

Heavy Neutral Leptons without Prejudice

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Neutrinos and Flavor: a stairway to new physics, Pisa, April 2025

Work done in collaboration with Nicolas Bernal and Kuldeep Deka

Outline

1. Motivation
2. Neutrino mass models
3. Long Lived Particles
4. Current experimental results
5. Projections for future collider experiments

Motivation

LHC results of Higgs discovery and measurement of physical properties.

First evidence of existence of a scalar particle with SM Higgs boson properties

Into the era of precision Higgs physics

No new elementary particles discovered with $m \sim 1\text{TeV}$

New physics via higher dimension operators or very weak interactions.

What is the NP scale?

eV

keV

MeV

GeV

TeV

Energy ranges that have been covered by accelerator experiments seem to imply the NP is weakly coupled (or phase space suppressed).

Accelerator Produced Detection

Search for very weakly coupled NP, and including for new long lived particles.

1

Colliders

2

Fixed target/Beam
dump

3

High intensity Beams

The basics and some history.....

Studying weak interaction with neutrinos..... Two neutrino types...

$$I = 5 \times 10^{12} \text{ protons/sec}$$

$I/10$ pions produced at target, $E > 2 \text{ GeV}$, 2 steradians

$$c\tau = 7.8\text{m} \Rightarrow d = 111\text{m}$$

10% of pions decay

$$N_{\text{flux}} = 10^{-9} I$$

events ~ 1 per hour

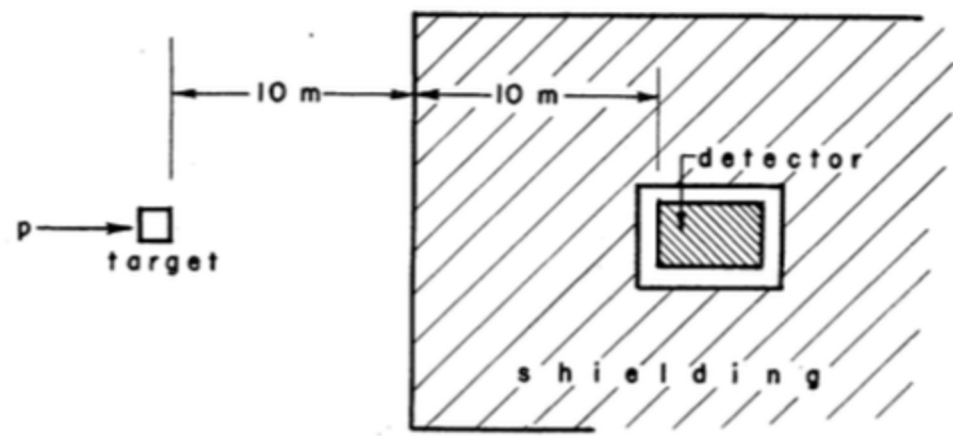


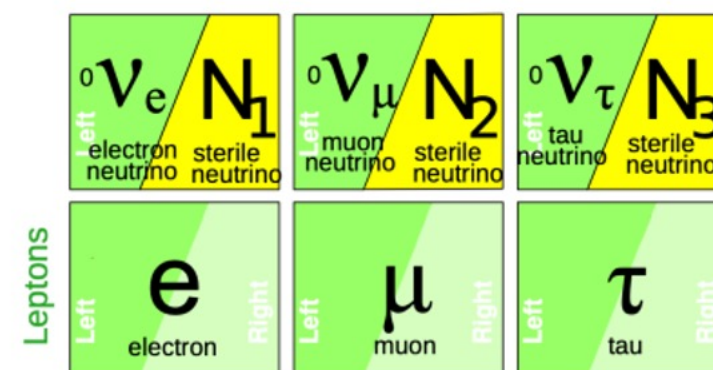
FIG. 1. Proposed experimental arrangement.

Pontecorvo 1959
Schwartz 1960
Danby et al 1962

Led to first neutrino beam and discovery of $\nu_e \neq \nu_\mu$

Why Heavy Neutral Leptons ?

- Neutrino masses
- Leptogenesis
- Dark matter candidate
- Neutrino anomalies



arXiv:1301.5516

Naturally introduces Yukawa-type couplings,
Connect HNLs with Higgs physics

However, it is completely unknown:

If the neutrino mass mechanism involves HNL what is the correct model?

Are there additional interactions of the HNL?

What is the nature of the HNL: Dirac or Majorana?

What is the HNL mass scale?

$m_N \sim \text{eV}$. for oscillation anomalies

$m_N \sim \text{keV}$ for warm dark matter

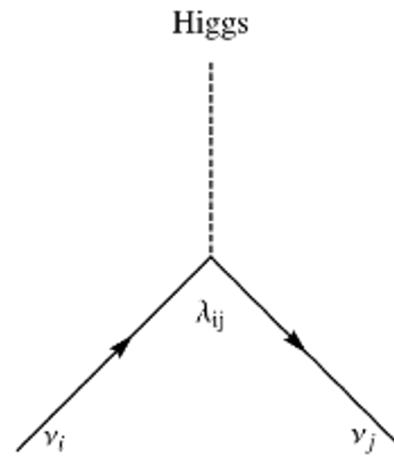
$m_N \sim \text{MeV- GeV}$ in deviations of SM

$m_N \sim \text{GeV-TeV}$ for BAU

Minimal neutrino mass models

$$-\mathcal{L}_{\text{Dirac}} = \bar{\nu}_L m_\nu \nu_R + h.c. \leftrightarrow \bar{L} \tilde{\Phi} \lambda \nu_R + h.c.$$

Dirac particles

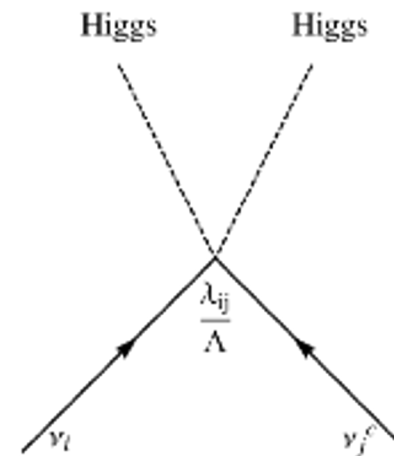


$$m_\nu \sim \lambda v$$

$$-\mathcal{L}_{\text{Majorana}} = \bar{\nu}_L m_\nu \nu_L^c + h.c. \leftrightarrow \bar{L} \tilde{\Phi} \alpha \tilde{\Phi} L^c + h.c.$$

Majorana particles

$$m_\nu \sim \lambda \frac{v^2}{\Lambda}$$

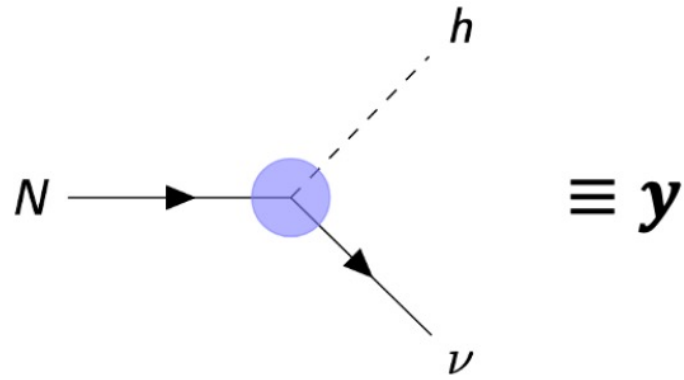


Minimal neutrino mass models

$$\mathcal{L} \supset y_{N\alpha} \bar{N} \tilde{H}^\dagger L_\alpha + h.c. + \dots$$

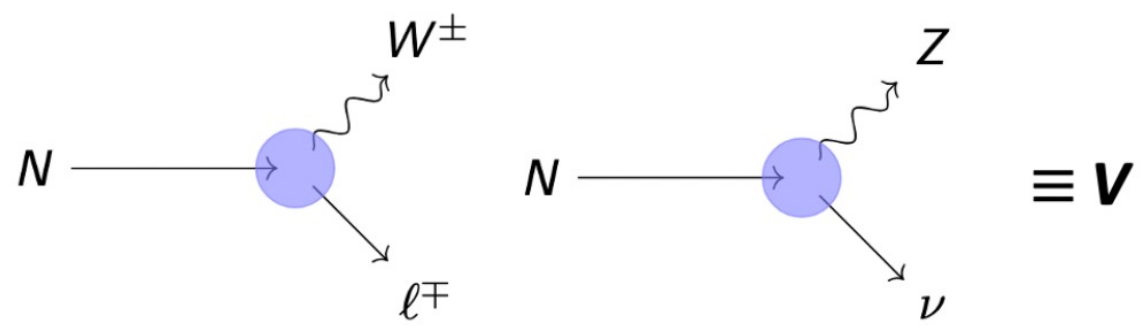
where $\tilde{H} = i\sigma^2 H^*$ and L_α are SM lepton doublets.

$$(M_D)_{N\alpha} = \frac{y_{N\alpha} v_h}{\sqrt{2}}$$



Diagonalisation of the mass matrix may result in mixing:

$$\nu_{L,\ell} = \sum_{\text{mass},i} U_{i,\ell} \nu_i + \sum_{\text{mass},j} V_{\ell,j} N_j$$

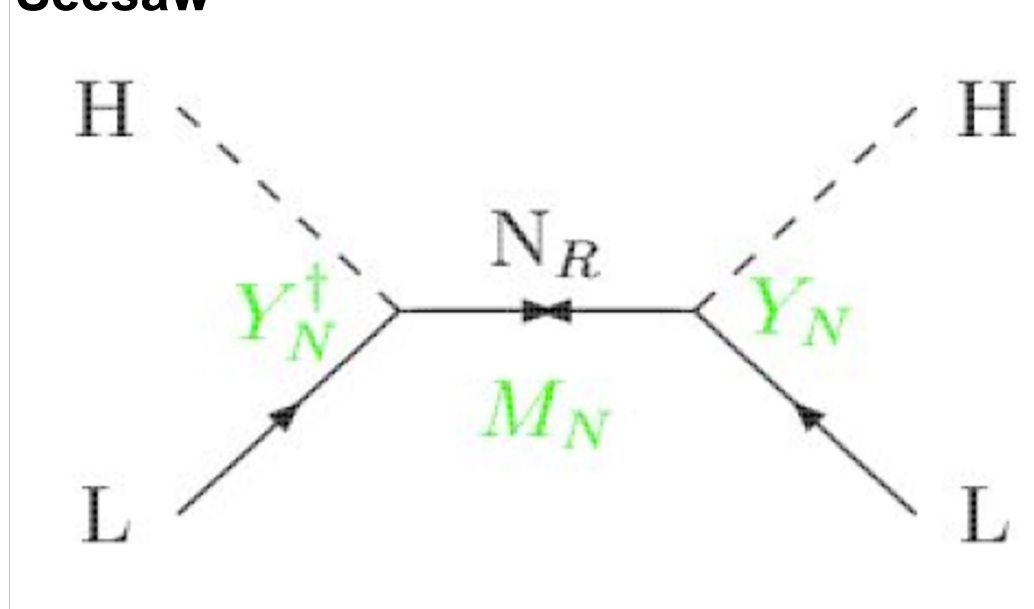


Type I Seesaw/Inverse Seesaw Mechanism

$$\mathcal{L} \supset y_{N\alpha} \bar{N} \tilde{H}^\dagger L_\alpha + h.c. + \dots$$

where $\tilde{H} = i\sigma^2 H^*$ and L_α are SM lepton doublets.

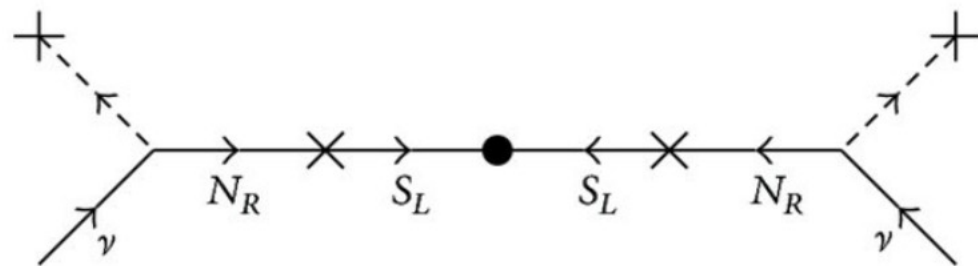
Type I Seesaw



$$m_\nu = \frac{\alpha v^2}{\Lambda} \equiv Y_N^T \frac{v^2}{M_N} Y_N$$

$$|V_{N\alpha}|^2 \simeq \left(|y_{N\alpha}| \frac{v}{m_N} \right)^2$$

Inverse Seesaw



$$|V_{N\alpha}|^2 \simeq \left(|y_{N\alpha}| \frac{v}{m_N} \frac{\mu}{m_N} \right)^2$$

Neutrino mass models - continued

Alternatively, consider a simple model with Dirac neutrinos and extra symmetries to decouple ν_L and N_R and still have non-zero masses.

Introduce the following fields: HNL given by (N_L, N_R) a right handed neutrino ν_R , and a complex scalar singlet.

$$\mathcal{L} \supset y \bar{L}_L \tilde{H} N_R + g \bar{N}_L \chi \nu_R + M \bar{N}_L N_R$$

arXiv:1411.5042

arXiv:1606.04543

Neutrino mass matrix

$$\mathcal{M} = \begin{pmatrix} 0 & m_1 \\ m_2 & M \end{pmatrix}$$

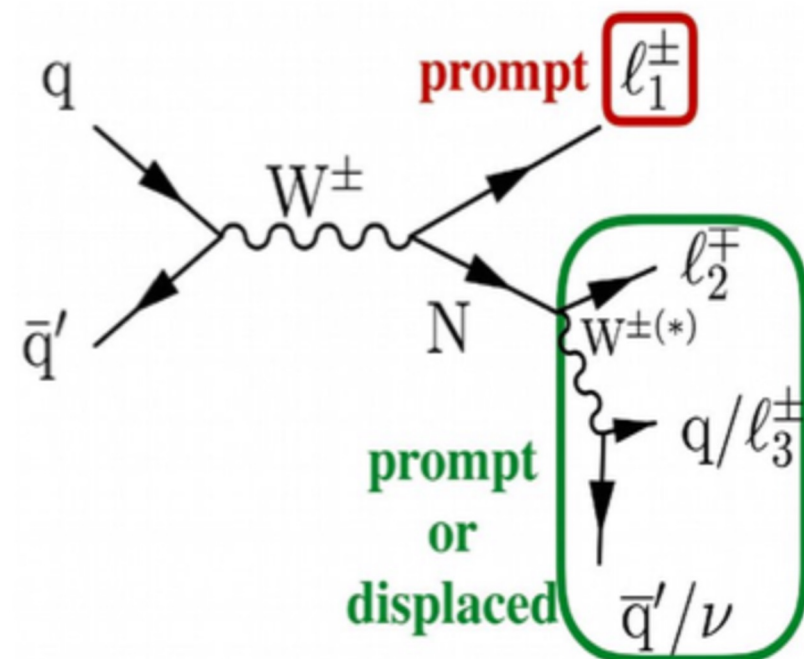
Can make mixing V to be zero taking $m_2 = 0$.

Phenomenologically we will just take a model independent approach such that the **mixing matrix**, **Yukawa coupling** and the **mass** as independent parameters

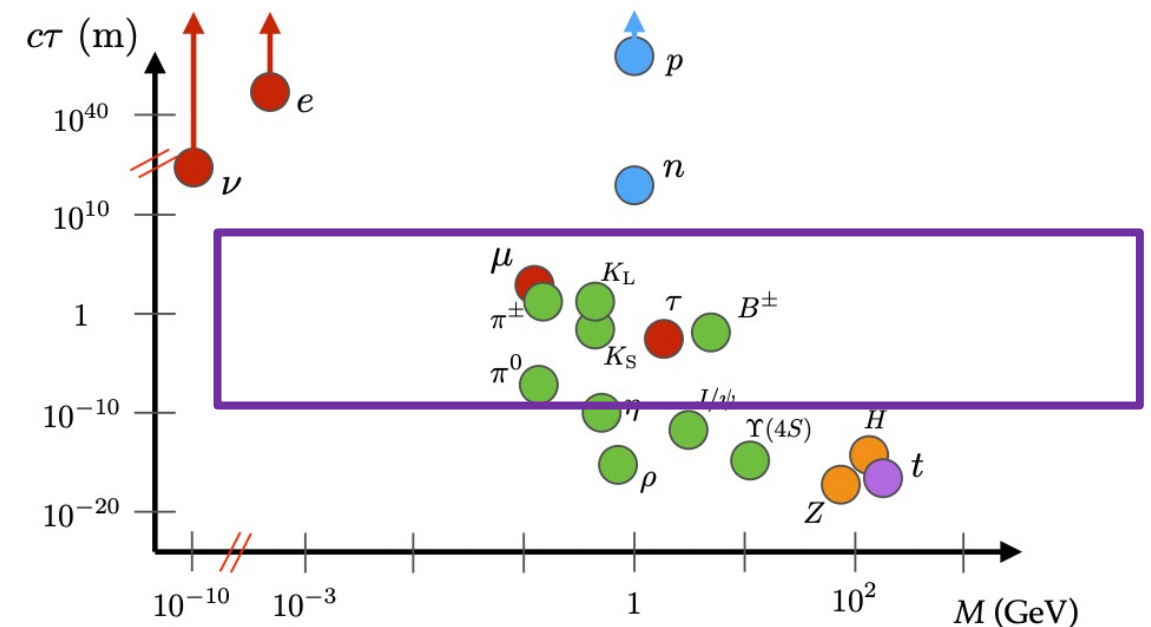
Long lived Particles

Displaced vertices in collider detectors

Longer lived LLPs use dedicated detectors



arxiv.org/pdf/1903.04497

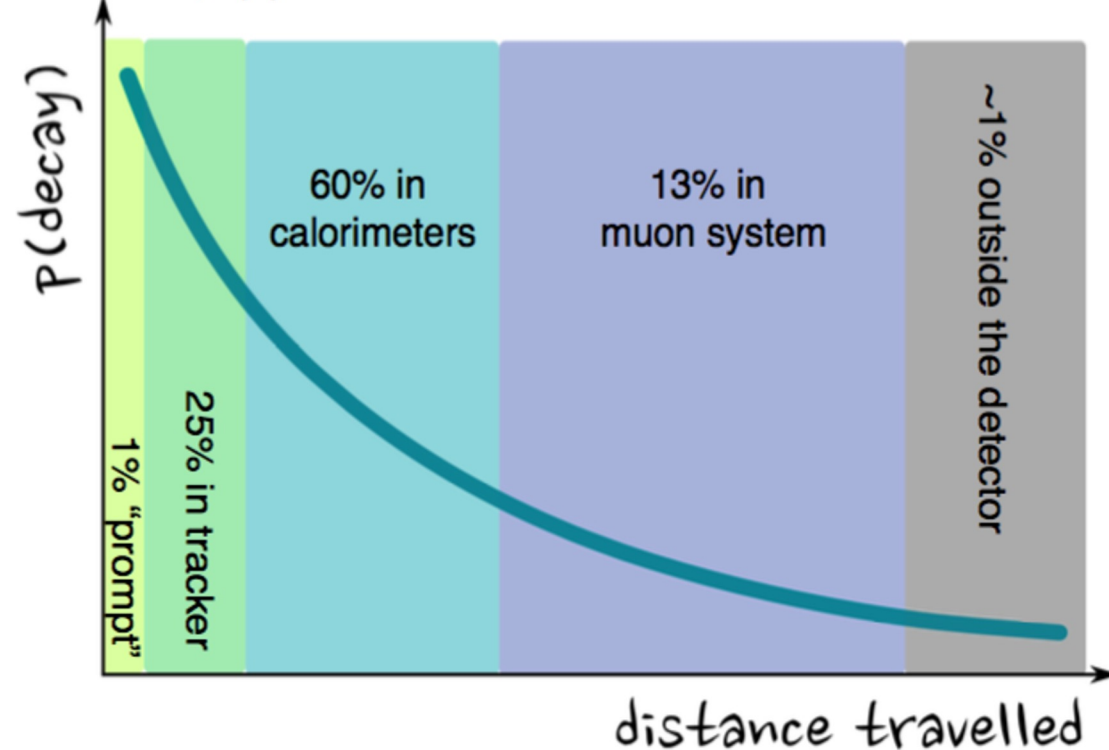


- Calculate the decay rate of the LLPs, Γ
- Corresponding lifetimes as a function of coupling and mass.
- Decay length: $d = p/M c\tau$

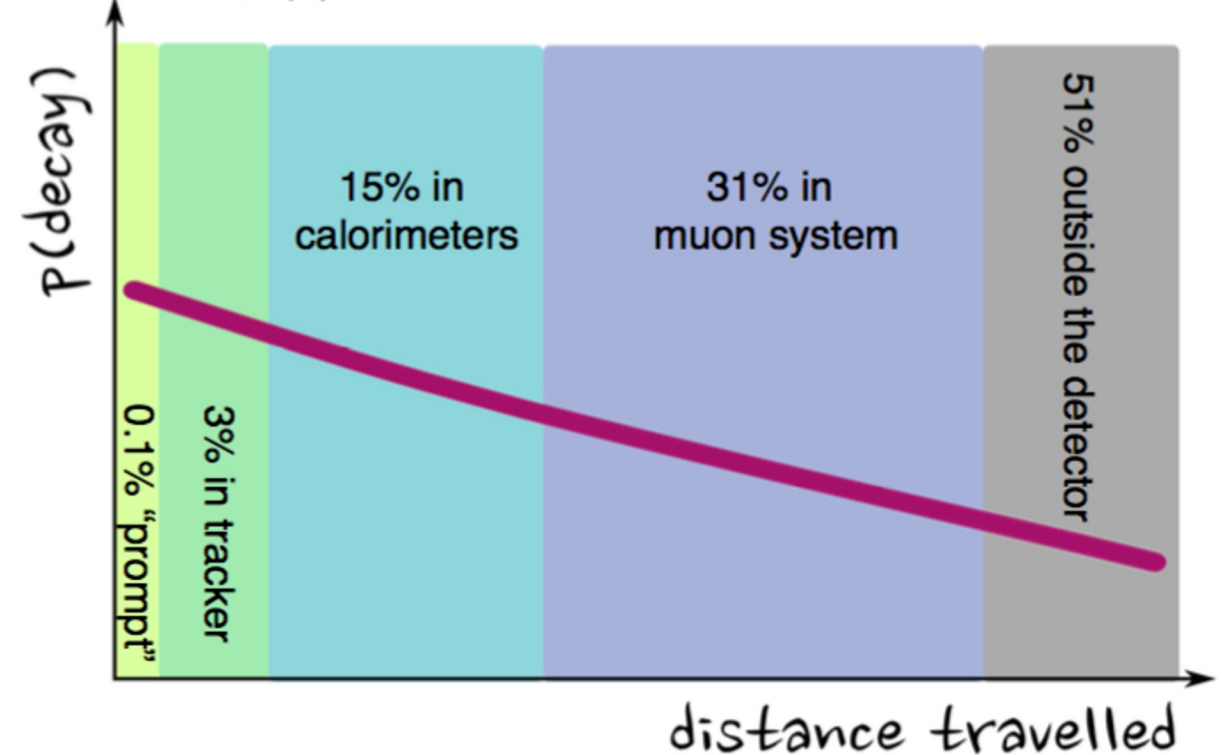
Sensitivity

Many factors affect signatures, exploit all detector components.....

e.g. for $c\tau = 5$ cm, $\langle\beta\gamma\rangle \sim 30$



e.g. for $c\tau = 50$ cm, $\langle\beta\gamma\rangle \sim 30$



H. Russell

$$N_{\text{events}} = \mathcal{L} \int d\sigma(\text{SM} \rightarrow \text{LPP}) \times BR(\text{LPP} \rightarrow \text{SM}) \times P_N(L) \times \epsilon_{\text{geo}}$$

Also will depend essentially on

Sensitivity will depend on:

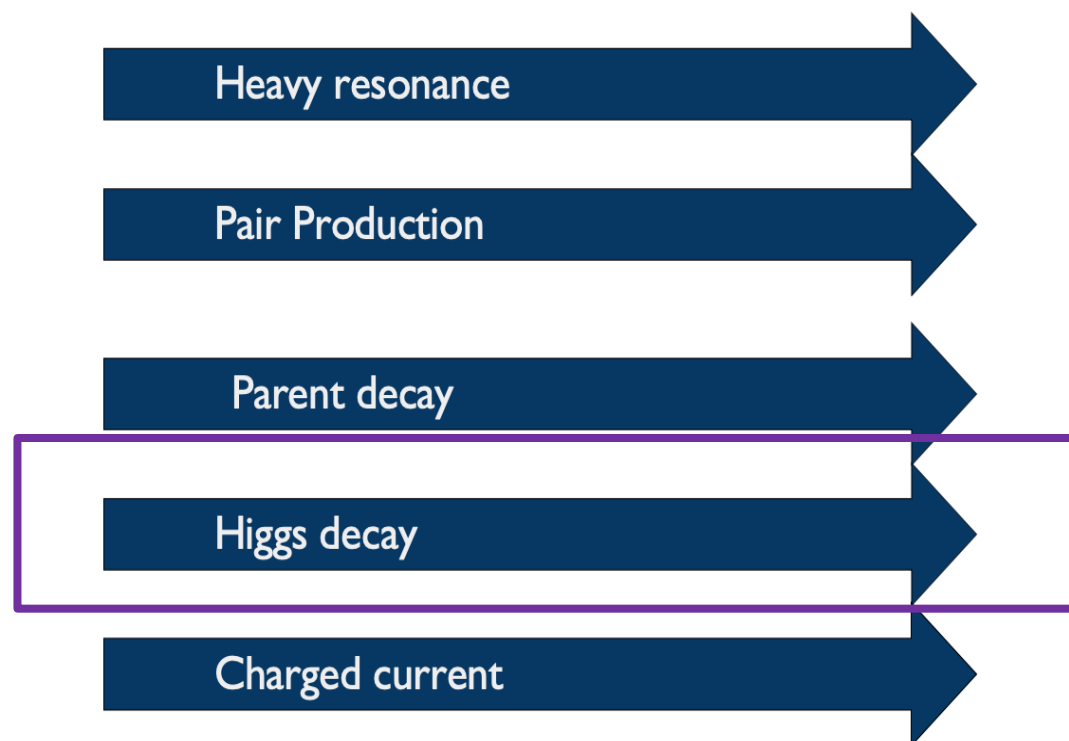
location, size, luminosity,...particle ID,

Mass of LLP

Interaction strength

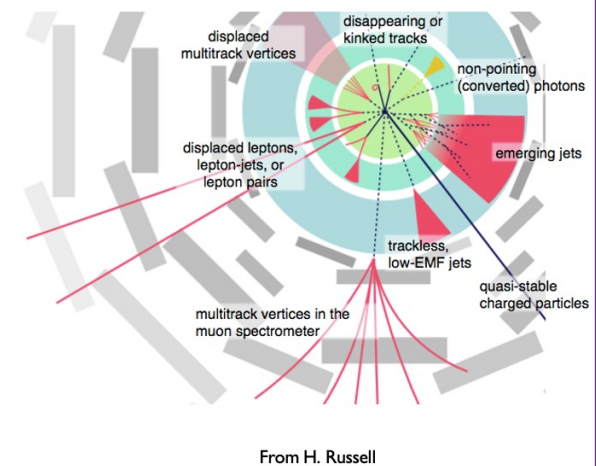
LLP Production and Decay

Colliders: LLP Production



Colliders: LLP decay modes and signatures

- Diphoton OR single photon
- Hadronic
- Semi-leptonic
- Leptonic
- Flavored leptonic
- Invisible decay



Intermediate lifetimes

LLP LHC White paper.

