Fundamental Physics with Exotic Atoms

Overview of Technological Challenges in Exotic Atom Precision Experiments

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Precision experiments with exotic atoms offer unique insights into fundamental physics, from testing quantum electrodynamics (QED) to probing the strong interaction (QCD), high precision test of BSQED

Achieving the necessary precision is fraught with significant technological challenges. These challenges primarily stem from the nature of exotic atoms themselves: their short lifetimes, the weak signals they produce, and the demanding experimental conditions required for their creation and study.



The key technological challenges:

Weak Signals and Low Event Rates:

- **Low Production Yields:** Creating exotic atoms often involves complex processes
- **Short Lifetimes:** time available for them to form exotic atoms and undergo de-excitation
- Dilute Targets: To minimize unwanted interactions and effects like the Stark effect, increase the stopping power
- Limited Beam Intensity:

Background Interference:

- **High Background Radiation:** accelerator facilities where a multitude of other particles and radiation are present
- Unwanted X-ray Emission: environment (e.g., from conventional atomic transitions in target materials, cosmic rays, or facility-related radiation).
- Charged Particle Contamination: Charged particles from the beam or secondary interactions
- **Spurious Signals:** Detector noise, electronic interference, and other spurious signals

Limitations in Detector Technology:

- Energy Resolution: detectors with high energy resolution are required.
- o Detection Efficiency: detectors must have very high quantum efficiency
- Timing Resolution: Good timing resolution fast trigger and veto capabilities
- Background Rejection Capabilities: pulse shape discrimination, spatial reconstruction, and coincidence measurements.
- Radiation Hardness: accelerator environments, detectors must withstand high levels of radiation
- Cryogenic Requirements: require extremely low operating temperatures (TES) (cryogenic conditions), adding complexity to the experimental setup and operation.
- Area vs. Resolution Trade-off: large active detection area (to maximize signal capture) and maintaining excellent energy resolution.



"The most important experiment to be carried out in low energy K-meson physics today is the definitive determination of the energy level shifts in the K-p and K-d atoms, because of their direct connection with the physics of KN interaction and their complete independence from all other kinds of measurements which bear on this interaction". R.H. Dalitz (1982)

Previous Kaonic X-ray Measurements Repulsive Attractive 300 200 2100 Zero Shift Ka energy 8150 100 × 50 ¥100 Davis et al. (1979 Energy (keV) E(Ka) - E(Ka(EM Only))

Carlo Guaraldo Hannes Zmeskal

the key initiators and drivers of kaonic atoms studies at the DA Φ NE, JPARC









SIDDHARTA-2 setup: cryogenics

hydrogen density of 1.3×10^{-3} g/cm³ Deuterium: ~1.5% of liquid deuterium density

liquid 4He / liquid 3He at JPARC



Kapton polyimide (C22H10N2O5) foils, with a thickness of 75 μm



Target + SDD cooling

support and cooling structure for target and SDDs

Kapton target wall with Alu support structure

Targets

SDD array

start counter T0

Leybold MD10 - 18 W @ 20 K target cell and SDDs are cooled via ultra pure aluminum bars T_{TC} = 20-30 K

 $T_{SDD} \sim 130 \text{ K}$ $P_{TC} = 1.5 - 2 \text{ bar}$



SIDDHARTA-2 setup: cryogenic targets

Selected materials in different configuration: vacuum entrance windows target walls cooling supports



windowless designs to enhance low-energy X-ray transmission and detection efficiency.

Detector System

Whole X-ray energy range Different detection methods Excellent features



















Characteristic	Charge Couple device (CCD)	Silicon Drift Detectors (SDDs)	Transition-Edge- Sensor (TES) Microcalorimeters	Cadmium-Zinc-Telluride (CdZnTe) Detectors
Energy Resolution (FWHM)	~140 eV at 6 keV ;	~150 eV at 6.4 keV ; <200 eV at 6 keV	< 10 eV at 6 keV Absolute calibration 0.04 eV uncertainty at 6.4 keV	Few % FWHM ; 6% at 60 keV; 2.2% at 511 keV
Timing Resolution	~30 sec Exposure time ~ 1 minutes	Below 1 µsec ; ~ 800 ns FWHM ; < 500 ns FWHM ;	fall times <100 μs rise times ~ tens μs	Few tens of ns
Material	Silicon (high purity)	Silicon (high purity)	Mo/Cu + Bi (absorber), Au 2um + Bi 4um, 20 mm ²	Cadmium-Zinc-Telluride (CdZnTe)
Quantum Efficiency	~ 80% up to 6 keV	> 95% efficiency in 1-10 keV ;	85% at 6 keV (NIST TES spectrometer)	> 98% theoretical absorption for 1.25 mm thick at 58.9 keV
Detector Area	22.5 x 22.5 um pixel size; 1242 x1152 pixels; 116 cm ² total area	SIDDHARTA: 144 SDDs, 1 cm ² ; SIDDHARTA-2: 384 SDDs, 246 cm ² ; J-PARC E57: 300 cm ² total	240 pixels, ~ 0.23 cm ² total collecting area	1 cm ² active surface (prototype)
Operating Temperature	~130 - 150 K ;	~170 K	~100 mK	Room temperature



