

ROUNTABLE WG4



Claudio Gatti

Marin Karuza

Deniz Aybas

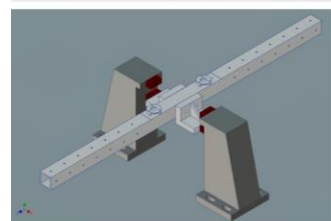
WG4 ACTIVITIES

172 MEMBERS

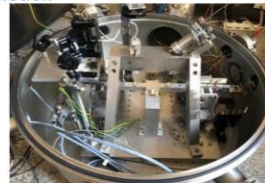
- Working Group Meeting – Feb Heidelberg
 - WG4 meeting – May --> The RES-NOVA neutrino observatory
 - Brainstorming meeting – April
 - General Meeting
 - Training School – hands on, CERN visit
 - White paper – finishing line (Claudio)
-

PARALLEL SESSION

The balance prototype



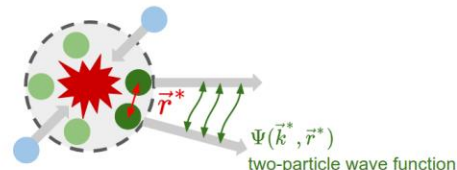
- 50 cm long arm with low momentum of inertia
- suspended through thin flexible joints (Cu-Be, 100 μm x 500 μm), similar to LIGO tiltmeters (Venkateswara et al. 2014)
- Balance center of mass positioned as close as possible to bending point ($\sim 10 \mu\text{m}$) to minimize the coupling with ground motion



- Resonance frequency $\sim 20\text{-}30$ mHz
- Installed at the Sar-Grav surface laboratories in Lula (NU) – Sos-Enattos mine



Femtoscropy
Koonin-Pratt equation



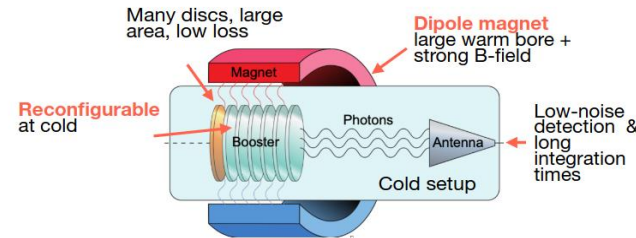
$$C(k^*) = \frac{N_{SE}(k^*)}{N_{ME}(k^*)} = \int S(r^*) \left| \Psi(\vec{k}^*, \vec{r}^*) \right|^2 d^3 r^* \xrightarrow{k^* \rightarrow \infty} 1$$

Relative distance and $\frac{1}{2}$ relative momentum evaluated in the pair rest frame

Lisa et al.
Ann. Rev. Nucl. Part. Sci. 55:357-402, 2005

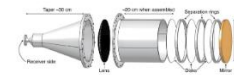
- Measure $C(k^*)$, fix $S(r^*)$, study the interaction.
- Detailed studies of the source in pp:
ALICE Coll. Phys. Lett. B 811 (2020) 135849
- CATS framework to evaluate the above integral Mihaylov et al. EPJC 78 (2018) 5, 394

On the path to full-scale MADMAX



➔ Prototyping phase since 2020 with physics runs @CERN (MORPURGO 1.6 T magnet)

ø200 mm closed prototype



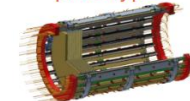
17-day run @18.5, 19.2 GHz with peak $|g_{ay}| \approx O(3 \times 10^{-11}) \text{ GeV}^{-1}$

ø100 mm closed prototype



20-hour cold run @ ≈ 19 GHz and $T < 10$ K under B field (Analysis ongoing)

ø300 mm open prototype



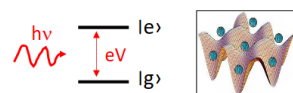
(2026) 3 discs $|g_{ay}| \approx O(10^{-12}) \text{ GeV}^{-1}$
(2027-2029) 20 discs $|g_{ay}| \approx O(10^{-13}) \text{ GeV}^{-1}$

ATOMIC CLOCKS & GRAVIMETERS

TIME

Optical atomic clocks

internal d.o.f., electronic states

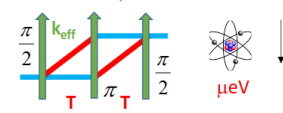


relative precision $< 10^{-20}$

GRAVITY

Atom interferometers (gravimeters/gradiometers)

external d.o.f., momentum states



relative precision $< 10^{-11} \text{ g}$

ultralight DM search

Table-top, low-energy experiments, high precision

High sensitivity to: **accelerations, spin precession, variation of fundamental constants**

10/09/2025

NICOLA.POLI@UNIFI.IT

3

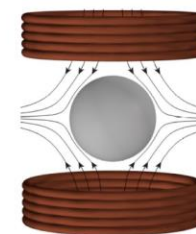
Towards a Hunt for ALPs Sourced by Pulsars

