# Report from SAND Calibration WG

P.Gauzzi
(Universita' La Sapienza e INFN – Roma)
for the SAND Calibration WG

Meeting annuale della Collaborazione DUNE Italia LNF - November 11, 2025







### **SAND Calibration WG**

- Calibration: from detector signals to physical variables
  - ECAL: energy, time and positions of the particles
  - GRAIN: tracks, time, energy, ....
  - Tracker: r-t relations, track momentum, dE/dx for PID, ....
  - Timing alignment among the subdetectors
- Define a strategy for each subdetector:
  - Sources: cosmics, particles from beam, ...
  - Choose suitable processes (given the expected fluxes of particles in the detector, e.g. for the ECAL: cosmic μ's as MIPs, MIPs from the beam, electrons and photons ....)
  - Set a calibration procedure (Which level of precision? How much time expected?)
    - People: ECAL P.Gauzzi R.D'Amico GRAIN: A.Surdo, Tracker: ......
- WG meetings generally every three weeks, on Thursday at 5 p.m. CET
- WG mailing list: <u>dune-nd-sand-calibration@fnal.gov</u>

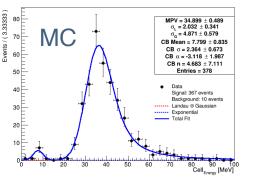




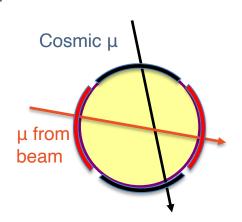


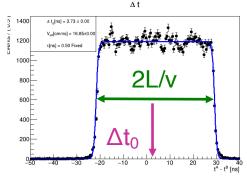
# ECAL calibration with MIPs

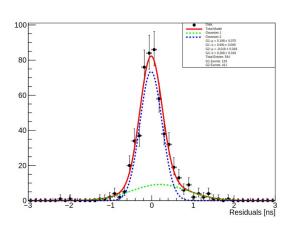
- ECAL: 4880 readout channels (PMTs), 2440 cells
- MIP peaks to equalize the Energy response of the ECAL cells



- Time calibration: 2 sets of constants to determine
- $t_0$ 's and  $\Delta t_0$ 's for timing alignment of the cells and for calibration of the coordinate along the fibres
- We need muons from beam and from cosmic rays
- MC study of calibration with muons
  - ⇒ see R.D'Amico's talk







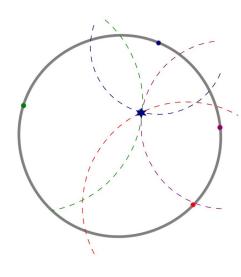






# **Energy scale calibration in SAND**

- $\gamma$ 's from  $\pi^0$  decays, invariant mass reconstruction (need a vertex from the tracker)
- $\gamma$  + electrons: ~ 30% of photons from  $\pi^0$  convert in the tracker
  - $\Rightarrow$  ~ 50% of  $\pi^0$  have at least one  $\gamma \rightarrow e^+e^-$  (from DUNE-doc-13262 A Near Detector for DUNE)
- High energy electrons from  $v_{\rm e}$  interactions  $\Rightarrow$  need the momentum measurement in the tracker
- Also exploit K<sup>0</sup>→π<sup>0</sup>π<sup>0</sup> →4γ
- From a naive rescaling of  $K^0 \rightarrow \pi^+\pi^- \Rightarrow O(10^5)$  evts in 5 years of FHC data-taking
- Reconstruct a vertex with the ECAL only, back-propagating each of the 4 photons, but the times of the ECAL cells must be very well aligned





