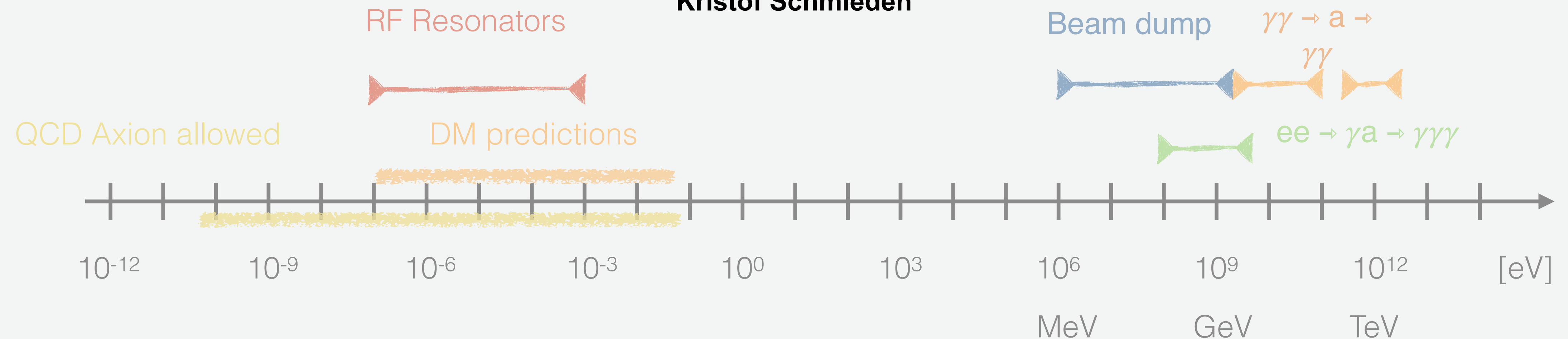


Stellar / Cosmological observations

# The case of PBH mergers: What's the reach?

GravNet meeting - 22/01/2025

Kristof Schmieden



# How many photons can we expect in a cavity?

- Signal strain:

$$h_0 = 9.77 \cdot 10^{-34} \left( \frac{f}{\text{GHz}} \right)^{\frac{2}{3}} \left( \frac{m_{PBH}}{10^{-12} M_\odot} \right)^{\frac{5}{3}} \left( \frac{d}{\text{kpc}} \right)^{-1} = 1.3 \cdot 10^{-22} \quad 2205.02153, \text{ eq. 2.29}$$

- RF power:

$$P_{sig} = \frac{1}{2\mu_0 c^2} Q (2\pi f)^3 V^{\frac{5}{3}} (\eta h_0 B)^2 = 1.3 \cdot 10^{-24} \text{ W} \quad 2303.06006, \text{ eq. 6}$$

- RF Energy deposited in cavity:

- Time of signal  $\Delta t$  within cavity bandwidth  $bw = f_0/Q = 50 \text{ kHz}$ :

$$\dot{f} = 4.62 \cdot 10^{11} \text{ Hz}^2 \left( \frac{f}{\text{GHz}} \right)^{\frac{11}{3}} \left( \frac{m_{PBH}}{10^{-9} M_\odot} \right)^{\frac{5}{3}} = 1.7 \cdot 10^4 \frac{\text{Hz}}{\text{s}} \quad 2205.02153, \text{ eq. 2.35}$$

$$\Delta t = \frac{bw}{\dot{f}} = \frac{f_0}{Q\dot{f}} = 3 \text{ s}$$

$$E_{sig} = P_{sig} \cdot \Delta t = 3.9 \cdot 10^{-24} \text{ J}$$

$$N_\gamma = \frac{E_{sig}}{hf} = \frac{3.9 \cdot 10^{-24}}{3.3 \cdot 10^{-24}} \approx 1$$

