Plasma Wakefield Acceleration

The long and winding road from proof-of-principle experiments to colliders

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Particle Accelerators

"Livingston" plot of evolution of accelerators



A. Seryi, Unifying physics of accelerators, lasers and plasma

- → New technology allows for a jump in energy and in applications
- \rightarrow Have we reached saturation of RF technology?
- → Can we reach higher energies in shorter distance? Next HEP machine will probably be:
 - Higgs factory (center of mass>250 GeV)
 - Discovery machine with center of mass >>TeV

 \rightarrow This is where plasma comes into play!

Plasma

What we talk about, when we talk about **plasma**:

- Ionized medium (normally a gas)
- Collisions can be (most of time) neglected
 → Electromagnetic interaction dominates
- Large number of particles → collective behavior
- Quasi-neutral $(n_{pe} \sim n_{pi})$



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neutral plasma		

Plasma – Space Charge Screening

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 - \rightarrow It tends to keep the charge and current neutrality:
 - Plasma electrons (m_{pi}>>m_{pe}) move to compensate for the disturbance
 - → Plasma screens electromagnetic fields



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When the equilibrium is perturbed:

- Electrons oscillate with angular frequency $\omega_{pe} = \sqrt{\frac{n_{pe}e^2}{m_e\varepsilon_0}}$
- Ions with $\omega_{pi} = \sqrt{\frac{n_{pi}e^2}{m_i\varepsilon_0}} \ll \omega_{pe}$ (ions considered immobile for short time-scales)



- Let's take a plasma with density n_{pe}
- Let's take a relativistic charged bunch (e.g. e⁻) with density $n_b << n_{pe}$ (roughly the same for laser pulses)
- lack of electrons

(inspired by P. Muggli's CAS lecture)

1. Transverse E field expels plasma electrons

- Let's take a plasma with density npe
- Let's take a relativistic charged bunch (e.g. e^{-}) with density $n_b < < n_{pe}$



1. Transverse E field expels plasma electrons

 Positively charged region behind the bunch head
 → restoring force

(inspired by P. Muggli's CAS lecture)

Linear regime: plasma electrons DO NOT cross longitudinal axis: PERTURBATION!! Blowout (non-linear) regime: electrons DO cross the axis

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- 3. Oscillation of plasma e⁻ with ω_{pe} \rightarrow periodic density variation

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$$\nabla \cdot \vec{E} = \frac{\rho}{\varepsilon_0}$$

→ Wakefields ←

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- Wake travels at driver's velocity (no dephasing using relativistic bunches) ٠



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→ Wakefields ←

 $W_{\perp}(\xi,r) = \frac{-n_{b0}q}{\varepsilon_0 k_{pe}} \int_{-\infty}^{\xi} n_{b||}(\xi') \sin(k_{pe}(\xi-\xi')) d\xi' \cdot \frac{dR(r)}{dr}$

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 \rightarrow Wakefields \leftarrow

$$E_z(\xi,r) = rac{n_{b0}q}{arepsilon_0} \int_{-\infty}^{\xi} n_{b||}(\xi') \cos(k_{pe}(\xi-\xi')) d\xi' \cdot R(r)$$
 $W_{\perp}(\xi,r) = rac{-n_{b0}q}{arepsilon_0 k_{pe}} \int_{-\infty}^{\xi} n_{b||}(\xi') \sin(k_{pe}(\xi-\xi')) d\xi' \cdot rac{dR(r)}{dr}$

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 $V_{\rm b} \sim c$



Accelerating Gradient

- Fields in plasmas are sustained by the charge separation
 - As high as the cold wave-breaking field: $E_{WB} = \frac{m_e c \, \omega_{pe}}{a} \rightarrow \text{oscillation length cannot exceed plasma wavelength}$
 - E.g. for $n_{pe} = (10^{14} 10^{18}) \text{ cm}^{-3}$, $E_{WB} \sim 100 \frac{V}{m} \sqrt{n_{pe} [cm^{-3}]} = (1 100 \text{ GV}/\text{m})$



