

Fundamental physics with the Pierre Auger Observatory

Dedicated to the memory of Aurelio Grillo (1945–2017),
from the L'Aquila Auger group

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Outline

1 Lorentz invariance violations

- Constraints from intergalactic UHECR propagation
- Constraints from air shower development

2 Searches for exotic particles

- Superheavy dark matter
- ANITA-like anomalous upward neutrinos
- Magnetic monopoles

UHECRs and new physics

- **Models of *nearly anything* about UHECRs** must assume that standard physics applies in **regimes where it has not otherwise been tested**.
 - e.g. interactions in centre-of-mass frames with **very large Lorentz factors** w.r.t. the lab frame ($\gtrsim 10^9$ in intergalactic propagation, $\gtrsim 10^6$ in air shower development)
 - even modest **violations of Lorentz invariance** might have **sizeable effects on UHECRs and EASs**.
- Other possible new physics we can probe using UHECR data includes **exotic particles**, for example super-heavy dark matter.
 - “Top-down” mechanisms **would produce lots of photons and/or neutrinos** in addition to nuclei.
 - Their air showers would be distinguishable from those of nuclei.

Modified dispersion relations

- **Special relativity** assumes **any process looks the same in any reference frame**, and that frames can be converted to each other via Lorentz transformations.
- Certain **candidate theories of quantum gravity** predict this is not exactly true:
 - **Lorentz invariance violation (LIV)**: Lorentz transformations stay the same, but background tensor fields pick a **privileged frame**.
 - **Deformed special relativity (DSR)**: Still no privileged frame, but **transformations between frames more complicated**
- It can be shown (Carmona et al. [PRD 86 \(2012\) 084032](#)) the effects are those of **modified dispersion relations (MDR)**,

$$E_i^2 = m_i^2 + (1 + \delta_i^{(0)})p_i^2 + \delta_i^{(1)}p_i^3 + \delta_i^{(2)}p_i^4 + \delta_i^{(3)}p_i^5 + \dots,$$

where i is the particle type (standard dispersion relations: $\delta_i^{(n)} = 0$ for all n).

- Alternatively, dimensionless $\eta_i^{(n)} = \delta_i^{(n)} M_{\text{pl}}^n$ or energies $E_{\text{QG},i}^{(n)} = 1/\sqrt[n]{\delta_i^{(n)}}$ can be used

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Effects on UHECR propagation

- MDRs can allow processes which would otherwise be kinematically forbidden, or *vice versa*, or alter the rates of processes which would be allowed either way.
 - Extensive review of possible quantum gravity effects in astroparticle physics (Addazi et al. [PrPNP 125 \(2022\) 103948](#)) and follow-up white paper (Alves Batista et al. [CQGra 42 \(2025\) 032001](#))
 - Major contributions from members of the Torino and L'Aquila Auger groups
 - [Publicly available database](#) of experimental bounds on quantum gravity effects
 - One of the main contributors: Caterina Trimarelli, then a PhD student in the L'Aquila Auger group
- The first example I'm going to discuss: superluminal LIV of hadrons ($\delta_{\text{had}}^{(n)} > 0$) could avert GZK interactions (Auger collab. [JCAP 01 \(2022\) 023](#)).
 - The situation is not as simple as you might have heard ...

“We observe cosmic rays with $E > E_{\text{GZK}}$, hence their propagation must be violating Lorentz invariance.”

(People said this in the AGASA days.)

“We observe many fewer cosmic rays with $E > E_{\text{GZK}}$ than the extrapolation from lower energies, hence their propagation cannot be violating Lorentz invariance too much.”

