### Seeding of cosmos structures, recent developments

### A.D. Dolgov

Novosibirsk State University, Novosibirsk, Russia Bogolyubov Laboratory of Theoretical Physics, JINR, Dubna, Russia The work was supported by the RSF Grant 23-42-00066 Università degli Studi dell'Aquila Italia March, 14, 2025

## Abstract

#### Resolution of cosmological conundrums with primordial black holes.

The works on possible explanations of the origin of the dense population of the early universe at redshifts near ten are critically discussed. It is argued that the suggested mechanisms suffer from serious, recently discovered inconsistencies.

- It is shown that multiple cosmological problems observed both in the early and the contemporary universe are uniquely resolved with a population of primordial black holes (PBHs).
- The model of PBH formation is described, that predicts the mass spectrum of PBHs in excellent agreement with observations.

• A by-product of the proposed mechanism leads to a noticeable population of our Galaxy with antimatter: positrons, anti-nuclei, and antistars, that are discovered during recent years.

# Which came first: the chicken or the egg?

Supermassive black holes (SMBHs) are observed at the centres of all large galaxies: in the contemporary 14.6 Gyr year old universe, as well as in the young one at the redshifts  $\sim$  10 or at the age from a billion down to 300 million years. It is commonly assumed that first came a galaxy and later the central SMBH was created by accretion of matter to the galactic centre. However this conjecture seems to be in strong contradiction both with observation and theoretical estimates.

# Seeding of galaxy formation by PBHs

Usually it is assumed that supermassive black holes (SMBHs) observed in large galaxies, are created by matter accretion to the galactic centers.

However, the necessary time is much larger than the universe age, even for the contemporary universe, with the age about 15 billion years, to say nothing of the 20 times younger universe at  $z \sim 10$ .

Contradiction between observations and canonical theory is neatly solved if the universe is populated by primordial black holes (PBH) that **seeded** formation of cosmic structures, as is suggested 30 years ago: A.Dolgov, J.Silk, PRD 47 (1993) 4244 (DS), "Baryon isocurvature fluctuations at small scale and baryonic **dark matter**"; A.Dolgov, M. Kawasaki, N. Kevlishvili (DKK), NPB807 (2009) 229, "Inhomogeneous baryogenesis, **cosmic antimatter**, and dark matter".

# Inverted mechanism of galaxy formation

In DS and DKK an **inverted** formation mechanism of galaxies and their central black holes is proposed. Namely, a primordial SMBH was formed firstly and later it **seeded** the galaxy formation.

The 30 year old DS idea of seeding, but of different kind, was rediscovered in several recent publications, under the pressure of the NASA Hubble Space Telescope (HST) and the NASA James Webb Telescope (JWST) observations of the early universe at redshifts z = 5 - 15.

Existence of black holes was ingeniously predicted in 1783 by John Michell, an English country parson, famous for many other discoveries in physics.

He noticed that there could be stellar bodies having the second cosmic velocity larger than the speed of light. Hence such objects neither shine nor reflect light, and it would be impossible to observe them directly.

Michell called such, not emitting radiation stars as **"dark stars"**. According to his idea a single dark star would be invisible, but if a double system of a dark and a usual star is formed, one may identify dark star observing the other one rotating around "nothing".

This is one of possible ways to observe black holes at the present time, among many others.

## Black hole observations

### Star rotating around nothing, exactly as envisaged by John Michel.

It appeared a consistent gravitational tug from a hitherto unknown companion was disrupting the star's motion. Known to weigh roughly 33 times our Sun, the cosmic behemoth is the heaviest stellar-mass black hole yet found in the Milky Way. **Quite unusual for astrophysical BH.** 

EAS2024, European Astronomical Society Annual Meeting, held 1-5 July, 2024 in Padova, Italy. Online at https://eas.unige.ch/EAS2024/.



An artist's conception shows the orbits of both the star (in blue) and the black hole (in red), dubbed Gala BH3, around their common center of mass. Credit: ESO/L. Calçada/Space Engine (spaceengine.org)

## Other methods of BH observations

- BHs evaporate and shine (Hawking radiation), though nobody yet saw it.
- Near-solar mass BHs are observed by X-rays from the accreting matter.
- Gravitational lensing by BH, rise of background star brightness, that's how  $0.5M_{\odot}$  MACHOs have been discovered. Could not be created by stellar collapse.
- Mass estimates via star motion around BH, as e.g. in our Galaxy.
- Quasars (QSO), supermassive black holes, that radiate as thousands galaxies,  $L_{QSO} = 10^{46} - 10^{47}$  erg/sec, i.e.  $10^{13}L_{\odot}$ , though they are practically point-like, their size is about  $10^9 - 10^{10}$  km, i.e. smaller than the Solar System. The only known mechanism of QSO radiation is the process of ultrarelativistic particle collision in the process of matter accretion. QSOs shine until they consume all "food" around and remains almost in desert. After that the shining fades out as e.g. BH in the center of our Galaxy.

• All these methods however only allow to measure the mass inside small central volume. According to General Relativity **theory** there must be a BH inside. So strictly speaking BH existence was not proven by "experiment". Now LIGO/Virgo/KAGRA registration of gravitational waves from BH binaries demonstrates the first direct "picture" of the Schwarzschild metric and presents strong evidence that the sources are PBHs.

- After creation of GR (Einstein, 1915) almost immediately exact solutions of GR equations describing all possible types of BHs have been found.
- BHs have only three types of "hairs" that may be observed outside BH:1. Gravitational field created by mass, M, asymptotically at large distances tending to the Newtonian limit.
- 2. Electric field created by charge Q, tending to the Coulomb one.
- 3. Field induced by rotation, absent in Newtonian gravity.

Schwarzschild (1916), BH with Q = J = 0. Reissner-Nordström (1916, 1918)  $J = 0, Q \neq 0$ . Kerr, (1963),  $Q = 0, J \neq 0$ . Kerr–Newman (1965)  $J \neq 0$ ,  $Q \neq 0$ . The space-time is not flat Minkowski one of special relativity with the interval:  $ds^2 = dt^2 - dr^2$ , but strongly curved (see next slide). If photon mass is non-zero, no matter how tiny, electric field of BH should completely vanish. No continuous limit to  $m_{\gamma} = 0$ . A.D. Dolgov, H. Maeda, T. Torii, "One more mechanism of electric charge nonconservation", hep-ph/0210267 [hep-ph]; A.D. Dolgov, K.S. Gudkova, "Massive photons and electrically charged black holes", Phys. Lett. B810 (2020) 135844 • e-Print: 2006.16332 [gr-qc]. Discovery of charged BH could present the absolute upper bound on  $m_{\gamma}$ . Possible way to create electric asymmetry of the universe.