Wave «breaks» when the maximum amplitude is reached

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Wave «breaks» when the maximum amplitude is reached

• RF cavities limited to 100MV/m by breakdown, caused e.g. by fatigue, pulse heating, etc..

one could dream of shrinking down the size of accelerators by orders of magnitude

Accelerating Gradient – Experimental Results

Laser Wakefield Acceleration (LWFA)

VOLUME 43, NUMBER 4 PHYSICAL REVIEW LETTERS

23 JULY 1979

Laser Electron Accelerator

T. Tajima and J. M. Dawson Department of Physics, University of California, Los Angeles, California 90024 (Received 9 March 1979)

Beam-Driven Plasma Wakefield Acceleration (PWFA)

VOLUME 54, NUMBER 7

PHYSICAL REVIEW LETTERS

18 FEBRUARY 1985

Acceleration of Electrons by the Interaction of a Bunched Electron Beam with a Plasma

Pisin Chen^(a) Stanford Linear Accelerator Center, Stanford University, Stanford, California 94305

and

J. M. Dawson, Robert W. Huff, and T. Katsouleas Department of Physics, University of California, Los Angeles, California 90024 (Received 20 December 1984)

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Driver: relativistic charged particle bunch





~ 42 GeV in 85 cm \rightarrow ~50GV/m $n_{pe} = 2.8 \times 10^{17} \text{ cm}^{-3}$



Accelerating Gradient – Experimental Results



PWFA – First Demonstration

First experimental demonstration: 1988, Argonne National Laboratory (US)





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 VOLUME 61, NUMBER 1
 PHYSICAL REVIEW LETTERS
 4 JULY 1988

 Experimental Observation of Plasma Wake-Field Acceleration

 J. B. Rosenzweig, D. B. Cline, ^(a) B. Cole, ^(b) H. Figueroa, ^(c) W. Gai, R. Konecny, J. Norem, P. Schoessow, and J. Simpson

 High Energy Physics Division, Argonne National Laboratory, Argonne, Illinois 60439 (Received 21 March 1988)

 We report the first experimental test of the physics of plasma wake-field acceleration performed at the Argonne National Laboratory Advanced Accelerator Test Facility. Megavolt-per-meter plasma wake fields are excited by a intense 21-MeV, multipiscosecond bunch of electrons in a plasma of density

 $n_r \approx 10^{13}$ cm⁻³, and probed by a low-intensity 15-MeV witness pulse with a variable delay time behind the intense bunch. Accelerating and deflecting wake-field measurements are presented, and the results

compared to theoretical predictions.



- Measurement of Witness energy as a function of delay
- Sinusoidal
- Linear regime
 - \rightarrow small gradient
 - ightarrow hard to preserve beam quality

PWFA — Non-linear Regime

Most of PWFA's work in the non-linear blowout regime:

- High gradient
- Linear focusing force



$$n_b \gg n_{pe}$$

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PWFA — Non-linear Regime

Most of PWFA's work in the non-linear *blowout* regime:

- High gradient
- Linear focusing force
- \rightarrow Main challenge: beam quality preservation (vital for applications)



Requirement:

 $n_b \gg n_{pe}$

Main challenge – Energy spread (longitudinal quality)





J. B. Rosenzweig et al., Phys. Rev. A 44, R6189(R) (1991)

- Accelerating field not uniform along the bubble
 → Energy spread increasing upon acceleration
- Solution: "*loading*" the wake with the presence of the witness bunch itself



Uniform accelerating field within the witness bunch

M. Tzoufras et al., PRL 101, 145002 (2008)

Main challenge – Energy spread (longitudinal quality)

→ Experimental Demonstrations:



Combination of beam loading and initial chirp to obtain final small energy spread

 $L_p = 3 \text{ cm}$

25

z (μm)