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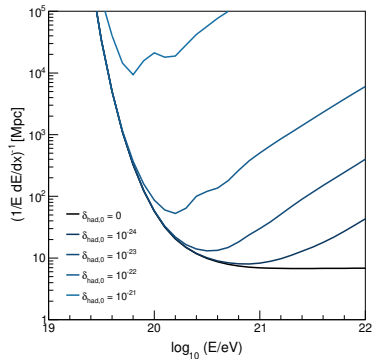
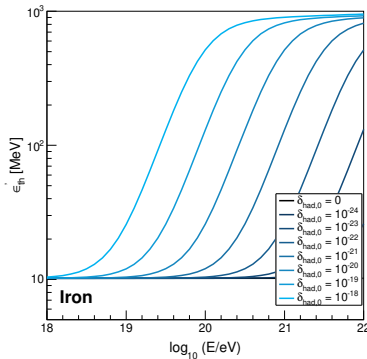
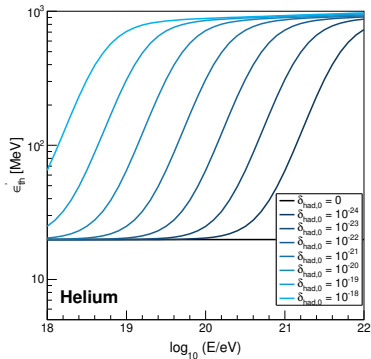
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LIV in the hadronic sector

Apologies for the font sizes in the plots

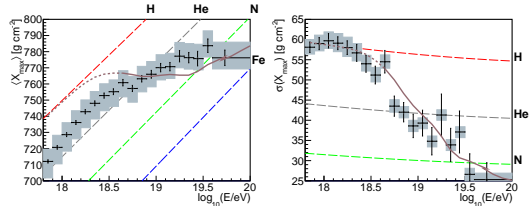
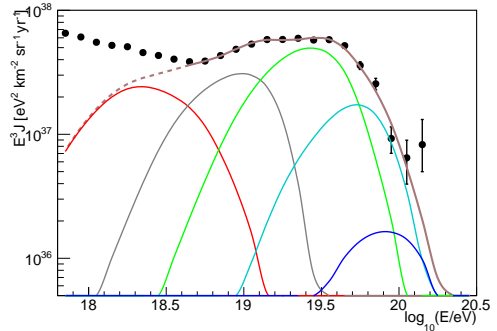
- Superluminal LIV in the hadronic sector would increase the threshold for pion production, allowing nuclei to propagate much farther.
- We're assuming the same LIV for all hadrons ($\delta_{\text{had}}^{(n)} = \delta_p^{(n)} = \frac{1}{2}\delta_\pi^{(n)} = A^n \delta_A^{(n)}$).



$$\delta_{\text{had}}^{(0)} = 0$$

- We can still get an acceptable fit with $\delta_{\text{had}}^{(0)} = 10^{-20}$.
- On the other hand, we can exclude $\delta_{\text{had}}^{(0)} \geq 10^{-19}$,
 $\delta_{\text{had}}^{(1)} \geq 10^{-38} \text{ eV}^{-1}$,
 $\delta_{\text{had}}^{(2)} \geq 10^{-57} \text{ eV}^{-2}$ at the 5σ level.

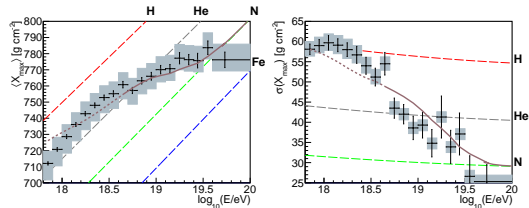
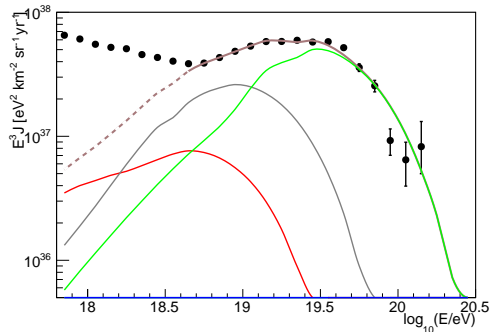
Analysis by the São Paulo and L'Aquila Auger groups
 (Auger collab. [JCAP 01 \(2022\) 023](#)).



