

One-loop gravitational action and the cosmological constant problem

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Introduction

Quantum Gravity as a QFT: perturbatively non-renormalizable

- EFT valid up to the physical cutoff scale Λ
- Non-perturbatively renormalizable: UV fixed point

For the RG analysis of QG see Contino talk

- * Contribution to vacuum energy **from quantum fluctuations** $\sim M_P^4$
- * Value inferred **from observed accelerated expansion** of the universe
 $\rho_{\text{vac}} \sim 10^{-123} M_P^4$

CC problem: **most severe naturalness problem** in physics

One-loop effective action

One-loop Effective Action

- * Euclidean action - Einstein-Hilbert truncation

$$S_{\text{grav}} = \frac{1}{16\pi G} \int d^4x \sqrt{g} (-R + 2\Lambda_{\text{cc}})$$

- * Gauge-invariant one-loop effective action, $\Gamma_{\text{grav}}^{1/} = S_{\text{grav}} + \delta S_{\text{grav}}^{1/}$
geometrical approach, Vilkovisky-DeWitt

- * Strategy put forward by Fradkin and Tseytlin / Taylor and Veneziano

- * **Background field** method: $g_{\mu\nu} = \bar{g}_{\mu\nu} + h_{\mu\nu}$

- * When $\bar{g}_{\mu\nu}$ has **spherical symmetry**:

one-loop VDW **effective action** = the **standard** one calculated with gauge-fixing term

$$S_{\text{gf}} = \frac{1}{32\pi G\xi} \int d^4x \sqrt{\bar{g}} \left[\nabla_\mu \left(h_\nu^\mu - \frac{1}{2} \delta_\nu^\mu h_\sigma^\sigma \right) \right]^2$$

after taking the limit $\xi \rightarrow 0$ at the end of the calculation

One-loop EA: spherical background

We calculate the **1-loop correction** $\delta S_{\text{grav}}^{1/}$

Take **spherical background** $\bar{g}_{\mu\nu} = g_{\mu\nu}^{(a)}$ (a radius of the sphere)

(coordinates $\times$ angles ; $g_{\mu\nu}^{(a)}$ goes like a^2 ; $\int d^4x \sqrt{g^{(a)}} = \frac{8\pi^2}{3} a^4$, $R(g^{(a)}) = \frac{12}{a^2}$)

Classical Action

$$S_{\text{grav}}^{(a)} = \frac{\pi \Lambda_{\text{cc}}}{3G} a^4 - \frac{2\pi}{G} a^2$$

and dS solution

$$a_{\text{dS}} = \sqrt{\frac{3}{\Lambda_{\text{cc}}}}$$

Add to $S_{\text{grav}} + S_{\text{gf}}$ the corresponding ghost action (v_μ vector ghost fields)

$$S_{\text{ghost}} = \frac{1}{32\pi G} \int d^4x \sqrt{g^{(a)}} g^{(a)\mu\nu} v_\mu^* \left(-\nabla_\rho \nabla^\rho - \frac{3}{a^2} \right) v_\nu$$

Identify 1-loop corrections to $\frac{\Lambda_{\text{cc}}}{G}$ and $\frac{1}{G}$ with coefficients of a^4 and a^2 in $\delta S_{\text{grav}}^{1/}$

$$(\Gamma_{\text{grav}}^{1/} = S_{\text{grav}} + \delta S_{\text{grav}}^{1/})$$

One-loop EA: spherical background

One-loop correction $\delta S_{\text{grav}}^{1/}$ to $S_{\text{grav}}^{(a)}$ given by

$$e^{-\delta S_{\text{grav}}^{1/}} = \lim_{\xi \rightarrow 0} \int \mathcal{D}\mu e^{-\delta S^{(2)}}$$

where

$$\delta S^{(2)} \equiv S_2 + S_{\text{gf}} + S_{\text{ghost}}$$

S_2 quadratic term in the expansion of $S_{\text{grav}}[g_{\mu\nu}^{(a)} + h_{\mu\nu}]$

$$S_2 \equiv \frac{1}{32\pi G} \int d^4x \sqrt{g^{(a)}} \left[\frac{1}{2} \tilde{h}^{\mu\nu} \left(-\nabla_\rho \nabla^\rho - 2\Lambda_{\text{cc}} + \frac{8}{a^2} \right) h_{\mu\nu} + \frac{h^2}{a^2} - \nabla^\rho \tilde{h}_{\rho\mu} \nabla^\sigma \tilde{h}_\sigma^\mu \right]$$

$$h \equiv g_{\mu\nu}^{(a)} h^{\mu\nu} \quad ; \quad \tilde{h}_{\mu\nu} \equiv h_{\mu\nu} - \frac{1}{2} g_{\mu\nu}^{(a)} h \quad ; \quad \text{indices raised with } g^{(a)\mu\nu} \quad ; \quad \text{covariant derivatives in terms of } g_{\mu\nu}^{(a)}$$

A delicate point ... the measure $\mathcal{D}\mu$

Recent discussion: C.Branchina, V. Branchina, F. Contino, R. Gandolfo, A. P. / Bonanno, Ferrero, Falls

Path integral measure : Fradkin-Vilkovisky

$$\mathcal{D}\mu \equiv \prod_x \left[g^{(a)00}(x) \left(g^{(a)}(x) \right)^{-1} \left(\prod_{\alpha \leq \beta} dh_{\alpha\beta}(x) \right) \left(\prod_{\rho} dv_{\rho}^*(x) \right) \left(\prod_{\sigma} dv_{\sigma}(x) \right) \right]$$

Liouville in phase space $\Rightarrow$ integrate over conj. momenta $\Rightarrow g^{(a)00}(x) \left(g^{(a)}(x) \right)^{-1}$

g^{00} : dangerous? Diffeomorphism invariance?

