

# **Tunable haloscope at 10 GHz probing up to KSVZ QCD axion models**

**Giosuè Sardo Infirri**  
on behalf of the QUAX collaboration

**20<sup>th</sup> PATRAS workshop**  
La Laguna (Tenerife), 22-26 September 2025

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# Agenda

1. Overview of the QUAX apparatuses
2. QUAX-LNL haloscope:
  - Dilution refrigerator
  - Low-noise electronics
  - Dielectrically loaded cavity
  - Cavity tuning
3. Improvements on automatic haloscope:
  - Data acquisition
  - Data analysis
4. Recent results
5. Upcoming improvements and research lines @ PD-LNL
6. Conclusions

# QUAX-ay haloscopes

	LNF	PD-LNL
Magnetic Field	9 T	8 T
Target frequency ranges	8 – 9 GHz	9 – 11 GHz
Base temperature	30 mK	100 mK
Cavity mode	TM <sub>010</sub>	TM <sub>030</sub>
Unloaded quality factor	~ 50000	~ 75000
Cavity effective volume	0.139 L	0.450 L
Cavity frequency (@ cold)	8.817 GHz	10.212 GHz
Scannable region	289 MHz	90 MHz
Scanned region	6 MHz	38 MHz
Sensitivity	~ 6 g <sub>ayy</sub>	~ 2 g <sub>ayy</sub>
Scan rate @ KSVZ	0.12 MHz/day	1.5 MHz/day



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$$\frac{\partial f}{\partial t} \propto f^{-4}$$

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Istituto Nazionale di Fisica Nucleare  
Laboratori Nazionali di Legnaro



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- Liquid-He to cool the magnets
- 100 mK @ Mixing Chamber (MC)

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- Working in persisting current mode @ L-He temperature



# Dilution refrigerator

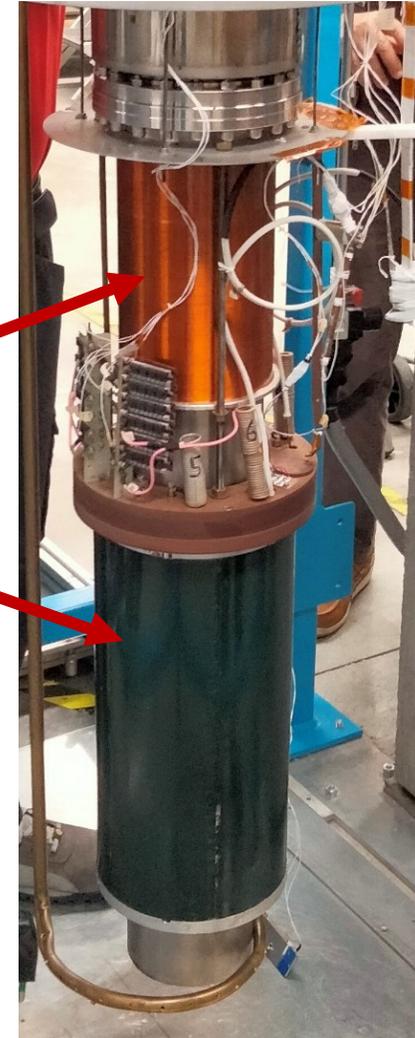
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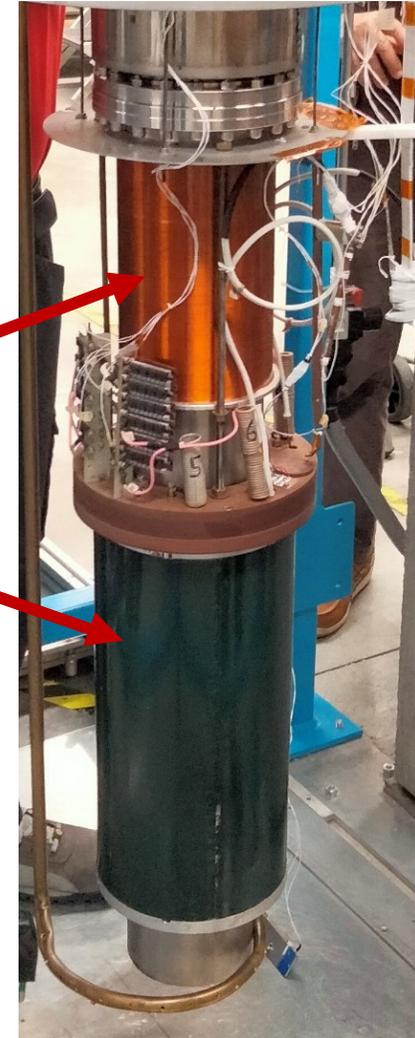
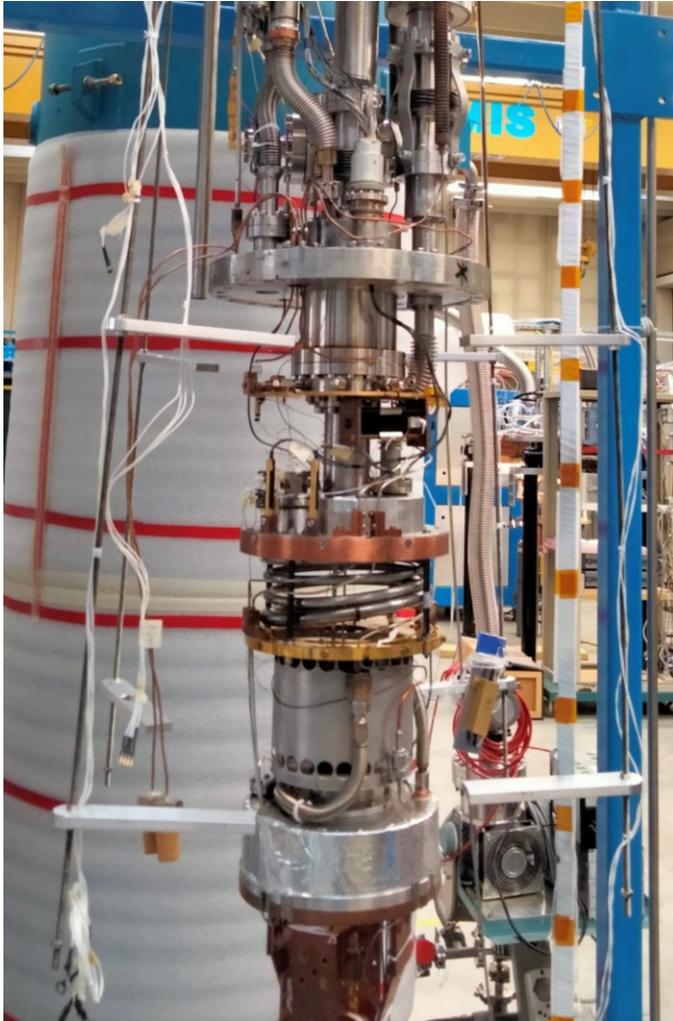
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Typical values for acquisition:

- 3500 L-He dewar  $\sim$  20 days run

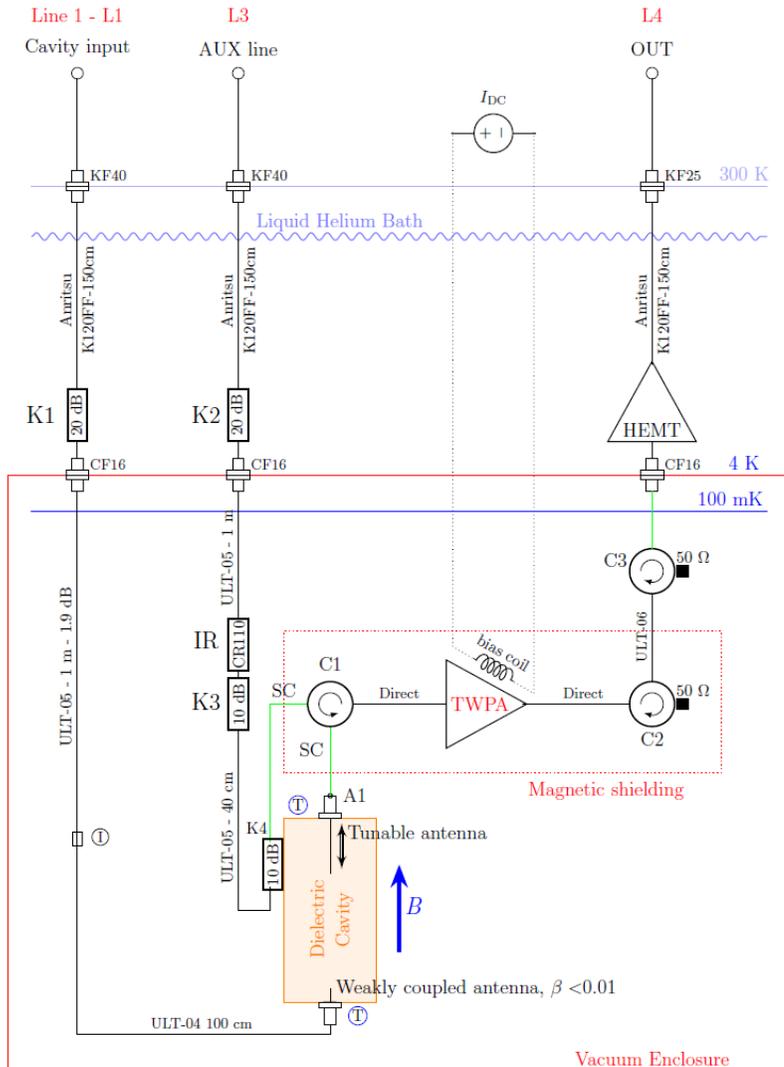


# Low-noise electronics

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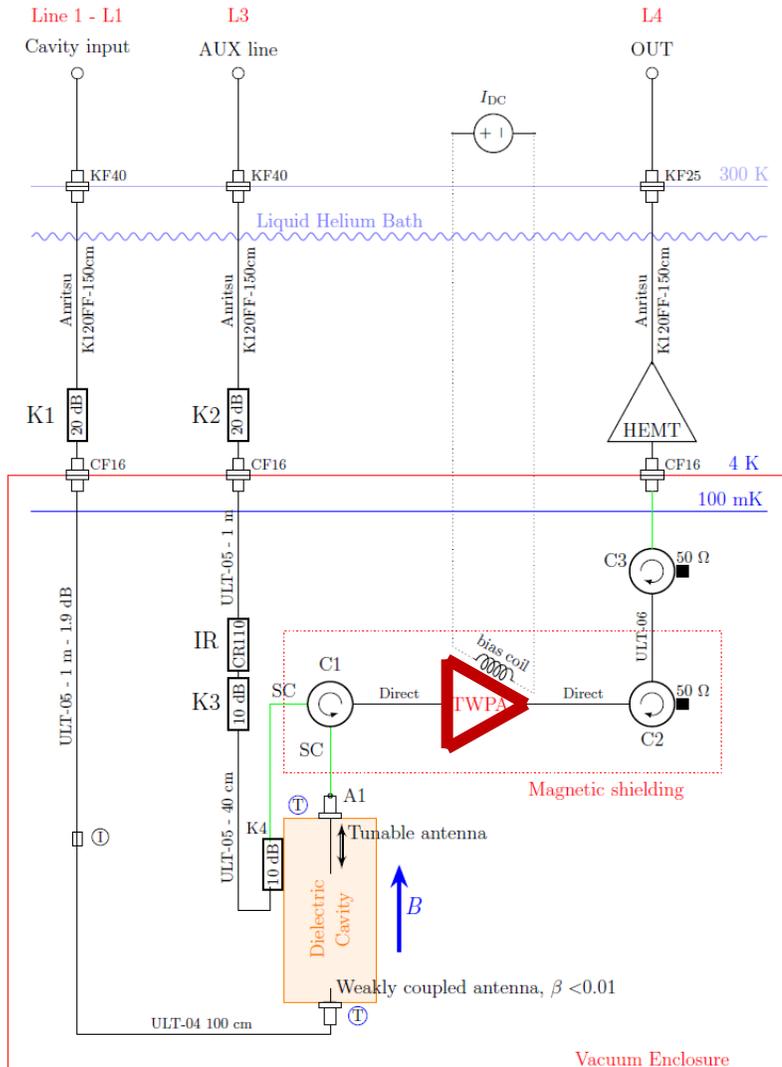
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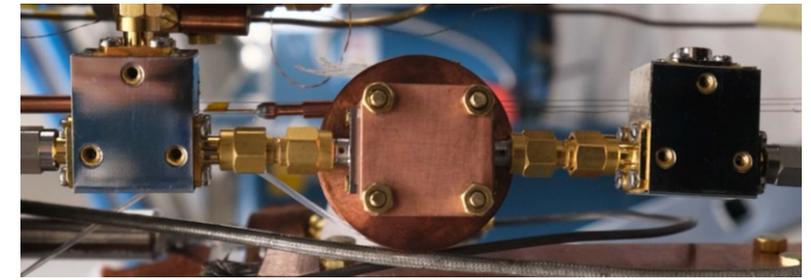
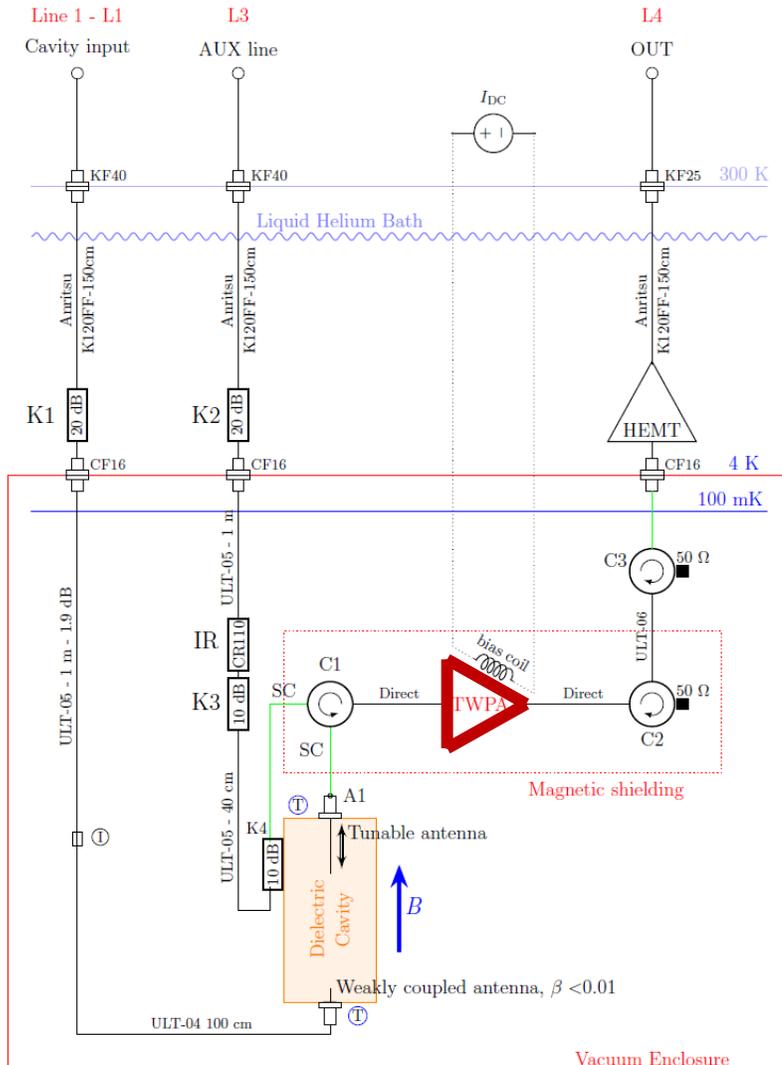
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Dispersion engineering Travelling Wave Parametric Amplifier **TWPA**:



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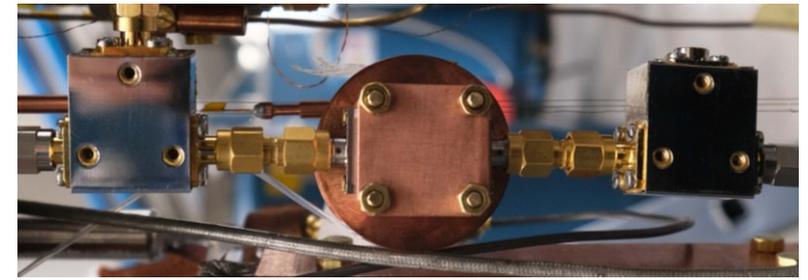


Ranadive, A., Esposito, M., Planat, L. *et al.*  
[Nat Commun \*\*13\*\*, 1737 \(2022\)](https://doi.org/10.1038/s41467-022-28111-1)

## Dispersion engineering Travelling Wave Parametric Amplifier **TWPA**:

- Developed by Nicholas Roch @ Grenoble
- Snail-based

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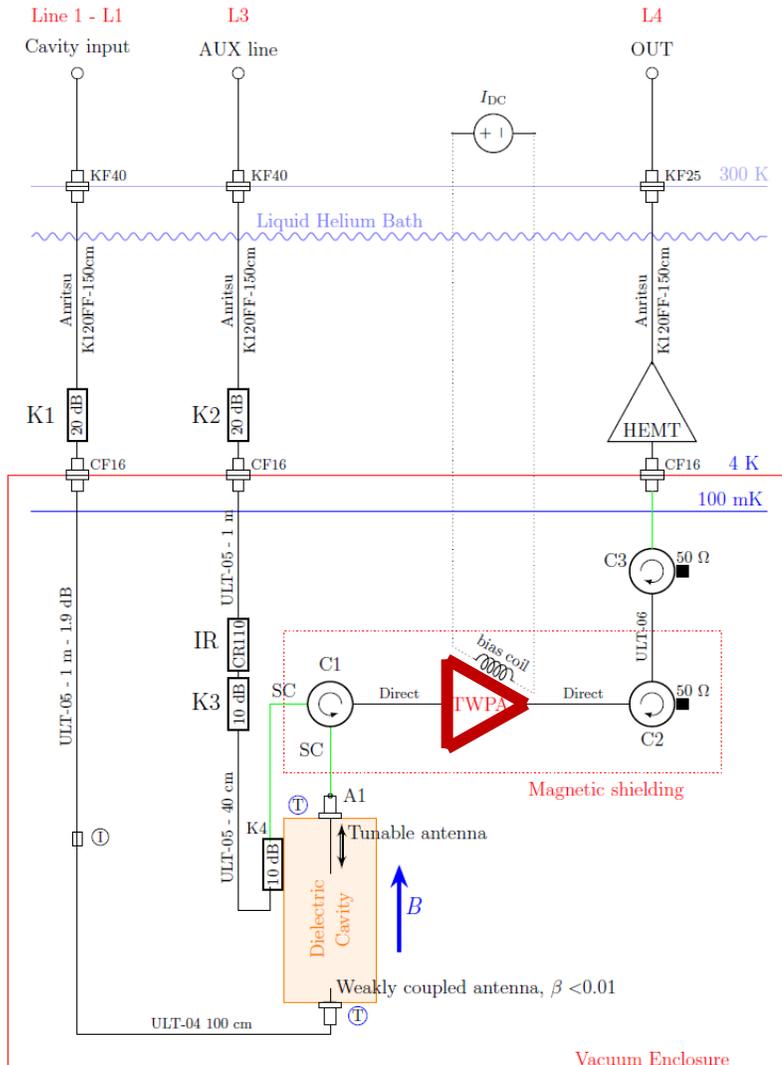


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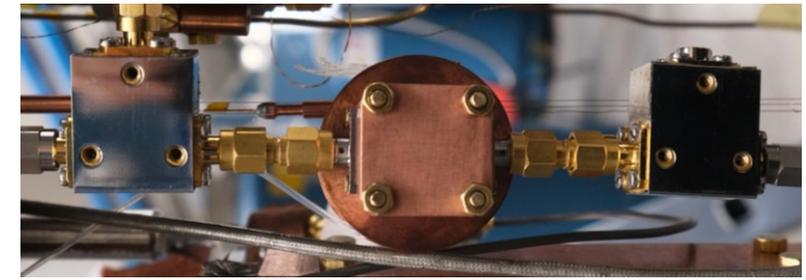
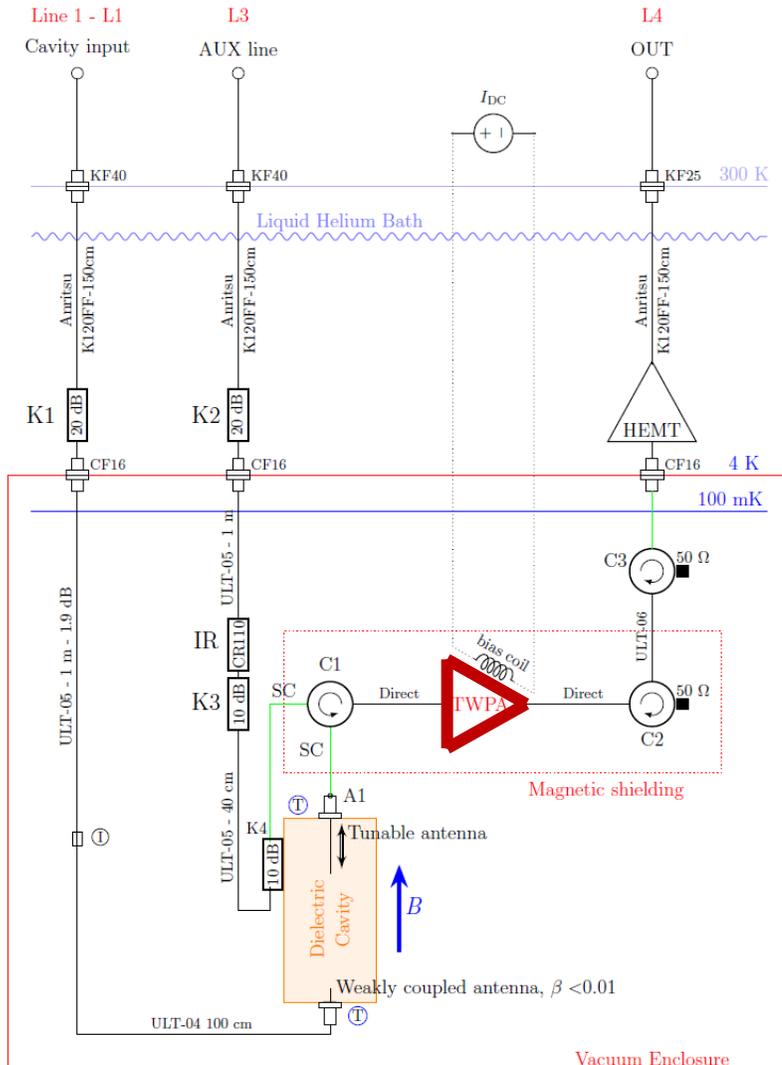
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### Typical values:



