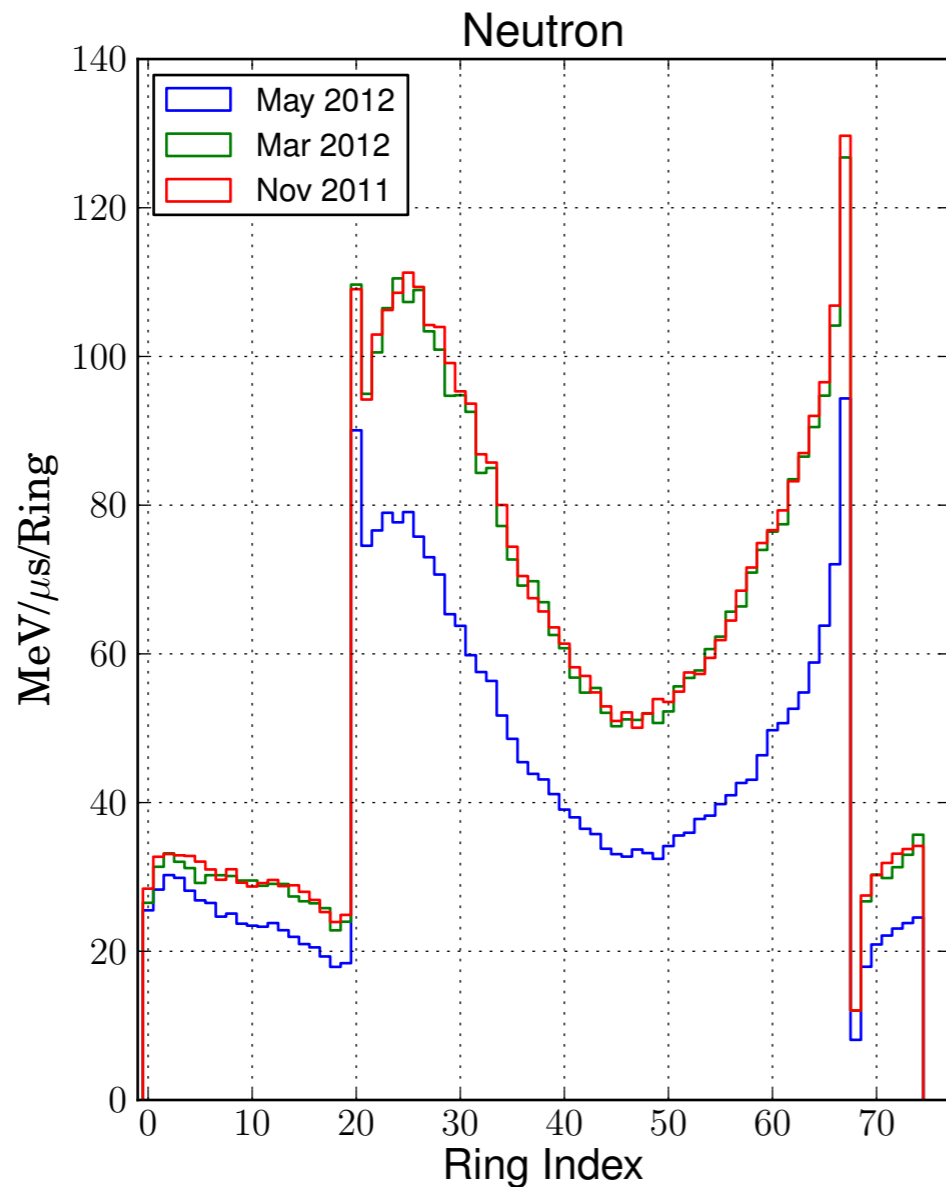


- ✓ Large performance improvement in fwd region
- ✓ Small effect in central barrel region
- ✓ More uniform Barrel performance across  $\theta$  angles

