Exploring CEvNS with NUCLEUS experiment & latest updates on the Cryogenic Outer Veto (COV)

Marco Giammei 08/10/2025







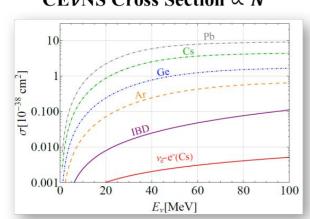
CEVNS

- The Coherent Elastic Neutrino-Nucleus Scattering (CEvNS)
 was predicted within the Standard Model in 1973
- It was firstly observed in 2017 by the COHERENT collaboration
- Although the cross section is several odm larger than other v
 process studied at low energies, the recoil energy (the observable)
 is very difficult to detect (~ few keV)

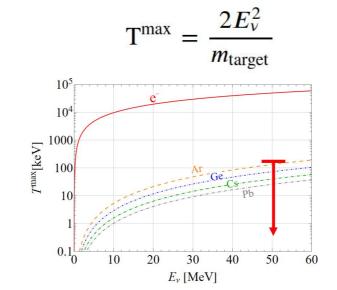
$$\sigma_{\text{CE}\nu \text{NS}} = \frac{G_F^2}{4\pi} F^2(q^2) Q_W^2 E_\nu^2$$

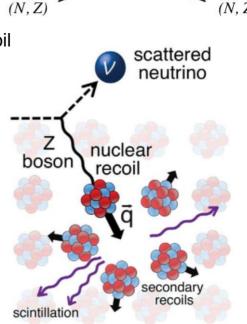
$$Q_W = N - Z(1 - 4\sin^2\theta_W) \sim N$$

CEVNS Cross Section $\propto N^2$



Observable: kinetic energy of nuclear recoil





 ν_{α}

The NUCLEUS Experiment

NUCLEUS Aim:

 Precision measurement of reactor CEvNS with cryogenic detectors at lowest energies

NUCLEUS Location:

- Measurements: Chooz nuclear power plant in the French Ardennes (operated by EDF)
- Detectors development and tests: Underground Laboratory (UGL) at TUM (Technical University of Munich)

NUCLEUS Collaboration:

















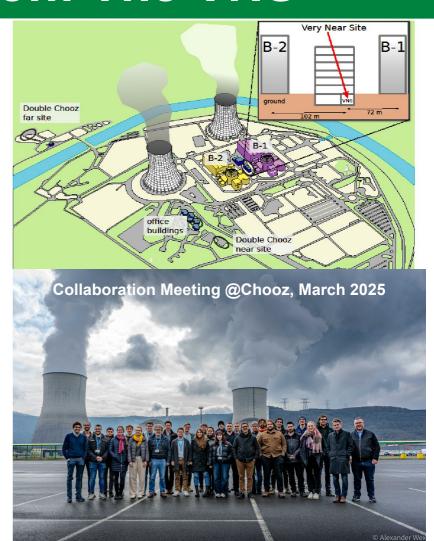


NUCLEUS Location: The VNS

- The experimental site (VNS) is located at 102 m and 72 m from the 2 reactors of the Chooz plant,
 at ≈ 3 m.w.e. depth
- Cosmic-ray induced neutrons will be challenging to mitigate with such a low overburden

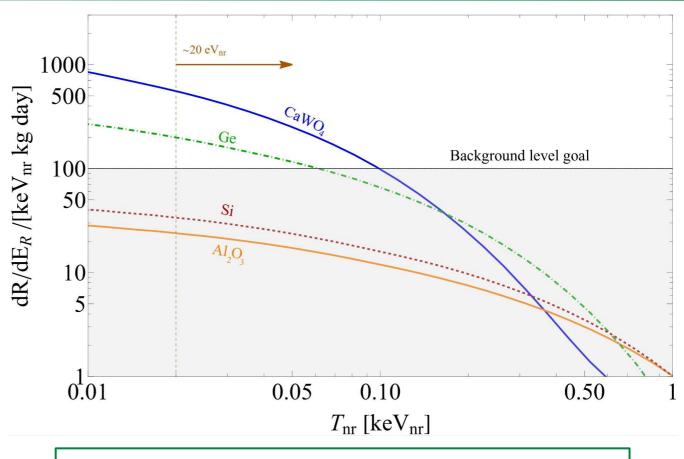
NUCLEUS Collaboration: Particle background characterization and prediction for the NUCLEUS reactor CEvNS experiment (arXiv:2509.03559)

- Reactor nominal thermal power:
 2 x 4.25 GWTh
- Expected v flux at the detectors: 1.7·10¹² v/s/cm²



CEVNS Signal at Nuclear Reactors

- ✓ Nuclear Reactors are intense source of anti-neutrinos, around 10²⁰ neutrinos per GW per second
- ✓ The majority of this anti-neutrinos have energy E_v < 10 MeV (fully coherent domain)</p>
- Induced nuclear recoils for CEvNS interaction are in sub-KeV range
- X Low threshold detectors and low background counting rate are required

