



Istituto Nazionale di Fisica Nucleare
SEZIONE DI ROMA TOR VERGATA



UNIVERSITÀ
DEGLI STUDI
DI PADOVA

PhD Research and Development Plan: 2nd Year Achievements and Next Year's Objectives

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Overview



AMS purpose is to perform high precision, high statistics, and log measurement of charged cosmic rays in .5GeV to 1.5 TeV

AMS 02 Launch Date : 16th May 2011 by NASA

400 km above the ground

3m*4m*6m in volume

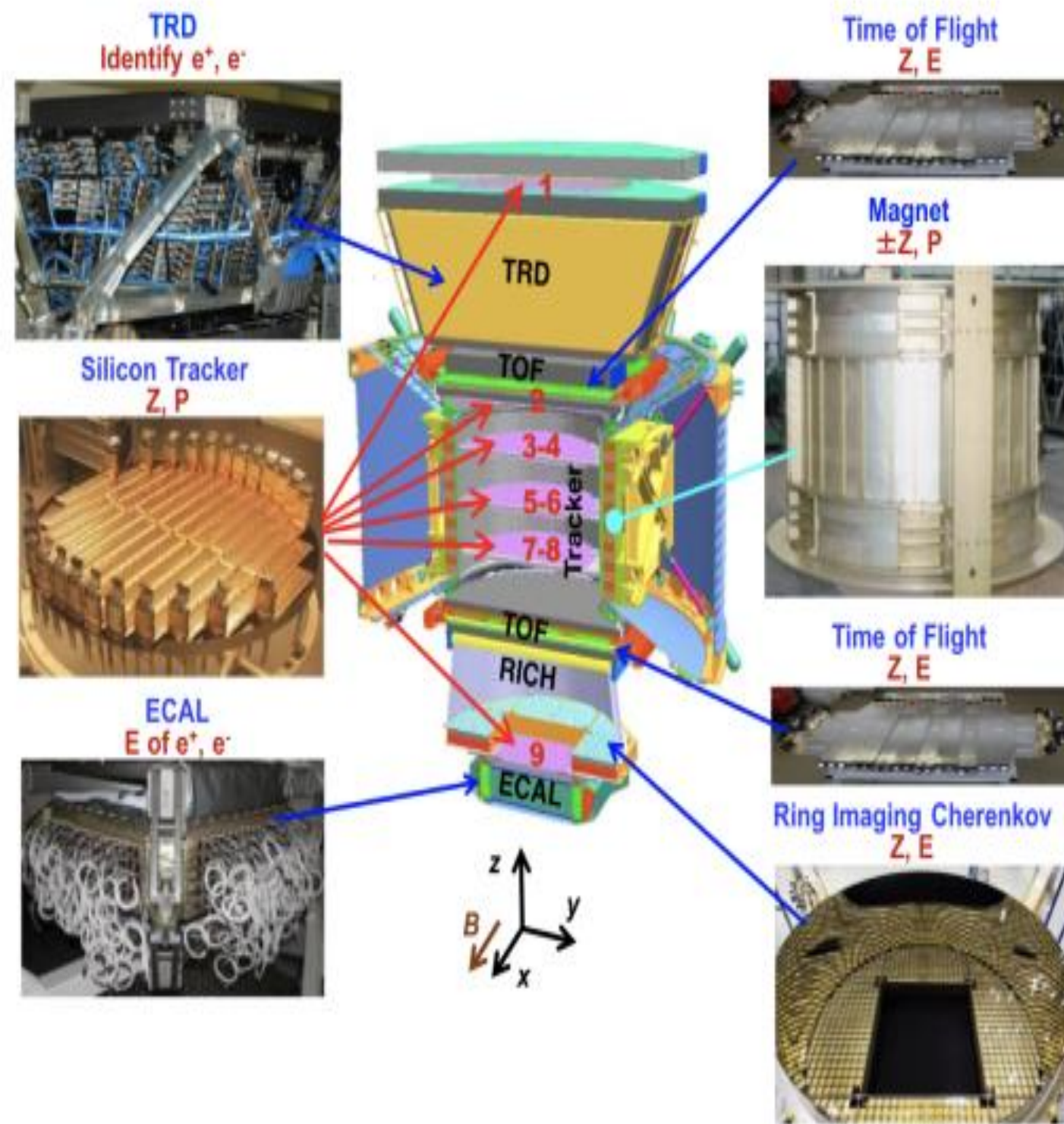


AMS 02 Collaboration is from 13 countries and 43 institutions.

AMS Scientific Goals are :

- Search for cosmic matter
- Search for Dark matter signatures
- Measurement of cosmic ray spectra, relative abundances, and isotopes
- Search for new form of matter
- Solar Activity and more

AMS 02 Sub-Detectors

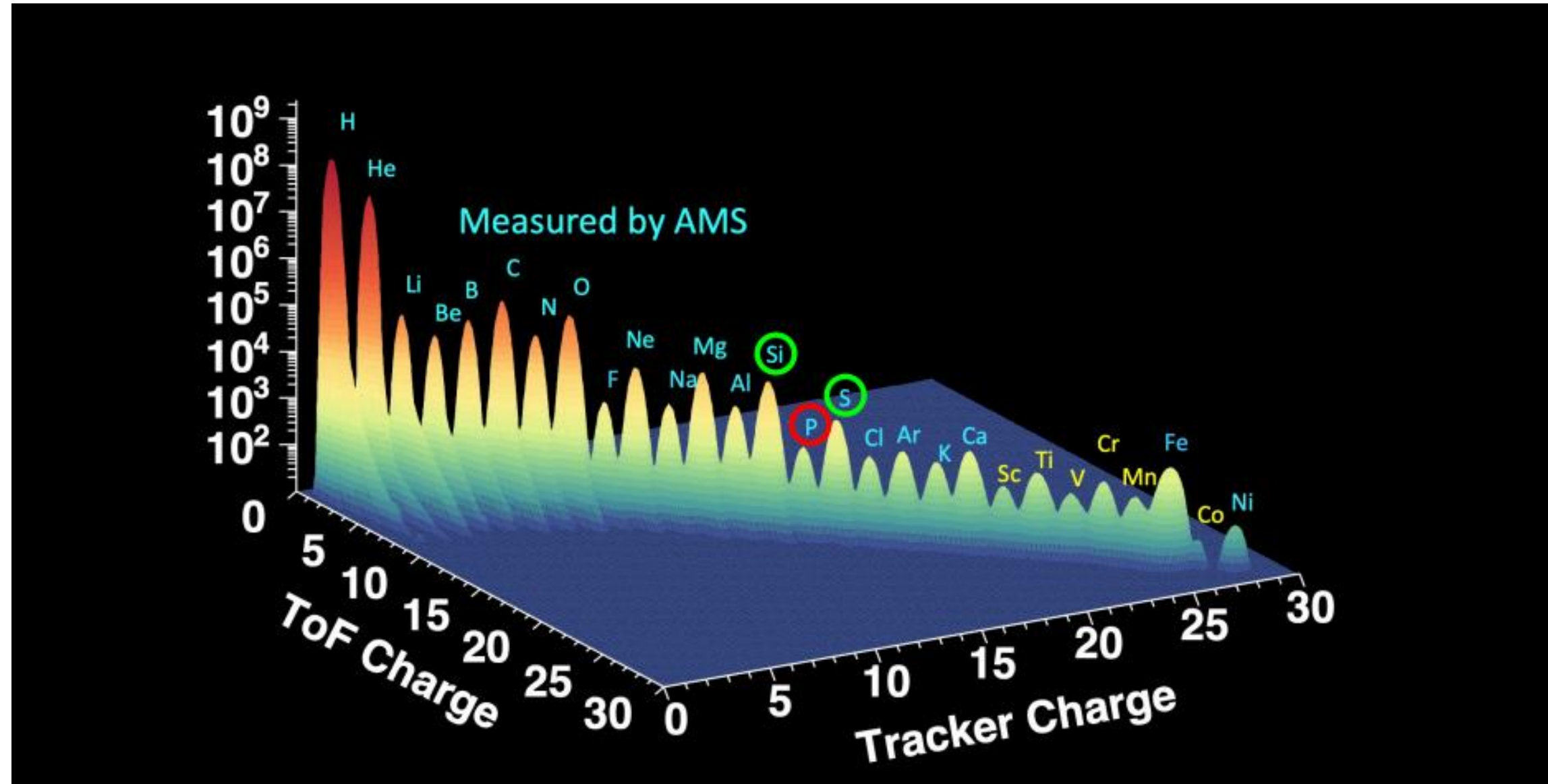


Schematic view of AMS 02 Detector

- Silicon Tracker & Magnet
- Transition Radiation Detector (TRD)
- Time Of Flight (TOF)
- Electromagnetic Calorimeter (ECAL)
- Ring Imaging Cherenkov (RICH)
- Anti-Coincidence Counters (ACC)

AMS 02 Data Analysis Pipeline

Phosphorus Flux



- Compared to nearby nuclei, phosphorus has a very low relative abundance
- Study and reproduce silicon ($Z=14$) and sulfur ($Z=16$) fluxes in order to proceed to phosphorus ($Z=15$)



AMS 02 Data Analysis Pipeline

Flux Formula:

