



UNIVERSITÀ
DEGLI STUDI
DI PADOVA



2nd Year Admission Presentation:

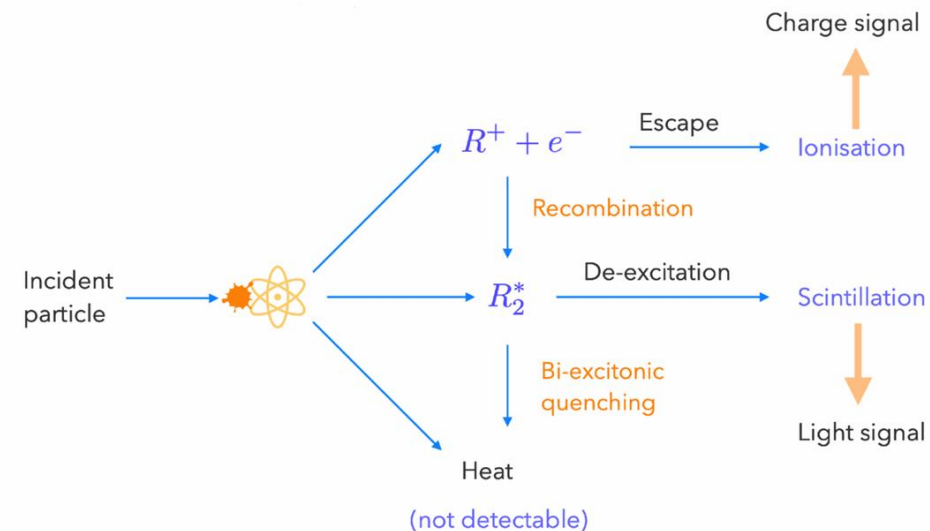
Background mitigation techniques and sensitivity studies for cryogenic noble-liquid detectors

Curriculum:
Detectors, lasers and optics

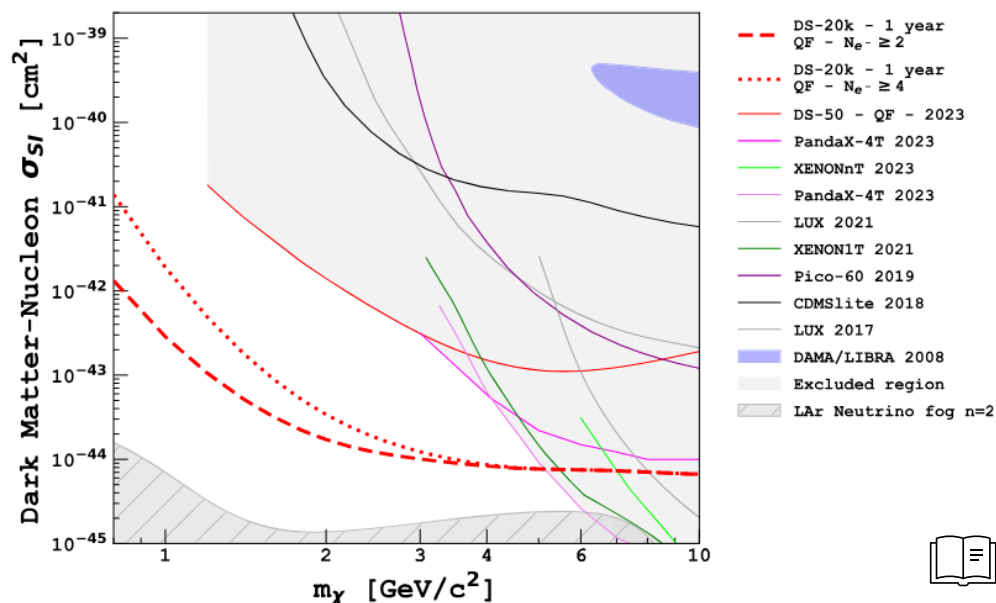
Riccardo Pavarani
Supervisors: Matteo Cadeddu,
Francesca Dordei

WIMP and Time Projection Chamber

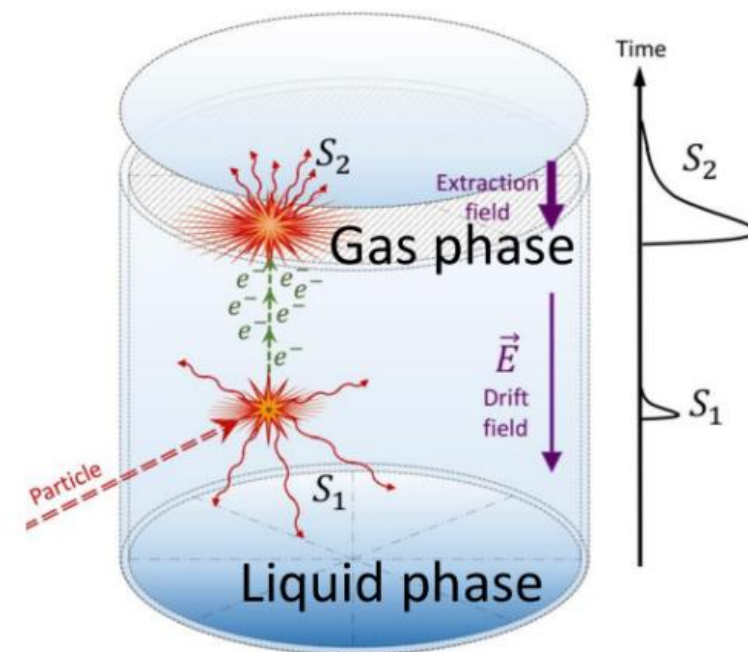
- Direct search of a dark matter candidate: **Weakly Interacting Massive Particle** (1-100 GeV)
- Dual-phase time projection chamber
- From DarkSide-50 to DarkSide-20k (DS-20k)



