



# CYGNO status and plans Feb 2025





# Collaboration Meeting 2024

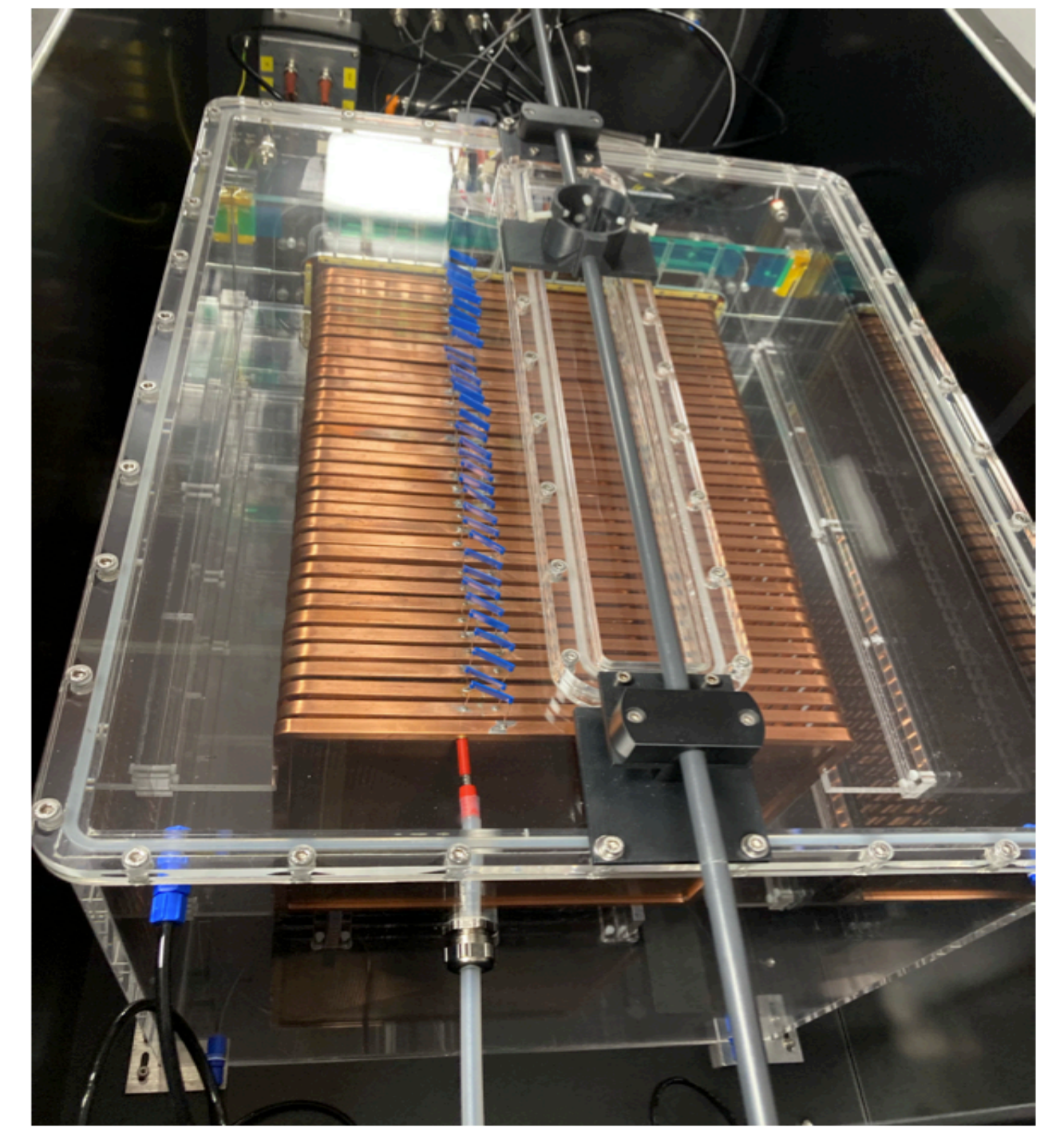
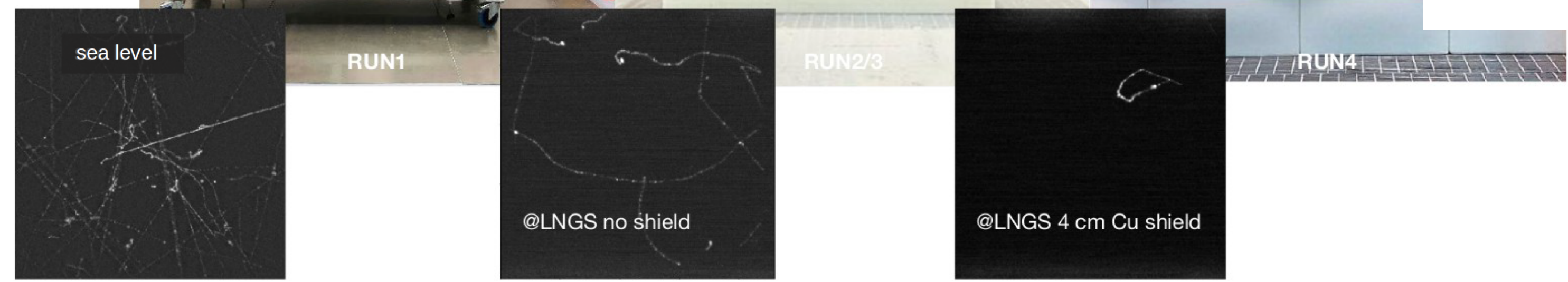