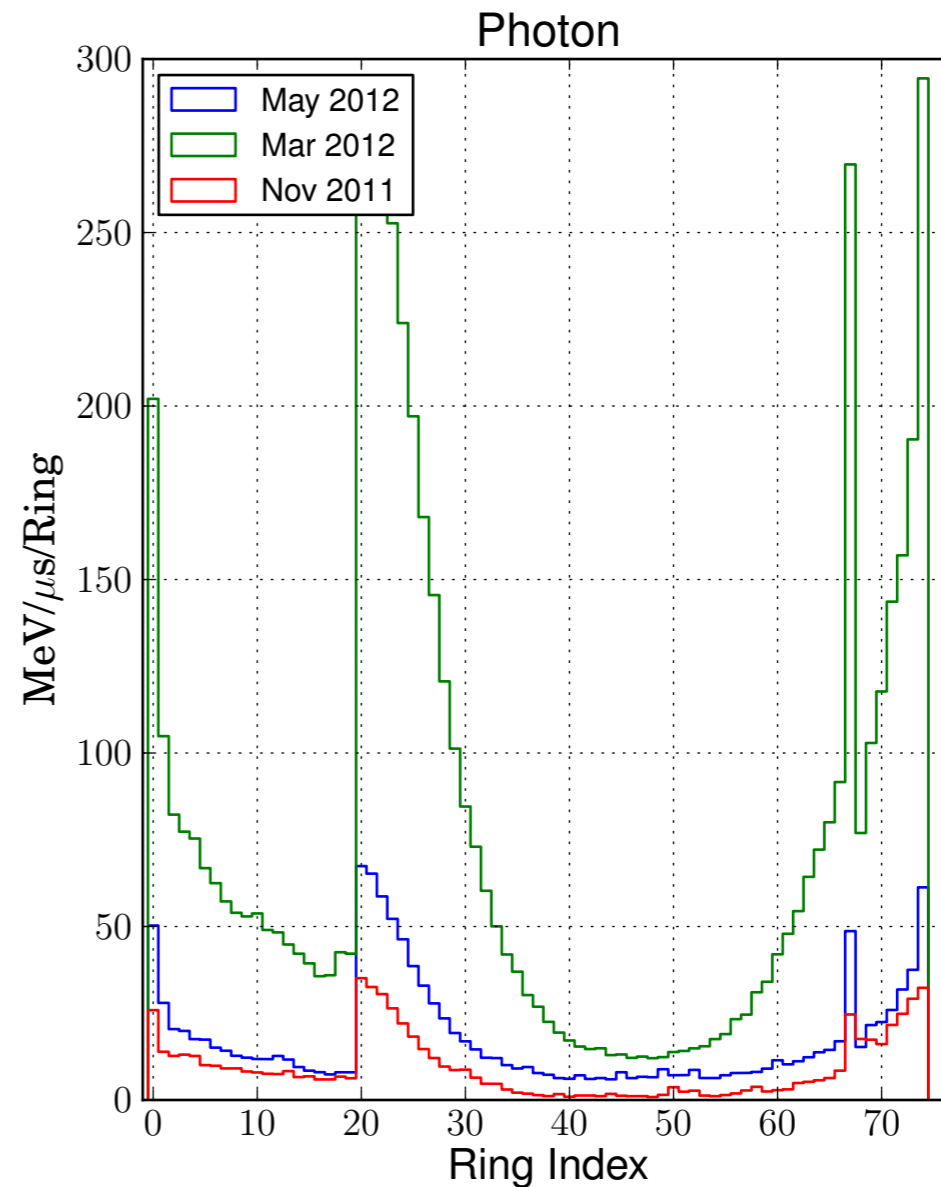




# Energy flux per ring



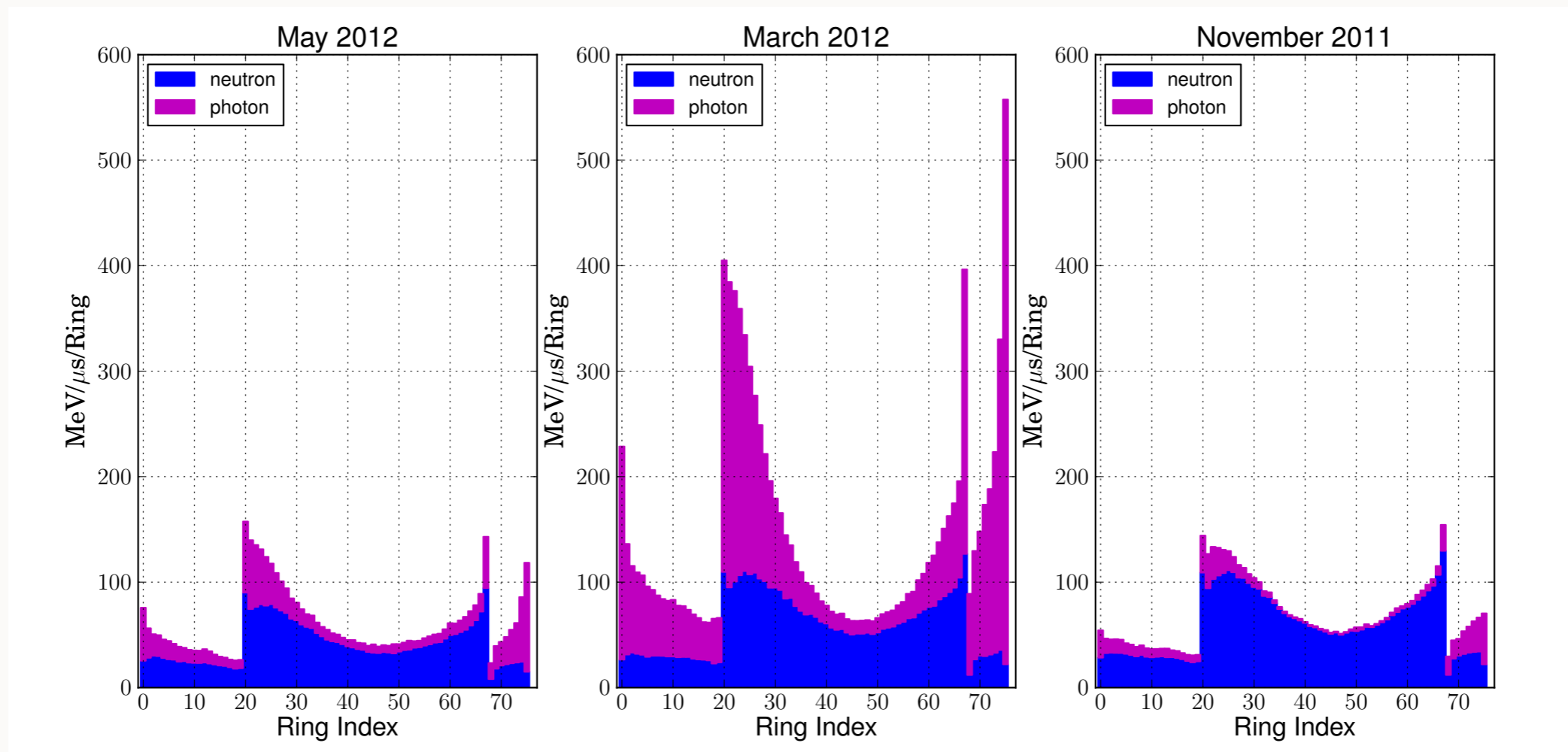
Neutron energy flux reduced by 20%  
–30% with new shielding



Photon energy flux reduced by ~4x  
with new shielding

# EMC PERFORMANCES

- Stefano is working on the determination of the EMC angular resolution in presence of background
- Energy resolution can be improved by reducing the neutron flux

