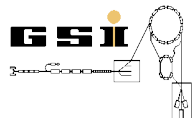


A new approach to detect hypernuclei and isotopes in the QMD phase space distribution at relativistic energies

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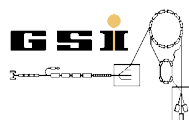
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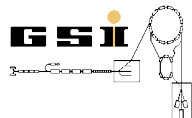
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- ➔ Because, apart from emitted elementary particles, **they carry the only information** that the experimental instruments can measure.
- ❖ Making clusters is **not an easy task**, because it involves, in a complex environment:
 - ▶ the fundamental nuclear properties,
 - ▶ quantum effects,
 - ▶ and variable timescales.





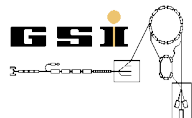
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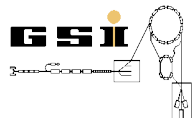




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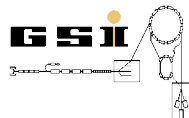


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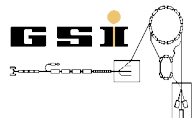


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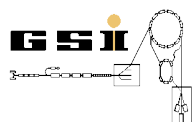


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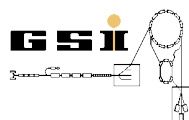


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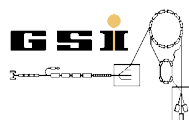
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Simulations show: Clusters chosen that way at early times are the pre-fragments of the final state clusters, because fragments are not a random collection of nucleons at the end but initial-final state correlations.



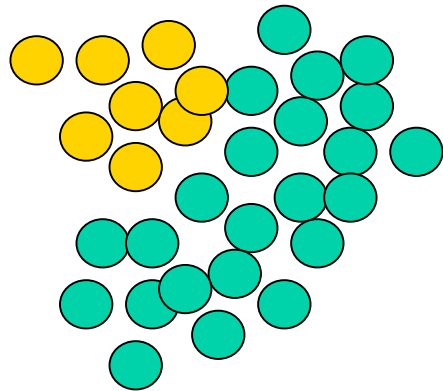


SACA: How does this work?

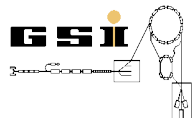
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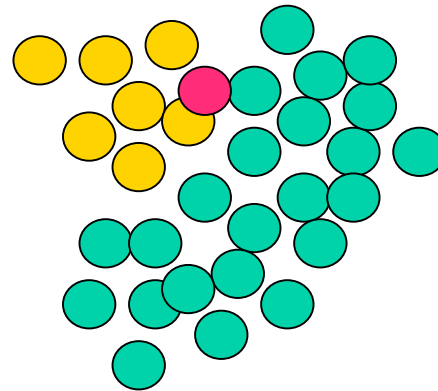
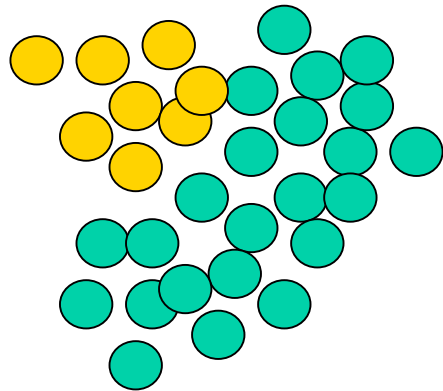


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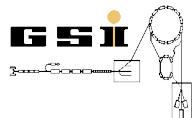
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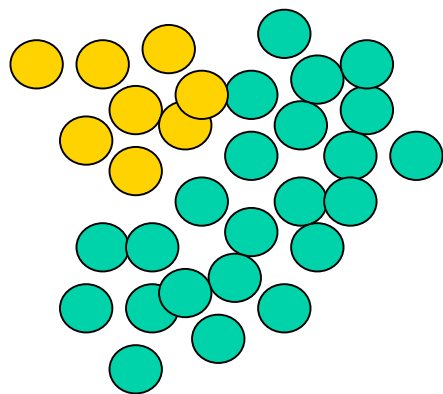


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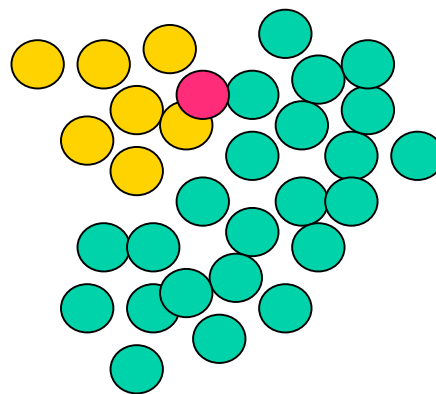
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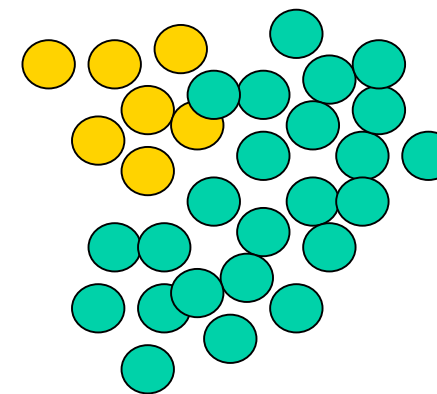
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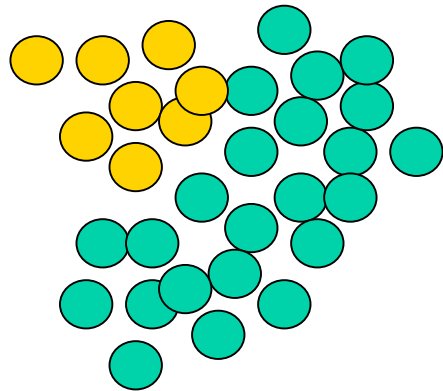


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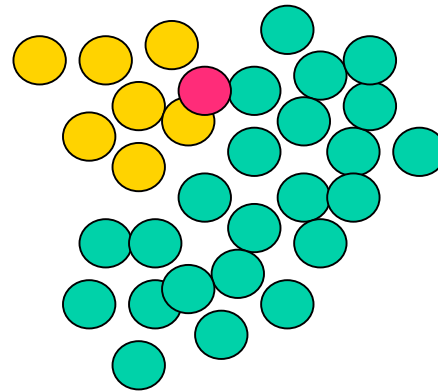
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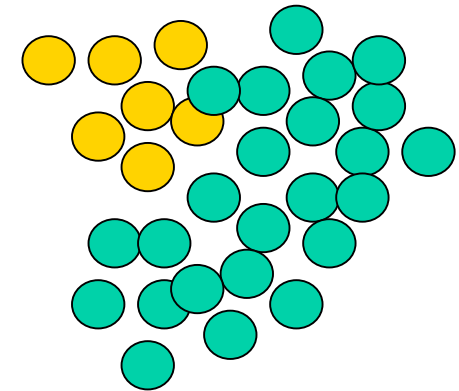
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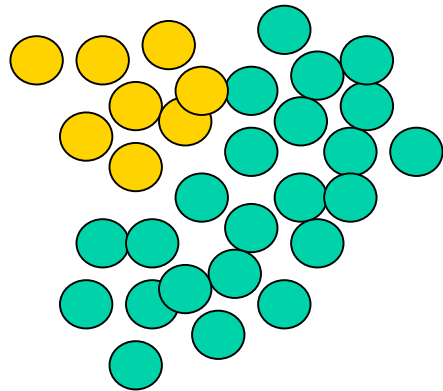


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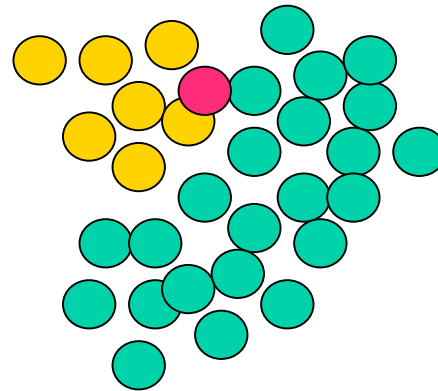
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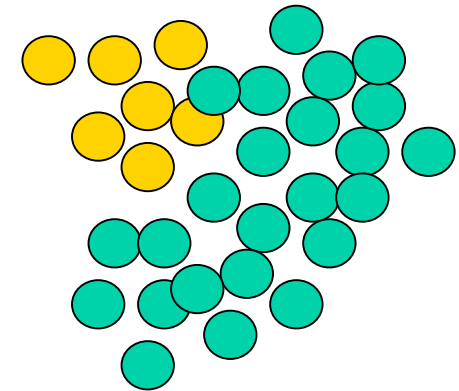
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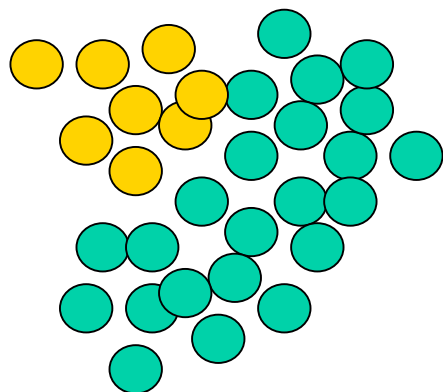


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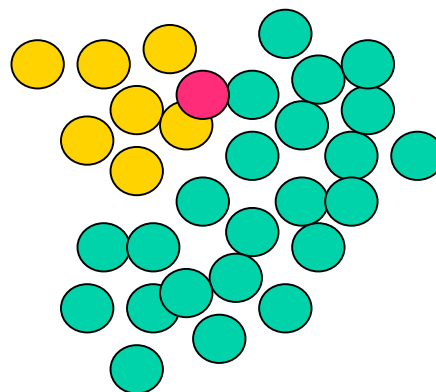
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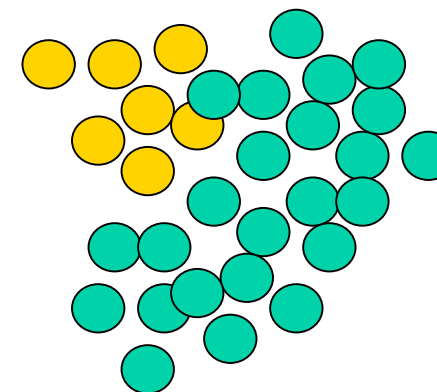
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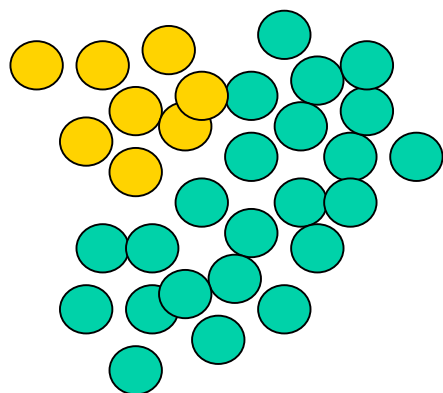


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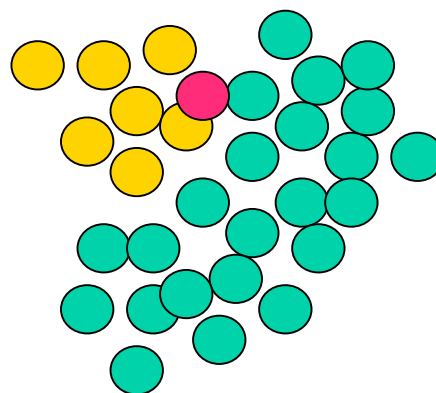
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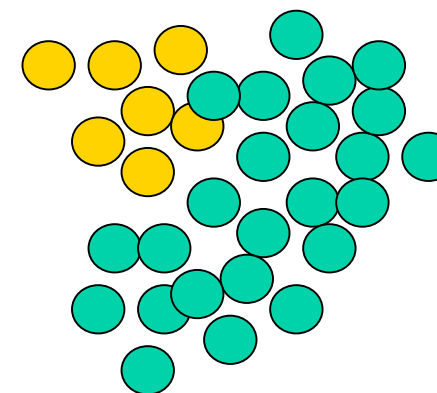
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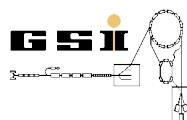


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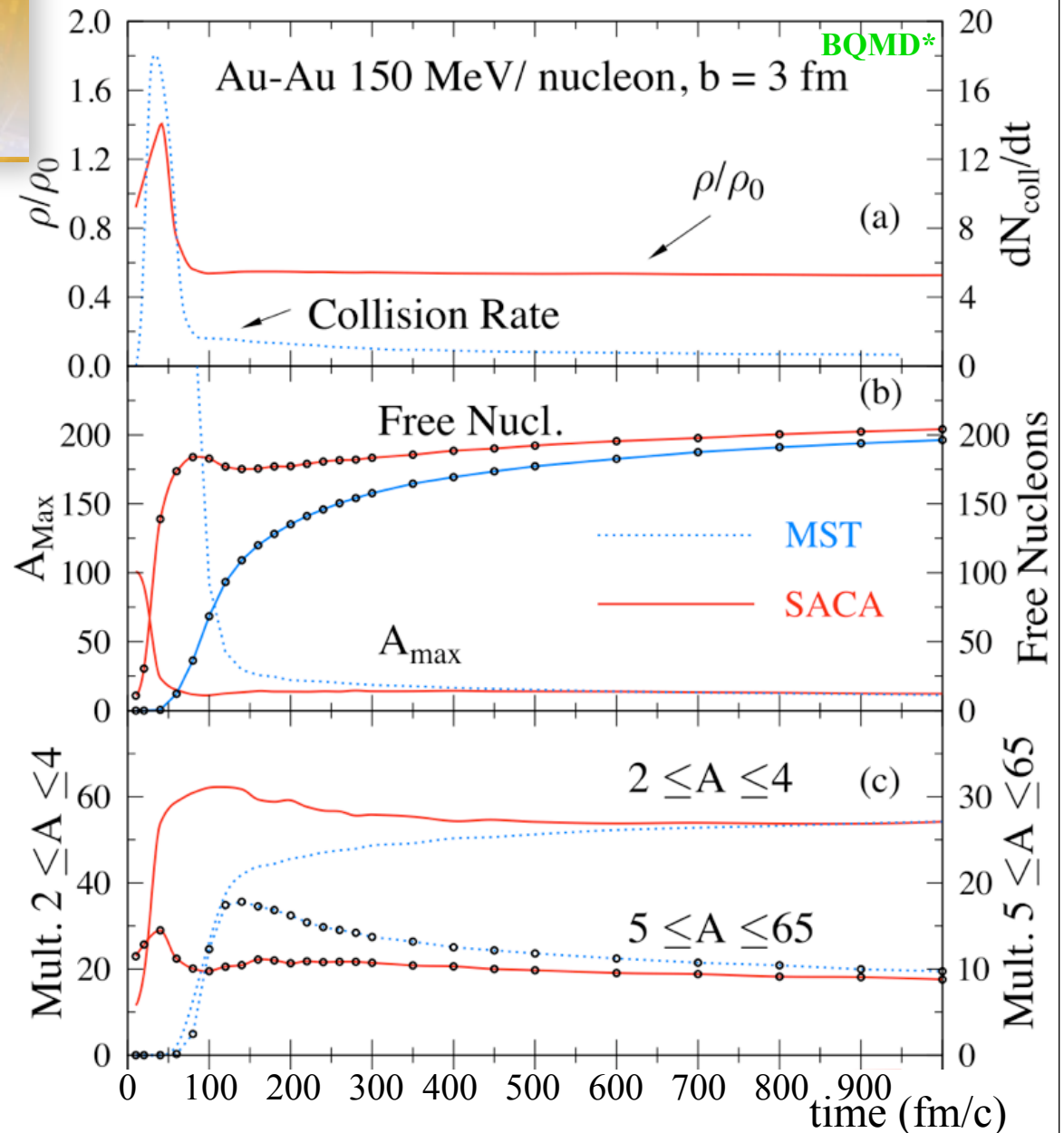
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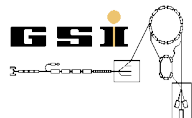
It leads automatically to the most bound configuration.



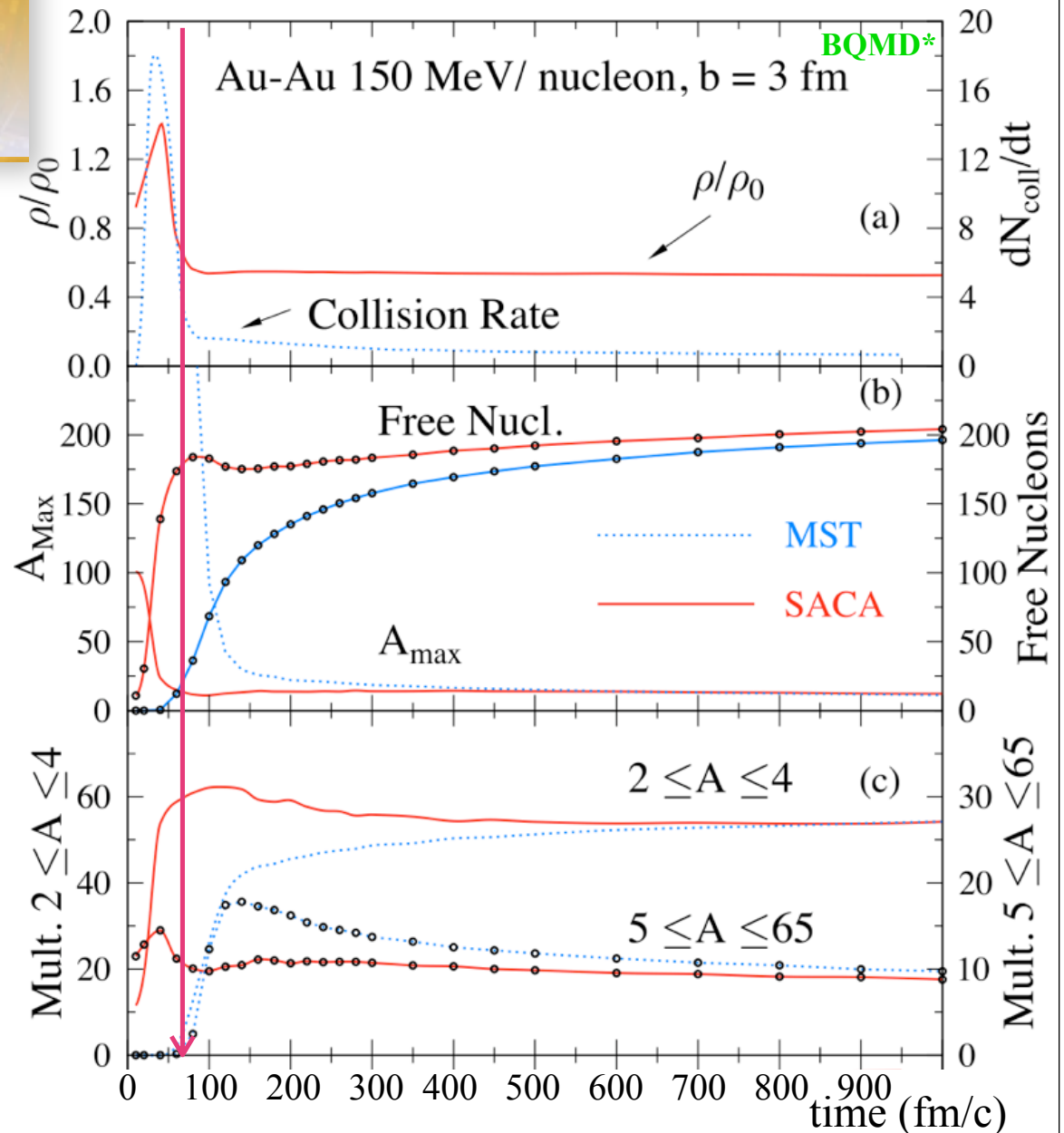
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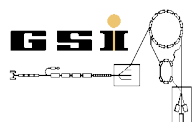
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Nuclear Physics A 619 (1997) 379-390



