

Running of the Newton and cosmological constant in quantum gravity

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Talk based on:

C. Branchina, V. Branchina, **FC**, A. Pernace, PRD 111 (2025) 10, 105018

C. Branchina, V. Branchina, **FC**, A. Pernace, PRD 111 (2025) 12, 125021

In part based on:

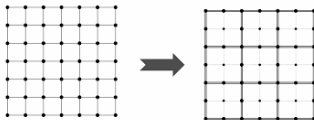
C. Branchina, V. Branchina, **FC**, R. Gandolfo, A. Pernace, PRD 112 (2025) 4, 045002

C. Branchina, V. Branchina, **FC**, R. Gandolfo, A. Pernace, arXiv:2505.07628



Wilson's Lesson

What is the Wilson's lesson all about?



Theory at Λ $\rightarrow$ Theory at $\Lambda/2$ $\rightarrow$...

S_Λ $\rightarrow$ $S_{\Lambda/2}$ $\rightarrow$...

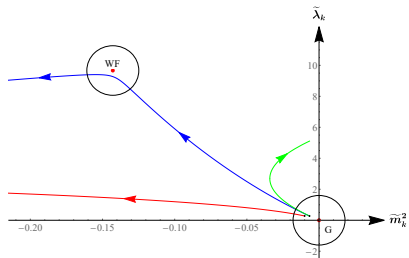
Effective Field Theory paradigm

Any QFT is an Effective Field Theory

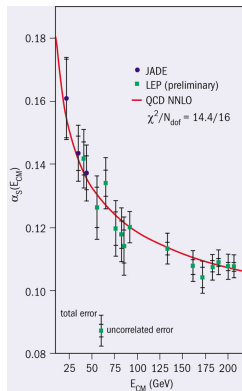
Steven Weinberg - Third Law of Progress in Theoretical Physics : you may use any degrees of freedom you like to describe a physical system, but if you use the wrong ones, you'll be sorry

... For theories in any dimension: ..., $d = 3$, $d = 4$, ...

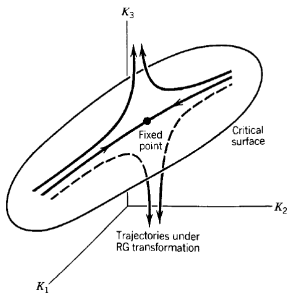
$d = 3$ dimensions : Wilson-Fisher



$d = 4$ dimensions : AF



RG flow



Renormalized theory: defined around a fixed point (critical surface)

Quantum Gravity

Perturbatively non-renormalizable

Is it **non-perturbatively renormalizable**?

Asymptotic safety scenario: Existence of a UV-attractive fixed point

One-loop effective action - Summary

Greetings to Arcangelo

One-loop Effective Action

- Euclidean action - Einstein-Hilbert truncation with spherical background

$$S_{\text{grav}} = \frac{1}{16\pi G} \int d^4x \sqrt{g} (-R + 2\Lambda_{\text{cc}}) \quad \Rightarrow \quad S_{\text{grav}}^{(a)} = \frac{\pi\Lambda_{\text{cc}}}{3G} a^4 - \frac{2\pi}{G} a^2$$

- One-loop effective action $\Gamma_{\text{grav}}^{1/} = S_{\text{grav}} + \delta S_{\text{grav}}^{1/}$

- **Diffeomorphism invariant** measure $\mathcal{D}\mu$: Fradkin-Vilkovisky

$$\mathcal{D}\mu \equiv \prod_x \left[g^{(a)00}(x) \left(g^{(a)}(x) \right)^{-1} \left(\prod_{\alpha \leq \beta} dh_{\alpha\beta}(x) \right) \left(\prod_{\rho} dv_{\rho}^*(x) \right) \left(\prod_{\sigma} dv_{\sigma}(x) \right) \right]$$

Related to **spacetime discretization** and Hamiltonian path integral

- Identify 1-loop corrections to $\frac{\Lambda_{\text{cc}}}{G}$ and $\frac{1}{G}$ with coefficients of a^4 and a^2 in $\delta S_{\text{grav}}^{1/}$

Taylor, Veneziano

One-loop correction to the action

$$\delta S_{\text{grav}}^{1/} = -\frac{1}{2} \log \frac{\det_1[-\tilde{\square}^{(1)} - 3] \det_2[-\tilde{\square}^{(0)} - 6]}{\det_0[-\tilde{\square}^{(2)} - 2a^2\Lambda + 8] \det_2[-\tilde{\square}^{(0)} - 2a^2\Lambda]} + \dots$$

- Argument of the log is **dimensionless**
- Eigenvalues $\lambda_n^{(s)}$ of **dimensionless** Laplace-Beltrami operator and degeneracies $D_n^{(s)}$

$$\lambda_n^{(s)} = n^2 + 3n - s \quad ; \quad D_n^{(s)} = \frac{2s+1}{3} \left(n + \frac{3}{2}\right)^3 - \frac{(2s+1)^3}{12} \left(n + \frac{3}{2}\right)$$

- Regularizing with hard cut / proper time ($N \gg 1$)

$$\delta S_{\text{grav}}^{1/} = - \left(\Lambda_{\text{cc}}^2 \log \frac{3\Lambda_{\text{cut}}^2}{\Lambda_{\text{cc}}} \right) a^4 + \left(-3\Lambda_{\text{cut}}^2 + 8\Lambda_{\text{cc}} \log \frac{3\Lambda_{\text{cut}}^2}{\Lambda_{\text{cc}}} \right) a^2 + \dots$$