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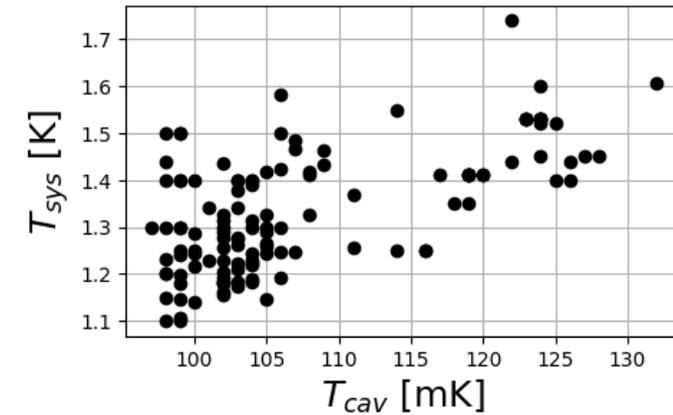
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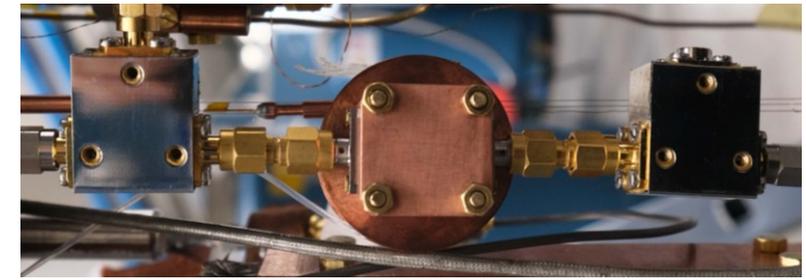
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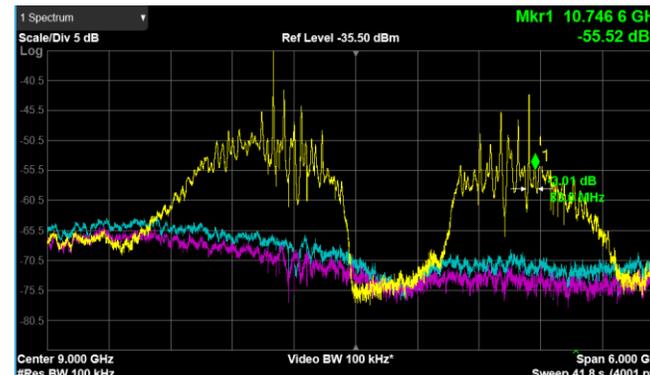
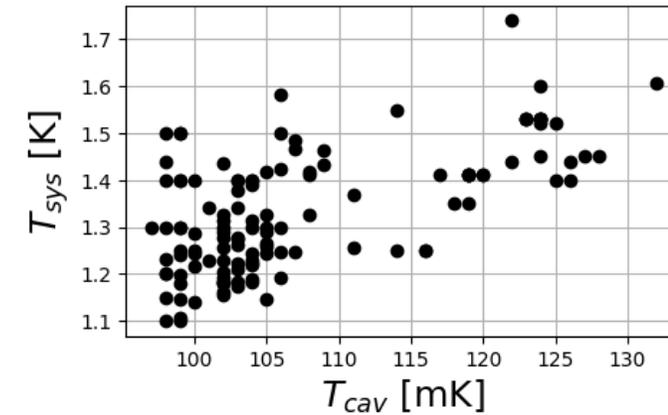
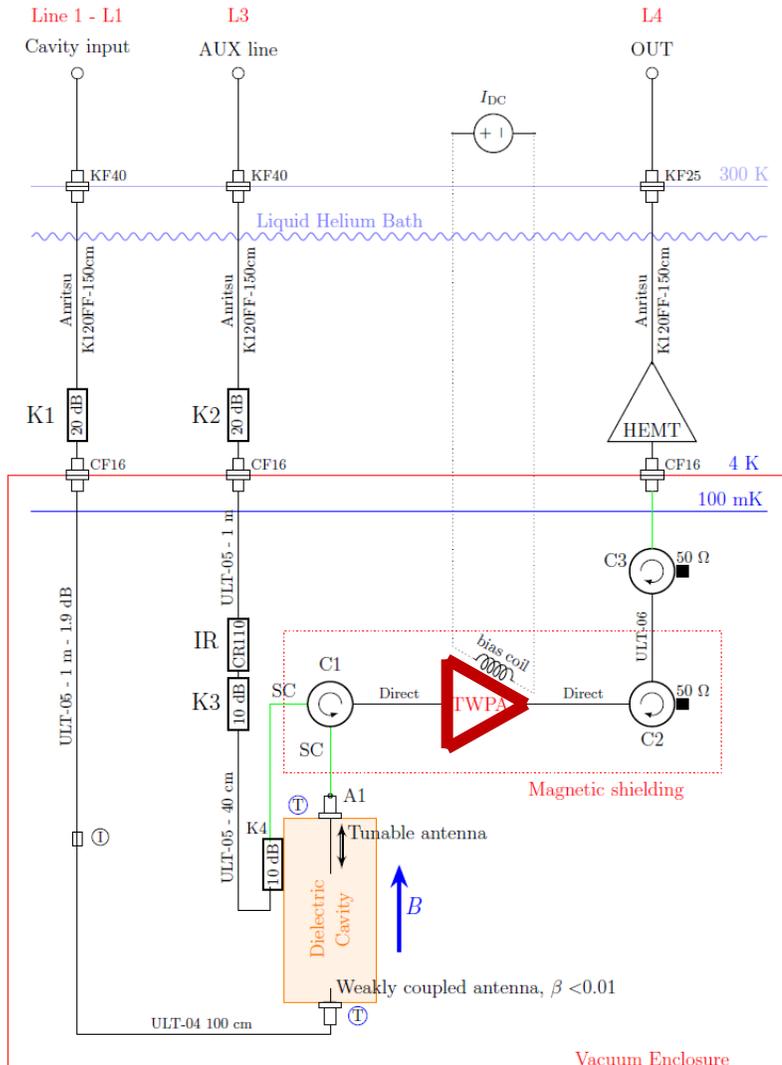
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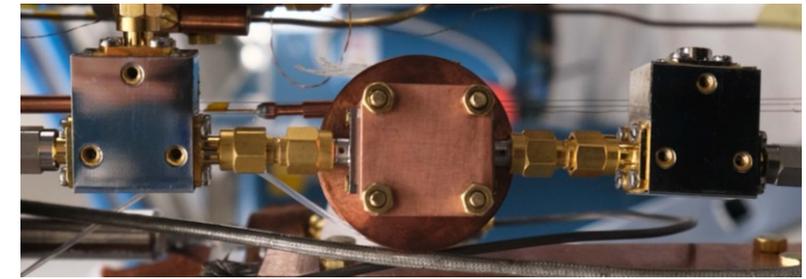
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- Wide-band amplification range (6-12 GHz)



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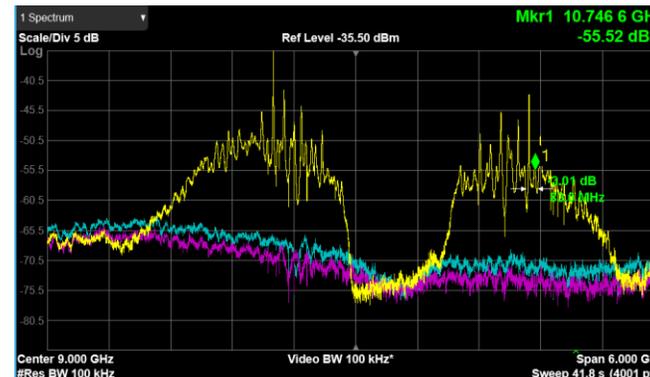
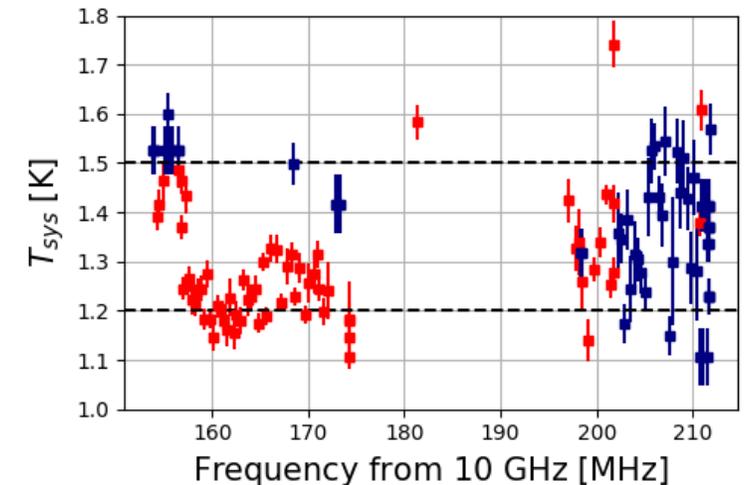
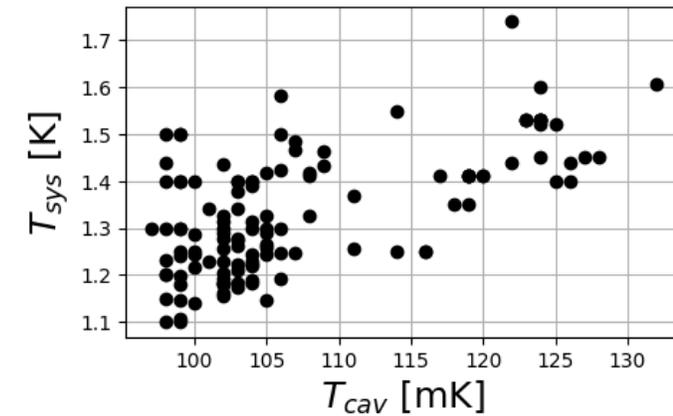
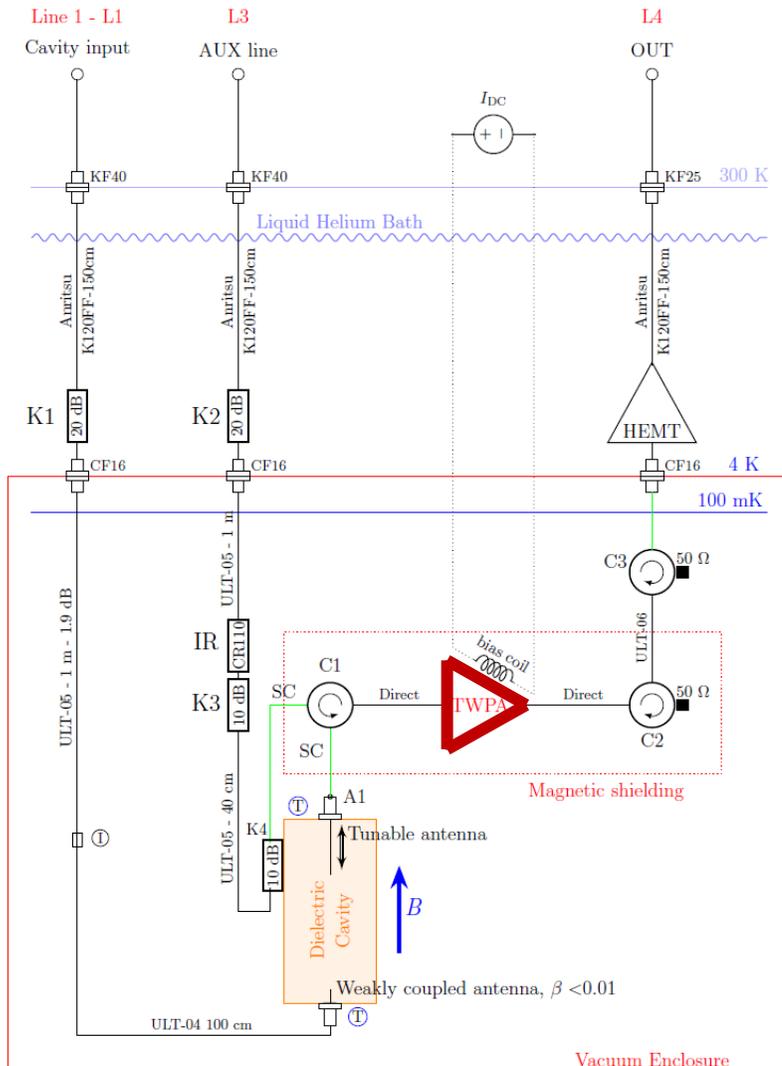
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- Working @ 100-110 mK
- Wide-band amplification range (6-12 GHz)
- Noise temperature  $T$  1.2 -1.5 K (2 SQL @ 10 GHz)



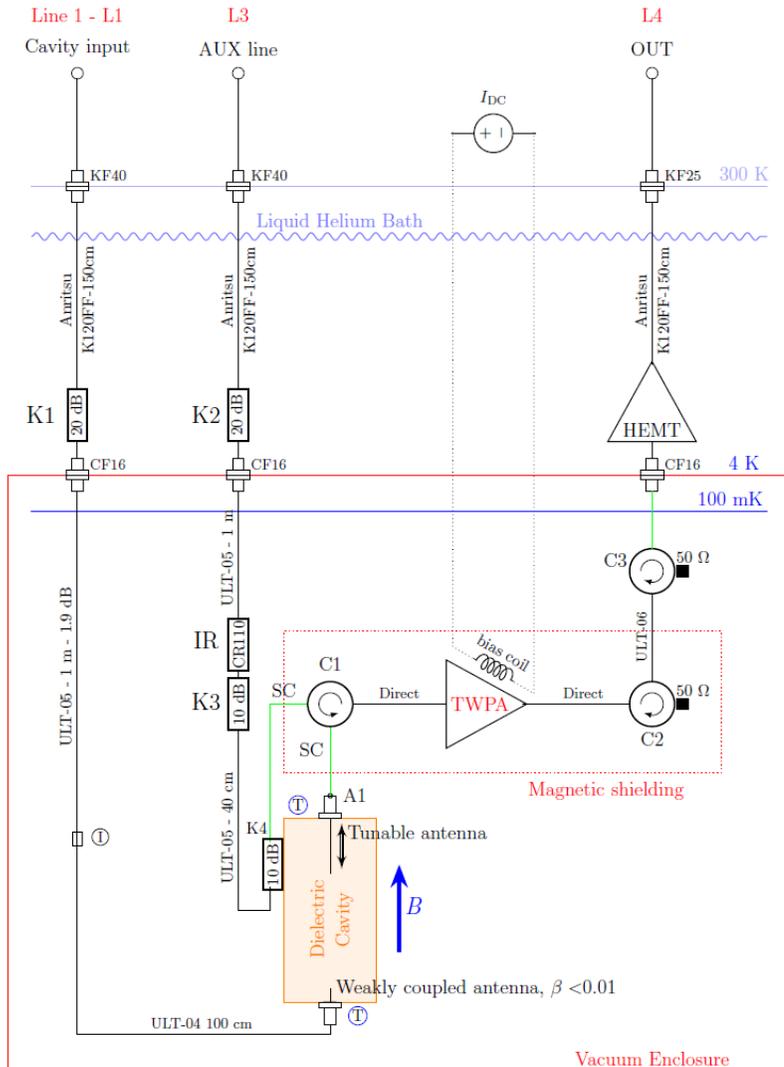
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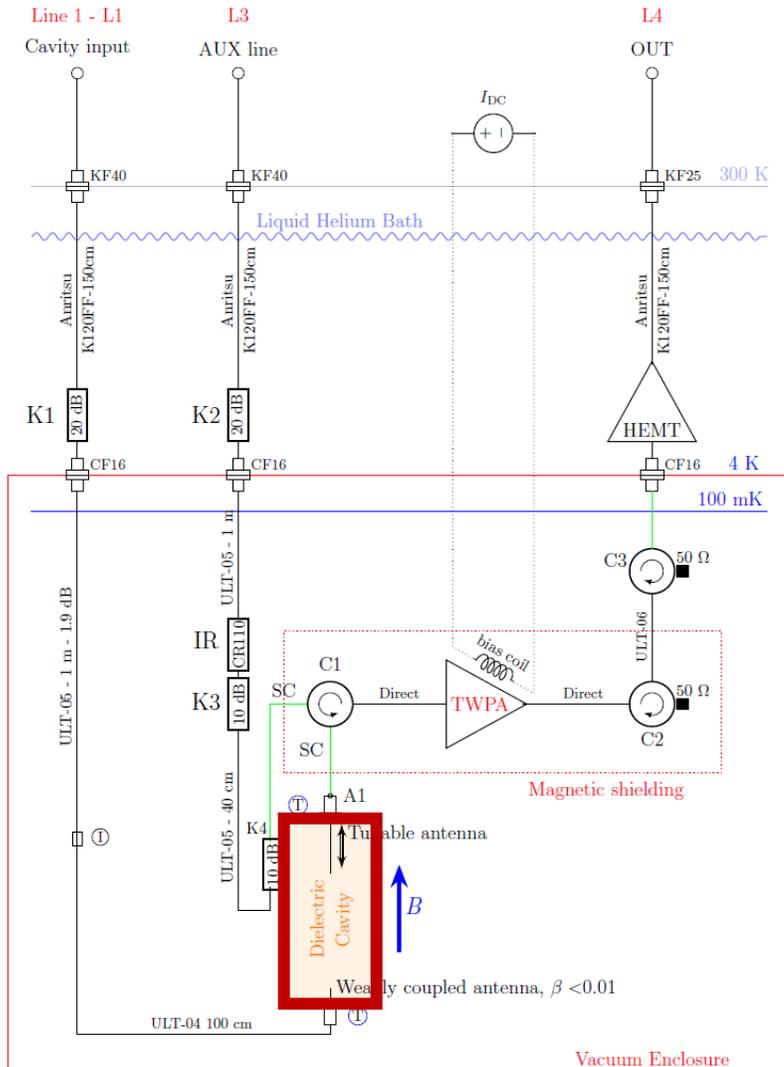
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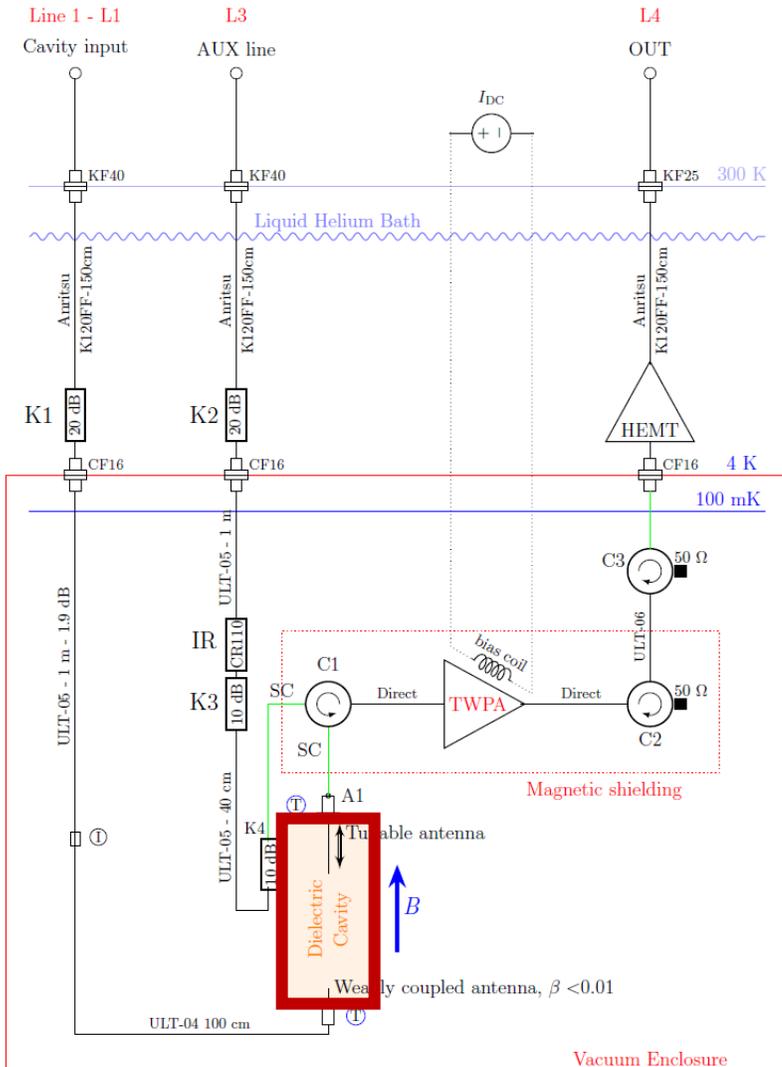
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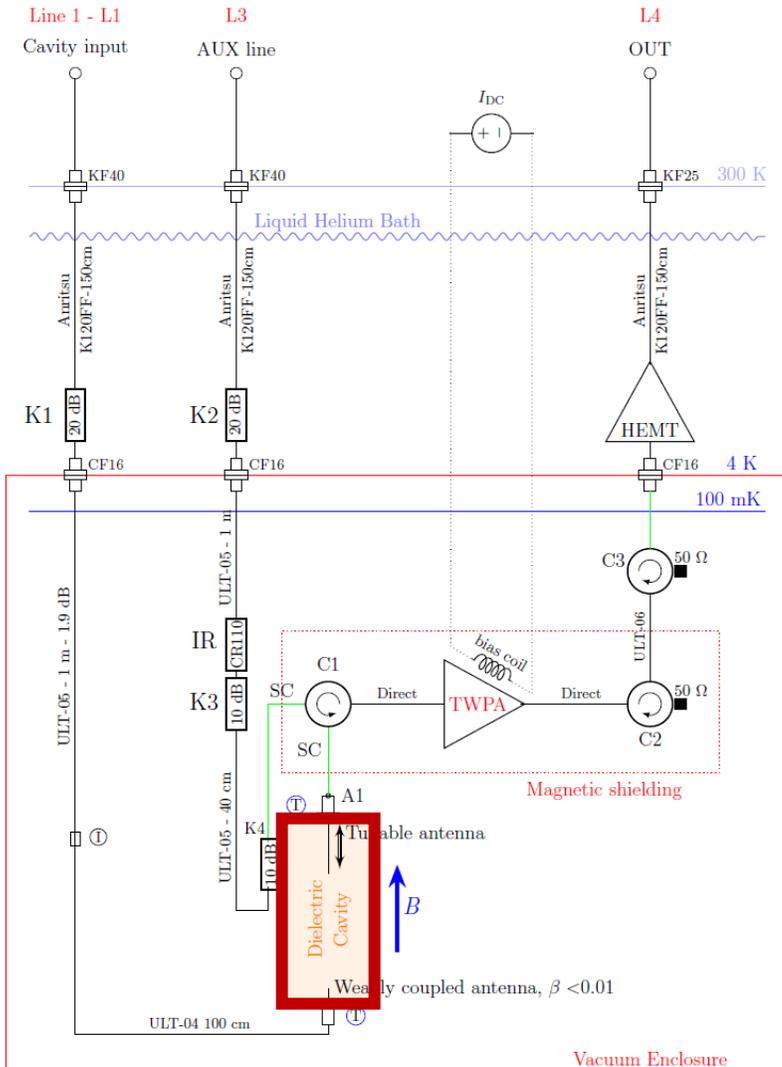