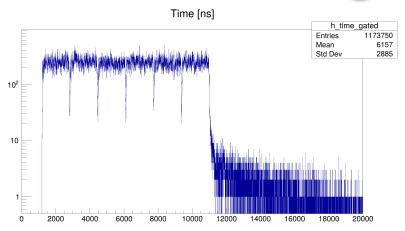


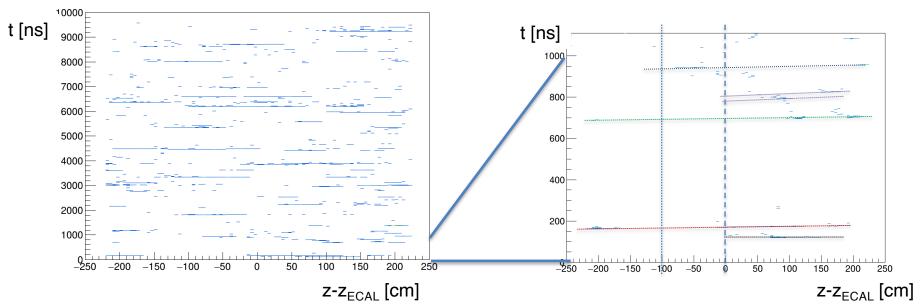


# ECAL global t<sub>0</sub>

$$t = \frac{1}{2}(t_A + t_B) - \frac{L}{2v} - t_0 - t_0$$

- Global t<sub>0</sub>, to be determined for each event
- The clusters will have the time distribution of the spill
- On average 88 v interactions/spill in the SAND volume (FHC mode, 1.2 MW)
- Cell time vs Z coordinate:







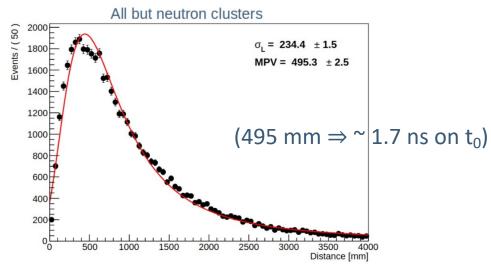


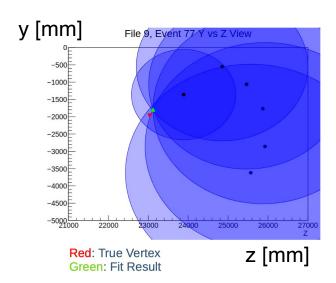


# ECAL: Global to

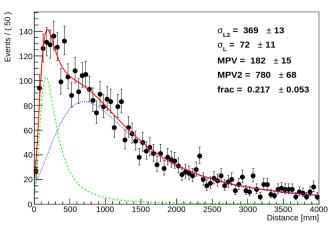
- A complementary approach is to find the vertex of the v interaction and compute the global t<sub>0</sub> from it
- Take the apex of each cluster as the center of a sphere, find the intersection of all the spheres with a fit.

 $5 \times 10^5 \, v_{\mu}$  interacting in the tracker volume: select events with at least 3 clusters (all particles except neutrons)





 $3 \times 10^4 v_{\mu}$  interacting in the whole SAND (all particles except neutrons)





 $(780 \text{ mm} \Rightarrow \sim 2.5 \text{ ns uncert. on } t_0)$ 





#### **GRAIN** calibration with muons

#### **Expected muon flux from the beam and CRs**

✓ Different contributions of the target masses in SAND for beam neutrinos

Table 1.34: Number of events per spill (9.6  $\mu s$ , 7.5 imes 10<sup>13</sup> POT) and selection efficiency for the signal from  $\nu_{\mu}$  CC in the front barrel ECAL and the backgrounds from rock muons and magnet events.

(from DUNE-doc-13262, A Near Detector for I	DUNE)
---------------------------------------------	-------

Table 1.29: Total number of  $(\nu_{\mu} + \bar{\nu}_{\mu} + \nu_{e} + \bar{\nu}_{e})$  CC+NC events expected within a single beam (9.6  $\mu$ s, 7.5  $\times$  10<sup>13</sup> POT) in the various detector components for both the FHC and RHC beam modes.

	ECAL		Rock muons		Magnet events	
Cut	Events	$\varepsilon$ (%)	Events	$\varepsilon$ (%)	Events	$\varepsilon$ (%)
No cut	2.23	100.0	1447.26	100.000	50.82	100.000
$\mu$ in ECAL FV	2.23	100.0	12.73	0.880	18.92	37.229
STT & ECAL hits	1.63	72.9	6.05	0.420	3.443	6.775
NN cut	1.56	95.5	0.10	0.007	0.07	0.136

• Further contribution from rock μ's (~ 1.7/spill) ...