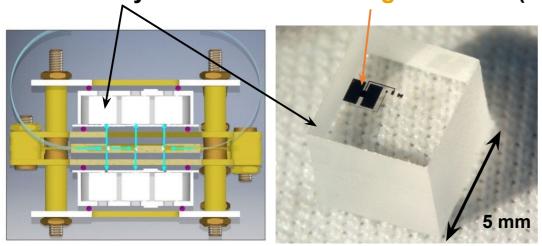


NUCLEUS expected signal:

~100 events per year in a gram-scale detector

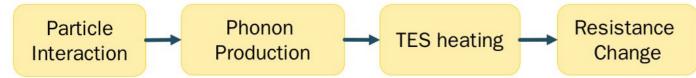
NUCLEUS Cryogenic Target Detector

Absorber crystals + Transition Edge Sensors (TES)

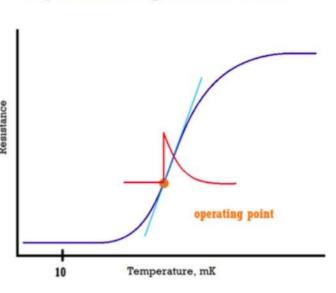


NUCLEUS-10g Cryogenic Detector: two 3x3 matrices of target detectors

- Baseline resolution well under 10 eV
- Threshold: 10-50 eV



Superconducting transition curve



Multitarget approach:

- Al₂O₃: used to probe background
- CaWO₄: similar background to Al₂O₃ but higher CEvNS rate

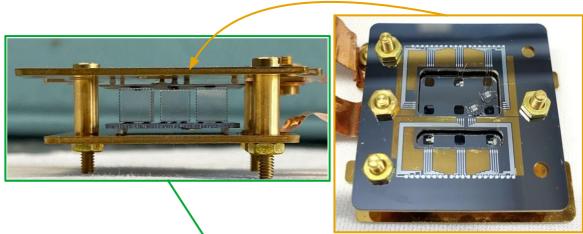
Inner and Outer Cryogenic Vetoes

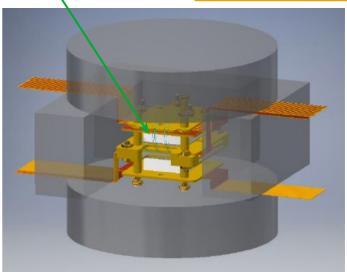
Inner Veto:

- Si wafer instrumented with TESs to hold and encapsulate the crystals (mechanical and thermal test concluded - in production)
- Veto against mechanical stress events, surface contamination
- Sub-keV cryogenic detector with TES readout

Cryogenic Outer Veto (COV):

- 6 HPGe crystals (2 cylindrical and 4 rectangular) of 2.5 cm thickness (under commissioning)
- Surround the inner detectors for active γ and neutron background rejection (almost 4π coverage)





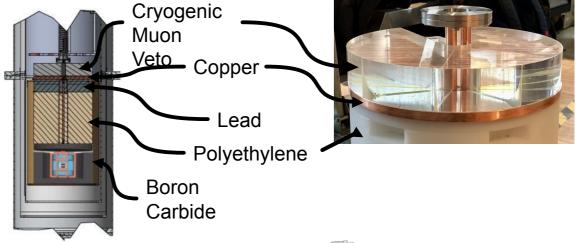
Active and Passive Shieldings

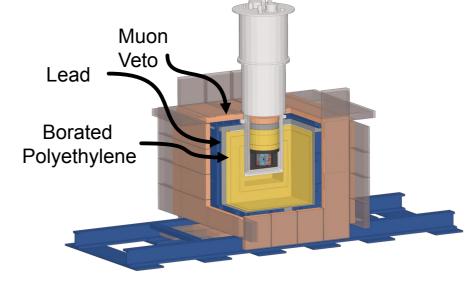
Inner Shielding:

- Lead, Copper support (acting also as a thermal contact), PE,
 B₄C for neutron reduction
- Thermalized at 800 mK

Outer Shielding:

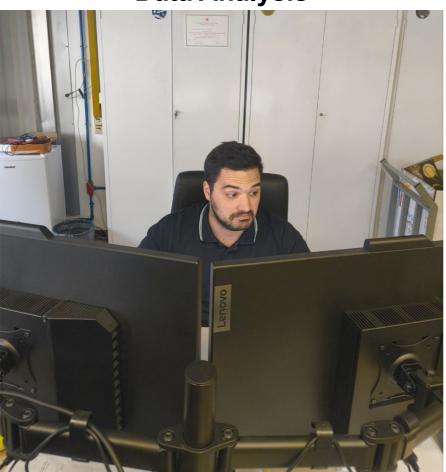
- Multilayer Passive Shield: Lead and Borated Polyethylene
- Muon Veto: plastic scintillators with SiPM & WLS-fiber readout
- Movable mechanical structure to allow easy opening/closing of the external shield
- Minimize µ induced neutrons