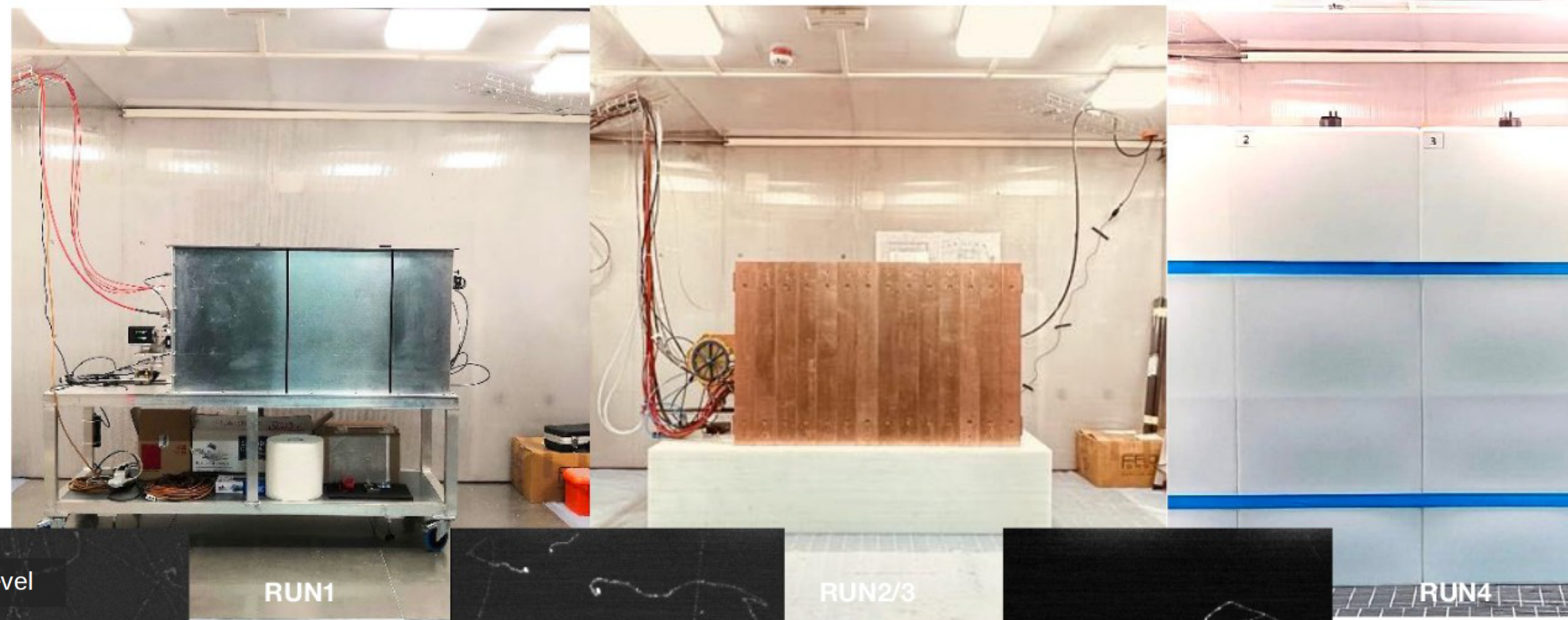
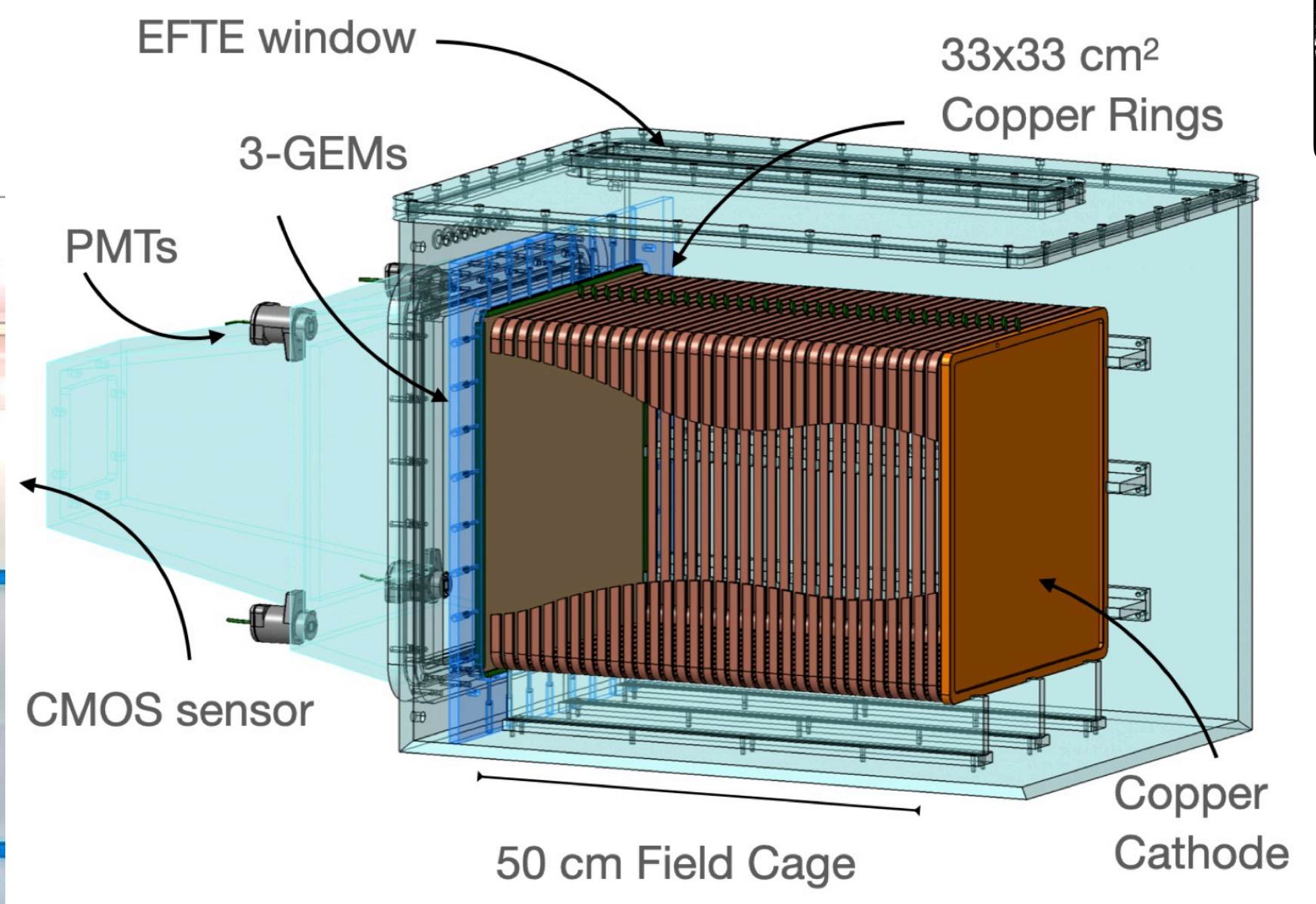
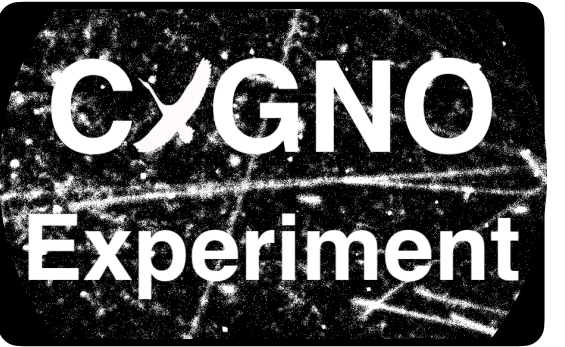