Credits: Laura Baudis

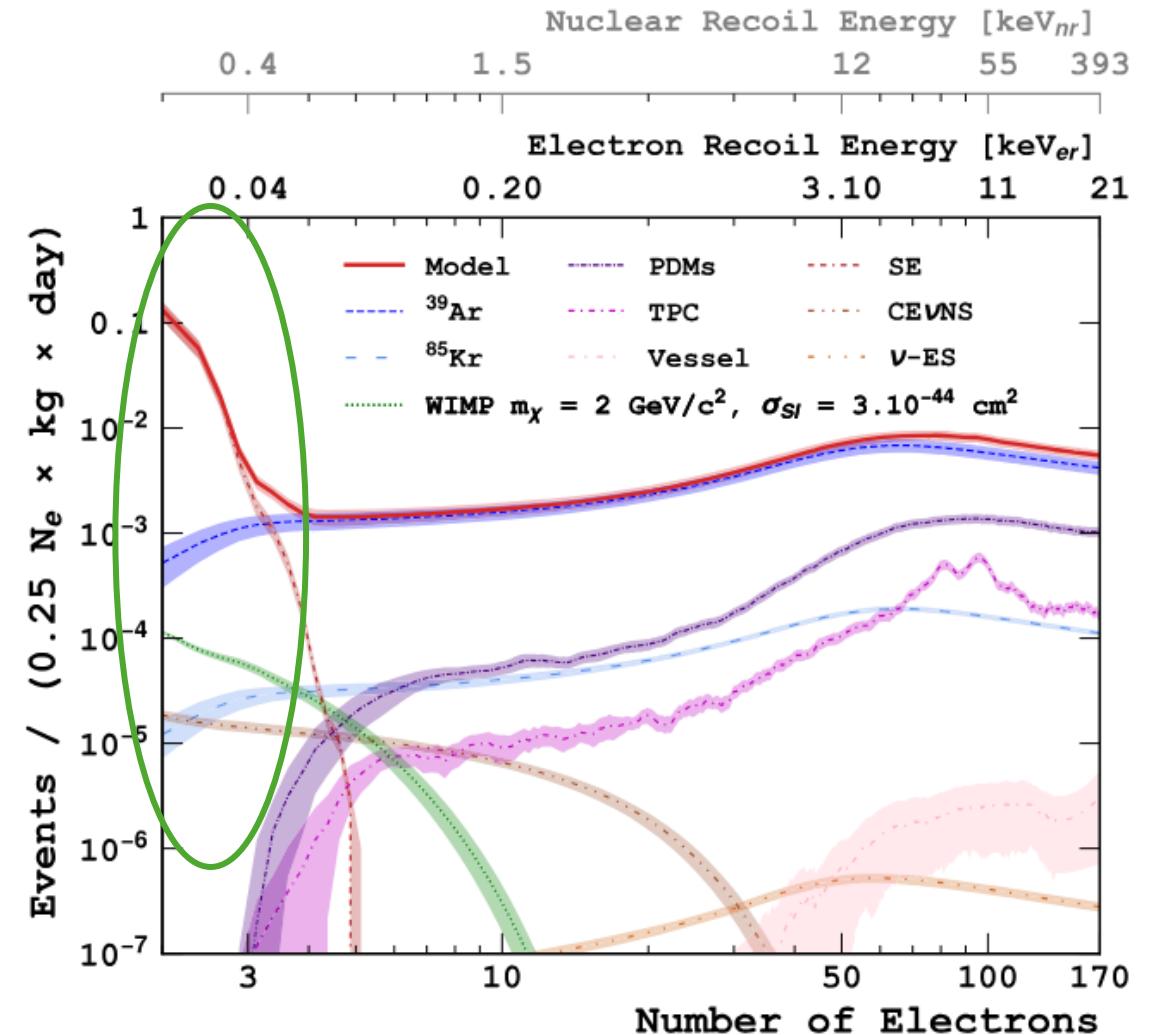


Commun.Phys. 7 (2024) 1, 422

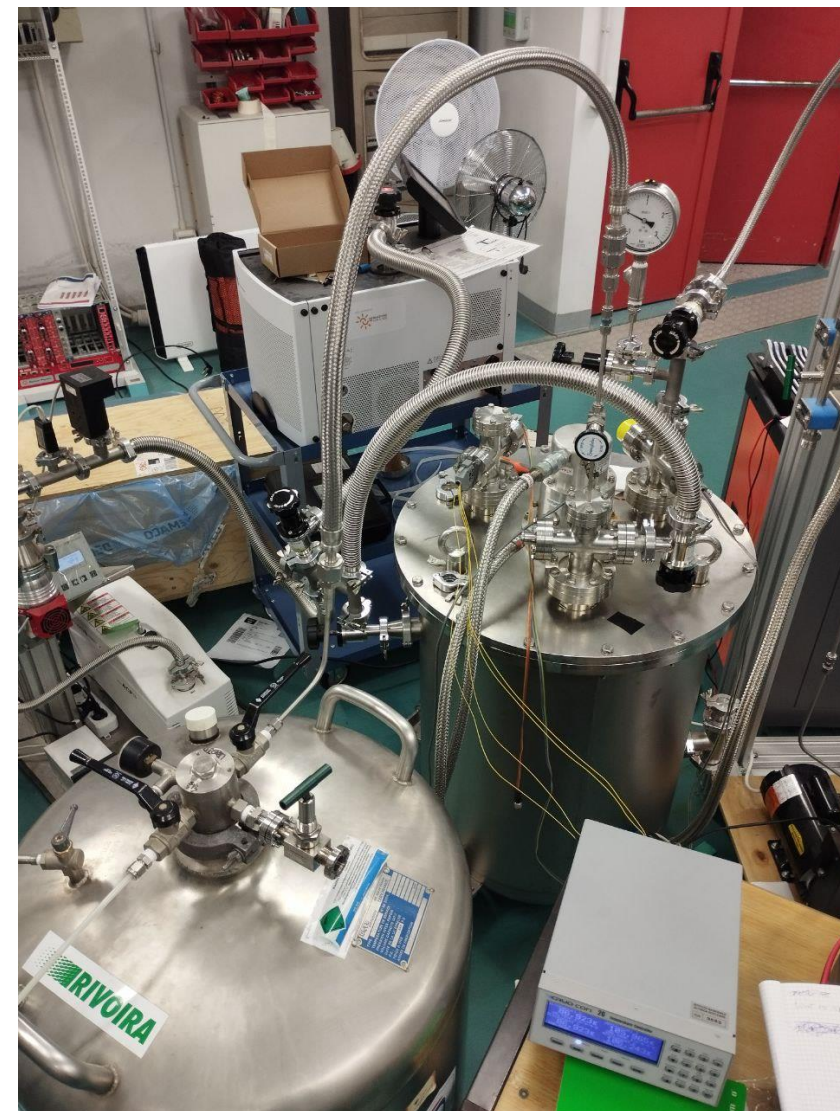
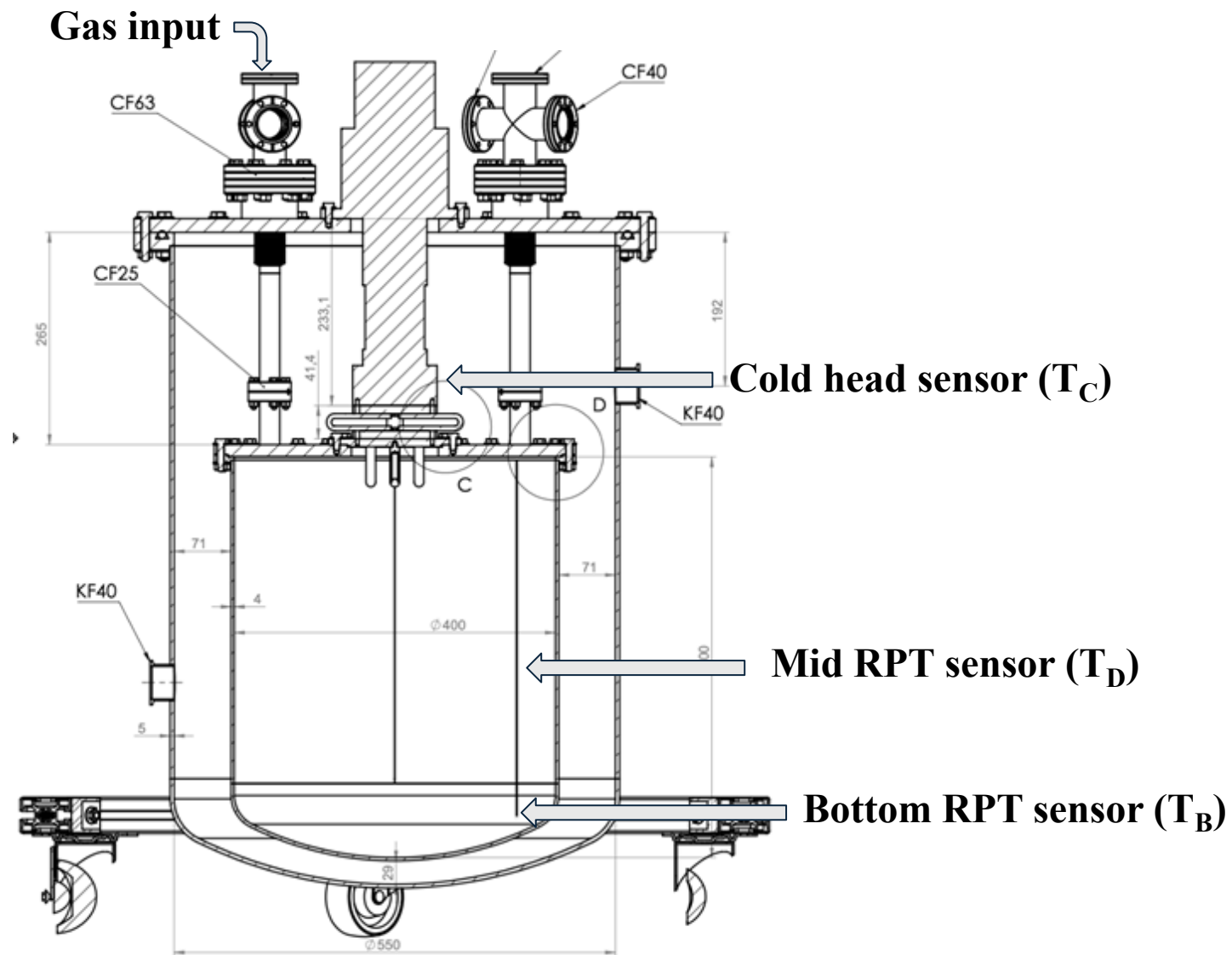


Research activity of my PhD

- Comprehensive understanding the **background** is fundamental to achieve the DarkSide-20k sensitivity
- Some backgrounds require further investigation (**single electron**, CEvNS)
- **Setup** to study innovative techniques to characterize and possibly reduce this spurious events

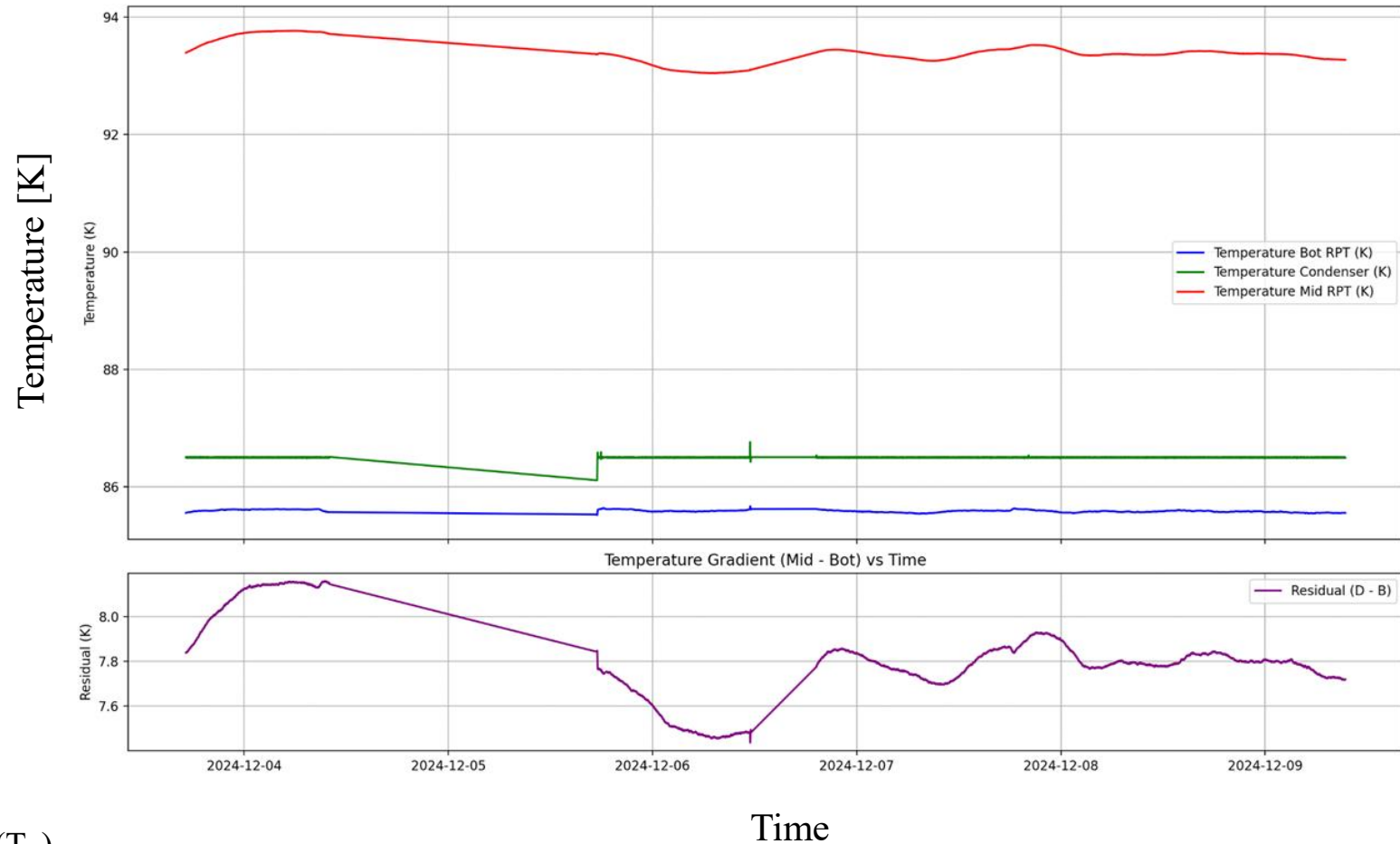
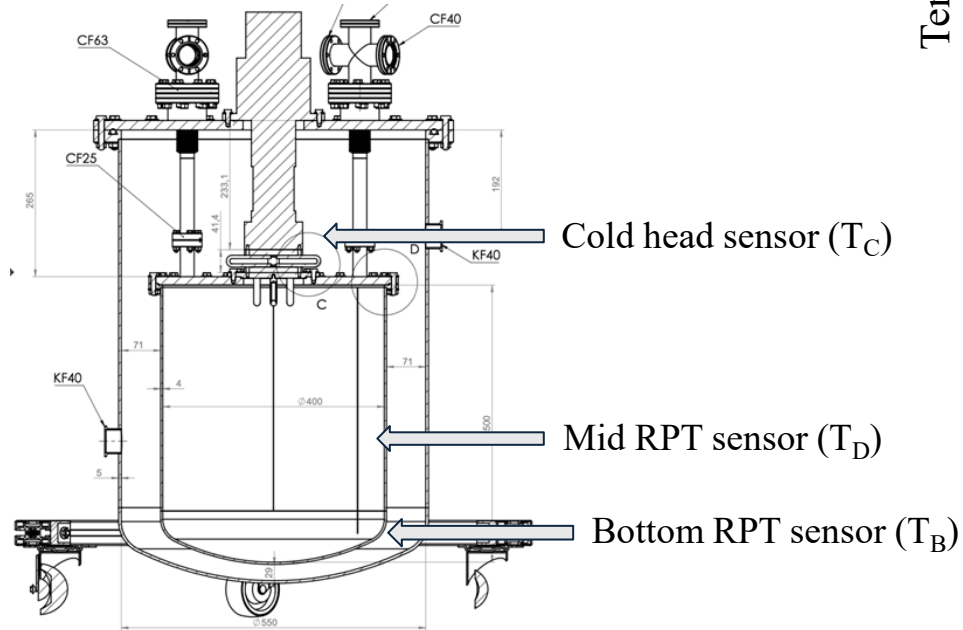