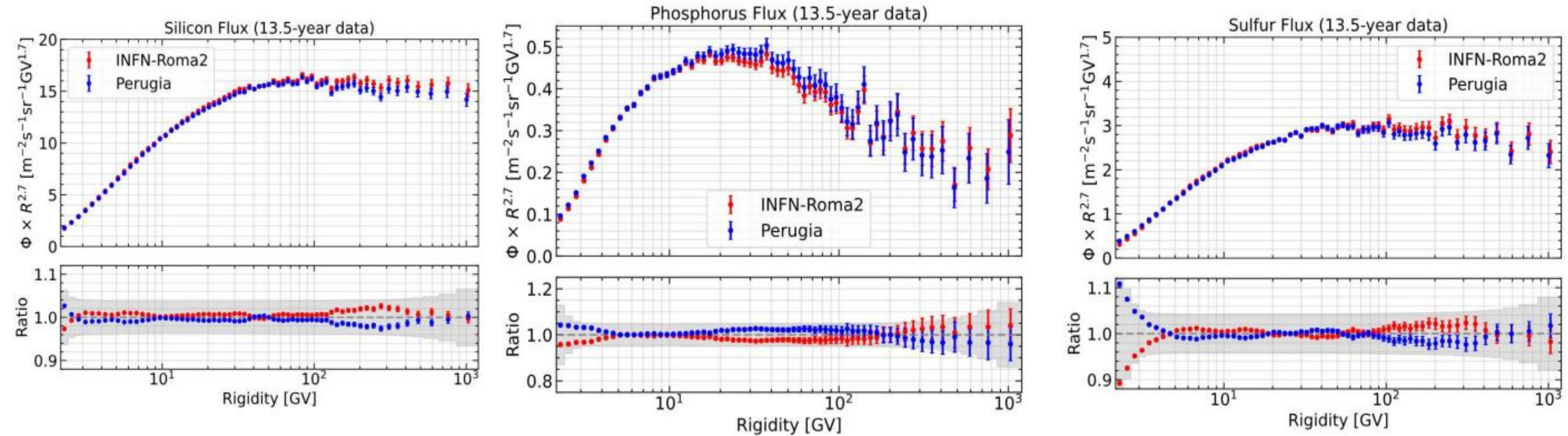
$$\Phi(R) = \frac{\Delta N}{A(R) \cdot T(R) \cdot \Delta R}$$

- **The Flux Equation:** The final flux depends on the raw counts (ΔN) and a large denominator representing the effective measurement area ($A(R) \cdot \Delta T(R) \cdot \Delta R$).
- **Exposure Time ($T(R)$):** "We meticulously track 13.5 years of good data, filtering by hardware status and orbital position to ensure every event is counted under nominal conditions."
- For the **Acceptance (A)**. We start with the theoretical **Raw Acceptance (A_{raw})** from our GEANT4 MC simulation. However, the real detector is imperfect.
- To account for this, the core of the pipeline is the **Efficiencies Step**. We calculate the **K Correction Factor** for every subsystem: $K = MC^{Efficiency} / Data^{Efficiency}$. This K factor is needed for all the Efficiencies (ToF, Trigger, Tracker etc)
- Effective Acceptance is achieved by running the **Efficiencies pipeline** to derive the unitless **K correction factors** (which quantify the difference between Data and MC efficiency for every detector step) and multiplying the Raw Acceptance by these factors.



AMS 02 Data Analysis Pipeline

Phosphorus Flux



- Comparison of Si-14, S-16 and P-15 flux with AMS-Italia Perugia group
- Study and reproduce silicon (Z=14) and sulfur (Z=16) fluxes in order to proceed to phosphorus (Z=15)



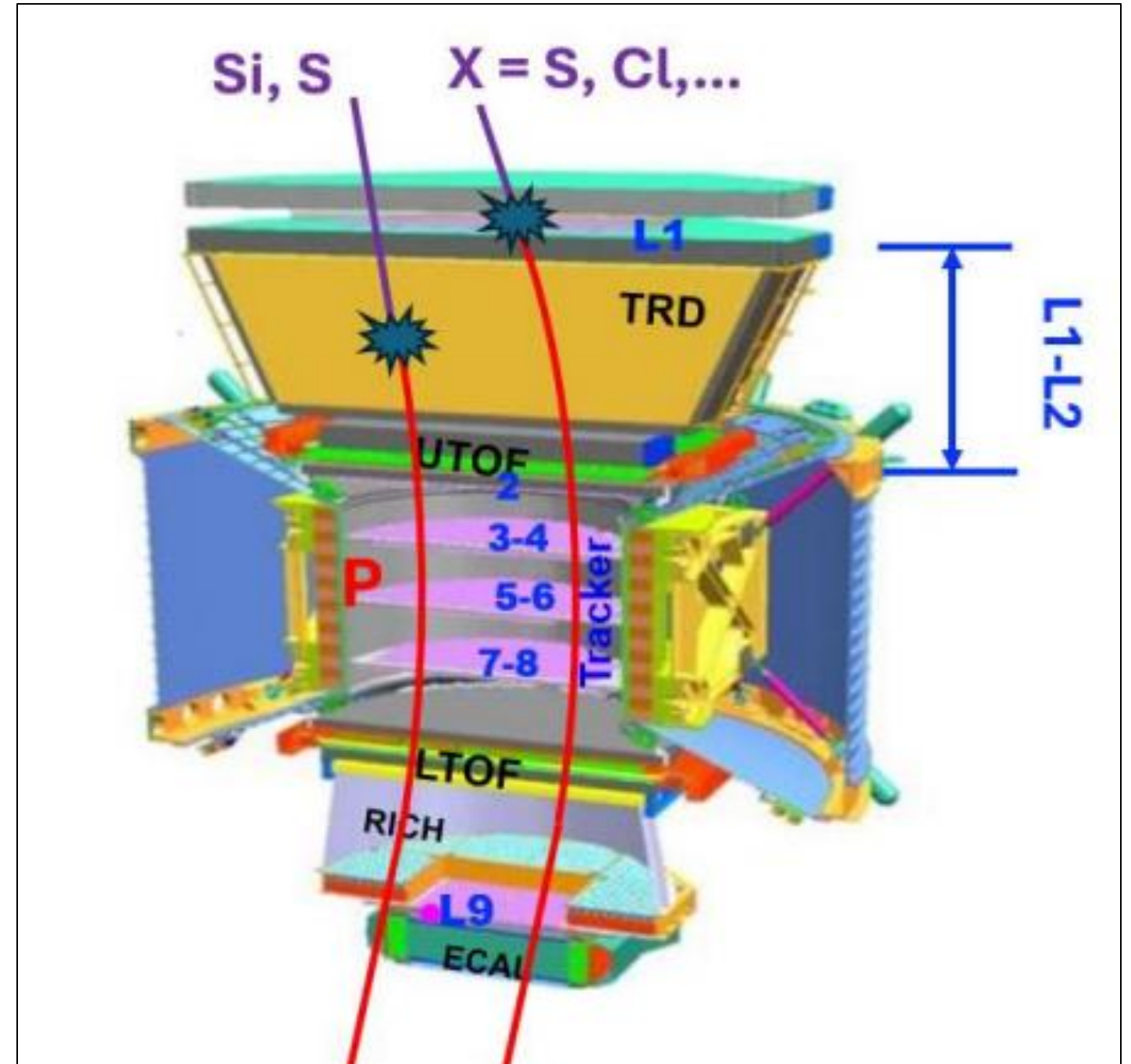
Contamination

Below L1 Contamination

- **Source:** Heavier nuclei fragment in the material **between Tracker L1 and L2** (TRD/Upper TOF).
- **Method:** **Data-Driven Template Fit** is used.

Content: Above L1 Contamination (TOI)

- **Source:** Heavier nuclei fragment in the **material above L1** (thin support structures made of carbon fiber and aluminum honeycomb). L1 cannot measure this composition.
- **Method:** **Monte Carlo (MC) Simulation** is mandatory.



Multimodal Domain Adversarial Network (M-DAN)

Motivation: Reduce P background after P event selection.