Three days of discussion about status and plans of:

- LIME Data **acquisition** and **analysis**;
- CYGNO04 **drawing** and **simulation**;
- Component **validation**;
- Project management and **financial status**;



# LIME: RUNs 1-5



- testing and optimisation of **ancillary systems** (HV, Gas, Slow Control, DAQ, Reco, Monitoring);
- development and optimisations of data **analysis algorithms** and performance study;
- development and validation of **detector** and **background simulation**;



# LIME: RUNs 1-5



	Time slot	Number of pictures	Event rate	Number of events
<b>RUN 1: No-shielding</b>	3 Nov 2023 - 15 Dec 2023	$4 \cdot 10^5$	35 Hz	$4 \cdot 10^6$
<b>RUN 2: 4 cm Cu shielding</b>	15 Feb 2023 - 15 March 2023	$4.5 \cdot 10^5$	3.5 Hz	$5 \cdot 10^5$
<b>RUN 3: 10 cm Cu shielding</b>	5 May 2023 - 16 Nov 2023	$1.6 \cdot 10^6$	1.5 Hz	$7.3 \cdot 10^5$
<b>RUN 4: 10 cm Cu + 40 cm water shielding</b>	30 Nov 2023 - 31 March 2024	$2 \cdot 10^6$	1.0 Hz	$6 \cdot 10^5$
<b>RUN 5: 10 cm Cu shielding (neutron flux measurements)</b>	17 May 2024 - 1 Dec 2024	$12 \cdot 10^6$	1.5 Hz	$5.4 \cdot 10^6$

## Special data takings

<b>AmBe for Nuclear Recoils</b>	2-4 Aug 2023	$2 \cdot 10^5$	0.04 Hz of NR	$2.5 \cdot 10^3$ NR
<b><math>^{241}\text{Am}</math> for Electron Recoils</b>	7-16 Nov 2023	$7 \cdot 10^5$	50 Hz	$10^6$
<b>AmBe for Nuclear Recoils</b>	5-15 Dec 2024	$6 \cdot 10^5$	0.04 Hz of NR	$7.0 \cdot 10^3$ NR



# LIME PLANS



With the **end of RUN4** in March 2024, LIME data taking for **PHASE\_0** is considered closed with **3 months delay** (including the delayed start due to civil works) in over **27 months of operations**;

The water shielding was then removed and in **May 2024** and **RUN5** has been taken for 6 months to measure the **flux of underground neutrons**, in the framework of PRIN “**Zero Radioactivity on future experiments**”;

Good exercise to **tune and test simulation and analysis algorithms** for **NR**, propaedeutical to DM search;

In **December 2024** a 10 days long **AmBe campaign** was performed

Description	Group	25-11-24	02-12-24	09-12-24	16-12-24	23-12-24	30-12-24	06-01-25	13-01-25	20-01-25	27-01-25	03-02-25	10-02-25	17-02-25	24-02-25	03-03-25	10-03-25	17-03-25	24-03-25
		1-12-24	08-12-24	15-12-24	22-12-24	29-12-24	5-1-25	12-01-25	19-01-25	26-1-25	02-02-25	09-02-25	16-2-25	23-02-25	02-03-25	9-3-25	16-03-25	23-03-25	30-3-25
End of RUN5	CYGNO Collab.	Green																	
AmBe Source [low gain]	CYGNO Collab.		Green	Green															
AmBe Source [high gain]	RM1				Green														
Pedestal studies	RM1					Green	Green	Green											
Gas Mixture Recovery	RM1							Green											
Scans in z, VGEM, VDRIFT	LNF, RM3								Green										
Test of new optical system	GSSI									Green	Green	Green	Green	Green	Green	Green	Green	Green	Green
Filters Tests	RM1												Green	Green	Green	Green	Green	Green	Green
Radon Monitor	RM1														Red	Red	Red	Red	Red
83Rb Source	GSSI																Red		
NID	GSSI																	Red	Red