Subtle point: $\prod_x \leftrightarrow$ discretization (given coordinate system)

Leutwyler / Fradkin, Vilkovisky

- Transition amplitude $\langle \text{in}, t_1 | \text{out}, t_2 \rangle$ requires a time ordering parameter
- Lattice $\mathcal{E}$: hypersurfaces $x^0 = \text{cte} \cap$ curves $x^i = \text{cte}$

Diffeomorphism invariance of the measure

Under a general coordinate transformation ($\hat{x} \rightarrow x$)

- change of time ordering parameter ($\hat{x}^0 \rightarrow x^0$)
- change of lattice ($\mathcal{E}_2 \rightarrow \mathcal{E}_1$)

$\Rightarrow \prod_{\hat{x}} (\mathcal{E}_2) \rightarrow \prod_x (\mathcal{E}_1)$ is **NOT** a trivial reshuffling

Careful calculation shows that

(scalar example)

$$e^{i\hat{F}[\hat{g}]} = \int e^{i\hat{S}[\hat{\phi}, \hat{g}]} \left[\prod_{\hat{x}} M(\hat{g}(\hat{x})) \right] \left[\prod_{\hat{x}} d\hat{\phi}(\hat{x}) \right] = \int e^{i\hat{S}[\hat{\phi}, \hat{g}]} \left[\overset{\text{B}}{\prod}_x M(\hat{g}(\hat{x}(x))) \right] \left[\overset{\text{A}}{\prod}_x d\hat{\phi}(\hat{x}(x)) \right]$$

NON-trivial factors **A** and **B** beyond those due to **fields transformation**

Accounting for this

- The measure turns out to be invariant
- Another typically used measure: Fujikawa¹. **No g^{00} , not invariant**
- g^{00} is **crucial** to ensure diffeomorphism invariance

¹Not from Liouville measure in phase space.

Back to the calculation of $\delta S_{\text{grav}}^{1/}$

Now : $g_{\mu\nu}^{(a)} = a^2 \tilde{g}_{\mu\nu}$ $\tilde{g}_{\mu\nu}$ dimensionless and a -independent metric

convenient (though not necessary) way to dig out the a -dependence from the background metric

Redefinition : $\hat{h}_{\mu\nu} = (32\pi G)^{-1/2} a^{-1} h_{\mu\nu}$, $v_\mu \rightarrow (32\pi G)^{\frac{1}{2}} v_\mu$

convenient but not necessary

$$\Rightarrow [\mathcal{D}\mu] \sim \prod_x \left[\left(\prod_{\alpha \leq \beta} d\hat{h}_{\alpha\beta}(x) \right) \left(\prod_\rho dv_\rho^*(x) \right) \left(\prod_\sigma dv_\sigma(x) \right) \right]$$

The radius “ a ” does not appear in the measure

In terms of $\hat{h}_{\mu\nu}$ and v_μ

$$S_2 + S_{\text{gf}} = \int d^4x \sqrt{g^{(1)}} \left[\frac{1}{2} \bar{h}^{\mu\nu} \left(-\nabla_\rho \nabla^\rho - 2a^2 \Lambda + 8 \right) \hat{h}_{\mu\nu} + \hat{h}^2 - \left(1 - \frac{1}{\xi} \right) \nabla^\rho \bar{h}_{\rho\mu} \nabla^\sigma \bar{h}_\sigma^\mu \right]$$

$\hat{h} \equiv g_{\mu\nu}^{(1)} \hat{h}^{\mu\nu}$, $\bar{h}_{\mu\nu} \equiv \hat{h}_{\mu\nu} - \frac{1}{2} g_{\mu\nu}^{(1)} \hat{h}$; indexes raised with $g^{(1)\mu\nu}$; covariant derivatives in terms of $g_{\mu\nu}^{(1)}$

$$S_{\text{ghost}} = \int d^4x \sqrt{g^{(1)}} g^{(1)\mu\nu} v_\mu^* \left(-\nabla_\rho \nabla^\rho - 3 \right) v_\nu$$

$\Rightarrow \delta S^{(2)} = S_2 + S_{\text{gf}} + S_{\text{ghost}}$ contains only dimensionless operators (no arbitrary scale μ)

Decomposing onto the eigenbasis

Consider the basis constructed with the eigenfunctions of

Dimensionless Laplace-Beltrami operator $-\tilde{\square}^{(s)} \equiv -a^2 \square_a^{(s)}$

Dimensionless eigenvalues $\lambda_n^{(s)}$ and corresponding degeneracies $D_n^{(s)}$

$$\lambda_n^{(s)} = n^2 + 3n - s; \quad D_n^{(s)} = \frac{2s+1}{3} \left(n + \frac{3}{2}\right)^3 - \frac{(2s+1)^3}{12} \left(n + \frac{3}{2}\right) \quad n = s, s+1, \dots$$