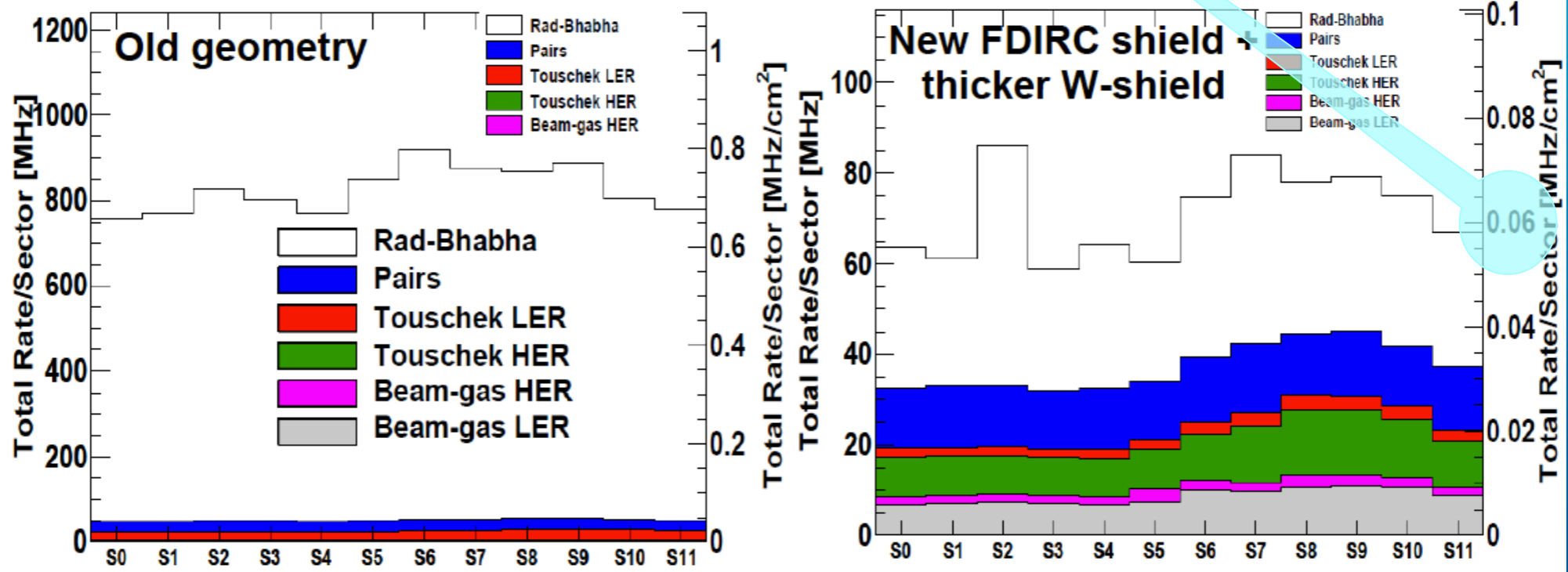


# FDIRC (D. ROBERTS)

FROM ALEJANDRO

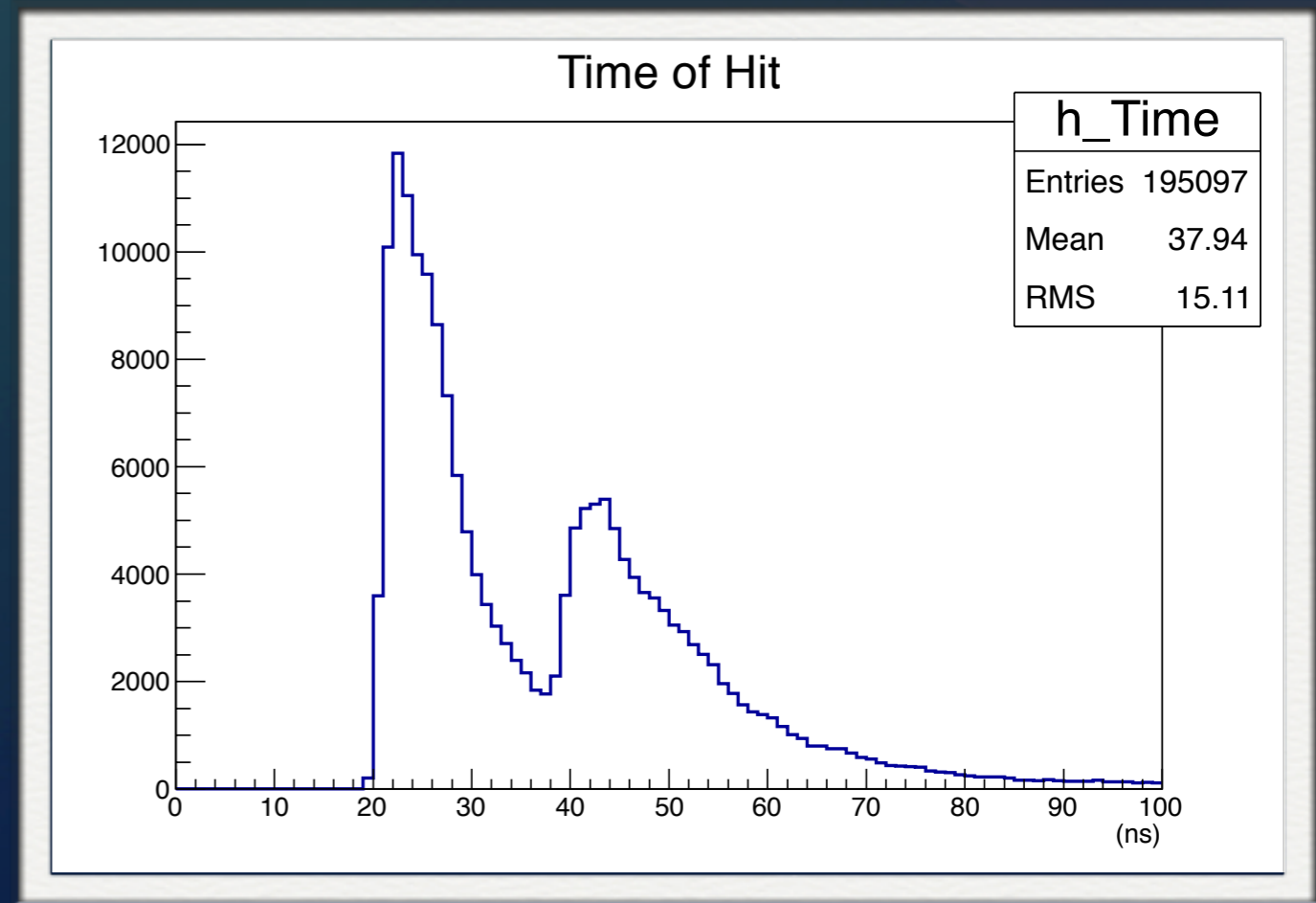
USE 0.06 MHz/cm<sup>2</sup> AS "NOMINAL" BACKGROUND RATE  
LOOK AT 1x, 2x, 5x, 10x, 20x NOMINAL

