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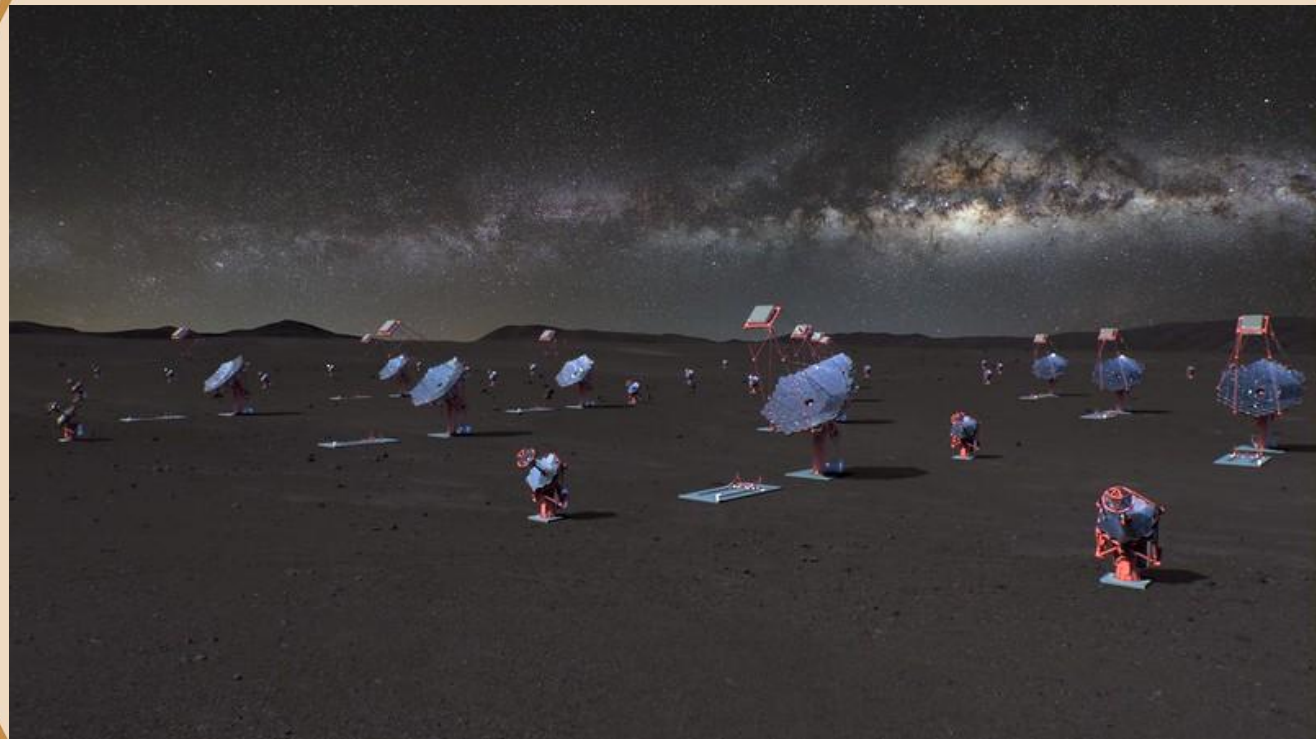
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The CTAO: status and perspectives

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Advances in Modeling High-Energy
Astrophysical Sources, Sexten, Italy
30 June - 4 July 2025





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Outline

- ❖ **CTAO : an astronomical observatory**
- ❖ **IACT: the present and the future**
- ❖ **Early science with CTAO**
- ❖ **Conclusions**



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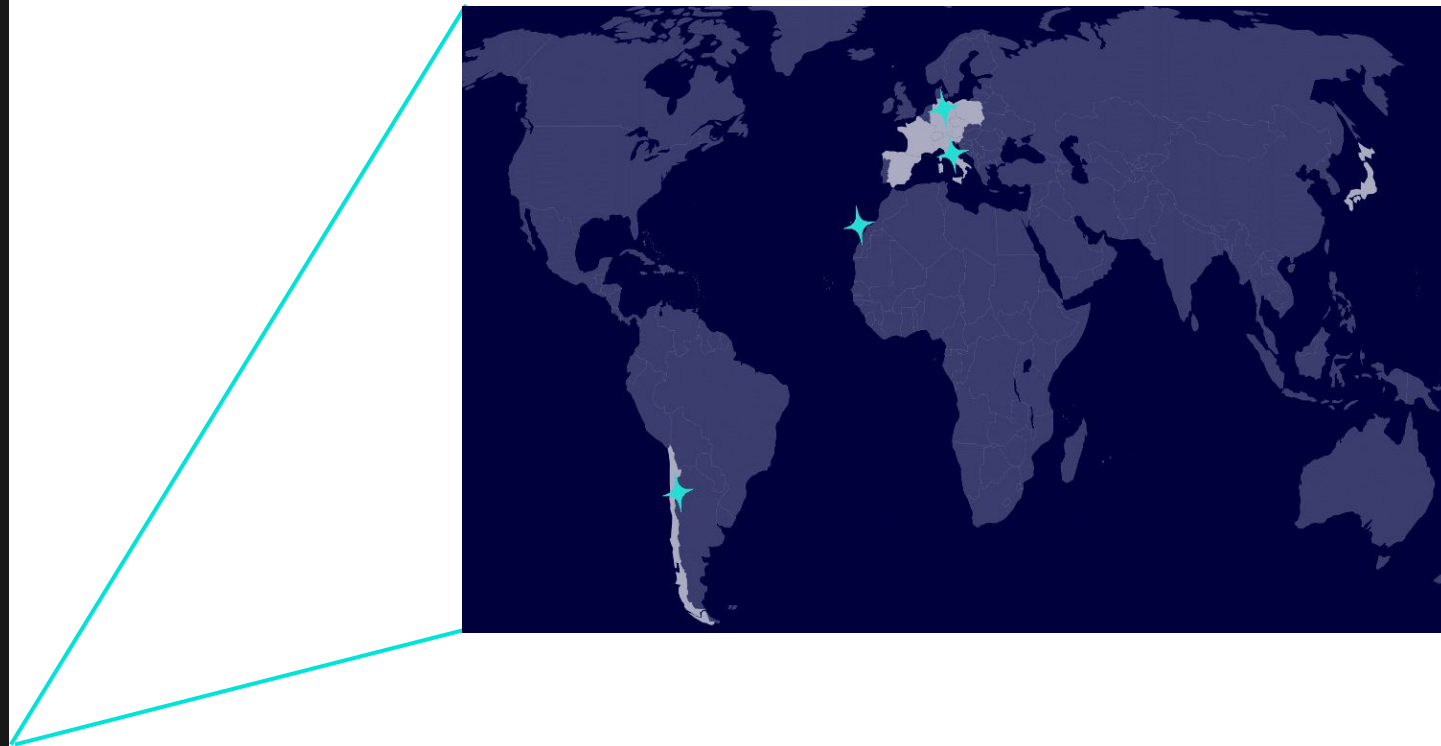
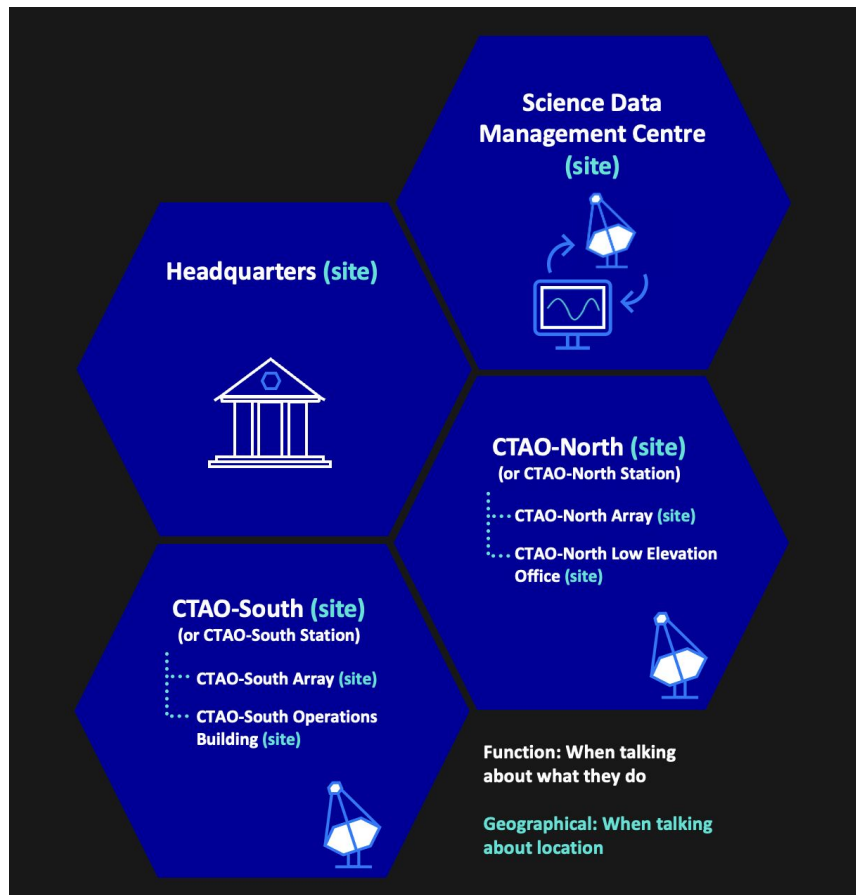


