



TECH-PFA PhD Program Admission to II year presentation

Multi-parametric thermal-mechanical modeling and analysis of high order active mirrors for space telescopes

Candidate: Chiara Scandaglia

Supervisors: Runa Briguglio, Marco Xompero

INAF - O.A. Arcetri

XL Cycle

Educational requirements

Attended courses

- Adaptive Optics for Astronomy I Dr. Carmelo Arcidiacono Passed
- Adaptive Optics for Astronomy II Dr. Kalyan Radhakrishnan Attended
- Fundamentals of system engineering and project management for large scientific projects Dr. Marco Xompero and Dr. Runa Briguglio Passed
- Coupled electrical-thermal-structural Finite Element Analyses Dr. Michele Ballan, Prof. Giovanni Meneghetti and Prof. Mattia Manzolaro *Passed*
- Methodologies and techniques for the analysis of experimental data Prof. Alexis Pompili *Attended*
- Random excitations and response of structures Prof. Giuseppe Petrone and Prof. Francesco Franco Ongoing lectures
- 12 total credits (10 credits required from educational obligations)
- Missing exams will be taken by the end of 2025

Educational requirements

Planned schools and conferences

- Workshop on Innovative Technologies for Space Optics (WITSO) 2025
 ESA/ESTEC, Noordwijk, The Netherlands 13-17 October 2025 Workshop Poster accepted
- SPIE Astronomical Telescopes + Instrumentation 2026 Copenhagen, Denmark 5-10 July 2026 Conference
- Observing with Adaptive Optics 2026 Observatoire de Haute-Provence, Saint-Michel-l'Observatoire, France - October 2026 -School

A **period of study and research abroad** during the second year is foreseen, but not yet defined. The **ESA OSIP program** is one of the proposed options.

Scientific context





REQUIREMENTS

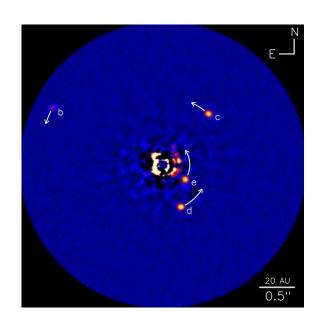


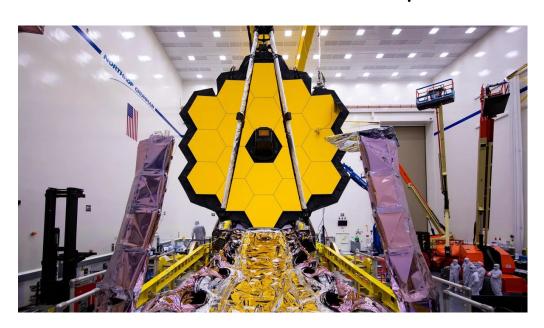
TECHNOLOGIES

- Earth-like exoplanets characterization
- Cosmological structures observation

- High resolution
- High contrast
- Ultra-stable wavefront
- Risks and costs mitigation

- Large primary apertures
- Deformable mirrors (w/ contactless actuators)
- Wavefront sensors
- Active optics control loops





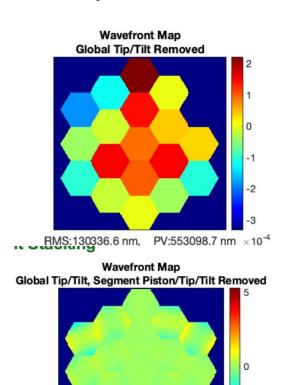
Why adaptive optics in space?

Features

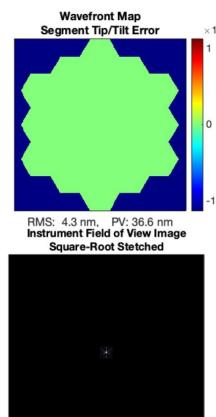
- Stable wavefront control
- Large primary aperture achievable
 - and correction directly on the primary
- Support vibration rejection through the use of particular actuators (e.g. voice coil)
 - Decoupling from the mirror
 - Use of lighter materials -> mass reduction
 - Looser tolerances
- Thermal load deformation correction
- **High contrast imaging** (e.g. if combined with coronagraph instruments)

Challenging topics

- New studies on adaptation of ground-based technologies in the space context
- New mission profiles taking into account deformable mirrors
- Cost effectiveness assessment



RMS: 24.2 nm, PV:317.6 nm



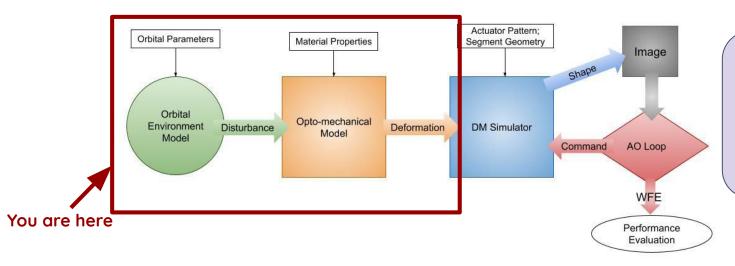
Main research activity An active mirror system parametric integrated model

TO

In what does it consists?