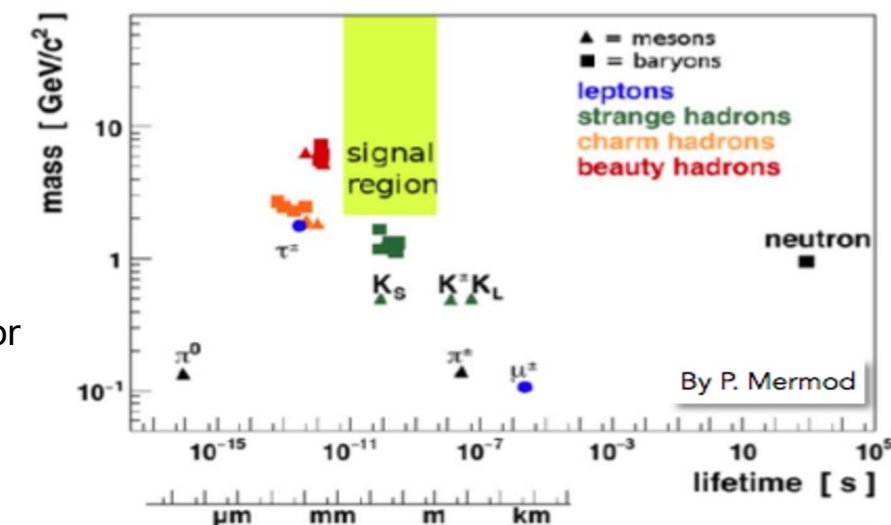
New objects not suitable to standard defs and algorithms
----**triggering!!!**

Unconventional backgrounds

Examples:

long lived particles in SM
hadronic detector interactions
particles from outside of collision
algorithmically induced fakes

Requiring more data-driven/AI techniques



The basics for HNL in type I SM + N

Width

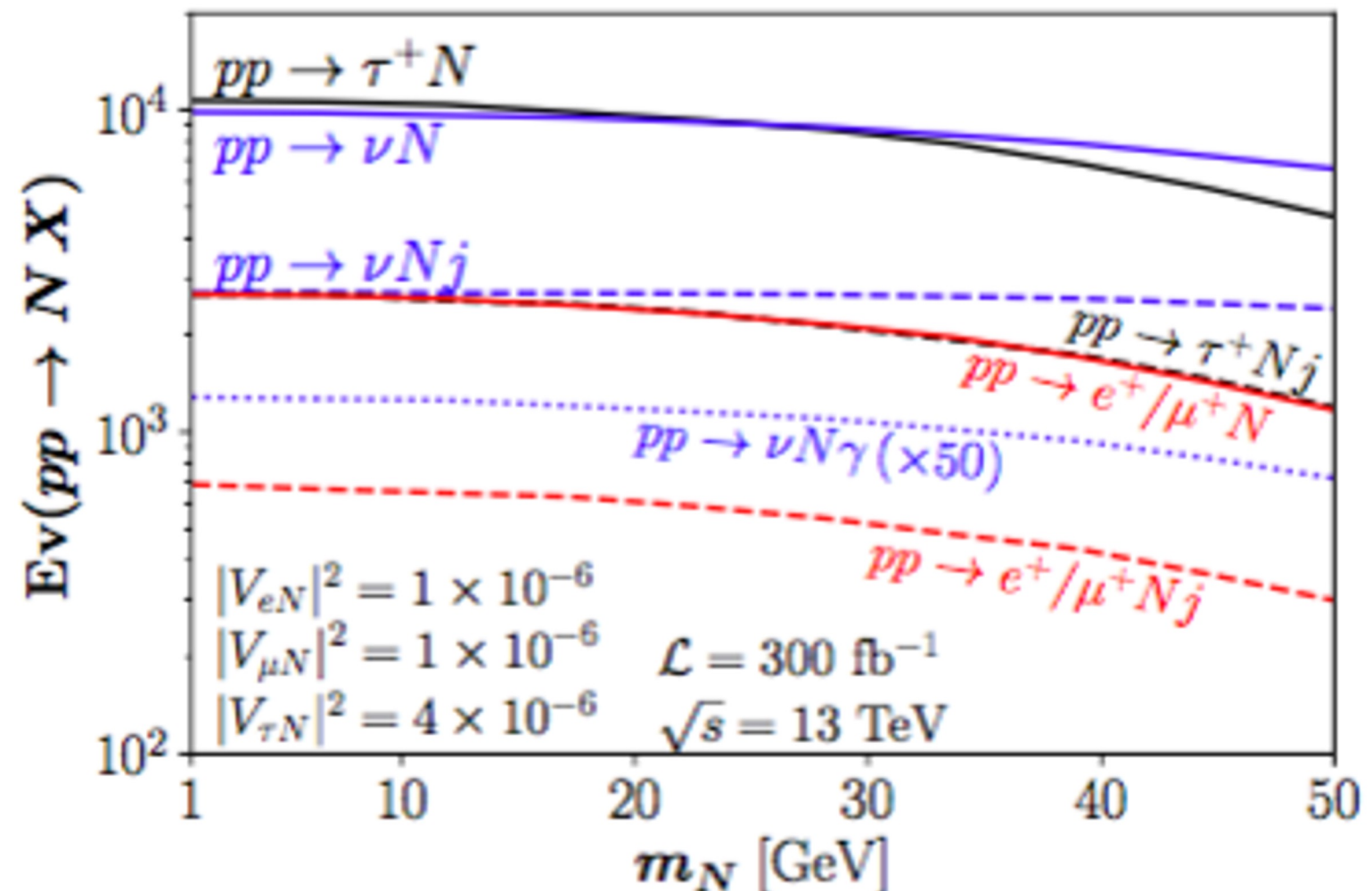
$$\text{BR}_{W^\pm \rightarrow \ell^\pm N} \propto |V_{\ell N}|^2$$

$$\text{BR}_{Z \rightarrow \nu N} \propto \sum_\ell |V_{\ell N}|^2, \quad \text{BR}_{H \rightarrow \nu N} \propto \left(\frac{m_N}{m_W}\right)^2 \sum_\ell |V_{\ell N}|^2$$

Production

Decay

Length



The basics for HNL in type I SM + N

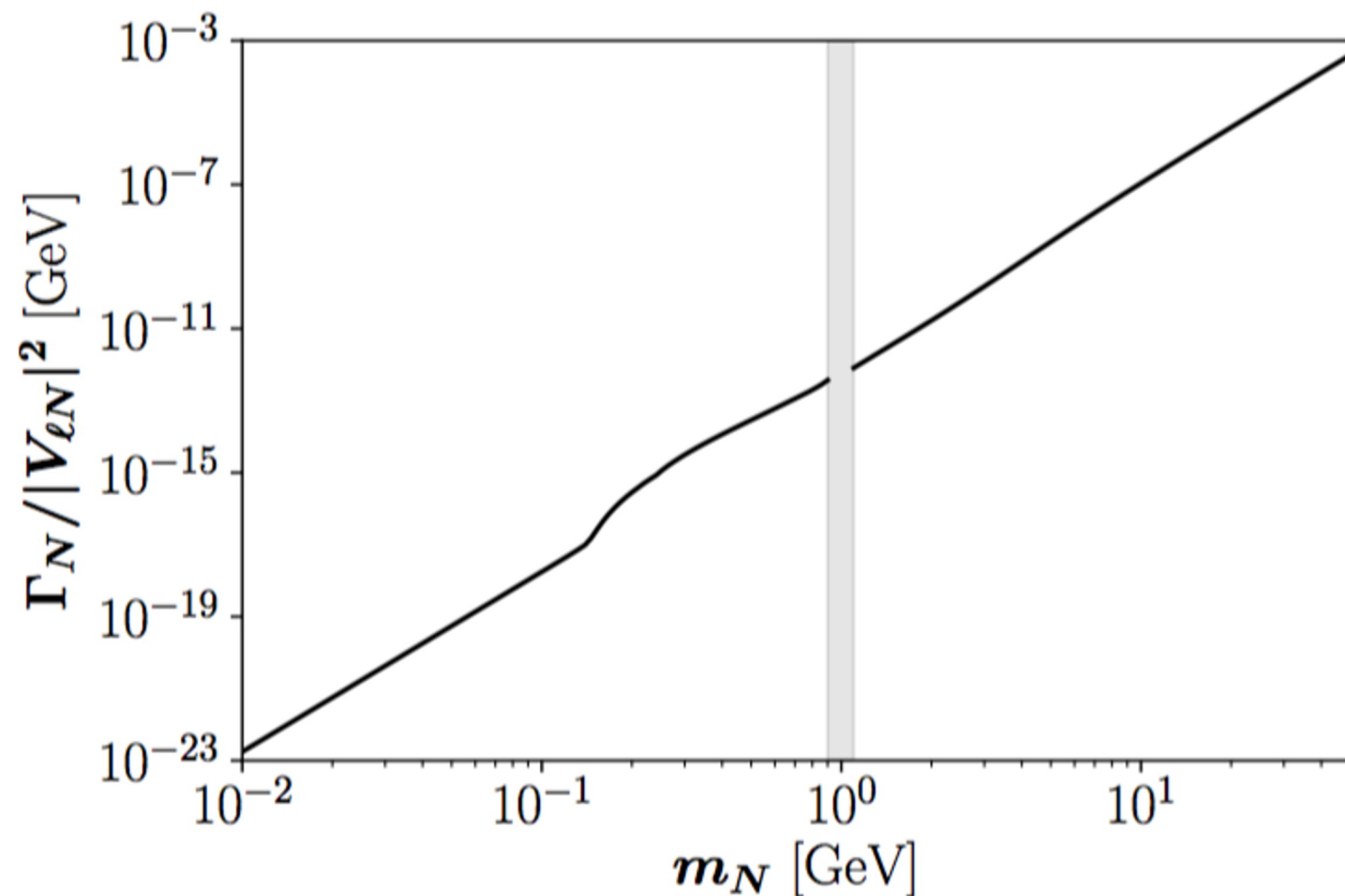
Width

$$\Gamma_N \propto G_F^2 m_N^5 \sum_{\ell=e,\mu,\tau} |V_{\ell N}|^2$$

Production

Decay

Length



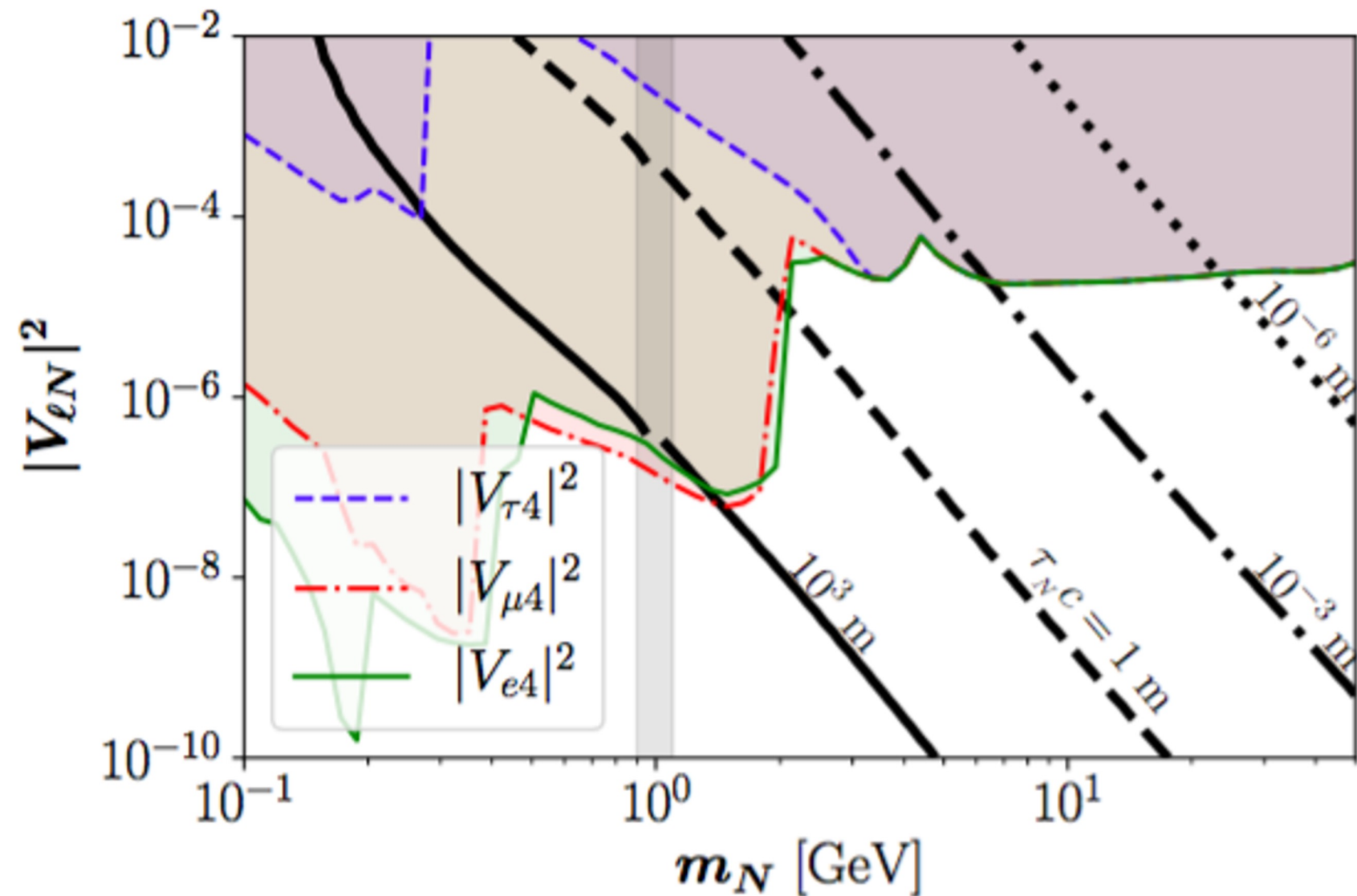
The basics for HNL in type I SM + N

Width

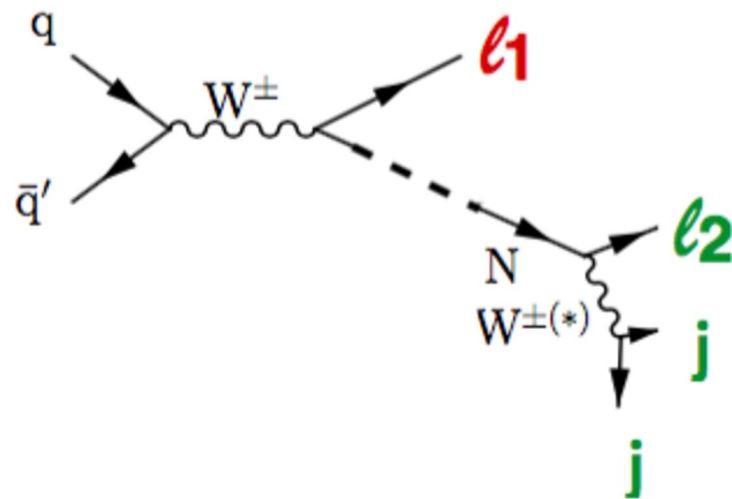
Production

Decay

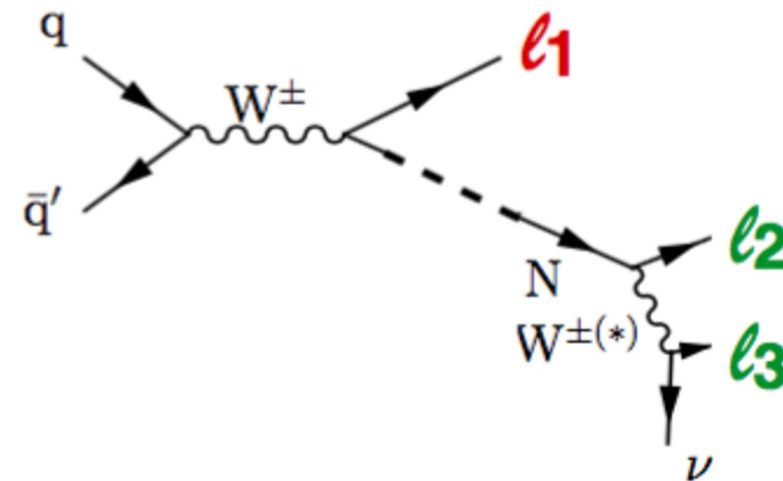
Length



Signatures



Dilepton + 2 jets
Invariant mass of N
Mostly sensitive to large m_N



No invariant N mass
Sensitive to low N

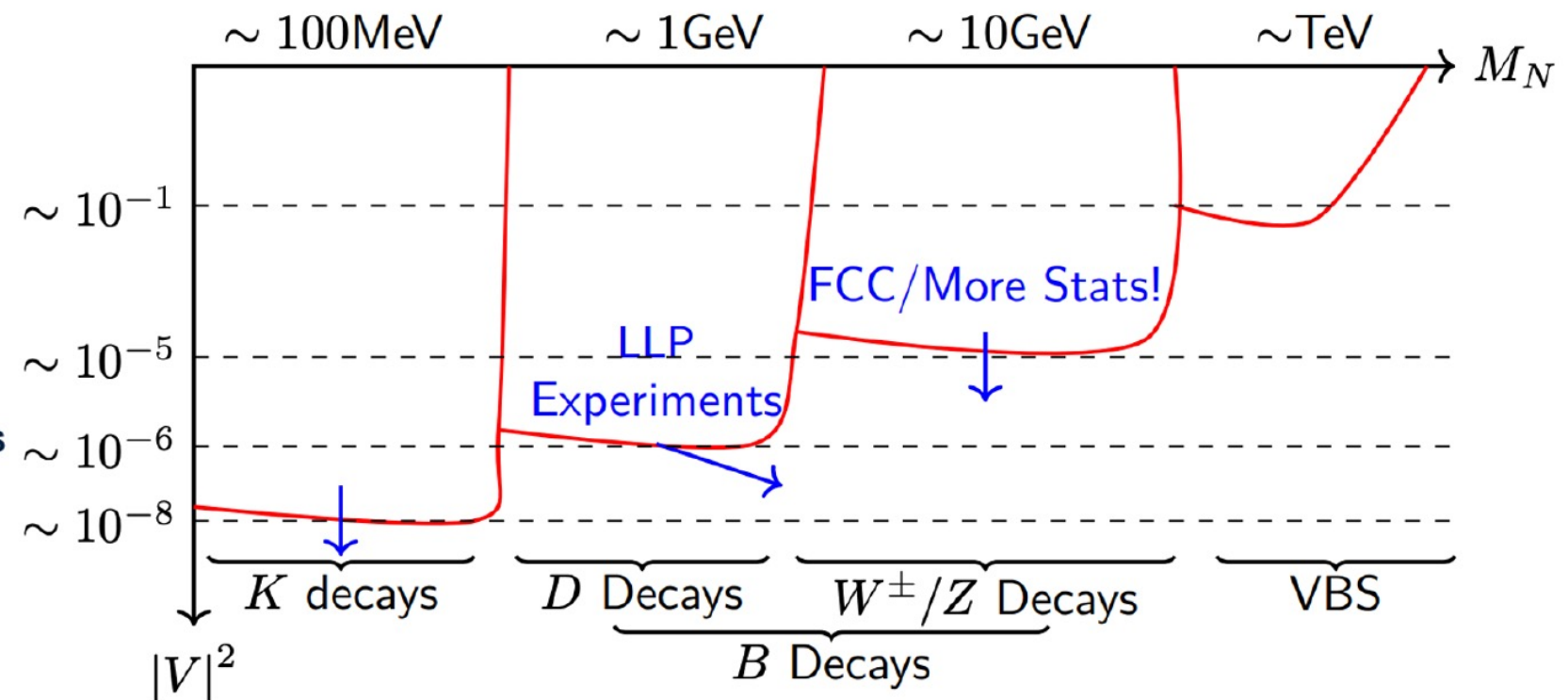
LNV and LNC, depending on OS and SS of the two higher p_t leptons

Experimental Results

Pushing the experimental frontiers

Collider constraints are usually shown in $V^2 - m_N$ plane

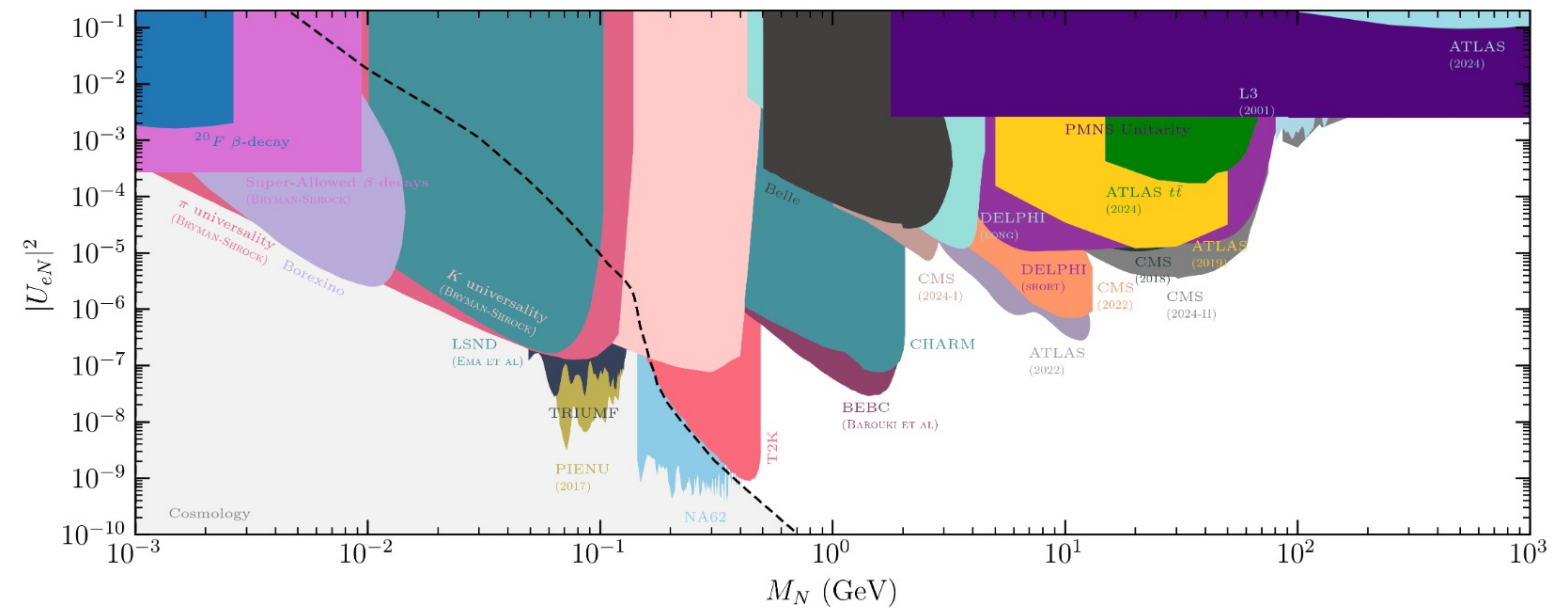
- **Below Kaon mass**
 - Kaon decays
 - e.g. NA62
- **Below B or D masses**
 - Heavy flavour decays
 - e.g. Belle, LHCb
- **Below W, Z masses**
 - Displaced and prompt searches
 - e.g. LEP, **LHC**
- **Above W,Z masses**
 - Decays to on shell bosons
 - e.g. **LHC**



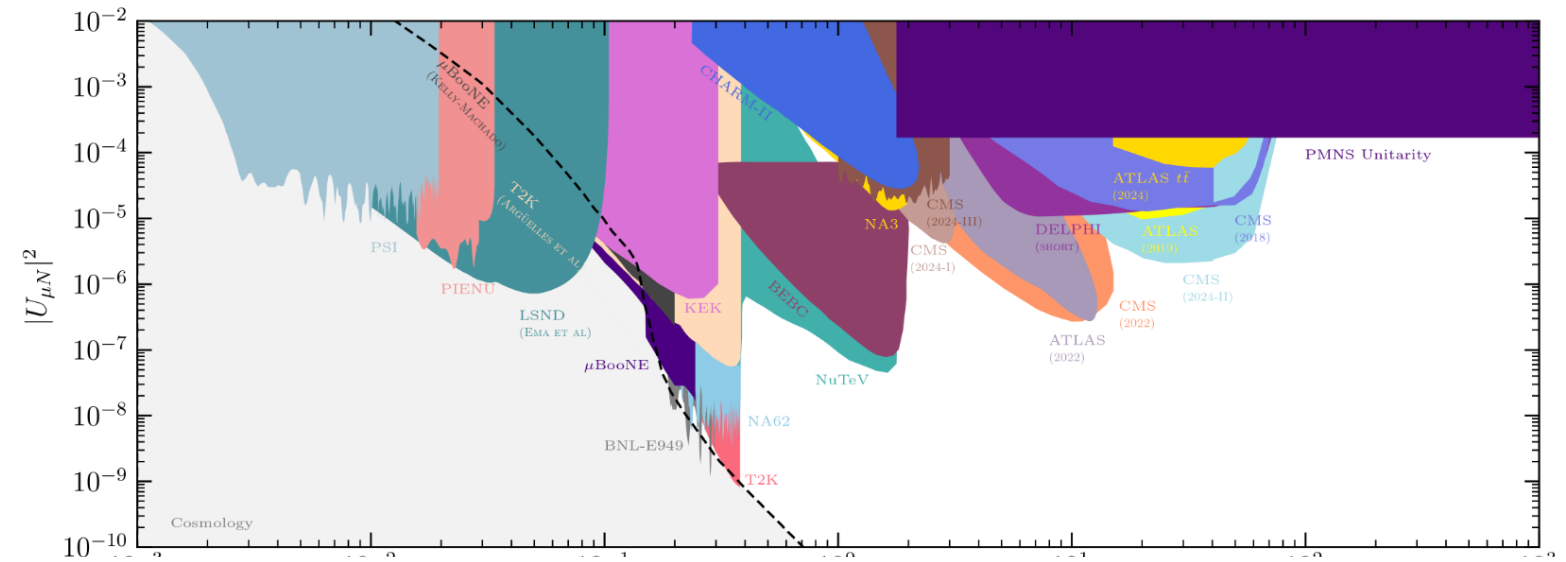
https://warwick.ac.uk/fac/sci/physics/research/epp/events/seminars/hnl_seminar_warwick_g_bird.pdf

