

Advanced Virgo

status & prospects

Massimiliano Razzano

(University of Pisa and INFN-Pisa)

DIECI ANNI DI ONDE GRAVITAZIONALI

24 September 2025

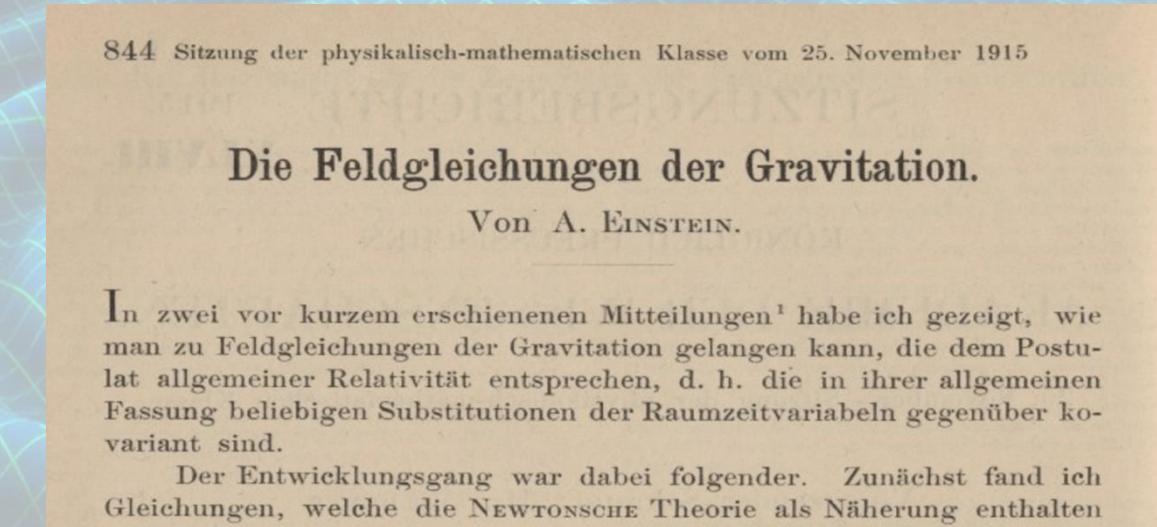
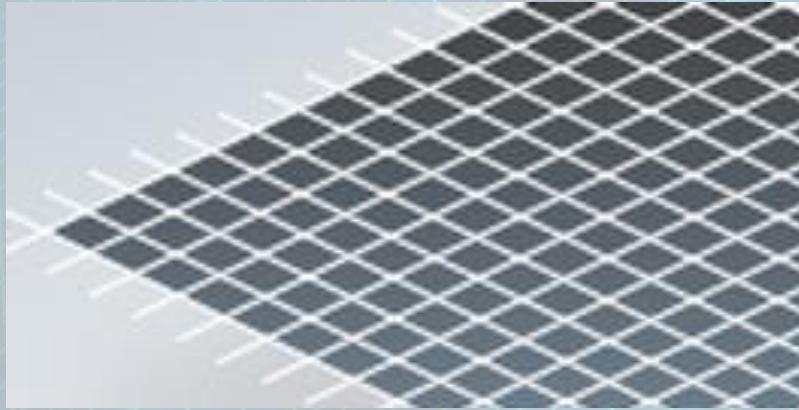


What are Gravitational Waves?

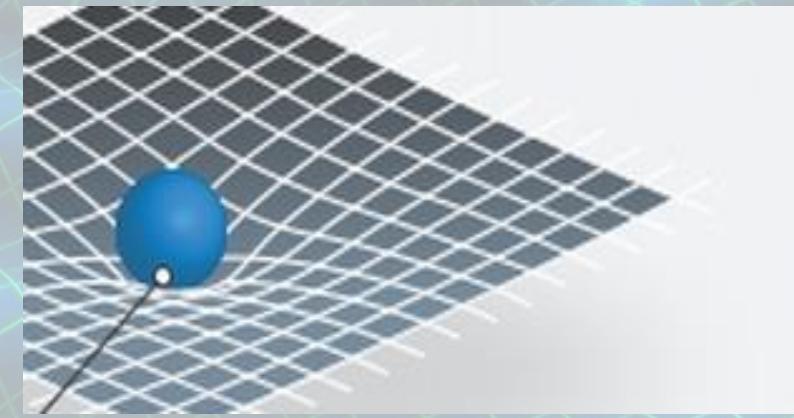
- A consequence of Einstein's General Relativity
 - Gravity as a manifestation of the geometry of the spacetime

*“Spacetime tells matter how to move;
matter tells spacetime how to curve”*

(J. Wheeler)



Credits: Preussische Akademie der Wissenschaften, Sitzungsberichte, 1915



Expected sources of Gravitational Waves

Transients

Non transients

- **Coalescence of compact binary systems (NSs and/or BHs)**

- Known waveforms (matched filter with template banks)
- Only source class detected so far

- **Core-collapse of massive stars**

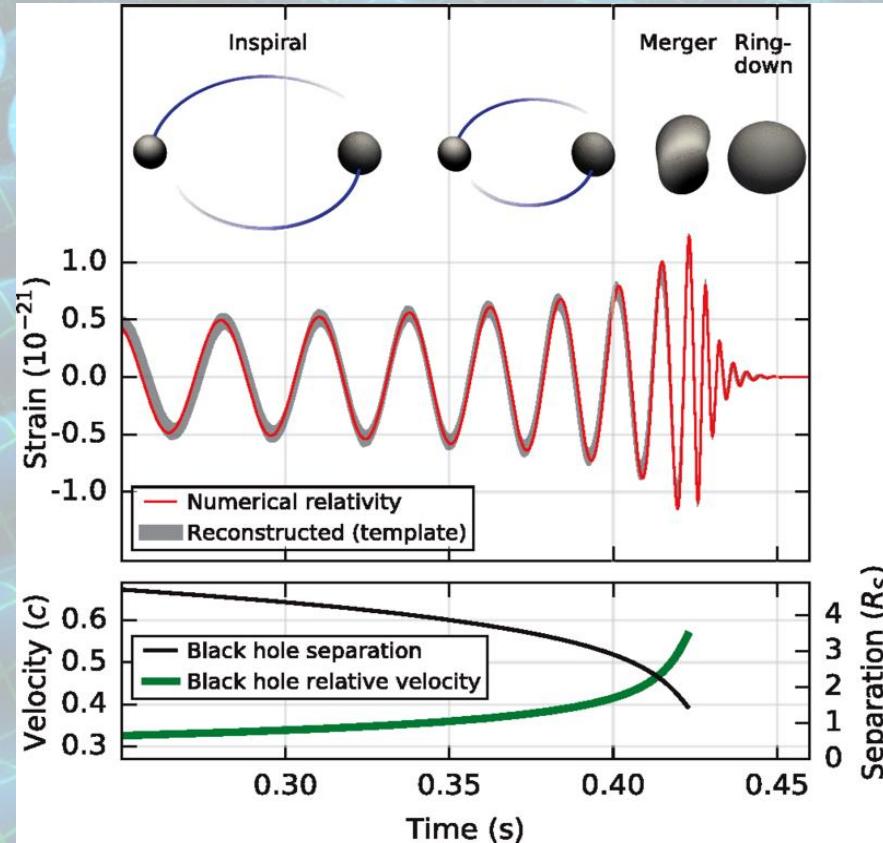
- Uncertain waveforms
- Unmodeled searches less sensitive than matched filter

- **Rotating neutron stars**

- Quadrupole emission from stellar asymmetry
- Continuous and periodic

- **Stochastic background**

- Continuous, due to unresolved sources/Big Bang relics



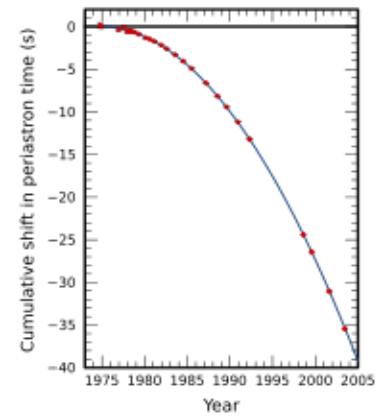
Abbott et al 2016, PRL, 116, 101103

The challenge of detecting GWs

1960's
J. Weber works
on resonant bars



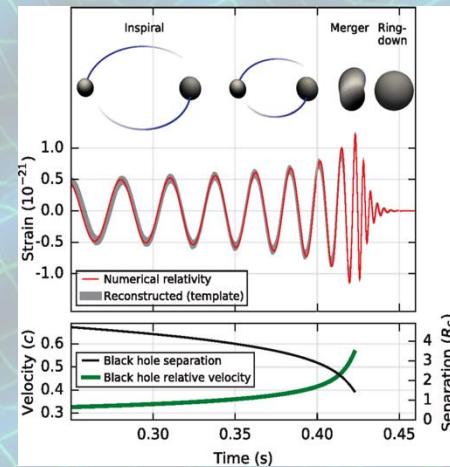
1974
Hulse-Taylor binary pulsar
Indirect evidence of GWs



1980s-1990s
First works on laser interferometers
(LIGO,Virgo)

2000s
First generation
(e.g. LIGO/Virgo/GEO600)
No detection 😞

2010s
Second «Advanced» generation
→ First detection!