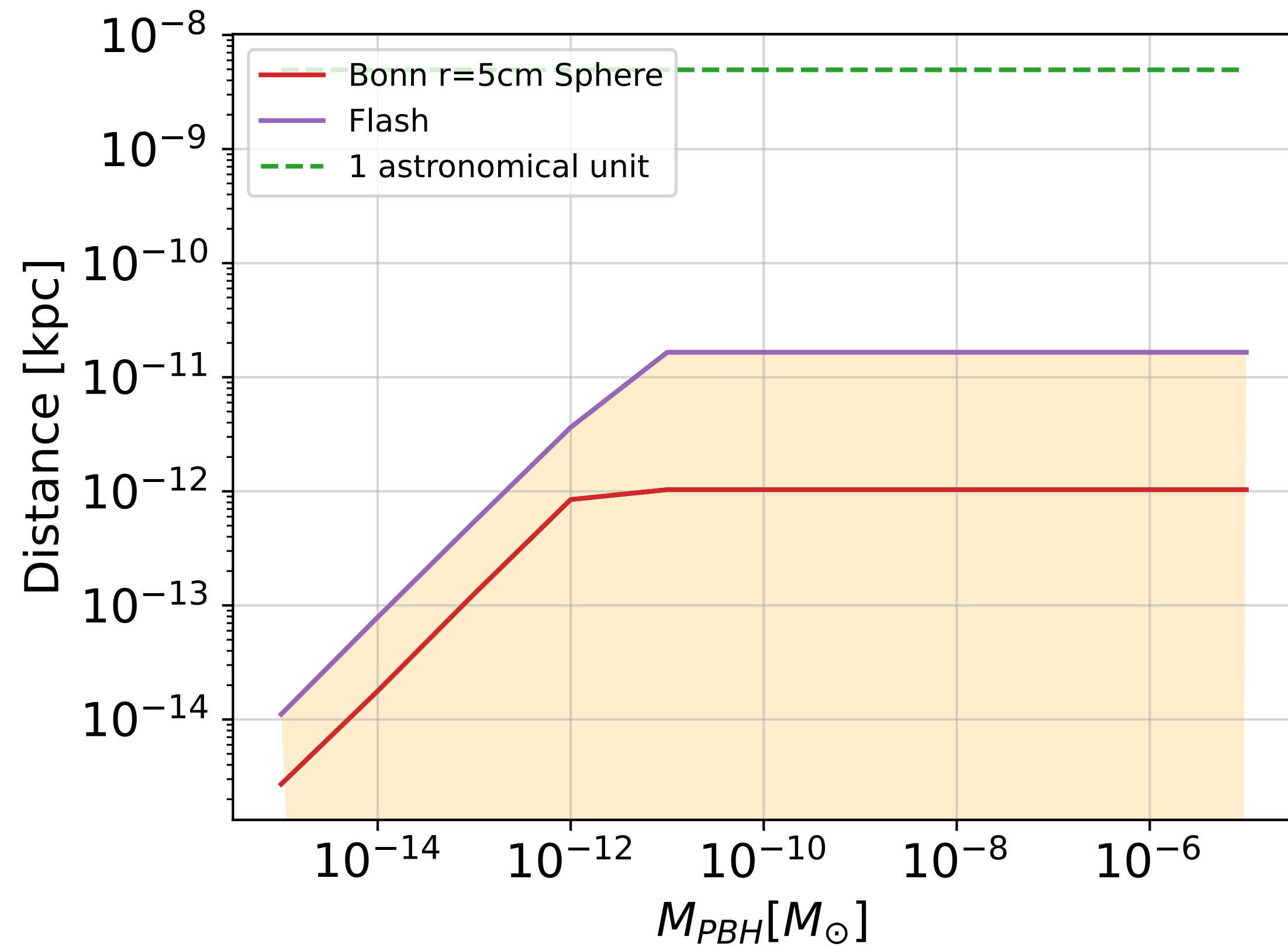
- Assumptions:**

- $m_{PBH} = 10^{-15} M_\odot$
- Distance  $d = 2.7 \cdot 10^{-15} \text{ kpc}$
- Cavity:  
 $f_0 = 5 \text{ GHz}$ ,  $V = 0.01 \text{ m}^3$ ,  $Q = 10^5$ ,  $\eta = 0.1$
- $B = 12 \text{ T}$

$$N_\gamma \propto d^{-2} B^2 V^{\frac{5}{3}} f^{\frac{5}{3}}$$

# How far can we look?

- Finding distance for which at least one photon is generated within the bandwidth of the cavity