25

40 30 20

50

Check for updat

→ Ion column provides linear focusing force

Radial electric field: $E_r(r) = \frac{en_{pe}}{2\varepsilon_0} r$ (Gauss' law on cylinder of ions)



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→ Plug it in envelope equation:
$$\sigma_r''(z) + \sigma_r(z) \begin{pmatrix} K & -\frac{\epsilon_g^2}{\sigma_r^4(z)} \end{pmatrix} = 0$$

Equilibrium
between focusing
force and
emittance
Matching condition

Too large

Too small

20

laser beam

dump

10

z [mm]

 $\sigma = 1.2\sigma_{matche}$

30

→ Ion column provides linear focusing force

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Equilibrium
between focusing =0:

force and Matching condition emittance

35

vacuum

window

-10

Matching Conditions:

•
$$\beta = \frac{\sigma^2(0)}{\epsilon_a} = \sqrt{\frac{2\epsilon_0 m_e c^2 \gamma}{n_{ne} e^2}}$$

• Injection at waist:
$$\sigma'(z=0)=0$$

30 25 20 [m] מ' [

(L. Verra et al 2020 J. Phys.: Conf. Ser. 1596 012007)

if: beam envelope is matched to the focusing force (which is extremely strong!)

 \rightarrow Possible emittance preservation

else: different energy slices rotate at different rates in transverse phase space

 \rightarrow slice emittance preserved (linear focusing)

 \rightarrow projected (i.e., overall) emittance grows!

Grating

• Direct experimental demonstration:



"Indirect" experimental proof:
 Quality good enough for free-electron lasing



Galletti et al., PRL 129, 234801 (2022) Pompili et al. Nature 605, 659–662 (2022)

Main Challenges

The main challenge remains:

• Do everything at the same time



Energy spread minimization



Main Challenges

The main challenge remains:

• Do everything at the same time



Energy spread minimization



• Do it many times \rightarrow high repetition rate



• etc..

EuPRAXIA@SPARC_LAB

Single-stage high-quality high energy gain, high repetition rate PWFA

- The most challenging application of single-stage PWFA: free-electron-laser
 - High charge
 - Low emittance
 - Low energy spread
 - High shot-to-shot reproducibility
 - Tunability
- Deliver radiation in water window (2-4 nm) for users

[®] User area

Undulators

Witness bunch generates radiation through FEL process

Plasma module

Witness boosted to 1 GeV (energy doubling)

X-band linac boost to 500 MeV

S-band injector: producing driver and witness bunches ~150 MeV

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First beam expected end of 2029!

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PWFA – Positrons

→ Positrons may be needed for future lepton colliders In principle, just a π phase difference in the wakefields



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Long Bunches: Head and center lose energy



B. E. Blue, ¹ C. E. Clayton, ¹ C. L. O'Connell, ² F.-J. Decker, ² M. J. Hogan, ² C. Huang, ¹ R. Iverson, ² C. Joshi, ¹ T. C. Katsouleas, ³ W. Lu, ¹ K. A. Marsh, ¹ W. B. Mori, ¹ P. Muggli, ³ R. Siemann, ² and D. Walz²



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Long Bunches: Head and center lose energy Tail gains energy



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- → Positrons may be needed for future lepton colliders In principle, just a π phase difference in the wakefields
- ightarrow Acceleration demonstrated also in the non-linear regime

doi:10.1038/nature14890

Multi-gigaelectronvolt acceleration of positrons in a self-loaded plasma wakefield

S. Corde^{1,2}, E. Adli^{1,3}, J. M. Allen¹, W. An^{4,5}, C. I. Clarke¹, C. E. Clayton⁴, J. P. Delahaye¹, J. Frederico¹, S. Gessner¹, S. Z. Green¹, M. J. Hogan¹, C. Joshi⁴, N. Lipkowitz¹, M. Litos¹, W. Lu⁶, K. A. Marsh⁴, W. B. Mori^{4,5}, M. Schmeltz¹, N. Vafaei-Najafabadi⁴, D. Walz¹, V. Yakimenko¹ & G. Yocky¹