# BH types by formation mechanisms

#### 1. Astrophysical black holes,

created by the collapse of a star which exhausted its nuclear fuel. The expected masses should start immediately above the neutron star mass, i.e. about  $3M_{\odot}$ , but noticeably below  $100M_{\odot}$ . Instead we observe that the BH mass spectrum in the galaxy has maximum at  $M \approx 8M_{\odot}$  with the width  $\sim (1-2)M_{\odot}$ . The result is somewhat surprising but an explanations in the conventional astrophysical frameworks is possible.

**Recently LIGO/Virgo discovered BHs with masses close to**  $100M_{\odot}$ . Their astrophysical origin was considered **impossible due to huge mass loss in the process of collapse.** Now some, quite exotic, formation mechanisms are specially invented.

No problem if  $100M_{\odot}$  BH is primordial.

## BH created by accretion

**2.** BH formed by accretion to the mass excess in the galactic center. In any large galaxy there exists a supermassive BH (SMBH) at the center, with masses varying from a few millions  $M_{\odot}$  (e.g. Milky Way) up to almost hundred billions  $M_{\odot}$ . However, the conventional accretion mechanisms are not efficient enough to create such monsters during the universe life-time,  $t_U \approx 14.6$  Gyr. At least 10-fold longer time is necessary, to say nothing about SMBH in 20 times younger universe, observed in impressive numbers by HST and JWST.

An estimate of the accretion efficiency in the Galaxy E.M. Murchikova, et al Nature 570, 83 (2019): Building up SMBH SgrA\* with the mass  $\sim 4 \times 10^6 M_{\odot}$  residing at the centre of our galaxy. within the  $\sim 10^{10}$  year lifetime of our galaxy would require a mean accretion rate of  $4 \times 10^{-4} M_{\odot}$  per year. At present, X-ray observations constrain the rate of hot gas accretion to  $\dot{M} \sim 3 \times 10^{-6} M_{\odot}$  per year and polarization measurements constrain it near the event horizon to  $\dot{M}_{hor} \sim 10^{-8} M_{\odot}/\text{yr}$ . The universe age is short by two orders of magnitude. to say nothing about 30 times younger universe.

Even worse, SMBH are observed in empty space or in tiny galaxies where is not enough matter to create such a monster.

# Primordial Black Holes

3. Primordial black holes (PBH) created during pre-stellar epoch

The idea of the primordial black hole (PBH) i.e. of black holes which could be formed the early universe prior to star formation was first put forward by Zeldovich and Novikov: "The Hypothesis of Cores Retarded During Expansion and the Hot Cosmological Model", Astronomicheskij Zhurnal, 43 (1966) 758, Soviet Astronomy, AJ.10(4):602–603;(1967).

According to their idea, the density contrast in the early universe inside the bubble with radius equal to the cosmological horizon might accidentally happen to be large,  $\delta \varrho / \varrho \approx 1$ , then that piece of volume would find itself inside its gravitational radius i.e. it became a PBH, decoupled from the cosmological expansion.

Elaborated later in S. Hawking, "Gravitationally collapsed objects of very low mass", Mon. Not. Roy. Astron. Soc. **152**, 75 (1971).

B. J. Carr and S. W. Hawking, "Black holes in the early Universe," Mon. Not. Roy. Astron. Soc. **168**, 399 (1974).

**PBHs are mostly rejected by astrophysical establishement. Why?!** Forbidden by law? There is the following conventional division of black holes by their masses:

- 1. Supermassive black holes (SMBH):  $M = (10^6 10^{11})M_{\odot}$ .
- 2. Intermediate mass black holes (IMBH):  $M = (10^2 10^5) M_{\odot}$ .
- 3. Solar mass black holes: masses from a fraction of  $M_{\odot}$  up to  $100 M_{\odot}$ .

The origin of most of these BHs is unclear in the conventional approach, except maybe of the BHs with masses of a few solar masses, that might be astrophysical. Highly unexpected was great abundance of IMBH which are being observed at the present day universe during last few years in large numbers.

The suggestion that (almost) all black holes in the universe are primordial strongly reduce or even eliminate the tension.

However, in earlier works the masses of PBHs were quite low,  $10^{20}$  g, but inflation permitted to create gigantic PBH.

# PBH and inflation

**Cosmological inflation allows for formation of PBH with very large masses.** Mass of perturbations inside cosmological horizon grows exponentially, that allows for creation even of supermassive PBHs. It was first applied to PBH production in A.Dolgov, J.Silk, PRD 47 (1993) 4244 (DS) "Baryon isocurvature fluctuations at small scale and baryonic **dark matter**".

- a year later in: B.J. Carr, J.H. Hilbert, J.E. Lidsey, "Black hole relics and inflation: Limits on blue perturbation spectra", Phys.Rev.D 50 (1994) 4853, astro-ph/9405027;
- and soon after in P. Ivanov, P. Naselsky, I. Novikov (May 10, 1994), "Inflation and primordial black holes as dark matter", PRD 50 (1994) 7173.
- An avalanche of papers on inflationary formation of PBH nowadays.
- Presently inflationary mechanism of PBH production is commonly used.
- However, almost always the calculated mass spectrum is multi-parameter one and quite complicated.
- The only exception is the log-normal spectrum predicted by DS, that is perfectly well confirmed by observations.

## Nuffield workshop, 1982, dedicated to inflation



### History of inflation

Inflationary cosmology is probably the most important step in understanding of the universe evolution during last 50 years. Solves problems of causality, flatness, perturbations,... creates supermassive PBH Pioneering papers:

A.A. Starobinsky, A new type of isotropic cosmological model without singulatiry Phys. Lett. 91B, 99 (1980) Received 11/01 1980.

D. Kazanas: Dynamics of the universe and spontaneous symmetry braking. ApJ, 241, L59 (1980), Received 05/05/1980

A. Guth, Inflationary universe: A possible solution to the horizon and flatness problems, PRD 23, 347 – Published 15 January 1981, Received 11/08/1980 A. D. Linde, "A new inflationary universe scenario: A possible solution of the horizon, flatness, homogeneity, isotropy and primordial monopole problems," Phys. Lett. B 108, 389 (1982). Published 04/02/1982.

A. Albrecht, P. J. Steinhardt, "Cosmology for grand unified theories with radiatively induced symmetry breaking," Phys. Rev. Lett. 48, 1220 (1982). Published 26/04/1982.

**Density perturbations, spectrum confirmed by observations:** V.F. Mukhanov, G.V. Chibisov, JETP Lett. 33 (1981) 532.

# Prehistory of inflation

There is significant "pre-inflationary" literature.

The idea that the universe avoided singularity and underwent exponential period during which the mass of the cosmological matter rose by tens orders of magnitude was discussed by E.B. Gliner ZhETF, 49 (1965) 542 and by E.B. Gliner and I.G. Dymnikova Astronomical Journal Letters, 1 (1975) 7. De Sitter like (exponentially expanding) non-singular cosmology was considered by V.Ts. Gurovich and A.A. Starobinsly Sov. Phys. JETP 50 (1979) 844; Zh. Eksp. Teor. Fiz. 77 (1979) 1683.

