

# Neutrino mass with the PTOLEMY project

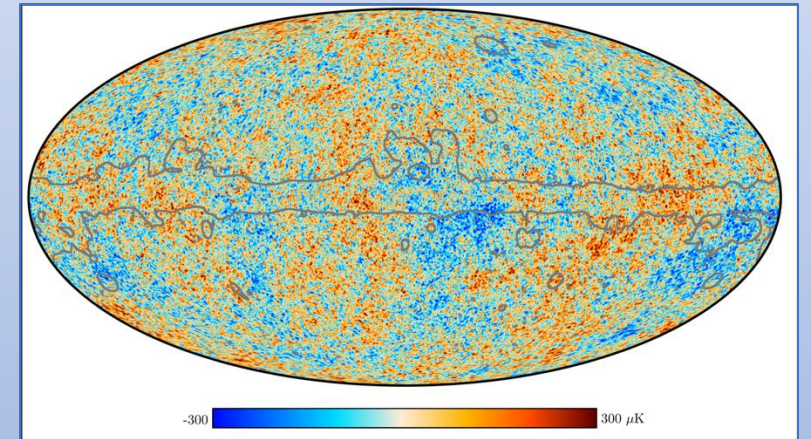
CSN II, zoom meeting 25 luglio 2025

M. Messina LNGS-INFN per conto della collaborazione PTOLEMY



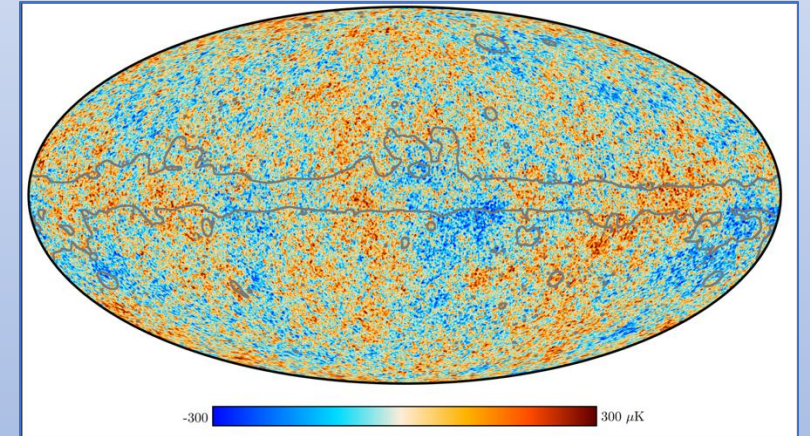
# Why we believe in the Big Bang?

1. Expansion of Universe
2. Light element abundances
3. Cosmic Microwave Background



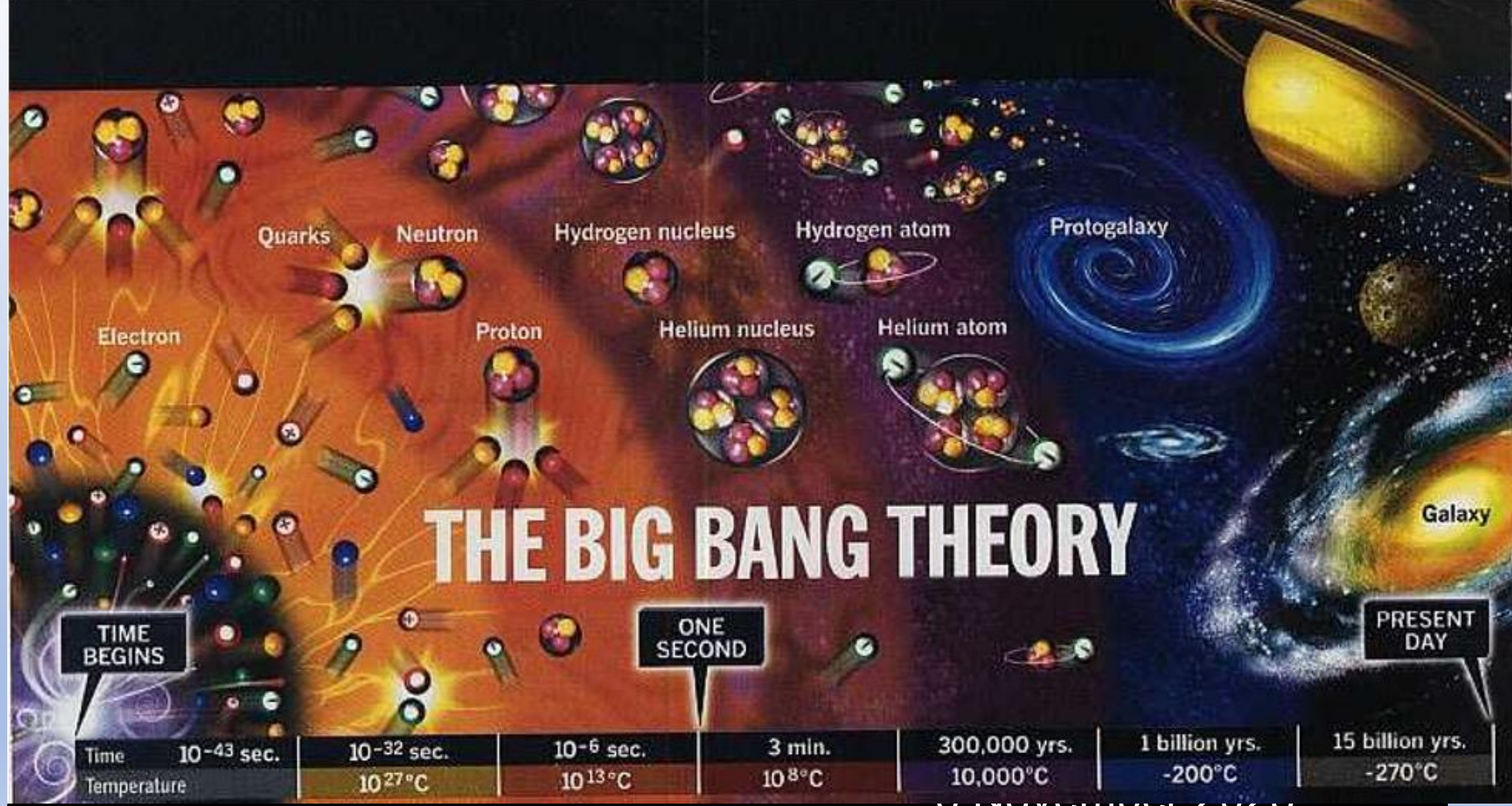
# Why believe Big Bang?

1. Expansion of Universe
2. Light element abundances
3. Cosmic Microwave Background
4. **Cosmic Neutrino Background**

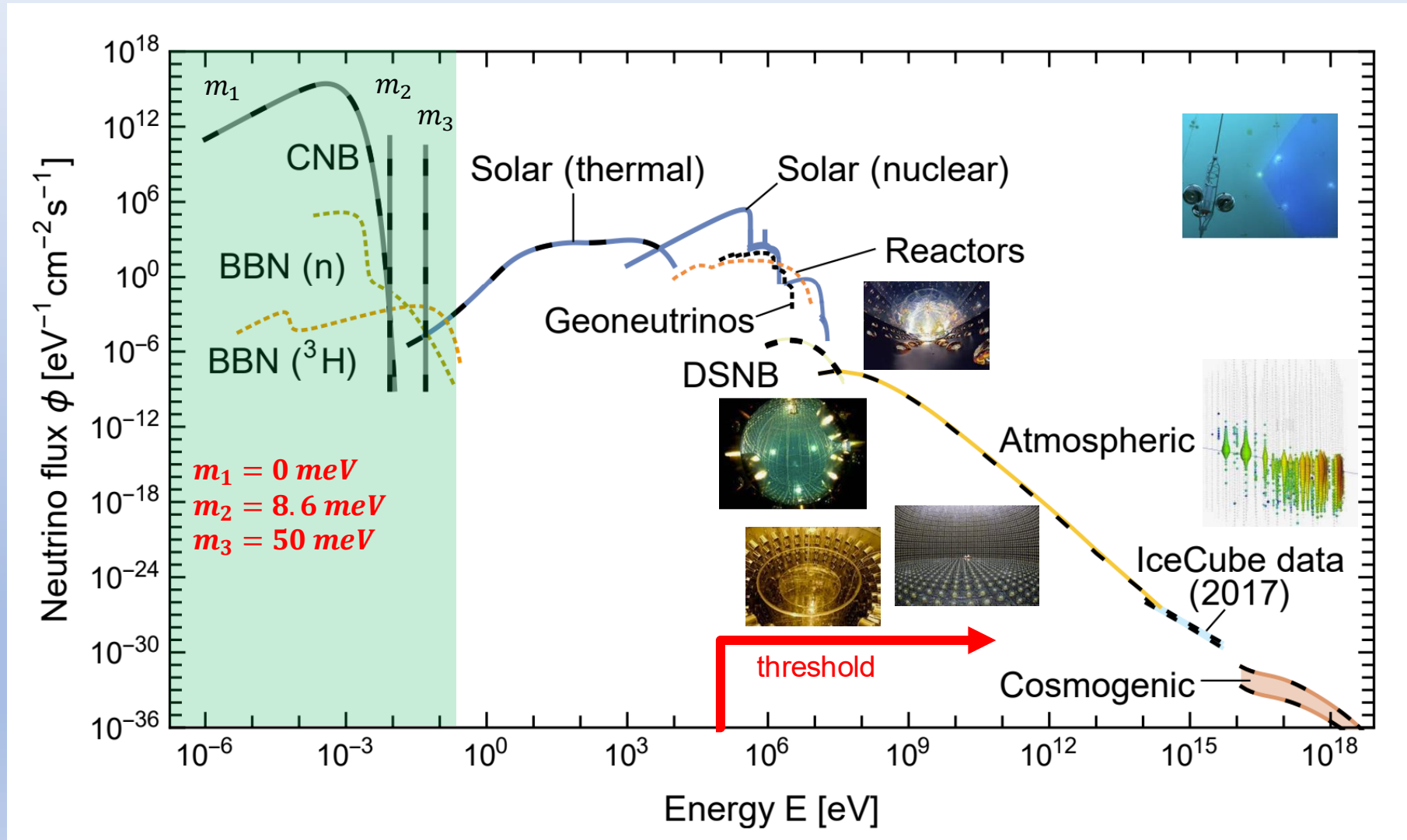




# THE BIG BANG THEORY



# Grand Unified Neutrino Spectrum





Where to look for  $C\nu B$ ?

- or -

Is it possible to detect 0.010 eV neutrinos?

# Induced beta decay

PHYSICAL REVIEW

VOLUME 128, NUMBER 3

NOVEMBER 1, 1962

## Universal Neutrino Degeneracy

STEVEN WEINBERG\*

*Imperial College of Science and Technology, London, England*

(Received March 22, 1962)

$$m_\nu = 0$$

Modern cosmological theories imply that the universe is filled with a shallow degenerate Fermi sea of neutrinos. In the steady state and oscillating models (and perhaps also the "big bang" theories) it can be shown rigorously that the proportion of filled neutrino levels (plus the proportion of filled antineutrino levels) is precisely one up to a finite Fermi energy  $E_F$ . The proof takes into account both absorption and the repulsive effects of already filled levels on neutrino emission. Experiment shows that  $E_F \leq 200$  eV for antineutrinos and  $E_F \leq 1000$  eV for neutrinos. The degenerate neutrinos could be observed (if  $E_F > 10$  eV) by looking for apparent violations of energy conservation in  $\beta^-$  decay. In the steady state and evolutionary cosmologies  $E_F$  is much too low to ever be observed, but in the oscillating cosmologies  $E_F \simeq 5R_c$  MeV, where  $R_c$  is the minimum radius of the universe in units of its present radius; thus experiment already shows that the universe will contract by a factor over  $10^3$ , if at all. Astronomical evidence plus Einstein's field equation (without cosmological constant) require in an oscillating cosmology that  $E_F < 2 \times 10^{-8}$  eV (so  $R_c < 10^{-9}$ ) and suggest that higher energy neutrinos may represent the bulk of the energy of the universe. A model universe incorporating this idea is constructed.

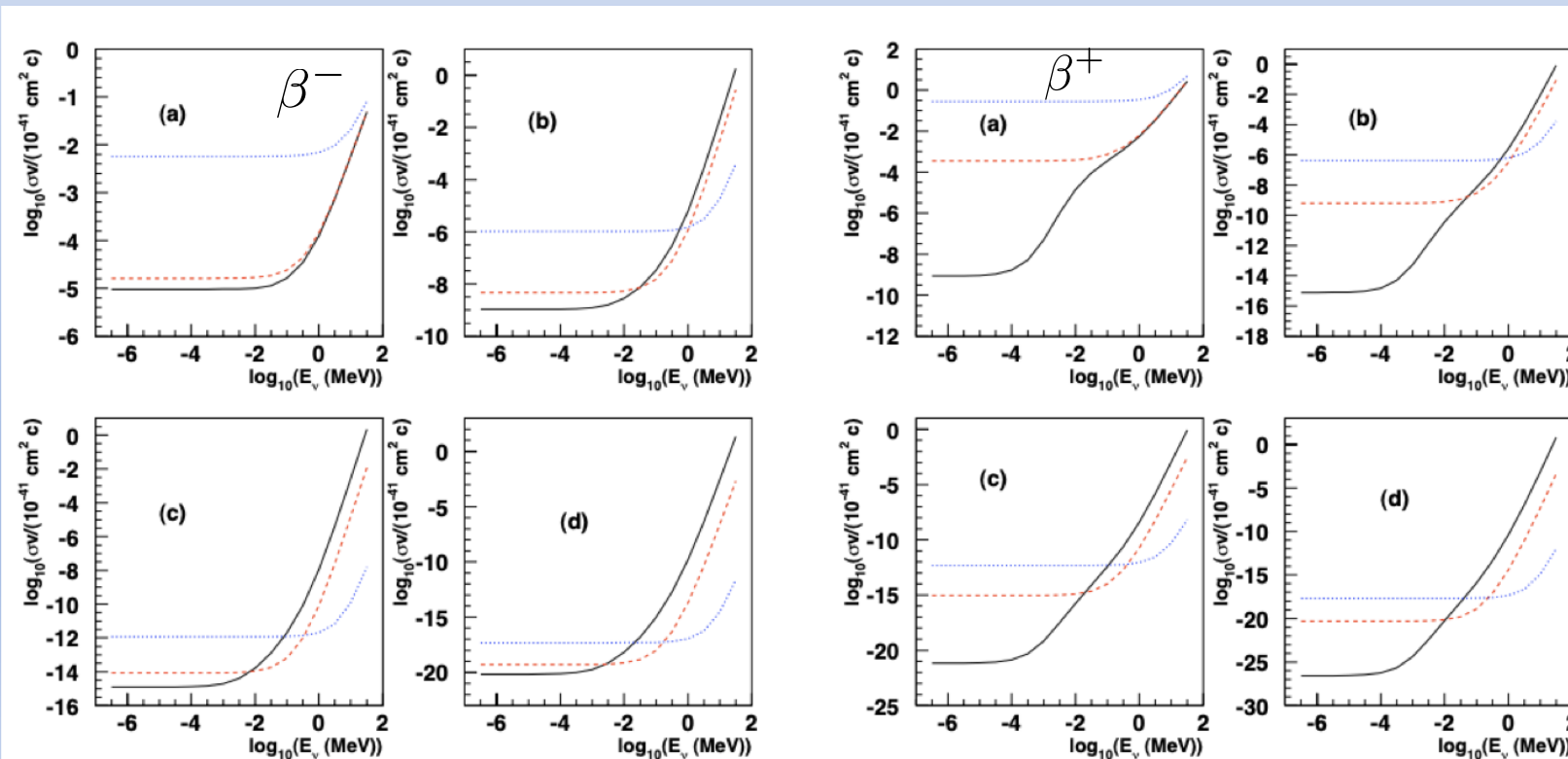
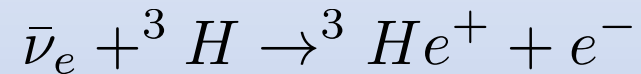
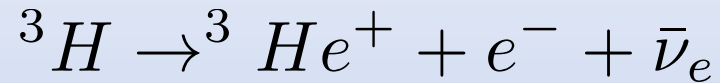
$$n_i = \frac{1}{e^{(\epsilon_i - \mu)/k_B T} + 1}$$
$$\frac{\mu}{k_B T} < 0.1$$

Cocco, Mangano, Messina calculated in 2007  
the interaction cross section and the rate with  $m_\nu \neq 0$  case

JCAP 0706:015,2007

# CROSS SECTIONS

Tritium has the **largest product of capture cross section and lifetime**



Detailed evaluation on 2007

**JCAP 06 (2007) 015**

of  $\sigma\chi\tau$  renewed the dormant discussion on relic neutrino detection and paved the view to a possible experiment.