Decomposing $\hat{h}_{\mu\nu}$, v_ρ^* and v_σ onto this basis we get

backup slides

$$\delta S_{\text{grav}}^{1/} = -\frac{1}{2} \log \frac{\det_1[-\tilde{\square}^{(1)} - 3] \det_2[-\tilde{\square}^{(0)} - 6]}{\det_0[-\tilde{\square}^{(2)} - 2a^2\Lambda + 8] \det_2[-\tilde{\square}^{(0)} - 2a^2\Lambda]} + \dots$$

* index i in $\det_i$ indicates that product of eigenvalues starts from $\lambda_{s+i}^{(s)}$

Calculation of the fluctuation determinants

Two different strategies

- 1 - **product of eigenvalues** of Laplace-Beltrami ops.
- 2 - **proper-time**

1 - Calculation in terms of the eigenvalues $\lambda_n^{(s)}$

$$\delta S_{\text{grav}}^{1/} = \frac{1}{2} \sum_{n=2}^{N-2} \left[D_n^{(2)} \log \left(\lambda_n^{(2)} - 2a^2 \Lambda + 8 \right) + D_n^{(0)} \log \left(\lambda_n^{(0)} - 2a^2 \Lambda \right) - D_n^{(1)} \log \left(\lambda_n^{(1)} - 3 \right) - D_n^{(0)} \log \left(\lambda_n^{(0)} - 6 \right) \right] + \dots$$

N : **numerical UV cutoff** on the number of eigenvalues

De Sitter solution for the classical action $a_{\text{ds}} = \sqrt{\frac{3}{\Lambda_{\text{cc}}}}$

a_{ds} **size of the universe** $\Rightarrow$ connection between N and **physical cutoff scale Λ** given by

$$\Lambda \sim M_P = \frac{N}{a_{\text{ds}}} = N \sqrt{\frac{\Lambda_{\text{cc}}}{3}}$$

Expanding for $N \gg 1$

$$\delta S_{\text{grav}}^{1/} = - \left(\Lambda_{\text{cc}}^2 \log N^2 \right) a^4 + \Lambda_{\text{cc}} \left(-N^2 + 8 \log N^2 \right) a^2 + \frac{N^4}{24} \left(-1 + 2 \log N^2 \right) + \frac{N^2}{36} \left(203 - 75 \log N^2 \right) - \frac{779}{90} \log N^2$$

$$\delta S_{\text{grav}}^{1/} = - \left(\Lambda_{\text{cc}}^2 \log \frac{3\Lambda^2}{\Lambda_{\text{cc}}} \right) a^4 + \left(-3\Lambda^2 + 8\Lambda_{\text{cc}} \log \frac{3\Lambda^2}{\Lambda_{\text{cc}}} \right) a^2 + \dots$$

2 - Calculation with proper-time

$(-\tilde{\square}^{(s)} - \alpha)$ dimensionless $\implies \det_i$ regularized with a **dimensionless proper-time** τ
(lower cut: $N \gg 1$) with kernel **kernel** $K_i^{(s)}(\tau)$

$$\det_i(-\tilde{\square}^{(s)} - \alpha) = e^{-\int_{1/N^2}^{+\infty} \frac{d\tau}{\tau} K_i^{(s)}(\tau)} \quad ; \quad K_i^{(s)}(\tau) = \sum_{n=s+i}^{+\infty} D_n^{(s)} e^{-\tau \left(\lambda_n^{(s)} - \alpha \right)}$$

After integration over τ , sum over n with **EML** sum formula. Expanding for $N \gg 1$

$$\delta S_{\text{grav}}^{1l} = - \left(\Lambda_{\text{cc}}^2 \log N^2 \right) a^4 + \Lambda_{\text{cc}} \left(-N^2 + 8 \log N^2 \right) a^2 - \frac{N^4}{12} + \frac{17}{3} N^2 - \frac{1859}{90} \log N^2 + \dots$$

$$\text{With } \Lambda \equiv \frac{N}{a_{\text{ds}}} = \sqrt{\frac{\Lambda_{\text{cc}}}{3}} N \implies$$

$$\delta S_{\text{grav}}^{1l} = - \left(\Lambda_{\text{cc}}^2 \log \frac{3\Lambda^2}{\Lambda_{\text{cc}}} \right) a^4 + \left(-3\Lambda^2 + 8\Lambda_{\text{cc}} \log \frac{3\Lambda^2}{\Lambda_{\text{cc}}} \right) a^2 + \dots$$

The two methods give the same result

One-loop corrections to $\frac{\Lambda_{\text{cc}}}{G}$ and $\frac{1}{G}$

Coefficients of a^4 and a^2 identify the one-loop corrections to $\frac{\Lambda_{\text{cc}}}{G}$ and $\frac{1}{G}$

$$\frac{\Lambda_{\text{cc}}^{1/}}{G^{1/}} = \frac{\Lambda_{\text{cc}}}{G} \left(1 - \frac{3G\Lambda_{\text{cc}}}{\pi} \log \frac{3\Lambda^2}{\Lambda_{\text{cc}}} \right) \quad \frac{1}{G^{1/}} = \frac{1}{G} \left[1 + \frac{G}{2\pi} \left(3\Lambda^2 - 8\Lambda_{\text{cc}} \log \frac{3\Lambda^2}{\Lambda_{\text{cc}}} \right) \right]$$