From **January 2025** we are taking **technical runs** intended as **pre-commissioning** of sub-parts of **CYGNO04**: **new camera + lens, gas filters, new calibration source**;

A **Negative Ion Drift** run is foreseen before the closure;

**No shifts** are required to collaboration, only sub-groups are responsible for these data taking;



# LIME DATA TAKING SUMMARY

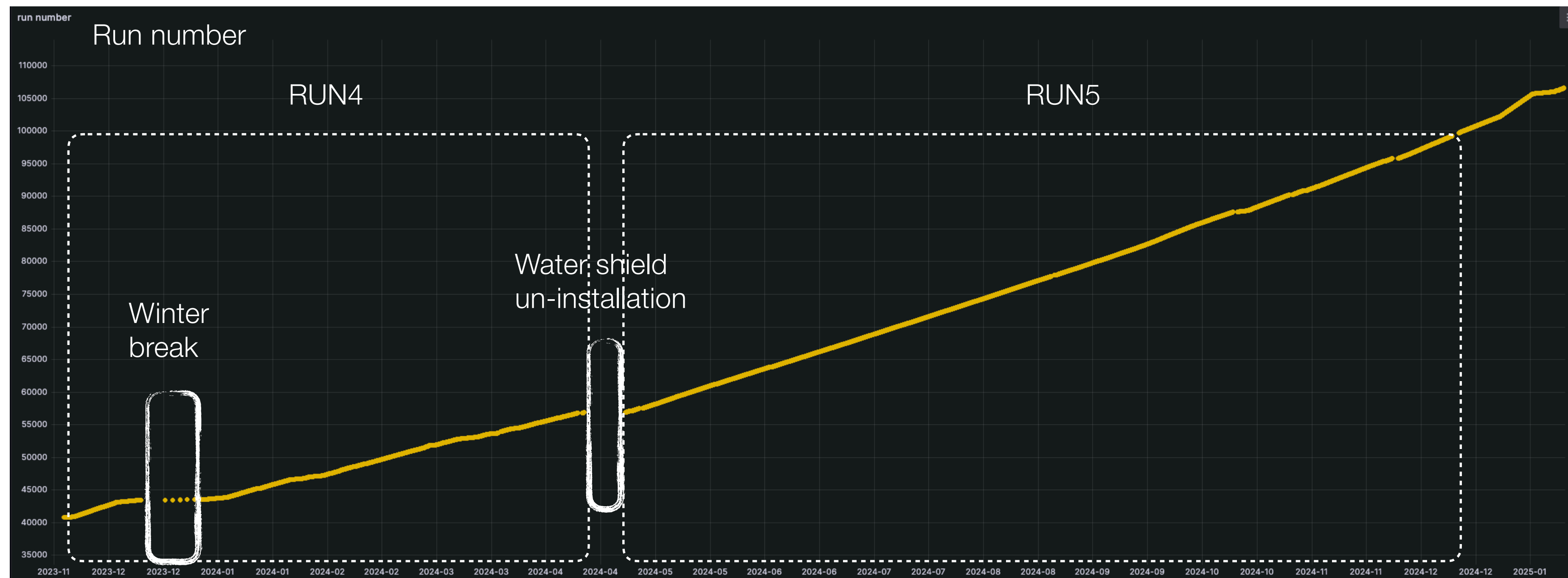


# Ancillary systems: DAQ, Slow control



Data taking started in October 2023 never stopped;

More than  $7 \times 10^4$  runs have been taken for a total of  $28 \times 10^6$  pictures;



All of them have been **promptly transferred** on the INFN-Cloud and **reconstructed**;

**Reco-files** are made available to the CYGNO users **few minutes after** the run was taken;

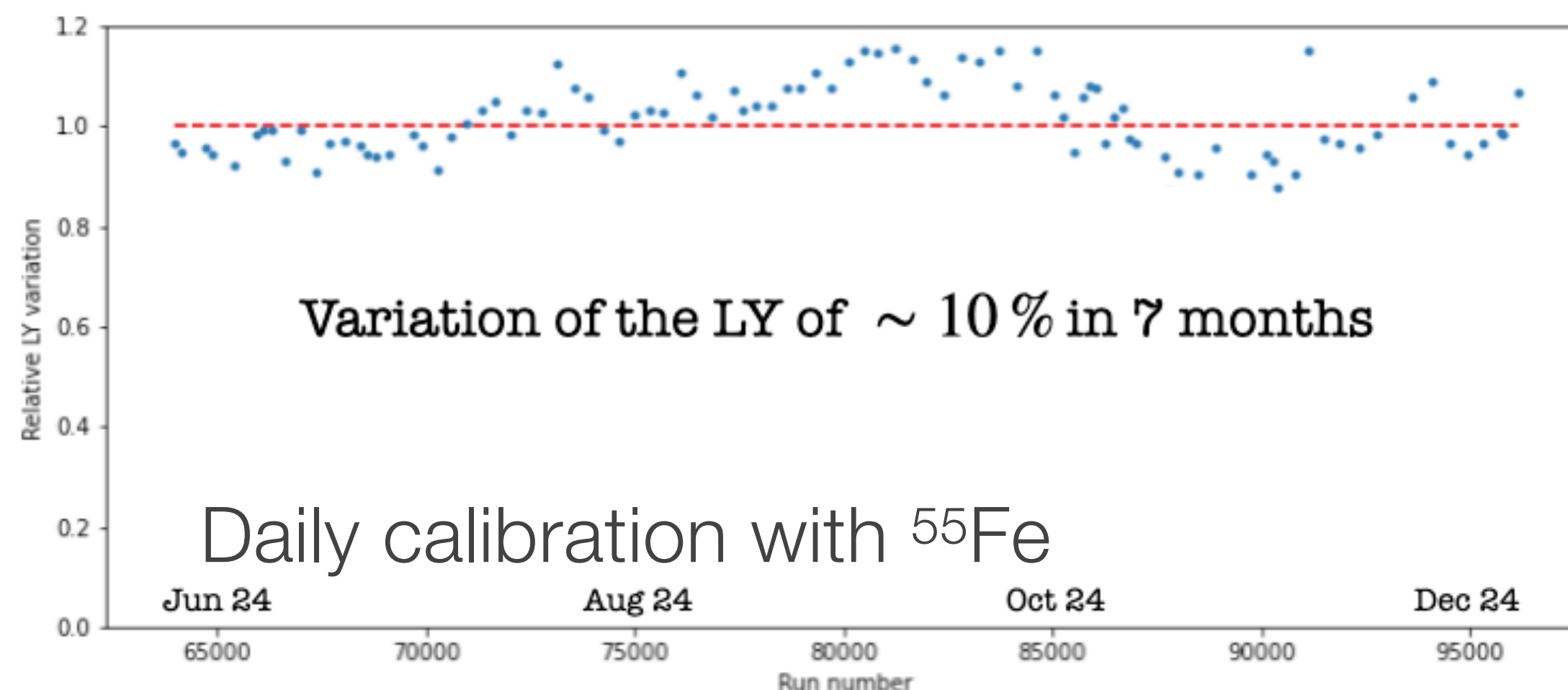
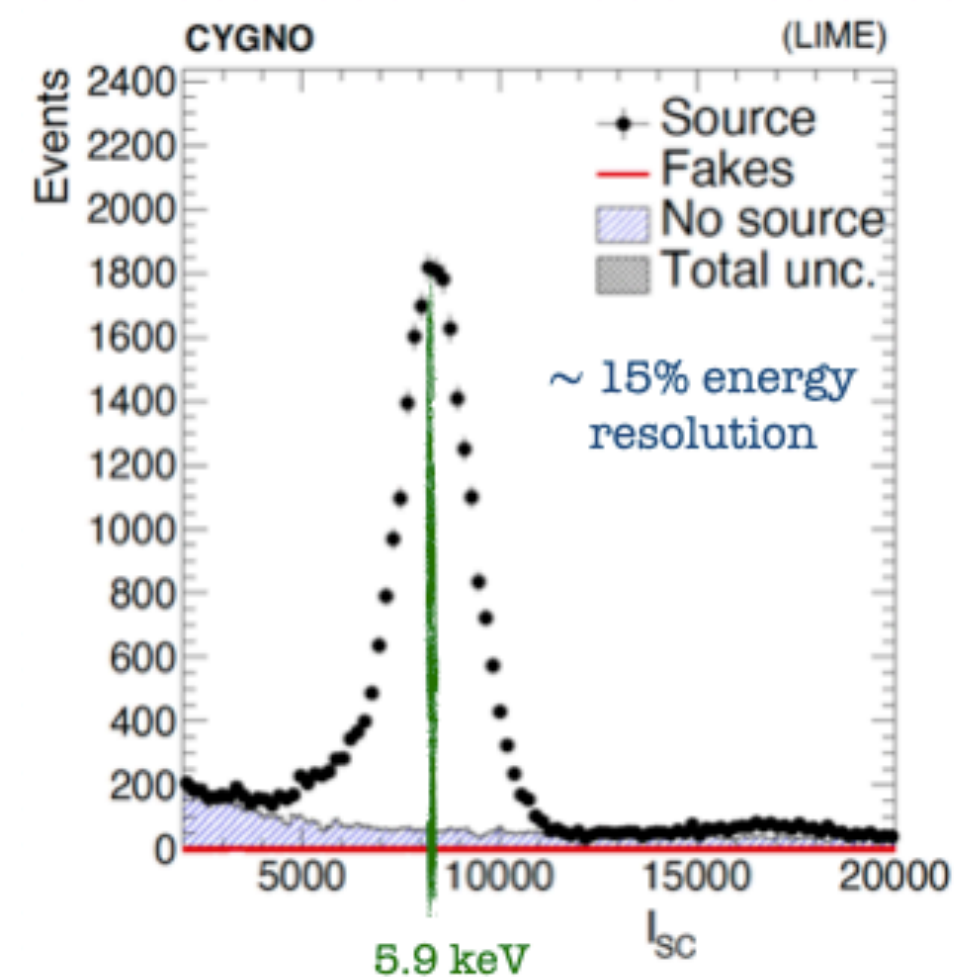
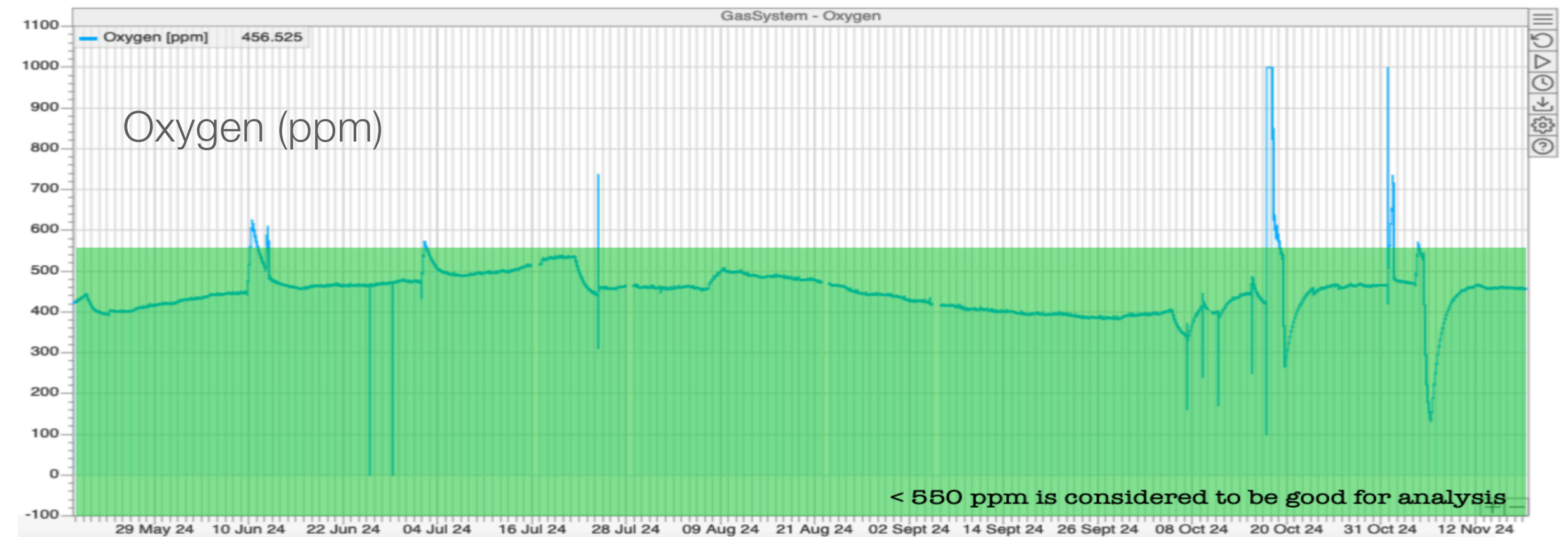