Several authors confirmed the cross section evaluation and added informations:

**J. Phys. G: Nucl. Part. Phys. 35 025001**


**JCAP 08 (2014) 038**



# Selection of target

“Highish” cross-section

“Longish” lifetime



Isotope	Decay	$Q_\beta$ (keV)	Half-life (sec)	$\sigma_{\text{NCB}}(v_\nu/c)$ ( $10^{-41}$ cm <sup>2</sup> )
<sup>3</sup> H	$\beta^-$	18.591	$3.8878 \times 10^8$	$7.84 \times 10^{-4}$
<sup>63</sup> Ni	$\beta^-$	66.945	$3.1588 \times 10^9$	$1.38 \times 10^{-6}$
<sup>93</sup> Zr	$\beta^-$	60.63	$4.952 \times 10^{13}$	$2.39 \times 10^{-10}$
<sup>106</sup> Ru	$\beta^-$	39.4	$3.2278 \times 10^7$	$5.88 \times 10^{-4}$
<sup>107</sup> Pd	$\beta^-$	33	$2.0512 \times 10^{14}$	$2.58 \times 10^{-10}$
<sup>187</sup> Re	$\beta^-$	2.64	$1.3727 \times 10^{18}$	$4.32 \times 10^{-11}$
<sup>11</sup> C	$\beta^+$	960.2	$1.226 \times 10^3$	$4.66 \times 10^{-3}$
<sup>13</sup> N	$\beta^+$	1198.5	$5.99 \times 10^2$	$5.3 \times 10^{-3}$
<sup>15</sup> O	$\beta^+$	1732	$1.224 \times 10^2$	$9.75 \times 10^{-3}$
<sup>18</sup> F	$\beta^+$	633.5	$6.809 \times 10^3$	$2.63 \times 10^{-3}$
<sup>22</sup> Na	$\beta^+$	545.6	$9.07 \times 10^7$	$3.04 \times 10^{-7}$
<sup>45</sup> Ti	$\beta^+$	1040.4	$1.307 \times 10^4$	$3.87 \times 10^{-4}$

# Expected rate: 100 gram-year exposure

JCAP 0706:015,2007

$m_\nu$ (eV)	FD (events yr <sup>-1</sup> )		NFW (events yr <sup>-1</sup> )		MW (events yrs <sup>-1</sup> )	
0.6	7.5	x0.5	90	x0.5	150	x0.5
0.3	7.5		23		33	
0.15	7.5		10		12	

Dirac

$m_\nu$ (eV)	FD (events yr <sup>-1</sup> )		NFW (events yr <sup>-1</sup> )		MW (events yrs <sup>-1</sup> )	
0.6	7.5	}	90	}	150	}
0.3	7.5		23		33	
0.15	7.5		10		12	

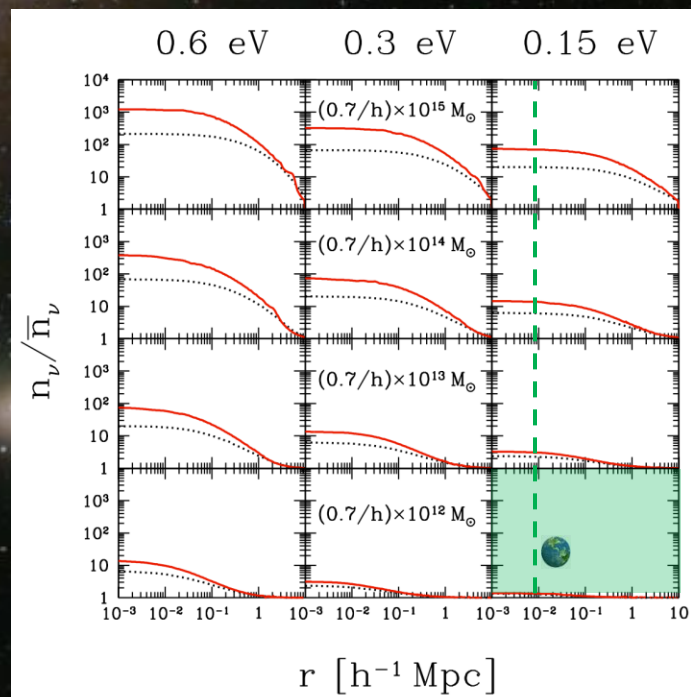
Majorana

JCAP 076:015, 2007  
JCAP 1408 (2014) 038

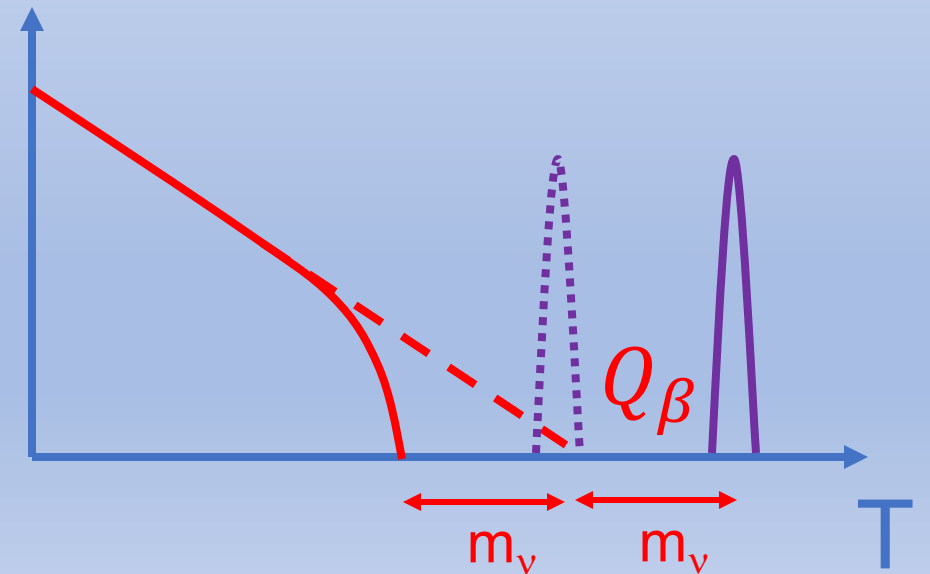
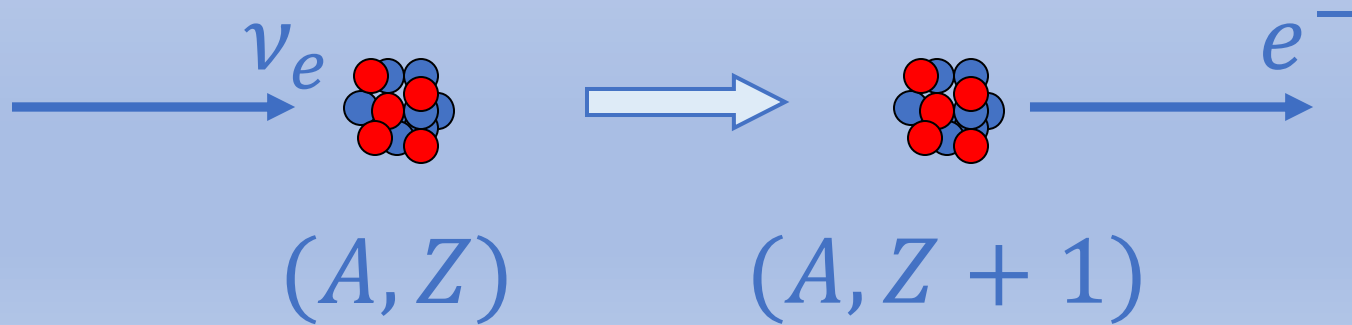
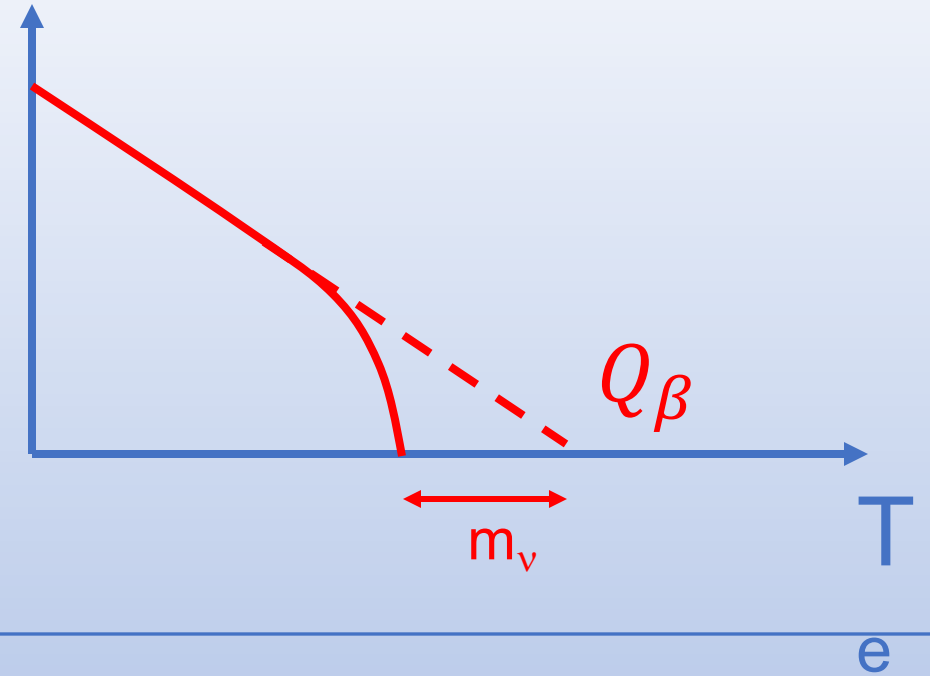
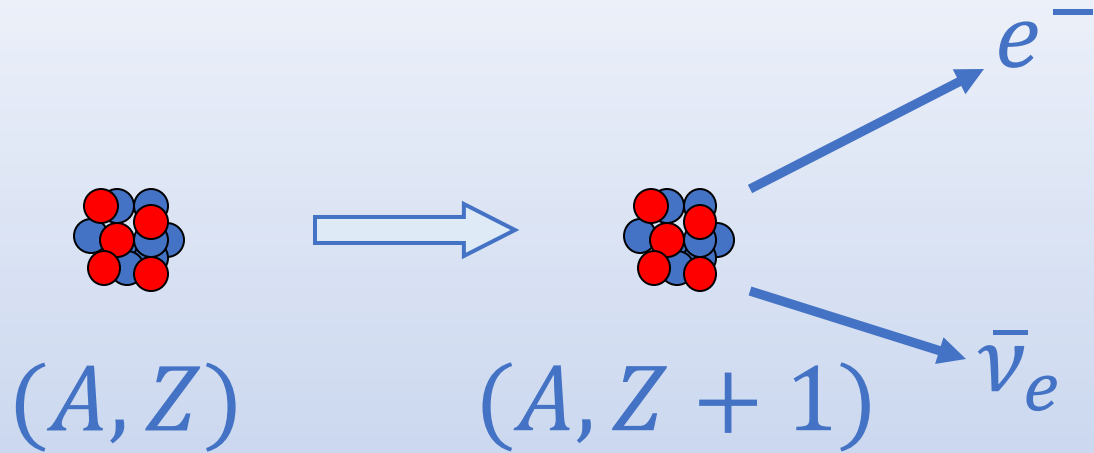
Exp



Gravitational clustering



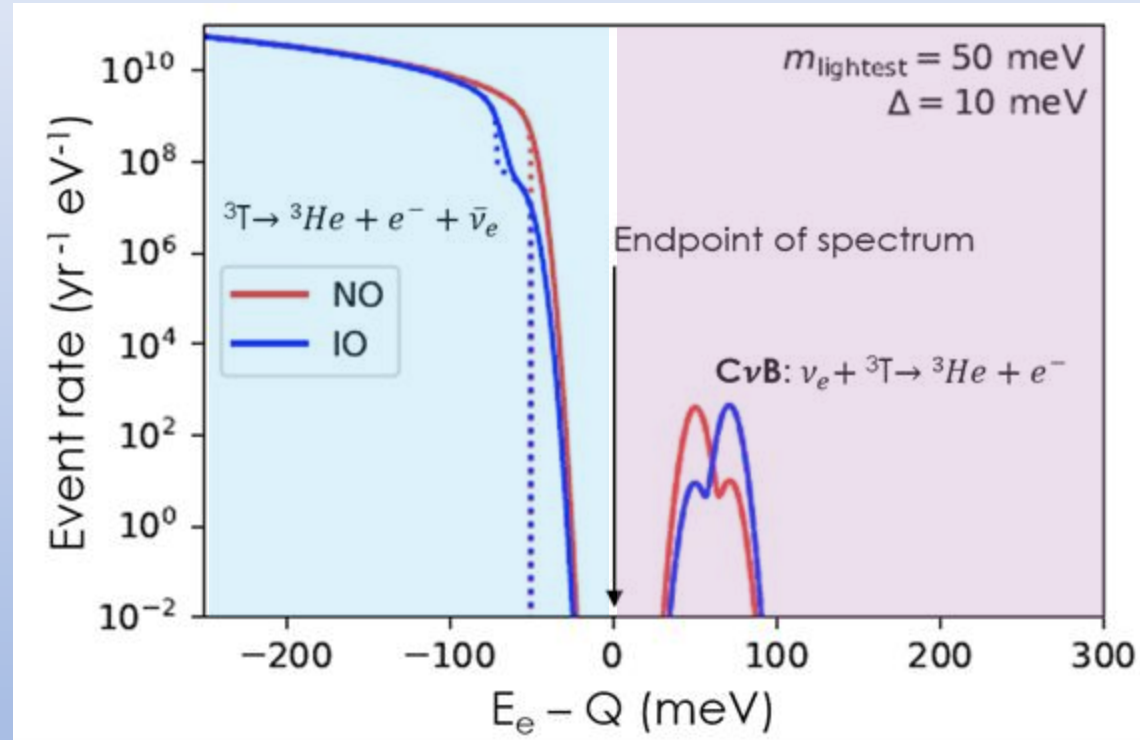
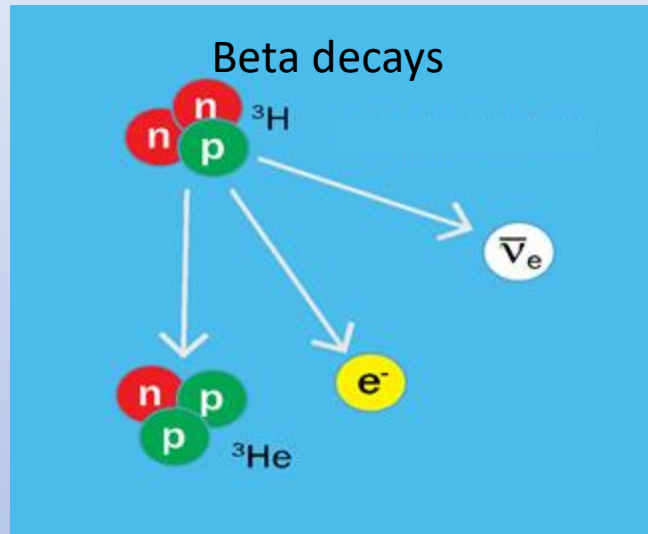
# Signature of relic neutrinos



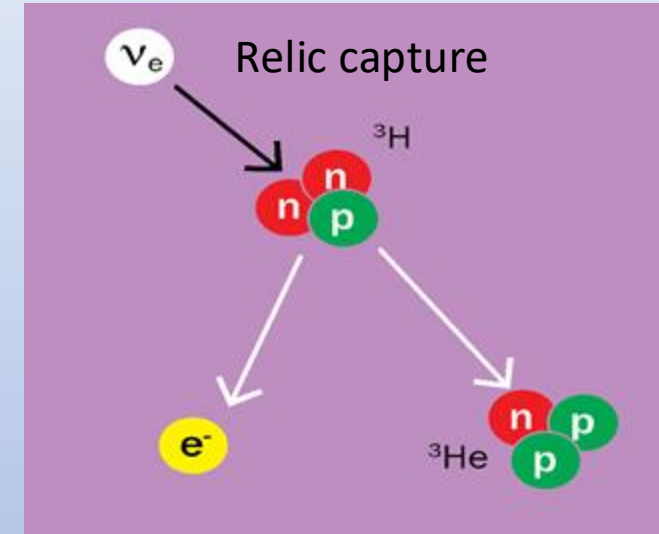


# Detection concept: Neutrino Capture on $\beta$ unstable nuclei

Proposal for their direct experimental detection in  
[AG.Cocco, G.Mangano, M.Messina JCAP 06(2007)015]



M.G. Betti et al JCAP  
07(2019)047



Electron flavor expected with  
 **$m \gtrsim 50 \text{ meV}$**   
from neutrino oscillations

CνB Detection Requires:  
energy resolution set by  $m_\nu$

Requirements on energy resolution for neutrino mass measurement are weaker than relic case (50 meV)

## Letter of Intent for the PTOLEMY project document for INFN CNS2

Data: 25/06/2025  
Rev. 0.2.1 - 25/06/2025

### Authors and approvals

Editors	Checked by	Approved by
G.Cavoto, M.Messina	G.Cavoto, M.Messina, F.Virzi, F.M. Pofi, A.G.Cocco	PTOLEMY collaboration

### Distribution

- Public document

### Revision history

Rev.	Data	Description of updates	Authors/Editors
0.2.1	25/06/25	First draft	G.Cavoto, M.Messina

# Lol:

2 years project (**Phase0**) to deliver a **CDR** for neutrino mass measurement.

We ask **support** above all to strengthen the collaboration and simultaneously address the remaining technical issue.