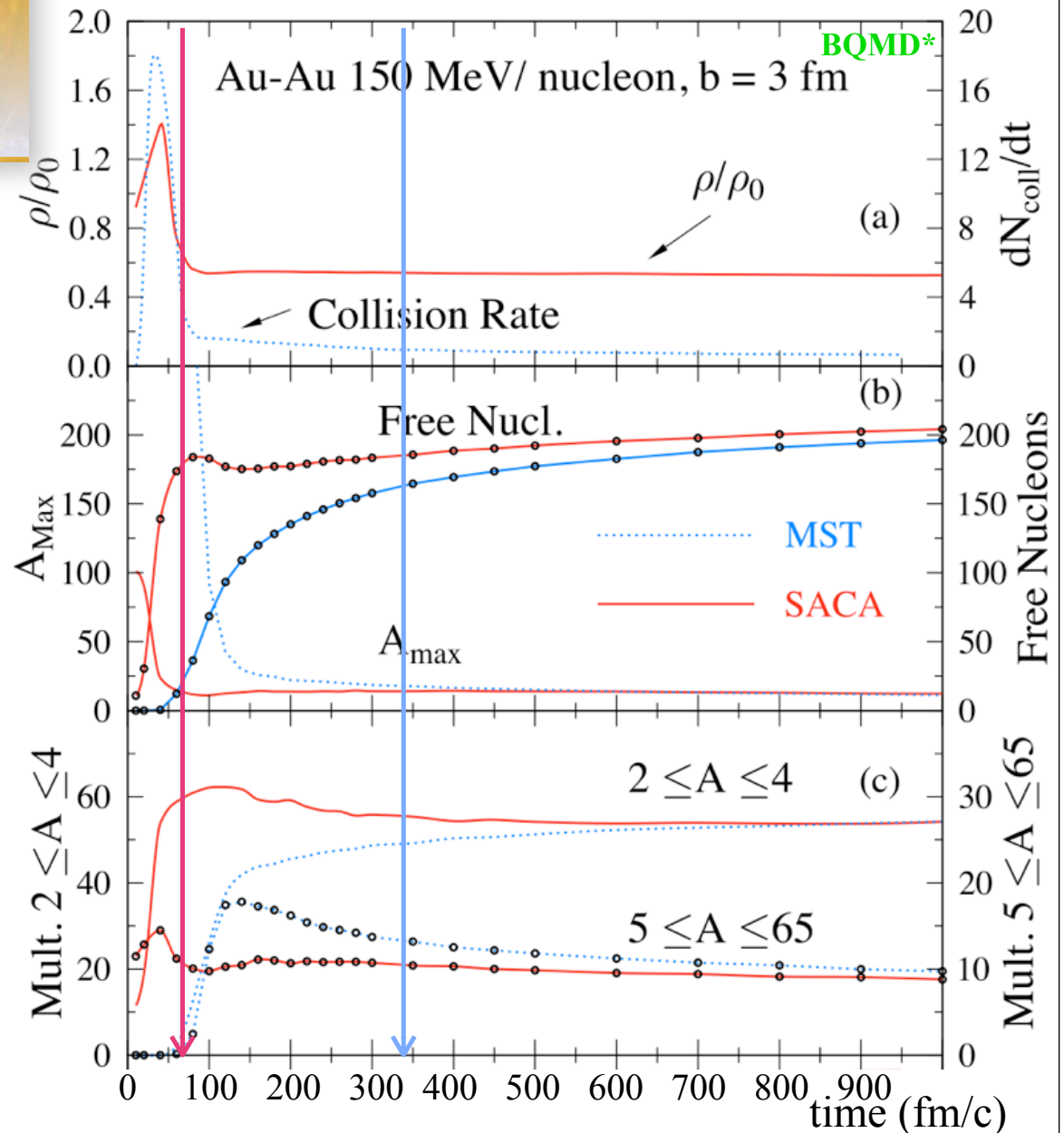
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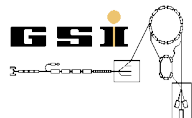
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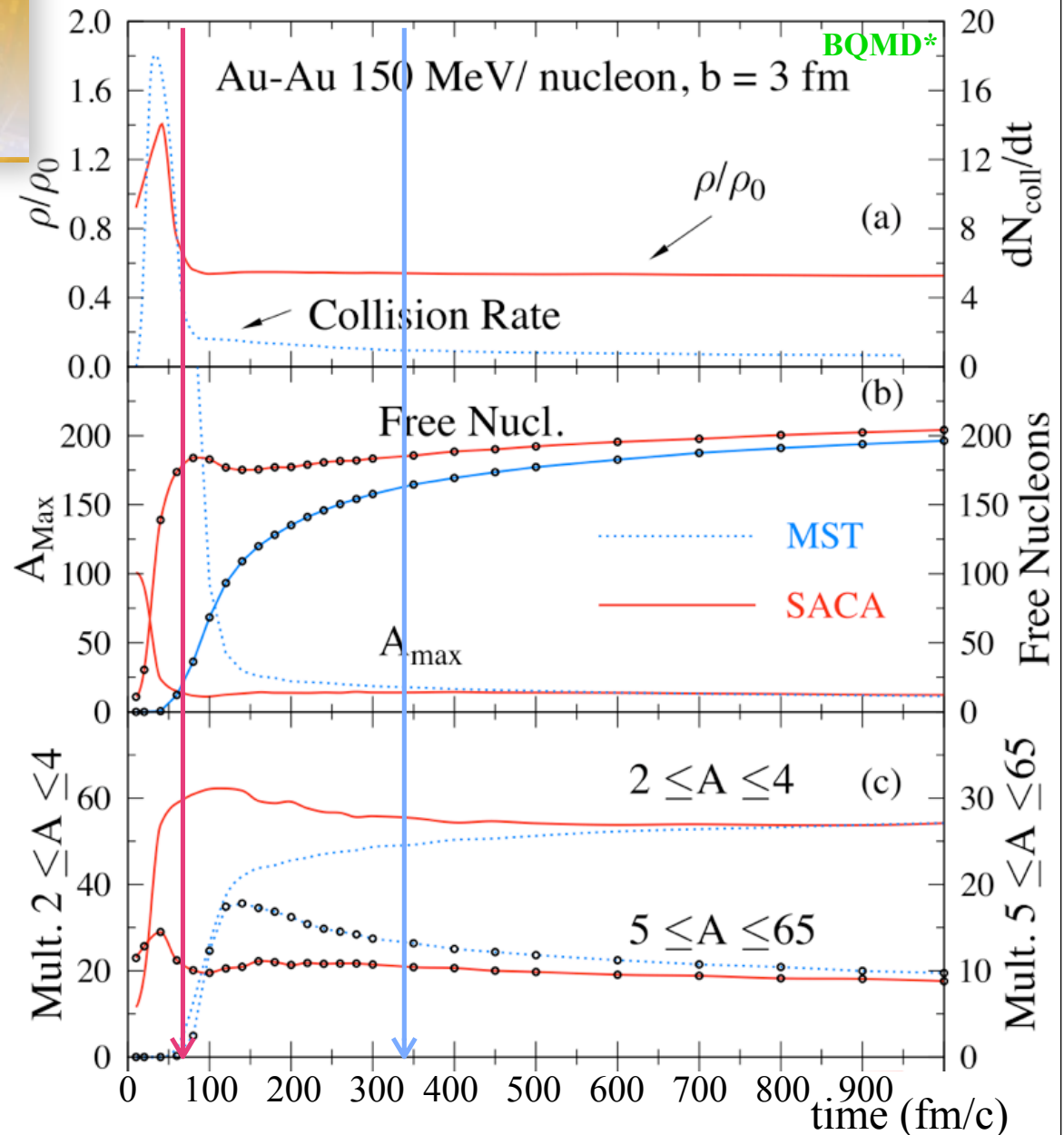


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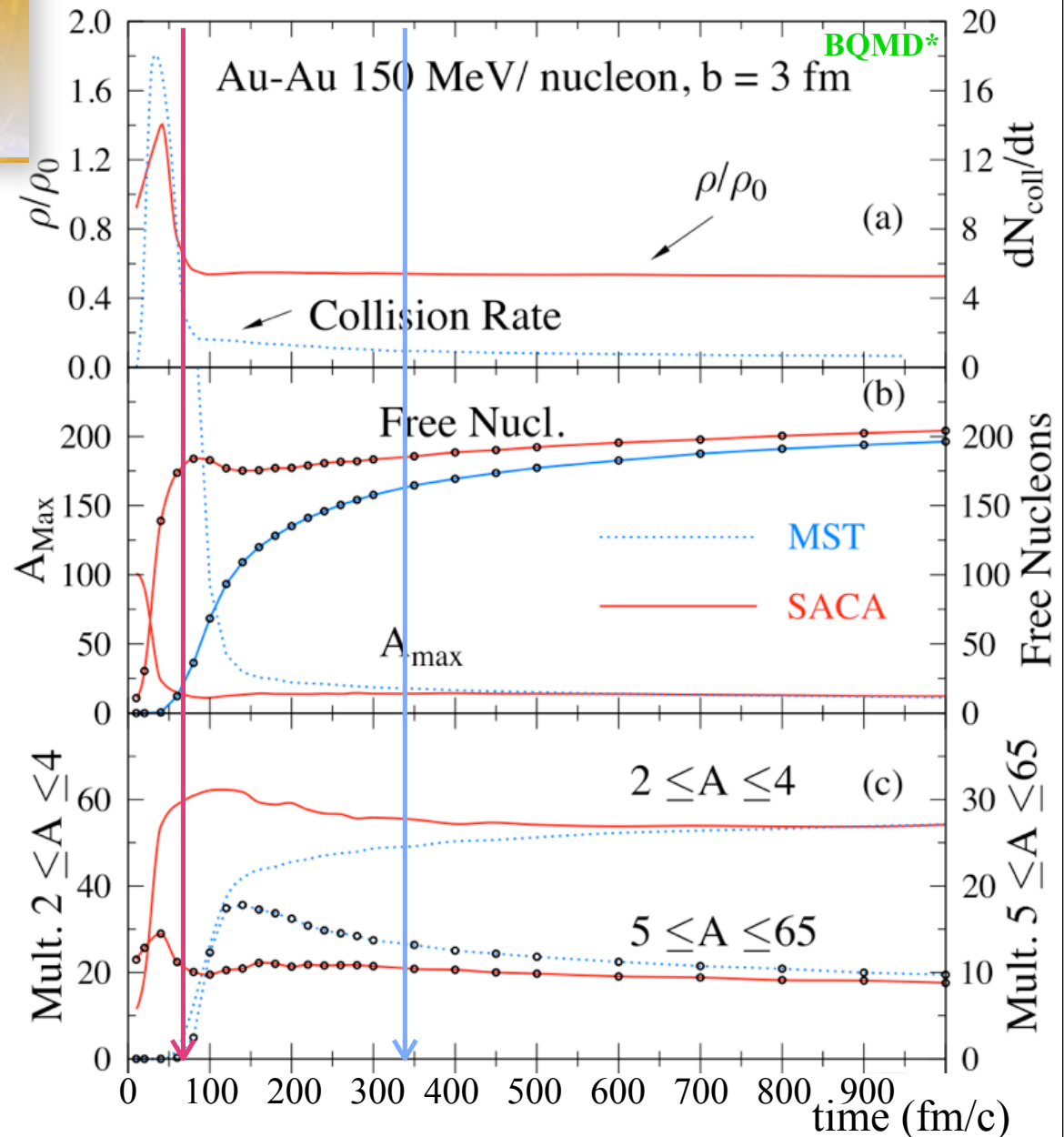


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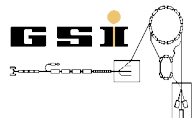
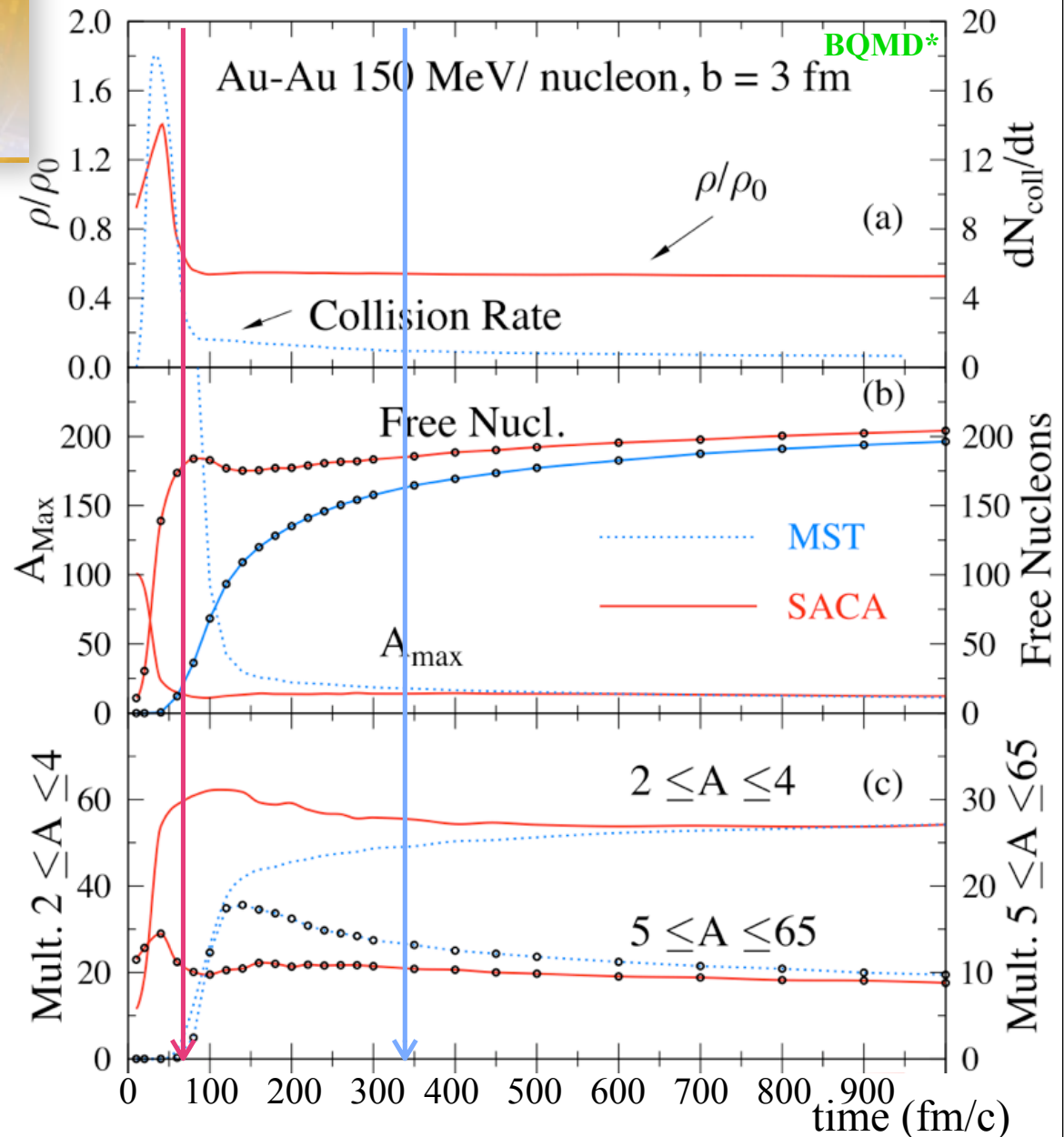


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- ➔ With MST, one has to consider necessarily later times (typically 200-400 fm/c), where the dynamical conditions are no longer the same.
- ▶ Advantage of SACA : the fragment partitions can reflect the early dynamical conditions (Coulomb, density, flow details, strangeness...).

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Toward the isotope yields... IQMD + new SACA

SACA is applied here on the IQMD transport model* calculations

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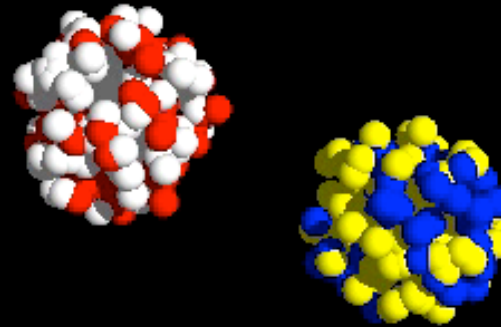


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Au+Au at 100 A.MeV - $b=7$ fm



An example of complex system accurately measured

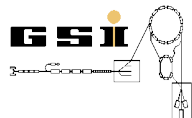
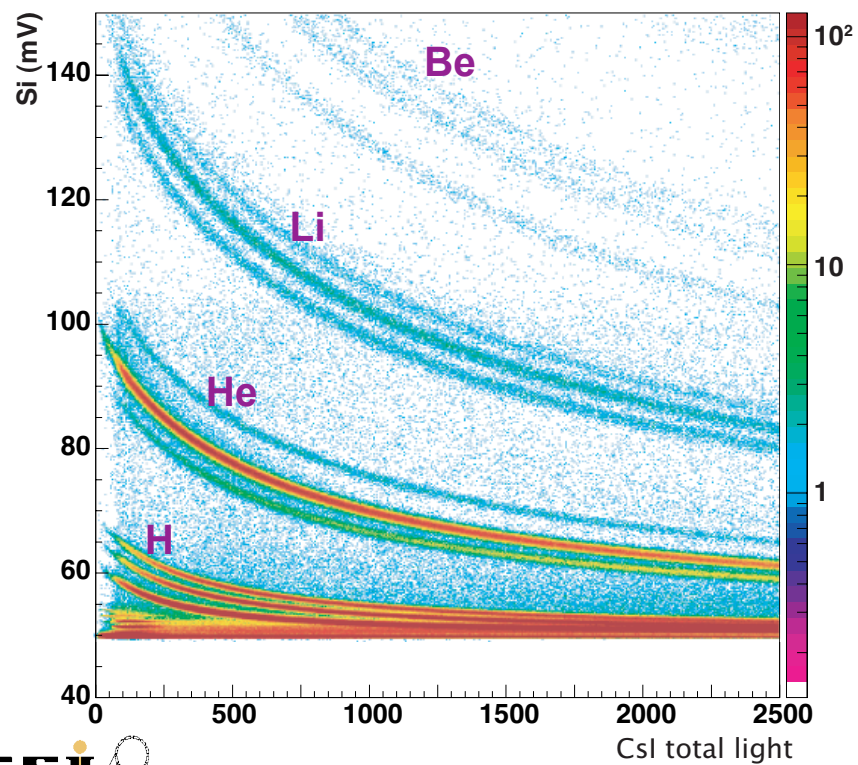
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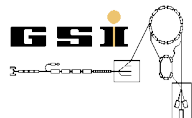
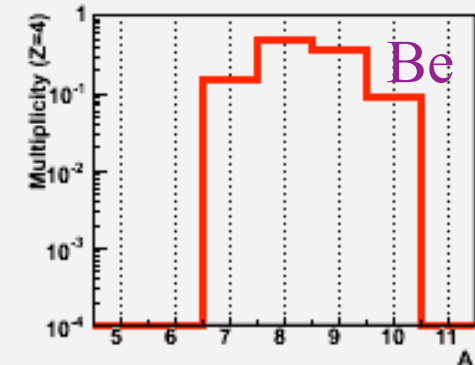
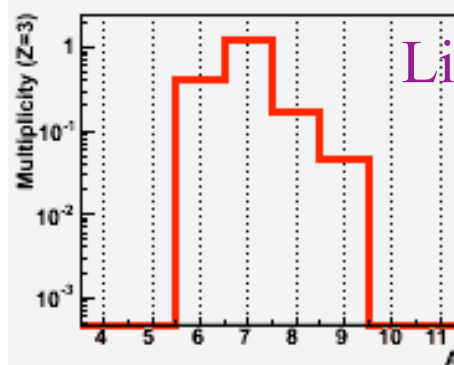
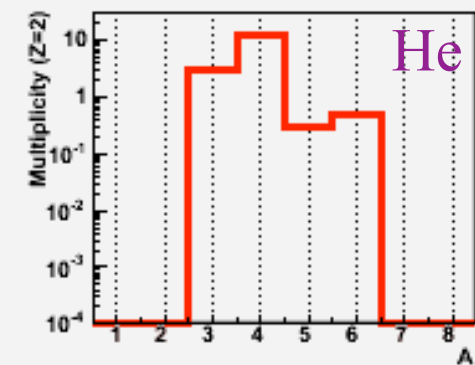
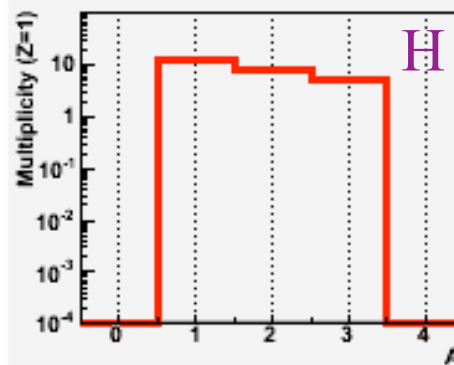
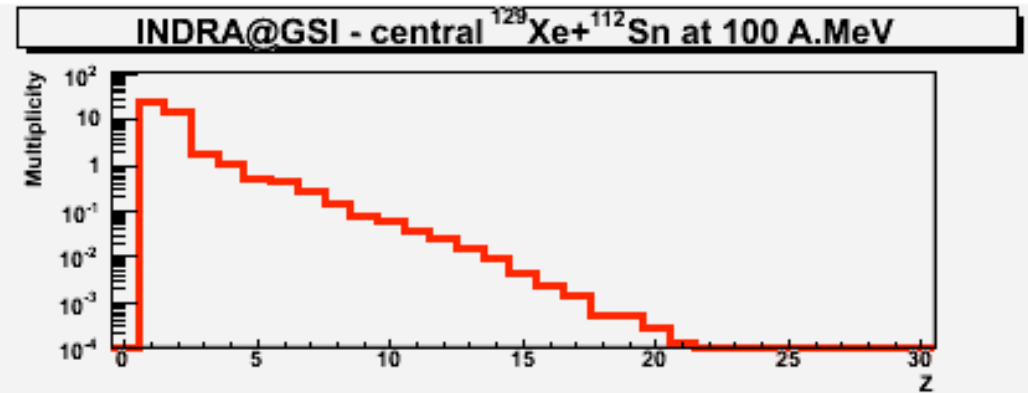
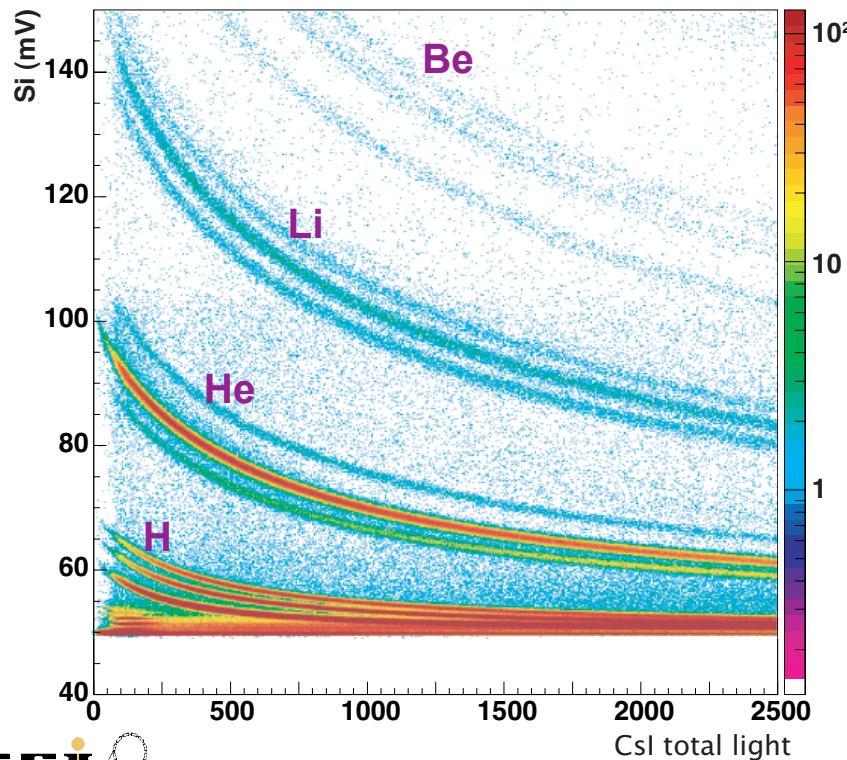
ALADiN - INDRA collaboration