- Telescope opto-mechanical model
 - w/ variable mirror configuration parameters
- Thermal-mechanical simulator
 - w/ variable orbital environment parameters

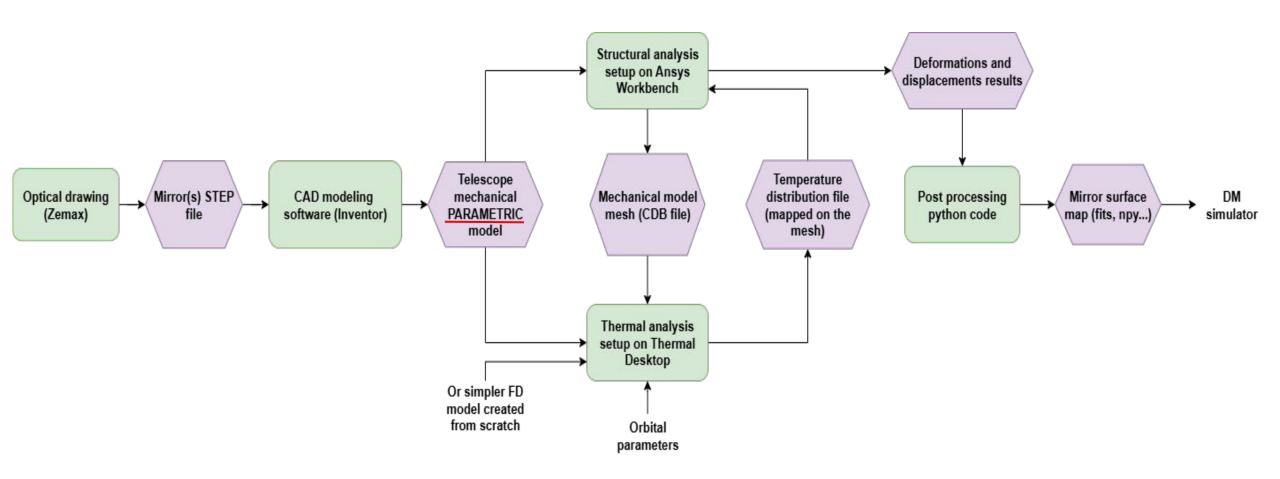
- Calculate deformations and displacements
- Compute wavefront correction
- Estimate **mirror performance** (fitting error)
 - Set trade-offs and optimize
 - mission profiles
 - mirror configurations
- Evaluate the system behaviour in different scenarios (i.e. costs)



Leading question: with regard to typical space mission profiles, costs and requirements, what could make this technology more attractive than others?

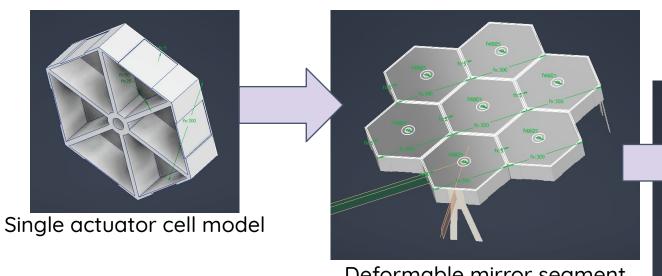
Part of the ULAOPS project, funded by ASI and in collaboration with PoliMi and industrial partners

Simulation pipeline



Simulation train meant to be as **automatized** as possible, to autonomously switch between the environmental condition and mirror configuration parameters and re-run the analysis

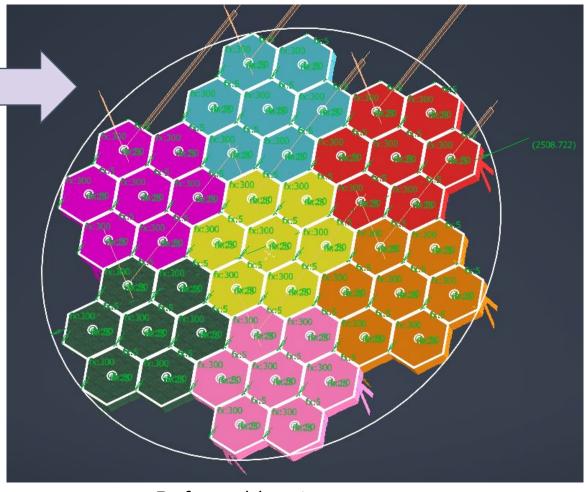
First implementations: structural parametric model