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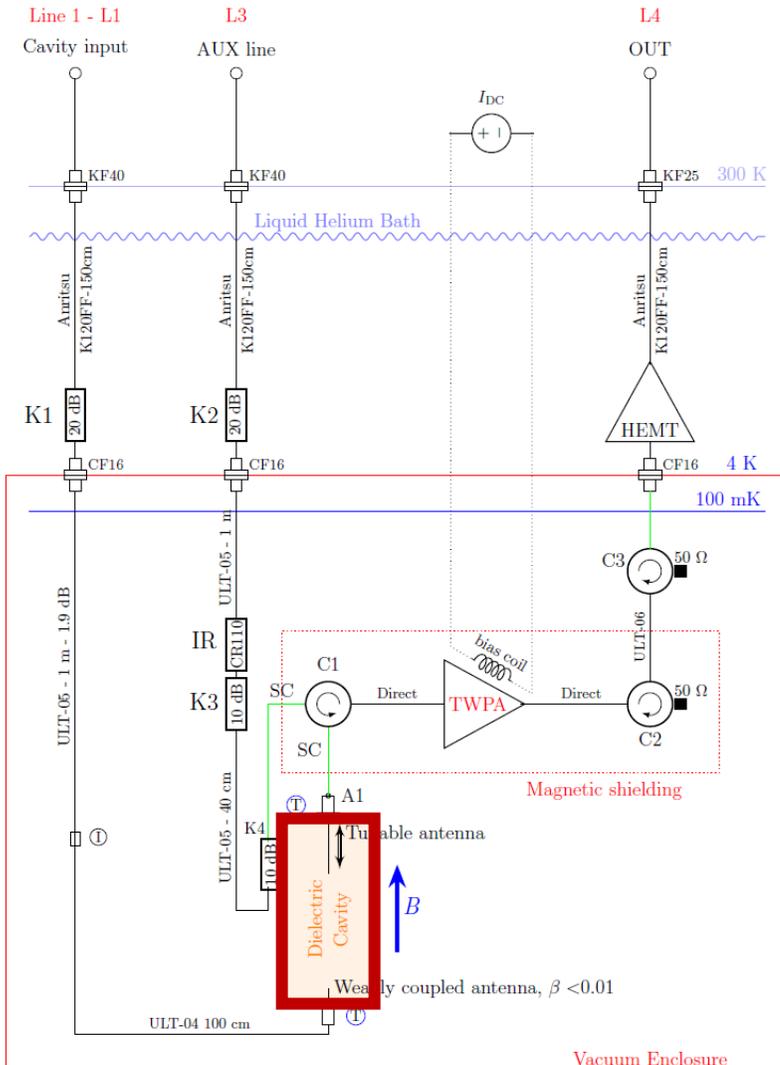


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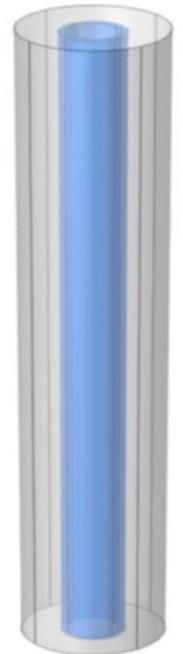
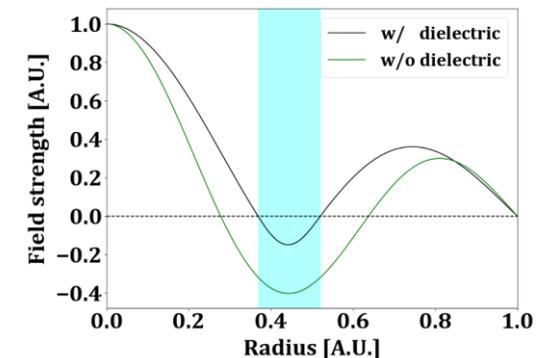
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## Dielectrically loaded cavity:

- Working frequency  $\nu_{030} \sim 10$  GHz
- Use of high-order mode  $TM_{030}$
- Cavity **field profile reshaped** using sapphire tube

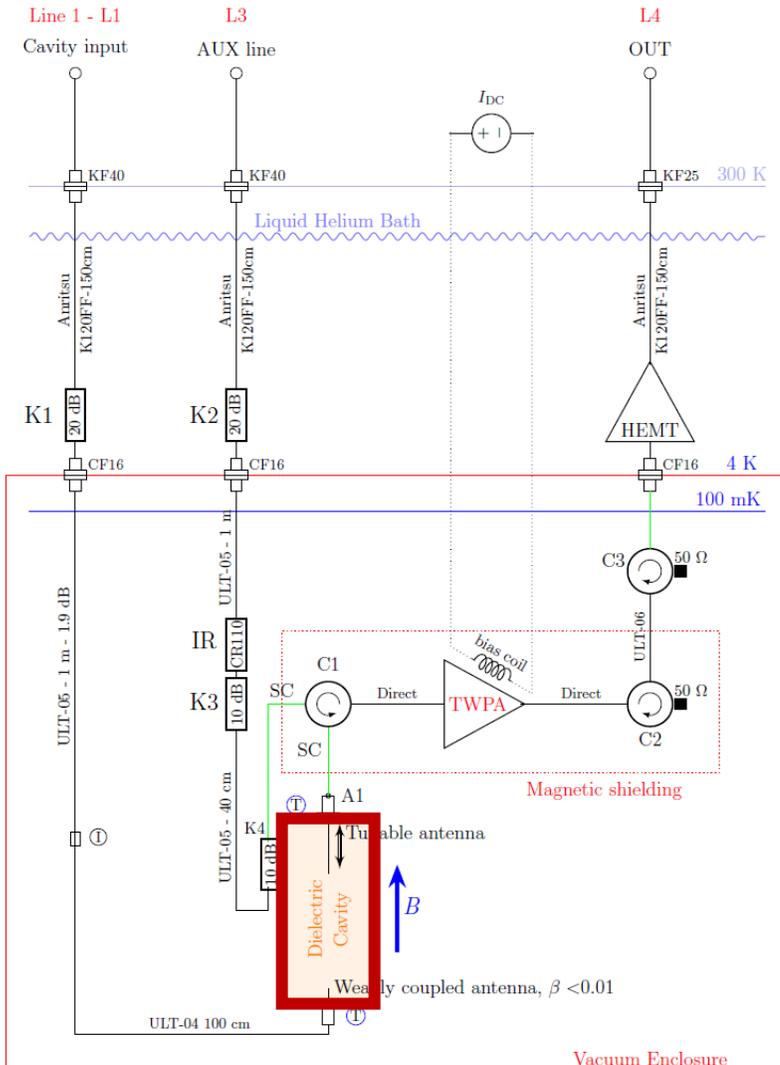


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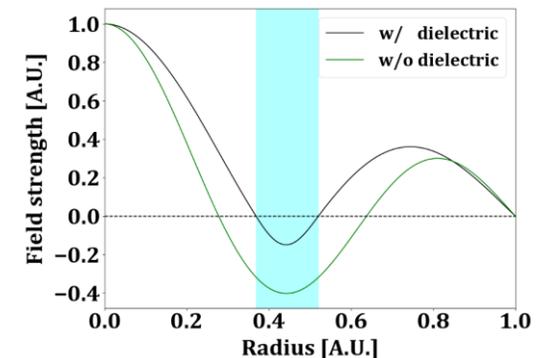
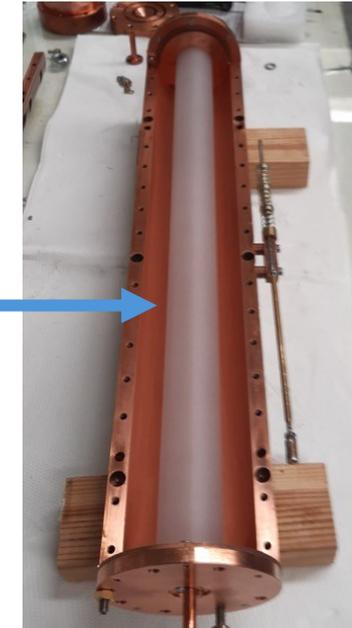
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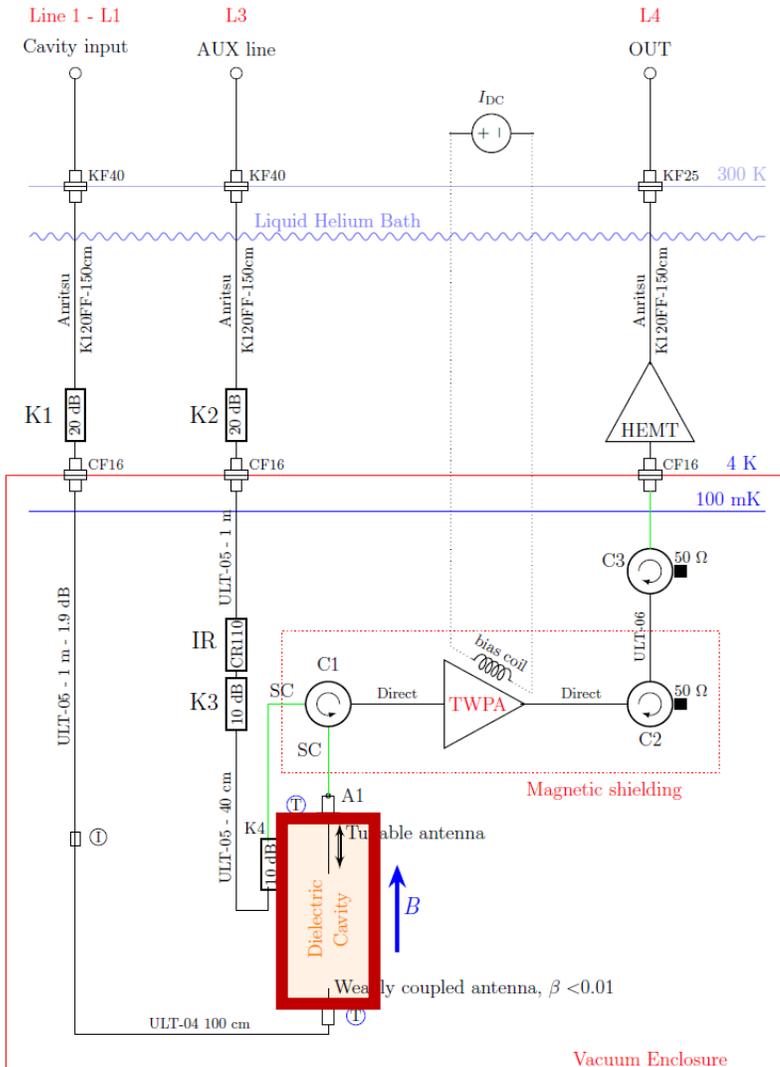


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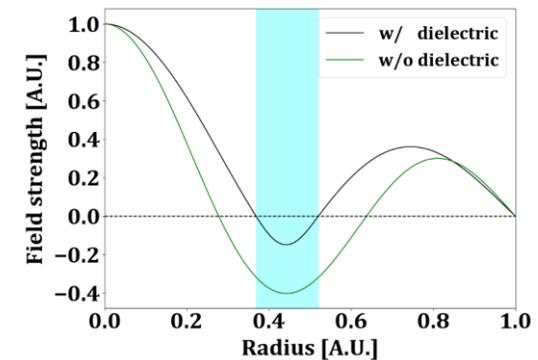
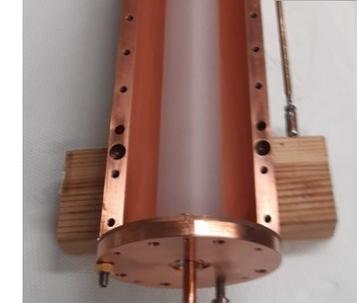
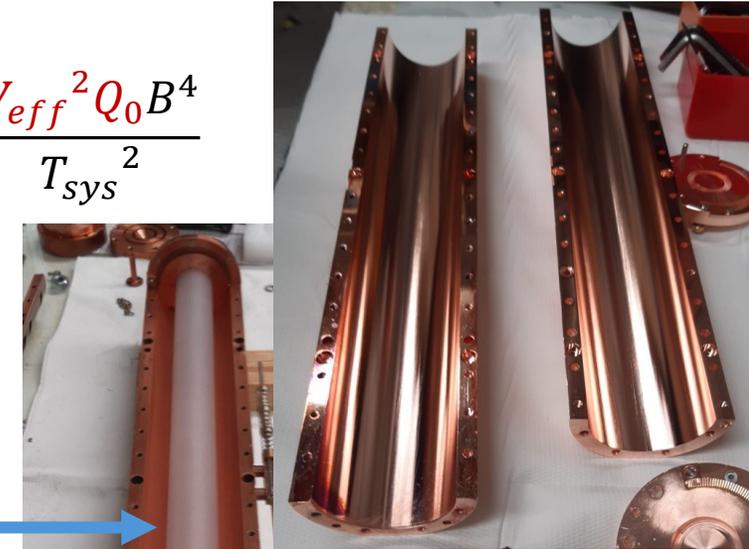
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- Dielectrically loaded cavity:**
- Working frequency  $\nu_{030} \sim 10$  GHz
  - Use of high-order mode  $TM_{030}$
  - Cavity **field profile reshaped** using **sapphire tube**
  - Comprised of two copper halves

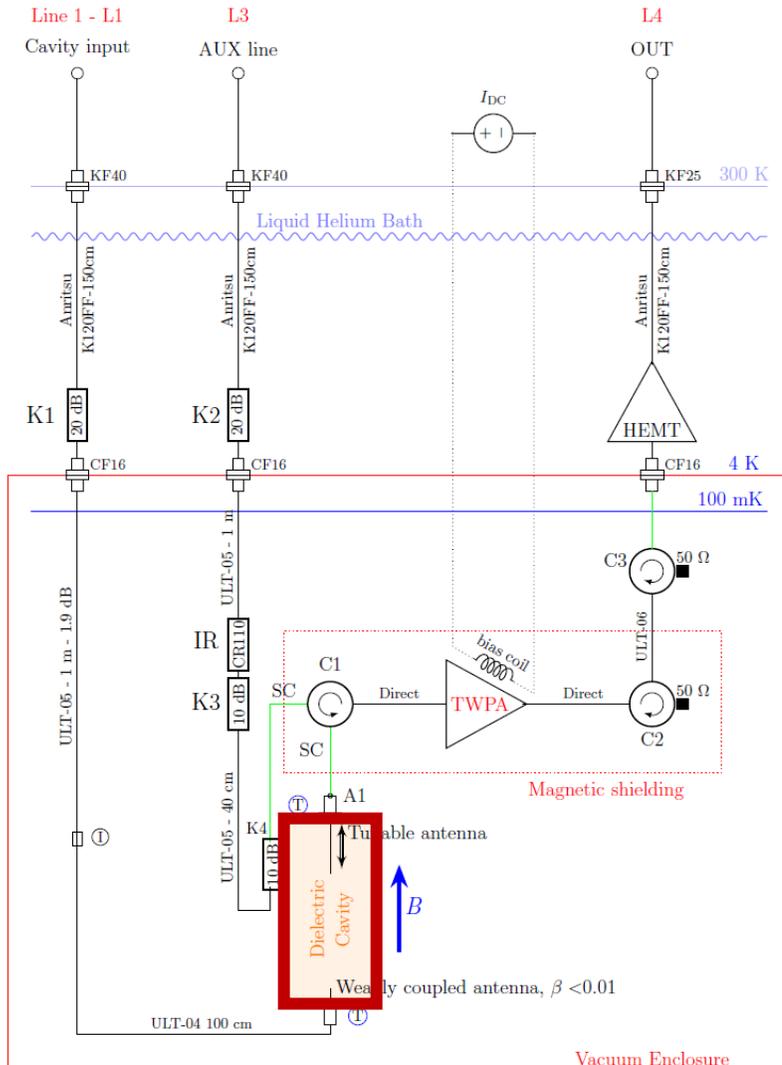


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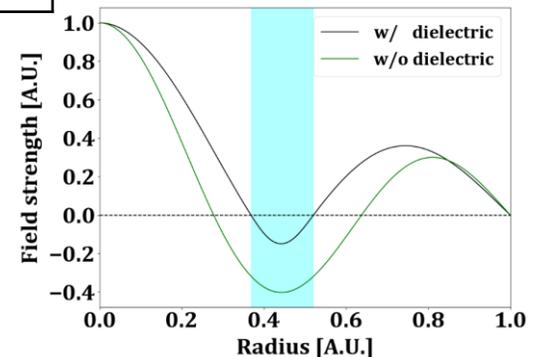
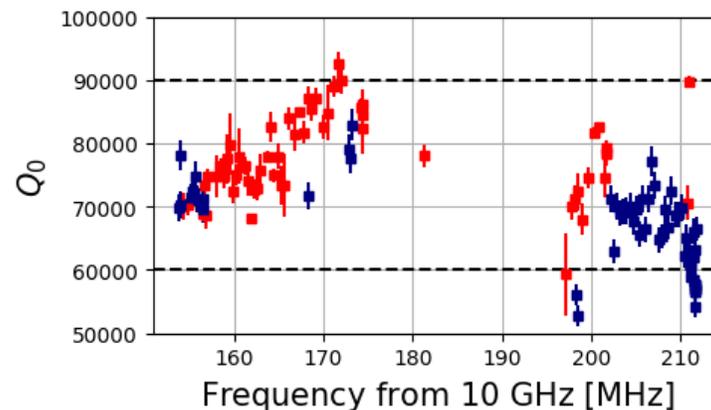
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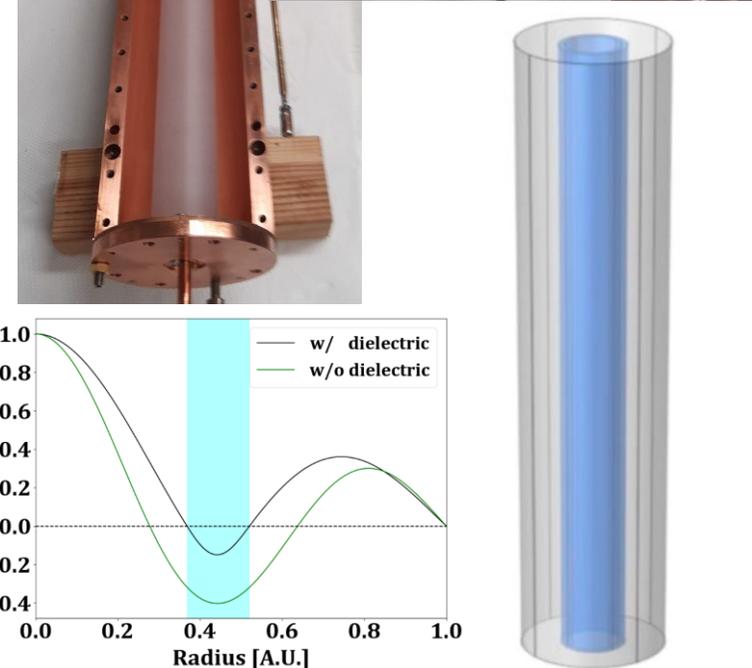
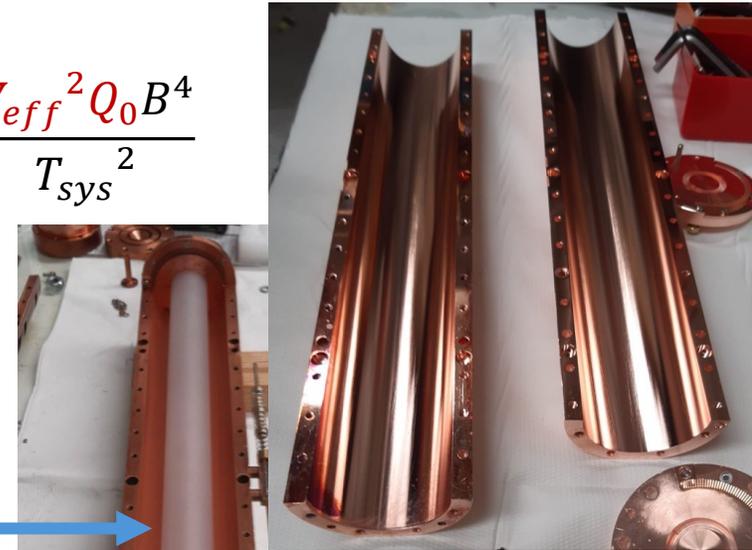
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Form factor	$C_{030} \sim 0.43$
Volume	$V \sim 1$ L ( $V_{eff} = V * C_{030}$ )
Quality factor	$Q_0 \sim 60000-90000$



