# All-optical, self-aligned Compton source using a PW-scale laser-wakefield accelerator and a plasma mirror



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25.09.2025











### Self-aligned Compton scattering at ELI-NP and ZEUS

All-optical Compton scattering

ELI NP: Self-reflecting Compton scattering Gerstmayr et al., arXiv:2506.23718

ZEUS: Self-reflecting Compton scattering with density tailoring









# All-optical Compton scattering









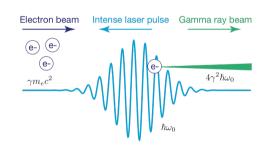
#### Inverse Compton scattering is a compact source of $\gamma$ -rays

#### Advantages:

- Length and time scales are matched and synchronised
- Co-location of accelerator and intense laser pulses

#### ► Challenges:

- Reproducibility of source
- Spatio-temporal overlap (drift and jitter)
- ► Beam quality





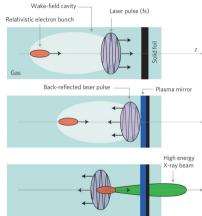






### Self-reflecting schemes enable higher collision rates

- Plasma mirror reflects driving laser pulse back onto electron beam
- High collision rate as self aligned (in time and space)
- ► Requires only one laser beam
  - $\rightarrow$  more accessible
- No independent control of scattering beam (typically lower intensities)
- Consider debris and back-reflection



Ta Phuoc et al., Nat. Photon. 6 (2012)

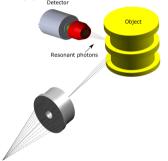


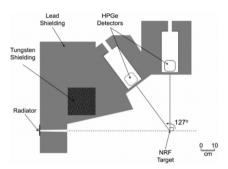






#### Examples of applications: Industry





Zilges et al., Prog. Part. Nucl. Phys 122 (2012); Quiter et al., PRC 86 (2012)

- ▶ Interrogation of hazardous materials (e.g. spent nuclear fuel)
- ► Transmutation of nuclear waste
- ► Imaging applications for high-Z materials









## Examples of applications: (Strong-field) QED

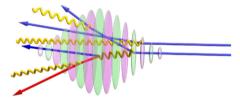
► Radiation Reaction



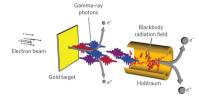
Blackburn, Rev. Mod. Plas. Phys. 5 (2020)

► Photon-Photon scattering

Pike et al., *Nat. Photon.* **8** (2014) Drebot et al., *PRAB* **20** (2017) ► QED Showers and Cascades



Blackburn et al., PRA 96 (2017)











# Self-reflecting Compton scattering

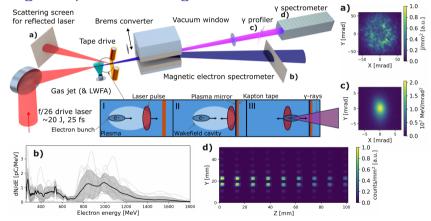








#### Self-reflecting Compton scattering at ELI-NP



 $\blacktriangleright$   $\langle E \rangle = 900 \pm 30$  MeV;  $\langle Q \rangle = 0.9 \pm 0.4$  nC; see arXiv:2506.23718 for details



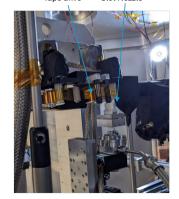


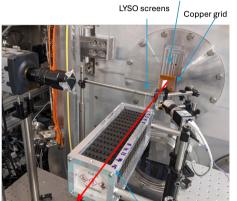




### Close-up of the diagnostics and interaction point

Tape drive Slot Nozzle





Al window (2 mm)

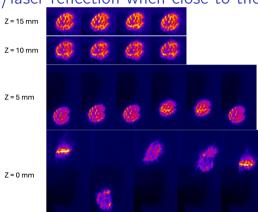








#### Gas flow affects tape/laser reflection when close to the jet



▶ Determine safe working distance at low power to avoid harmful back-reflections

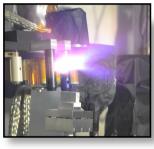








#### Plasma mirror in action







- ► Reproducible reflections using tape drive (Scitech/CLF, UK)
- ▶ Occasional EMP issues resolved by maximising distance to gas nozzle
- ► Shots 1/40 s due to gate valve/to check tape spooled (can operate at 1 Hz)

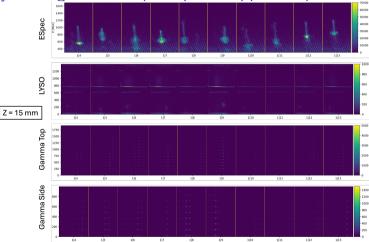








Tape 'far' away from gas nozzle (Compton suppressed)



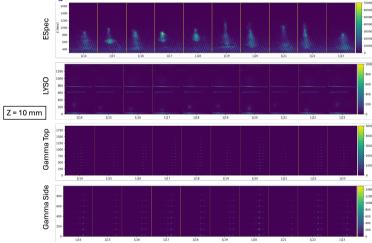








Moving closer to the jet



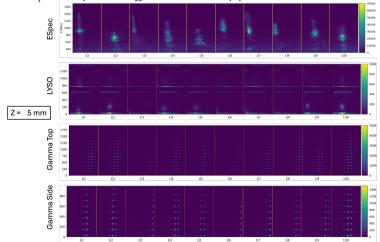








Increasing gamma/Compton signal as we approach



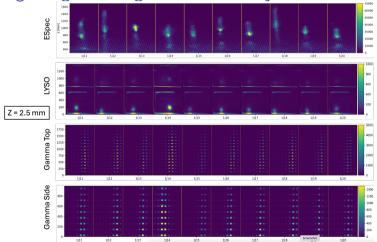








Significantly brighter gamma signal close to the jet



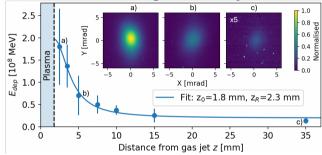








#### Gamma signal increases when moving closer to jet



- Decrease in gamma yield consistent with ICS
- ► Simple fit of Gaussian waist consistent with input parameters
- ▶ Brems (kapton) and betatron signal negligible at close distance
- lacktriangle Divergence of photon beam is  $\sim 4.0 \pm 0.5$  mrad FWHM