## CDR:

will contain the **Demonstrator design** aiming at demonstrating the technology

**REMARK:**

**demonstrator = detector for a competitive neutrino mass measurement**

# PTOLEMY first physics goal: neutrino mass

- Once the CDR will be delivered, in two years from now, the schedule for neutrino mass measurement will be clarified.
- Novelty elements of the PTOLEMY detector:
  - Novel concept of a **compact dynamic electromagnetic filter**
  - Atomic tritium on a graphene support (monoatomic layer of C atoms)
  - Differential energy measurement with unprecedented energy resolution

# PTOLEMY Demonstrator

aims at showing the capability of the PTOLEMY collaboration to master technologies

**As described in the Lol we focus on a Phase0 to demonstrate the technology and this is the topic on which we ask the CSN II to approve the project:**

M1: Experimental Demonstration of the Filter Concept

M2: Feasibility study of tritium on graphene production

M3: Realistic design of 1 mug tritium target

M4: RF detection

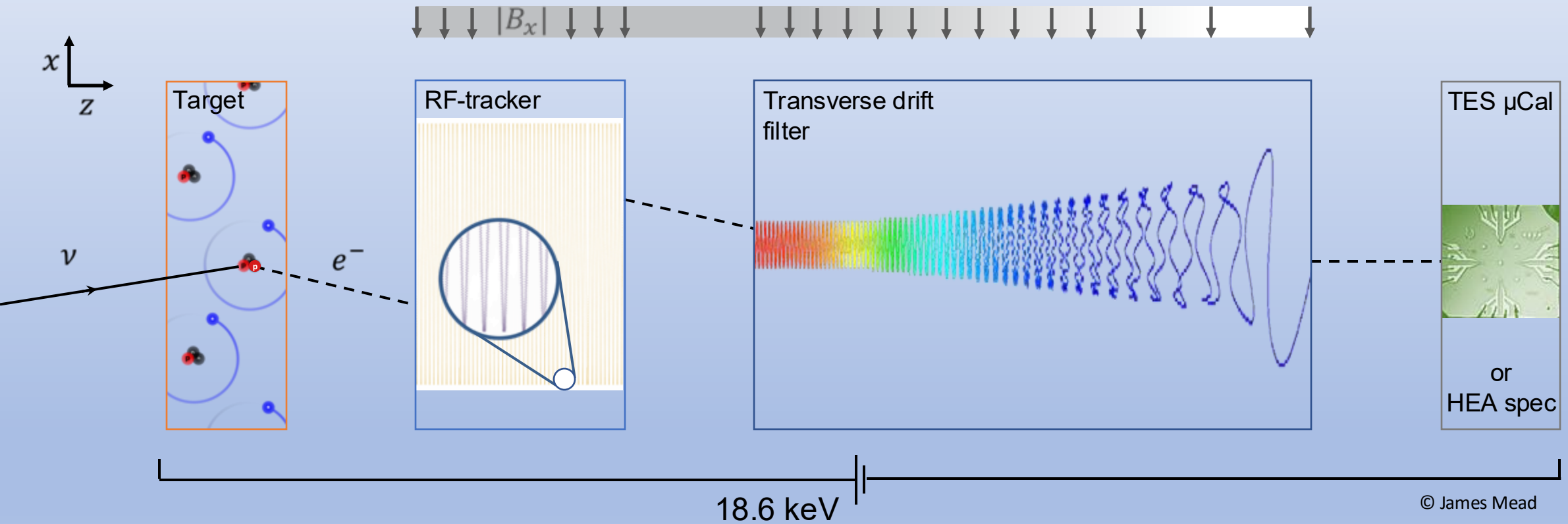
M5: Full simulation of a filter for neutrino mass experiment

M6: Conceptual design report ready (Detailing Phase1 time scale )



# PTOLEMY detector scheme

PonTecorvo / PrinceTon Observatory for Light Early-universe Massive-neutrino Yield

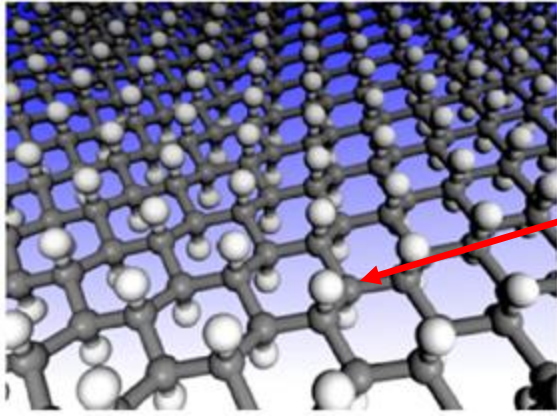


© James Mead

$$E_{total} = q(V_{TES} - V_{target}) + E_{RF} + E_{cal}$$

# Hydrogen and Deuterium loading on graphene at Roma1 and Roma3

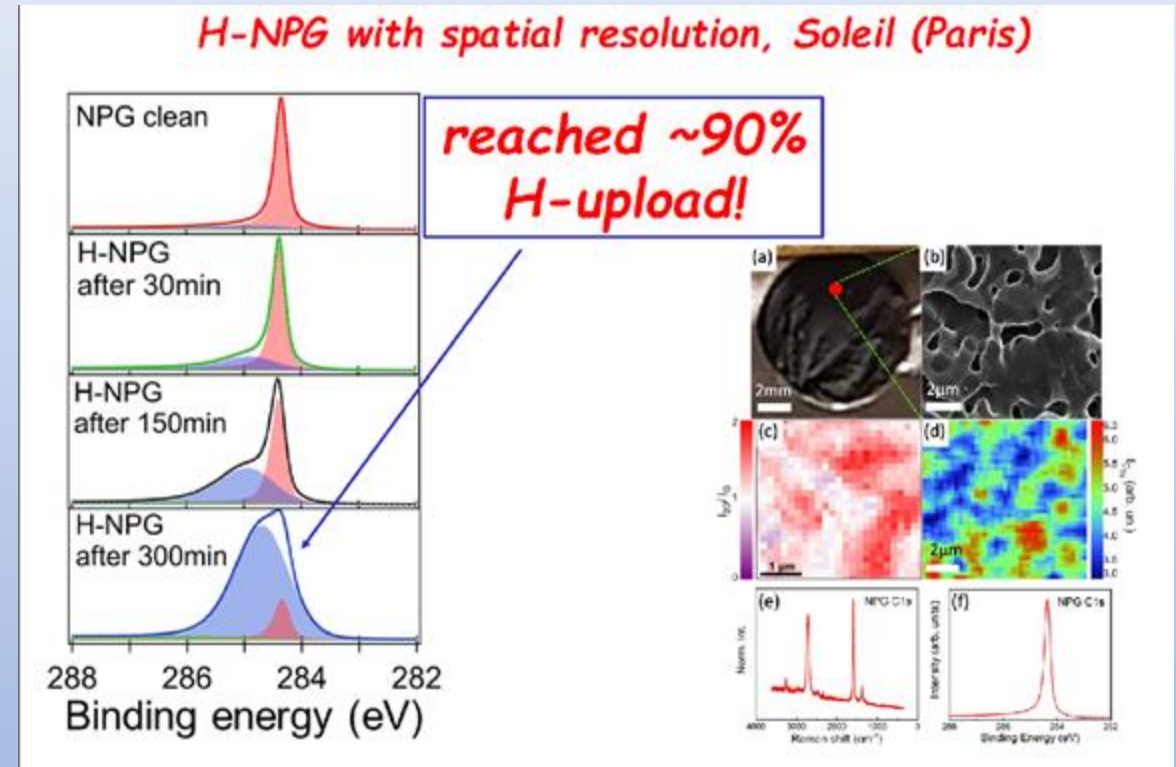
atomic H as a tool to '*pinch*' the  $sp^2$  bonds towards a  $sp^3$  configuration while maintaining the planar nature of graphene



T-chamber in Rome side view:



H on Nanoporous Graphene (NPG):

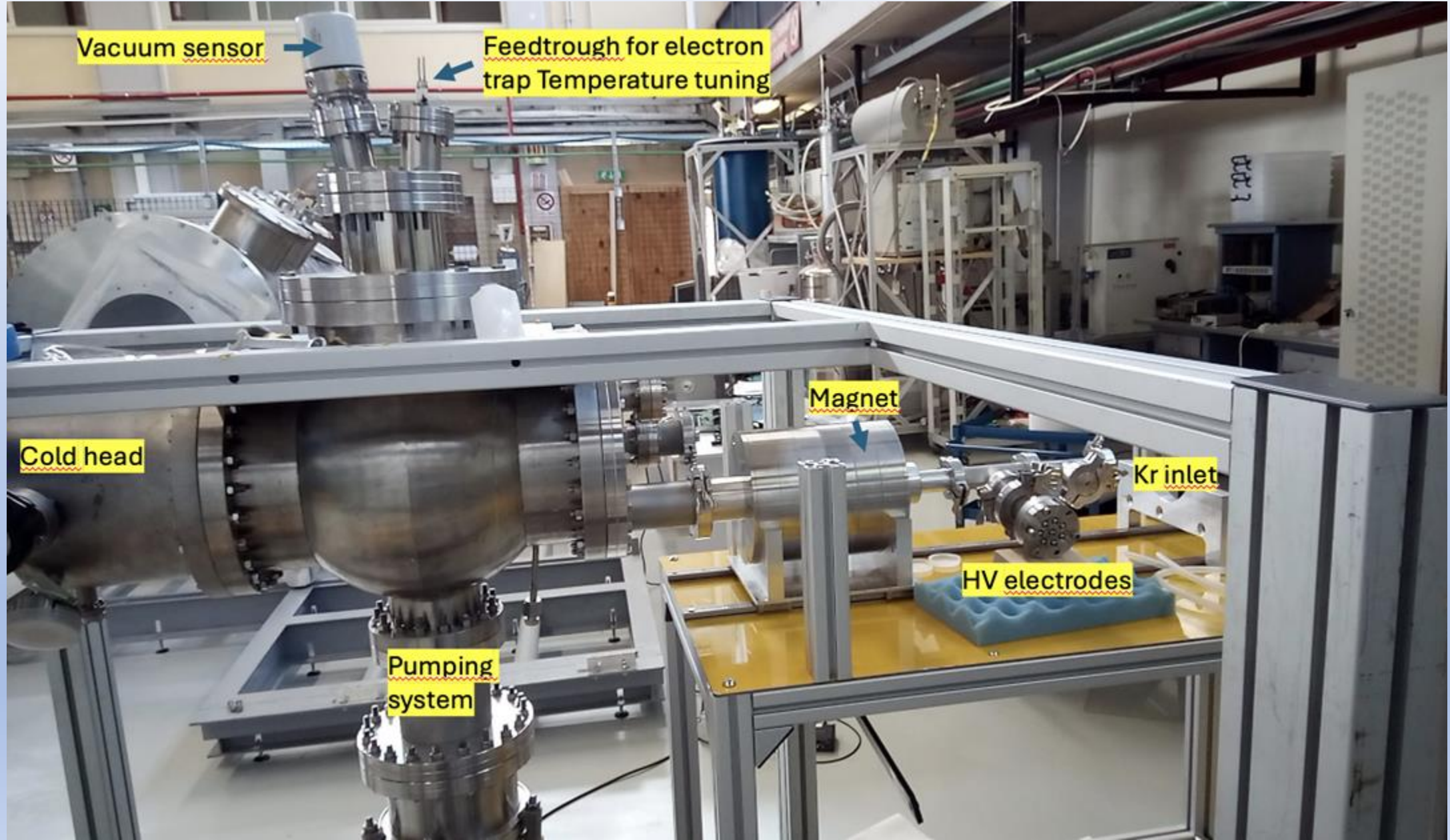


UKAEA's Active Gas Handling System (tritium for JET, EU Tokamak) for **feasibility study & design requirement** of a new T loading chamber

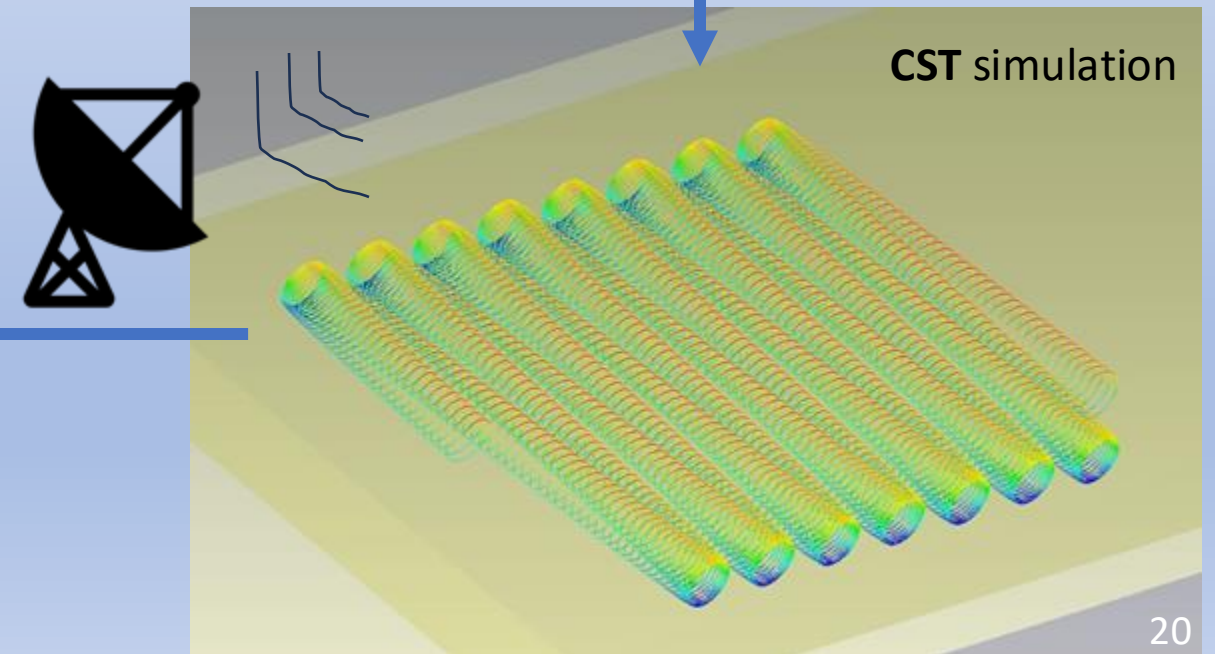
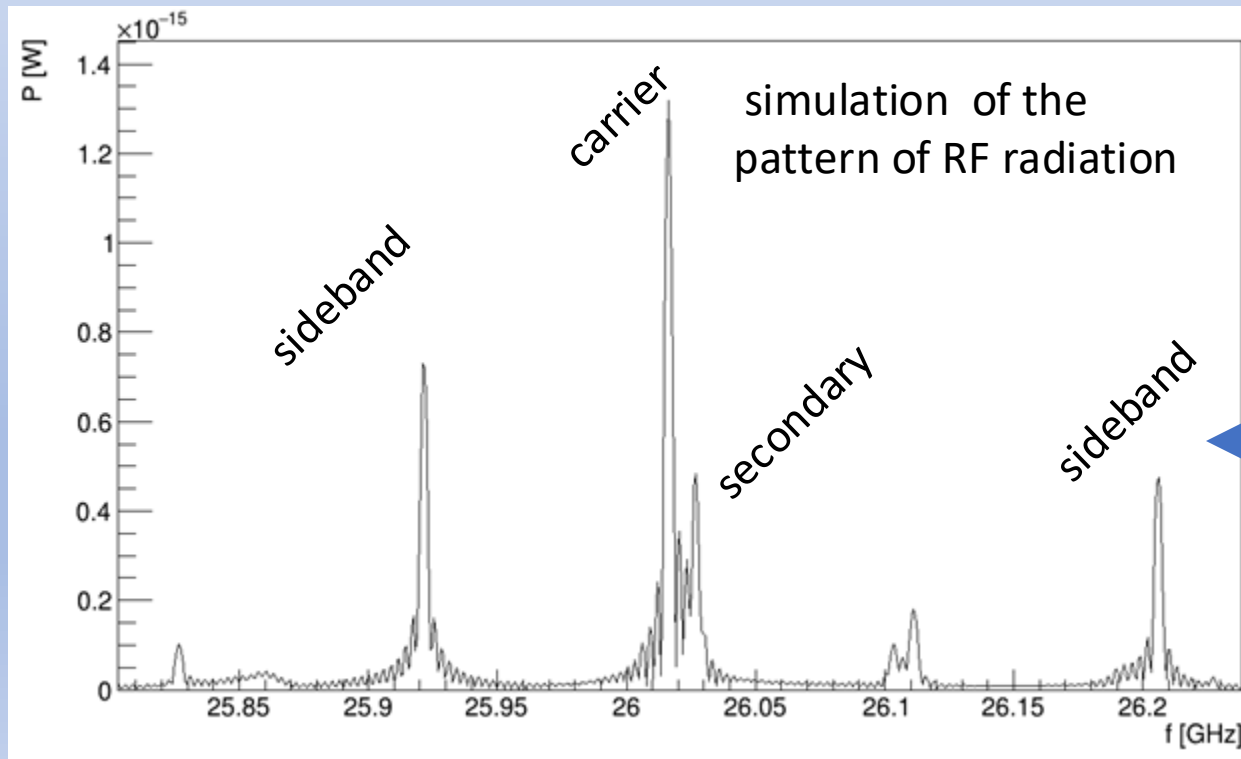
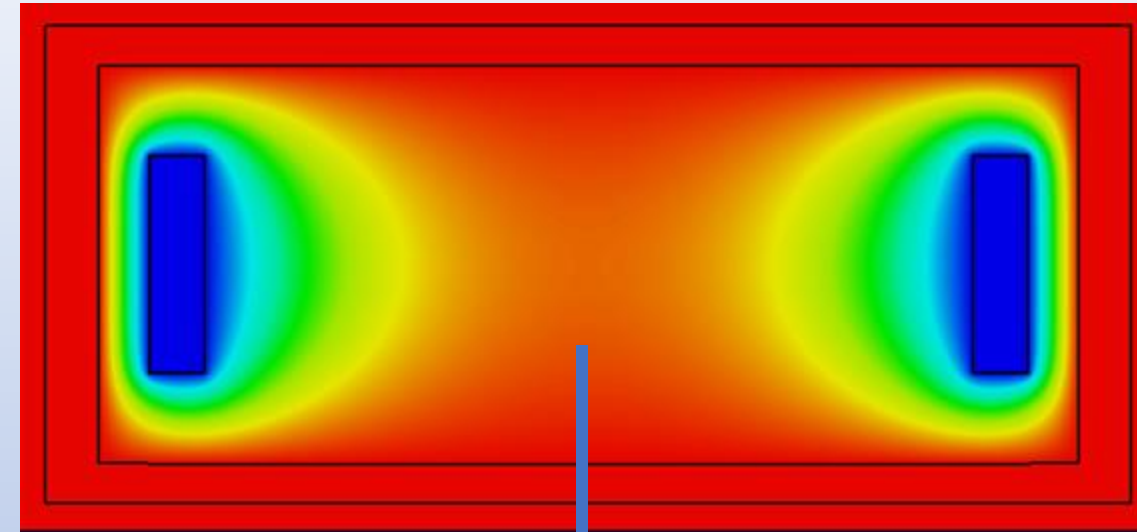
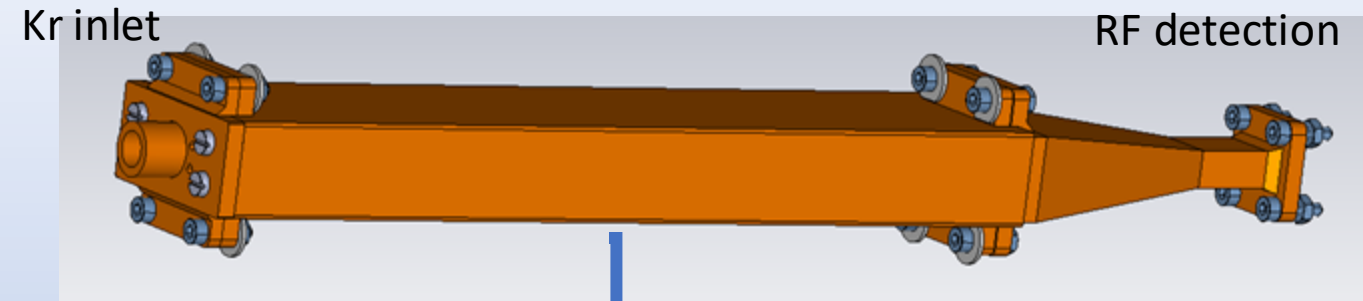
Quote received few months ago