In the last paper inflation was supposed to be induced by radiative corrections to the Einstein equations that could in particular lead to an addition of  $R^2$  term to the classical Hilbert-Einstein action.

The calculation of these corrections were first performed by V.L. Ginzburg, D. A. Kirzhnits, and A. A. Lyubushin, 1971, Zh. Eksp. Teor. Fiz. 60, 451.

### Log-normal mass spectrum

The log-normal mass spectrum of PBHs was predicted by Dolgov-Silk and further worked out in A.Dolgov, M. Kawasaki, N. Kevlishvili (DKK), NPB807 (2009) 229, "Inhomogeneous baryogenesis, cosmic antimatter, and dark matter":

$$\frac{dN}{dM} = \mu^2 \exp\left[-\gamma \ln^2(M/M_0)\right],$$

confirmed by LiGO/Virgo/Kagra with very good precision.

According to DS and DKK scenario initial large perturbations of baryonic number (isocurvature perturbations, since quarks are massless) are generated at inflationary stage, transformed into density perturbations when massless quarks turned into heavy protons and neutrons. Hence  $M_0$  should be equal to horizon mass at the QCD phase transition:  $M_0 \sim 10M_{\odot}$ , as predicted in A.Dolgov, K.Postnov, "Why the mean mass of primordial black hole distribution is close to  $10M_{\odot}$ ". JCAP 07 (2020) 063. The horizon mass at QCD p.t. is close to  $10M_{\odot}$ , for  $\mu = 0$ . At larger chemical statements of the perturbation of the pe

potential  $T_{pt}$  is smaller and  $M_{hor}$  is larger, namely,  $\approx 17 M_{\odot}$ , according to lattice calculations (I. Arefyeva), exactly what is observed, see below.

Conventional ACDM cosmology encounters serious difficulties describing astronomical observations during all the history of the universe, starting from our time with the universe age about 15 billion years, and back to the past down to  $\sim$  300 million years discovered recently by HST, JWST, and ALMA at high redshifts  $z\sim10.$ 

The inconsistenies discovered earlier are reviewed in 2018: A.D. "Massive and supermassive black holes in the contemporary and early Universe and problems in cosmology and astrophysics", Phys. Usp. 61 (2018) 2, 115.

The tension that existed during all life-time of the universe between the canonical theory and observations is neatly solved if the universe is populated by primordial black holes (PBH) that **seeded** formation of cosmic structures, **as is suggested in DS and DKK**.

# Confirmation by "experiment"

#### Observational confirmation of DS/DKK mechanism:

- The calculated mass spectrum of PBH very well agrees with "experiment".
- Noticeable antimatter population of the Galaxy is anticipated and confirmed by the observations of **positrons, antinuclei, and antistars.**
- The early galaxy formation observed by HST and JWST is explained if galaxies are SEEDED by BHs, as it is also rediscovered in several papers of the last years.
- Discovery of IMBH, with  $M \sim (10^3 10^5) M_{\odot}$  in dwarfs and globular clusters, predicted in AD & K. Postnov. "Globular Cluster **Seeding** by Primordial Black Hole Population", JCAP 04 (2017) 036, e-Print: 1702.07621 [astro-ph.CO].
- Recent observation of antistar gamma burster possible source star-antistar collision!!!

# Problems of the contemporary universe. Summary.

- 1. Supermassive BH (SMBH) in all large galaxies. Too short time (15 billion year) for their formation through the conventional accretion mechanism. 2. Huge SMBH in small galaxies and even in (almost) EMPTY space. No material for their creation. Pushed out of large galaxies? Wandering BHs? 3. Too old stars, older than the Galaxy and maybe one **older** than the universe? 4. MACHOs, non-luminous objects,  $M \sim 0.5 M_{\odot}$ , observed through microlensing. 5. Problems of BH mass spectrum in the Galaxy with  $(M = 7.8 \pm 1.2) M_{\odot}$ . 6. Origin and properties of the sources of the observed gravitational waves. 7. Origin of intermediate mass BH (IMBH) with masses  $(10^3 - 10^5)M_{\odot}$ . Plentv of them are observed everywhere in the universe, in particular in dwarfs and globular clusters, as predicted by AD & K. Postnov. 8. Discovery of BH with  $M \approx 100 M_{\odot}$ , that is strictly forbidden but
- nevertheless observed by LIGO/Virgo.
- 9. Strange stars in the Galaxy, too fast and with unusual chemistry.

## Problems of the early universe

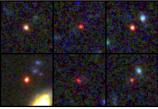
Serious problems, similar to those recently found by JWST, are known already for many years. HST discovered that the early universe, at z = 6 - 7 is too densely populated with quasars, alias SMBH, supernovae, gamma-bursters and it is very dusty. No understanding in conventional cosmology how all these creature were created in such a short time.

"Hubble" sees the universe up to z = 6 - 7, but accidentally a galaxy at  $z \approx 12$  has been discovered for which both Hubble and Webb are in good agreement. Huge BHs in small galaxies discovered in the early universe (as well as in the contemporary one). Such huge BHs could not be created by the accretion of matter to galactic center, since the amount of material is too small. Quite a few SMBH are observed in practically empty space!!!

All the problems are neatly solved if the universe is populated by primordial black holes (PBH) and the astrophysical large bubbles with very high baryonic density, according to DS and DKK.

### Impossible galaxies

I. Labbé, P. van Dokkum, E. Nelson, *et al*, "A population of red candidate massive galaxies 600 Myr after the Big Bang", Nature, 616, 7956, (2022), arXiv:2207.12446. Six candidate massive galaxies (stellar mass > 10<sup>10</sup> solar masses) at  $7.4 \lesssim z \lesssim 9.1$  500–700 Myr after the Big Bang, one galaxy with a possible stellar mass of ~  $10^{11} M_{\odot}$ , too massive to be created in so early universe. According to the authors it is **impossible** to create so well developed galaxies. NB: "May be they are supermassive black holes of the kind never seen before. That might mean a revision of usual understanding of black holes." Well agrees with our suggestion of PBHs.



The six candidate galaxies identified in the JWST data. (NASA, ESA, CSA, I. Labbe/Swinburne University of Technology)

# Ultra-massive early QSO observed by ALMA

ALMA (Atacama Large Millimeter Array) confirmation of an obscured hyper-luminous radio-loud AGN at z = 6.853 associated with a dusty starburst in the 1.5 deg2 COSMOS field,

R. Endsley et al, MNRAS, 250, 4609 (2023).

VIRCam and IRAC photometry perhaps suggests that COS-87259 is an extremely massive reionization-era galaxy with  $M_*=1.7\times10^{11}M_\odot$ 

Such a very high AGN luminosity suggests that this object is powered by

 $\sim 1.6 \times 10^9 M_{\odot}$  black hole if accreting near the Eddington limit.

BH mass is about 1% of the stellar mass, 100 times larger than usually. Normally impossible, but PBH could seed such monster.

The low stellar mass confirms idea of seeding by PBH: the galaxy is still in the process of formation.