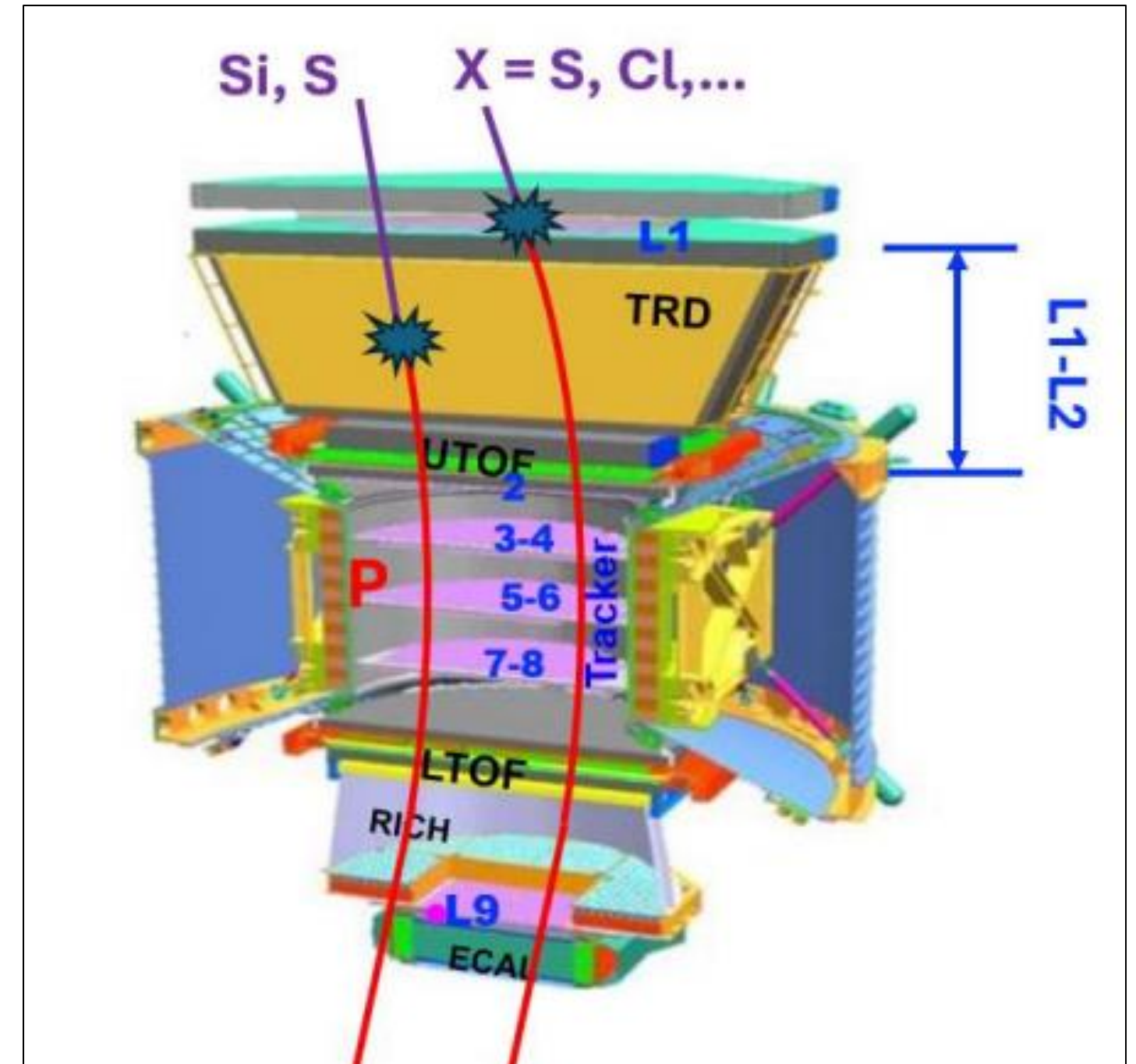
Pre-Processing

- Combining low-level, physics-relevant variables from distinct AMS-02 subsystems e.g ToF, tracker, to enhance signal separation.
- Use variables that have similar distributions in both MC and Data.

Trained from MC sample Signal

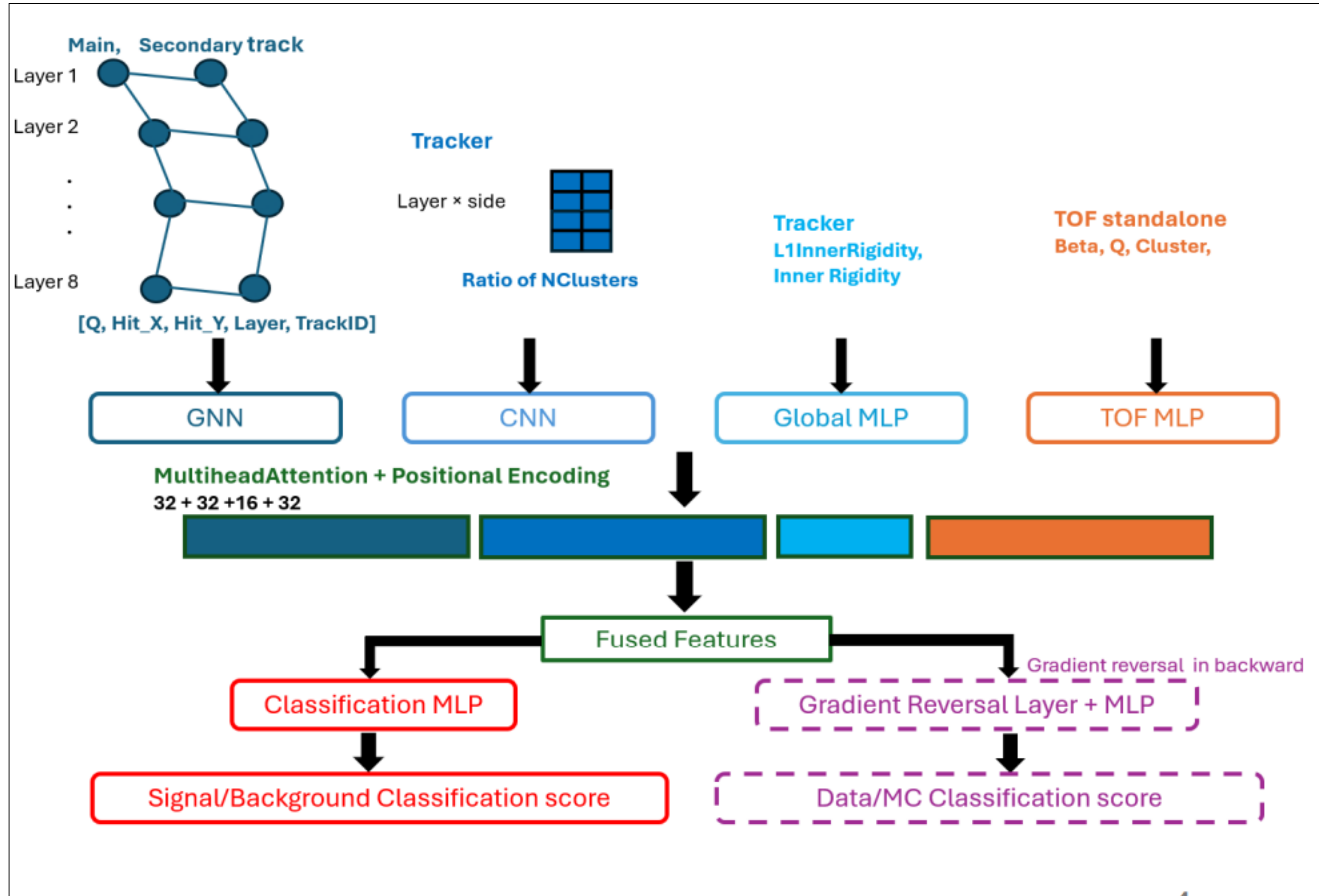
- MC Phosphorus
- Background: MC Si and S
- Data

Use domain-adversarial training to learn domain-invariant features, to achieve model generalization across MC and Data



Multimodal Domain Adversarial Network (M-DAN)

- MDAN Architecture:
The Key to Robustness is Adversarial Training.
- The network has two competing goals: maximize signal separation (Classification) while minimizing the distinction between simulated and real events (Adversarial).



STAGE 1: INPUT - Multi-Modal Features

SINGLE PARTICLE EVENT (Z-Nucleus)



GNN

[N × 5]

Graph Topology

- Charge, X, Y
- Layer, Track ID

Captures:

Spatial relationships & trajectory

CNN

[8 × 2]

Layer Sequence

- Charge ratio
- N clusters/layer

Captures:

Longitudinal energy pattern

GLOBAL

[4]

Rigidity

- Rigidity_{IL1} (x,y)
- Rigidity_{ln} (x,y)

Captures:

Momentum/Charge ratio

TOF

[7]

Timing

- Beta (v/c)
- Charge, clusters

Captures:

Mass via timing

Different detectors provide complementary physics information

STAGE 2: BRANCH PROCESSING

GNN

GCNConv (5 → 32)
Local patterns

Dropout 10%

GCNConv (32 → 32)
2-hop patterns

Global Mean Pool

Output: [32]

CNN

Conv1d (2 → 16, k=3)
Layer patterns

BatchNorm + ReLU

Conv1d (16 → 32, k=3)
Higher patterns

AdaptiveAvgPool1d

Output: [32]

GLOBAL

Linear (4 → 16)

ReLU

BatchNorm

Dropout 10%

Output: [16]

TOF

Linear (7 → 32)

ReLU

BatchNorm

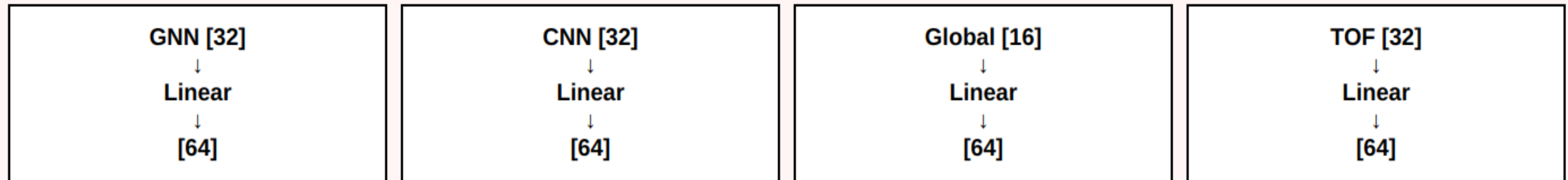
Dropout 10%

Output: [32]

Four specialized networks extract complementary representations

STAGE 3: ATTENTION FUSION

Step A: Projection to 64-d



Step B: Add Positional Encoding

$[64] + \text{Pos}_0 \rightarrow \text{GNN [64]}$
 $[64] + \text{Pos}_1 \rightarrow \text{CNN [64]}$
 $[64] + \text{Pos}_2 \rightarrow \text{Global [64]}$
 $[64] + \text{Pos}_3 \rightarrow \text{TOF [64]}$

Stack: [batch, 4, 64]



Step C: Multi-Head Attention (4 heads)

Q, K, V = Linear projections
Attention = softmax($Q \cdot K^T / \sqrt{64}$)
Output = Attention $\times$ V