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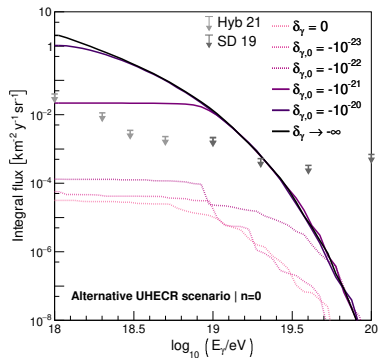
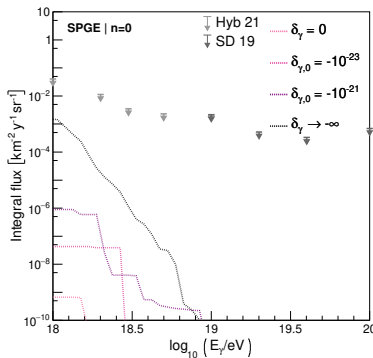
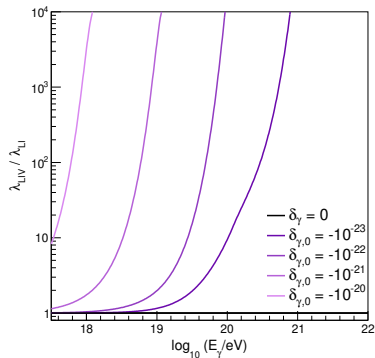
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LIV in the photon sector ([ibid.](#))

- Subluminal LIV of photons ($\delta_\gamma^{(n)} < 0$) could prevent $\gamma_{\text{UHE}} + \gamma_{\text{bg}} \rightarrow e^+ + e^-$.
- In the best fit to the current Auger data, not many photons are produced in the first place (due to heavy UHECR composition).
- If there were a proton admixture, we'd be able to set limits on photon LIV.



Examples of other studies

Using Auger data, but not authored by the Auger collaboration

- L.A. Anchordoqui et al., “New test of Lorentz symmetry using ultrahigh-energy cosmic rays”, [*Phys. Rev. D* **97** \(2018\) 043010](#)
- M.D.C.Torri et al., “Predictions of Ultra-High Energy Cosmic Ray Propagation in the Context of Homogeneously Modified Special Relativity”, [*Symm.* **12** \(2020\) 1961](#)
- Z. Berezhiani, “Extreme energy cosmic rays: a quest for new physics?”, [talk at UHECR 2022](#)
- L.A. Anchordoqui, “Dark dimension, the swampland, and the origin of cosmic rays beyond the Greisen-Zatsepin-Kuzmin barrier”, [*Phys. Rev. D* **106** \(2022\) 116022](#)
- N.T. Noble et al., “Probing the dark dimension with Auger data”, [*Phys. Dark Univ.* **42** \(2023\) 101278](#)