My Main Contributions inside the Experiment

Data Analysis

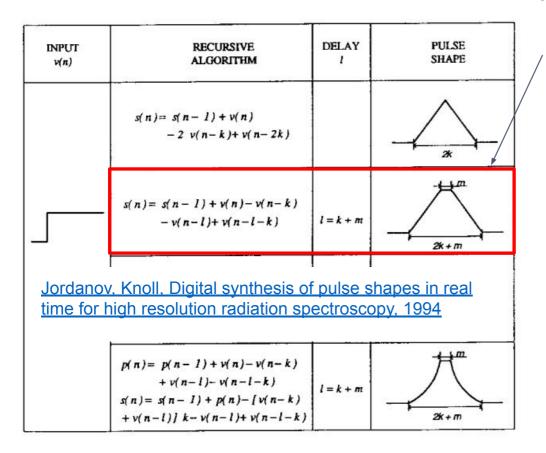


Hardware

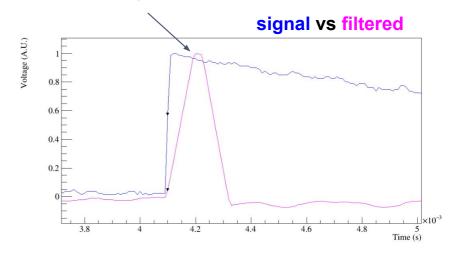


Trapezoidal filter implementation

- The optimal tool used for TES data (Matched Filter) is not suitable to analyze COV data
- I decided to implement a Trapezoidal filter algorithm



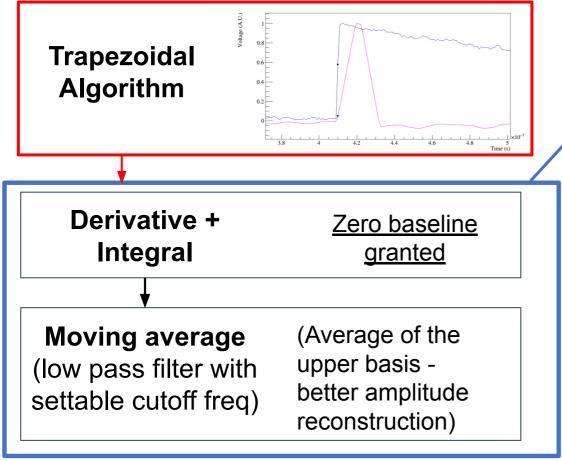
- The algorithm used is the one highlighted
- Amplitude is the maximum point on the top basis of the trapezium



N.B.: this is normalized to 1 for both pulses, the real gain of the filter is not 1

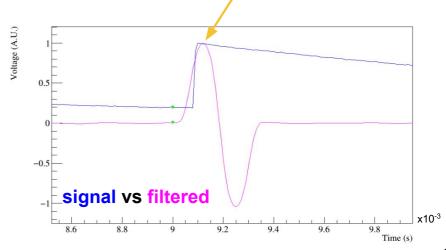
Trapezoidal filter upgrade

• I improved the standard trapezoidal filter to obtain better results from our data