Detector element	Mass	FHC	RHC
Magnet	511 t	68.9	36.6
ECAL	100 t	13.5	7.2
LAr+STT	8.2 t	1.1	0.59
STT fiducial volume	5.5 t	0.74	0.39
Total	619.2	83.5	44.39

• From the interaction rate /spill in Magnet yoke and ECAL, a quite low number of clean muons are expected to cross GRAIN per spill (  $\leq$  1  $\mu$  / spill)

#### ✓ Contribution from Cosmic Rays ...

CR Muon flux at surface  $\sim 0.01 \,\mu/(s \,cm^2) + underground reduction of <math>\sim 100$  Effective area of GRAIN for  $<60^{\circ}$  CR muons:  $\sim 3\times 10^4 \,cm^2 \Rightarrow \sim 3 \,\mu/s$  are expected to cross GRAIN

Drawback: smaller acceptance by the tracker for a precise track reconstruction

Main contribution only if inter-spill DAQ were ON



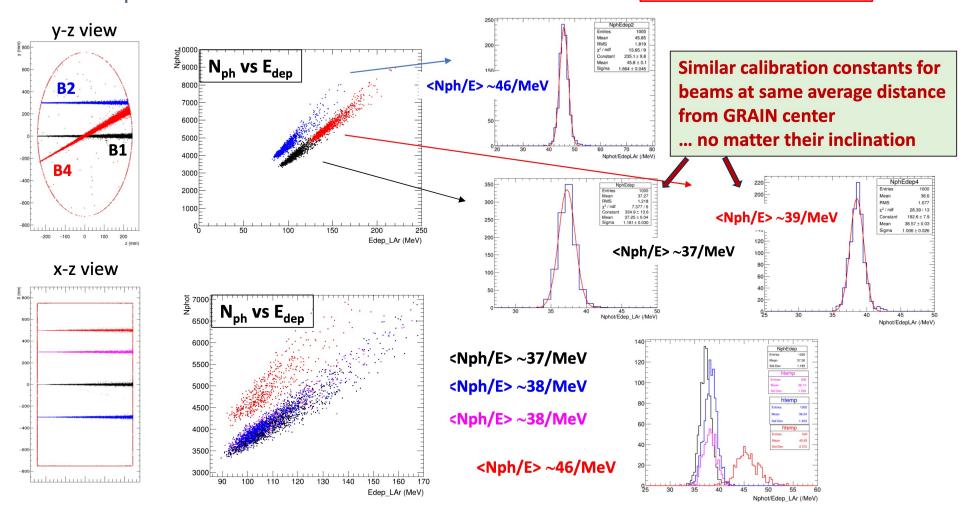




#### Muon beams with different positions and inclinations

MC samples of monochromatic muons of 1 GeV

See talk by A.Surdo









#### **Conclusions**

#### **ECAL**

- Calibration with muons (see Riccardo's presentation)
- First MC tests for a global t<sub>0</sub> determination
   (with ECAL alone, the method will improve by using the info from the tracker)
- Next steps:
  - Start the study of the energy scale calibration
  - Define the number of constants and the amount of data to be saved in a DB

#### **GRAIN**

Studying the calibration with muons (see Antonio's talk)

Strategy of time alignment among the subdetectors has to be studied







# **SPARES**

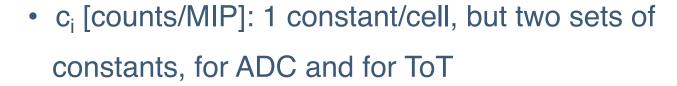




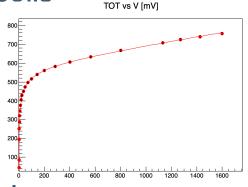
#### **ECAL** calibration constants

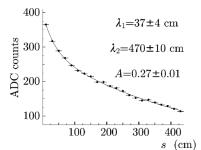
• ECAL: 4880 readout channels (PMTs), 2440 cells

- Pedestals: 1 constant/readout channel
- ToT calibration: 5<sup>th</sup> order polynomial,
   6 constants/readout channel
- Attenuation curve: 3 constants/readout channel



K: absolute energy scale (few values for the whole ECAL)