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(P. Muggli, CAS 2014)

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→ BUT: In the blowout regime, e⁺ witness bunches need to be placed very close to the singularity → Some creative solutions were proposed:



→ Acceleration in hollow plasma (avoiding focusing force on axis) Gessner et al., Nat. Comm. 7, 11785 (2016)



→ But tight alignment tolerance to avoid transverse instabilities Lindstrøm et al., PRL 120, 124802 (2018)

→ BUT: In the blowout regime, e⁺ witness bunches need to be placed very close to the singularity
 → Some creative solutions were proposed:



PWFA – Muons

Acceleration of muons is possible in principle, but:

- if injected with v<<c: need for a slow (non-relativistic) driver
 → dephasing upon acceleration (witness gets closer to the driver)
- Possible solution: tapered plasma density profile

- Initial down ramp to slow down the wake
- Constant density region after $v_{\mu} \sim v_{d}$
- Up ramp region to mitigate dephasing



C. Badiali (IST), EAAC 2023

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C. Badiali (IST), EAAC 2023

But: need a muon beam to test..

PWFA — Very high energy

• The maximum net energy gain of the witness bunch is in general limited to:

 $\Delta E \sim 2 \times E_d$ (and for energy conservation $Q_W \Delta E_W \leq Q_D \Delta E_D$)

e.g: 1 GeV drive bunch drives 1 GV/m accelerating field

- → Depleted after 2m (500MV/m decelerating field)
- → Maximum energy gain of witness: 2GeV



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To reach VERY high energy (e.g. 5 TeV)

• One highly energetic driver to accelerate witness in a single stage (afterburner):

Assuming energy tripling of witness bunch → need a 1.7 TeV driver and witness couple → CLIC-scale main linac





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Many low-energy (cheaper) drivers in multiple stages

Using e.g. 30 GeV drive bunches (requiring a > 300 m-long linac)

- $\rightarrow \Delta E < 60 \text{ GeV}$ (R<2 for single symmetric driver)
- → Need > 83 stages



PWFA – Staging

• The most outstanding challenge in PWFA:



J. Rosenzweig et al., NIM A 410 (1998) 532—543

In each of them:

- Inject the driver bunch
- Inject and match the witness bunch with appropriate beam loading
- Accelerated with high gradient and high net energy gain
- Extract (i.e., reduce divergence)
- Dispose of spent driver

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From HALHF design Foster et al 2023 New J. Phys. **25** 093037

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- ightarrow Preserve the beam quality all along

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Basically, repeating EuPRAXIA n times..

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Basically, repeating EuPRAXIA n times..

SPARC_LAB and EuPRAXIA could play a major role on this topic!

PWFA — Avoiding staging

- Alternative: single long stage with extremely energetic driver:
 → Proton bunches from synchrotrons
- Available proton bunches carry large amounts of energy:
 - CERN SPS proton bunch: $3 \cdot 10^{11}$ p⁺ at 400 GeV/c \rightarrow 19.2 kJ
 - CERN LHC proton bunch: $1 \cdot 10^{11} \text{ p}^+$ at 7 TeV/c \rightarrow 112 kJ
 - SLAC FFTB electron bunch: $7 \cdot 10^9 e^-$ at 40 GeV/c \rightarrow 40 J

⇒ Drive wakefields over very long distance! no need for staging





A. Caldwell et al., Nature Phys. 5, 363–367 (2009)

Parameters: single proton bunch $\sigma_z = 100 \,\mu\text{m},$ E = 1 TeV, population: 1.10¹¹ particles per bunch (16nC)