Jackson Fowler
Magnetic Levitation Group

CA21106 General Meeting
Sofia, Bulgaria
10.09.2025

MBI Marietta Blau
Institute for
Particle Physics

ÖAW AUSTRIAN
ACADEMY OF
SCIENCES



PIFNARY SESSION

Preliminary ALPs sensitivity projections for the RES-NOVA Observatory

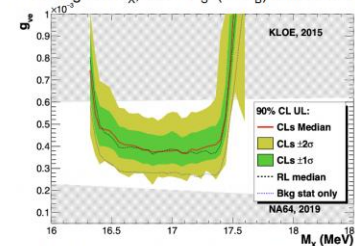


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PADME Expected sensitivity

- Evaluate expected 90% CL UL in absence of signal
- Modified frequentist approach, LEP-style test statistic
- Likelihood fits performed for the separate assumptions of **signal + background vs background only**, define Q statistic based on Likelihood ratio: $Q = LS + B(g_{\gamma\gamma}, M_X) / LB$. The likelihood includes terms for each nuisance parameter pdf
- For a given M_X , $CLs = P_S / (1 - P_B)$ is used to define the UL on $g_{\gamma\gamma}$



Source	Uncertainty [%]
N ₂	0.6
B	0.35
N _{POT}	0.55
TOTAL on g _{γγ}	0.88
TOTAL on K(s)	2.1

Pseudo data (SM background) is generated accounting for the expected uncertainties of nuisance parameters + statistical fluctuations

Probing New Physics with Slow and Trapped Molecules

Steven Hoekstra, University of Groningen and Nikhef, The Netherlands



Universität Hamburg
DER FORSCHUNG | DER LEHRE | DER BILDUNG

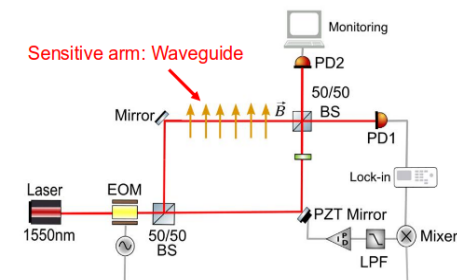
CLUSTER OF EXCELLENCE
QUANTUM UNIVERSE



WISPFi (WISP searches on a Fiber Interferometer)

- $\gamma \rightarrow a$ conversion inside a **waveguide** measuring **photon disappearance** in the presence of B_{ext} [4].
- Axion conversion probability scales with [5]:
 $\text{For } P_{\gamma \rightarrow a} \ll 1: P_{\gamma \rightarrow a} \propto g_{a\gamma\gamma}^2 (BL)^2$
- Mach-Zehnder interferometer (MZI)** with the sensing arm inside the magnetic field.
- Light guiding & **resonant detection**.
- Signal: **amplitude reduction & phase shift**.

- No ρ_{DM} dependence.
- Room temperature operation



TECHNOLOGY FORUM: HIGH MAGNETIC FIELDS FOR WISPS SEARCHES

23–24 FEB 2023

LABORATORI NAZIONALI DI FRASCATI (ROME), ITALY

Introduction

Aula Touschek, Laboratori Nazionali di Frascati (Rome), Italy

14:30 - 14:40

The Babylaxo Magnet

Matthias Mentink



Aula Touschek, Laboratori Nazionali di Frascati (Rome), Italy

14:45 - 15:15

The MADMAX Magnet

Walid Abdel Maksoud



Aula Touschek, Laboratori Nazionali di Frascati (Rome), Italy

15:20 - 15:50

The GrAHal project within the context of the Grenoble high magnetic field facility