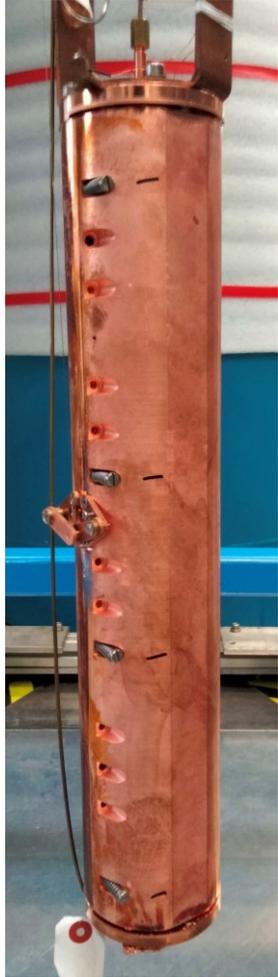
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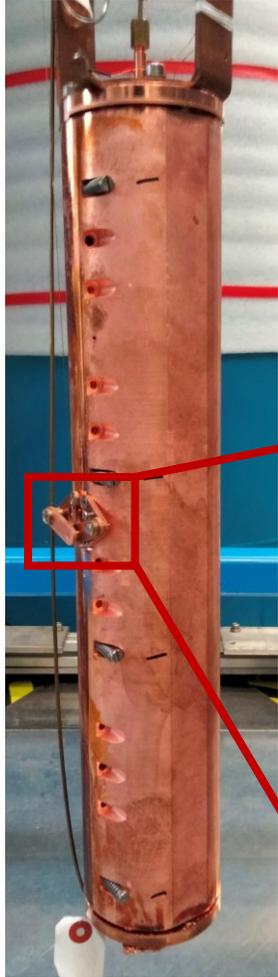
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Cavity tuning system:

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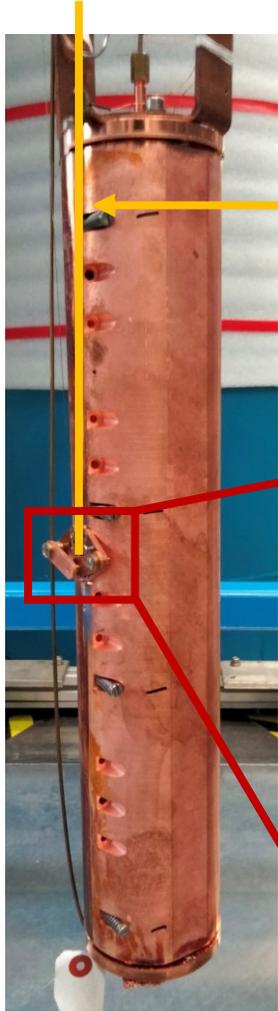


Cavity tuning system:

- A **pantograph** is connected via **wire** to room temperature stepper motor

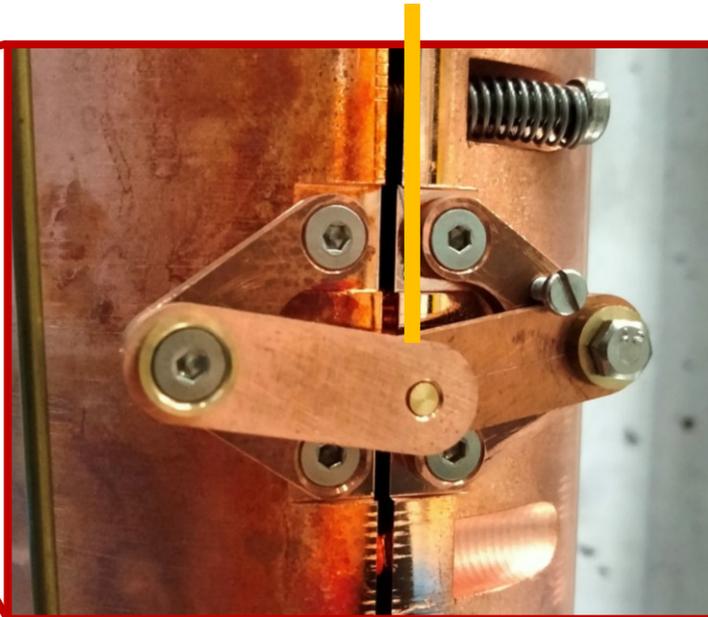


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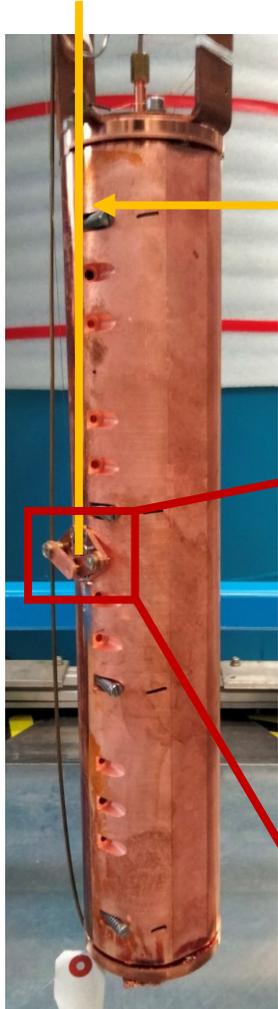


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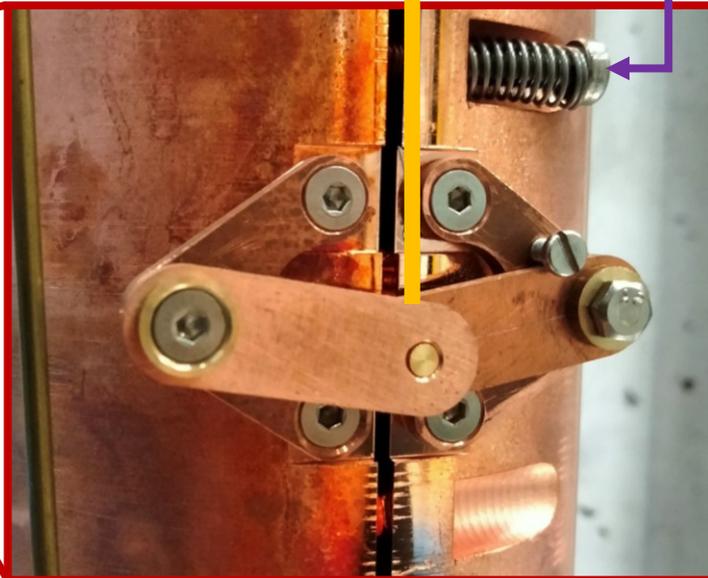


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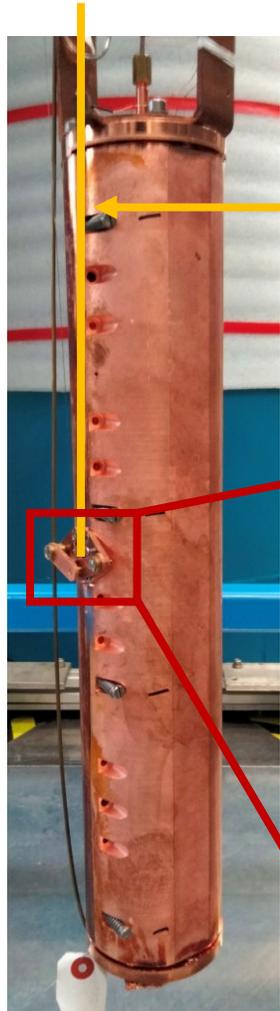


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- A **pantograph** is connected via **wire** to room temperature stepper motor
- Restoring **springs**

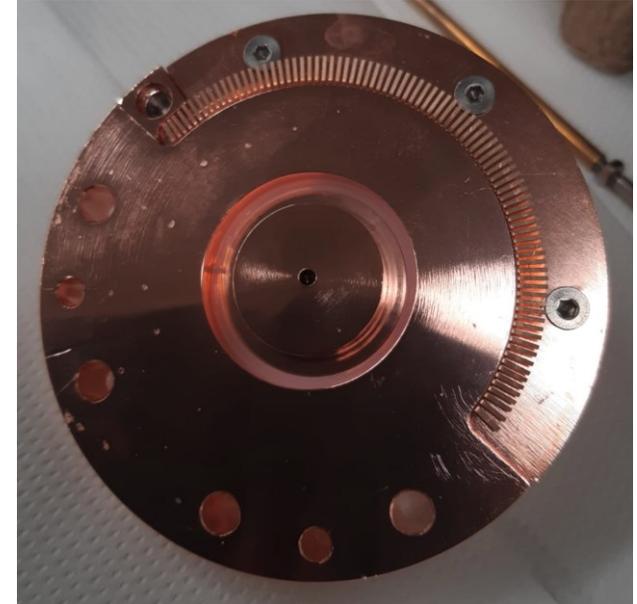
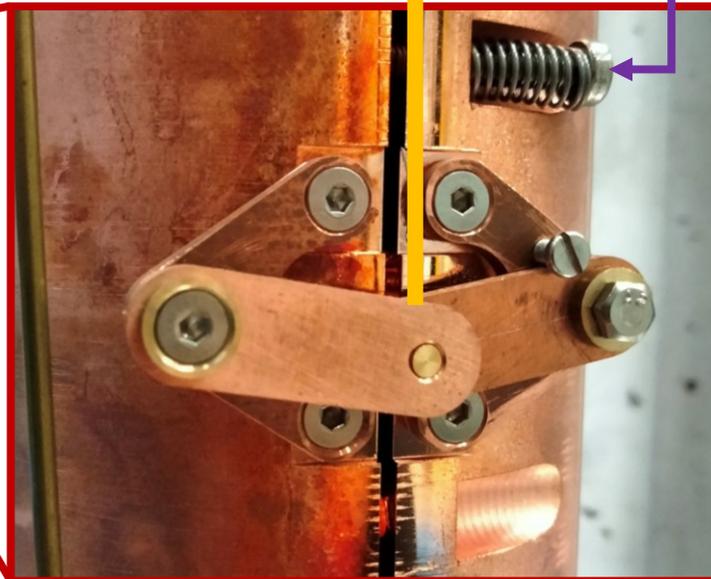


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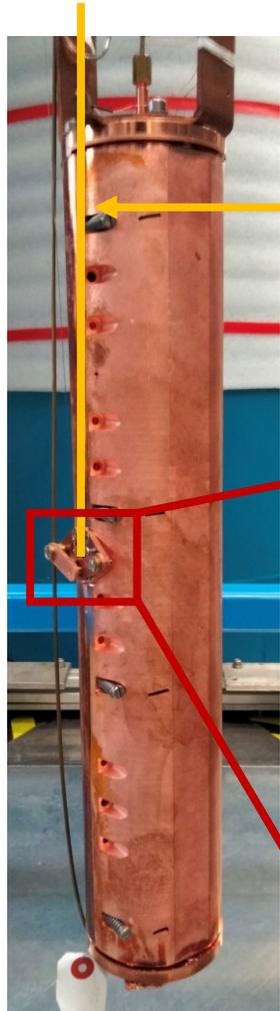


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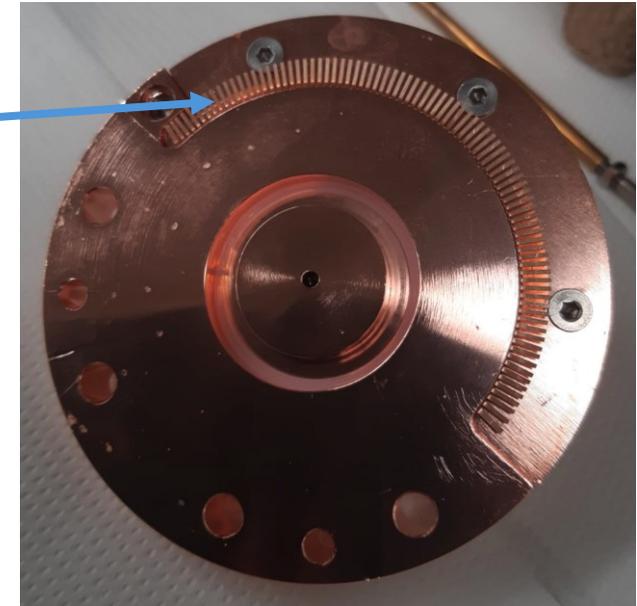
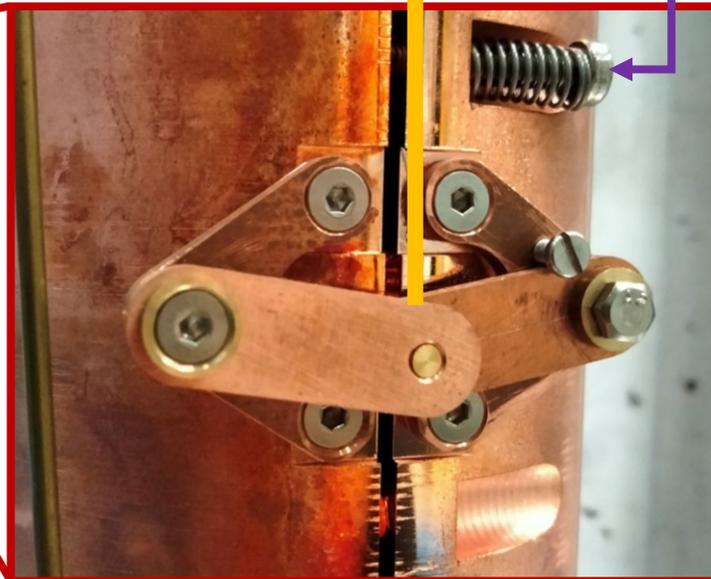


# Cavity tuning

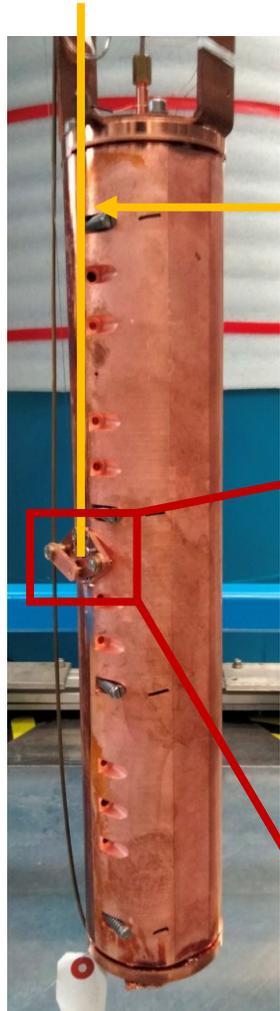


## Cavity tuning system:

- A **pantograph** is connected via **wire** to room temperature stepper motor
- Restoring **springs**
- Several **springs** to guarantee rf contact on endcaps

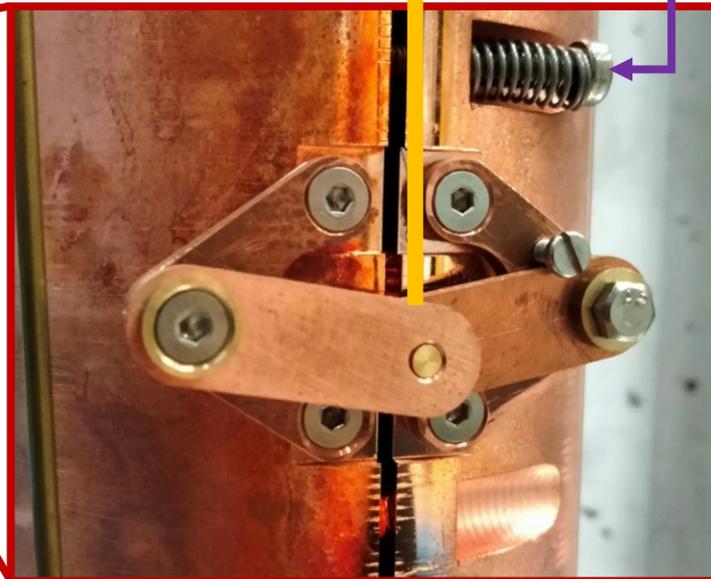


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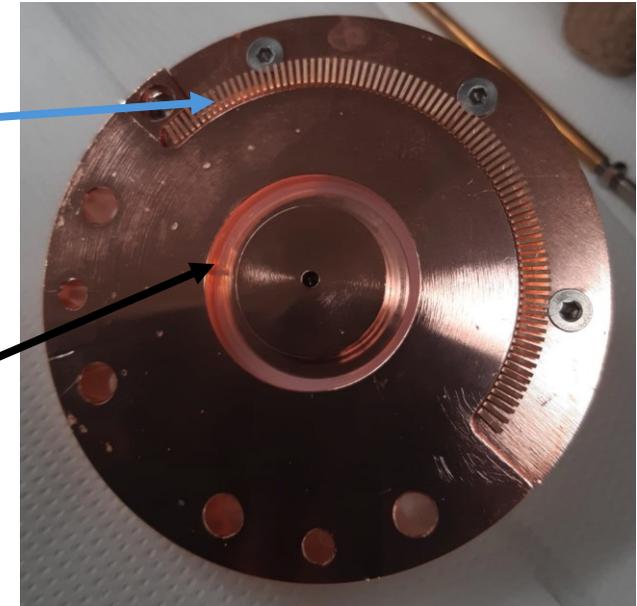


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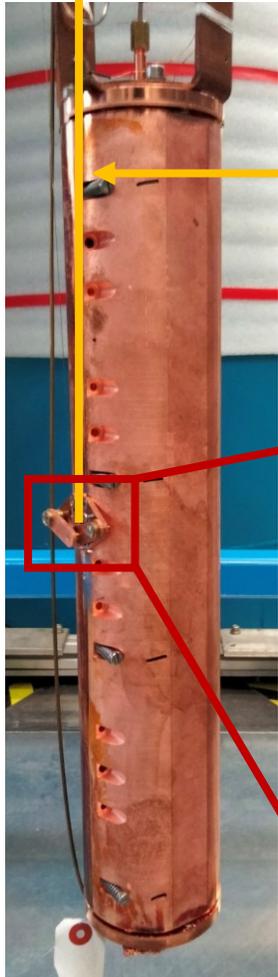
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- Annular **grooves** to hold the sapphire on endcaps

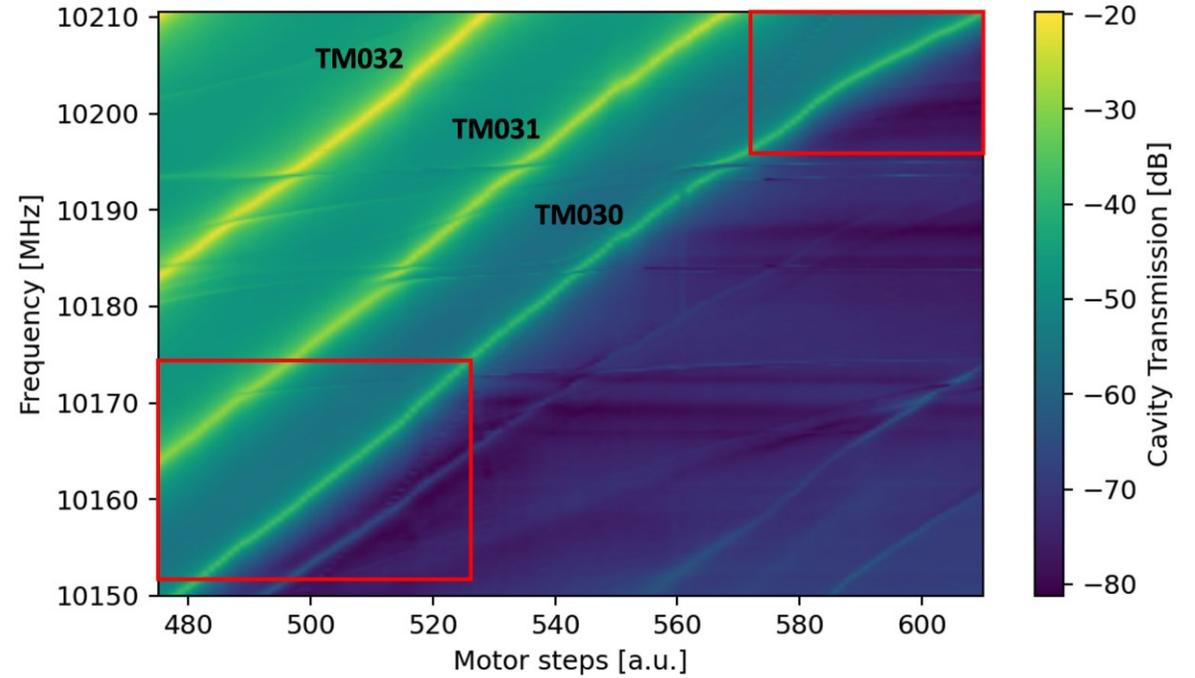
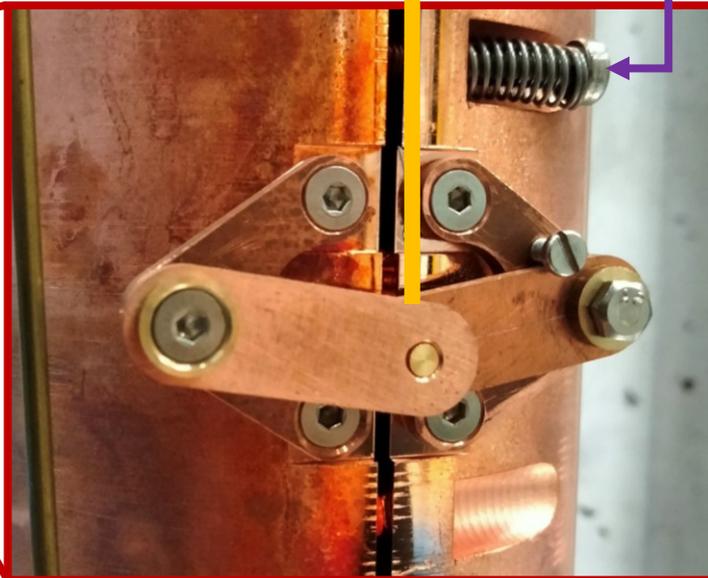