Summarising: "recent observations have found a large number of supermassive black holes already in place in the first few hundred million years after Big Bang. The channels of formation and growth of these early, massive black holes are not clear, with scenarios ranging from heavy seeds to light seeds, experiencing bursts of high accretion rate".

# Seven little red dots

"Environmental Evidence for Overly Massive Black Holes in Low Mass Galaxies and a Black Hole - Halo Mass Relation at  $z \sim 5$ ", J. Matthee, R.P. Naidu, G. Kotiwale, *et al* arXiv:2412.02846 [astro-ph.GA].

At z = 4 - 5 seven BHs with masses  $\sim (4 - 15) \times 10^6 M_{\odot}$  are observed. The black holes seem too massive compared to the mass of the stars in the galaxies that host them, about 10% of the stellar mass in the galaxies.

Milky Way, supermassive black holes tend to have masses equal to around 0.01% of the stellar mass of their host galaxy.

Supermassive black holes in 'little red dot' galaxies are 1,000 times larger than they should be.

Quote: "...and astronomers don't know why. One possible explanation for this is that supermassive black holes in the early universe managed to form and grow much more efficiently than those in the present-day universe."

Primordial BHs seeding galaxy formation easily resolve this conundrum. This is not that BHs are too heavy, but the galaxies are too light, since they are still in the process of formation.

## Superfast accretion or PBH observation

It was announced in ref. H. Suh, J. Scharwächter, E.P. Farina, *et al* "Feeding Hidden Monsters: a Super-Eddington accreting Black Hole 1.5 Gyr after the Big Bang", Nature Astronomy, Advanced Online Publication, Pub Date: November 2024, arXiv:2405.05333 that the supermassive black hole LID-568 appears to be feeding on matter at a rate 40 times its Eddington limit, **theoretically impossible**?

According to the authors such **impossibly huge** accretion rate is one of the possible explanations for heavy black holes formation so early in the universe. To this end some smaller black hole "seeds," are necessary which current theories suggest arise either from the death of the universe's first stars (light seeds) or the direct collapse of gas clouds (heavy seeds).

The authors claim that this observation is a **discovery** that it's possible for a black hole to grossly exceed its Eddington limit. How could it be achieved?

But it could be a proof of existence of supemassive PBH and no breaking of sacred principles is necessary.

## Strong blow to the super-Eddington accretion

"A dormant, overmassive black hole in the early Universe", I. Juodzbalis, R. Maiolino, W.M. Baker, *et al*, Nature, Volume 636, Issue 8043, pp. 594-597, 2403.03872: Recent observations have found a large number of SMBHs already in place in the first few hundred million years after Big Bang. The channels of formation and growth of these early, massive black holes are not clear, with scenarios ranging from heavy seeds to light seeds experiencing bursts of high accretion rate.

The detection, from the JADES survey, of broad H $\alpha$  emission in a galaxy at z = 6.68, which traces a black hole with mass of  $\sim 4 \times 10^8 M_{\odot}$  and accreting at a rate of **only 0.02 times the Eddington limit.** 

The black hole to stellar mass ratio is 0.4, i.e.  $10^3$  times above the local relation. Huge BH in small galaxies, see more examples below.

This object is most likely the tip of the iceberg of a much larger population of dormant black holes around the epoch of reionization. Its properties are consistent with scenarios in which short bursts of **super-Eddington accretion have resulted in black hole overgrowth and massive gas expulsion from the accretion disk; in between bursts, black holes spend most of their life in a dormant(?!) state.** 

## Direct collapse out !?

To save direct collapse of cold gas in the early universe is necessary to assume that gas is not heated in the process of compression???!!! **Impossible**.

M.J. Hayes, J.C. Tan, R.S. Ellis, *et al* "Glimmers in the Cosmic Dawn: A Census of the Youngest Supermassive Black Holes by Photometric Variability", The Astrophysical Journal Letters, Volume 971, Issue 1, id.L16.

The variability estimate of  $n_{SMBH}$  at  $z = (6 - 7) \ge 8 \times 10^{-3} \text{ Mpc}^{-3}$  places constraints on  $d_{iso}$  to be << 100 kpc. It requires the halo mass threshold for SMBH formation to be  $<< 3 \times 10^{10} M_{\odot}$ . However, such models begin to have more severe tension by having significantly greater abundance than the z = 0 estimate of  $n_{SMBH}$ .

Finally, the estimate of the necessary value of  $n_{SMBH}$  is about 100 times greater than the direct collapse prediction of Chon et al. (2016) and at least  $10^4$  times greater than that of Wise et al. (2019).

# Huge black holes in tiny galaxies at high z

Black holes in high z universe are too massive w.r.t. the expectations based on observations of BHs in contemporary large galaxies - PBHs solve the problem.

F. Pacucci, B. Nguyen, S. Carniani *et al* "JWST CEERS and JADES Active Galaxies at z = 4 - 7 Violate the Local  $M_* - M_{BH}$  Relation at  $> 3\sigma$ : Implications for Low-mass Black Holes and Seeding Models", The Astrophysical Journal Letters, 957:L3 (10pp), 2023 November 1.

Black holes are overmassive by factor 10 - 100 compared to their low-z counterparts in galactic hosts of the same stellar mass.

M. Volonteri, M. Habouzit, M. Colpi. "What if young z > 9 JWST galaxies hosted massive black holes?" Monthly Notices of the Royal Astronomical Society, Volume 521, Issue 1, pp.241-250. Only MBHs overmassive relative to expected galaxy scaling relations, accreting at high Eddington rates, would be detectable. Their discovery would point to the presence of heavy MBH **seeds**, but care is needed to exclude the existence of lighter seeds as only overmassive MBHs. **Possible resolution: (primordial)BH seeds operated too little time to create so massive galaxies as those in the contemporary universe.** 

## Confirmation of the inverted galaxy formation

The paper: M.A. Marshall, M. Yue, A-Ch. Eilers, *et al* "GA-NIFS & EIGER: A merging quasar host at z = 7 with an overmassive black hole", arXiv:2410.11035v2 [astro-ph.GA] 17 Oct 2024 presents a strong argument in favour of inverted mechanism of galaxy formation, as suggested by DS. Indeed, the measured mass of the central black hole is  $M_{BH} \approx 1.4 \times 10^9 M_{\odot}$ , that is only twice smaller than the host stellar mass equal to  $M_* \approx 2.6 \times 10^9 M_{\odot}$ . It is hard to imagine that the central BH was created by accretion of matter to the galactic center, as is commonly assumed. For comparison, the stellar mass of the Milky Way is about 60 billion solar masses and the mass of the central BH is about 5 million solar masses. A natural conclusion is that the observed young galaxy is still in the process of formation seeded by supermassive primordial black hole.

The authors note that the mystery of how these huge black holes grew so big so soon after the Big Bang remains unsolved - it is solved if they are primordial.