Research activity of my PhD



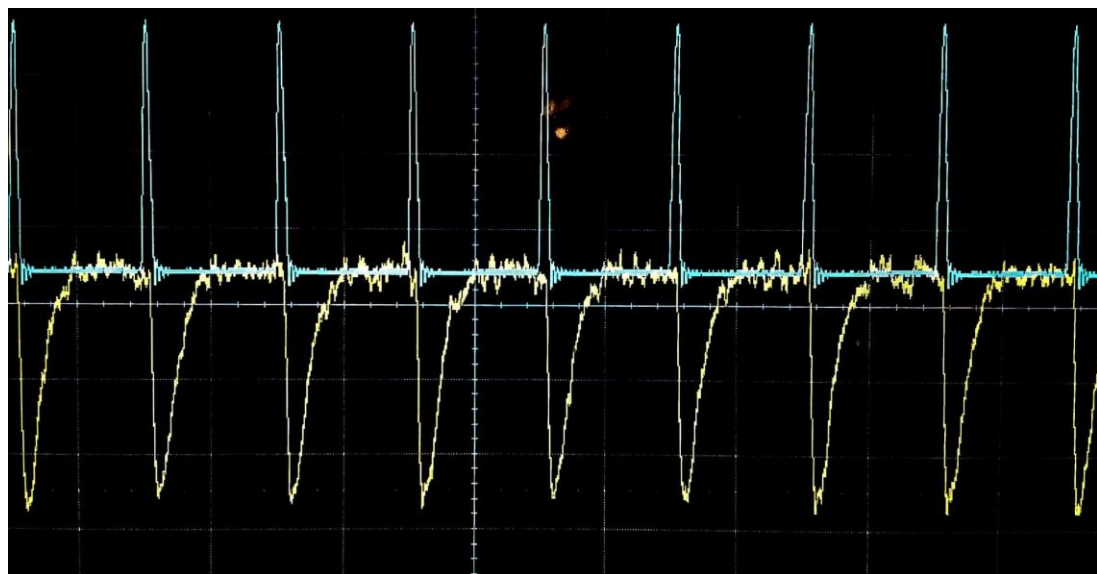
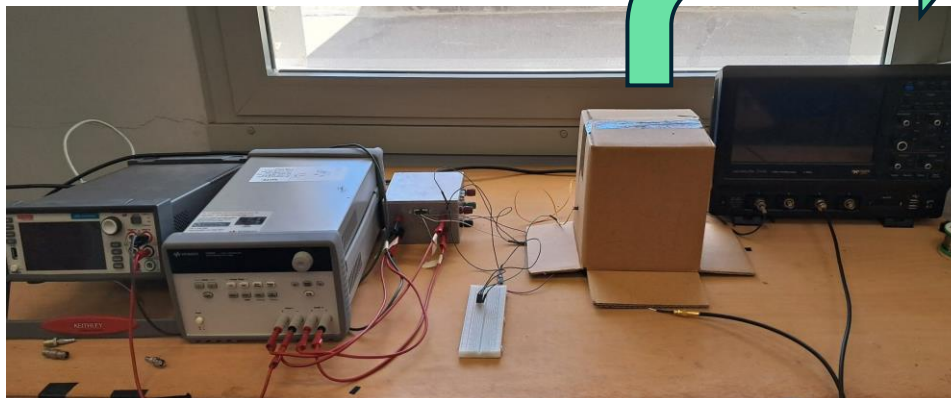
First run test

- Argon condensation inside the inner chamber
- About 7 L of liquid condensed
- Tested the stability (~ 1 week)



DarkSide-20k and DArT SiPM tiles warm test

- Circuit to power the tile
- Pulse generator to power LED
- Black box to shield external photons
- Waveform output signal readout through the oscilloscope



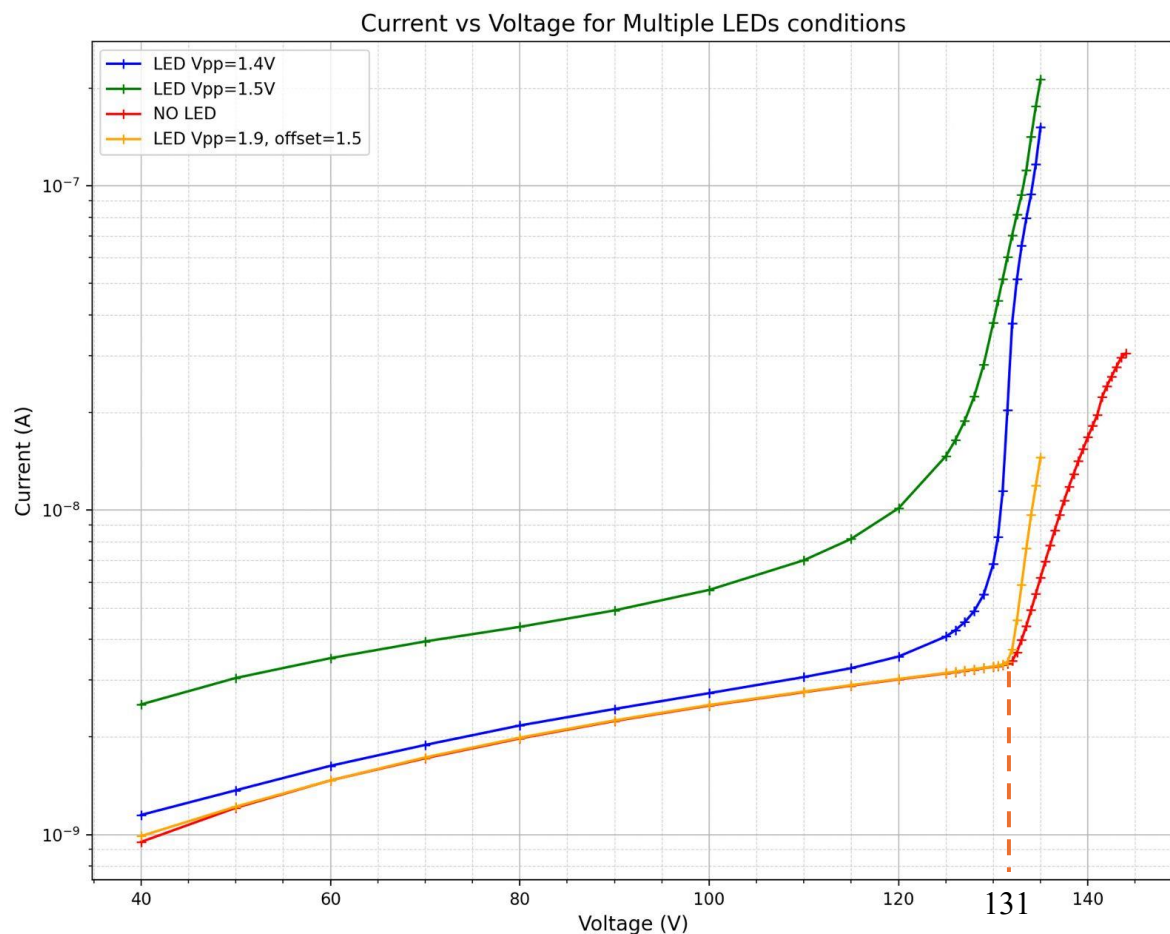
— Input Pulse to the LED
— SiPM output signal

DArT SiPM signal

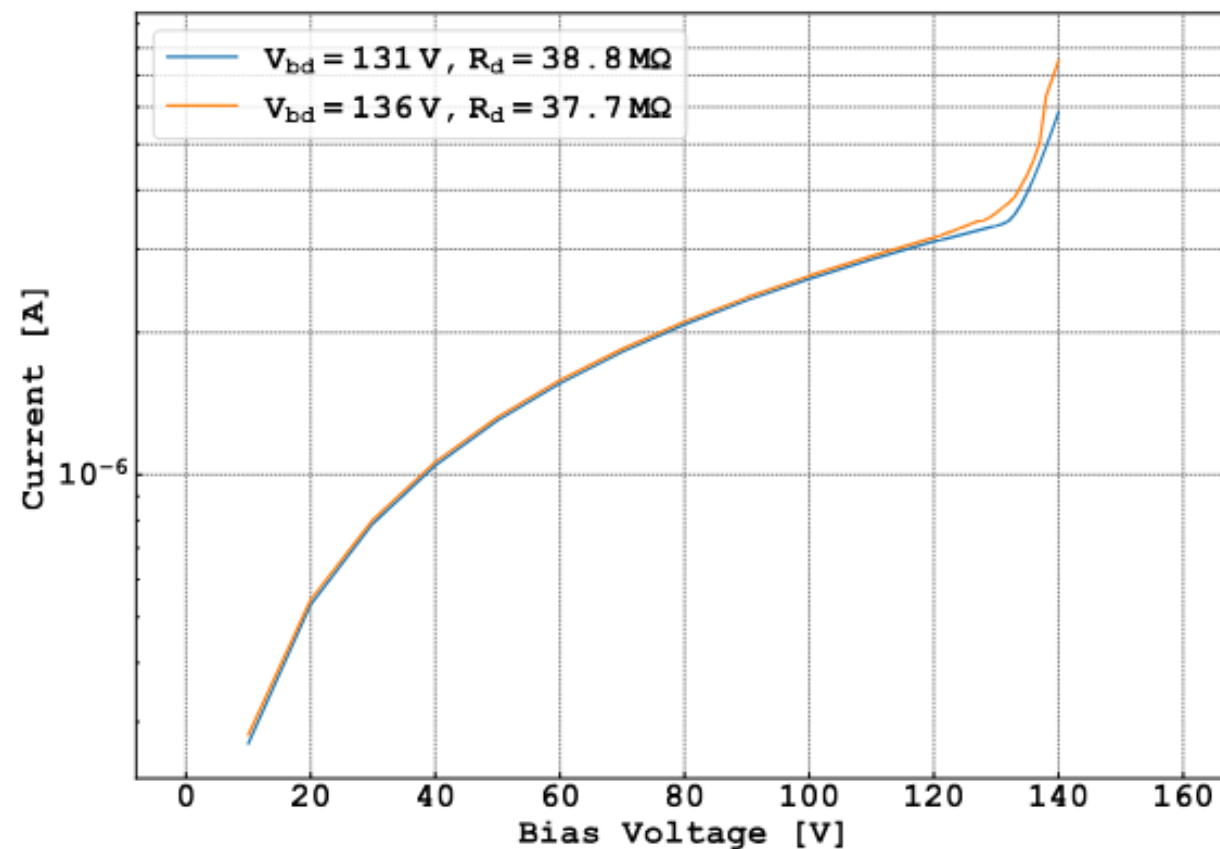
- SiPMs conditions:
 - $V_{\text{bias}} = 31\text{ V} + 1.5\text{ V overvoltage}$
 - Operational Amplifier range (-1.4V , +3.6V)
- LED conditions:
 - $V_{\text{pp}} = 1.9\text{ V}$
 - Offset = 1.5 V
 - Pulse Frequency: 3MHz