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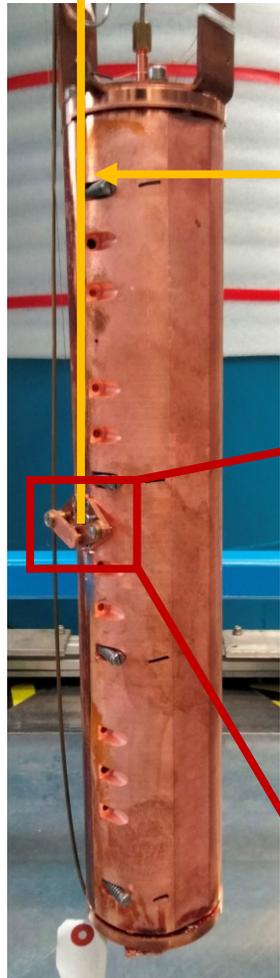
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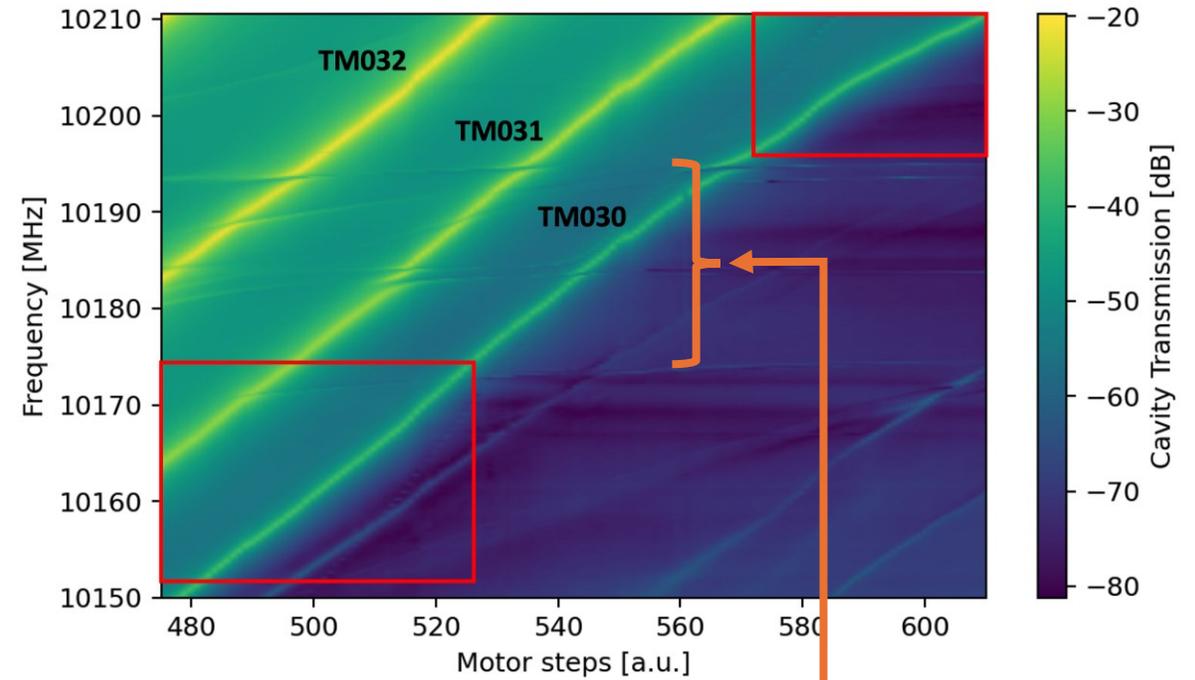
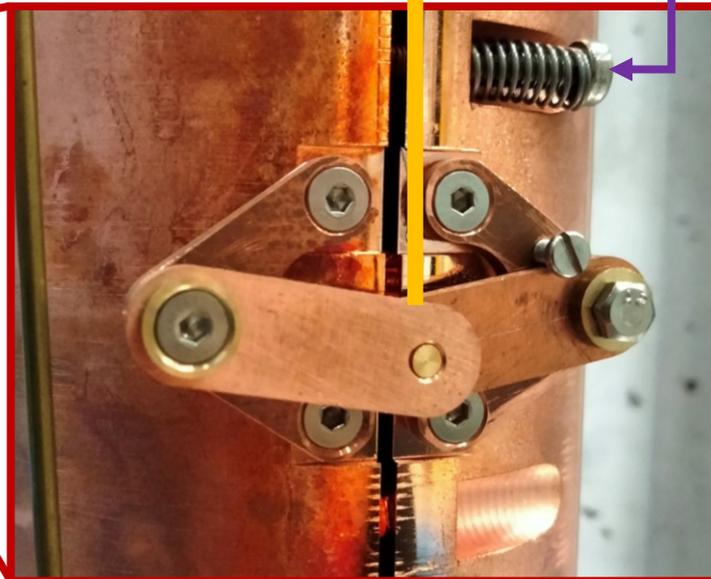
Cavity physical tuning:  $\sim 200$  MHz  
Allowed by tuning mechanism:  $\sim 90$  MHz  
Without **spurious modes**:  $\sim 80$  MHz  
**Probed** as of today:  $\sim 38$  MHz (not contiguous)

# Cavity tuning



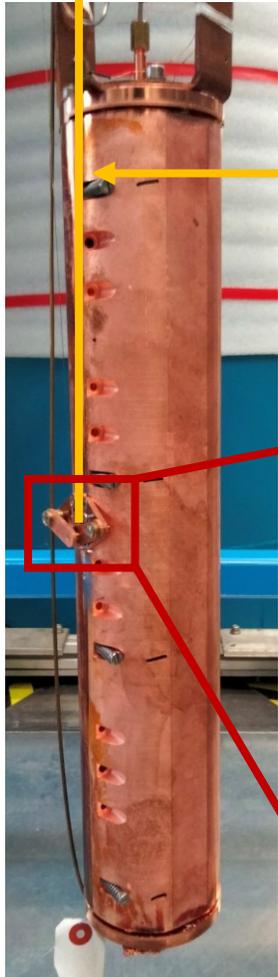
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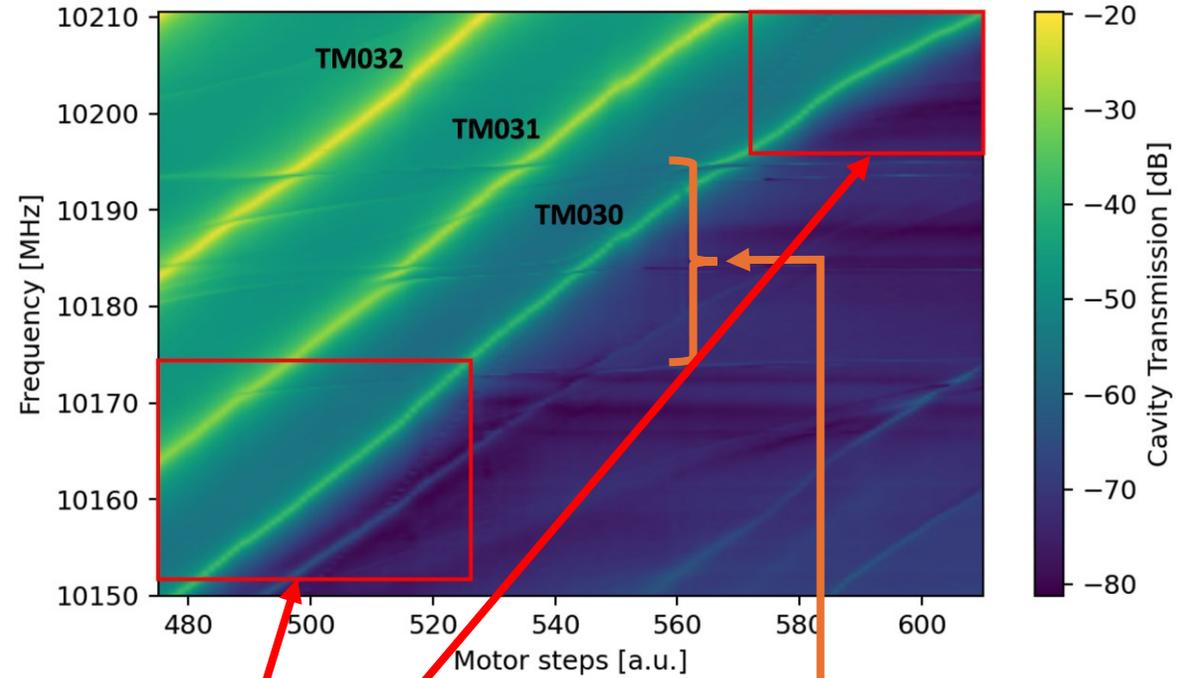
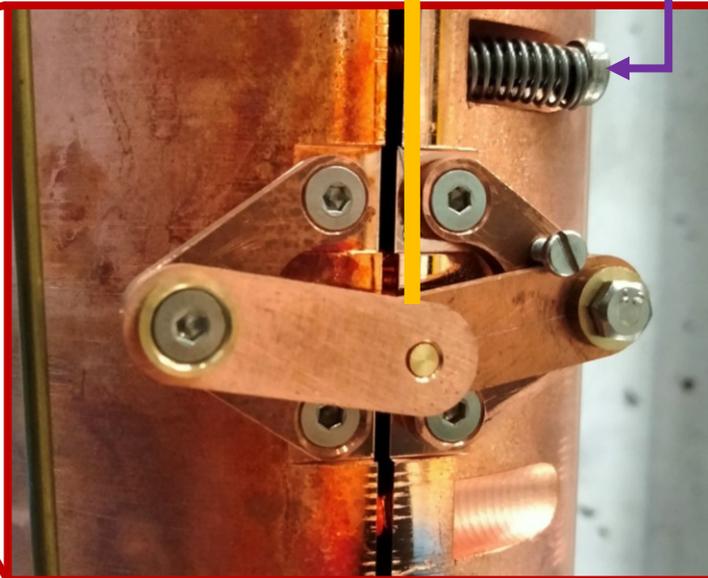
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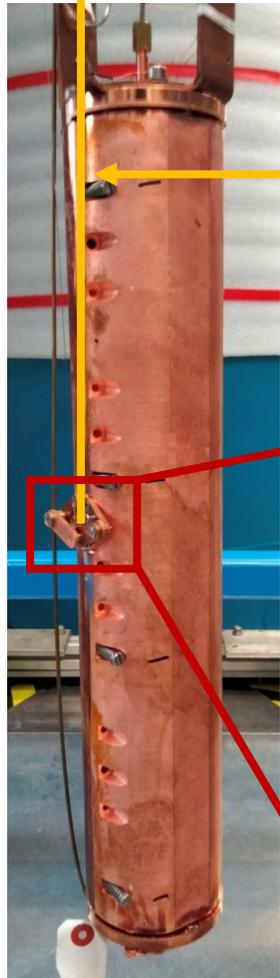
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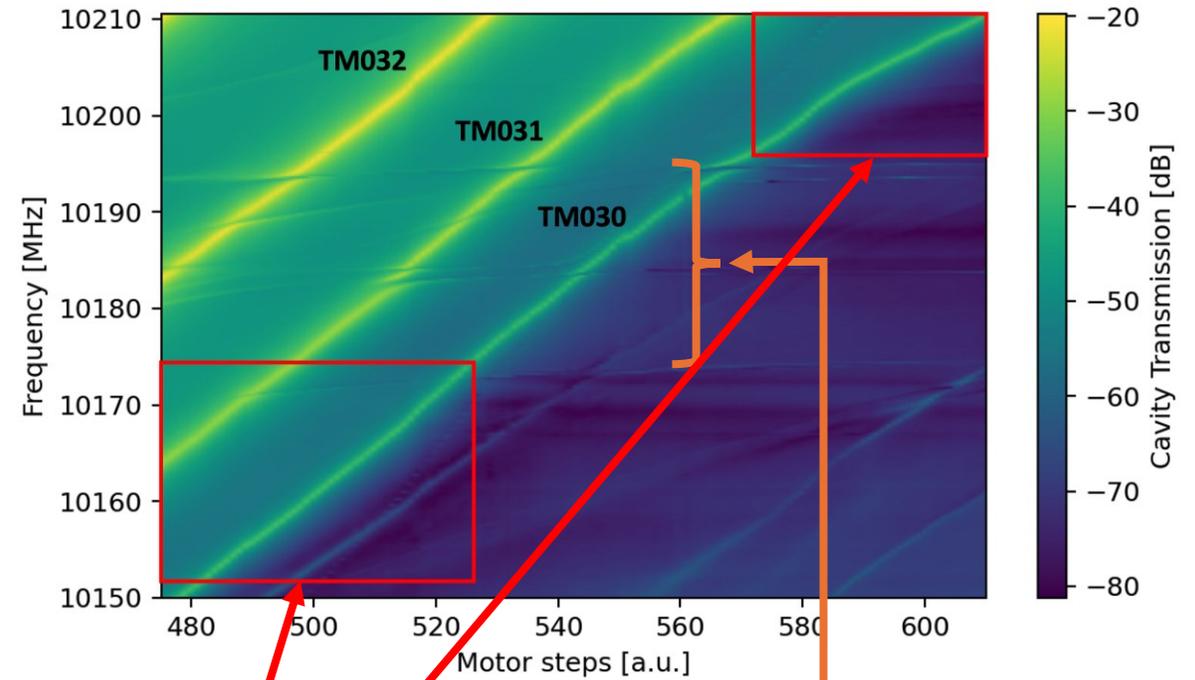
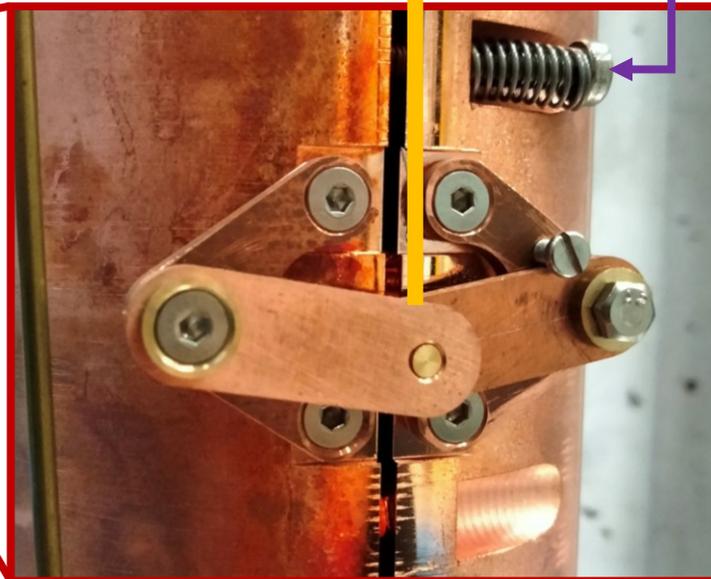
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# Cavity tuning



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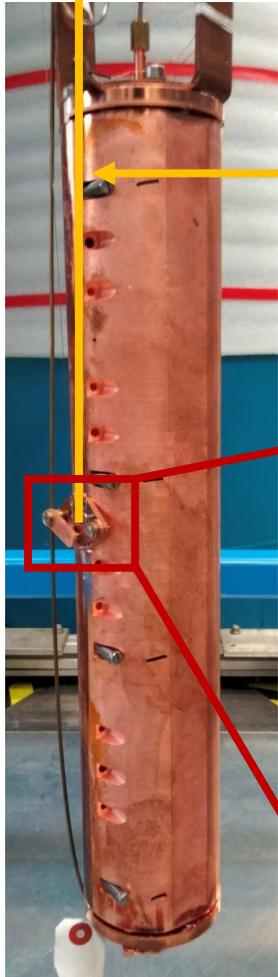
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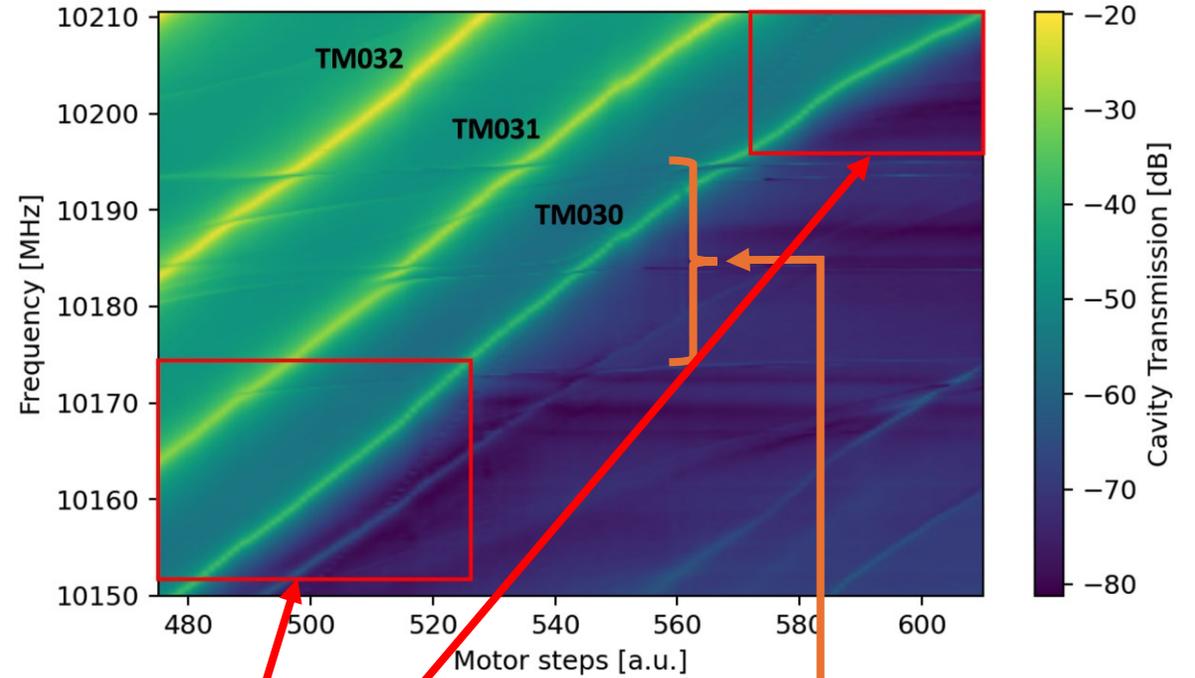
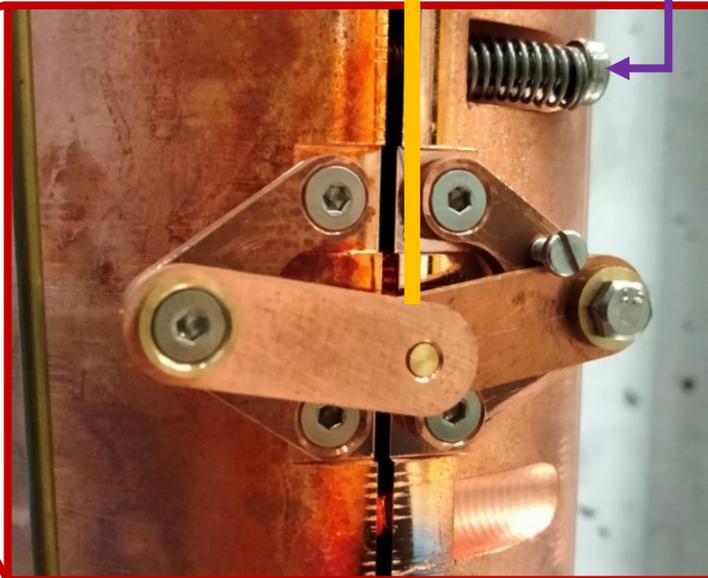
**BUT**