Pixelated detectors for a wide energy range





oject

Characteristic	Value
Pixel Format/size	1152 (H) x 1242 (V) pixels 22.5 x 22.5µm
Active Area	7.24 cm ²
Depletion Depth	~30 µm
Energy Resolution	~140 eV at 6 keV
Timing Capability	Non-triggerable; ~10 s readout time
Background Rejection Mechanism	Event Topology (Spatial Pattern Analysis)







Silicon Drift Detectors (first large area) with integrated JFET

PNSens@r



(b) -150electron potential [V]

back contact -100 -50 [m] utop anode 20 radius [mm] 0.1

New x-ray detectors specially designed as well as readout electronics

energy range < 20 keV energy resolution ~ 150 eV (FWHM) at 6 keV; stability and linearity better than 10⁻⁴; Fast detector - > trigger system at the level of 1µs; **Operating in high radiation environment;** Custom topology – large active area (~ cm² / channel)



SDD with external CUBE preamplifier

Monolithic SDDs arrays developed by **Fondazione Bruno Kessler**

new technology, lower production cost

- 2x4 matrix SDD units (0.64 cm²)
- active/total surface ratio of 0.75



CUBE



A CMOS low-noise charge sensitive preamplifier (CUBE operate at lower cryogenic temperature (up to 50k) SIDDHARTA-2 Ceramic carrier



Thick SDD: extending the working range





See talk of F. Cloza

X-ray Energy (keV)	Eff. (0.45 mm Si)	Eff. (1.0 mm Si)
10	~100 %	~100 %
15	~90 %	~90 %
20	~37 %	~65 %
25	~18 % (≈)	~45 %
30	~10 % (≈)	~28 %



Counts in SDD 1 (BUS 1)



Transition Edge Sensors (TES)

Response function ×10⁴ CoK_{a1} intrinsic E62 setup Thermometer w/o tail 1.0 X-ray energy : fit @J-PARC K1.8BR data 0.8 Absorber Heat capacity : Counts / 1 eV CoK_{a2} 0.6 Thermal conductance : G Lig. Helium rget Cryosta 0.4 Low temperature heat sink ~5.8 eV FWHM 0.2 a few mK Resistance Pb shield 0.0 6880 6900 6920 6940 X-ray generator Energy (eV) supernormal dE counter & MWDC Detector response is well described 900 MeV/c Cu degrader conducting conducting by a gaussian and a low-energy exponential tail state state Groundbreaking high-resolution measurement of K_{3,4}He isotopic ~ 100 mK Temperature 0 shift on 2p level Muonic Ne atom $5 \rightarrow 4$ 35 ӟНе ⁴He Sync. 50 K³He Lα Phys. Rev. Lett. 130, 173001 (2023) 30 Async. Async. BG C⁴He 25 Energy resolution 40µNe 5-4 → negligible Counts / 2 eV 0.9 atm ΔE = 5.2 ~ 5.5 eV 20 (c.f. ΔE (off beam) = 5.0 ~ 5.2 eV @ Co Ka) 30 - χ^2/ndf Kα χ^2/ndf 5~6 eV FWHM =91.3/96=97.0/7115 Fe Counts / 2.0 eV 20 10² 10 Cu Ka µNe 7-5 **Co** Kα Main BG come from 10 stopped K absorption Cr Ka 5 µBe 3-2 ┡╍╋╍┫┓┓┫┙┛┥ 6250 6500 6150 6200 6400 6450 10 X-ray energy (eV) X-ray energy (eV) $E_{3d \to 2p}^{K^{-3}\text{He}} = 6224.5 \pm 0.4 \text{(stat)} \pm 0.2 \text{(syst) eV}$ $E^{K^{-4}\text{He}}_{3d \rightarrow 2p}$ $= 6463.7 \pm 0.3(\text{stat}) \pm 0.1(\text{syst}) \text{ eV}$ $\Gamma_{2n}^{K^{-4}\text{He}} = 1.0 \pm 0.6(\text{stat}) \pm 0.3(\text{syst}) \text{ eV}$ $= 2.5 \pm 1.0 (\text{stat}) \pm 0.4 (\text{syst}) \text{ eV}$ 5000 5500 6000 6500 7000 7500 8000 Energy (eV)

See talk of T. Hashimoto

Cadmium Zinc Telluride (CZT)

High atomic number Good absorption efficiency

- 1mm thick detectors >98% at 60keV
- 10mm thick detectors >86% at 200keV Optimal band gap *Room Temperature Operation* High energy resolution:
- 1.5 % at 50 keV
- 0.82 % at 660 keV
- Fast detector response: down to 50 ns