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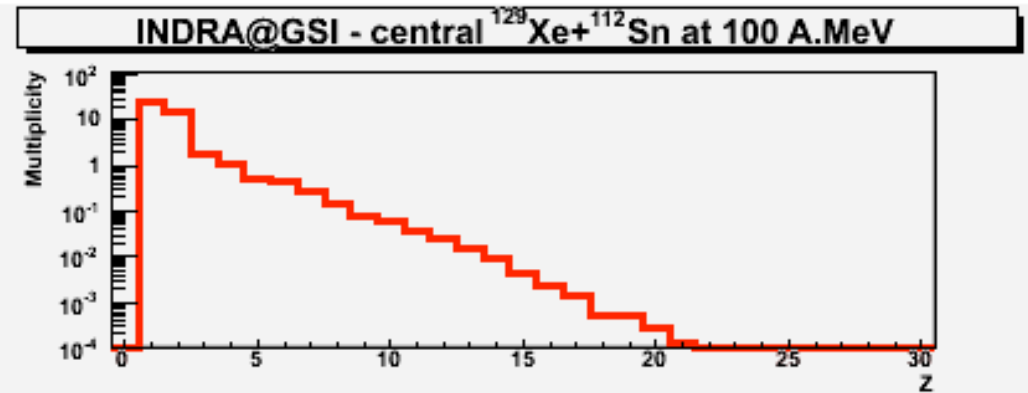
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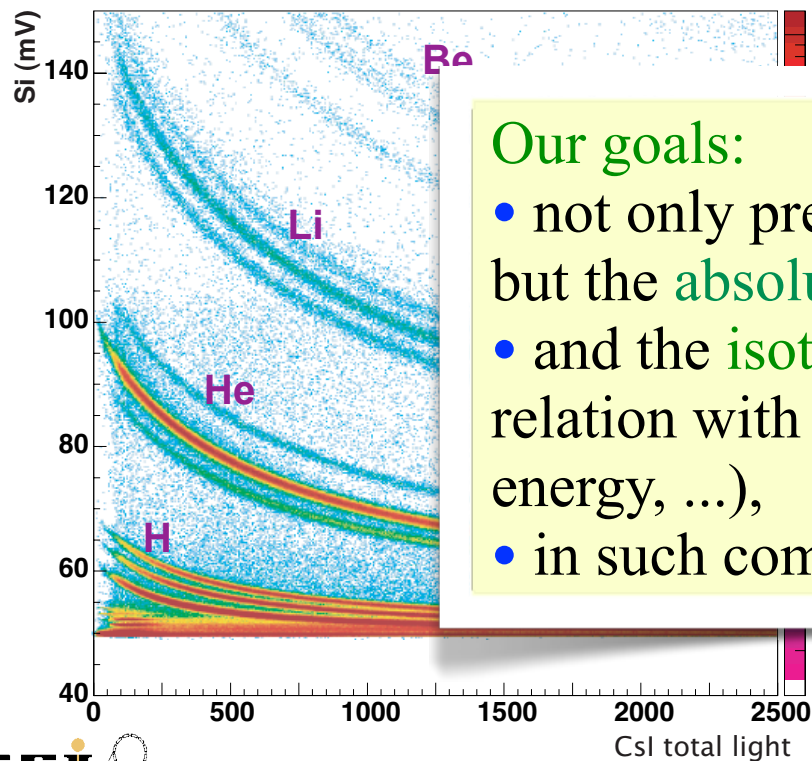


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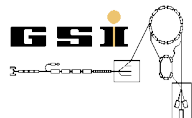
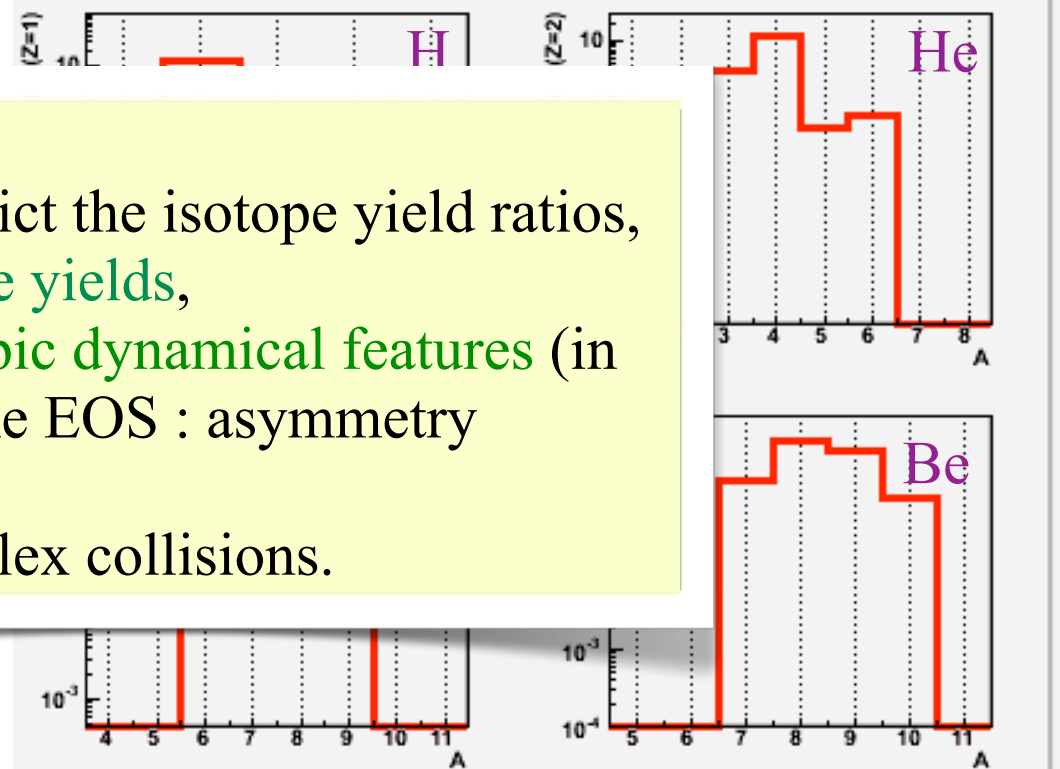


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Our goals:

- not only predict the isotope yield ratios, but the **absolute yields**,
- and the **isotopic dynamical features** (in relation with the EOS : asymmetry energy, ...),
- in such complex collisions.



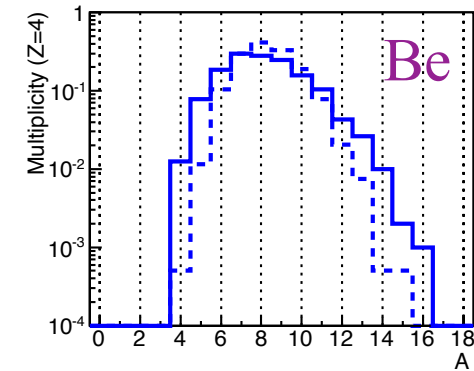
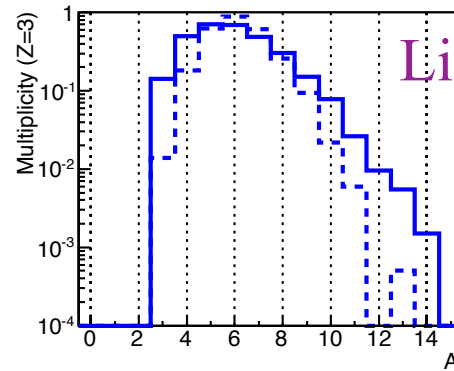
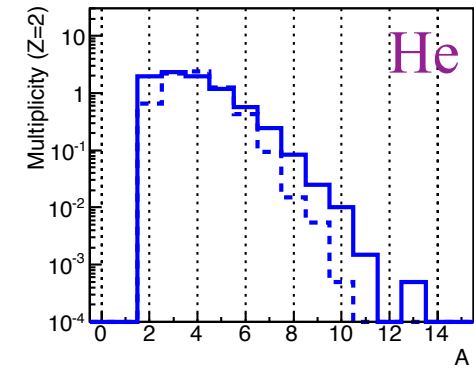
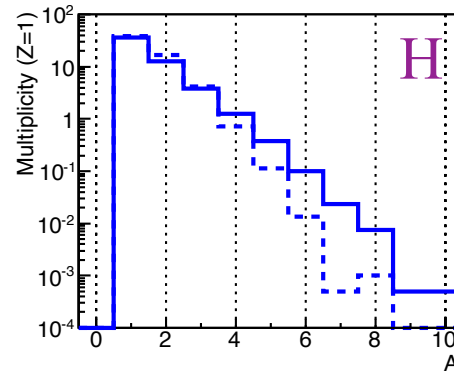
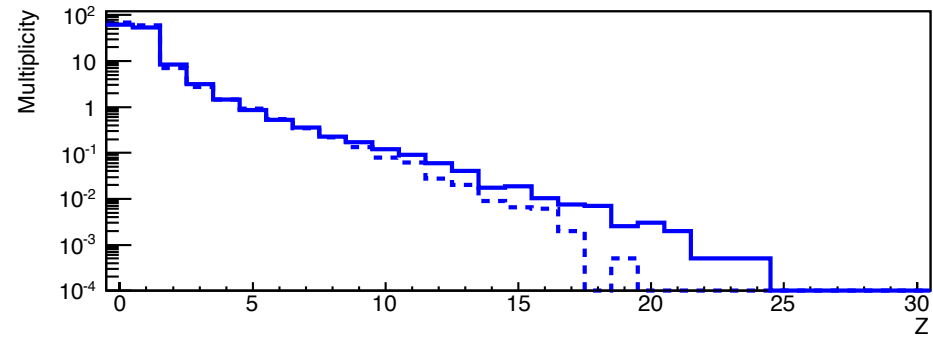
SACA versus coalescence (Minimum Spanning Tree)

IQMD $^{136}\text{Xe} + ^{112}\text{Sn}$ at 100 A.MeV, $b=1$ fm, $t_{\text{SACA}} = 60$ fm/c

SACA version:

--- MST only (200 fm/c)

— $E_{\text{asy}}=0$, no pairing



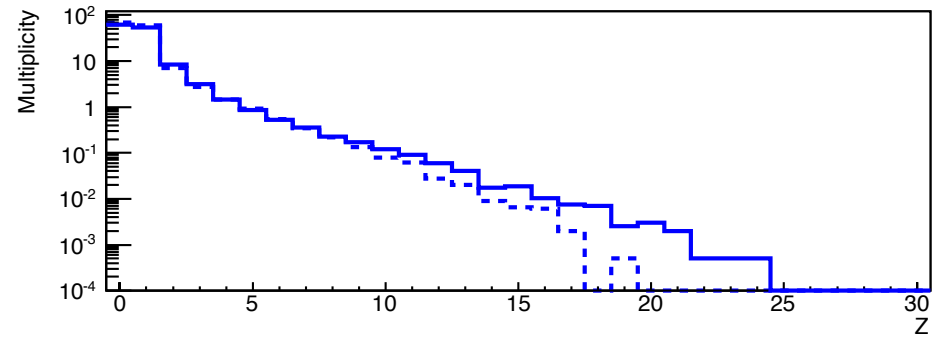
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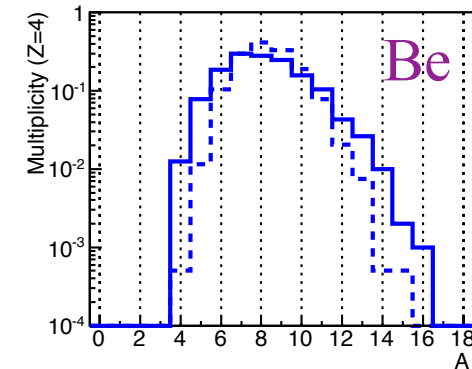
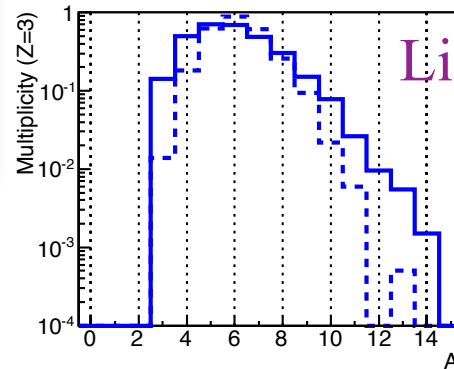
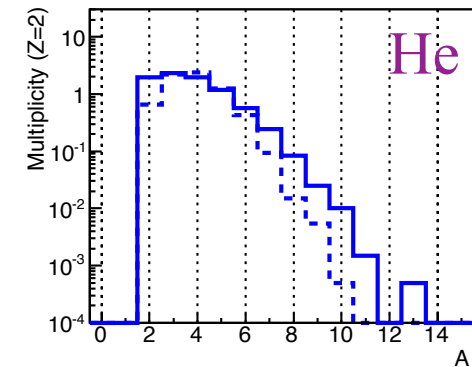
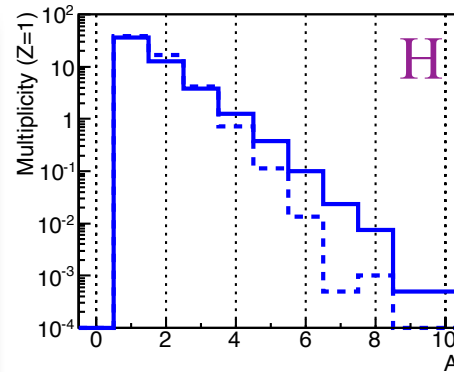
----- MST only (200 fm/c)

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At this stage, SACA contains as ingredients of the potential making the binding energy of the clusters :

- ① volume component:
mean field (Skyrme, dominant)
- ② correction of surface effects:
Yukawa

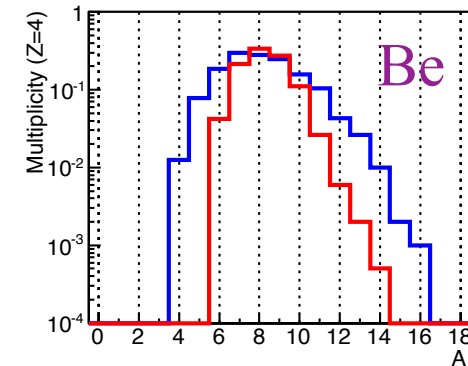
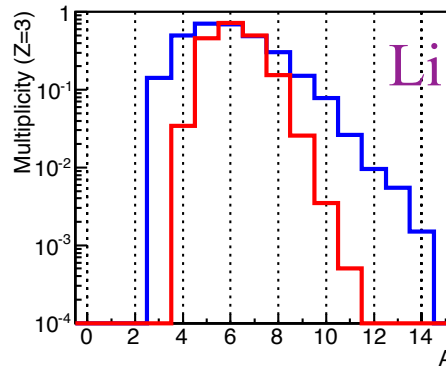
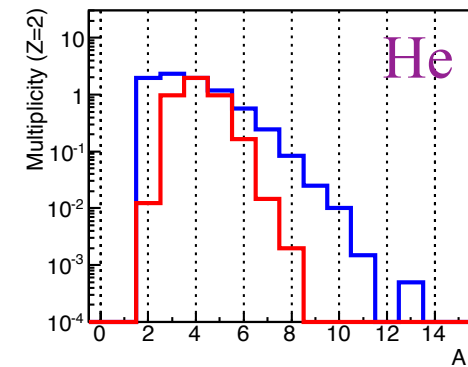
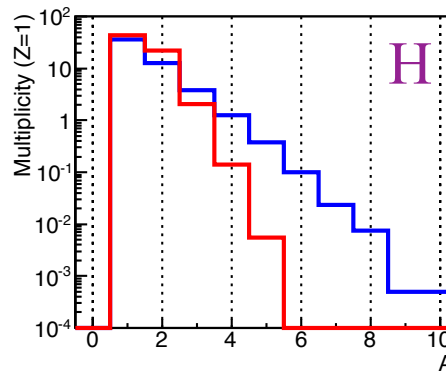
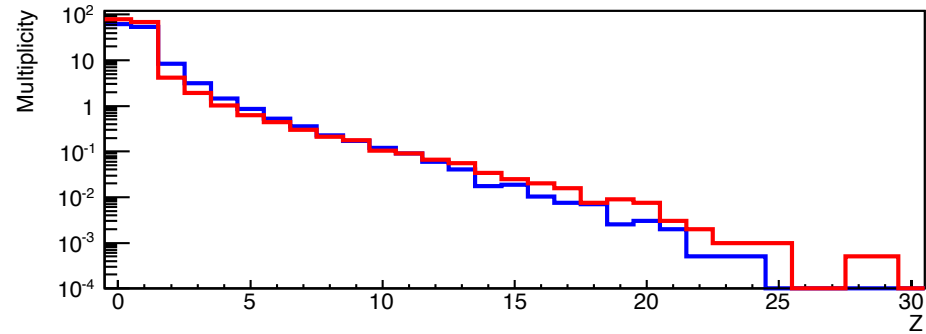


SACA with asymmetry energy

IQMD $^{136}\text{Xe} + ^{112}\text{Sn}$ at 100 A.MeV, $b=1$ fm, $t_{\text{SACA}} = 60$ fm/c

SACA version:

- $E_{\text{asy}}=0$, no pairing
- $E_{\text{asy}}=32$ MeV ($\gamma=1$), no pairing

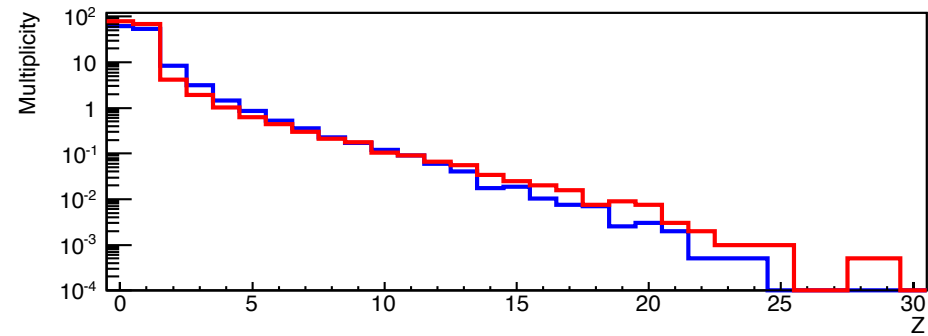


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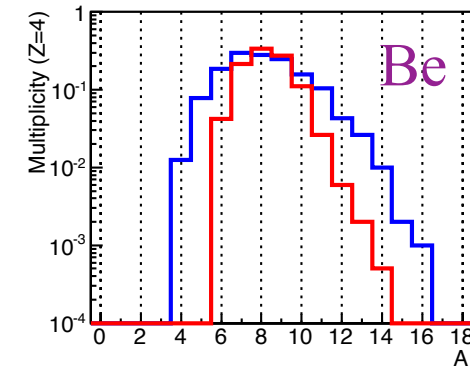
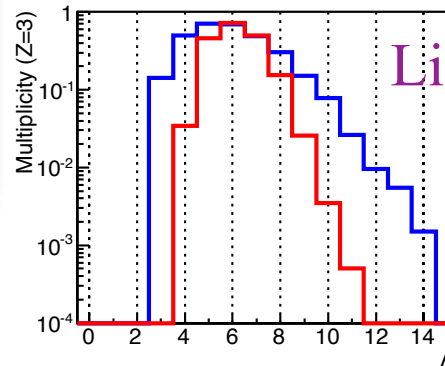
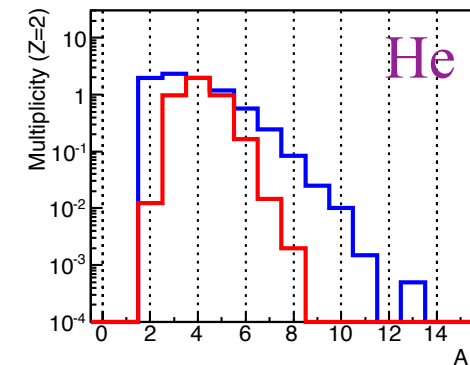
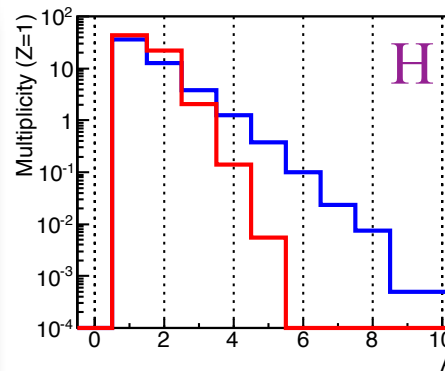


Here, in IQMD and SACA, we adopt the following asymmetry energy parametrisation:

$$E_{\text{asy}} = E_0 \cdot (\langle \rho_B \rangle / \rho_0)^{(\gamma-1)} \cdot (\langle \rho_n \rangle - \langle \rho_p \rangle) / \langle \rho_B \rangle$$

with $E_0 = 32$ MeV, $\gamma = 1$ («stiff»)

- ➔ Z and A yields not strongly modified
- ➔ Isotope yields shrink onto the $N=Z$ line
- ➔ Still not fully realistic: shell, odd-even effects (pairing) still absent.