# Cavity tuning



## Cavity tuning system:

- A **pantograph** is connected via **wire** to room temperature stepper motor
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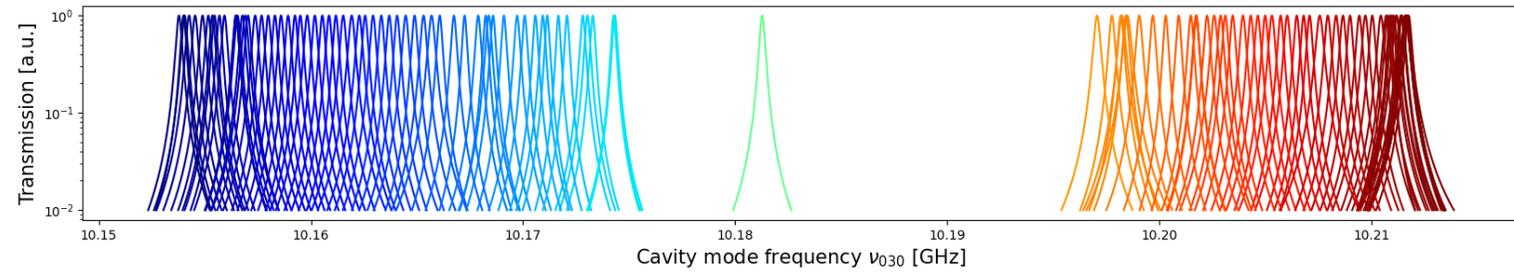


Cavity physical tuning:  $\sim 200$  MHz  
Allowed by tuning mechanism:  $\sim 90$  MHz  
Without **spurious modes**:  $\sim 80$  MHz  
**Probed** as of today:  $\sim 38$  MHz (not contiguous)

**BUT** New cavity tuning range:  $\sim 400$  MHz

# Data acquisition

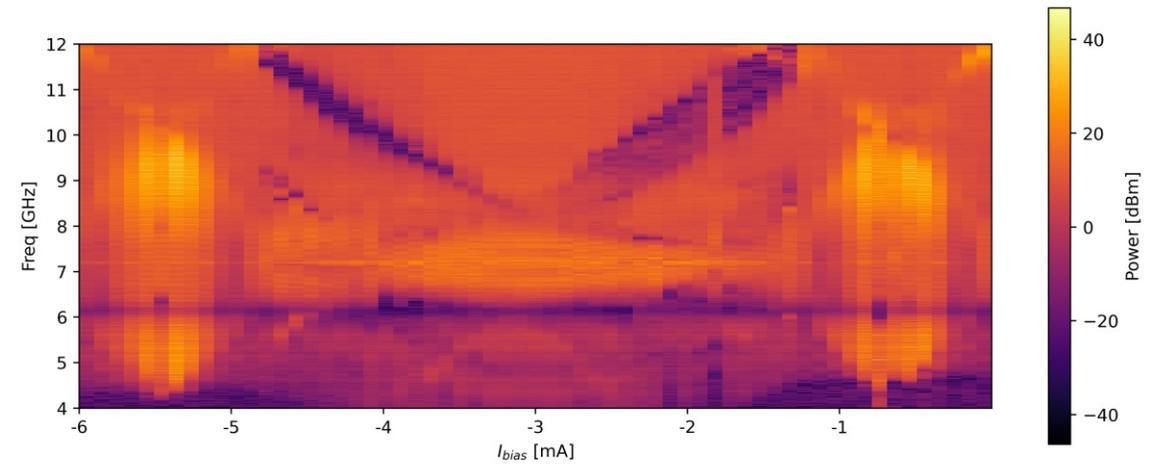
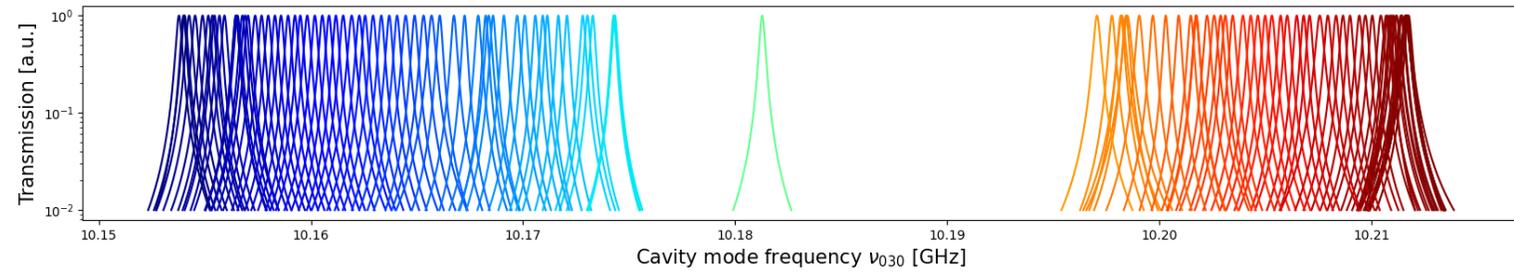
# Data acquisition



- Setting cavity frequency

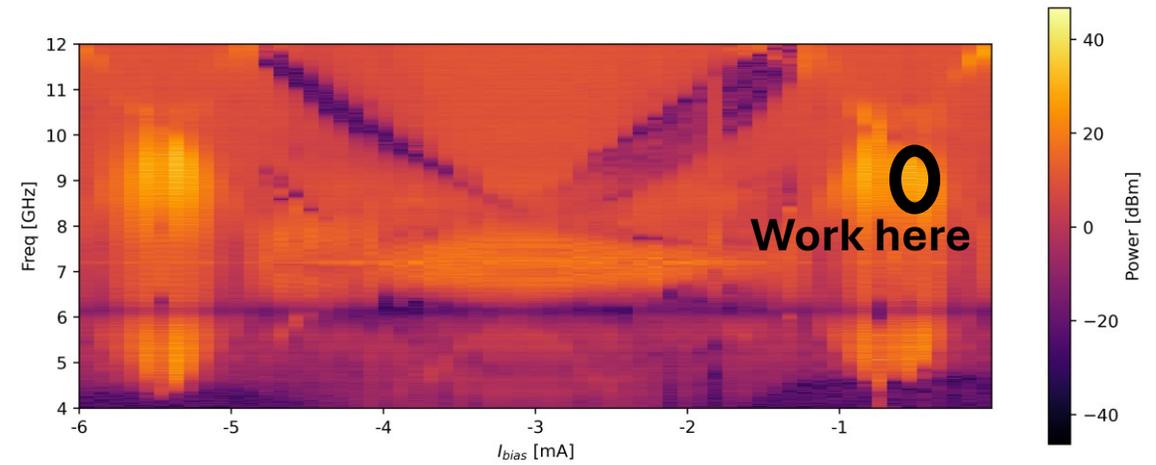
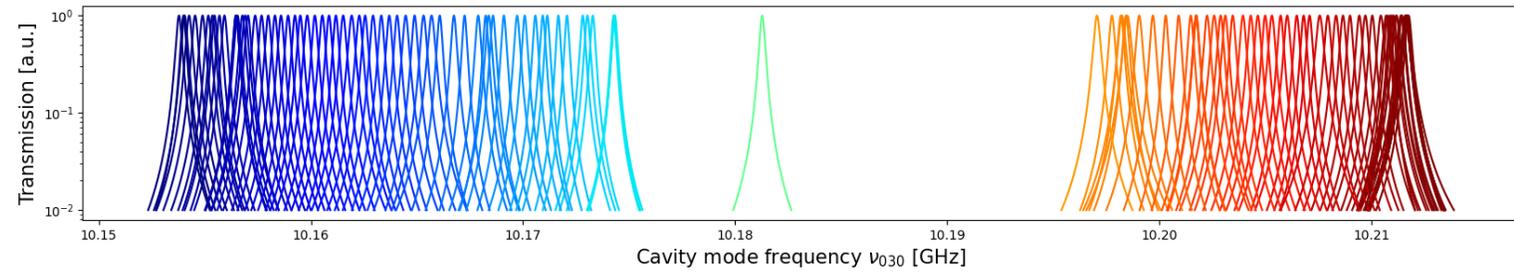
# Data acquisition

- Setting cavity frequency
- TWPA optimization



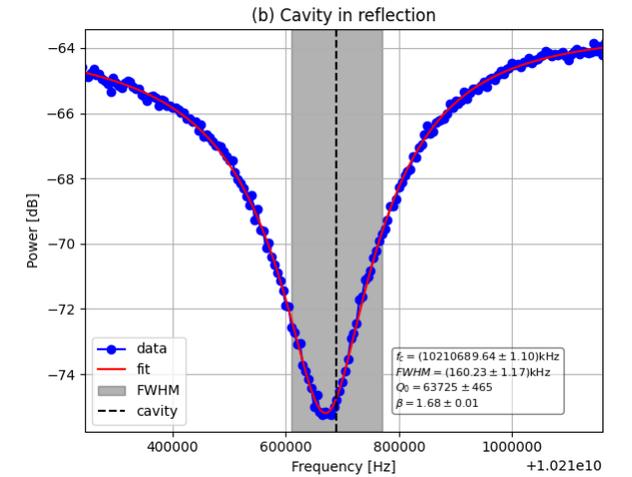
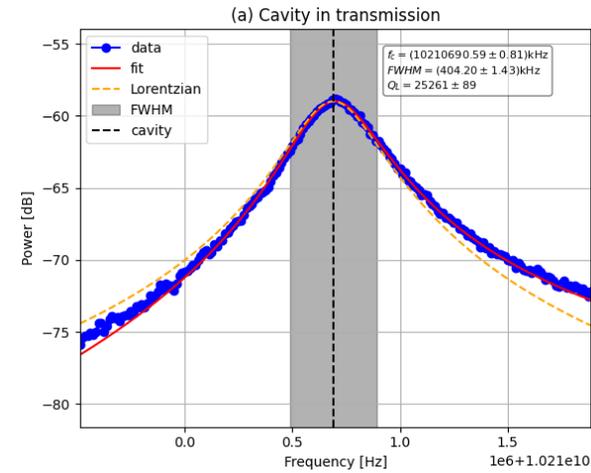
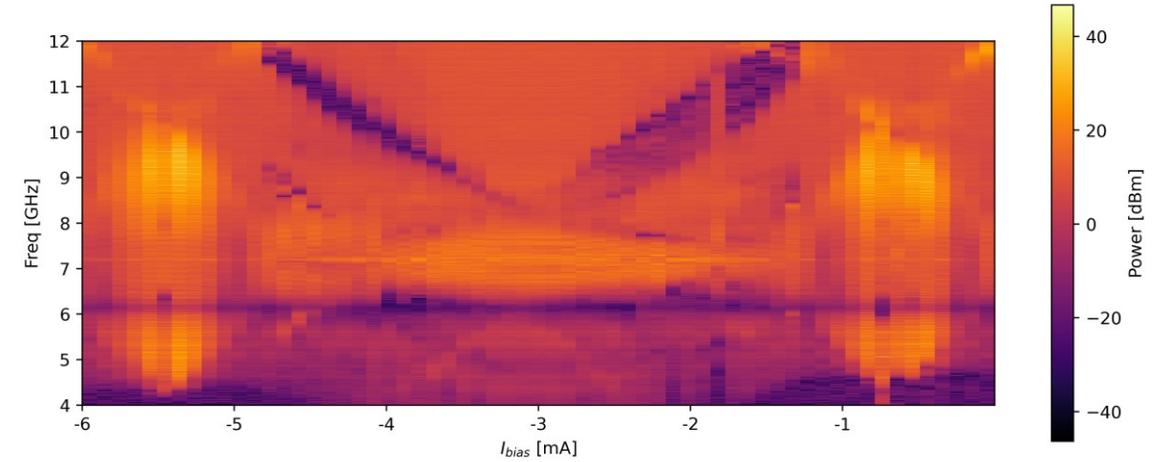
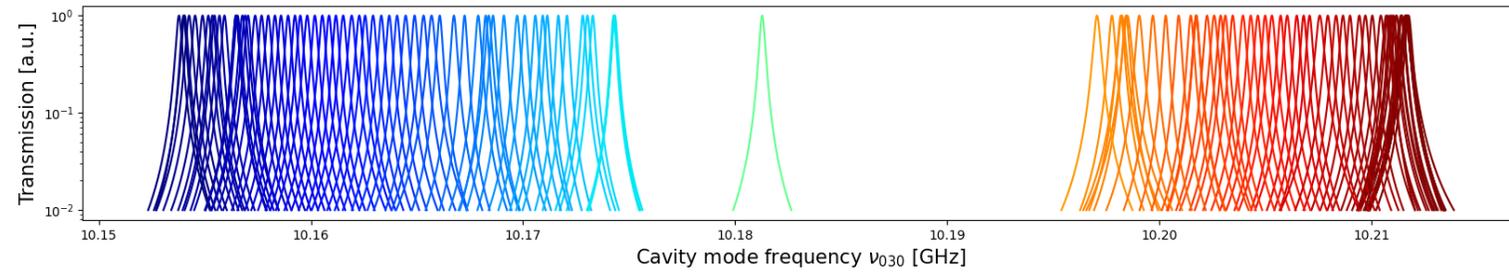
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- Setting cavity frequency
- TWPA optimization



# Data acquisition

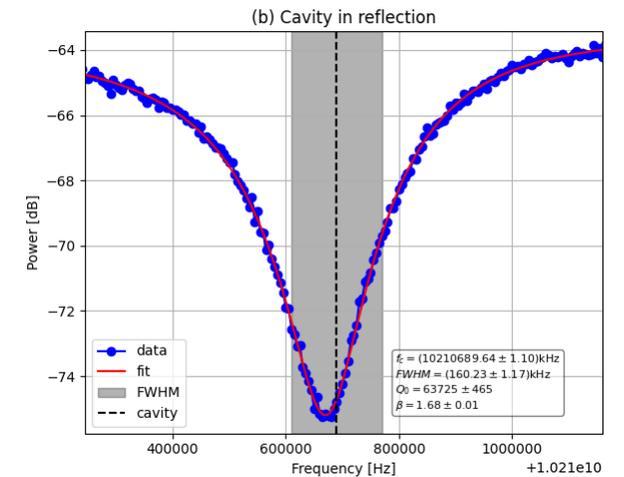
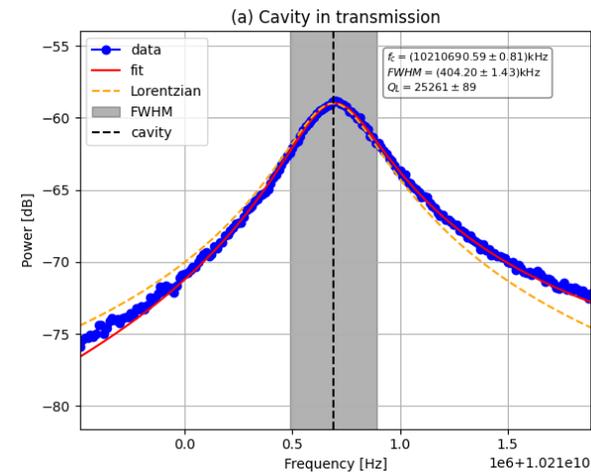
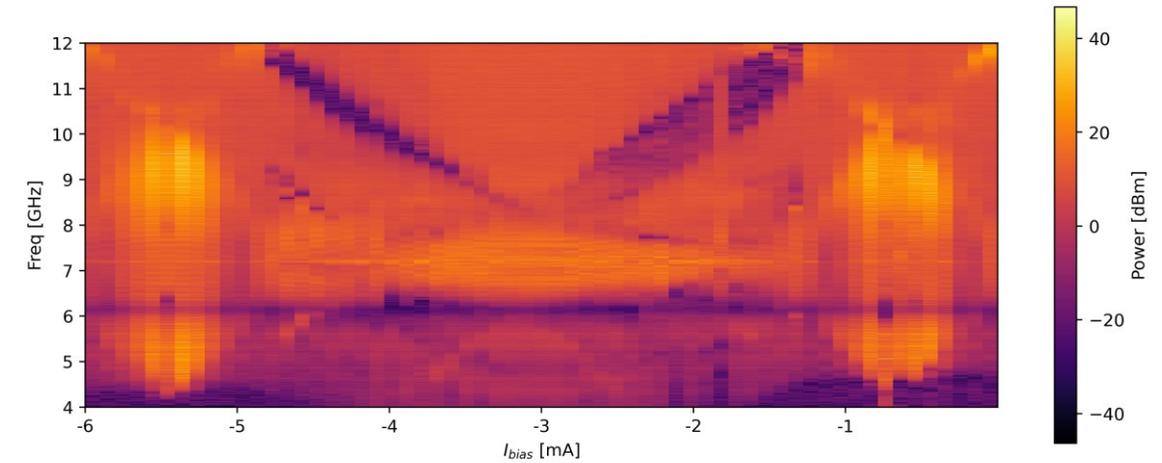
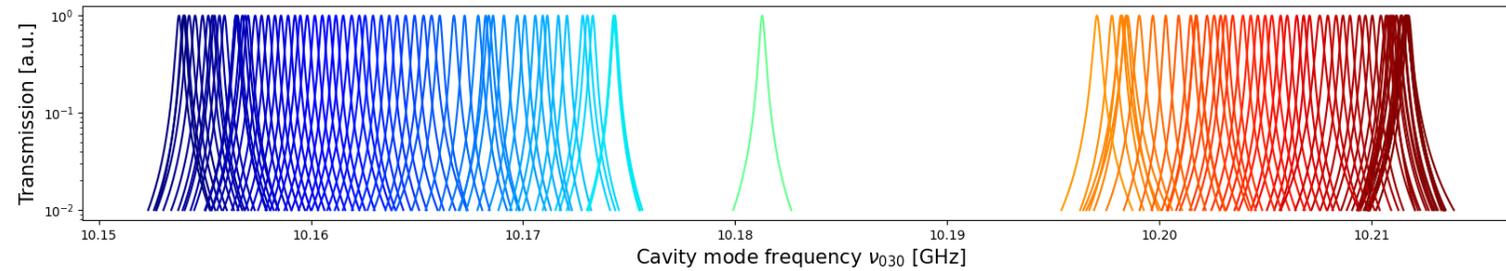
- Setting cavity frequency
- TWPA optimization
- Cavity parameter estimation



# Data acquisition

- Setting cavity frequency
- TWPA optimization
- Cavity parameter estimation
- Calibration of the lines and noise temperature estimation

C. Braggio et al.  
[Rev. Sci. Instrum. 93, 094701 \(2022\)](#)



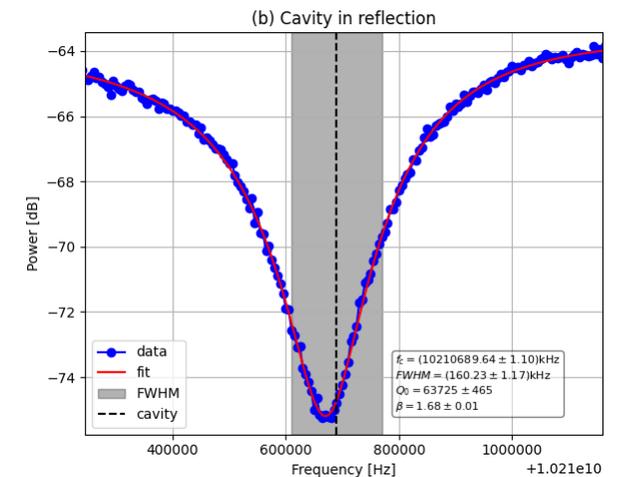
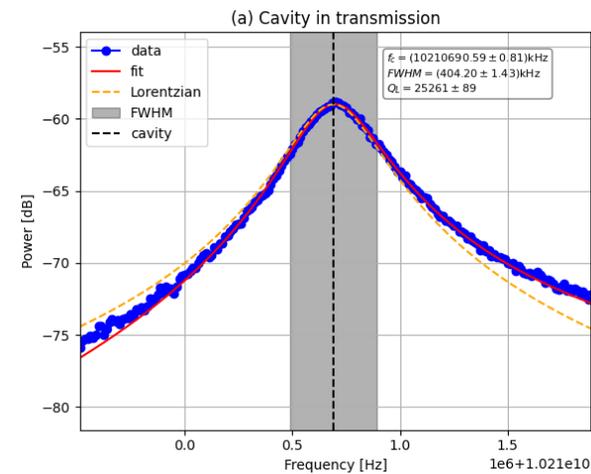
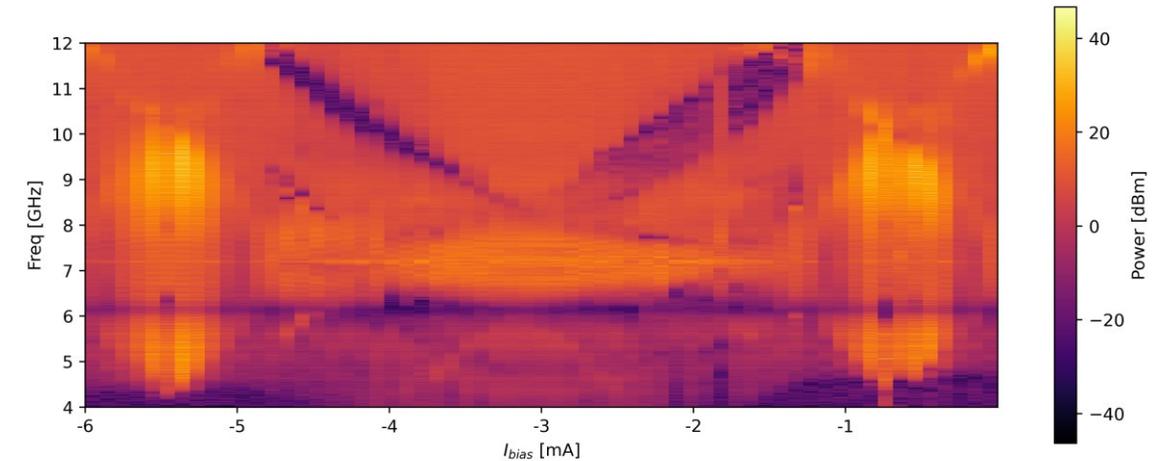
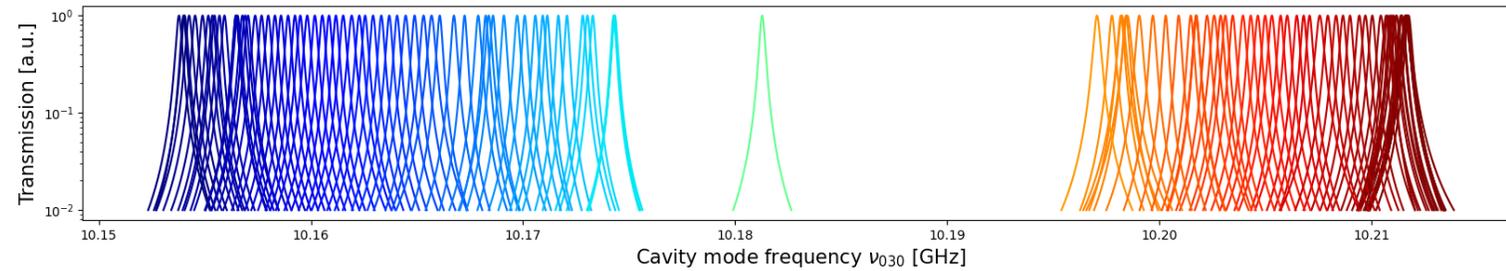
# Data acquisition