See talk of F. Artibani



Compound	Si	Ge	GaAs	CZT	CdTe
Mean atomic number	14	32	32	49.1	50
Bandgap (eV)	1.12	0.66	1.42	1.57	1.5
electrons (cm²/V)	2-5	5	10-4	10-2	10-3
holes (cm²/V)	1-2	2	10-5	3 10 ⁻⁵	5 10-4
Resistivity (Ωcm)	2.3 105	47	10 ⁸	5 10 ¹⁰	10 ⁸ -10 ⁹
Thickness to absorb 90% of 60keV incident radiation (cm)	130	2.6	2.6	0.5	0.5



UniPa DiFC (Palermo) IMEM-CNR (Parma)

CZT does not need complex cryogenic cooling systems, like those needed for high-purity germanium (HPGe)



SIDDHARTA-2 setup: kaon trigger and luminosity monitor



Kaon Trigger consists of two plastic scintillators read by PMT's placed above and below the IR.



The **Luminosity monitor** consists of two plastic scintillators in the horizontal plane

Kaon Trigger



The combined used of Kaon Trigger and SDDs drift time allows to reduce the asynchronous background by a factor $\sim 2\cdot 10^4$



Synchronous background associated to kaon absorption on materials nuclei, or to other Φ decay channels.

Veto-1 14 plastic scintillator read by PMTs to select the events occurring in the gas target, rejecting the X-ray background corresponding to K- stopped in the solid elements of the setup



Veto-2 48 plastic scintillator read by SiPMs to suppress the background induce by particles produced by kaon absorption, passing through the SDDs.







Beyond SIDDHARTA-2 - EXKALIBUR



EXtensive Kaonic Atoms research: from Lithium and Beryllium to URanium

 proposal to perform fundamental physics at the strangeness frontier at DAΦNE presented in varies Scientific Committee and INFN commissions

C. Curceanu et al., Front.in Phys. 11 (2023) 1240250

Kaon mass by Kaonic neon measurement

- Use the present setup
- Minimal modification
- new calibration system
- 300 pb-1 of integrated luminosity

Light kaonic atoms (LHKA)

- Use the present setup
- solid target Li, Be, B
- Minimal modification target
- integration 1mm SDD
- 200 pb-1 of integrated luminosity

Intermediate kaonic atoms (IMKA)

In parallel we plan dedicated runs for kaonic atoms (*O*, *AI*, *S*) with different types of detectors: CdZnTe detectors - 200 -300 pb-1 of integrated luminosity



Light kaonic atoms (LHKA)

Integration of 1 mm SDD

Lithium-6		Lithium-7	Beryllium-9		Boron-10	Boron-11
Transition	Energy (keV)	Energy (keV)	Transition	Energy (keV)	Energy (keV)	Energy (keV)
$3 \rightarrow 2$	15.085	15.261	$3 \rightarrow 2$	27.560	43.568	43.768
$4 \rightarrow 2$	20.365	20.603	$4 \rightarrow 3$	9.646	15.156	15.225
$5 \rightarrow 2$	22.809	23.075	$5 \rightarrow 3$	14.111	22.171	22.273
$4 \rightarrow 3$	5.280	5.341	$5 \rightarrow 4$	4.465	7.015	7.047
$5 \rightarrow 3$	7.724	7.814	$6 \rightarrow 4$	6.890	10.826	10.875
$5 \rightarrow 4$	2.444	2.472	$6 \rightarrow 5$	2.425	3.811	3.828



Solid target system

Kaonic boron test measurement successfully achieved Construction of new support system conical shape to maximize the solid angle MC simulations ongoing



Intermediate kaonic atoms (IMKA)

In parallel we plan dedicated runs for kaonic atoms (*O*, *AI*, *S*) with different types of detectors: CdZnTe detectors

- 200 - 300 pb-1 of integrated luminosity



Kaonic Oxygen: key role in the description of the nuclear-matter density distribution which enters in the formula for the density-dependent optical potentials
Kaonic Aluminium: 3->2 QCD – never measured;
Kaonic Sulphur: 4->3 past measurements are inconsistent

We are developing an optimized CdZnTe based setup

Larger active area: 32 detectors instead of 8 Optimized geometry and shielding to reduce the background



EXKALIBUR DEAR **SIDDHARTA Kaonic Atoms** 2002 2009 measurements ALL TRUNKING CAL at Da Φ ne antikaon **Kaonic Neon** (Kaon mass) First kaonic deuterium Hydrogen Nitrogen (2023 - 2024).00794 14 0007 X-ray X-ray **Kaonic Neon** (2023) kicked-out electron Helium Neon **Kaonic Helium-4** 4.003 20,180 S external (2022)stimulation Sta Deuterium (2011)2.01410 NA: 0.0115% **Kaonic Nitrogen** VISOTOPE.CO (2002)





Silicon Drift Detectors spectroscopy response

The energy response is linear within few (<5 eV between 4 keV and 14 keV) Excellent energy and time resolution @ 140 K