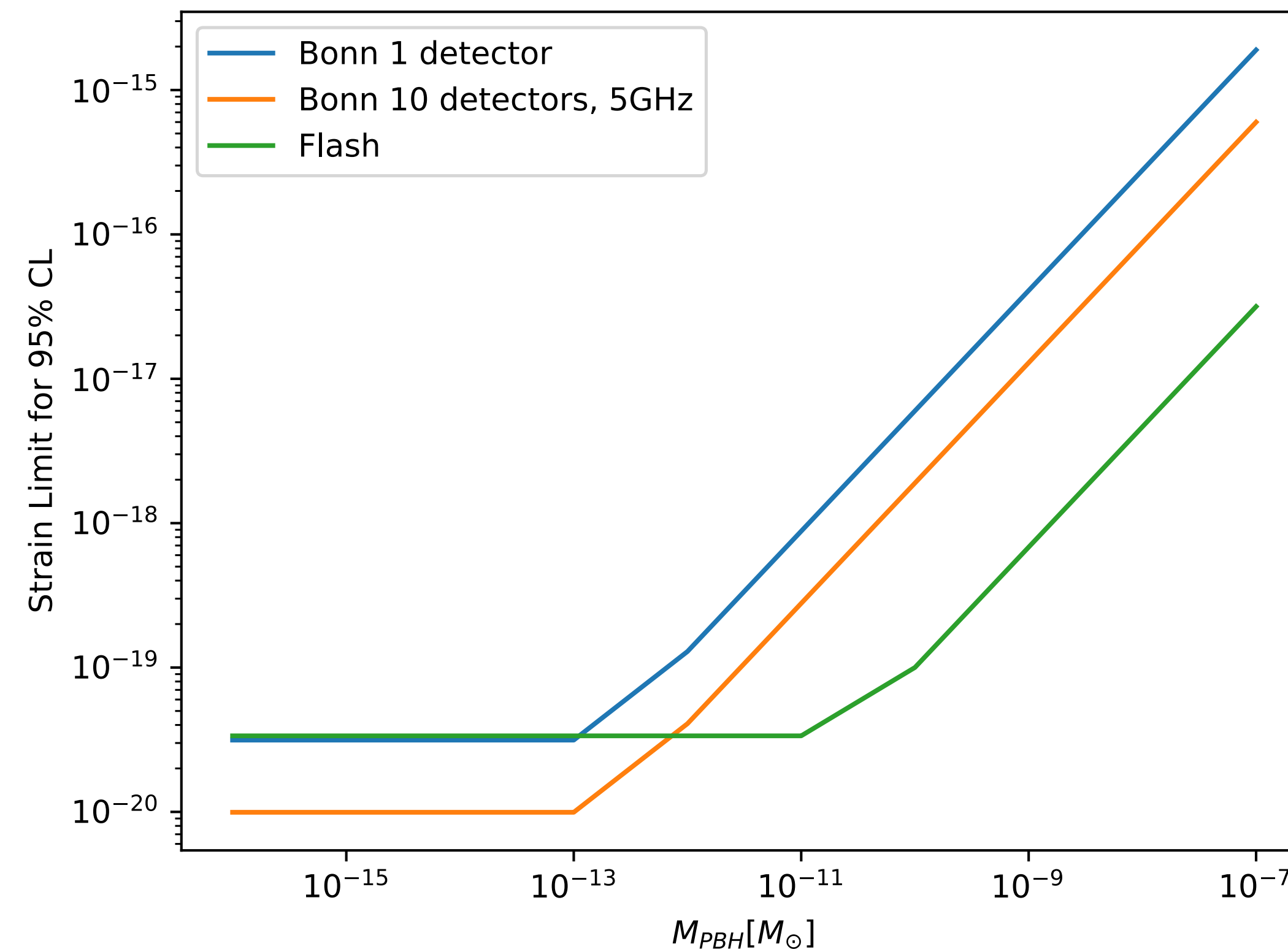


- Effective quality factor for fast transients:
  - $Q_{eff} = \min(Q, N_{cycles})$

New plots

# Whats the limit on the strain that can be reached?

- Assuming only thermal noise
- Given a readout / integration time, what strain can be detected with a 95% CL given 1 event per year?
  - This is just using number of photons, assuming a detector with  $\epsilon = 1$ , capable of counting photons
- At least 1 photon is forced -> const. region in plot



# Whats the limit on the strain that can be reached?

- Assuming only thermal noise
- Given a readout / integration time, what strain can be detected with a 95% CL given 1 event per year?
  - This is just using number of photons, assuming a detector with  $\epsilon = 1$ , capable of counting photons
- For long integration times  $> 1$  second: SNR = 1 is chosen as figure of merit, in line with existing publications