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CTAO : a distributed facility

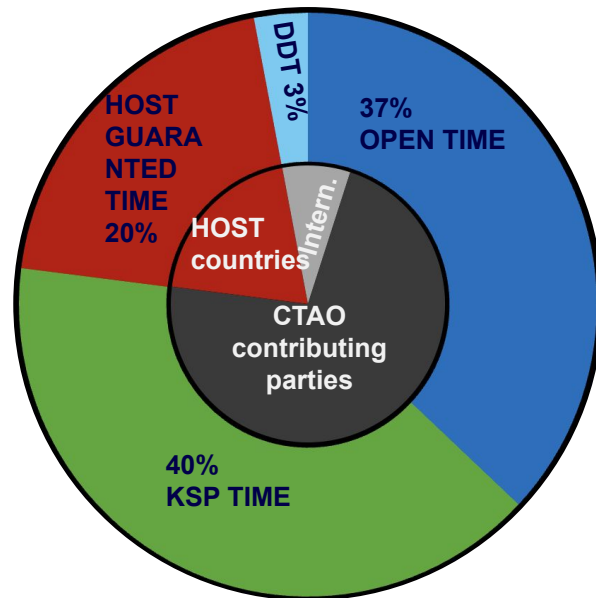




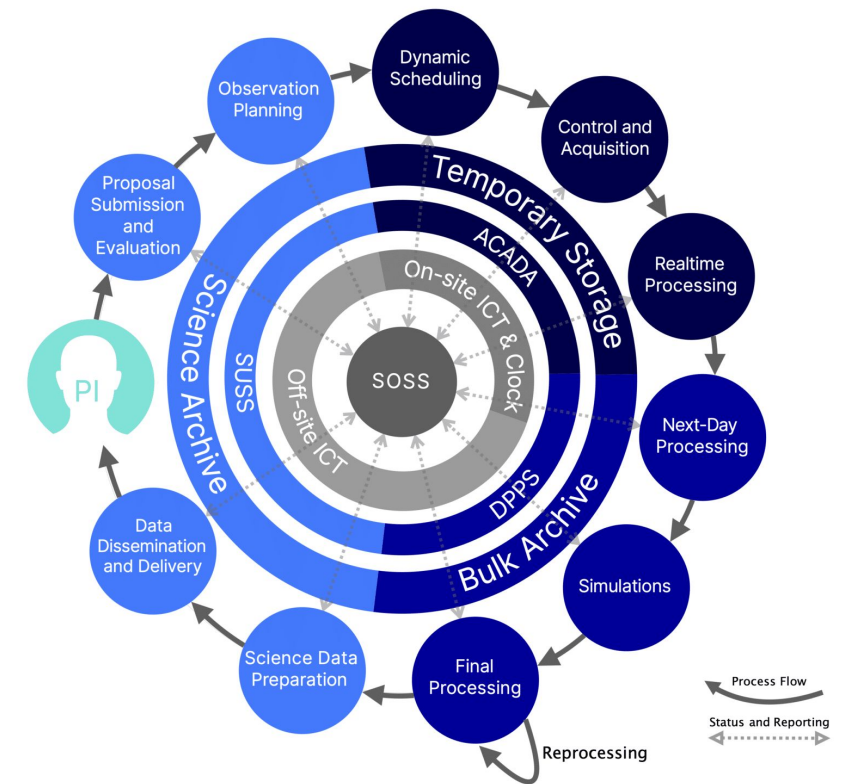
An Astronomical observatory

An open proposal-driven observatory

- ❖ Proposals will be evaluated only on their scientific merit
- ❖ Data with a proprietary period of 1 yr after that fully open



integrated over 10 yr



Today

The second generation of IACTs (since 2004)



Today

The second generation of IACTs & of particle array detectors (since 2012)





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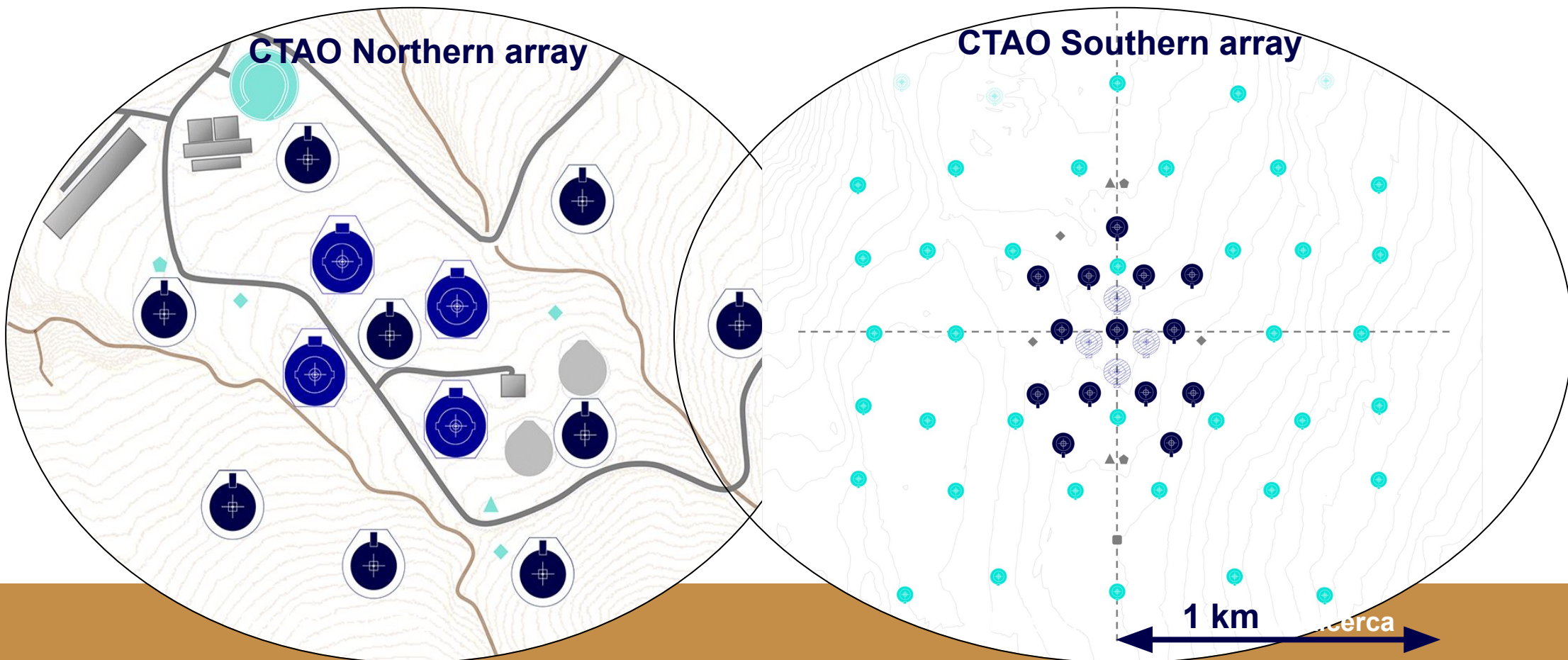
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Two observation stations, one unique observatory