## Total bkg rates on FDIRC



# SIMULATION PARAMETERS

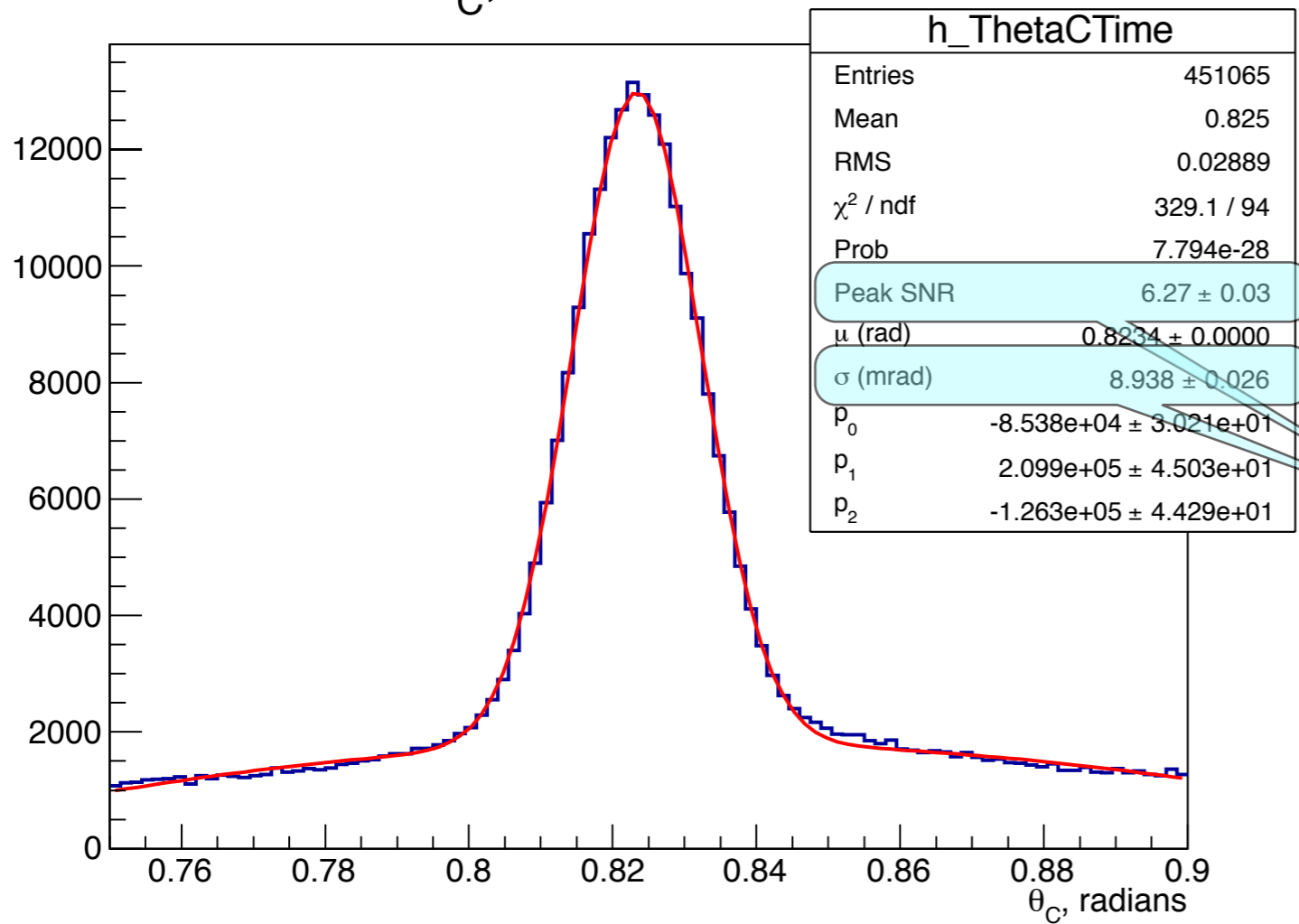
- BACKGROUND HITS:
  - RANDOMLY DISTRIBUTED OVER DETECTOR PLANE IN SPACE
  - UNIFORMLY RANDOM IN TIME FROM -25NS TO 500NS
- HITS "MASKED" IF THERE IS ANOTHER HIT <25NS EARLIER IN THE SAME PIXEL
- SIMULATED 10,000 COSMIC TRACKS
- USED "ABS( $k_x$ ), ABS( $k_y$ )" VERSION OF PHOTON DICTIONARY
- FULL FDIRC (48 TUBES IN A SECTOR)