* Taking $G \sim M_P^{-2}$: quantum corrections **do not spoil the naturalness** of this relation $G \sim G^{1/}$

* **Usual result:** $\rho \sim M_P^4 \implies$ bare value of $\rho \sim M_P^4$ with coefficient that must be **enormously fine-tuned** to cancel the one-loop M_P^4 correction

* **Our result:** **only** mild logarithmic corrections to $\rho = \frac{\Lambda_{\text{cc}}}{8\pi G}$: **unexpected!**

In pure gravity **no naturalness problem** arises:
the bare cosmological constant Λ_{cc} does **not** need to be $\sim M_P^2$.

We may **naturally** have $\Lambda_{\text{cc}} \ll M_P^2$, and so

$$\Lambda_{\text{cc}}^{1/} \sim \Lambda_{\text{cc}}$$

Peeking behind the curtain: what is going on

...Usually we have **power-law “divergences”** in ρ ...

Why don't we see them? What is going on?

- Consider $\delta S_{\text{grav}}^{1/}$

$$\delta S_{\text{grav}}^{1/} = - \left(\Lambda_{\text{cc}}^2 \log N^2 \right) a^4 + \Lambda_{\text{cc}} \left(-N^2 + 8 \log N^2 \right) a^2 - \frac{N^4}{12} + \frac{17}{3} N^2 - \frac{1859}{90} \log N^2$$

- Connect N and Λ through $\Lambda = \frac{N}{a}$ (rather than $\Lambda = \frac{N}{a_{\text{ds}}}$):

* Corresponds to identify Λ with the **maximal eigenvalue** $\tilde{\lambda}_{\text{max}}^{(s)}$ of $-\square_a^{(s)}$

$$\delta S_{\text{grav}}^{1/} = - \left[\frac{\Lambda^4}{12} + \Lambda_{\text{cc}} \Lambda^2 + \Lambda_{\text{cc}}^2 \log \left(\Lambda^2 a^2 \right) \right] a^4 + \left[\frac{17}{3} \Lambda^2 + 8 \Lambda_{\text{cc}} \log \left(\Lambda^2 a^2 \right) \right] a^2 - \frac{1859}{90} \log \left(\Lambda^2 a^2 \right)$$

known result found with **proper time regularization** using heat-kernel expansion

Dependence on $g_{\mu\nu}^{(a)}$ **artificially** altered in $\delta S_{\text{grav}}^{1/}$

Usual calculations implement $\Lambda = N/a$

- * Using Fujikawa measure (**not diffeomorphism invariant**)

$$\delta S_{\text{grav}}^{1/} = -\frac{1}{2} \log \frac{\det_1[-\square_a^{(1)} - 3/a^2] \det_2[-\square_a^{(0)} - 6/a^2]}{\det_0[-\square_a^{(2)} - 2\Lambda + 8/a^2] \det_2[-\square_a^{(0)} - 2\Lambda]} + \dots$$

- * Trace usually regulated as

$$\text{Tr} \log \left(\frac{-\square_a^{(s)} + \alpha}{\mu^2} \right) \rightarrow - \int_{1/\Lambda^2}^{+\infty} \frac{d\tau}{\tau} \text{Tr} e^{-\tau (-\square_a^{(s)} + \alpha)} = - \int_{1/\Lambda^2}^{+\infty} \frac{d\tau}{\tau} K(\tau)$$

- * τ **dimensionful** proper-time (mass) $^{-2}$ and

$$K(\tau) = \sum_{n=0}^{+\infty} D_n e^{-\tau (\tilde{\lambda}_n^{(s)} + \alpha)}$$

with $\tilde{\lambda}_n^{(s)}$ the **a-dependent** eigenvalues of $-\square_a^{(s)}$ and D_n the corresponding degeneracies

$$\tilde{\lambda}_n^{(s)} = \frac{n^2 + 3n - s}{a^2} \quad ; \quad D_n^{(s)} = \frac{2s+1}{3} \left(n + \frac{3}{2} \right)^3 - \frac{(2s+1)^3}{12} \left(n + \frac{3}{2} \right)$$

- * Cut $1/\Lambda^2$ in the proper-time integral **effectively restricts the sum** to the N eigenvalues of $-\square_a^{(s)}$ such that $\tilde{\lambda}_n^{(s)} \lesssim \Lambda^2 \implies$ highest eigenvalue $\tilde{\lambda}_N^{(s)} \sim \frac{N^2}{a^2} \sim \Lambda^2$

Scalar field on spherical background

Free real scalar field ϕ of mass m_ϕ on gravitational background $g_{\mu\nu}^{(a)}$

see Gandolfo talk for the interacting case

$$S = \frac{\pi\Lambda_{\text{cc}}}{3G} a^4 - \frac{2\pi}{G} a^2 + \int d^4x \sqrt{g^{(a)}} \left[\frac{1}{2} g^{(a)\mu\nu} \partial_\mu \phi \partial_\nu \phi + \frac{1}{2} m_\phi^2 \phi^2 \right]$$