# RF detection setup at LNGS: electron trap



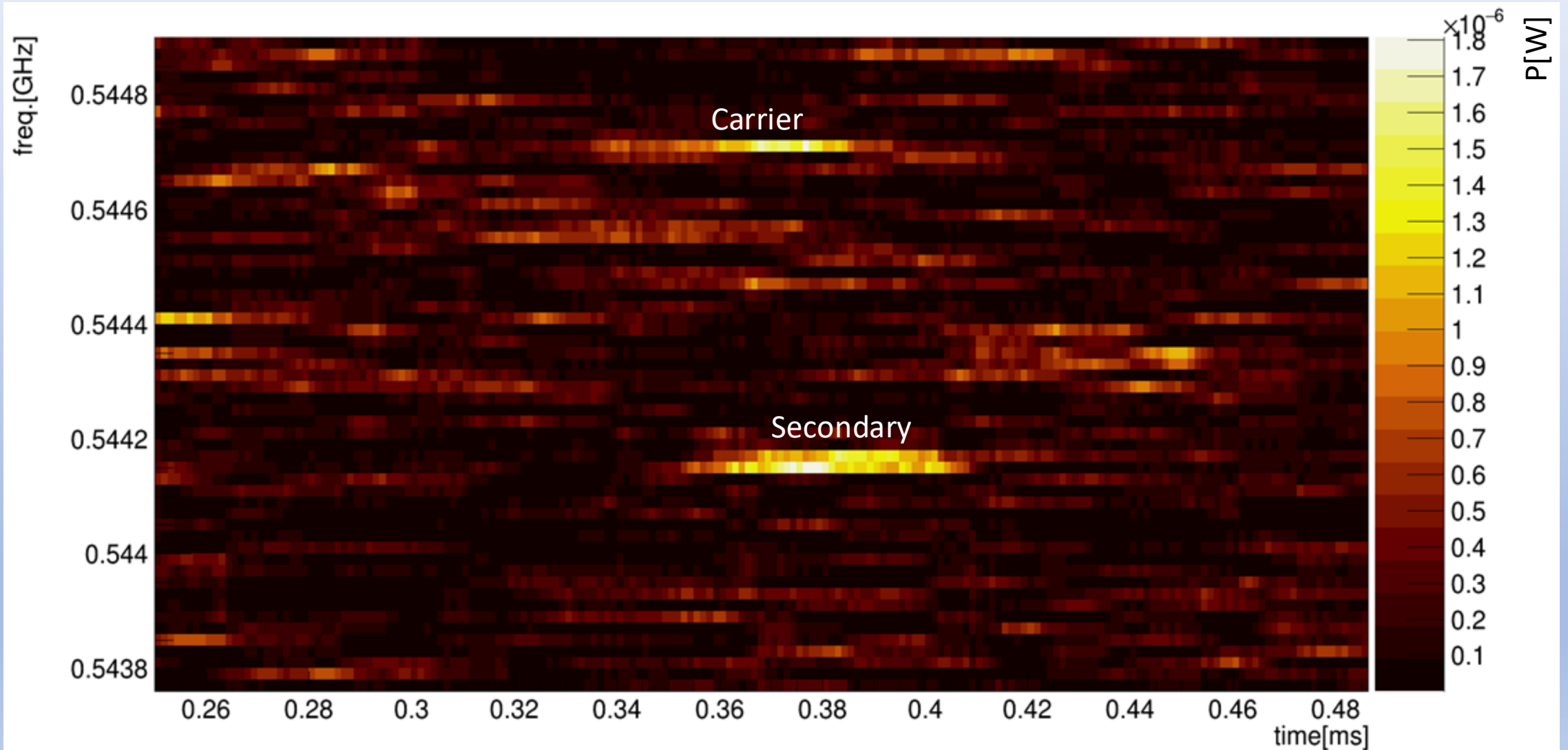
# Electron cyclotron RF radiation studies





# Candidate electron events

RF emission frequency vs time of single electron in the permanent magnet

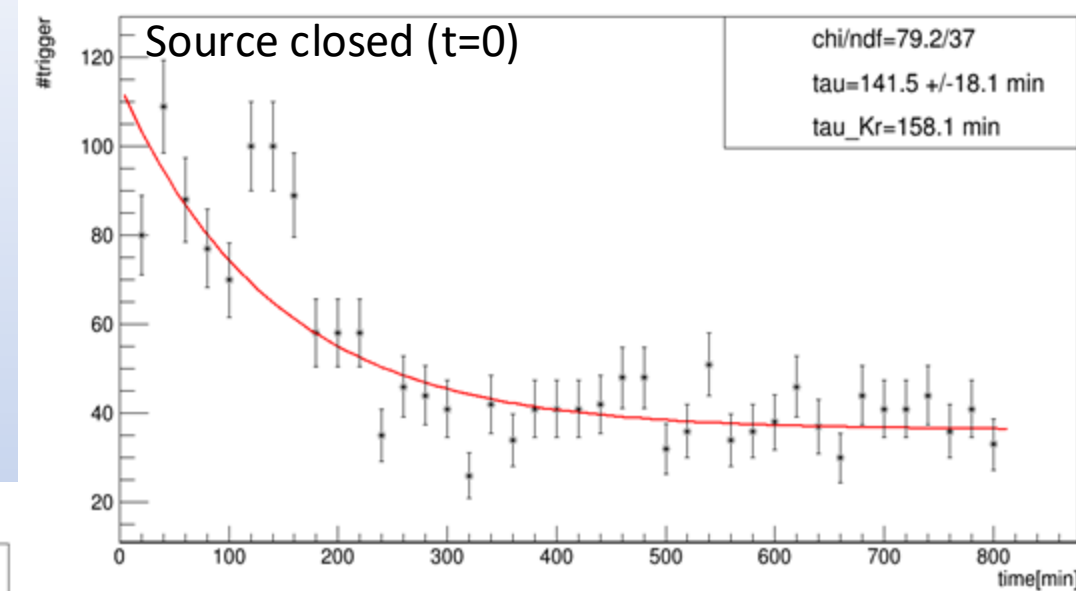
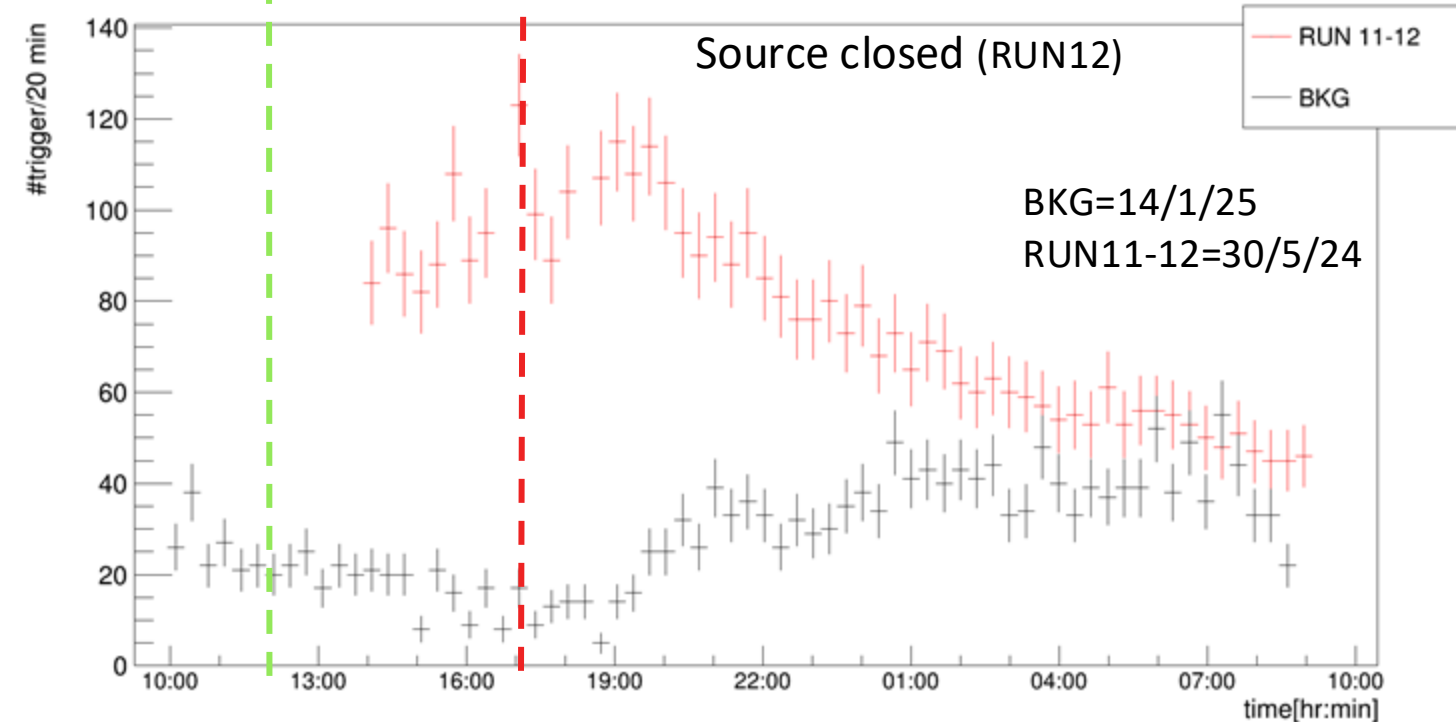


# Electron signature: trigger rate (data)

Source:  $^{85}\text{Rb} \rightarrow ^{83\text{m}}\text{Kr}$   
Lifetime Kr = 158.1 min

Source open (RUN11)      Source closed (RUN12)

Trigger Rate versus time



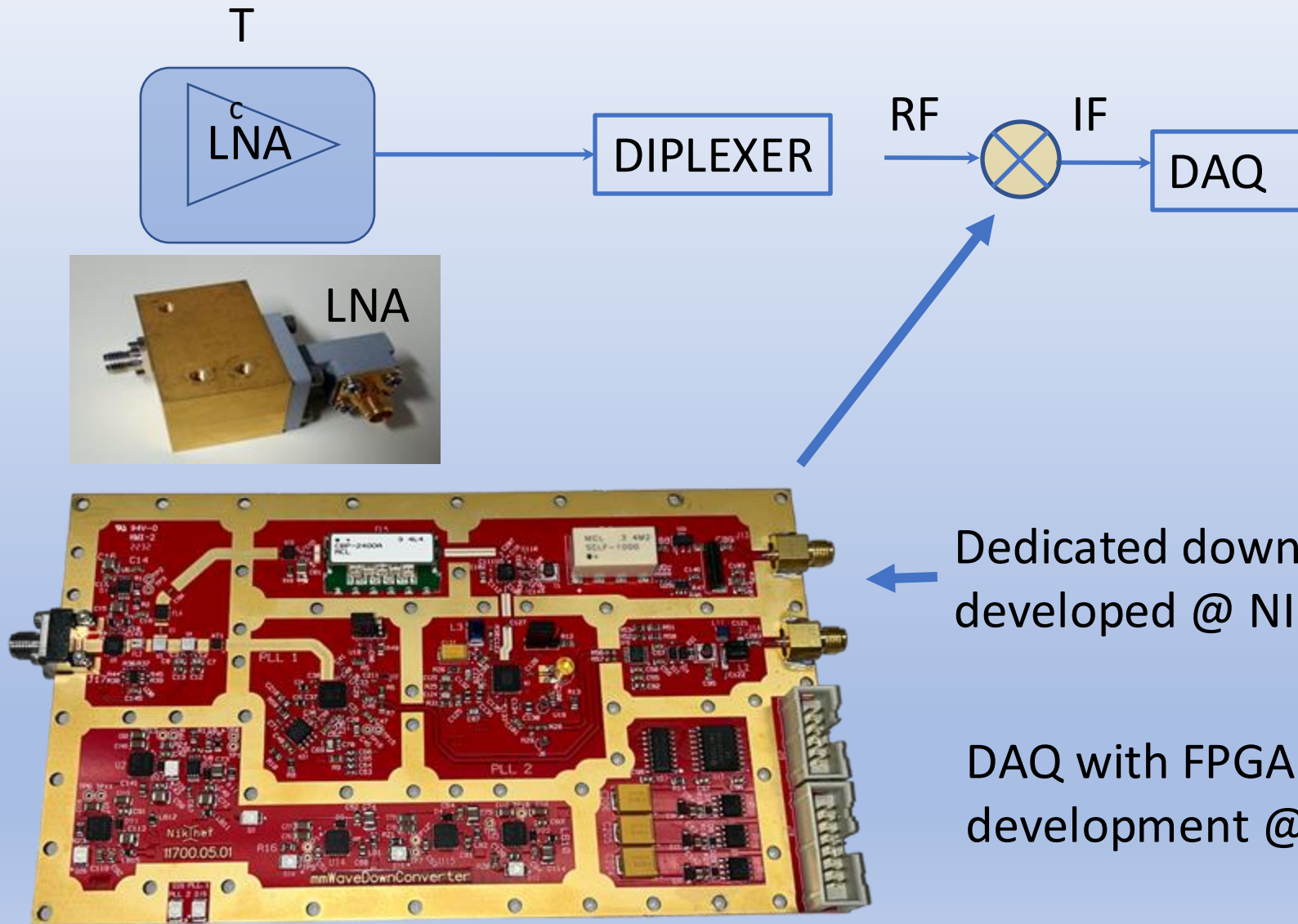
•  $\text{Freq}_{\text{RF}} = 26$   
GHz

•  $P_{\text{RF}} = 1$  fW

**Electronic chain:**  
Low noise amplification +  
down-conversion

•  $\text{Freq}_{\text{IF}} = 550$  MHz  
•  $P_{\text{RF}} = 2$   $\mu\text{W}$

# RF readout electronics



RTO64 used for DAQ

Dedicated downconverter developed @ NIKHEF

DAQ with FPGA trigger under development @ NIKHEF

# PTOLEMY filter: years of development

Particle trajectories can be calculated/predicted analytically:

- Given the following field configuration:

$$V(x, y, z) = T'_\perp \sin\left(\frac{y}{\lambda}\right) e^{-z/\lambda}$$

$$E_x = 0$$

$$E_y = \frac{T'_\perp}{\lambda} \cos\left(\frac{y}{\lambda}\right) e^{-z/\lambda}$$

$$E_z = -\frac{T'_\perp}{\lambda} \sin\left(\frac{y}{\lambda}\right) e^{-z/\lambda}$$

$$B_x = B_0 \cos\left(\frac{x}{\lambda}\right) e^{-z/\lambda}$$

$$B_y = 0$$

$$B_z = -B_0 \sin\left(\frac{x}{\lambda}\right) e^{-z/\lambda}$$

- With  $x \ll \lambda$ ,

$$\frac{E_z}{B_x} \hat{y} = -\frac{\mu}{qB_x} \frac{dB_x}{dz} \hat{y}$$

$$\approx \frac{E_z}{B} = -\frac{\mu}{qB} \frac{\partial B_x}{\partial z}$$

$$-\frac{T'_\perp}{B_0 \lambda} \sin\left(\frac{y_0}{\lambda}\right) = \frac{\mu}{q\lambda}$$