	GNN	CNN	Global	TOF
GNN	0.45	0.25	0.15	0.15
CNN	0.20	0.50	0.20	0.10
Global	0.30	0.15	0.40	0.15
TOF	0.10	0.25	0.20	0.45



Step D: Mean Pooling

FUSED FEATURES [batch, 64]

Attention learns dynamic weighting - not equal mixing

STAGE 4: DUAL CLASSIFICATION HEADS

Classification Head

Linear (64 → 32) + ReLU

Dropout 10%

Linear (32 → 16) + ReLU

Dropout 10%

Linear (16 → 1)

Sigmoid → [0,1]

>0.5 = Signal
<0.5 = Background

Domain Adversarial Head

Gradient Reversal Layer

$\lambda: 0.0 \rightarrow 1.0$

Forward: x

Backward: $-\lambda \cdot \text{grad}$

Linear (64 → 128) + ReLU + BN

Linear (128 → 64) + ReLU + BN

Linear (64 → 32) + ReLU

Linear (32 → 1)

Sigmoid → [0,1]

Goal: ~50% Accuracy
(Domain-invariant)

GRL creates adversarial game: Classify well BUT domain-invariant features

$$\text{Total Loss} = \text{Classification Loss} + \text{Domain Loss}$$

Classification Loss

$\text{BCE}(\text{pred}, \text{true_label})$

Example: $\text{BCE}(0.73, 1.0) = 0.31$

Goal: Minimize

Target: High accuracy (>80%)

Domain Loss

$\text{BCE}(\text{pred}, \text{domain_label})$

Example: $\text{BCE}(0.51, 0/1) \approx 0.69$

Goal: ~ 0.69 (random)

Target: $\sim 50\%$ accuracy

↓ BACKWARD PASS ↓

Classification Gradient

"Make features discriminative"

↓
Normal backprop

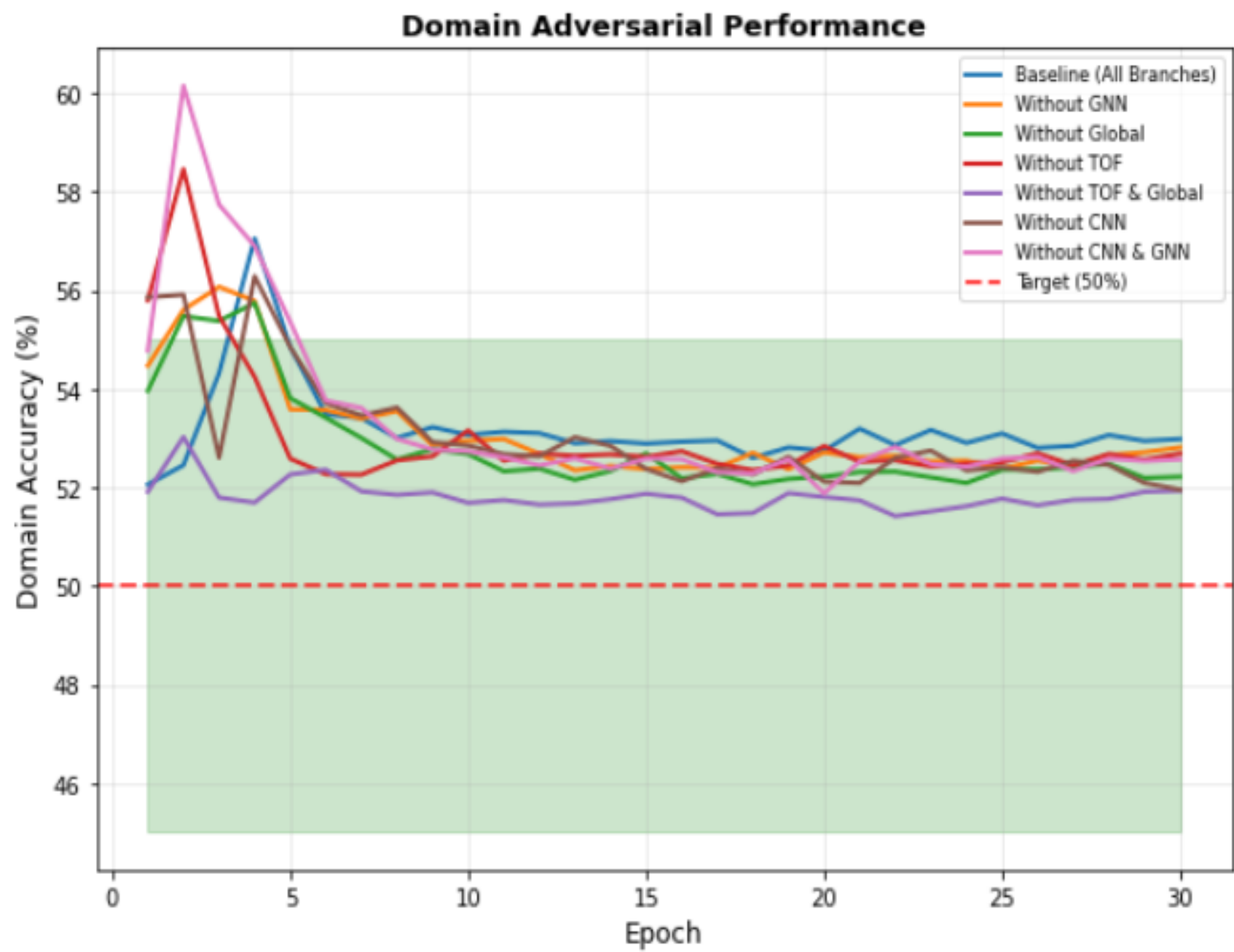
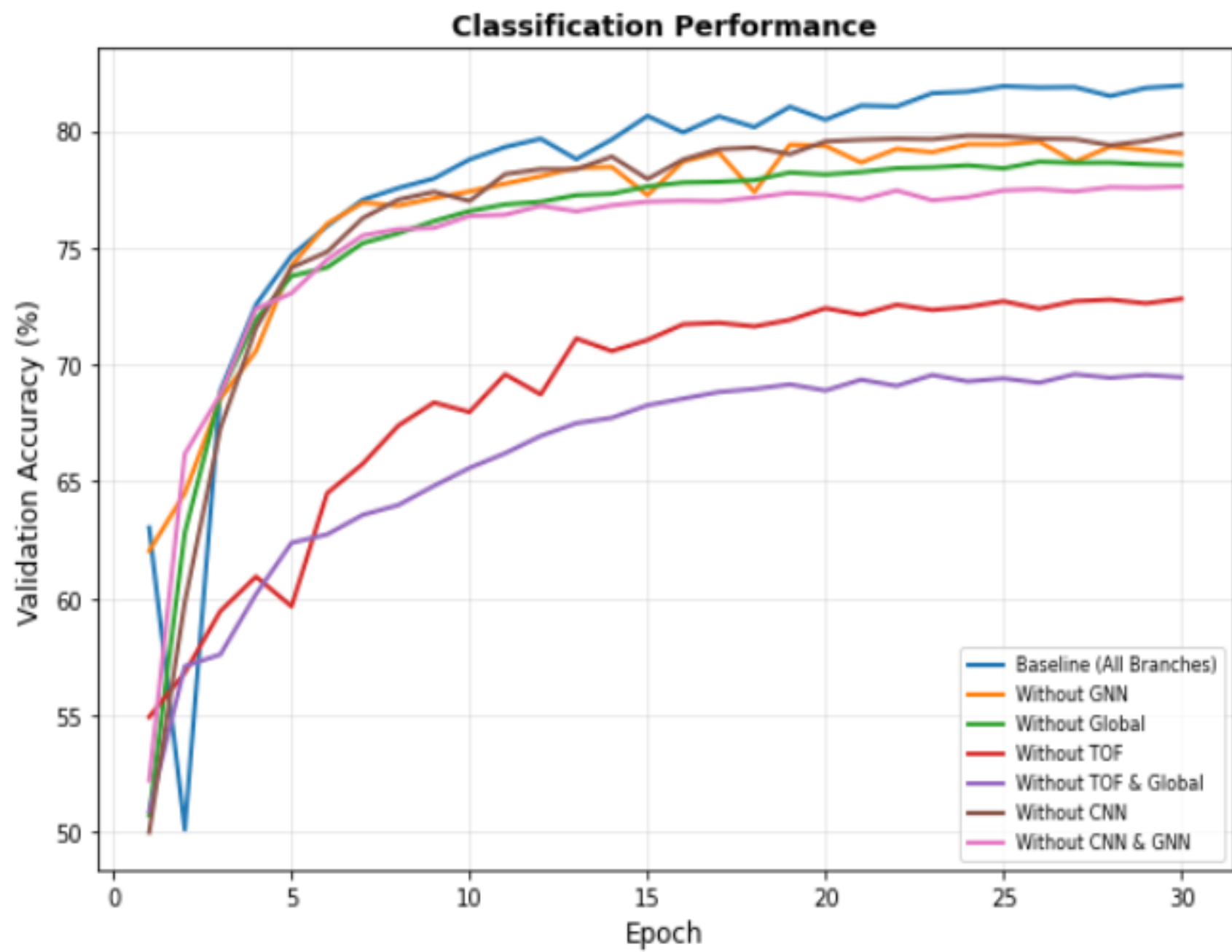
Domain Gradient (via GRL)

