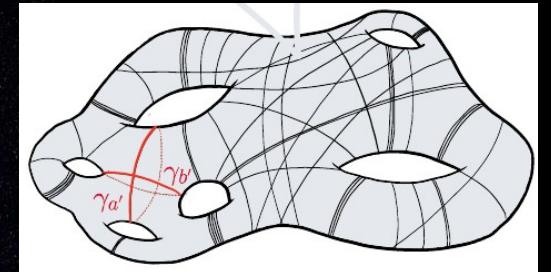
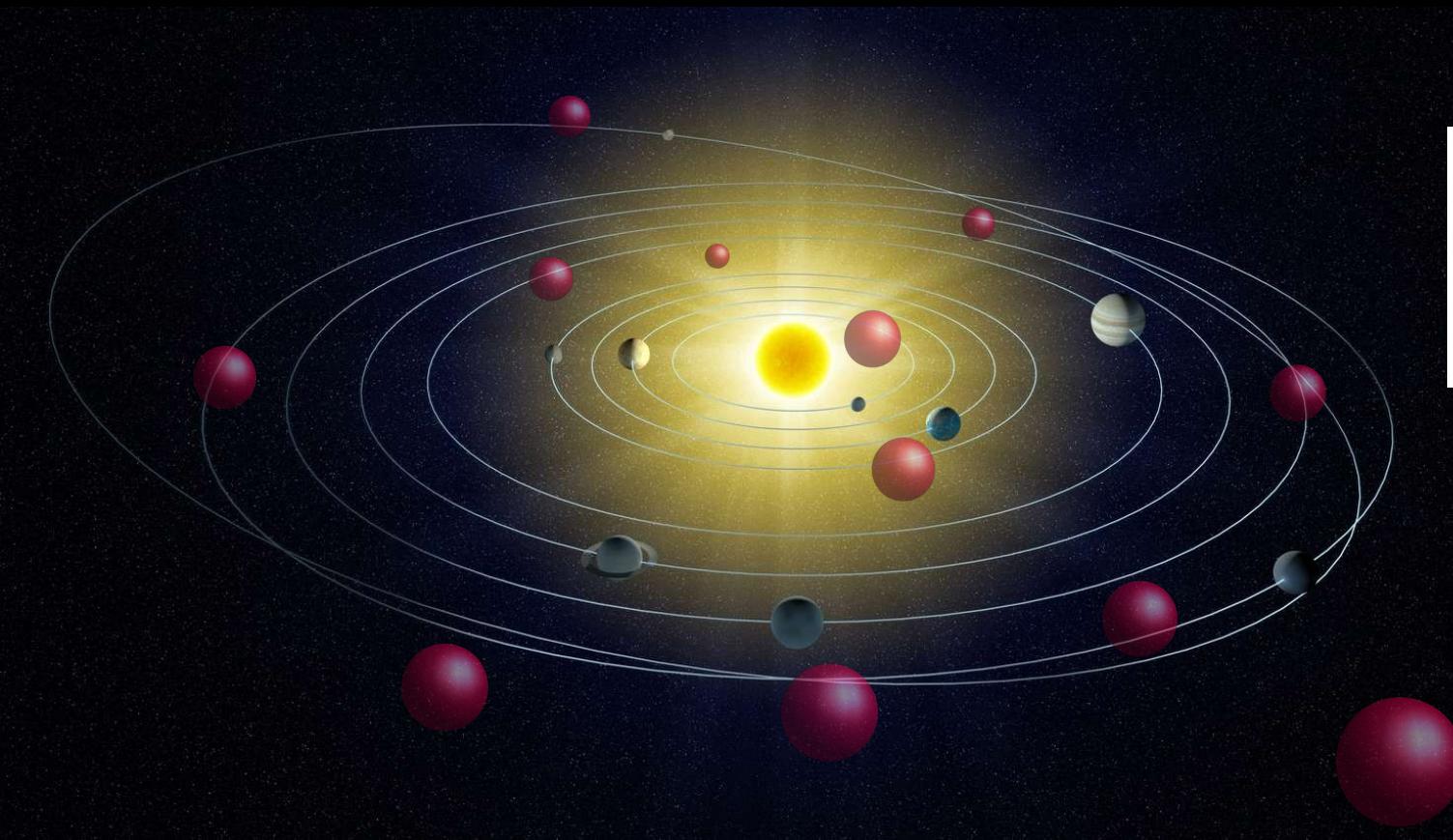


# Open String Axiverse

Giovanni Villadoro  
ICTP

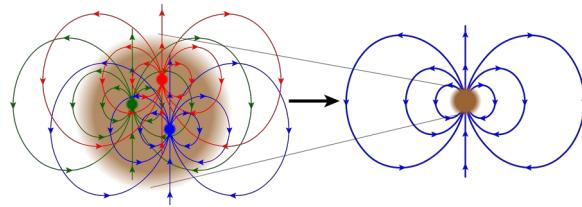


with  
Rudin Petrossian-Byrne  
2503.16387

# The Strong CP problem

# The Strong CP ~~problem~~ puzzle

No nEDM

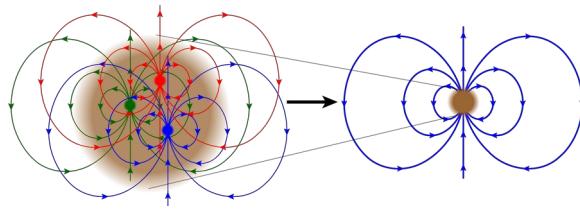


$$\theta \equiv \arg \left\{ \det (Y_u Y_d) e^{i\theta_0} \right\} \lesssim 10^{-10}$$

$$\xrightarrow{\quad} \frac{\theta_0}{32\pi^2} G\tilde{G} \in \mathcal{L}_{\text{SM}}$$

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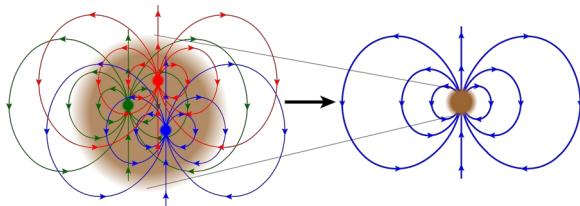
$\downarrow$

$$\frac{\theta_0}{32\pi^2} G \tilde{G} \in \mathcal{L}_{\text{SM}}$$

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \frac{1}{2} (\partial a)^2 + \frac{a}{f_a} \frac{\alpha_s}{8\pi} G \tilde{G} + \dots \quad \Rightarrow \quad m = \frac{\chi_{\text{top}}^{1/2}}{f_a} \simeq 0.57 \text{ meV} \left( \frac{10^{10} \text{ GeV}}{f_a} \right)$$

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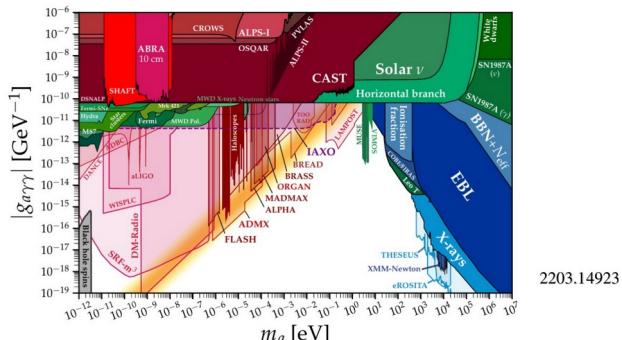


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Testable:



Dark Matter:



# Cosmological Scenarios:

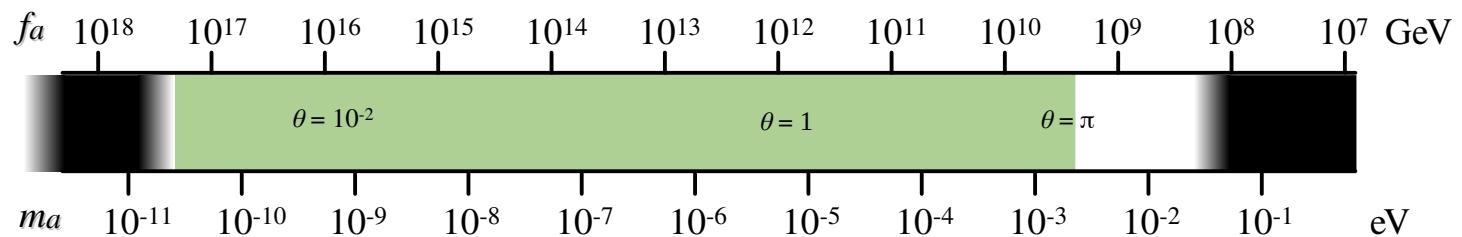
1. “pre-inflation”:  
(standard misalignment)

2. “post-inflation”:  
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$$\Omega_a \simeq \theta_0^2 \left( \frac{f_a}{10^{12} \text{ GeV}} \right)^{1.2} \Omega_{\text{DM}}$$

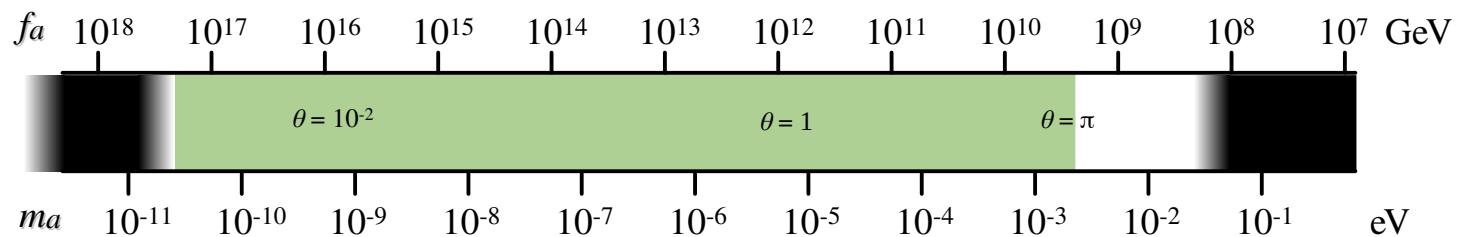


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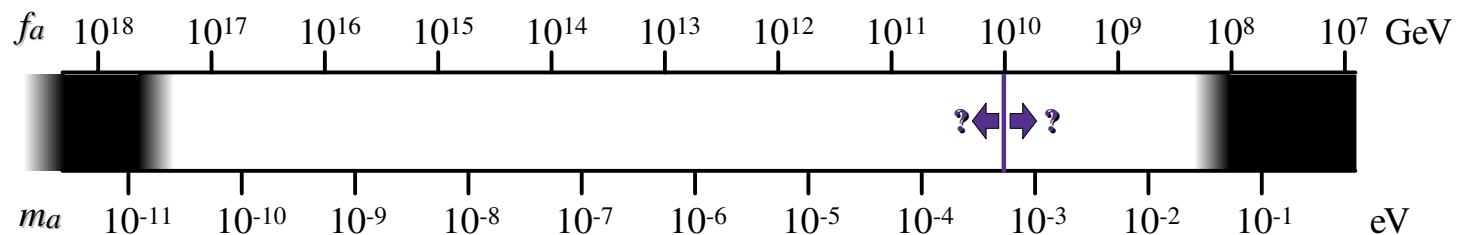
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$$\Omega_a \simeq \left( \frac{f_a}{? \cdot 10^{10} \text{ GeV}} \right)^{1.2} \Omega_{\text{DM}}$$



## PQ Quality:

$$V(a) = V_{\text{QCD}}(a) + V_{\text{PQ}}$$

$$\begin{aligned}\theta &\lesssim 10^{-10} \\ \Rightarrow\end{aligned}$$

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Quantum Gravity → No Global Symmetries...