Alpha configuration





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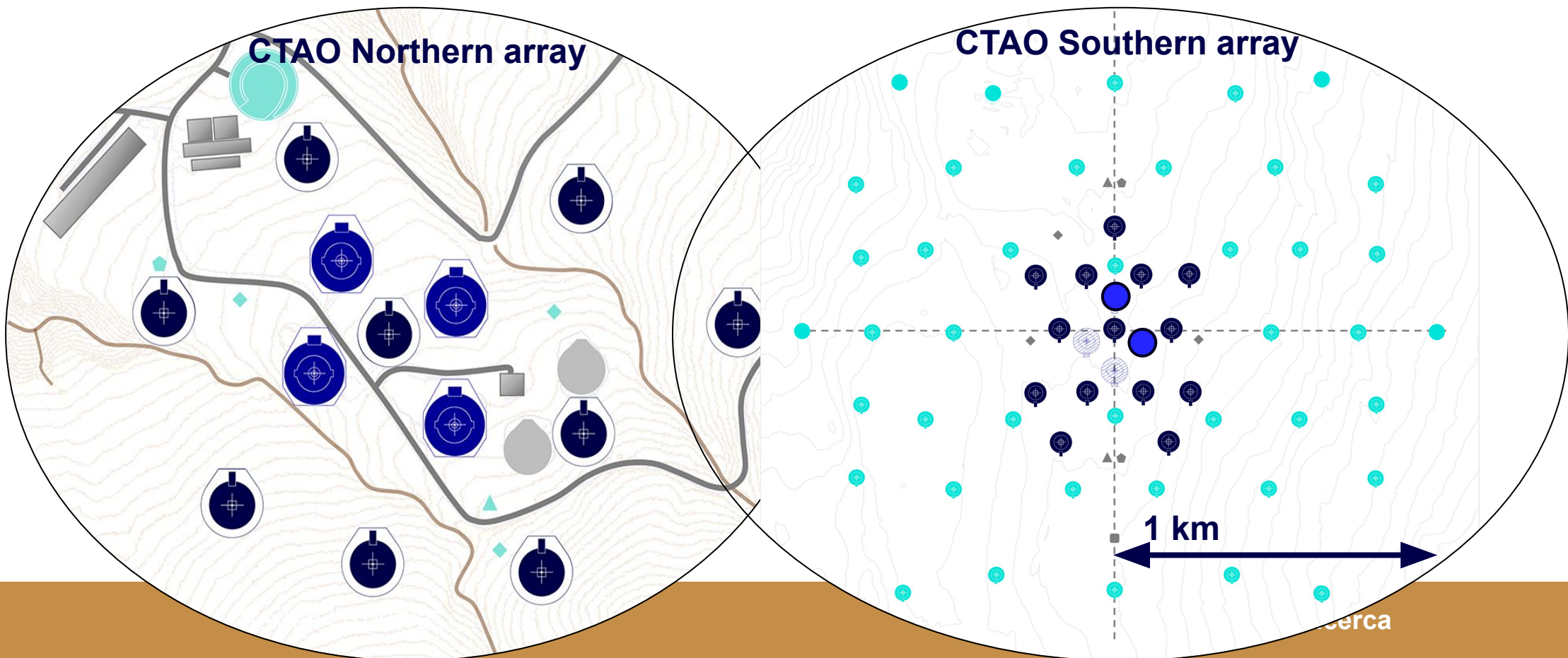
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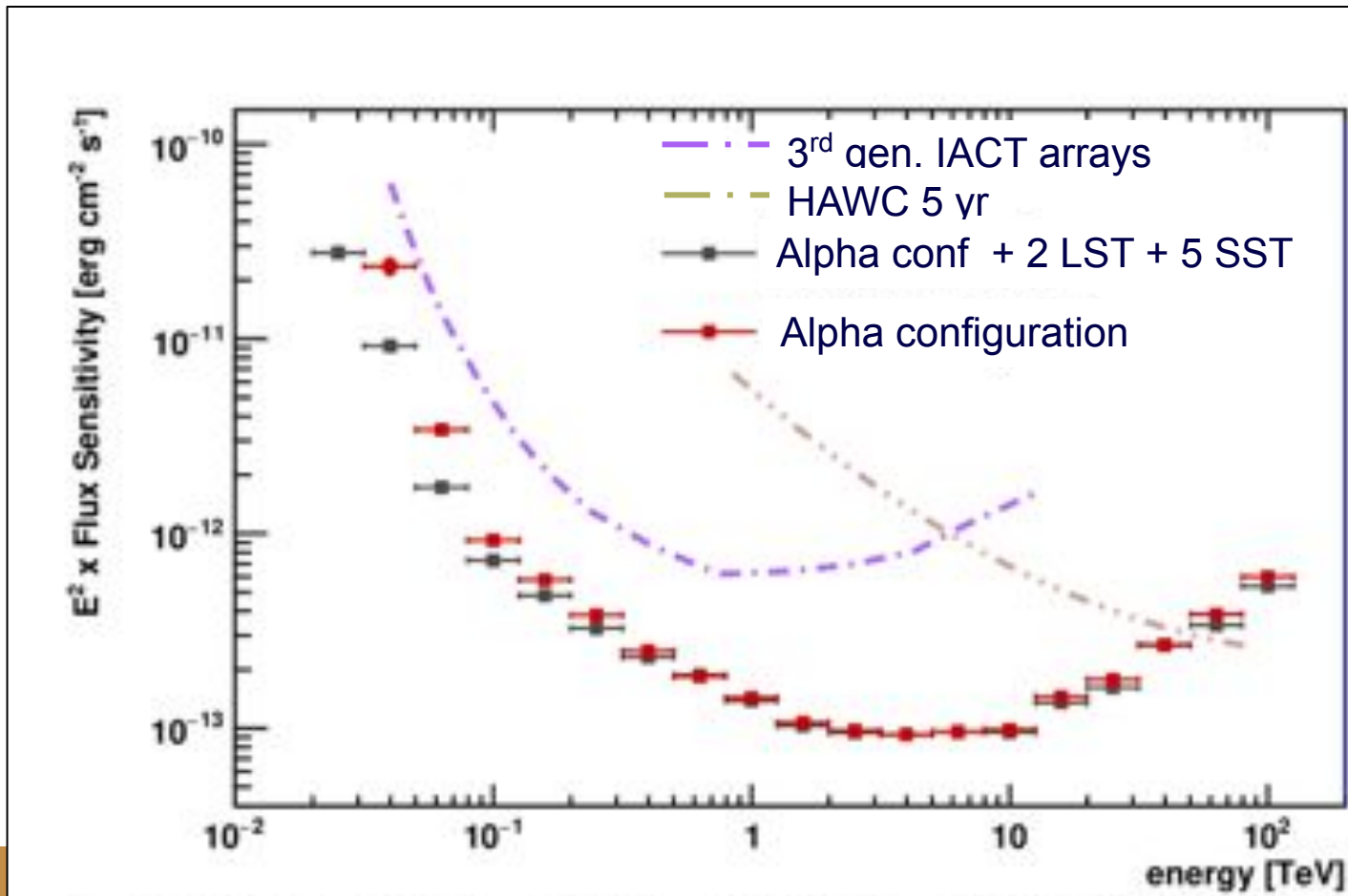
Two observation stations, one unique observatory