PWFA — Avoiding staging

- Alternative: single long stage with extremely energetic driver:
 → Proton bunches from synchrotrons
- → Acceleration of electrons demonstrated at AWAKE (using 400 GeV p+ from SPS and 10-m-long plasma source)
- \rightarrow But: rely on beam-plasma instability to drive large amplitude wakefields





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L. Verra et al. (AWAKE Collaboration) Phys. Rev. Lett. 129, 024802 (2022)

- ightarrow Difficult to achieve collider-quality beams
- ightarrow Applications for fixed target experiments



- Need to reach high center-of-mass energy (100's GeV or multi-TeV) (relevant for the target application/process to investigate)
 - Beam Species \rightarrow impact on the physics and on the statistics



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- Center-of-mass energy (relevant for the target application/process to investigate)
 - Beam Species \rightarrow impact on the physics and on the statistics
- Luminosity

(gathering enough data in reasonable amount of time):



- Center-of-mass energy (relevant for the target application/process to investigate)
 - Beam Species \rightarrow impact on the physics and on the statistics
- Luminosity (gathering enough data in reasonable amount of time):
- Luminosity per power
 - Better metrics to quantify the luminosity one can "buy"

Efficiency:

- Wall-plug → Drive beam (Klystrons, etc..)
 ~55% (CLIC)
 M. Aicheler et al., CLIC CDR (2012)
- Drive Beam → Plasma ~60% with Gaussian bunch F. Peña et al., Phys. Rev. Res. 6, 043090 (2024)

Plasma → Witness ~22% preserving quality

Lindstrøm et al., Nat. Comm. 15, 6097 (2024)

Which reads:

- High quality
- High efficiency

 $4\pi\sigma_x\sigma_v E_b$

 P_{tot}

Bunch energy

- Center-of-mass energy (relevant for the target application/process to investigate)
 - Beam Species \rightarrow impact on the physics and on the statistics
- Luminosity (gathering enough data in reasonable amount of time):
- Luminosity per power
 - Better metrics to quantify the luminosity one can "buy"



Which reads:

- High quality
- High efficiency

$$\frac{\mathcal{L}}{P_{tot}} = \frac{\eta N}{4\pi\sigma_x\sigma_y E_b}$$

HALHF: a hybrid, asymmetric collider concept

Plasma acceleration for electrons + RF acceleration for positrons

> Solving the plasma positron problem by accelerating positron with RF linacs.



> Length dominated by the beam-delivery system. Cost dominated by the RF linac.

UNIVERSITY OF OSLO 27 May 2025 | Carl A. Lindstrøm | 10 TeV collider monthly meeting

An asymmetric collider: can it work?

The more asymmetric, the better



UNIVERSITY OF OSLO 2

Lessons learned from HALHF

Based on Bayesian optimisation of a detailed collider physics+cost model

- > Cost of power dominates (not length)
 - > Must design the plasma accelerator for the driver, not vice versa
- > Lower density is greatly favoured
 - > Little need for gradients beyond 1 GV/m (for sub-TeV machines)
 - > Suppresses further beam ionisation of plasma
 - > Requires high-charge beams (multi-nC)
- > Maximise the *effective* transformer ratio (transformer ratio × number of stages)
- > Ion motion is required for transverse stability
- > HALHF bunch trains have the potential to heat plasma to O(100 keV) temps

> The next key R&D issue will be plasma heating, cooling, and confinement

OXFORD 19th Feb 2025 | R. D'Arcy, C.A. Lindstrøm | 10 TeV Plasma Linac WG Monthly Meeting

Cost estimation, to be compared with:

- FCC-ee (~20B€)
- ILC-CLIC (~7-12B€)

<u>arXiv:2503.19880</u>

 Response to 2023 Snowmass P5 Report (analogue to ESPP in USA)

Recommendation 4: Support a comprehensive effort to develop the resources—theoretical, computational, and technological—essential to our 20-year vision for the field. This includes an aggressive R&D program that, while technologically challenging, could yield revolutionary accelerator designs that chart a realistic path to a <u>10 TeV pCM</u> collider.