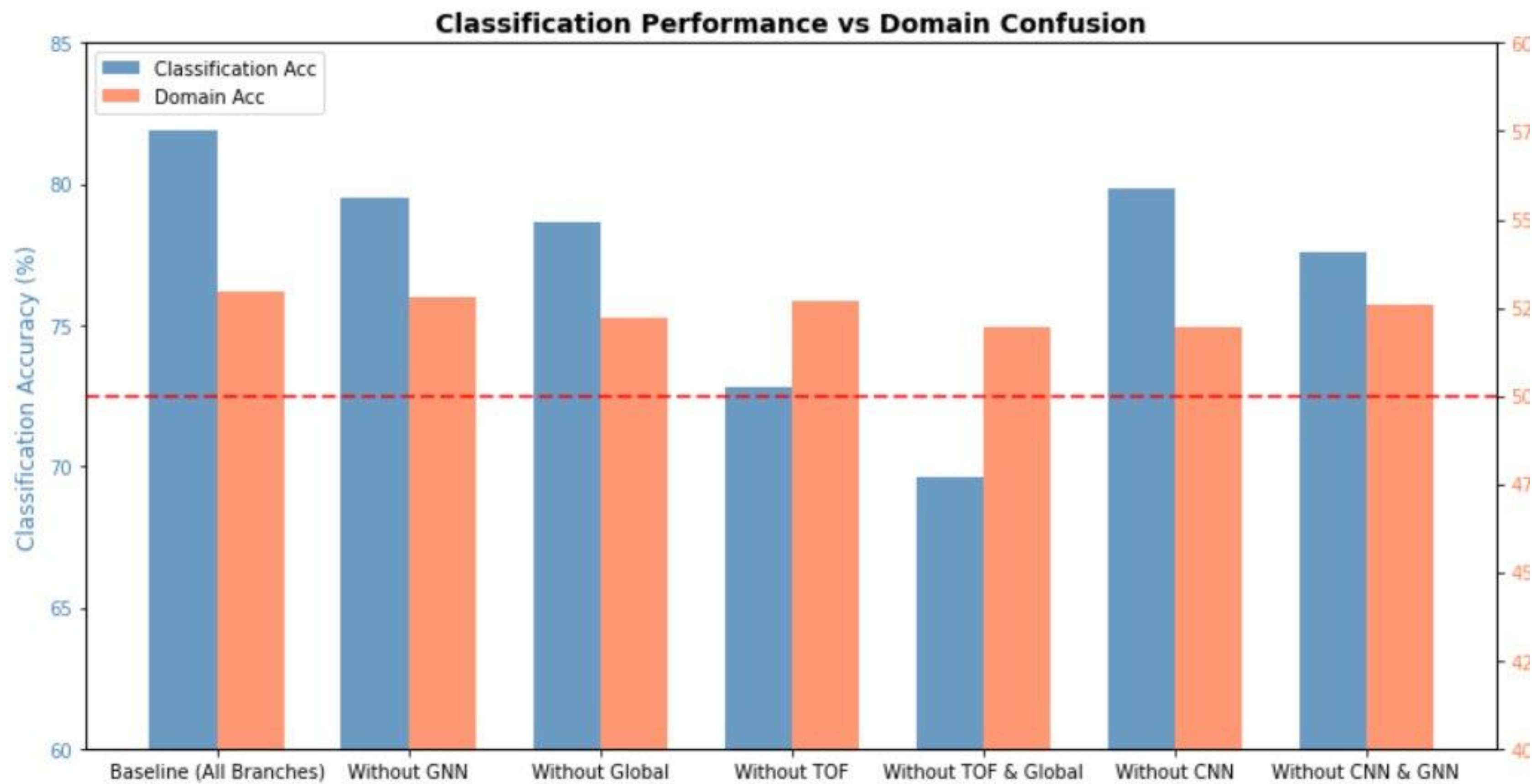
"Make features domain-invariant"

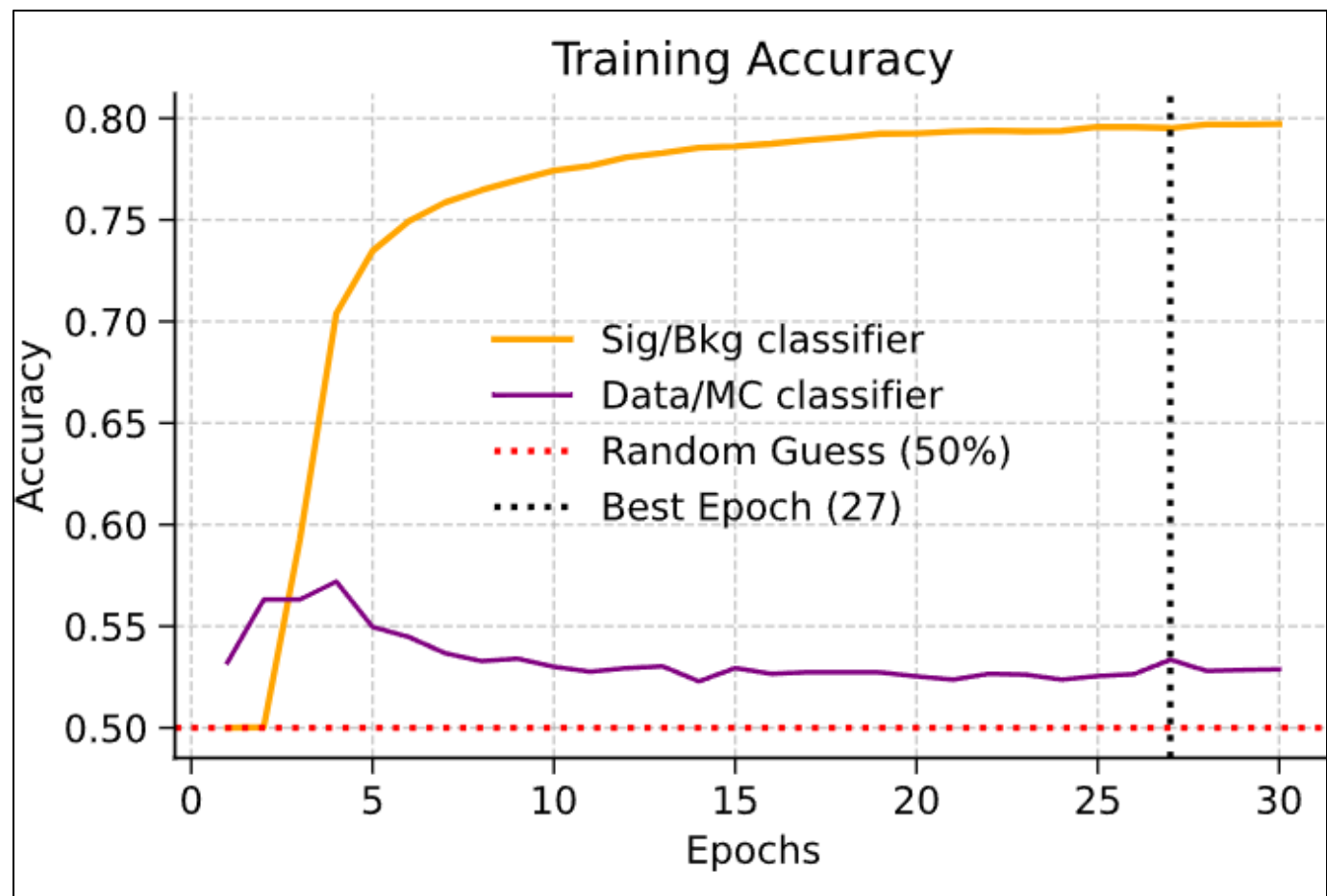
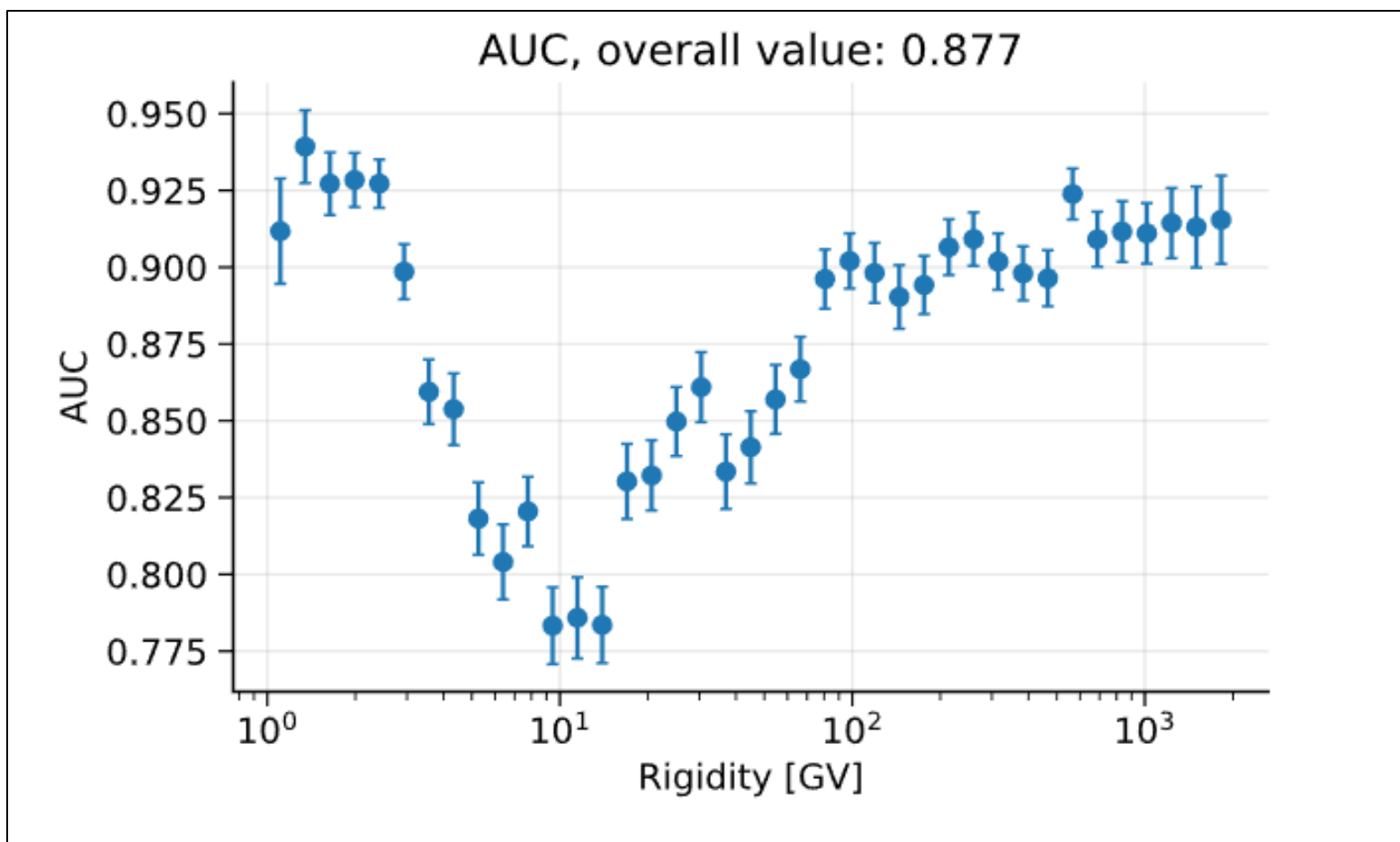
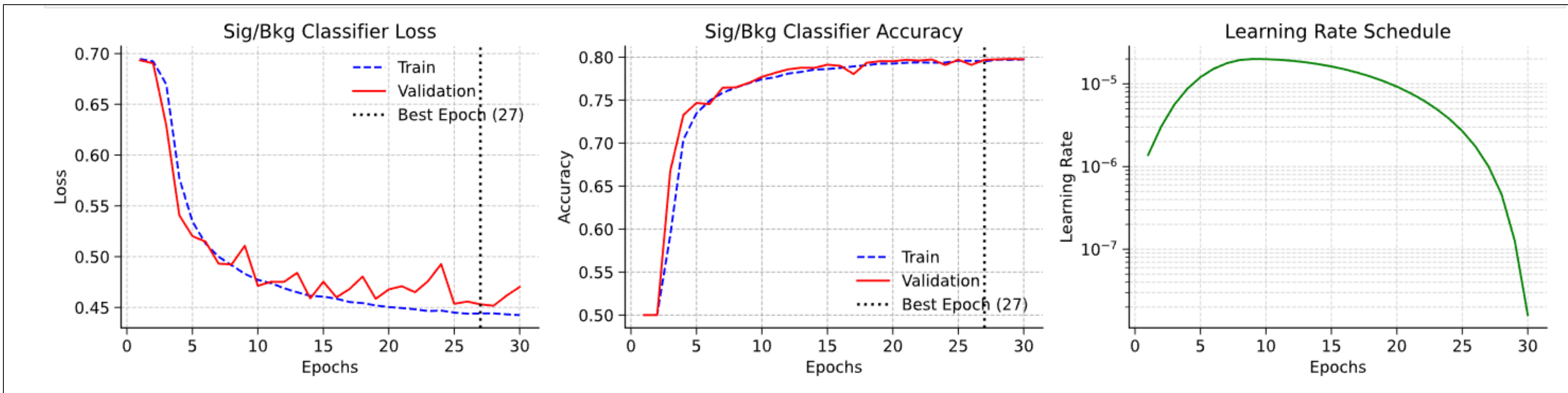
↓
Reversed: $\times(-\lambda)$