Alpha configuration



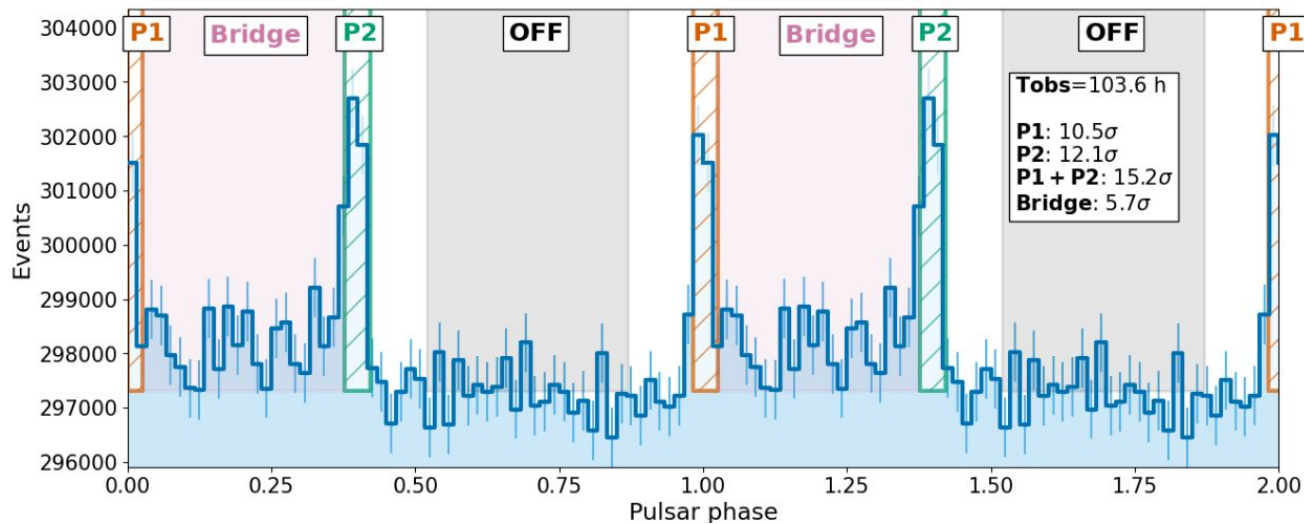


Improved Alpha configuration

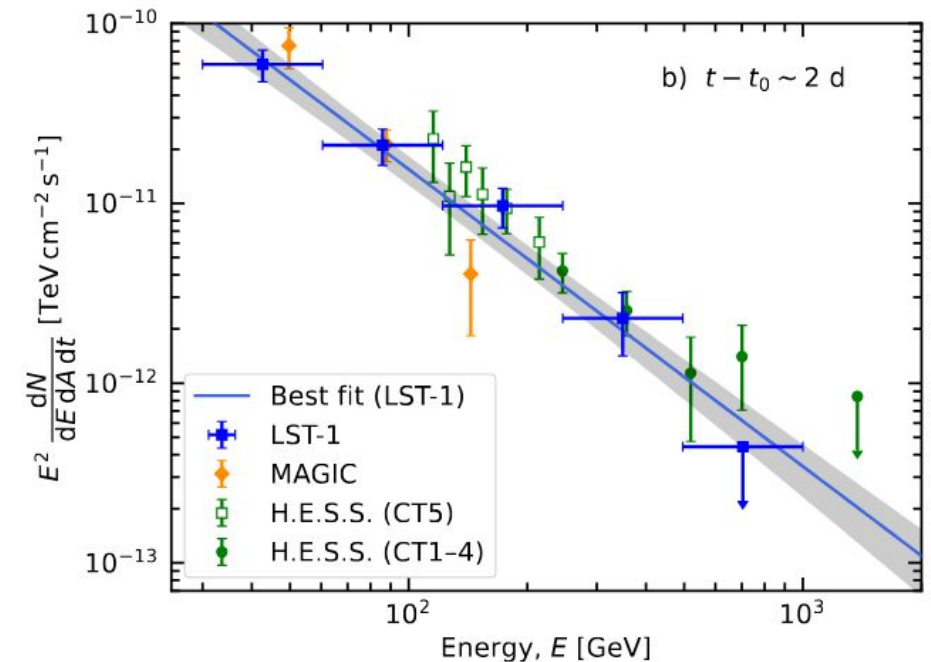




Improved Alpha configuration



LST coll., 2024: Phaseogram of the Crab pulsar



LST coll., 2025: VHE detection of the nova RS Oph with LST1



Towards the first CTAO data

Intermediate array configurations

CONCEPT

Intermediate array configurations: incremental array configurations that become progressively operative

- array elements fully integrated with the intermediate releases of the software packages
- array elements include telescopes but also calibration devices and atmospheric characterization instruments

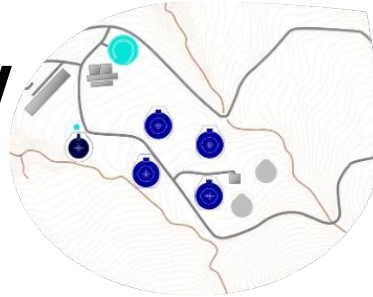
MID-PERIOD PLAN (3 yr long) BASED ON THE CONSTRUCTION SCHEDULE
built accounting for the inputs of the in-kind contribution teams



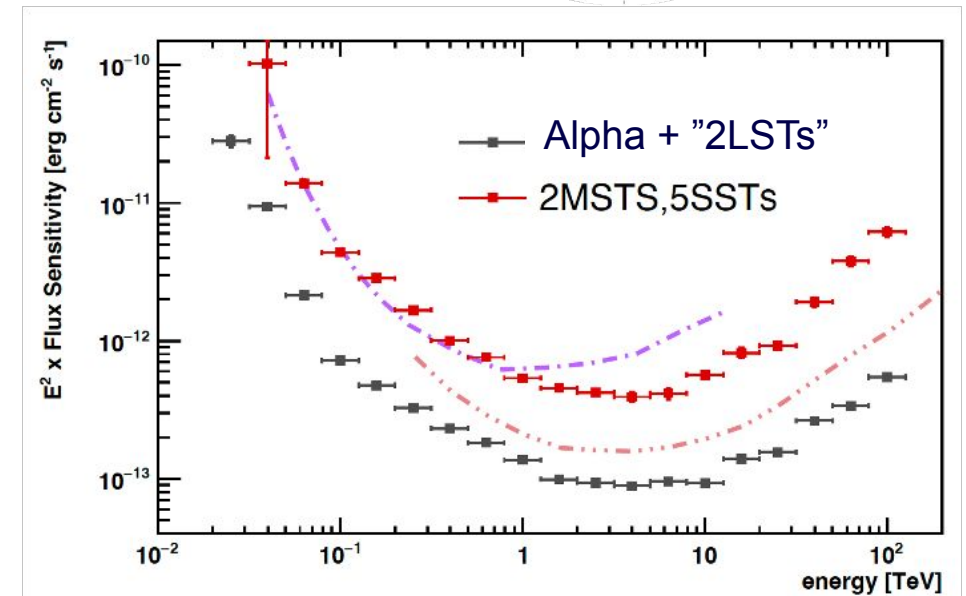
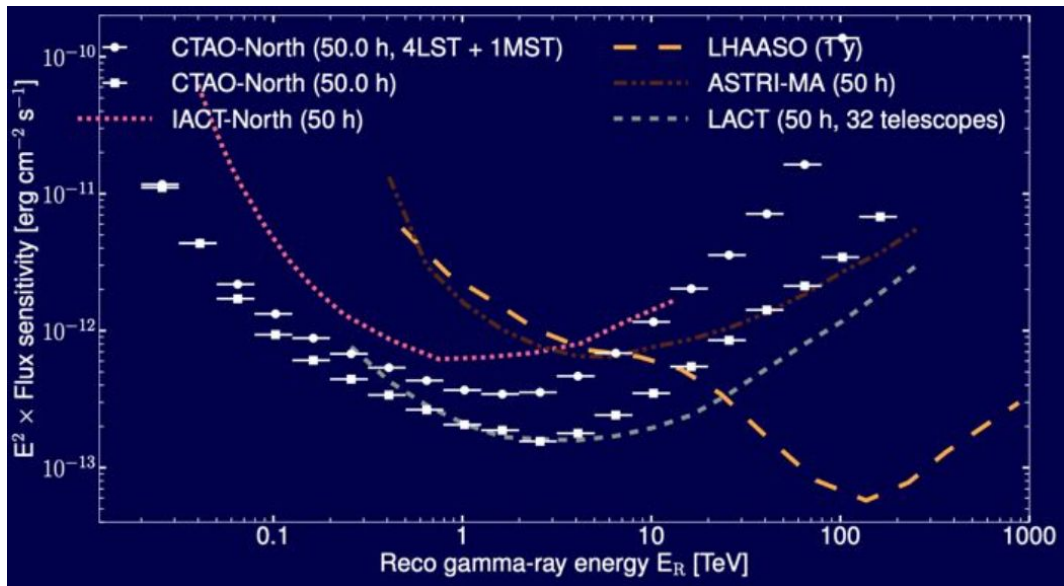
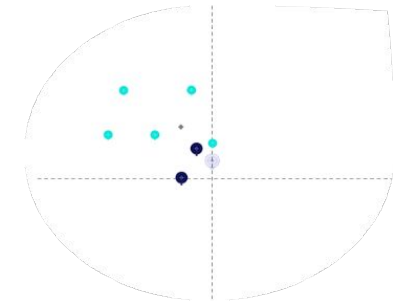
First intermediate configuration