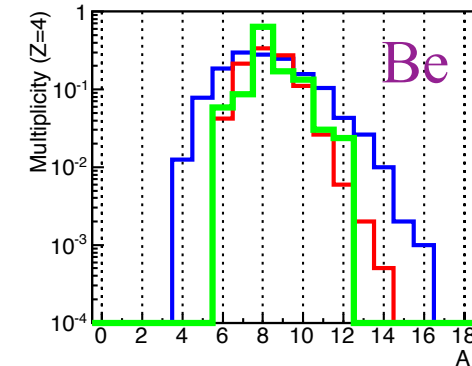
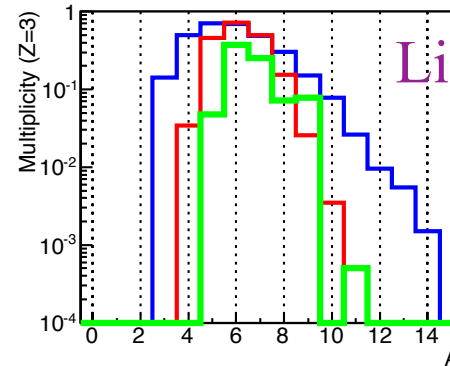
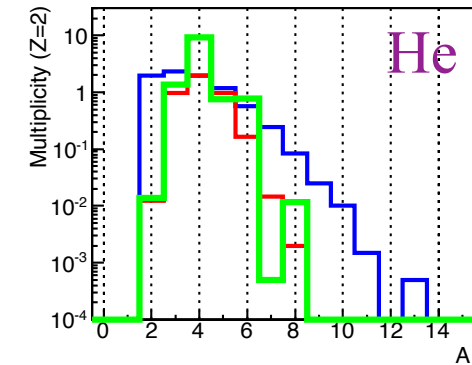
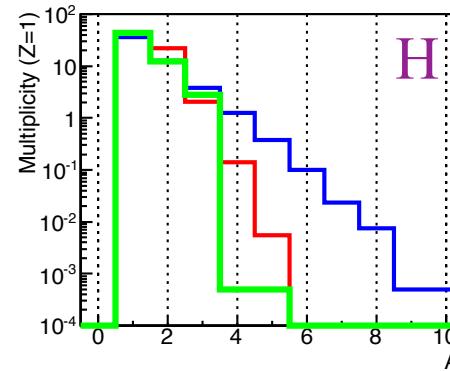
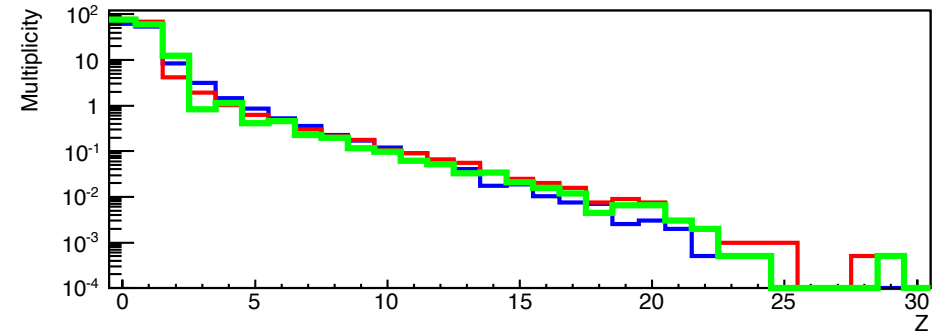


SACA with asymmetry energy and pairing

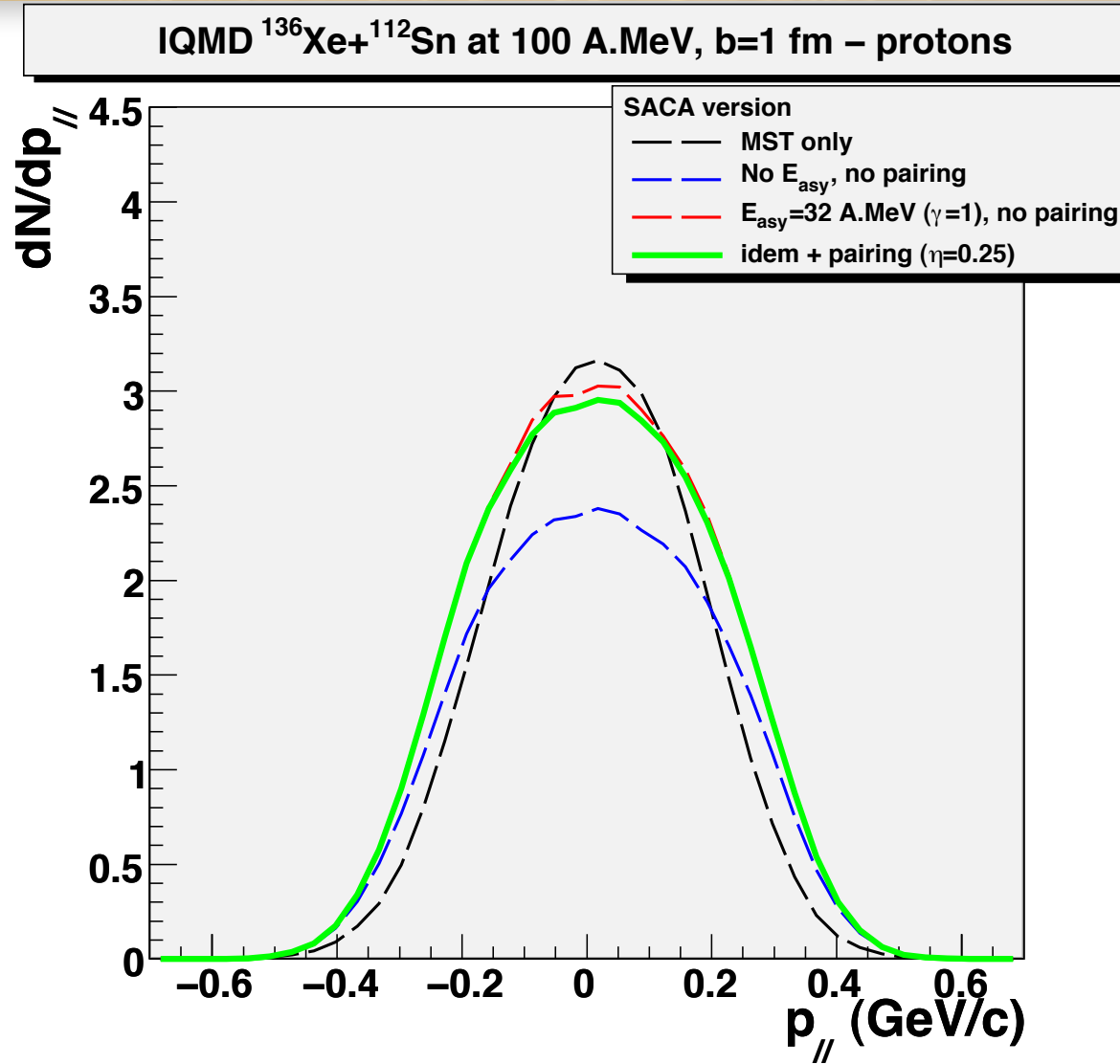
INDRA@GSI - central $^{129}\text{Xe} + ^{112}\text{Sn}$ at 100 A.MeV

SACA version:

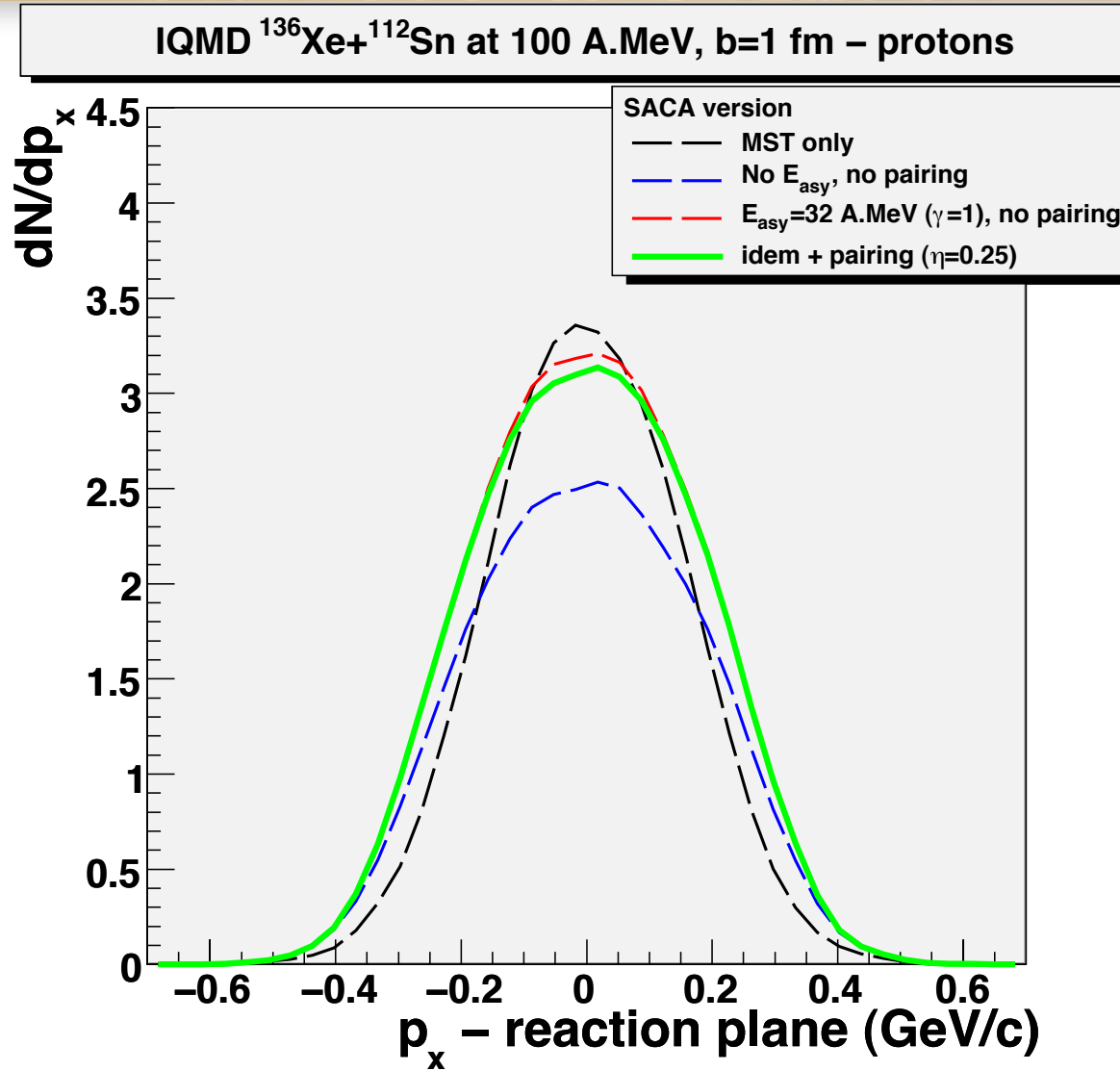
- $E_{\text{asy}}=0$, no pairing
- $E_{\text{asy}}=32$ MeV ($\gamma=1$), no pairing
- $E_{\text{asy}}=32$ MeV ($\gamma=1$) + $\eta_{\text{pairing}}=0.25$



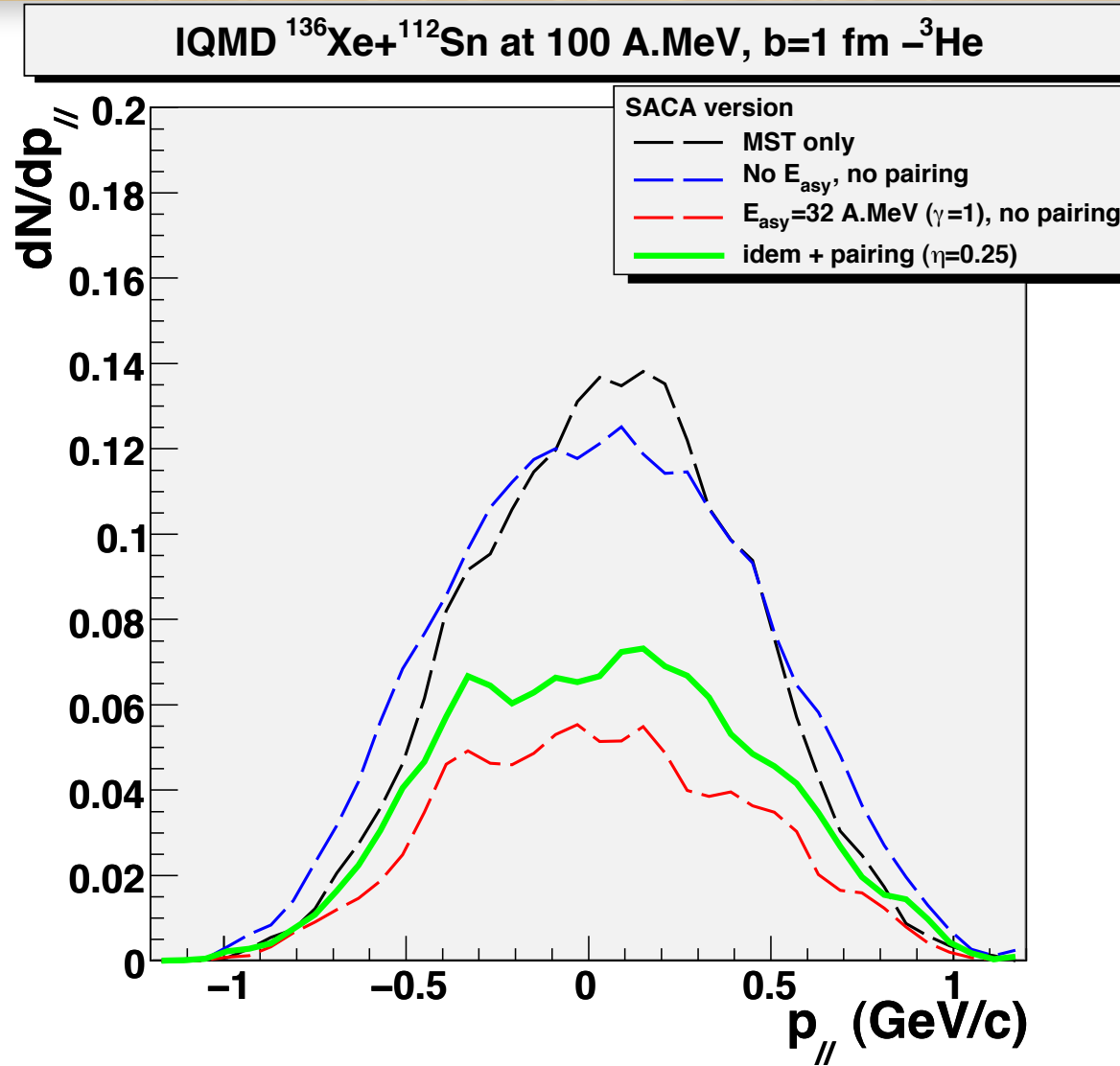
How the dynamical patterns of isotopes are affected



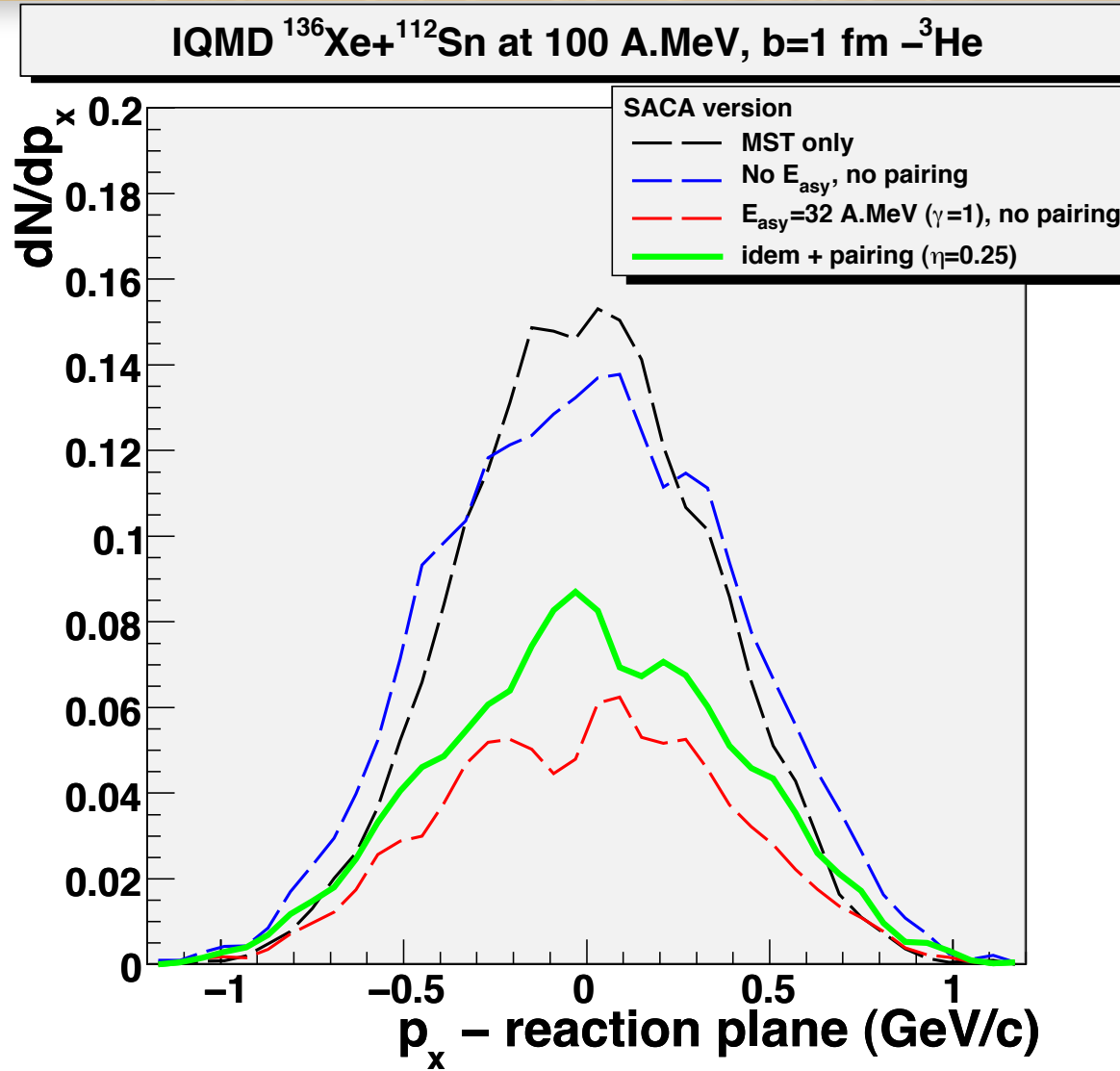
How the dynamical patterns of isotopes are affected



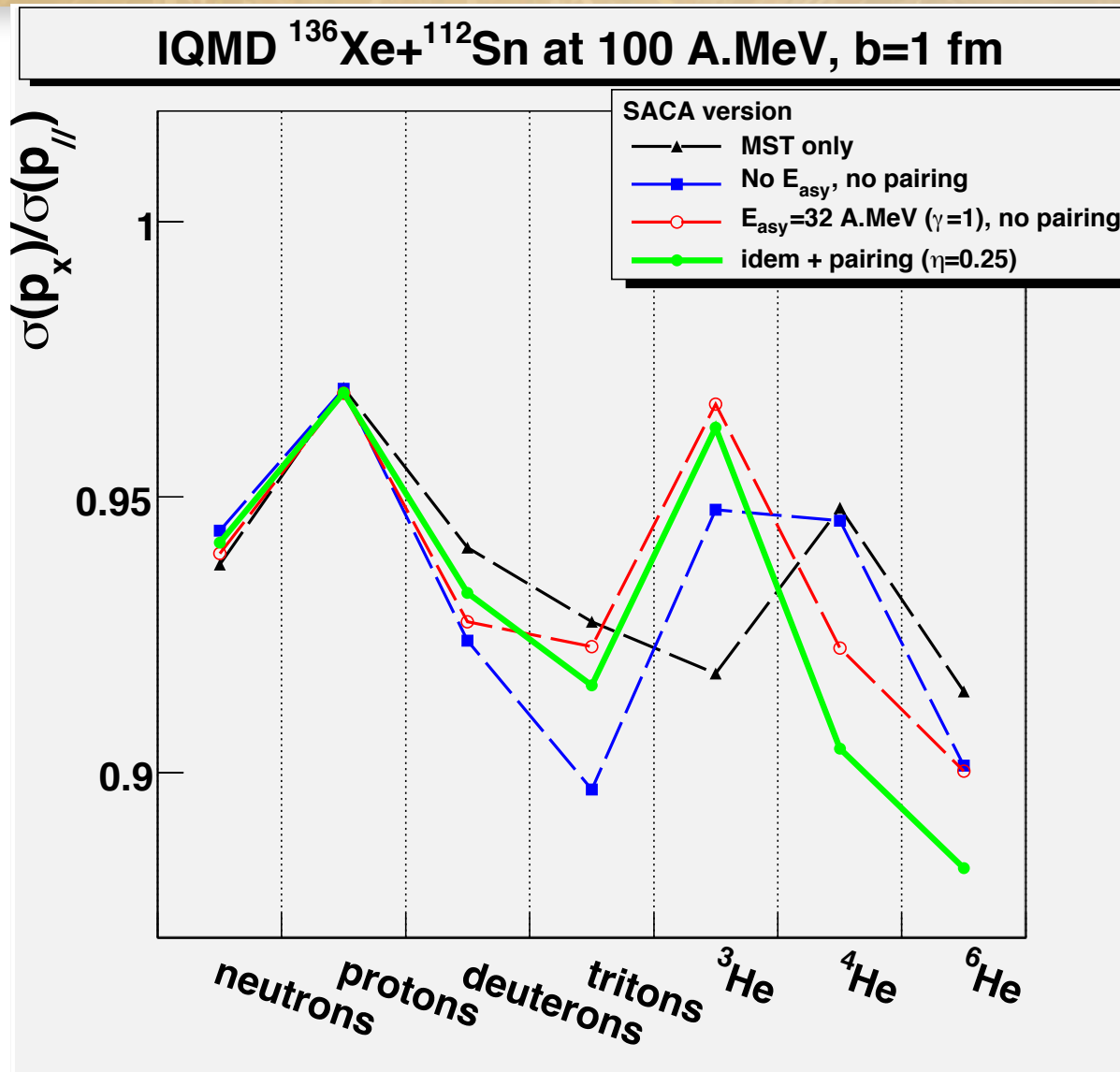
How the dynamical patterns of isotopes are affected