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## Accidental PQ

→ forbids PQ operators up to dim.  $\sim 10 \div 40$  (for  $f_a \sim 10^{10} \text{ GeV} \div \text{GUT}$ )

Irreducible QG symm. breaking is non-perturbative (BH, wormholes, ...)

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NOTE: QCD-breaking  $\rightarrow \chi_{\text{top}} \propto e^{-\frac{\#}{g_s^2}}$

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High Quality PQ  $\rightarrow$  QG is weakly coupled

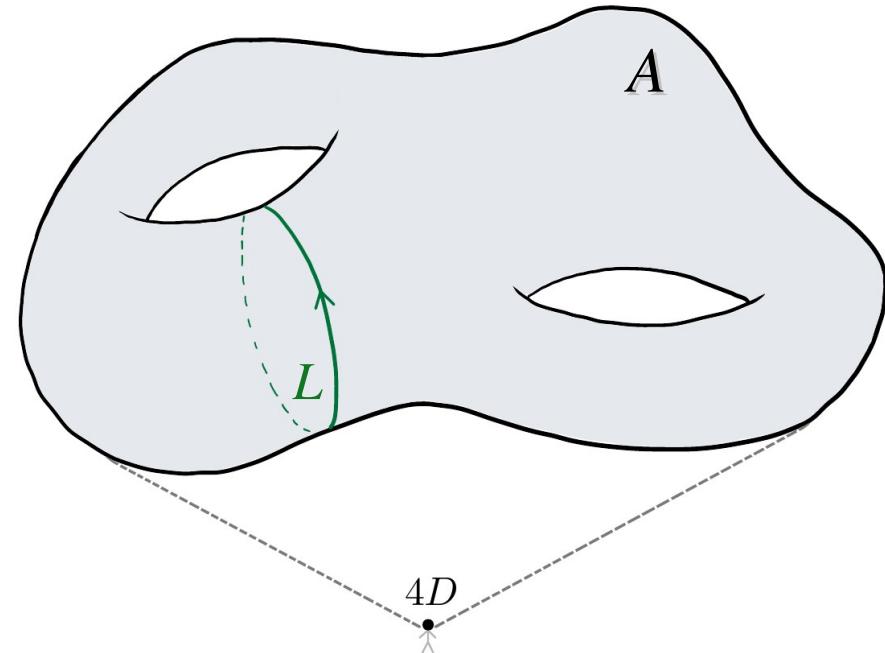
i.e. worry about PQ quality only in explicit UV completion of QG:  
e.g. string theory

# String Axiverse

Gauge Fields in Extra Dimensions

Choi '03

$$\frac{a}{v} = \arg \left( e^{i \oint A} \right)$$



# String Axiverse

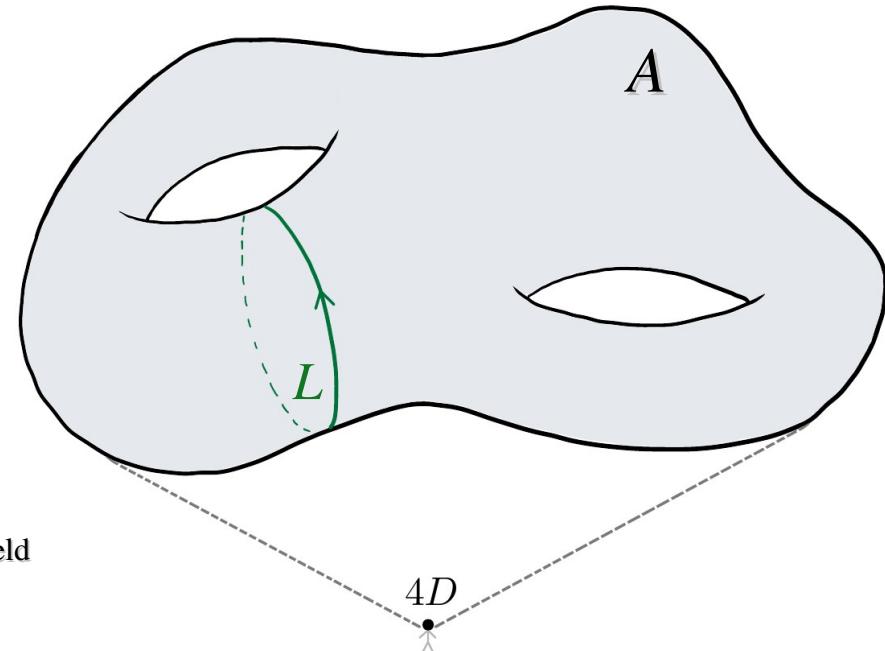
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$$V_{\text{PQ}} \propto e^{-ML}$$

lightest charged bulk field



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$$V_{\text{PQ}} \propto e^{-\text{Vol}/g_s} \sim e^{-8\pi^2/g_4^2}$$

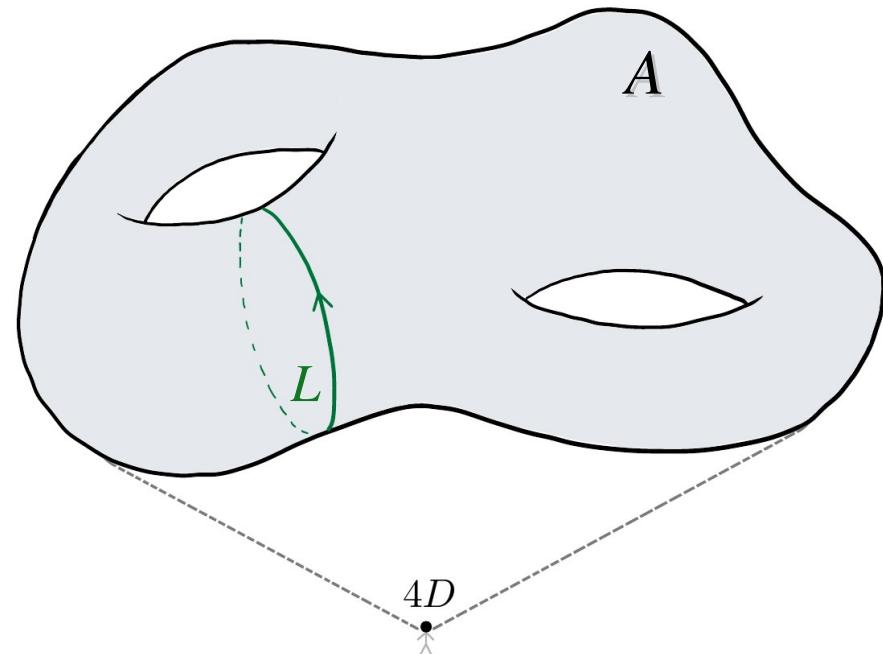
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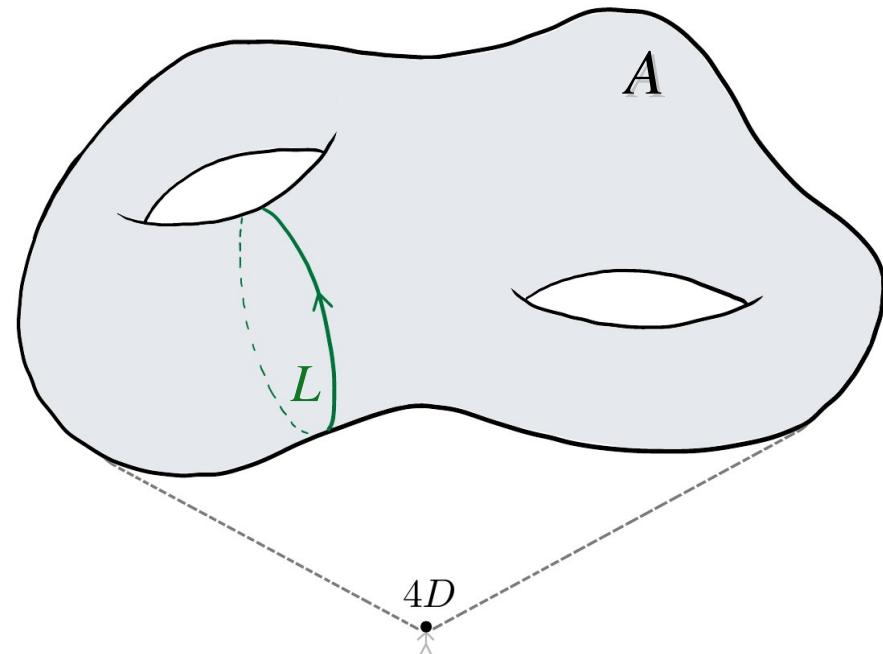
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- ✓ pre-inflationary and large  $f_a$  axions



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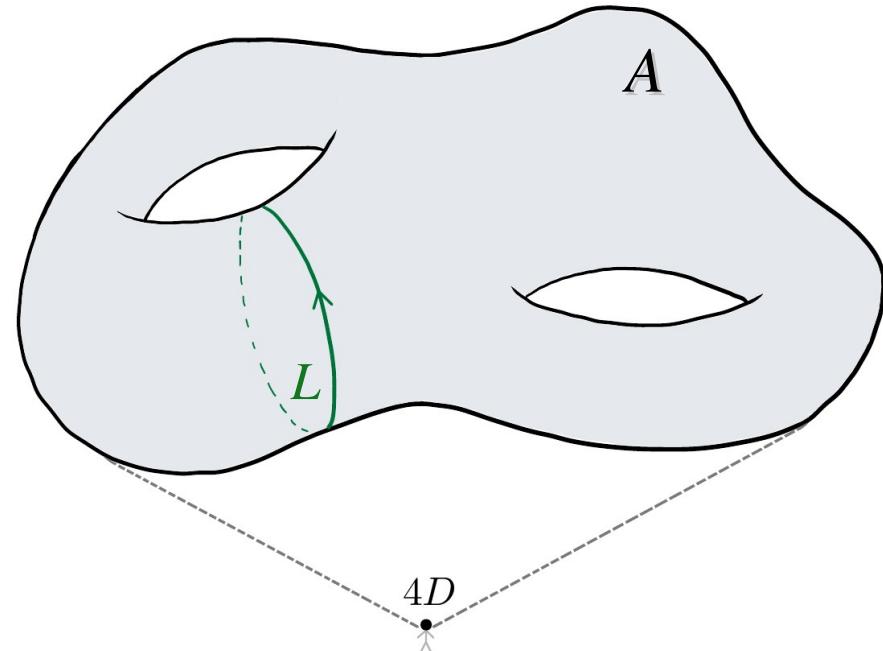
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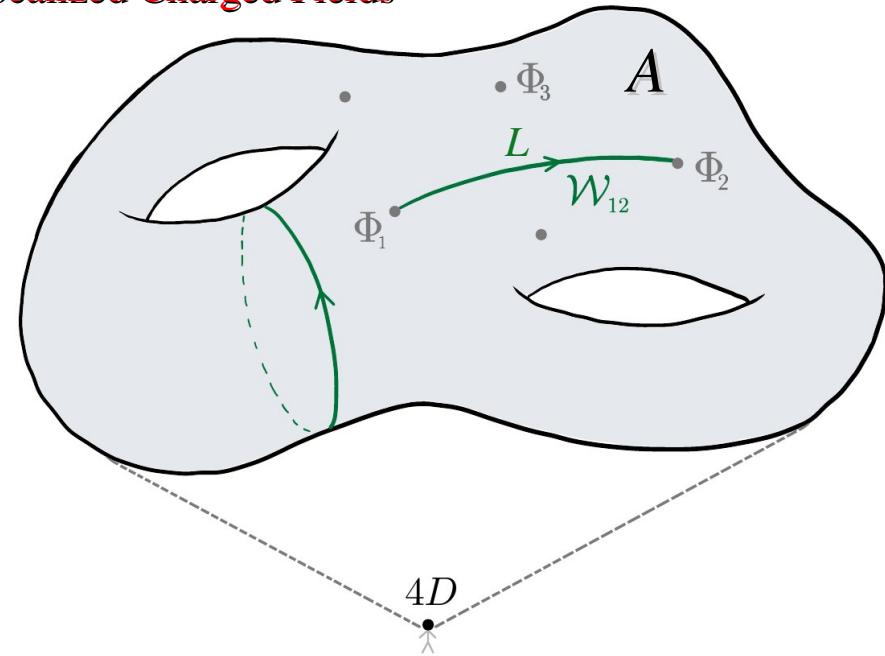
✓ pre-inflationary and large  $f_a$  axions

✗ post-inflationary axions  
 (PQ-phase restoration  $\rightarrow$  D-branes after inflation...)