LIGO & Virgo : origins

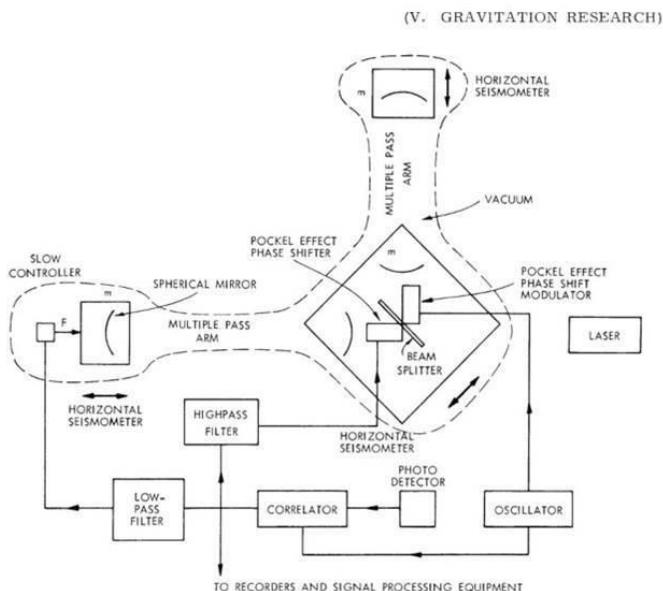
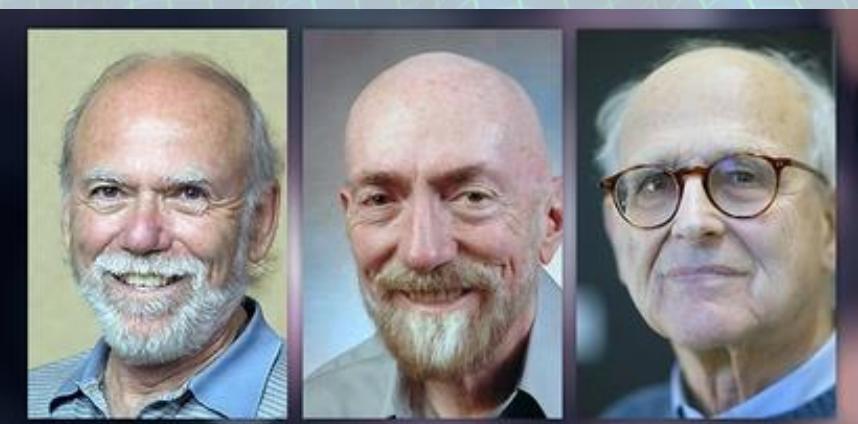
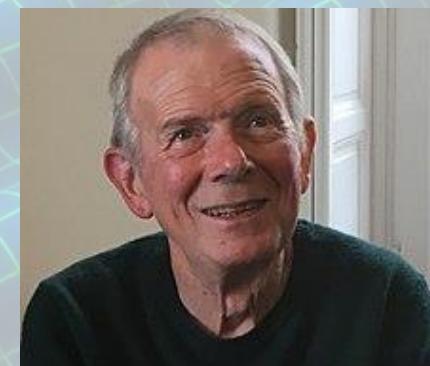
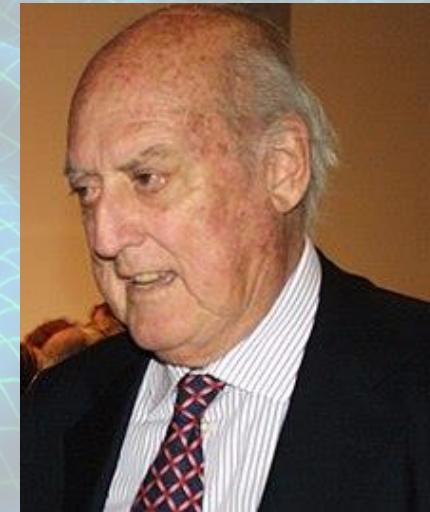
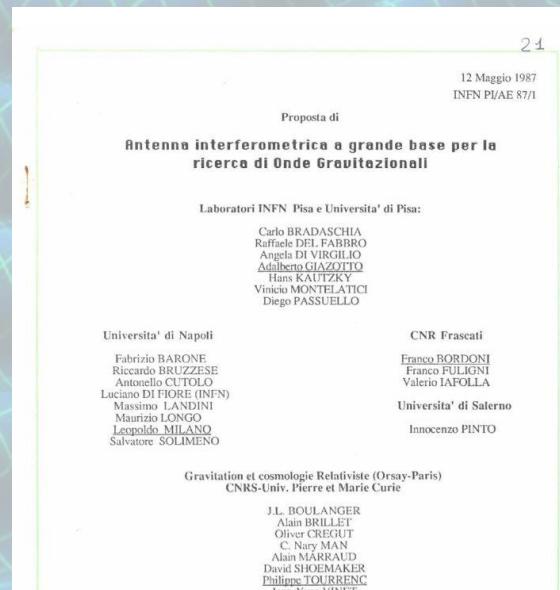


Fig. V-20. Proposed antenna.