- $N \longrightarrow$ **physical cutoff Λ** through the De Sitter solution $a_{\text{ds}} = \sqrt{\frac{3}{\Lambda_{\text{cc}}}}$

$$\Lambda_{\text{cut}} \sim M_P = \frac{N}{a_{\text{ds}}} = N \sqrt{\frac{\Lambda_{\text{cc}}}{3}}$$

One-loop corrections to $\frac{\Lambda_{\text{cc}}}{G}$ and $\frac{1}{G}$

Coefficients of a^4 and a^2 identify the one-loop corrections to $\frac{\Lambda_{\text{cc}}}{G}$ and $\frac{1}{G}$

$$\frac{\Lambda_{\text{cc}}^{1/}}{G^{1/}} = \frac{\Lambda_{\text{cc}}}{G} \left(1 - \frac{3G\Lambda_{\text{cc}}}{\pi} \log \frac{3\Lambda_{\text{cut}}^2}{\Lambda_{\text{cc}}} \right) + \text{finite}$$

$$\frac{1}{G^{1/}} = \frac{1}{G} \left[1 + \frac{G}{2\pi} \left(3\Lambda_{\text{cut}}^2 - 8\Lambda_{\text{cc}} \log \frac{3\Lambda_{\text{cut}}^2}{\Lambda_{\text{cc}}} \right) \right] + \text{finite}$$

Loop corrections $\rightarrow$ only mild (log) correction to $\rho \implies$

In pure gravity **no naturalness problem arises**:
the bare cosmological constant Λ_{cc} does not need to be $\sim M_P^2$.

We may **naturally** have $\Lambda_{\text{cc}} \ll M_P^2$, and so

$$\Lambda_{\text{cc}}^{1/} \sim \Lambda_{\text{cc}}$$

Renormalization group equations

Two methods to obtain **RG equations** for Λ_{cc} and G :

- **Hard cut** on sum over **eigenvalues**
- **Proper time**

Sum over eigenvalues

UV action

$$S_{\text{grav}}^{\text{UV}}[g_{\mu\nu}] \equiv S_N[g_{\mu\nu}] = \frac{1}{16\pi G_N} \int d^4x \sqrt{g} (-R + 2\Lambda_N) \quad (N \text{ integer})$$

As before: physical cutoff $\Lambda_{\text{cut}} (\sim M_P) = N/a_{\text{dS}}$

Wilsonian action: $S_L[g_{\mu\nu}^{(a)}] \quad (L \text{ integer}, L < N ; \delta L \ll L)$

RG equation: $S_{L-\delta L}[g_{\mu\nu}^{(a)}] = S_L[g_{\mu\nu}^{(a)}] + \delta S_L \equiv S_L[g_{\mu\nu}^{(a)}] + \sum_{n=L-\delta L}^L f_L(n)$

where

$$\delta S_L = -\frac{1}{2} \log \frac{\det_1[-\widetilde{\square}^{(1)} - 3] \det_2[-\widetilde{\square}^{(0)} - 6]}{\det_0[-\widetilde{\square}^{(2)} - 2a^2\Lambda_L + 8] \det_2[-\widetilde{\square}^{(0)} - 2a^2\Lambda_L]}$$

$\Updownarrow$

$$f_L = D_n^{(2)} \log(\lambda_n^{(2)} - 2a^2\Lambda_L + 8) + D_n^{(0)} \log(\lambda_n^{(0)} - 2a^2\Lambda_L) - D_n^{(1)} \log(\lambda_n^{(1)} - 3) - D_n^{(0)} \log(\lambda_n^{(0)} - 6)$$

Differential form: $\frac{\partial}{\partial L} S_L = - \left(\frac{\partial}{\partial L} \sum_{n=2}^{L-2} f_L(n) \right)_{\Lambda_L, G_L} \quad (\text{note } L_{\text{min}} = 4)$

Proper time

● Proper time method

$$\det_i(-\widetilde{\square}^{(s)} - \alpha) = e^{-\int_{1/L^2}^{1/(L-\delta L)^2} \frac{d\tau}{\tau} K_i^{(s)}(\tau)} \quad ; \quad K_i^{(s)}(\tau) = \sum_{n=s+i}^{+\infty} D_n^{(s)} e^{-\tau(\lambda_n^{(s)} - \alpha)}$$

Differential RG equation:

$$\frac{\partial}{\partial L} S_L[g_{\mu\nu}^{(a)}] = - \left(\frac{\partial}{\partial L} \delta S_L^{1l} \right)_{\Lambda_L, G_L}$$

Leads to **same result** obtained with **direct sum**

● Computing the r.h.s. (**direct sum** or **heat kernel expansion**) ($L \gg 1$):

$$L \frac{\partial}{\partial L} S_L = 2\Lambda_L^2 a^4 + 2\Lambda_L (L^2 - 8) a^2 + \dots$$

From which we obtain the **RG equations**

$$\begin{aligned} L \frac{d}{dL} \left(\frac{\Lambda_L}{G_L} \right) &= \frac{6}{\pi} \Lambda_L^2 \\ L \frac{d}{dL} \left(\frac{1}{G_L} \right) &= -\frac{\Lambda_L}{\pi} (L^2 - 8) \end{aligned} \quad \Rightarrow \quad \begin{aligned} L \frac{d\Lambda_L}{dL} &= \frac{G_L \Lambda_L^2}{\pi} (L^2 - 2) \\ L \frac{dG_L}{dL} &= \frac{G_L^2 \Lambda_L}{\pi} (L^2 - 8) \end{aligned}$$