Integrating $\phi(x) \implies$ Effective gravitational action $S_{\text{grav}}^{\text{eff}}$

$$S_{\text{grav}}^{\text{eff}} = \frac{\pi\Lambda_{\text{cc}}}{3G} a^4 - \frac{2\pi}{G} a^2 + \delta S_{\text{grav}}$$

with δS_{grav} given by

$$e^{-\delta S_{\text{grav}}} = \int \prod_x \left[\left(g^{(a)00}(x) \right)^{\frac{1}{2}} \left(g^{(a)}(x) \right)^{\frac{1}{4}} d\phi(x) \right] e^{-\int d^4x \sqrt{g^{(a)}} \left[-\frac{1}{2} \phi \square \phi + \frac{1}{2} m_\phi^2 \phi^2 \right]}$$

$\left(g^{(a)00}(x) \right)^{\frac{1}{2}} \left(g^{(a)}(x) \right)^{\frac{1}{4}}$ from integration over conjugate momenta

Fradkin, Vilkovisky

Using $g_{\mu\nu}^{(a)} = a^2 \tilde{g}_{\mu\nu}$ and introducing $\hat{\phi} = a\phi$ (convenient but not necessary)

$$e^{-\delta S_{\text{grav}}} = \mathcal{N} \int \prod_x [d\hat{\phi}(x)] e^{-\int d^4x \sqrt{\tilde{g}} \left[-\frac{1}{2} \hat{\phi} \left(\tilde{\square}^{(0)} \right) \hat{\phi} + \frac{1}{2} a^2 m_\phi^2 \hat{\phi}^2 \right]}$$

Fluctuation determinant

Performing Gaussian integrations

$$S_{\text{grav}}^{\text{eff}} = \frac{\pi \Lambda_{\text{cc}}}{3G} a^4 - \frac{2\pi}{G} a^2 + \frac{1}{2} \text{Tr} \log \left(-\tilde{\square}^{(0)} + a^2 m_\phi^2 \right)$$

To evaluate the “Tr log”

$$\text{Proper time } \int_{1/N^2}^{\infty} \frac{d\tau}{\tau} K^{(0)}(\tau), \quad K^{(0)}(\tau) = \sum_{n=0}^{+\infty} D_n^{(0)} e^{-\tau \left(\lambda_n^{(0)} + a^2 m_\phi^2 \right)}$$

⊕

Expansion for $N \gg 1$:

$$S_{\text{grav}}^{\text{eff}} = \frac{\pi}{3} \left(\frac{\Lambda_{\text{cc}}}{G} - \frac{m_\phi^4}{8\pi} \log N^2 \right) a^4 - 2\pi \left[\frac{1}{G} - \frac{m_\phi^2}{24\pi} (N^2 + 2 \log N^2) \right] a^2 + \dots$$

* Similarity with the result obtained in the pure gravity case evident

* Once again : if N properly related to $\Lambda \sim M_P$ through $\Lambda \sim \frac{N}{a_{\text{ds}}} = N \sqrt{\frac{\Lambda_{\text{cc}}}{3}}$

$\Rightarrow \rho = \frac{\Lambda_{\text{cc}}}{8\pi G}$ receives only mild logarithmically divergent correction

$$\delta \left(\frac{\Lambda_{\text{cc}}}{G} \right) = -\frac{m_\phi^4}{8\pi} \log \frac{3\Lambda^2}{\Lambda_{\text{cc}}}$$

Usual result

Again: if we were to perform the mapping $N(\Lambda)$ as $N = a\Lambda$

$\Rightarrow$ spurious quartically and quadratically divergent terms $\Leftarrow$

For instance:

$$-\frac{N^4}{48} \longrightarrow -\frac{\Lambda^4}{48} a^4$$

Quartically divergent contribution to $\frac{\Lambda_{cc}}{G}$

$$\frac{N^2}{12} m_\phi^2 a^2 \longrightarrow \frac{\Lambda^2}{12} m_\phi^2 a^4$$

Quadratically divergent contribution to $\frac{\Lambda_{cc}}{G}$

Conclusions

We have seen that when

- quantum fluctuations are calculated in curved spacetime
- diffeomorphism invariant measure (FV) used
- physical cutoff properly identified

Power-law divergences are **absent in the vacuum energy**

Same is true even in the presence of matter fields

THANKS FOR YOUR ATTENTION

ADDITIONAL SLIDES

Diffeomorphism invariance of the measure

Gravitational action $S_g[g]$ in the presence of a scalar field with action $S_m[\phi, g]$ (sign. $(-, +, +, +)$)

$$S[\phi, g] = S_g[g] + S_m[\phi, g] = S_g[g] + \int d^4x \mathcal{L}_m(\phi(x), \partial_\mu \phi(x), g(x))$$

where $\mathcal{L}_m(\phi(x), \partial_\mu \phi(x), g(x))$ is the matter Lagrangian density

$$\mathcal{L}_m(\phi(x), \partial_\mu \phi(x), g(x)) = -\frac{1}{2} \sqrt{-g(x)} \left(g^{\mu\nu}(x) \partial_\mu \phi(x) \partial_\nu \phi(x) + m^2 \phi^2(x) \right)$$

Effective gravitational action $\Gamma[g]$ given by

$$e^{i\Gamma[g]} = \int e^{iS[\phi(x), g(x)]} \prod_x \left[M(g(x)) d\phi(x) \right]$$