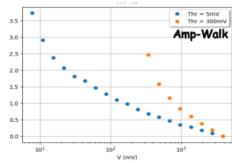
# **ECAL** calibration constants (II)

$$t_{A,B}$$
 [ns] =  $c_{A,B}$  [ns/(TDC count)] $(T_{A,B} - T_{A,B}^{(P)})$  [TDC counts]

- T<sub>P</sub>, time offsets due to cables, electronics: 1 constant/readout channel
- C<sub>A,B</sub>[ns/count]: 1 constant/readout channel
- Time-walk corrections: polynomial fit, few constants/readout

channel

- t<sub>0</sub>'s from MIPs: 1 constant/cell
- $\Delta t_0$ 's : 1 constant/cell
- Global T<sub>0</sub>: 1 value/event





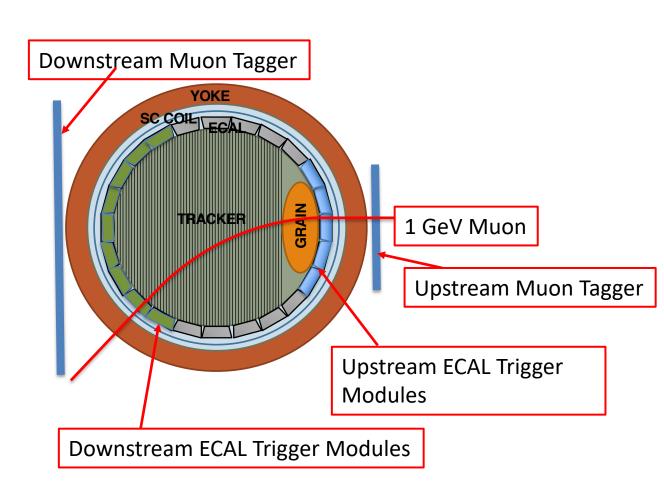




12

#### **Scheme**

- Upstream ECAL
  - In-time hits in 3 out of 4 layers
    - ~10ns window
  - Ignore outer (upstream)
     layer
- Downstream ECAL
  - In-time hits in 4 out of 5 layers
    - ~10ns window
- Coincidence
  - Upstream left and right
  - Downstream left and right
  - Pick out 100ns window for event-builder





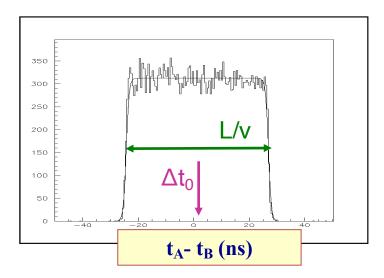




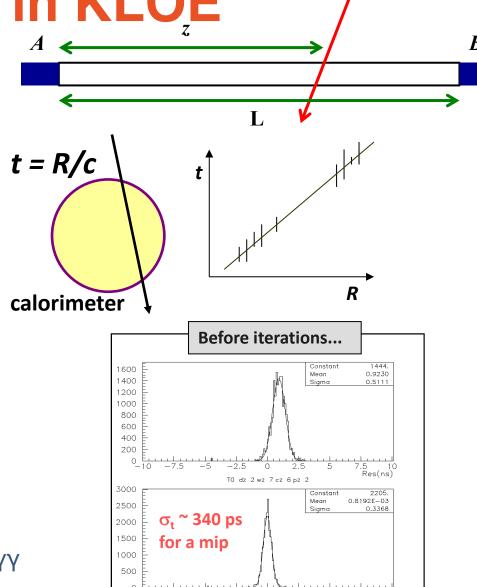
### Time calibration in KLOE

 MIPs: uniform illumination of the ECAL for time and z-coord. calibration

$$t = \frac{1}{2}(t_A + t_B) - \frac{L}{2v} - t_0 - t_G^0$$
$$z = \frac{1}{2}v(t_A - t_B) - \Delta t_0$$

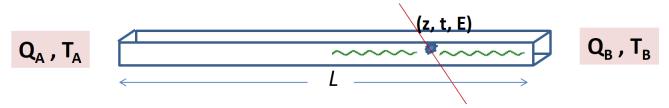


 Fine determination with Bhabha and γγ t<sub>0</sub>'s at ~ 20 ps



... After iterations

#### **Energy calibration in KLOE**



$$E_i^{(A,B)}[\text{MeV}] = \frac{(Q_i^{(A,B)} - P_i^{(A,B)})[\text{ADC counts}]}{C_i[\text{ADC counts/MIP}]} K \times f_{MIP2MeV}[\text{MeV/MIP}]$$