RG equations - Solutions

In the $L^2 \gg 1$ approximation we have

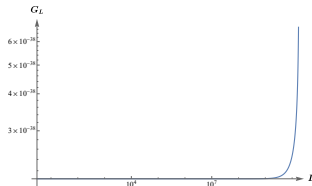
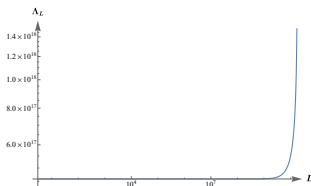
$$(\Lambda_{\text{cc}} \equiv \Lambda_N, G \equiv G_N)$$

$$\Lambda_L = \frac{\Lambda_{\text{cc}}}{\sqrt{1 + \frac{G\Lambda_{\text{cc}}}{\pi}(N^2 - L^2)}} \quad \longleftrightarrow \quad \Lambda_{\text{cut}} = N\sqrt{\frac{\Lambda_{\text{cc}}}{3}}$$

$$G_L = \frac{G}{\sqrt{1 + \frac{G\Lambda_{\text{cc}}}{\pi}(N^2 - L^2)}}$$

$$\Lambda_L = \frac{\Lambda_{\text{cc}}}{\sqrt{1 + \frac{G}{\pi}(3\Lambda_{\text{cut}}^2 - \Lambda_{\text{cc}}L^2)}}$$

$$G_L = \frac{G}{\sqrt{1 + \frac{G}{\pi}(3\Lambda_{\text{cut}}^2 - \Lambda_{\text{cc}}L^2)}}$$



- Sign of Λ_L and G_L fixed
- Λ_L monotonically increasing with L

From $L \rightarrow$ physical running scale k

$$L \frac{d\Lambda_L}{dL} = \frac{G_L \Lambda_L^2}{\pi} (L^2 - 2)$$

$$L \frac{dG_L}{dL} = \frac{G_L^2 \Lambda_L}{\pi} (L^2 - 8)$$

Physical running scale k

$$k \equiv \frac{L}{\bar{a}_L} = L \sqrt{\frac{\Lambda_L}{3}} \quad (k_{\text{IR}} \leq k \leq \Lambda_{\text{cut}} ; k_{\text{IR}} = (\frac{16}{3} \Lambda_4)^{1/2})$$

The RG equations become $(\Lambda_k \equiv \Lambda_L, G_k \equiv G_L)$

$$k \frac{d\Lambda_k}{dk} = \frac{3G_k}{\pi} \frac{\Lambda_k (k^2 - \frac{2}{3}\Lambda_k)}{1 + \frac{3G_k}{2\pi} (k^2 - \frac{2}{3}\Lambda_k)}$$

$$k \frac{dG_k}{dk} = \frac{3G_k^2}{\pi} \frac{k^2 - \frac{8}{3}\Lambda_k}{1 + \frac{3G_k}{2\pi} (k^2 - \frac{2}{3}\Lambda_k)}$$

Solutions

$$k \frac{d\Lambda_k}{dk} = \frac{3G_k}{\pi} \frac{\Lambda_k \left(k^2 - \frac{2}{3}\Lambda_k\right)}{1 + \frac{3G_k}{2\pi} \left(k^2 - \frac{2}{3}\Lambda_k\right)}$$

$$k \frac{dG_k}{dk} = \frac{3G_k^2}{\pi} \frac{k^2 - \frac{8}{3}\Lambda_k}{1 + \frac{3G_k}{2\pi} \left(k^2 - \frac{2}{3}\Lambda_k\right)}$$

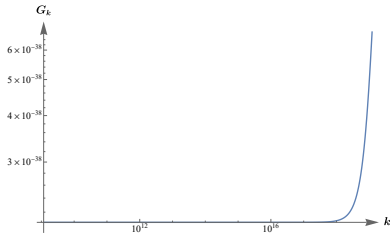
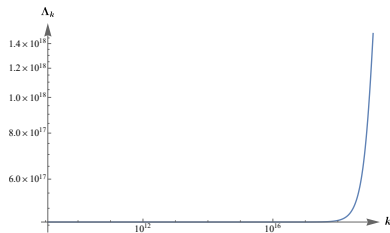
$$\frac{\Lambda_k}{k^2} = \frac{3}{L^2} \ll 1 \rightarrow$$

$$k \frac{d\Lambda_k}{dk} = \frac{3G_k}{\pi} \frac{\Lambda_k k^2}{1 + \frac{3G_k}{2\pi} k^2}$$

$$k \frac{dG_k}{dk} = \frac{3G_k^2}{\pi} \frac{k^2}{1 + \frac{3G_k}{2\pi} k^2}$$

$$\Lambda_k = \frac{k^2 \Lambda_{\text{cut}}}{2 \left(\Lambda_{\text{cut}}^2 + \frac{\pi}{3G} \right)} \left[1 + \sqrt{1 + \frac{4\pi}{3G} \left(\Lambda_{\text{cut}}^2 + \frac{\pi}{3G} \right) \frac{1}{k^4}} \right]$$

$$G_k = \frac{k^2 G}{2 \left(\Lambda_{\text{cut}}^2 + \frac{\pi}{3G} \right)} \left[1 + \sqrt{1 + \frac{4\pi}{3G} \left(\Lambda_{\text{cut}}^2 + \frac{\pi}{3G} \right) \frac{1}{k^4}} \right]$$



Naturalness

Taking for G the natural value $G \sim M_P^{-2}$

For $\sim$ any value of Λ_{cc}

$$G_{\text{IR}} \sim G \sim M_P^{-2}$$

$$\Lambda_{\text{IR}} \sim \Lambda_{\text{cc}}$$

As for the one-loop result

In pure gravity **no naturalness problem arises**:
the bare cosmological constant Λ_{cc} does not need to be $\sim M_P^2$.