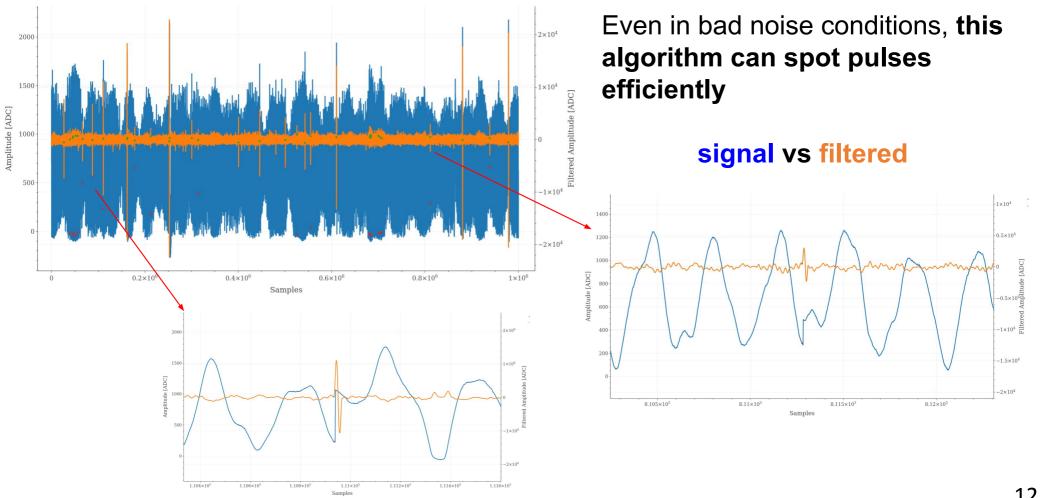


Parabolic interpolation

From 6% to 1.5% resolution on ¹³⁷Cs peak (662 keV) on average on COV crystals

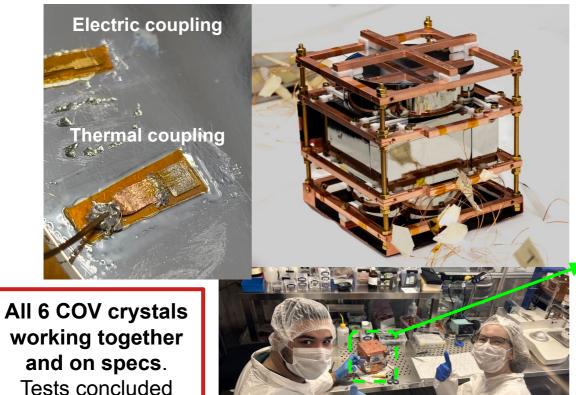


Trapezoidal Filter example - bad noise

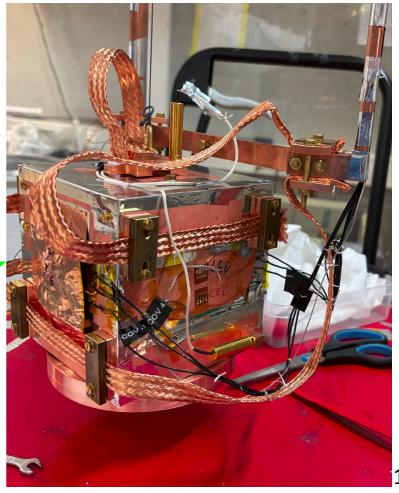


COV: from 1 to 6 HPGe Crystals

COV Cage V1 while being mounted from scratch

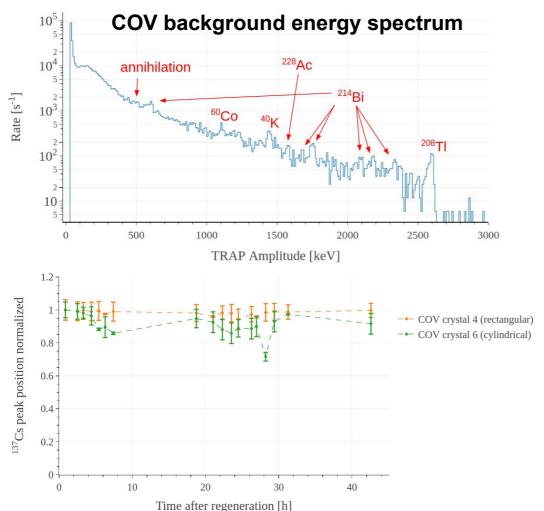


COV Cage V1 inside the cryostat



with no detector module inside

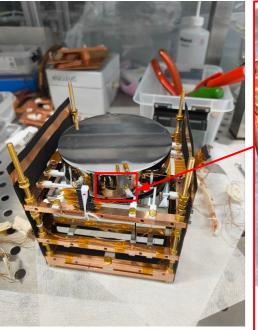
Results of Cage V1 tests

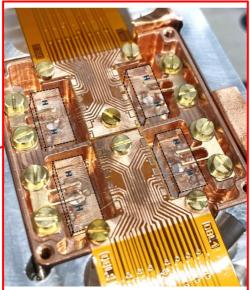


- All 6 COV crystals reached nominal performances i.e. less than 2 keV of baseline resolution
- From ~2h to >36h crystals
 operational time needed between regeneration
- Changed crystal regeneration procedure
 - From ~3h with LEDs to 20 min using bias voltage flips
- To mitigate the effect of mechanical vibrations, a new version of the Cage was designed

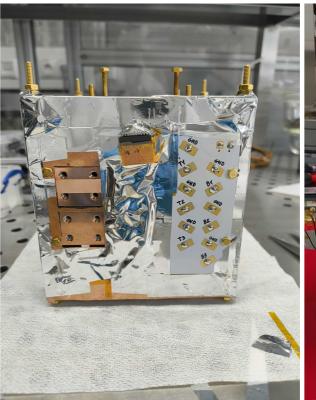
Ongoing: COV Cage V2

COV Cage V2: ready to be tested





Detector module assembled inside the new COV cage: 4 CaWO₄ cubes with double-TES

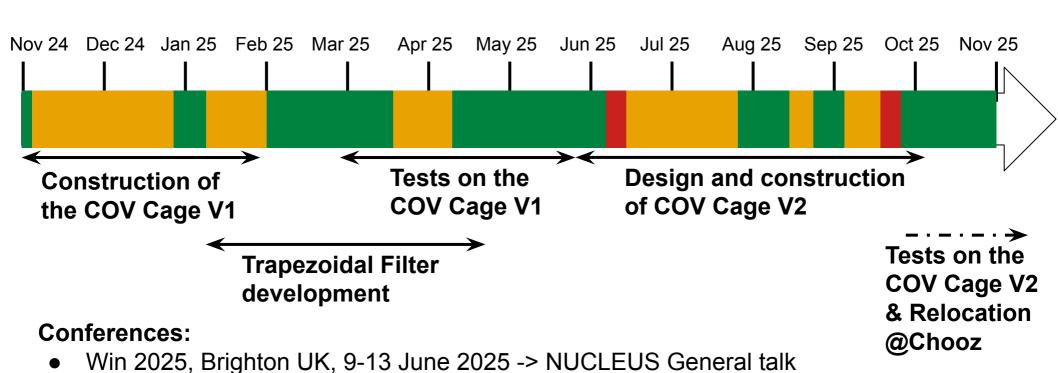




COV V2 cage fully assembled and ready to be tested (ongoing)