IV curves of Darkside-20k tile

- Measurement of a DarkSide tile IV curves
(this work)



- Measurement of a DarkSide tile IV curves
(DarkSide collaboration)



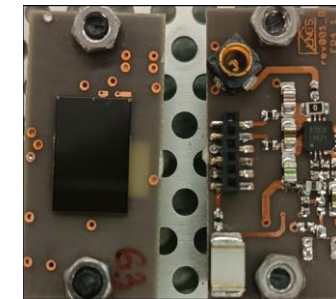
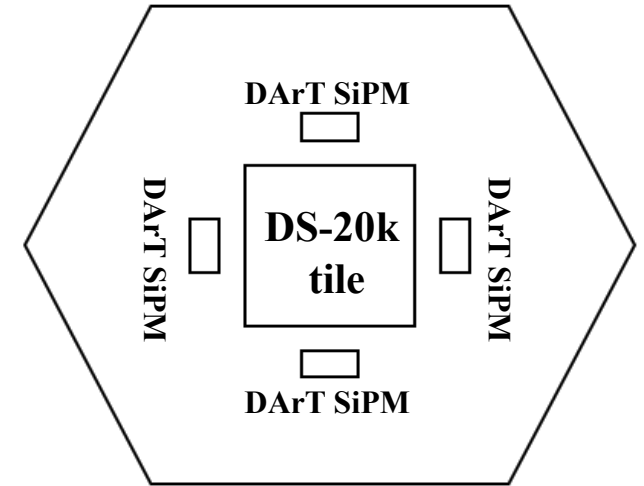
Test at cryogenic temperature

Characterization of the electronic response at cryogenic temperature
In Liquid Argon (87 K) or Liquid Nitrogen (77K)

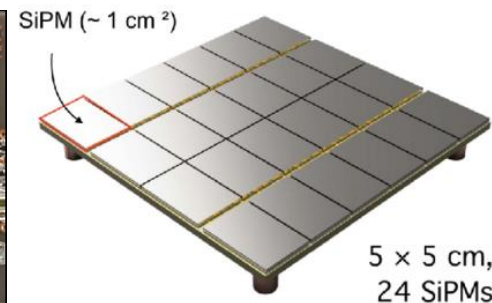
Design of the support structure to be placed inside of the inner chamber

- 3 supporting bars
- Exagonal plate (~10 cm from the bottom)
- 1 DarkSide tile and 4 DArT SiPM
- Black box
- LED
- Additional heater and temperature sensor (PT100)

Optical plane



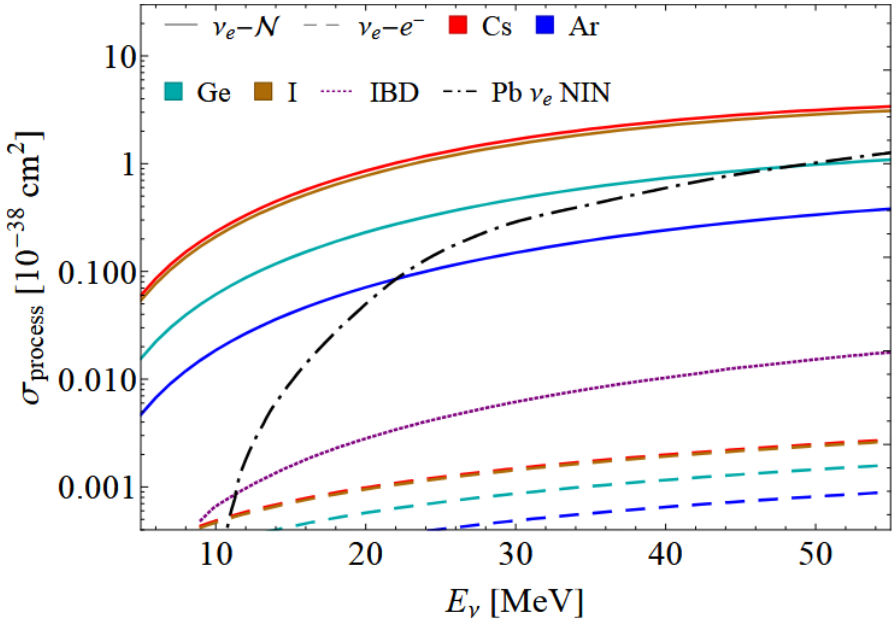
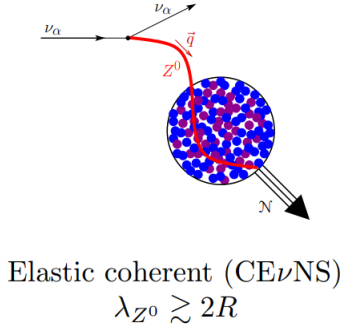
DArT SiPM



DS-20k tile

Research activity of my PhD

- Coherent Elastic neutrino Nucleus Scattering will be the dominating background of DS-20k in the high mass region
- Higher cross section with respect to other neutrino processes at low energy

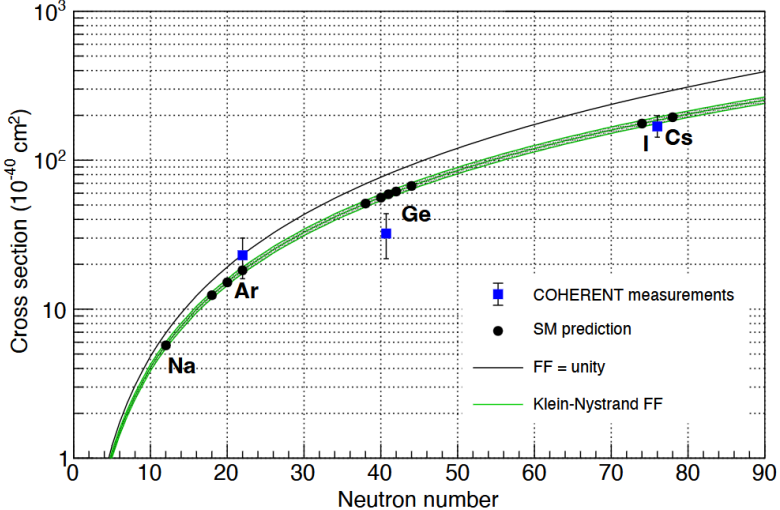
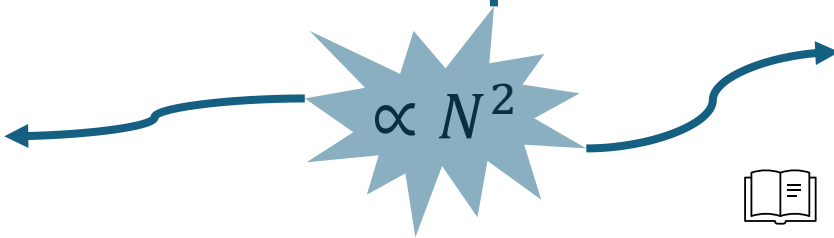


$$\frac{d\sigma}{dE_r} = \frac{G_F^2 m_N}{\pi} \left(1 - \frac{m_N E_r}{2E_\nu^2} \right) \underbrace{\left[g_V^p (\sin^2 \vartheta_W) Z F_Z(q^2) + g_V^n N F_N(q^2) \right]^2}_{Q_W^2}$$

Weak mixing angle Nuclear form factor

Neutrino-nucleon tree level couplings:

- $g_V^p = \frac{1}{2} - 2\sin^2(\vartheta_W) \cong 0.02274$
- $g_V^n = -\frac{1}{2} = -0.5$