- Solving for  $T'_\perp$  and rewriting  $\mu = \frac{T_{\perp 0}}{B_0 e^{-z_0/\lambda}}$ ,

$$T'_\perp = -\frac{T_{\perp 0} e^{z_0/\lambda}}{q \sin\left(\frac{y_0}{\lambda}\right)} = -\frac{T_{\perp 0} e^{z_0/\lambda}}{\sin\left(\frac{y_0}{\lambda}\right)} [\text{eV}]$$

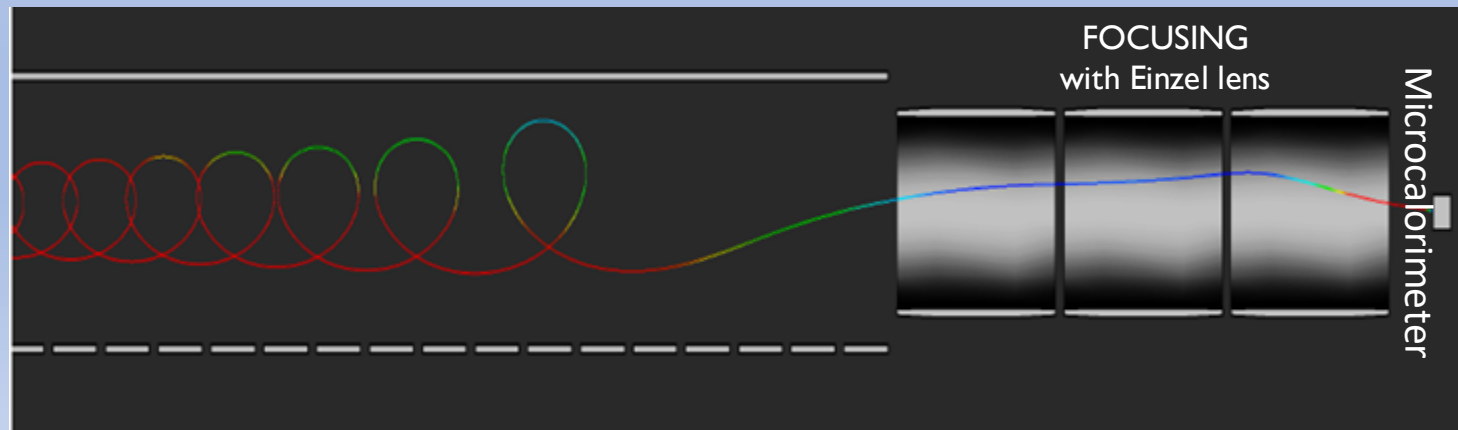
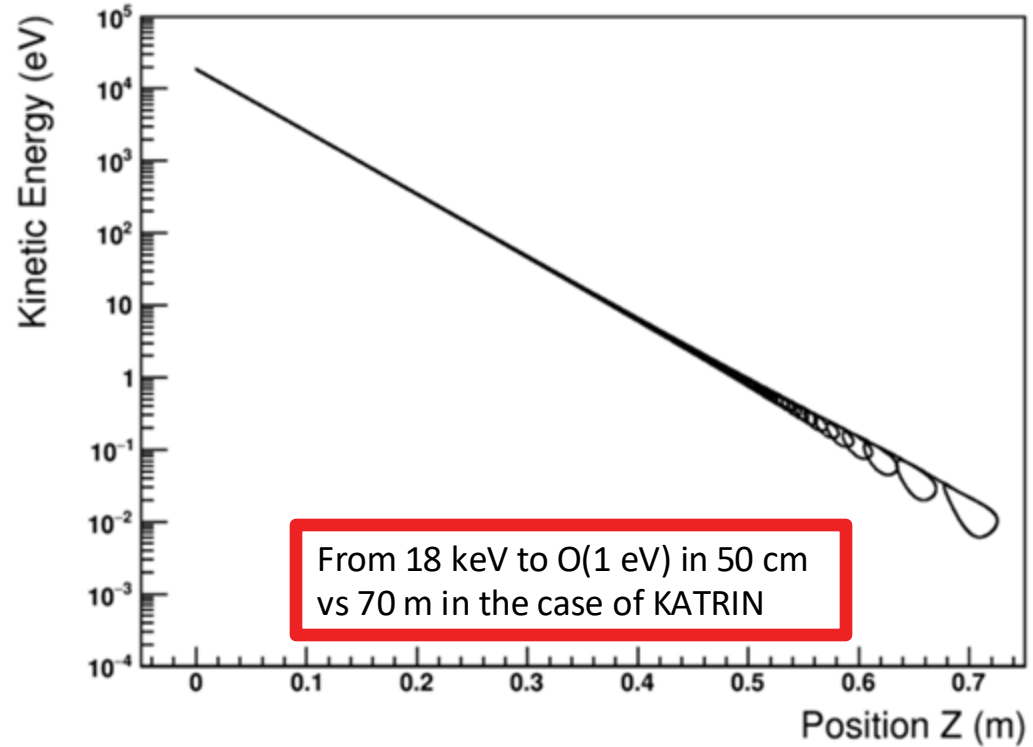
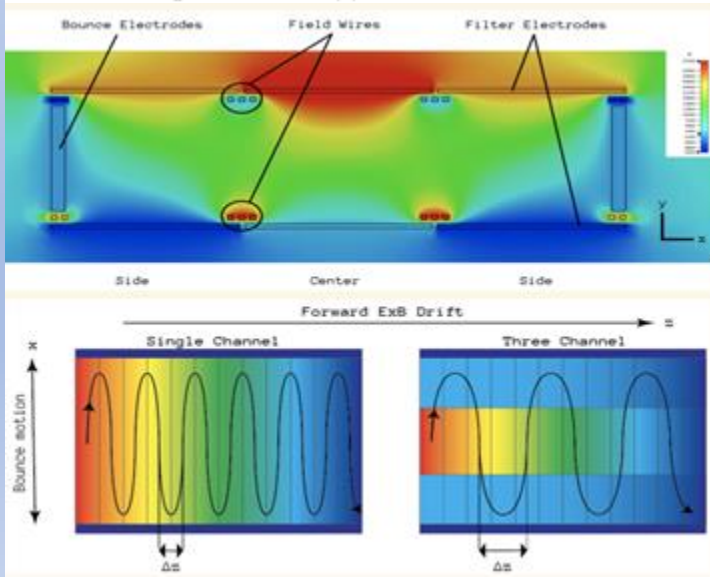
- The drift balancing condition along z is:

$$V_{E \times B}^y(z) = V_{\nabla B}(z)$$

$$\frac{E \times B}{B_x^2} = -\frac{\mu \times \nabla_\perp B(z)}{qB(z)}$$

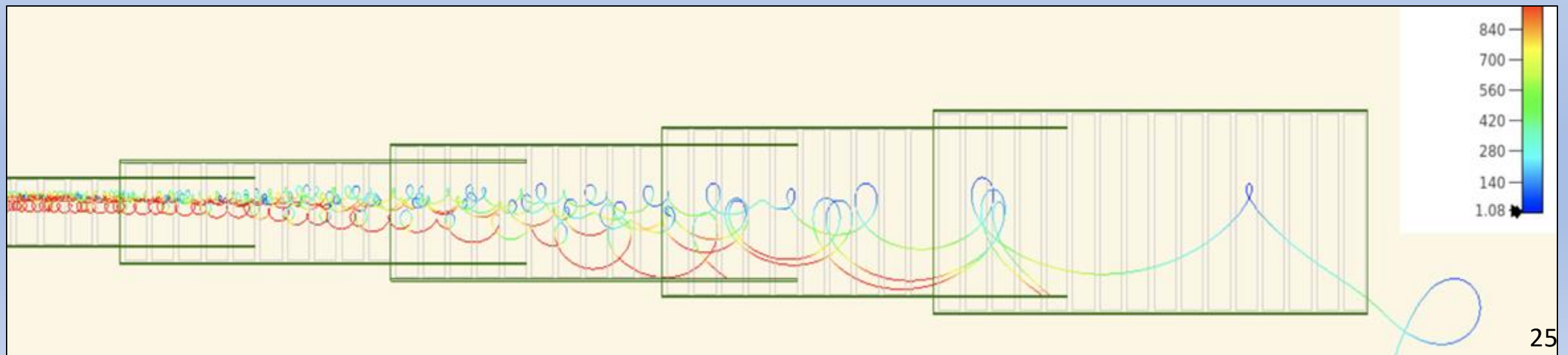
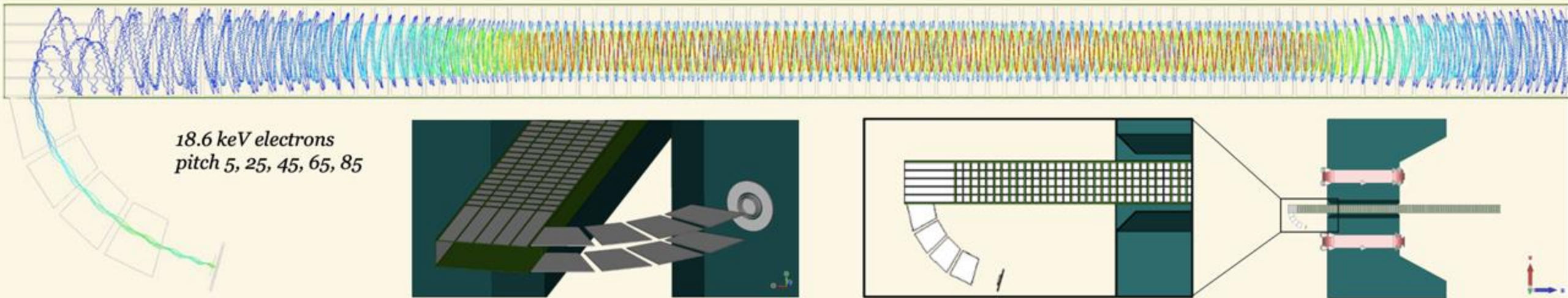
- Rearranging,

$$T_{\perp 0} = -T'_\perp \sin\left(\frac{y_0}{\lambda}\right) e^{-z_0/\lambda}$$



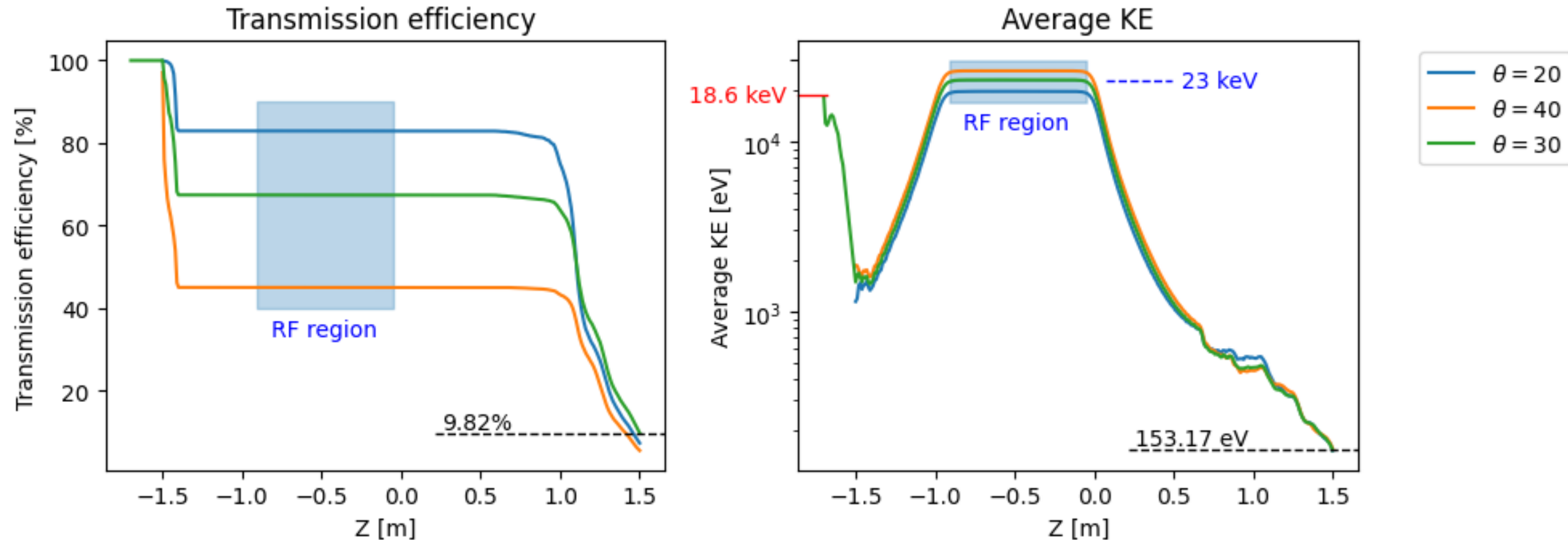


# Simulation (Princeton) of 100 ng target mass



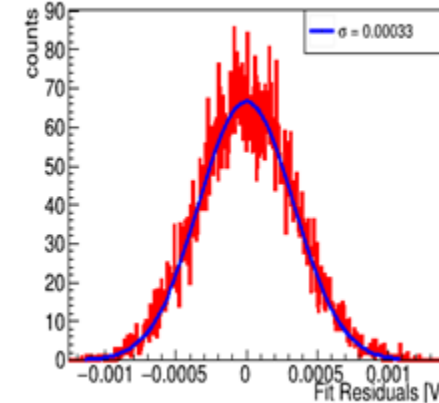
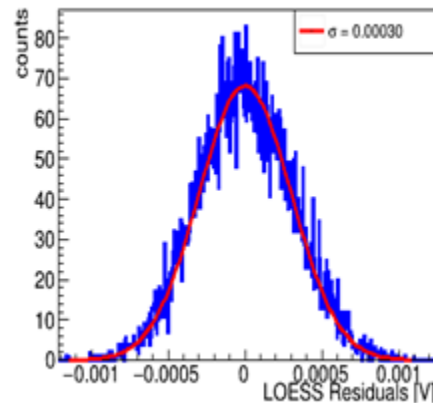
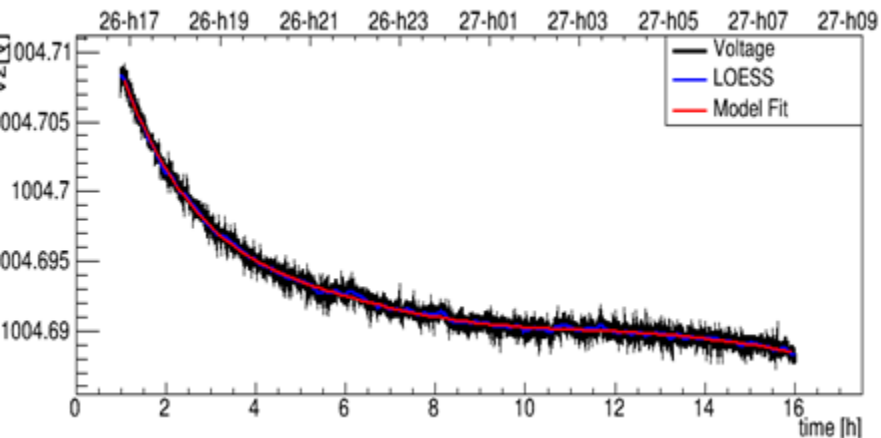
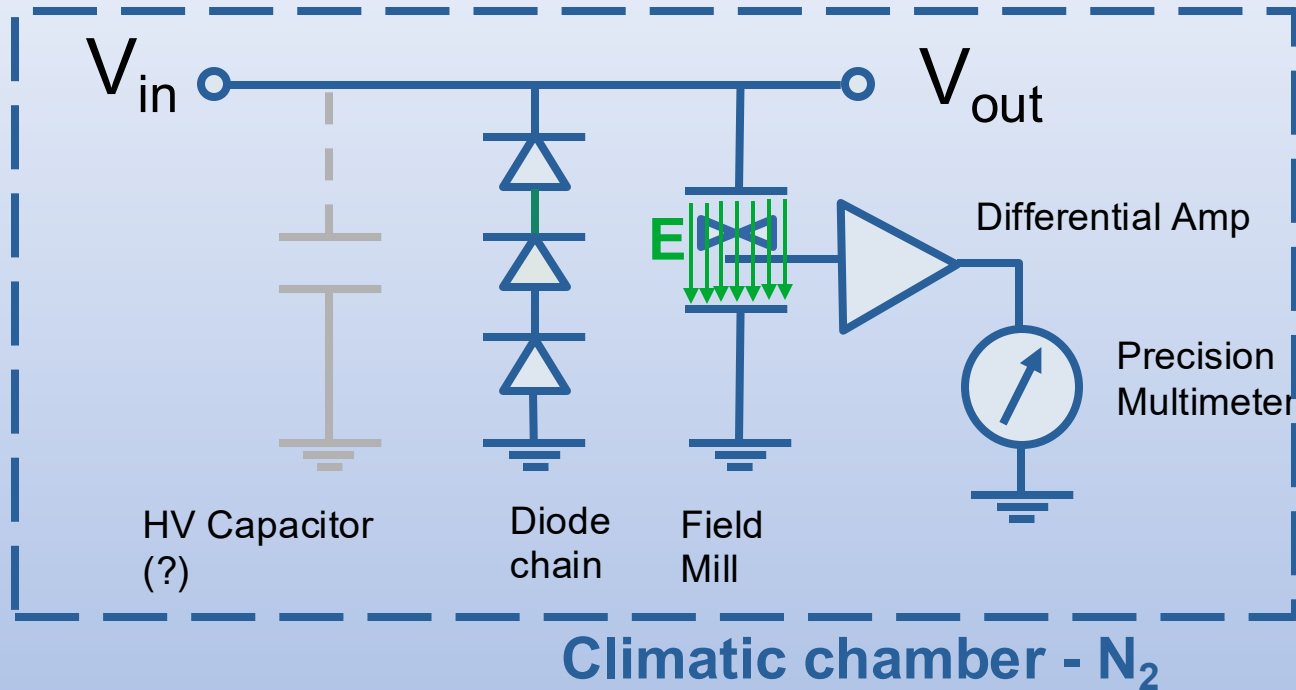
Showing  $\theta = 20 \pm 5, 30 \pm 5, 40 \pm 5$   
averaged over  $\Phi$  bins