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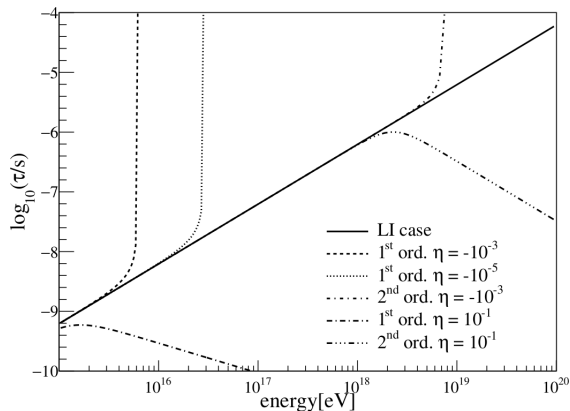
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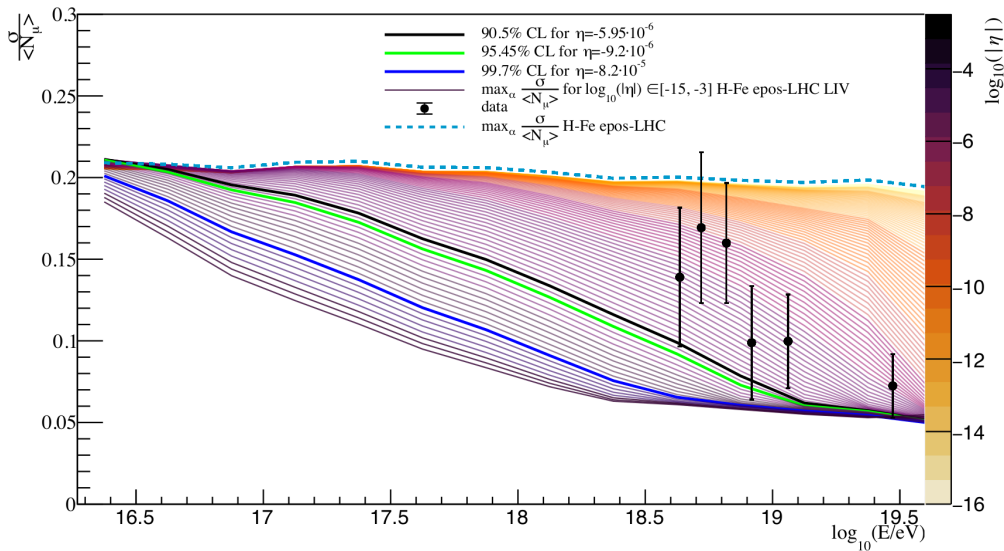
Neutral pion LIV in air showers

- **Subluminal LIV of neutral pions** would **delay or stop their decay**, allowing them to continue the hadronic cascade rather than starting electromagnetic subshowers.
- This would **decrease shower-to-shower fluctuations** in the number of muons.
(It would also increase the average, but that's too model-dependent to be useful.)

Analysis by the L'Aquila Auger group (C. Trimarelli for the Auger collab. [PoS 395 \(ICRC2021\) 340](#)).

(Superluminal LIV would speed up their decay, but it's already much faster than all other relevant timescales in the LI case anyway.)



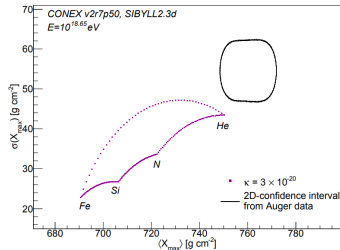
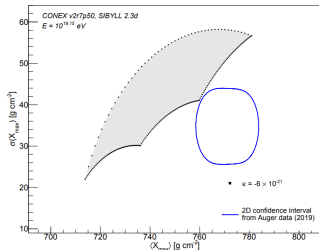


- We can use this to exclude $\eta_\pi^{(1)} < -6 \times 10^{-6}$ at the 90.5% CL ([ibid.](#)).
- Full-author-list collaboration paper in preparation by the L'Aquila Auger group

Limits on isotropic nonbirifrengent LIV in the photon sector

Using Auger data, but not authored by the full Auger collaboration

- **Superluminal photon LIV** would allow **photon decay** into electron–positron pairs
→ **shallower air showers** (Duenkel et al. [PRD 104 \(2021\) 015010](#))
 - Using Auger X_{\max} data, we can exclude $\kappa \leq -6 \times 10^{-21}$ at the 98% CL.
- **Subluminal photon LIV** would allow charged particles to emit **vacuum Cherenkov radiation** → **no nuclei above a certain energy** (near-instantaneous losses);
shallower air showers (Duenkel et al. [PRD 107 \(2023\) 083004](#))
 - Combining Auger energy and X_{\max} data, we exclude $\kappa \geq 3 \times 10^{-20}$ at the 98% CL.