Current Bounds

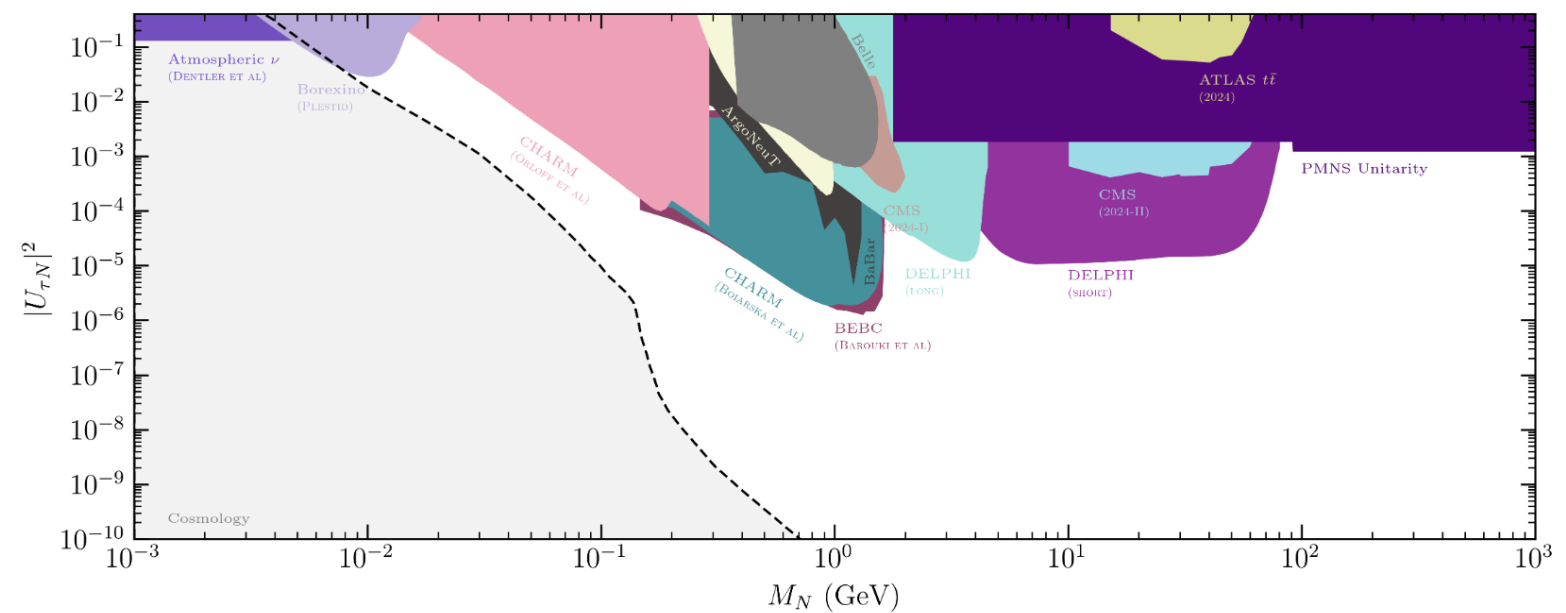
Electron mixing



Muon mixing

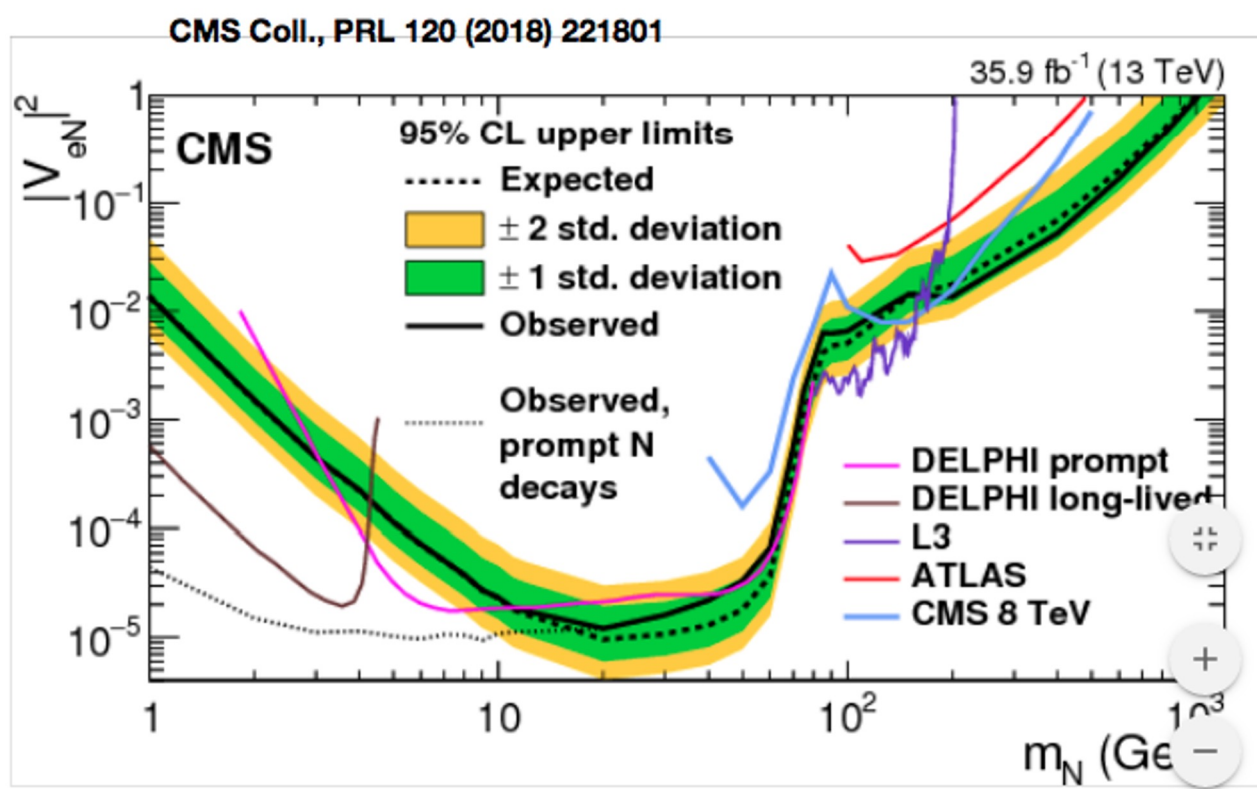
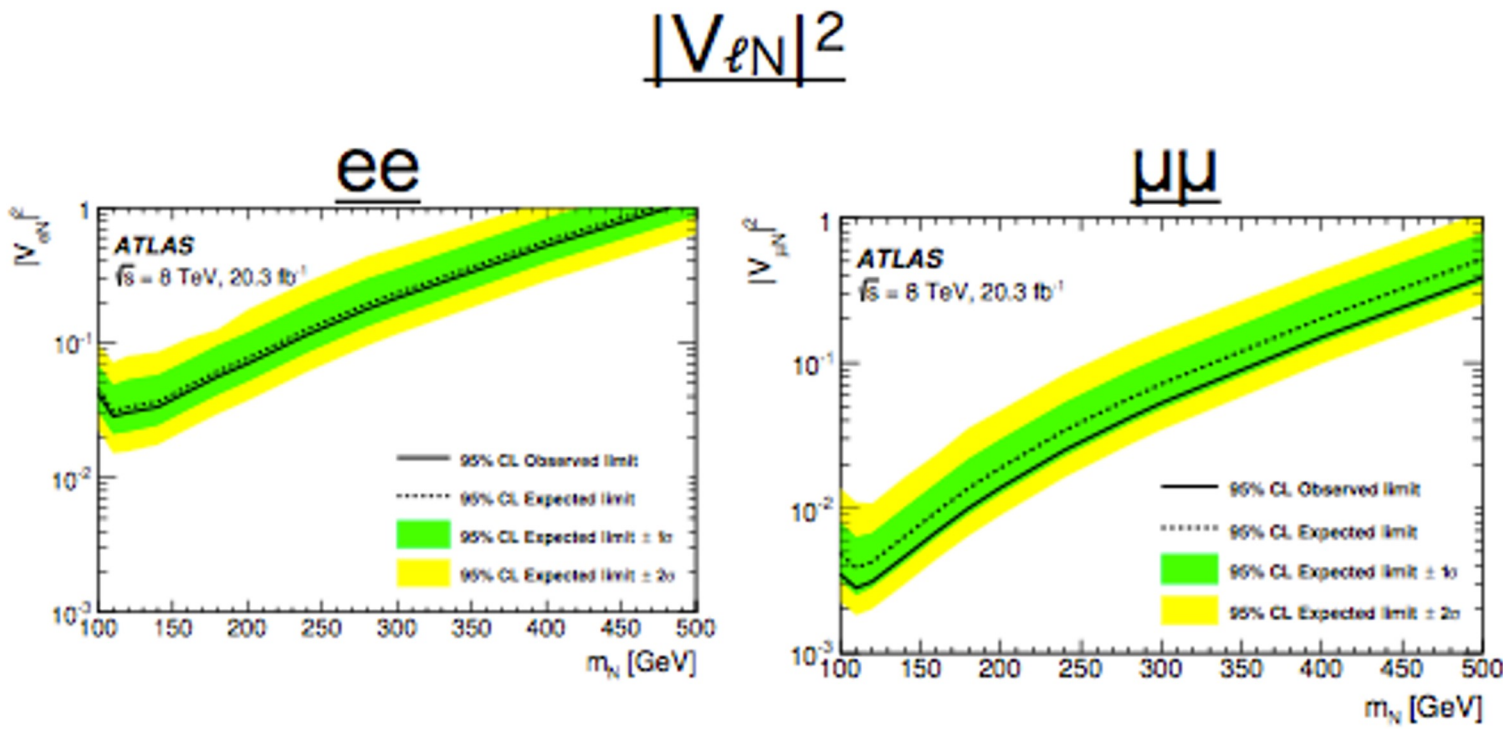
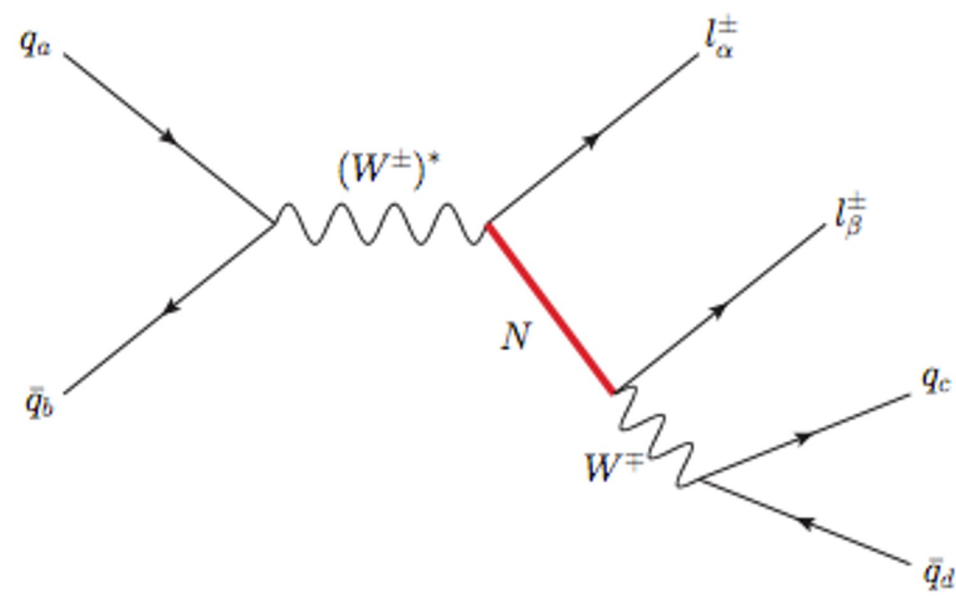


Tau mixing



LHC Results: Prompt HNL ATLAS, CMS searches

$$M_N > M_Z$$

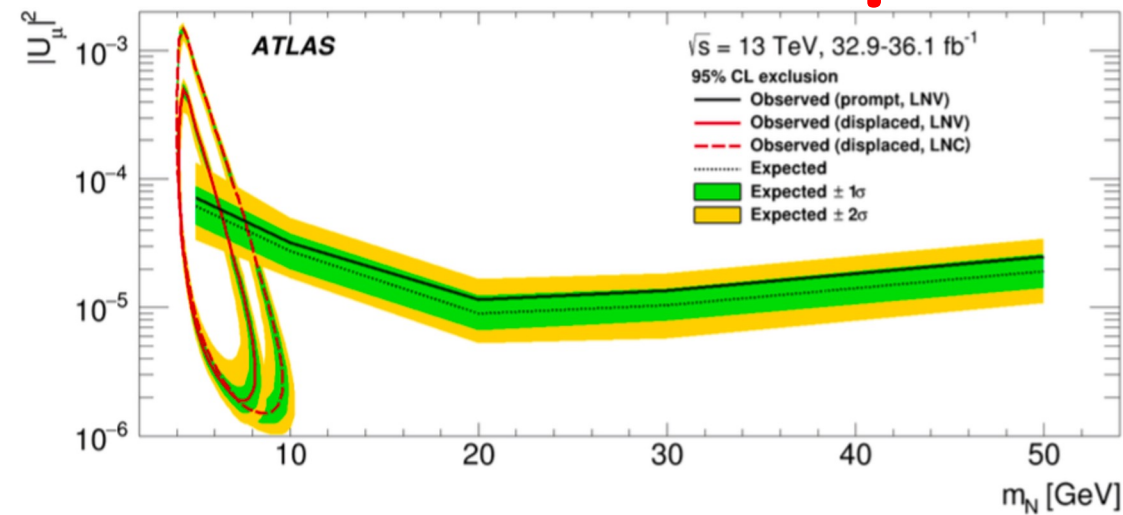


ATLAS/CMS DV

1905.09787

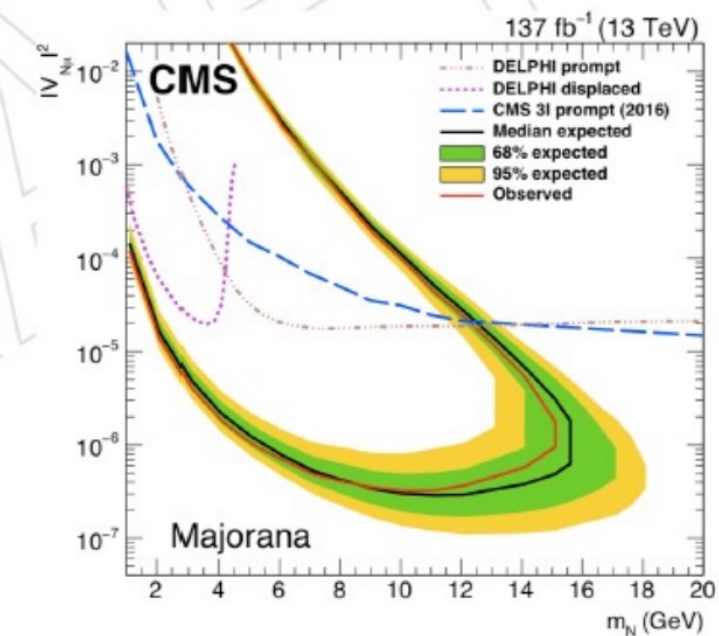
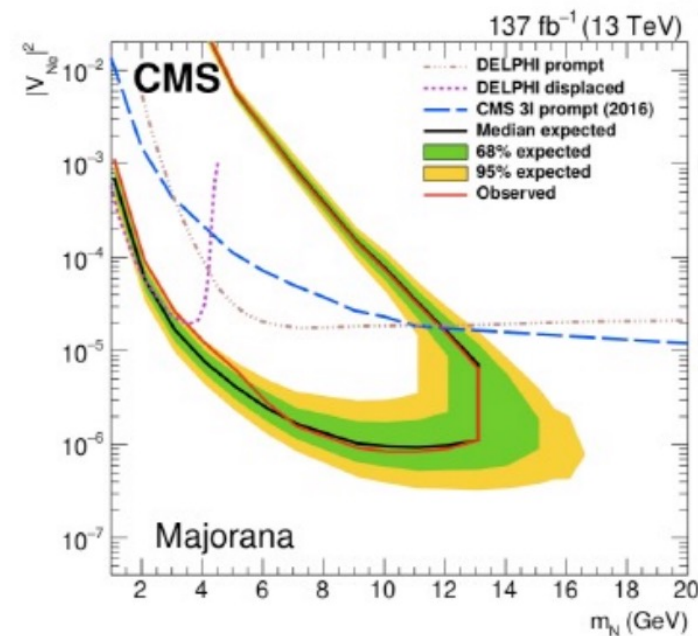
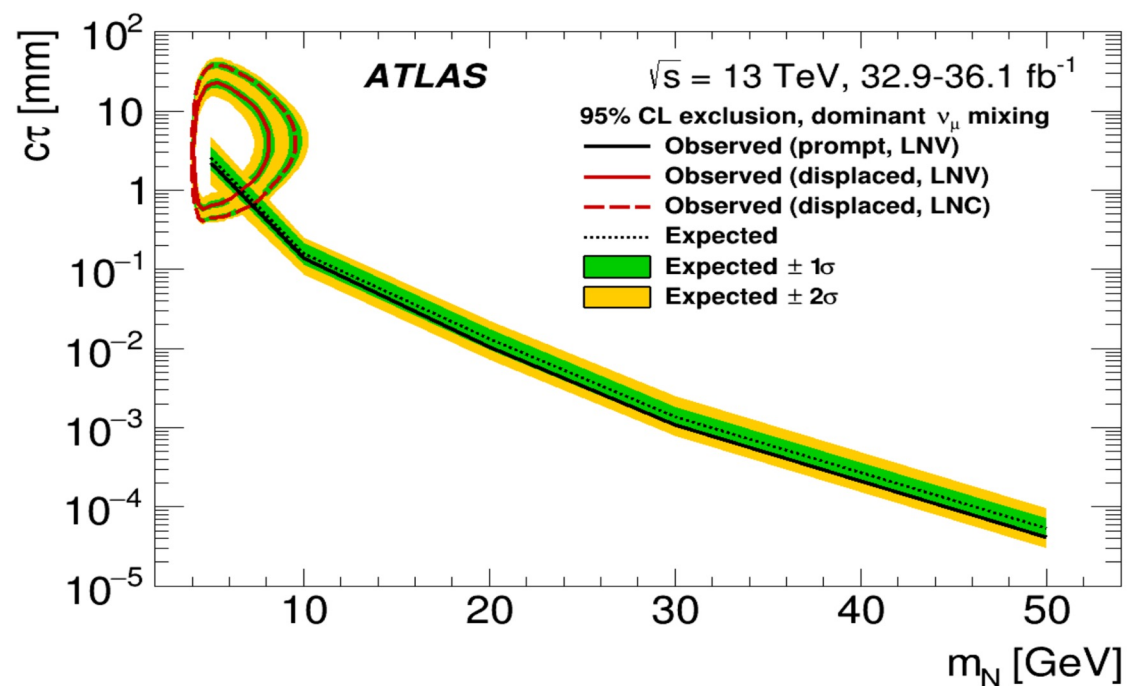


$$M_N < M_W$$



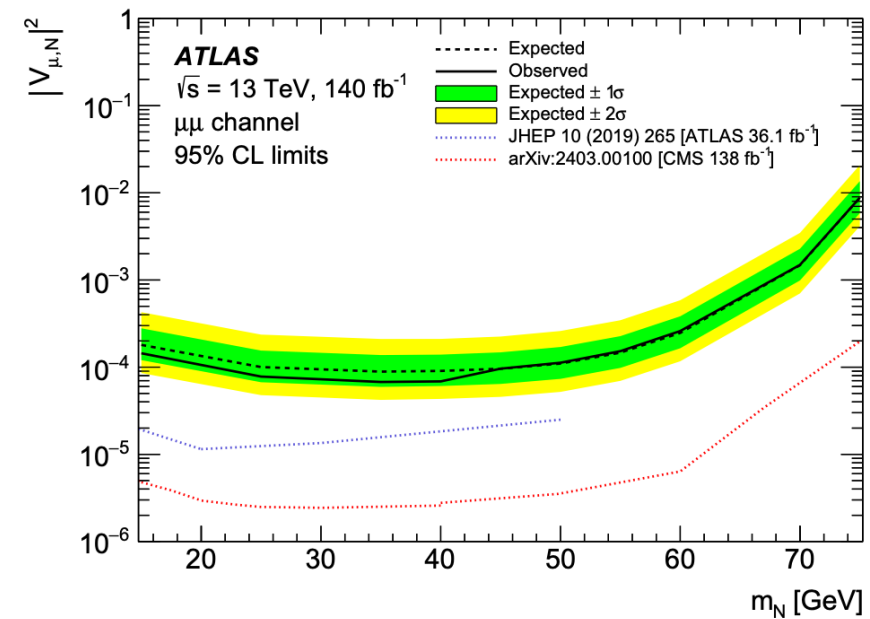
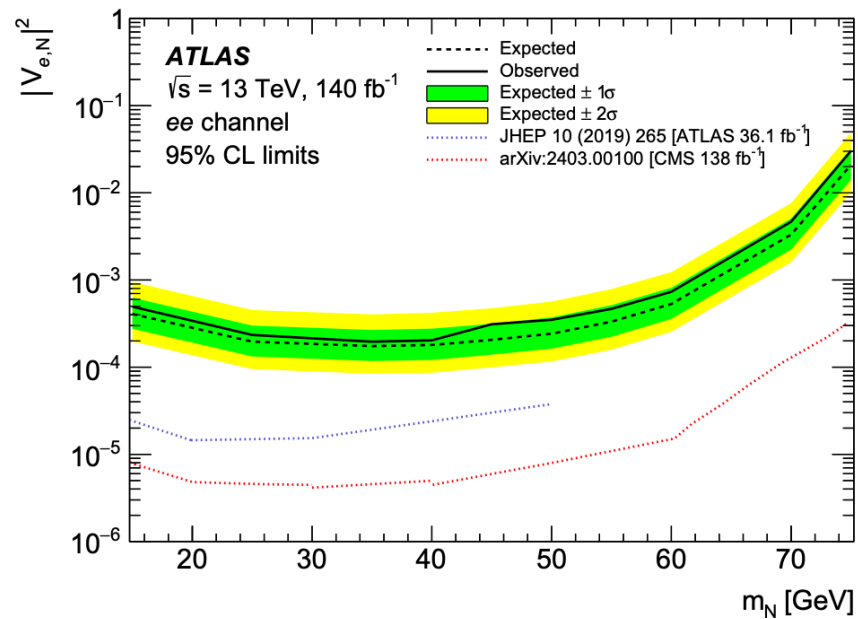
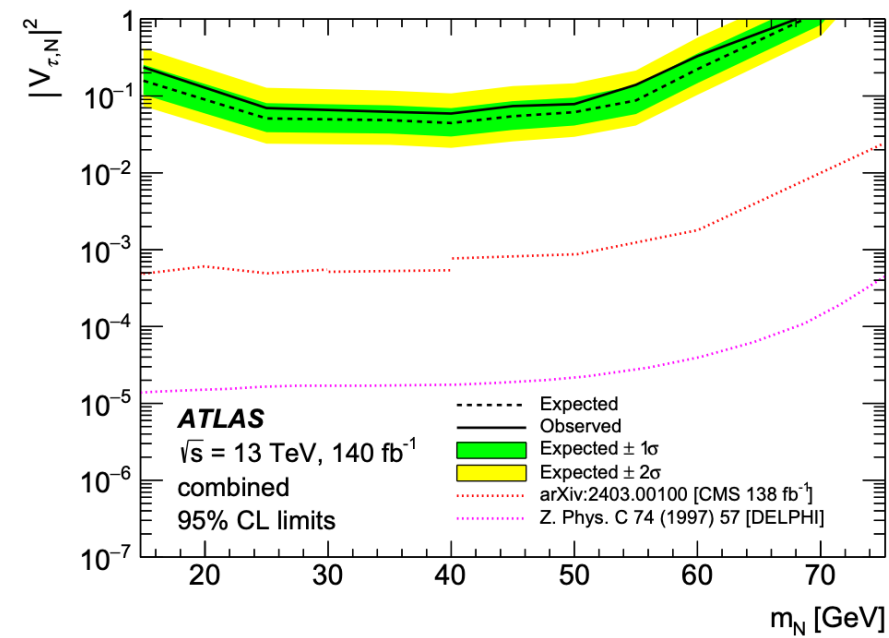
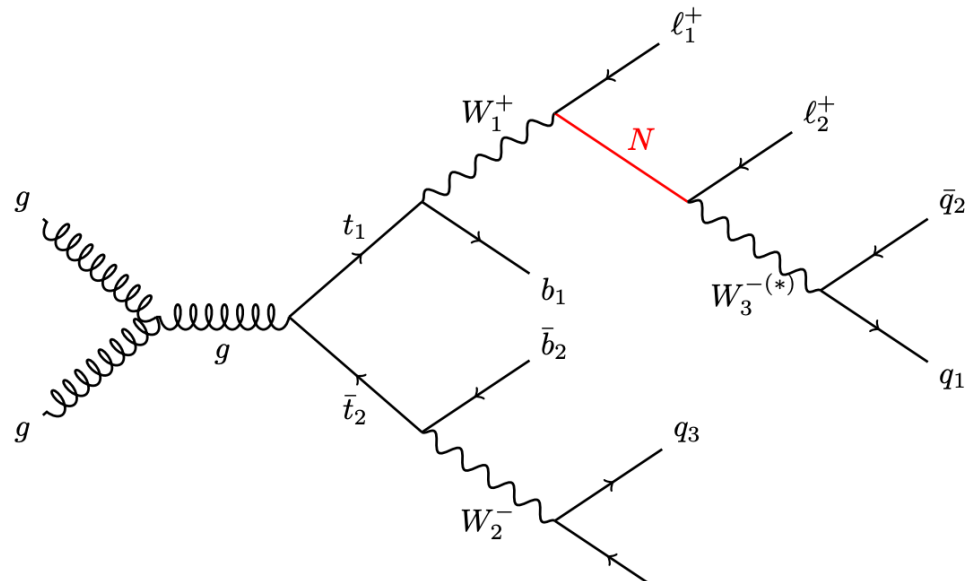
Electrons and Muons

2201.05578.