$M(g(x))$ non-trivial measure term

FV measure obtained from the Liouville measure in phase space $\prod_x \left[d\pi(x) d\phi(x) \right]$

Diffeomorphism invariance of the measure - 3

Under a general coordinate transformation

- change of time ordering parameter
- change of lattice

$$\implies \prod_x (\mathcal{E}_1) \rightarrow \prod_{\hat{x}} (\mathcal{E}_2) \text{ is } \mathbf{NOT} \text{ a trivial reshuffling}$$

To realize the bridge $\hat{\Gamma}[\hat{g}]$ (in $\hat{\Sigma}$) $\rightarrow$ $\Gamma[g]$ (in Σ):

1 - In $\hat{\Sigma}$, move from $\mathcal{E}_2 \rightarrow \mathcal{E}_1$ (from $\hat{x}^0 \rightarrow x^0$)

$$e^{i\hat{\Gamma}[\hat{g}]} = \int e^{i\hat{S}[\hat{\phi}, \hat{g}]} \left[\prod_{\hat{x}} M(\hat{g}(\hat{x})) \right] \left[\prod_{\hat{x}} d\hat{\phi}(\hat{x}) \right] = \int e^{i\hat{S}[\hat{\phi}, \hat{g}]} \left[\textcolor{blue}{B} \prod_x M(\hat{g}(\hat{x}(x))) \right] \left[\textcolor{red}{A} \prod_x d\hat{\phi}(\hat{x}(x)) \right]$$

with $\hat{S}$ written as (change of integration variables from $\hat{x}$ to x through $\hat{x} = \hat{x}(x)$)

$$\begin{aligned} \hat{S}[\hat{\phi}, \hat{g}] &= \hat{S}_g[\hat{g}] + \hat{S}_m[\hat{\phi}, \hat{g}] \\ &= \hat{S}_g[\hat{g}] - \frac{1}{2} \int d^4x J \sqrt{-\hat{g}(\hat{x}(x))} \left(\hat{g}^{\mu\nu}(\hat{x}(x)) \frac{\partial x^\rho}{\partial \hat{x}^\mu} \frac{\partial x^\sigma}{\partial \hat{x}^\nu} \partial_\rho \hat{\phi}(\hat{x}(x)) \partial_\sigma \hat{\phi}(\hat{x}(x)) + m^2 \hat{\phi}^2(\hat{x}(x)) \right) \\ &\equiv \hat{S}_g[\hat{g}] + \int d^4x \tilde{\mathcal{L}}_m(\hat{\phi}(\hat{x}(x)), \partial_\mu \hat{\phi}(\hat{x}(x)), \hat{g}(\hat{x}(x))) \quad J \equiv \left| \det \frac{\partial \hat{x}}{\partial x} \right| \end{aligned}$$

Diffeomorphism invariance of the measure - 4

2 - Move from $\hat{\phi}(\hat{x}(x))$ and $\hat{g}_{\mu\nu}(\hat{x}(x))$ to $\phi(x)$ and $g_{\mu\nu}(x)$

$$\begin{aligned} e^{i\hat{\Gamma}[\hat{g}]} &= \int e^{i\hat{S}[\hat{\phi}, \hat{g}]} \left[\textcolor{violet}{B} \prod_x M(\hat{g}(\hat{x}(x))) \right] \left[\textcolor{violet}{A} \prod_x d\hat{\phi}(\hat{x}(x)) \right] \\ &= \int e^{iS[\phi, g]} \left[\textcolor{violet}{B} \textcolor{violet}{C} \prod_x M(g(x)) \right] \left[\textcolor{violet}{A} \textcolor{violet}{E} \prod_x d\phi(x) \right] \end{aligned}$$

* Factor C arises when we express $M(\hat{g}(\hat{x}(x)))$ in terms of $g_{\mu\nu}(x)$

* E Jacobian of $\hat{\phi}(\hat{x}(x)) \rightarrow \phi(x)$

To have $\hat{\Gamma}[\hat{g}] = \Gamma[g]$, it is necessary that $\textcolor{violet}{A} \textcolor{violet}{B} \textcolor{violet}{C} \textcolor{violet}{E} = 1$

Transformation factors

Factor E - ϕ scalar field, $\hat{\phi}(\hat{x}(x)) = \phi(x)$; E Jacobian of $\hat{\phi}(\hat{x}(x)) \rightarrow \phi(x) \implies E = 1$.

Factor A - In phase space

$$e^{i\Gamma_{\mathcal{E}_1}[g]} = e^{iS_g[g]} \int e^{i \int d^4x [\partial_0 \phi(x) \pi(x) - \mathcal{H}(\pi(x), \phi(x))]} \left[\prod_x f(g(x)) d\pi(x) d\phi(x) \right]$$

with $\mathcal{H}(\pi(x), \phi(x))$ the Hamiltonian density, $\pi(x)$ the momentum conjugate to $\phi(x)$

$$\pi(x) = \frac{\partial \mathcal{L}_m(\phi(x), \partial_\mu \phi(x), g(x))}{\partial(\partial_0 \phi(x))},$$

and $\mathcal{L}_m(\phi(x), \partial_\mu \phi(x), g(x))$ the Lagrangian density in Σ

$f(g(x)) = 1$ FV measure ; $f(g(x)) = (g^{00}(x))^{-1/2}$ Fujikawa measure.