- Setting cavity frequency
- TWPA optimization
- Cavity parameter estimation
- Calibration of the lines and noise temperature estimation
- Acquisition and saving of both temporal data and 4 ms Fourier Transforms (FT)

C. Braggio et al.

[Rev. Sci. Instrum. 93, 094701 \(2022\)](#)

All data (temporal and FT) are saved on Cloud (~ 60 GB/hour) for further analysis



# Data acquisition

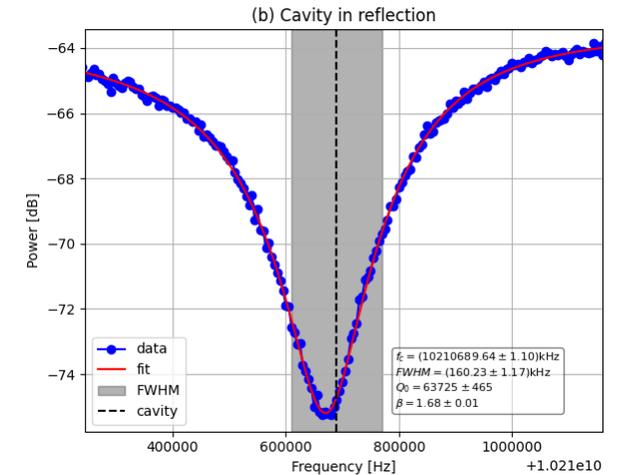
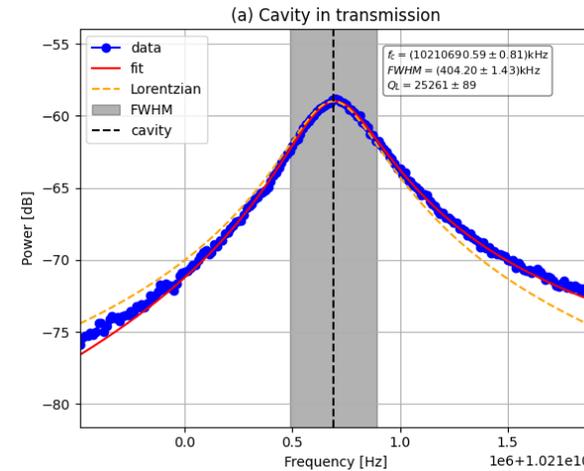
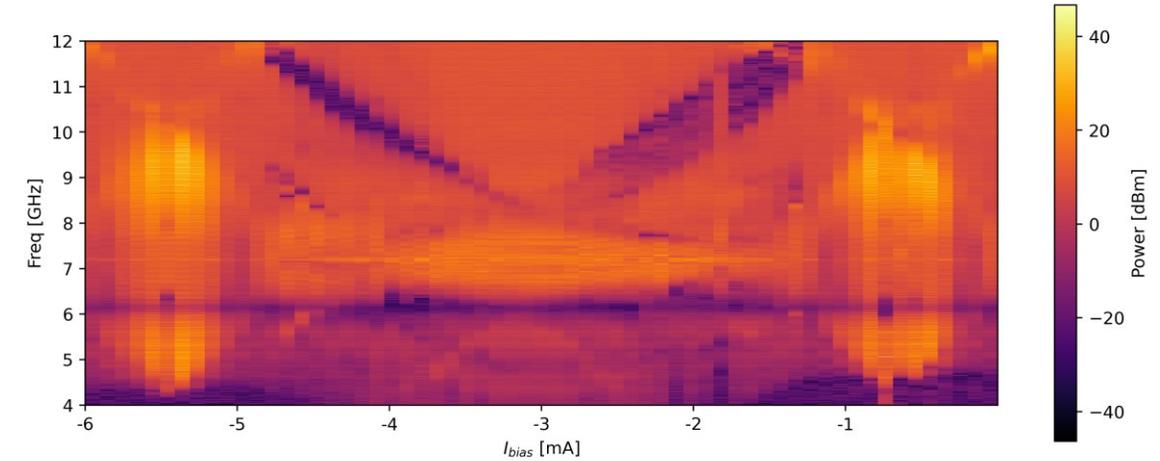
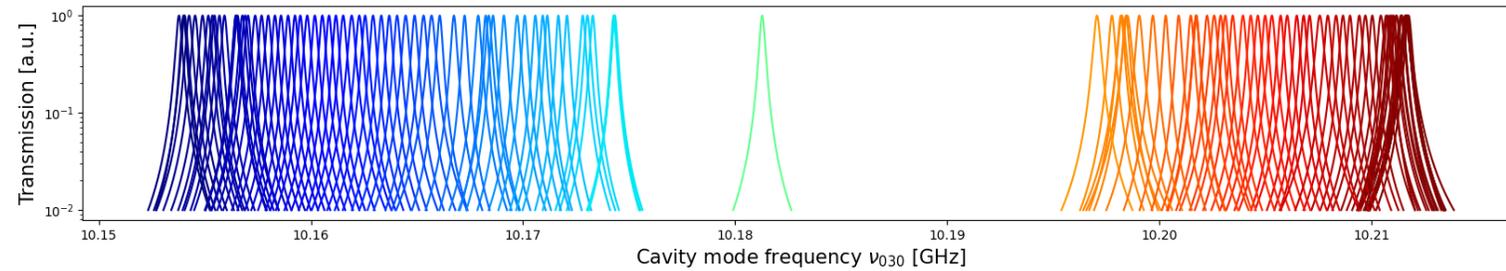
## Automation:

- ✗ • Setting cavity frequency
- ✗ • TWPA optimization
- ✓ • Cavity parameter estimation
- ✓ • Calibration of the lines and noise temperature estimation
- ✓ • Acquisition and saving of both temporal data and 4 ms Fourier Transforms (FT)

C. Braggio et al.

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# Data analysis

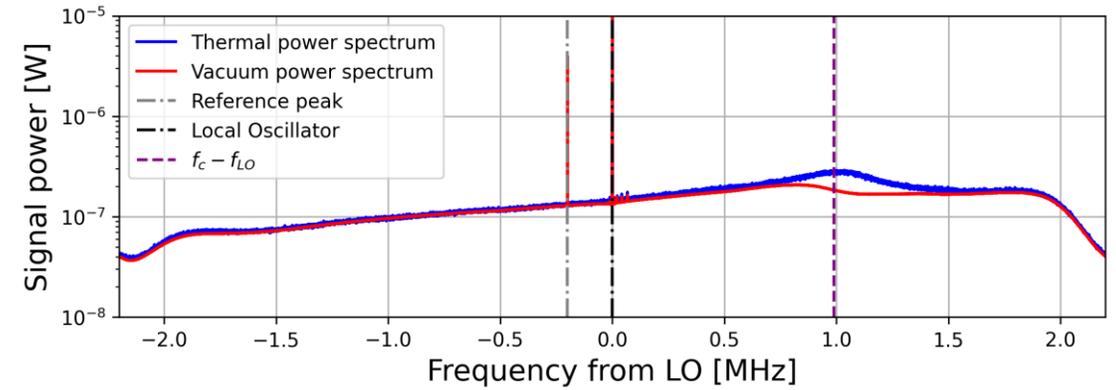
# Data analysis

Broad frequency region **analysis**:

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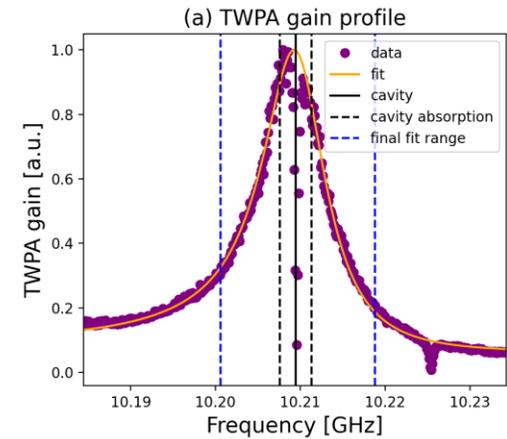
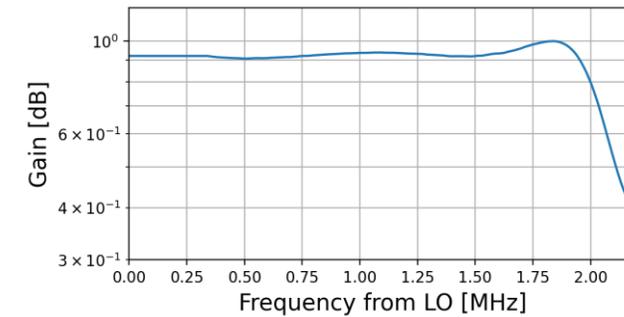
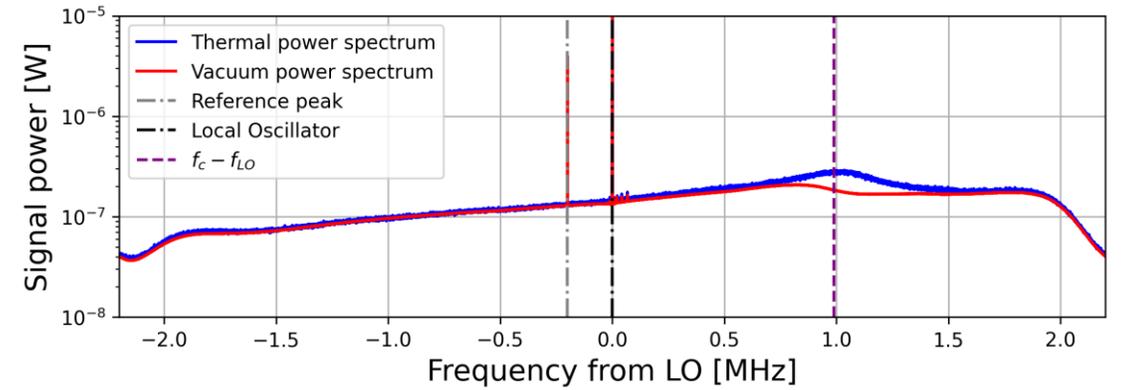
- Evaluate spectrum for each cavity aperture



# Data analysis

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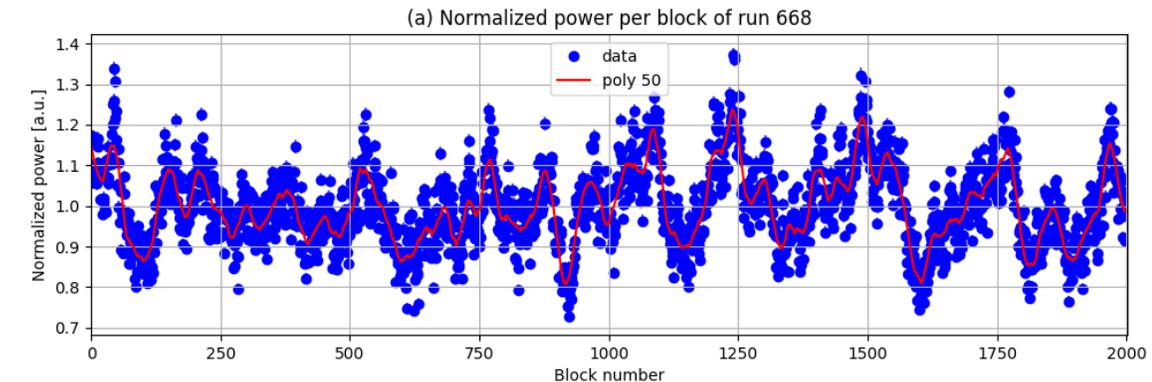
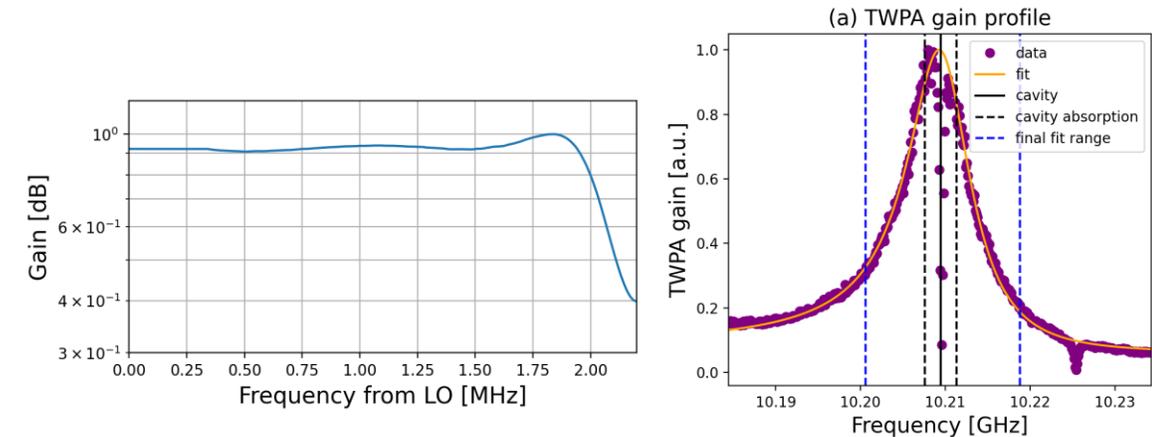
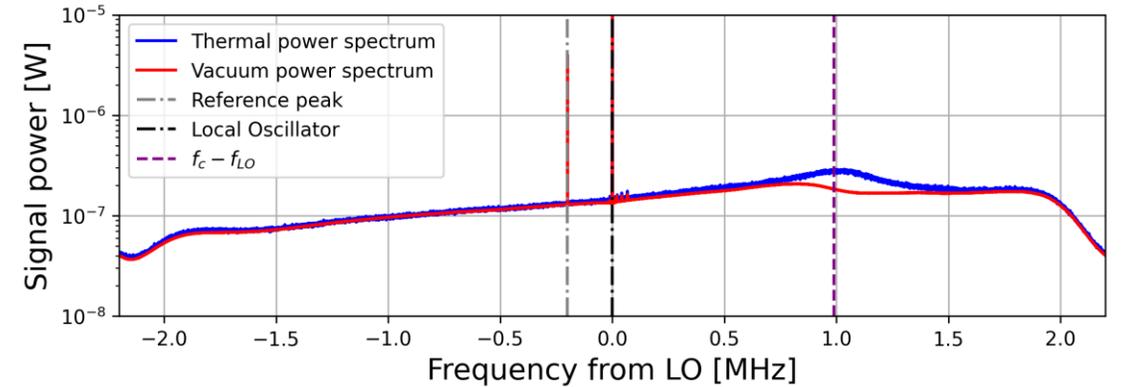
- Evaluate spectrum for each cavity aperture
- Removing frequency dependent gains of the lines



# Data analysis

Broad frequency region **analysis**:

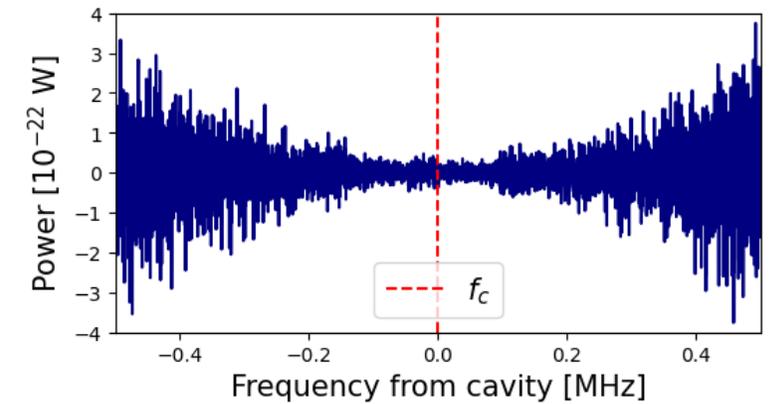
- Evaluate spectrum for each cavity aperture
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- Correct with time-dependent TWPA gain fluctuations



# Data analysis

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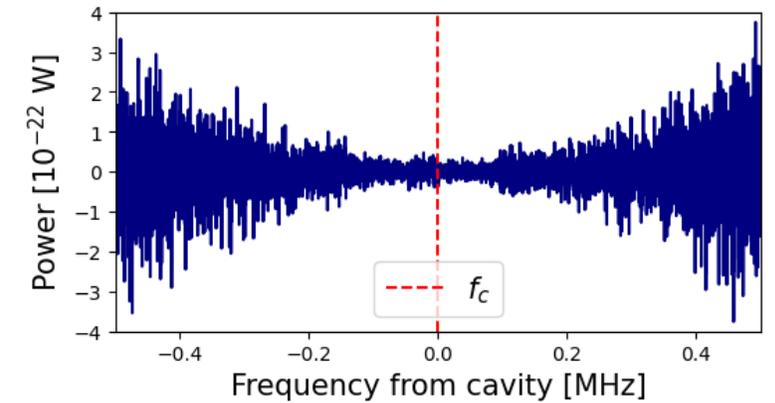
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- Remove Lorentzian cavity line shape