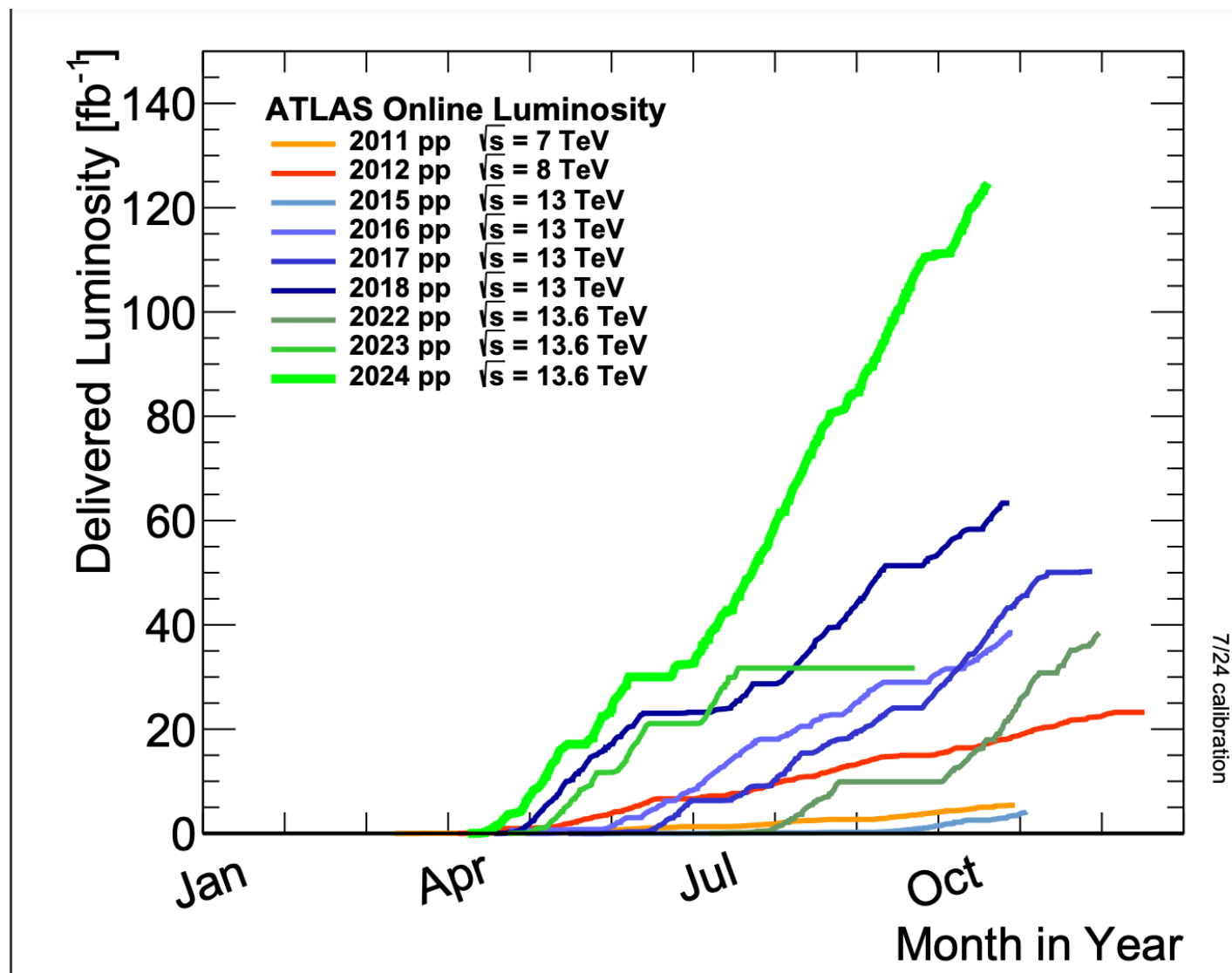


ATLAS $t\bar{t}b\bar{b}$ analysis

2408.05000



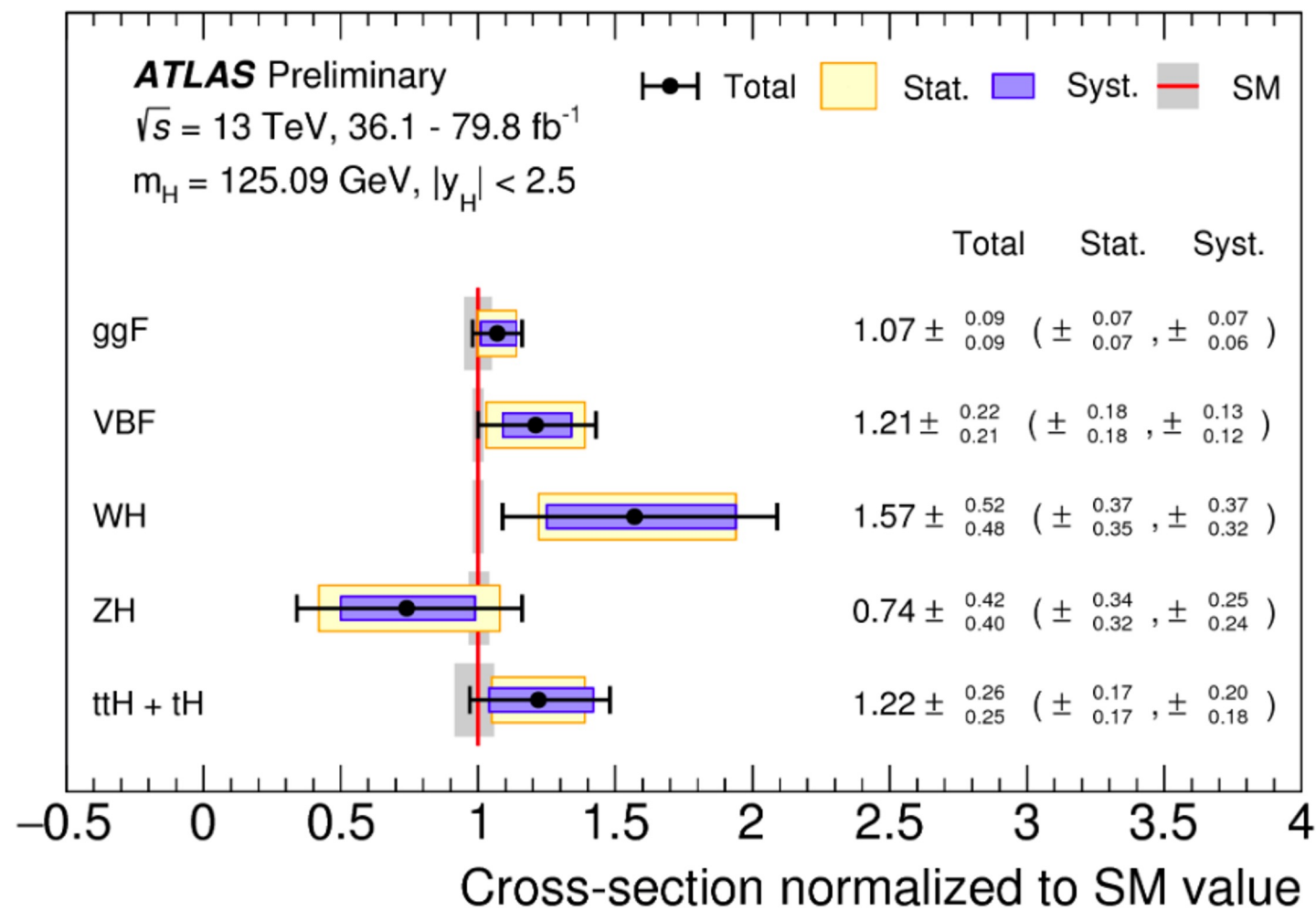
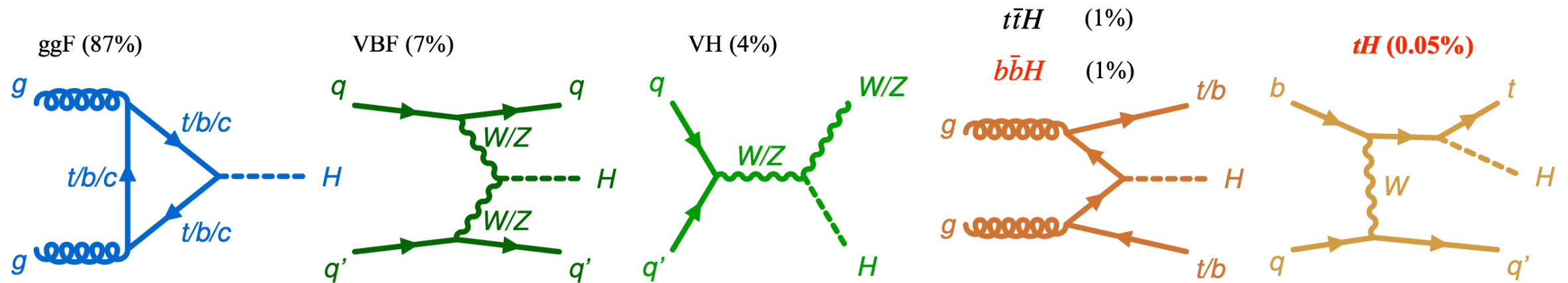
LHC delivery for 13.6 TeV an integrated luminosity of 195 fb⁻¹



Still only about 10% of data that will be produced.

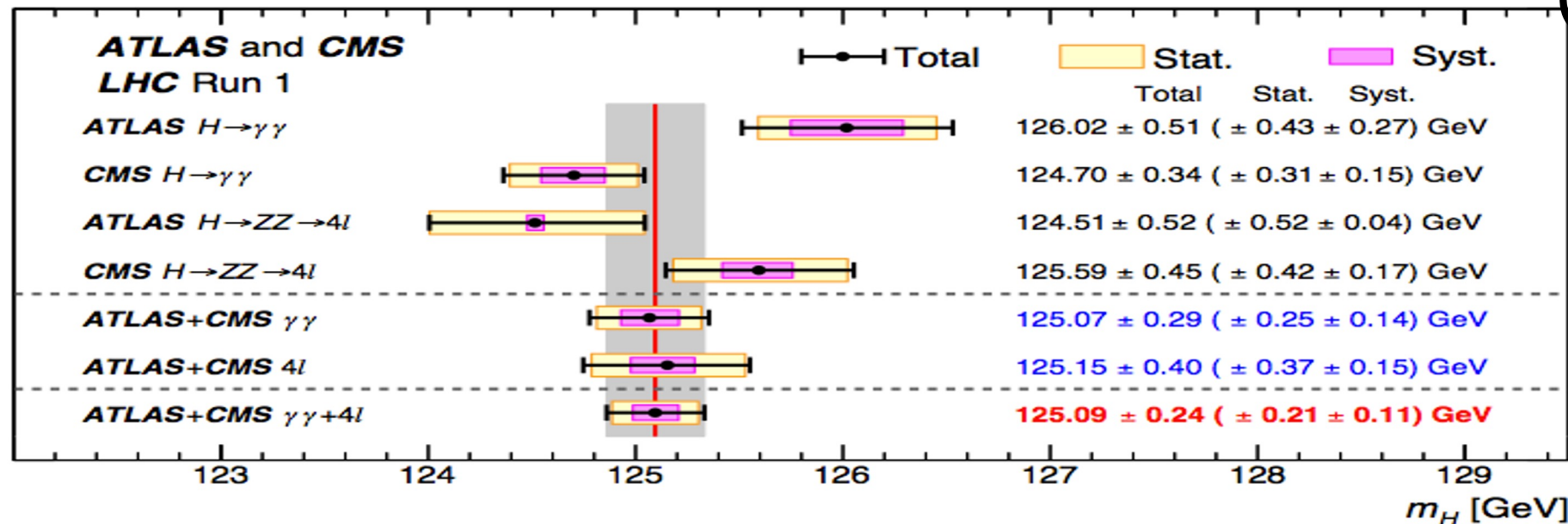
Key parameter is the integrated luminosity $N_{\text{events}} = \sigma \int L dt$

Higgs production



Mass measurement Run I

@ $m_H = 125 \text{ GeV}$



$$\hat{m}_H = 125.09 \pm 0.21(stat) \pm 0.11(syst) \text{ GeV}$$

$$H \rightarrow WW \quad 22\%$$

$$H \rightarrow ZZ \quad 2.9\%$$

$$H \rightarrow bb \quad 58\%$$

$$H \rightarrow cc \quad 2.8\%$$

$$H \rightarrow \gamma\gamma \quad 0.23\%$$

$$H \rightarrow \gamma Z \quad 0.16\%$$

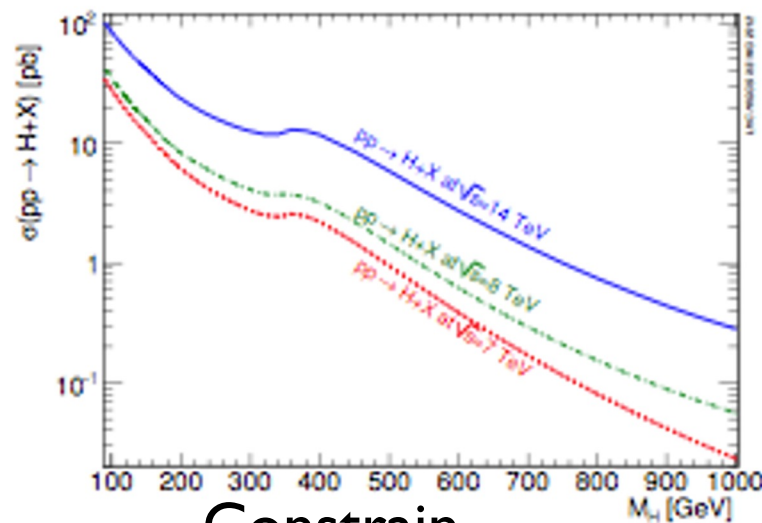
Currently (PDG):

$$m_H = (125.20 \pm 0.11) \text{ GeV}$$

$$m_t = (172.57 \pm 0.29) \text{ GeV}$$

With known mass then everything is fixed

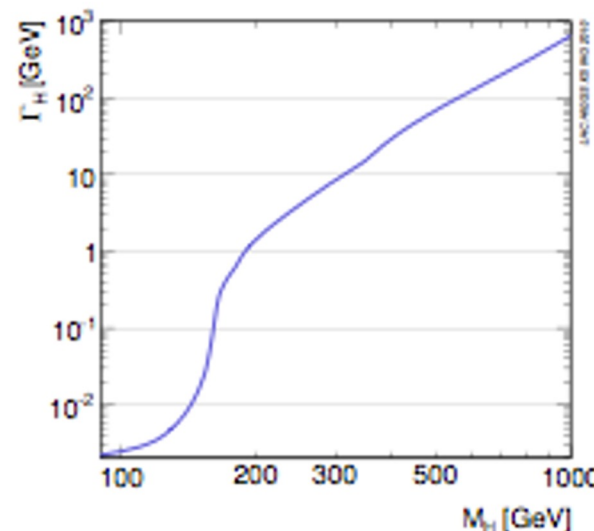
cross-section



Constrain

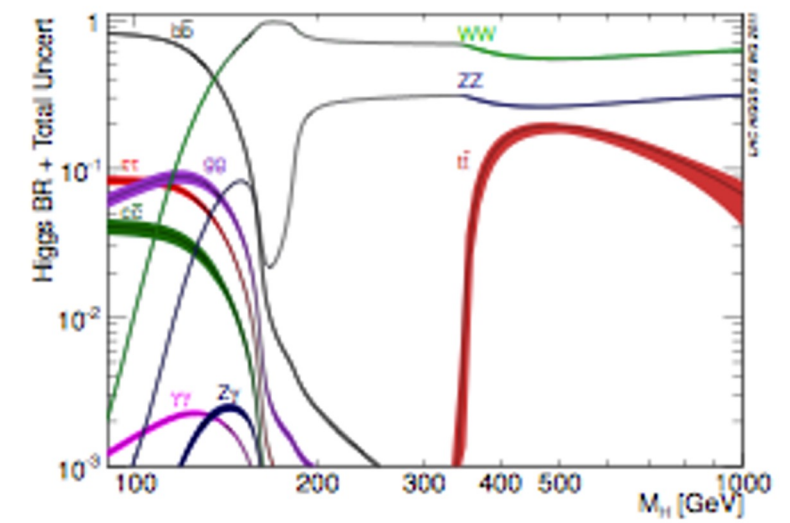
B_{BSM}

decay width



$$\Gamma \approx 4.1 \text{ MeV}$$

branching ratios



$$\lambda = (m_H/v)^2 = 0.131$$

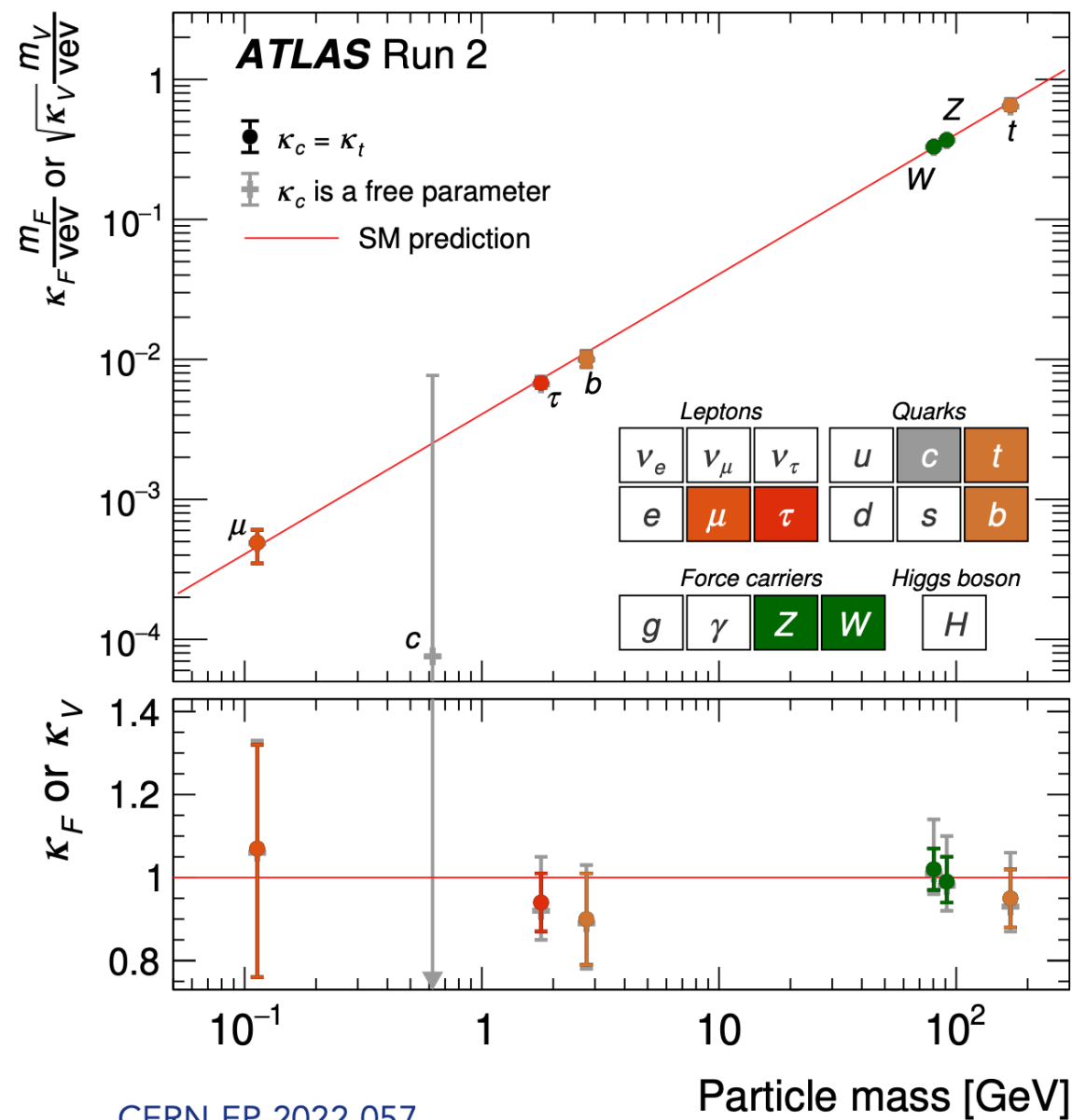
ATLAS
 $Br(h \rightarrow inv) < 10.7\%$

ATLAS: $4.5^{+3.3}_{-2.5} \text{ MeV}$,
CMS: $2.9^{+2.3}_{-1.7} \text{ MeV}$

$$\Gamma_h = 3.2^{+2.4}_{-1.7} \text{ MeV}$$

Run II Probe Yukawa couplings

Cross section 4 X higher @ 13 TeV wrt 8 TeV



CERN-EP-2022-057

2018 : direct observation of top quark Yukawa

2020: evidence for muon Yukawa

2022:: hints of Higgs responsible for c-quark mass

Increasing precision way above expectation

Until now all consistent with the SM !!

Future Experiments and Colliders

HL-LHC
FCC-ee
FCC-hh
CEPC
CLIC

FASER/FASER-2

Codex-b

Mathusla

milliQan

NA62

SeaQuest

SHIP

DUNE-ND

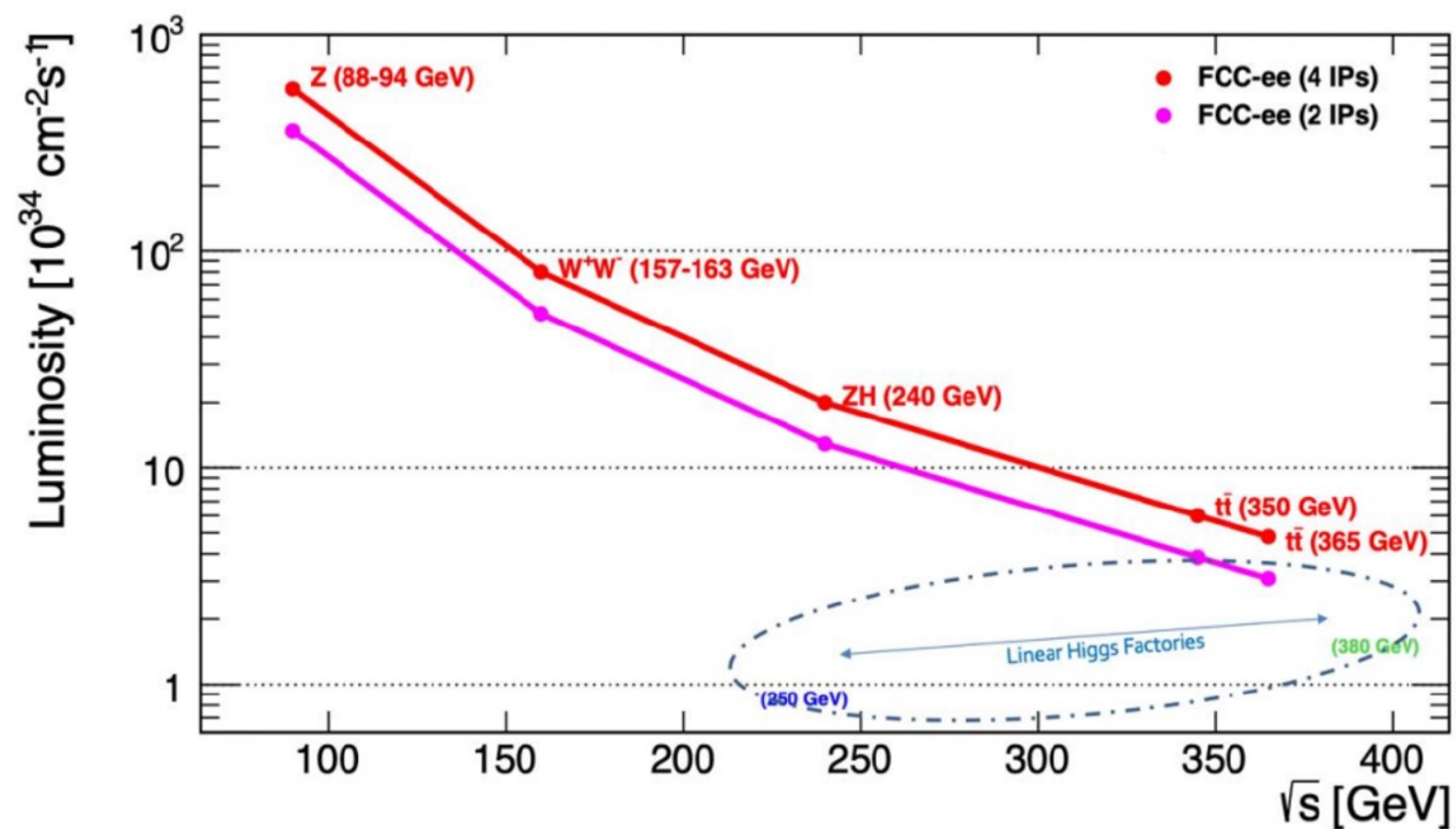
....