# Ancillary systems: Gas and HV



Less than one **spark**-like every 3 days was recorded in 2024 (12 times lesser than previous best limit ...)

Gas **humidity**, **oxygen** and other **contaminants** were **under control** for the whole **2024**



Detector calibration with  $^{55}\text{Fe}$  is now a **completely automated** procedure and showed a **stability with 7% RMS** over 6 months



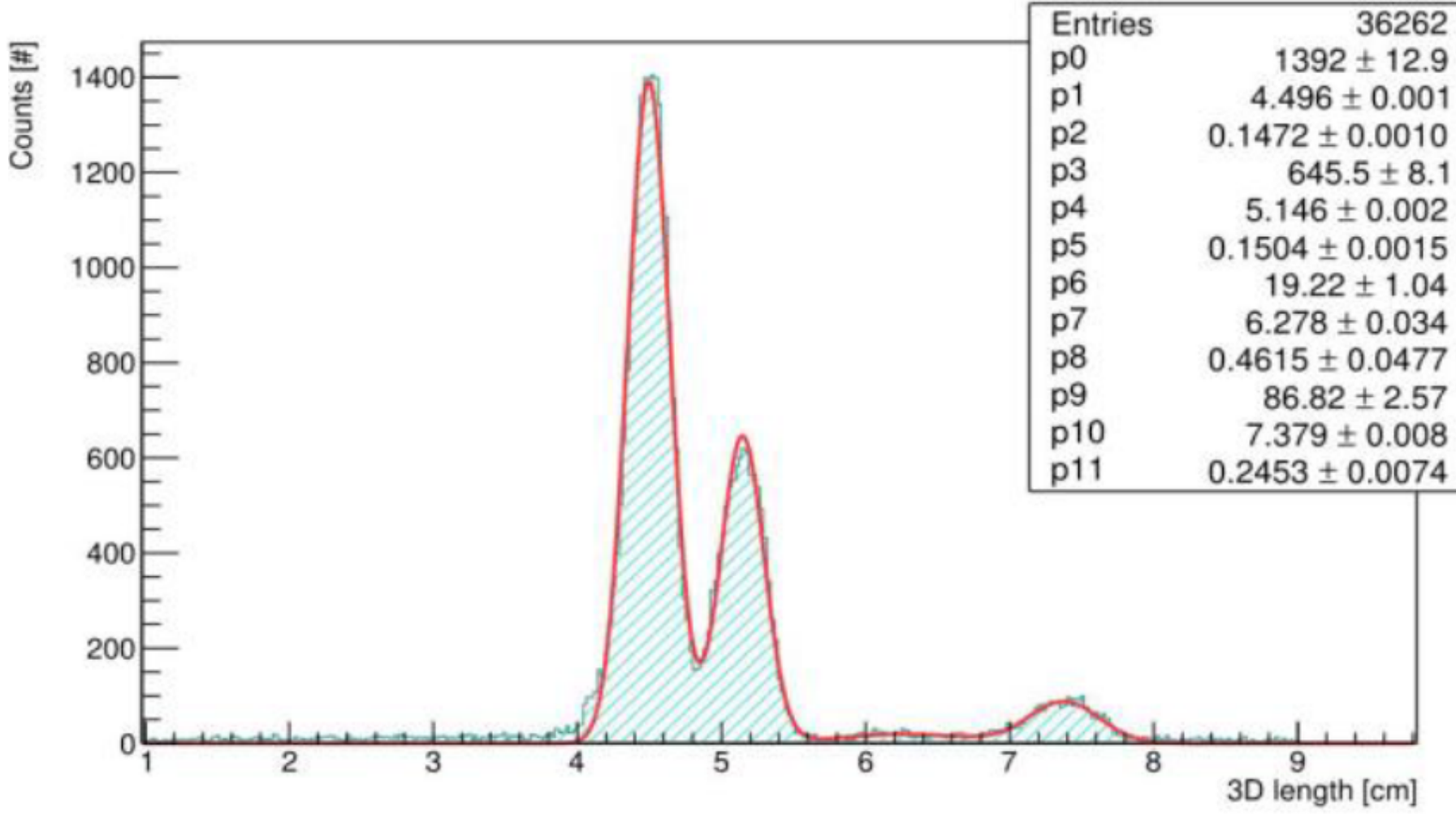
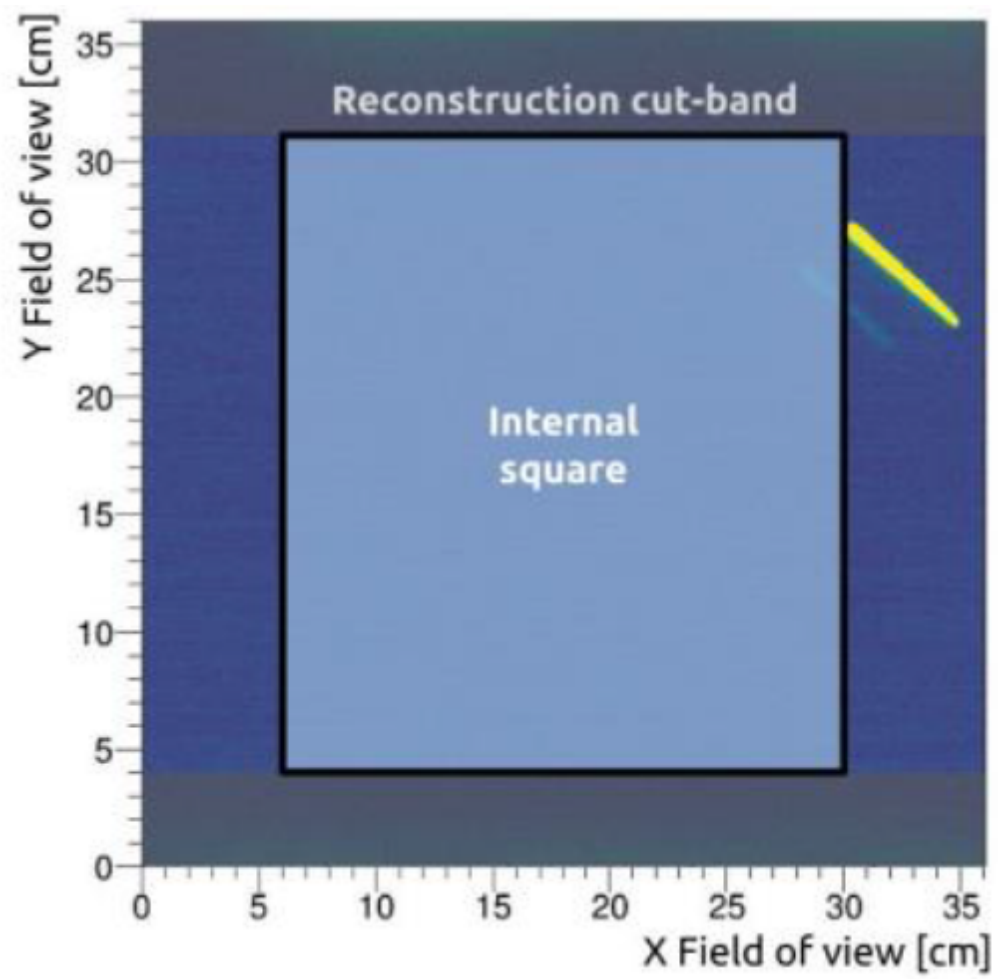
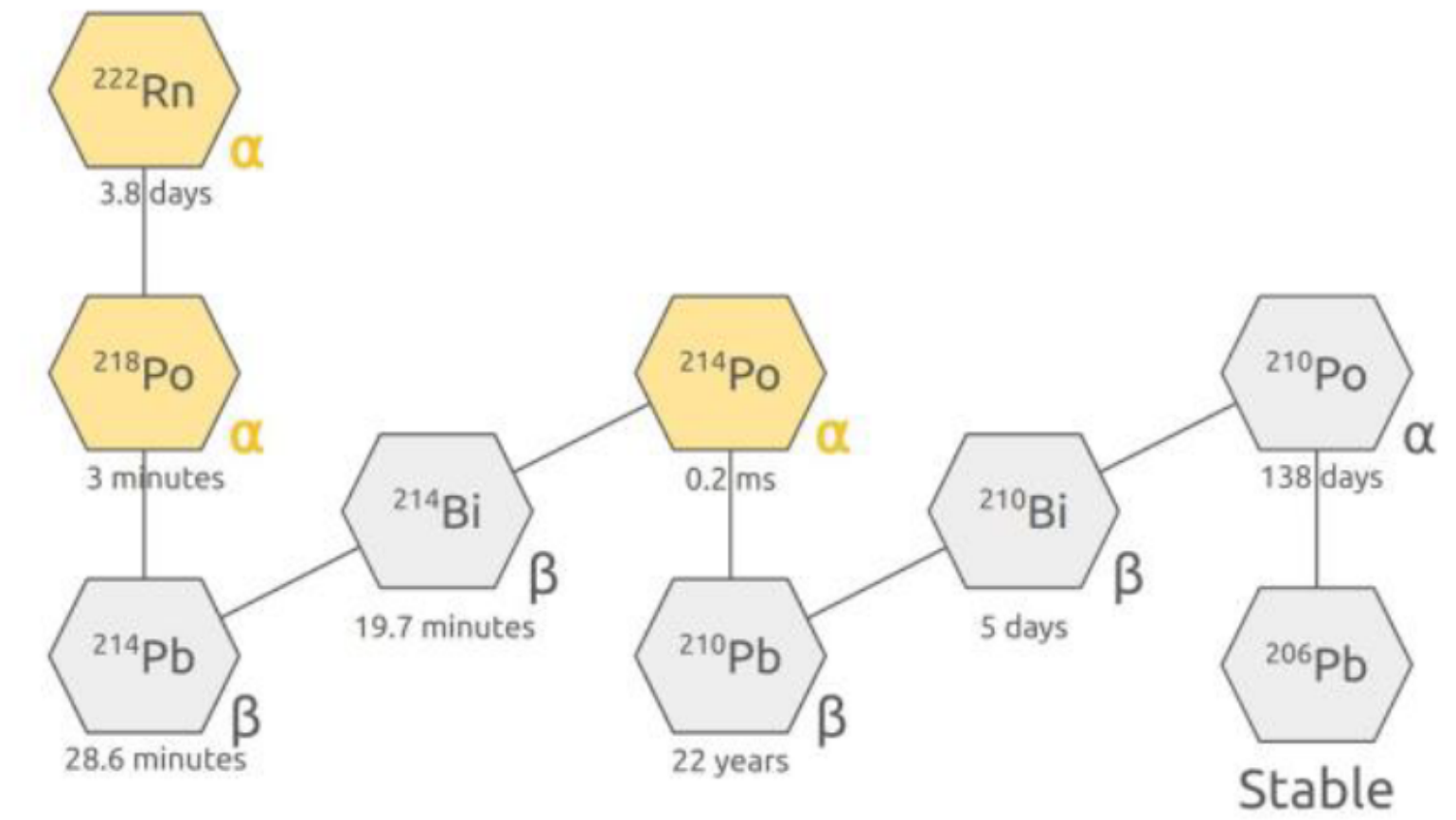
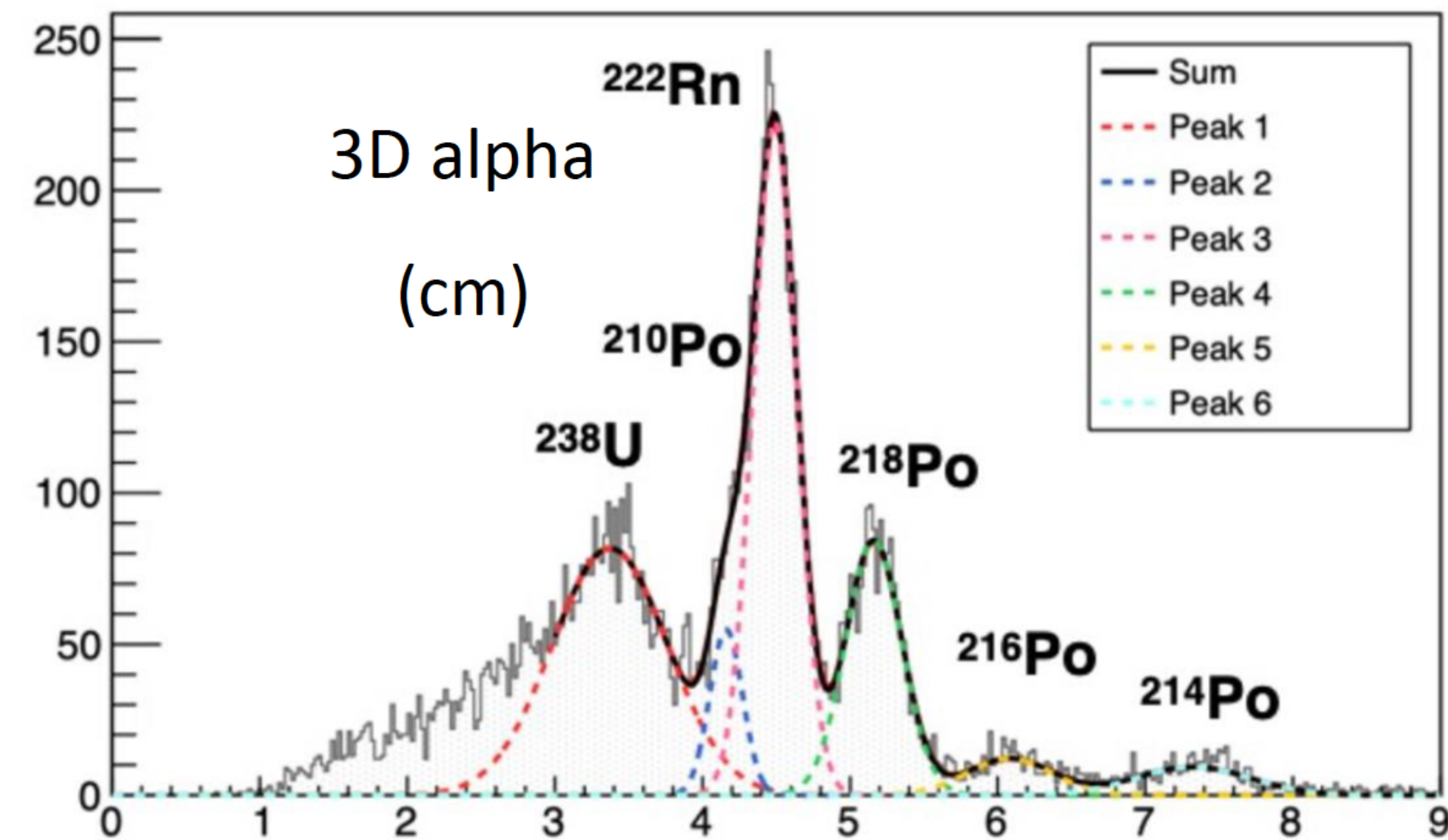
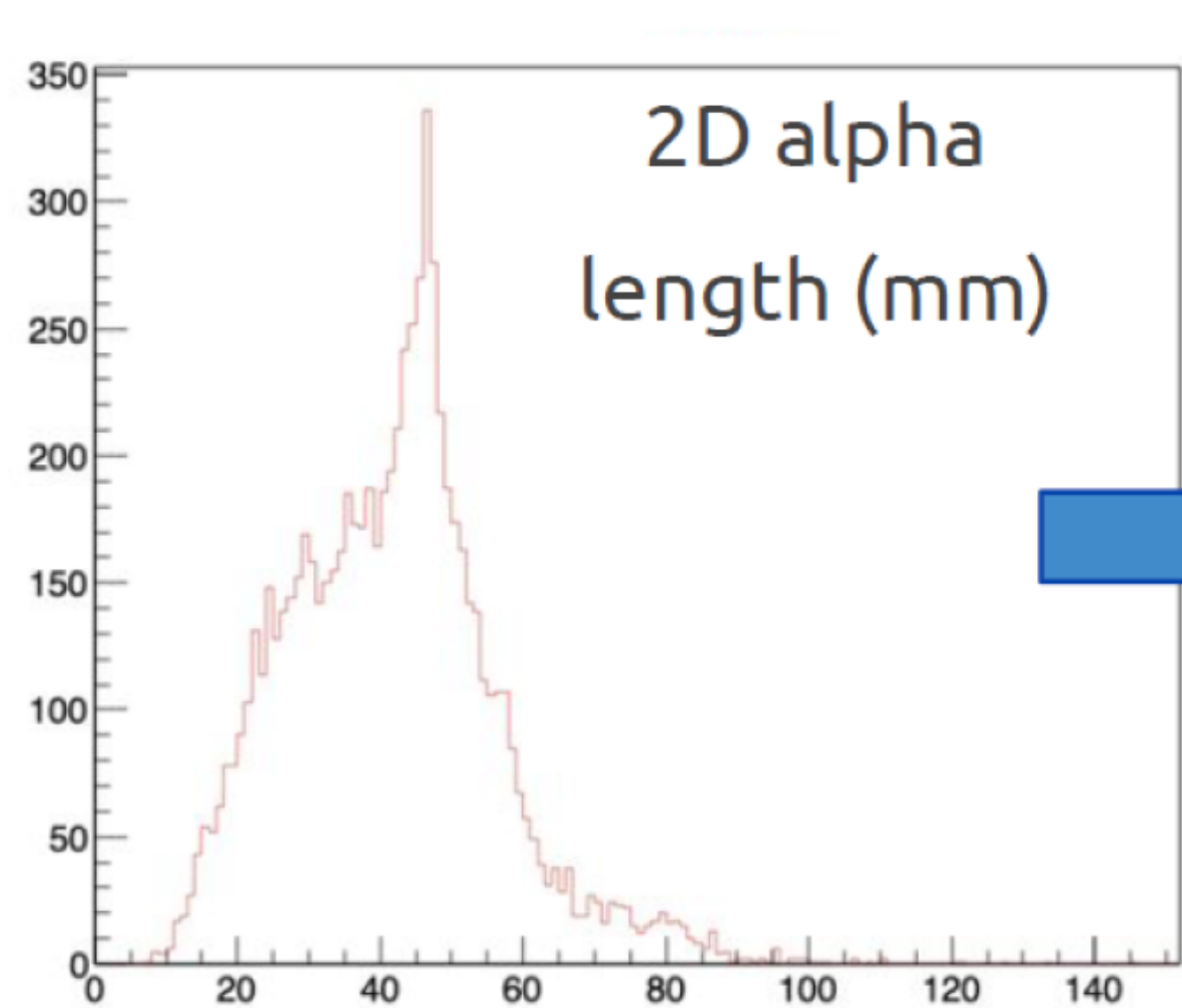
# LIME DATA ANALYSIS SUMMARY