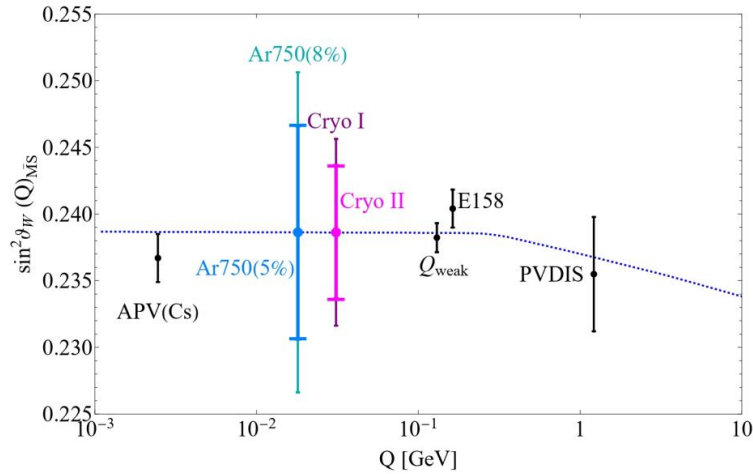


EPL 143 (2023) 3, 34001



Sensitivity of COH CsI and LAr detectors

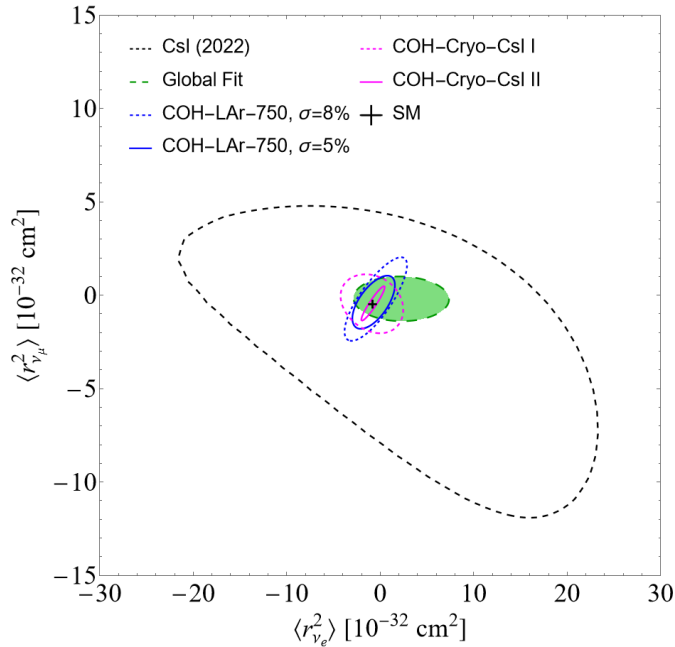
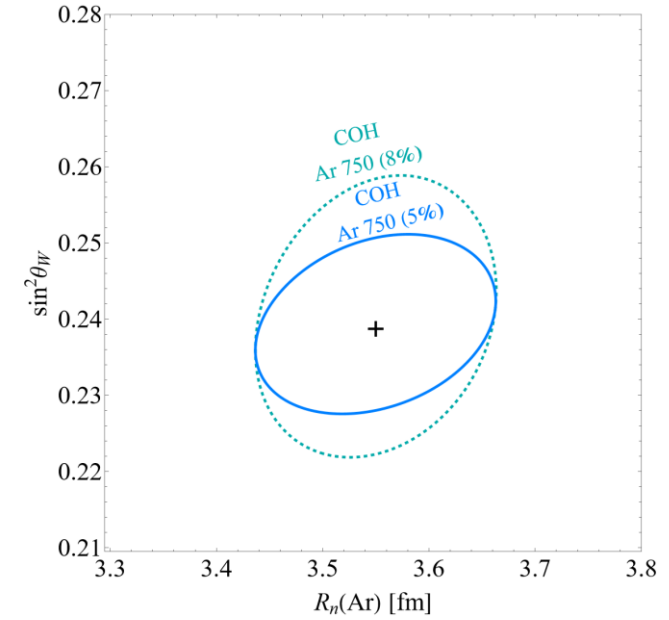
R. Pavarani et al. arXiv:2509.0420



Weak mixing angle

Neutron distribution

CEvNS



Neutrino charge radius

Beyond Standard Model

Toward precision physics tests with future COHERENT detectors

M. Atzori Corona^{1,a}, M. Cadeddu^{2,b}, N. Cargioli^{2,c},
F. Dordei^{2,d}, C. Giunti^{3,e} and R. Pavarani^{2,4,f}

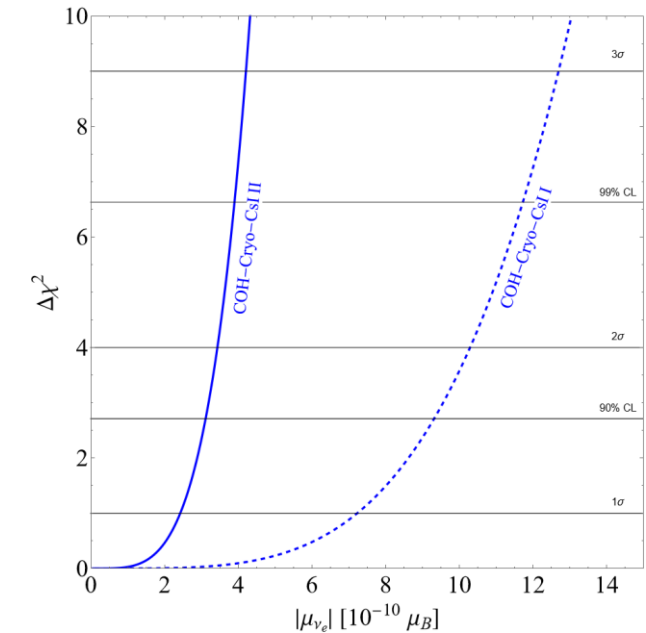
¹Istituto Nazionale di Fisica Nucleare (INFN), Sezione di Roma Tor Vergata,
Via della Ricerca Scientifica, I-00133 Rome, Italy

²Istituto Nazionale di Fisica Nucleare (INFN), Sezione di Cagliari,
Complesso Universitario di Monserrato - S.P. per Sestu Km 0.700, 09042 Monserrato (Cagliari), Italy

³Istituto Nazionale di Fisica Nucleare (INFN),
Sezione di Torino, Via P. Giuria 1, I-10125 Torino, Italy

⁴Università degli Studi di Padova, Dipartimento di Fisica e Astronomia "Galileo Galilei",
Via Francesco Marzolo 8, 35131 Padova, Italy

We present a comprehensive sensitivity study of future CEvNS detectors, focusing on a cryogenic cesium iodide detector and a tonne-scale liquid argon one, currently being developed by the COHERENT Collaboration. These setups will enable precision measurements of the weak mixing angle at low energies and allow accurate extraction of the neutron nuclear distribution radius. We also demonstrate that next-generation detectors will place constraints on the neutrino charge radius comparable to or better than current global fits. In addition, we explore the sensitivity to non standard neutrino electromagnetic properties, such as magnetic moments and millicharges, as well as new mediators. These findings reinforce the role of CEvNS experiments in the upcoming precision era, with future detectors playing a key role in advancing our understanding of neutrino interactions and electroweak physics at low energies.



Courses, schools and other activities

Courses:

- DarkSide Young Academy (exam passed): dark matter theory, Aria and argon purification, DarkSide 20k status and update, cryogenics of DarkSide-20k and my **talk**
- Rare event search with noble liquids (attended)



Other activities:

- TECH-FPA PhD retreat in LNGS, L'Aquila
- **Poster** at Incontri di Fisica delle alte energie (IFAE), Cagliari
- **Poster** and **talk** at Multi Aspect Young ORiented Advanced Neutrino Academy (MAYORANA) – International School, Modica (exam passed)
- Shift DEAP3600
- Darkside Masterclass, Cagliari
- Gran Sasso Hands-on 2025 PhD summer school on experimental astroparticle physics, L'Aquila



Next activities of my PhD

- Development of **data acquisition system (DAQ)**
- Testing **SiPMs at cryogenic temperatures**
- Design and construction of **dual-phase TPC**
- Improvement of **slow control system** (LabView?)
- **Simulations** for the final system (Geant4)

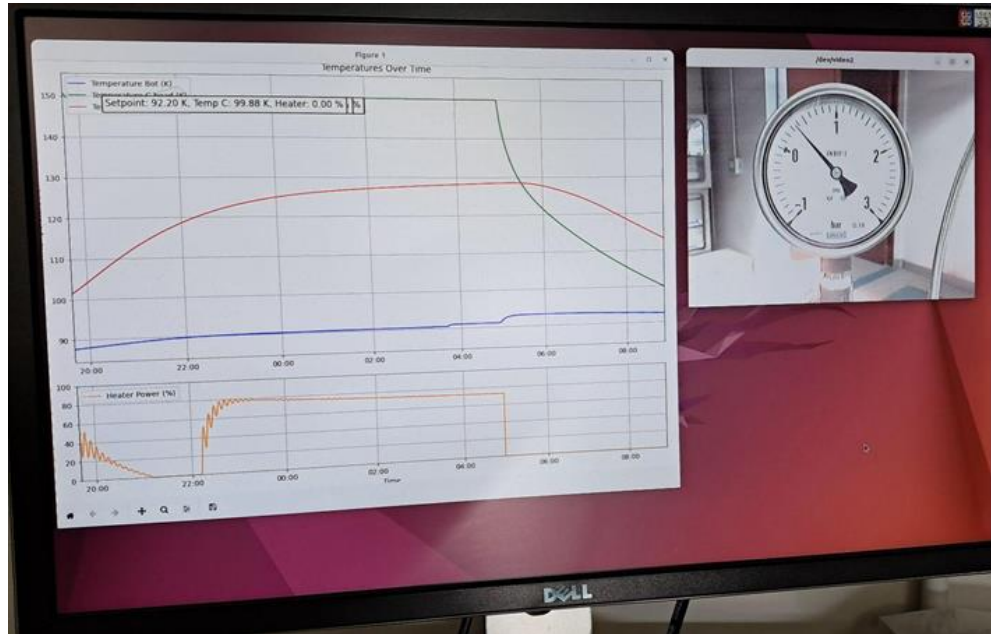
Next activities of my PhD

- Development of a **data acquisition system (DAQ)**
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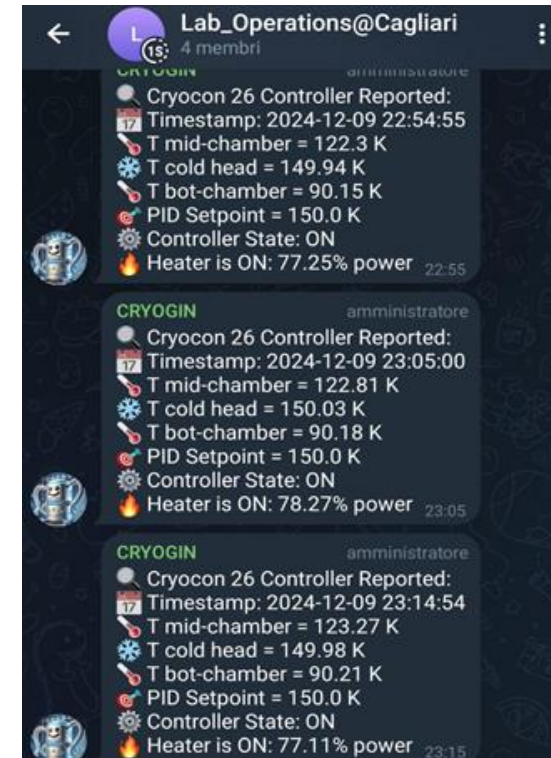
THANK
YOU!