Deformable mirror segment (module) support model with 7 actuator cells

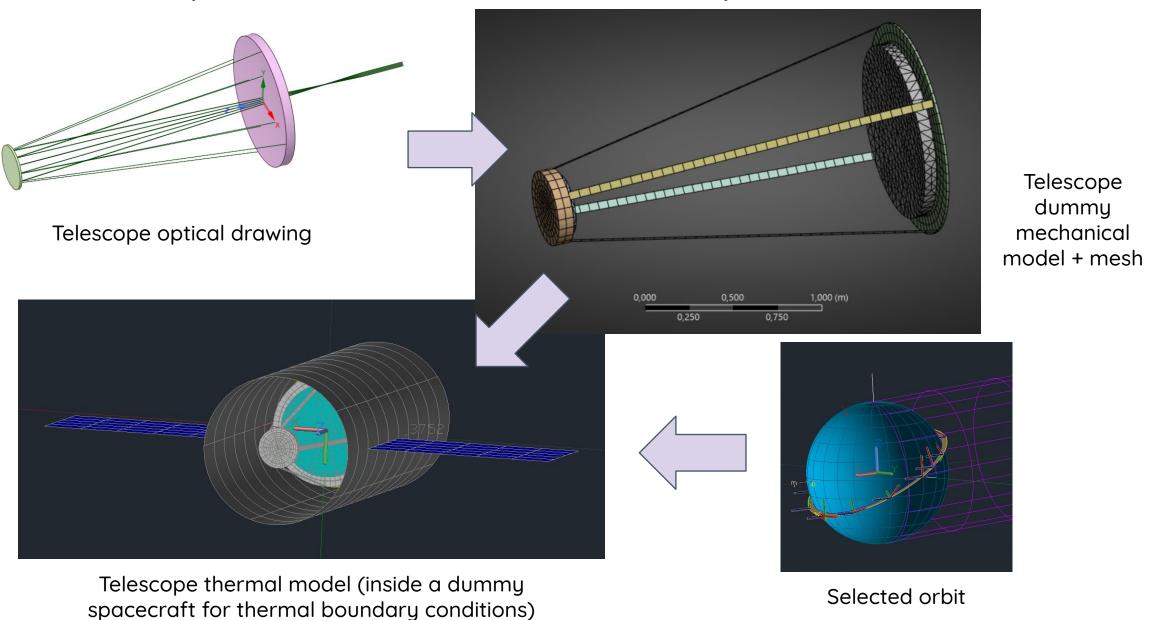
t	5	mm	parametri cellula
side	400	mm	6111 (a.e.)
hole	25	mm	
height	80	mm	
n_pat1	3	ul	parametri modulo
d_pat1	402,88	mm	
delt_T	100		
CTE	2,4E-05	mm	
bipod_h	200	mm	structure
bipod_t	2	mm	
bipod_a	80	deg	