How the dynamical patterns of isotopes are affected



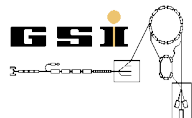
How the dynamical patterns of isotopes are affected





Excitation energy of the primary fragments

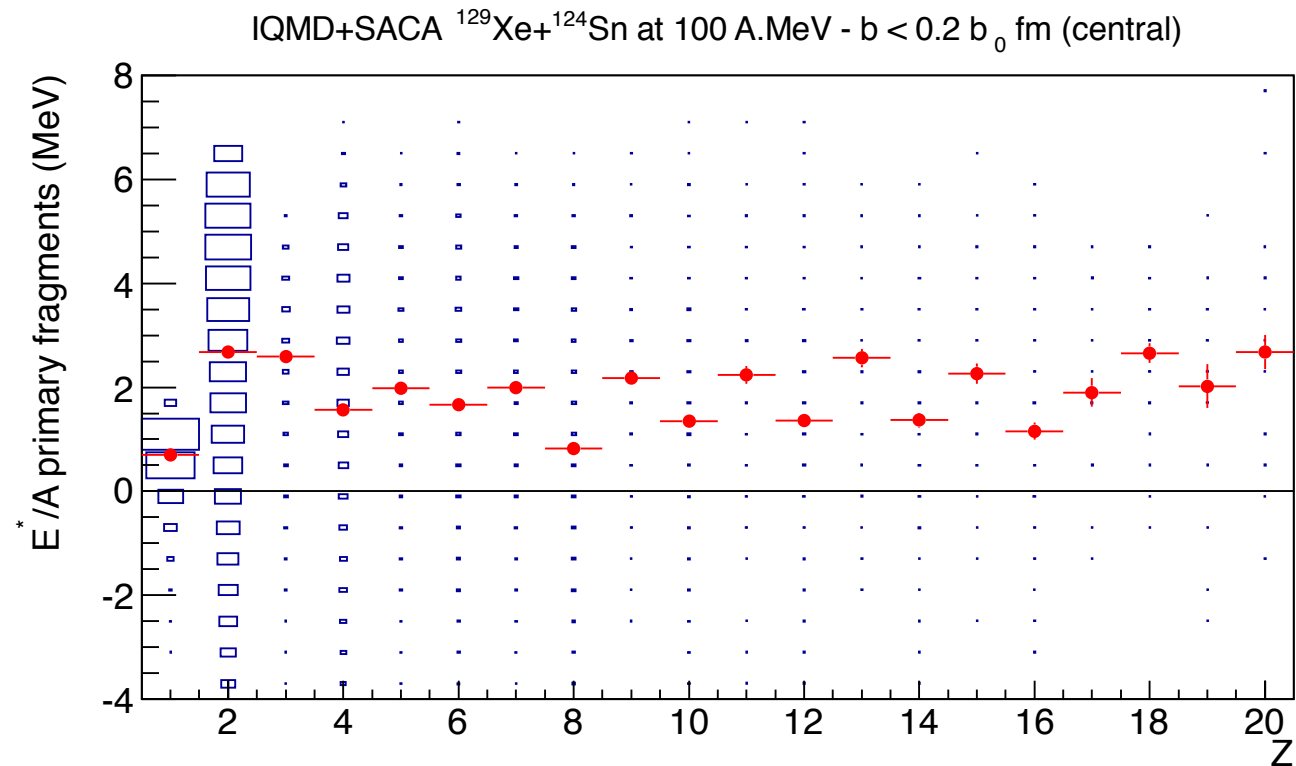
$$E^* = E_{g.s.} - E_{bind}$$





Excitation energy of the primary fragments

$$E^* = E_{\text{g.s.}} - E_{\text{bind}}$$



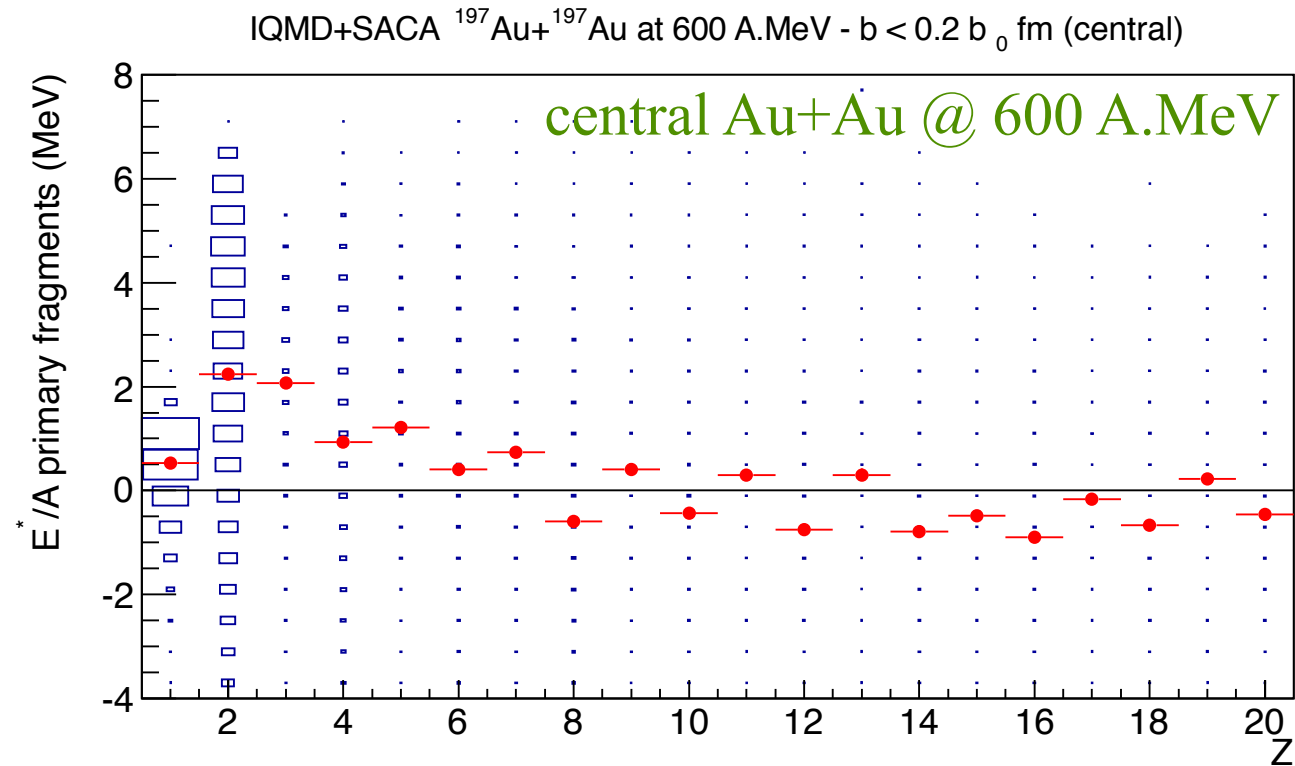
On Average, **2 A.MeV** of excitation energy. Corresponds to findings of S. Hudan et al. (INDRA collaboration), PRC 67, 064613 (2003).
=> secondary decay (GEMINI) justified here.





Excitation energy of the primary fragments

$$E^* = E_{\text{g.s.}} - E_{\text{bind}}$$



At relativistic energies, in the participant-spectator regime, heavy primary clusters are produced colder on average.

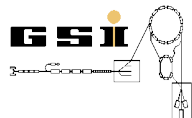




What can we learn from the isotope yields regarding the asymmetry energy?

Mean radius of primary clusters

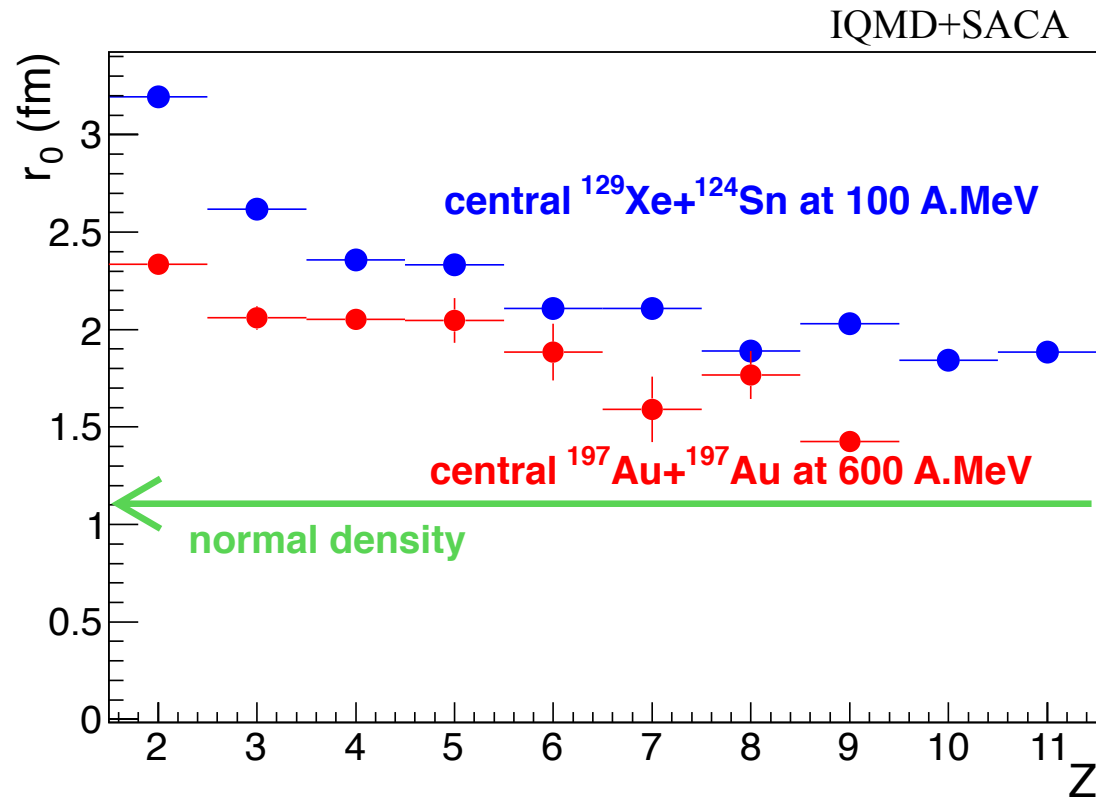
IQMD+SACA





What can we learn from the isotope yields regarding the asymmetry energy?

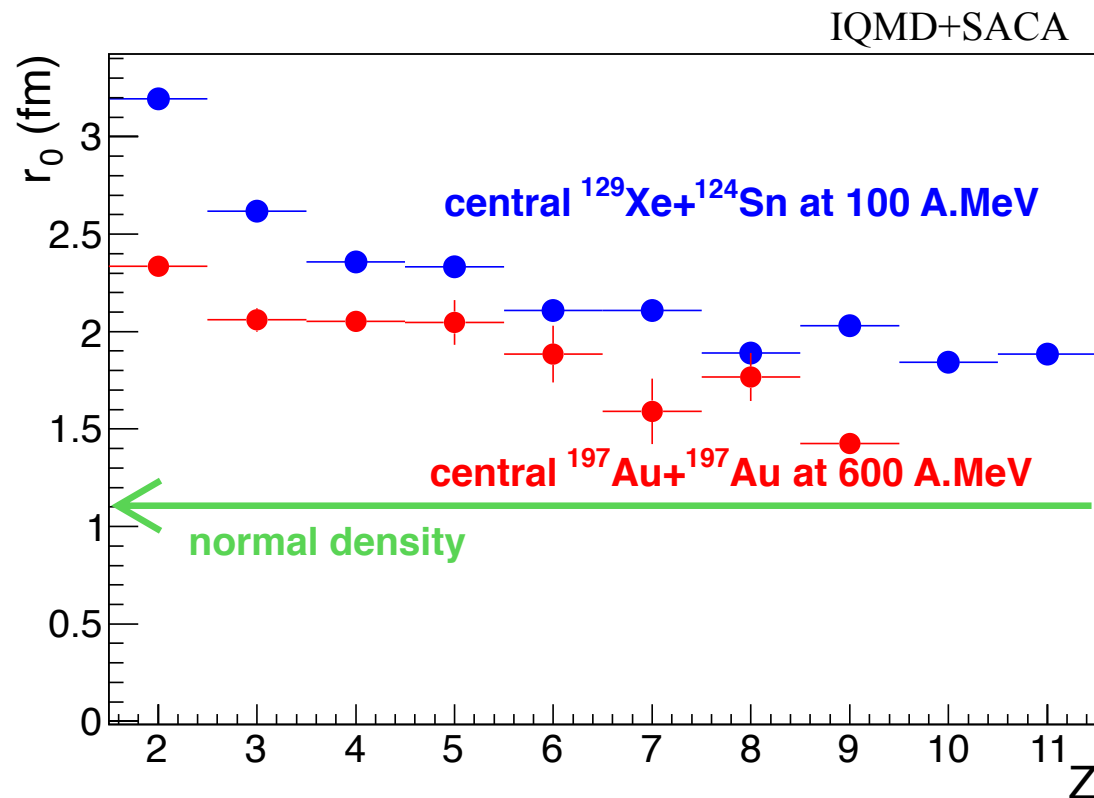
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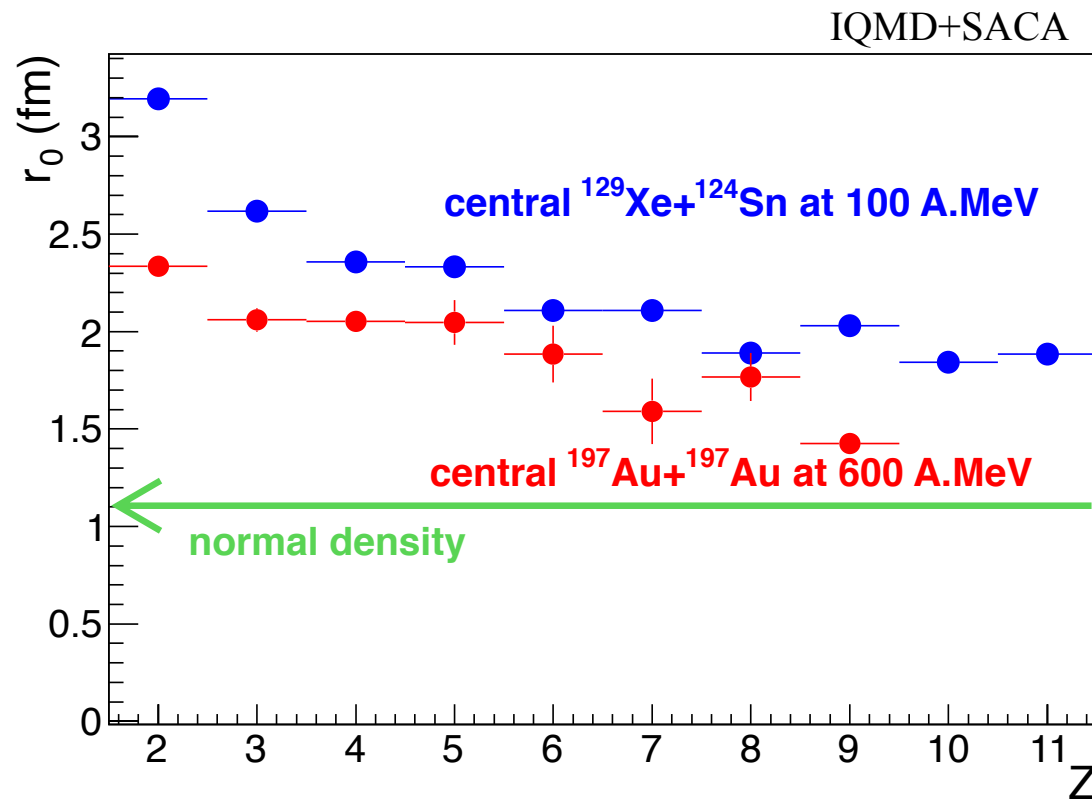


Though the medium is dense at this early stage, the dense clusters are disfavoured, because they would correspond to nucleons flowing against each other, hence with too high relative momenta to make a cluster.



What can we learn from the isotope yields regarding the asymmetry energy?

Mean radius of primary clusters



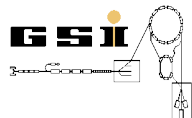
Though the medium is dense at this early stage, the dense clusters are disfavoured, because they would correspond to nucleons flowing against each other, hence with too high relative momenta to make a cluster.

=> The isotope yields can only inform on the low density dependence on the asymmetry energy.





Asymmetry energy influence versus system energy

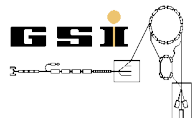




Asymmetry energy influence versus system energy

W. Reisdorf and the FOPI Collaboration

Nuclear Physics A 848 (2010) 366–427

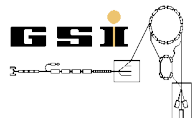
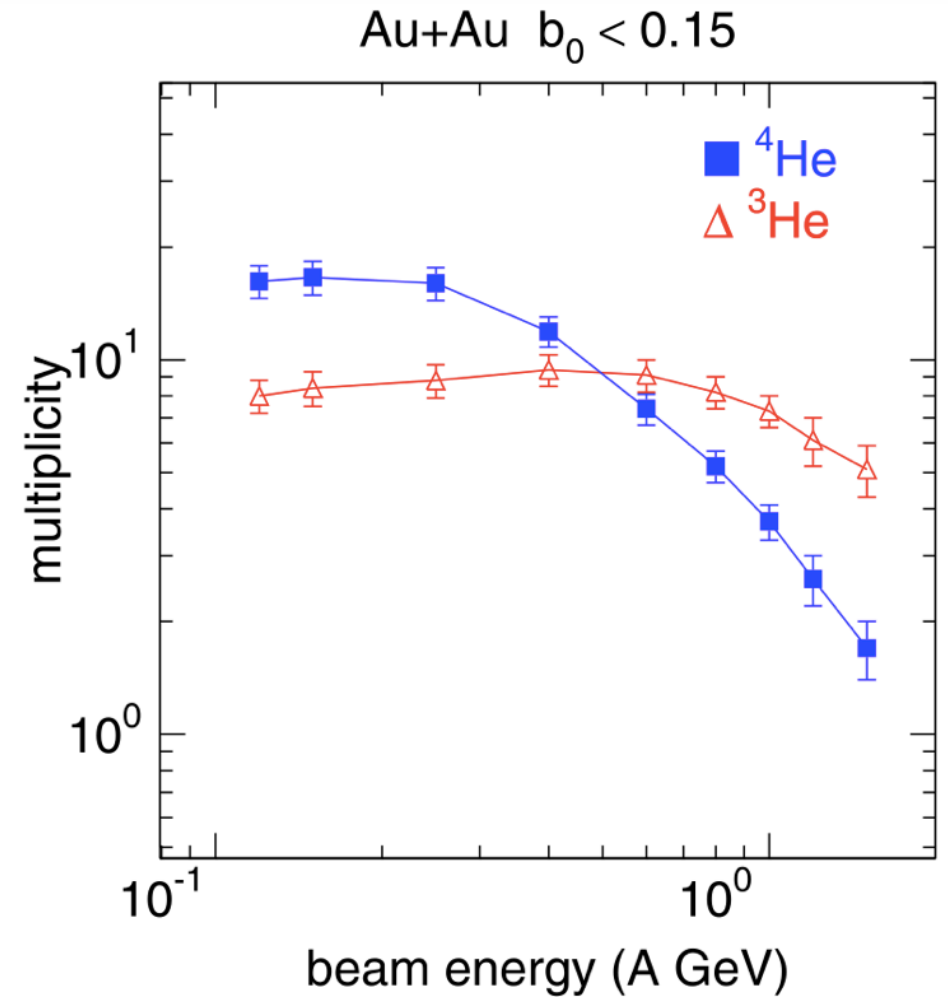