In 1972 Rainer Weiss (MIT) proposed a GW detector exploiting interferometry

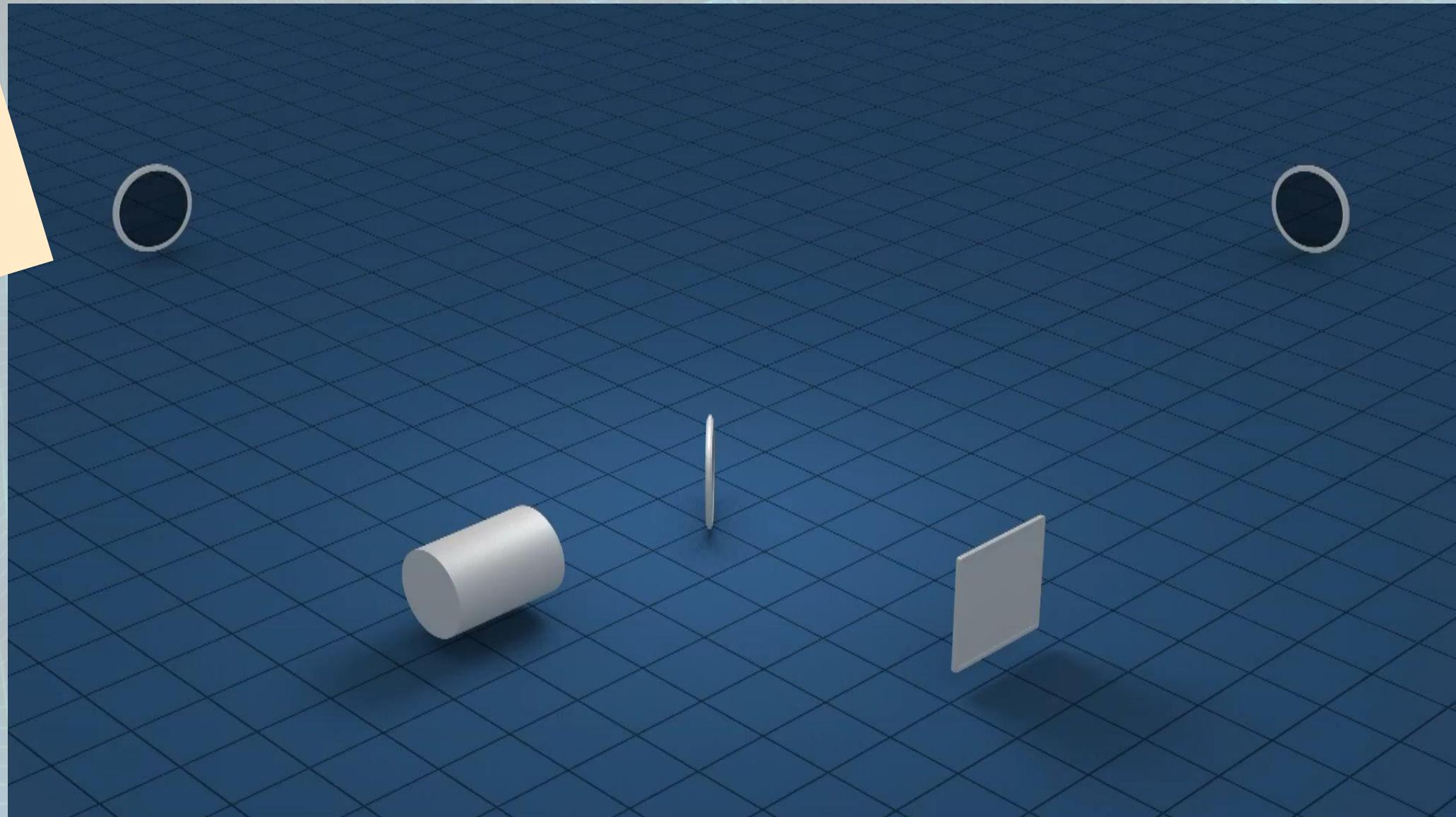
In mid-80s Giazotto&Brillet discussion on a GW detector focusing on low frequencies



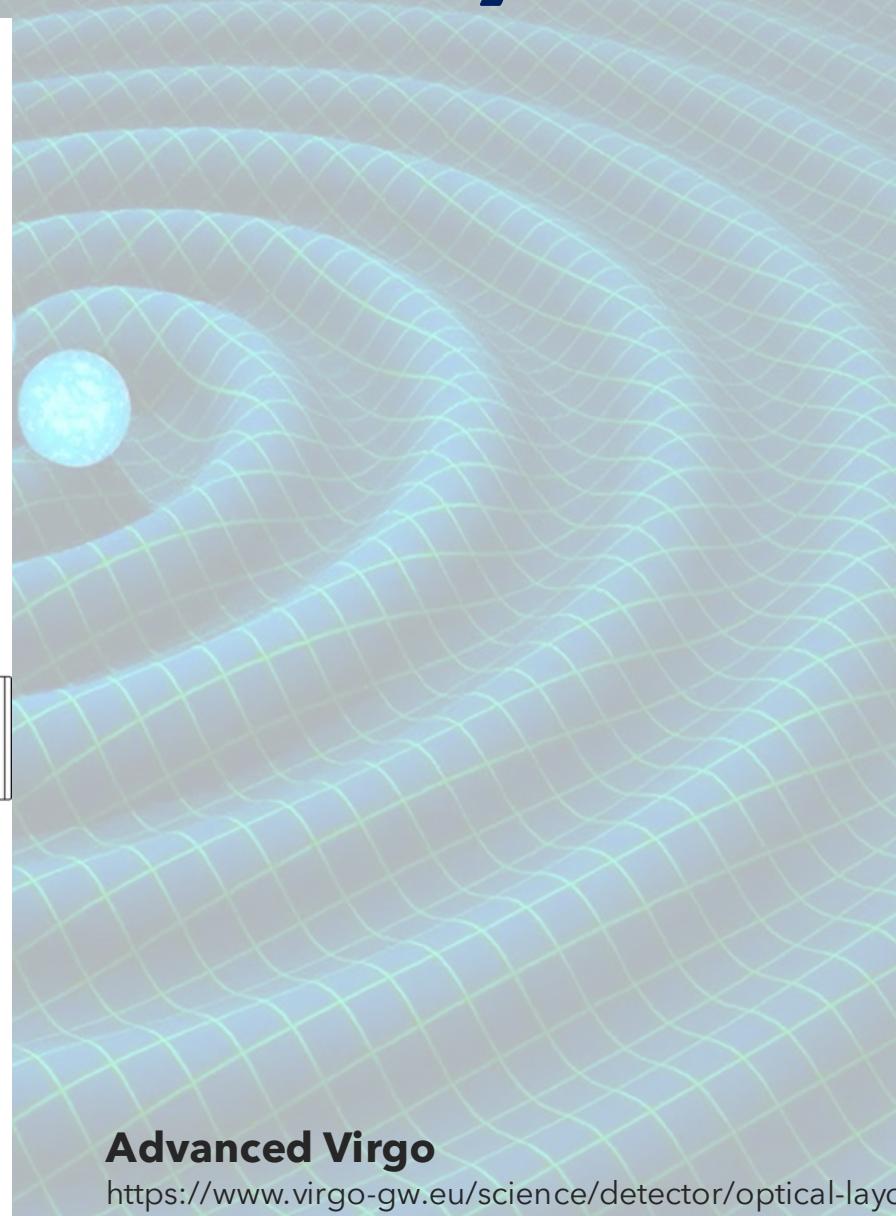
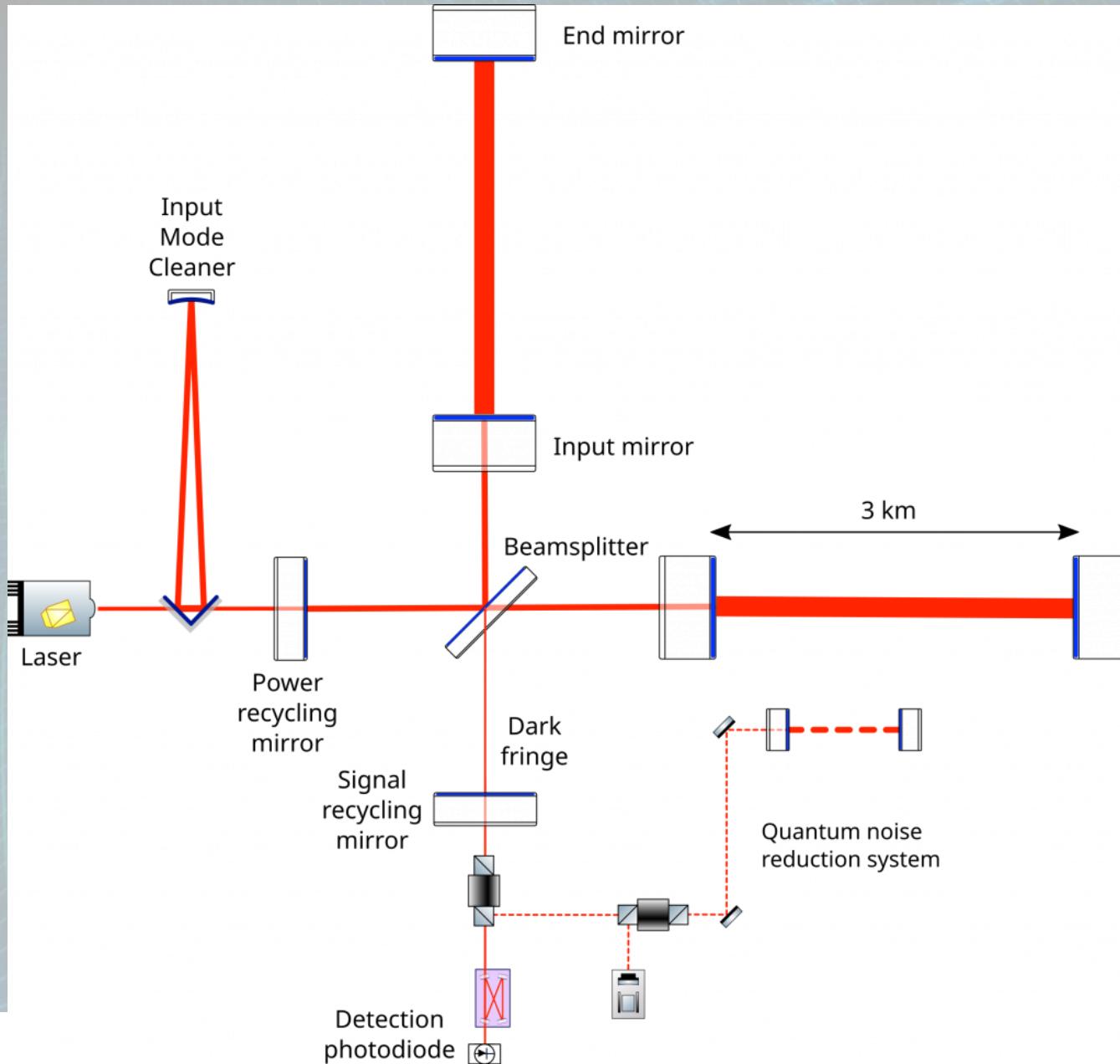
2017 Nobel Prize in Physics