# Open String Axiverse:

Gauge Fields in Extra Dimensions + **Localized Charged Fields**

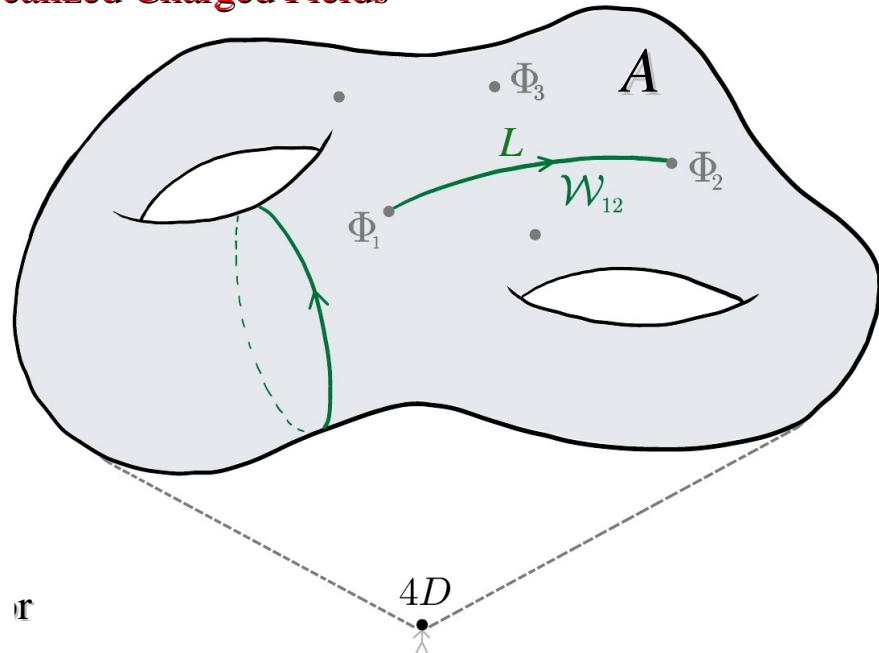


# Open String Axiverse:

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$$\mathcal{W} = \Phi_+ e^{i \int_{y_+}^{y_-} A} \Phi_-$$

gauge invariant:  $\begin{cases} A \rightarrow A + \partial \Lambda \\ \Phi_\pm \rightarrow e^{\pm i \Lambda} \Phi_\pm \end{cases}$



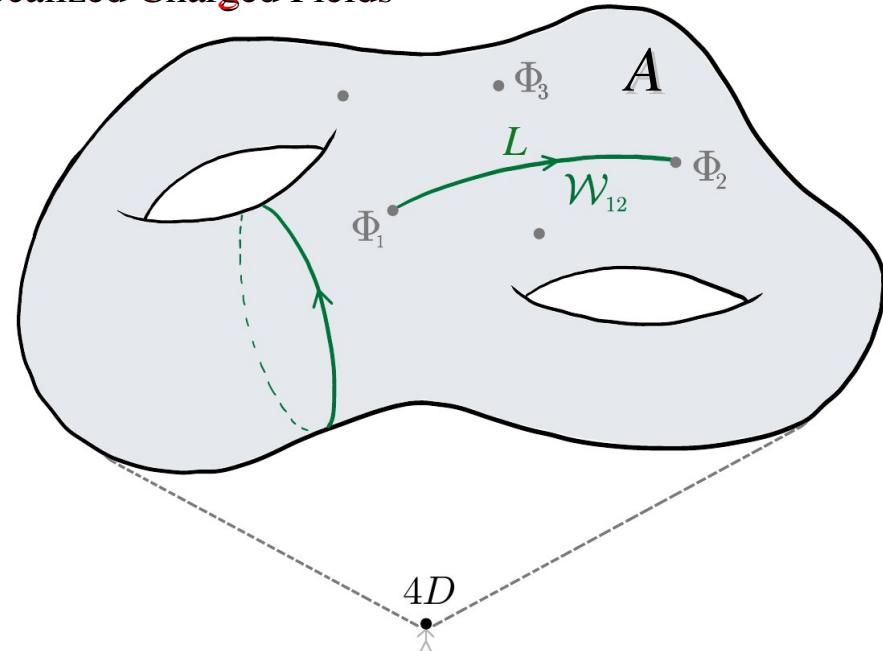
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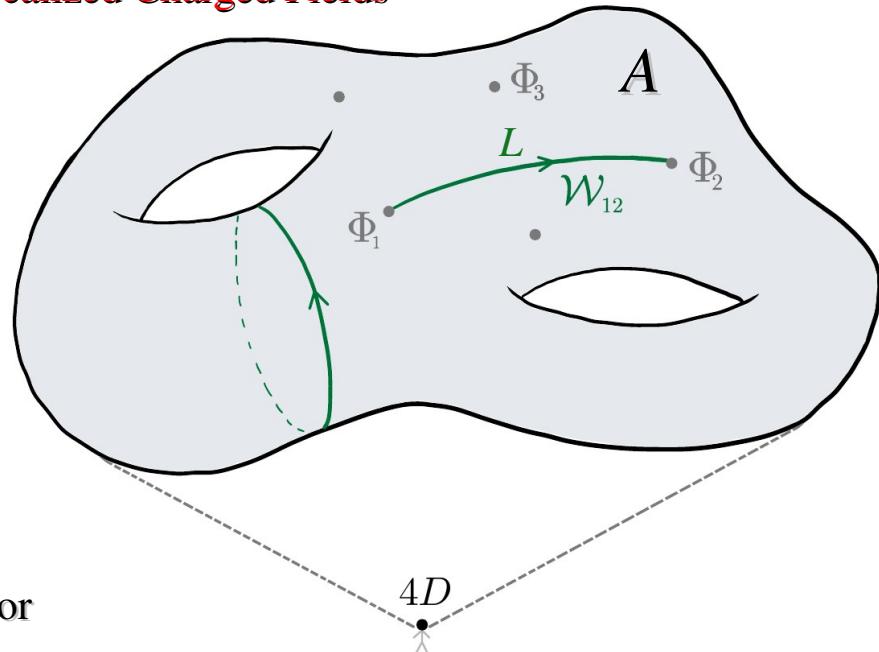
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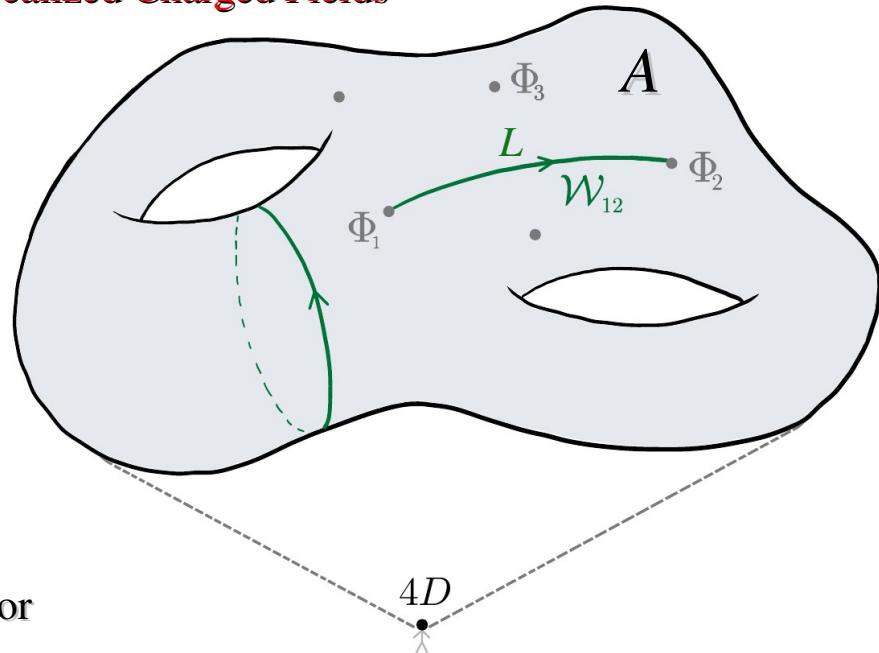
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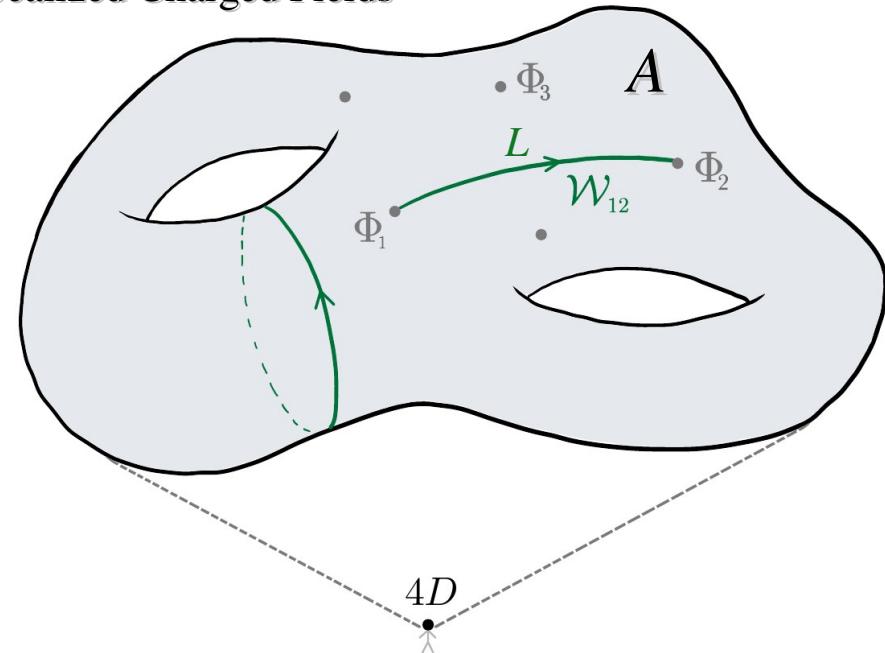
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# Open String Axiverse:

Gauge Fields in Extra Dimensions + Localized Charged Fields

- With  $N$  localized fields  $\rightarrow N - 1$  global U(1)s



# Open String Axiverse:

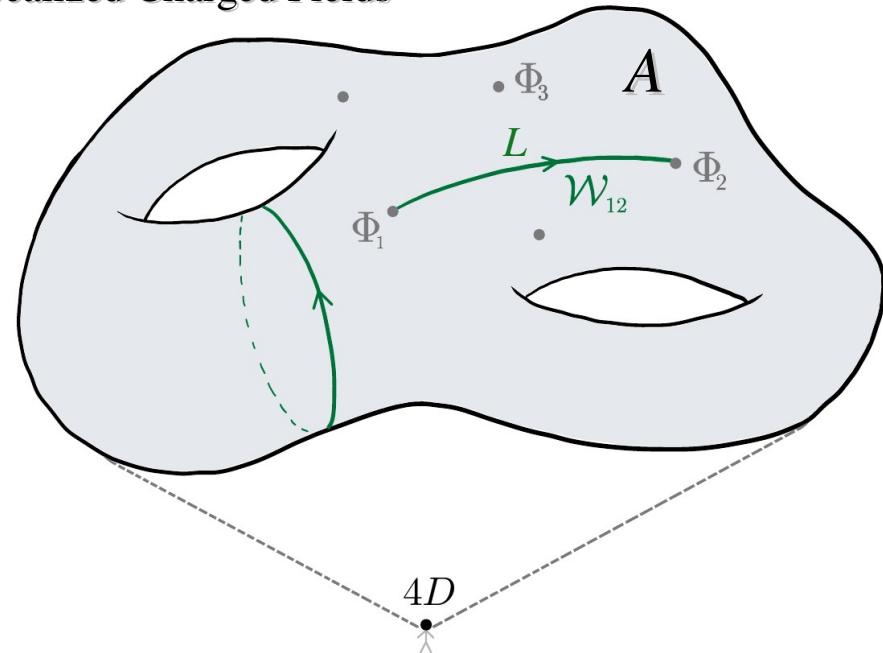
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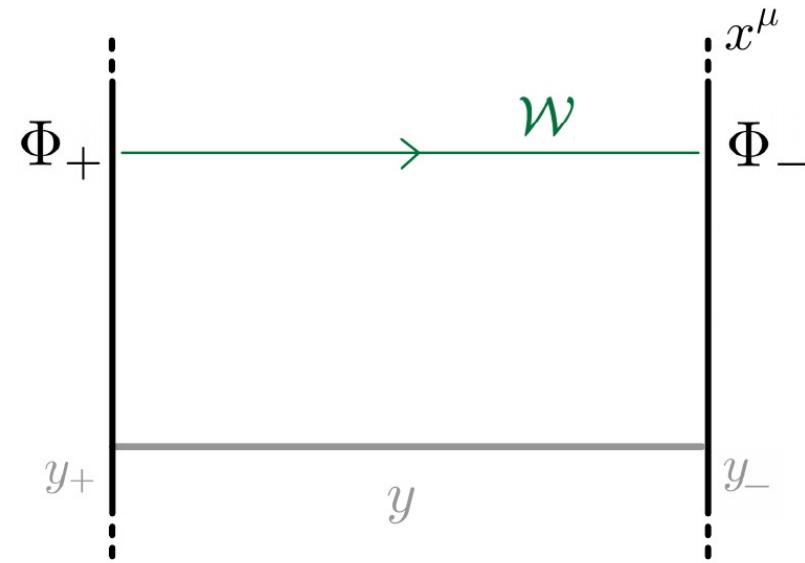
– PQ is linearly realized in 4D:

$$\frac{a}{v} = \arg \left( \Phi_+ e^{i \int_{y_+}^{y_-} A} \Phi_- \right)$$

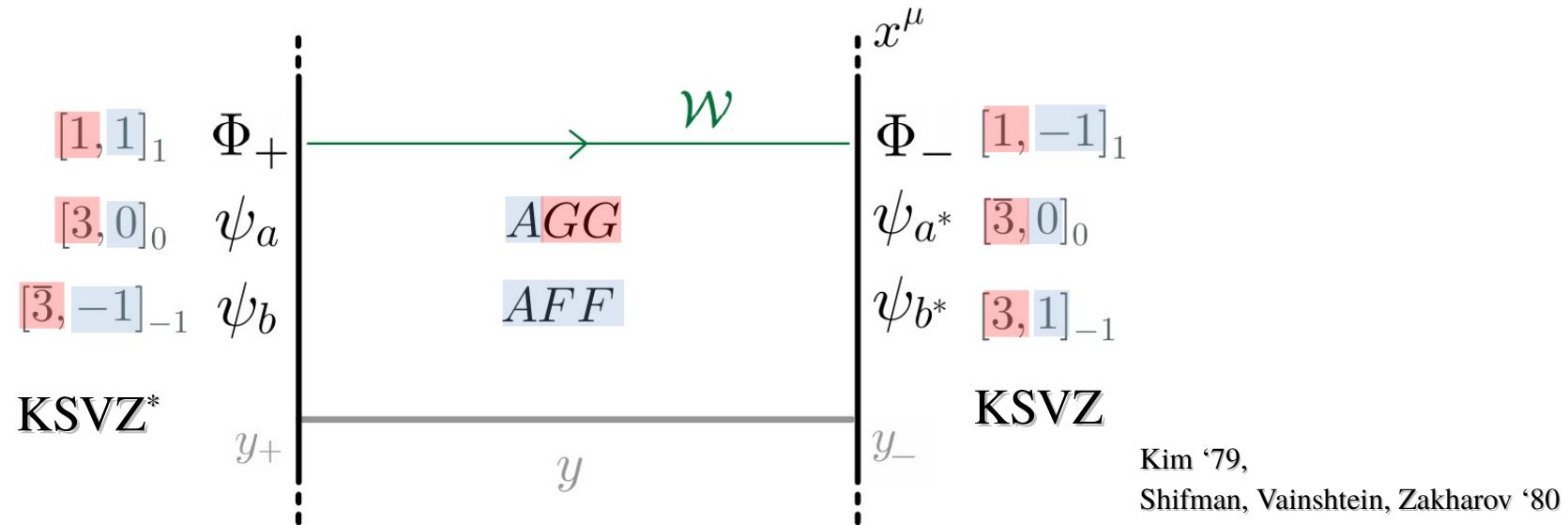
$$v \ll \frac{1}{L} \quad \text{if} \quad \min \langle \Phi_\pm \rangle \ll \frac{1}{L}$$



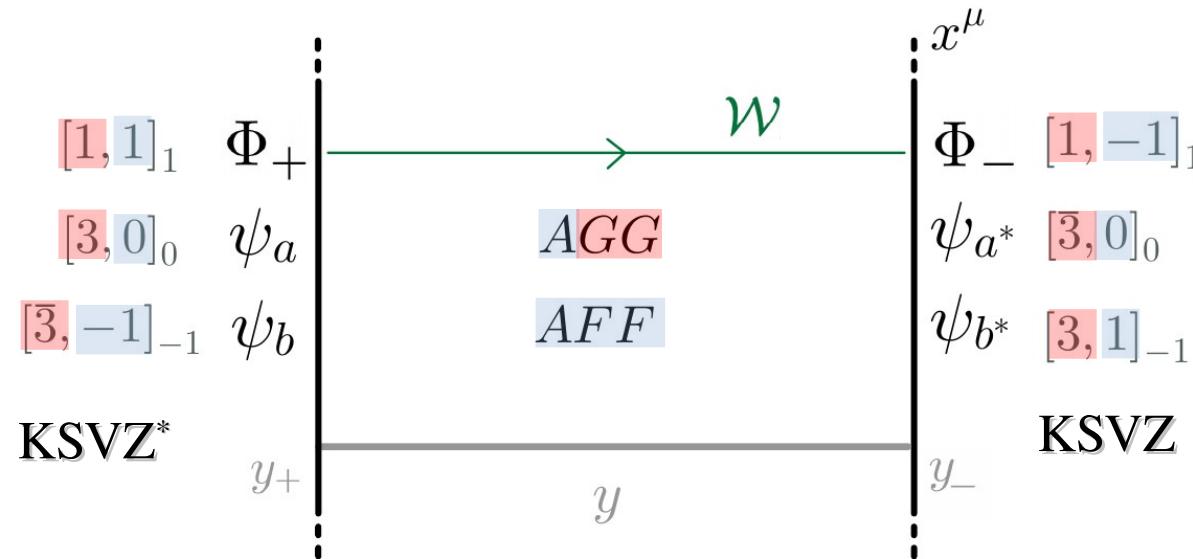
## A 5D toy model:



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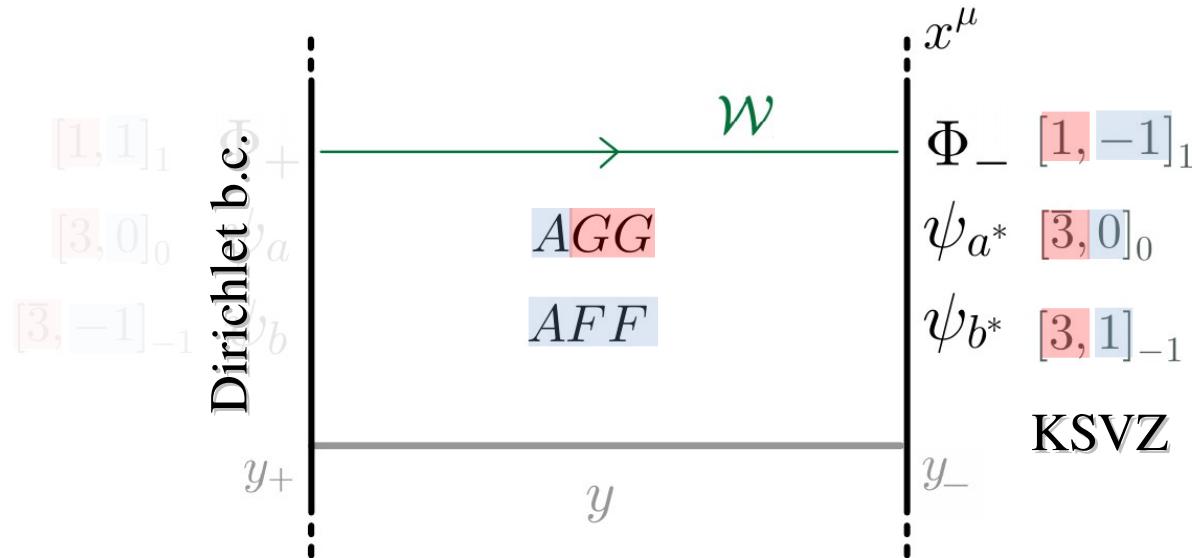


Kim '79,  
Shifman, Vainshtein, Zakharov '80

- In the limit  $v_- \ll (v_+, 1/L)$  minimal ( $N_{\text{DW}} = 1$ ) KSVZ model recovered:
- $L$  can be small so that all 4D KSVZ extensions work as usual: SUSY, GUT, exotics, ...