Recap of my 2nd year of PhD

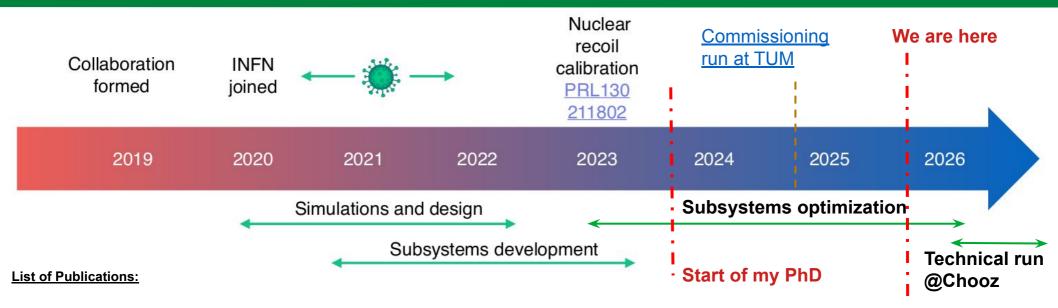
<u>Legend</u>: Tor Vergata (~6.5 months), <u>Munich</u> (~5 months), <u>Conferences</u> (2 weeks)



SIF 111 Congress, Palermo IT, 22-26 Sept 2025 -> NUCLEUS General talk

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NUCLEUS Timeline



NUCLEUS: recent results and perspectives, M. Giammei for NUCLEUS Collaboration, PoS ICHEP2024 (2025) 228, DOI:10.22323/1.476.0228

Decoupling pulse tube vibrations from a dry dilution refrigerator at milli-Kelvin temperatures, A. Wex, J. Rothe, L. Peters et al, JINST 20 (2025) 05, P05022, DOI:10.1088/1748-0221/20/05/P05022

Sub-keV Electron Recoil Calibration for Macroscopic Cryogenic Calorimeters using a Novel X-ray Fluorescence Source, NUCLEUS Collaboration, J.Low Temp.Phys. (2025), DOI:10.1007/s10909-025-03330-2

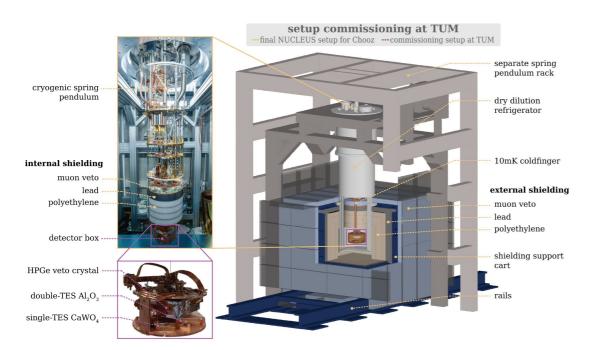
Commissioning of the NUCLEUS Experiment at the Technical University of Munich, NUCLEUS Collaboration, American Physical Society (PRD), DOI:10.1103/c95p-8kh2

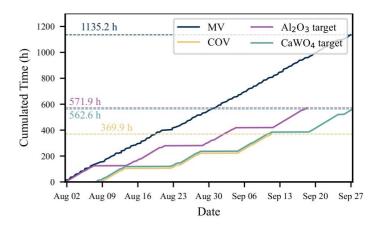
Particle background characterization and prediction for the NUCLEUS reactor CEvvNS experiment, NUCLEUS Collaboration, DOI:10.48550/arXiv.2509.03559

Thank you for your attention!

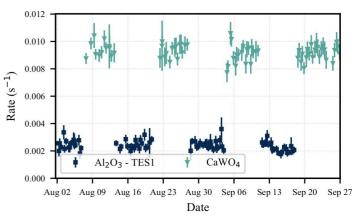
Backup: Commissioning Run at TUM

- Concluded in summer 2024 in the UGL (TUM);
 it's a milestone for the experiment
- Essential version of the experiment tested
 - Limited by the old DAQ of the experiment at the time
 - B₄C to be tested soon (crucial for neutron reduction)

