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 [accelerating expansion of the universe]

Cosmology The "Nightmare Scenario"

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Outline of this talk

 High energy collider experiments can measure far more than we can compute

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High energy EW: showering

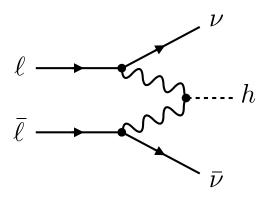
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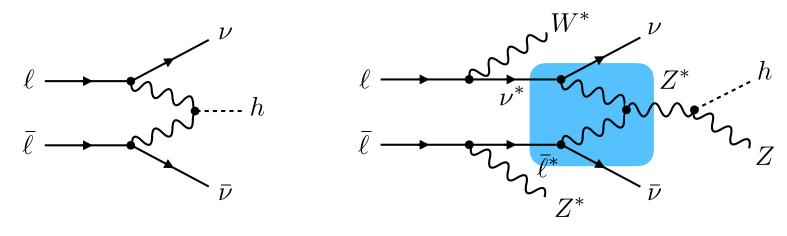
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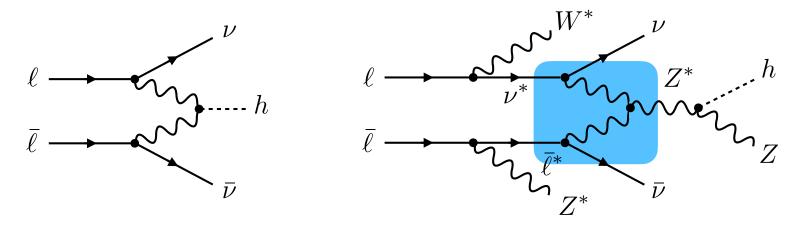
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At PeV collider, expect $\sim 10^4~W/Z/h$ in typical EW event

QCD: Factorization

Standard approach: focus on partially inclusive quantities

$$\begin{array}{l} \ell \bar{\ell} \to X \\ pp \to W + X \\ pp \to \ell \bar{\ell} + X \\ \vdots \end{array} \qquad X = \text{hadrons}$$

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Factorization theorems:

$$\sigma(p_1p_2 \to \ell\bar{\ell} + X) \sim \sum_{a,b} \int_0^1 dx_1 \int_0^1 dx_2 \underbrace{f_{1/a}(x_1) f_{2/b}(x_2)}_{\text{soft}} \times \underbrace{\sigma(ab \to \ell\bar{\ell} + X)}_{\text{hard}}$$

Energy correlators

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Basham, Brown, Ellis, Love (1978) ...
Hoffman, Maldacena (2008) ...
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Intuitive

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Quantum simulation

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Factorization approach is important even if it is not the only game in town

- Intuitive
- Wilsonian (compute 'one scale at a time')
- # of possible observables scales with energy

$$\mu^+\mu^- \to W_1W_2\cdots W_n + X$$

EW Factorization?

Extending factorization to high energy EW processes is not straightforward

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 $SU(3)_C$ gauge invariant

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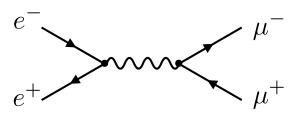
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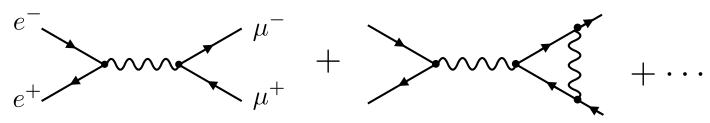
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- EW gauge invariance broken in IR
- Unbroken EW gauge governs underlying hard process

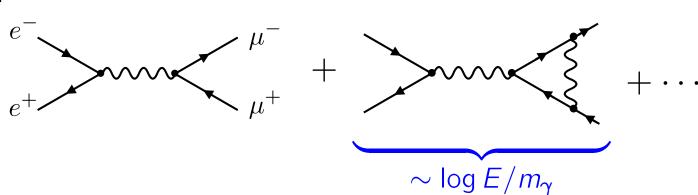
QED:



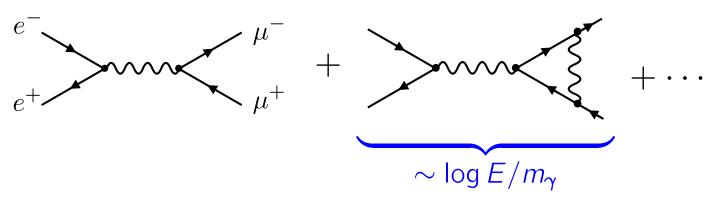
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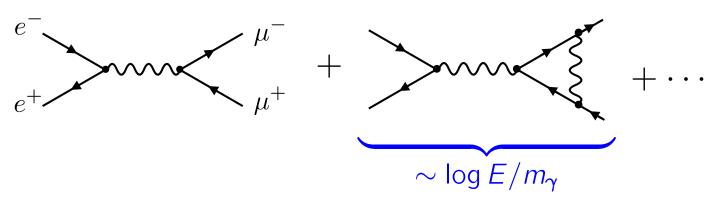
QED:



Bloch, Nordsiek (1937!): QED amplitudes are finite if we sum over soft photons in final state

$$\sum_{n=0}^{\infty} \int d\Phi(\gamma_1) \cdots d\Phi(\gamma_n) \Theta(E_{\gamma} < \delta) \times \sigma(e^+e^- \to \mu^+\mu^- + \gamma_1 \cdots \gamma_n) < \infty$$

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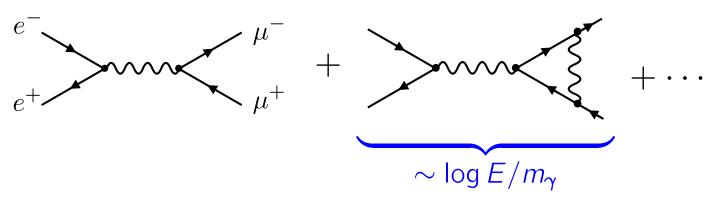
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But what about the initial state?

QFT amplitudes are finite if we sum over initial <u>or</u> final states within an energy range

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Can we obtain a pricipled understanding of the relation between 'parton' states and initial/final states in experiments?



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Dark matter, matter-antimatter asymmetry, inflation...

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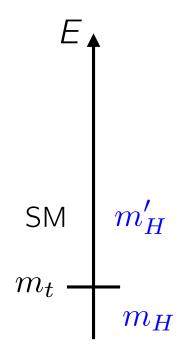
Naturalness Calculability problem

^{*}Except for the strong CP problem

• Given SM parameters, we can make predictions

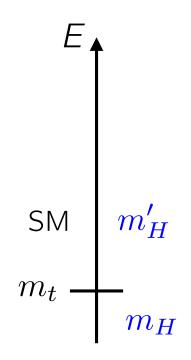
• Given SM parameters, we can make predictions *E.g.* integrate out top quark

$$m_H^2(\mu = m_t) = m_H'^2(\mu = m_t) + \frac{N_c y_t^2}{16\pi^2} m_t^2 + \cdots$$



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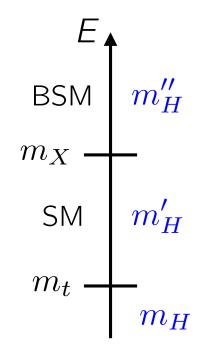


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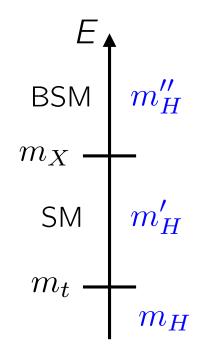


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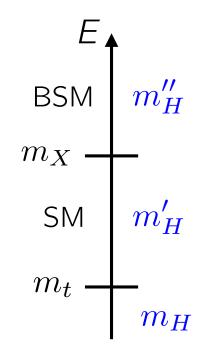
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E.g. heavy
$$\nu_R$$
: $\Delta \mathcal{L}_{\mathrm{BSM}} = y_{\nu} \bar{L}_L H \nu_R + m_{\nu_R} \nu_R \nu_R$ $m_{\nu_L} \sim 0.1 \ \mathrm{eV} \ y_{\nu_R}^2$ for $m_{\nu_R} \sim 10^{14} \ \mathrm{GeV}$