BACKUP

Research activity of my PhD



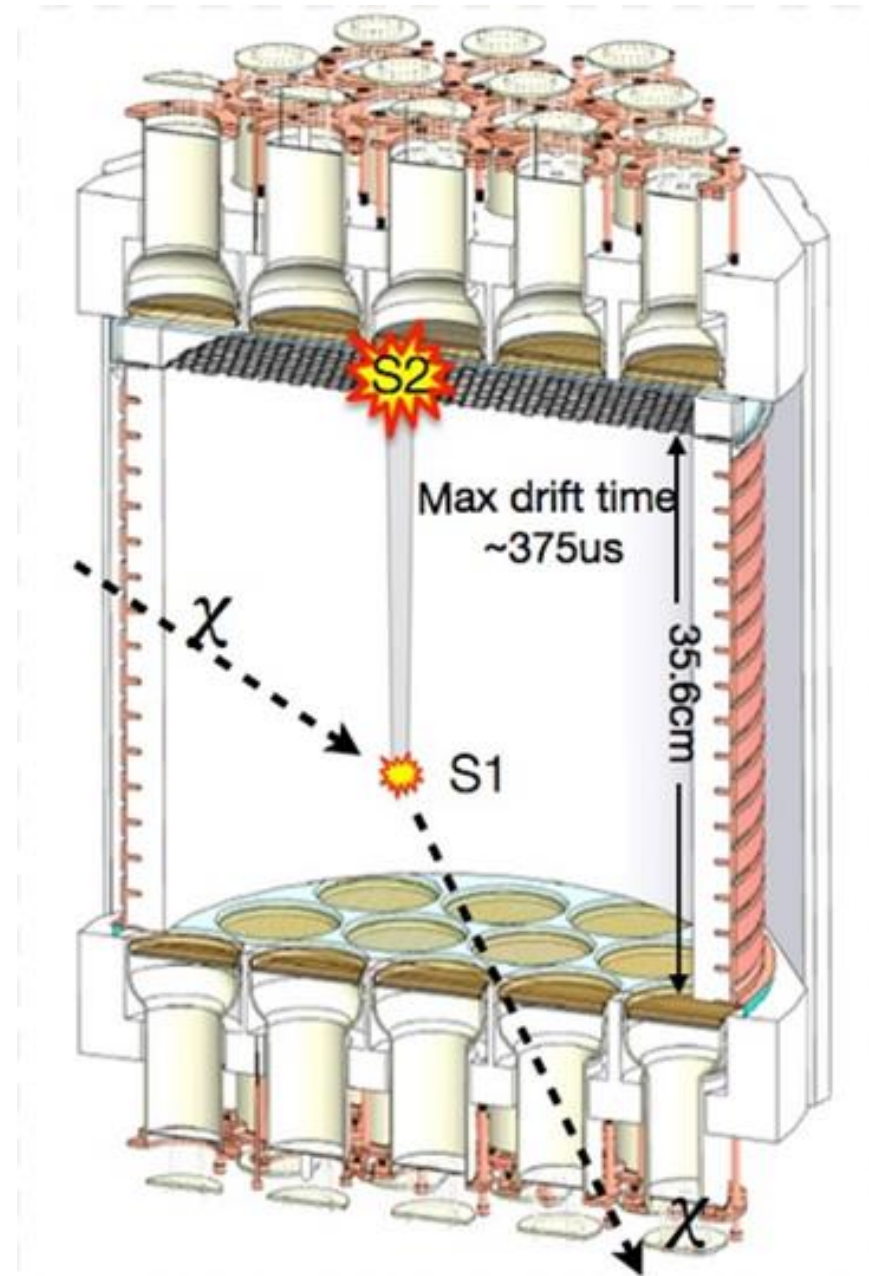
- Live plot of temperature and heaters power
- Live time of analog pressure
- Heaters can be controlled from the PC setting a temperature



- Anydesk remotely
- Telegram bot reporting about the status of the system every 10 minutes

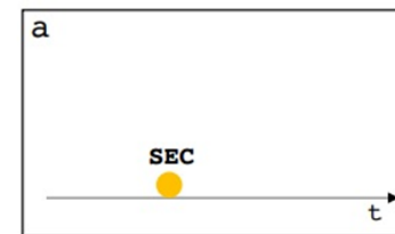
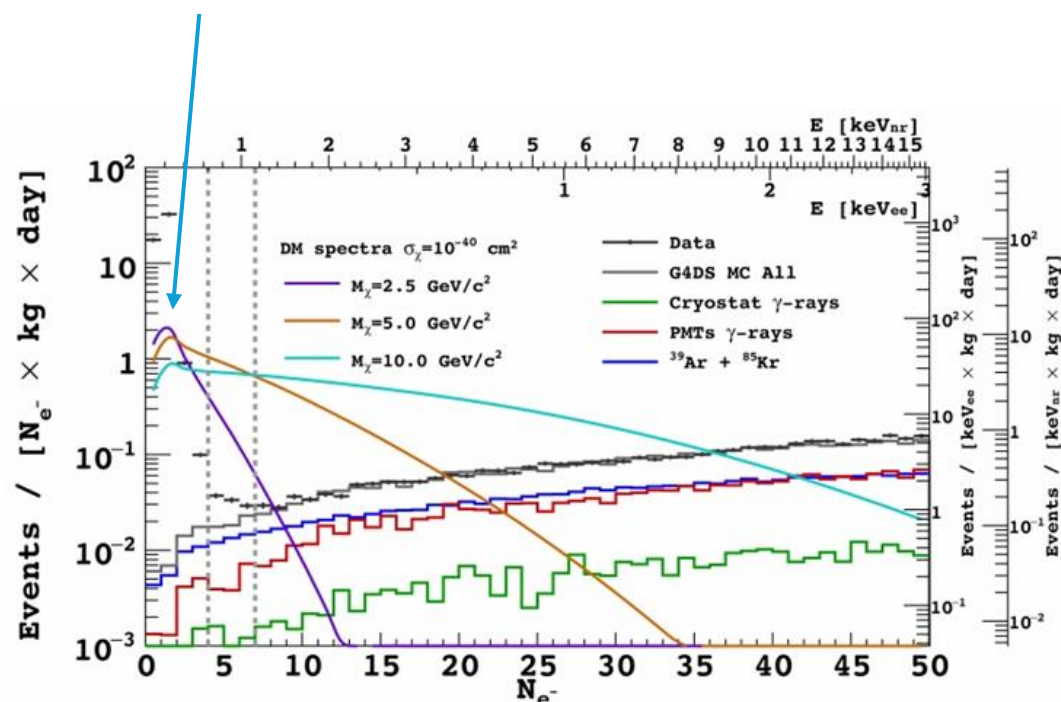
Research activity of my PhD

Time Projection Chamber

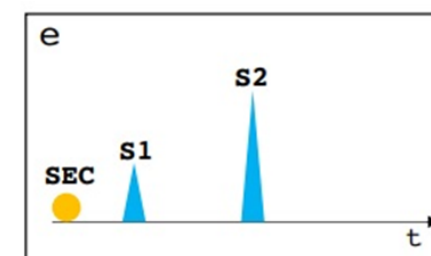
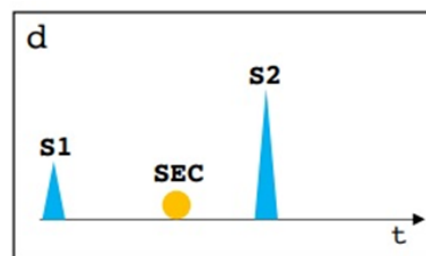
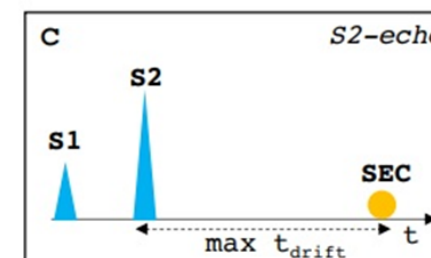
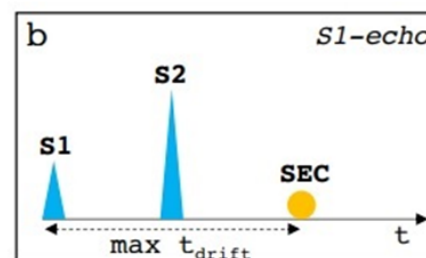


Research activity of my PhD

The physics goal is to build a setup to investigate the so-called **single electron background**