in 2 - 2.5 yr from now

CTAO - North

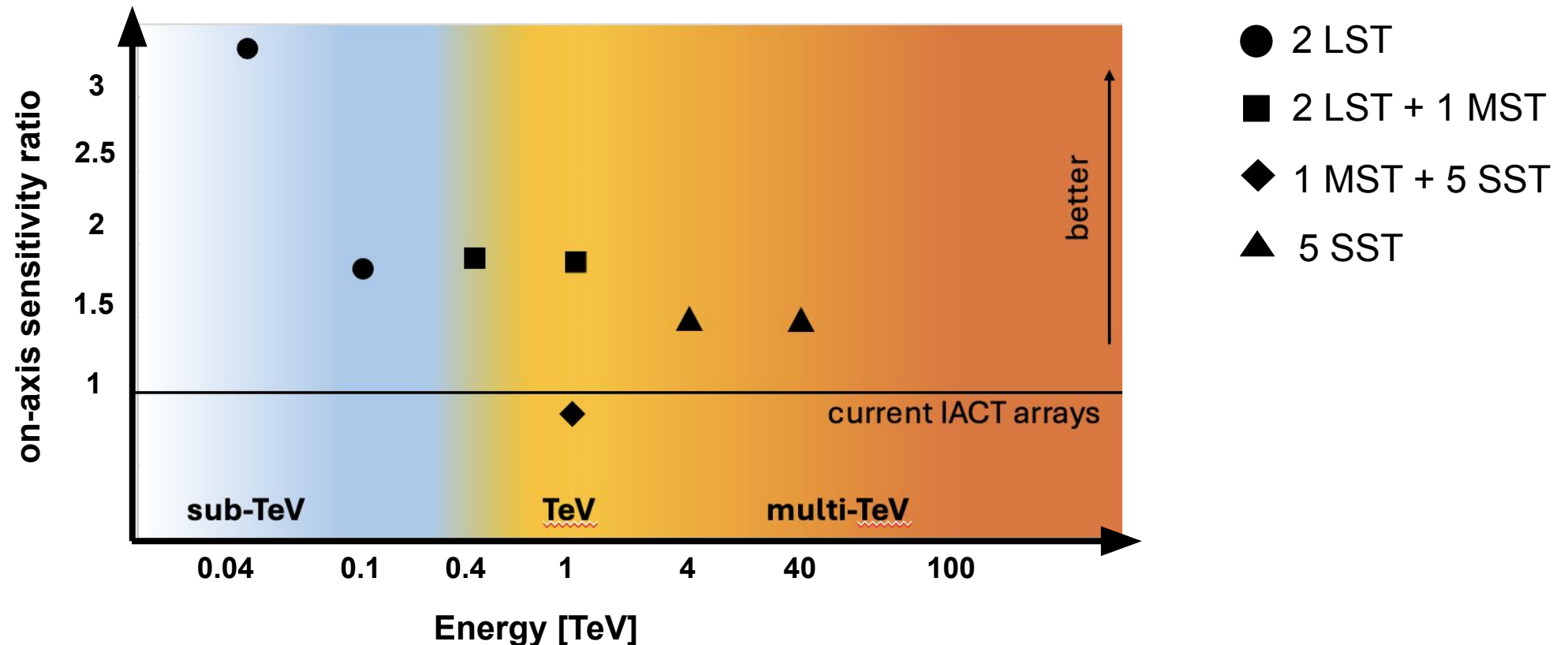


CTAO - South



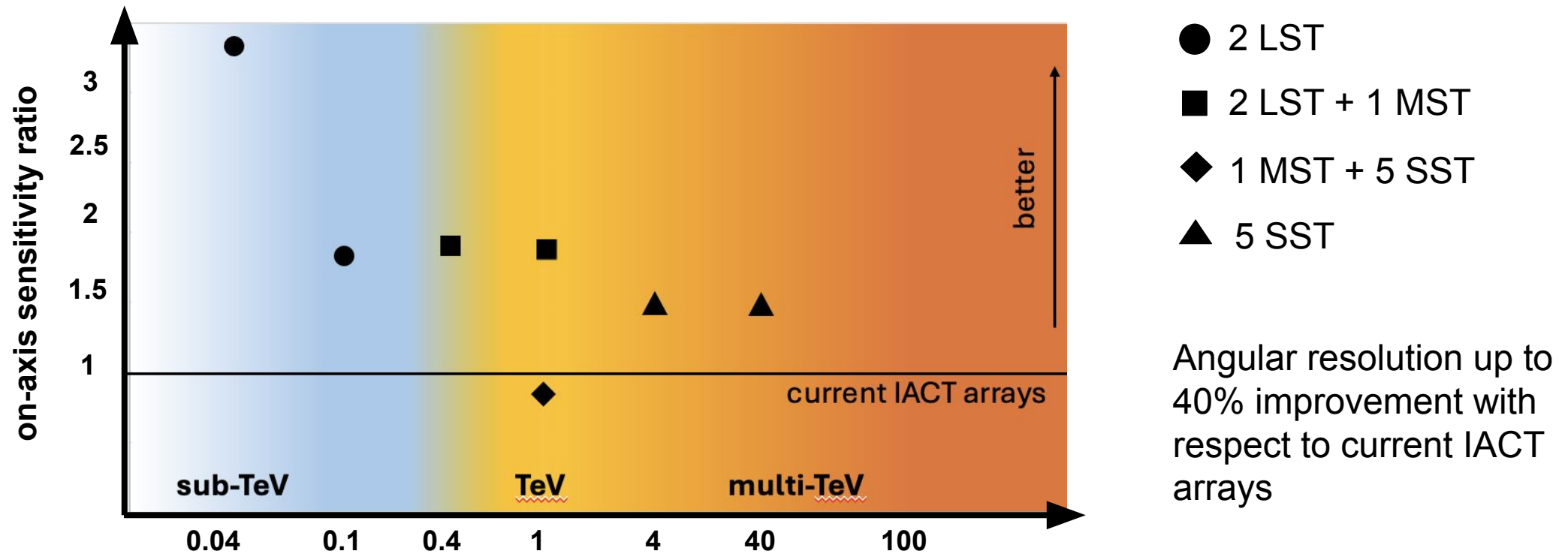


When will the scientific impact begin?





When will the scientific impact begin?



To get first high impact results we shall focus on science cases needing sensitivity more than angular resolution

A bright upcoming future

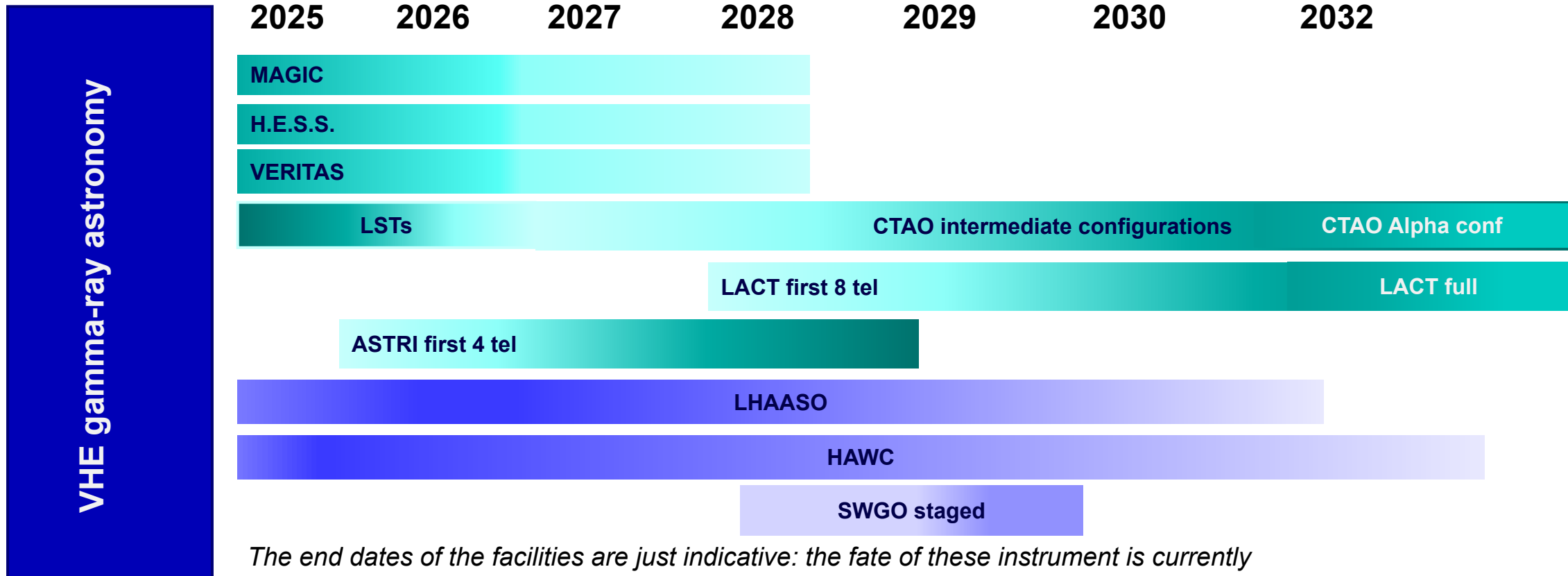
ASTRI - CTAO - SWGO - ALPACA



Galactic equator



Status of the VHE astronomy



The end dates of the facilities are just indicative: the fate of these instrument is currently under discussion