# Early (impossible!) quasars in empty space

**Final blow:** six early quasars in empty space. Anna-Christina Eilers *et al*, "EIGER. VI. The Correlation Function, Host Halo Mass, and Duty Cycle of Luminous Quasars at  $z \gtrsim 6$ ", the Astrophysical Journal, Volume 974, Number 2. The data indicate that

(a) luminous quasars do not necessarily reside within the most overdense regions in the early Universe,

(b) the UV-luminous duty cycle of quasar activity at these redshifts is  $f_{duty} \ll 1$ . Such short quasar activity timescales challenge our understanding of early supermassive black hole growth.

Using the James Webb Space Telescope, astronomers have peered back 13 billion years to discover **surprisingly lonely** supermassive black hole-powered quasars. Eilers: **"Some of them seem to be sitting in the middle of nowhere** It's difficutl **(impossible?)** to explain how these quasars could have grown so big if they appear to have nothing to feed from.

"Contrary to previous belief, we find, on average, these quasars are not necessarily in those highest-density regions of the early universe." Surely they are primordial SMBH.

# Huge SMBH in tiny early galaxy

Madeline A. Marshall, Minghao Yue, Anna-Christina Eilers, *et al*, "GA-NIFS & EIGER: A merging quasar host at z = 7 with an overmassive black hole", arXiv:2410.11035 [astro-ph.GA].

The black hole powering the quasar has a record mass relative to the stars of the host galaxy.  $M_{BH} \approx 1.4 \times 10^9 M_{\odot}$  while  $M_* \approx 2.6 \times 10^9 M_{\odot}$  It weighs in at 54 percent of its galaxy's stellar mass, versus only about 0.1 percent for central black holes in modern giant galaxies.

Simplest explanation: SMBH is possibly primordial or it has been seeded by PBH that in turn is seeding the galaxy, that have not yet reached its final mass value.

# Summary of the impossibles in the early universe

- Impossible (super) Eddington accretion?
- Impossible direct collapse.
- Impossible quasars in empty space.
- Impossible creation of supermassive BH in very young universe.
- PBHs make the impossible possible.

# IMBHs in globular clusters today

Primordial IMBHs with masses of a few thousand solar mass explain formation of globular clusters (GCs) and dwarf galaxies, otherwise the formation is not well understood, even mysterious.

AD+K.Postnov prediction "Globular Cluster **Seeding** by Primordial Black Hole Population", JCAP 04 (2017) 036, e-Print: 1702.07621 [astro-ph.CO] seems to be recently confirmed:

The observation of IMBH in Omega-Centauri (the most massive globular cluster of the Milky Way): M. Häberle, N. Neumayer, A. Seth *et al*, Nature, **631**, 285–288 (2024):  $M_{IMBH} \gtrsim 8200 M_{\odot}$ ;

F. Pei $\beta$ ker, M. Zajaček, M. Labaj, *et al*, "The Evaporating Massive Embedded Stellar Cluster IRS 13 Close to Sgr A\*. II. Kinematic Structure". Discovery of IMBH with  $M \approx 3 \times 10^4 M_{\odot}$ , 2024, ApJ 970 74.

The origin of IMBH is mysterious, if they are not primordial.

# Huge BHs in dwarf galaxies today

The seeding of dwarfs by intermediate mass BHs is confirmed by the recent data, e.g. in the dwarf galaxy SDSS J1521+1404 the BH is discovered with the mass  $M \sim 10^5 M_{\odot}$ .

Two Candidates for Dual AGN in Dwarf-Dwarf Galaxy Mergers, M. Mićić, et al, arXiv:2211.04609 [astro-ph.GA]. For the first time, astronomers have spotted evidence of a pair of dwarf galaxies featuring GIANT black holes on a collision course with each other. In fact, they haven't just found just one pair – they've found two.

Intermediate-mass black holes: finding of episodic, large-scale and powerful jet activity in a dwarf galaxy SDSS J090613.77+561015.2. Jun Yang et al, e-Print: 2302.06214 [astro-ph.GA,astro-ph.HE]. Discovery of an intermediate-mass black hole (IMBH) with a mass of  $M_{BH} = 3.6^{+5.9}_{-2.3} \times 10^5 M_{\odot}$ , that surely cannot be created by accretion but might seed the dwarf formation.

Critical prediction: IMBHs in Magellanic Clouds?

# Unexpectedly light BH

"Observation of Gravitational Waves from the Coalescence of a 2.5 – 4.5  $M_{\odot}$  Compact Object and a Neutron Star",

Authors: The LIGO Scientific Collaboration, the Virgo Collaboration, the KAGRA Collaboration Report-no: LIGO-P2300352, arXiv:2404.04248:

(2.5-4.5)  $M_{\odot}$  and (1.2-2.0)  $M_{\odot}$ 

The masses do not fit any conventional hypothesis, if they are not primordial.

In the paper T. Jayasinghe, K.Z. Stanek, T.A. Thompson, *et al.* "A unicorn in monoceros: the  $3M_{\odot}$  dark companion to the bright, nearby red giant V723 Mon is a non-interacting, mass-gap black hole candidate",

MNRAS, **504**, (2021), Issue 2, p.2577 a promising but puzzling candidate for one of the lightest black holes ever detected, the system that **shouldn't exist**. It falls right in the middle of the mass gap, weighing at about three solar masses with the mass  $M_{(BH?)} = (3.04 \pm 0.06) M_{\odot}$ .

It is impossible to explain existence of such a light BH by stellar evolution, light stars turn into neutron stars but not into BHs.

# Gravitational waves from BH binaries

• GW discovery by LIGO strongly indicate that the sources of GW are PBHs, see e.g. S. Blinnkov, A.D., N. Porayko, K. Postnov, JCAP 1611 (2016), 036 "Solving puzzles of GW150914 by primordial black holes,"

1. Origin of heavy BHs (~  $30M_{\odot}$ )? There appeared much more striking problem of BH with  $M \sim 100M_{\odot}$ . To form so heavy BHs, the progenitors should have  $M > 100M_{\odot}$  and a low metal abundance to avoid too much mass loss during the evolution. Such heavy stars might be present in young star-forming galaxies but they are not observed in the necessary amount. A possible way out was proposed in J. Ziegler, K. Freese, "Mechanism for nonnuclear energy to fill in the black hole mass gap", Phys.Rev.D 109 (2024) 10, 103042 • e-Print: 2212.13903 based on assumption of dark matter annihilation inside heavy stars.

Formation of BH binaries from the original stellar binaries. Recoil momentum?
 Low spins of the coalescing BHs .

PBHs with the observed by LIGO masses may be easily created with sufficient density.