Mixing with **electrons**: NA62-dump, NA62 K⁺ decays, FASER, PIONEER, SHADOWS, DarkQuest, SHiP, DUNE, T2K, Hyper-K

Mixing with **muons**: NA62-dump, NA62 K⁺ decays, FASER, DarkQuest, SHADOWS, PIONEER, SHiP, DUNE, Hyper-K

Mixing with **taus**: NA62-dump, FASER, SHADOWS, DUNE, DarkQuest, and SHiP

FCC



| Working point | Z, years 1-2 | Z, later | WW, years 1-2 | WW, later | ZH | $t\bar{t}$ |
|--|----------------------|----------|----------------------|-----------|--|---|
| \sqrt{s} (GeV) | 88, 91, 94 | | 157, 163 | | 240 | 340–350 365 |
| Lumi/IP ($10^{34} \text{ cm}^{-2} \text{ s}^{-1}$) | 70 | 140 | 10 | 20 | 5.0 | 0.75 1.20 |
| Lumi/year (ab^{-1}) | 34 | 68 | 4.8 | 9.6 | 2.4 | 0.36 0.58 |
| Run time (year) | 2 | 2 | 2 | – | 3 | 1 4 |
| Number of events | 6×10^{12} Z | | 2.4×10^8 WW | | 1.45×10^6 ZH + 45k WW \rightarrow H | 1.9×10^6 $t\bar{t}$ +330k ZH +80k WW \rightarrow H |

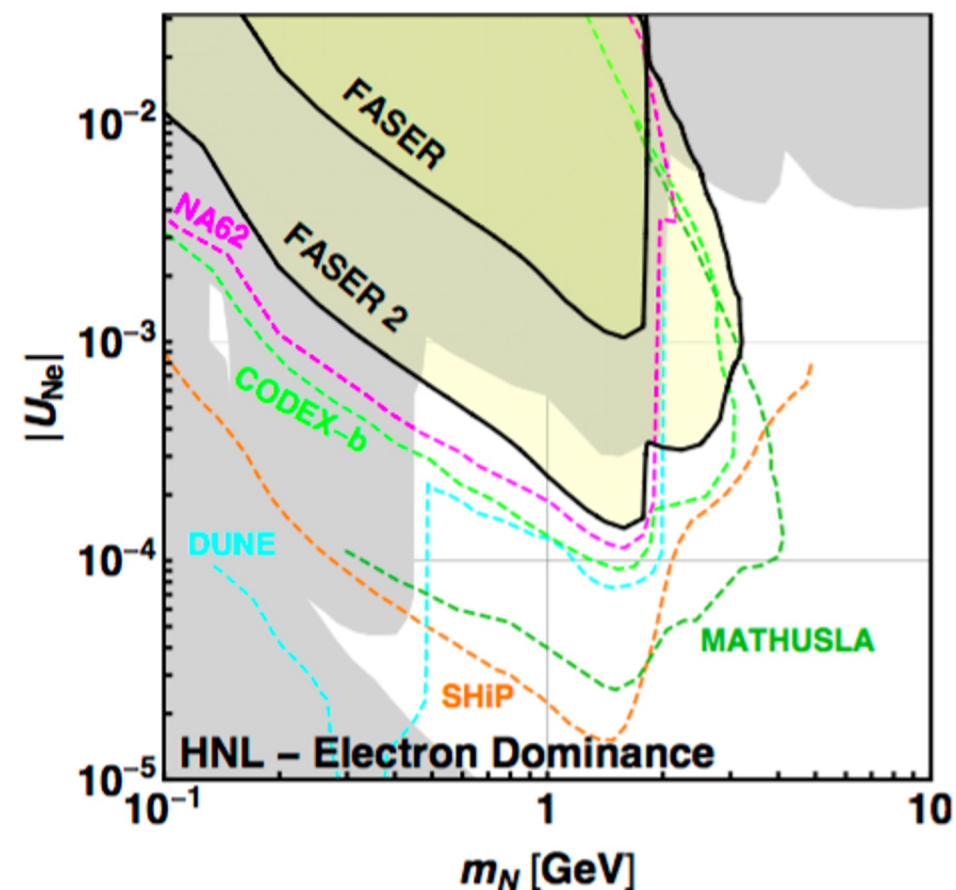
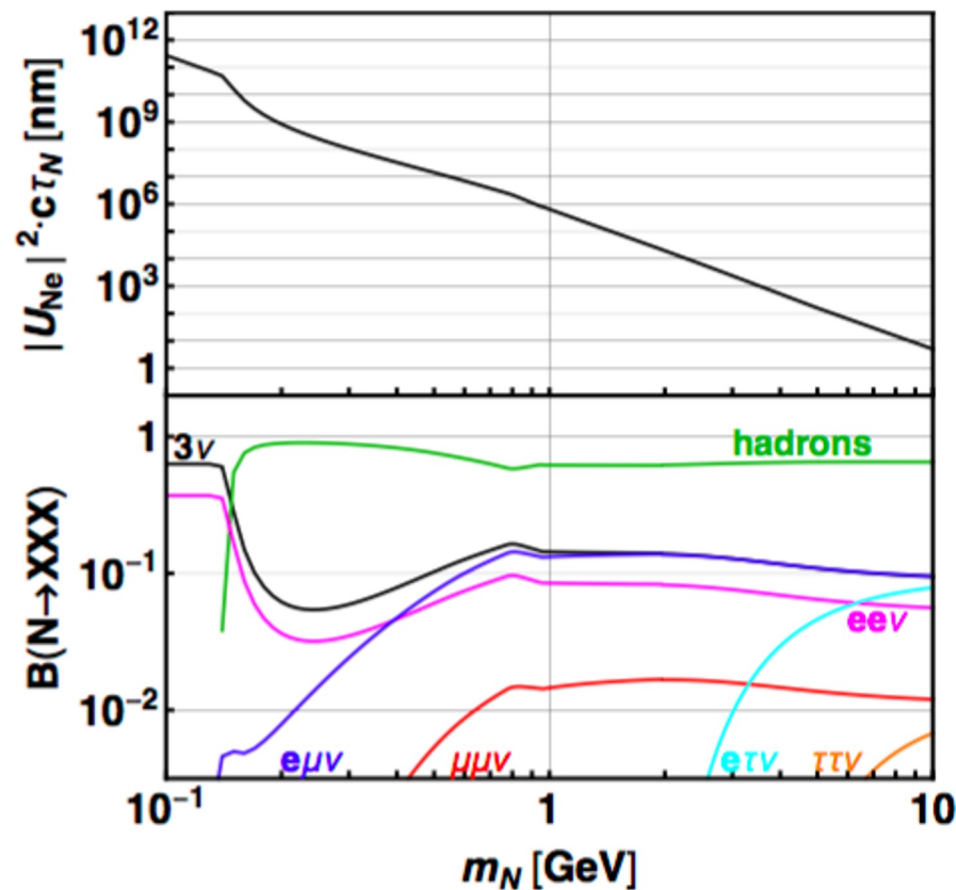
FASER/FASER2

Inelastic cross section producing for 150 fb^{-1}

$$N_{\pi^0} \approx 2.3 \times 10^{17}, \quad N_{\eta} \approx 2.5 \times 10^{16}, \quad N_D \approx 1.1 \times 10^{15}, \quad \text{and} \quad N_B \approx 7.1 \times 10^{13}$$

FASER-2

Larger detector for total integrated luminosity of 3 ab^{-1}



Higgs decays in Type I

For $L = 3000 \text{ fb}^{-1}$

$\sim 10^{11}$ W will be produced

$\sim 10^8$ H will be produced

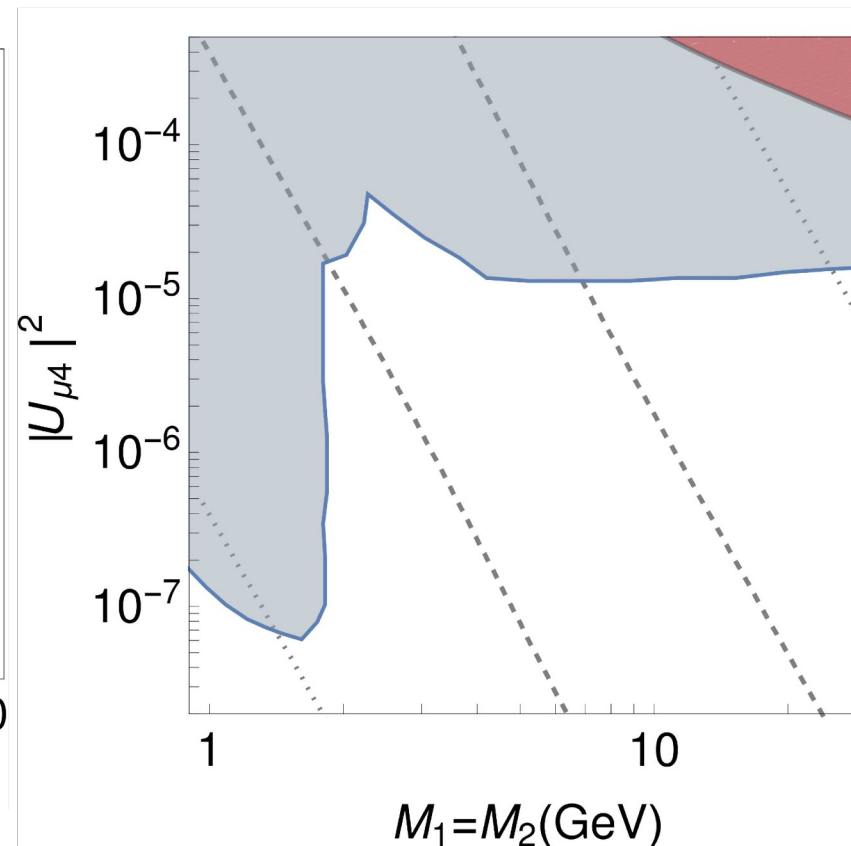
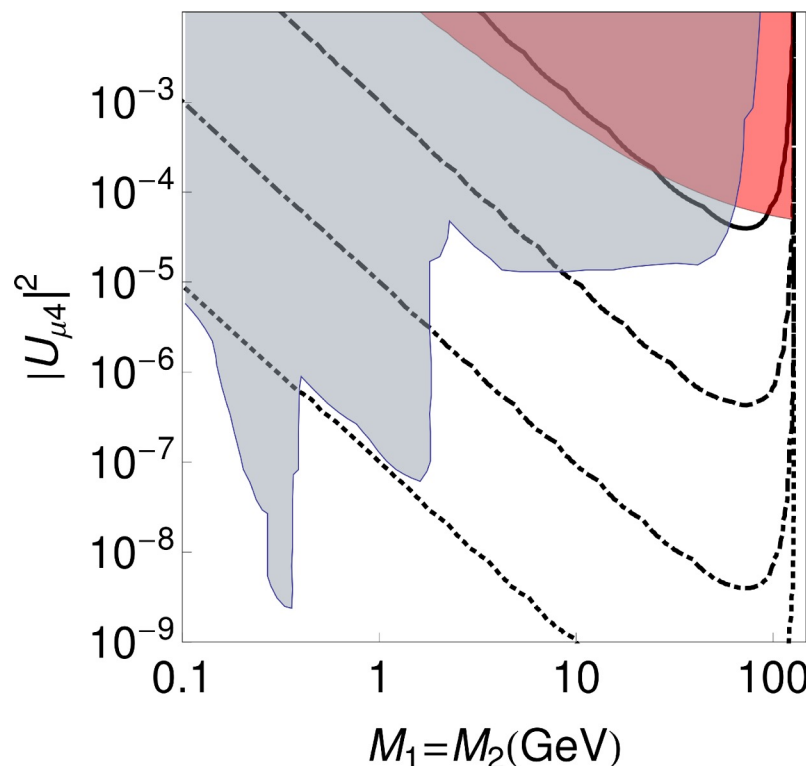
Consider ggF production mode, only 2 HNL

Directly test seesaw :

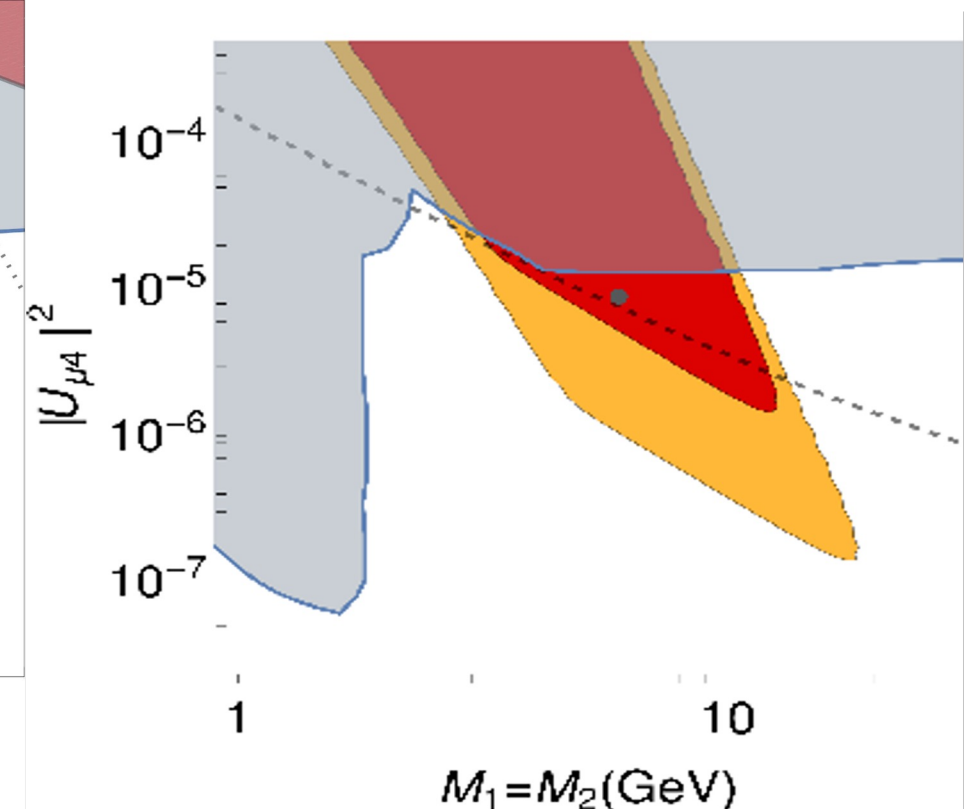
From Higgs production and decay explore heavy neutrinos with large lifetimes, which lead to a displaced vertex with leptonic and semi-leptonic decays

250 events with 300 fb^{-1}

w/o cuts

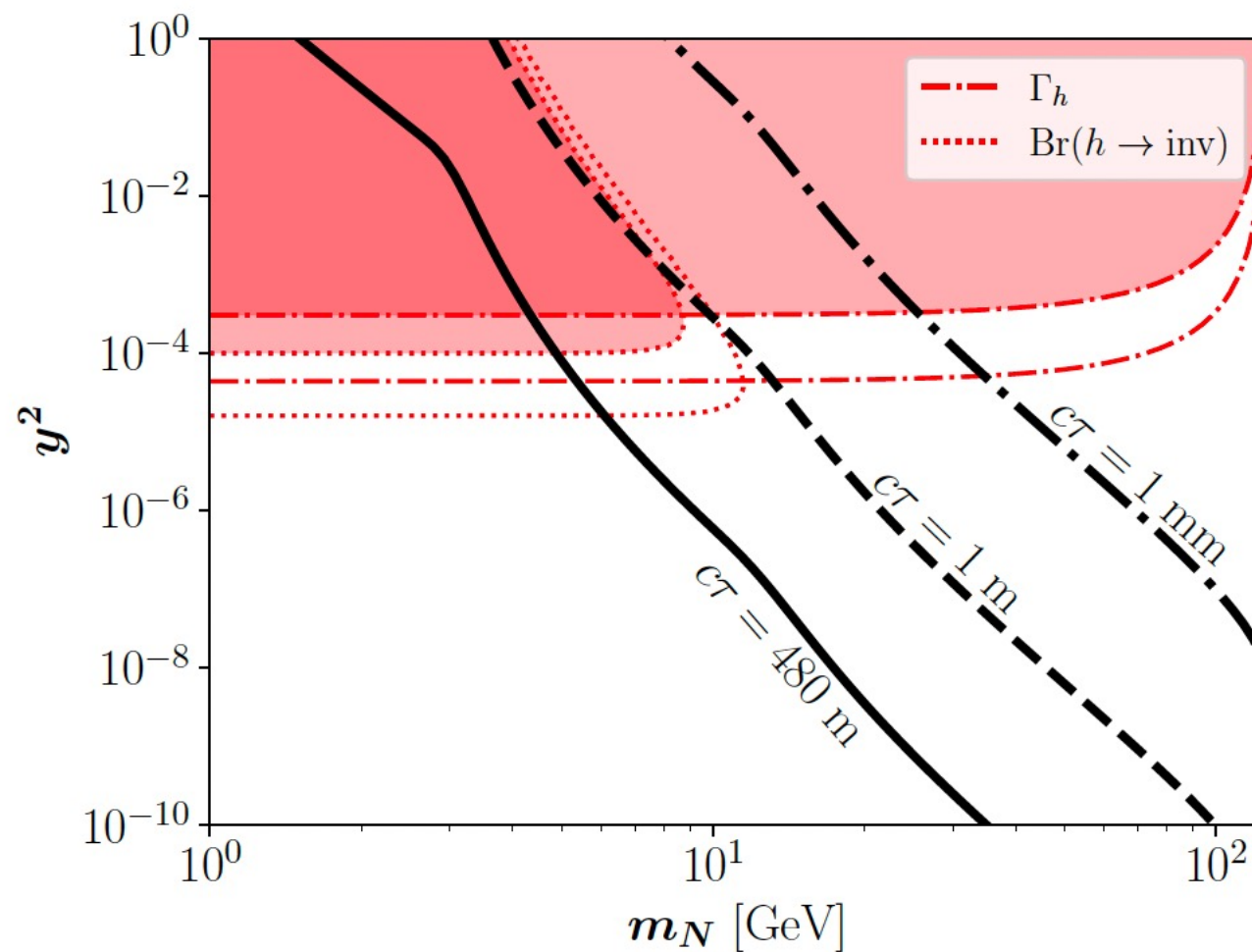


Gago, Jones, Hernández, Losada, Moreno



Higgs Decays: Only Yukawa coupling: Lifetime + BR

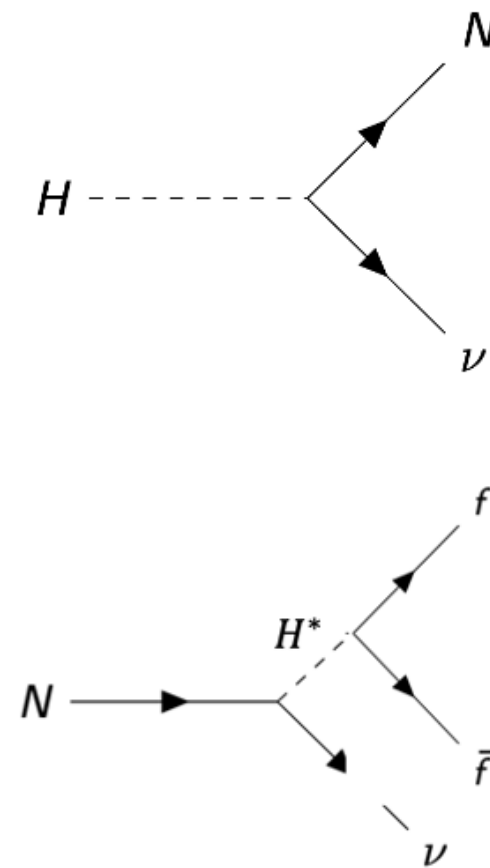
Take mixing to zero. Only have the Yukawa coupling, what are the experimental constraints and sensitivity?



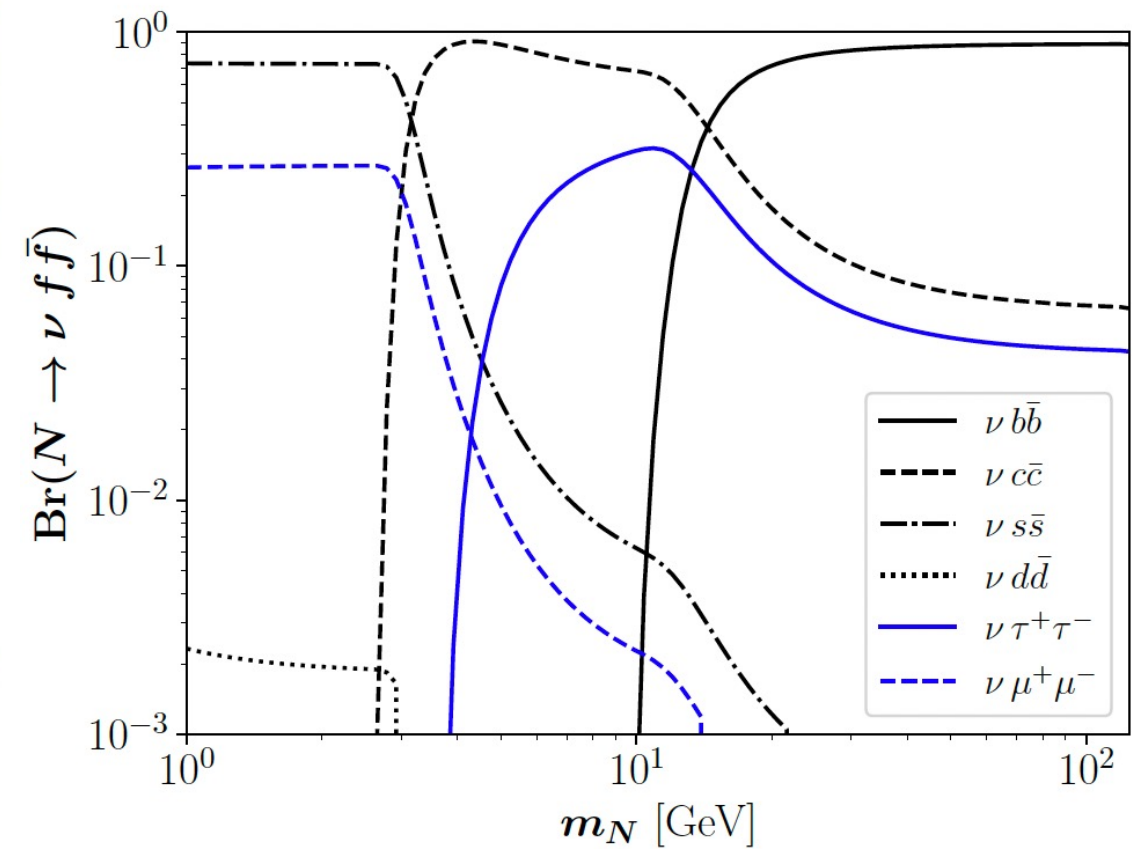
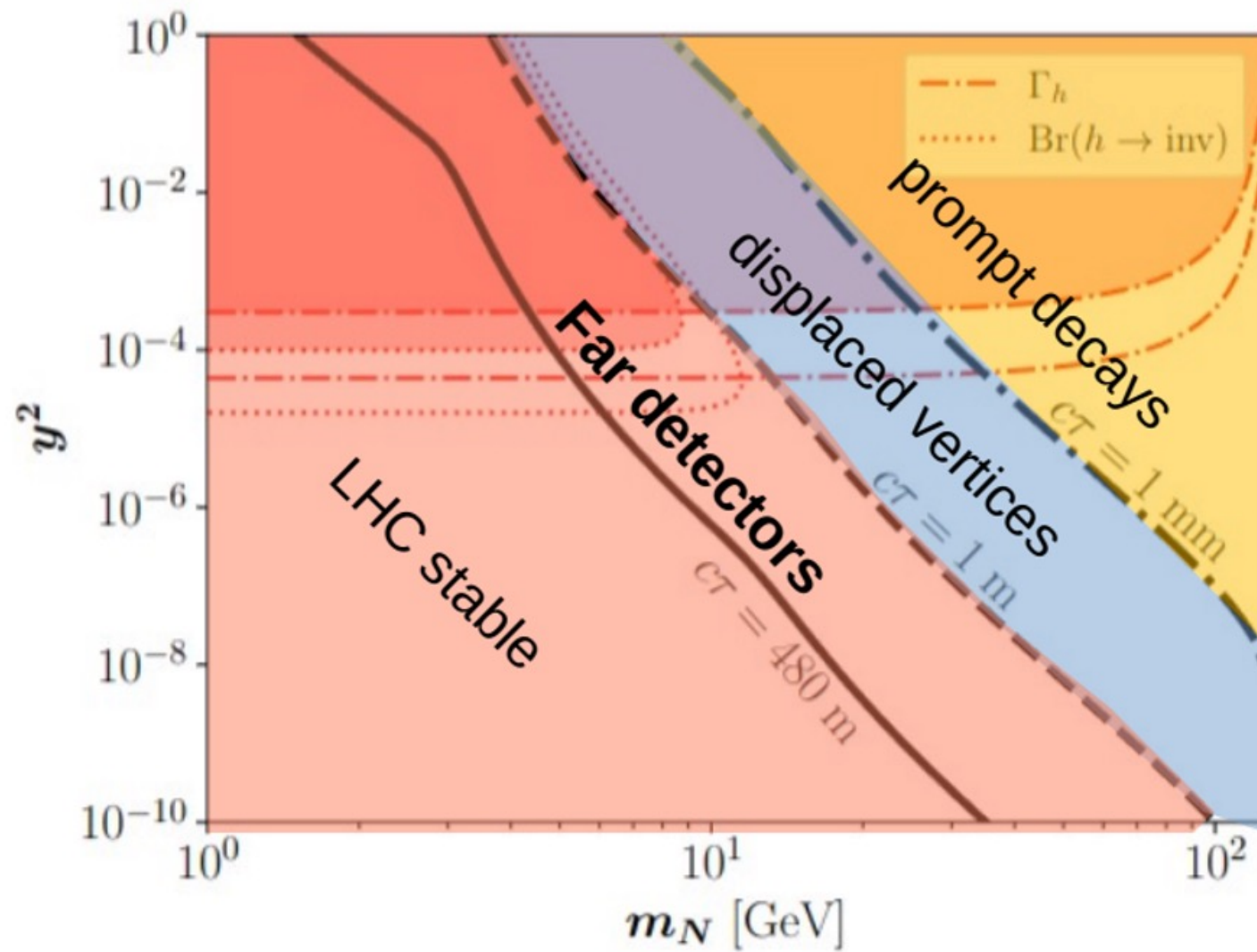
Bernal, Deka, Losada 2023

HNL Production through Higgs decays

$$\Gamma(h \rightarrow N\nu) = \frac{y^2}{8\pi} m_h \left[1 - \left(\frac{m_N}{m_h} \right)^2 \right]^2$$



