

Dark Matter and BSM physics: recipe to search it at colliders

Deborah Pinna

(University of Wisconsin-Madison)

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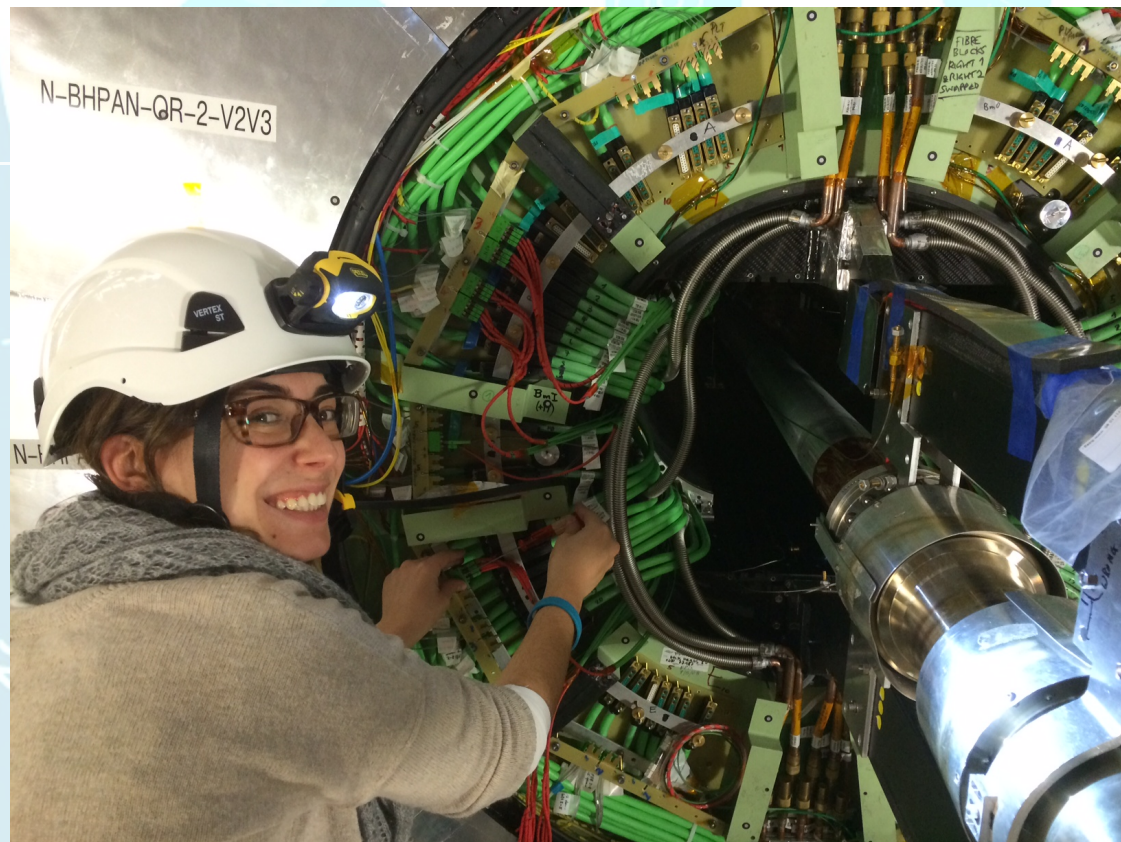
Dark matter hunter ...

- * Msc Thesis (NAGO)



- * Scientist at UW
- * Post-doc at Boston University
- * PhD at University of Zurich
- * Erasmus at ETH Zurich

- * CERN summer student 2010



Deborah Pinna

- * Scientist at University of Wisconsin-Madison
- * PhD at University of Zurich
- * Msc and Bsc: Università di Cagliari



Research

- * Dark matter hunter :), new physics with top quarks and di-Higgs signatures

The Standard Model ...

u UP	c CHARM	t TOP	g GLUON	H HIGGS
d DOWN	s STRANGE	b BOTTOM	γ PHOTON	
e ELECTRON	μ MUON	τ TAU	Z Z BOSON	
ν_e ELECTRON NEUTRINO	ν_μ MUON NEUTRINO	ν_τ TAU NEUTRINO	W W BOSON	

- ▶ Remarkable accuracy and predictive power
- ▶ LHC/CMS extended sensitivity beyond expectations
 - * H couplings: precision already $<10\%$ for most couplings with $\sim 5\%$ of expected HL-LHC dataset
 - * HH : end of Run2, we reached the precision that HL-LHC studies predicted for 1000 fb^{-1}

The Standard Model ... how do we modify it?

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+

Dark sector

Dark matter

New heavy resonances

Leptoquarks

Long lived particles

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- * direct searches for BSM particles coupling to SM sector
- * indirect searches from deviations between precision measurements and SM

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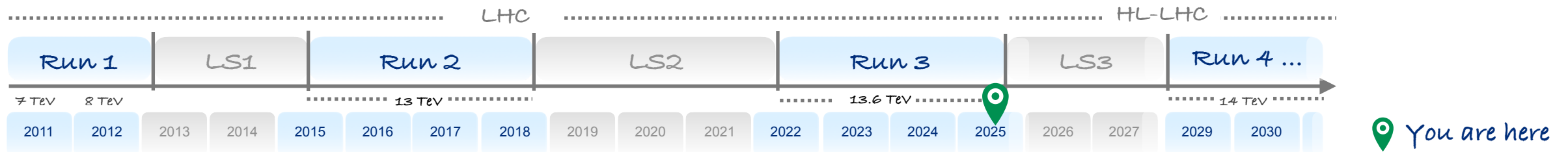
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(biased) focus on dark matter and CMS

The journey towards new physics



► LHC/CMS sensitivity beyond expectations

- * first phase of LHC program to be completed soon
- * already >300 fb⁻¹ (Run 2 + Run 3)

► ... but the future is already now

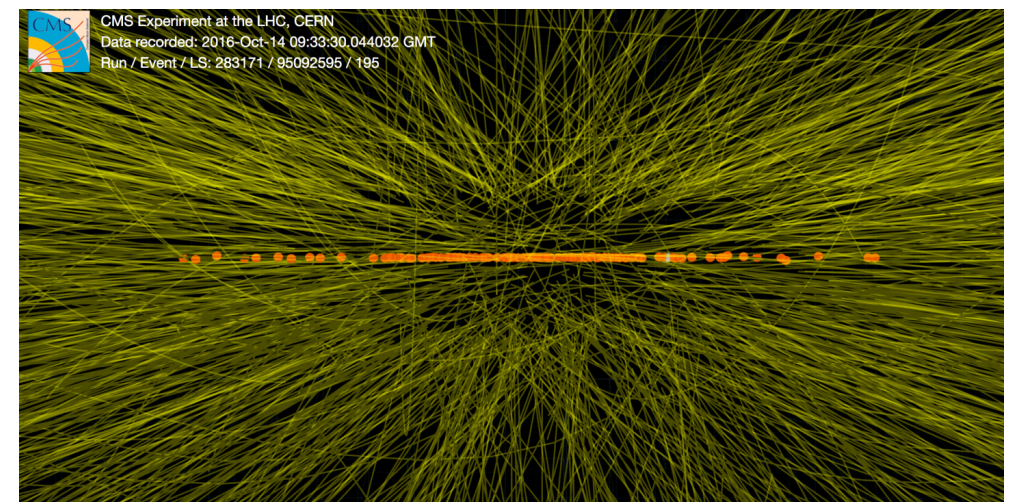
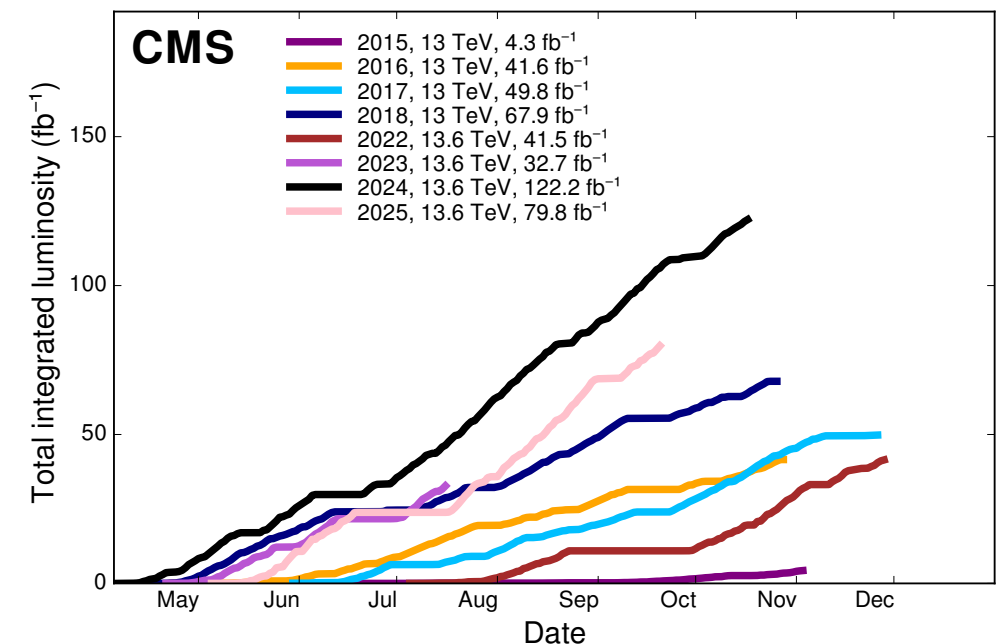
- * pushing the detector beyond its limits
- * recording up to 63 simultaneous collisions/event (2.5x CMS design, 45% of HL-LHC)
- * collecting data @7 kHz (70% of HL-LHC, 7x Run 2 normal operations)

► Pushing physics boundaries across multiple frontiers

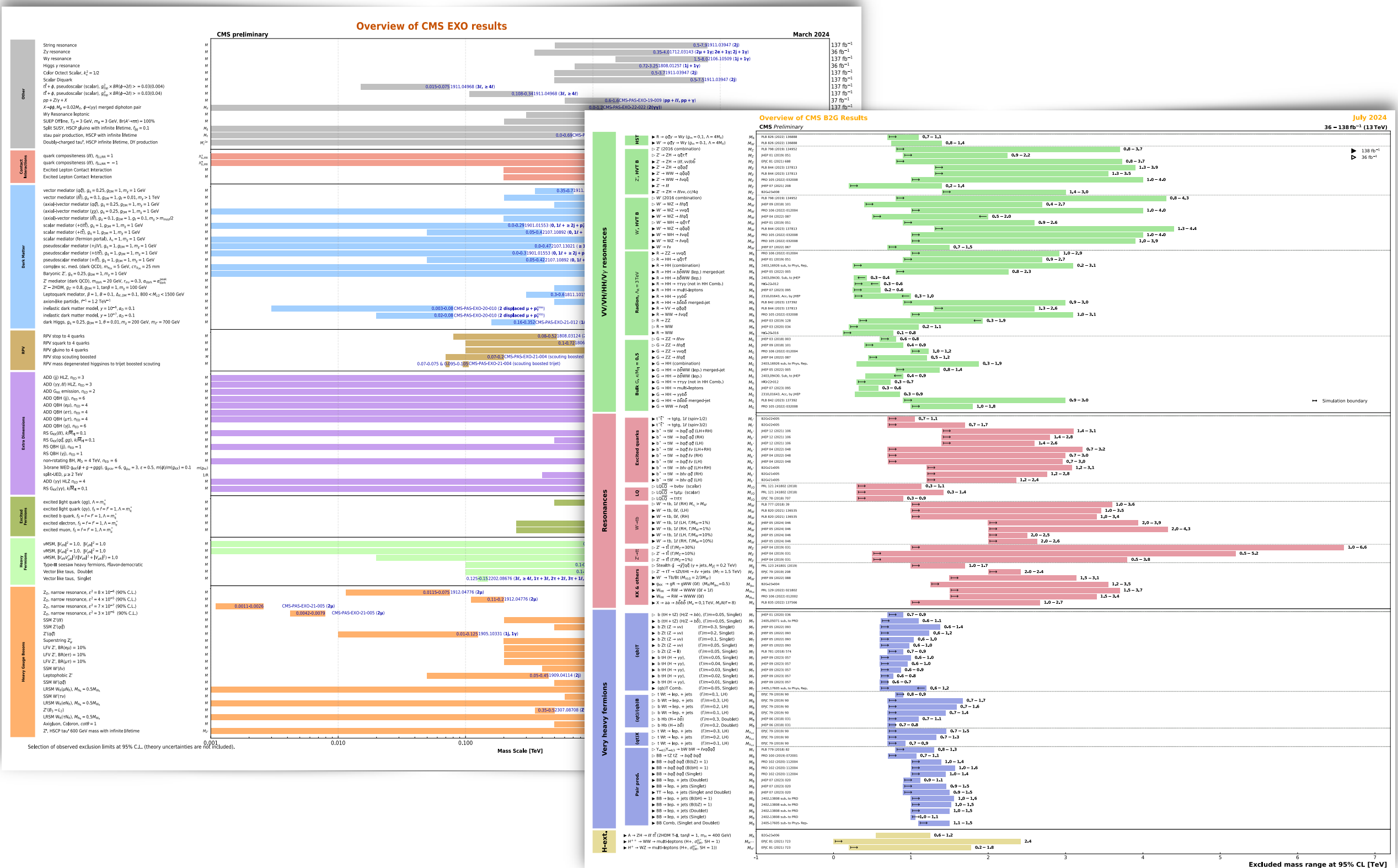
- * searches, precision-, flavour-physics, ...

► A key technology driver

- * multiple data-taking strategies, artificial Intelligence, ...



Rich new physics program at CMS

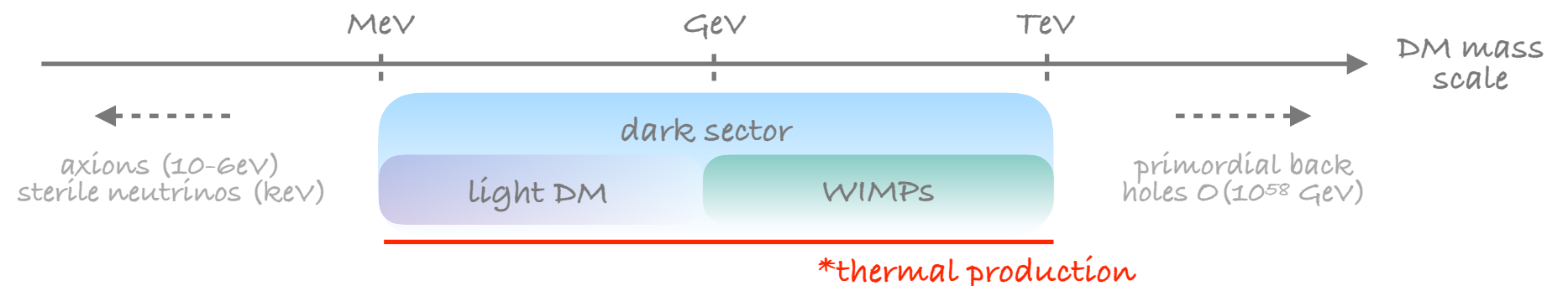


Hunt for dark matter - identikit

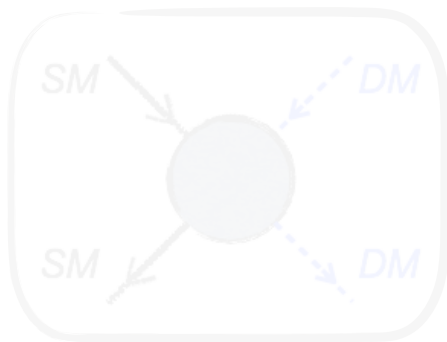
DM evidence



- Empirical evidence of DM from astrophysical observations at different scales
 - interacts gravitationally, long lived and neutral
 - no information about its nature (what DM could be?)
- Only measured quantitative property is DM mass abundance
 - very large set of possible masses can account for observed relic density



DM production



- DM must have some non-gravitational interactions with SM particles

MeV-GeV:

- avoid DM overproduction with new mediator below weak scale
- suff. small SM coupling for consistency with collider searches

GeV-TeV:

- WIMP models, DM has small couplings to SM particles
- most minimal scenario, one DM and one mediator

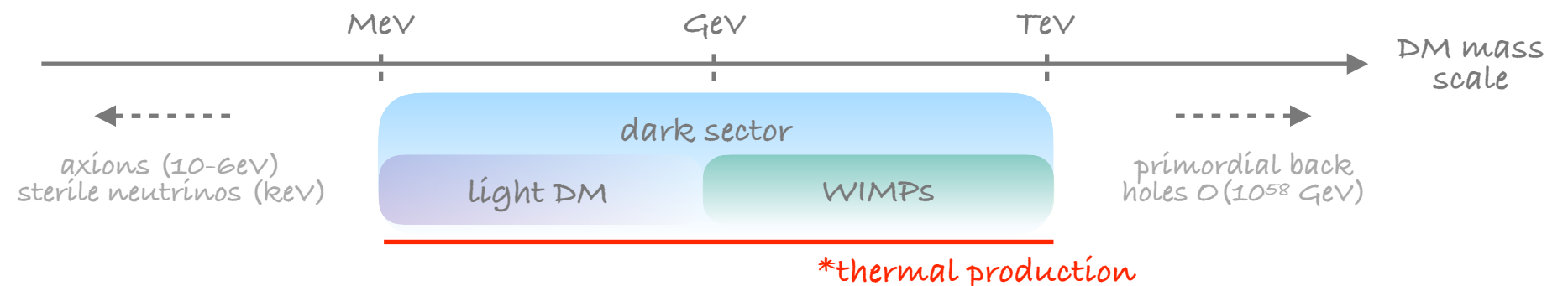
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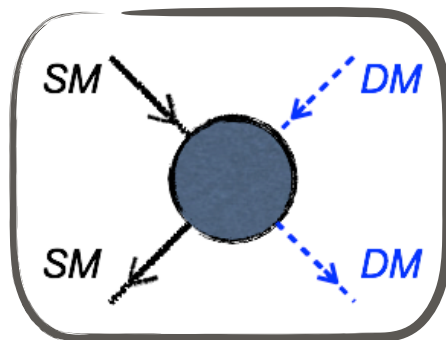


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assume
interactions
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Hunt for dark matter - where and how to study it?

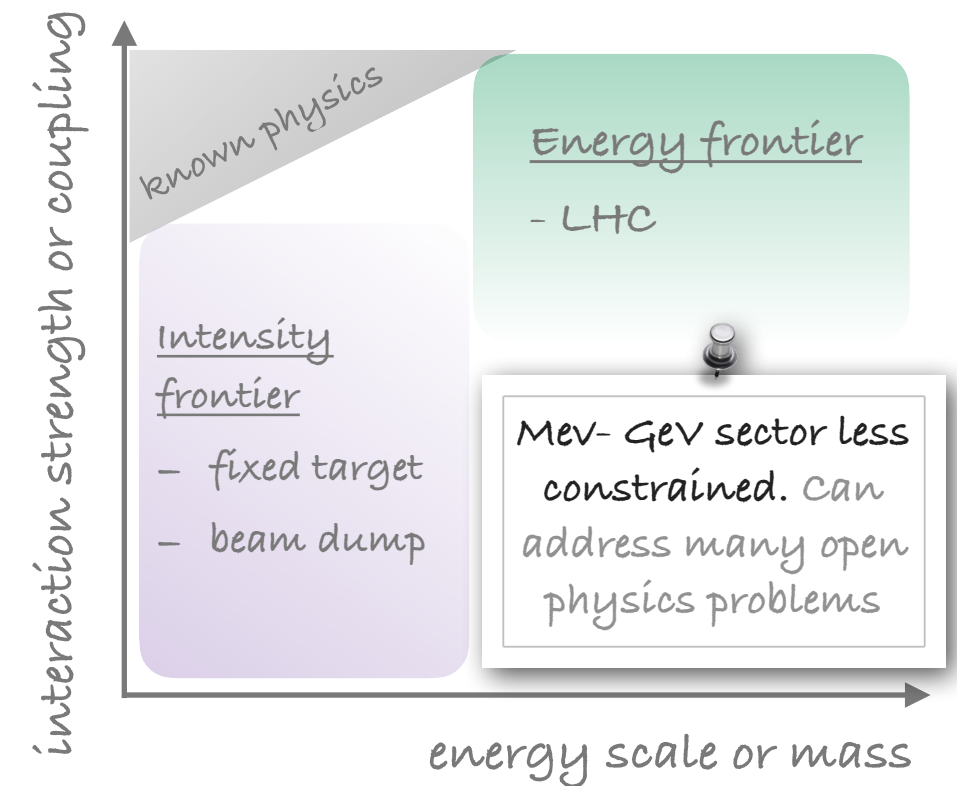
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- 📌 high-intensities

GeV-TeV:

- extensions of the SM at the GeV-TeV scale
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* **Dark sector:** could be heavy, light, and/or long-lived



► Assuming DM-SM interactions enables different searches

- indirect detection, products from DM annihilation
- direct detection, nuclear recoil from DM-nuclei scattering
- colliders: DM production



Complementarity between experiments essential!

eg. info about lifetime if DM discovered at colliders ($\sim 10^{-7}$ s), particle properties compared with cosmological constraints



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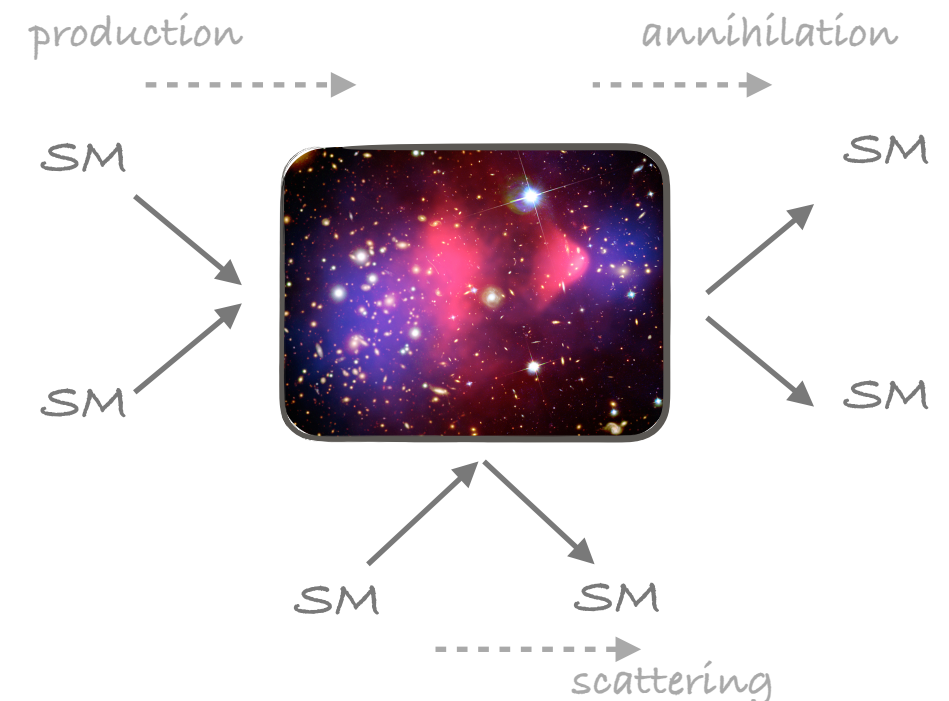
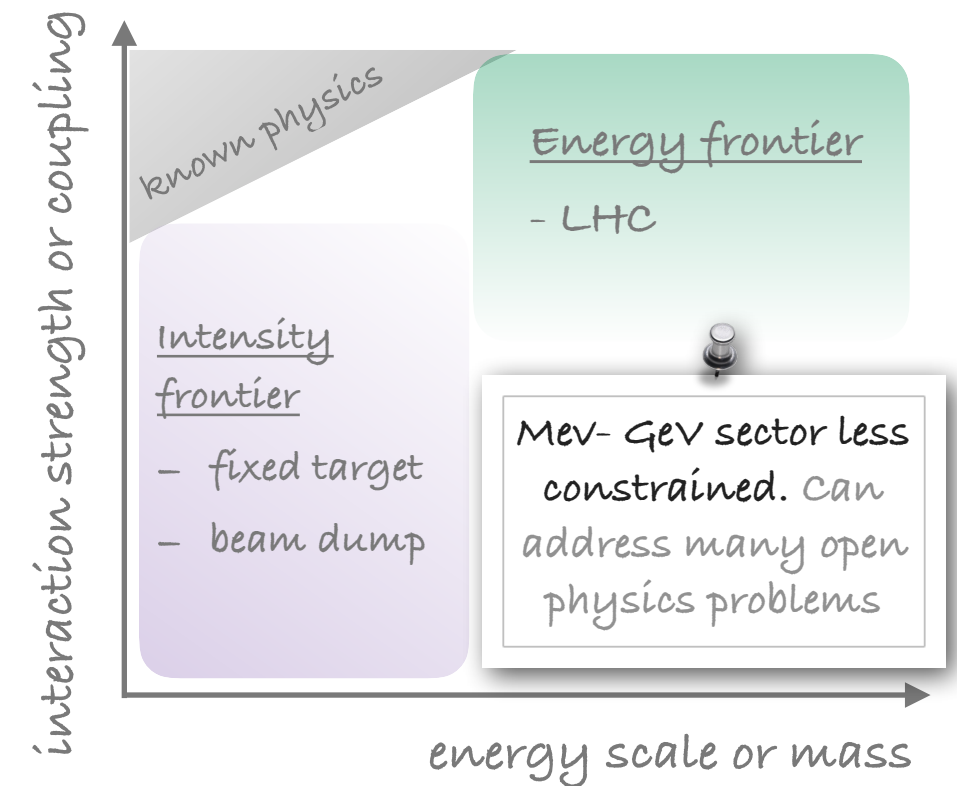
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Hunt for dark matter - investigation tactics



- We do not have information about the DM nature, how to hunt it?
 - remain very general and make very little assumptions, eg. "is it a 2D shape?"
 - more assumptions and tests more specific models, eg. "is it a 2D shape, yellow color and with only 90° angles?"

Hunt for dark matter - phenomenology

Intensity frontiers: MeV-GeV

- small couplings to SM, challenging to detect but guide to relevant interactions
- * **vector**: dark vector boson (dark photon) mediator that mixes with SM photon
- * **scalar**: dark Higgs mediator, mixing with SM Higgs, which leads to couplings to fermions
- * **fermion**: heavy neutral lepton (HNL) mediator, that can mix with SM neutrinos
- * **pseudo-scalar**: axion-like particle can have couplings to SM fermions or bosons

Energy frontiers: GeV-TeV



Signatures

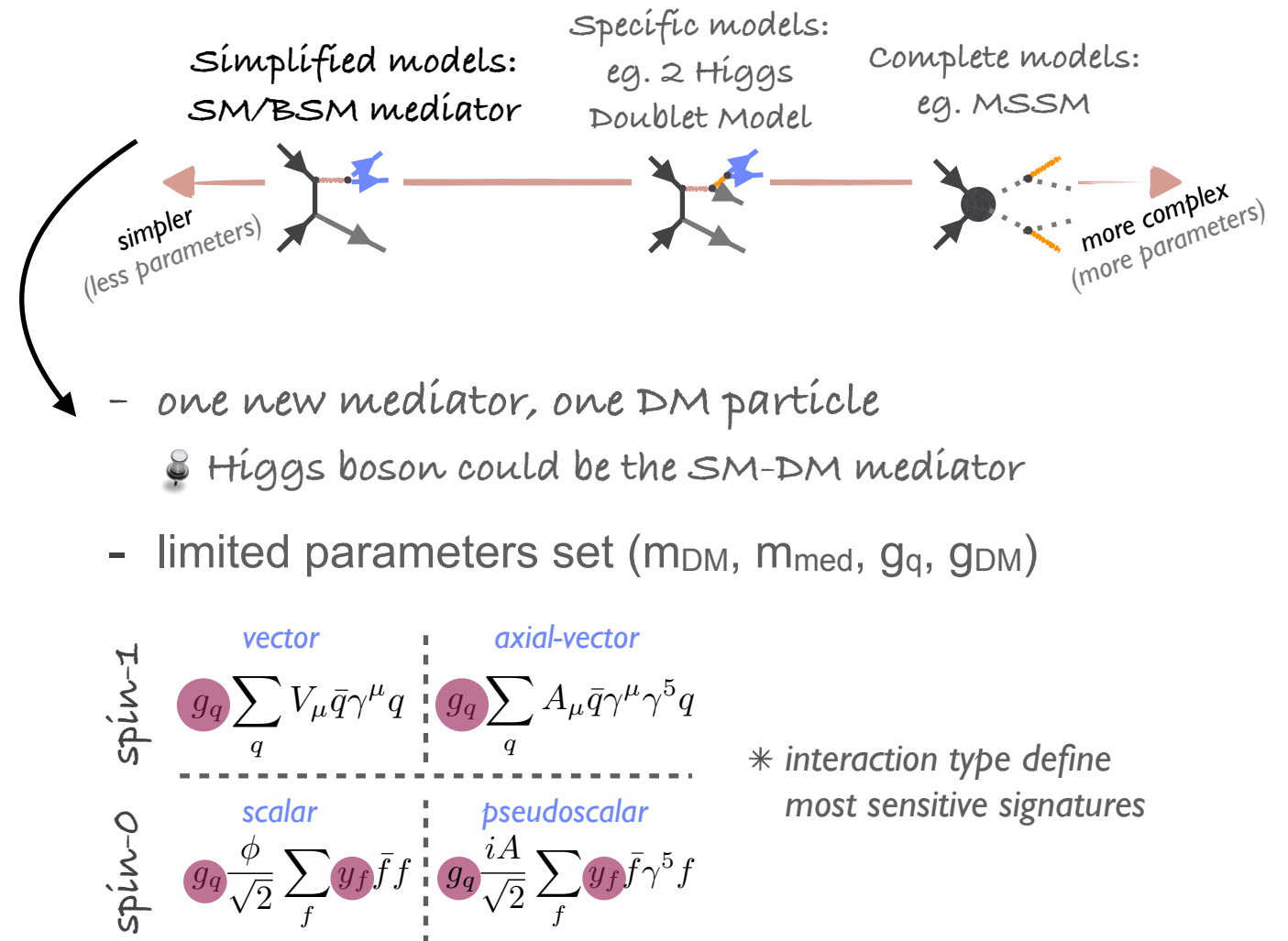
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- **visible**: dark mediator particles can decay back to SM particles (especially if sector's lightest state)
- **displaced (long-lived)**: production of dark sector particle with significant lifetime that decays visibly to SM