Average over  $\phi = [0, 2\pi], N = 81600$



# HV High precision stability (LNGS)

$$E_{total} = q(V_{TES} - V_{target}) + E_{RF} + E_{cal}$$



Single board  
(1 kV)

$$\sigma = 0.3 \text{ mV}$$

# Reference source

Total number of  $\beta$  emissions from tritium:

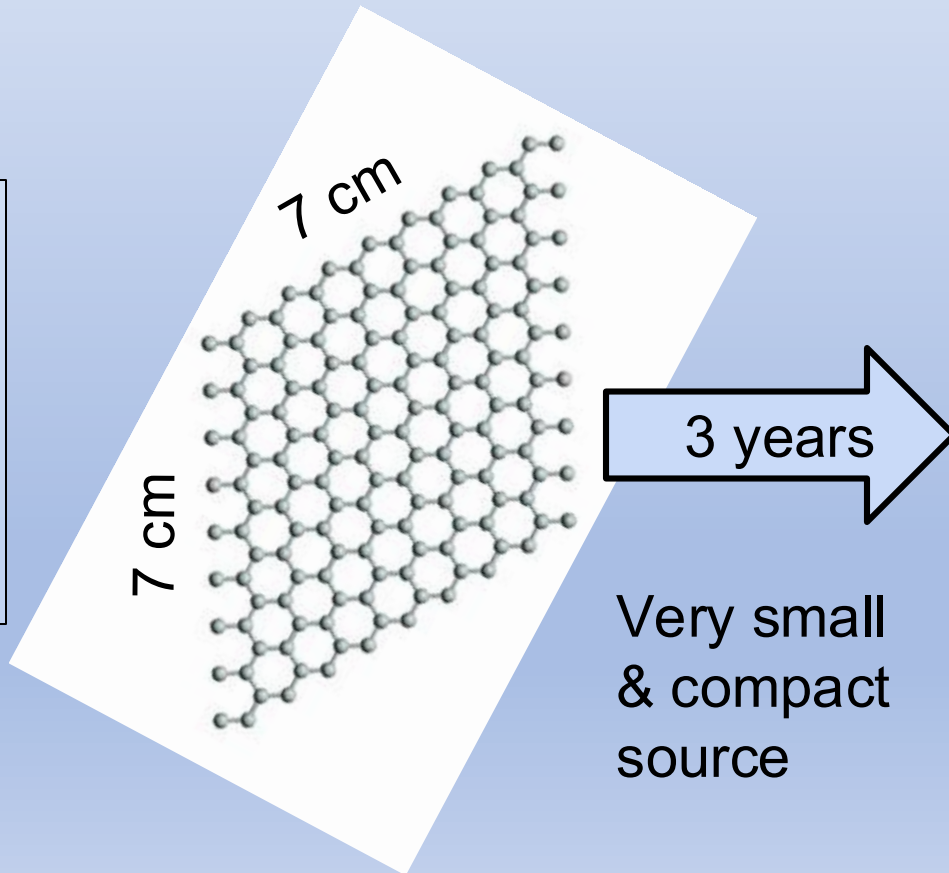
$$N_{dec} = \left( \frac{m_{\text{source}} \mathcal{N}_A}{A({}^3\text{H})} (1 - e^{-t_{\text{expo}}/\tau_{{}^3\text{H}}}) \right) \times 0.5 \simeq 2.2 \cdot 10^{16} \text{ events}$$

**Exposure:**

3 years

**Efficiency:**

50%



**Source:**

$\rho = 0.2 \text{ mg/m}^2$   
(full loading)

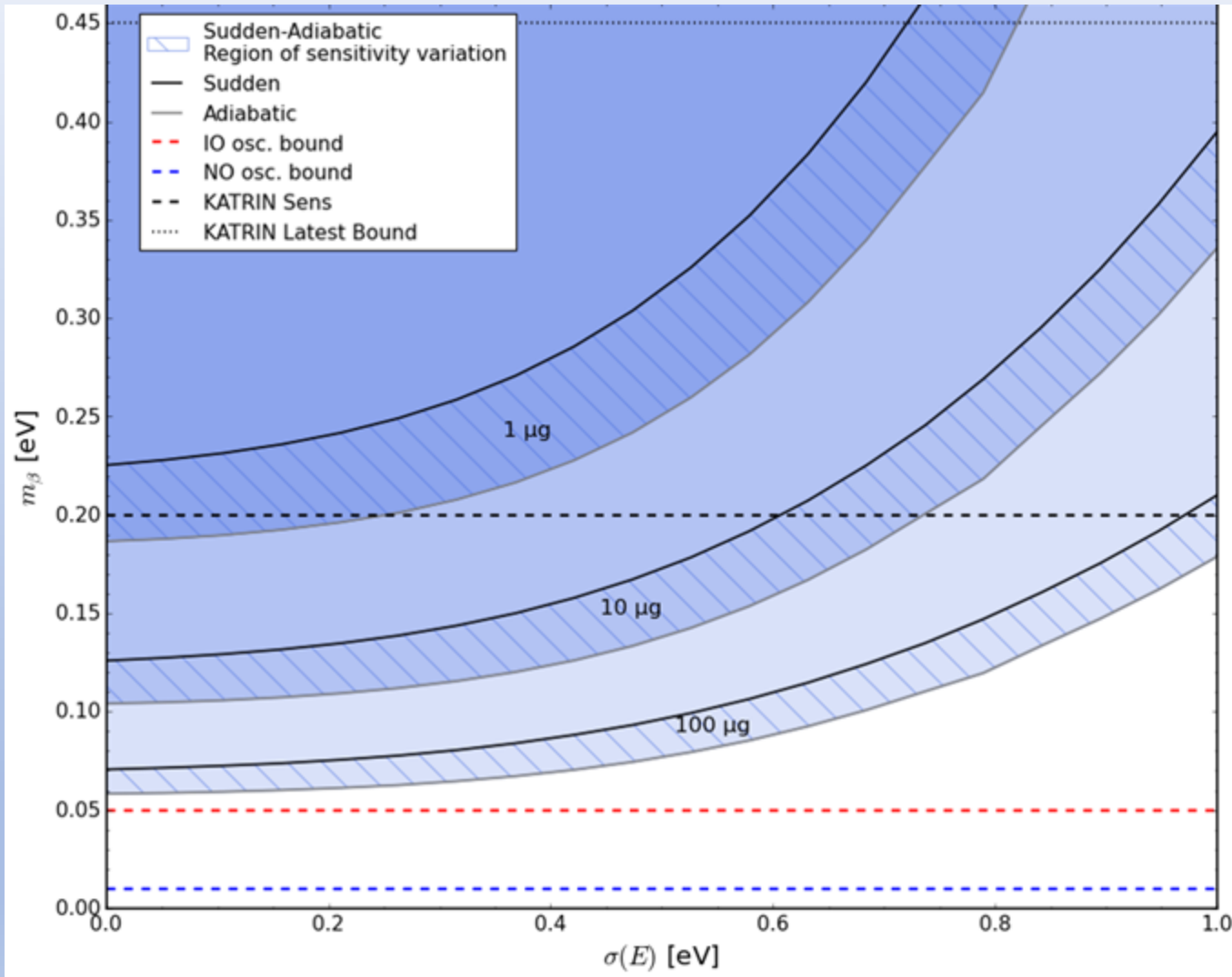
**1  $\mu\text{g}$**  (50 cm<sup>2</sup>)

716 MBq  
(19.3 mCi)



# Sensitivity for different tritium masses

50 % eff. and 3 years data-taking



## Remarks:

- sensitivity weakly dependent upon energy resolution (400 meV is already a good starting point)
- 1  $\mu\text{g}$  potentially comparable or even better than projection of best technology on the market
- 100  $\mu\text{g}$  (0.5 m<sup>2</sup>) close to probe the IO scenario

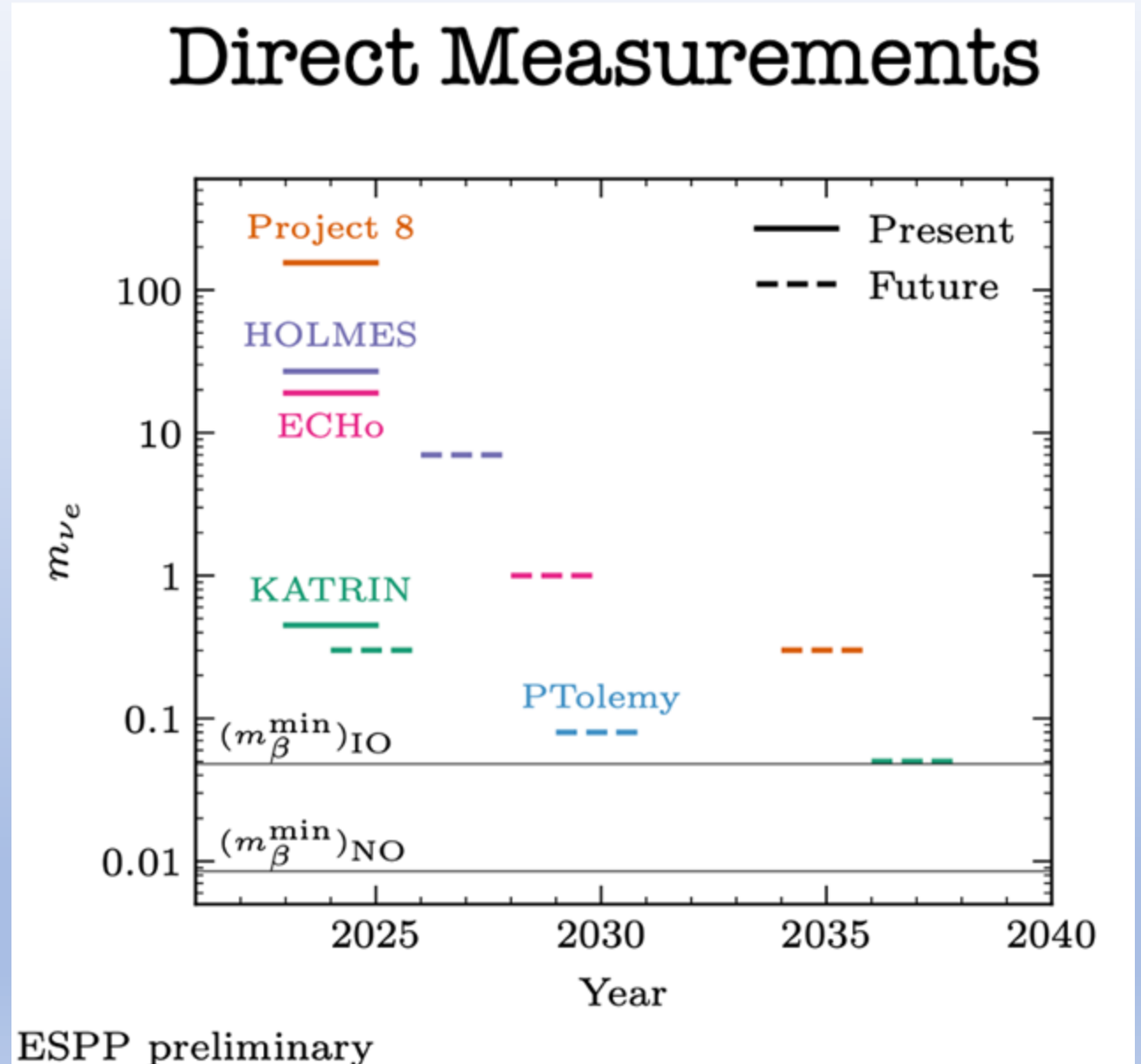


# Neutrino Mass Future

**PTOLEMY unique feature:**  
same detector  
but scalable target



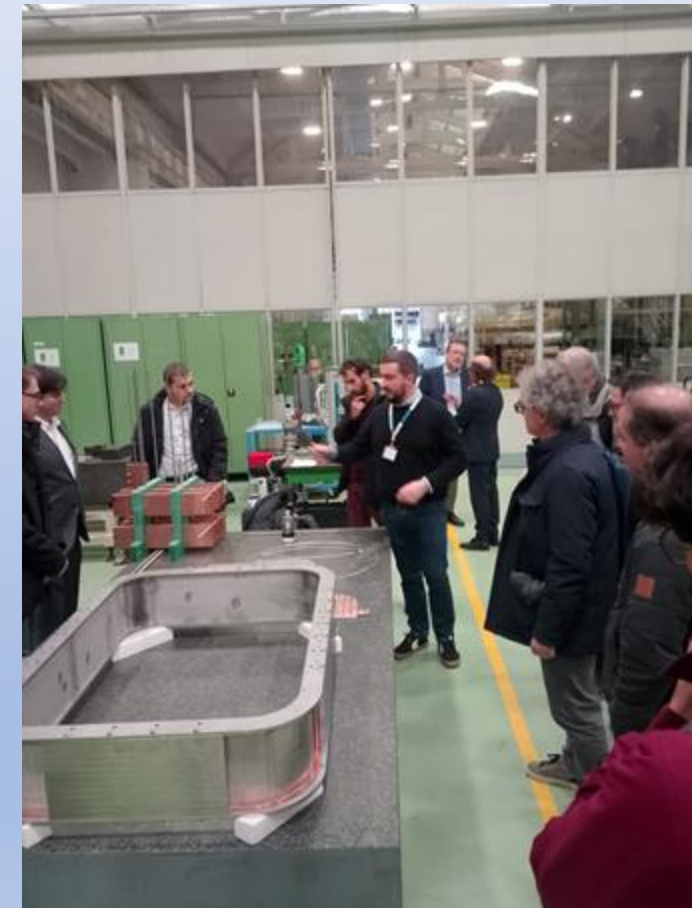
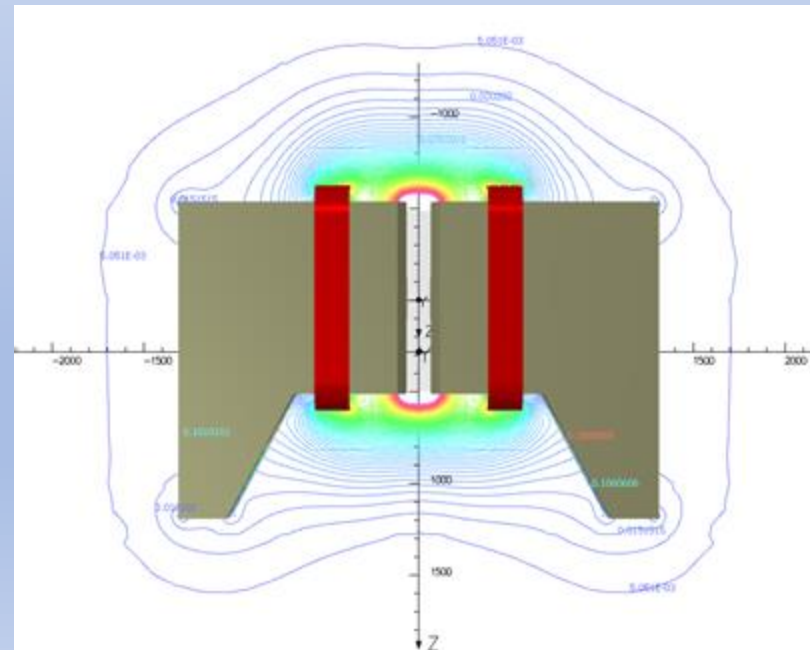
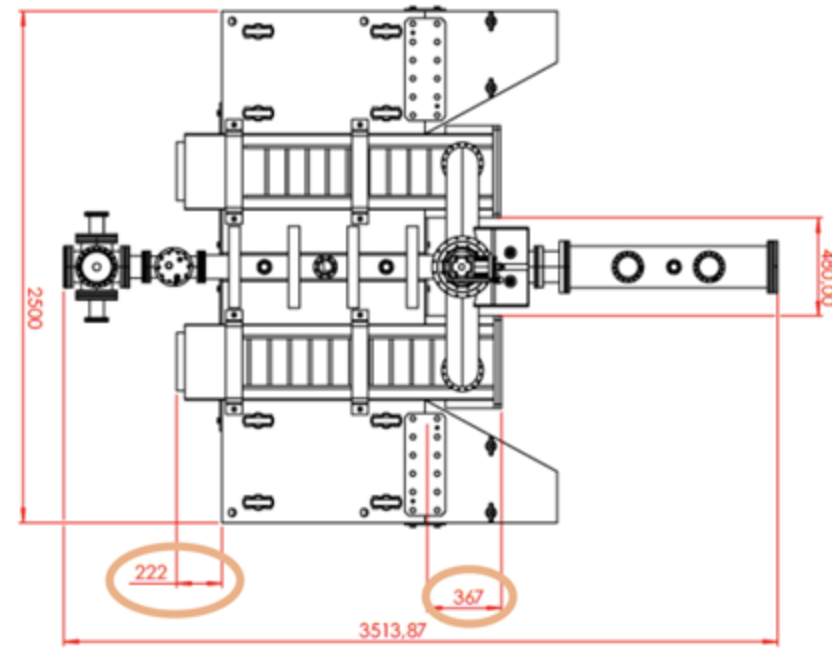
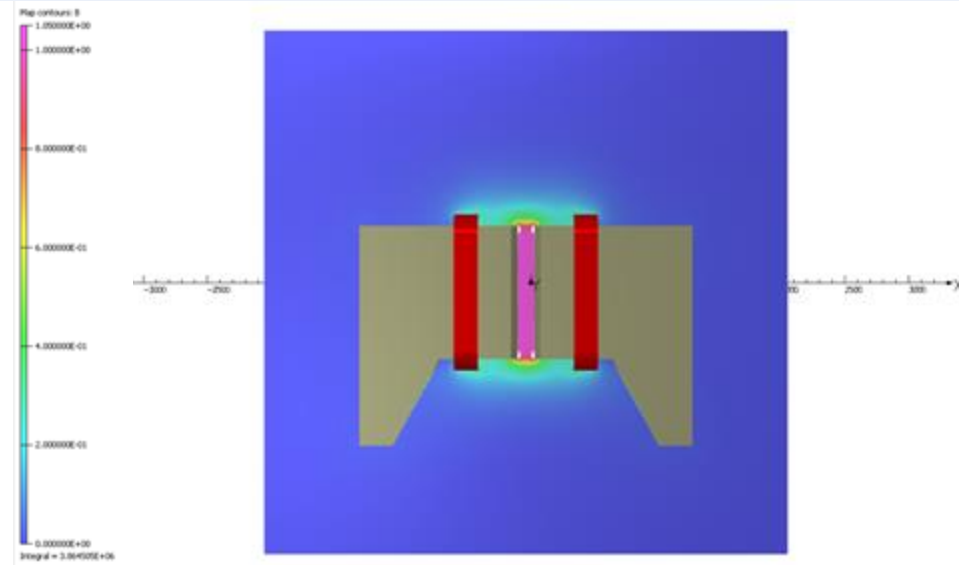
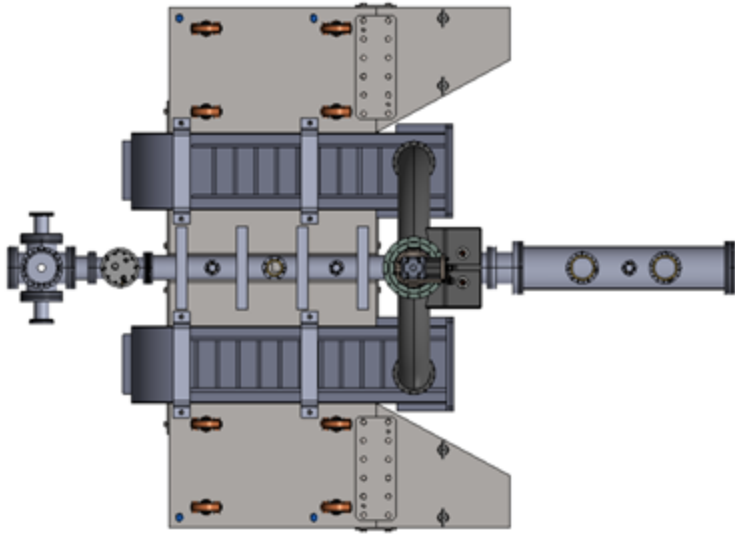
further limit lowering in future  
years