Zero mixing: Lifetime+ BR

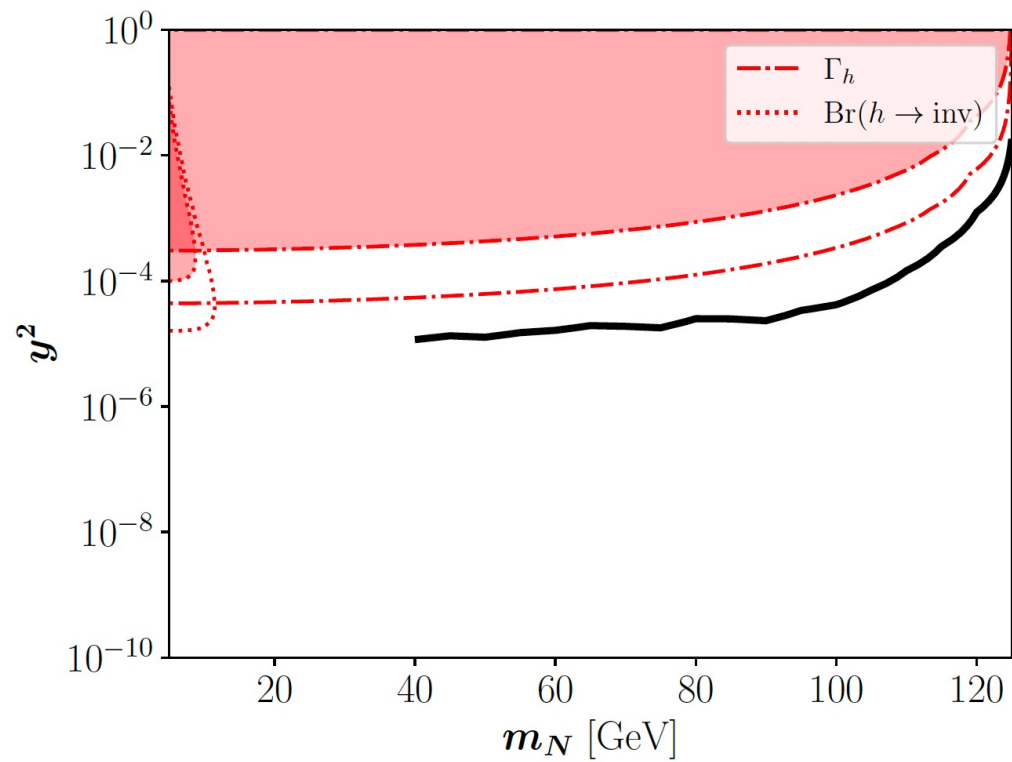


$$\Gamma(N \rightarrow \nu f \bar{f}) \propto y^2 \frac{m_f^2 m_N^5}{m_h^6}$$

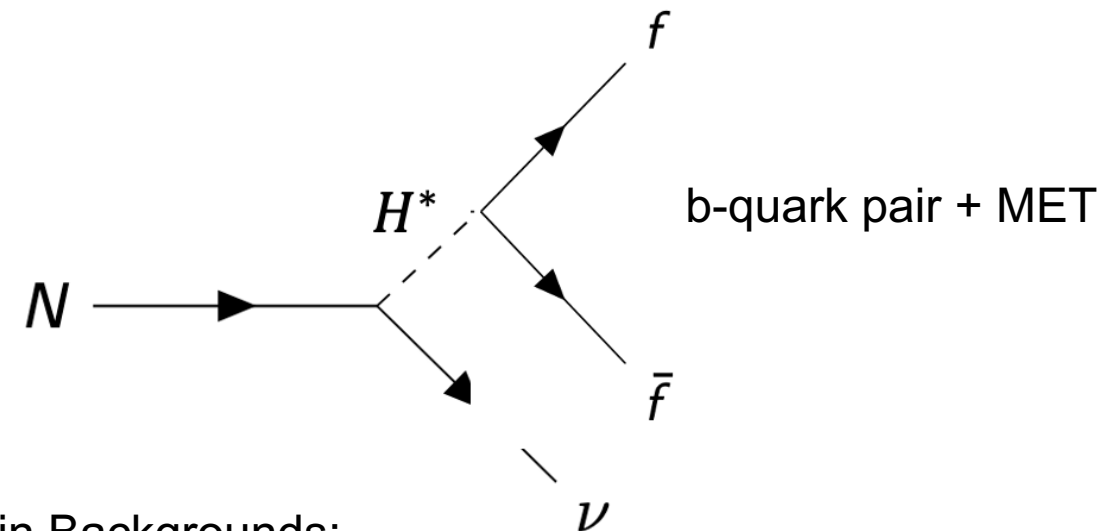
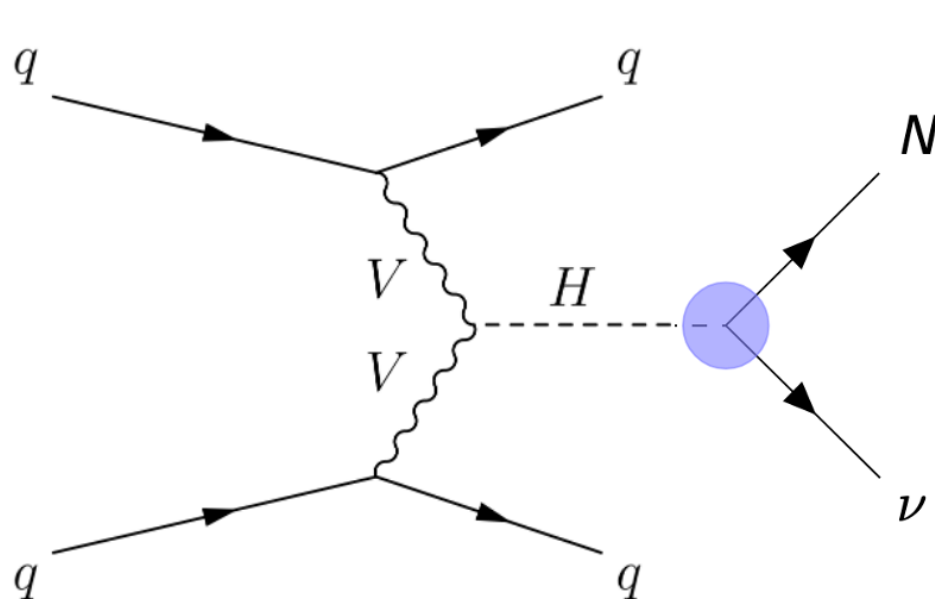
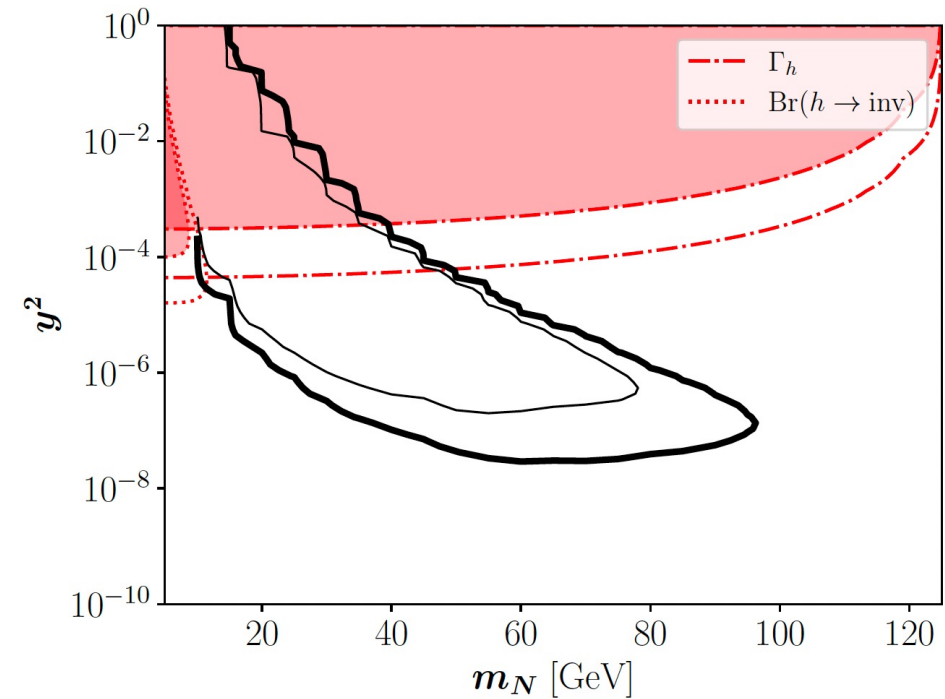
HL-LHC Results – no mixing

$$\sqrt{s} = 14 \text{ TeV and } \mathcal{L} = 3 \text{ ab}^{-1}$$

Prompt



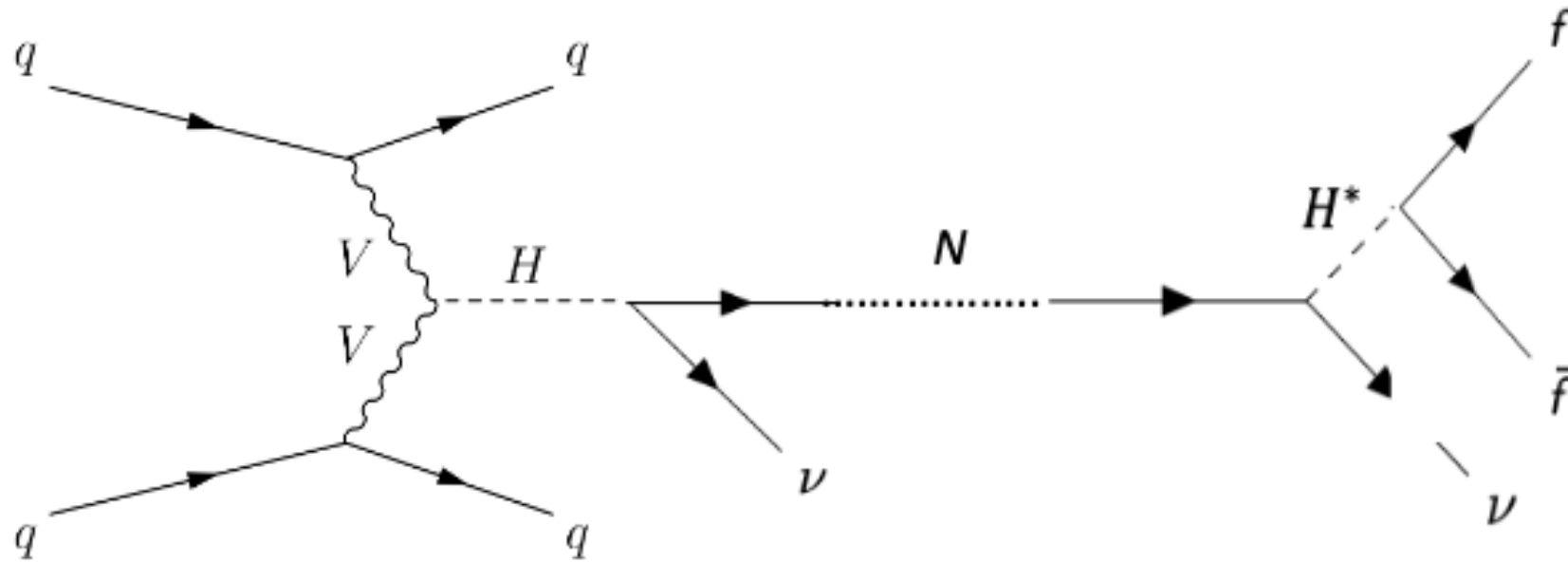
Displaced in ATLAS/CMS detector



Main Backgrounds:

$h \rightarrow b\bar{b}$, $t\bar{t}$ + jets, and $b\bar{b}$ + jets

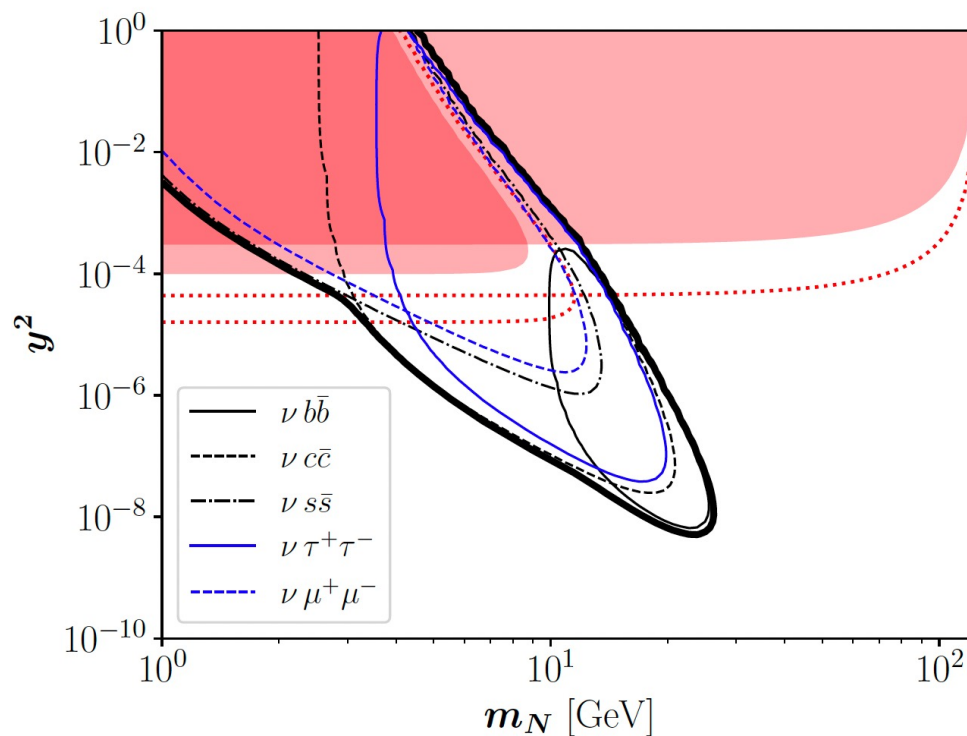
vLLP Results – no mixing



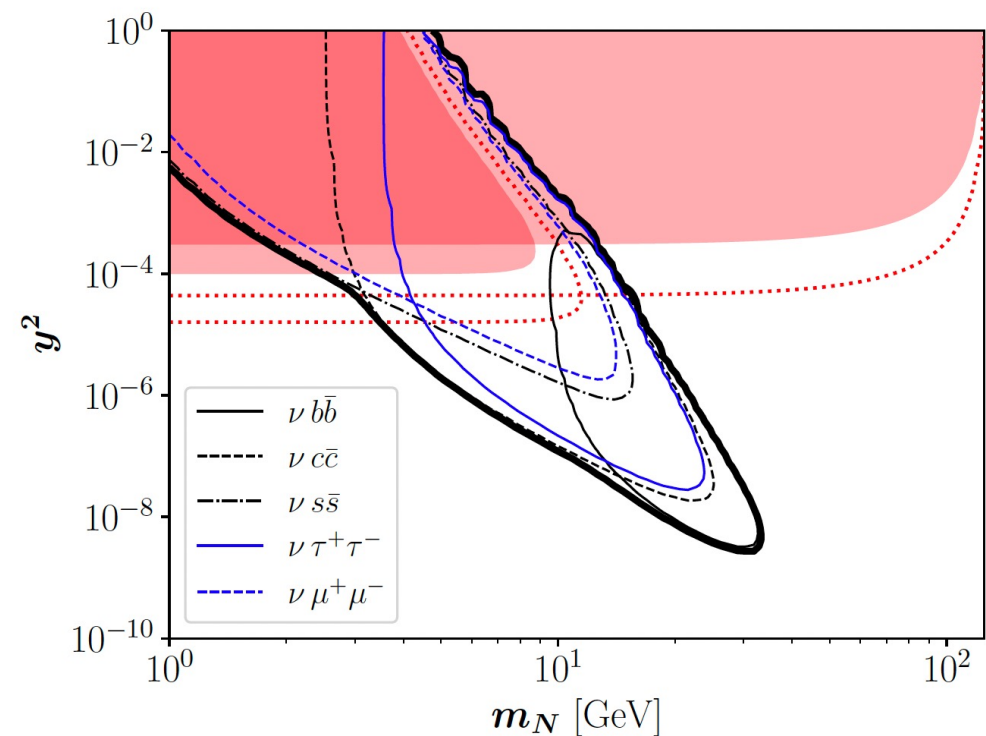
FASER-2 $\Delta = 10$ m, $R = 1$ m
Probability of HNL decay inside the detector

$$\mathcal{P} = \left[e^{-(L-\Delta)/d} - e^{-L/d} \right] \Theta[R - L \tan \theta]$$

$$d = c\tau \beta \gamma = c\tau \frac{|\vec{p}|}{m_N}$$

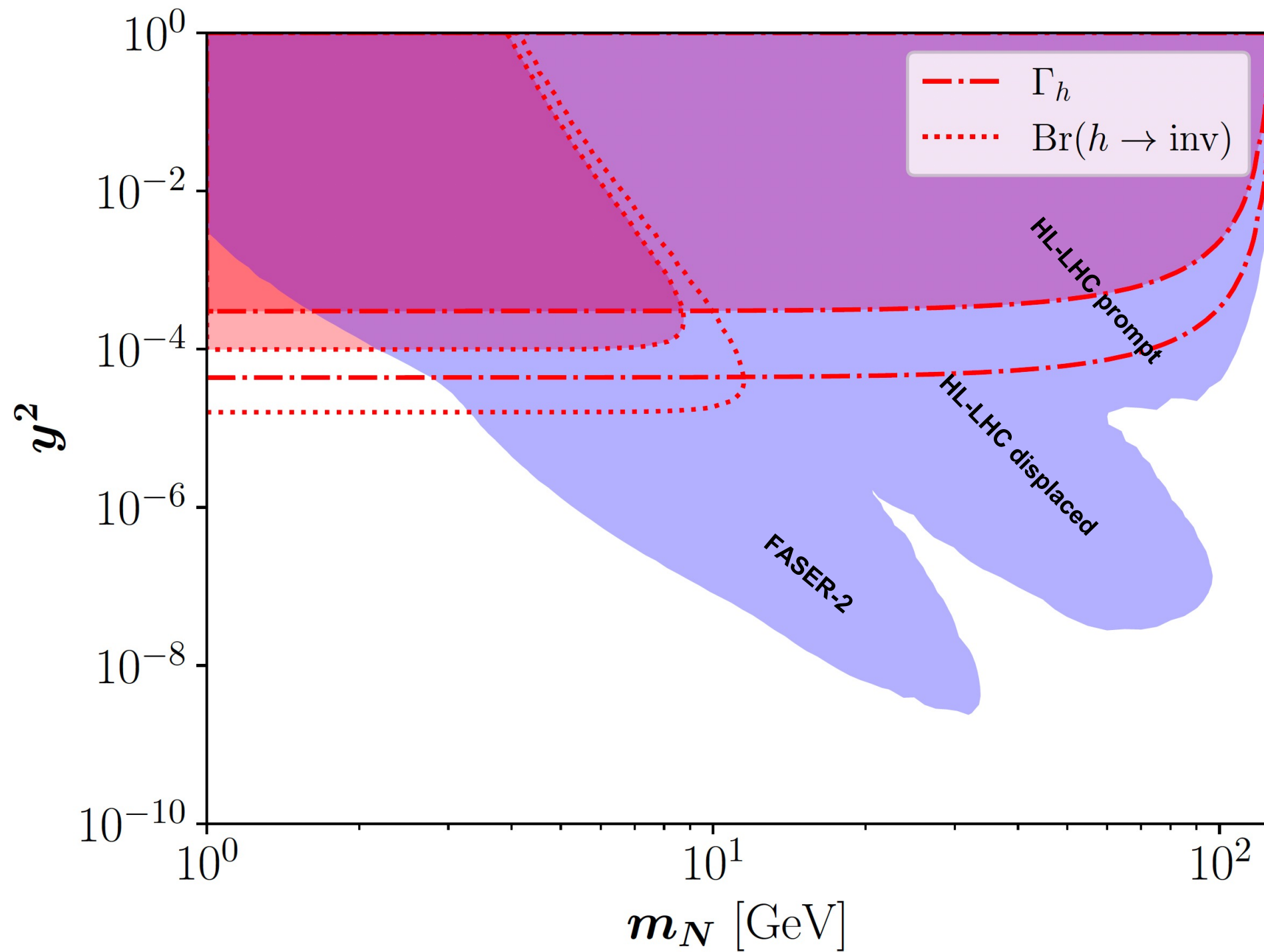


FASER-2



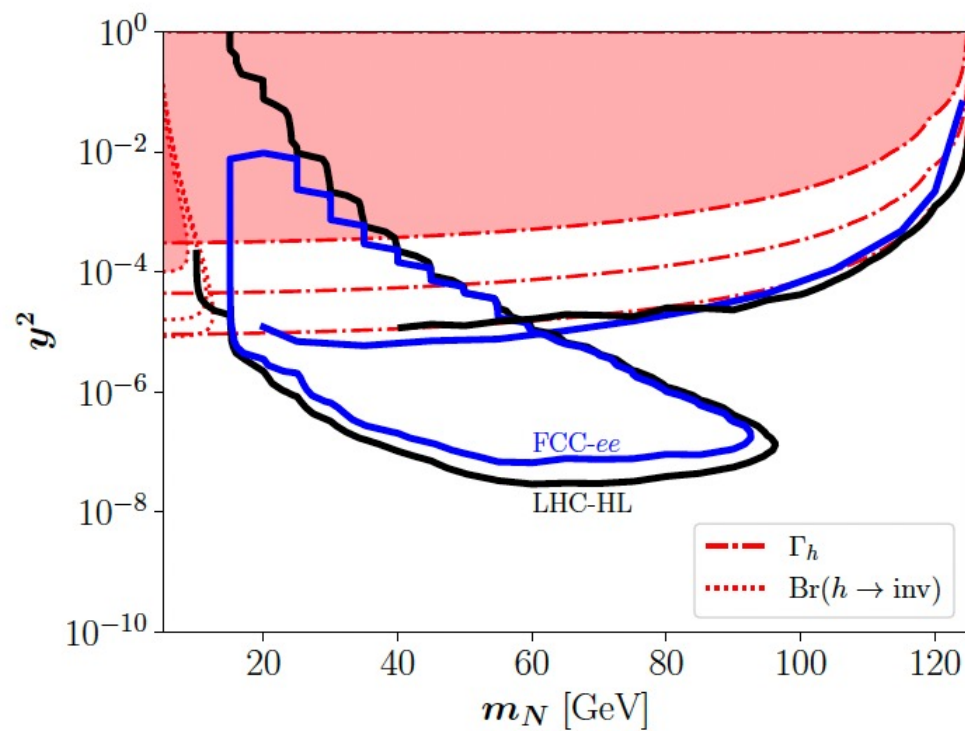
MATHUSLA

LHC Sensitivity – no mixing

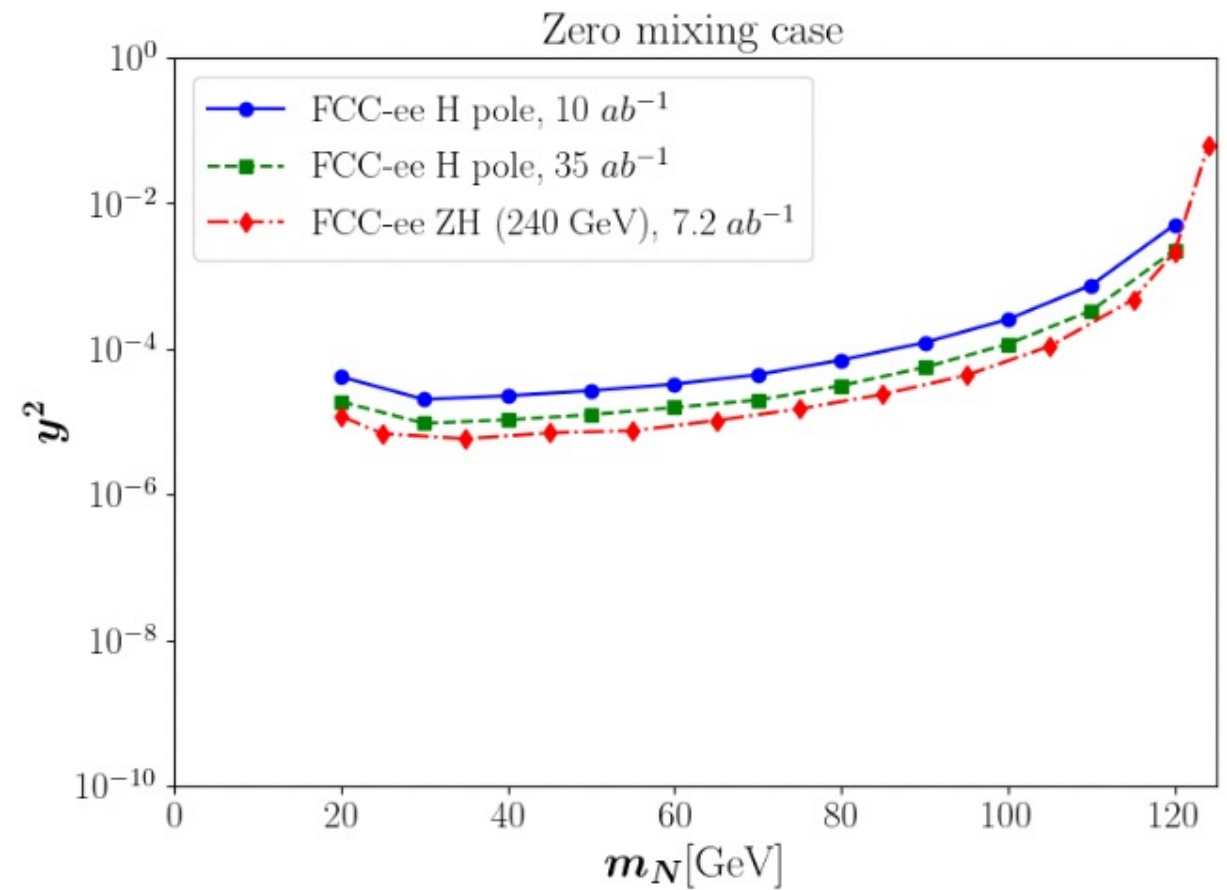


HL-LHC and FCC-ee with no mixing

$$\sqrt{s} = 240 \text{ GeV}, \mathcal{L} = 7.2 \text{ ab}^{-1}.$$

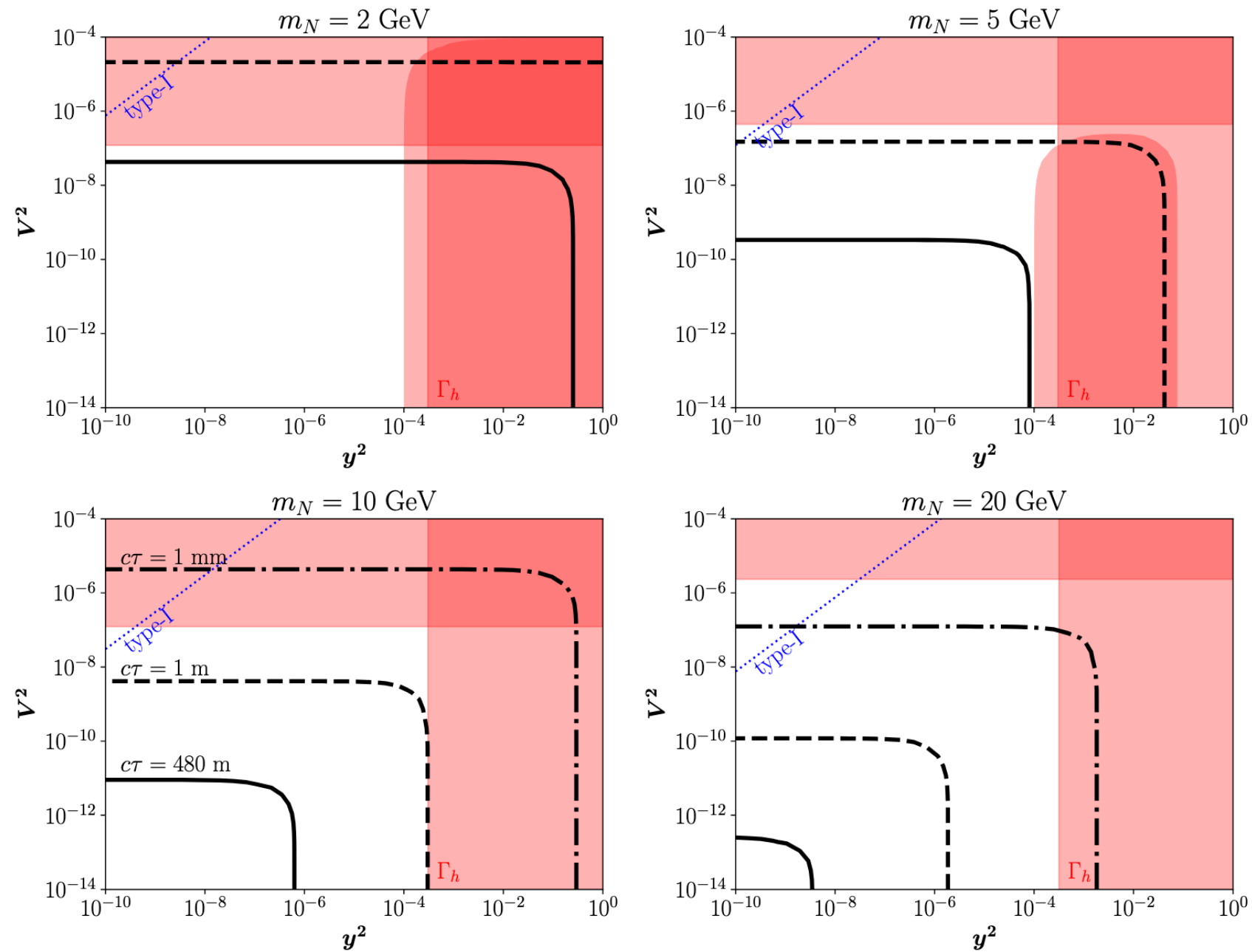


No major improvement in sensitivity with FCC-ee in this case.