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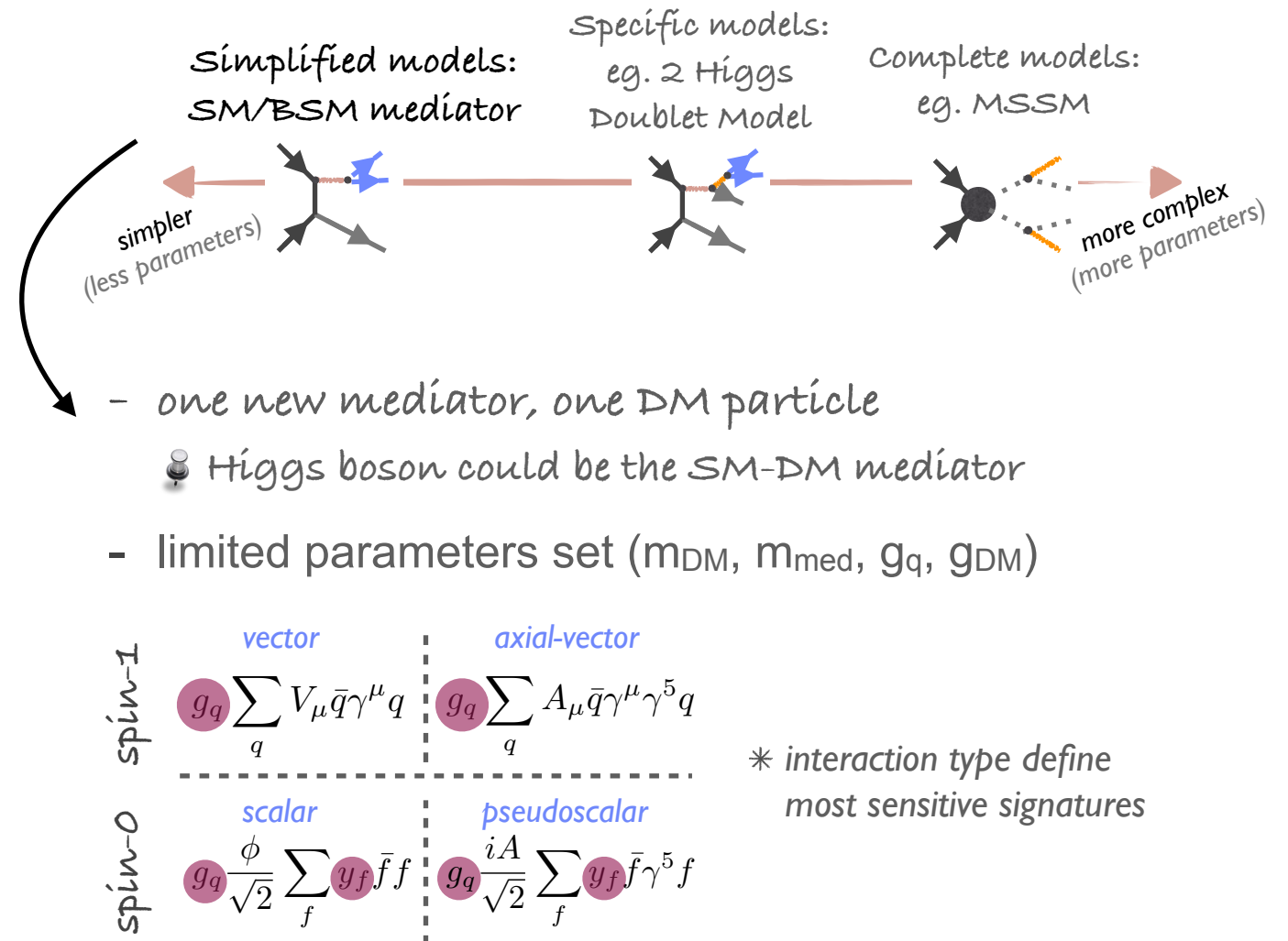
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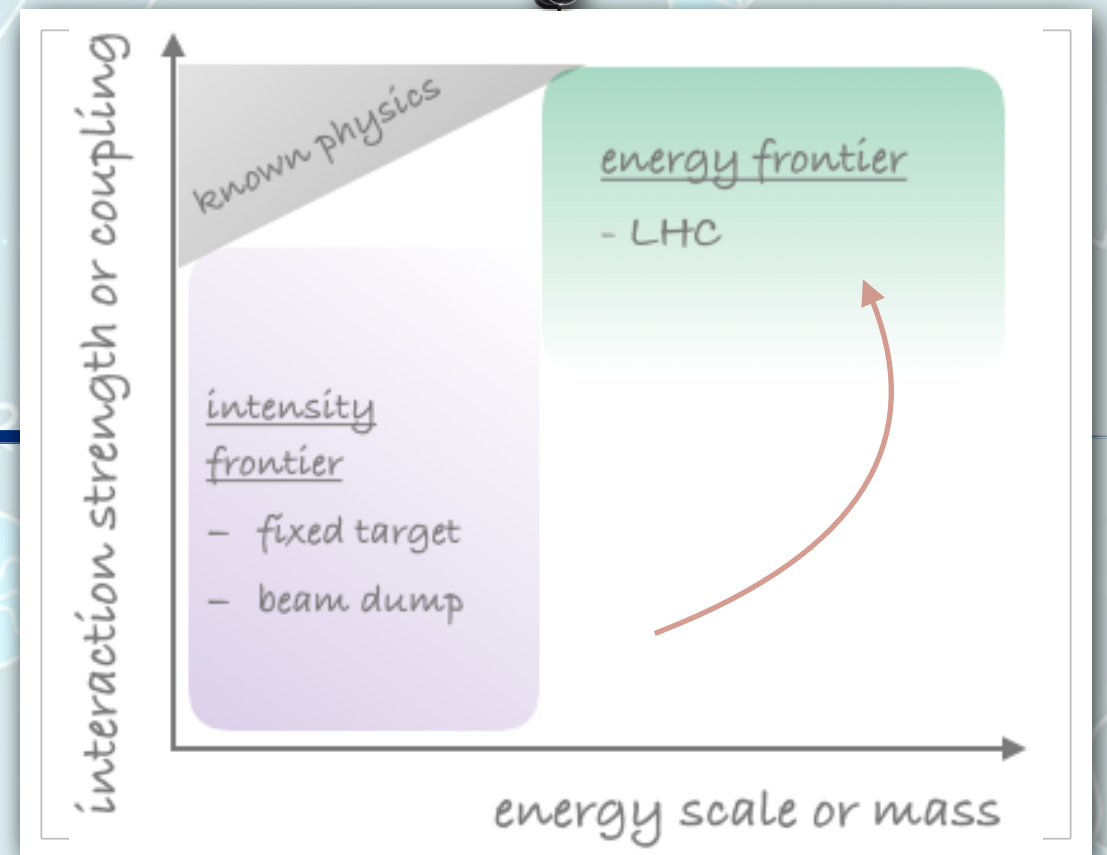
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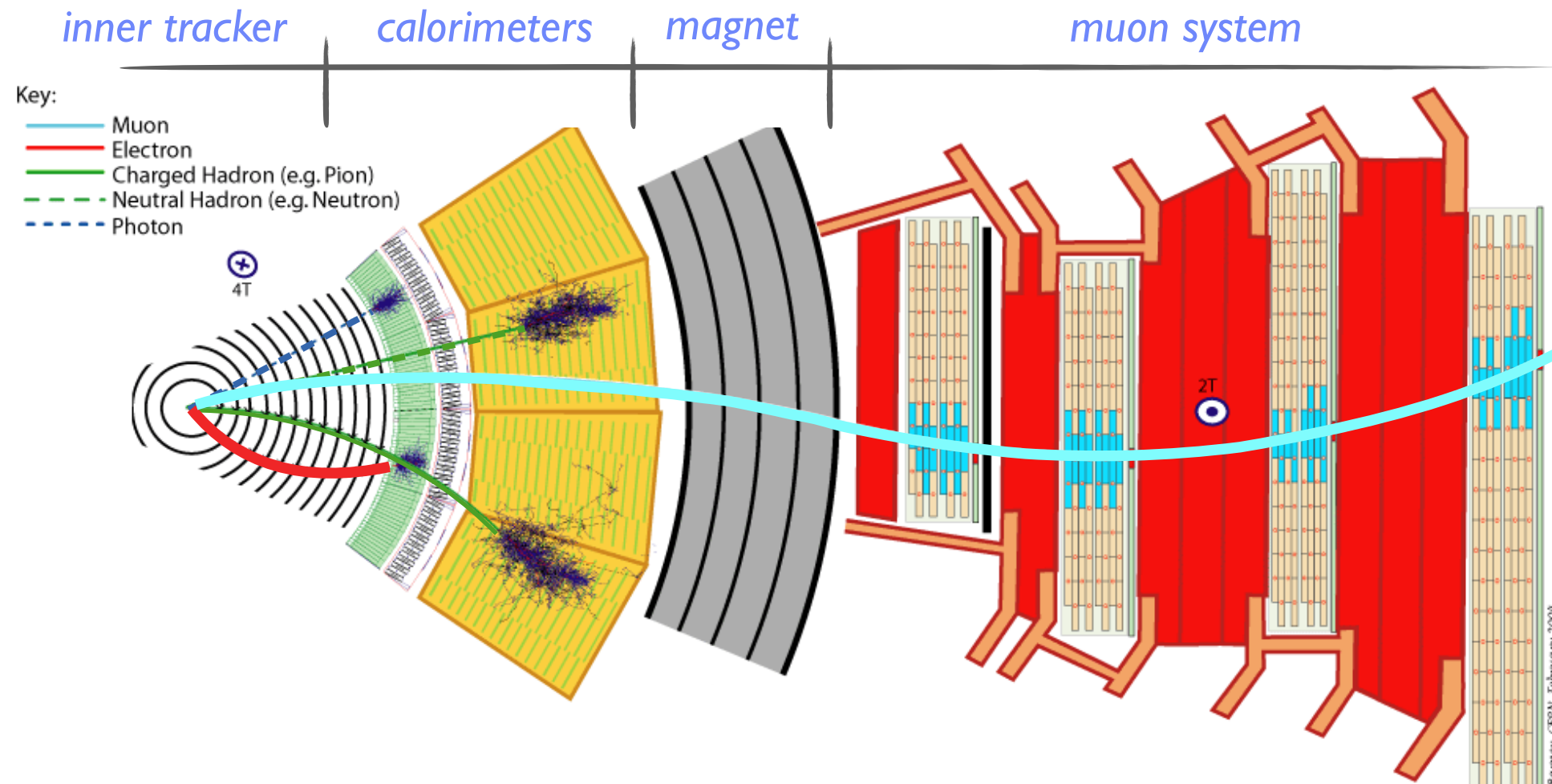
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The energy frontier

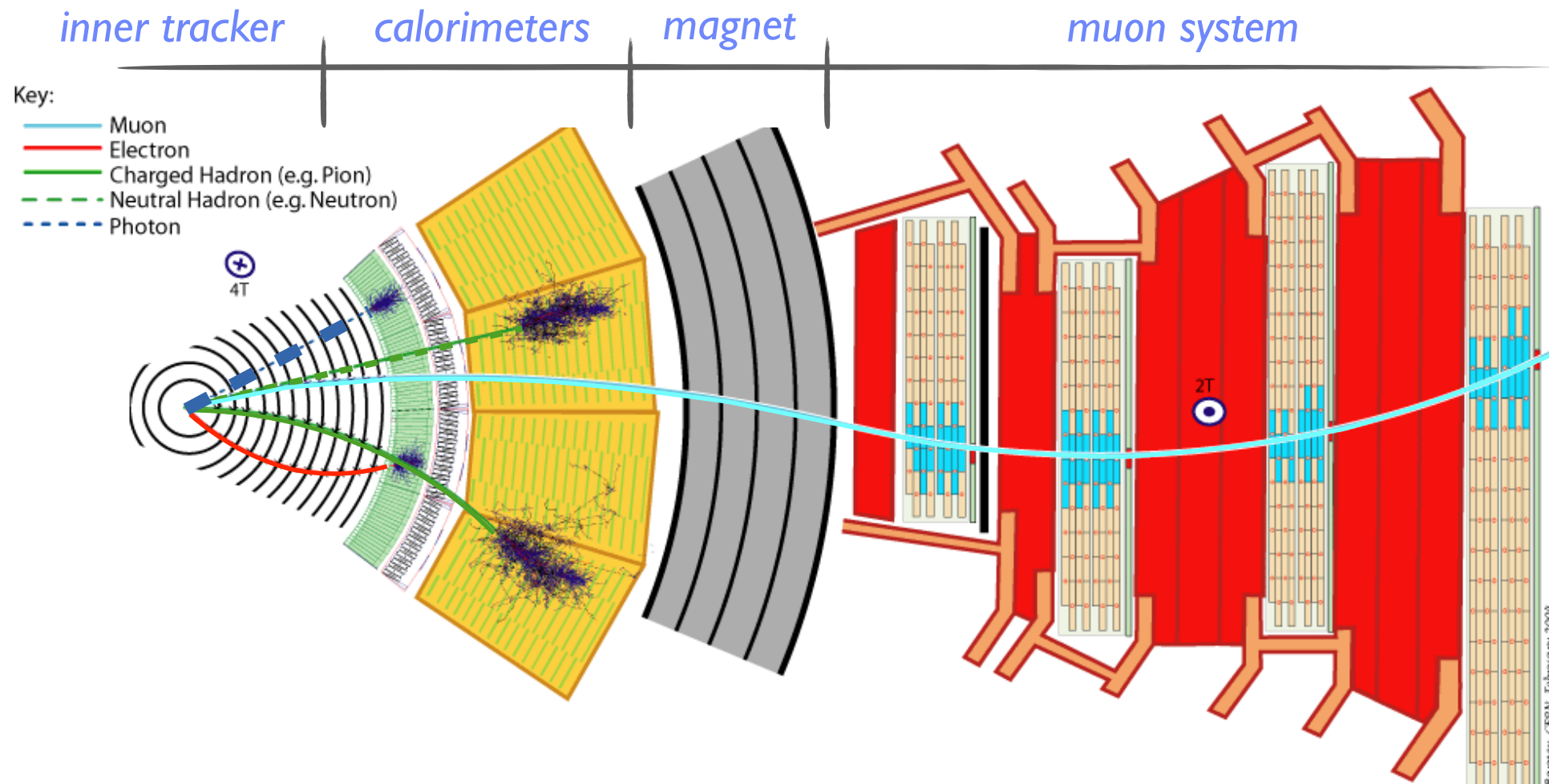


Hunt for dark matter - how to detect it?



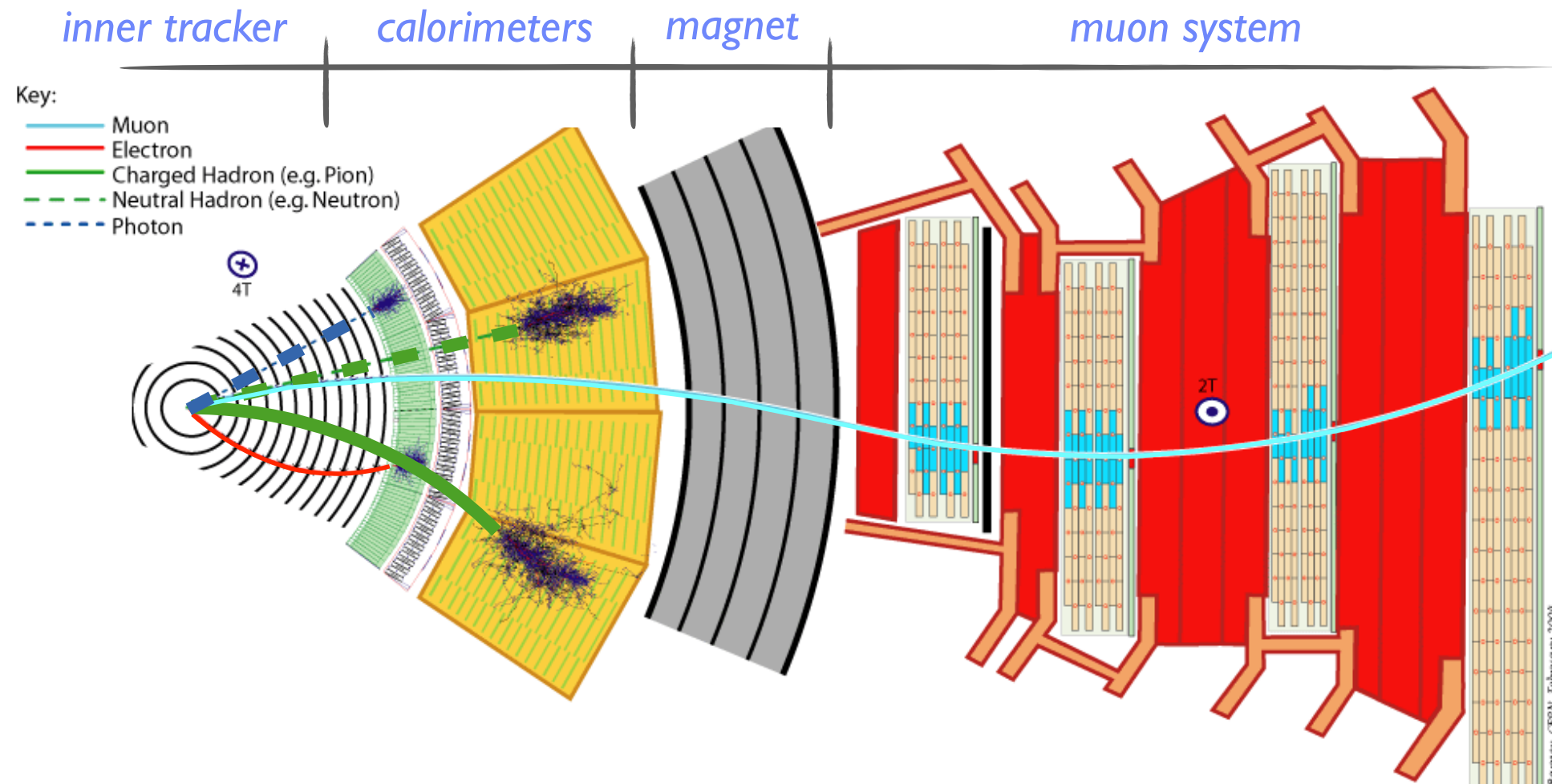
- **electron** (**muon**), from inner tracker tracks and energy in calorimeter (track in muon spectrometer)
- photon, from energy deposits in electromagnetic calorimeter
- jets from quarks and gluons, partons hadronize in colour-neutral particles groups, so-called jet.
Parton energy and momentum reconstructed clustering all particles from hadronization

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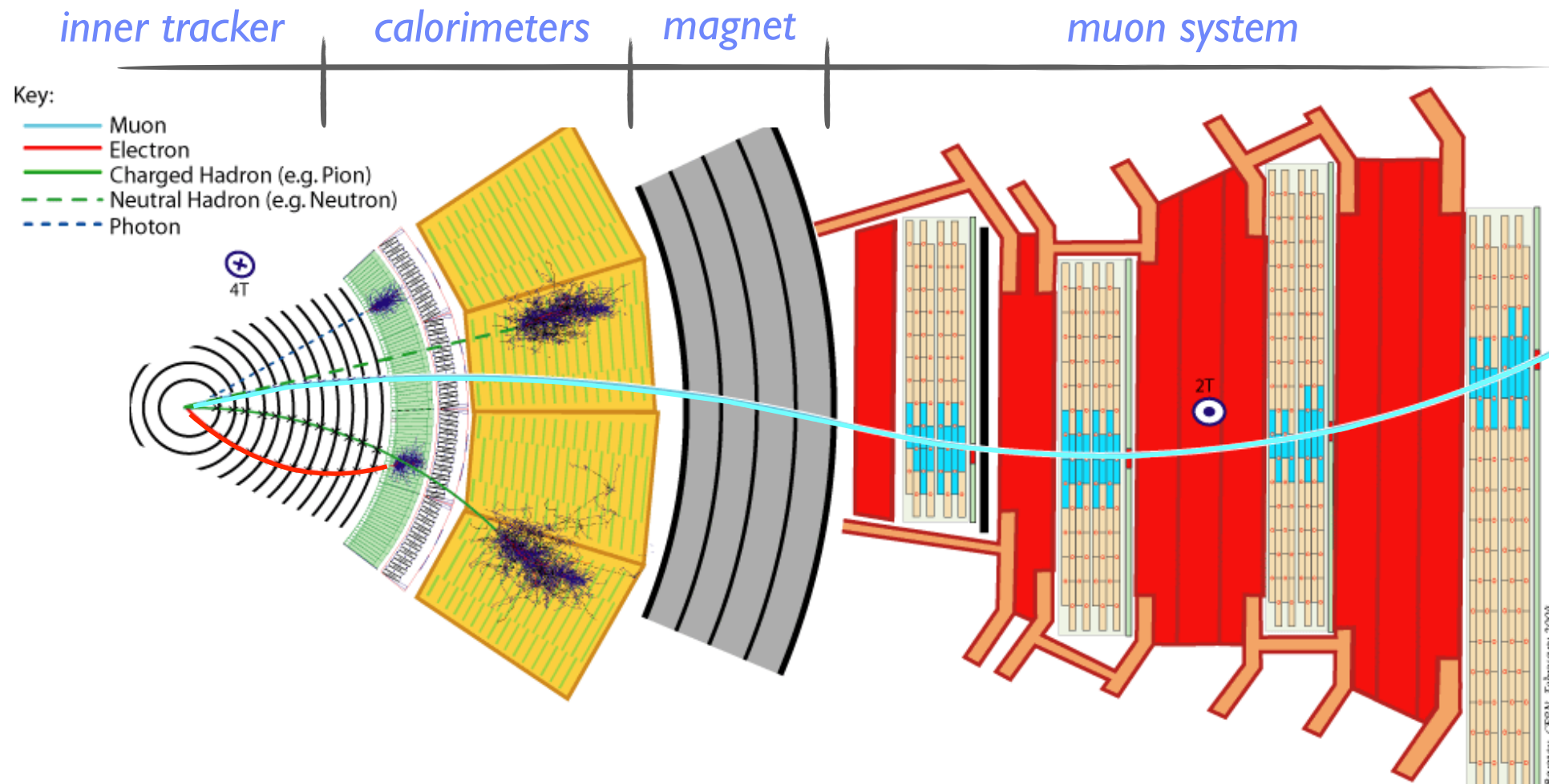
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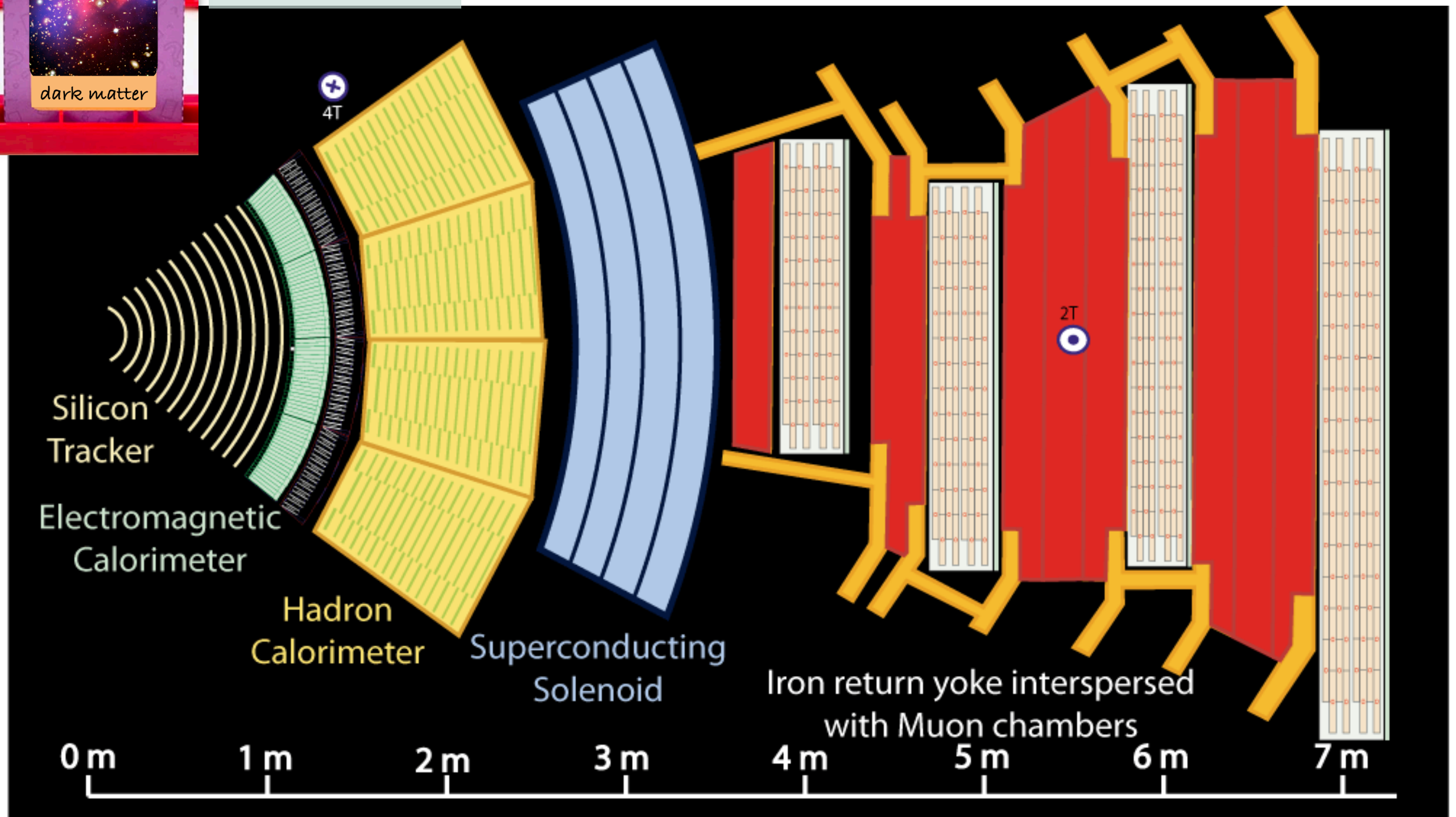
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 *Dark matter?*

Hunt for dark matter - how to detect it?



- stable
- electrically neutral
- massive



Hunt for dark matter - signature at colliders

- DM could be produced at colliders (rare process)

no direct trace in the detector, but could create a p_T imbalance (**MET**)

conservation of momentum

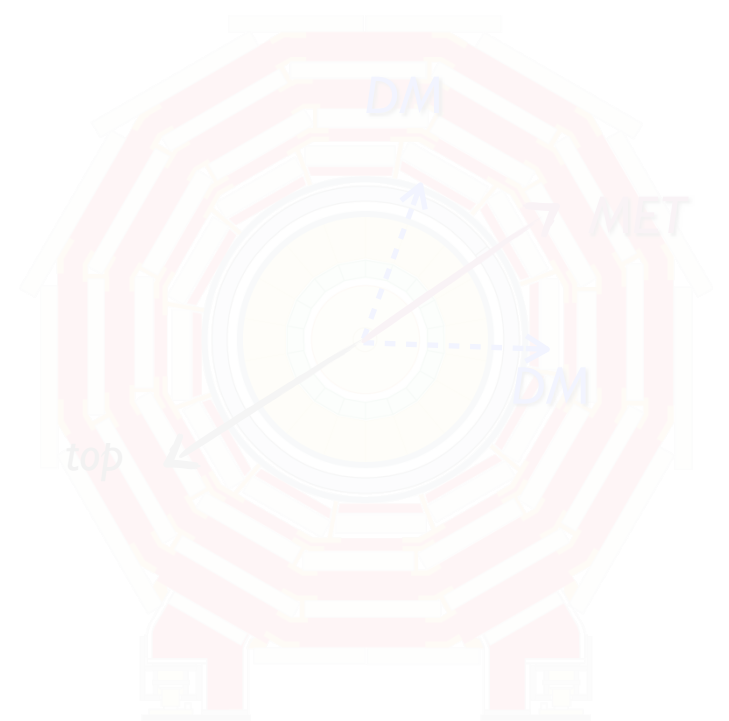
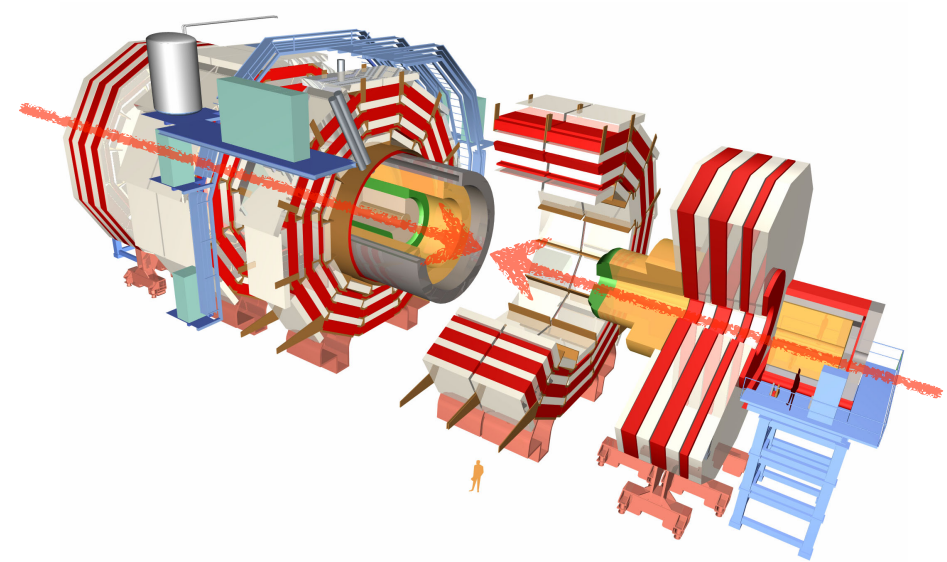
- no info on longitudinal momentum of colliding partons
- but total initial parton $p_T = 0$

- need to be conserved after collision $\sum \vec{p}_T = 0$

- $\sum \vec{p}_T \neq 0$ particles escaped detector carrying

$$\vec{E}_T^{miss} = - \sum \vec{p}_T$$

📍 $|\vec{E}_T^{miss}|$ = missing transverse energy (MET)



To see the invisible we need the visible ...

- need visible particle to which DM particle recoils against
- “mono- X searches”: X includes jets, vector bosons, top, ...