How to detect Gravitational Waves

Expected $h(t) \sim 10^{-21}$
Since $h(t) \sim dL/L$
 $\rightarrow dL \sim 10^{-18} \text{ m}$



A more complex interferometer layout



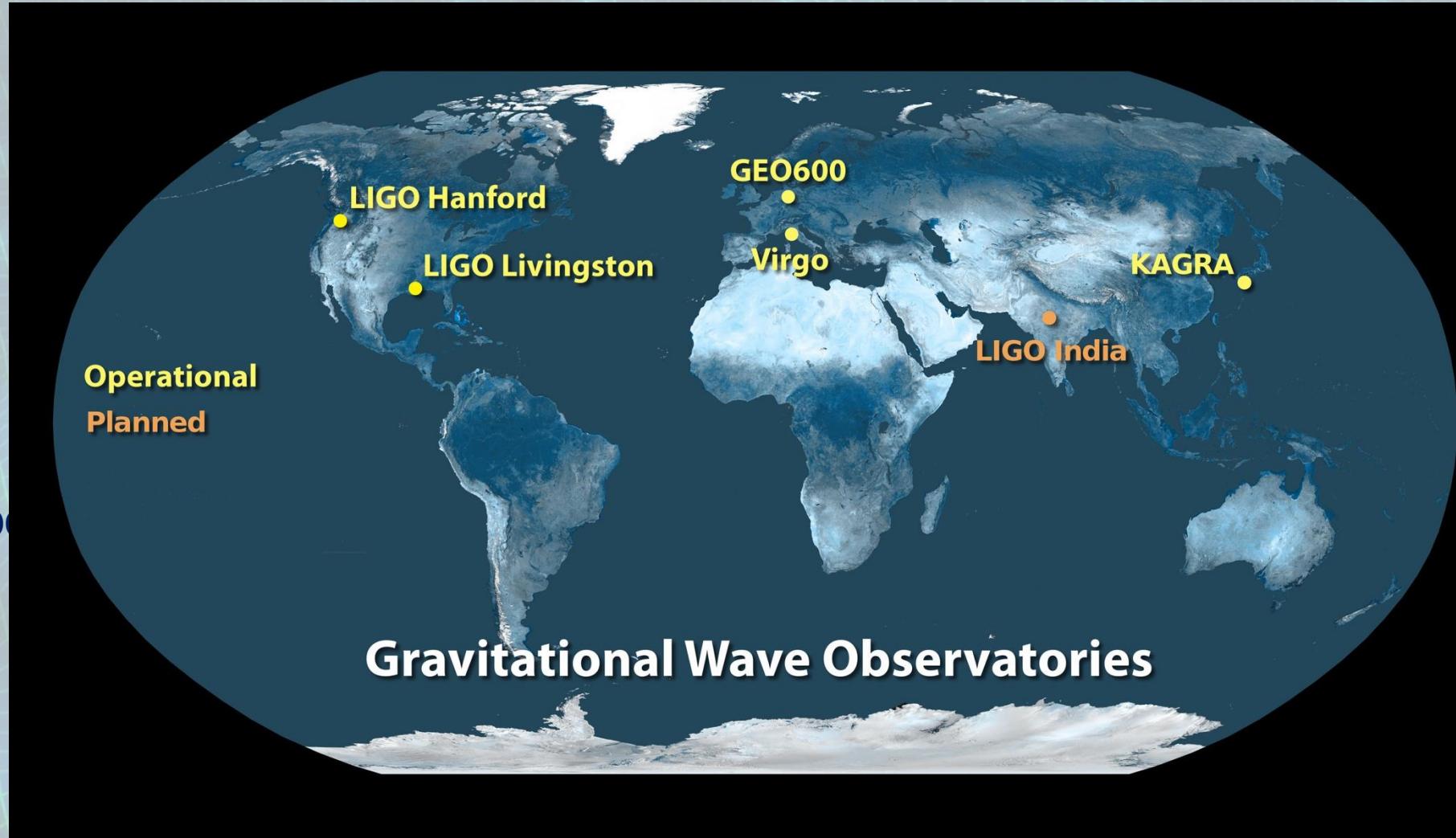
Advanced Virgo

<https://www.virgo-gw.eu/science/detector/optical-layout>

An international Network

Better sensitivity

- ~10x wrt previous generation (2002-2011)
- ~1000x more volume → ~1000 higher rates



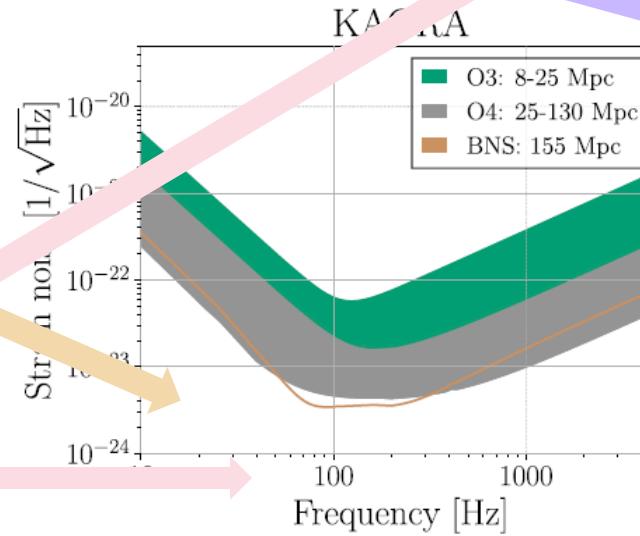
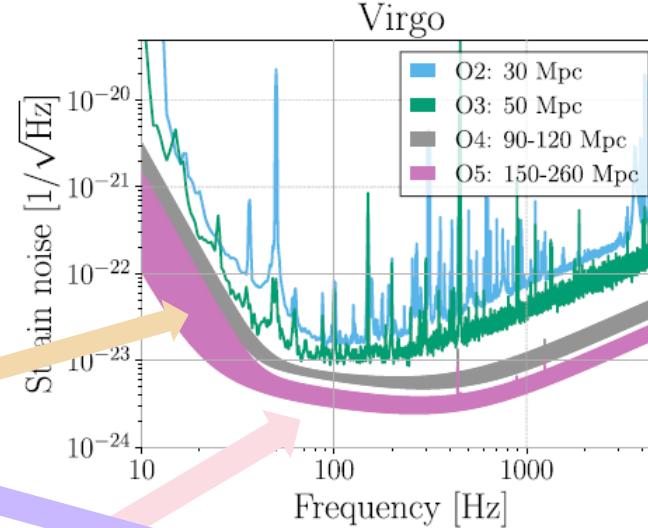
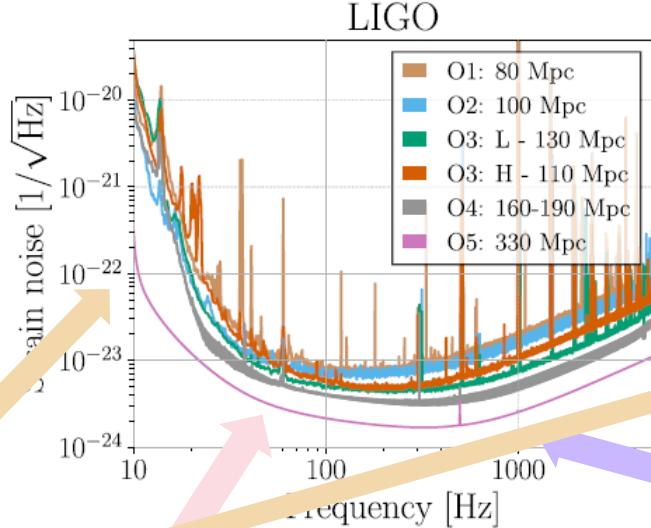
Sensitivity Curves