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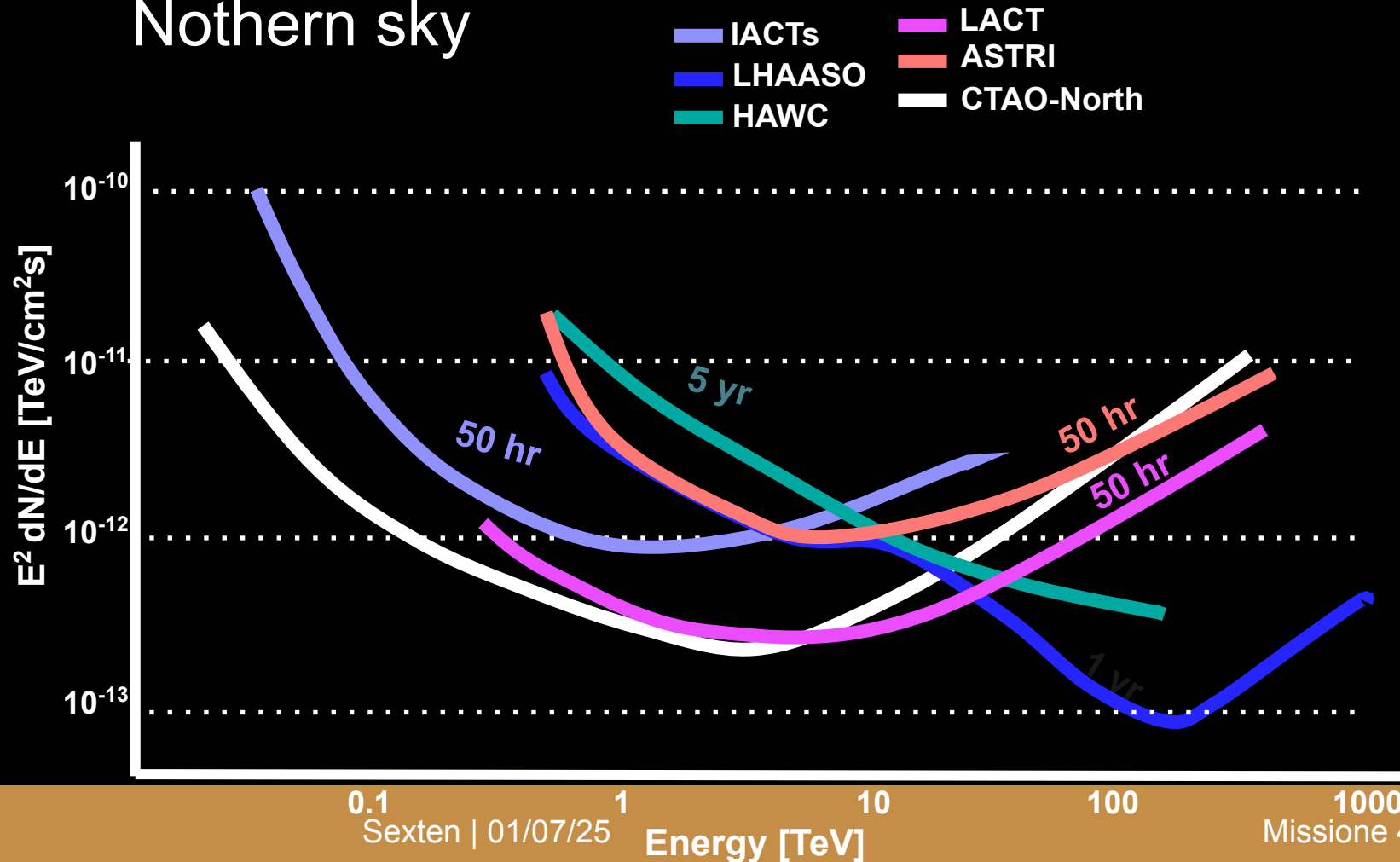
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Steady source performance