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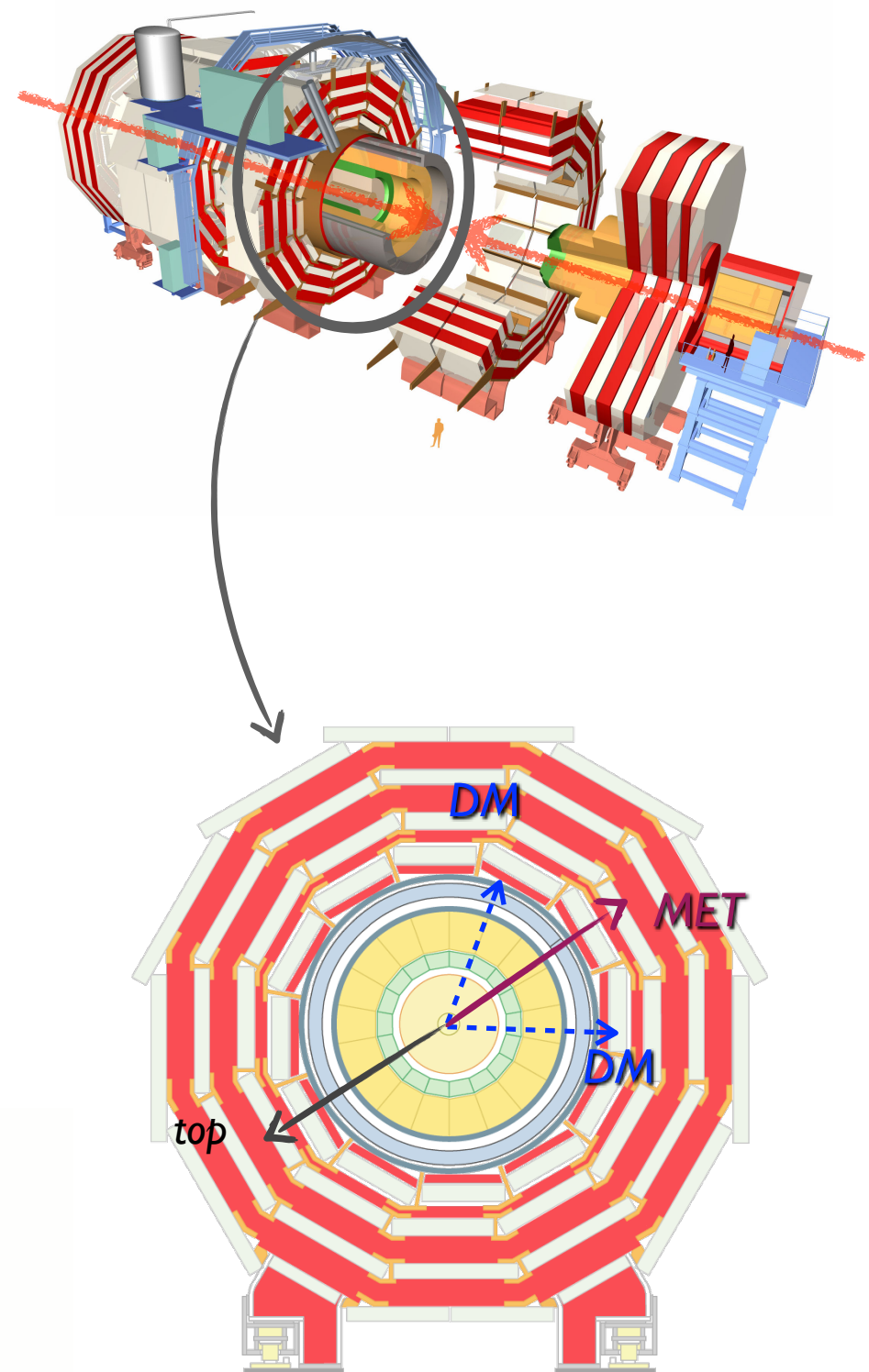
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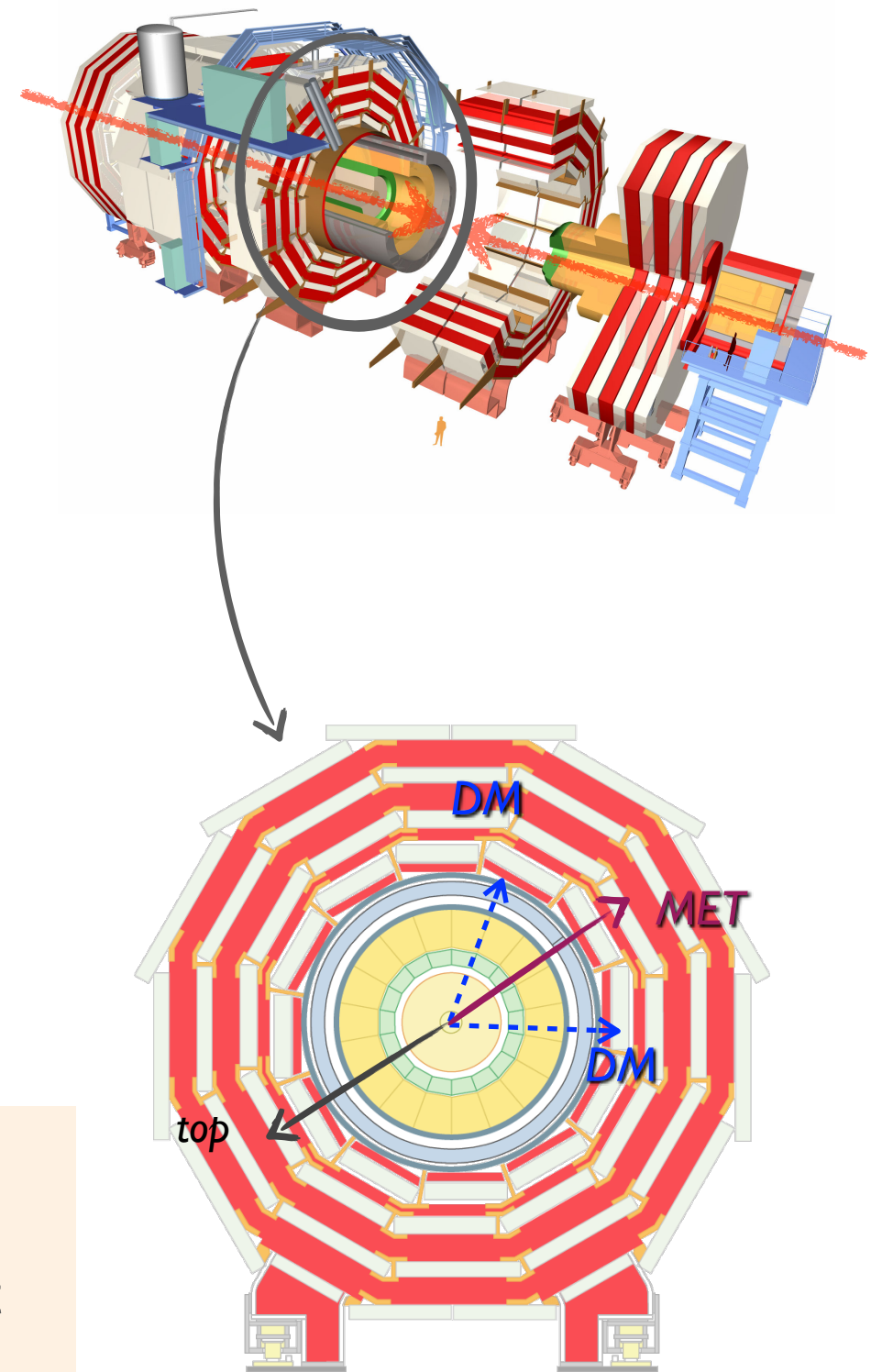
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Hunt for dark matter - how to search it at colliders?

1- *Selection*: DM appears as events excess in MET tail wrt SM

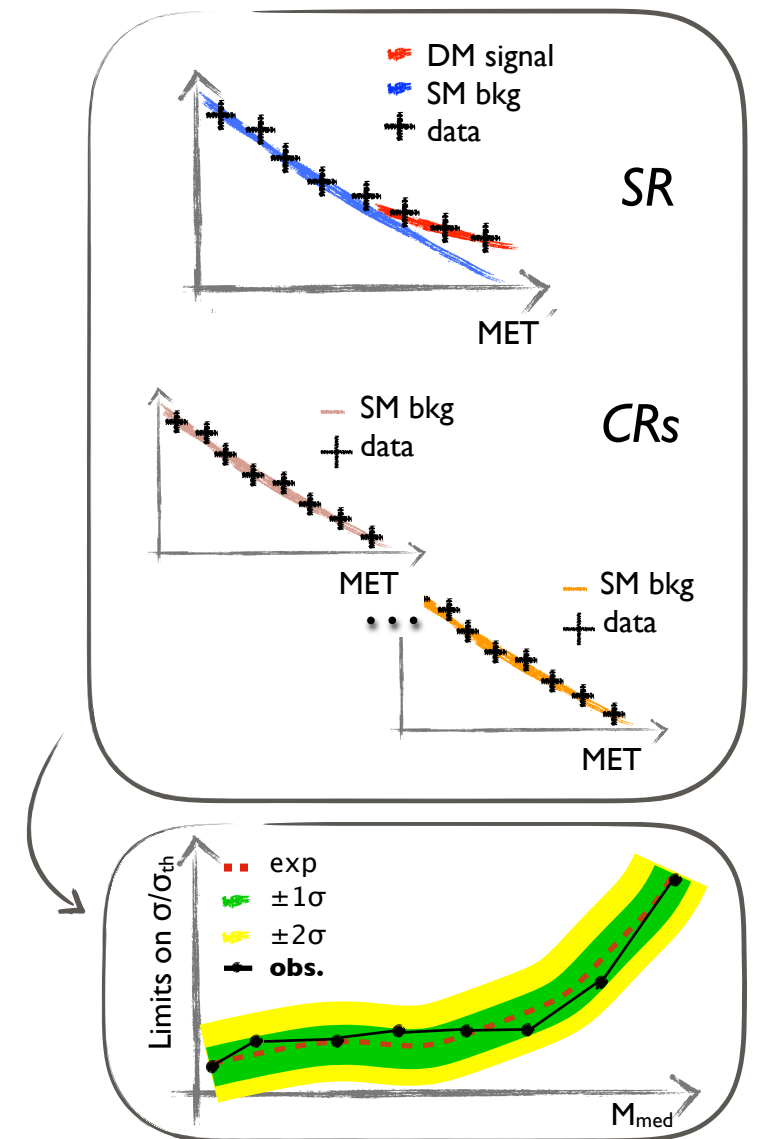
- no striking signature, eg. mass peak, m_T kinematic endpoint
- look for excess in region enriched in signal (*signal region* - SR)

2- *Bkg*: precise evaluation of SM processes in SR essential

- achieved through use of multiple control regions (CRs)

3- *Results*: compare SM predictions with data

- *Excess of events in data*, did we find DM?
- *No excess*, interpret result in terms of theory model parameters



Experimental challenges for invisible signatures

- * accurate E calibration/resolution of visible objects (*"fake" MET from mis-measured jets*)
- * precise particle reconstruction and identification
- * mitigate effects from additional pp collisions (pile-up)
- * MET thresholds affected by trigger (very high collision rates)

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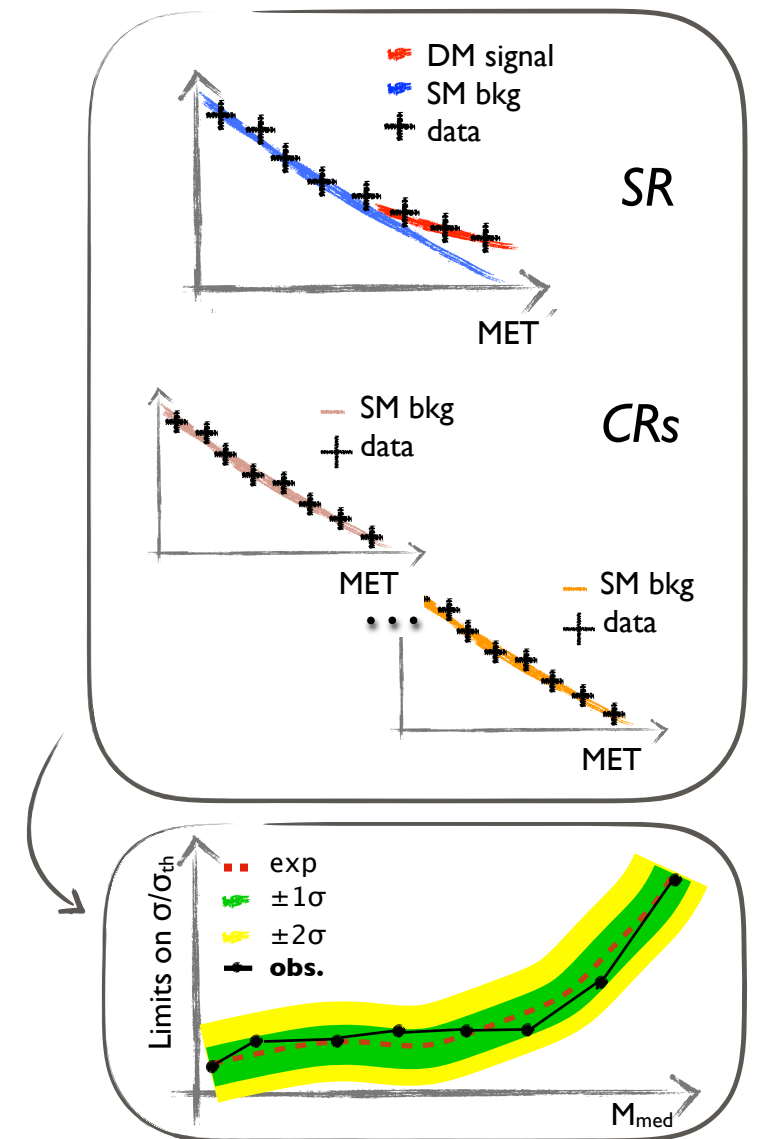
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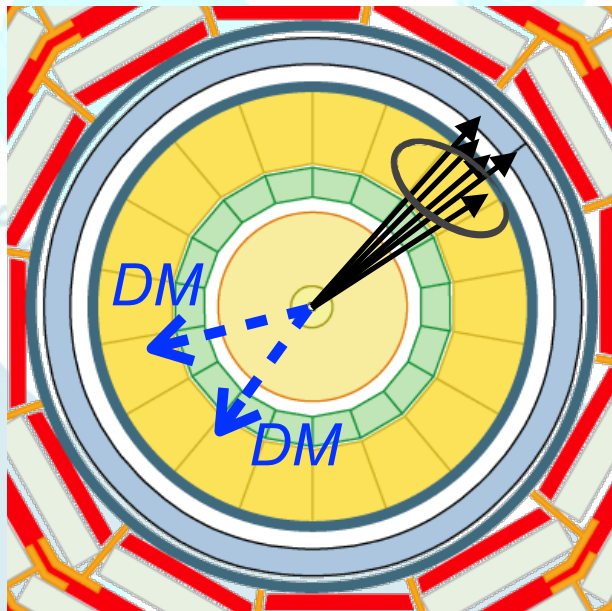


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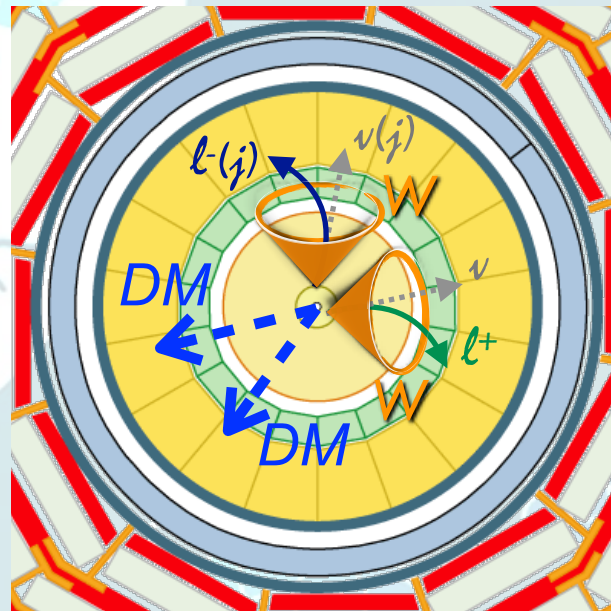
Invisible signatures

mono-jet/ $V(=W,Z)$



* CMS: EXO-20-004

mono- WW



* CMS: EXO-21-012

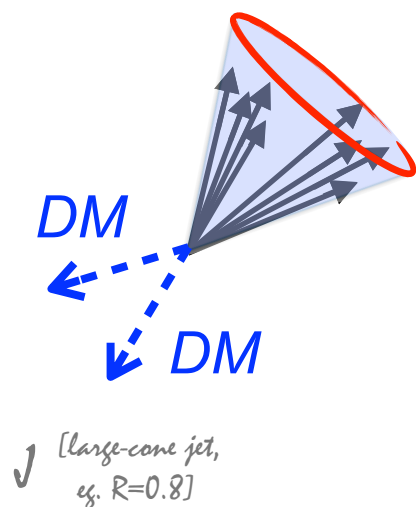


* ATLAS: PLB842(2023)

* CMS: HIG-21-007

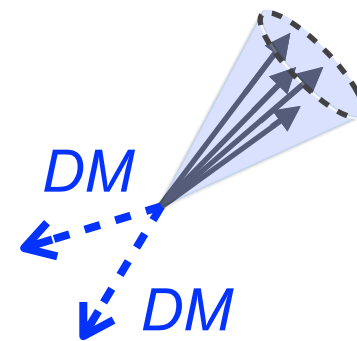
Hunt for dark matter - DM+jet/V search

- 1 - Selection: events categorized based on jet nature



mono-V

- * MET > 250 GeV
- * ≥ 1 jets, $p_T(j_1) > 250$ GeV
- * ML technique to identify V hadronic decays
- * jet mass consistent with V



mono-jet

- * not selected as mono-V
- * ≥ 1 jets, $p_T(j) > 100$ GeV
- * b-tagged jets veto

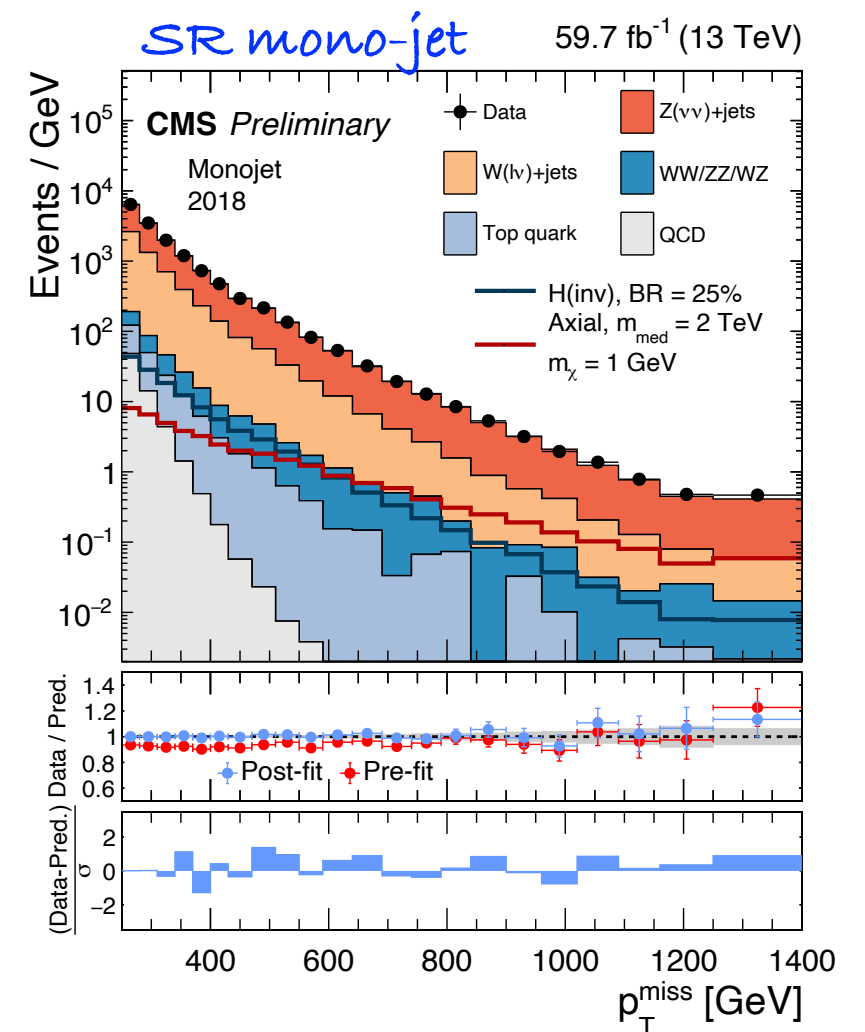
- 2- Bkg:

- Z($\nu\nu$)+jets and W(lv)+jets from CRs

- 3- Results: combined fit of SRs, CRs

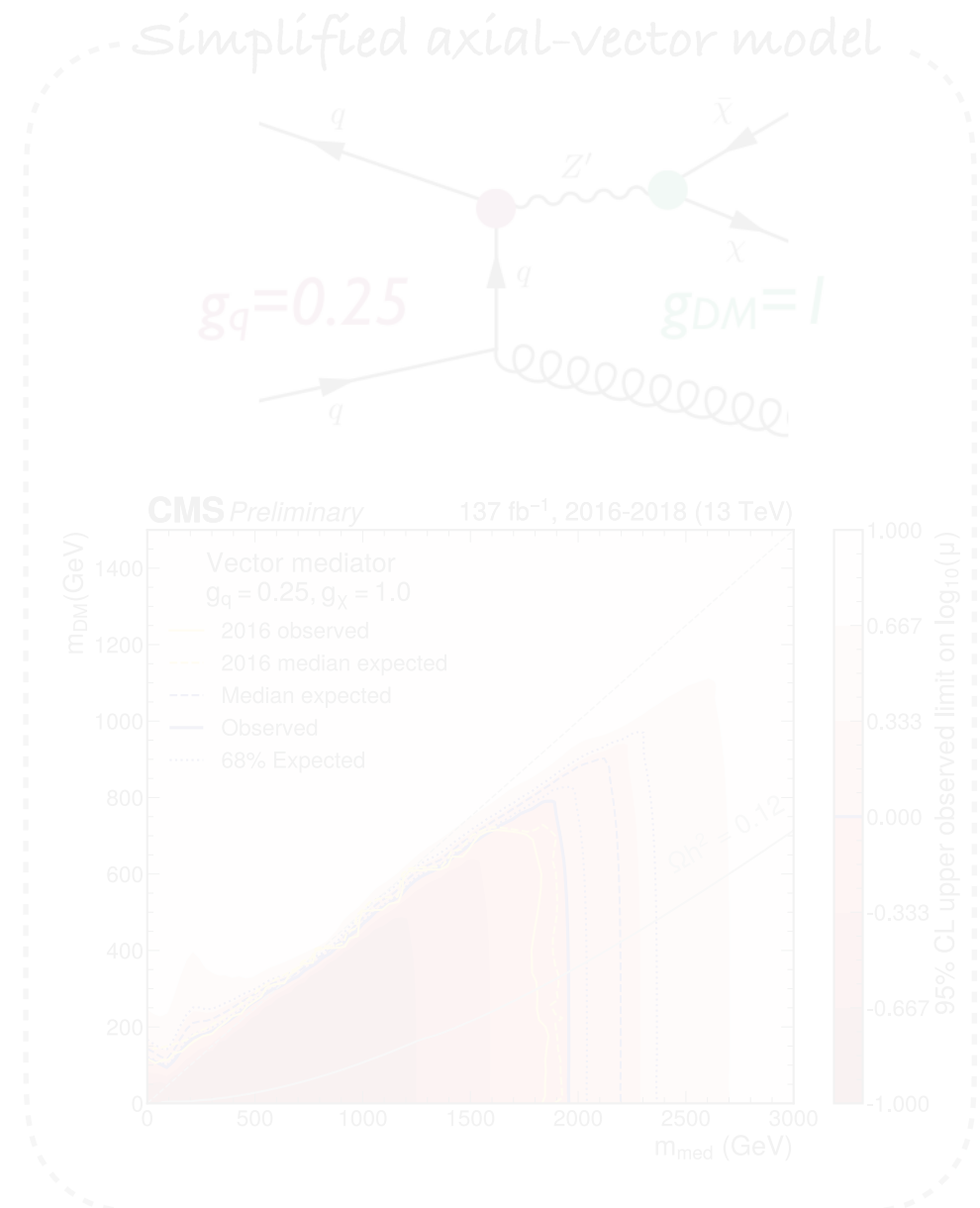
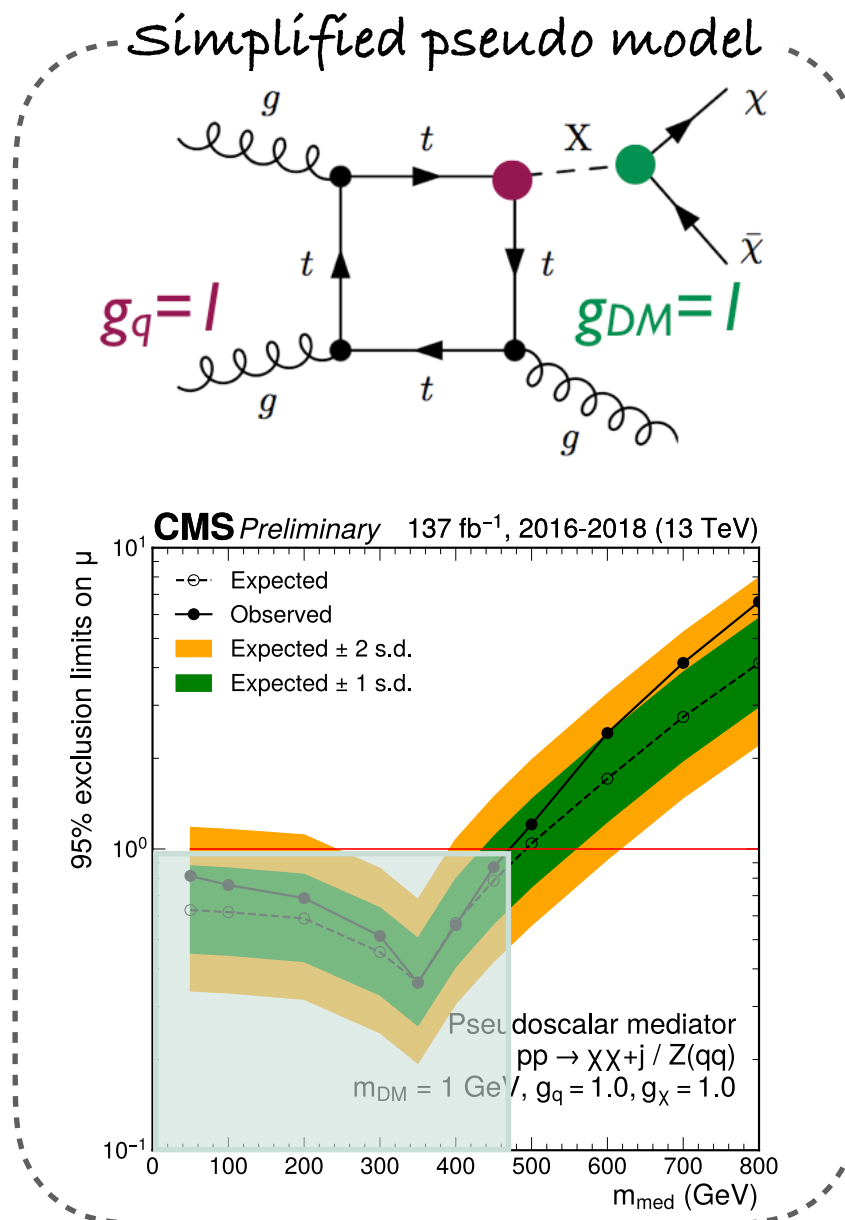
- systematic unc. included as nuisance parameters

- * prefit: as from simulation
- * post-fit: after allowing simulation to vary within unc and the scaling factors from CRs



Hunt for dark matter - DM+jet/V search

► **3- Results:** interpretation in terms of DM model, upper limits at 95% CL on cross section