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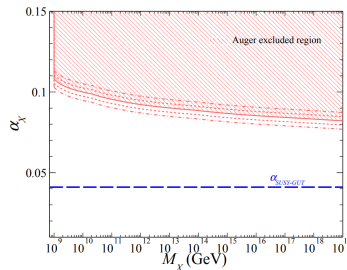
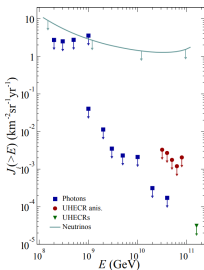
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- Superheavy dark matter particles would be subject to instanton-induced decays.
- The decay products of SHDM particles in the Galactic halos — photons, (anti-)neutrons, (anti-)protons — could be detected in Auger data.
- Assuming the observed dark matter density is entirely due to SHDM, for each possible mass we can set limits on the reduced coupling constant.

(Auger collab. [PRL 130 \(2023\) 061001](#)) (Auger collab. [PRD 107 \(2023\) 042002](#))



- In the case of SHDM coupled to sterile neutrinos, we can also set limits on their mixing angle with active neutrinos (Auger collab. [PRD 109 \(2024\) L081101](#))

Major contribution from Roberto Aloisio from the L'Aquila Auger group

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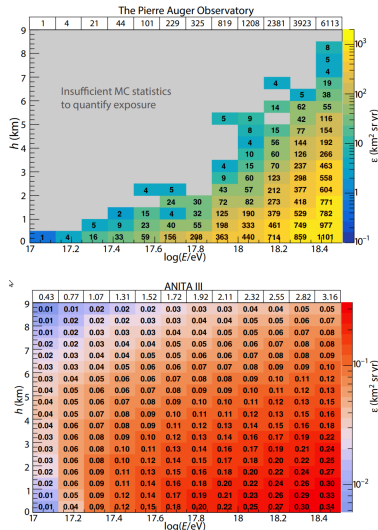
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Search for upgoing events in Auger (Auger collab. [PRL 134 \(2025\) 121003](#))

- ANITA detected events looking like upward-going ν_τ -initiated air showers, at angles that should not be possible according to the Standard Model.
- Auger FD data were searched for similar showers.
- After filtering misreconstructed downgoing showers and calibration laser shots out, **one event** survived (compatible with the expected background).
- This fact and the exposure (computed via MC) were used to set limits to upgoing τ leptons.

Major contributions from the Lecce and L'Aquila Auger groups



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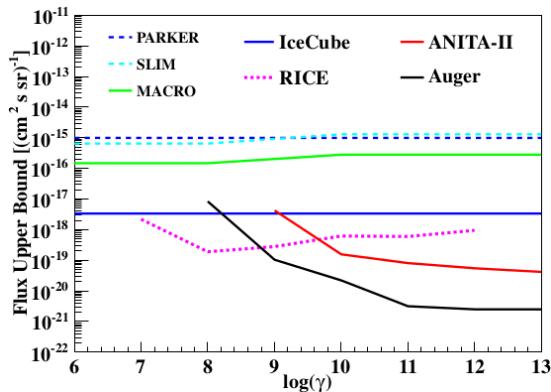
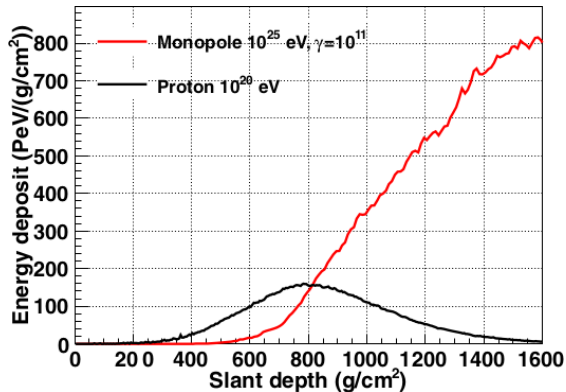
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Search for ultrarelativistic magnetic monopoles

(Auger collab. [PRD 94 \(2016\) 082002](#))

- Ultrarelativistic monopoles would produce very bright, very deep air showers.
- No such showers were found in Auger data.
- This was used to set limits on the flux of magnetic monopoles.





That's all Folks!