Low Frequencies

Seismic & Newtonian
Noise

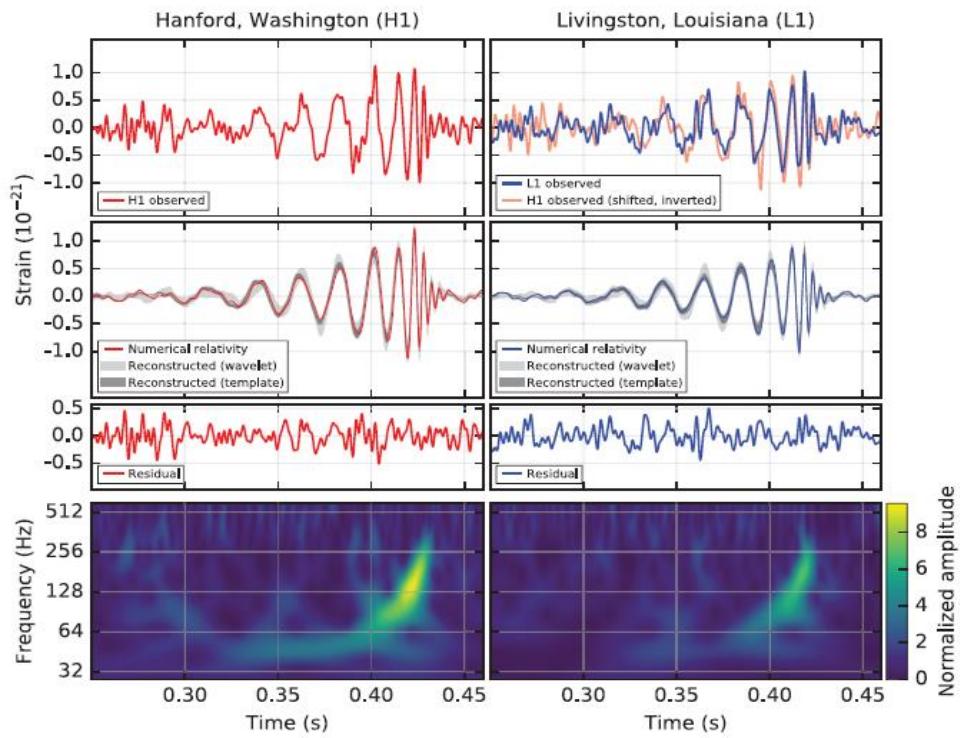
Mid Frequencies

Thermal Noise



High Frequencies
Shot Noise

10 years ago



GW150914
Abbott+16, PRL116,6

Modeled as coalescence of two black holes

PRL 116, 061102 (2016) Selected for a Viewpoint in Physics
PHYSICAL REVIEW LETTERS week ending
12 FEBRUARY 2016

Observation of Gravitational Waves from a Binary Black Hole Merger

B. P. Abbott *et al.*
(LIGO Scientific Collaboration and Virgo Collaboration)
(Received 21 January 2016; published 11 February 2016)

On September 14, 2015 at 09:50:45 UTC the two detectors of the Laser Interferometer Gravitational-Wave Observatory simultaneously observed a transient gravitational-wave signal. The signal sweeps upwards in frequency from 35 to 250 Hz with a peak gravitational-wave strain of 1.0×10^{-21} . It matches the waveform predicted by general relativity for the inspiral and merger of a pair of black holes and the ringdown of the resulting single black hole. The signal was observed with a matched-filter signal-to-noise ratio of 24 and a false alarm rate estimated to be less than 1 event per 203 000 years, equivalent to a significance greater than 5.1σ . The source lies at a luminosity distance of 410^{+160}_{-180} Mpc corresponding to a redshift $z = 0.09^{+0.03}_{-0.04}$. In the source frame, the initial black hole masses are $36^{+5}_{-4} M_{\odot}$ and $29^{+4}_{-4} M_{\odot}$, and the final black hole mass is $62^{+4}_{-4} M_{\odot}$, with $3.0^{+0.5}_{-0.5} M_{\odot} c^2$ radiated in gravitational waves. All uncertainties define 90% credible intervals. These observations demonstrate the existence of binary stellar-mass black hole systems. This is the first direct detection of gravitational waves and the first observation of a binary black hole merger.

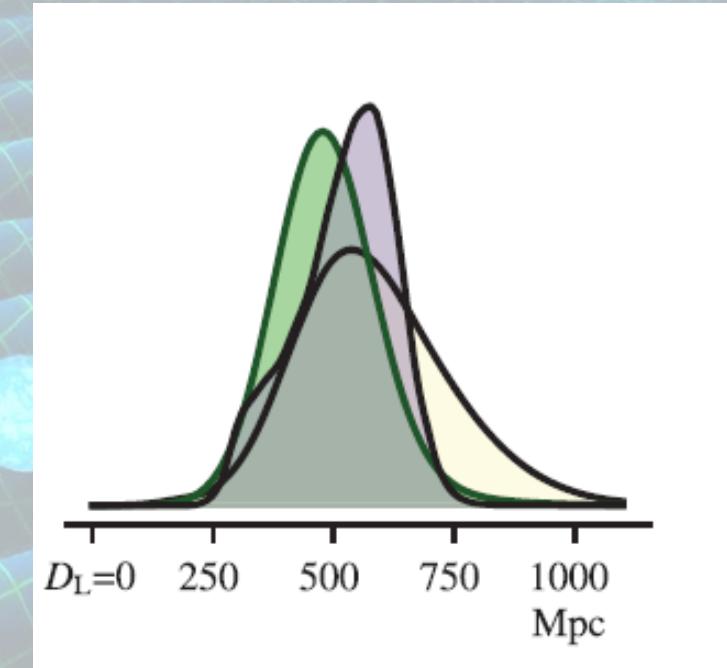
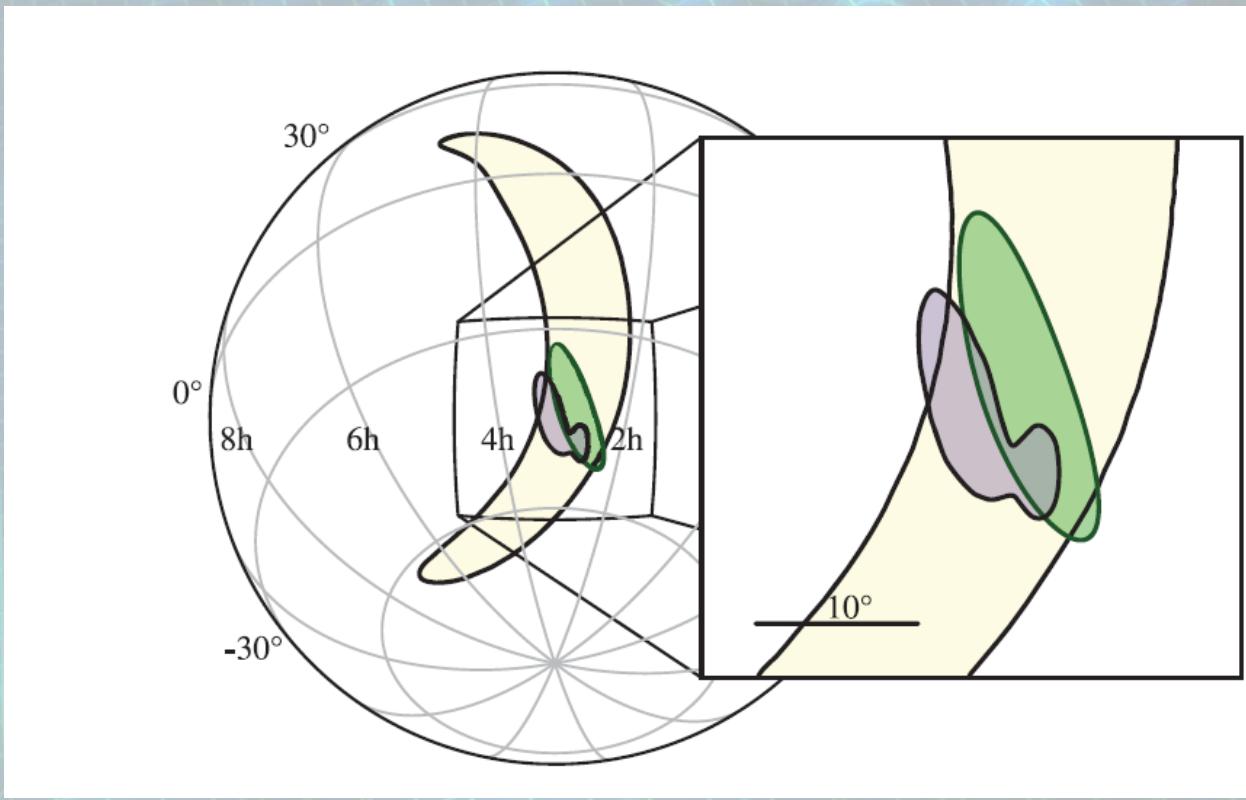
DOI: 10.1103/PhysRevLett.116.061102