We may **naturally** have $\Lambda_{\text{cc}} \ll M_P^2$, and so

$$\Lambda_{\text{cc}}^{1/4} \sim \Lambda_{\text{cc}}$$

Fixed points

Dimensionful $\rightarrow$ Dimensionless equations

$$k \frac{d\Lambda_k}{dk} = \frac{3G_k}{\pi} \frac{\Lambda_k \left(k^2 - \frac{2}{3}\Lambda_k\right)}{1 + \frac{3G_k}{2\pi} \left(k^2 - \frac{2}{3}\Lambda_k\right)}$$

$$k \frac{dG_k}{dk} = \frac{3G_k^2}{\pi} \frac{k^2 - \frac{8}{3}\Lambda_k}{1 + \frac{3G_k}{2\pi} \left(k^2 - \frac{2}{3}\Lambda_k\right)}$$

Now: $t = \ln \frac{k}{k_0} \quad ; \quad \lambda_t = \frac{\Lambda_k}{k^2} \quad ; \quad g_t = k^2 G_k$

Dimensionless RG equations

$$\frac{d\lambda_t}{dt} = -2\lambda_t + \frac{3g_t}{\pi} \frac{\lambda_t \left(1 - \frac{2}{3}\lambda_t\right)}{1 + \frac{3g_t}{2\pi} \left(1 - \frac{2}{3}\lambda_t\right)} \equiv \beta_\lambda(\lambda, g)$$

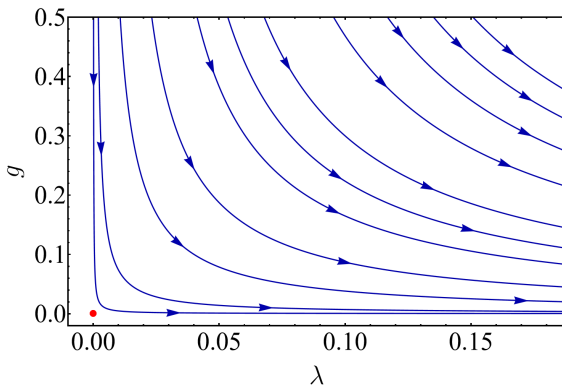
$$\frac{dg_t}{dt} = 2g_t + \frac{3g_t^2}{\pi} \frac{1 - \frac{8}{3}\lambda_t}{1 + \frac{3g_t}{2\pi} \left(1 - \frac{2}{3}\lambda_t\right)} \equiv \beta_g(\lambda, g)$$

Fixed points from $\beta_\lambda = 0$; $\beta_g = 0$

$$(\lambda, g)_1 = (0, 0)$$

$$(\lambda, g)_2 = (0, -\pi/3)$$

RG flows and Fixed points



$(\lambda, g)_1 = (0, 0)$ axes $\lambda = 0$, $g = 0$ UV-repulsive/attractive respectively

$(\lambda, g)_2 = (0, -\pi/3)$ **unphysical** UV-attractive fixed point

NO sign of any **physical UV-attractive** fixed point (**AS**)

What generates AS behaviour?

$$L \frac{\partial}{\partial L} S_L = 2\Lambda_L^2 a^4 + 2\Lambda_L \left(L^2 - 8 \right) a^2 + \frac{L^4}{3} - \frac{34L^2}{3} + \frac{1859}{45}$$

Identify the running scale k as

$$k = L/a$$

(rather than $k = L/\bar{a}_L$)

$$k \frac{d}{dk} \Lambda_k = \frac{G_k}{\pi} \left[k^4 + 6\Lambda_k \left(k^2 + \Lambda_k \right) - \Lambda_k \frac{34k^2 + 48\Lambda_k}{6} \right] ; \quad k \frac{d}{dk} G_k = -G_k^2 \frac{34k^2 + 48\Lambda_k}{6\pi}$$

Dimless RG equations

$$\frac{d\lambda_t}{dt} = -2\lambda_t + \frac{g_t}{\pi} \left[1 + 6\lambda_t (1 + \lambda_t) - \lambda_t \frac{34 + 48\lambda_t}{6} \right] \equiv \beta_\lambda (\lambda_t, g_t)$$

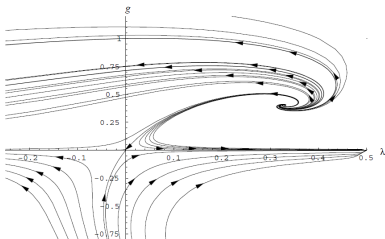
$$\frac{dg_t}{dt} = 2g_t - g_t^2 \frac{34 + 48\lambda_t}{6\pi} \equiv \beta_g (\lambda_t, g_t)$$

Fixed points

$$(\lambda, g)_1 = (0, 0) \quad (\lambda, g)_2 = \left(-\frac{8 + \sqrt{154}}{30}, \frac{2\pi}{23} (53 + 4\sqrt{154}) \right)$$