# A 5D toy model:

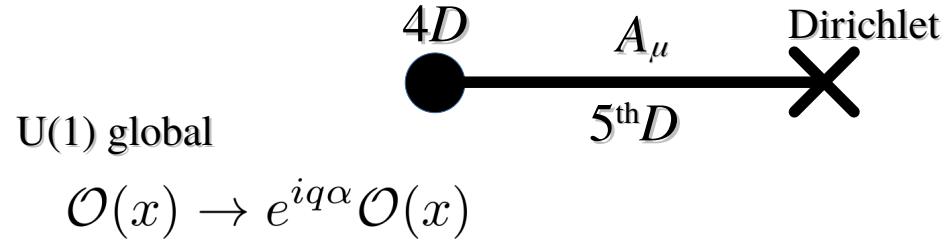
NOTE: in the limit  $v_+ \rightarrow \infty$



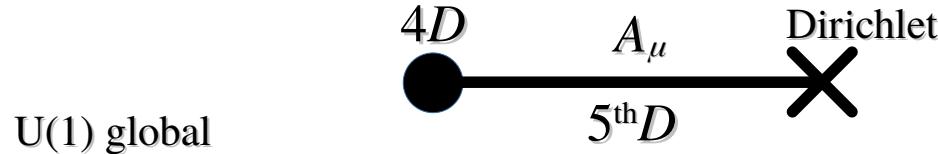
# KK lifting of symmetries

$$\begin{array}{c} 4D \\ \bullet \\ \text{U(1) global} \\ \mathcal{O}(x) \rightarrow e^{iq\alpha} \mathcal{O}(x) \end{array}$$

# KK lifting of symmetries



## KK lifting of symmetries

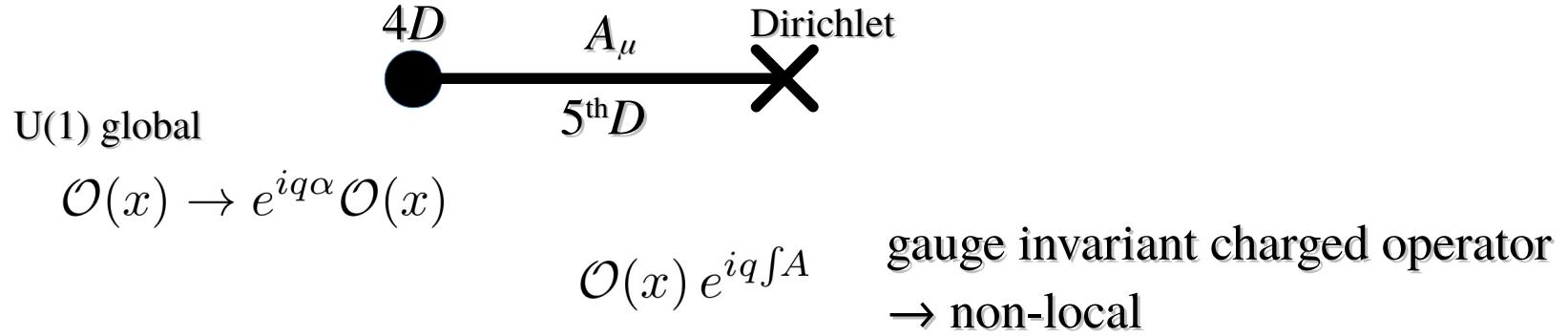


$$\mathcal{O}(x) \rightarrow e^{iq\alpha} \mathcal{O}(x)$$

$$\mathcal{O}(x) e^{iq\int A}$$

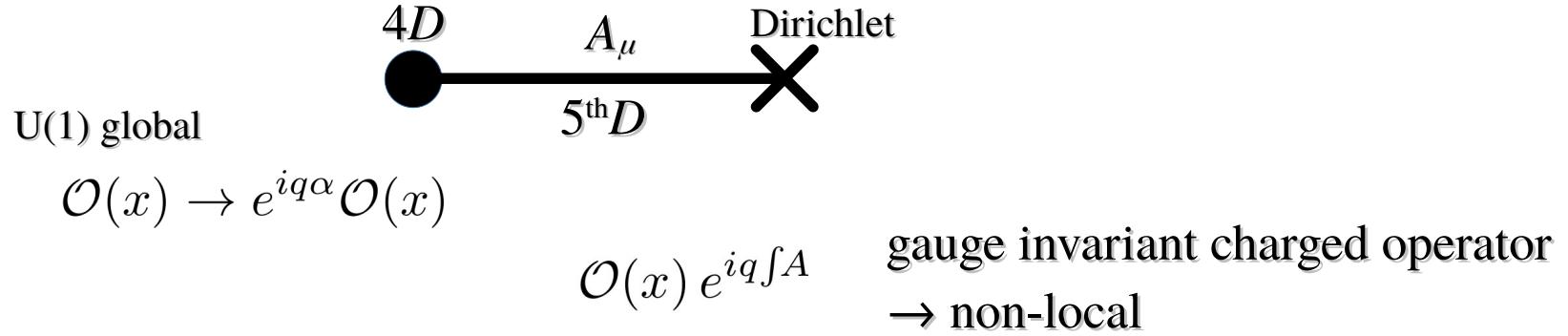
gauge invariant charged operator  
→ non-local

# KK lifting of symmetries



Breaking  $\rightarrow e^{-ML}$

# KK lifting of symmetries

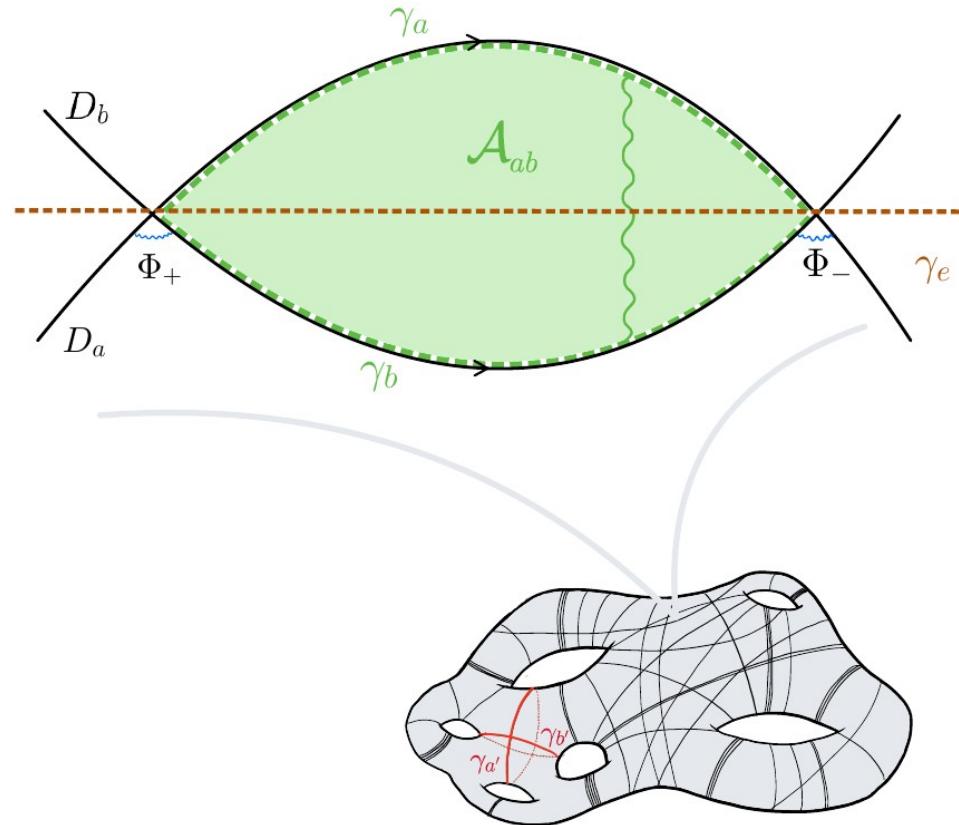


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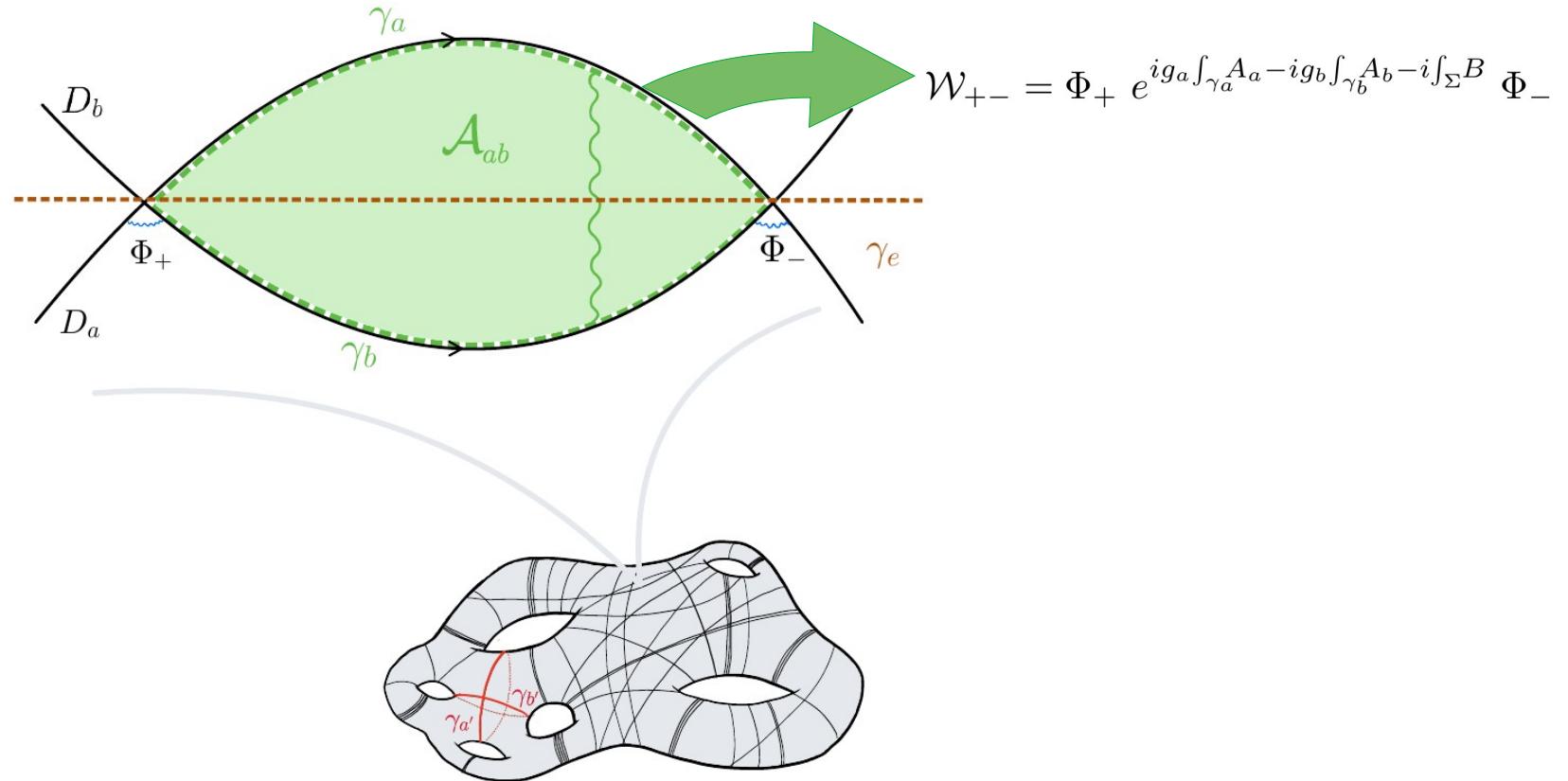
NOTE:

in 4D the global symmetry is not *accidental* but exponentially good

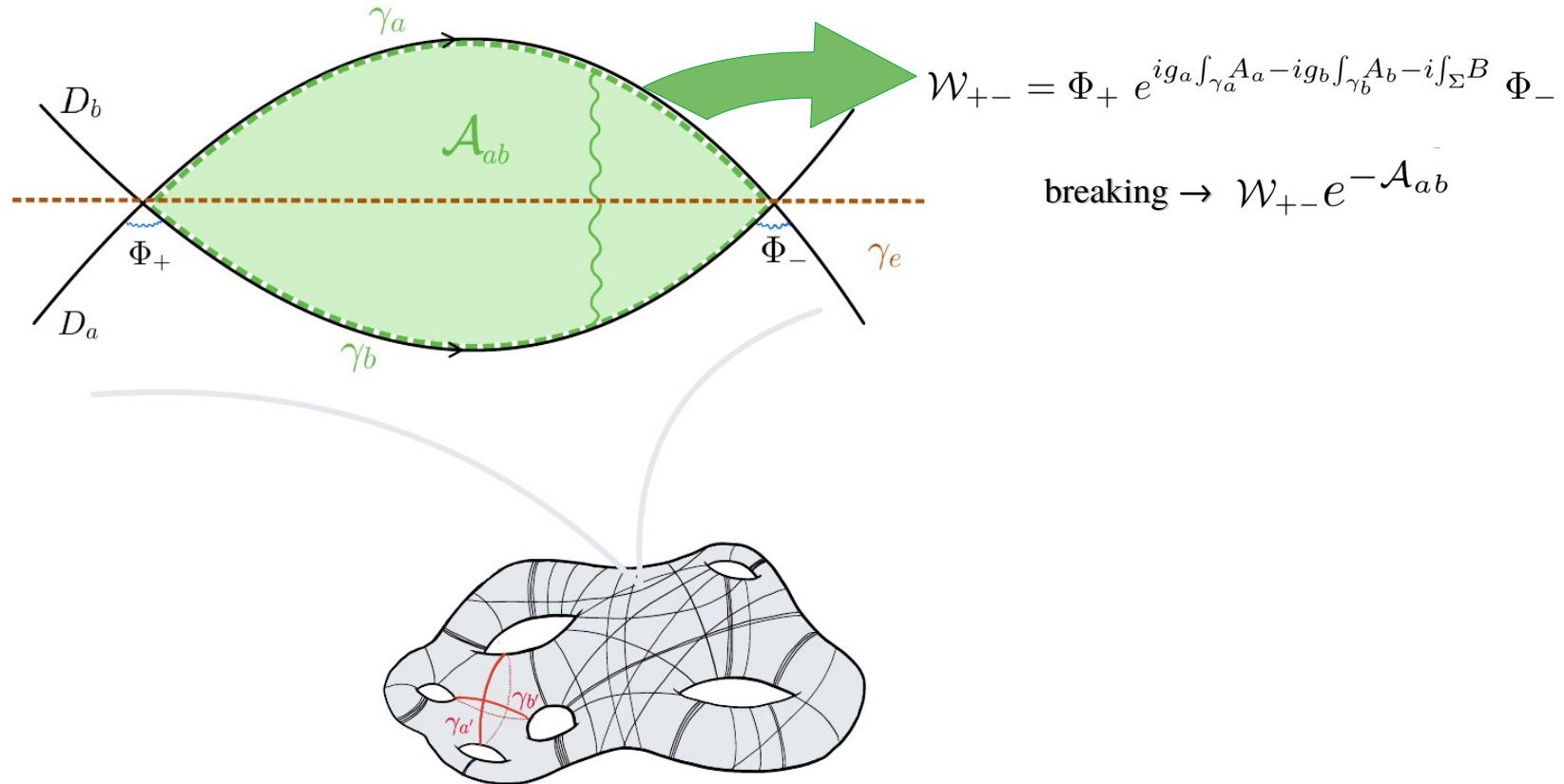
# A simple string theory embedding



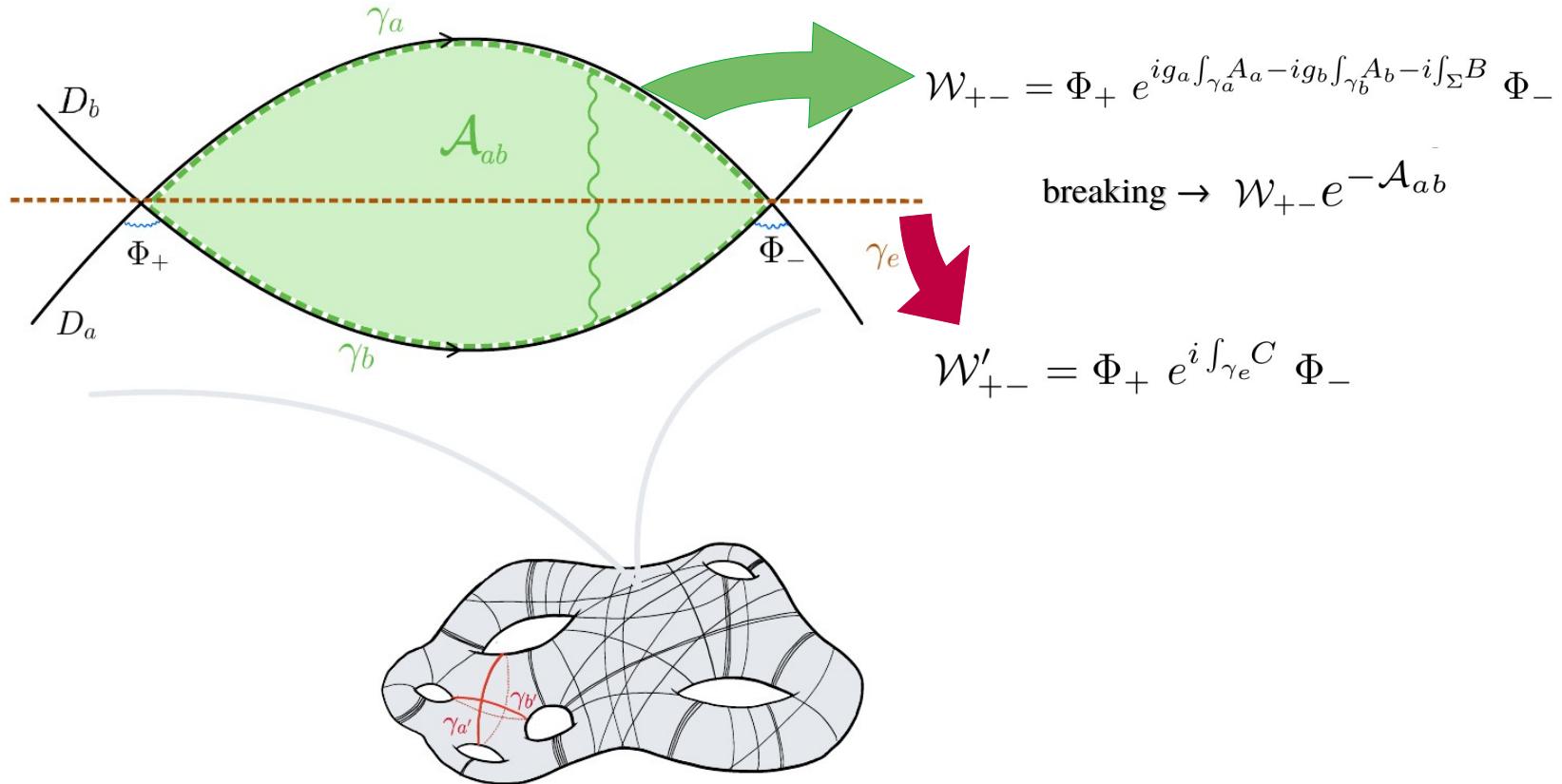
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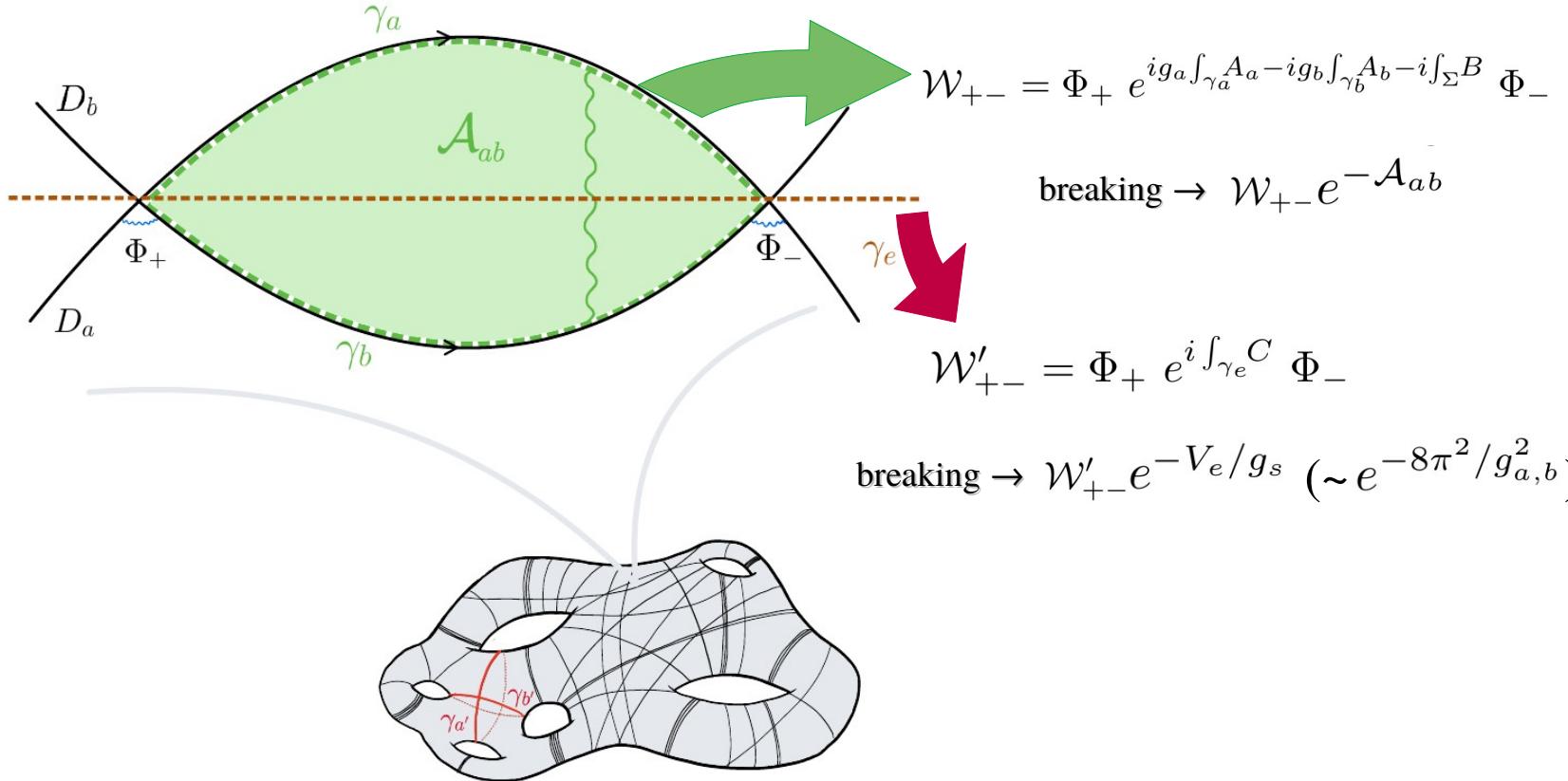
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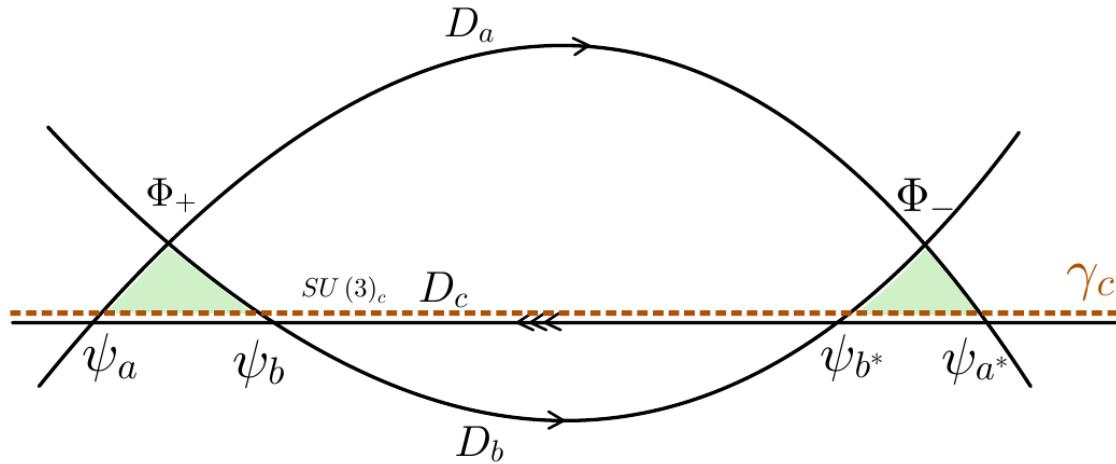
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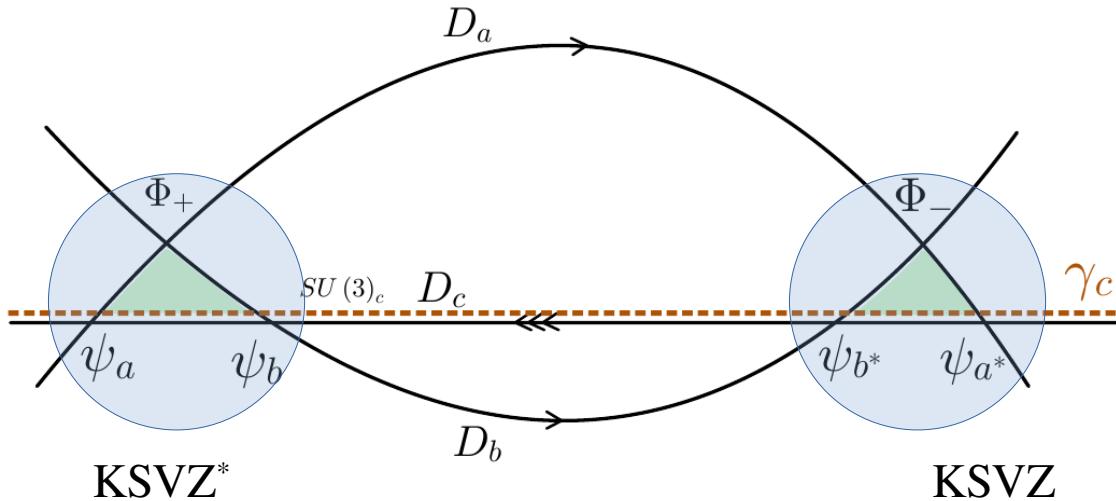
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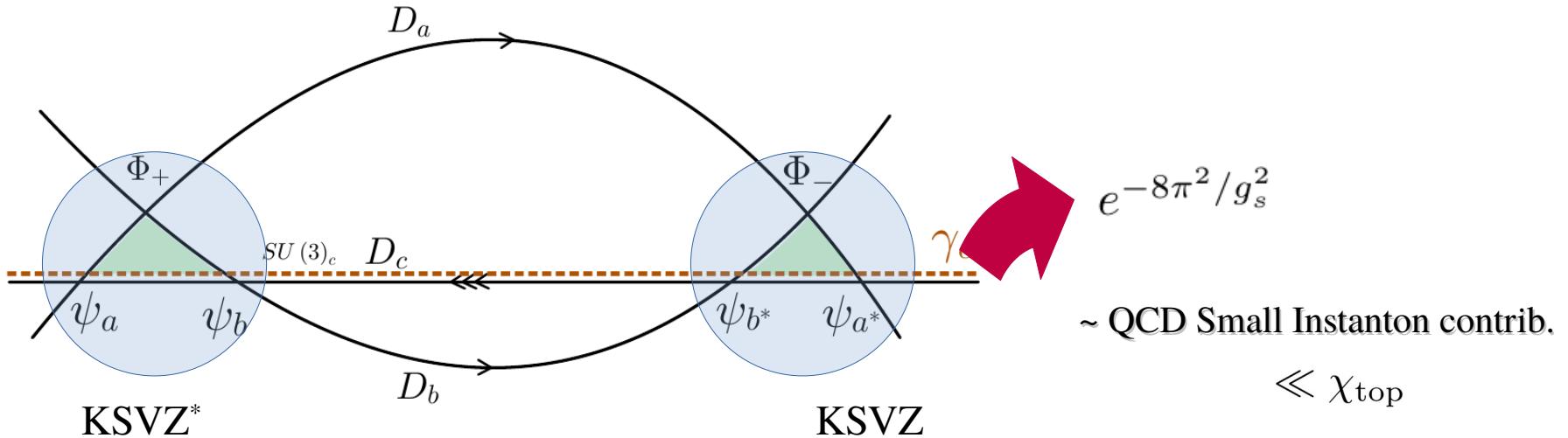
# A simple string theory embedding of minimal KSVZ