# Chirp mass

Two rotating gravitationally bound massive bodies are known to emit gravitational waves. In quasi-stationary inspiral regime, the radius of the orbit and the rotation frequency are approximately constant and the GW frequency is twice the rotation frequency. The luminosity of the GW radiation is:

$$\mathbf{L} = \frac{32}{5} \, \mathbf{m}_{\mathbf{P}\mathbf{I}}^2 \left( \frac{\mathbf{M}_c \, \boldsymbol{\omega}_{orb}}{\mathbf{m}_{\mathbf{P}\mathbf{I}}^2} \right)^{10/3} \,,$$

where  $M_1$ ,  $M_2$  are the masses of two bodies in the binary system and  $M_c$  is the so called chirp mass:

$$M_c = \frac{(M_1 M_2)^{3/5}}{(M_1 + M_2)^{1/5}},$$

and

$$\omega_{orb}^2 = \frac{M_1 + M_2}{m_{PI}^2 R^3} \,.$$

A.D. Dolgov, A.G. Kuranov, N.A. Mitichkin, S. Porey, K.A. Postnov,
O.S. Sazhina, I.V. Simkine On mass distribution of coalescing black holes, JCAP 12 (2020) 017, e-Print: 2005.00892.

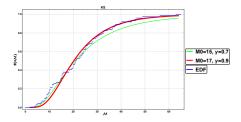
The available data on the chirp mass distribution of the black holes in the coalescing binaries in O1-O3 LIGO/Virgo runs are analyzed and compared with theoretical expectations based on the hypothesis that these black holes are primordial with log-normal mass spectrum.

The inferred best-fit mass spectrum parameters,  $M_0 = 17 M_{\odot}$  and  $\gamma = 0.9$ , fall within the theoretically expected range and shows excellent agreement with observations.

On the opposite, binary black hole formation based **on massive binary star evolution** require additional adjustments to reproduce the observed chirp mass distribution.

# Chirp mass distribution

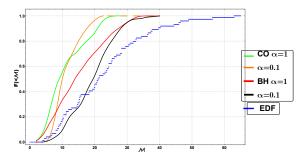
Model distribution  $F_{PBH}(< M)$  with parameters  $M_0 \approx 17 M_{\odot}$  and  $\gamma \sim 1$  for two best Kolmogorov-Smirnov tests. EDF= empirical distribution function.



Similar value of the parameters are obtained in M. Raidal et al, JCAP.,2019. Feb. V. 2019, no. 2. P. 018. arXiv:1812.01930 and L. Liu, et al arXiv:2210.16094. See also K. Postnov and N. Mitichkin, e-Print: 2302.06981.

# Chirp mass distribution

Cumulative distributions F(< M) for several astrophysical models of binary BH coalescences.



Conclusion: PBHs with log-normal mass spectrum perfectly fit the data. Astrophysical BHs seem to be disfavoured.

## Black Dark Matter

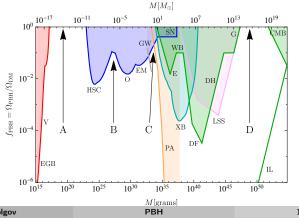
The first suggestion PBH might be dark matter "particles" was made by S. Hawking in 1971 "Gravitationally collapsed objects of very low mass", Mon. Not. R. astr. Soc. (1971) 152, 75 and repeated later by G. Chapline in 1975 who noticed that low mass PBHs might be abundant in the present-day universe with the density comparable to the density of dark matter. G.F. Chapline, Nature, 253, 251 (1975) "Cosmological effects of primordial black holes". Assumed flat mass spectrum in log interval:

 $dN = N_0 (dM/M)$ 

with maximum mass  $M_{max} \lesssim 10^{22}$  g, which hits the allowed mass range. The next one: DS (Mar 13, 1992), Baryon isocurvature fluctuations at small scales and baryonic dark matter, with more realistic masses. First paper with inflation applied to PBH formation, so PBH masses as high as  $10^6 M_{\odot}$ , and even higher can be created, log-normal mass spectrum was predicted.

#### Black Dark Matter

Constraints on PBHs - B.Carr, F. Kuhnel "Primordial Black Holes as Dark Matter: Recent Developments", arXiv:2006.02838, June 2020 Primordial black holes as dark matter candidates B. Carr, F. Kuhnel SciPost Phys.Lect.Notes 48 (2022), e-Print: 2110.02821 [astro-ph.CO] For monochromatic mass spectrum of PBHs (model-dependent and have caveats).



A.D. Dolgov

#### Figure caption

Constraints on f(M) for a **monochromatic** mass function, from evaporations (red), lensing (blue), gravitational waves (GW) (gray), dynamical effects (green), accretion (light blue), CMB distortions (orange) and large-scale structure (purple). Evaporation limits from the extragalactic gamma-ray background (EGB), the Voyager positron flux (V) and annihilation-line radiation from the Galactic centre (GC). Lensing limits from microlensing of supernovae (SN) and of stars in M31 by Subaru (HSC), the Magellanic Clouds by EROS and MACHO (EM) and the Galactic bulge by OGLE (O). Dynamical limits from wide binaries (WB), star clusters in Eridanus II (E), halo dynamical friction (DF), galaxy tidal distortions (G), heating of stars in the Galactic disk (DH) and the CMB dipole (CMB). Large scale structure constraints(LSS). Accretion limits from X-ray binaries (XB) and Planck measurements of CMB distortions (PA). The incredulity limits (IL) correspond to one PBH per relevant environment (galaxy, cluster, Universe). There are four mass windows (A, B, C, D) in which PBHs could have an appreciable density.

# Lifting bounds for extended mass spectrum

N. Bellomo, J.L. Bernal, A. Raccanelli, L. Verde, "Primordial Black Holes as Dark Matter: Converting Constraints from Monochromatic to Extended Mass Distributions", JCAP 01 (2018) 004 • e-Print: 1709.07467 [astro-ph.CO]. As it is well known, because of all the theoretical uncertainties, all constraints on the maximum allowed  $f_{PBH}$  have to be considered as order of magnitudes rather than exact numbers.

Similarly the behaviour of this window, where  $f_{PBH} \sim 1$  is allowed, for the considered Extended Mass Distribution (EMD) should not be thought as general, in fact it could well become wider and allow a larger  $f_{PBH}$  for other EMDs. A window that is closed for a Monochromatic Mass distribution (MMD) can open for an EMD, in particular for log-normal mass spectrum.

# Intermediate summary and antimatter in the Galaxy

The mechanism of AD and DKK solves the problem of the observed population of the universe at high redshifts by SMBH (QSO), galaxies, SN, and of a large amount of dust.

The predicted log-normal spectrum of PBH is tested and confirmed by the observations (the only one existing in the literature).

The existence of IMBH in GCs is confirmed.

The crazy by-product of AD and DKK mechanism, namely prediction of antimatter in the Galaxy seems to come true as well.

Astronomical data of the several recent years present strong evidence in favour of notieable antimatter population in our Galaxy including:

• Observation of gamma-rays with energy 0.511 MeV, which surely originate from electron-positron annihilation at rest.

• Very large flux of anti-helium nuclei, observed at AMS.

• Several stars are found which produce excessive gamma-rays with energies of several hundred MeV which may be interpreted as indication that these stars consist of antimatter.