Single pulse events

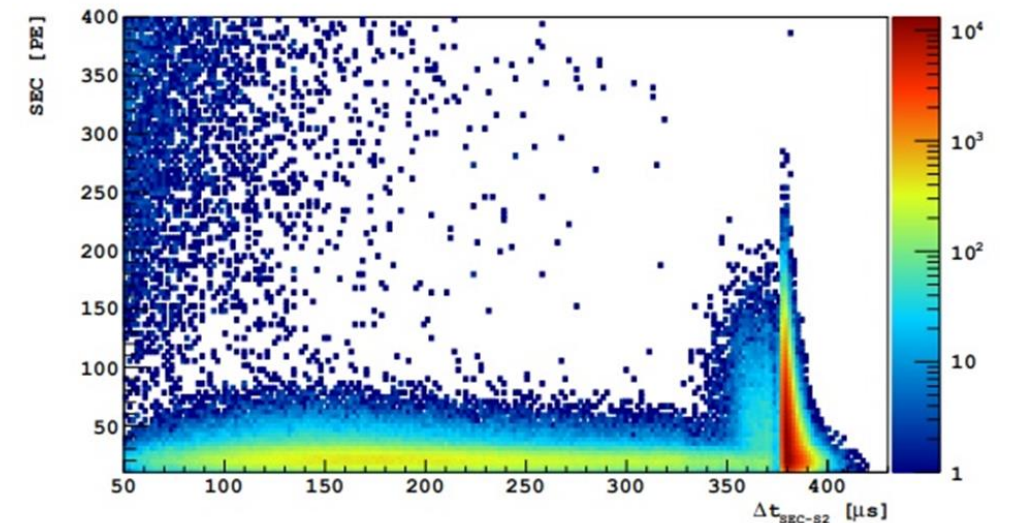
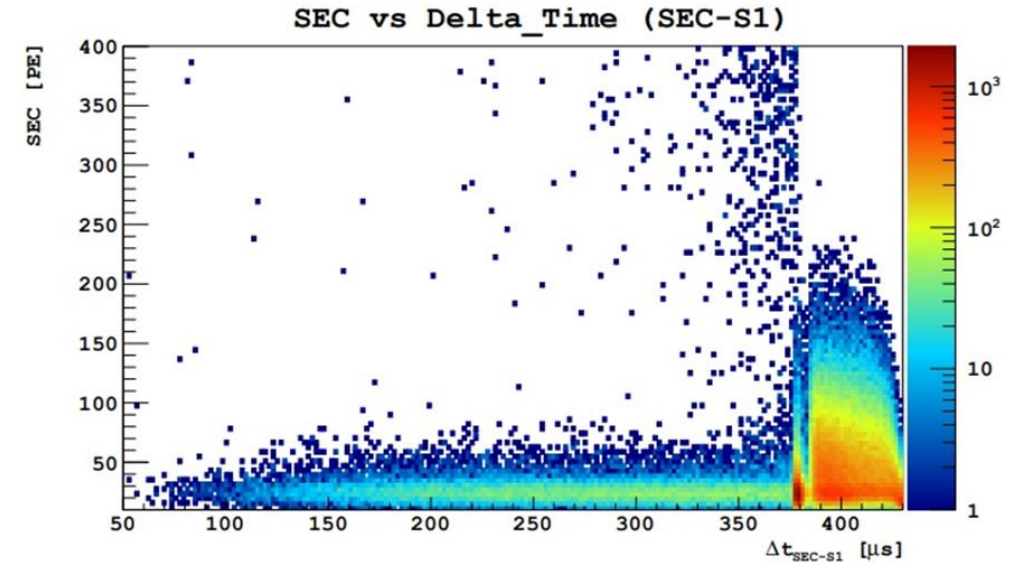
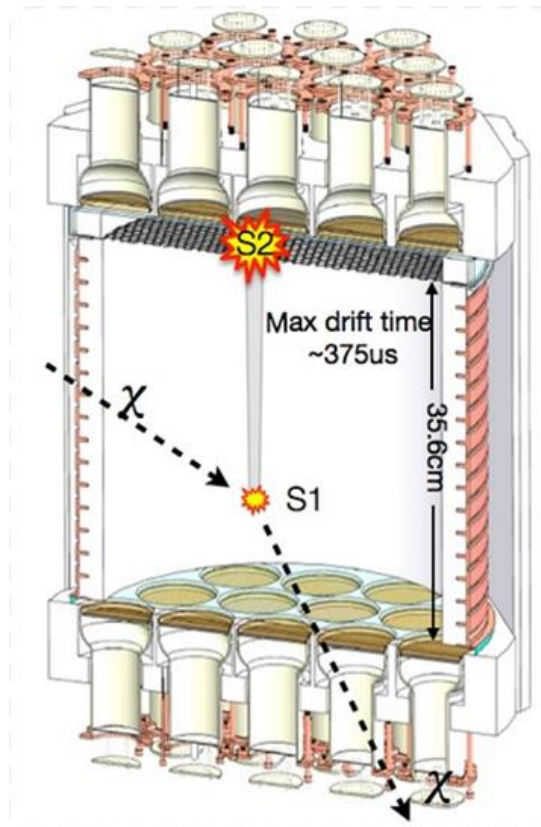


Three pulse events

Research activity of my PhD

Explanations among others:

- Photoionization from the cathode
- Impurities



Research activity of my PhD

Cryo I: $\sigma_{\sin^2\vartheta_W} \sim 3\%$

Cryo II: $\sigma_{\sin^2\vartheta_W} \sim 2\%$

Ar750 ($\sigma=8\%$): $\sigma_{\sin^2\vartheta_W} \sim 5\%$

Ar750 ($\sigma=5\%$): $\sigma_{\sin^2\vartheta_W} \sim 3.5\%$

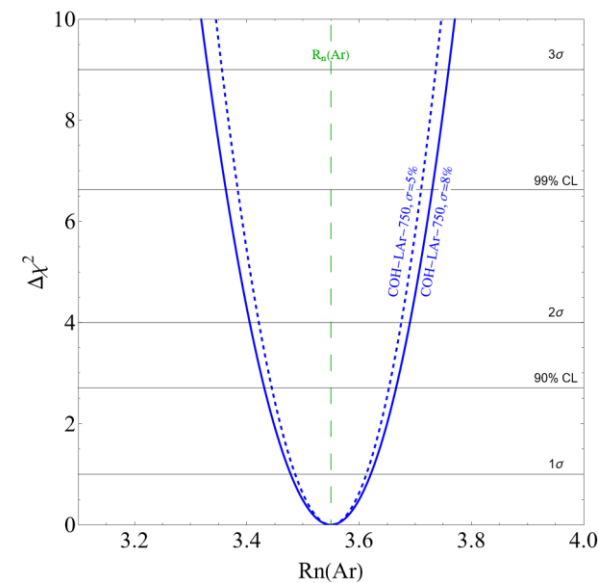
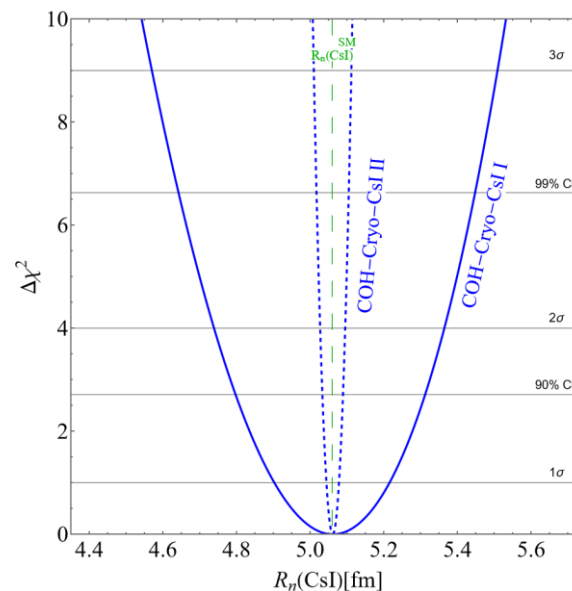
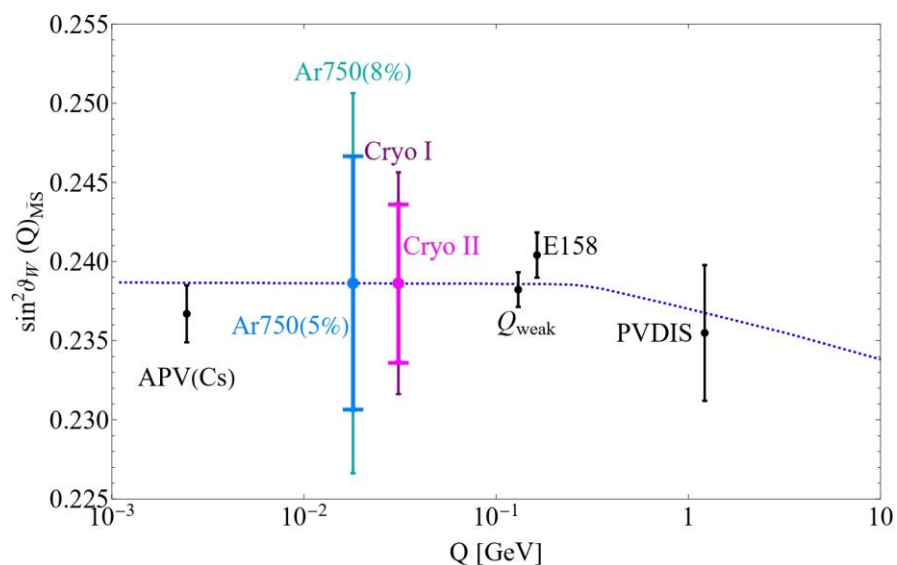
CEvNS allows one to test the SM through the **weak mixing angle**, and to access to the **weak nuclear distributions**.

Cryo I: $\sigma_{R_n} \sim 3\%$

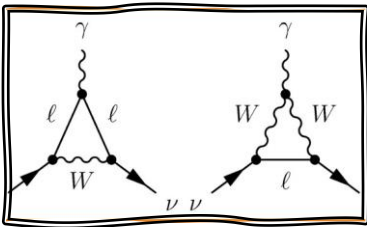
Cryo II: $\sigma_{R_n} \sim 0.4\%$

Ar750 ($\sigma=8\%$): $\sigma_{R_n} \sim 2\%$

Ar750 ($\sigma=5\%$): $\sigma_{R_n} \sim 2\%$



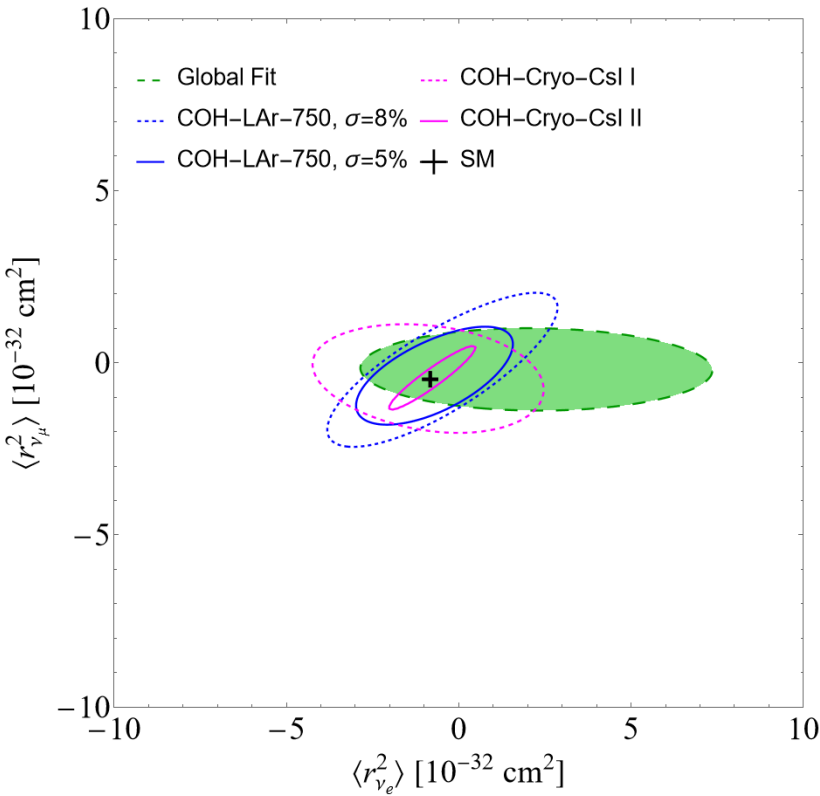
Research activity of my PhD



Although neutrinos are electrically neutral particles, they can interact with other charged particle through photon exchange in loop diagrams via the **neutrinos charge radius** (NCR). The NCR introduces a **flavor dependence** in the cross section