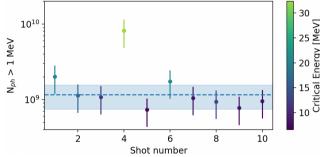








#### Shot-to-shot variations in performance (close shots)



- ▶ Forward fitting using  $dN_{\gamma}/d\epsilon_{\gamma} \sim \epsilon^{-2/3}e^{-\epsilon_{\gamma}/\epsilon_{crit}}$
- ▶ Average photon energy is few MeV ( $\epsilon_{crit}/3$ )
- $\triangleright$  Combine charge and brems calibration with  $\epsilon_{crit}$  for  $N_{ph}$
- Outlier: potentially variations in laser/plasma properties

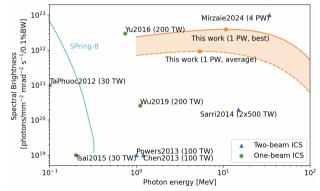








#### Photon source in context



- ▶ Combine measured  $\epsilon_{crit}$ ,  $N_{ph}$  and  $\theta$ , with simulated source size
- ► Source not fully optimised yet (energy, divergence...)
- Photons in bandwidth and average flux might be more important than brightness

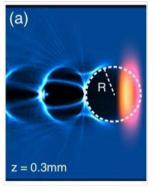








#### Simple model describes ICS yield for a given laser energy



$$\begin{array}{l} \text{Total} \\ \text{radiated} = \begin{array}{l} \text{Classical} \\ \text{radiated power} \times \begin{array}{l} \text{Beam} \\ \text{charge} \end{array} \times \begin{array}{l} \text{Interaction} \\ \text{duration} \end{array}$$

$$W_{ICS} pprox rac{2}{3} rac{e^2 m_e^2 c^3}{4\pi\epsilon_0 \hbar^2} \chi^2 N_b rac{ au}{2}$$

$$W_{ICS} \approx \left[ \frac{1024}{405\pi^2} \frac{r_e^2}{m_e c^4} \right] \omega_0^2 W_L^2 \frac{2a_0^2}{(a_{g,0})^3}$$

$$a_{g,0}=4$$
  $W_{ICS}\approx 4\times 10^{-4}W_L^2$   $a_0=\frac{a_{g,0}}{\sqrt{2}}$  Efficiency  $\eta_{ICS}=4\times 10^{-4}W_L$ 

- ► Assuming matched bubble regime (Lu et al. *PRSTAB* **10** (2007))
- ▶ Improve efficiency by enhancing  $a_0$ , e.g. density tailoring  $\rightarrow$  arXiv:2408.13238



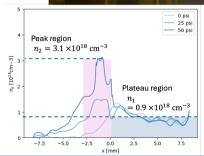




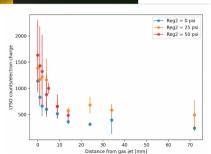


ZEUS Experiment to tailor plasma density and increase scattering intensity









Assuming only self-focusing affected



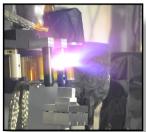


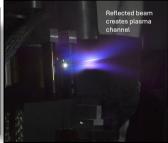




#### All-optical, self-aligned Compton source at ELI-NP and ZEUS

- ► High yield nonlinear inverse Compton scattering at ELI-NP
- ▶ Self-guiding at constant a0 in LWFA gives rise to simple scaling law
- $\triangleright$  Density tailoring can improve efficiency by increasing  $a_0$  at the interaction
- see arXiv:2506 23718 for details.







# Acknowledgements

Conor McAnespie, Gianluca Sarri, Matthew Streeter Centre for Light-Matter Interactions, Queen's University, Belfast, UK

Brendan Kettle, Laurence Bradley, Stuart Mangles The John Adams Institute for Accelerator Science, Imperial College London, UK

Yong Ma, Brandon Russel, Anatoly Maksimchuk, John Nees, Tanner Nutting, Paul Campbell, Alec Thomas, Karl Krushelnick Center for Ultrafast Optical Science, University of Michigan, USA

Paul Gellersen, Chris Murphy, Christopher Ridgers, Christopher Arran\* York Plasma Institute, University of York, UK

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## Backup slides

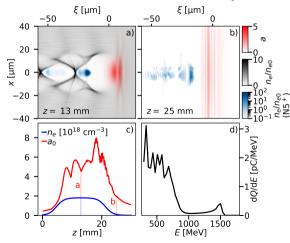








### Simulations to check consistency of results



- ► FBPIC using measured density profile (offline) and focal spot to match electron beam performance
- Predicted intensity  $a_0 \sim 2$  ( $a_0 \sim 4$  in plasma)
- ightharpoonup Electron beam size  $\sim 5 \mu m$
- ► Significant redshifting  $\langle \lambda \rangle \approx 1 \mu m$
- ► Ptarmigan simulations (Russell) match *a*<sub>0</sub> < 2
- A lot of variables, need more diagnostics