*Doug*

# SINGLE PHOTON RESOLUTION

$\theta_C$ , After Time Cut

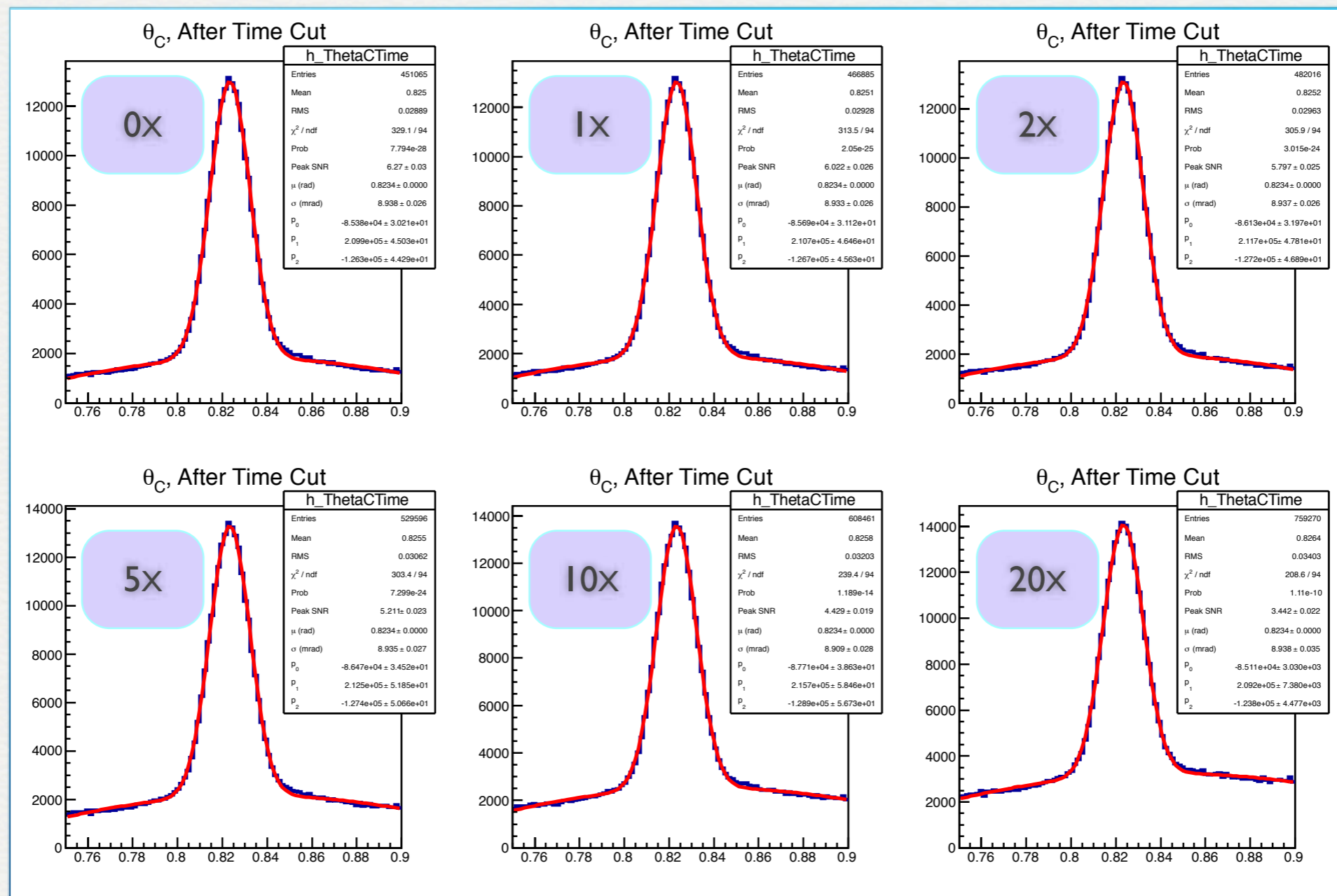


- NO BACKGROUND
- FIT TO A GAUSSIAN PLUS 2<sup>ND</sup> ORDER POLYNOMIAL
- LOOK AT  $\sigma$  AND PEAK SIGNAL-TO-NOISE RATIO
- TAILS COME FROM AMBIGUITIES