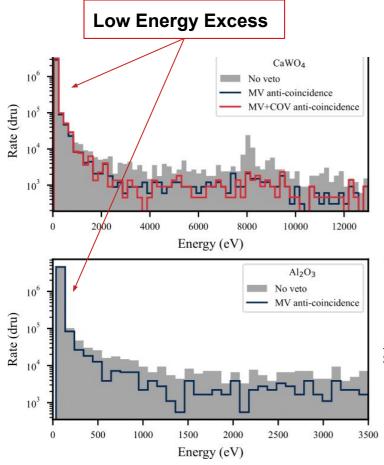




Cumulated time and TESs stability over time measured during the run



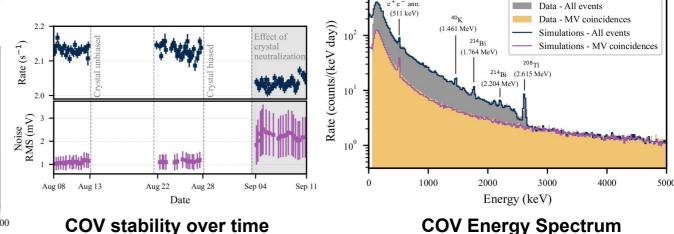
BU: Commissioning Run at TUM: results



TESs energy spectra with vetoes coincidences

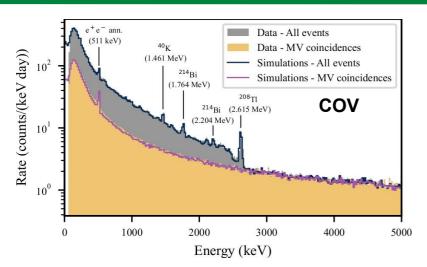
Main results:

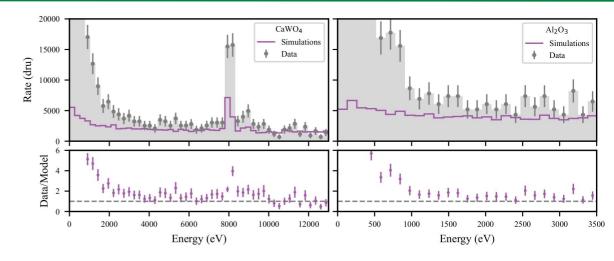
- ✓ All installed parts of the detector worked together
- Quite stable operation of all parts for the entire duration of the run (~2 months)
- ✓ Validation of shielding strategy
- Good agreement with simulations (above 1 keV)



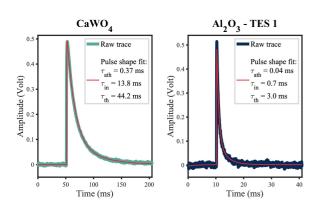
NUCLEUS collaboration: Commissioning of the NUCLEUS Experiment at the Technical University of Munich (arXiv:2508.02488)

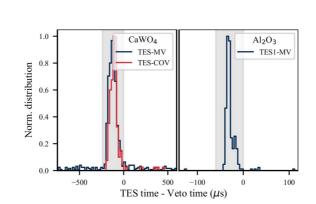
Backup: Commissioning Run at TUM

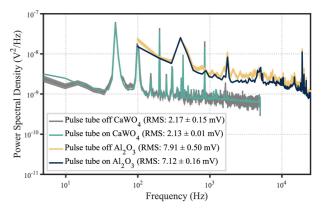




NB: UGL is at ≈ 15 m.w.e. depth



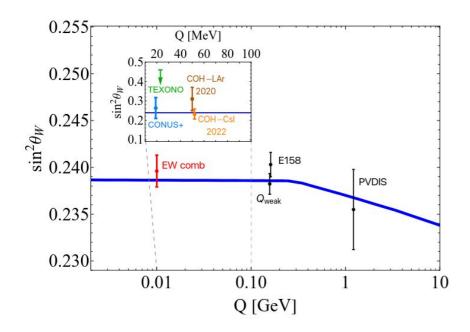




Backup: CEvNS Physics reach

$\sin^2 \theta_{\scriptscriptstyle W}$ at low momentum transferred

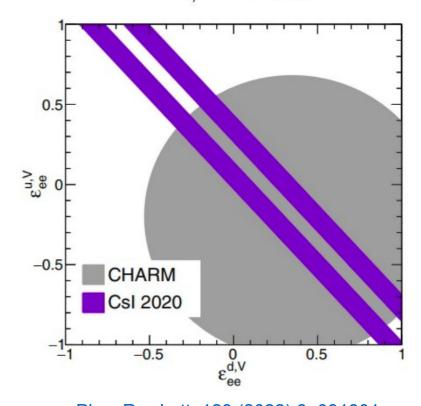
$$\sigma_{\text{CE}\nu\text{NS}} = \frac{G_F^2}{4\pi} F^2(q^2) Q_W^2 E_\nu^2 \quad Q_W = N - Z(1 - 4\sin^2\theta_W) \sim N$$



Phys.Rev.D 112 (2025) 1, 015007

Non standard interactions (NSI)

$$\mathcal{L}^{\rm NSI} = -\,\epsilon_{\alpha\beta}^{\,qV} 2\sqrt{2} G_F(\bar{\nu_\alpha}\gamma_\mu\nu_\beta)(\bar{q}\gamma^\mu q)$$