$\prod_x \iff$ lattice $\mathcal{E}_1$ and time ordering $x^0 \implies \pi(x)$ correctly defined above

Transformation factors

$$A = \frac{\prod_x \gamma(\hat{g}(\hat{x}(x)))}{\prod_{\hat{x}} W(\hat{g}(\hat{x}))}$$

To write A in a convenient form, consider an infinitesimal coordinate transformation (only first order in ε)

$$\hat{x}^\mu = x^\mu + \varepsilon^\mu(x),$$

and switch from $\prod_x$ to $\prod_{\hat{x}}$ in the numerator. Observing that $(\delta^{(4)}(0))$ from $\sum_x \rightarrow \int d^4x$

$$\prod_x J^{1/2} = \exp\left(\frac{\delta^{(4)}(0)}{2} \int d^4x \log(1 + \partial_\mu \varepsilon^\mu)\right) = \exp\left(\frac{\delta^{(4)}(0)}{2} \int d^4x \partial_\mu \varepsilon^\mu\right) = 1$$

where we have used $d^4\hat{x} = d^4x J \equiv d^4x \left| \det \frac{\partial \hat{x}}{\partial x} \right| = d^4x (1 + \partial_\mu \varepsilon^\mu(x))$, and observing also that

$$-\hat{g}^{\mu\nu}(\hat{x}(x)) \hat{\partial}_\mu x^0 \hat{\partial}_\nu x^0 = -\hat{g}^{00} + 2 \hat{g}^{0\mu} \partial_\mu \varepsilon^0 = -\hat{g}^{00} \left(1 - 2 \frac{\hat{g}^{0\mu}}{\hat{g}^{00}} \partial_\mu \varepsilon^0\right)$$

we have

$$\begin{aligned} \prod_x \gamma(\hat{g}(\hat{x}(x))) &= \left[\prod_{\hat{x}} W(\hat{g}(\hat{x})) \right] \exp\left(-\frac{\delta^{(4)}(0)}{2} \int d^4\hat{x} \left[\partial_\mu \varepsilon^\mu \log\left((- \hat{g}(\hat{x}))^{1/2} (- \hat{g}^{00}(\hat{x}))\right) + 2 \frac{\hat{g}^{0\mu}(\hat{x})}{\hat{g}^{00}(\hat{x})} \partial_\mu \varepsilon^0 \right]\right) \\ \Rightarrow A &= \exp\left(-\frac{\delta^{(4)}(0)}{2} \int d^4x \left[\partial_\mu \varepsilon^\mu \log\left((-g(x))^{1/2} (-g^{00}(x))\right) + 2 \frac{g^{0\mu}(x)}{g^{00}(x)} \partial_\mu \varepsilon^0 \right]\right) \end{aligned}$$

Transformation factors

Factor B - We begin by observing that $M(\hat{g}(\hat{x}))$ is nothing but the product $f(\hat{g}(\hat{x}))W(\hat{g}(\hat{x}))$. We then have

$$\begin{aligned} \prod_{\hat{x}} M(\hat{g}(\hat{x})) &= \prod_{\hat{x}} [f(\hat{g}(\hat{x}))W(\hat{g}(\hat{x}))] = \exp\left(\delta^{(4)}(0) \int d^4\hat{x} \log(f(\hat{g}(\hat{x}))W(\hat{g}(\hat{x})))\right) \\ &= \exp\left(\delta^{(4)}(0) \int d^4x \partial_\mu \varepsilon^\mu \log(f(g(x))W(g(x)))\right) \prod_x [f(\hat{g}(\hat{x}(x)))W(\hat{g}(\hat{x}(x)))] \\ &= \exp\left(\delta^{(4)}(0) \int d^4x \partial_\mu \varepsilon^\mu \log(M(g(x)))\right) \prod_x M(\hat{g}(\hat{x}(x))). \end{aligned}$$

Recalling that $\prod_{\hat{x}} M(\hat{g}(\hat{x})) = B \prod_x M(\hat{g}(\hat{x}(x)))$

$$B = \exp\left(\delta^{(4)}(0) \int d^4x \partial_\mu \varepsilon^\mu \log(M(g(x)))\right)$$

Transformation factors

Factor C - Defined by $\prod_x M(\hat{g}(\hat{x}(x))) = C \prod_x M(g(x))$. At first order in ε

$$\begin{aligned} M(\hat{g}^{\mu\nu}(\hat{x}(x))) &= M\left(g^{\mu\nu}(x) + g^{\mu\sigma}(x)\partial_\sigma\varepsilon^\nu(x) + g^{\nu\sigma}(x)\partial_\sigma\varepsilon^\mu(x)\right) = \\ &= M(g) + \frac{\partial M(g)}{\partial g^{\alpha\beta}(x)} (g^{\alpha\sigma}(x)\partial_\sigma\varepsilon^\beta(x) + g^{\beta\sigma}(x)\partial_\sigma\varepsilon^\alpha(x)) \end{aligned}$$

where we used

$$\hat{g}^{\mu\nu}(\hat{x}) = \frac{\partial \hat{x}^\mu(x)}{\partial x^\rho} \frac{\partial \hat{x}^\nu(x)}{\partial x^\sigma} g^{\rho\sigma}(x) = g^{\mu\nu}(x) + g^{\mu\sigma}(x)\partial_\sigma\varepsilon^\nu(x) + g^{\nu\sigma}(x)\partial_\sigma\varepsilon^\mu(x)$$