# Demonstrator parts under constructions: magnet, order issued

Magnet: Key subsystem

ASG (Genova)/Suprasys(Bilbao)





# Photo gallery



# Demonstrator parts under constructions: vacuum chamber

Cylindrical volume hosting  
future electron detector  
(not drawn)

HV filter electrodes feedthrough

Vacuum gauge

ALCA technology (Vicenza)  
**order issued**

Antenna signal FT  
top or bottom or both

Cold source  
For vacuum  
environment

Volume hosting G-T sample  
with standard CF flanges with  
possibility to make it longer and  
connected to vacuum suitcase  
for G-T transfer. Possible adapter  
to be added afterword

200 mm OD

Vacuum  
station

120 cm OD

Vacuum  
gauge  
(not visible)



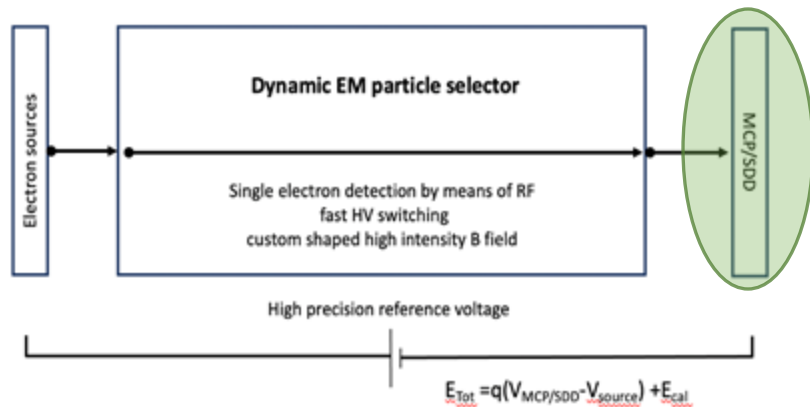
# Novelty elements of pahse1

# Calorimetry at INRiM



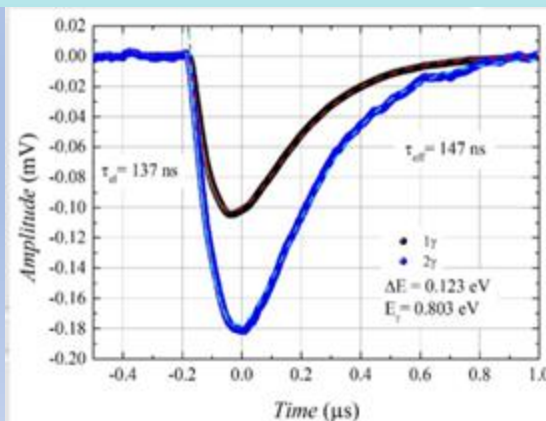
Based on the expertise of the INRiM  
electron detection with TES!

Key elements: high quality TES and new e-source  
based on nanostructures



On  $\gamma$ : 0.11 eV @ 0.8 eV and 106 mK and  $10 \times 10 \mu\text{m}^2$   
TiAuTi 90nm [ Ti(45nm) Au(45nm) ] ( $\tau \sim 137$  ns)

Design Goal (PTOLEMY):  $\Delta E_{\text{FWHM}} = 0.05$  eV @ 10 eV  
translates to  $\Delta E \propto E^\alpha$  ( $\alpha \leq 1/3$ )  
 $\Delta E_{\text{FWHM}} = 0.022$  eV @ 0.8 eV

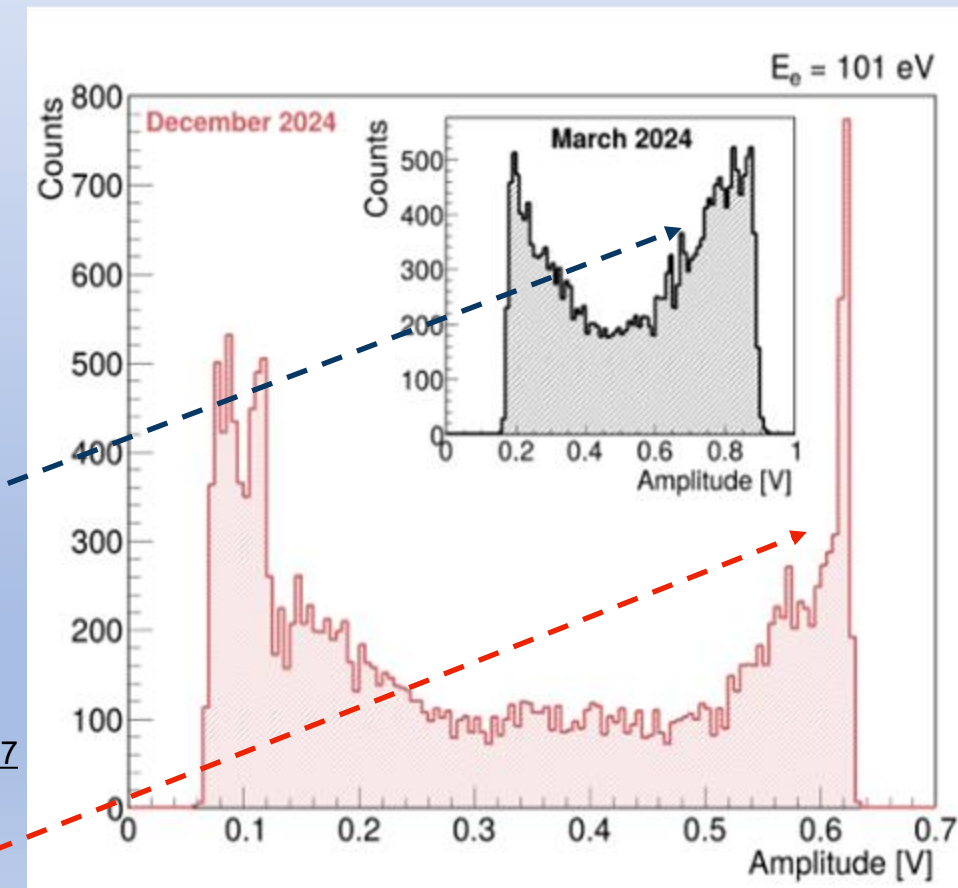


to  $e^-$   
detection

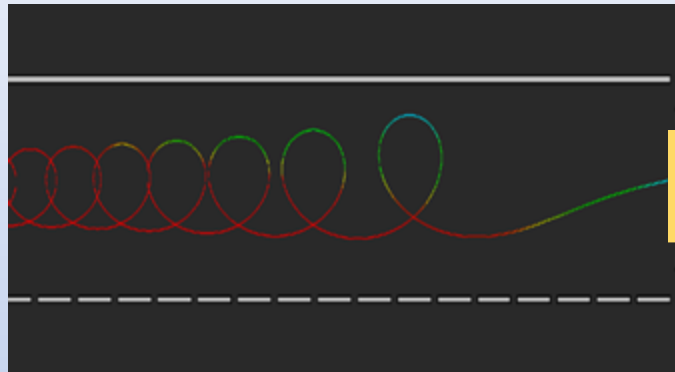
First measurement  
of electrons  
@ 100 eV with  
resolution  
~1-1.5 eV

*Phys.Rev.Applied* 22 (2024) 4, L041007

FWHM reduced by factor  
30 in last results!

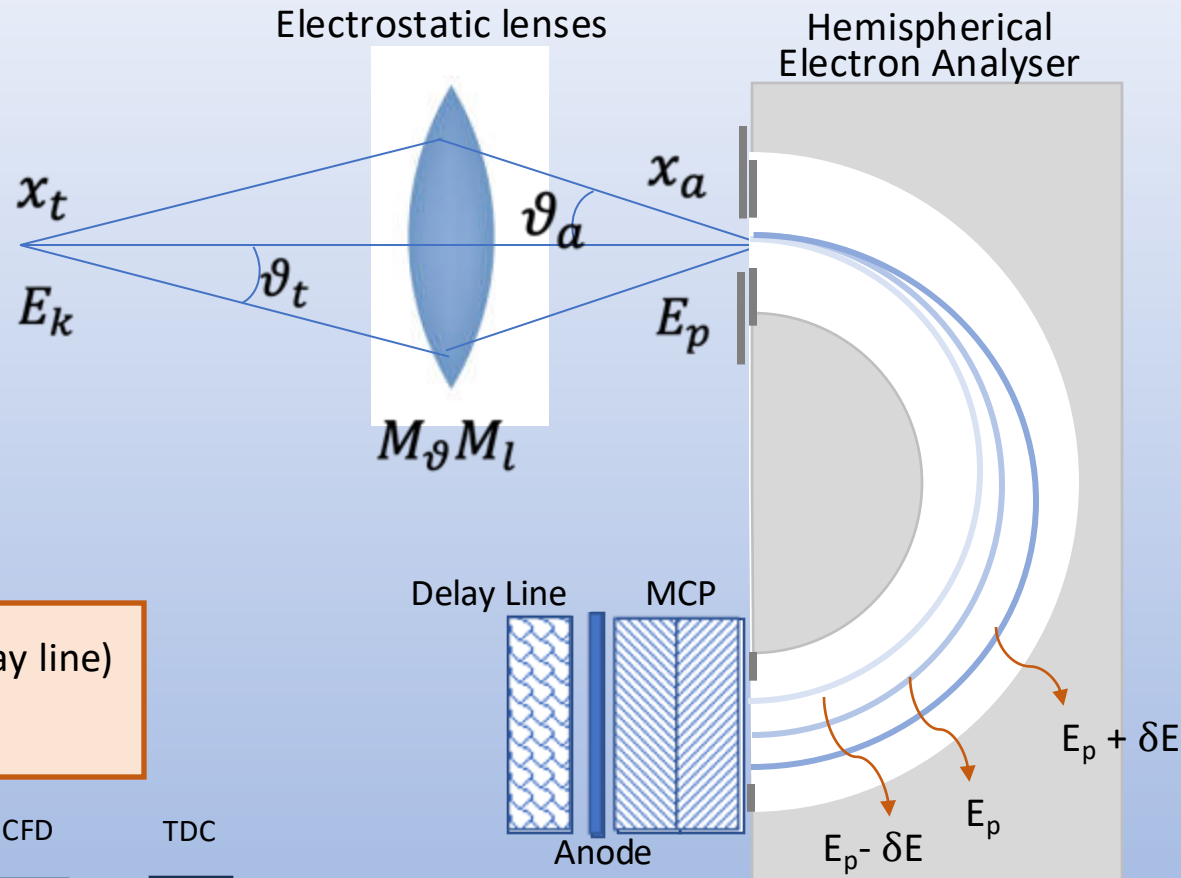


# Alternative energy measurement: Electrostatic analyser



PTOLEMY e.m. filter

Position sensitive detector (MCP+delay line)  
allows parallel acquisition



Electron optic basic equation

Helmholtz – Lagrange law

$$x_t \vartheta_t \sqrt{E_k} = x_a \vartheta_a \sqrt{E_p}$$

Two concentric hemispheres  
for the energy selection

Energy resolution  
(bandpass energy):