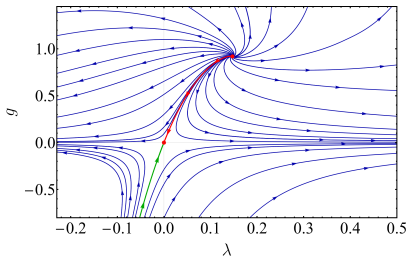
$$(\lambda, g)_3 = \left(\frac{\sqrt{154} - 8}{30}, \frac{2\pi}{23} (53 - 4\sqrt{154}) \right)$$

What generates the AS behaviour? - 2



M.Reuter, Phys.Rev.D 57 (1998) 971-985

M.Reuter and F.Saueressig, Phys.Rev.D 65 (2002) 065016



Our result with the identification $k = \frac{L}{a}$

$$(\lambda, g)_3 = (0.147, 0.918)$$

UV-attractive fixed point (AS)

Identification of k with a seems to be at the root of AS flow

... More on this point ...

Wetterich-Reuter equation

● Effective Average Action $\Gamma_k[g, \bar{g}]$

$$k \partial_k \Gamma_k[g, \bar{g}] = \frac{1}{2} \text{Tr} \left[\left(\kappa^{-2} \Gamma_k^{(2)}[g, \bar{g}] + R_k^{\text{grav}}[\bar{g}] \right)^{-1} k \partial_k R_k^{\text{grav}}[\bar{g}] \right] \\ - \text{Tr} \left[\left(-\mathcal{M}[g, \bar{g}] + R_k^{\text{gh}}[\bar{g}] \right)^{-1} k \partial_k R_k^{\text{gh}}[\bar{g}] \right]$$

M. Reuter, C. Wetterich

$\mathcal{M}[g, \bar{g}]$ kinetic term of the **ghosts**

$$\mathcal{M}[g, \bar{g}]^\mu{}_\nu = \bar{g}^{\mu\rho} \bar{g}^{\sigma\lambda} \bar{D}_\lambda (g_{\rho\nu} D_\sigma + g_{\sigma\nu} D_\rho) - \bar{g}^{\rho\sigma} \bar{g}^{\mu\lambda} \bar{D}_\lambda g_{\sigma\nu} D_\rho$$

● Consider $\bar{g}_{\mu\nu} = g_{\mu\nu}^{(a)}$ **spherical background** of radius **a**

$R_k^{\text{grav}}[\bar{g}]$ and $R_k^{\text{gh}}[\bar{g}]$ have the form

$$R_k[\bar{g}] = \mathcal{Z}_k k^2 R^{(0)}(-\square/k^2) \quad , \quad \square \equiv \bar{g}^{\mu\nu} \bar{D}_\mu \bar{D}_\nu$$

Modes $p^2 > k^2$ integrated out ; modes $p^2 < k^2$ suppressed by R_k :

$\Rightarrow$ “Tr” effectively restricted to eigenmodes of $-\square$ with eigenvalues

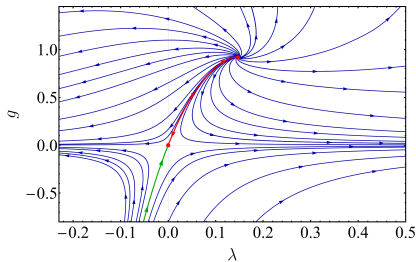
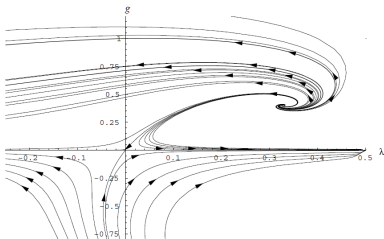
$$p^2 \sim \frac{n^2}{a^2} \sim k^2$$

RG flows from Wetterich/Reuter equation

$$p^2 \sim \frac{n^2}{a^2} \sim k^2$$

Eigenvalues of $-\square$ (not $-\tilde{\square}$) used to introduce sliding scale k $\lambda_n \sim \frac{n^2}{a^2}$

$\Rightarrow$ **spurious** k^4 terms $\Rightarrow$ **“AS flow”**



M.Reuter, Phys.Rev.D 57 (1998) 971-985

Our result with the identification $k = \frac{L}{a}$

M.Reuter and F.Saueressig, Phys.Rev.D 65 (2002) 065016

Proper time RG

With Fujikawa measure (**not diffeomorphism invariant**)

$$\delta S_{\text{grav}}^{1l} = -\frac{1}{2} \log \frac{\det_1[-\square_a^{(1)} - 3/a^2] \det_2[-\square_a^{(0)} - 6/a^2]}{\det_0[-\square_a^{(2)} - 2\Lambda + 8/a^2] \det_2[-\square_a^{(0)} - 2\Lambda]}$$

Wilsonian **RG strategy** implemented introducing an **IR regulator k**

$$\int_{1/\Lambda_{\text{cut}}^2}^{+\infty} ds \longrightarrow \int_{1/\Lambda_{\text{cut}}^2}^{1/k^2} ds$$

Equivalently introducing a **smooth function** $f_k(s)$ ($t \equiv \log \frac{k}{k_0}$)

$$\partial_t \widehat{S}_k[g, \bar{g}] = -\frac{1}{2} \text{Tr} \int_0^{+\infty} \frac{ds}{s} \partial_t f_k(s) \left[e^{-s \widehat{S}_k^{(2)}} - 2 e^{-s S_{\text{ghost}}^{(2)}} \right]$$