GW151226
Abbott+16, PRL116,24

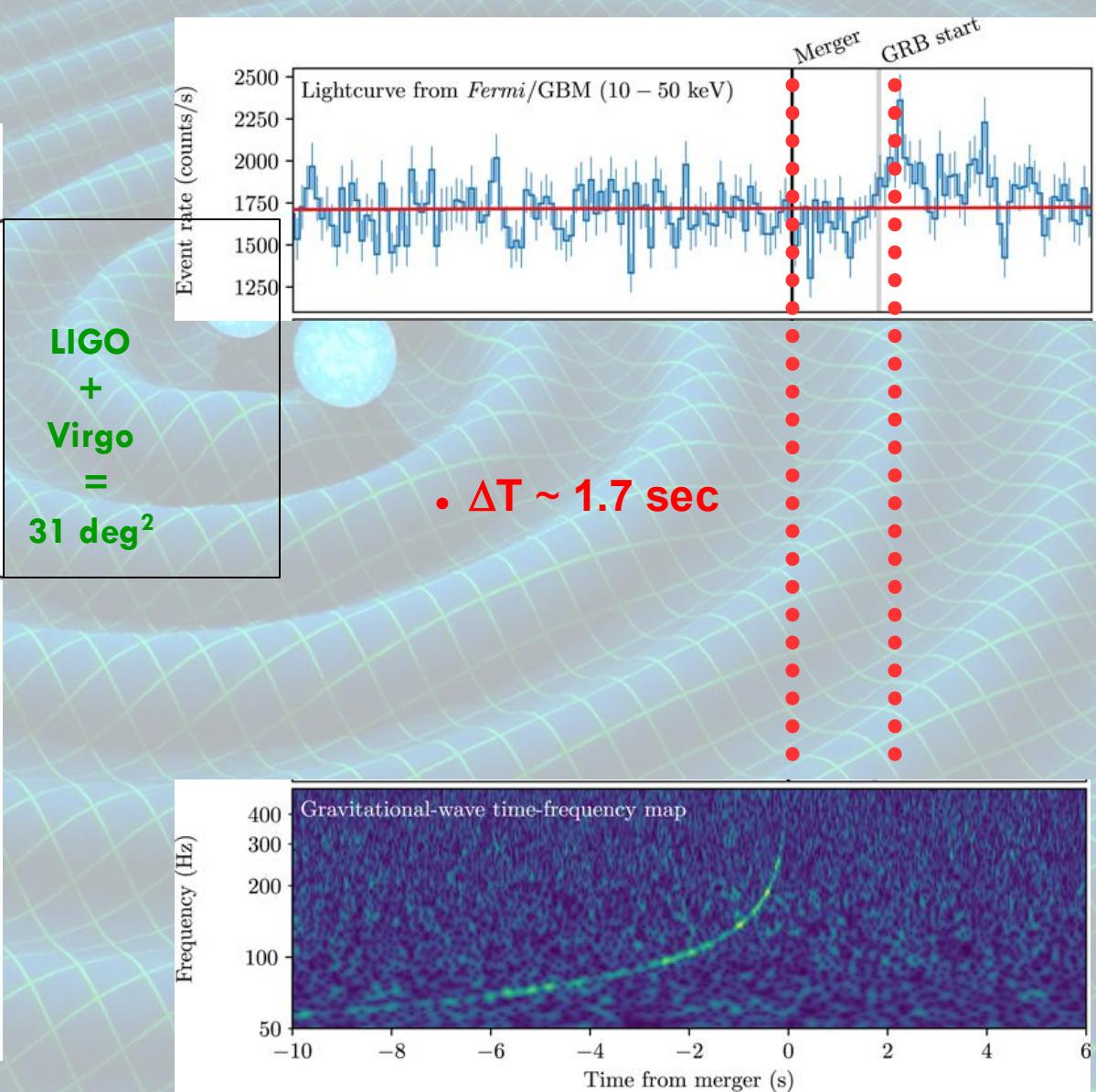
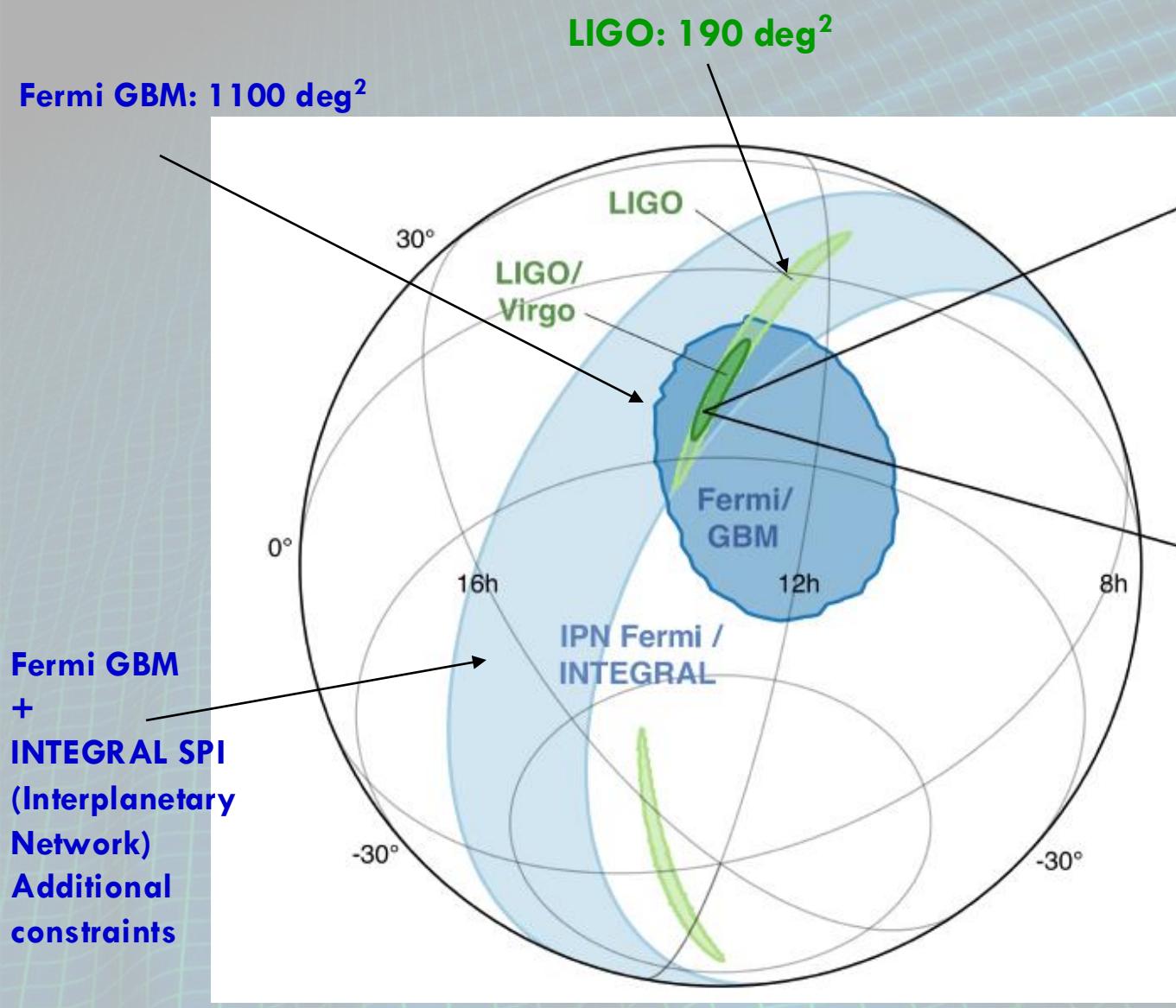
GW170814: the first «triple» detection