Investing in the future of the field to fulfill this vision requires the following:

a. Support vigorous R&D toward a cost-effective 10 TeV pCM collider based on proton, muon, or possible wakefield technologies, including an evaluation of options for US siting of such a machine, with a goal of being ready to build major test facilities and demonstrator facilities within the next 10 years (sections 3.2, 5.1, 6.5, and Recommendation 6).

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L. Xu, T. Opferkuch, I. Savoray, C. Scherb, S. Chigusa, SK in progress

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- Goal: present an end-to-end design concept with consistent parameters
- Evaluating the best acceleration technique and investigating new aspects as:
 - beam-beam interaction
 - \rightarrow e.g. flat or round beams?
 - environmental impact
 - Beam delivery system
 - STAGING
 - Full-program cost \rightarrow is it really competitive?

- System integration and optimization
- Beam sources (incl. damping rings)
- Drivers
 - Laser
 - Beams SWFA
 - Beams PWFA
- Linacs
 - LWFA
 - SWFA
 - PWFA
- Beam delivery system
- Beam-beam interactions
- Beam diagnostics
- Machine-detector interface
- HEP detector
- HEP physics case
- Environmental impact
- Simulations/computing/Al

Contributions in any of these areas are welcome!

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• Submitted a contribution to ESPP Update

https://arxiv.org/abs/2503.20214

10TeV WFA indico page: https://indico.slac.stanford.edu/category/138/

Advertised by Nature:

https://www.nature.com/articles/d41586-025-01181-1

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• Starting working on the "details" to design a real machine:

e.g. energy loss due to synchrotron radiation in chicanes between each stage

- → Reducing the effective gradient (energy / full length of the linac)
- \rightarrow Length of optics increases with energy

- ightarrow Lower B field for longer optics
- → W_{eff} quickly goes <0.5GV/m if length of optics is increased</p>

Summary

• Plasma Wakefield Acceleration is a vibrant research field

- Long promised revolutionary applications
- Time to deliver!
 - Single-stage: FEL (EuPRAXIA)
 - Multi-stage: fixed target, SFQED, colliders
 - Lots of physics to investigate along the way

ALEGRO Workshop 2026

- The ALEGRO (Advanced LinEar collider study GROup) Workshop gathers the advanced and novel accelerator community and reflects the global ambition towards ultra-high energy colliders driven by wakefield acceleration technology, while <u>also seeking engagement from HEP Theorists and Experimentalists.</u>
- We will host the next iteration at LNF in 2026! https://agenda.infn.it/event/47329/
- Presentations and discussions on:
 - Beam Physics
 - Collider Physics Case
 - Other applications
 - etc..
- You are all invited!

Thank you for listening!
Backup Slides

List of not mentioned issues

Instabilities Jitter Tolerances Rep rate Heat resistance plasma sources (plasma generation + beam power deposition)

PWFA — Non-linear Regime

When the electric field of the bunch is strong enough to expel ALL plasma electrons
 → BUBBLE of plasma electrons around a column of pure ions





Requirement:

Along ξ :

- Periodic ''Steepened'' accelerating field
- Uniform focusing field

PWFA — Non-linear Regime

When the electric field of the bunch is strong enough to expel ALL plasma electrons
 → BUBBLE of plasma electrons around a column of pure ions



Along r (behind the bunch):

- Uniform accelerating field \rightarrow uniform acceleration
- Linear focusing force \rightarrow possible emittance preservation



Requirement:

Along ξ :

- Periodic "Steepened" accelerating field
- Uniform focusing field

PWFA — Beam Loading

→ BEAM LOADING:

The presence of the witness bunch affects the wakefields

Linear regime



Non-linear regime

The point is: compromise on accelerating gradient \rightarrow smaller energy spread \rightarrow beam quality