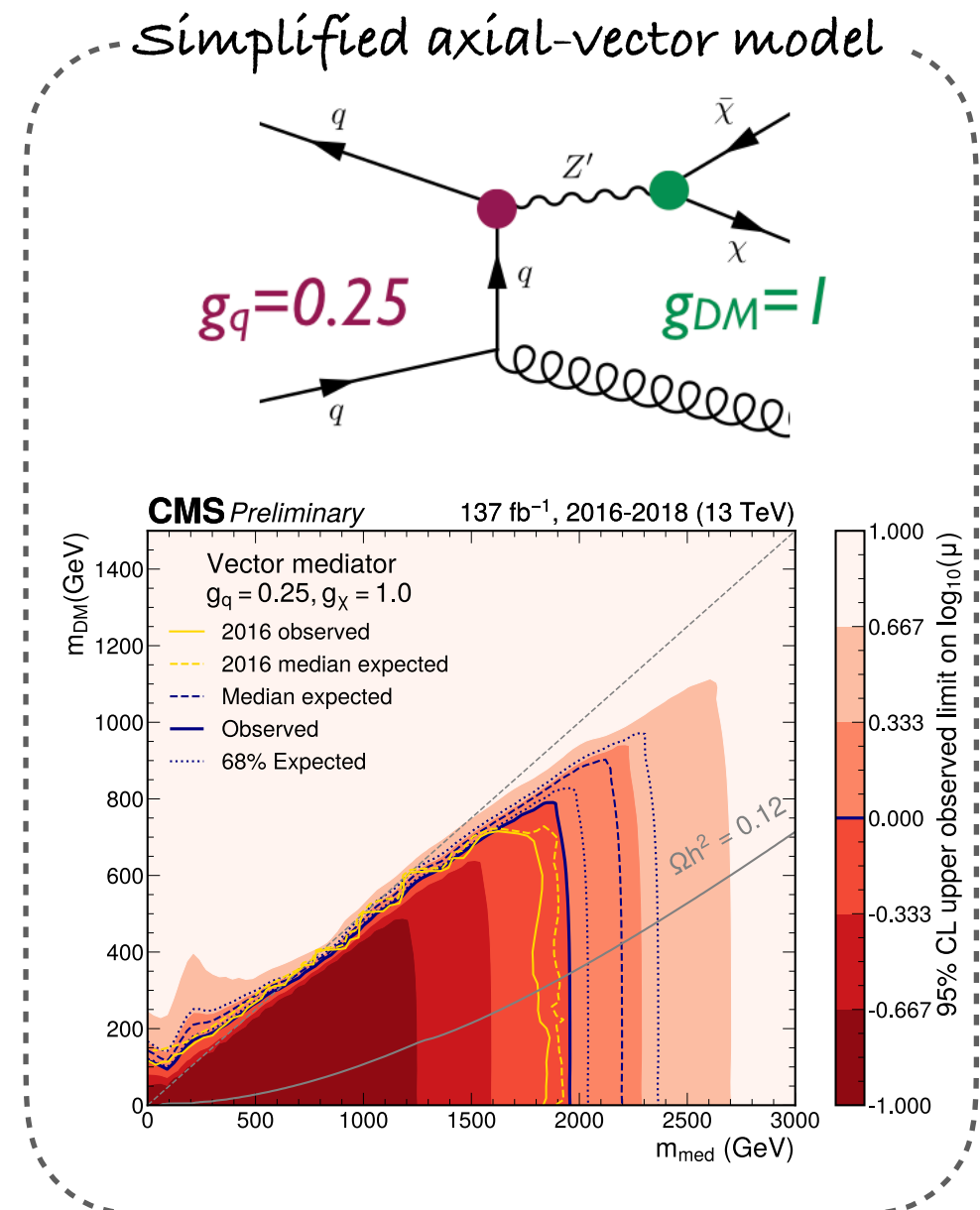
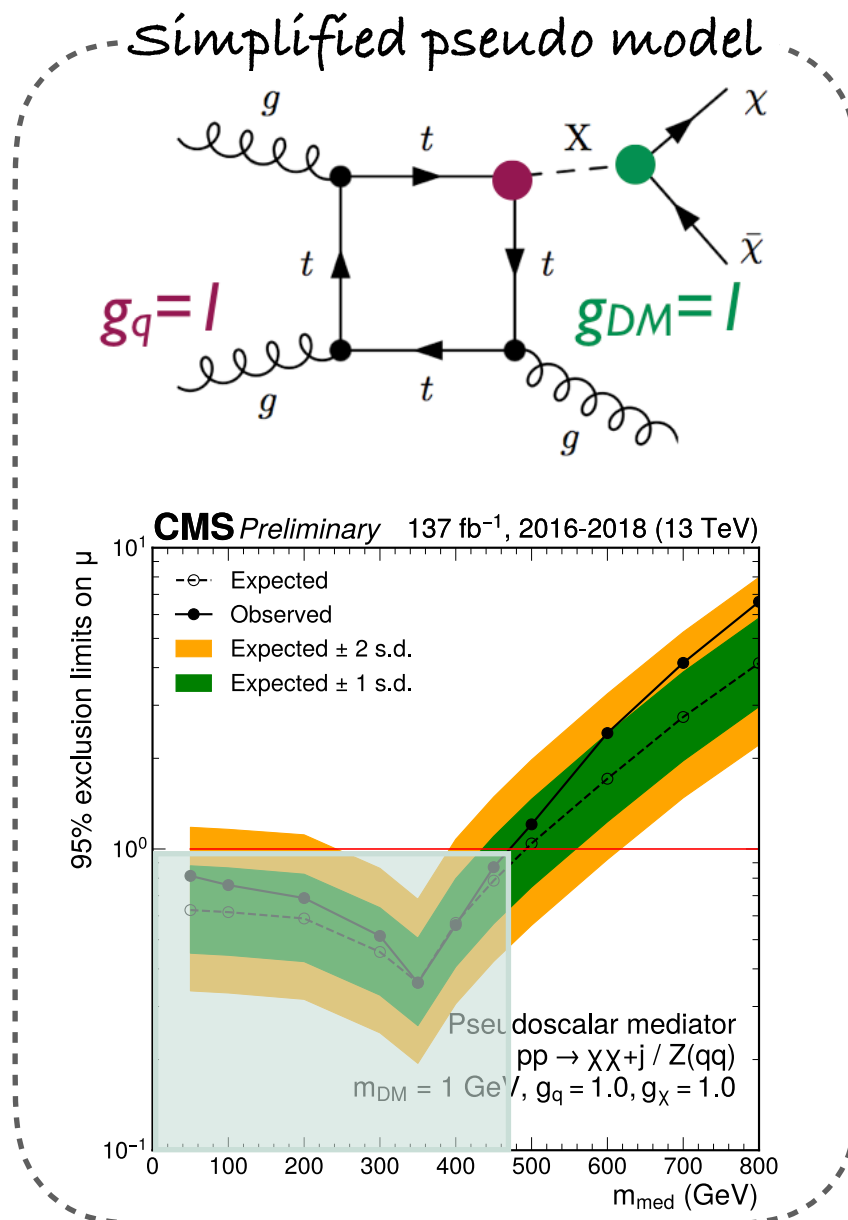


- * $\mu = \sigma / \sigma_{th}$, $\mu = 1$ exclude the theory value, $\mu < 1$ exclude below theory value, $\mu > 1$ does not exclude theory value
- * parameter: cannot scan all parameters at once. Fixed ones only affect xsec but not kinematic (selection)

- * here $\mu = \sigma / \sigma_{th}$ in on z axis (notice here $\log_{10}(\mu)$)
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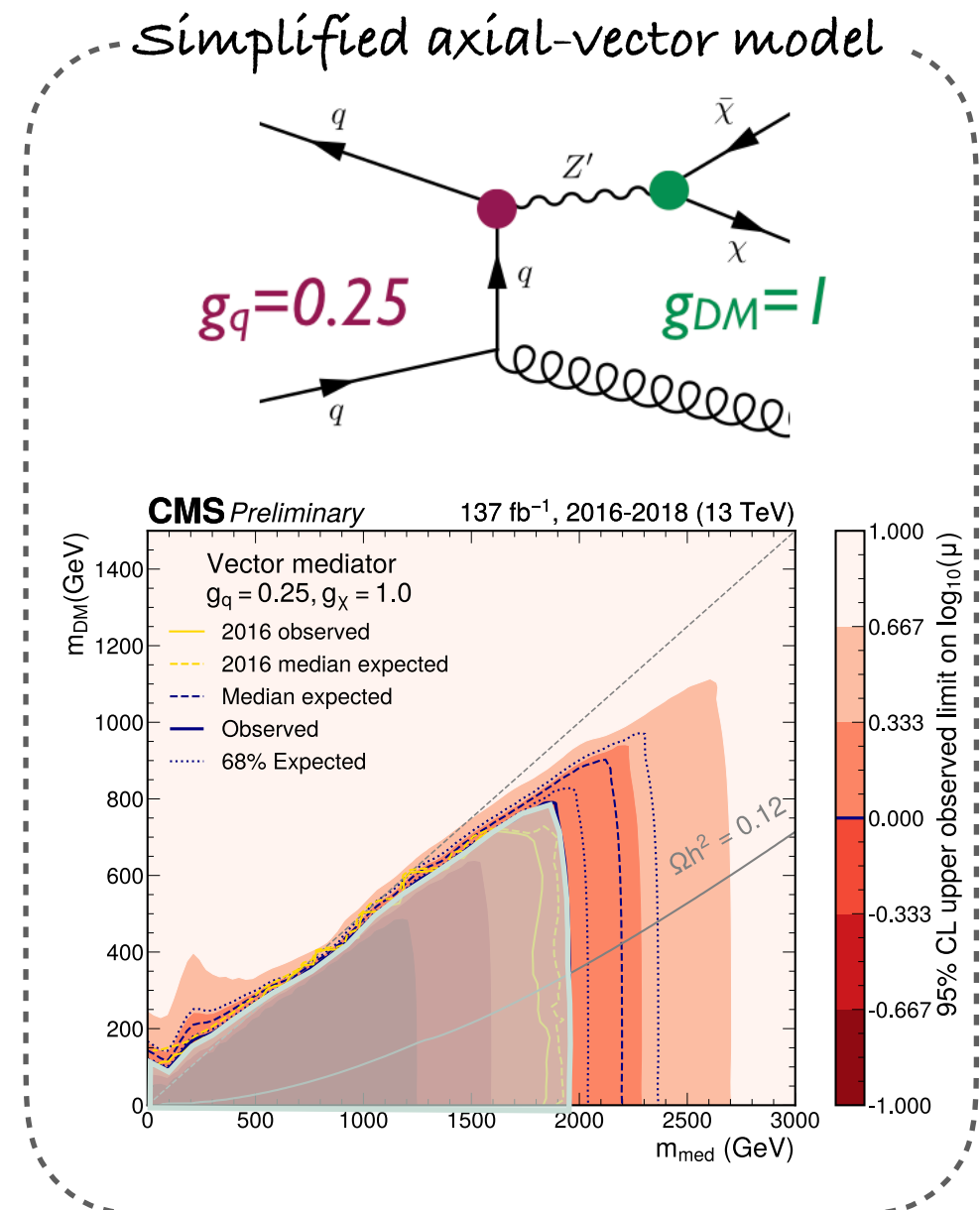
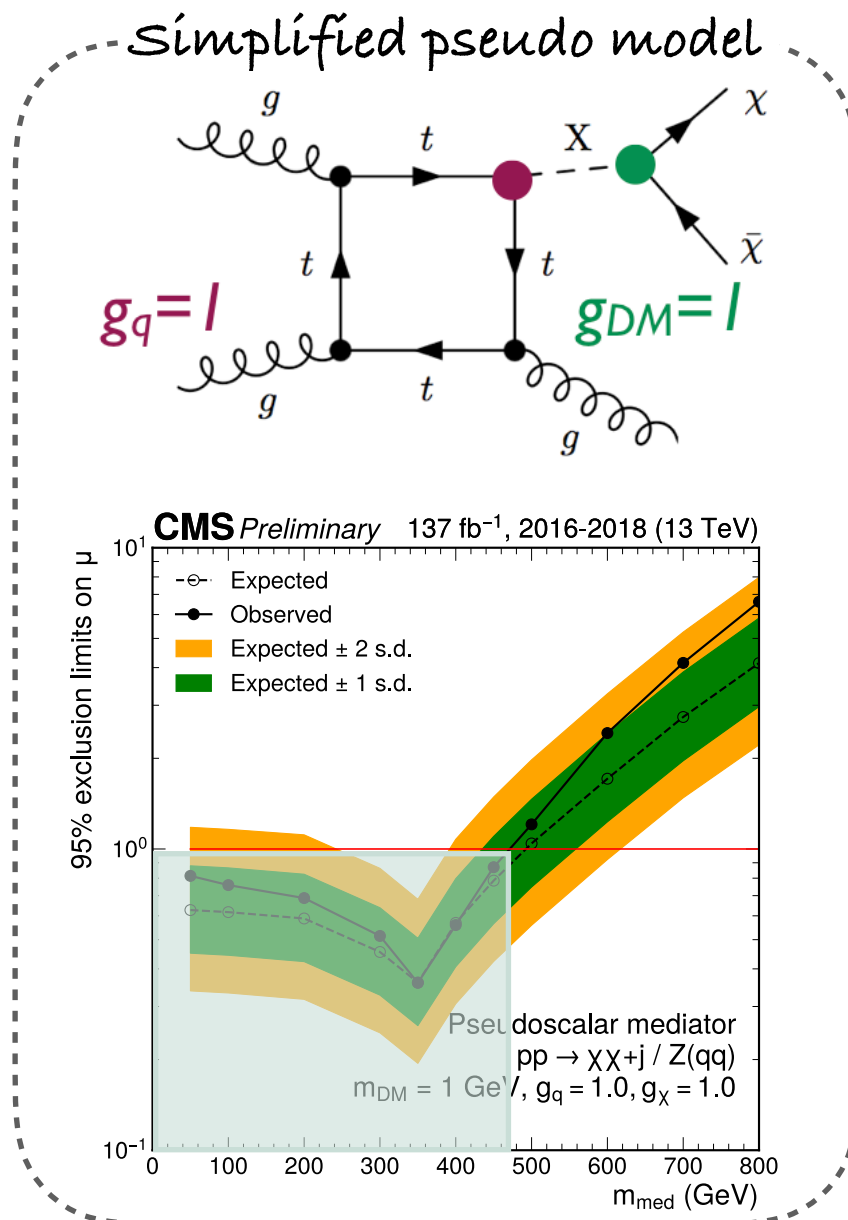


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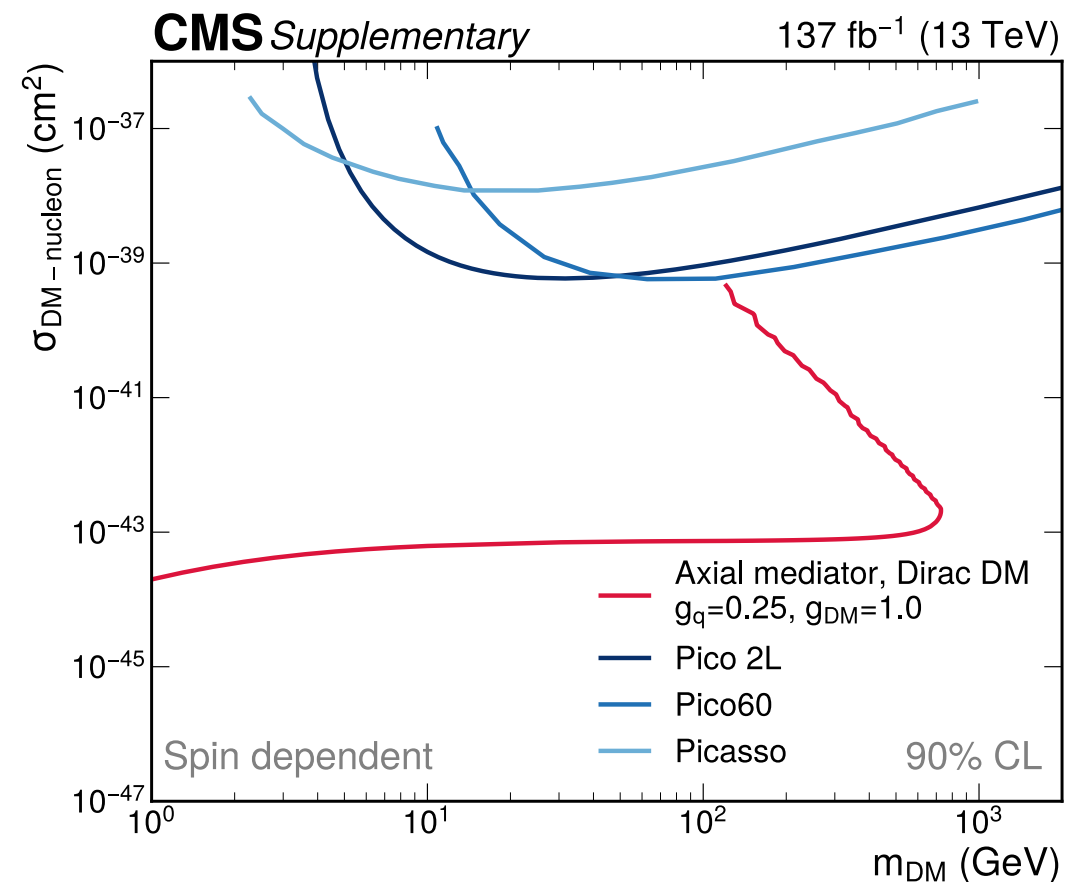
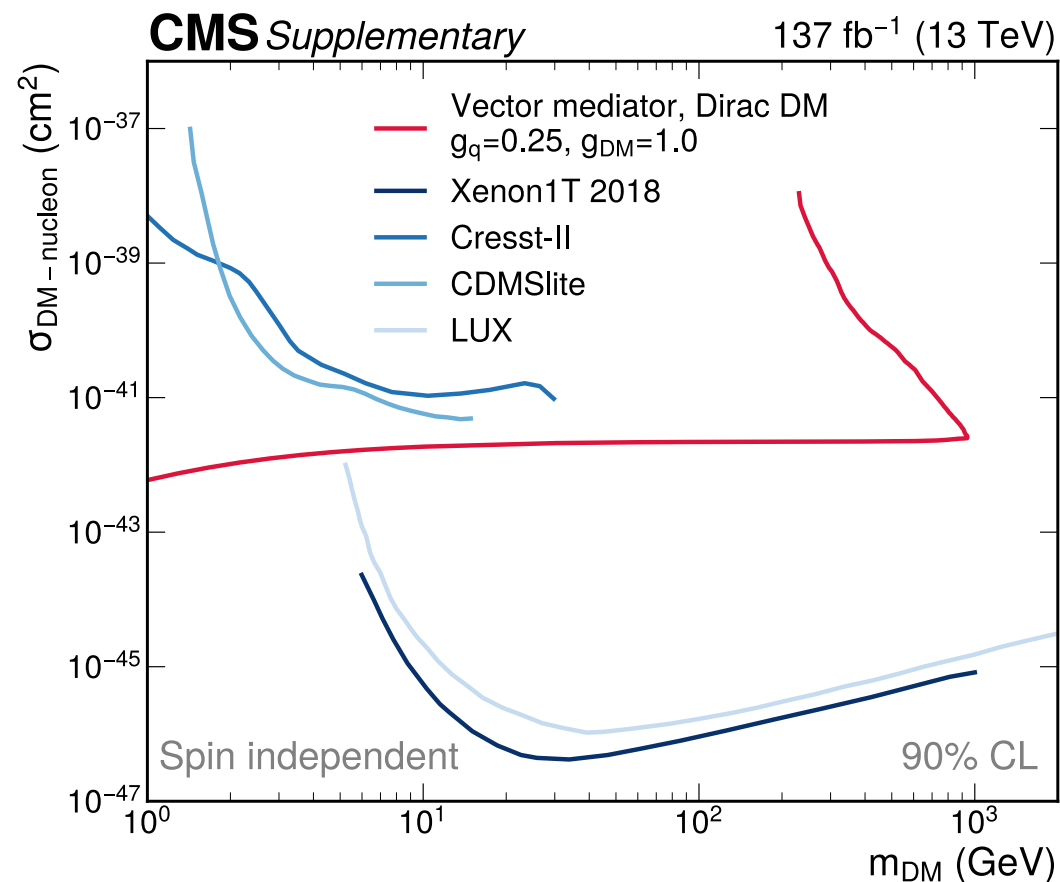


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Hunt for dark matter - interplay with direct detection

► **3- Results:** 90% CL lower limits on interaction xsec between DM and nuclei

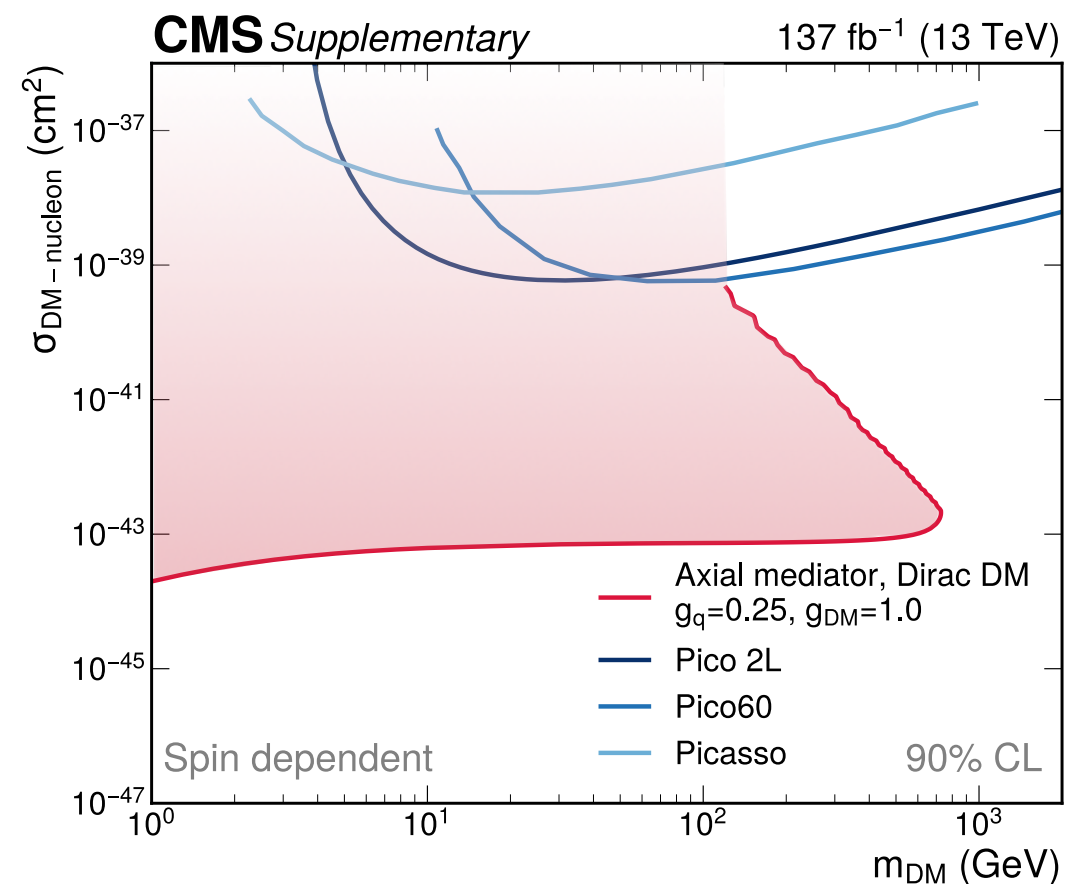
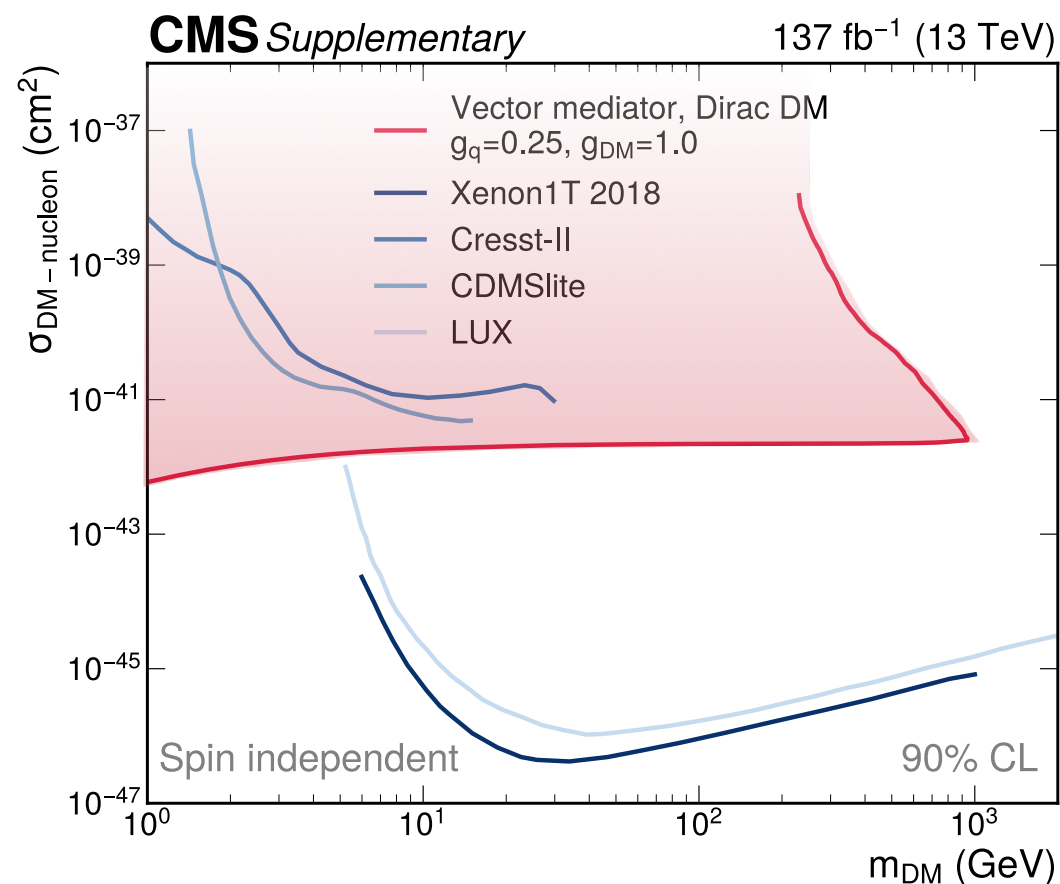


Comparison recommendations [[arXiv:1603.04156](https://arxiv.org/abs/1603.04156)]

- * DM non-relativistic: dominant DM-nuclei interactions spin-independent and spin-dependent scattering
 - * vector/scalar mediator lead to a SI interaction
 - * axial-vector/pseudo-scalar lead to SD interaction
- * Very model dependent
 - * DD bounds may be valid for multiple models, LHC limits hold exclusively for considered models

Hunt for dark matter - interplay with direct detection

► **3- Results:** 90% CL lower limits on interaction xsec between DM and nuclei

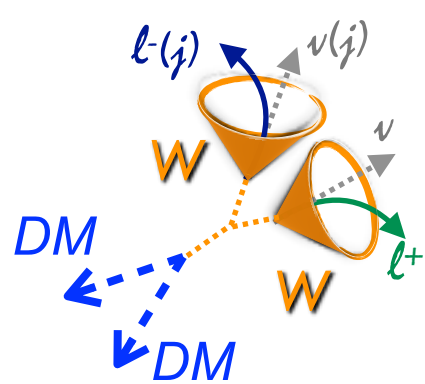


Comparison recommendations [[arXiv:1603.04156](https://arxiv.org/abs/1603.04156)]

- * DM non-relativistic: dominant DM-nuclei interactions spin-independent and spin-dependent scattering
 - * vector/scalar mediator lead to a SI interaction
 - * axial-vector/pseudo-scalar lead to SD interaction
- * Very model dependent
 - * DD bounds may be valid for multiple models, LHC limits hold exclusively for considered models

Hunt for dark matter - DM+WW search

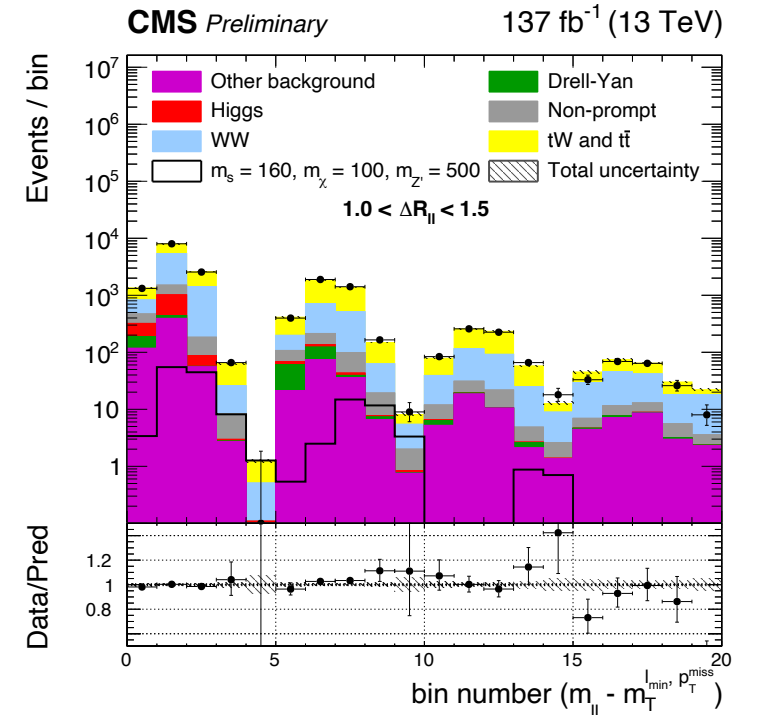
► 1- Selection: #lep categorization



- 1l** * 1 lep, ≥ 2 jets ($m_{jj} \sim W$ mass)
- * 0 b-tagged jets, MET > 60 GeV
- * MVA to recover sensitivity
- 2l** * 1 opposite-charge $\ell\ell$ pair
- * 0 b-tagged jets, MET > 20 GeV
- * events categ. on dark-H boost

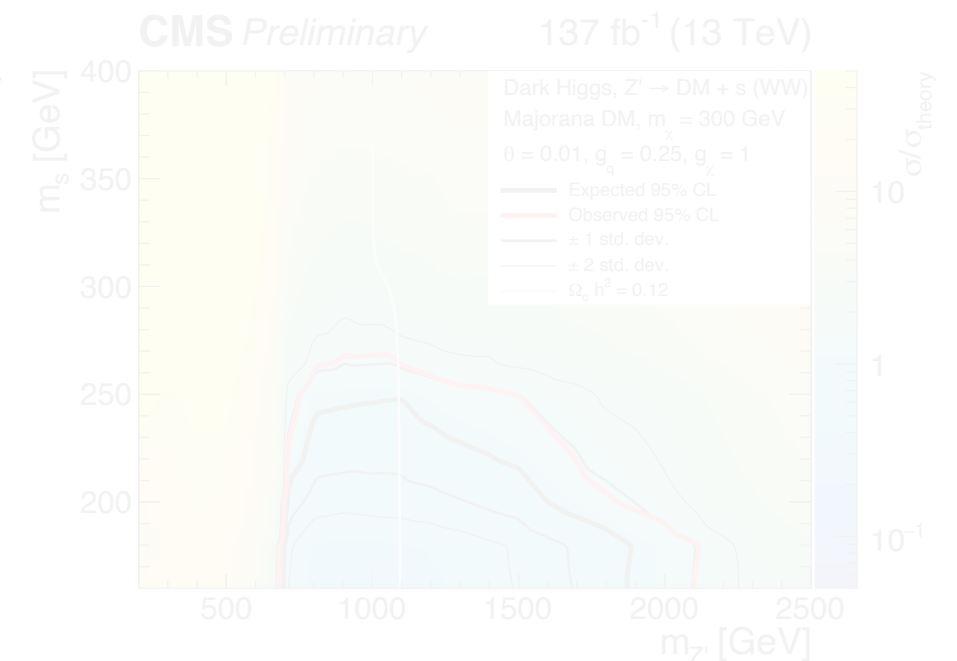
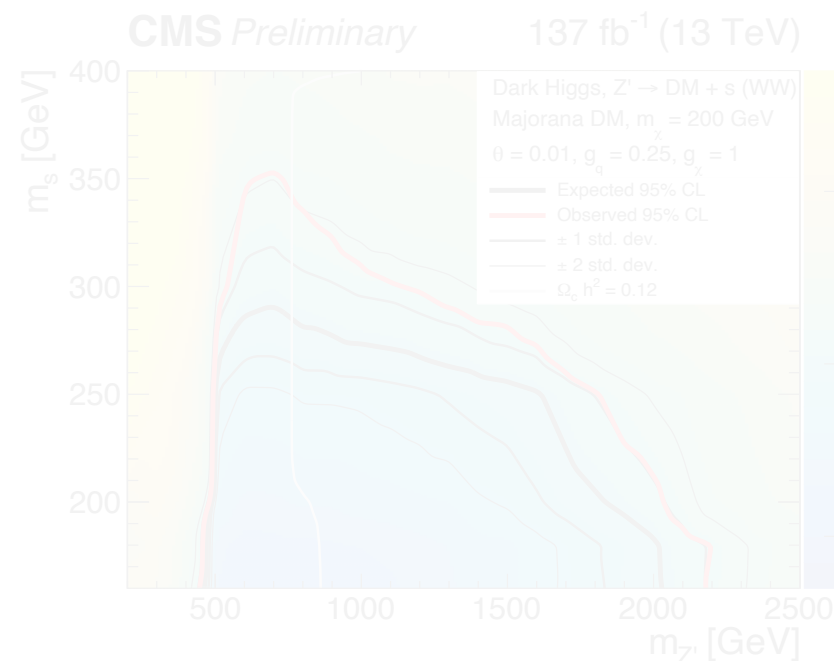
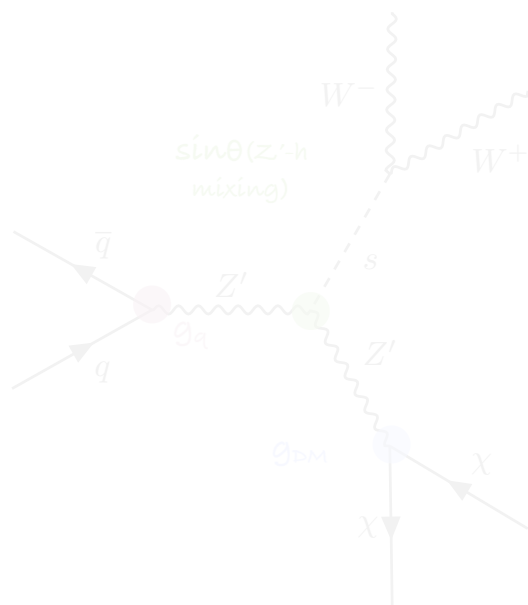
► 2- Bkg:

- tt , tW , W/Z +jets, WW
main bkg, from CRs



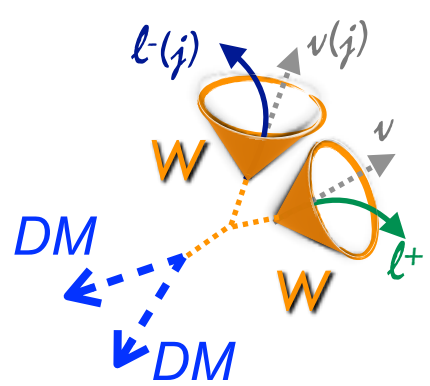
► 3- Results: upper limits at 95% CL on cross section

Dark-Higgs model: dark Higgs (s) can be lighter than DM, observed relic can be attained



Hunt for dark matter - DM+WW search

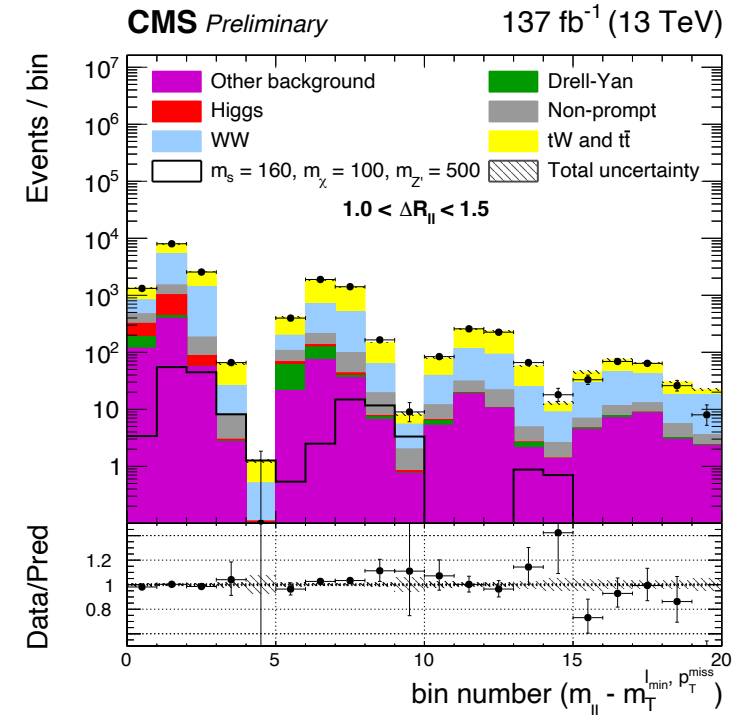
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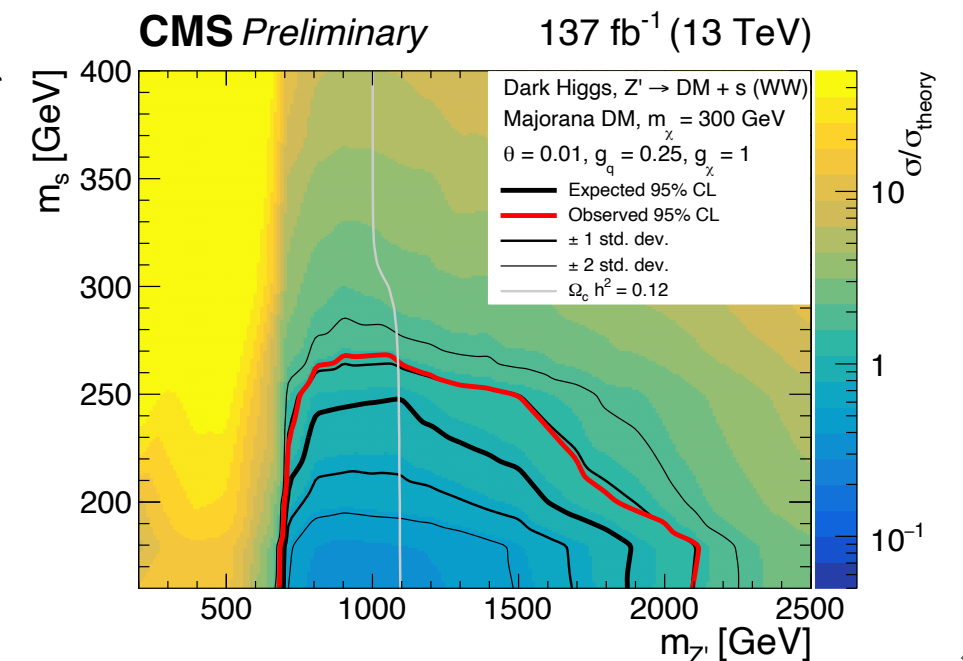
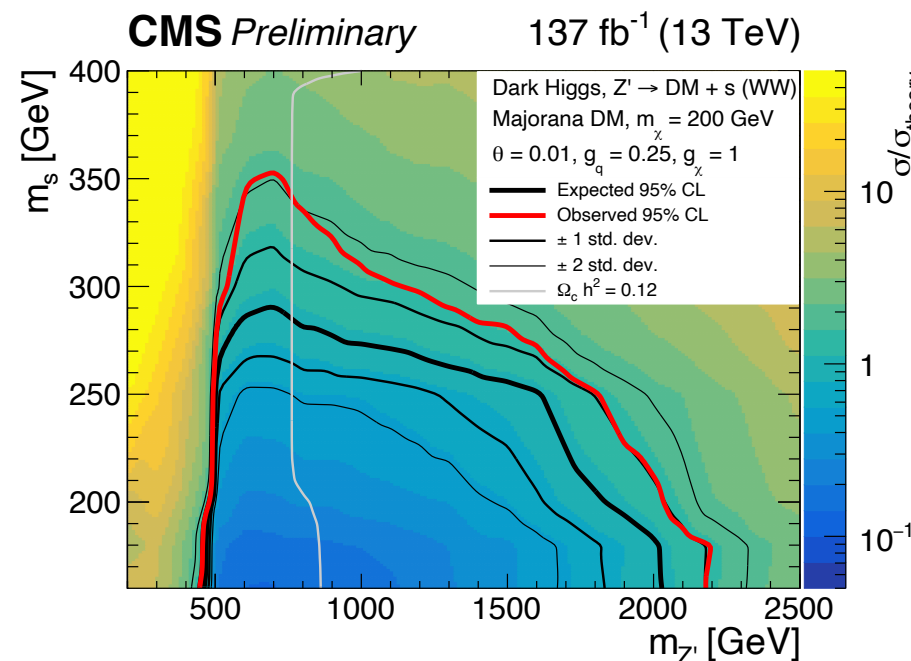
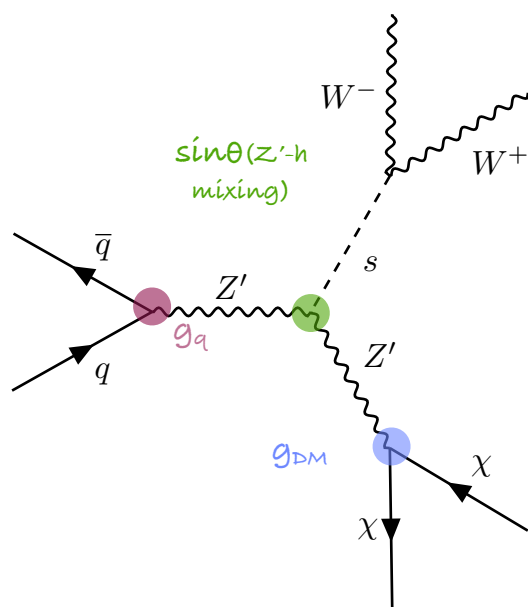
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Hunt for dark matter - Higgs can be a portal?

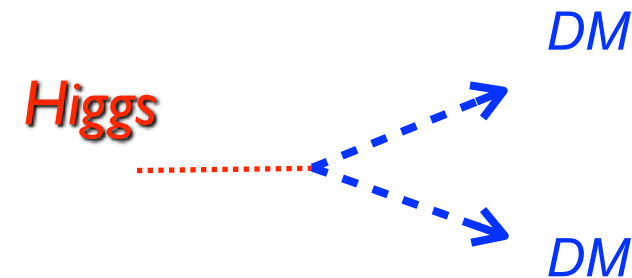
ATLAS: [PLB842\(2023\)](#)

CMS: [HIG-21-007](#)

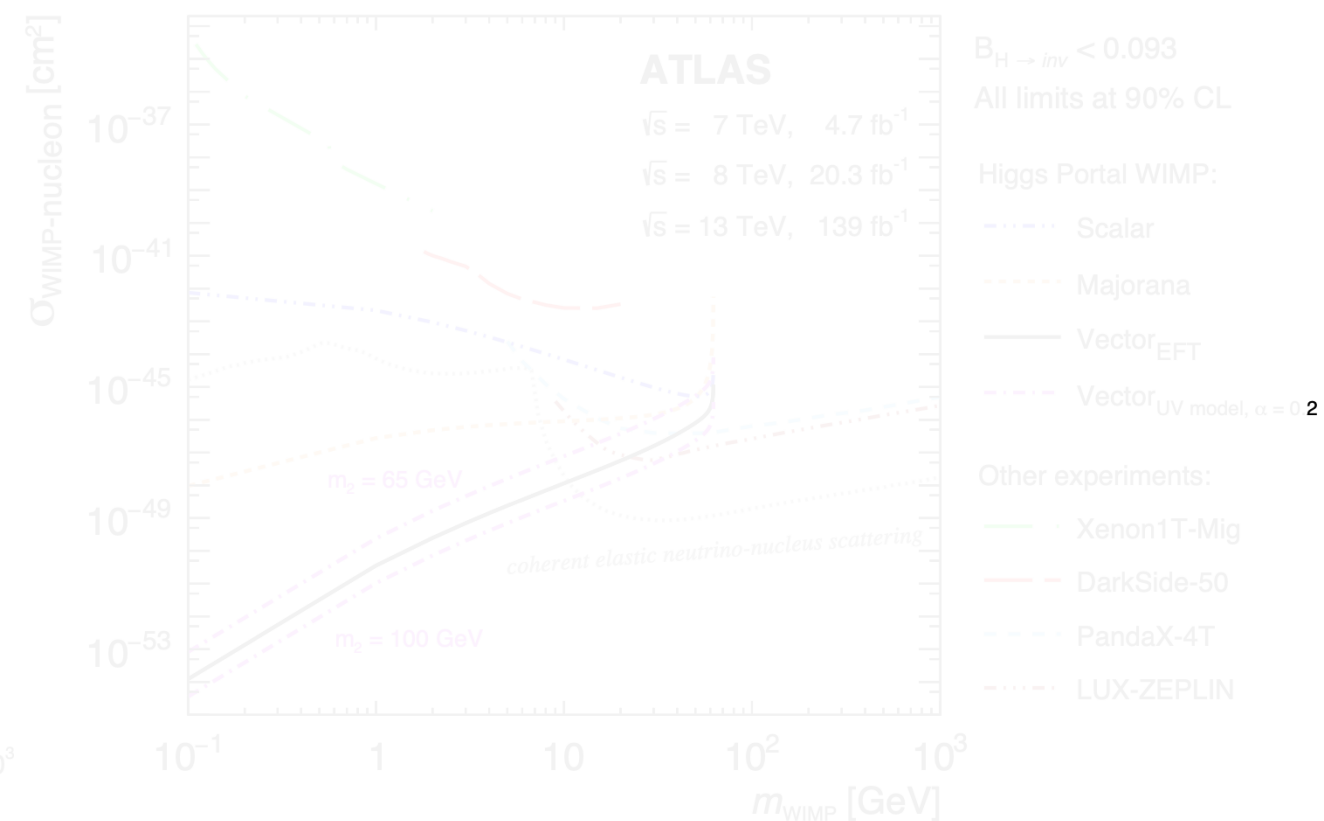
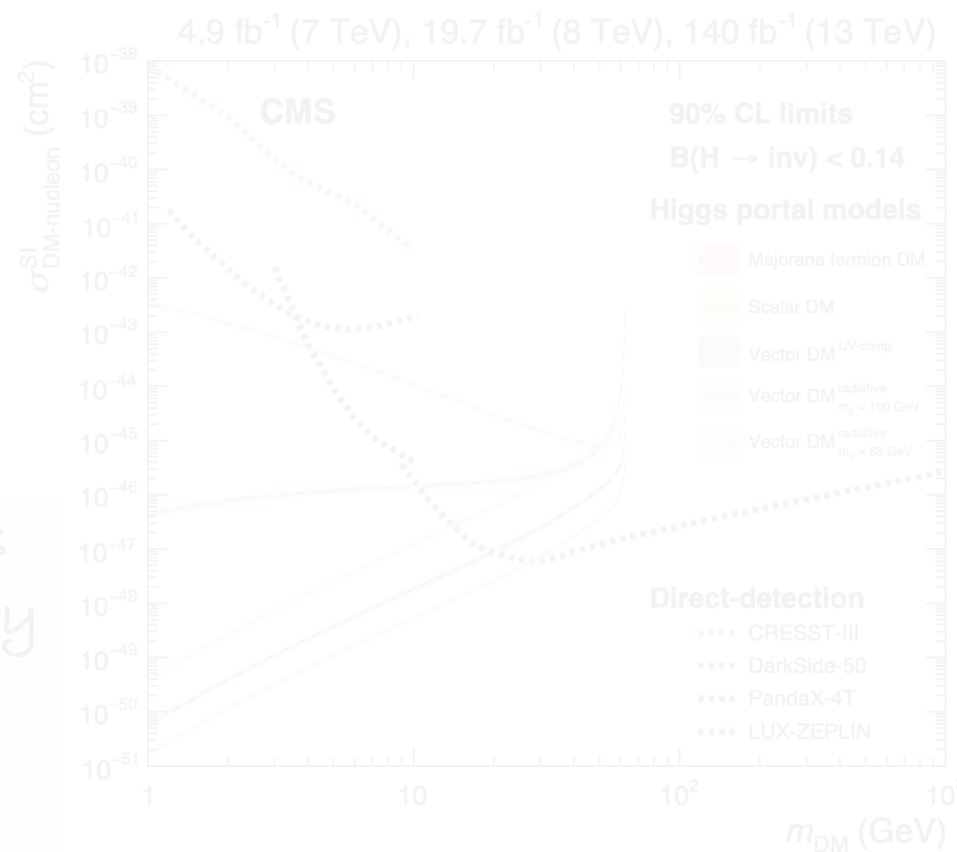
► **DM-SM interactions mediated by Higgs boson:** DM coupling enhance H invisible BR (SM $\sim 0.1\%$)

► **Higgs production as in SM**

- gluon fusion (MET+j)
- associated VH (MET+V), ttH (MET+tt)
- vector-boson fusion (MET+2jets)



► **3- Results:** translated in spin-independent DM-nucleon elastic scattering xsec limit ($m_{DM} < m_H/2$)



LHC searches
complementary
to direct
detection
(under model
assumptions)

$BR(H \rightarrow inv) < 0.15(0.08)$ obs(exp.)

$BR(H \rightarrow inv) < 0.11(0.08)$ obs(exp.)

Hunt for dark matter - Higgs can be a portal?

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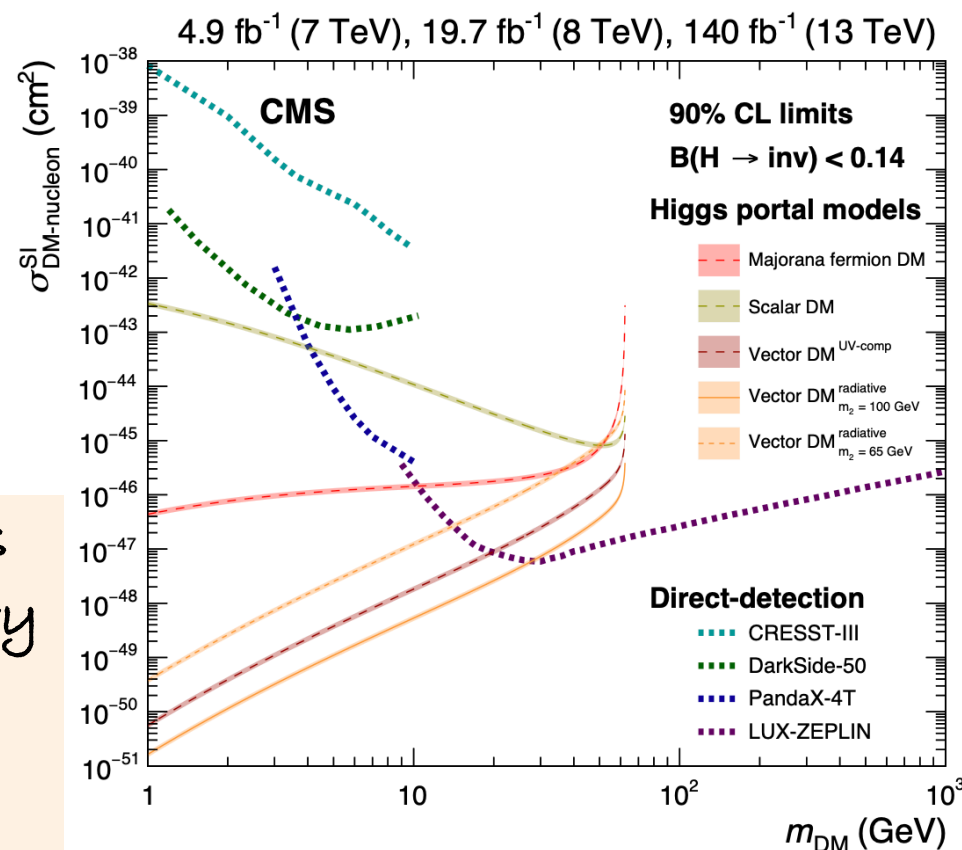
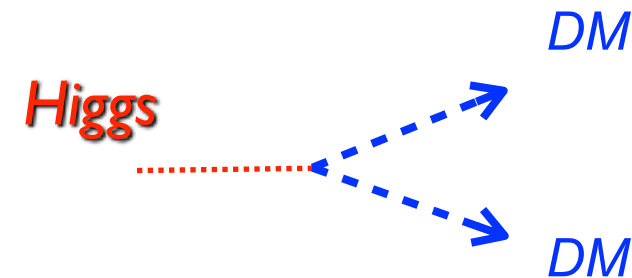
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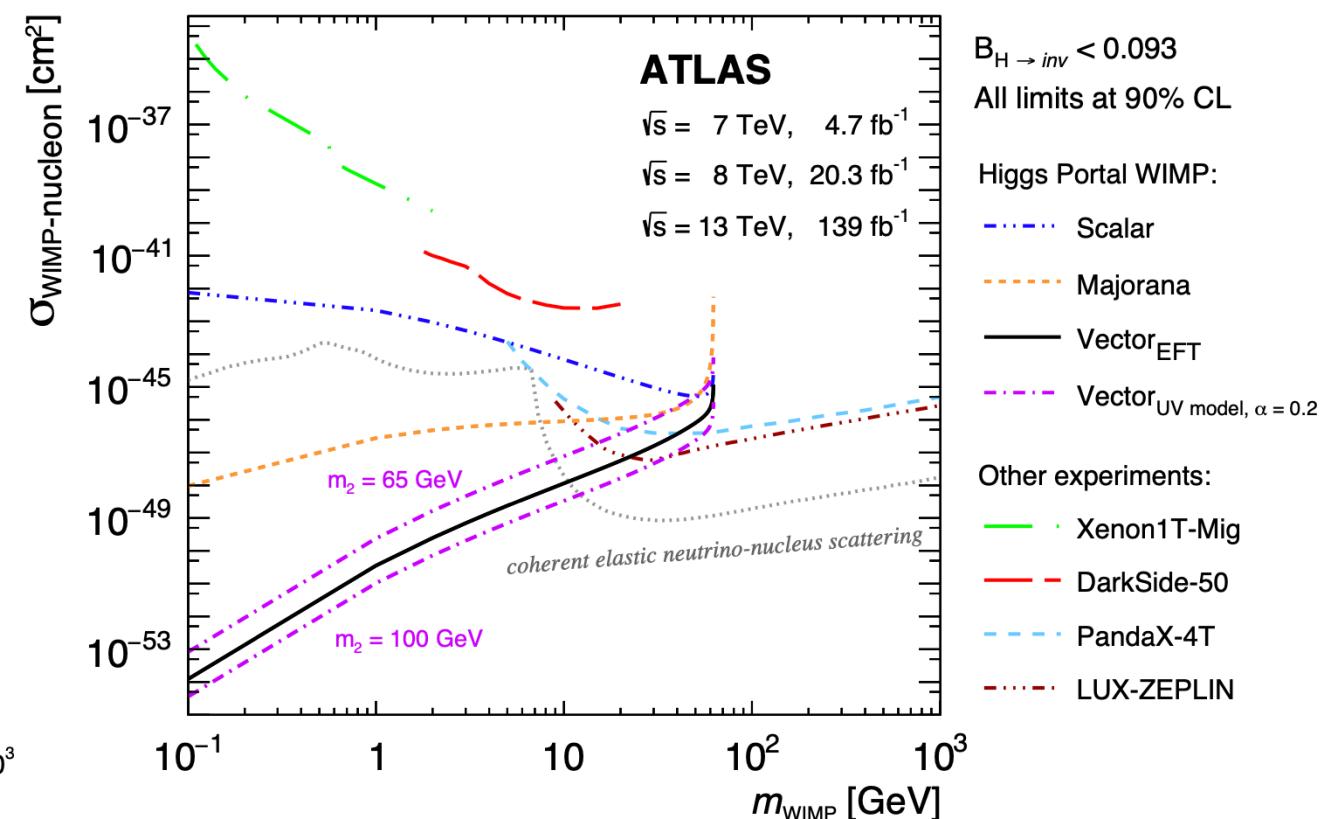
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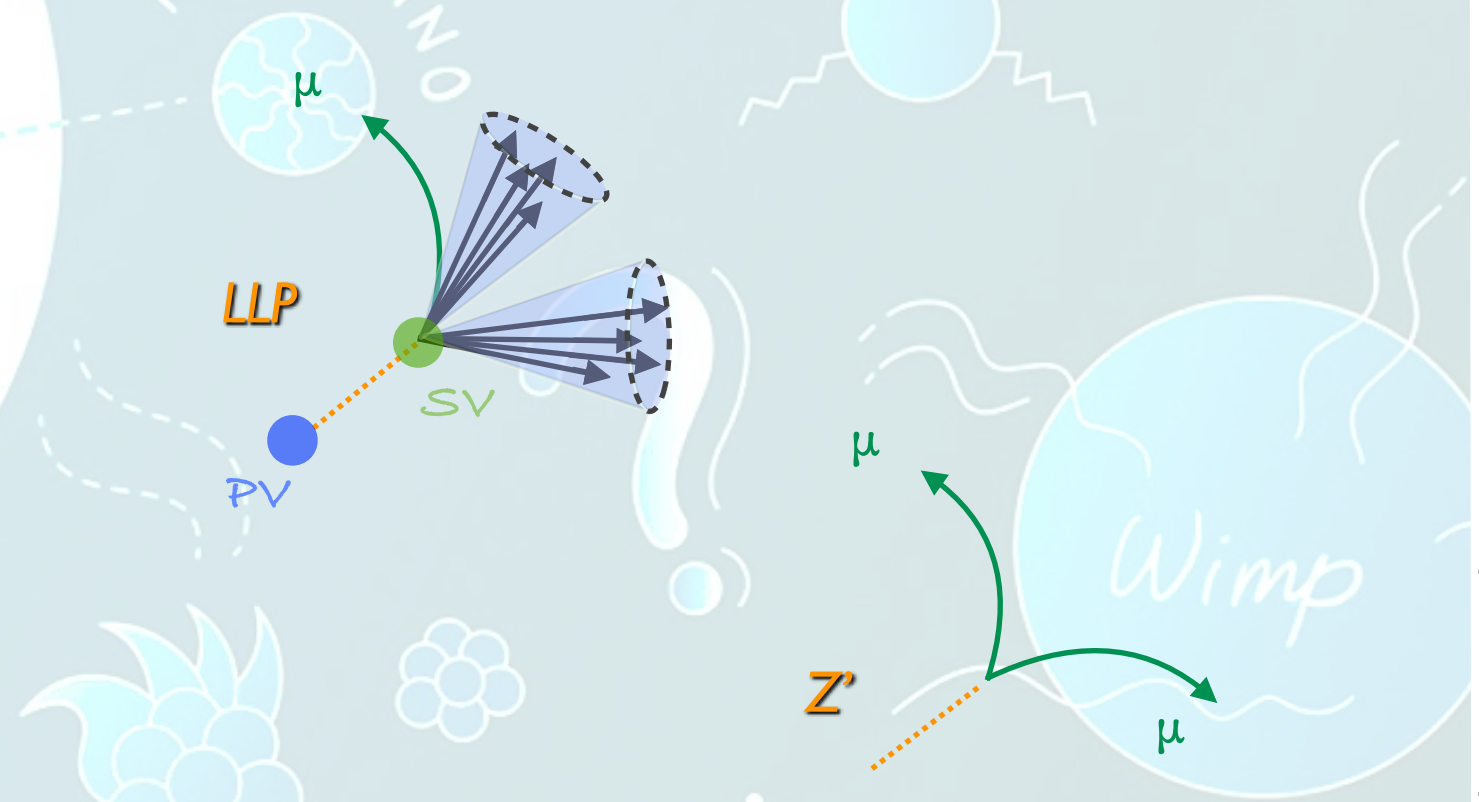
Visible signatures

- massive long-lived particles:

* LHCb: [EPJC373\(2022\)](#)

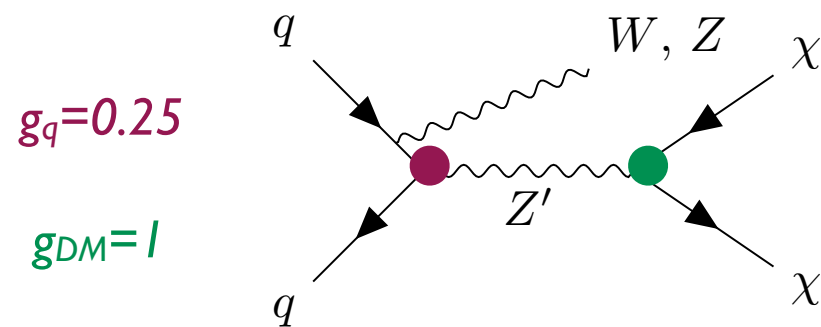
- di-muon resonances:

* CMS: [EXO-21-005](#)



Hunt for dark matter - hints from indirect searches

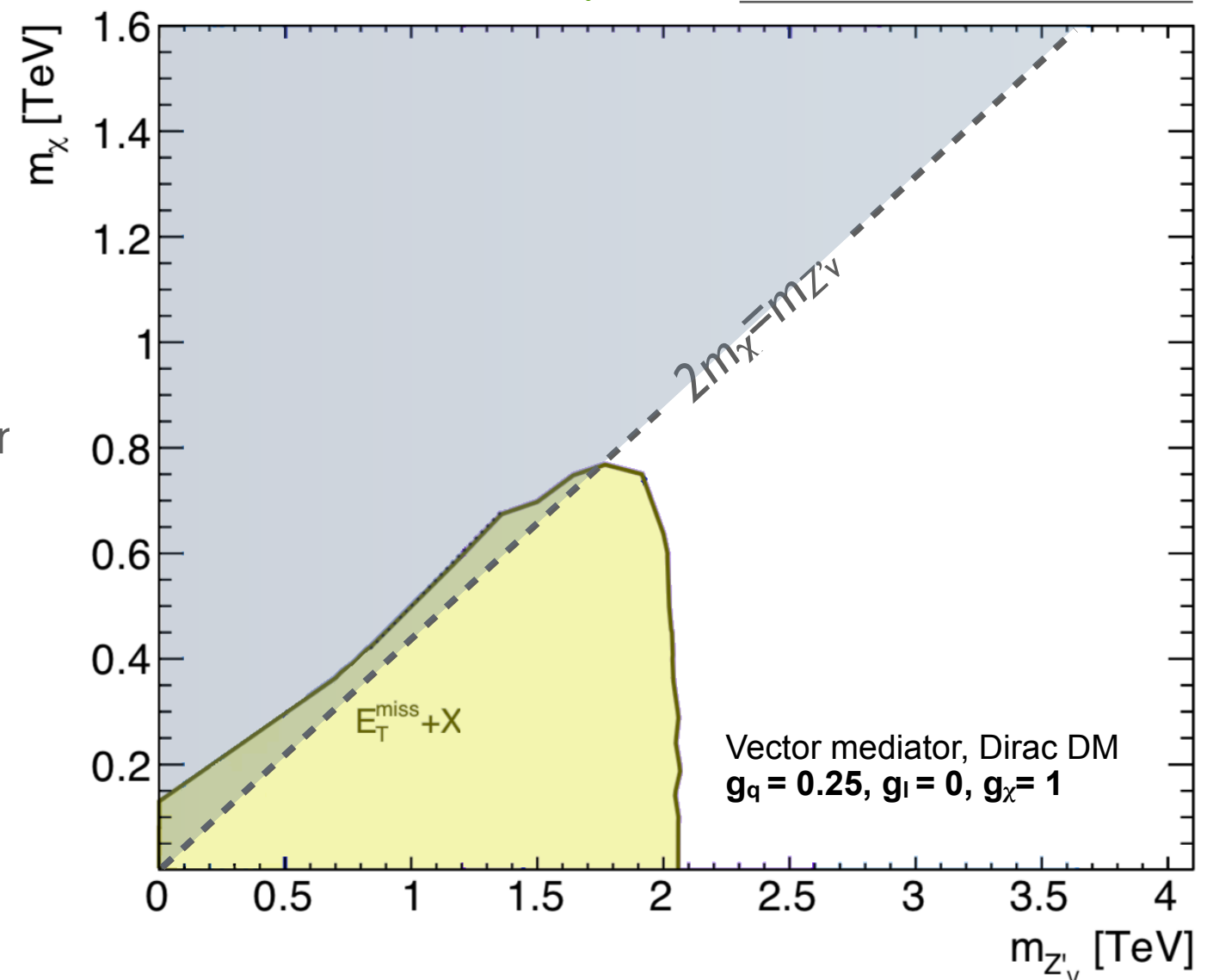
Simplified vector model



► low sensitivity to off-shell region due to strong reduction of production cross-section

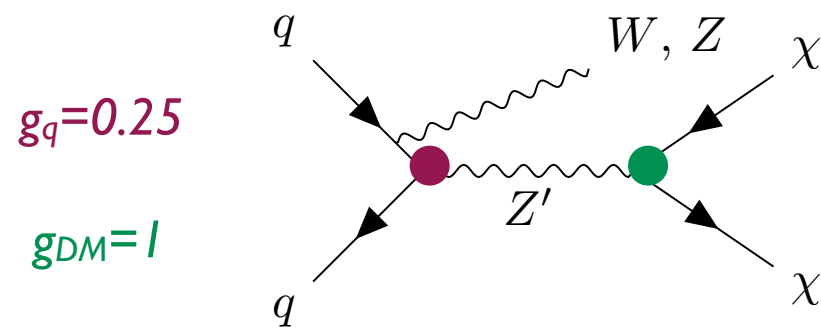
► Can we recover the sensitivity?