↓
FUSED FEATURES receive BOTH signals
Result: Discriminative yet domain-invariant

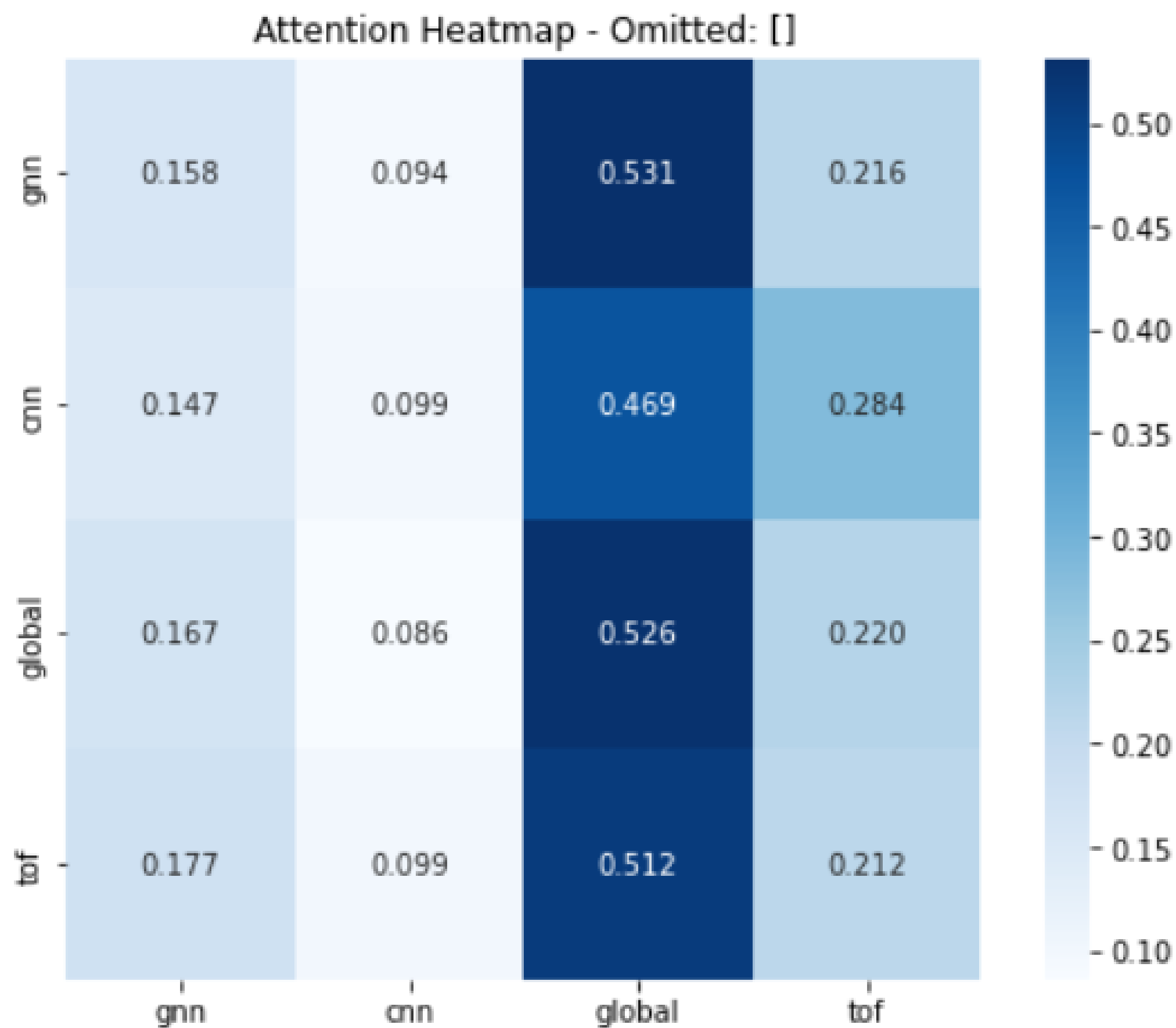
Equilibrium: High classification accuracy + $\sim 50\%$ domain accuracy