We then get

$$\begin{aligned} \prod_x M(\hat{g}(\hat{x}(x))) &= \exp\left(\delta^{(4)}(0) \int d^4x \log \left[M(g) + \frac{\partial M(g)}{\partial g^{\alpha\beta}(x)} (g^{\alpha\sigma}(x)\partial_\sigma\varepsilon^\beta(x) + g^{\beta\sigma}(x)\partial_\sigma\varepsilon^\alpha(x)) \right]\right) \\ &= \exp\left(\delta^{(4)}(0) \int d^4x [M(g)]^{-1} \frac{\partial M(g)}{\partial g^{\alpha\beta}(x)} (g^{\alpha\sigma}(x)\partial_\sigma\varepsilon^\beta(x) + g^{\beta\sigma}(x)\partial_\sigma\varepsilon^\alpha(x))\right) \prod_x M(g(x)) \\ &\Rightarrow C = \exp\left(\delta^{(4)}(0) \int d^4x [M(g)]^{-1} \frac{\partial M(g)}{\partial g^{\alpha\beta}(x)} (g^{\alpha\sigma}(x)\partial_\sigma\varepsilon^\beta(x) + g^{\beta\sigma}(x)\partial_\sigma\varepsilon^\alpha(x))\right) \end{aligned}$$

Fradkin-Vilkovisky versus Fujikawa measure

Note: B and C depend on the specific form of $M(g(x))$, A does not and $E = 1$. We have then to calculate B and C for $M_{\text{FV}}(g(x))$ and $M_{\text{Fuj}}(g(x))$, respectively.

Fradkin-Vilkovisky measure - Using $M(g(x)) = M_{\text{FV}}(g(x)) = (-g^{00}(x))^{1/2} (-g(x))^{1/4}$ we get

$$B_{\text{FV}} = \exp \left(\frac{\delta^{(4)}(0)}{2} \int d^4x \partial_\mu \varepsilon^\mu \log[(-g(x))^{1/2} (-g^{00}(x))] \right)$$

$$C_{\text{FV}} = \exp \left(\delta^{(4)}(0) \int d^4x \frac{g^{0\mu}(x)}{g^{00}(x)} \partial_\mu \varepsilon^0 \right)$$

from which

$$B_{\text{FV}} C_{\text{FV}} = A^{-1}$$

which shows that for the FV measure

$$A B_{\text{FV}} C_{\text{FV}} E = 1$$

and therefore that in this case: $\hat{\Gamma}[\hat{g}] = \Gamma[g]$.

Comparison with Becker-Reuter, PRD 102 (2020) 12 - PRD 104 (2021) 12

and Ferrero-Percacci, e-Print: 2404.12357

Consider the **modified Einstein equation** at one-loop

Becker, Reuter PRD 104 (2021) 12

$$\frac{3G}{\pi} a \frac{d}{da} \Gamma_{\text{grav}}^{1/}(a) = 4\Lambda a^4 - 12a^2 + \frac{3G}{\pi} a \frac{d}{da} \delta S_{\text{grav}}^{1/}(a) = 0 \quad (g_{\mu\nu} = g_{\mu\nu}^{(a)})$$

* **BRST-invariant Fujikawa measure** (or measure not fully considered)

$$\left[\mathcal{D}u(h) \mathcal{D}v_{\rho}^* \mathcal{D}v_{\sigma} \right] \equiv \prod_x \left[\underbrace{\left(g^{(a)}(x) \right)^{-2}}_{\propto a^{-16}} \left(\prod_{\alpha \leq \beta} dh_{\alpha\beta}(x) \right) \left(\prod_{\rho} dv_{\rho}^*(x) \right) \left(\prod_{\sigma} dv_{\sigma}(x) \right) \right]$$

$$\Rightarrow \delta S_{\text{grav}}^{1/}(a) \sim N^4 \log(a\mu) \Rightarrow \text{solution } \bar{a}_N \sim N \sqrt{M_P^{-1} \Lambda_{\text{cc}}^{-1}}$$

* **Fradkin-Vilkovisky measure**

$$\left[\mathcal{D}u(h) \mathcal{D}v_{\rho}^* \mathcal{D}v_{\sigma} \right] \equiv \prod_x \left[\underbrace{\left(g^{(a)00}(x) \right) \left(g^{(a)}(x) \right)^{-1}}_{\propto a^{-10}} \left(\prod_{\alpha \leq \beta} dh_{\alpha\beta}(x) \right) \left(\prod_{\rho} dv_{\rho}^*(x) \right) \left(\prod_{\sigma} dv_{\sigma}(x) \right) \right]$$

$$\Rightarrow \delta S_{\text{grav}}^{1/}(a) \text{ does not contain } N^4 \log(a\mu) \Rightarrow \text{NO } \bar{a}_N \text{ solution}$$

Similar considerations apply to the results found in Becker-Reuter, PRD 102 (2020) 12, and Ferrero-Percacci, JHEP 09 (2024) 074