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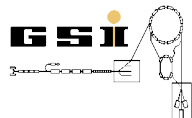
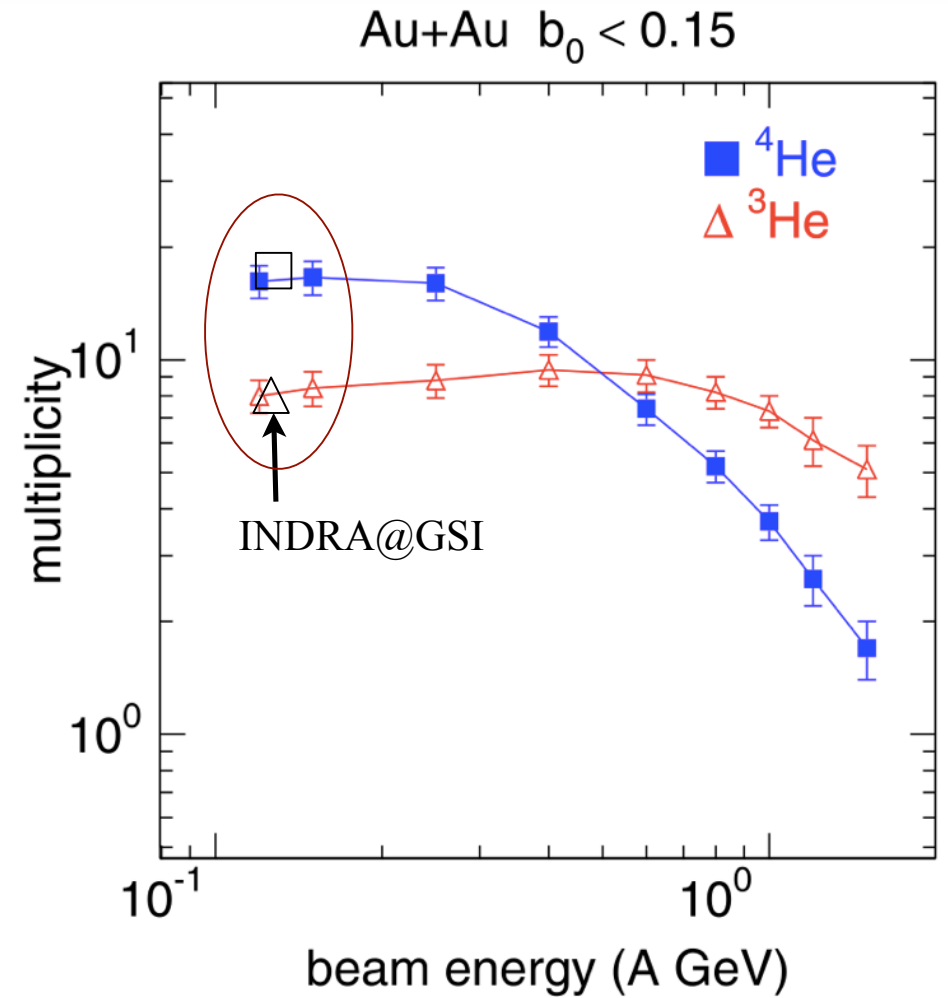




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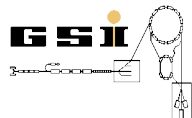
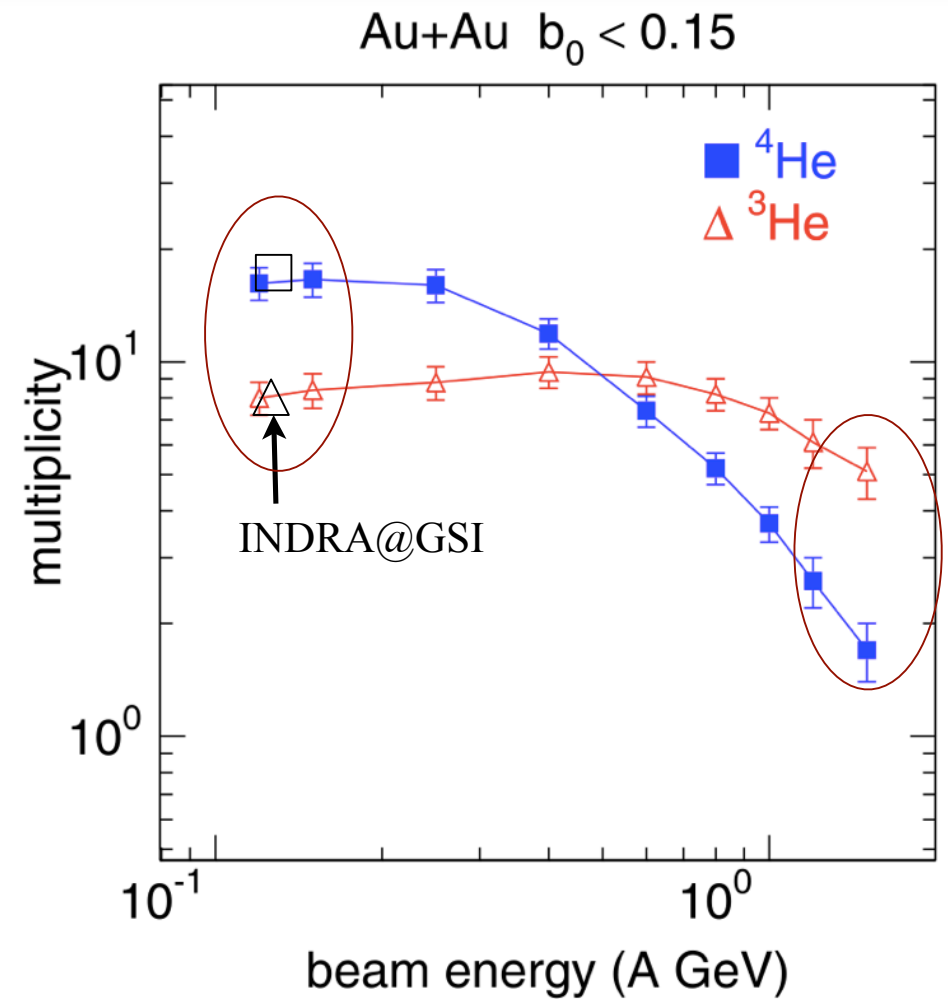




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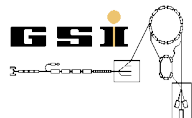
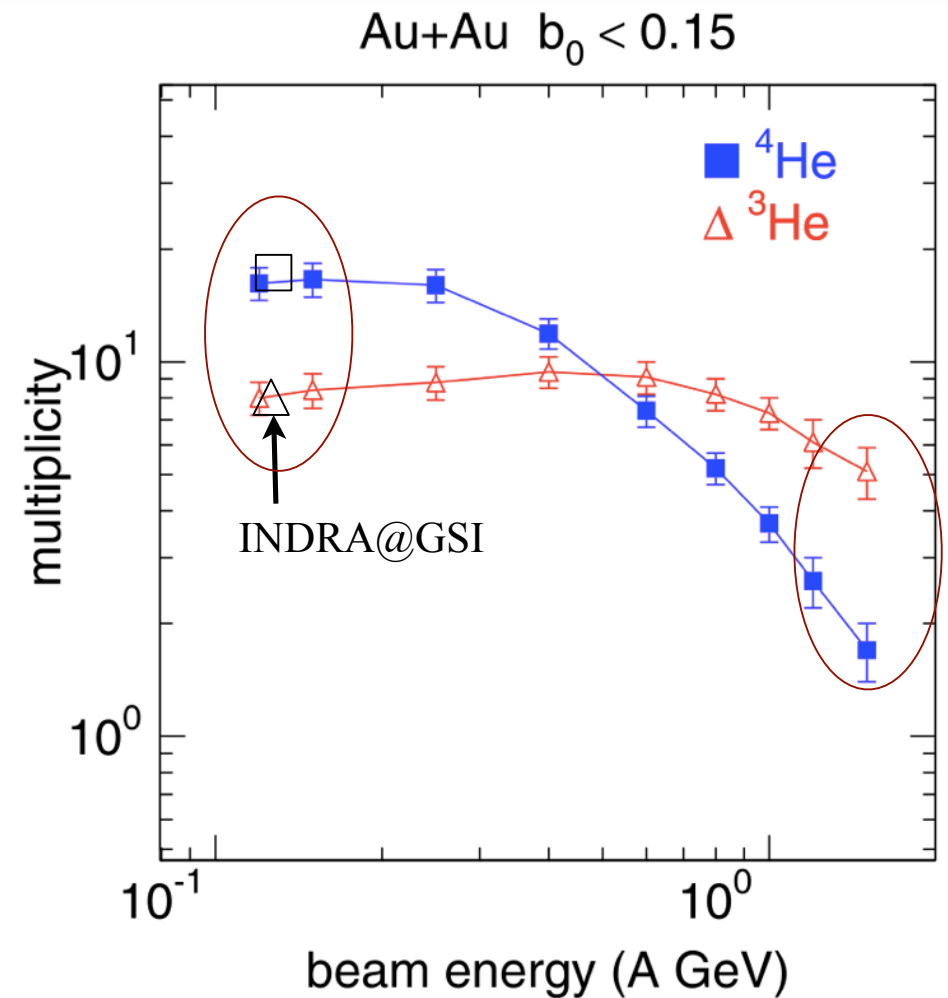


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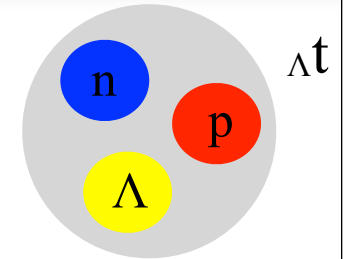
Nuclear Physics A 848 (2010) 366–427

=> At high energy, the asymmetry energy effect on clusters seems to vanish. Timescale effect? Non-linear dependence on the density?





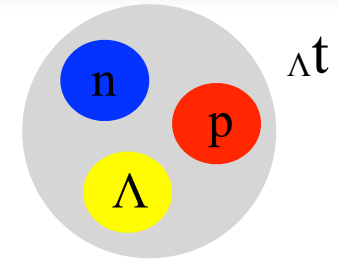
Another application of SACRA : hypernuclei production





Another application of SACA : hypernuclei production

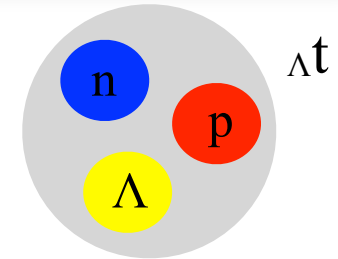
A hypernucleus is a nucleus which contains at least one hyperon (Λ (uds), ...) in addition to nucleons.





Another application of SACA : hypernuclei production

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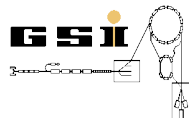


Extending SACA for clusterising hadrons with hyperons (lambdas,...) for making **hypernuclei** is straightforward:

- ❖ one replaces V_{n-p} by $V_{\Lambda-p}$ and V_{n-n} by $V_{\Lambda-n}$
- ❖ and applies with these modifications the SACA algorithm.

If in a final fragment there is a lambda, a hypernucleus should be created.

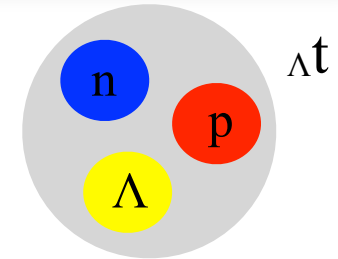
As a first approach, we have adopted $V_{\Lambda-N} = 2/3 V_{n-N}$; further refinements are possible.





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A hypernucleus is a nucleus which contains at least one hyperon (Λ (uds), ...) in addition to nucleons.



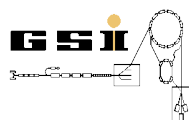
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As a first approach, we have adopted $V_{\Lambda-N} = 2/3 V_{n-N}$; further refinements are possible.

Many ways of producing lambdas: $K N \rightarrow \Lambda \pi$, $\pi^+ n \rightarrow \Lambda K^+$, $\pi^- p \rightarrow \Lambda K_0$, $p p \rightarrow \Lambda X$
 \Rightarrow influence of the EOS, in medium-properties, etc.





Another application of SACA : hypernuclei production

IQMD+SACA

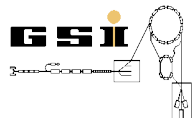
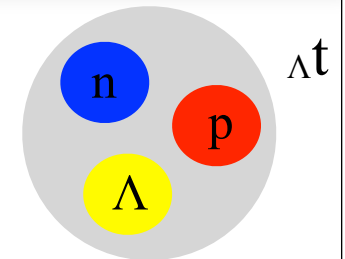
$^{58}\text{Ni}+^{58}\text{Ni}$

at

1.91 A.GeV

($b < 6$ fm) -

$t_{\text{cluster.}}=20$ fm/c

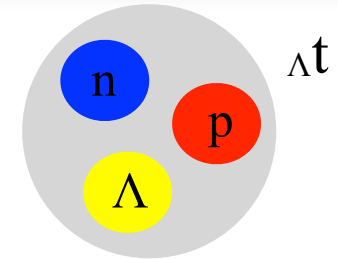
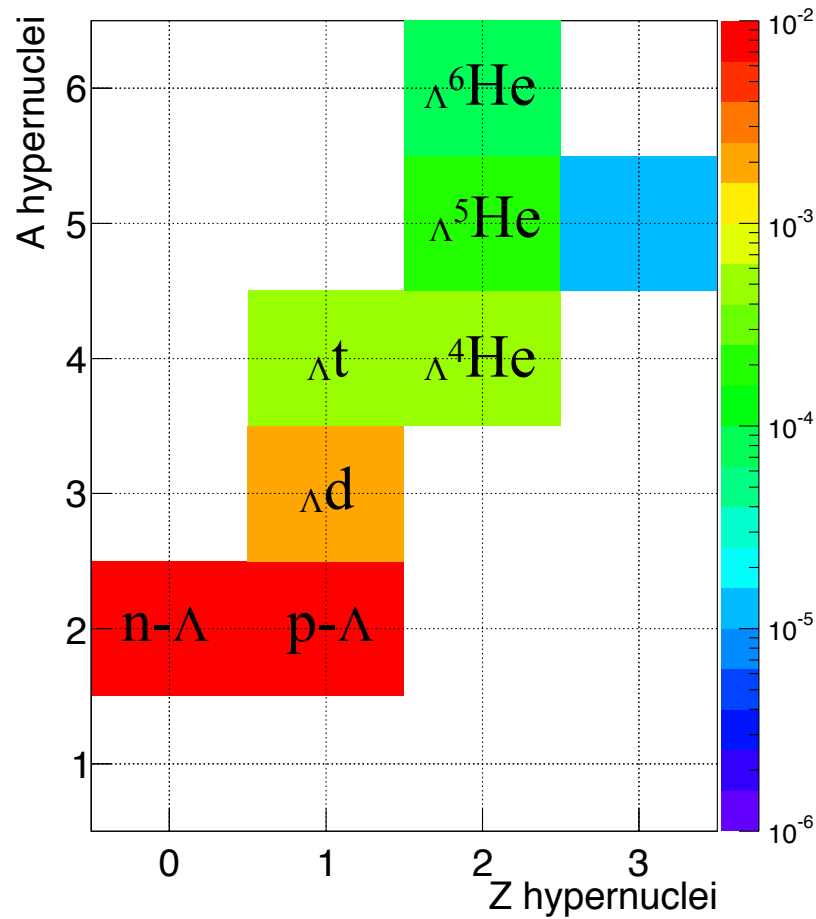




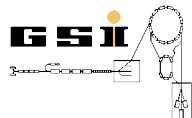
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 $t_{\text{cluster.}} = 20$ fm/c

IQMD+SACA $^{58}\text{Ni}+^{58}\text{Ni}$ at 1.93 A.GeV ($b < 6$ fm, $t_{\text{cluster.}} = 20$ fm/c) - soft no mdi, kaon pot.



Soft EOS
 no m.d.i.
 with Kaon pot.

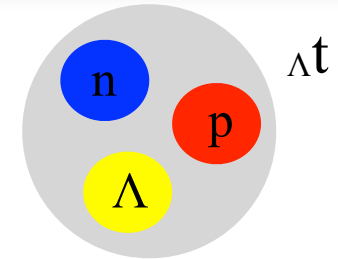
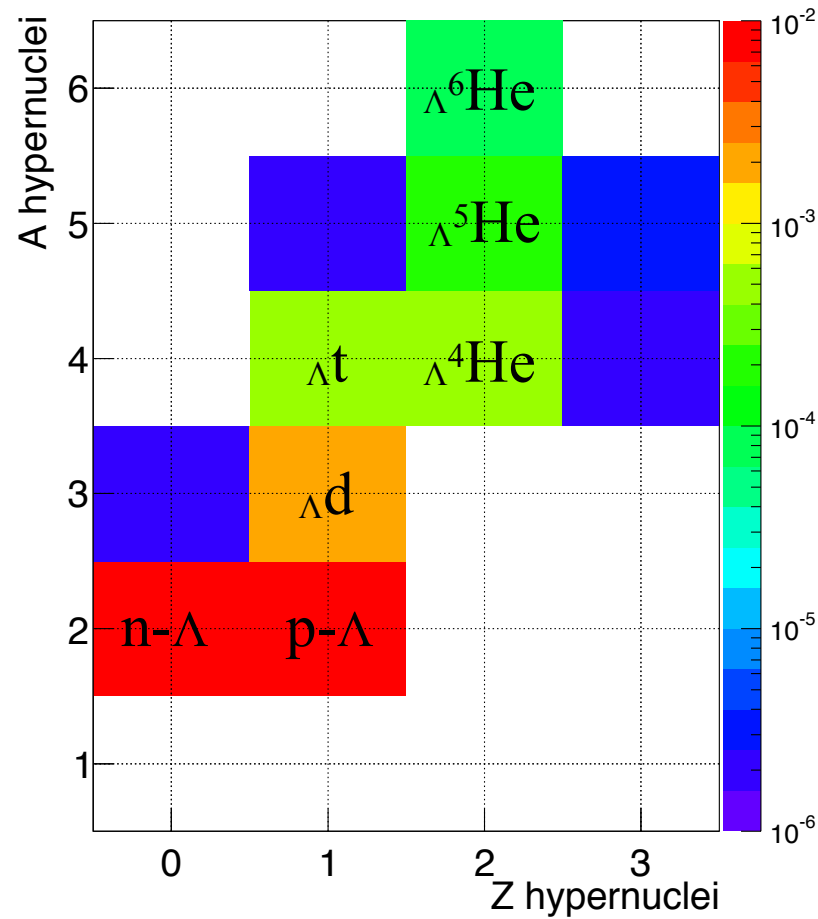




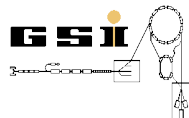
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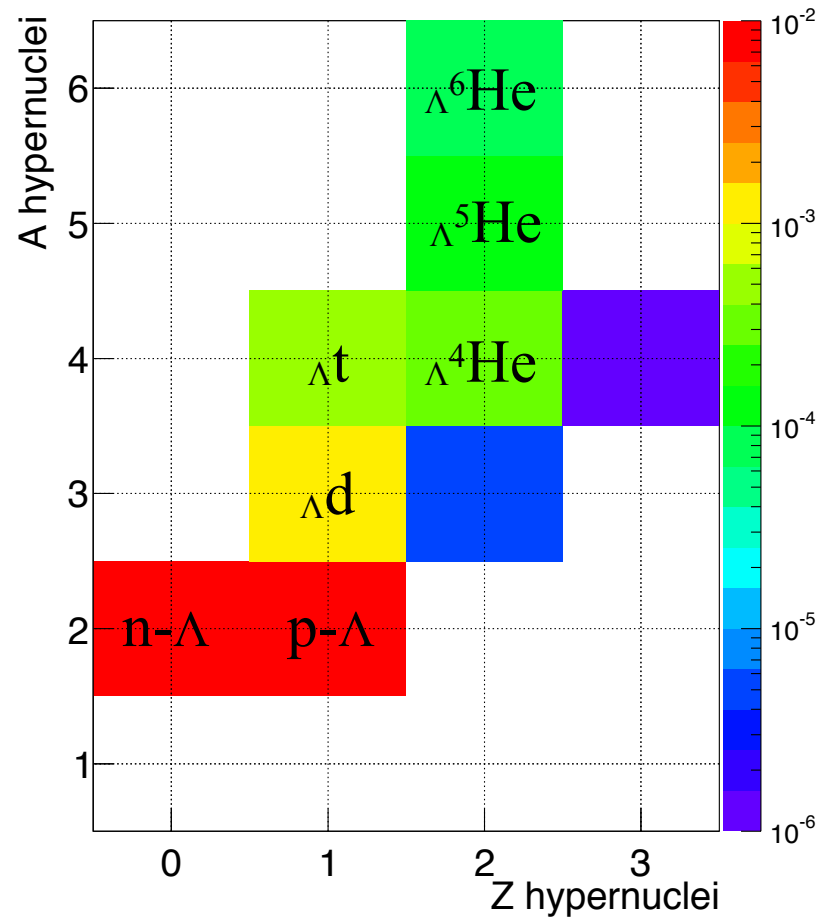
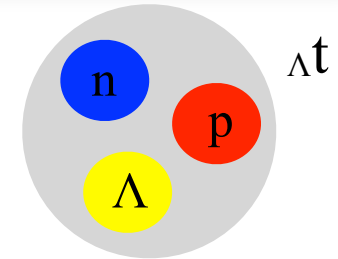




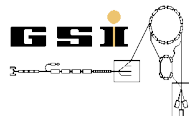
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IQMD+SACA
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IQMD+SACA $^{58}\text{Ni}+^{58}\text{Ni}$ at 1.93 A.GeV ($b < 6$ fm, $t_{\text{cluster.}} = 20$ fm/c) - soft+mdi, kaon pot.



Soft EOS
 with m.d.i.
 with Kaon pot.