- $\widehat{S}_k^{(2)}$ contains the dimensionful operators $-\square_a$: eigenvalues $\widehat{\lambda}_n \sim \frac{n^2}{a^2}$
- $\partial_t f_k(s)$ effectively selects eigenmodes with $\widehat{\lambda}_n \sim k^2$

Again k identified through $k = L/a$

Conclusion: UV-attractive fixed point of **AS generated by $k = L/a$**

Conclusions

- Sliding cutoff scale imposed on the **dimless Laplace Beltrami operator**: **cut on the number of modes**
- Wilsonian RG equations: **absence of quartic and quadratic contributions to ρ !**
 $\Rightarrow$ **no naturalness problem arises in pure gravity!**
- When **diffeomorphism invariant (FV) measure** used and **physical running scale properly identified**

$\Rightarrow$

Only Gaussian fixed point

no sign of the UV-attractive fixed point of the **AS** scenario

- AS fixed point generated if we impose the cutoff as a cut on $\sim n^2/a^2$ rather than $n \dots$ **spurious powers of a**

... Is Gravity non-perturbatively renormalizable? ...

Diffeomorphism invariance of the measure - 3

Under a general coordinate transformation

- change of time ordering parameter
- change of lattice

$$\Rightarrow \prod_x (\mathcal{E}_1) \rightarrow \prod_{\hat{x}} (\mathcal{E}_2) \text{ is } \mathbf{NOT} \text{ a trivial reshuffling}$$

To realize the bridge $\hat{\Gamma}[\hat{g}]$ (in $\hat{\Sigma}$) $\rightarrow \Gamma[g]$ (in Σ):

1 - In $\hat{\Sigma}$, move from $\mathcal{E}_2 \rightarrow \mathcal{E}_1$ (from $\hat{x}^0 \rightarrow x^0$)

$$e^{i\hat{\Gamma}[\hat{g}]} = \int e^{i\hat{S}[\hat{\phi}, \hat{g}]} \left[\prod_{\hat{x}} M(\hat{g}(\hat{x})) \right] \left[\prod_{\hat{x}} d\hat{\phi}(\hat{x}) \right] = \int e^{i\hat{S}[\hat{\phi}, \hat{g}]} \left[\textcolor{blue}{B} \prod_x M(\hat{g}(\hat{x}(x))) \right] \left[\textcolor{red}{A} \prod_x d\hat{\phi}(\hat{x}(x)) \right]$$

with $\hat{S}$ written as (change of integration variables from $\hat{x}$ to x through $\hat{x} = \hat{x}(x)$)

$$\begin{aligned} \hat{S}[\hat{\phi}, \hat{g}] &= \hat{S}_g[\hat{g}] + \hat{S}_m[\hat{\phi}, \hat{g}] \\ &= \hat{S}_g[\hat{g}] - \frac{1}{2} \int d^4x J \sqrt{-\hat{g}(\hat{x}(x))} \left(\hat{g}^{\mu\nu}(\hat{x}(x)) \frac{\partial x^\rho}{\partial \hat{x}^\mu} \frac{\partial x^\sigma}{\partial \hat{x}^\nu} \partial_\rho \hat{\phi}(\hat{x}(x)) \partial_\sigma \hat{\phi}(\hat{x}(x)) + m^2 \hat{\phi}^2(\hat{x}(x)) \right) \\ &\equiv \hat{S}_g[\hat{g}] + \int d^4x \tilde{\mathcal{L}}_m(\hat{\phi}(\hat{x}(x)), \partial_\mu \hat{\phi}(\hat{x}(x)), \hat{g}(\hat{x}(x))) \quad J \equiv \left| \det \frac{\partial \hat{x}}{\partial x} \right| \end{aligned}$$

Diffeomorphism invariance of the measure - 4

2 - Move from $\hat{\phi}(\hat{x}(x))$ and $\hat{g}_{\mu\nu}(\hat{x}(x))$ to $\phi(x)$ and $g_{\mu\nu}(x)$

$$\begin{aligned} e^{i\hat{\Gamma}[\hat{g}]} &= \int e^{i\hat{S}[\hat{\phi}, \hat{g}]} \left[\textcolor{blue}{B} \prod_x M(\hat{g}(\hat{x}(x))) \right] \left[\textcolor{red}{A} \prod_x d\hat{\phi}(\hat{x}(x)) \right] \\ &= \int e^{iS[\phi, g]} \left[\textcolor{blue}{B} \textcolor{red}{C} \prod_x M(g(x)) \right] \left[\textcolor{red}{A} \textcolor{red}{E} \prod_x d\phi(x) \right] \end{aligned}$$

* Factor C arises when we express $M(\hat{g}(\hat{x}(x)))$ in terms of $g_{\mu\nu}(x)$

* E Jacobian of $\hat{\phi}(\hat{x}(x)) \rightarrow \phi(x)$

To have $\hat{\Gamma}[\hat{g}] = \Gamma[g]$, it is necessary that $\textcolor{red}{A} \textcolor{red}{B} \textcolor{red}{C} \textcolor{red}{E} = 1$

Transformation factors

Factor E - ϕ scalar field, $\hat{\phi}(\hat{x}(x)) = \phi(x)$; E Jacobian of $\hat{\phi}(\hat{x}(x)) \rightarrow \phi(x) \implies E = 1$.