Pierre Pognat



Aula Touschek, Laboratori Nazionali di Frascati (Rome), Italy

15:55 - 16:25

IRIS

Lucio Rossi



Aula Touschek, Laboratori Nazionali di Frascati (Rome), Italy

16:30 - 17:00

HIGH MAGNETIC FIELDS



Finanziato
dall'Unione europea
NextGenerationEU



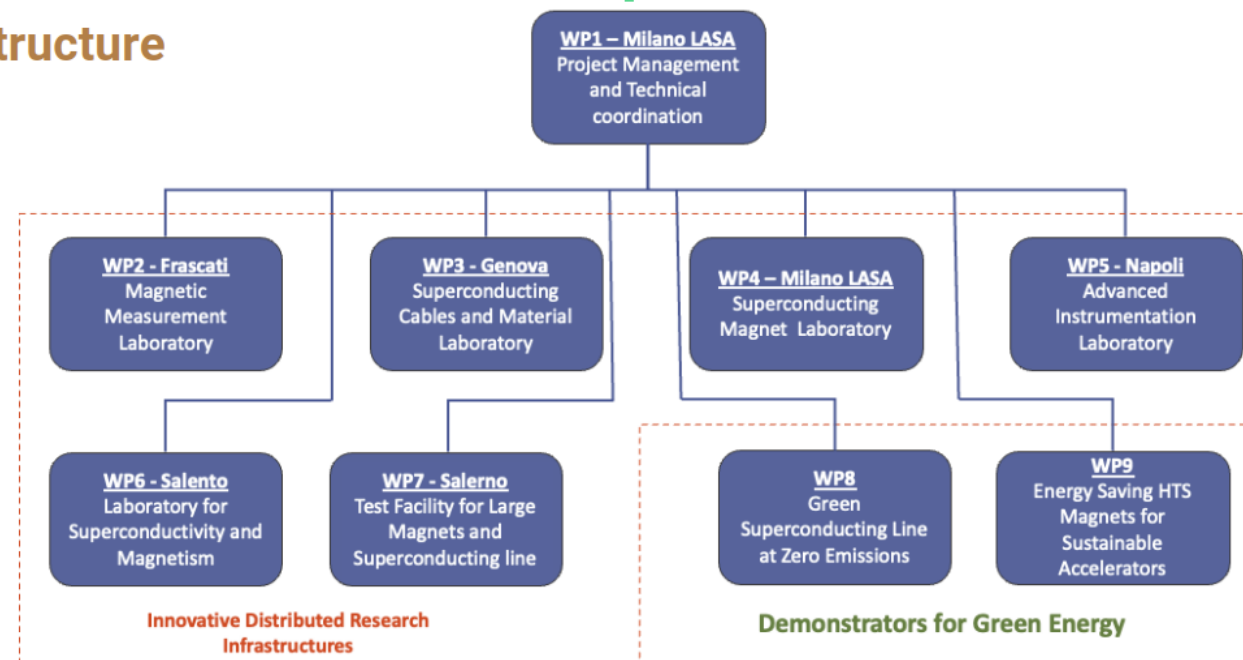
Ministero
dell'Università
e della Ricerca



Italiadomani
PIANO NAZIONALE
DI RIPRESA E RESILIENZA



IRIS Structure



24-Feb-2023

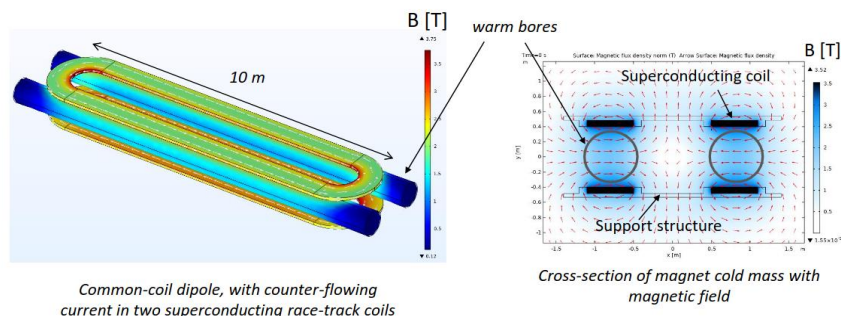
5

IRIS - L. Rossi - Cosmic Wispers

HIGH MAGNETIC FIELD - EXPERIMENTS



Winding layout



Proposed winding layout:

- Common-coil superconducting dipole with a coil length of 10 meters
- Two race-track coils with counter-flowing currents
- Produces a transverse magnetic field over the two 11-meter-long bores with a free diameter of 0.7 m
- To be operated at $T \leq 5$ K featuring Nb-Ti-based superconducting coils with about 2 T in the bore

CA21106 Kick-off meeting Feb 2023 4

General specification of the magnet

Design	Physics specification	Value
Magnetic design	"Ideal" FoM	$\sim 100 \text{ T}^2\text{m}^2$
Mechanical design	Booster length 'L _b '	1.3 m
Conductor design	Bore diameter	1.35 m
Manufacturing aspects	Dipole Field	$\sim 9 \text{ T}$
Winding issues	Homogeneity (z=0)	$\pm 10\%$
Casing issues	Maximum overall length	6.9 m
Conductor R&D	Maximum overall weight	200 t
Copper profile shape		
Cold work		
Quench propagation R&D		
Design		
Manufacturing & integration		
Test results		

Magnet specification	Value
Superconductor	NbTi
Operating temperature	1.8 K
Load Line margin	10 %
Temperature margin	1 K
Allowable VM stress	180 MPa
Hot spot	100 K
Discharge voltage	$\pm 1 \text{ kV}$
Maximum current	30 kA