- Using only LIGO detectors → 1160 deg^2
- Adding Virgo → 100 deg^2
- Full analysis → 60 deg^2
- Credible volume (and # of galaxies)
 $71 \times 10^6 \text{ Mpc}^3 \rightarrow 2.1 \times 10^6 \text{ Mpc}^3$



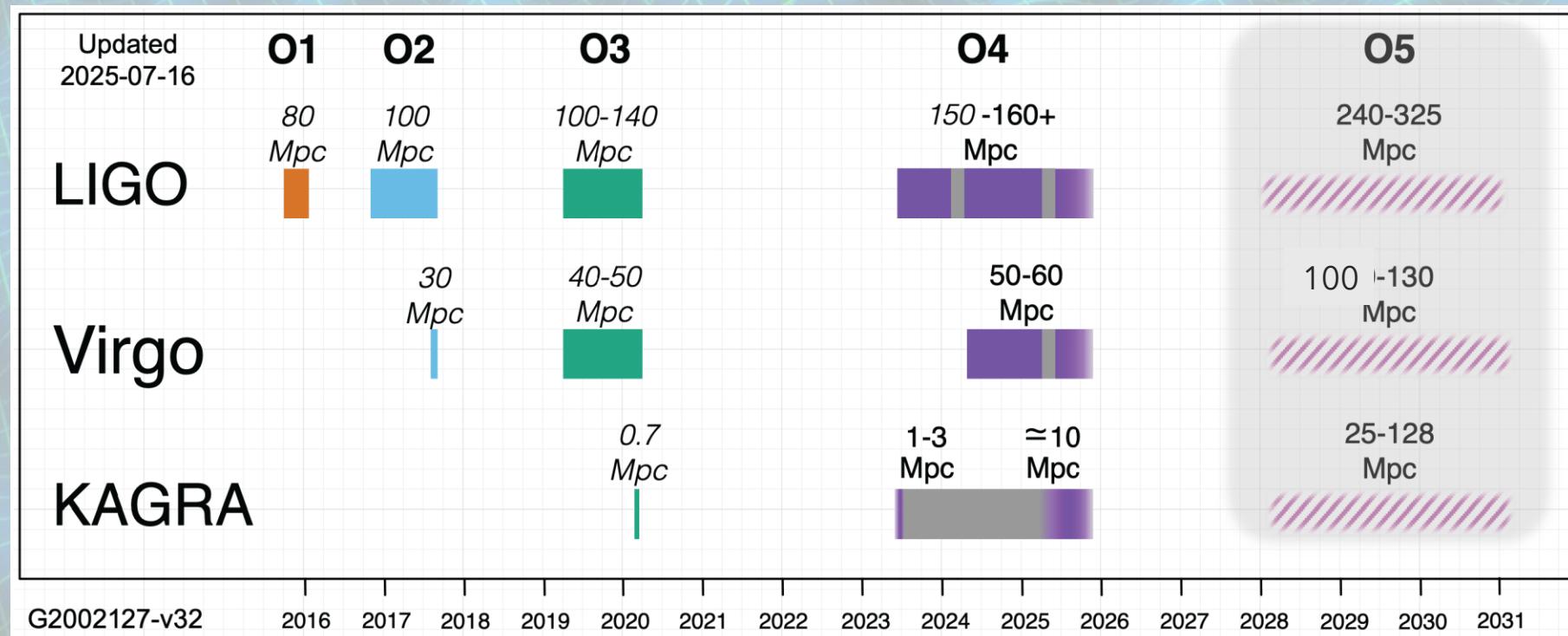
Abbott et al 2017, PRL, 119, 141101

The GW170817 event



The story so far

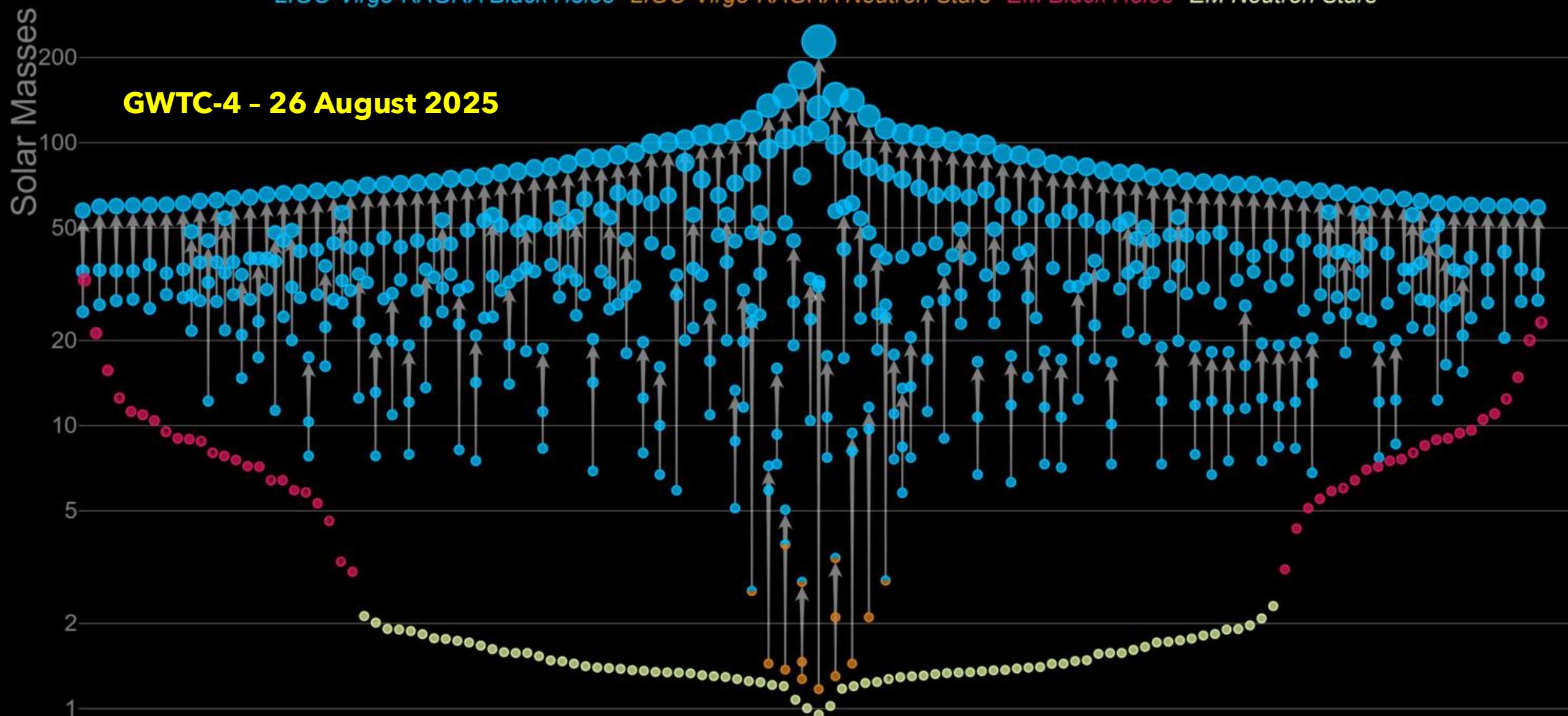
- Joint LIGO-Virgo-KAGRA runs
- O1 (H1+L1) - Sep 12, 2015 - Jan 19, 2016
- O2 (H1+L1+V1) - Nov 30, 2016 - Aug 25, 2017
- O3a (H1+L1+V1) - Apr 1 - Oct 1, 2019
- O3b (H1+L1+V1) - Nov 1, 2019 - Mar 27, 2020
- O4a (H1+L1) – May 24, 2023 – Jan 16, 2024
- O4b (H1+L1+V+K*) – Apr 10, 2024 – Jan 28 2025
- O4c (H1+L1+V1) - Jan 28, 2025 – Nov 18, 2025 (maintenance & commissioning break Apr 1 – Jun 11)