Factor A - In phase space

$$e^{i\Gamma_{\mathcal{E}_1}[g]} = e^{iS_g[g]} \int e^{i \int d^4x [\partial_0 \phi(x) \pi(x) - \mathcal{H}(\pi(x), \phi(x))]} \left[\prod_x f(g(x)) d\pi(x) d\phi(x) \right]$$

with $\mathcal{H}(\pi(x), \phi(x))$ the Hamiltonian density, $\pi(x)$ the momentum conjugate to $\phi(x)$

$$\pi(x) = \frac{\partial \mathcal{L}_m(\phi(x), \partial_\mu \phi(x), g(x))}{\partial(\partial_0 \phi(x))},$$

and $\mathcal{L}_m(\phi(x), \partial_\mu \phi(x), g(x))$ the Lagrangian density in Σ

$f(g(x)) = 1$ FV measure ; $f(g(x)) = (g^{00}(x))^{-1/2}$ Fujikawa measure.

$\prod_x \iff$ lattice $\mathcal{E}_1$ and time ordering $x^0 \implies \pi(x)$ correctly defined above

Transformation factors

Factor B - We begin by observing that $M(\hat{g}(\hat{x}))$ is nothing but the product $f(\hat{g}(\hat{x}))W(\hat{g}(\hat{x}))$. We then have

$$\begin{aligned}\prod_{\hat{x}} M(\hat{g}(\hat{x})) &= \prod_{\hat{x}} [f(\hat{g}(\hat{x}))W(\hat{g}(\hat{x}))] = \exp\left(\delta^{(4)}(0) \int d^4\hat{x} \log(f(\hat{g}(\hat{x}))W(\hat{g}(\hat{x})))\right) \\ &= \exp\left(\delta^{(4)}(0) \int d^4x \partial_\mu \varepsilon^\mu \log(f(g(x))W(g(x)))\right) \prod_x [f(\hat{g}(\hat{x}(x)))W(\hat{g}(\hat{x}(x)))] \\ &= \exp\left(\delta^{(4)}(0) \int d^4x \partial_\mu \varepsilon^\mu \log(M(g(x)))\right) \prod_x M(\hat{g}(\hat{x}(x))).\end{aligned}$$

Recalling that $\prod_{\hat{x}} M(\hat{g}(\hat{x})) = B \prod_x M(\hat{g}(\hat{x}(x)))$

$$B = \exp\left(\delta^{(4)}(0) \int d^4x \partial_\mu \varepsilon^\mu \log(M(g(x)))\right)$$

Fradkin-Vilkovisky versus Fujikawa measure

Note: B and C depend on the specific form of $M(g(x))$, A does not and $E = 1$. We have then to calculate B and C for $M_{\text{FV}}(g(x))$ and $M_{\text{Fuji}}(g(x))$, respectively.

Fradkin-Vilkovisky measure - Using $M(g(x)) = M_{\text{FV}}(g(x)) = \left(-g^{00}(x)\right)^{1/2} (-g(x))^{1/4}$ we get

$$B_{\text{FV}} = \exp \left(\frac{\delta^{(4)}(0)}{2} \int d^4x \partial_\mu \varepsilon^\mu \log[(-g(x))^{1/2} (-g^{00}(x))] \right)$$

$$C_{\text{FV}} = \exp \left(\delta^{(4)}(0) \int d^4x \frac{g^{0\mu}(x)}{g^{00}(x)} \partial_\mu \varepsilon^0 \right)$$

from which

$$B_{\text{FV}} C_{\text{FV}} = A^{-1}$$

which shows that for the FV measure

$$A B_{\text{FV}} C_{\text{FV}} E = 1$$

and therefore that in this case: $\hat{\Gamma}[\hat{g}] = \Gamma[g]$.

Fradkin-Vilkovisky versus Fujikawa measure

Fujikawa measure - Using $M(g(x)) = M_{\text{Fuji}}(g(x)) = \mu (-g(x))^{1/4}$ (μ arbitrary mass scale) we get

$$B_{\text{Fuji}} = \exp \left(\delta^{(4)}(0) \int d^4x \partial_\mu \varepsilon^\mu \log \left((-g(x))^{1/4} \mu \right) \right)$$

$$C_{\text{Fuji}} = 1$$

so that

$$A B_{\text{Fuji}} C_{\text{Fuji}} E = \exp \left(\delta^{(4)}(0) \int d^4x \left[\partial_\mu \varepsilon^\mu \log \left(\frac{\mu}{(-g^{00})^{1/2}} \right) - \frac{g^{0\mu}(x)}{g^{00}(x)} \partial_\mu \varepsilon^0 \right] \right)$$

which means that in this case: $\hat{\Gamma}[\hat{g}] \neq \Gamma[g]$.

Contrary to what is claimed in JHEP 05 (2025), 164, the **effective action** calculated using the **Fujikawa measure** *is not* diffeomorphism invariant