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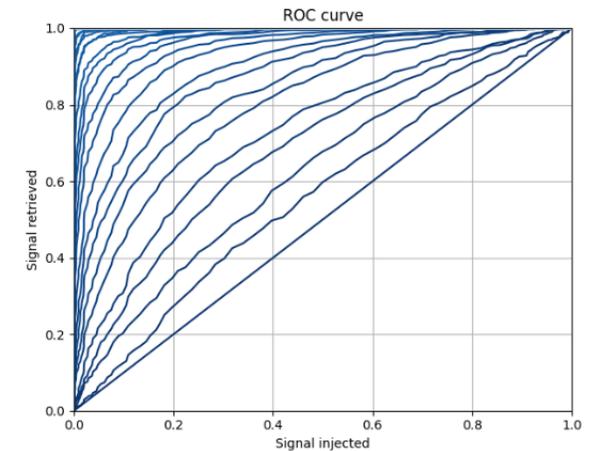
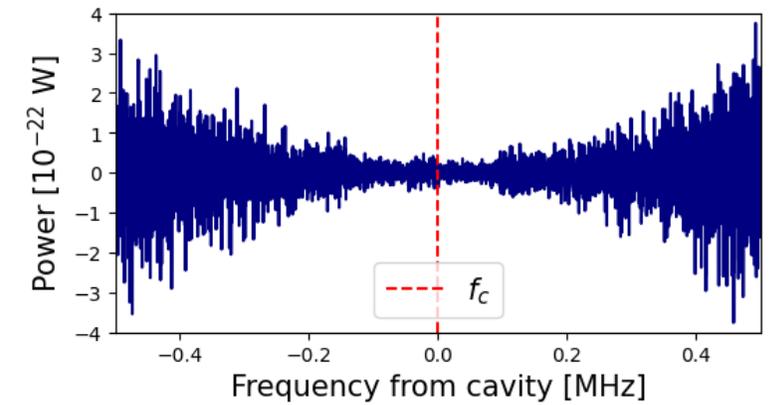
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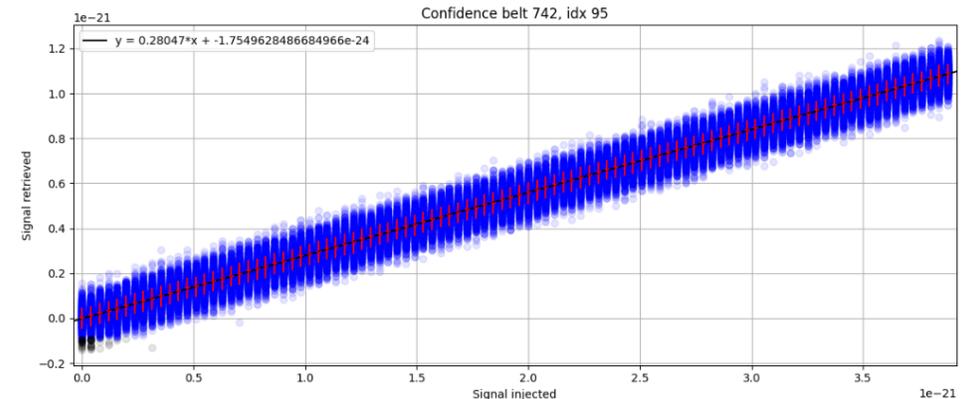
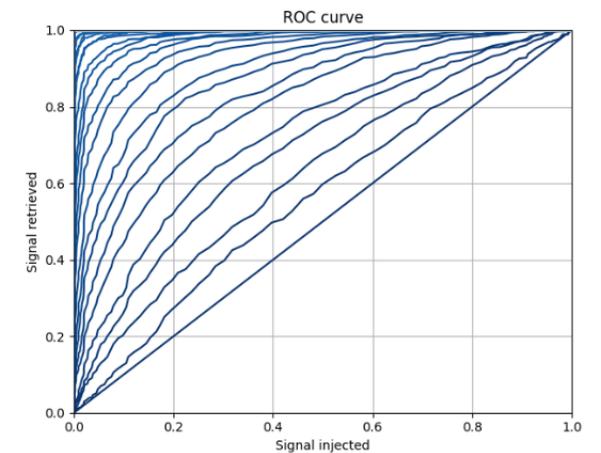
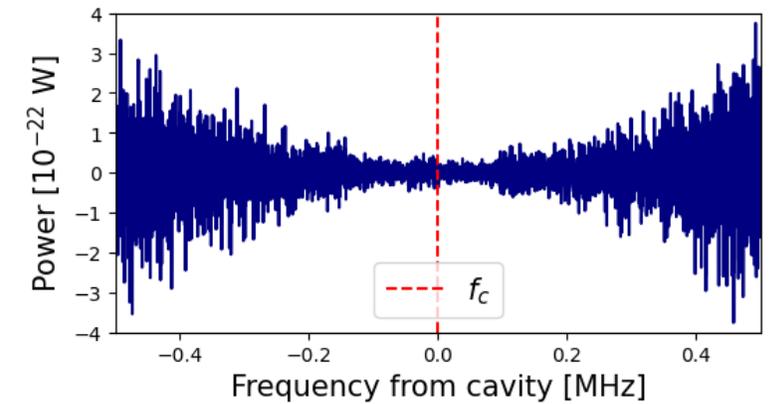
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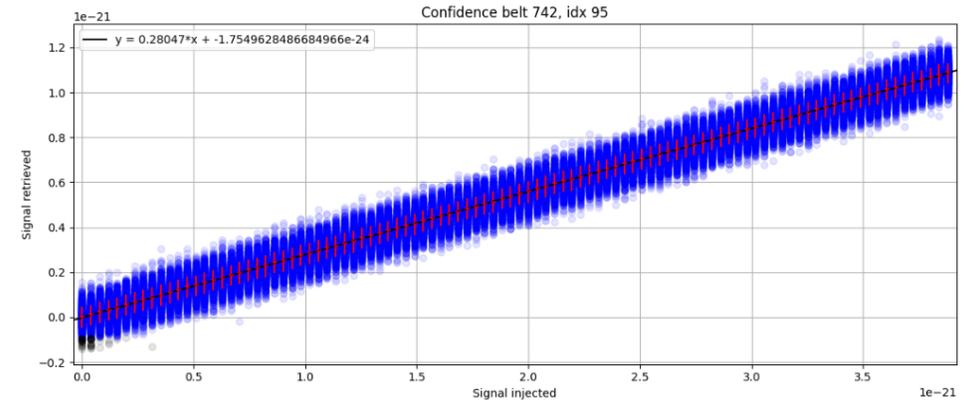
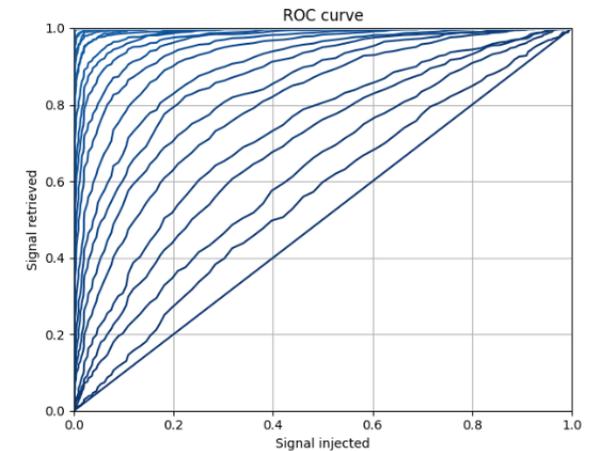
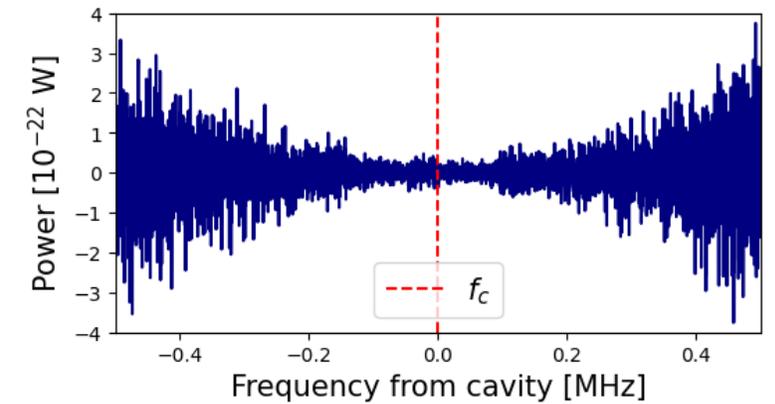
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- Combine results on single spectra to look for candidates



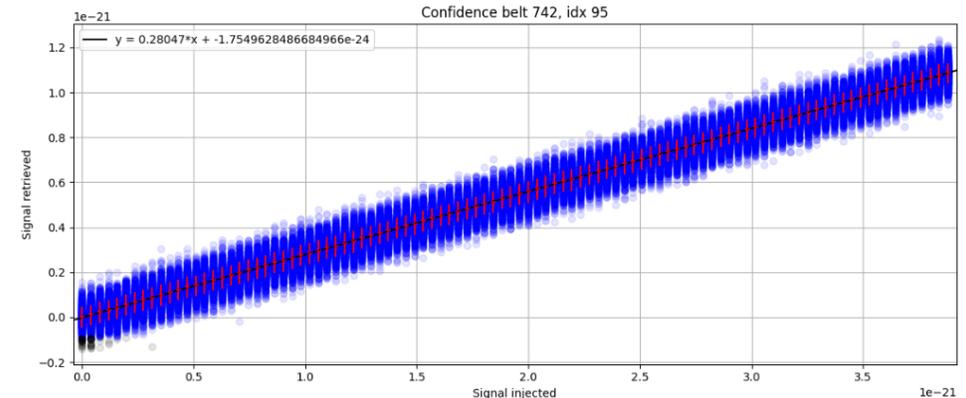
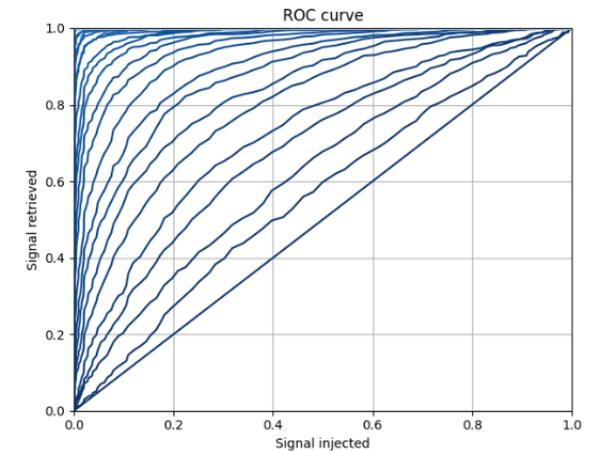
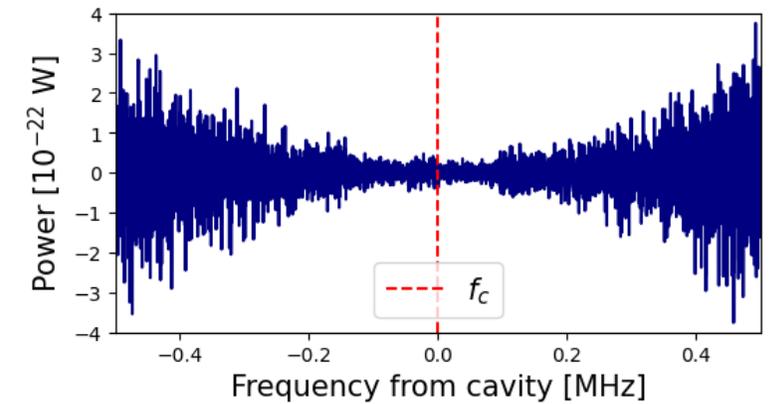
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- Correct with time-dependent TWPA gain fluctuations
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- Monte Carlo simulations to determine threshold (ROC) for axion sensitivity (Confidence Belt)
- Combine results on single spectra to look for candidates

Expectation with threshold set to  $4.5 \sigma$ :

- **One** axion candidate **per month**
- **No found axion candidate** in the 225 hours long acquisition

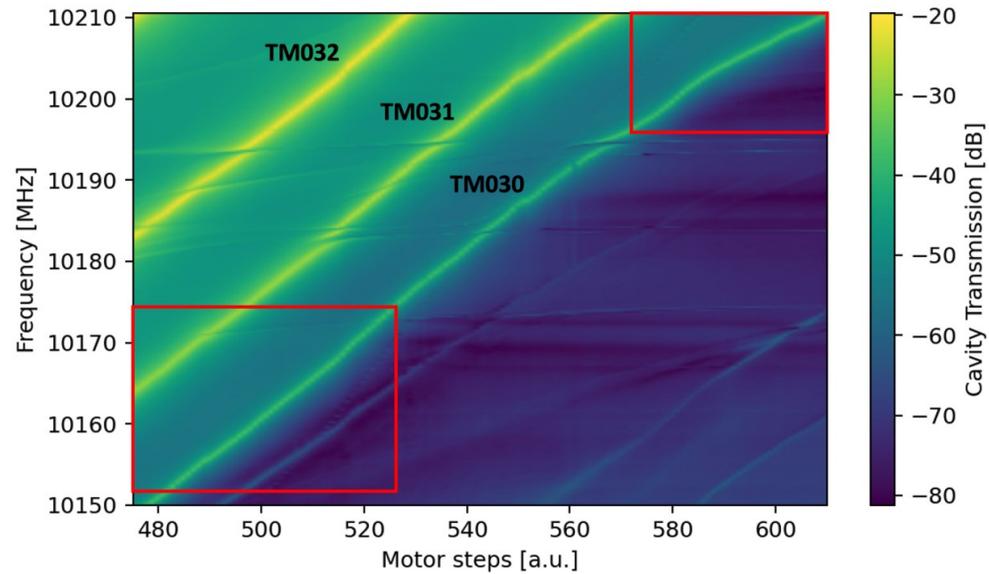


# Recent results

# Recent results

## Cavity tuning:

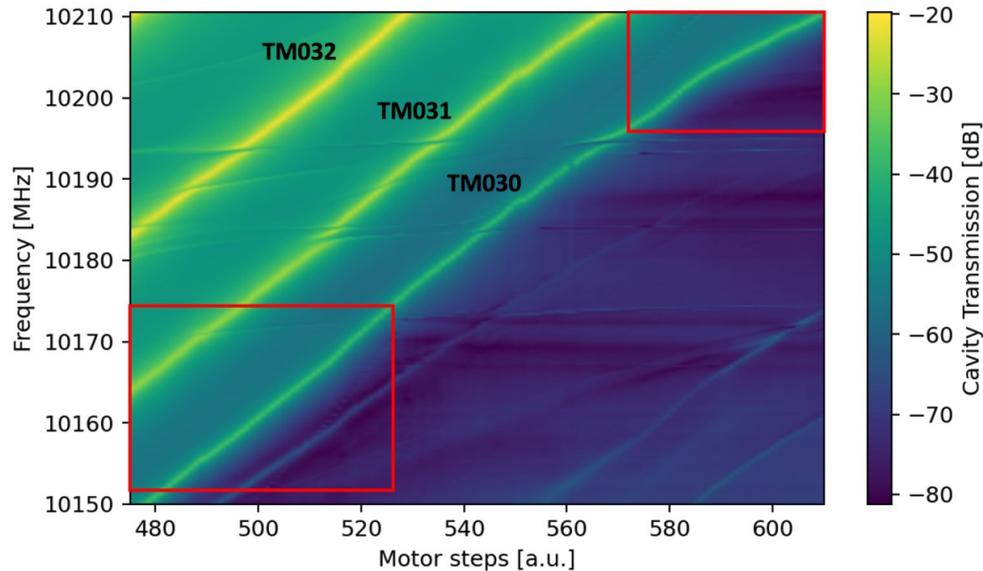
- Scanned  $\sim 38$  MHz
- $\sim 40$  MHz in upcoming run



# Recent results

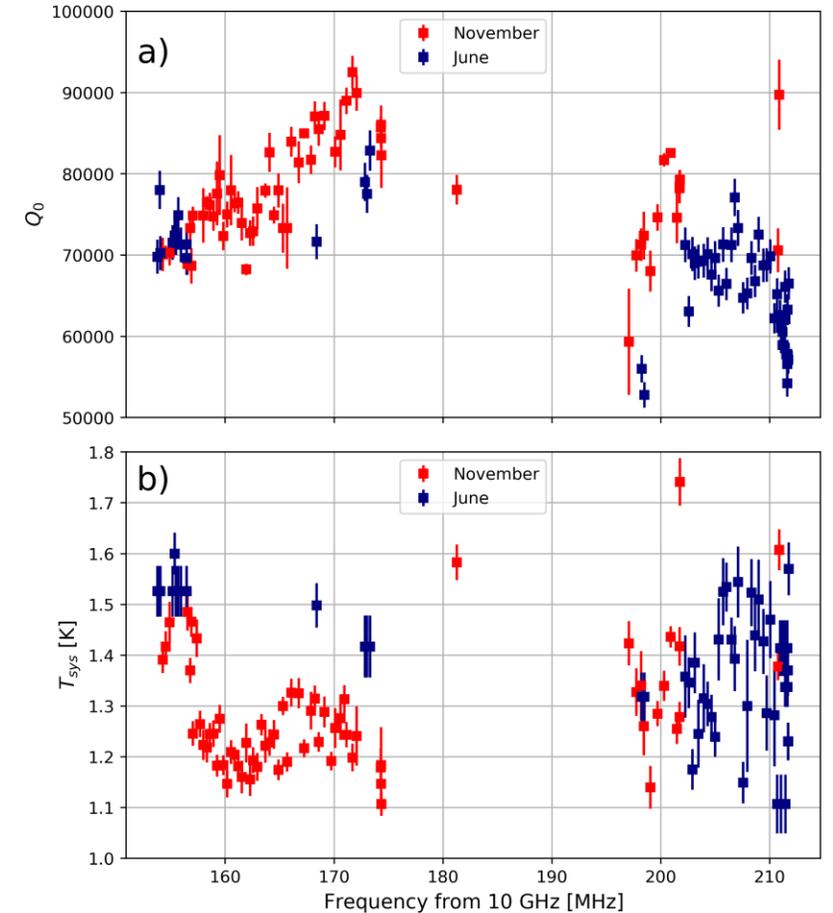
## Cavity tuning:

- Scanned  $\sim 38$  MHz
- $\sim 40$  MHz in upcoming run



## Stability:

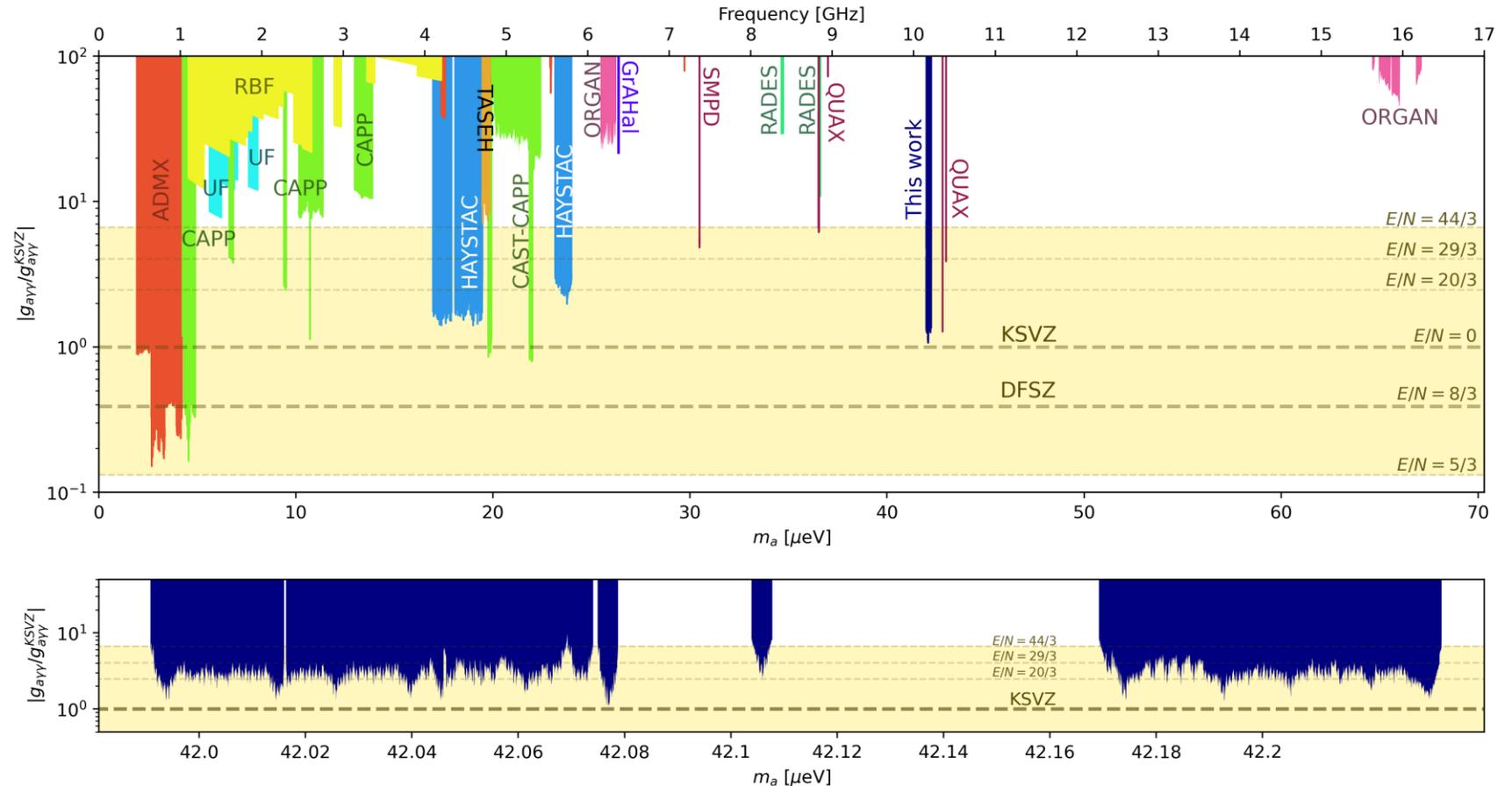
- Cavity quality factor not significantly impacted by tuning
- Noise temperature slightly above 2 SQL



# Recent results

## Exclusion limits:

- No rescan candidate  
=> All data used for exclusion plot
- Largest frequency interval ever probed  
@ QUAX-LNL: **we are ready to become an observatory!**



G. Sardo Infirri, G. Ruoso et al.  
[arXiv:2506.11589](https://arxiv.org/abs/2506.11589)

# Upcoming improvements and other research lines @ QUAX-LNL

## On main project:

- 14 T magnet installed in dry refrigerator (in an external re-liquefier system)
- New dielectrically loaded cavity with 400 MHz tuning
- Autopilot for TWPA optimization and cavity tuning
- Hardware axion signal injection

## Other ongoing QUAX projects:

- Single Microwave Photon Detector (SMPD) @ ~7.4 GHz
- Use of SMPD in a-e and a- $\gamma$  coupling based haloscopes

*C. Braggio et al.*  
[\*Phys. Rev. X\* \*\*15\*\*, 021031](#)

# Conclusions

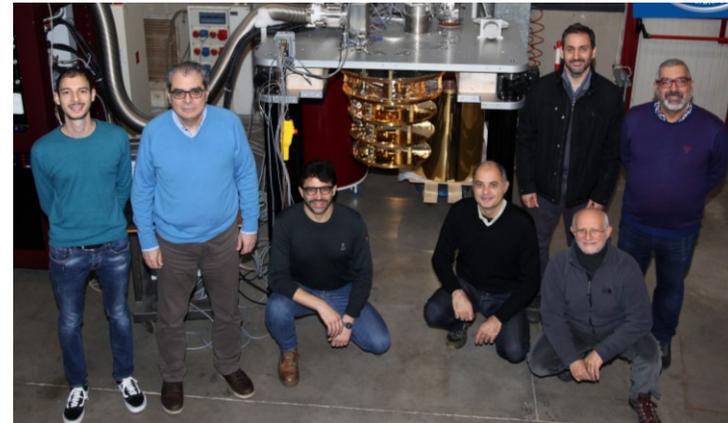
1. Overview of the QUAX apparatuses
2. Improvements in PD-LNL QUAX haloscope:
  - Dilution refrigerator stability
  - Low-noise electronics stable at  $\sim 2$  SQL noise temperature
  - Dielectrically loaded cavity with  $V \sim 1\text{L}$  and  $Q_0 \sim 75000$
  - Cavity tuning reliable in an 80 MHz region (soon expanded to 200 MHz)
3. Data acquisition and analysis:
  - Quasi-automated procedure
  - Simulations to set threshold
4. Results:
  - No rescan candidate
  - Exclusion plot in 38 MHz non-contiguous region with  $\sim 2 g_{\text{a}\gamma\gamma}$  sensitivity
5. Future improvements also concerning SMPD

# This is QUAX!



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# Thank You!

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**20<sup>th</sup> PATRAS workshop**  
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