~ 6 m

9T in 1.35 m

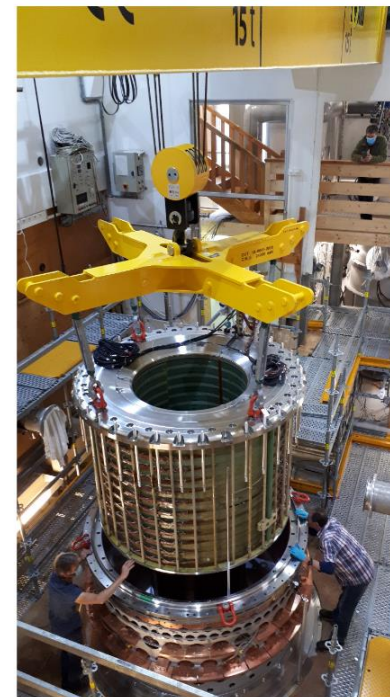
Outline

► Focus on high field facilities & magnets

- International context
- Highest DC field magnets worldwide
- 43-T Hybrid Magnet
- Presentation of LNCMI

► GrAHal-QUAX-CAPP

- Scientific Motivations
- Principle of the Experiment (Sikivie Haloscope)
- Current Status of GrAHal & First Results
 - Cu and dielectric RF cavity 4-7 GHz (ANR French funding agency, 2022-2026)
- Objectives and plans
 - GrAHal-QUAX for 7-10 GHz (dielectric RF cavity)
 - GrAHal-CAPP for 200-600 MHz (thin Cu RF cavity)
- Upgrade phases under study
 - Within a 40 T all superconducting magnet under construction
 - Toward a 60 T Hybrid magnet (?)



Pierre.Pugnat@lncmi.cnrs.fr

TECHNOLOGY FORUM ON QUANTUM TECHNOLOGIES FOR WISPS

FEB 1 – 2, 2024
DESY HAMBURG

SQUIDs and their applications	<i>Michael Mueck</i>	
<i>Main Auditorium, DESY Hamburg</i>	15:00 - 15:25	
Development of superconducting circuits in Italy	<i>Federica Mantegazzini</i>	
<i>Main Auditorium, DESY Hamburg</i>	15:25 - 15:50	
Quantum Sensing and the Dark Quantum project	<i>Takis Kontos</i>	
<i>Main Auditorium, DESY Hamburg</i>	15:50 - 16:15	

QUANTUM TECHNOLOGIES FOR WISPS

SQUIDs and Their Applications

Michael Mück, ez SQUID Mess- und Analysegeräte

1. Introduction
2. The dc SQUID
3. 'Traditional' Applications of SQUIDs (Medical Diagnostics, Nondestructive Testing of Material)
4. Using a dc SQUID as a high-frequency amplifier
5. A SQUID amplifier with microstrip input coupling (MSA)
6. Gain and Noise Temperature of Microstrip SQUID Amplifiers
7. Some Recent Applications of MSAs

Development of superconducting circuits in Italy

2 February 2024

Federica Mantegazzini

FONDAZIONE BRUNO KESSLER

SENSORS AND DEVICES

fmantegazzini@fbk.eu
www.fbk.eu

Networks

1. Help extending existing networks
2. Create links for new networks
3. Create opportunities for joint proposal for axion experiments with multi detectors

What is a GNOME?¹

- Global Network of Optical Magnetometers for Exotic physics searches
- Looking for transient dark matter signals
- Sensitive to Axion-fermion coupling:

$$H_{\text{int}} = -\frac{\hbar c^{3/2}}{f_{\text{int}}} \frac{S_i}{|S_i|} \cdot \nabla a$$



¹Phys. Dark Univ. 22 (2018), 162-180
Daniel Gavrila, Martin

See talk by Daniel Gavilán



<https://www.pi.uni-bonn.de/gravnet/en>

TECHNOLOGY FORUM: ITC CONTRIBUTIONS

SOFIA 8 – 12 SEP 2025

Advances in optomechanics

Nenad Kralj

Aula, Sofia University

15:30 - 16:00

Satellites & Space Missions

Lubomir Toshev et al.

Aula, Sofia University

16:00 - 16:30

Magnetometry: from single measurements to network of magnetometers

Boyan Benev

Aula, Sofia University

16:30 - 17:00

ITC CONTRIBUTION

Advances in Optomechanics

Nenad Kralj, Faculty of Physics, University of Rijeka

3rd General Meeting, COST Action "Cosmic WISPer" (CA21106)
11/09/2025



SATELLITES AND SPACE MISSIONS

THE ENDUROSAT STORY



Lubomir Toshev,
Principal Software Architect

Privileged and confidential

3rd General Meeting of COST Action COSMIC WISPer (CA21106)
9-12 September, 2025, Sofia, Bulgaria

Magnetometry: from single measurements to network of magnetometers



Boyan Benev
Benev Science & Technology Ltd.
Institute of Astronomy and NAO,
Bulgarian Academy of Sciences