$$\frac{\Delta E}{E_p} = \frac{x}{2R_0} + \frac{\vartheta^2}{4}$$

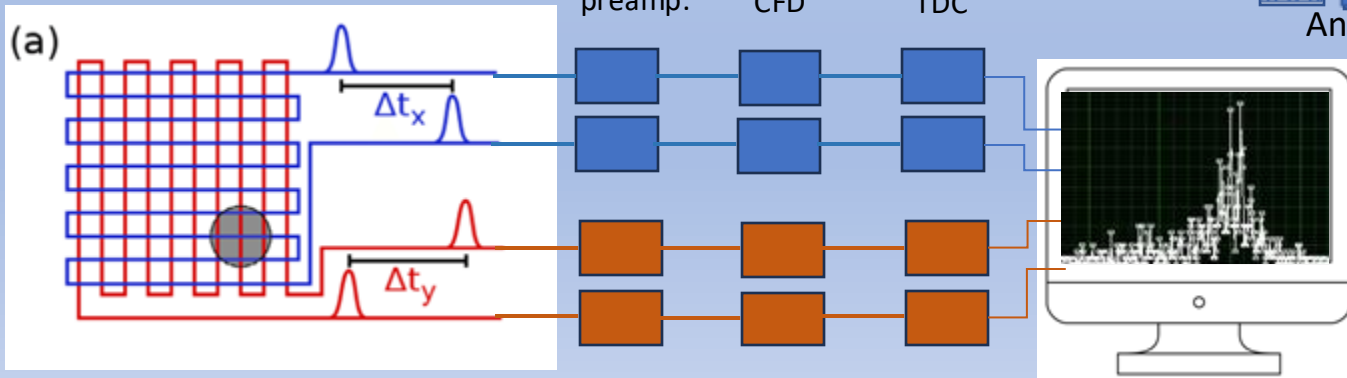
$\Delta E$ : energy resolution

$E_p$ : pass energy

$x$ : slit width

$R_0$ : mean radius

$\theta$ : accepted angle



# Commercial Electron Analyser

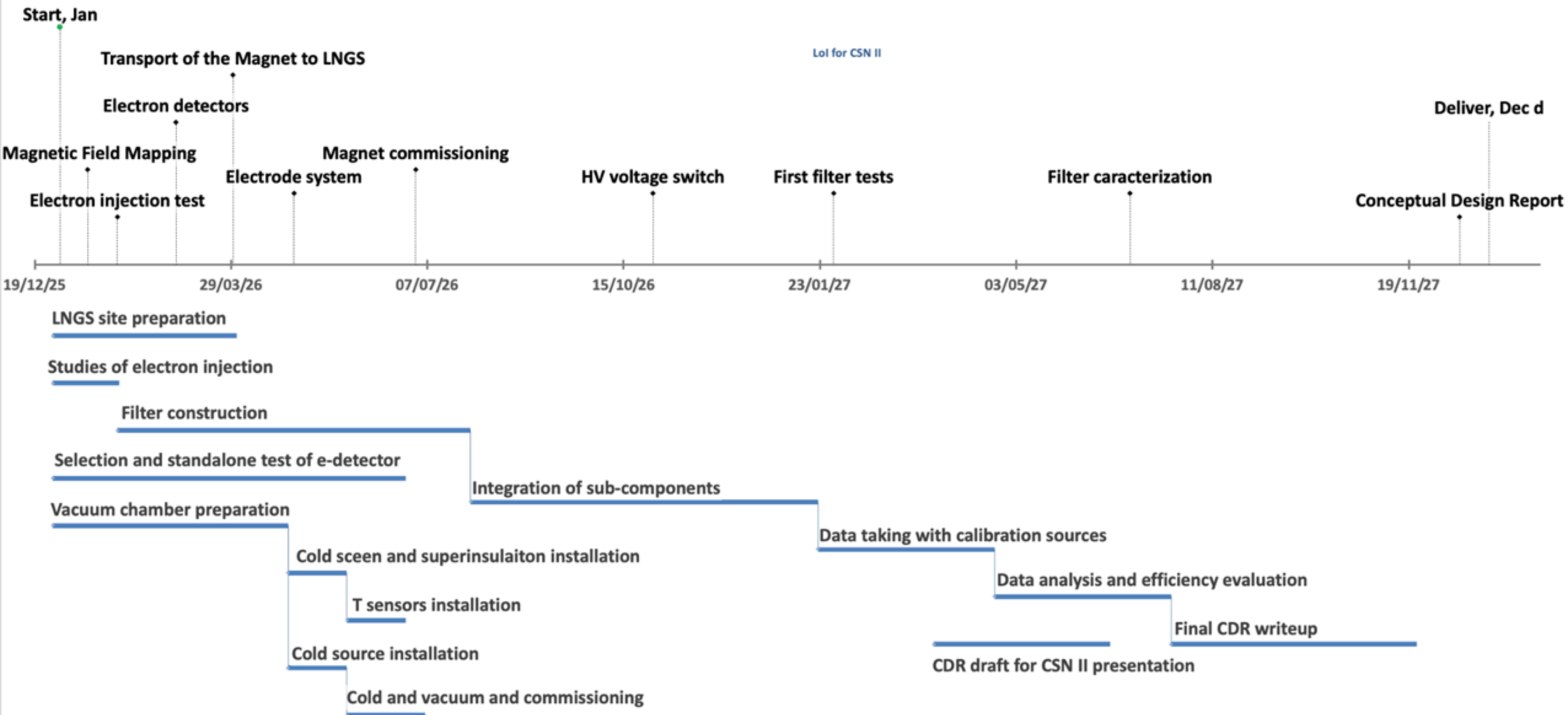
from Scientaomicron



<b>Detector</b>	Multi-channel (snapshot acquisition) Particle counting
<b>Energy resolution</b>	From <1.0 meV FWHM
<b>Transmission mode energy range</b>	0.5 eV – 1 500 eV
<b>Deflection &amp; Angular modes</b>	$\pm 3.5^\circ$ , $\pm 7^\circ$ , $\pm 15^\circ$
<b>Angular mode energy range</b>	From 0.5 eV – 1 500 eV
<b>Deflection mode energy range</b>	From 0.5 eV – 200 eV
<b>Deflection mode Spin scan</b>	Yes
<b>Operating pressure</b>	< $10^{-5}$ mbar
<b>Mounting flange</b>	DN200CF (10" O.D.)
<b>Electrostatic 3D Focus Adjustment</b>	Yes



# GANTT



# Detailed list of Researchers

## Non Italians

Name	FTE	Institute	Task / Responsibility
PI A.P. Colijn		University of Amsterdam and Nikhef	Analogical and Digital electronics
P. Bos			
J. Mead			
M. Naafs			
G. Visser			
PI Prof. C.G. Tully		Princeton University	Filter design/simulation and tests, Magnet, RF Antenna
A. Tan			
W. Chung			
M. Farino			
A. Langella			
PI N. de Groot		University of Radboud	Graphene studies and measurement
I. van Rens			
PI C. Perez de los Heros		Uppsala University	Editorial board

# Detailed list of Researchers Italians

PI M. Cadeddu	30%	INFN-CA	T handling, safety procedure and authorizations
L. Ferro (PhD)	30%		
R. Pavarani	30%		
A. Santucci	30%		
M. Sestu (PhD)	30%		
PI E. Celasco	20%	University of Genova and INFN-GE	Magnetic field measurement and cryogenics
S. Farinon	5%		
PI G.P. Mangano	20%	University of Napoli and INFN-NA	Theorist
O. Pisanti	20%		theorist
PI M. Messina	60%	INFN-LNGS	Infrastructure, Magnet, Calibration, RF
A.G. Cocco	60%		
N. D'Ambrosio	20%		
A. Ferella	30%		
N. Rossi	40%		
S. Nahorny	50%		
F.M. Pofi (PhD)	100 %		
F. Virzi (PhD)	100%		

# Detailed list of Researchers Italians II

## Some Numbers:

- 12 groups from 4 countries
- 8 INFN units participating
- 52 scientists in the Collaboration
  - Out of which 39 italians with tot ~**14. FTE**
- 77 scientists supporting the Collaboration

PI V. Tozzini	20%	INFN-PI and NANO-CNR Pisa	Theorist
L.E. Marcucci	20%		theorist
G. Menichetti	20%		theorist
M. Viviani	20%		theorist
PI G.L. Cavoto	30%	Sapienza University and INFN-Roma1	Tritium loading on graphene
M.G. Betti	20%		
B. Corcione (PhD)	100%		Electron detection
G. De Bellis	30%		HV developments
A. Esposito	20%		Theorist
L. Ficcadenti	30%		
G. Galbato Muscio (PhD)	30%		HV developments
L. Ginevra	20%		
C. Mariani	30%		Tritium loading on graphene
F. Pandolfi	40%		Electron detection
A.D. Polosa	10%		Theorist
I. Rago	30%		Graphene scale-up
PI G. Salina	30%	INFN-Roma2	DAQ and FPGA programming
Roberto Ammendola	10%		
Cosmin Marin(PhD)	50%		
V. Narcisi	30%		
PI A. Ruocco	40%	Roma Tre University and INFN-Roma3	Graphene target scale-up
A. Apponi	20%		CNS5 grant synergic (GREEAT)
O. Castellano (PhD)	40%		Nanostructures hydrogenation
S. Ritarossi	40%		



# Financial Request

## Costs 2026

INFN unit	FTE per INFN unit	Travel (kE)	hardware (kE)
CA	1.5	5	
GE	0.25	2.5	10 (SJ) magnet transport
NA	0.4	2	
LNGS	4.8	8.	
PI	0.8	2	
Roma1	3.9	10.5	6 e-detector
Roma2	1.2	3	
Roma Tre	1.5	5	5 graphene test production
Tot	14.35	37	21

Table 2: Residual costs for year 2026

## Costs 2027 (expected)

INFN unit	FTE per INFN unit	Travel	hardware
CA	1.5	5	
GE	0.25	2.5	
NA	0.4	2	
LNGS	4.8	8.	
PI	0.8	2	
Roma1	3.9	10.5	
Roma2	1.2	3	
Roma Tre	1.5	5	
Tot	14.35	38	

Table 3: Residual costs for year 2027 (expected)

# **To conclude: activities ongoing in 2025**

- 1) Graphene hydrogenation demonstrated, to be ported to tritium - no scientific show-stopper.  
Collaboration with UKAEA ongoing
- 2) Magnet at CERN by fall 2025 for field map measurement, electrodes design ongoing
- 3) RF radiation detection well advanced - to be integrated in the demonstrator
- 4) Several electron sources for calibration, HV stability close to requirement
- 5) First e-detection with standard detector.

# Outlooks

- First magnet operation in 2026
- CDR in preparation for July 2027 CSN II meeting
- Define a T0 for a possible physics data-taking of 3 years

# Papers production

Filter papers plus:

1. Low energy electron measurement
2. Stability of H/D on graphene (hints on T stability)
3. First RF measurement of single e- in a PTOLEMY setup (in preparation)
4. HV reference voltage (written)
5. Demonstrator setup (editing ongoing)
6. Theory paper on beta-decay endpoint
7. Sensitivity curves based on novel theory models on tritium-carbon bound state.
8. Lol where Phase0 and Phase1 contents are defined



Backup

# Systematics

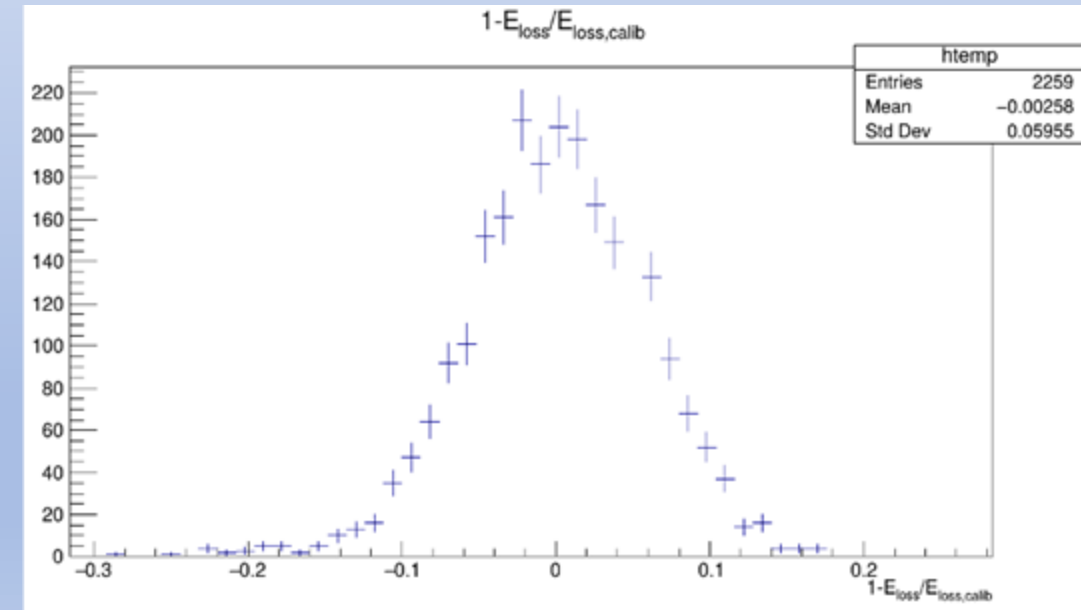
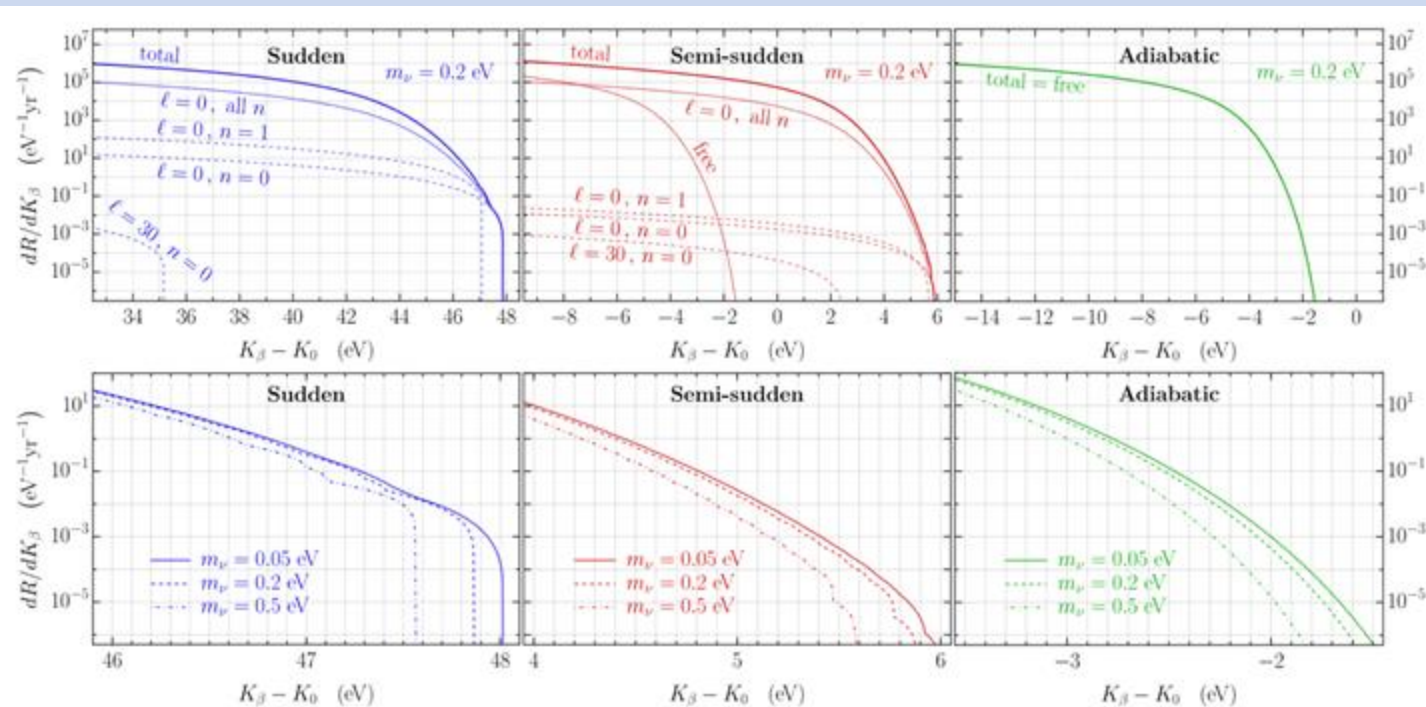
- 1) Theory aspects:  
further investigation of e ground level is needed

A. Casale, A. Esposito G. Menichetti, V. Tozzini

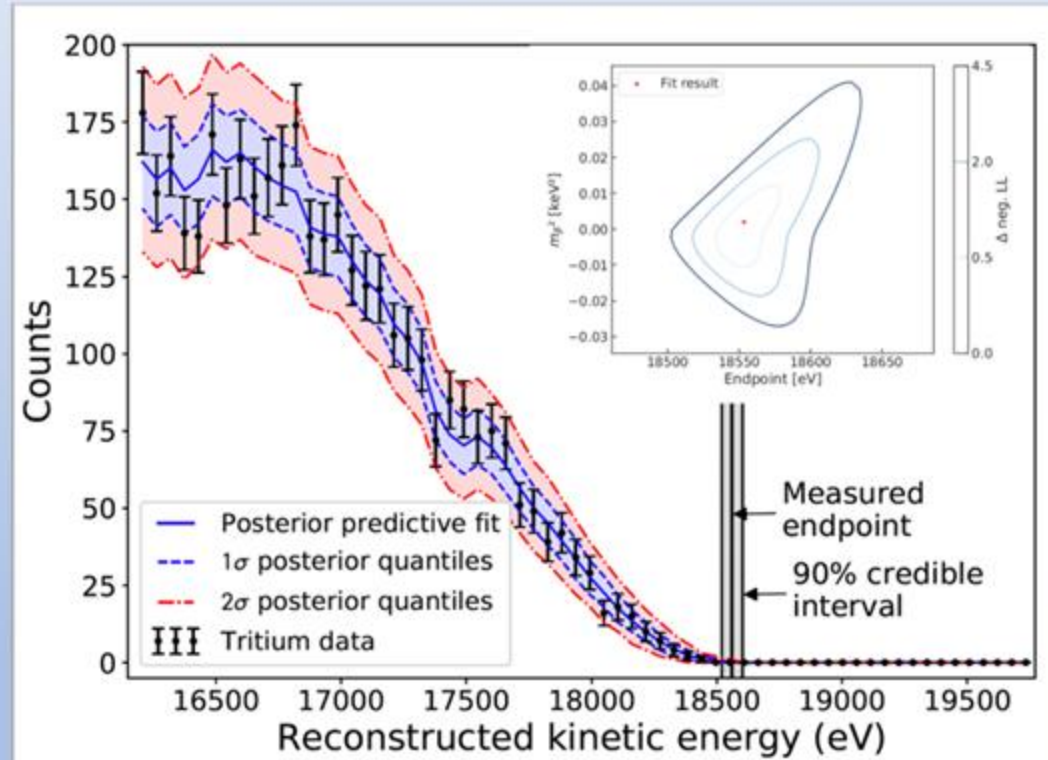
<https://arxiv.org/pdf/2504.13259>

- 2) Overall energy scale known at better than 5 meV

- 3) RF correction known at better than 10%



# Recent Project 8 Tritium RF Measurement



RF measurement  
background levels  
extremely low.

No events observed above  
endpoint,  
Setting upper limit on  
background rate

$< 3 \times 10^{-10}$  /eV/s (90% CL)

→ **Background Rate**  
 **$< 1$  event per eV**  
**in 100 years!**

<https://arxiv.org/abs/2203.07349>