- $C_i$  = peak of the MIP distribution
- Corrections to the  $C_i$  with the Bhabha scattering events  $(e^+e^-\rightarrow e^+e^-)$ : showers of 510 MeV
- Absolute energy scale K fixed at cluster level with the  $e^+e^- \rightarrow \gamma\gamma$  events

$$\Rightarrow$$
 Calib. Const.  $=\frac{K}{C_i}$ 



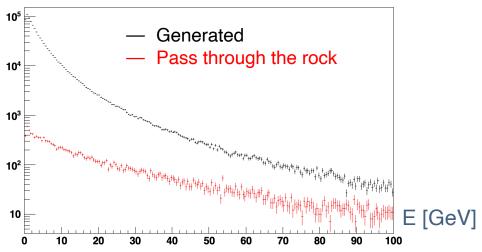




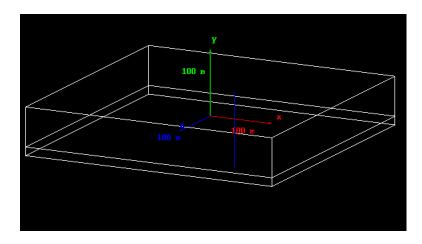
### Cosmic µ with GEANT4

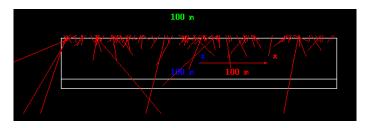
#### Very simple model:

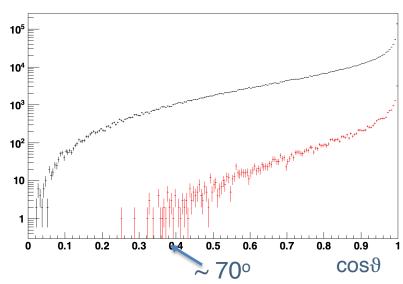
- Box of Standard Rock, 60 m thick (52% O, 27% Ca, 12% C, 9% Mg, density 2.65 g/cm<sup>3</sup>)
- Box of air 10 m thick
- 10<sup>6</sup> muons generated on top of the rock box



Events surviving	16965	1.7%
MIPs ( $p_{\mu} > 300 \text{ MeV}$ )	16583	1.7%
p <sub>μ</sub> > 1 GeV	15983	1.6%
Straight tracks (p <sub>µ</sub> > 6 GeV)	12228	1.2%







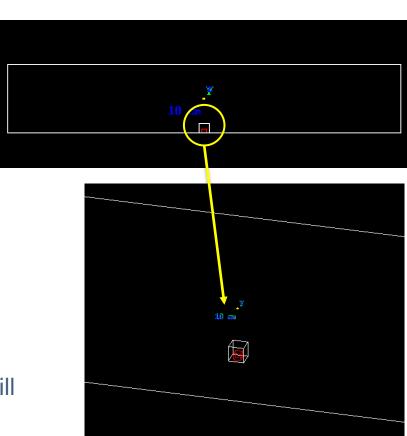






### Cosmic µ simulation with GEANT4

- Simplified model of the hall (a cube of air of 10 × 10 × 10 m<sup>3</sup> embedded in the rock)
- Red box representing the SAND:
   cube 4.5 × 4.5 × 4.5 m<sup>3</sup> inside the air cube
- Generated 10<sup>6</sup> events on a surface of 50 × 50 m<sup>2</sup>
- 95 events on SAND volume (~ 10<sup>-4</sup>)
- Rescaling the muon flux at the surface  $(\sim 0.02 \ \mu/(s \ cm^2)) \Rightarrow ~40 50 \ \mu/s \ on \ ECAL$
- To collect cosmics we need data out of the spill



Next step: Interface with Edepsim and SAND geometry







### Cosmic µ with GEANT4

#### Very simple model:

- Box of Standard Rock, 60 m thick (52% O, 27% Ca, 12% C, 9% Mg, density 2.65 g/cm<sup>3</sup>)
- Box of air 10 m thick
- Muons generated on top of the rock box