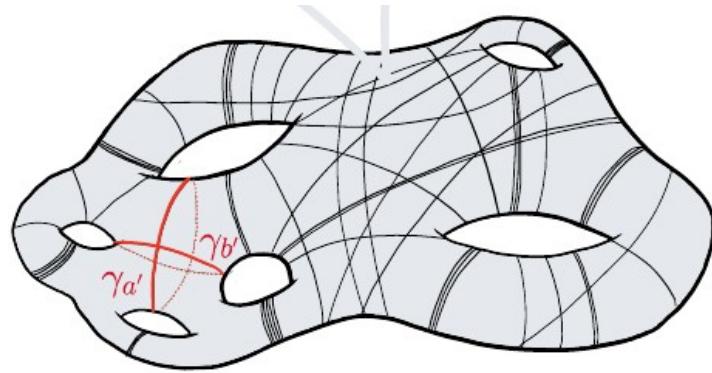
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# Open String Axiverse



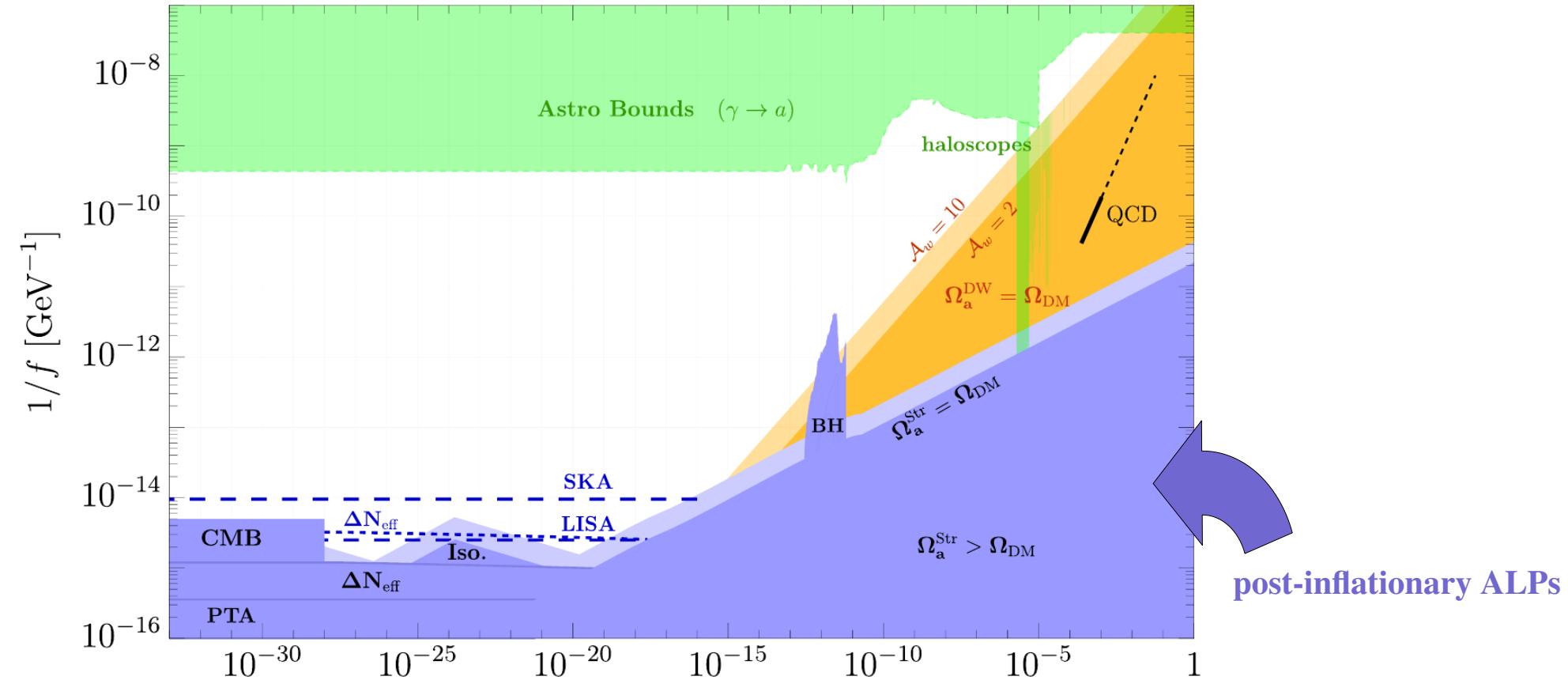
In realistic compactifications branes have multiple intersections

→ potentially multiple post-inflationary ALPs

see e.g. works of:

Allahverdi, Berenstein, Cicoli, Diaz, Dutta, Guidetti, Honecker, Klevers, Krippendorf, Mayrhofer, Perkin, Quevedo, Rummel, Sinha, Staessens, Valandro ...