* ATLAS:ATL-PHYS-PUB-2021-006



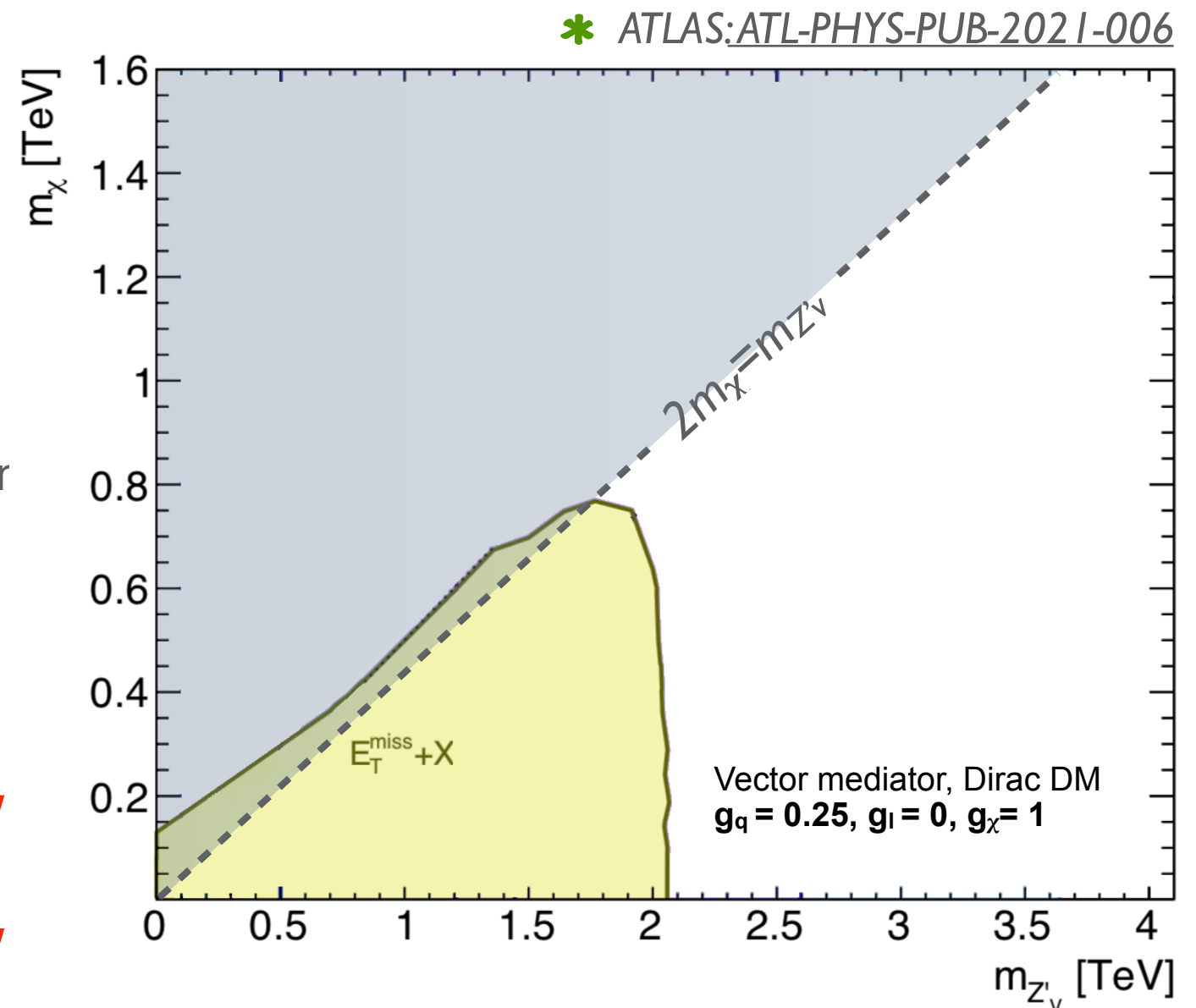
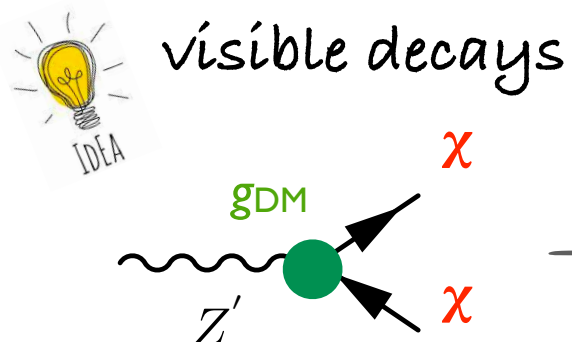
Hunt for dark matter - hints from indirect searches

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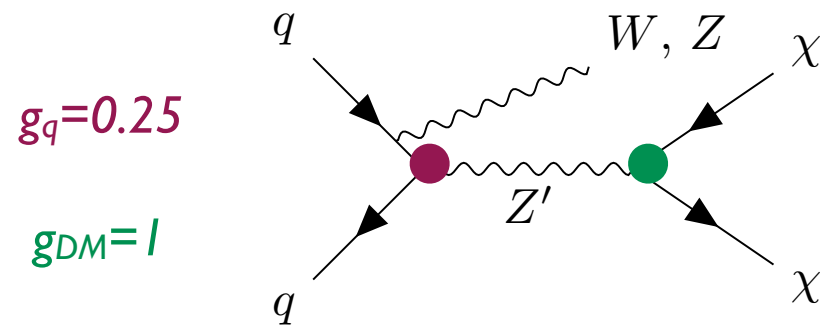
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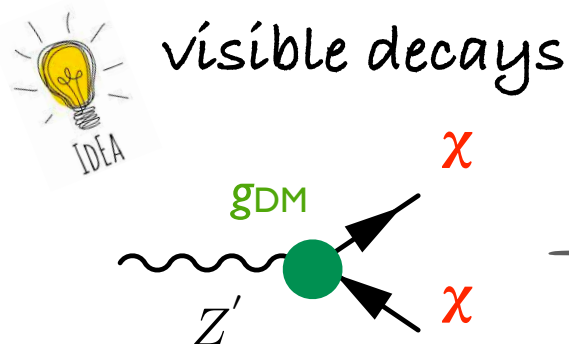
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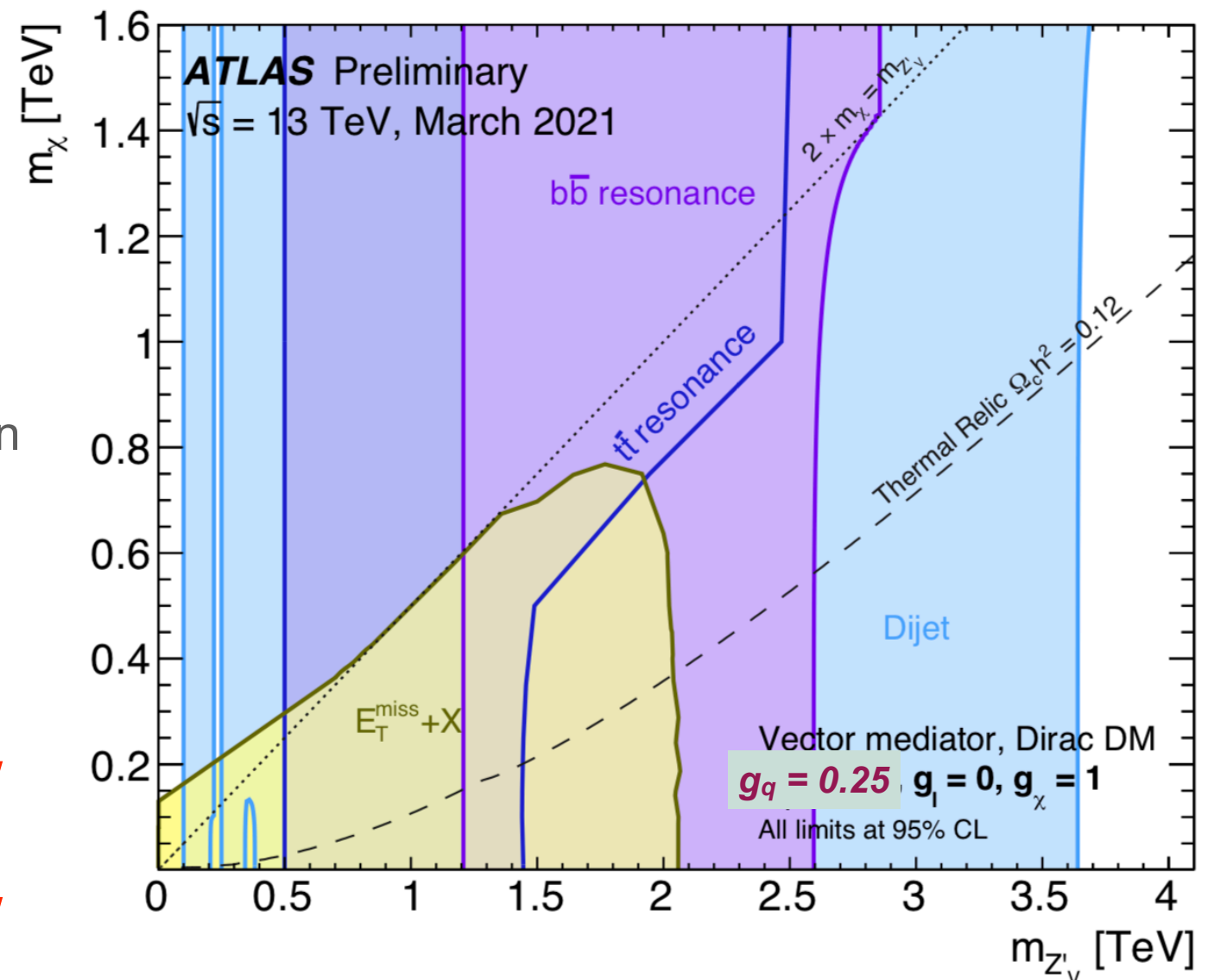
* mediator

$g_q =$

- narrow resonance
- wide resonance

Did we exclude already everything?

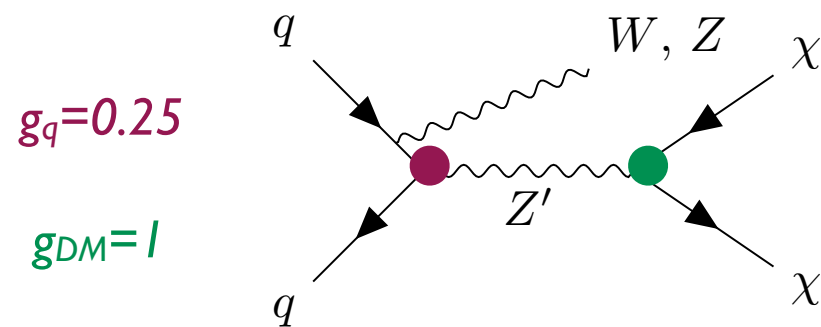
* ATLAS:ATL-PHYS-PUB-2021-006



Dijet	$t\bar{t}$ resonance	$b\bar{b}$ resonance	$E_T^{\text{miss}} + X$
Dijet, 139 fb ⁻¹ JHEP 03 (2020) 145	36.1 fb ⁻¹ EPJC 78 (2018) 565	139 fb ⁻¹ JHEP 03 (2020) 145	$E_T^{\text{miss}} + \text{jet}$, 139 fb ⁻¹ arXiv:2102.10874
Dijet TLA, 29.3 fb ⁻¹ PRL 121 (2018) 081801			$E_T^{\text{miss}} + \gamma$, 139 fb ⁻¹ arXiv:2011.05259
Dijet+ISR, 79.8 fb ⁻¹ PLB 795 (2019) 56			$E_T^{\text{miss}} + V(\text{had})$, 36.1 fb ⁻¹ JHEP 10 (2018) 180
Boosted dijet+ISR, 36.1 fb ⁻¹ PLB 788 (2019) 316			$E_T^{\text{miss}} + Z(\ell\ell)$, 36.1 fb ⁻¹ PLB 776 (2017) 318
Boosted di- b +ISR, 80.5 fb ⁻¹ ATLAS-CONF-2018-052			

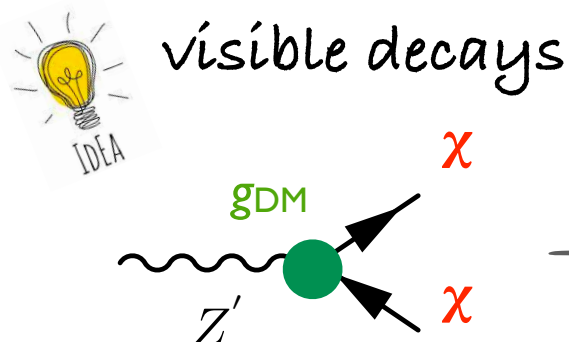
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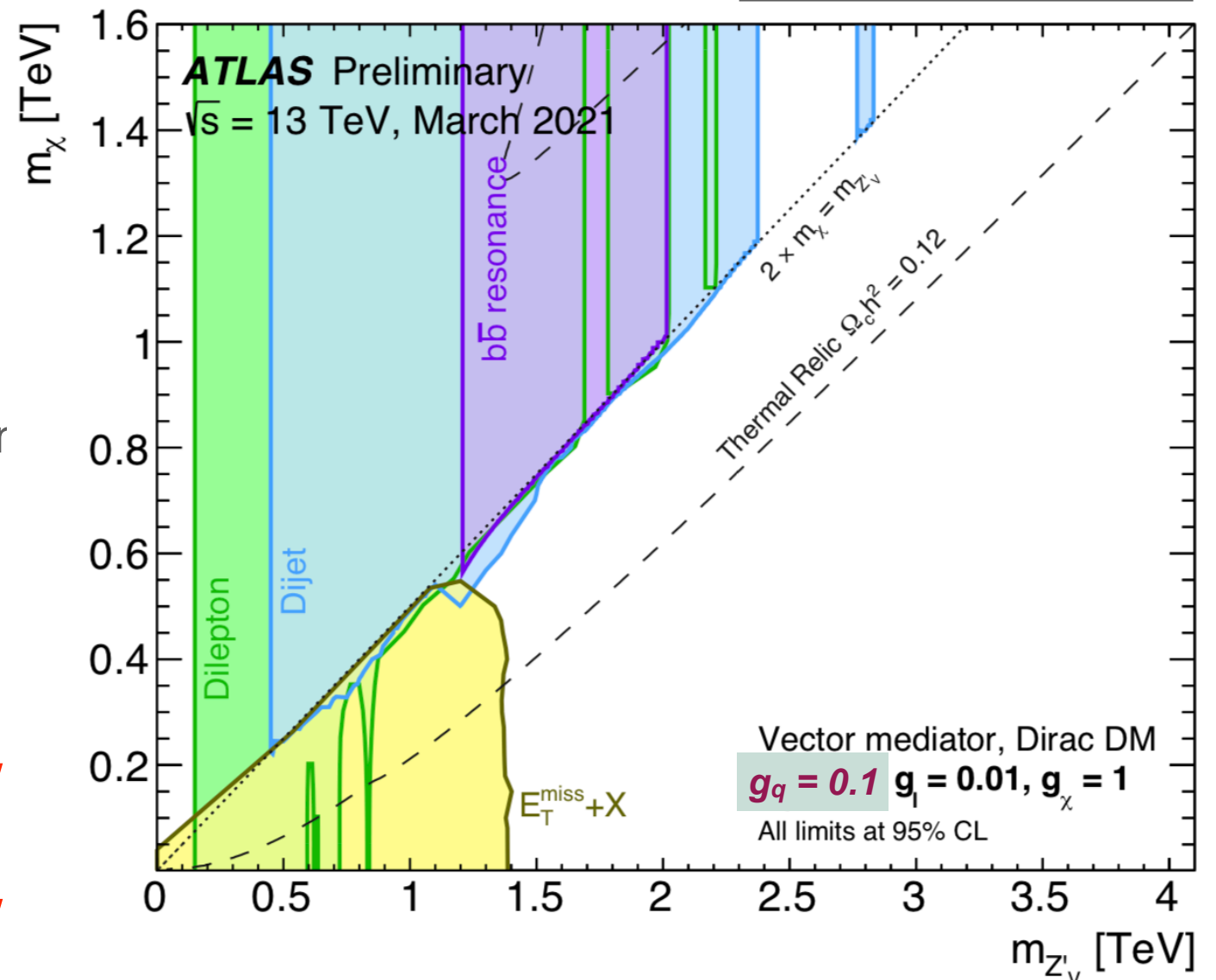


$g_q =$

- * mediator
 - narrow resonance
 - wide resonance
- * interplay changes

Did we exclude
already
everything? NO!

* ATLAS:ATL-PHYS-PUB-2021-006



Dilepton
139 fb⁻¹
PLB 796 (2019) 68

Dijet
Dijet, 139 fb⁻¹
JHEP 03 (2020) 145
Dijet TLA, 29.3 fb⁻¹
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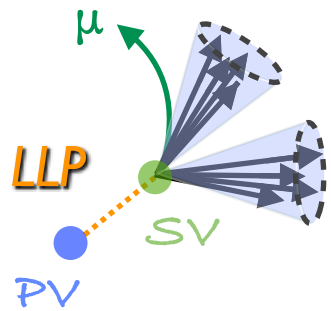
$b\bar{b}$ resonance
139 fb⁻¹
JHEP 03 (2020) 145

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Hunt for dark matter - LLPs $\rightarrow \mu + \text{jets}$ search

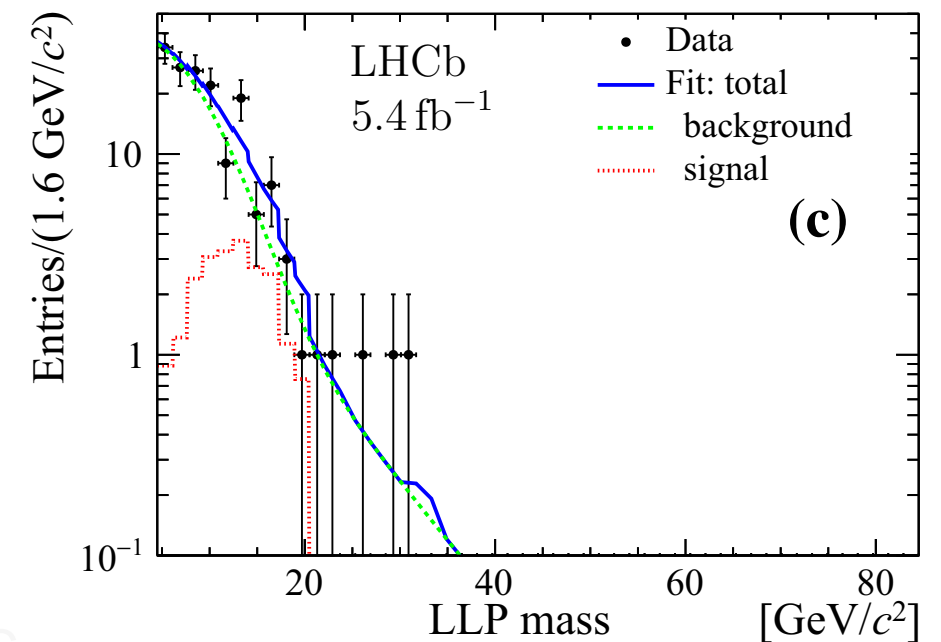
► 1- Selection:

- * 1 displaced vertex from any PV containing 1μ
- * ≥ 1 LLP candidates : ≥ 3 tracks (1μ) with inv mass above B resonances
- * MVA techniques to recover sensitivity



► 2- Bkg:

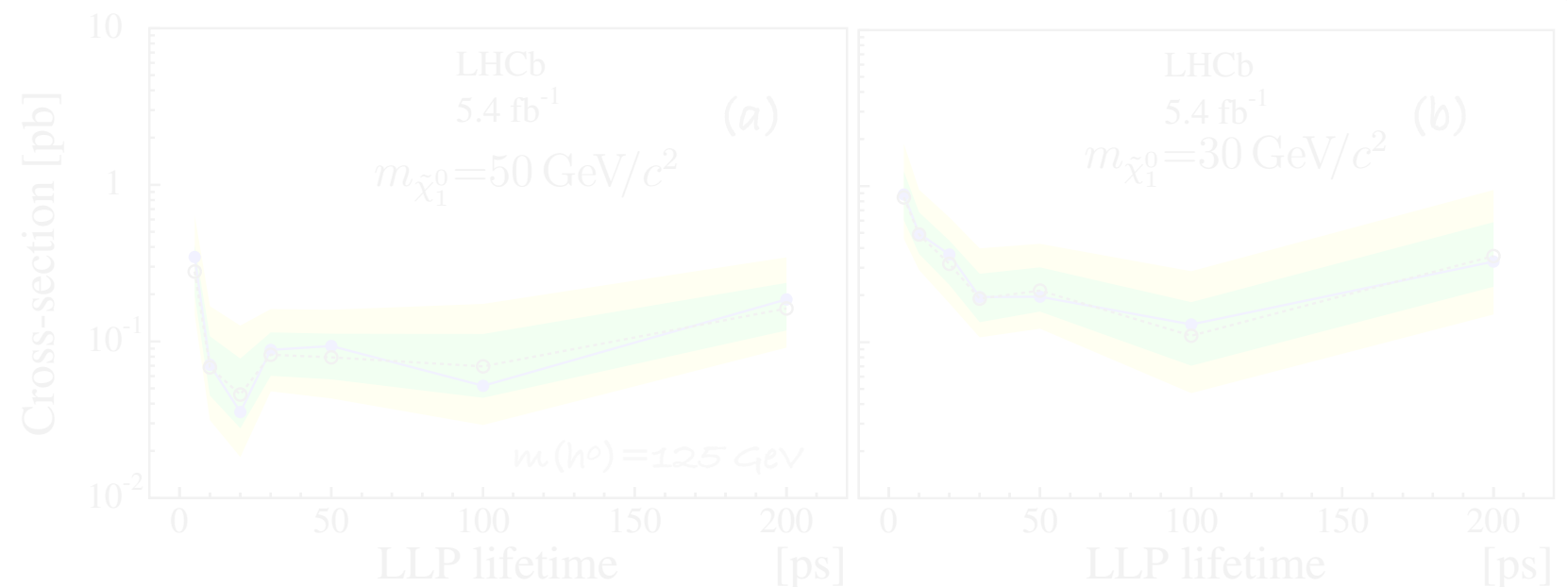
- *bb processes and material interactions*



► 3- Results: combined fit to LLP reconstructed mass in SRs and C Upper limits at 95% CL on cross section

Massive long-lived particles (LLP)

- (a) Higgs-like particle h^0 produced by ggF, decays into two LLPs
- (b) direct LLP production from quark interactions

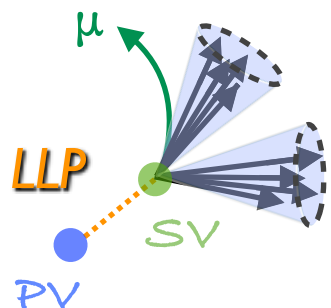


heavier particles \rightarrow lower lifetime/boost

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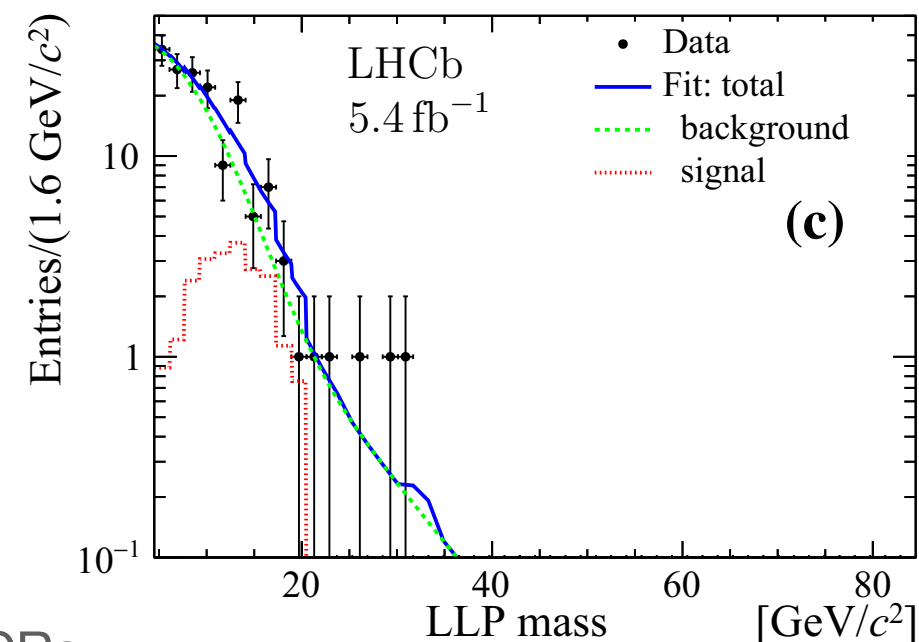
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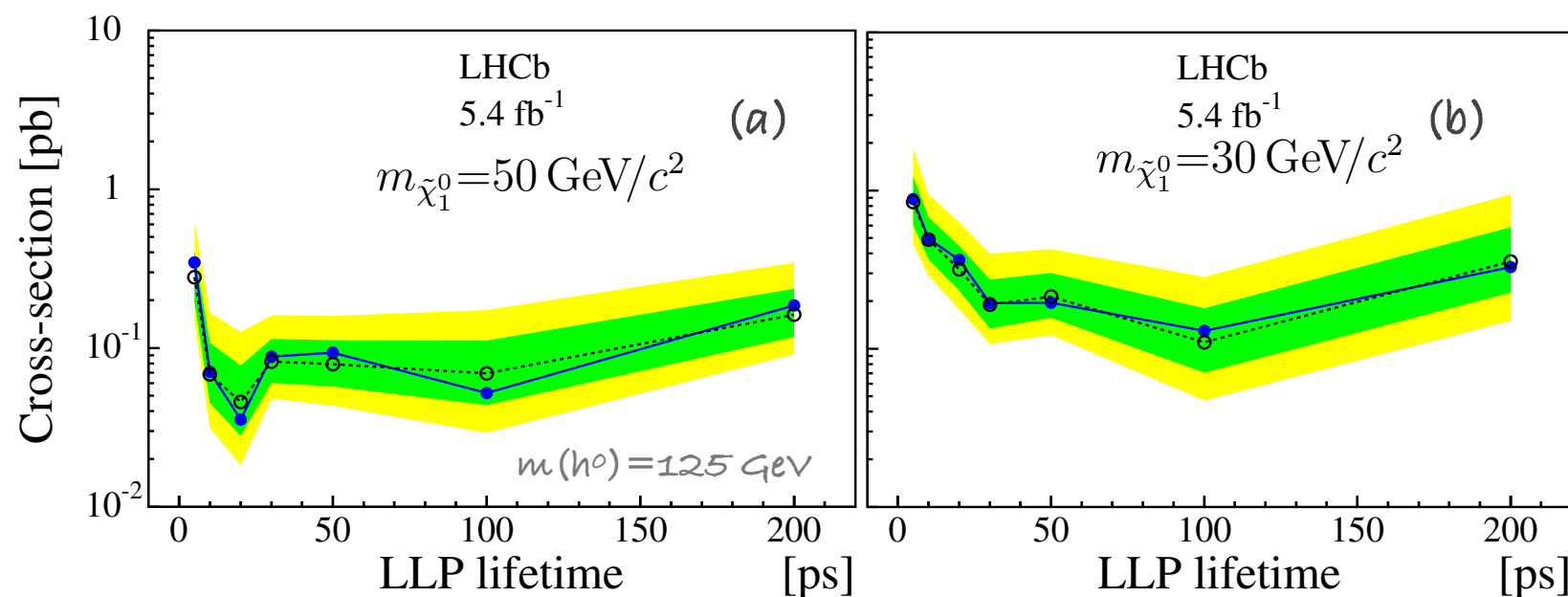
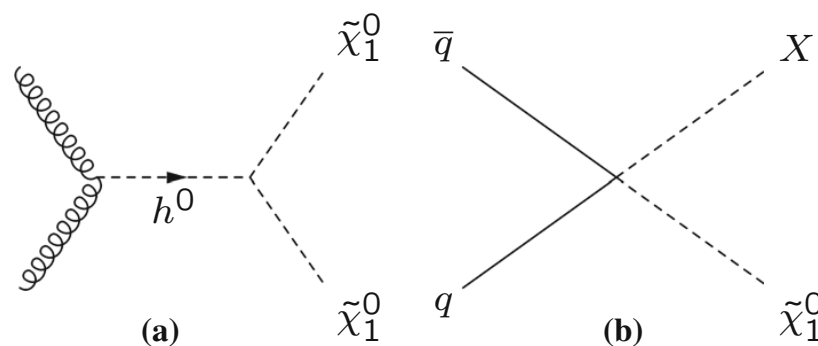
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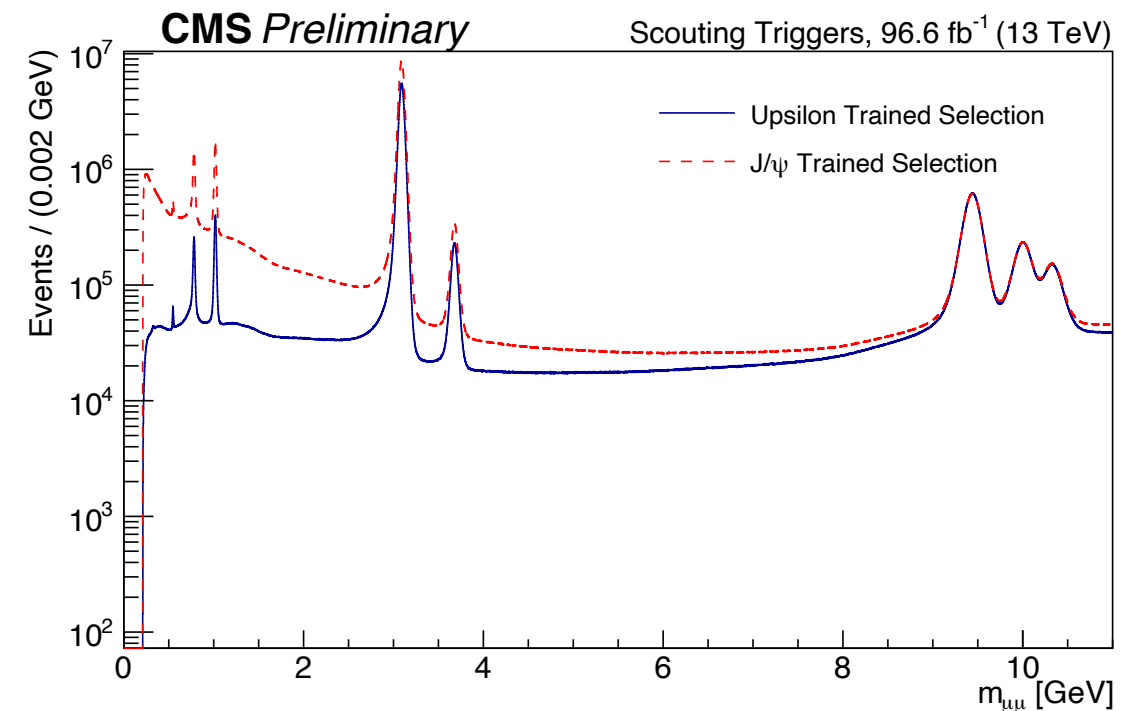
Hunt for dark matter - di-muon low-mass resonances