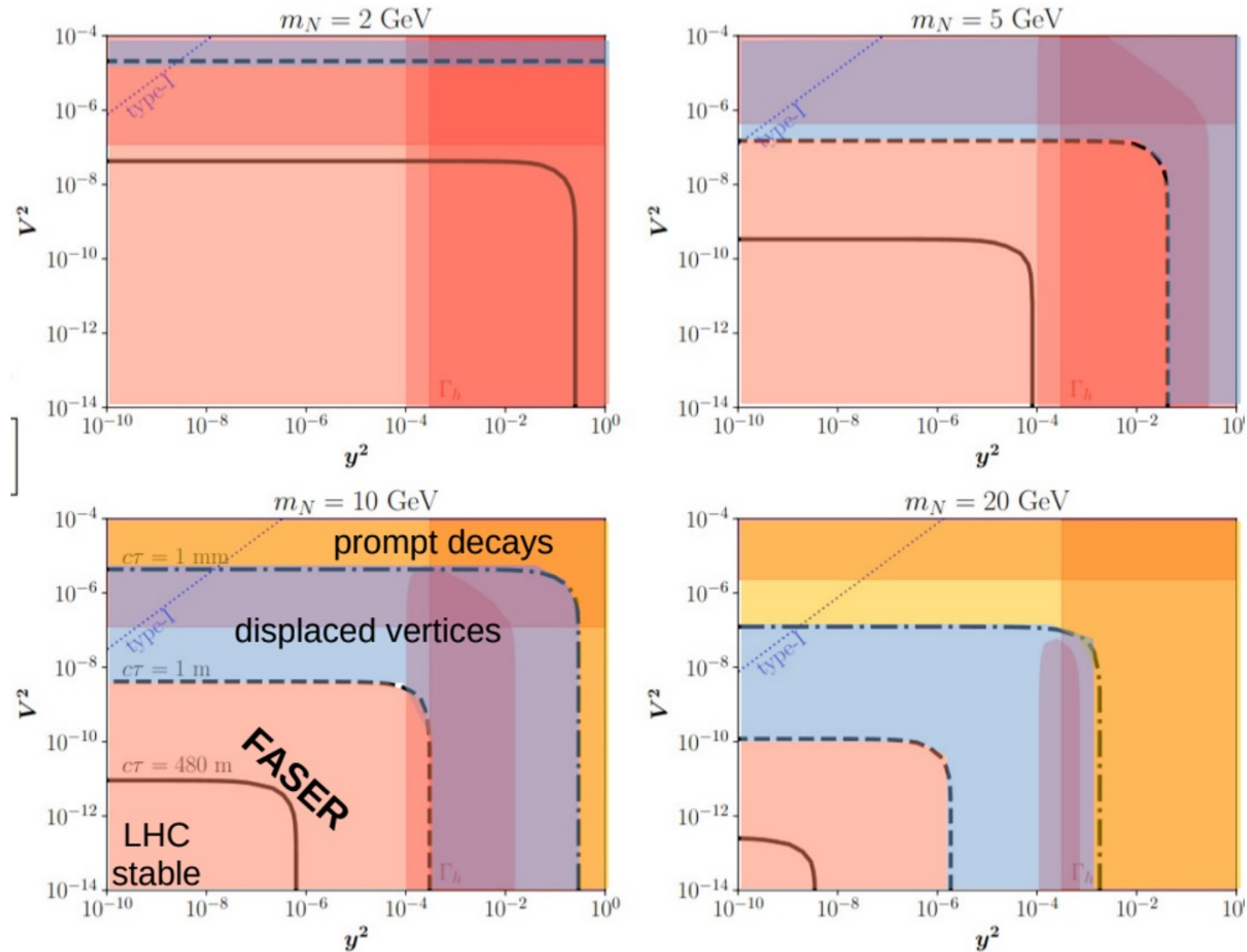


No major improvement on H pole

Yukawa + Non-zero mixing: Lifetimes



Yukawas + Non-zero mixing: Lifetimes

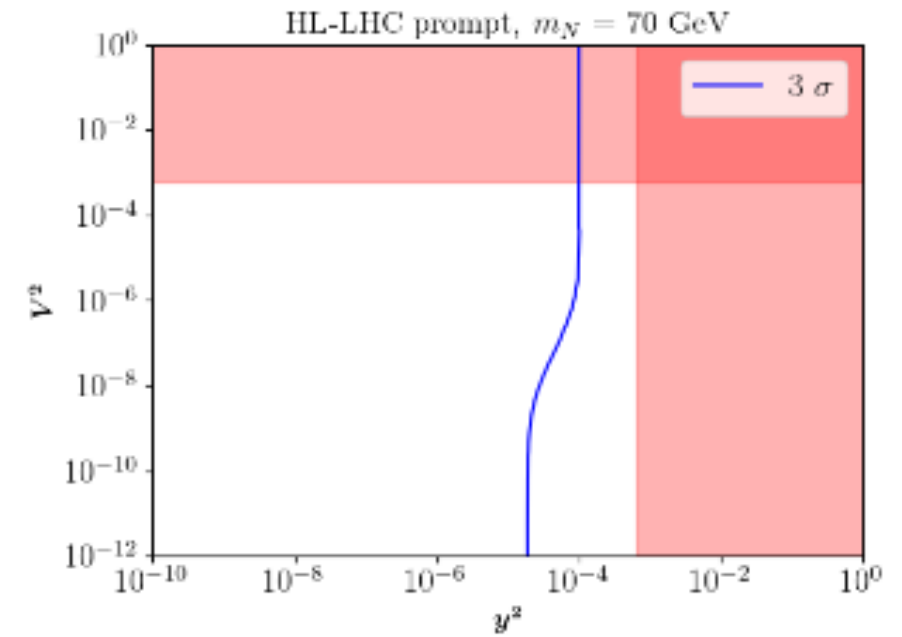
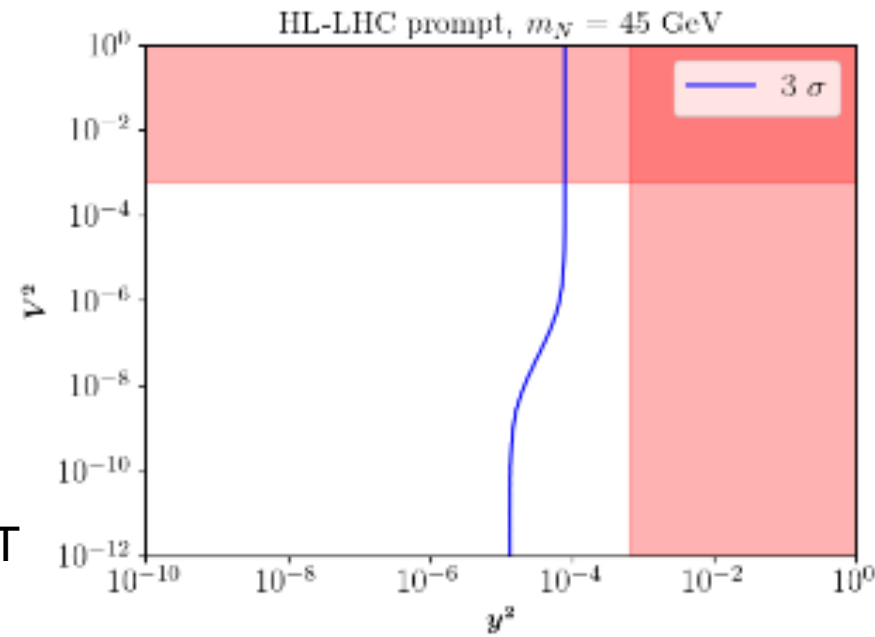


HL-LHC results with active-sterile mixing

$\sqrt{s} = 14 \text{ TeV}$ and $\mathcal{L} = 3 \text{ ab}^{-1}$

Prompt

VBF



Again use VBF topology
for HNL via Higgs decay

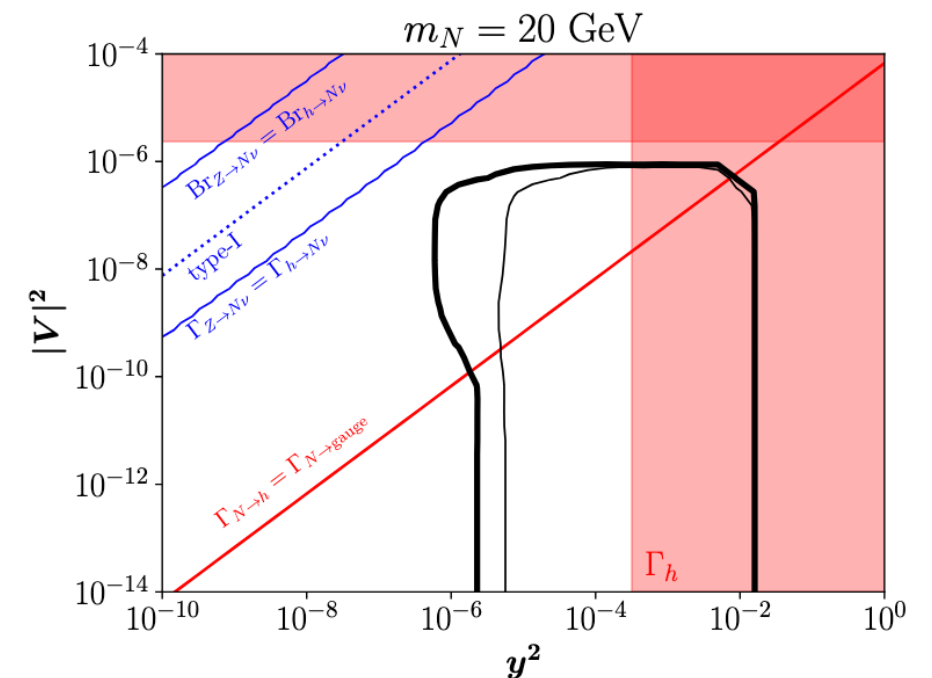
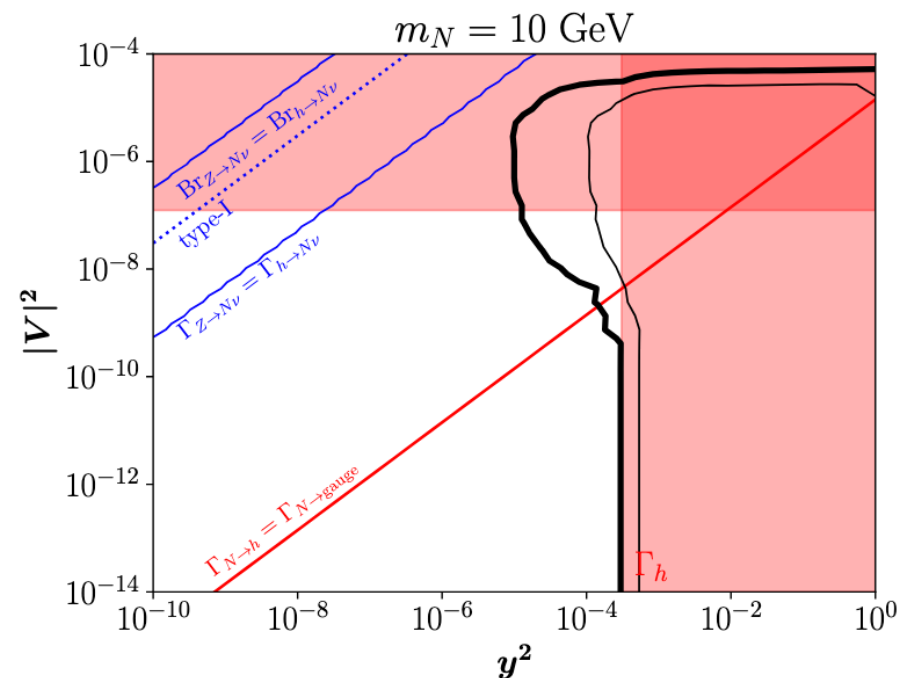
Final state: b-quark pair + MET

Main Backgrounds:

$h \rightarrow bb, tt + \text{jets}, \text{ and } bb + \text{jets}$

Displaced vertex

ggF+VBF

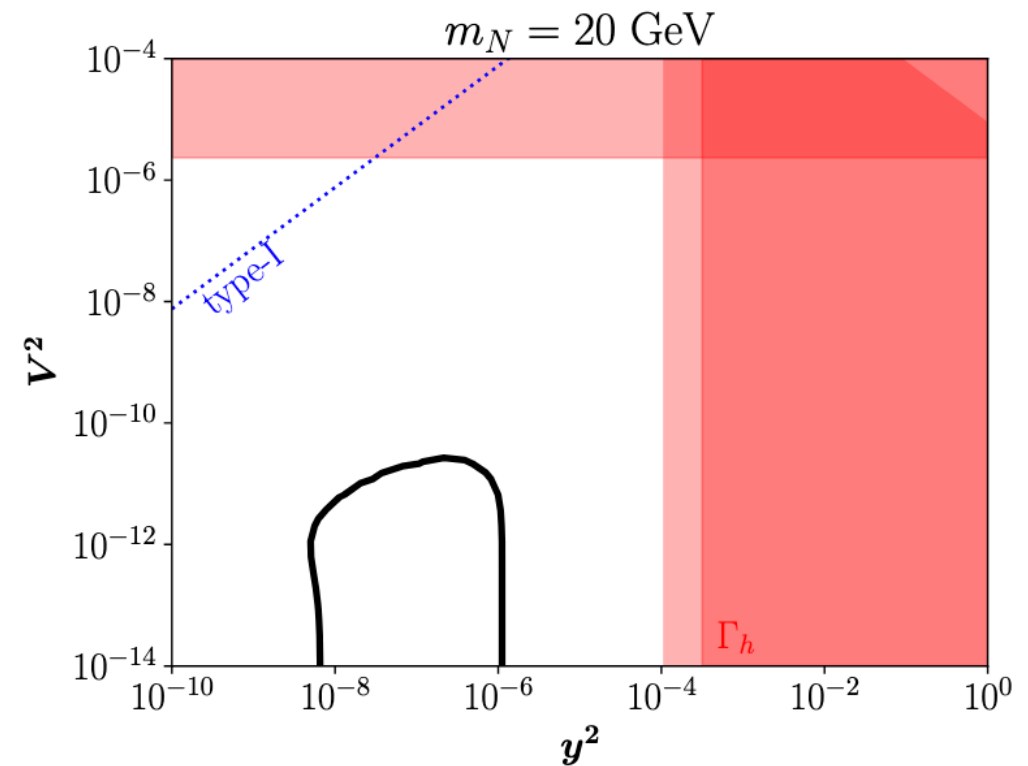
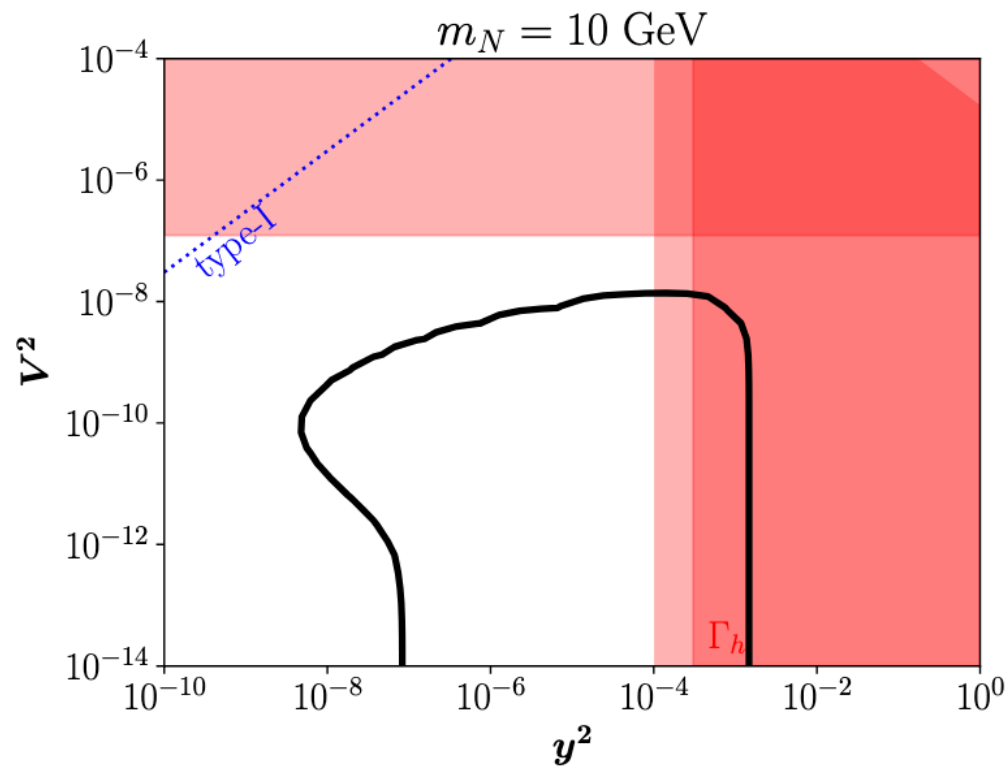
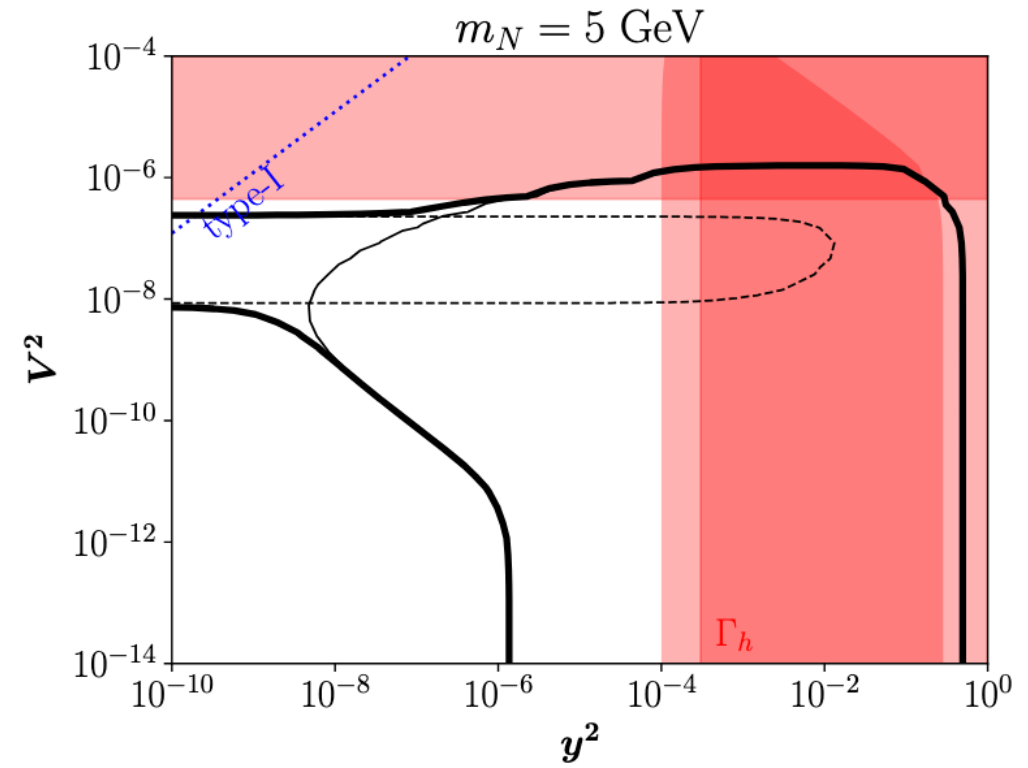
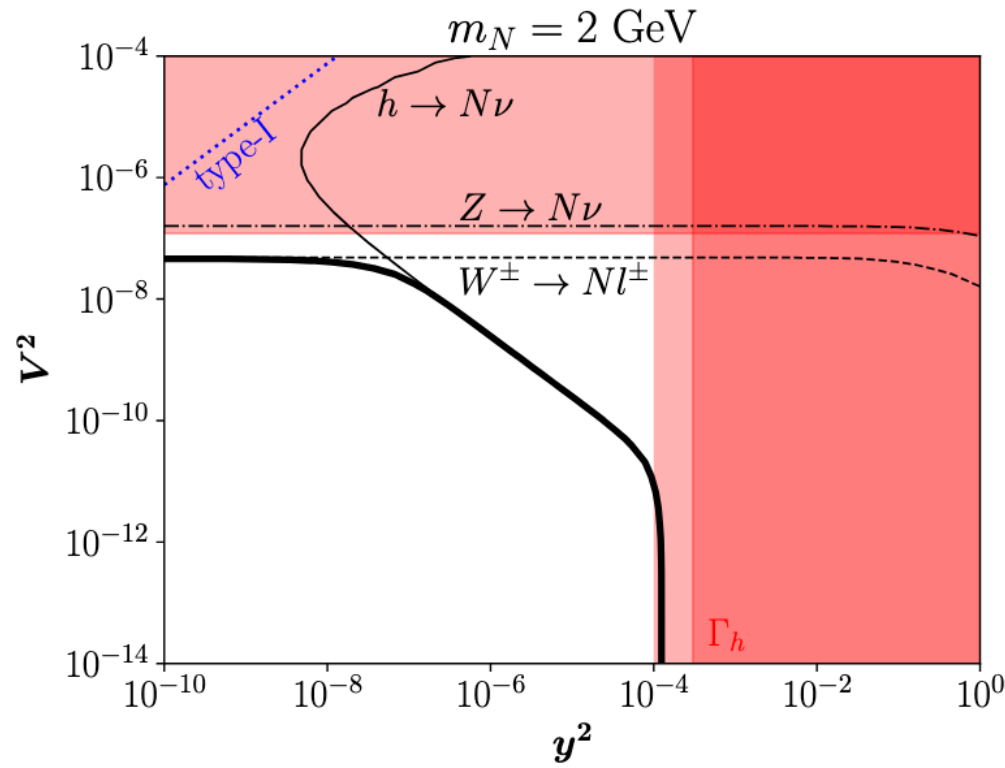


FASER

FASER-2 $\Delta = 10$ m, $R = 1$ m
Probability of HNL decay inside the detector

$$\mathcal{P} = \left[e^{-(L-\Delta)/d} - e^{-L/d} \right] \Theta[R - L \tan \theta]$$

$$d = c\tau \beta \gamma = c\tau \frac{|\vec{p}|}{m_N}$$

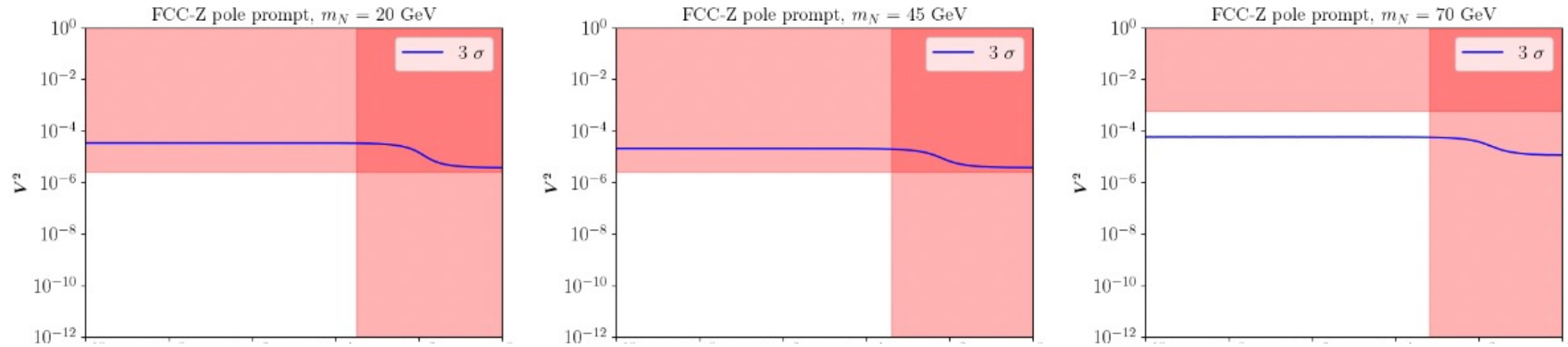


FCC with active-sterile mixing: prompt

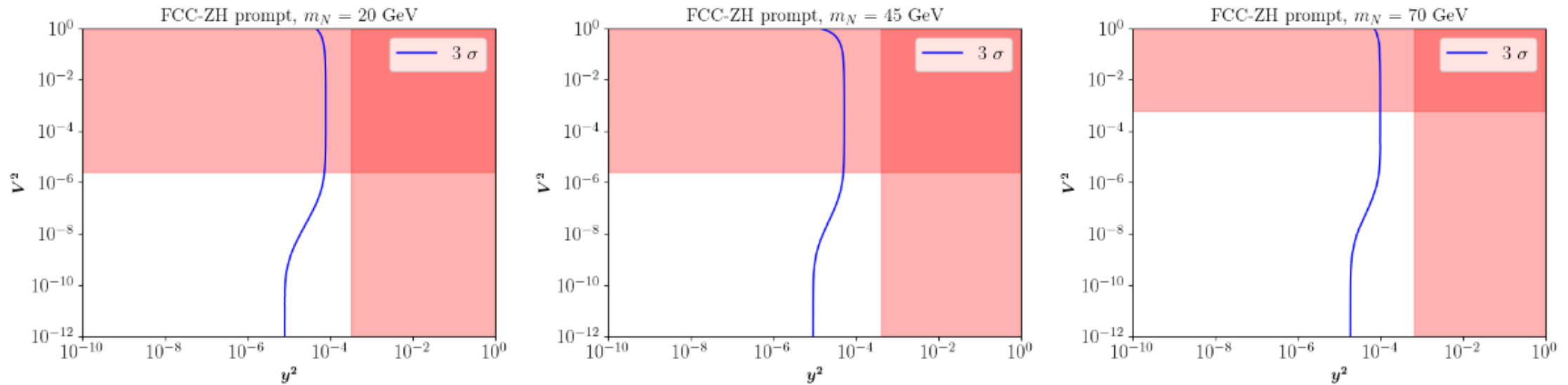
Z pole

$$\mathcal{L} = 204 \text{ ab}^{-1}$$

$$\Gamma(Z \rightarrow N\nu) = \frac{e^2 V^2}{96\pi c_W^2 s_W^2} m_Z \left[2 - 3 \left(\frac{m_N}{m_Z} \right)^2 + \left(\frac{m_N}{m_Z} \right)^6 \right]$$