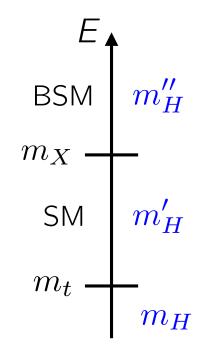
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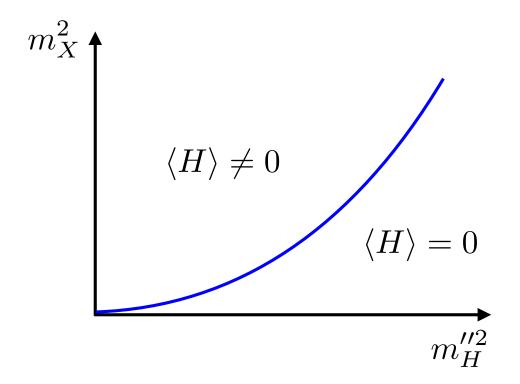
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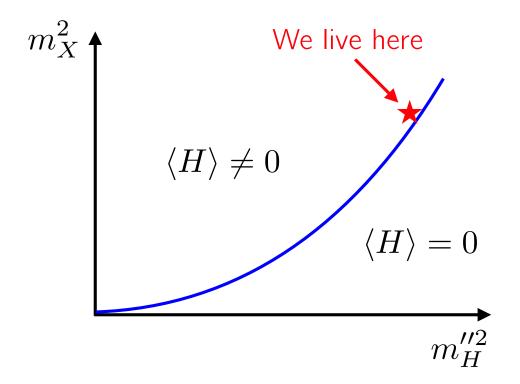
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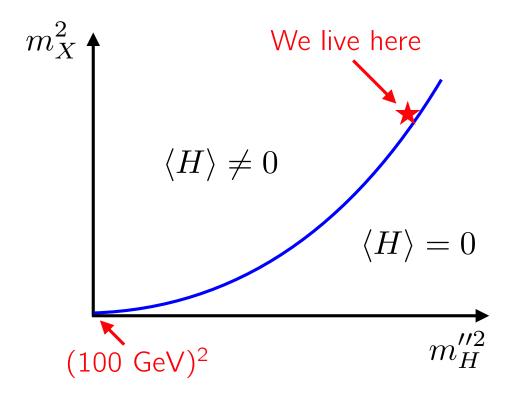
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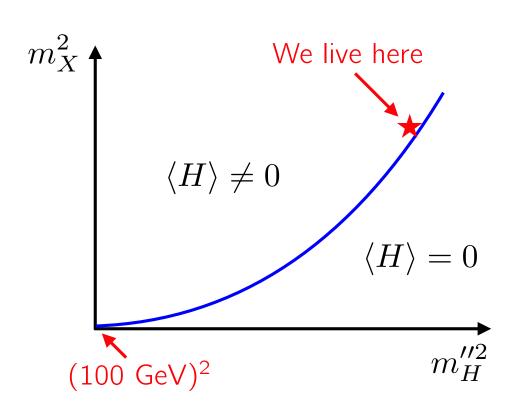
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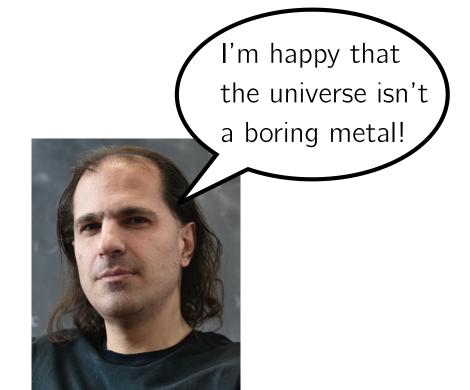


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No Conclusions

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I hope we keep exploring...