# 3D and Radon



3D reconstruction allowed a precise measurement of the length of the alpha tracks with a consequent assessment of their energies

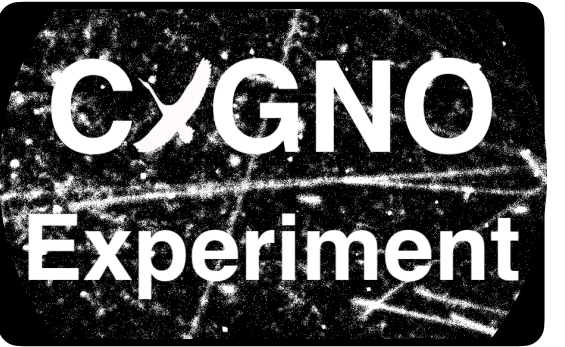


Isotope	Energy (MeV)	Expected length (mm)	Measured length (mm)
222Rn	5,50	45,7	44,3
218Po	6,00	51,0	51,2
214Po	6,79	71,0	72,3

222Rn contamination **confirmed**

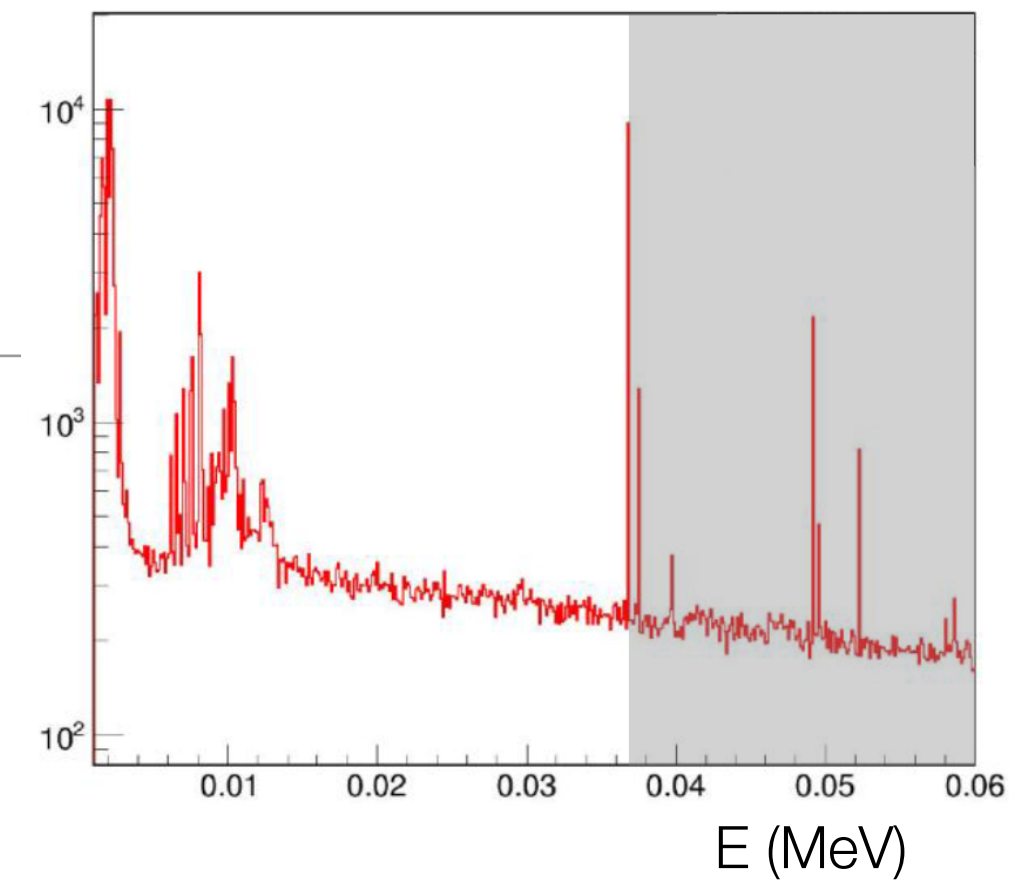


# 3D and Radon



Is there any correlation between the  $^{222}\text{Rn}$  quantity (evaluable from the rate of the high energy tracks) and the low energy part of the spectrum?

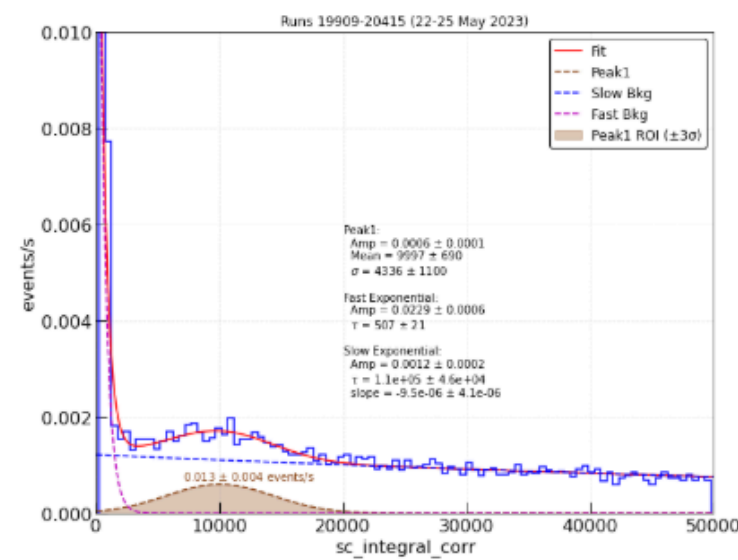
Energy deposits from  $^{222}\text{Rn}$  chain  
GEANT



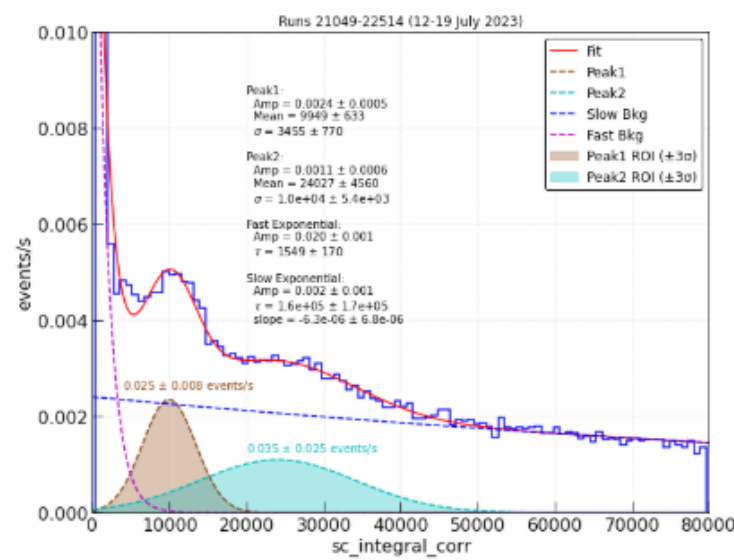
low energy

high energy

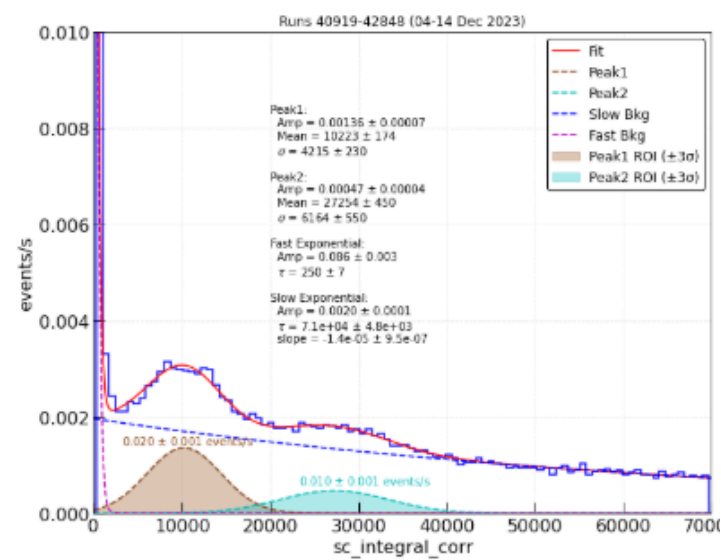
**RUN3: May 2023**  
only fresh gas



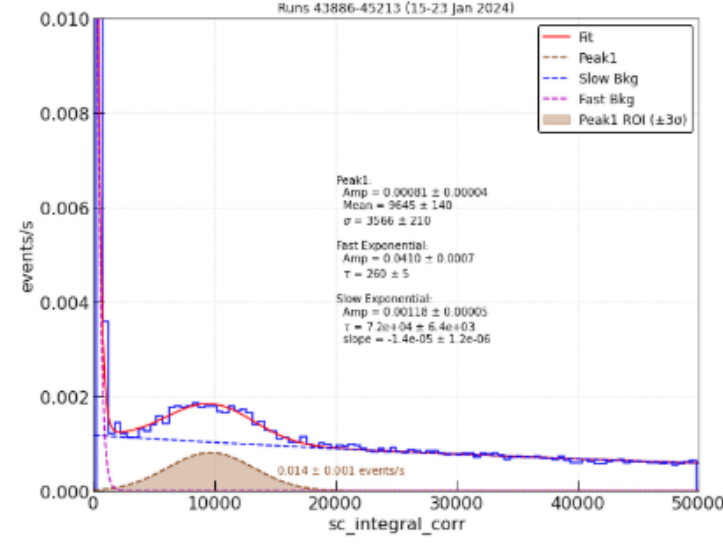
**RUN3: July 2023**  
high contamination