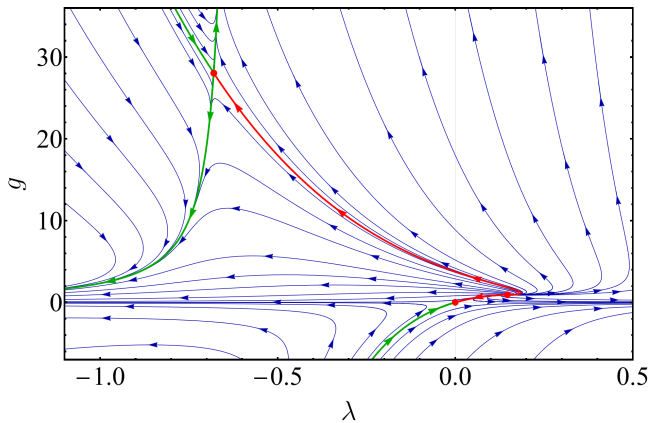
$$\begin{aligned}
64\pi G \left(S_2 + S_{\text{gf}} \right) &= \sum_{n=2}^{\infty} a_n^2 \left[\lambda_n^{(2)} - 2a^2 \Lambda + 8 \right] \\
&+ \sum_{n=2}^{\infty} b_n^2 \left[\xi^{-1} \left(\lambda_n^{(1)} - 3 \right) - 2a^2 \Lambda + 6 \right] \\
&+ \sum_{n=2}^{\infty} c_n^2 \left[\xi^{-1} \left(\frac{3}{4} \lambda_n^{(0)} - 6 \right) - \frac{\lambda_n^{(0)}}{2} - 2a^2 \Lambda + 6 \right] \\
&+ \sum_{n=0}^{\infty} e_n^2 \left[\frac{-3 + \xi^{-1}}{2} \lambda_n^{(0)} + 2a^2 \Lambda \right] \\
&+ \sum_{n=2}^{\infty} 2e_n c_n (\xi^{-1} - 1) \left[\lambda_n^{(0)} \left(\frac{3}{4} \lambda_n^{(0)} - 3 \right) \right]^{\frac{1}{2}} \\
32\pi G S_{\text{ghost}} &= \sum_{n=1}^{\infty} g_n^* g_n \left(\lambda_n^{(1)} - 3 \right) + \sum_{n=1}^{\infty} f_n^* f_n \left(\lambda_n^{(0)} - 6 \right) .
\end{aligned}$$

Therefore, the functional measure can be written as (defined as)

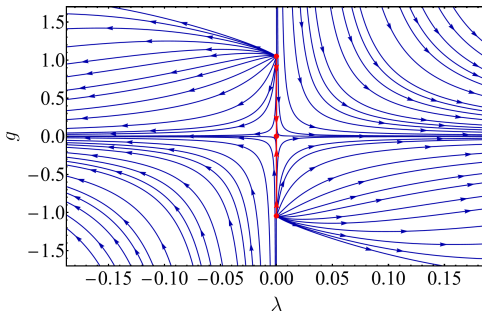
$$\widehat{\mathcal{D}h_{\mu\nu}} \mathcal{D}v_{\rho}^* \mathcal{D}v_{\sigma} \equiv \frac{1}{V_{SO(5)}} \prod_{n=2}^{\infty} da_n \prod_{n=2}^{\infty} db_n \prod_{n=2}^{\infty} dc_n \prod_{n=0}^{\infty} de_n \prod_{n=2}^{\infty} dg_n^* \prod_{n=2}^{\infty} dg_n \prod_{n=1}^{\infty} df_n^* \prod_{n=1}^{\infty} df_n ,$$

Notice that there is no integration over the zero modes g_1^* and g_1 of S_{ghost} . The corresponding ghost fields are proportional to the ten Killing vectors ξ_1^{μ} . These zero eigenmodes correspond to residual gauge invariances which are not eliminated by gauge fixing in the presence of an $SO(5)$ spherical symmetry. Overcounting has been compensated by inserting the explicit group-volume factor $V_{SO(5)}$

$$K = L/a \text{ full}$$



Flow - all quadrants



$\lambda > 0$

$$\frac{d\lambda}{dt} = -2\lambda + \frac{2g\lambda(3-2\lambda)}{2\pi + g(3-2\lambda)}$$

$$\frac{dg}{dt} = 2g + \frac{2g^2(3-8\lambda)}{2\pi + g(3-2\lambda)}$$

$\lambda < 0$

$$\frac{d\lambda}{dt} = -2\lambda - \frac{2g\lambda(3+2\lambda)}{2\pi - g(3+2\lambda)}$$

$$\frac{dg}{dt} = 2g - \frac{2g^2(3+8\lambda)}{2\pi - g(3+2\lambda)}$$

Fixed points

$(\lambda, g)_1 = (0, -\pi/3)$

$(\lambda, g)_2 = (0, 0)$

$(\lambda, g)_3 = (0, \pi/3)$

The case $\Lambda = 0 - 2$

$$\begin{aligned}\partial_t g_k = \beta_g &= [d - 2 + \eta(k)]g_k, \\ \partial_t \lambda_k = \beta_\lambda &= -(2 - \eta)\lambda_k + \frac{1}{2}g_k(4\pi)^{1-d/2} \left\{ 2d(d+1)\Phi_{d/2}^1(-2\lambda_k) \right. \\ &\quad \left. - d(d+1)\eta(k)\widetilde{\Phi}_{d/2}^1(-2\lambda_k) - 8d\Phi_{d/2}^1(0) \right\}.\end{aligned}$$

Souma

$$\begin{aligned}\frac{d\lambda_t}{dt} &= -2\lambda_t + \frac{g_t}{\pi} \left[1 + 6\lambda_t(1 + \lambda_t) - \lambda_t \frac{34 + 48\lambda_t}{6} \right] \equiv \beta_\lambda(\lambda_t, g_t) \\ \frac{dg_t}{dt} &= 2g_t - g_t^2 \frac{34 + 48\lambda_t}{6\pi} \equiv \beta_g(\lambda_t, g_t)\end{aligned}$$

Our equations with the identification $k = L/a$

λ is generated ... inconsistent to study g equation alone