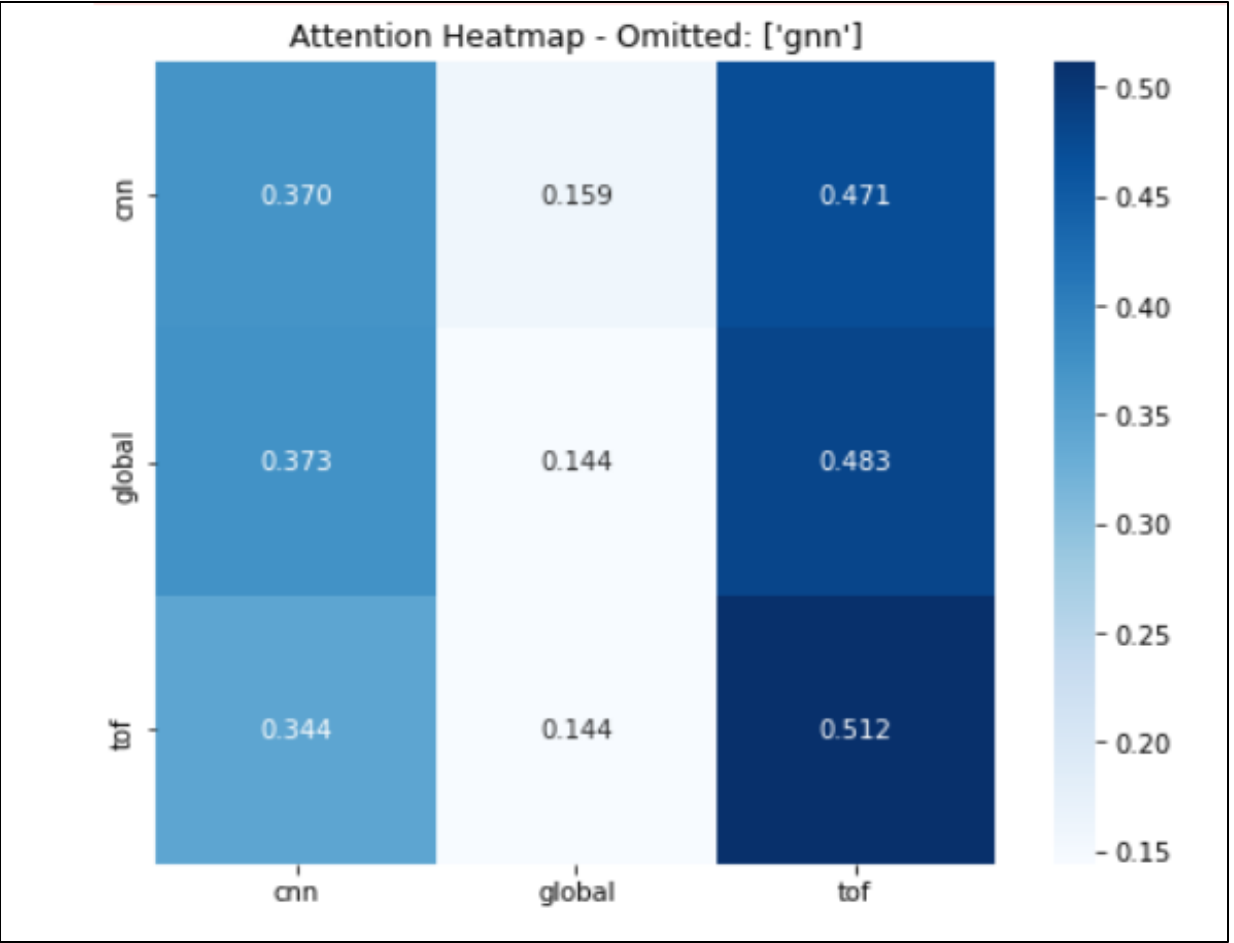
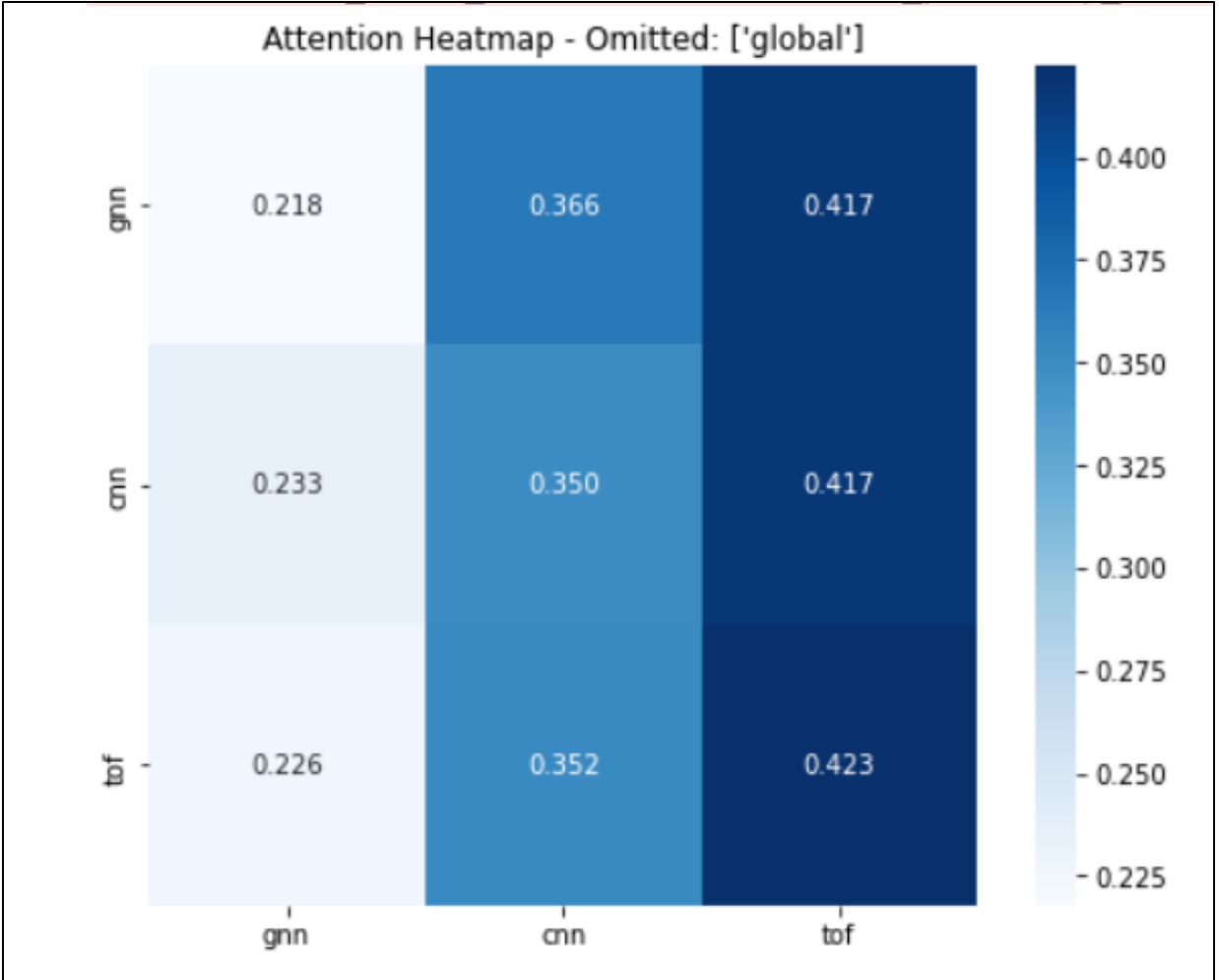
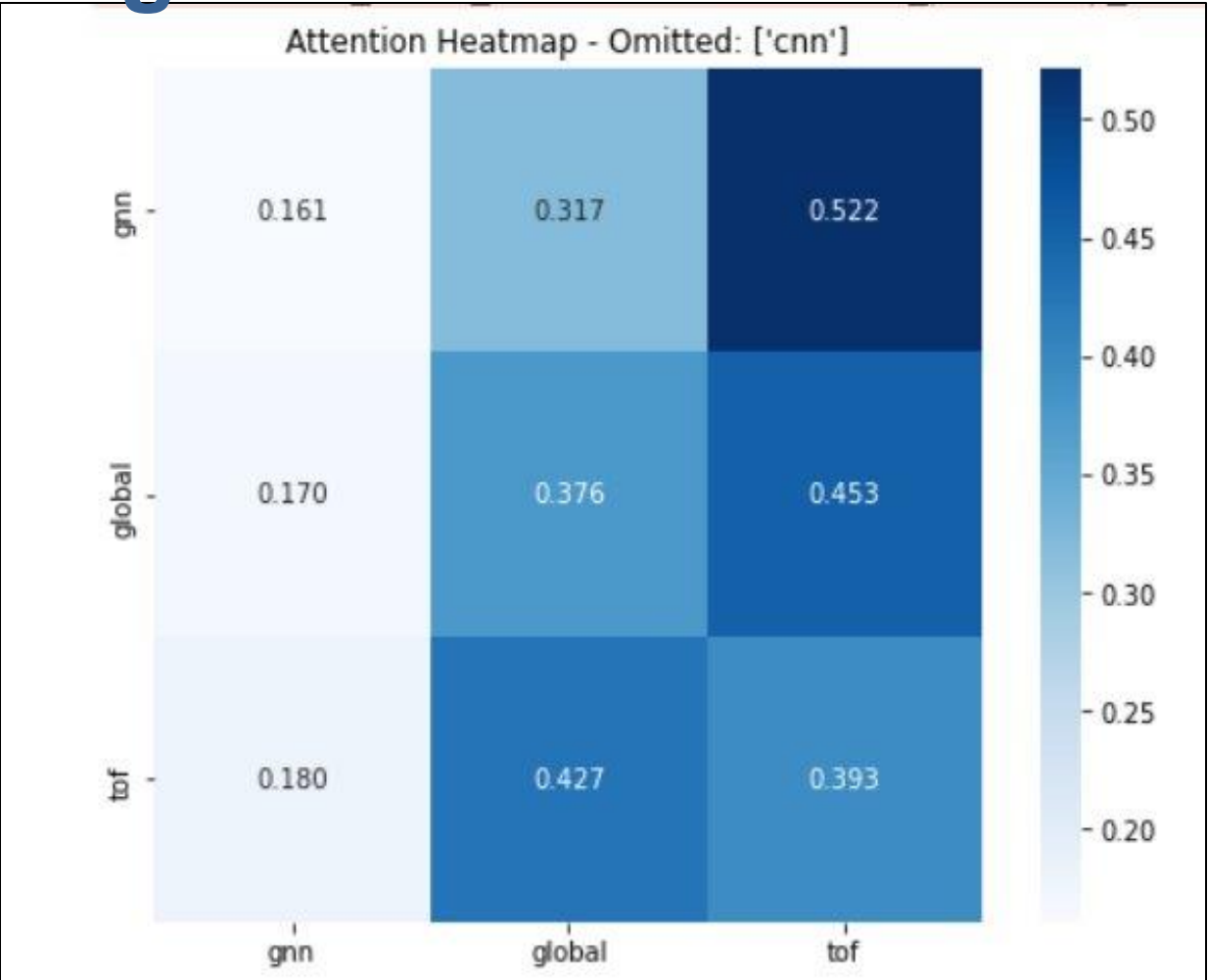
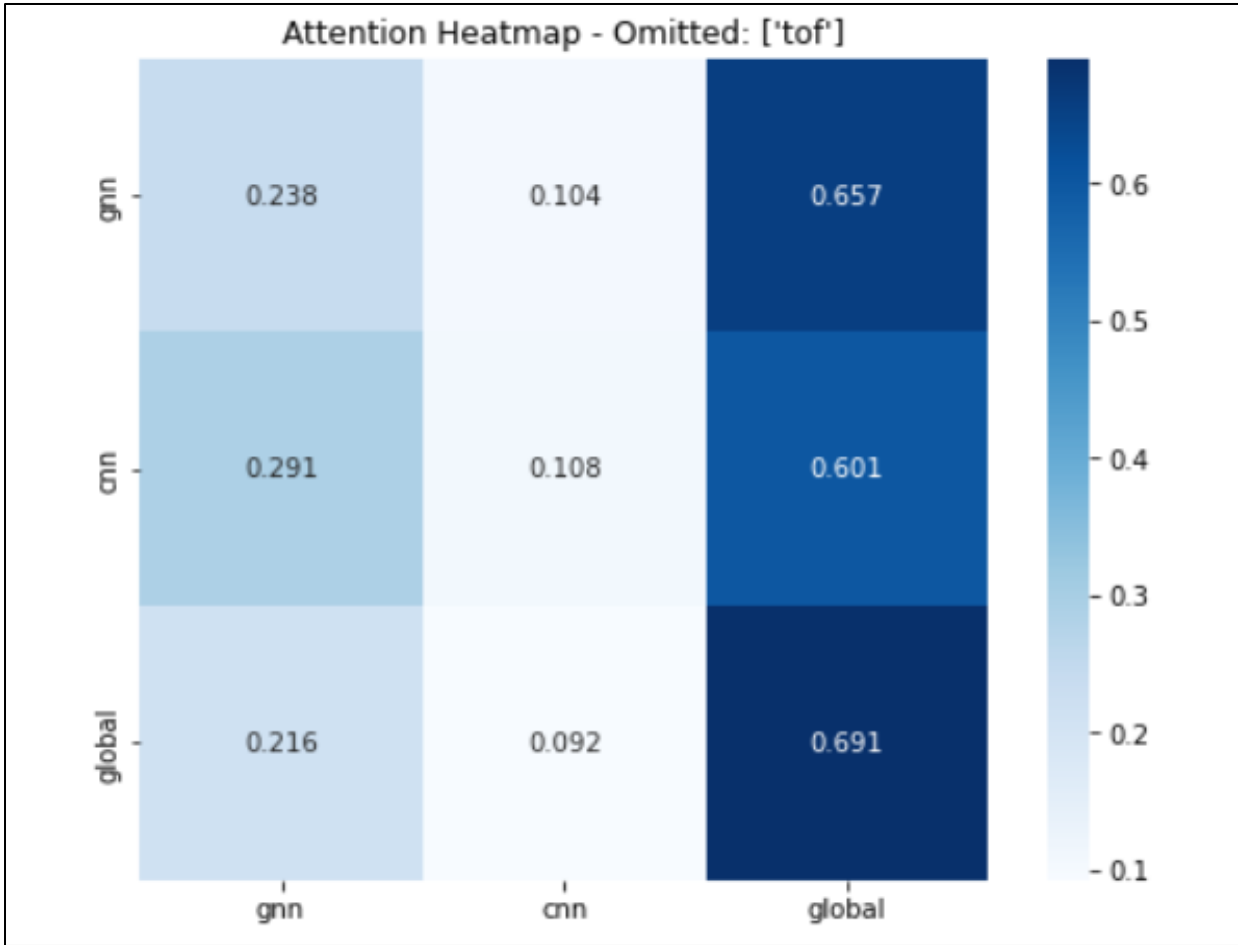