1 - Selection: resonance as peak wrt SM invariant mass

- * 1 opp.-sign μ pair, categories on $p_T(\mu\mu)$
 - * dedicated $\mu\mu$ trigger: low p_T thresholds, high rate, retain only 4-momentum, isolation, track information
 - * muon identification based on MVA techniques

2 - Bkg:

- known resonances, D meson decays to kaons (from CR)

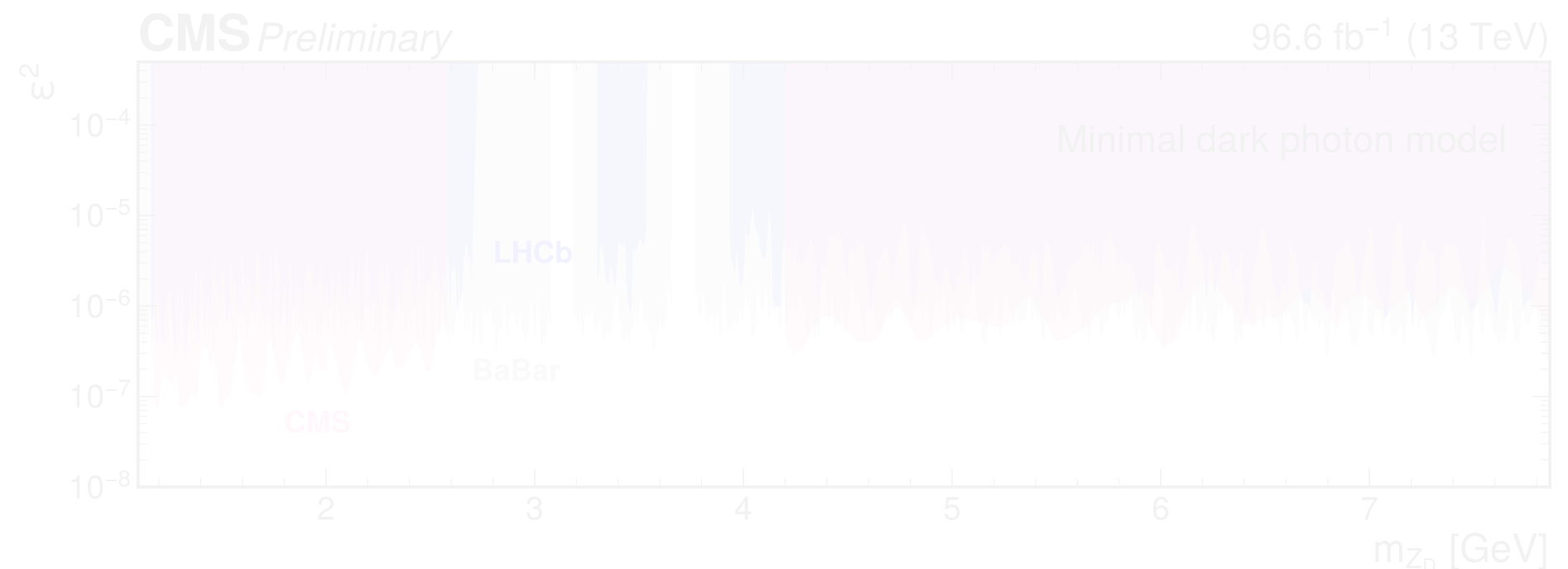


3- Results: fit to $\mu\mu$ invariant mass, upper limits at 90% CL on mixing coefficient ϵ

Dark photon: dark/SM sectors interaction through dark photon Z_D , with kinetic mixing ϵ



Higher sensitivity at low masses



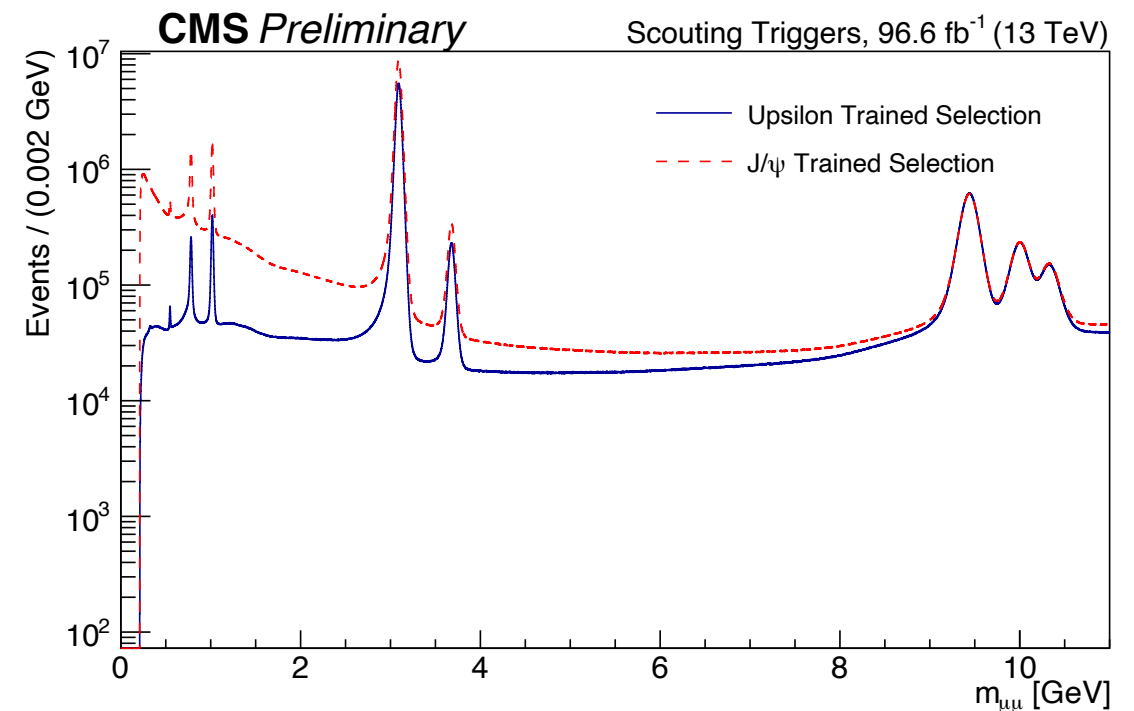
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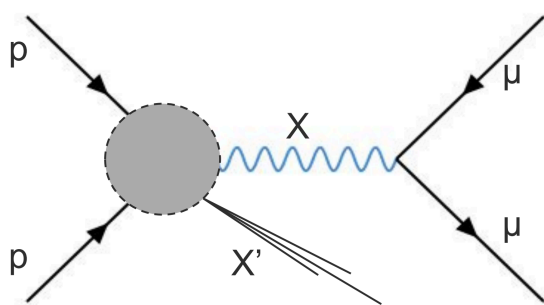
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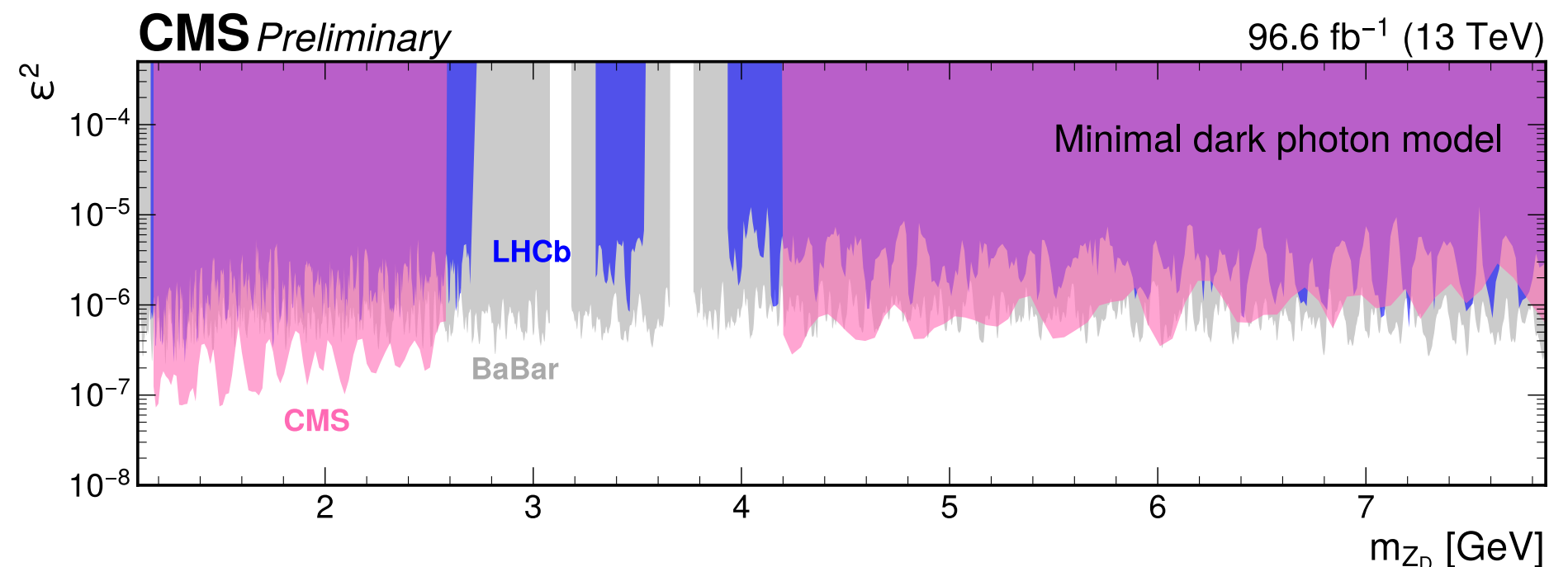


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Higher sensitivity at low masses



... *connecting the dots results*



Hunt for dark matter - complementarity

- * **Luminosity**: determines size of coupling that can be probed
- * **Energy**: determines probed mass range
- * **Invisible/visible**: directly probe DM or mediator

Complementarity essential!

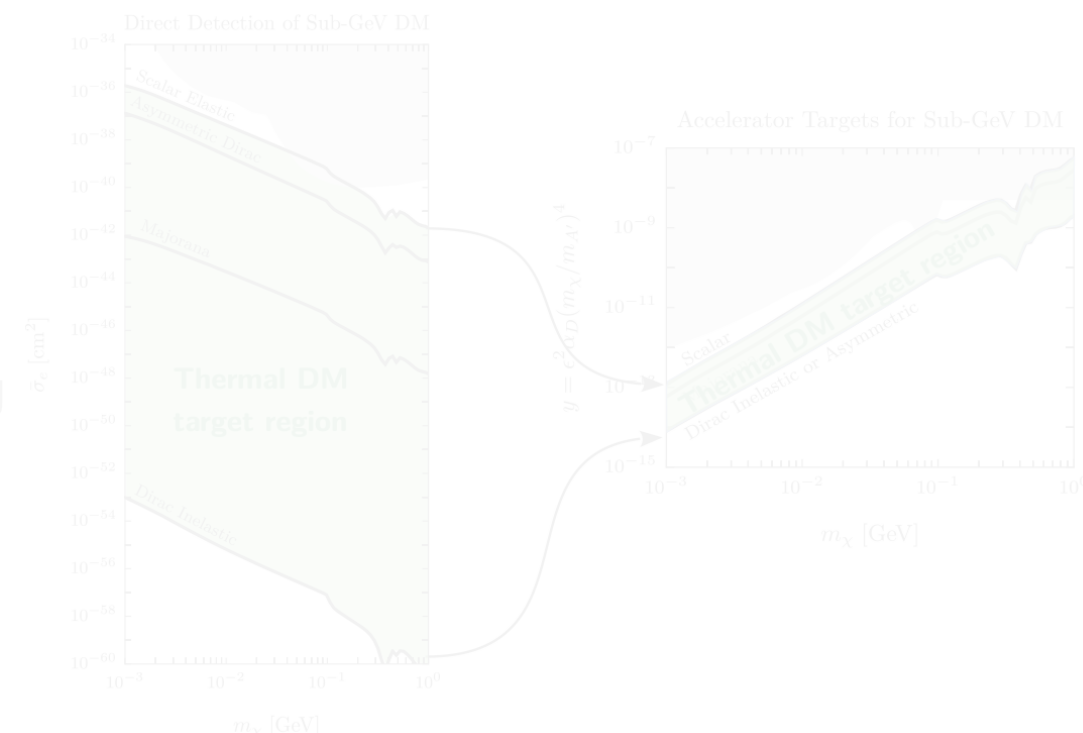
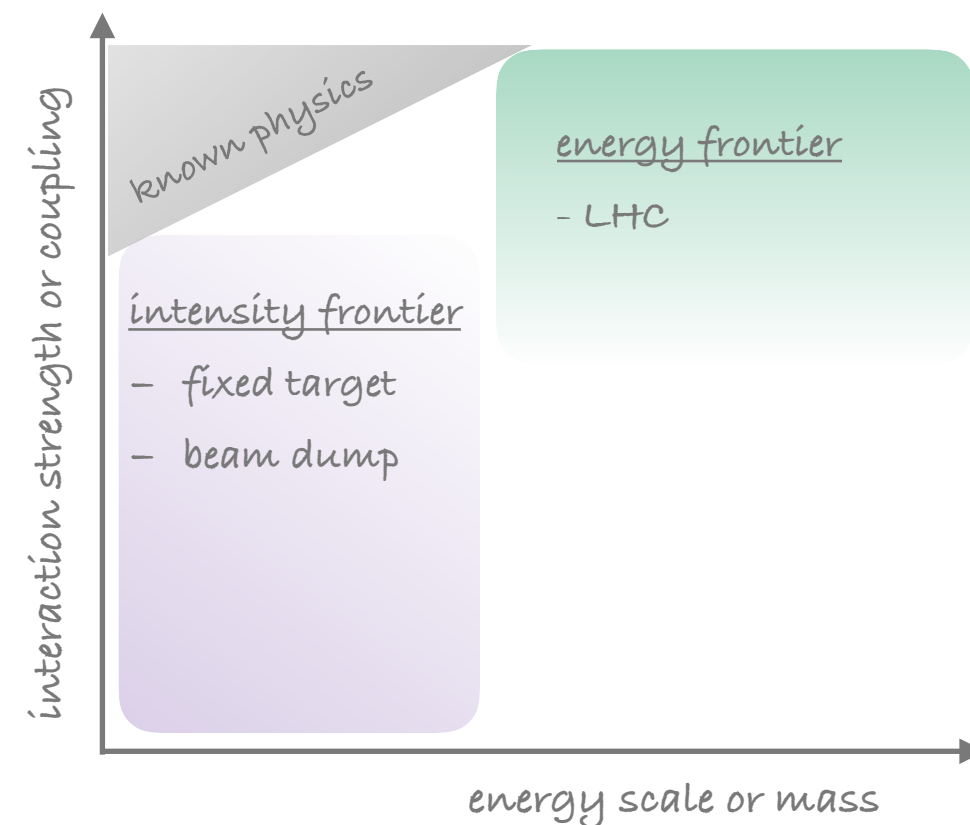
- results often presented in terms of different mediator masses, notation and model assumptions

🧠 Concerted effort to align models and representation to exploit complementary of energy- and intensity-frontiers to discover DM/dark-sector

... also with direct-detection

- accelerator-based directly characterize DM particle properties, explore relativistic DM production
- direct detection explores a combination of DM properties with their cosmological abundance, probe non-relativistic scattering

🧠 Different type of interactions can be suppressed or enhanced based on velocity → complementarity



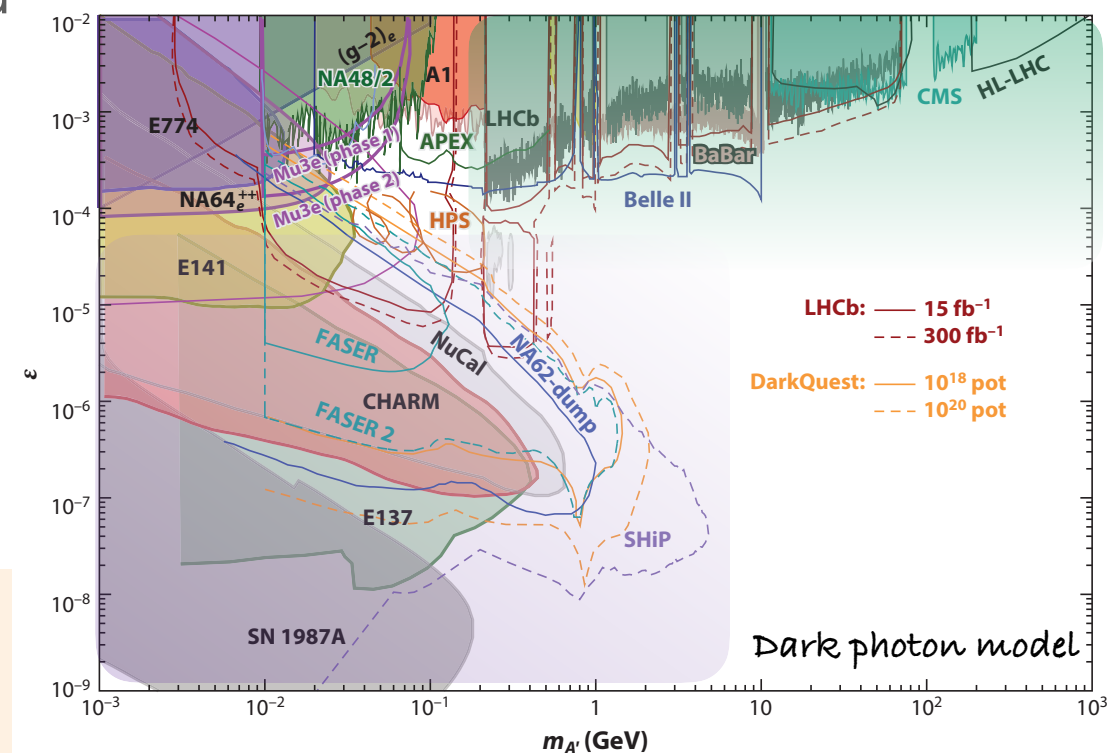
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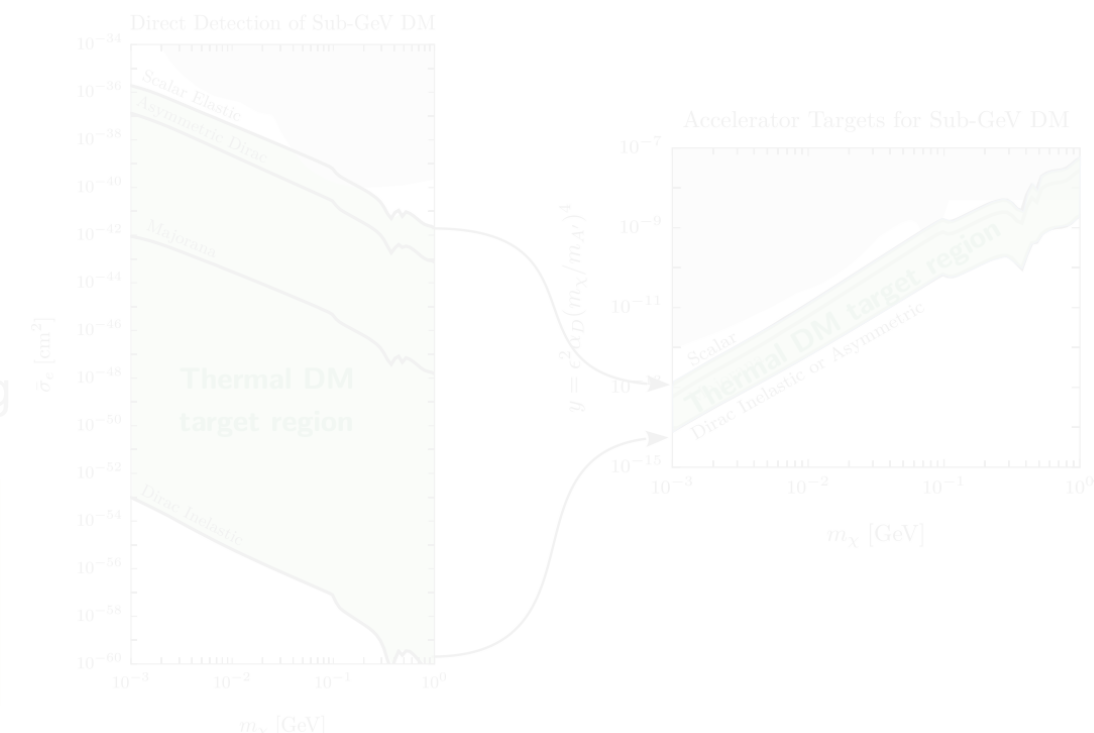
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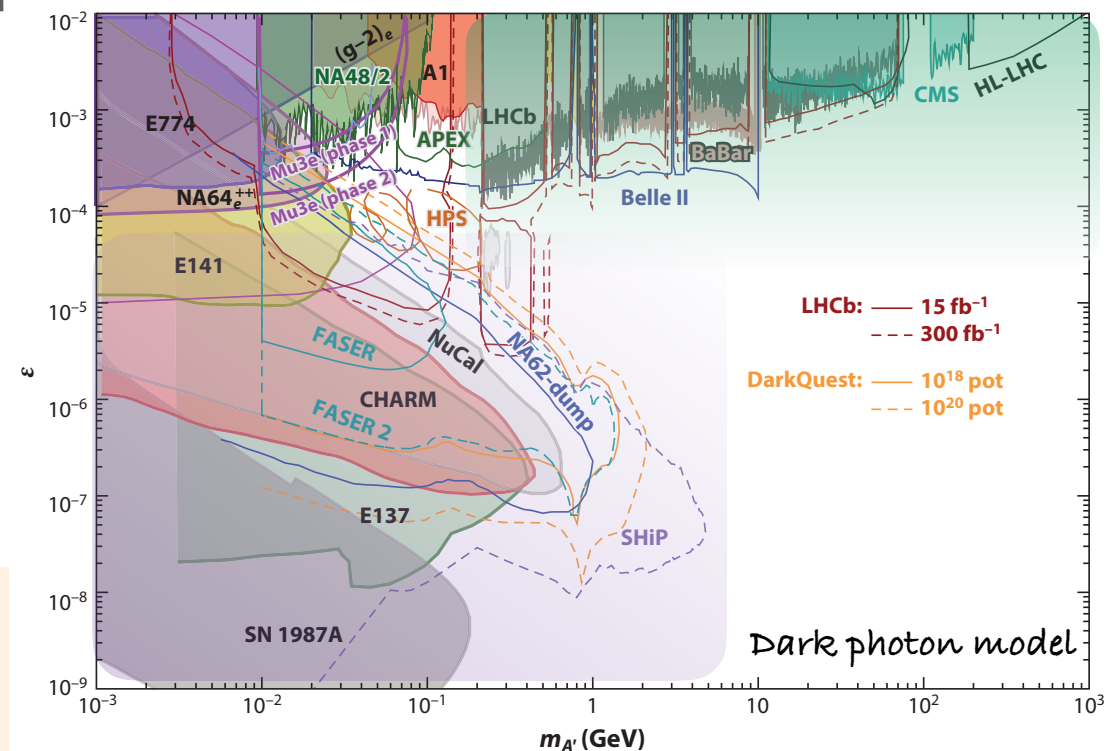
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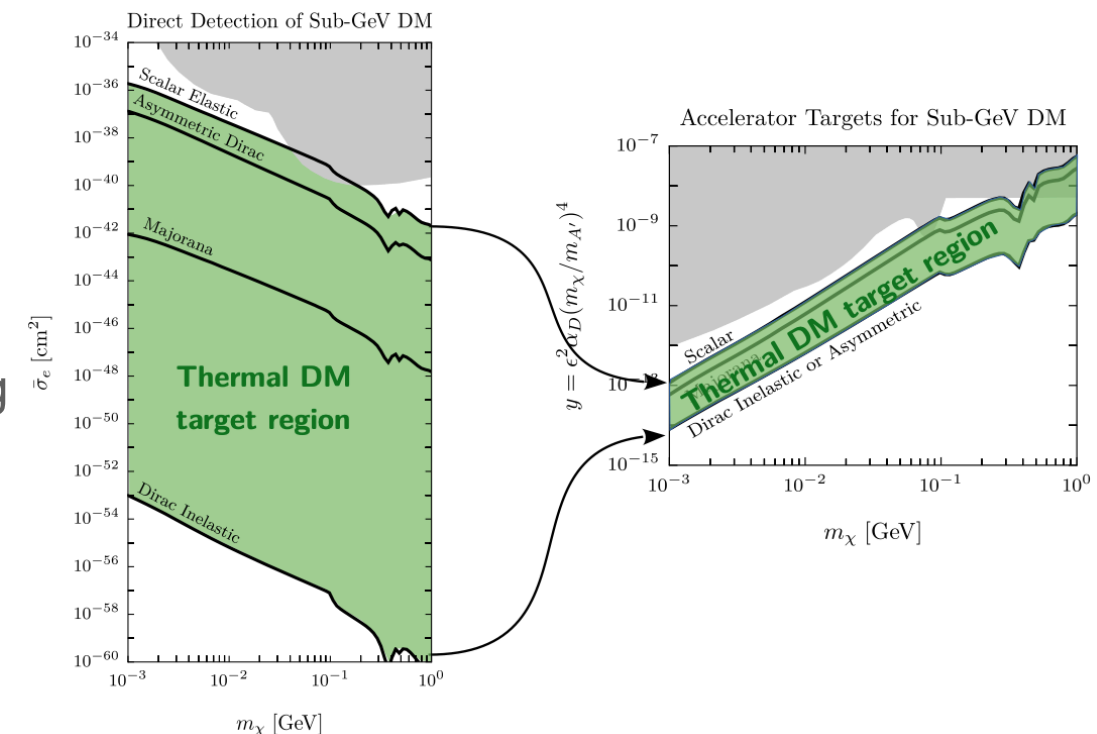
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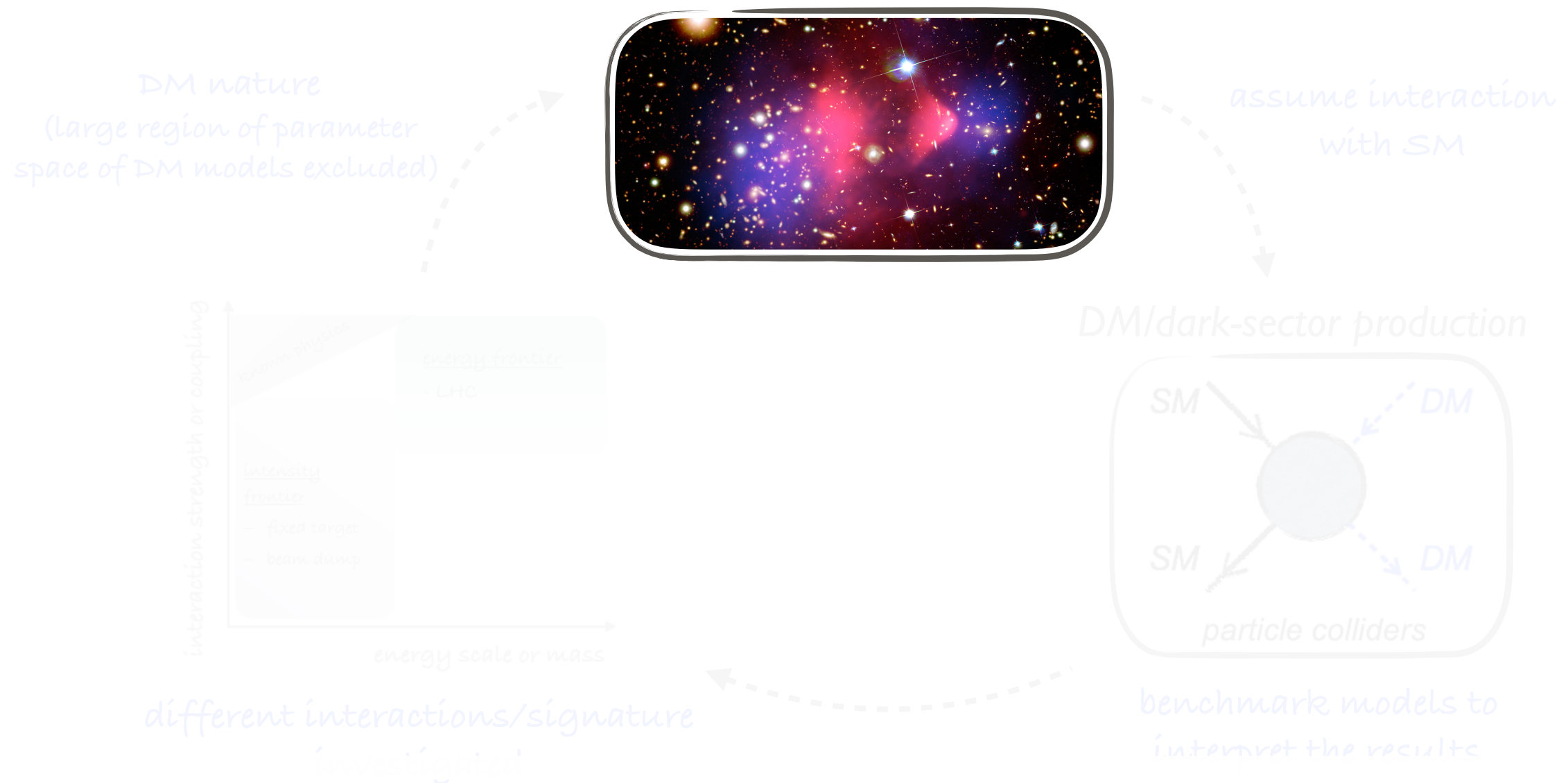
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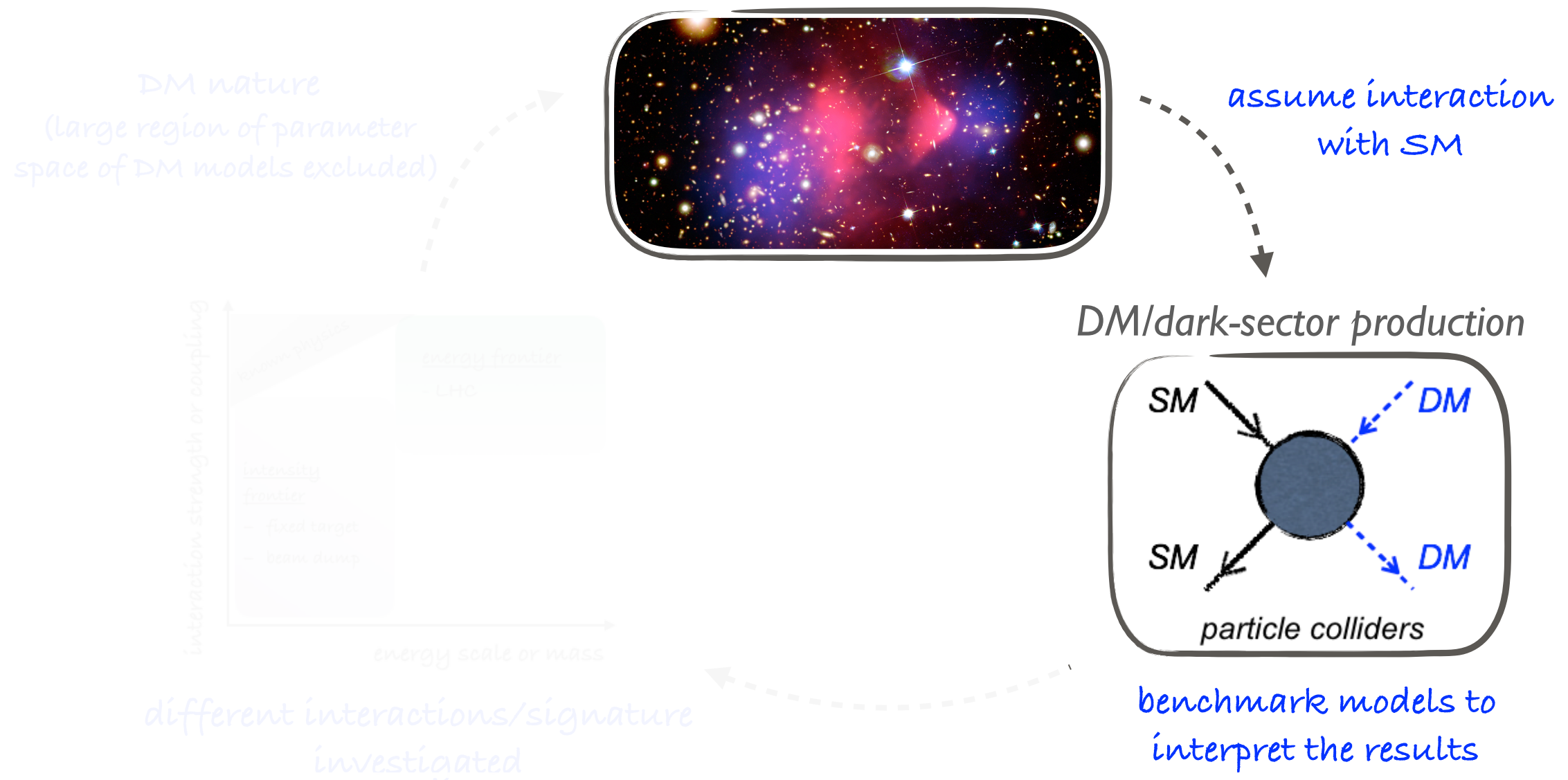