ZH production

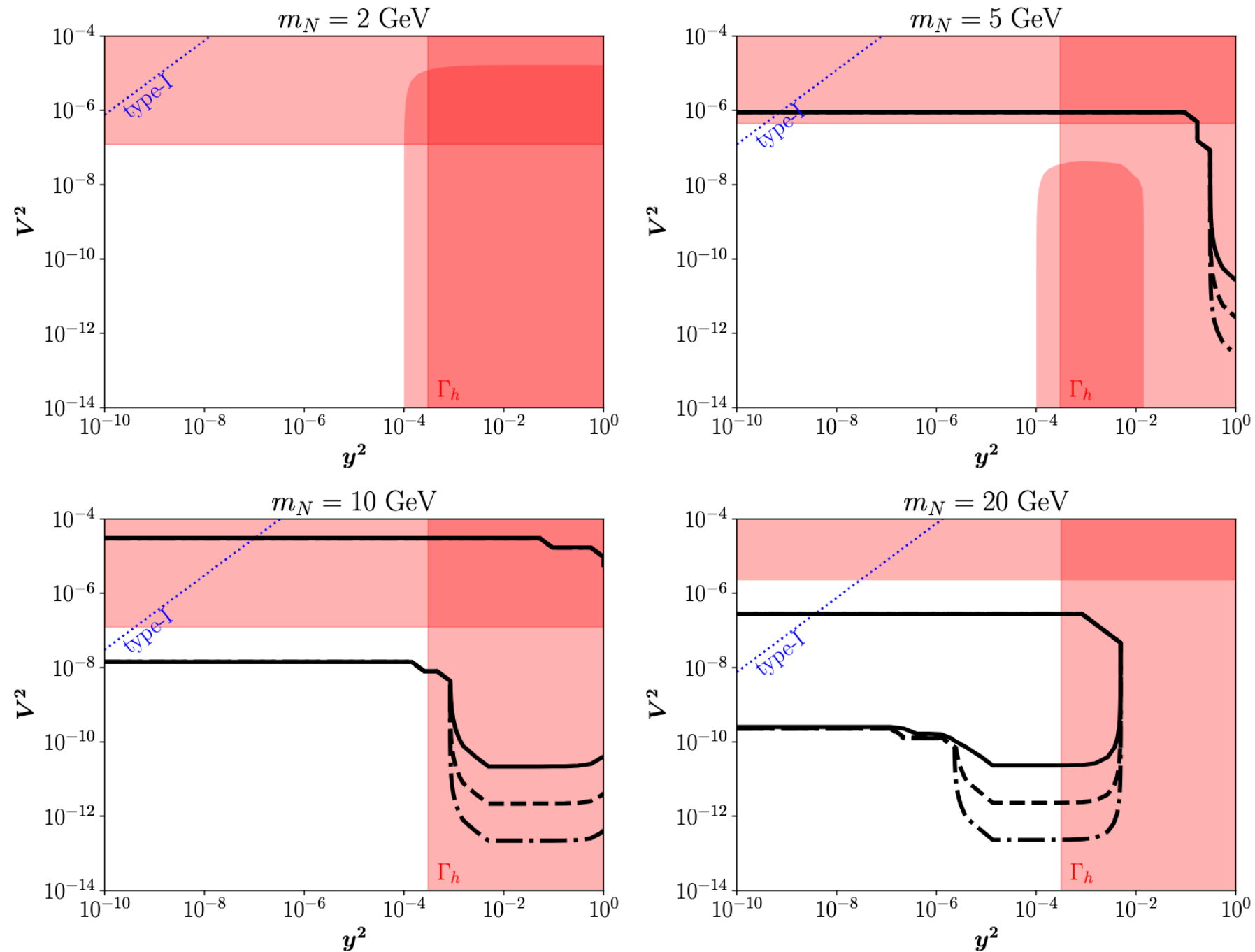


$\sqrt{s} = 240 \text{ GeV}$ and $\mathcal{L} = 2.4 \text{ ab}^{-1}$ per year, for 3 years: total integrated luminosity $\mathcal{L} = 7.2 \text{ ab}^{-1}$. $\sigma(e^+e^- \rightarrow Zh) = 0.2403 \text{ pb}$.

FCC results with active-sterile mixing

Z pole

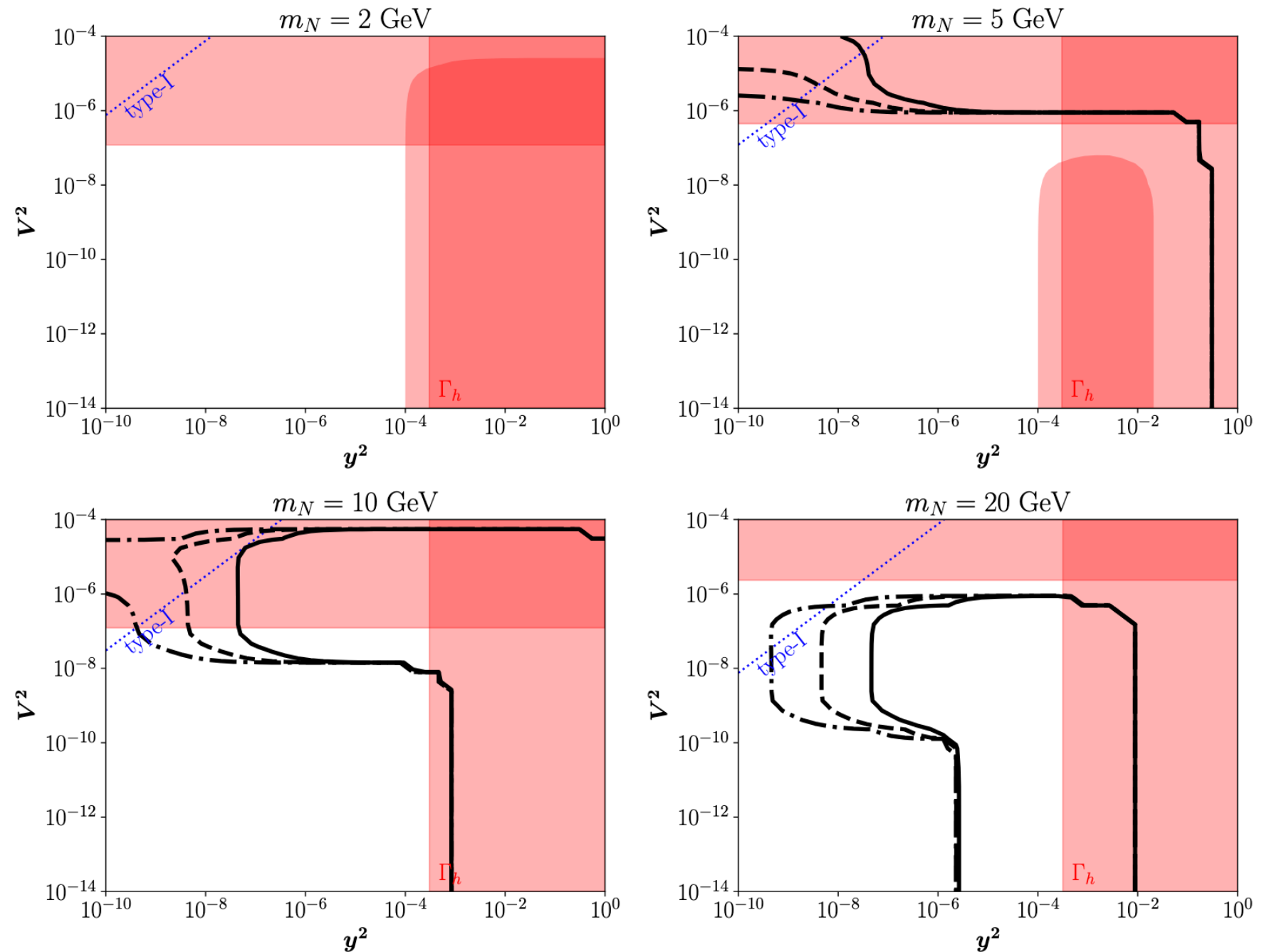
Displaced vertex



FCC results with active-sterile mixing

Combined Z+H @ 240 GeV center of mass energy

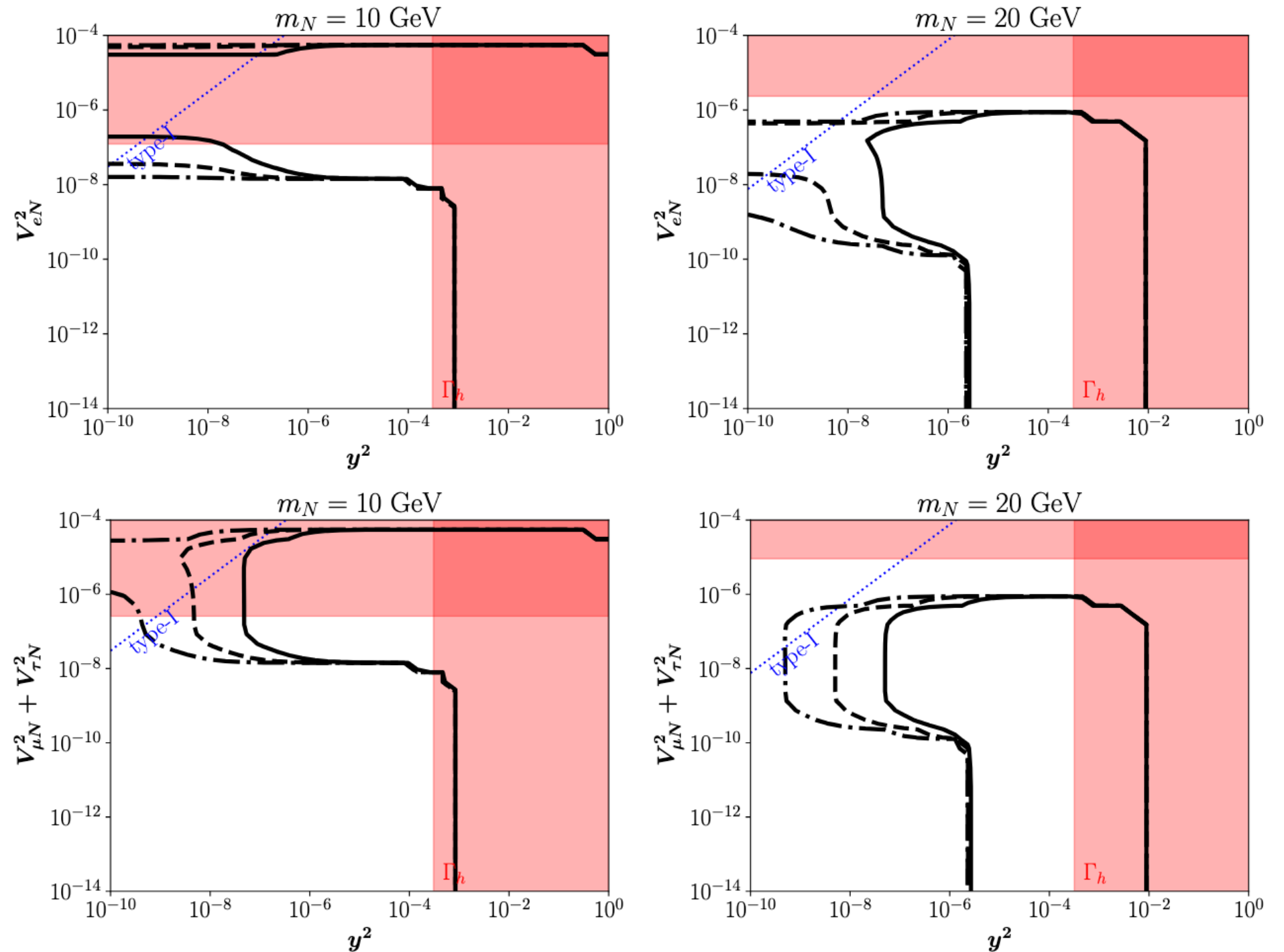
Displaced vertex



FCC-ee with active-sterile mixing

Addition of t channel W exchange,
Gives enhancement for the electron channel

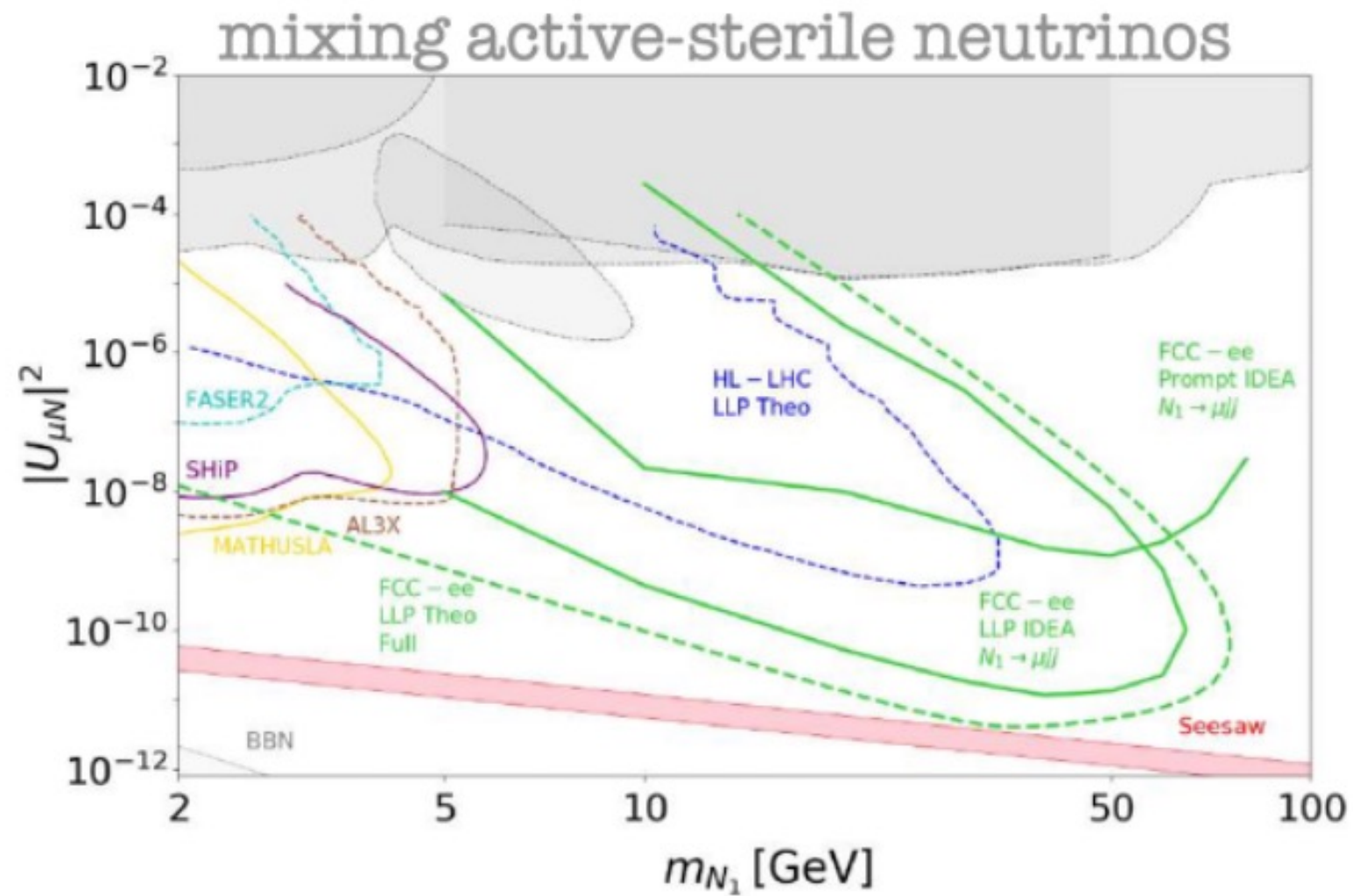
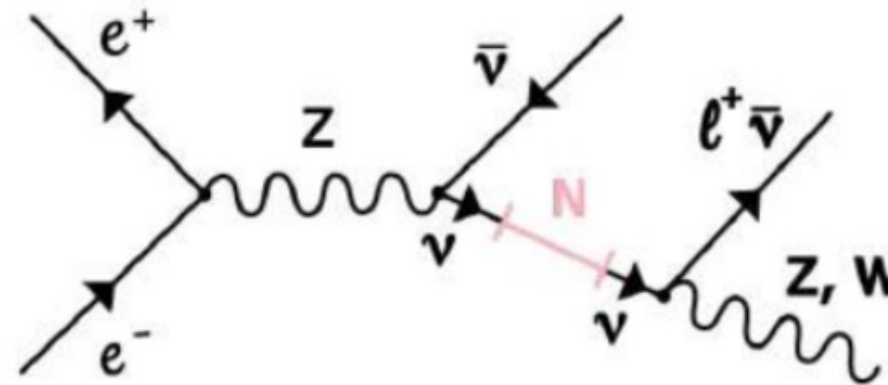
Displaced vertex



FCC-ee

<https://indico.cern.ch/event/1307378/contributions/5720989/attachments/2789031/4879011/Grojean.pdf>

Direct observation
in Z decays
from LH-RH mixing

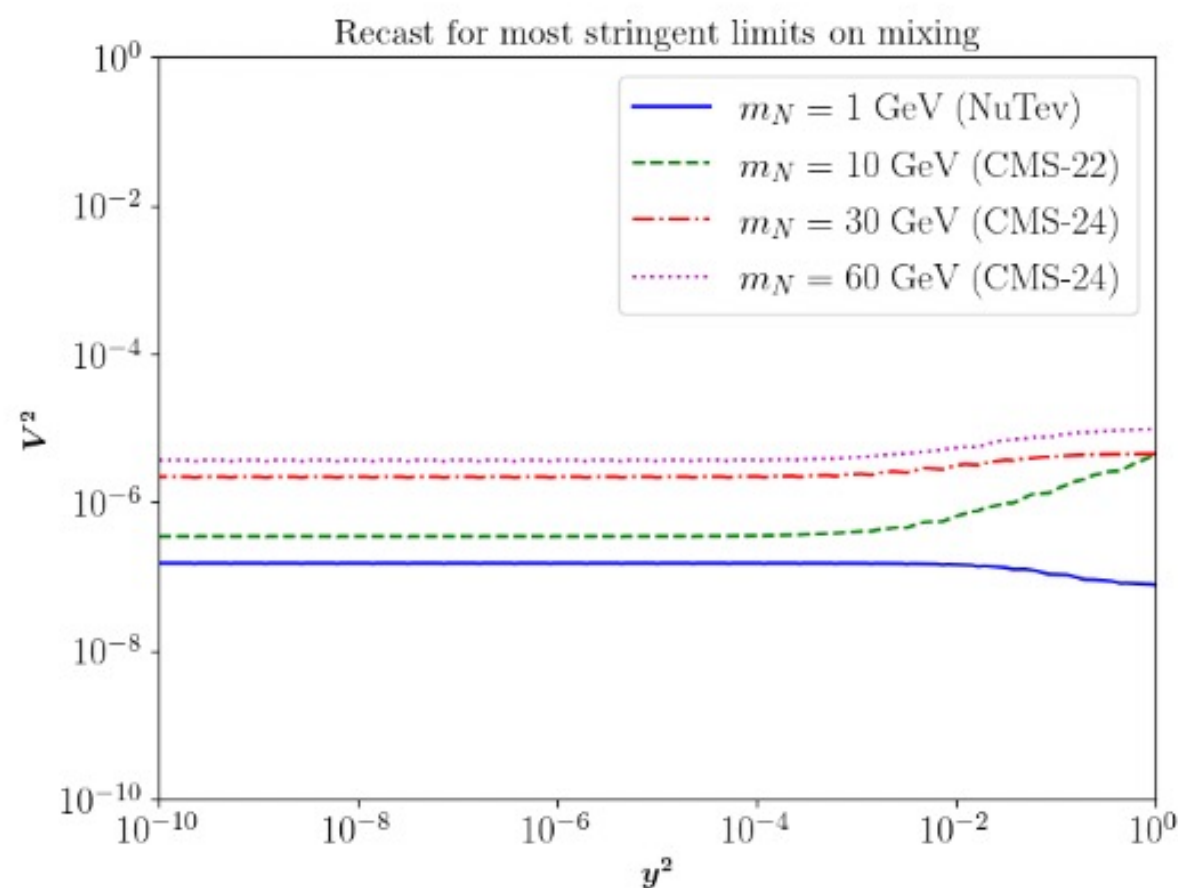
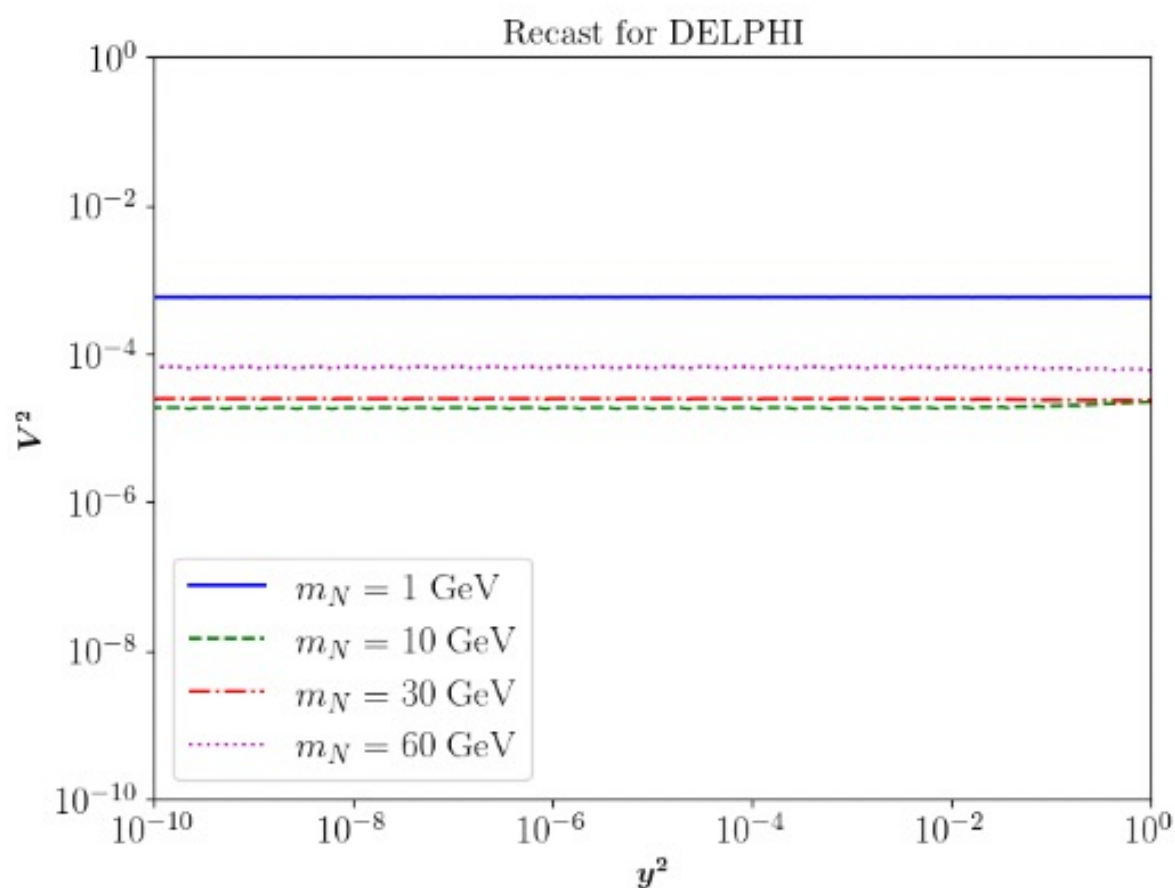


Conclusions

1. There are compelling motivations for HNLs to solve fundamental questions related to neutrino masses, matter-antimatter asymmetry, etc.
2. In the regime of HNLs MeV-TeV a clear and more complete picture of experiments (running or proposed) can further explore in depth parameter space.
3. In the case of no active-sterile mixing there is not much of an increase in sensitivity from Higgs physics processes in the production and decay of HNLs when comparing HL-LHC and FCC.
4. We've shown when you are directly sensitive to the HNL Yukawa. As soon as non-zero mixing is turned on it quickly dominates the sensitivity of the NP search.
5. With FCC a much increased sensitivity to active-sterile mixing is obtained at the Z pole.
6. Current and future constraints on the Higgs width provide the most relevant constraints on the HNL Yukawa coupling in particular for the case of zero mixing.

Thank you

Turn on mixing: Recasting previous results



$$V^2(m_N) \cdot BR_{N \rightarrow \nu f \bar{f}}^{(Z^*/W^*/h^*)}(V^2, y^2, m_N) = V_{\text{exp}}^2(m_N) \cdot BR_{N \rightarrow \nu f \bar{f}}^{(Z^*/W^*)}(m_N),$$

| Experiment | final states |
|------------|---|
| NuTeV | Production through W (looks at electrons or muons from HNL) |
| DELPHI | Production through Z (looks at jets coming from HNL) |
| CMS22 | Production through W (looks at electrons or muons from HNL) [90] |
| CMS24 | Production through W (looks at electrons, muons or taus from HNL) [136] |