FEATURE ENGINEERING – ATTENTION HEATMAP



FEATURE ENGINEERING by Omitting sub-Models



Courses & Schools 2nd Year

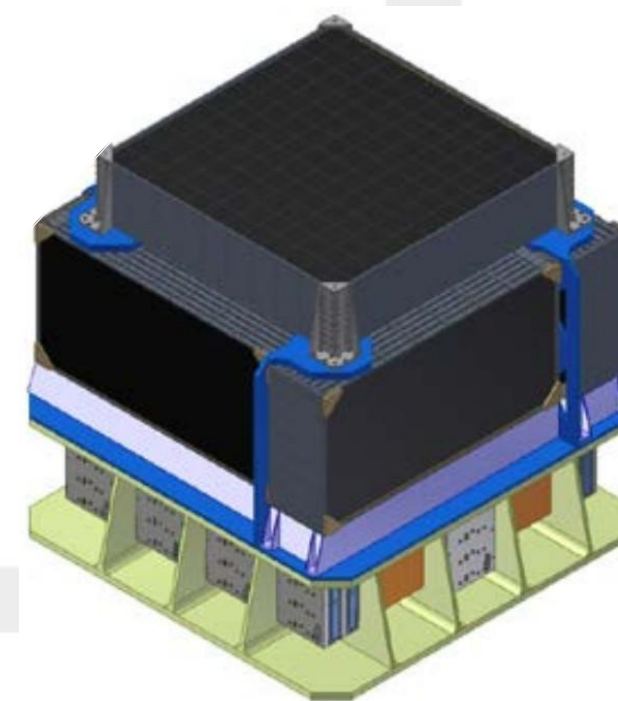
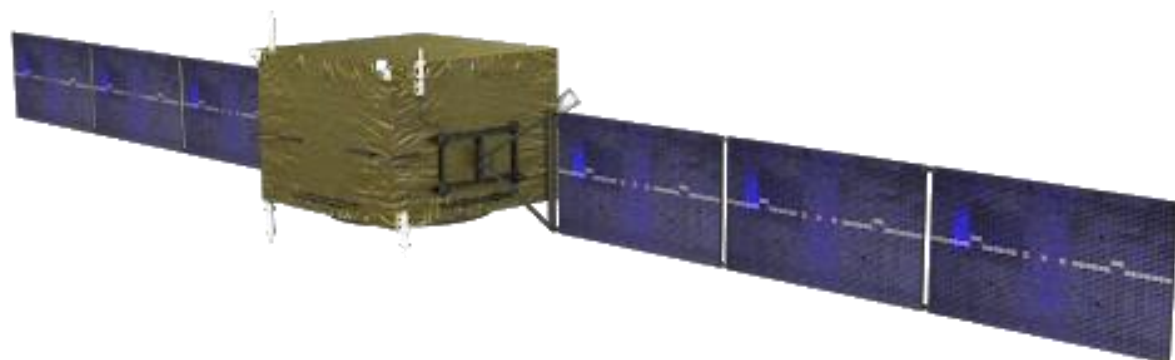
- ❖ Introduction to Neuromorphic Computing (Material submitted)
- ❖ Big Data Modeling and Learning (Project & Technical Report is submitted)
- ISAPP School: "High Energy Cosmic Rays: From Space to Ground" in Lecce, Italy, offering intensive summer school coverage of cosmic ray physics from space-based detection to ground-based observation (June 9-20).
- AMS Italia Meeting in Bologna, Italy, focused on strategic meetings regarding the Alpha Magnetic Spectrometer (AMS) Italian collaboration (November 27-29).
- NAIA Data Framework Use: Hands-on Training (Part 1) in Bologna, Italy, providing practical training on the specific data processing framework used for AMS-02 analysis (November 29).
- 2025:
 - NAIA Data Framework Use: Hands-on Training (Part 2) in Rome, Italy, continuing the practical data framework training (February 03).

Courses & Schools 2nd Year

- TECH-FPA PHD Retreat 2025 at LNGS, L'Aquila, Italy, focusing on technology, funding, and career paths in particle astrophysics (February 17-21).
- Authentic Impact Workshop by Barbara Van for Team Development at LNGS, L'Aquila, Italy, focusing on maximizing authentic professional impact and collaboration (TBD).
- AMS Nuclei Meeting at SDIAT, Jinan, China, specifically dedicated to the analysis, spectra, and interpretation of cosmic ray nuclei data from AMS-02 (September 6-12).
- 1st PhyStat School of Statistics: "Statistics in the era of ML" in the Netherlands, providing specialized training on advanced statistical methods and the application of Machine Learning (ML) in physics (November 17-21).
- TEE Shifts in POCC Data Center for AMS 02 , Geneva, Switzerland (28/03/2025 - 05/04/2025) & (11/05/2025 - 18/05/2025 & 20/09.2025 - 25/09/2025)

Third Year

- Complete feature engineering optimization and finalize thesis Chapter 1 (Cosmic Ray Physics fundamentals).
- Initiate flux measurement comparison study with collaborating research groups to validate model predictions.
- Complete thesis Chapter 2 (AMS-02 Detector systems and data processing).
- Analyze test beam data for tracker performance validation and systematic uncertainty quantification.
- Integrate test beam calibration results into production analysis pipeline.
- Complete thesis experimental chapters and begin final synthesis.



Thank You