Hunt for dark matter - in a nutshell



- * Rich DM physics program at high-energy and high-intensity experiments
 - essential complementarity among experiments and with non-collider searches
 - inputs from various signatures (mono-X, resonances, $H \rightarrow \text{inv}$, visible, ...)
- * Many new results/experiments expected in the near future!

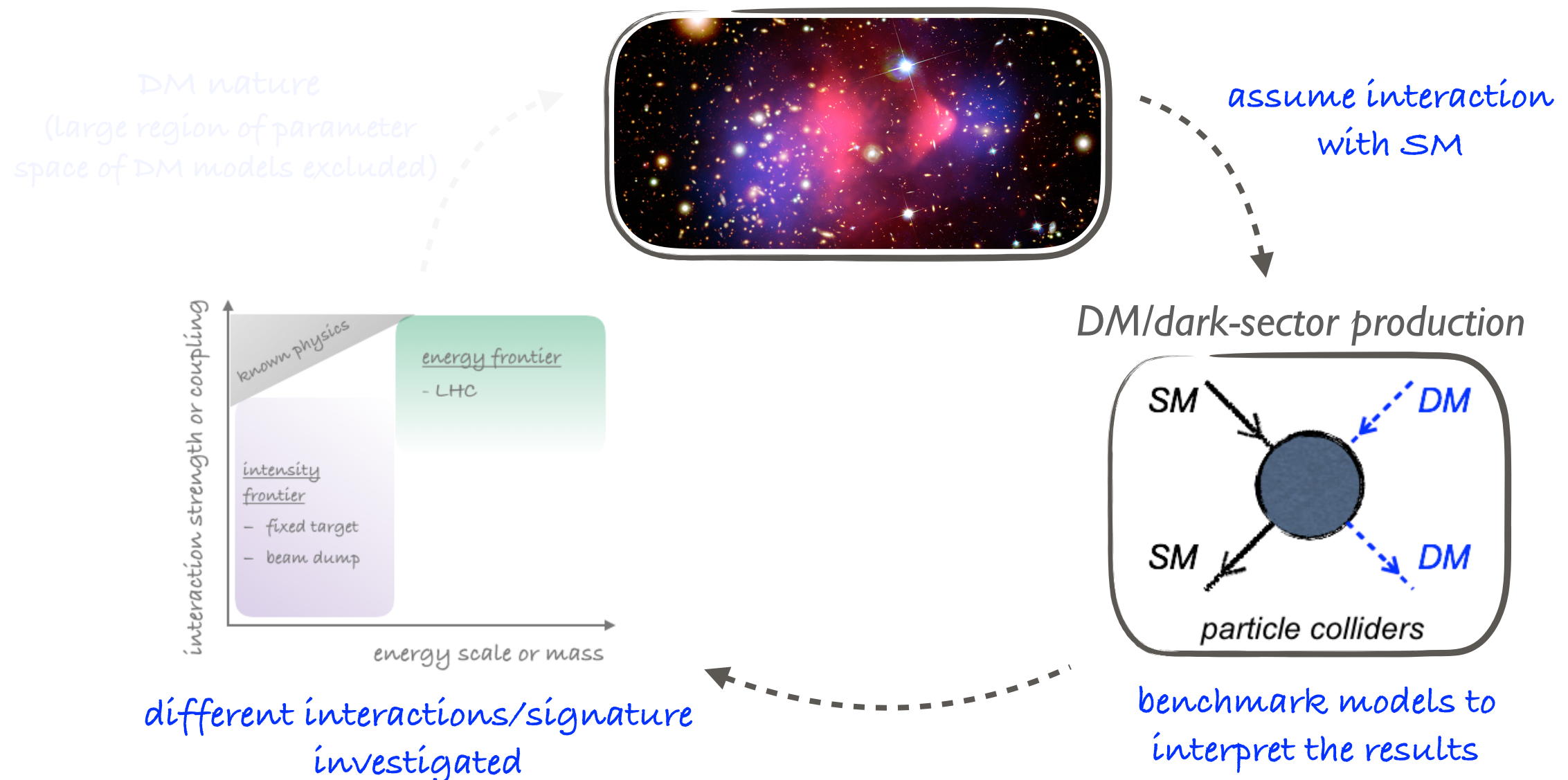
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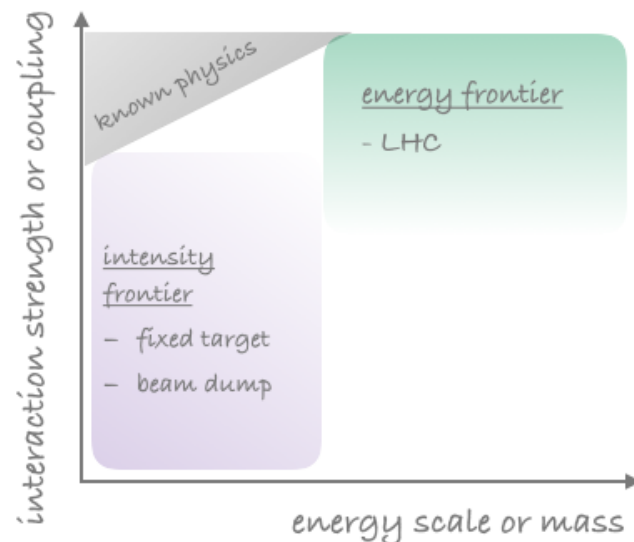
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Hunt for dark matter - in a nutshell

DM nature
(large region of parameter
space of DM models excluded)

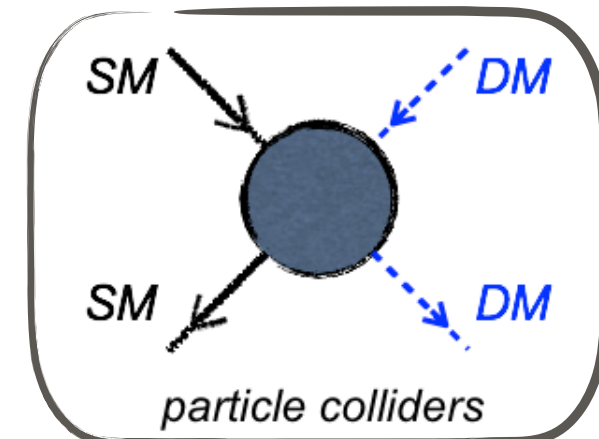


assume interaction
with SM



different interactions/signature
investigated

DM/dark-sector production

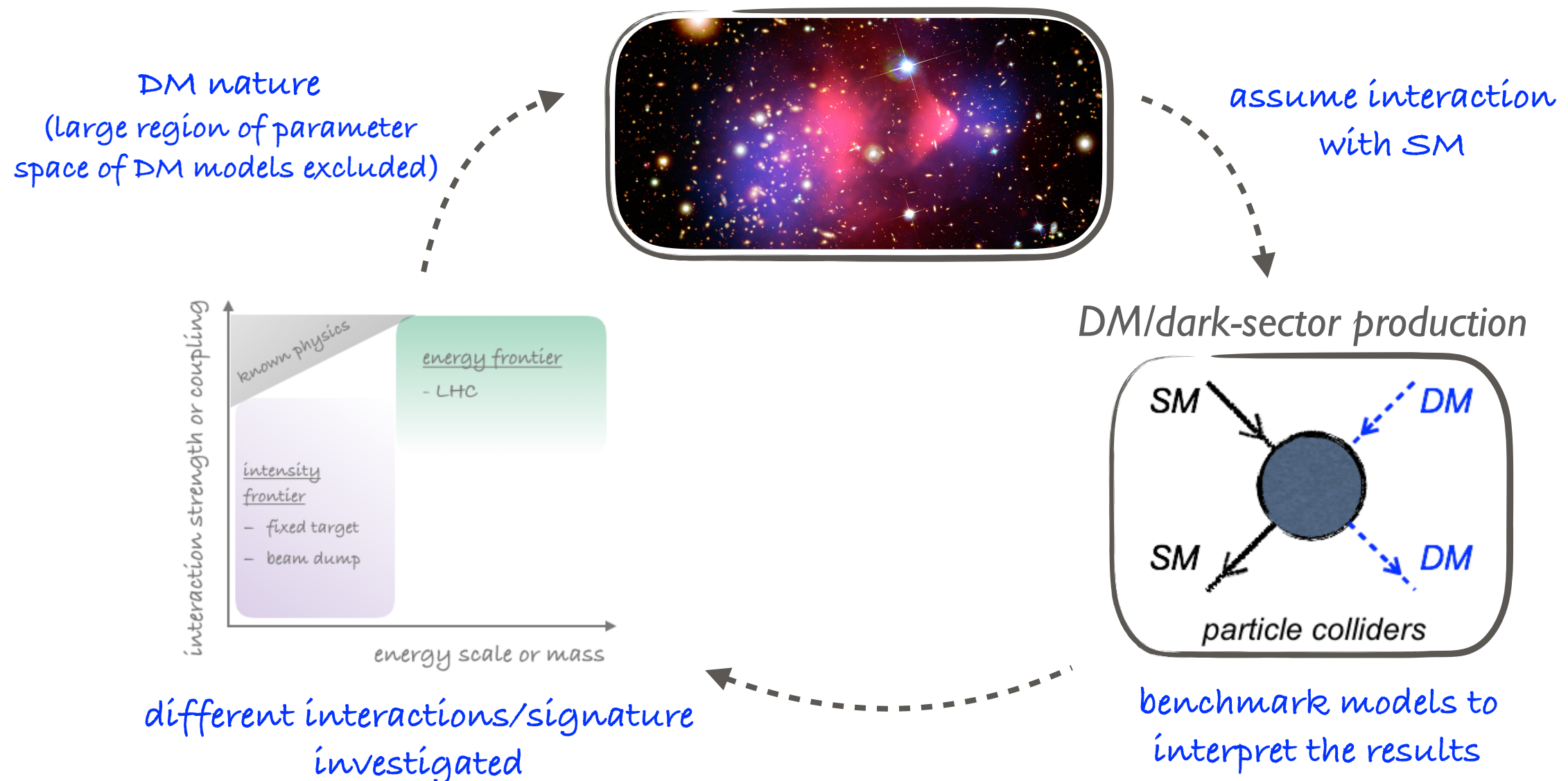


benchmark models to
interpret the results

- * Rich DM physics program at high-energy and high-intensity experiments
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Hunt for dark matter - in a nutshell



*Rich DM physics program at LHC

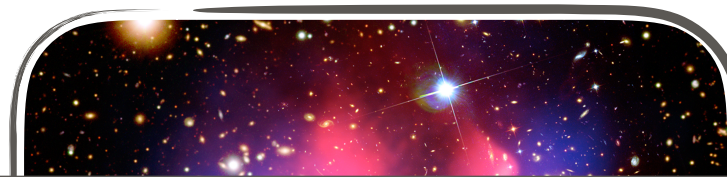
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Hunt for dark matter - in a nutshell



re
parameters



assume interaction

An eye towards the (near) future

New physics working group (LHC BSM WG)

guidelines and recommendations for the
benchmark models, interpretation, and
characterisation of BSM searches at the LHC

first general meeting in 10-13 November 2025
(indico)



inputs from various signatures (mono-X, resonances, $tt \rightarrow tt\nu$, visible, ...)

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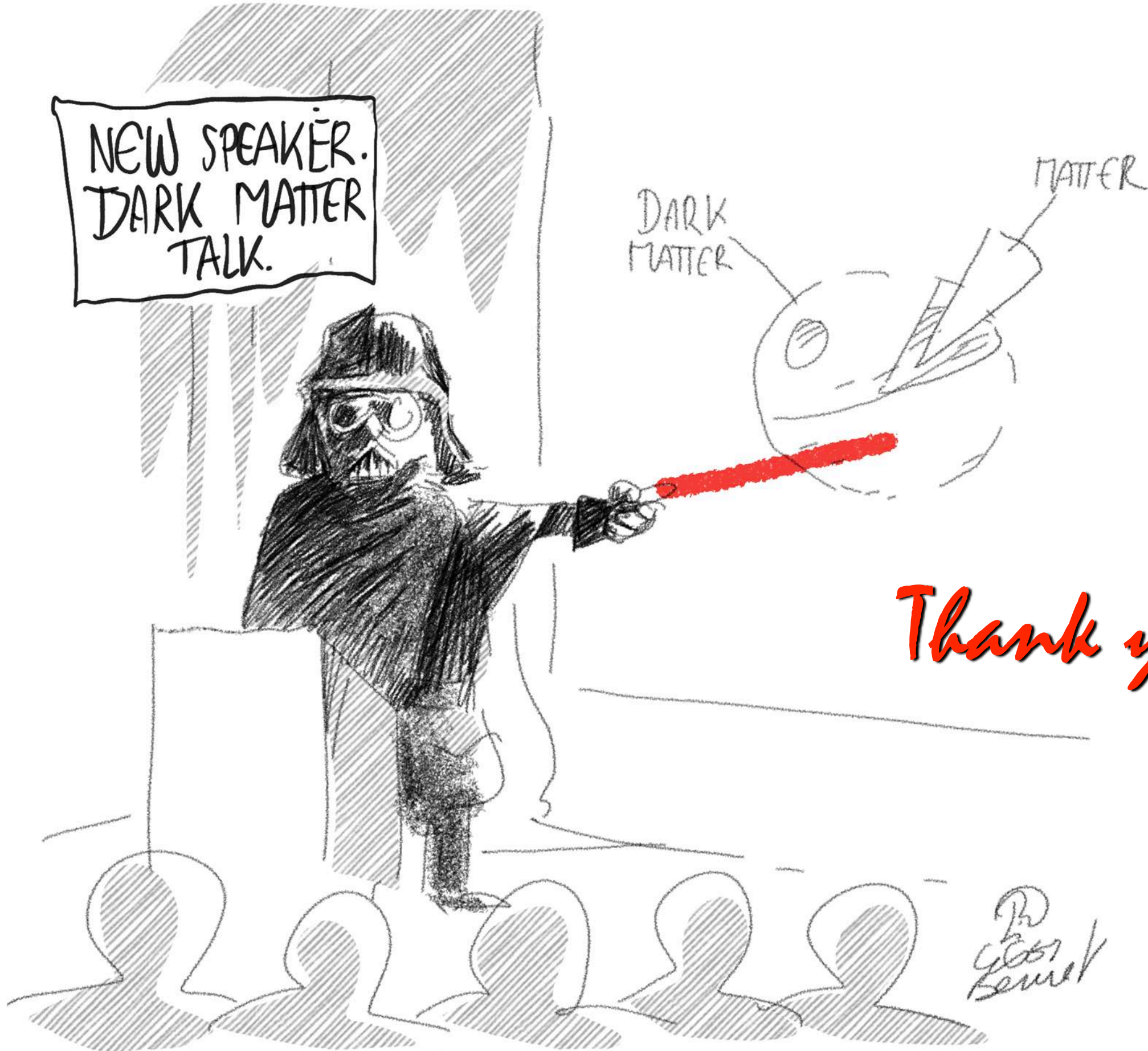
NEW SPEAKER.
DARK MATTER
TALK.

DARK
MATTER

MATTER

Thank you!

Dr
Gibson
Bennet



DARK MATTER

Backup

PROTON

PHOTON

NEUTRINO

MUON

QUARK

NEUTRON

VISIBLE MATTER

NEUTRALINO

AXION

AXION

GRAVITINO

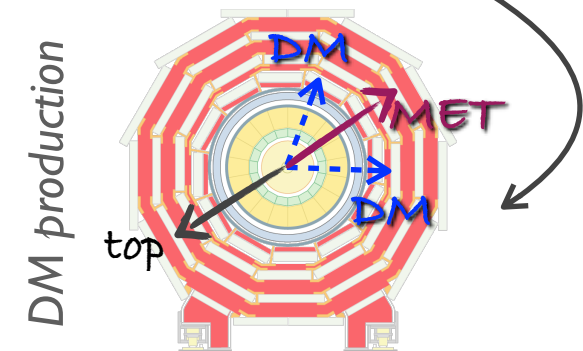
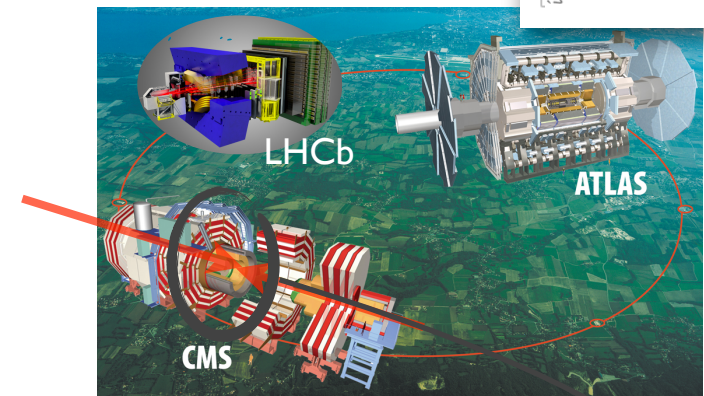
Wimp

Collider experiments: ATLAS, CMS, LHCb

ATLAS and CMS multipurpose detectors at LHC

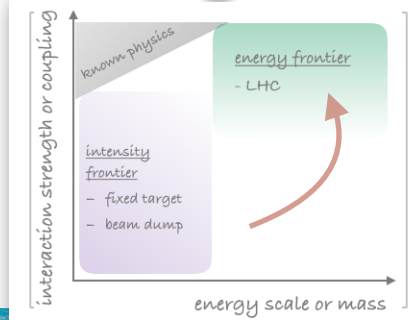
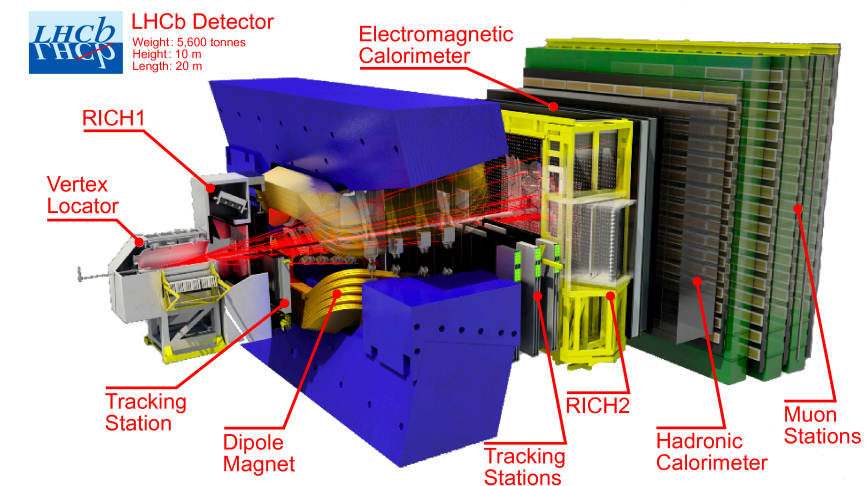
- goals: precision test of SM, search for new physics
- particle identification, energy and momenta measurements
- trigger system: select events interesting for physics analysis
- *pp collisions*
 - Run2: 13 TeV, $\sim 140 \text{ fb}^{-1}$ | Run3: 13.6 TeV, already collected $\sim 70 \text{ fb}^{-1}$
 - more than 8.5 million Higgs boson produced!

- *DM could be produced at colliders (rare process):*
 - invisible signature: *no direct trace in the detector, but ...*
 - can be inferred from p_T imbalance (**MET**)
 - *need visible particle* to which DM particle recoils against “mono-X”



LHCb at LHC

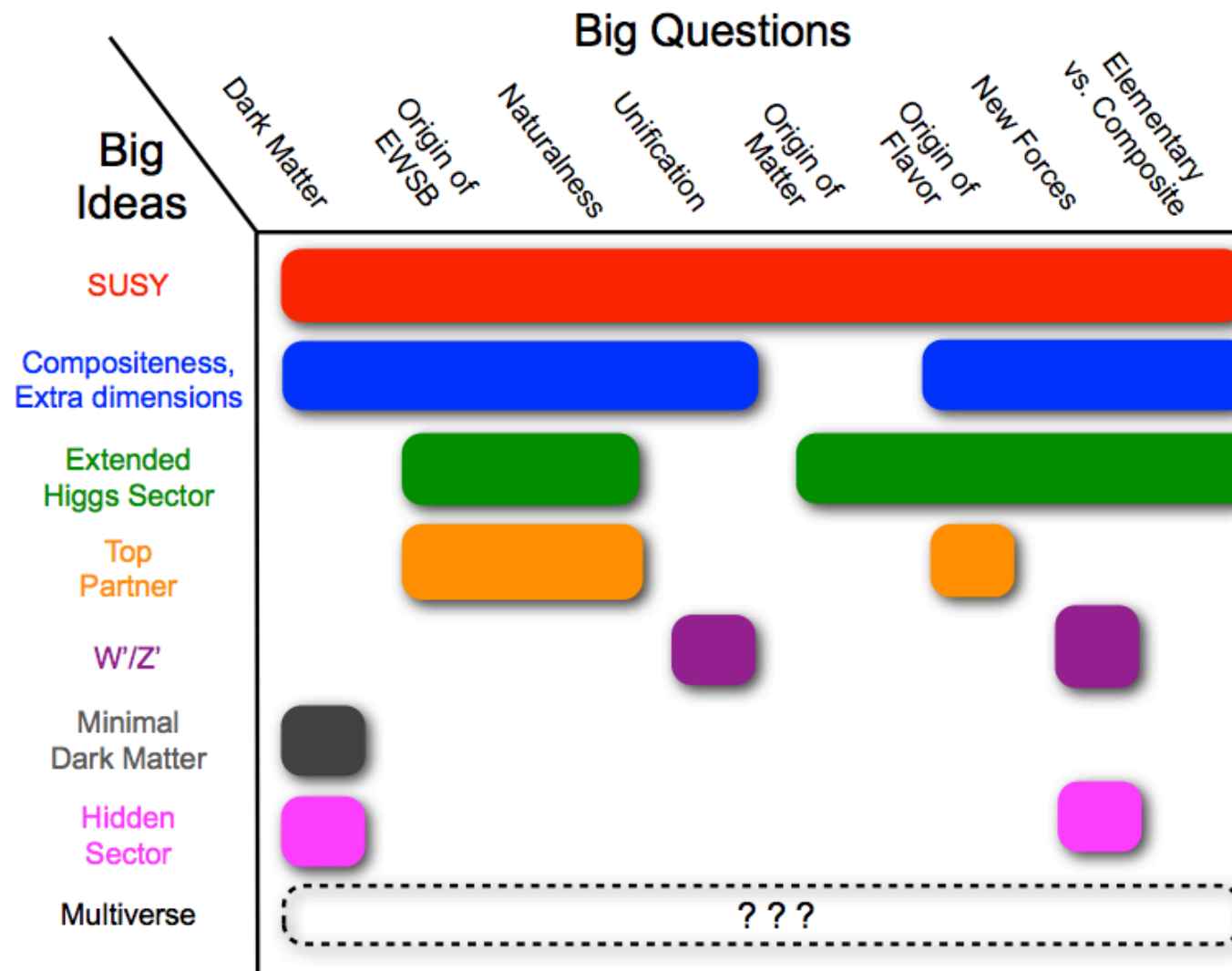
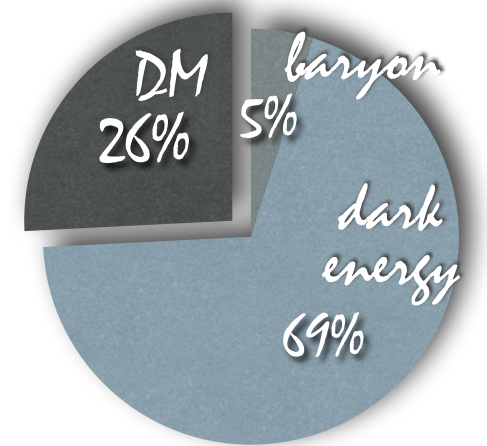
- single-arm spectrometer originally devoted to heavy flavour physics, now a general purpose experiment
- triggers with low p_T thresholds, probes rapidity region only partially accessible to other LHC experiments
- excellent vertex, mass and lifetime resolution, particle identification
- can operate in collider and fixed target mode
- *pp collisions*: $\sim 1/20$ ATLAS/CMS, reduced luminosity by offset beam collisions



Rich new physics program at ATLAS and CMS

► Despite the accuracy of the SM and its predictive power many open outstanding questions, eg.:

- matter-antimatter asymmetry
- hierarchy problem
- describes only ~5% of the universe, explanations for DM are not provided
- gravitational force cannot be included in the current theoretical framework



[arXiv:1311.0299]