Masses in the Stellar Graveyard

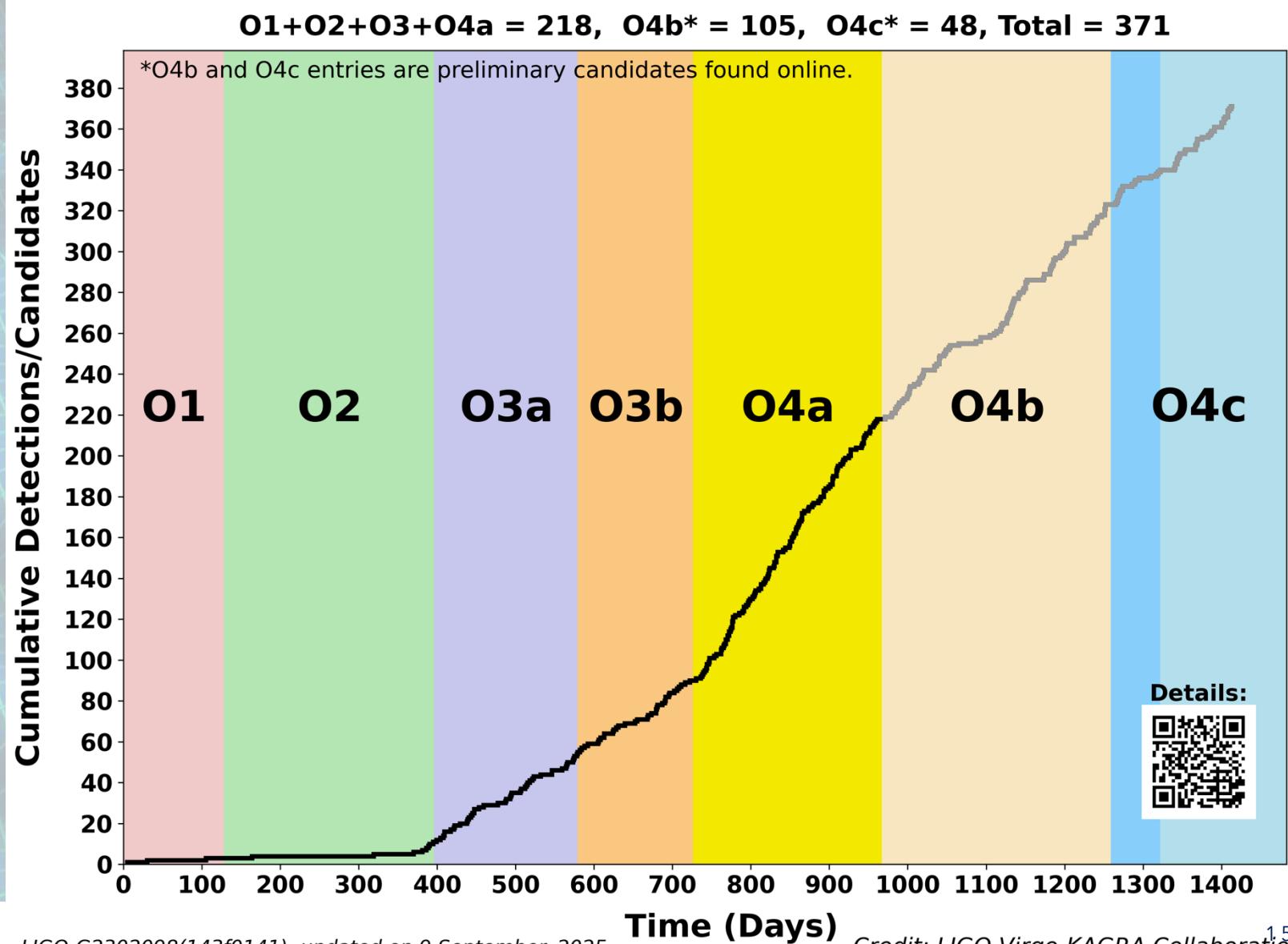
LIGO-Virgo-KAGRA Black Holes LIGO-Virgo-KAGRA Neutron Stars EM Black Holes EM Neutron Stars

GWTC-4 - 26 August 2025



O4 run - detections so far

Upgrades from **O3** to **O4**



Alerts in O4 (so far)

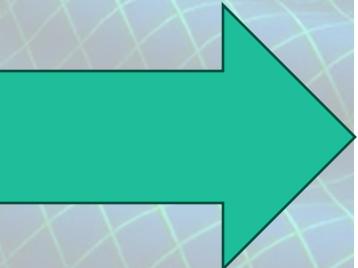
O4a

- 81 significant (FAR <1/6 mo) alerts, 11 retracted
- 1610 Low-significance

O4c*

- 51 significant (FAR <1/6 mo) alerts, 7 retracted
- 1394 Low-significance

Updated data from <https://gracedb.ligo.org/>

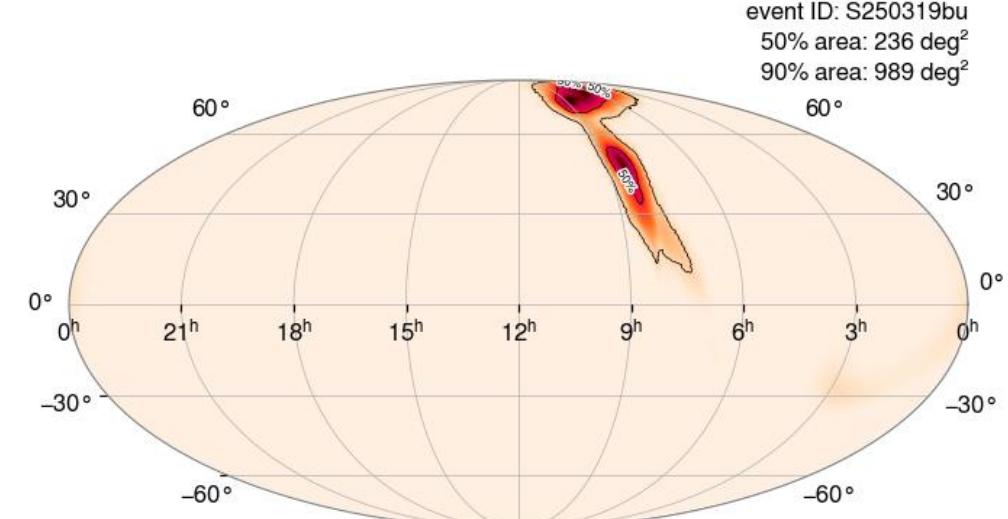


O4b

- 105 significant (FAR <1/6 mo) alerts, 9 retracted
- 1706 Low -significance

RECORD DETECTION OF 200 GRAVITATIONAL WAVES IN THE CURRENT RUN OF LIGO, VIRGO AND KAGRA

Mar 20, 2025



*as of Sep 24,2025

Conclusions

- Gravitational waves have opened a new windows on the Universe
- 3 runs successfully concluded
- O4 ongoing, extended until Nov 18, 2025
- Ca 300 detections so far (including O4 alerts)
- Still... lots of open questions
- Plans for future upgrades to further improve sensitivity in O5
- Many years of great science ahead!

Thank you for your attention !