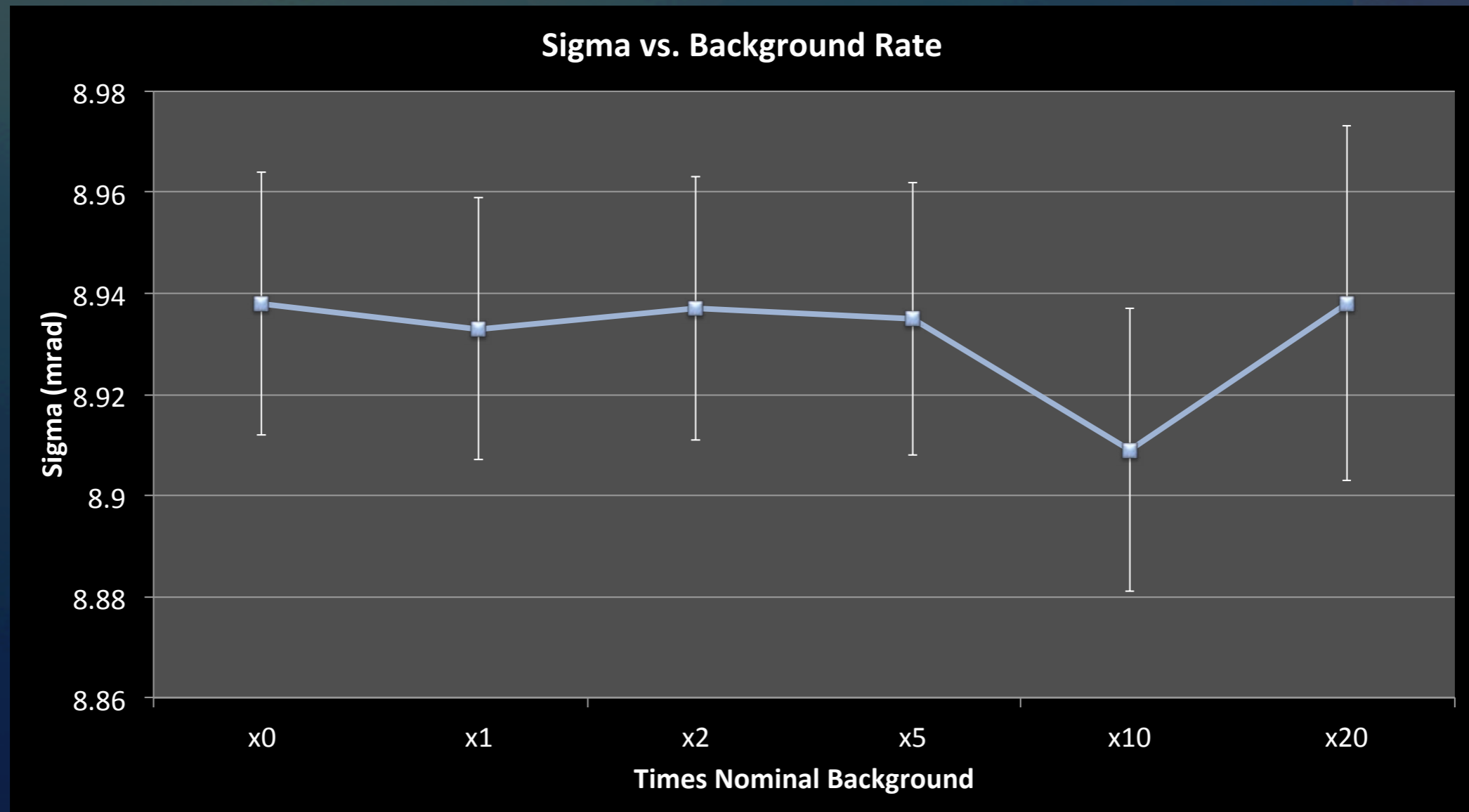
*Doug*

# LINE SHAPE VS. BACKGROUND

NO EFFECT ON SIGMA  
 BEGIN TO NOTICE EFFECT ON SNR AT 10X AND 20X  
 NOMINAL RATE



# SIGMA VS. BACKGROUND RATE



# PEAK SNR vs. BACKGROUND RATE