Configuration parameters

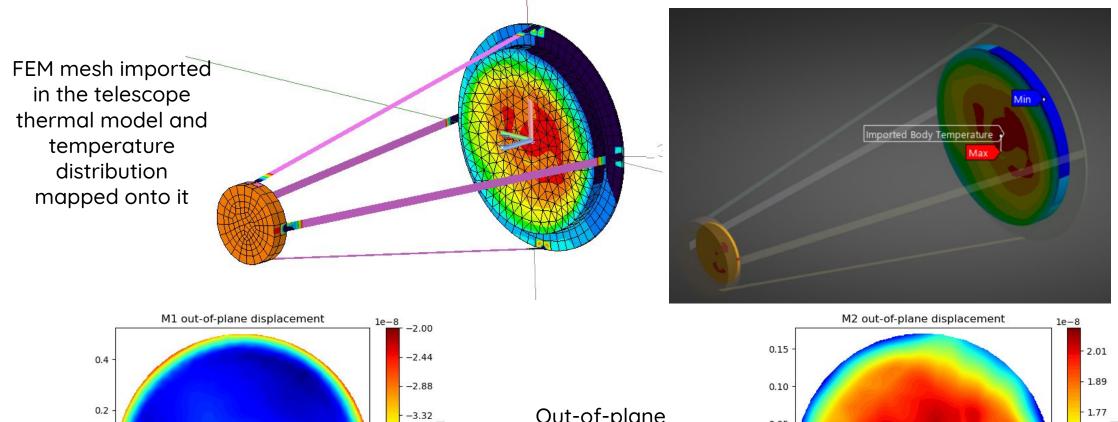


Deformable mirror support model with 7 modules

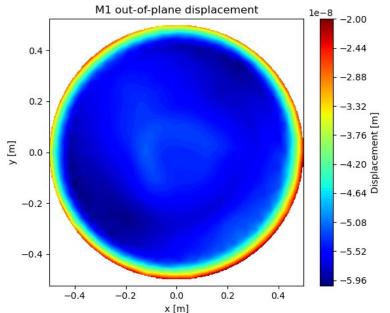
First implementations: thermal inputs on FEM model



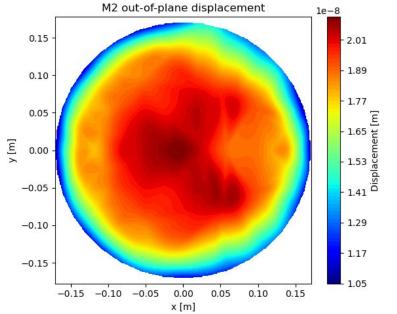
First implementations: thermal inputs on FEM model



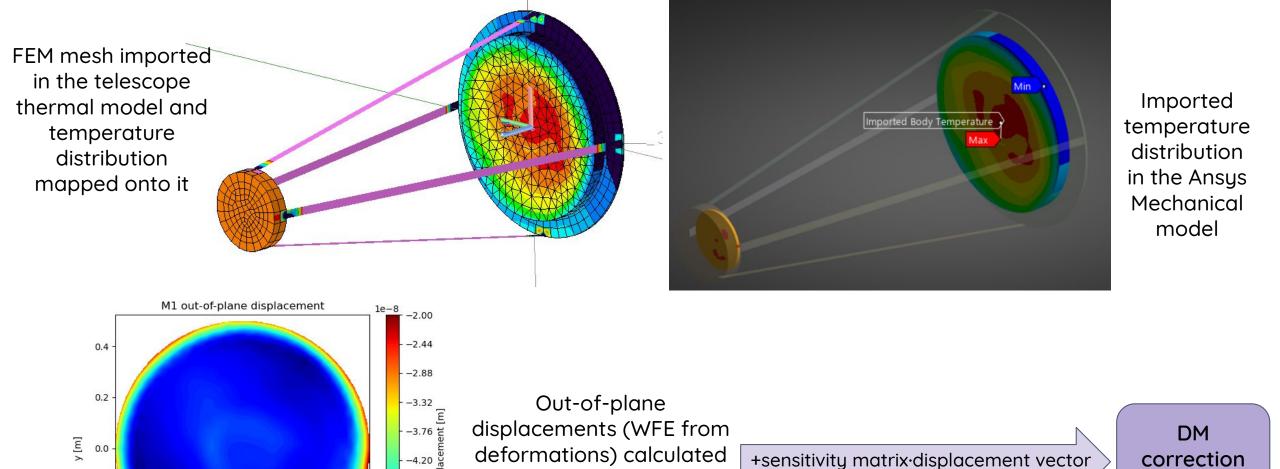
Imported temperature distribution in the Ansys Mechanical model



Out-of-plane displacements (WFE from deformations) calculated by Ansys and exported as python maps



First implementations: thermal inputs on FEM model



by Ansys and exported as

python maps

-5.08

-5.52

-5.96

0.2

x [m]