Northern sky





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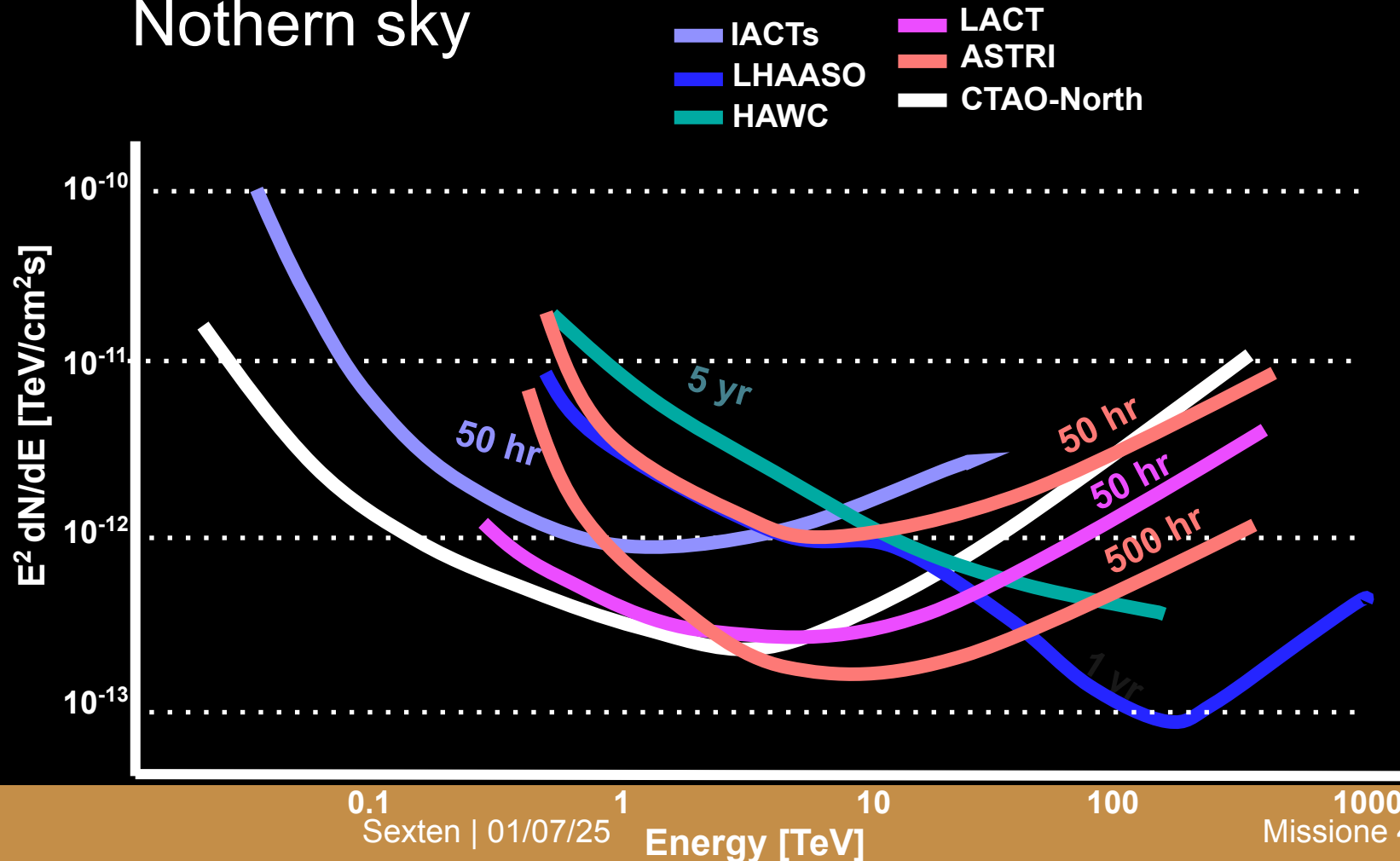
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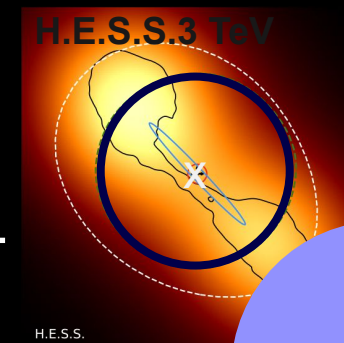
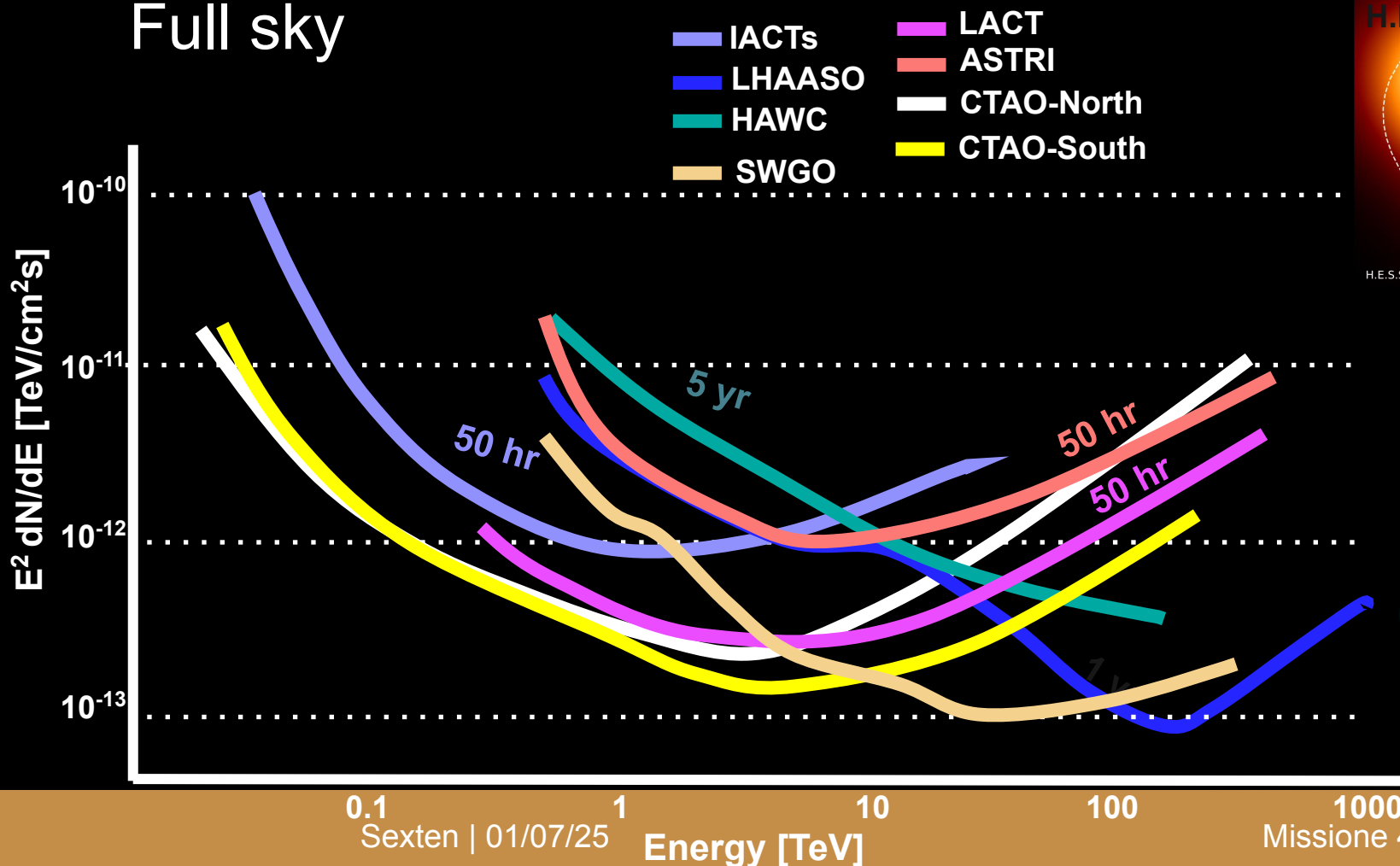
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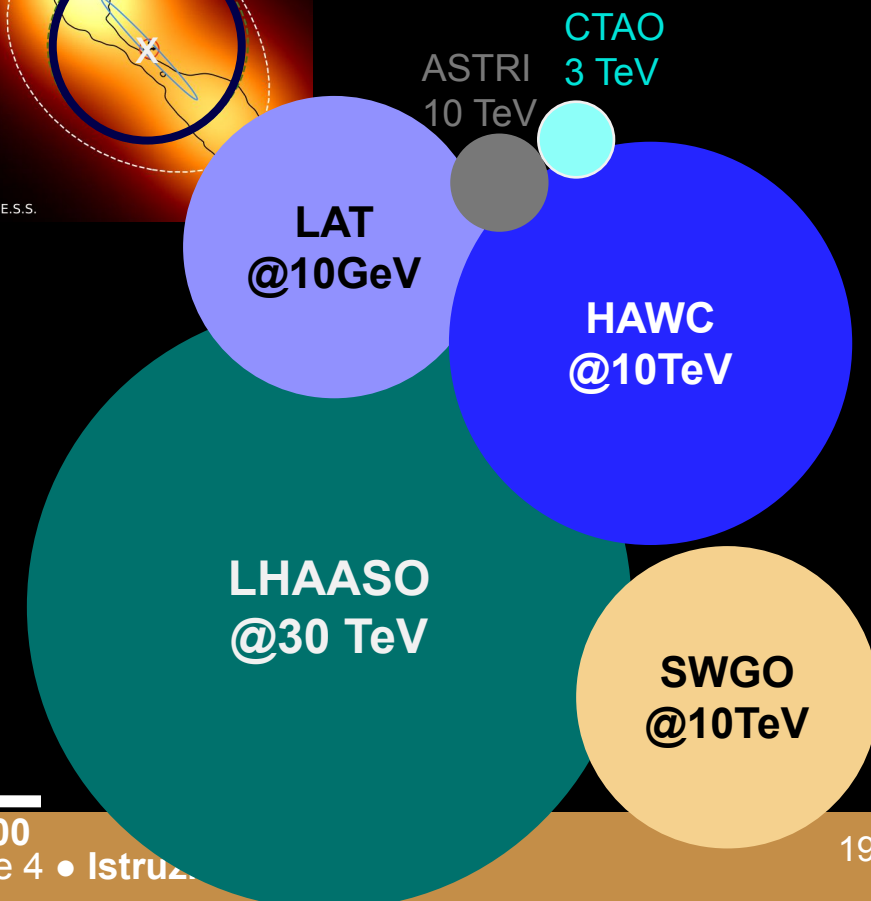
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Steady source performance

Full sky



H.E.S.S.