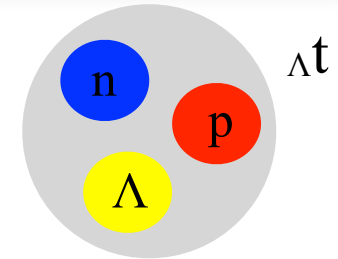
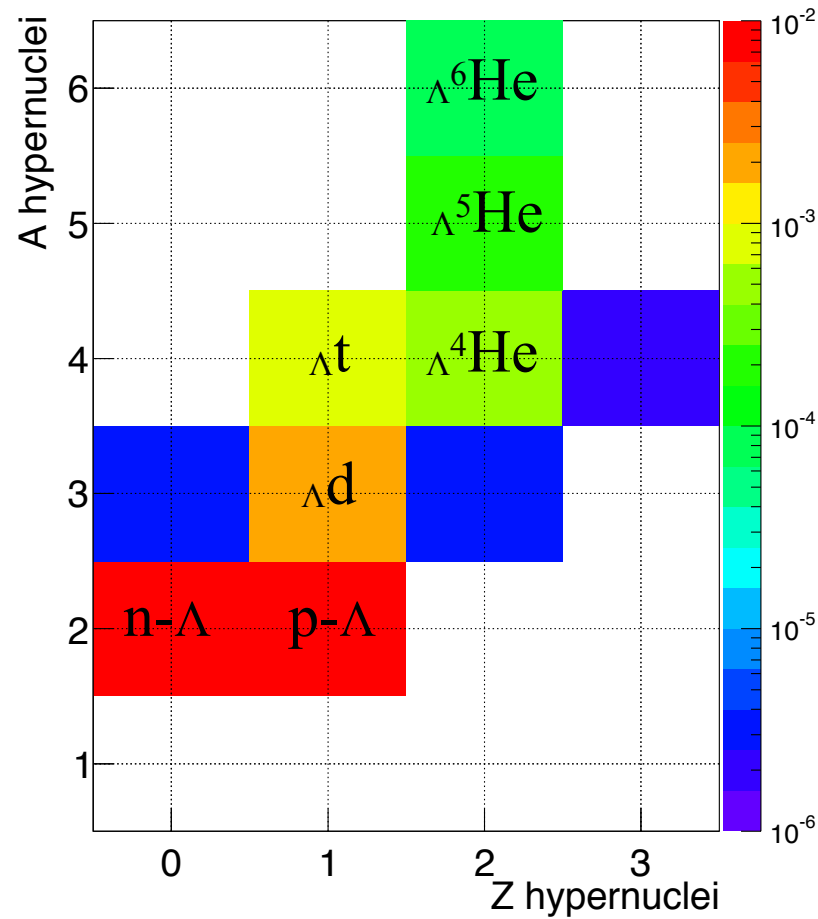




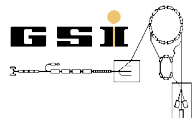
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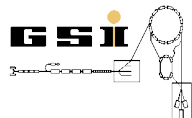
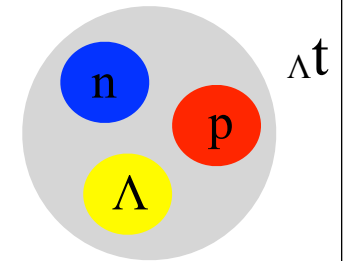


Soft EOS
with m.d.i.
no Kaon pot.



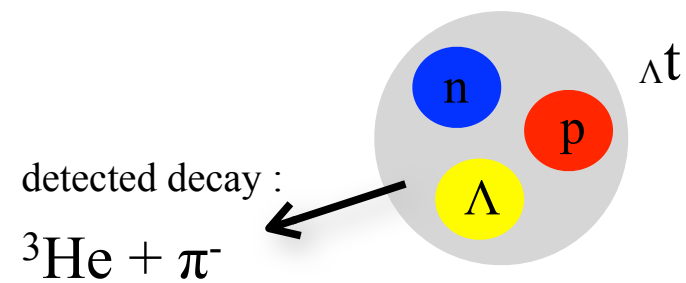


Strong phase space constraints





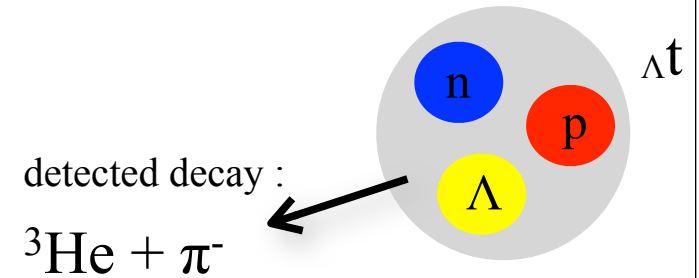
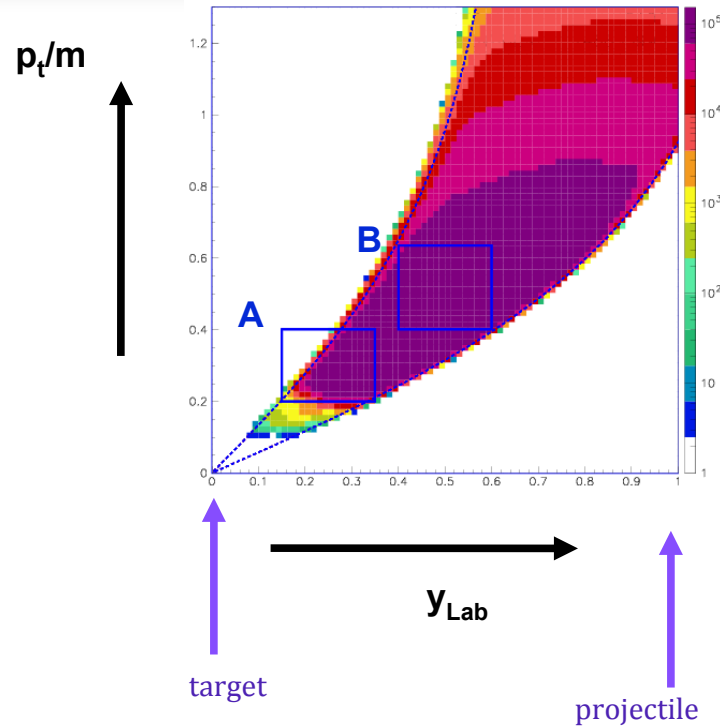
Strong phase space constraints





Strong phase space constraints

FOPI Coll.
Y. Zhang, Heidelberg



Ni+Ni @ 1.91 A.GeV

Preliminary

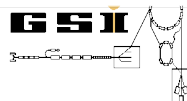
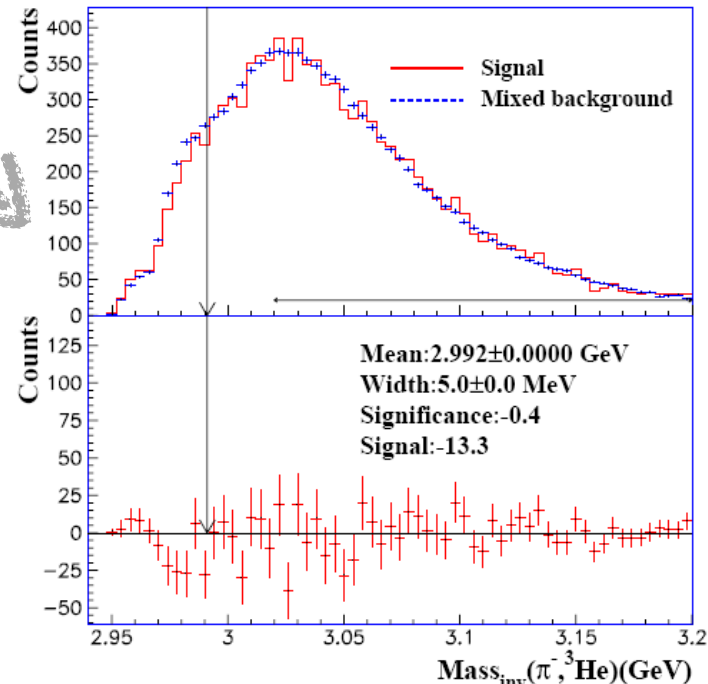
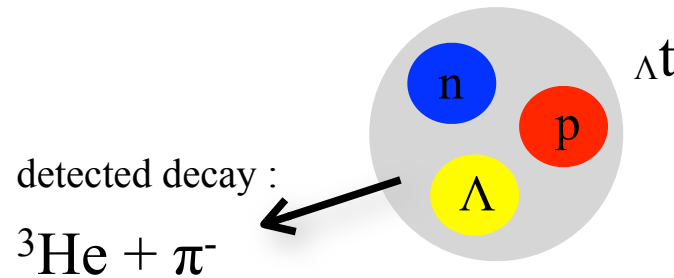
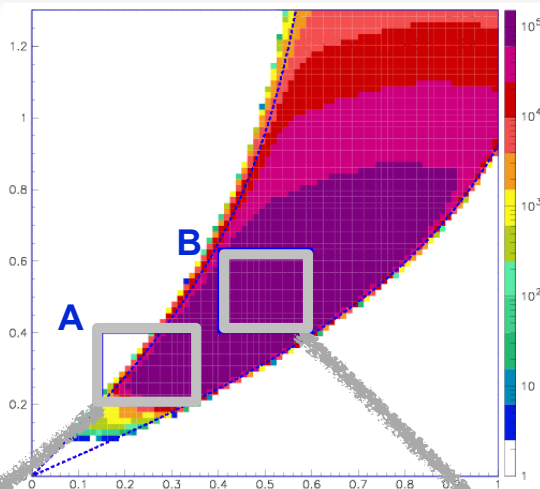
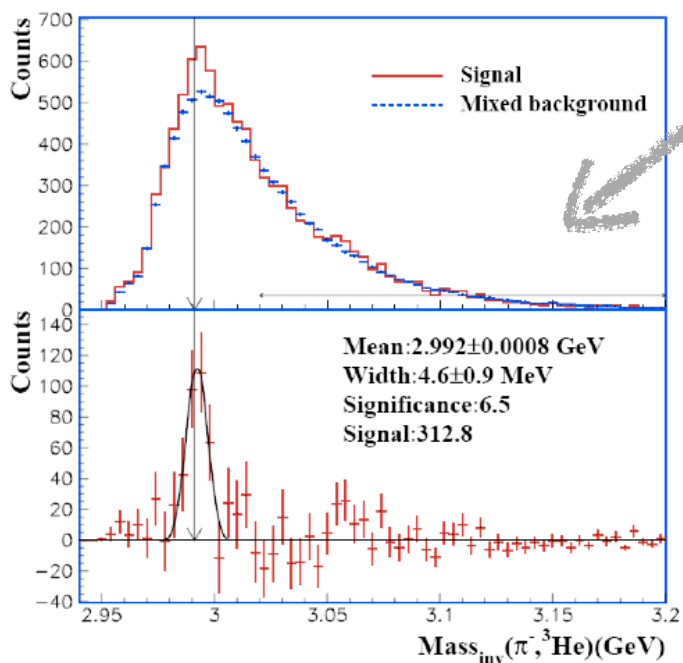




Strong phase space constraints

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Excess over combinatorial background only in region A





Strong phase space constraints

IQMD+SACA

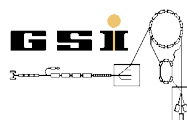
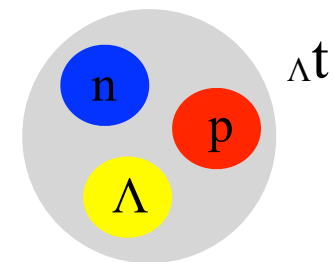
$^{58}\text{Ni}+^{58}\text{Ni}$

at

1.91 A.GeV

($b < 6$ fm) -

$t_{\text{cluster.}} = 20$ fm/c





Strong phase space constraints

IQMD+SACA

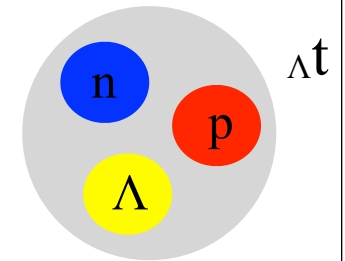
$^{58}\text{Ni}+^{58}\text{Ni}$

at

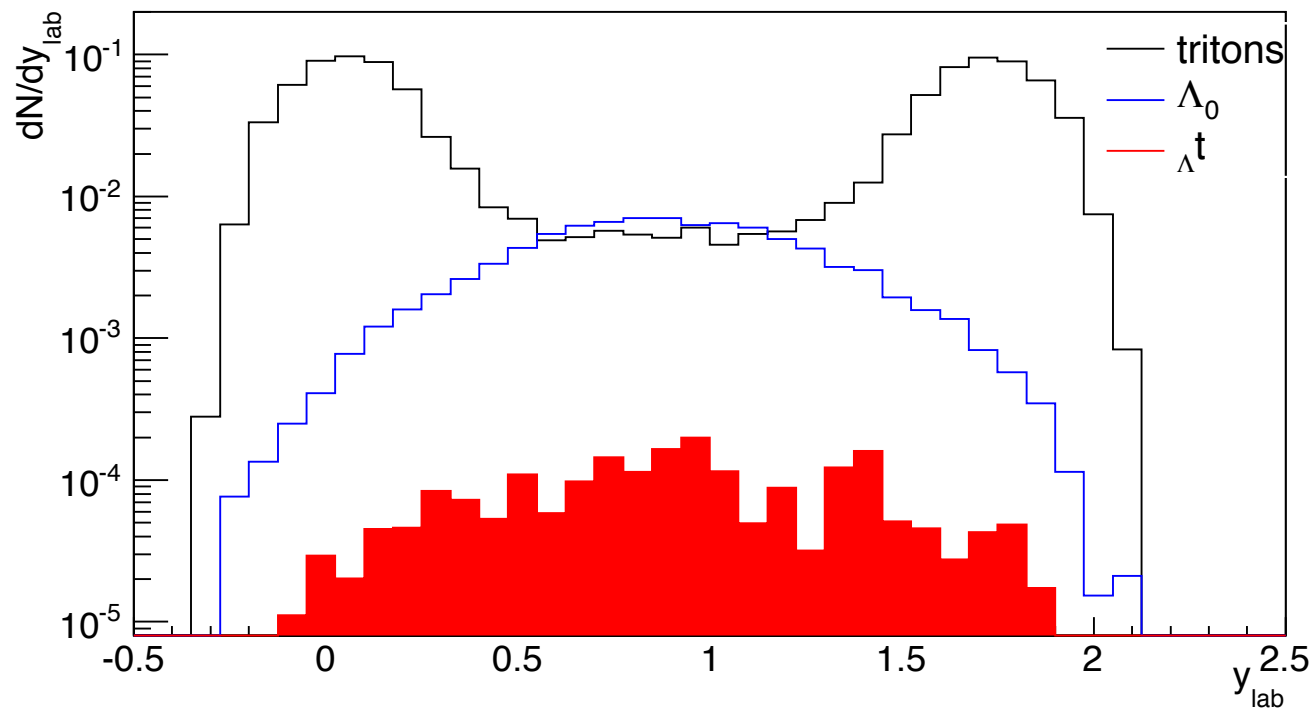
1.91 A.GeV

($b < 6$ fm) -

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Soft EOS, no m.d.i., with Kaon pot.