# Antimatter history

Search for galactic antimatter

B.P. Konstantinov, et al Cosmic Research, 4, 66 (1968);

B.P. Konstantinov, et al Bulletin of the Academy of Sciences of the USSR. Physical series, 33, No,11, 1820 (1969). Strogly criticised by Zeldovich.

Antimatter int the universe:

F. W. Stecker, et al Possible Evidence for the Existence of Antimatter on a Cosmological Scale in the Universe, Phys. Rev. Letters 27, 1469 (1971);
F. W. Stecker, Grand Unification and possible matter-antimatter domain structure in the the universe. Tenth Texas Symposium on Relativistic Astrophysics, p. 69 (1981),

Summary of the situation presented at 2002:

F. W. Stecker, "The Matter-Antimatter Asymmetry of the Universe (keynote address for XIVth Rencontres de Blois)" arXiv:hep-ph/0207323.

A.D. Dolgov, "Cosmological matter antimatter asymmetry and antimatter in the universe", keynote lecture at 14th Rencontres de Blois on Matter - Anti-matter Asymmetry • e-Print: hep- ph/0211260.

Paul A.M. Dirac: "Theory of electrons and positrons", Nobel Lecture, December 12, 1933: "It is quite possible that... these stars being built up mainly of positrons and negative protons. In fact, there may be half the stars of each kind. The two kinds of stars would both show exactly the same spectra, and there would be no way of distinguishing them by present astronomical methods." It seems that now we know ways to distinguish stars from an antistars by observations from the Earth. A.D. Dolgov, V.A. Novikov, M.I. Vysotsky, "How to see an antistar" JETP Lett. 98 (2013) 519, e-Print: 1309.2746 The spectra are not exactly the same, even if CPT is unbroken and the polarization of radiation could be a good indicator or the type of emitted neutrinos/antineutrinos from supernovae.

# Antimatter history

Dirac was the second person to talk about antimatter. In 1898, 30 years before Dirac and one year after discovery of electron (J.J. Thomson, 1897) Arthur Schuster (another British physicist) conjectured that there might be other sign electricity, ANTIMATTER, and supposed that there might be entire solar systems, made of antimatter, INDISTINGUISHABLE from ours.

Schuster's wild guess: matter and antimatter are capable to annihilate and produce VAST energy.

He believed that they were gravitationally repulsive having negative mass. Two such objects on close contact should have vanishing mass!?

A. Schuster, Nature, 58 (1898) 367. Potential Matter. Holiday Dream.

"When the year's work is over and all sense of responsibility has left us, who has not occasionally set his fancy free to dream about the unknown, perhaps the unknowable?"

"Astronomy, the oldest and yet most juvenile of the sciences, may still have some surprises in store. May antimatter be commended to its case".

### Antimatter in the Galaxy

Based on the conventional approach no antimatter object is expected to be in the Galaxy.

However, it was predicted in 1993 and elaborated in 2009 that noticeable amount of antimatter, even antistars might be in the Galaxy and in its halo:

A. Dolgov, J.Silk, PRD 47 (1993) 4244 "Baryon isocurvature fluctuations at small scale and baryonic dark matter.

A.Dolgov, M. Kawasaki, N. Kevlishvili, Nucl.Phys. B807 (2009) 229, "Inhomogeneous baryogenesis, cosmic antimatter, and dark matter". Bounds on the density of galactic antistars are rather loose, because the annihilation proceeds only on the surface of antistars as analyzed in: C.Bambi, A.D. Dolgov, "Antimatter in the Milky Way", Nucl.Phys.B 784 (2007) 132-150 • astro-ph/0702350,

A.D. Dolgov, S.I. Blinnikov, "Stars and Black Holes from the very Early Universe", Phys.Rev.D 89 (2014) 2, 021301 • 1309.3395,

S.I.Blinnikov, A.D., K.A.Postnov, "Antimatter and antistars in the universe and in the Galaxy", Phys.Rev.D 92 (2015) 023516 • 1409.5736.

# Anti-evidence: cosmic positrons

Observation of intense 0.511 line, a proof of abundant positron population in the Galaxy. In the central region of the Galaxy electron–positron annihilation proceeds **at a surprisingly high rate**, creating the flux:

$$\Phi_{511 \text{ keV}} = 1.07 \pm 0.03 \cdot 10^{-3} \text{ photons cm}^{-2} \text{ s}^{-1}.$$

The width of the line is about 3 keV. Emission mostly goes from the Galactic bulge and at much lower level from the disk,

### "Great Annihihilator" in the Galactic bulge.

- G. Weidenspointner et al., Astron. Astrophys. 450, 1013 (2006);
- J. Knodlseder et al., Astron. Astrophys. 441, 513 (2005);
- P. Jean et al., Astron. Astrophys. 445, 579 (2006).

Until recently the commonly accepted explanation was that  $e^+$  are created in the strong magnetic fields of pulsars but the recent results of AMS probably exclude this mechanism, since the spectrum of  $\bar{p}$  and  $e^+$  at high energies are identical. L'Aquila Joint Astroparticle Colloquium, 10th November, 2021 by S. Ting.

# Anti-evidence: cosmic antinuclei

Registration of anti-helium: In 2018 AMS-02 announced possible observation of six  $\overline{He}^3$  and two  $\overline{He}^4$ .

A. Choutko, AMS-02 Collaboration, "AMS Days at La Palma, La Palma, Canary Islands, Spain," (2018).

S. Ting, Latest Results from the AMS Experiment on the International Space Station. Colloquium at CERN, May, 2018.

Recent registration of more events L'Aquila Joint Astroparticle Colloquium, 10th November by S. Ting; and COSPAR 2022, 16-24 July:

7  $\overline{D}$  ( $\lesssim 15$  GeV) and 9  $\overline{He}$ , ( $\sim 50$  GeV). fraction  $\overline{He}/He \sim 10^{-9}$ , too high.

Secondary creation of  $\overline{He}^4$  is negligibly weak.

Nevertheless S. Ting expressed hope to observe  $\overline{Si}$  !!!

It is not excluded that the flux of anti-helium is even much higher because low energy  $\overline{He}$  may escape registration in AMS.

# Deuterium/Helium problem

There is noticeable discrepancy between the large fraction of D with respect to He. In the case of the standard BBN this ratio should be much smaller than unity, but the observed one is practically 1.

It is assumed that the abundances of D and He are determined by BBN with large  $\beta$  (or  $\eta$ ). However if  $\beta \sim 1$  there is no primordial D. On the other hand in our scenario formation of primordial elements takes place inside non-expanding compact stellar-like objects with fixed temperature. If the temperature is sufficiently high, this so called BBN may stop before abundant He formation with almost equal abundances of D and He. One can see that looking at abundances of light elements at a function of temperature. Is it is so, antistars may have equal amount of  $\overline{D}$  and  $\overline{He}$ !!!