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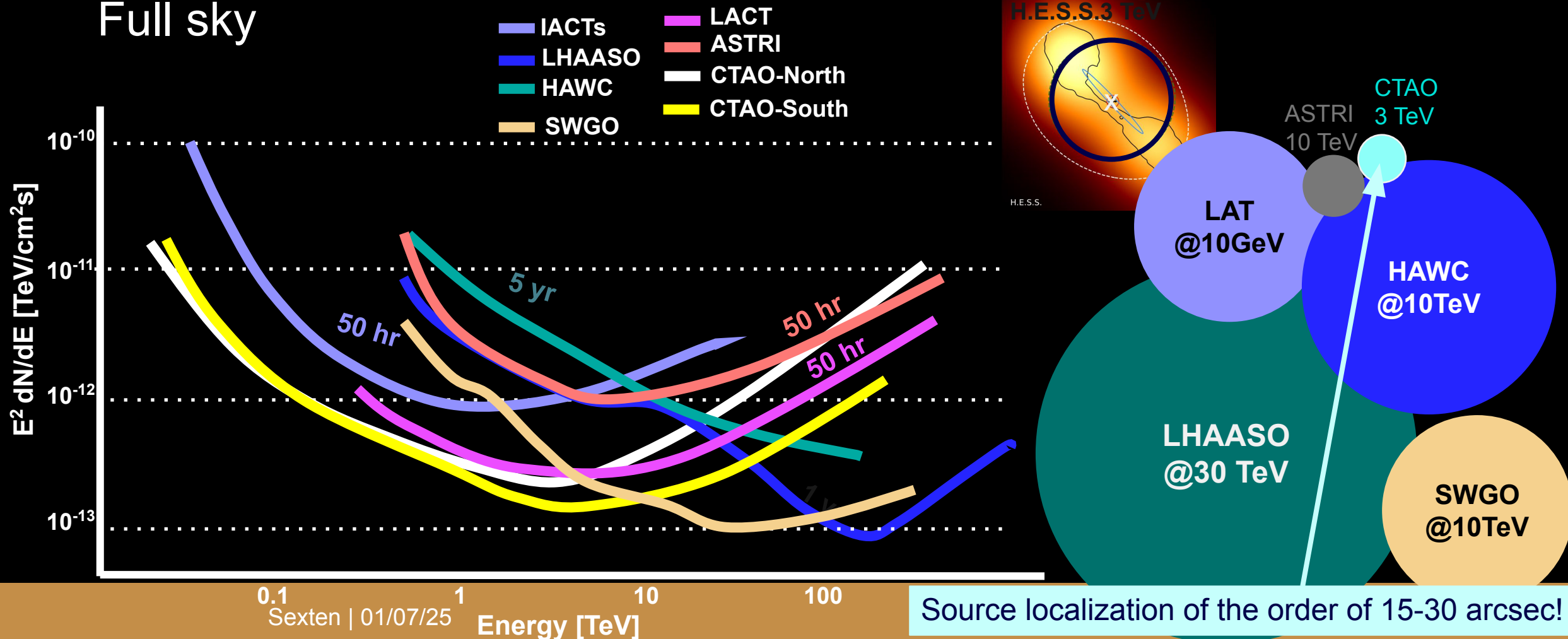
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Steady source performance

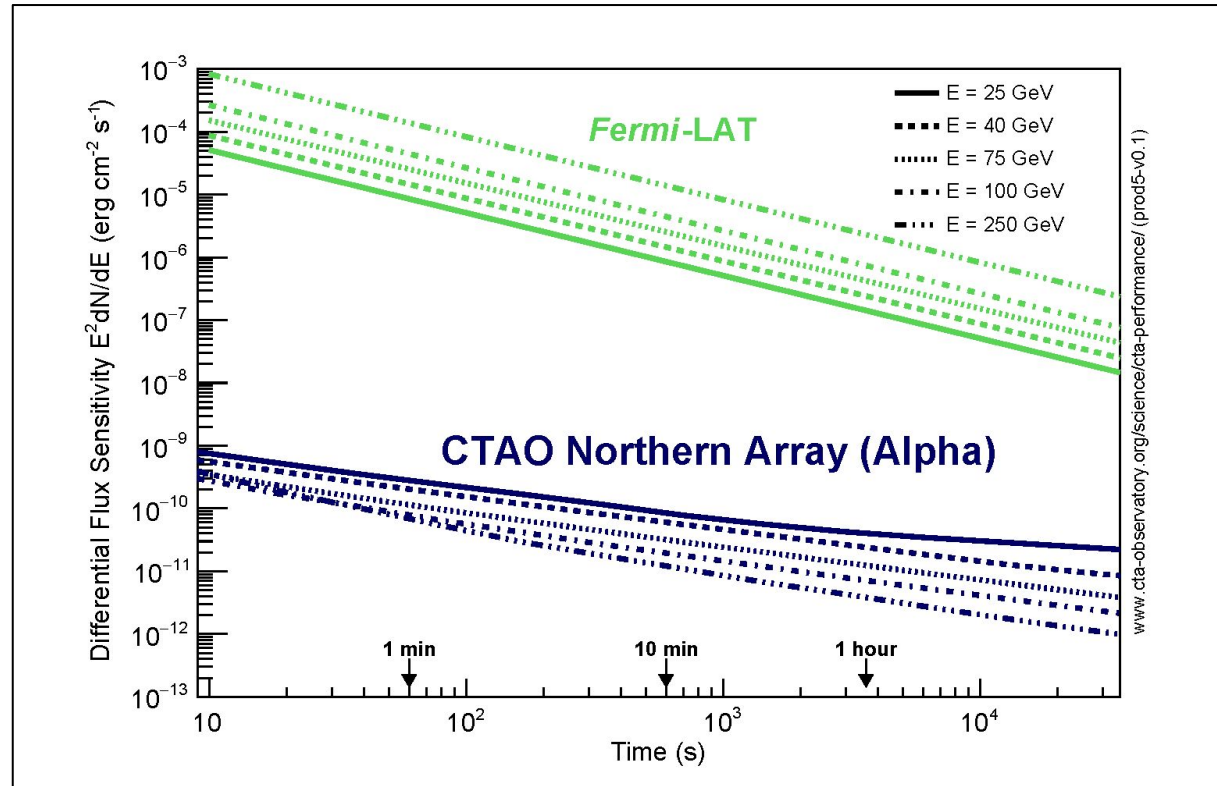
Full sky





Variable-flux-phenomena sensitivity

alias short-term observation sensitivity



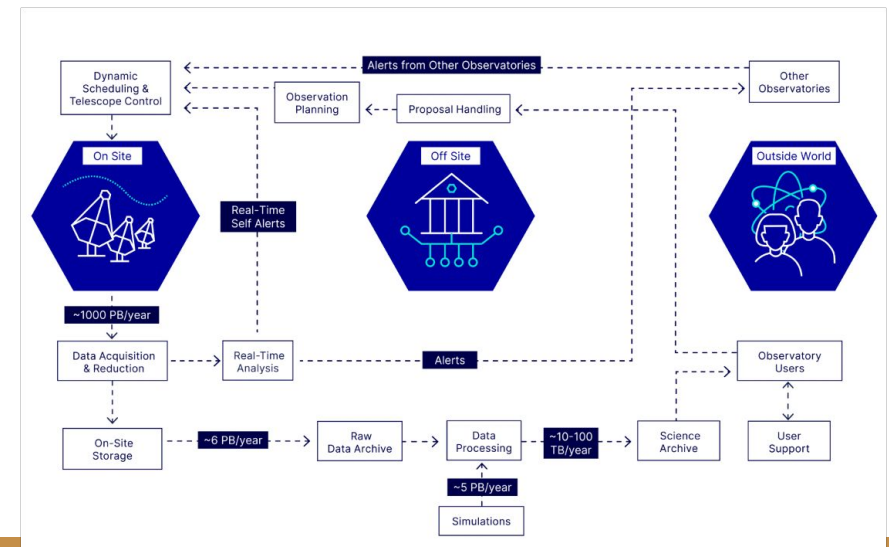


Transient Synergies

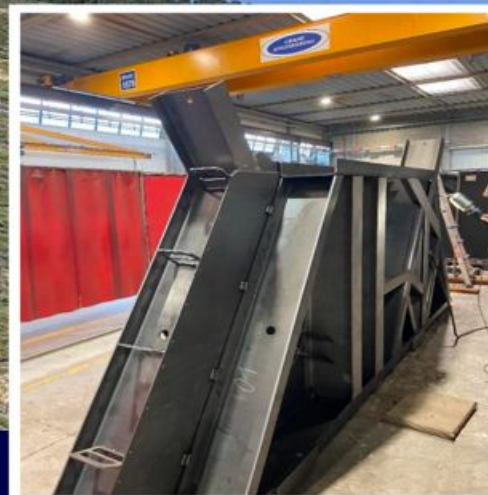
- **Arrays of IACTs are optimal instruments for the study of transient/variable phenomena** (especially at low-energies)
.... **however they require observational triggers**
- **Particle array detectors are good triggering instruments** for nearby sources or super strong phenomena thank to their large FoV (serendipitous discoveries) and 100% duty cycle

To guarantee the maximum exploitation of the CTAO follow-up capabilities

- CTAO telescopes (LSTs) can re-pointing any point in the sky in less than 20"
- CTAO telescope can process the received alert in less than 50"



Construction has started





Conclusions

- **The future of the VHE gamma ray astronomy in the next decade is very bright**
- **CTAO first data is at reach within the next 3 years with performance capabilities that are already a factor 2 better than the existing facilities**
- **Transient phenomena can be explored at maximum capabilities already in 3 years from now**
- **We can start exploring the highest energies also in the Southern sky thanks to the first intermediate configuration**