0.4

-0.2

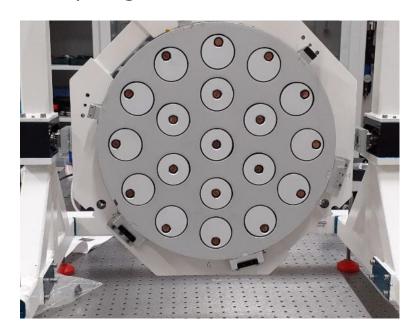
-0.4

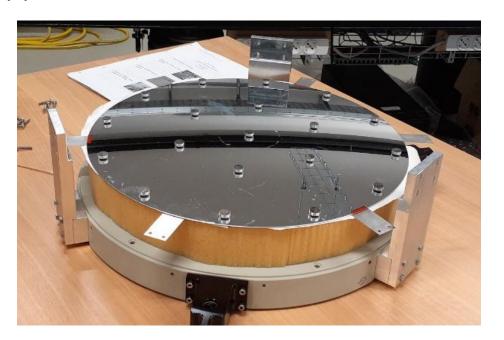
-0.4

loop

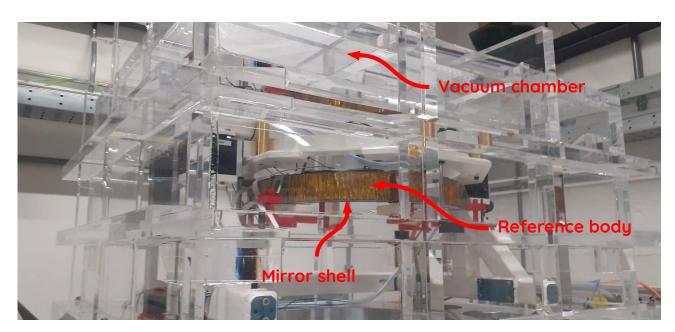
Other activities: introducing LATT

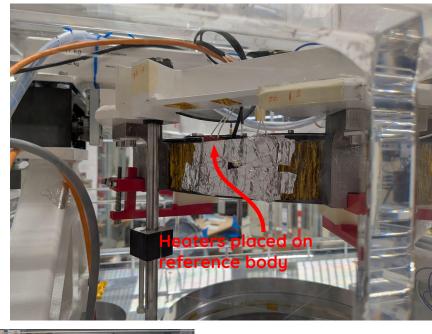
- Deformable segment of a primary mirror prototype
- 40 cm thin Zerodur deformable shell (mirror)
- Aluminium honeycomb reference body (support)
- Low mass density
- 19 voice-coil actuators and 19 capacitive sensors (contactless technology)
 - When set, the shell floats at a given gap from the reference body
 - Decoupling between mirror and support!

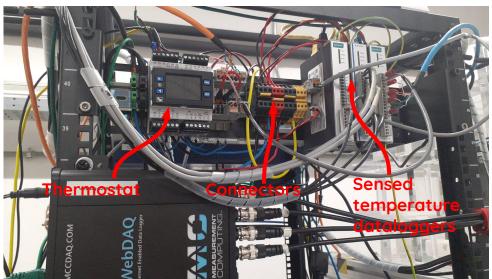


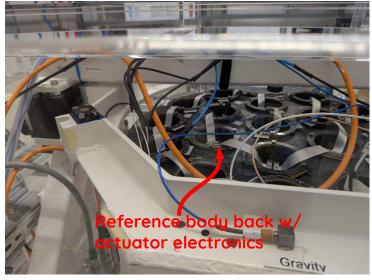


Other activities: LATT thermal tests



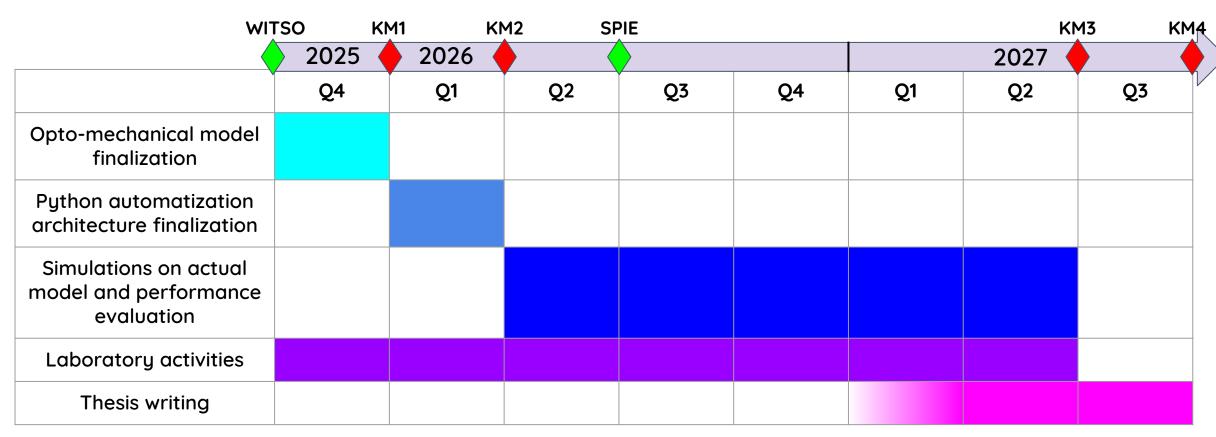






Goal: assess effects of reference body deformations and resulting actuator commands on mirror correction

Roadmap for the next years



MILESTONES

- WITSO @ ESA Simulation scheme finalized
- KM1 Model completed
- KM2 Architecture code completed

- SPIE @ Copenhagen Paper ready
- KM3 Activities completed
- KM4 Thesis completed

THANK YOU FOR YOUR ATTENTION!