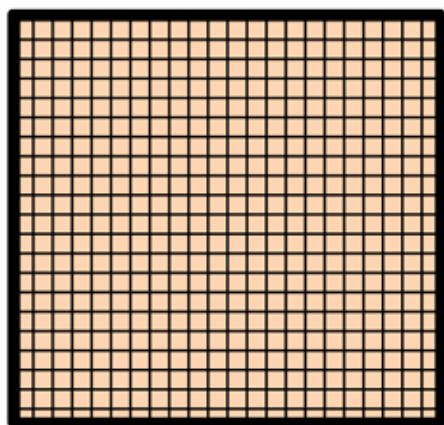
$$\langle r_{\nu_\alpha}^2 \rangle_{\text{SM}} = -\frac{G_F}{2\sqrt{2}\pi^2} \left[3 - 2 \log \left(\frac{m_\alpha^2}{m_W^2} \right) \right] \rightarrow \begin{cases} \langle r_{\nu_e}^2 \rangle_{\text{SM}} = -8.2 \times 10^{-33} \text{cm}^2 \\ \langle r_{\nu_\mu}^2 \rangle_{\text{SM}} = -4.8 \times 10^{-33} \text{cm}^2 \\ \langle r_{\nu_\tau}^2 \rangle_{\text{SM}} = -3.0 \times 10^{-33} \text{cm}^2 \end{cases}$$

Method	Experiment	Limit [10^{-32} cm^2]	C.L.	Year
Reactor $\bar{\nu}_e e^-$	Krasnoyarsk	$ \langle r_{\nu_e}^2 \rangle < 7.3$	90%	1992
	TEXONO	$-4.2 < \langle r_{\nu_e}^2 \rangle < 6.6^a$	90%	2009
Accelerator $\nu_e e^-$	LAMPF	$-7.12 < \langle r_{\nu_e}^2 \rangle < 10.88^a$	90%	1992
	LSND	$-5.94 < \langle r_{\nu_e}^2 \rangle < 8.28^a$	90%	2001
Accelerator $\nu_\mu e^-$	BNL-E734	$-5.7 < \langle r_{\nu_\mu}^2 \rangle < 1.1^{a,b}$	90%	1990
	CHARM-II	$ \langle r_{\nu_\mu}^2 \rangle < 1.2^a$	90%	1994
CEvNS [arXiv:2205.09484]	COHERENT	$-7.1 < \langle r_{\nu_e}^2 \rangle < 11.2$	90%	2022
	+ Dresden-II	$-8.1 < \langle r_{\nu_\mu}^2 \rangle < 4.3$		

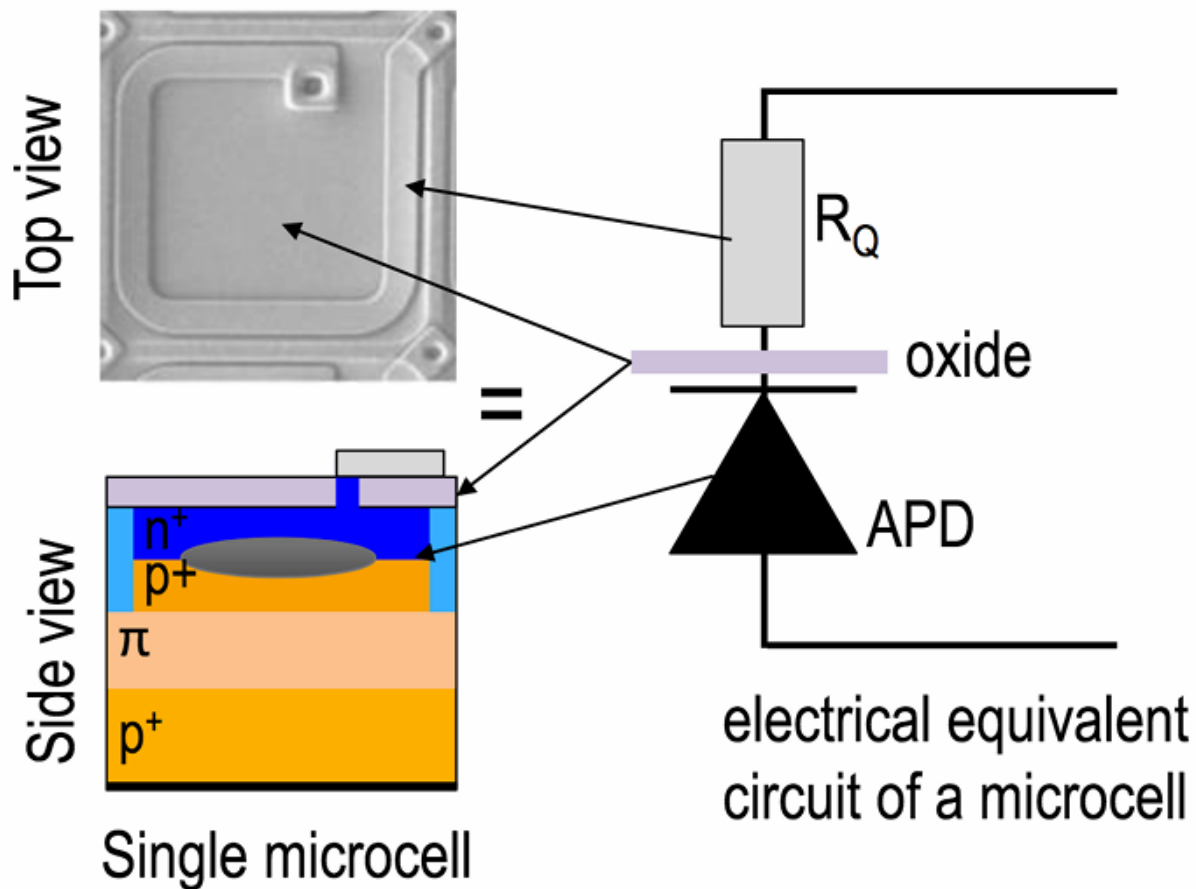


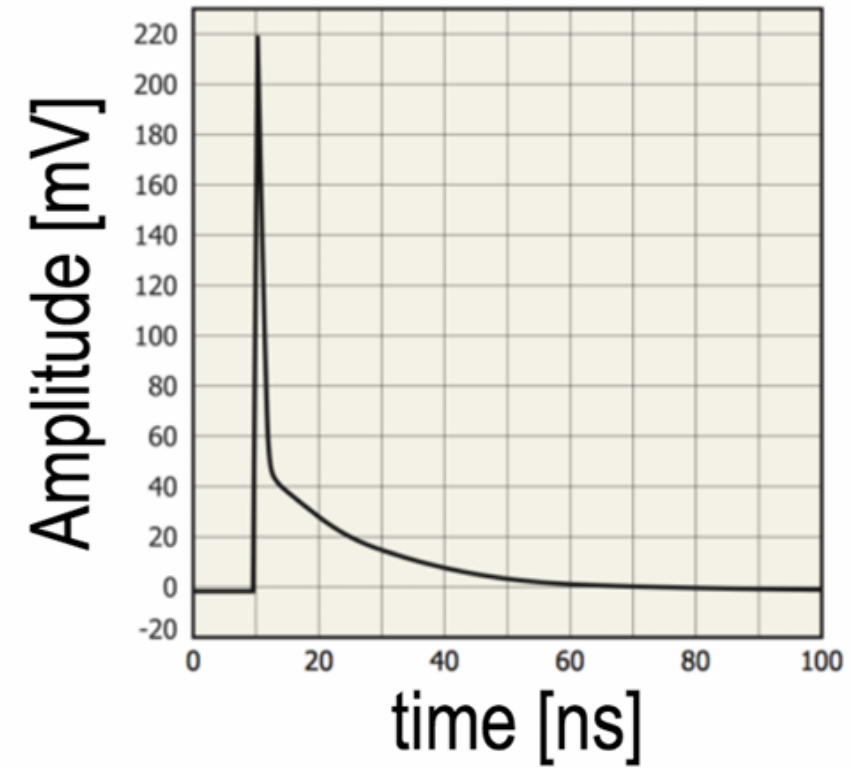
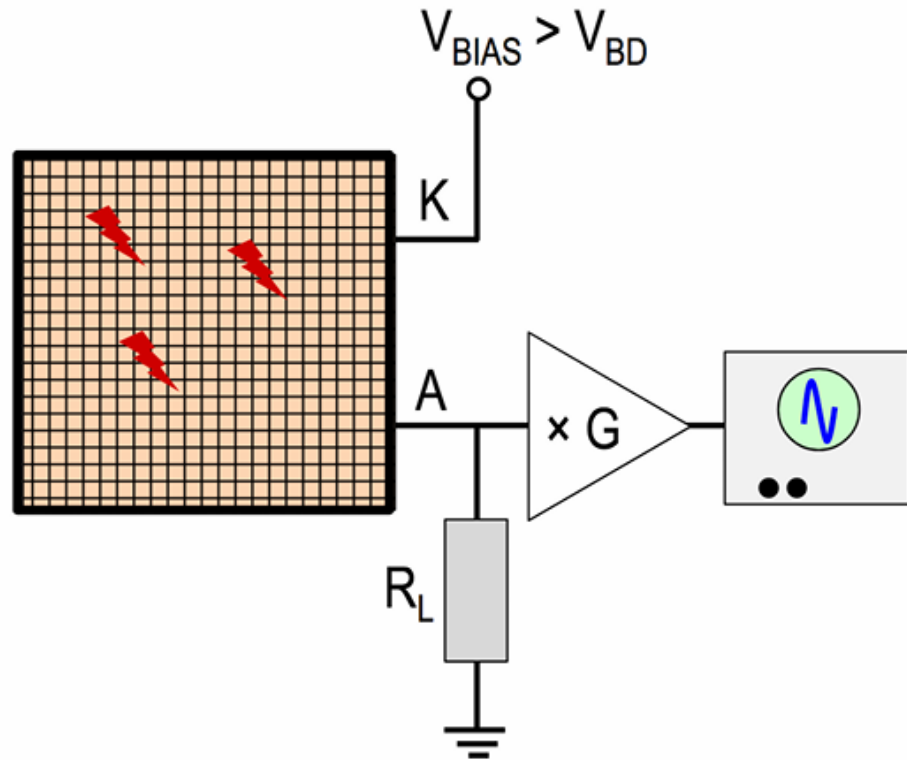
M. Atzori Corona et al., JHEP 09, 164 (2022)





SiPM is an array of microcells





Singolo fotoelettrone (1 p.e.)
Guadagno = area sotto la curva in
elettroni