# Open String Axiverse Phenomenology

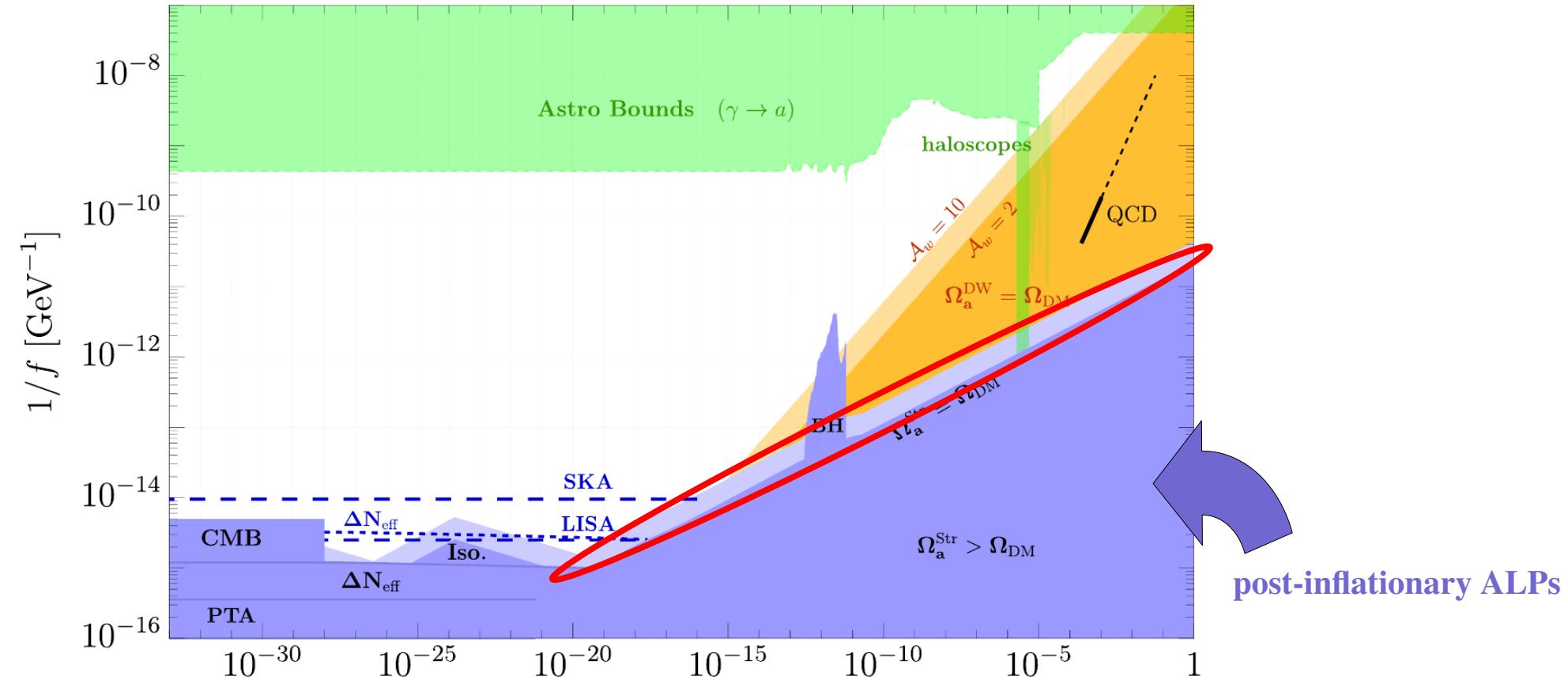


see also

Gorghetto, Hardy, Nicolaescu '21

Gorghetto, Hardy '22

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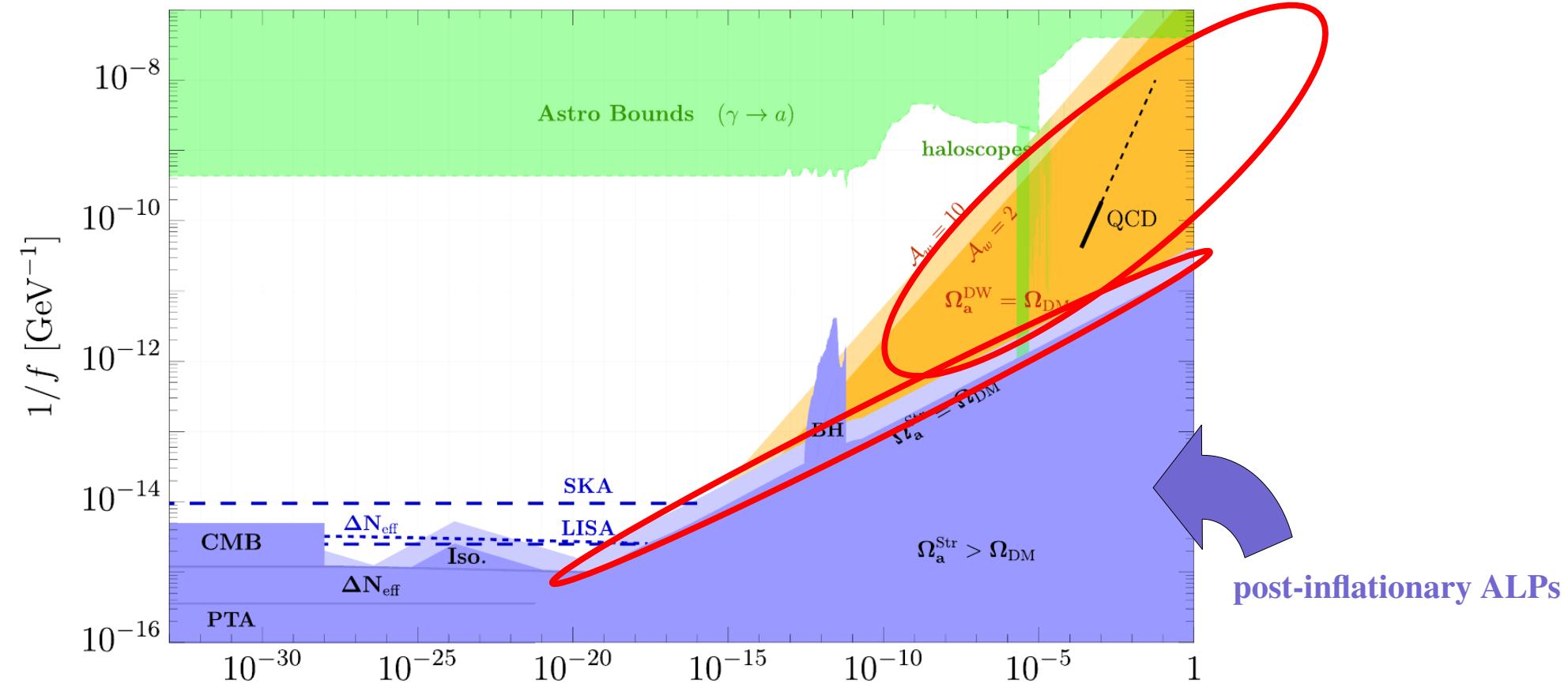


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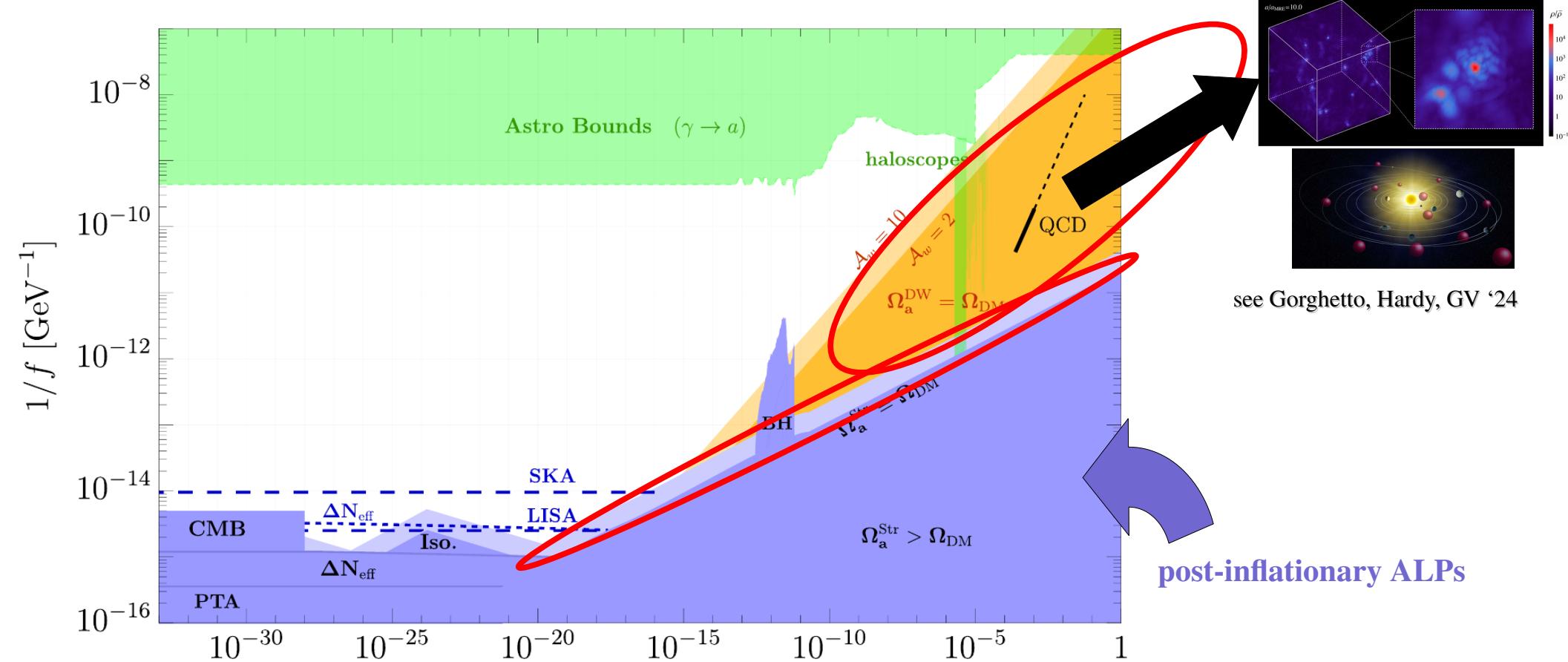


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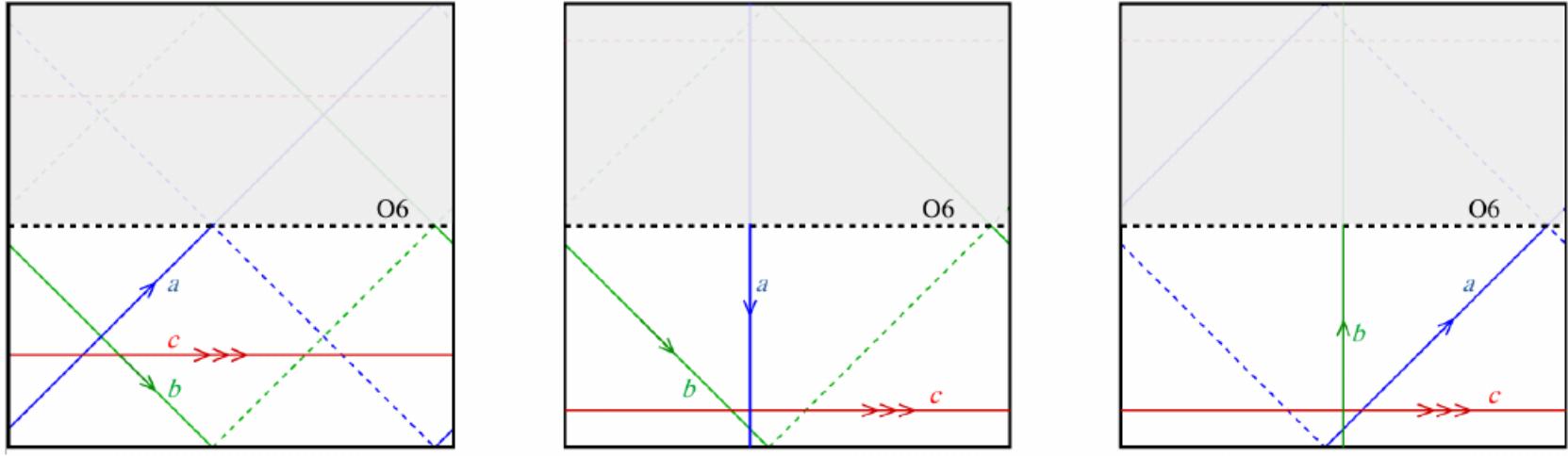
Gorgetto, Hardy '22

## *Executive Conclusions*

1. No problem between Quantum Gravity and HQ Symmetries
2. No fundamental obstruction to post-inflation QCD axions/ALPs
3. In fact it's generic → “open string axiverse”
4. New phenomenological opportunities

*Thank You!*

*Back up*



Example: SUSY cycles on  $T^6$  compatible with (approx.) SUSY moduli stabilization