II. Non-linear Regime – Beam Loading

→ BEAM LOADING:

The presence of the witness bunch affects the wakefields Linear regime



Particle Accelerators, 1987, Vol. 22, pp. 81–99 Photocopying permitted by license only © 1987 Gordon and Breach Science Publishers, Inc. Printed in the United States of America

BEAM LOADING IN PLASMA ACCELERATORS

T. KATSOULEAS, S. WILKS, P. CHEN,[†] J. M. DAWSON and J. J. SU Department of Physics, University of California, Los Angeles, CA 90024



• Short Gaussian placed at the right phase can work 78

Ob. Beam Physics

- Relativistic particle bunches:
 - Propagate at v_b~c
 - Not affected by index of refraction
 - Large inertia (γm >> m)

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Ob. Beam Physics

- Relativistic particle bunches:
 - Propagate at v_b~c
 - Not affected by index of refraction
 - Large inertia (γm >> m)

- In Lab frame, space-charge electric field is almost purely transverse:
 - Effectively sets in motion the plasma electrons



$$\sigma'' = \frac{\epsilon_{rms}^2}{\sigma^3}$$

$$\sigma(z) = \sqrt{\left(\sigma_0 + \sigma_0'(z - z_0)\right)^2 + \frac{\epsilon_{rms}^2}{\sigma_0^3}(z - z_0)^2}.$$



EuPRAXIA@SPARC_LAB

Radiation Parameter	Unit	PWFA	Full X-band	iul	Electron Beam Parameter	Unit	PWFA	Full X-band
Wavelength	nm	3-5	4		Electron Energy	GeV	1-1.2	1.2
Pulse length	fc	10.0		-	Bunch Charge	рС	30 - 50	200-500
(fwhm)	15	10.0	-		Peak Current	kA	~ 2.2	1-2
Photons per Pulse	$\times 10^{12}$	0.1- 0.25	1		RMS Energy Spread	%	< 1	0.1
Photon	%	0.1	0.5	1	RMS Bunch Length	μ m	3-6	24-20
Bandwidth Undulator Area	m	2	6		RMS norm. Emittance	μ m	0.7 – 1.2	1
Length $ ho(1D/3D)$	× 10 ⁻³	1	1		Slice Energy Spread	%	≤0.05	≤0.05
Photon	$\begin{pmatrix} s mm^2mrad^2 \\ bw(0.1\%) \end{pmatrix}$	$1-2 \times 10^{28}$	1×10^{27}		Slice norm Emittance	mm-mrad	0.5 – 0.8	0.5
shot					Energy jitter	%	< 1	0.1

• Bold values indicate the main working point

Application Area	Scientific Focus	Techniques	Key Impact
Renewable Energy	 Charge transport in solar cells & catalysts Photocatalytic H₂ production PFAS/PCIAS analysis 	XAS, XES, PI-MS	 Efficient solar materials Clean hydrogen production Environmental remediation
Warm Dense Matter (WDM)	 Extreme temperature/density Astrophysics & fusion 	Time-resolved XAS, TR-XES	Stellar/planetary modelingFusion research support
Battery Technology	 Ion migration & interfaces Solid-state battery study 	XAS, non-linear X-ray spectroscopy	Safer, longer-life batteriesHigh energy density
Structural Biology	 Live-cell imaging Cellular processes (stress, DNA damage) 	CDI, XAS, CEI	 Real-time biomolecular studies Disease mechanism insight
Health & Radiobiology	 DNA damage by radiation 	X-ray pump– probe, photo- fragmentation	 Radioprotection Cancer treatment improvement
Atmospheric Chemistry	VOC/NOx oxidationAerosol dynamics	PI-MS, ion spectroscopy	Pollution modelsClimate studies
Astrochemistry	 Radiation chemistry in space Organic molecule formation 	XAS, photochemistry	 Space molecular evolution









Courtesy F. Stellato



Courtesy M. Del Franco