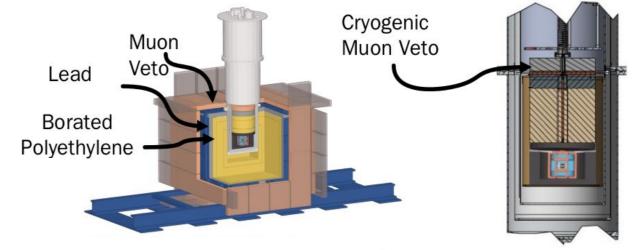


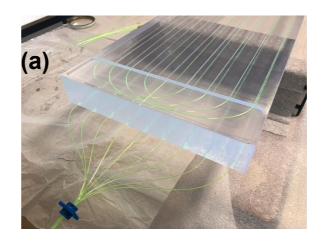
Phys.Rev.Lett. 129 (2022) 8, 081801

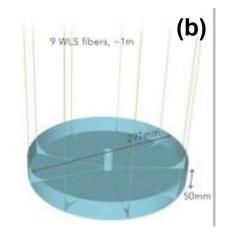
Backup: Muon Veto

Muon Veto:

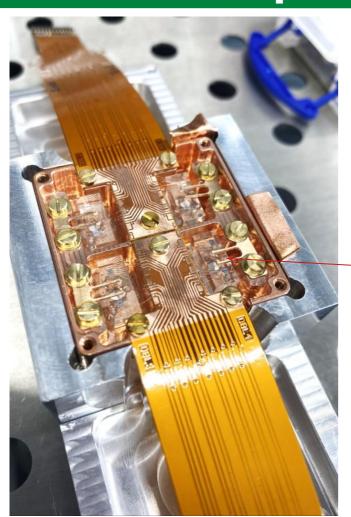
- Made of 5 cm thick
 plastic scintillators with
 SiPM & WLS-fiber readout
- **High efficiency** for muon detection (>99%)
- Expected μ counting rate of
 ~325 Hz (induced dead time on target detectors <10%)
- 4π coverage of the set-up
- Consists in 2 different parts:
 - Warm part: 28 rectangular plates (a)
 - Cold part: 1 cylindrical plate (b)



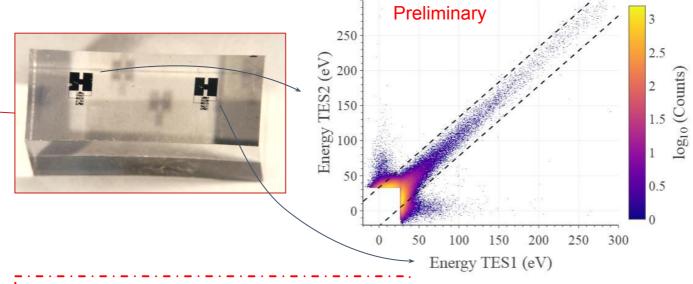




Backup: Minimal Detector Module

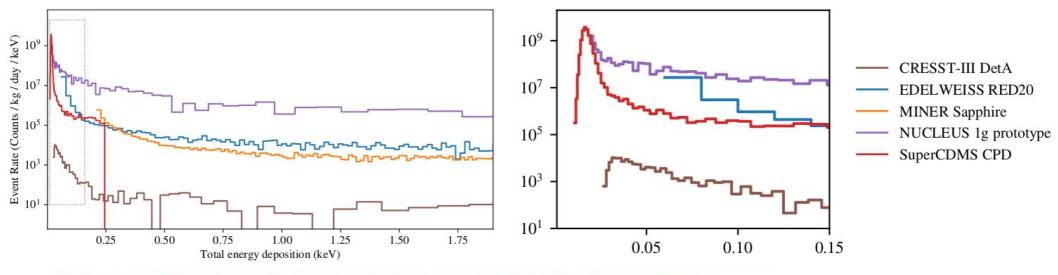


 4 x CaWO₄ detectors with double-TES, inside minimized copper encapsulation (compatible with technical run @Chooz)



Baseline resolution is 3-6 eV (for the first detectors and the first test)

Backup: Excess Background



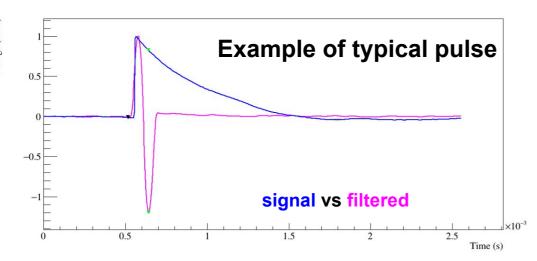
P. Adari, et al.: EXCESS workshop: Descriptions of rising low-energy spectra SciPost Phys. Proc. 9 (2022) 001

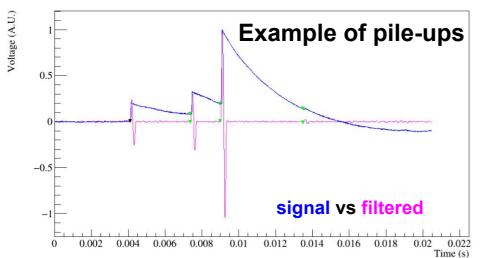
Not understood excess background rising at low energies:

- → Phonon bursts (crystal-support friction) ?
- → Phonon leakage from interactions in the supports ?
- → Lattice relaxations after cool down ?
- → Neutrons (cosmic ray induced, radioactivity) ? (excluded)

- This background limits NUCLEUS sensitivity
- Paper in preparation with the latest results

Backup: Examples Trapezoidal Filter



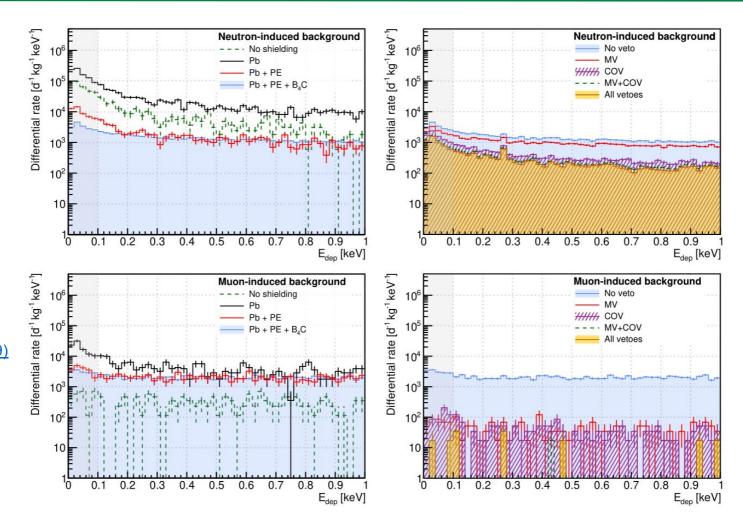


Improvements wrt standard Trapezoidal filter algorithm:

- Zero baseline granted
- From 6% to 1.5% resolution on ¹³⁷Cs peak (662 keV) on average on COV crystals

Backup: Neutron and muon background

NUCLEUS
Collaboration:
Particle
background
characterization
and prediction for
the NUCLEUS
reactor CEvNS
experiment
(arXiv:2509.03559)



Backup: Gamma induced background

NUCLEUS Collaboration: Particle background characterization and prediction for the NUCLEUS reactor CEvNS experiment (arXiv:2509.03559)

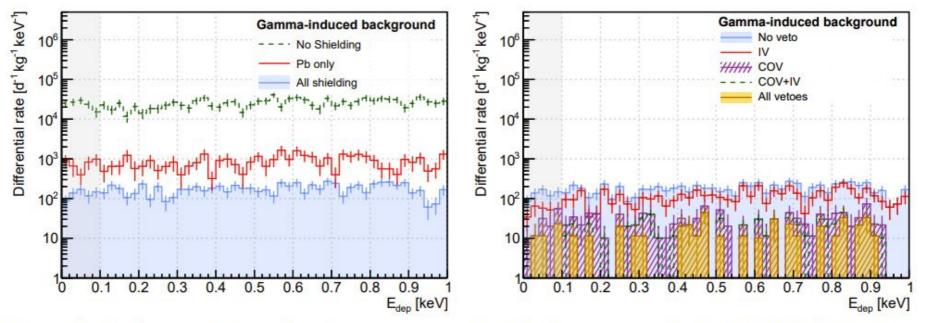


Figure 9: Reduction of the ambient gamma ray-induced backgrounds with the NUCLEUS shielding. Each histogram shows the rate of events with deposited energy between 0 and 1 keV in the CaWO₄ array of target detectors. Uncertainties are statistical only. The impact of the NUCLEUS setup passive materials is shown on the right panel. The left panel illustrates the complementary rejection brought by the use of the IV and COV. The $CE\nu NS$ RoI is indicated by the light gray filled area.

Backup: NUCLEUS Sensitivity

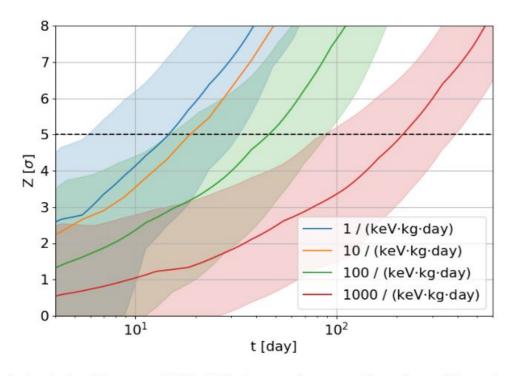


Figure 11: Statistical significance of $\text{CE}\nu\text{NS}$ observation as a function of live time for NUCLEUS-10g, for different background indices, using the expected neutrino flux at the VNS and an energy threshold of 10 eV. For each background index, the median line and 90% probability bands (computed from 600 simulated random spectra at each point) are shown. For the background index of 10 counts/(keV·kg·day), only the median line is shown for clarity.