• Antistars: Several (14?) stars are found which produce excessive gamma-rays with energies of hundreds MeV which may come from  $p\bar{p}$  annihilation. S. Dupourqué, L. Tibaldo, P. von Ballmoos, Phys Rev D.103.083016 103 (2021). Very powerful gamma-ray burster (GRB) was reported in M.E. Ravasio, O.S. Salafia, G. Oganesyan, *et al*, SCIENCE, 25 Jul 2024, **385**, Issue 6707, p. 452; 2303.16223 [astro-ph.HE]. This event got the nickname: the Brightest Of All Time or the **BOAT**. A bright megaelectronvolt emission line was observed, Usually the gamma-ray spectra of GRBs consist of a smooth continuum without absorption or emission lines. The authors interpret this line as a result of the annihilation of electron-positron pairs within the relativistic jet produced by the GRB possibly emerging from star-antistar annihilation !?

# X-ray signatures of antistars

#### X-ray signature of antistars in the Galaxy A.E. Bondar, S.I. Blinnikov, A.M. Bykov, A.D. Dolgov, K.A. Postnov e-Print: 2109.12699 [astro-ph.HE], JCAP, Sep 26, 2021,

In astrophysically plausible cases of the interaction of neutral atmospheres or winds from antistars with ionised interstellar gas, the hadronic annihilation **will be preceded by the formation of excited**  $p\bar{p}$  and  $He\bar{p}$  atoms. These atoms rapidly cascade down to low levels prior to annihilation giving rise to a series of narrow lines which can be associated with the hadronic annihilation gamma-ray emission. The most significant are L (3p-2p) 1.73 keV line (yield more than 90%) from  $p\bar{p}$  atoms, and M (4-3) 4.86 keV (yield ~ 60%) and L (3-2) 11.13 keV (yield about 25%) lines from  $He^4\bar{p}$  atoms. These lines can be probed in dedicated observations by forthcoming sensitive X-ray spectroscopic missions XRISM and Athena and in wide-field X-ray surveys like SRG/eROSITA all-sky survey.

# Antihelium and antistars

A.M. Bykov, K.A. Postnov, A.E. Bondar, S.I. Blinnikov, A.D. Dolgov, "Antistars as possible sources of antihelium cosmic rays", JCAP 08 (2023) 027 • e-Print: 2304.04623 [astro-ph.HE].

Possible sources of antinuclei in cosmic rays from antistars which are predicted in a modified Affleck-Dine baryogenesis scenario by DS (1993) are discussed. The expected fluxes and isotopic content of antinuclei in the GeV cosmic rays produced in scenarios involving antistars are estimated. It is shown that the flux of antihelium cosmic rays reported by the AMS-02 experiment can be explained by Galactic anti-nova outbursts, thermonuclear anti-SN Ia explosions, a collection of flaring antistars, or an extragalactic source with abundances not violating existing gamma-ray and microlensing constraints on the antistar population. SUSY motivated baryogenesis, Affleck and Dine (AD).

SUSY predicts existence of scalars with  $B \neq 0$ . Such bosons may condense along flat directions of the quartic potential:

 $U_{\lambda}(\chi) = \lambda |\chi|^4 (1 - \cos 4\theta)$ 

and of the mass term,  $U_m = m^2 \chi^2 + m^{*\,2} \chi^{*\,2}$ :

$$\boldsymbol{U}_{\boldsymbol{m}}(\boldsymbol{\chi}) = \boldsymbol{m}^2 |\boldsymbol{\chi}|^2 [1 - \cos\left(2\theta + 2\alpha\right)],$$

where  $\chi = |\chi| \exp(i\theta)$  and  $m = |m|e^{\alpha}$ . If  $\alpha \neq 0$ , C and CP are broken. In GUT SUSY baryonic number is naturally non-conserved - non-invariance of  $U(\chi)$  w.r.t. phase rotation. Initially (after inflation)  $\chi$  is away from origin and, when inflation is over, starts to evolve down to equilibrium point,  $\chi = 0$ , according to Newtonian mechanics:

$$\ddot{\chi} + 3H\dot{\chi} + U'(\chi) = 0.$$

Baryonic charge of  $\chi$ :

$$B_{\chi} = \dot{\theta} |\chi|^2$$

is analogous to mechanical angular momentum.  $\chi$  decays transferred baryonic charge to that of quarks in B-conserving process.

AD baryogenesis could lead to baryon asymmetry of order of unity, much larger than the observed  $10^{-9}$ .

# Creation Mechanism

If  $m \neq 0$ , the angular momentum, B, is generated by a different direction of the quartic and quadratic valleys at low  $\chi$ . If CP-odd phase  $\alpha$  is small but non-vanishing, both baryonic and antibaryonic domains might be formed with possible dominance of one of them.

Matter and antimatter objects may exist but globally  $B \neq 0$ .

Affleck-Dine field  $\chi$  with CW potential coupled to inflaton  $\Phi$  (AD and Silk; AD, Kawasaki, Kevlishvili):

$$U = g|\chi|^2 (\Phi - \Phi_1)^2 + \lambda|\chi|^4 \ln\left(\frac{|\chi|^2}{\sigma^2}\right)$$
$$+\lambda_1(\chi^4 + h.c.) + (m^2\chi^2 + h.c.).$$

Coupling to inflaton is the general renormalizable one. When the window to the flat direction is open, near  $\Phi = \Phi_1$ , the field  $\chi$  slowly diffuses to large value, according to quantum diffusion equation derived by Starobinsky, generalized to a complex field  $\chi$ . If the window to flat direction, when  $\Phi \approx \Phi_1$  is open only during a short period, cosmologically small but possibly astronomically large bubbles with high  $\beta$  could be created, occupying a small fraction of the universe, while the rest of the universe has normal  $\beta \approx 6 \cdot 10^{-10}$ , created by small  $\chi$ . The mechanism of massive PBH formation quite different from all others. The fundament of PBH creation is build at inflation by making large isocurvature fluctuations at relatively small scales, with practically vanishing density perturbations.

Initial isocurvature perturbations are in chemical content of massless quarks. Density perturbations are generated rather late after the QCD phase transition. The mechanism is very much different from other conventionl ones.

The emerging universe looks like a piece of Swiss cheese, where holes are high baryonic density objects occupying a minor fraction of the universe volume.

### Results

- PBHs with log-normal mass spectrum confirmed by "experiment"!
- Strange stars in the Galaxy with unusual chemistry and velocity.
- Disperse hydrogen and helium clouds with (much) higher than average n<sub>B</sub> density.
- $\beta$  may be negative leading to creation of (compact?) antistars which could survive annihilation with the homogeneous baryonic background.
- Extremely old stars would exist even, "older than universe star" is found; the older age is mimicked by the unusual initial chemistry. Several too old stars are observed.

The mechanism of PBH creation pretty well agrees with the data on the mass spectrum and on existence of antimatter in the Galaxy, especially of antistars. So we may expect that it indeed solves the problems discovered by HST and JWST.

Primordial black holes remove the tension between ∧CDM cosmology and astronomical data, that exists during all the history of the universe.

The model of their formation predicts and explains the origin of the observed antimatter population of the Milky Way