Strong phase space constraints

IQMD+SACA

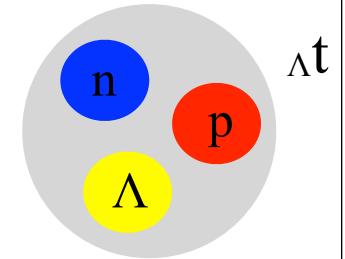
$^{58}\text{Ni}+^{58}\text{Ni}$

at

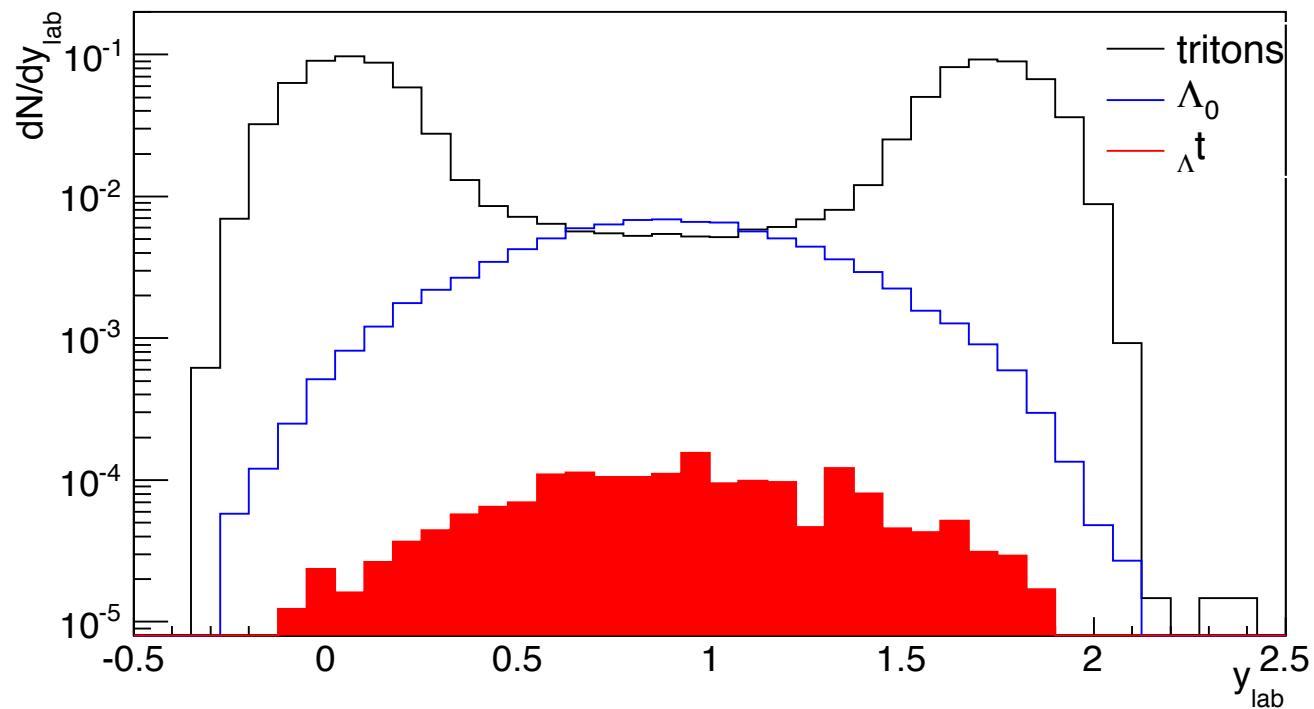
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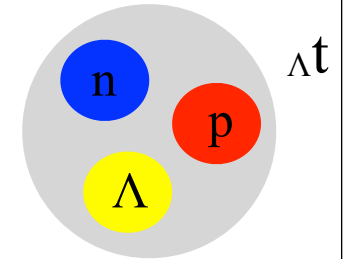
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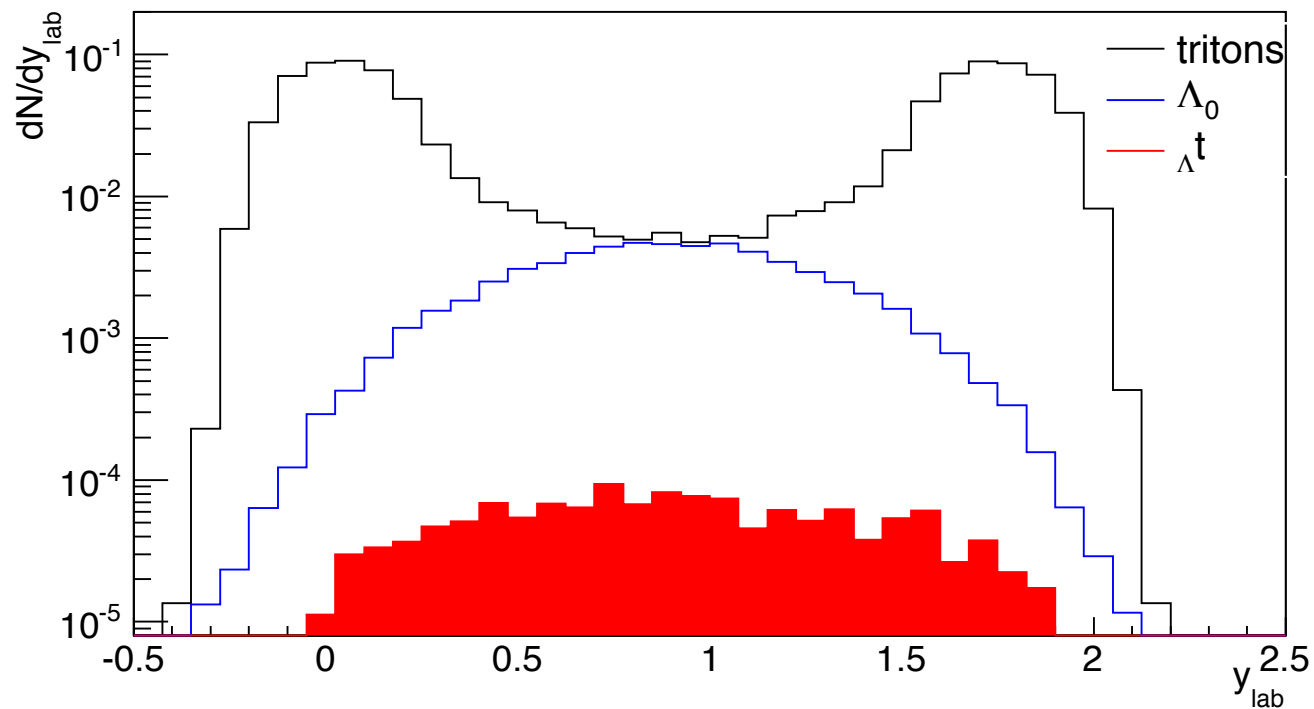
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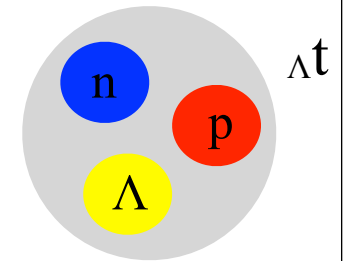
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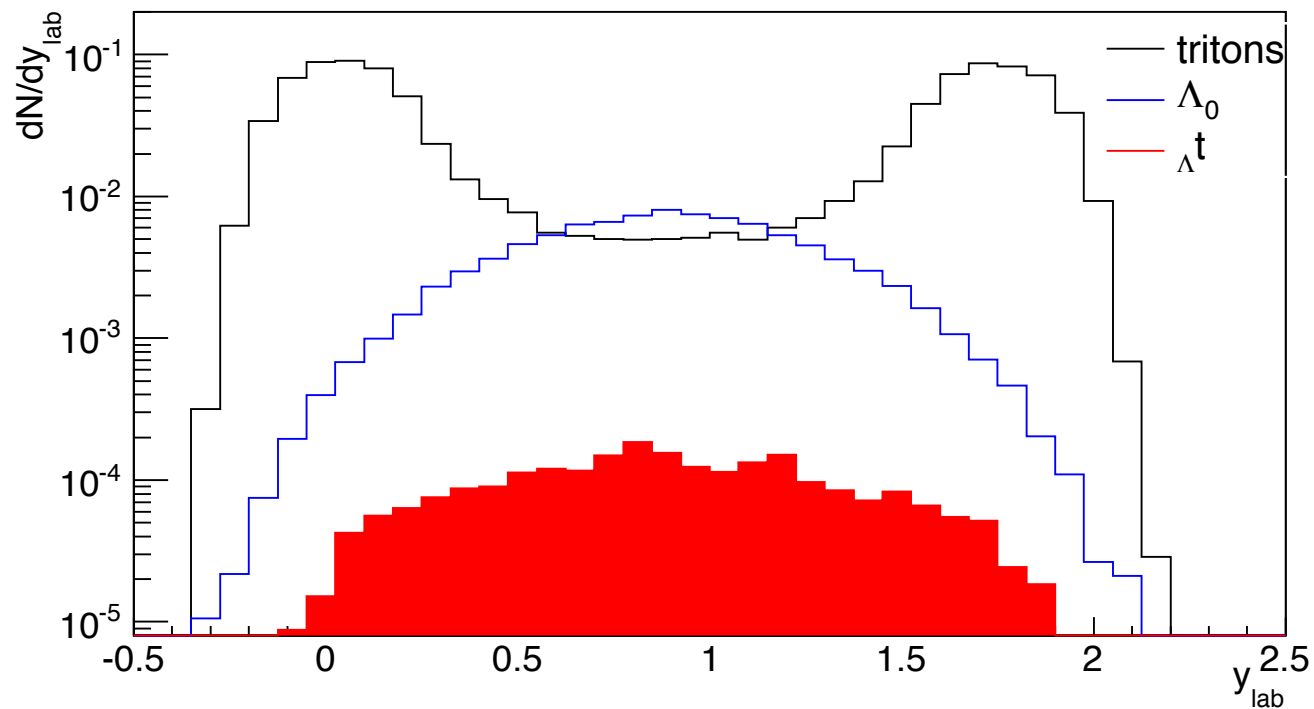
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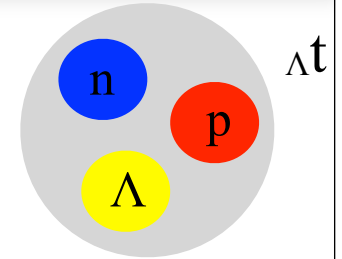
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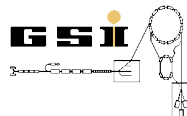
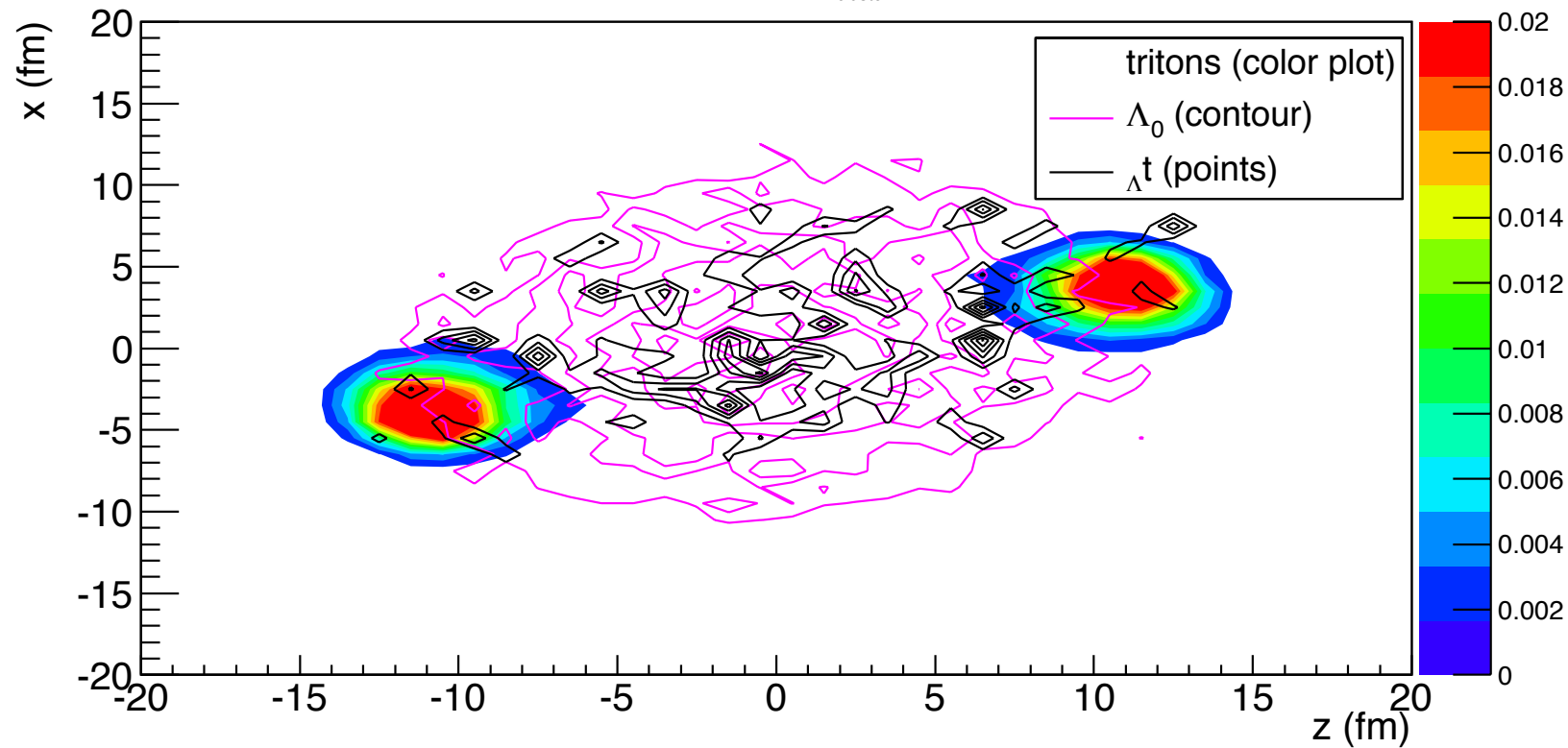
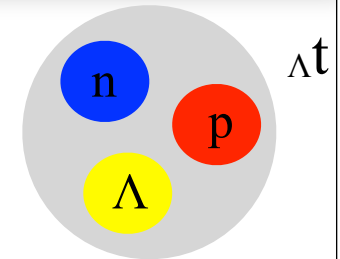


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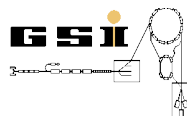
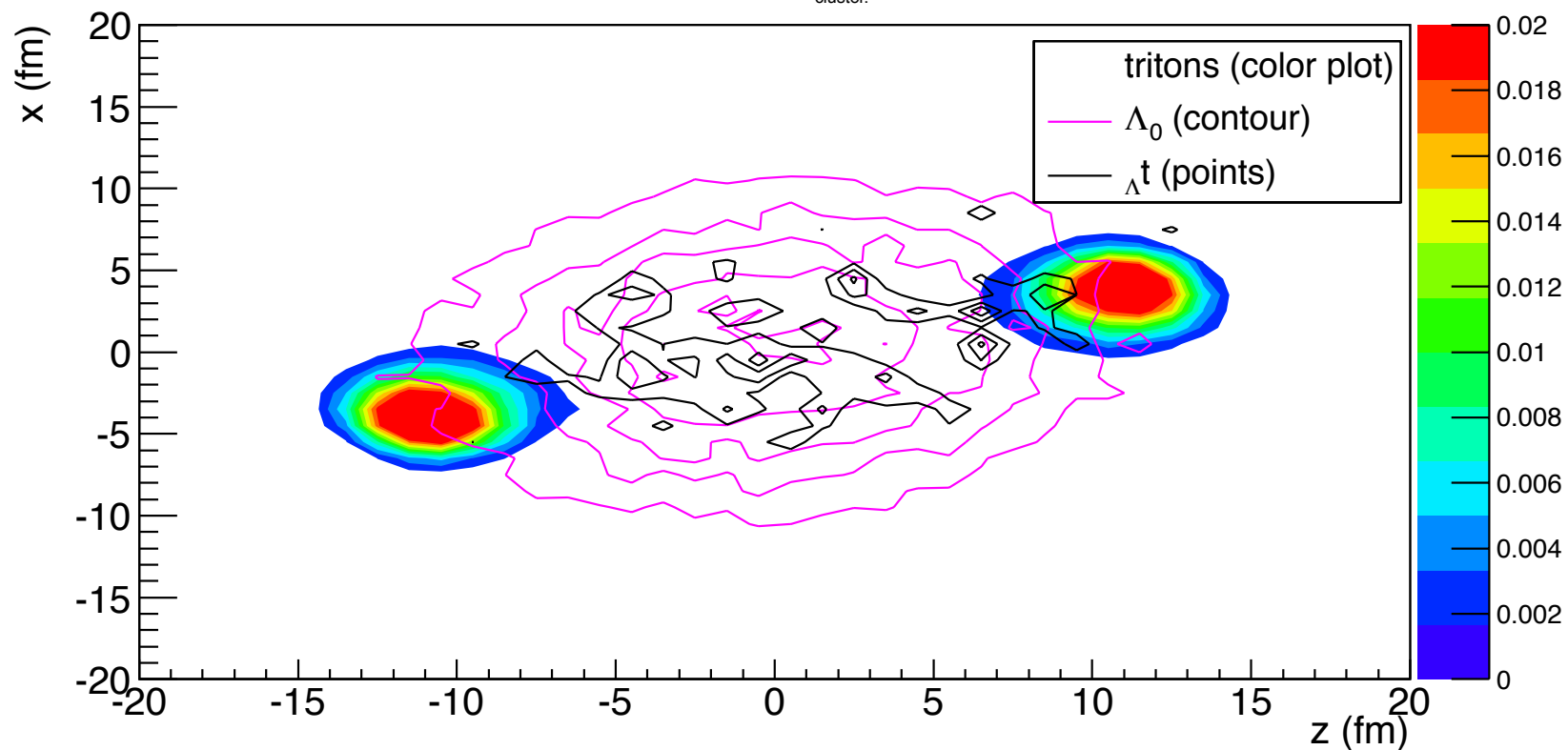
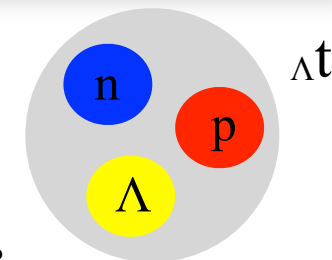


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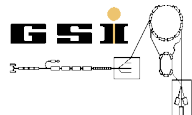
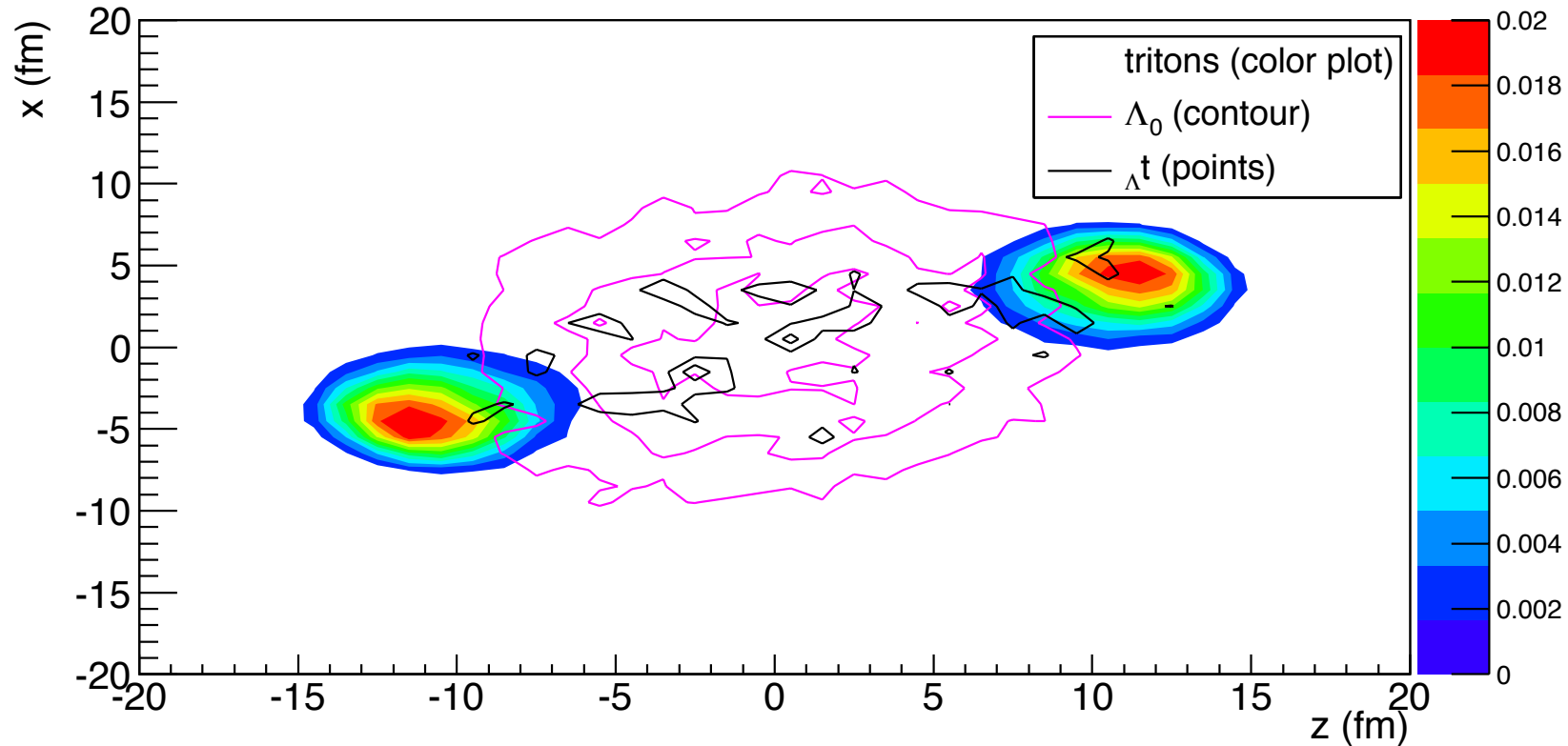
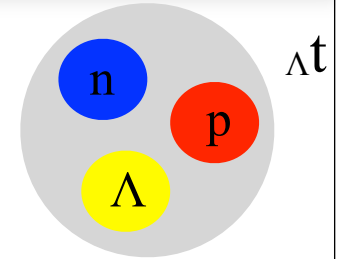


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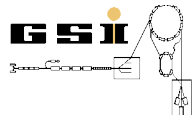
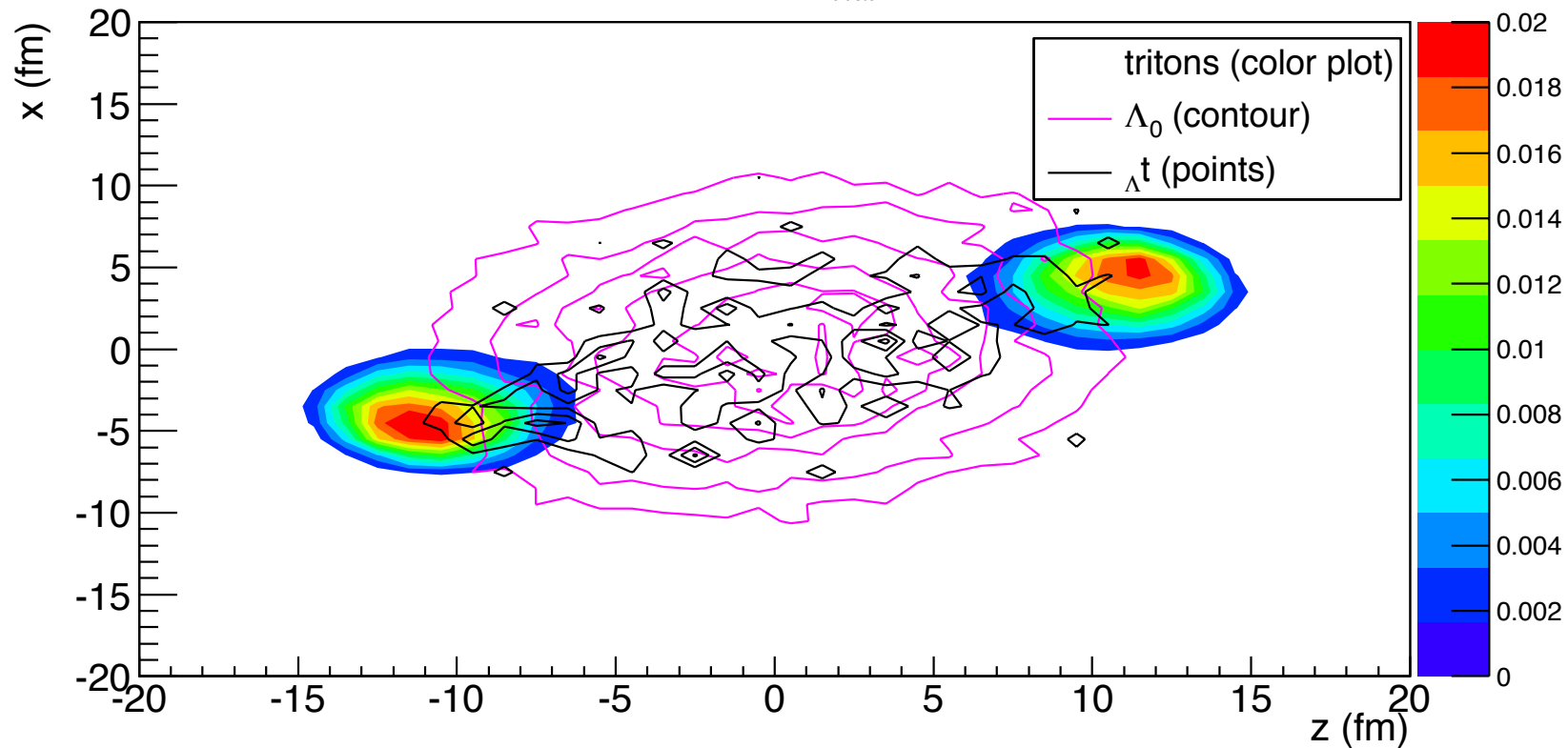
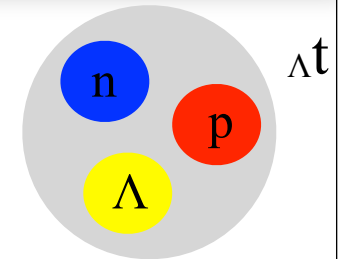


Strong phase space constraints

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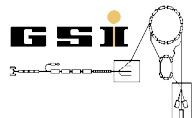
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Summary and perspectives

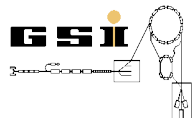




Summary and perspectives

Summary:

- ❖ Supplying SACA with a more precise description of nuclei binding energy at abnormal density allows promising, realistic predictions of absolute isotope yields, and hypernuclei.
- ❖ The asymmetry energy has a strong influence on the anisotropy (apparent stopping power) for some isotopes (^3He , ^4He , ...).
- ❖ Within this model, isotope yields cannot inform on the high density dependence of the asymmetry energy. \Rightarrow better look at n / p , K^+ / K^- , π^+ / π^- yields/flows for that purpose.





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Further developments:

After processing SACA, proceed:

- ❖ further decay of primary unstable isotopes like ^8Be , ^5He , etc., which lifetime do not allow to detect them still bound in the detectors,
- ❖ allow early $^3\text{He} + n \rightarrow ^4\text{He}$ according to its particularly high cross-section.
- ❖ secondary decay (evaporation code like GEMINI) of still excited clusters. Particularly relevant at intermediate energies (E_{beam} 100 A.MeV down to the Fermi regime)
- ❖ for hypernuclei formation, refine lambda-N potential in SACA or EOS/Kaon potential in IQMD in order to predict reasonably the measured cross-sections, and momentum distributions, which are very constraining.