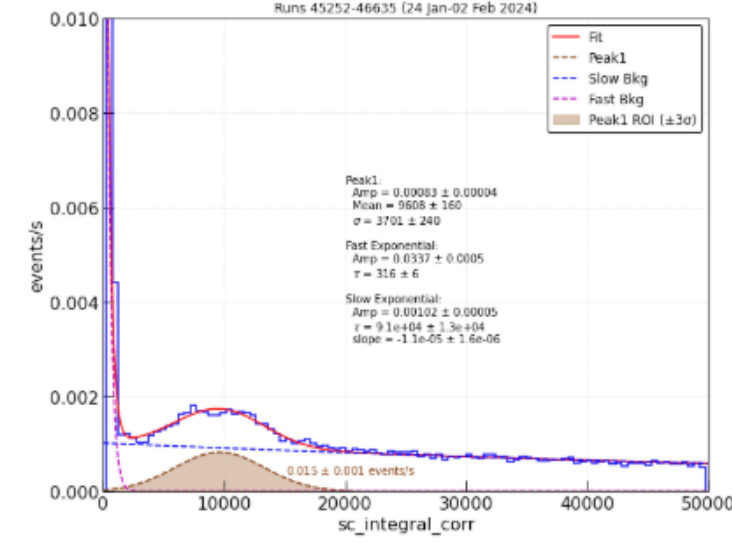
**RUN4: Dec 2023**  
part-unfiltered



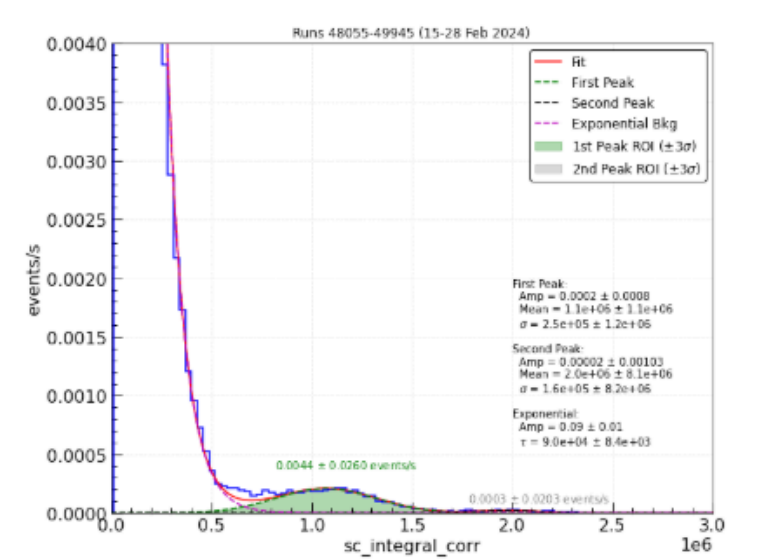
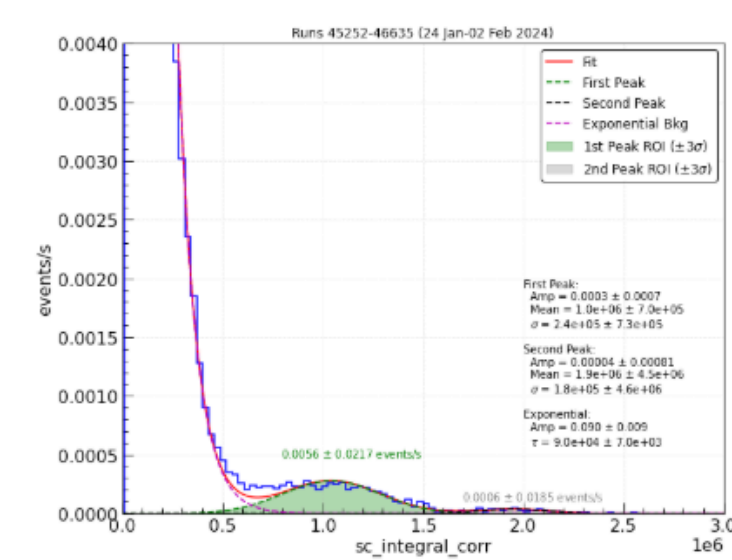
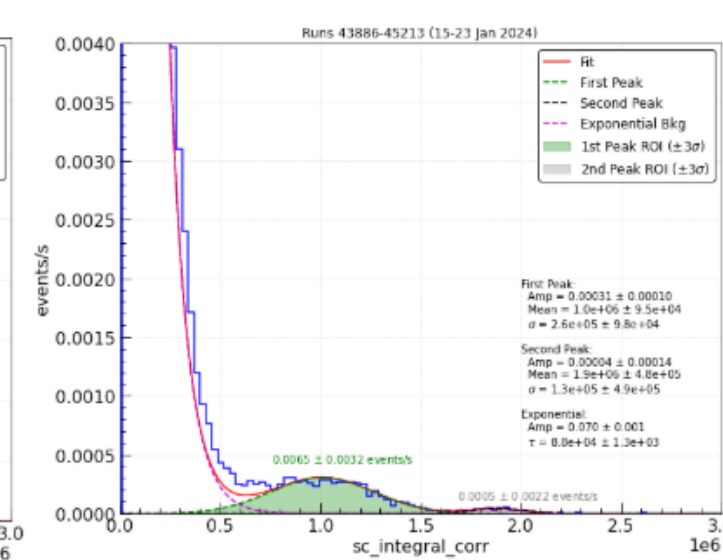
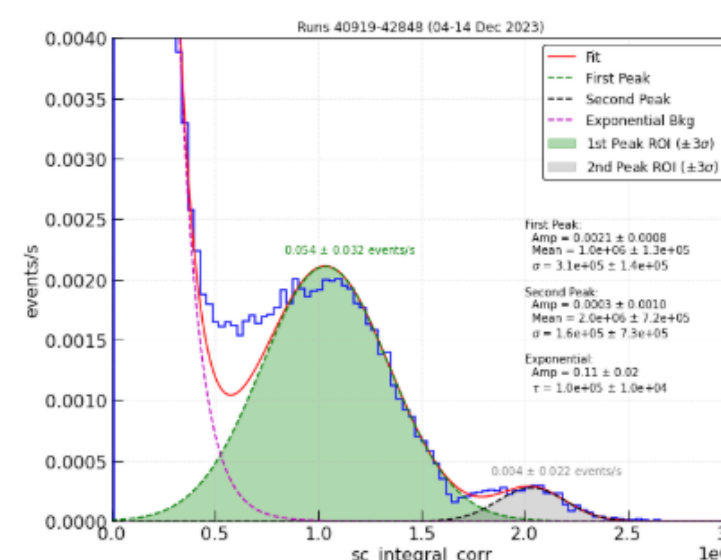
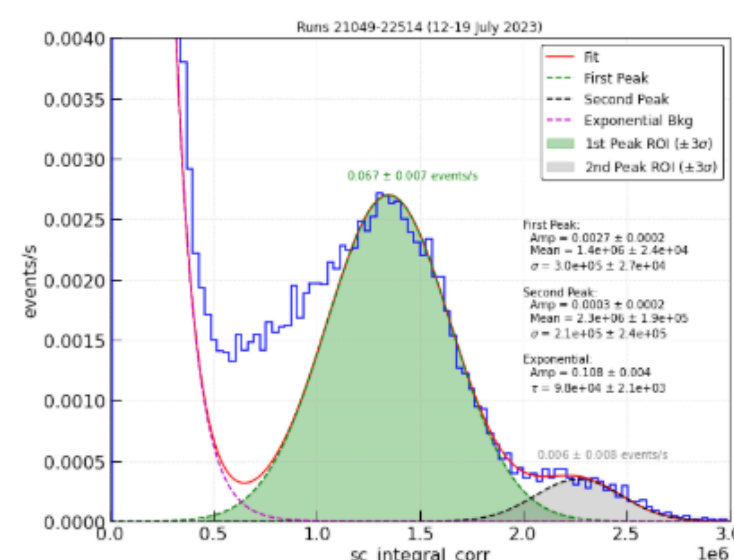
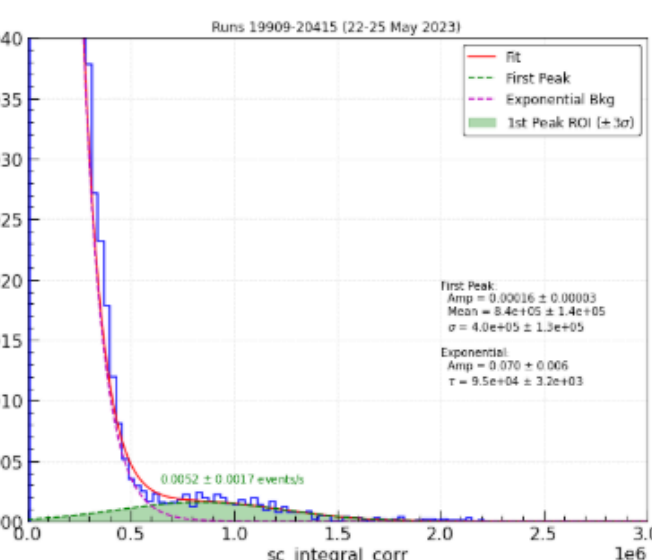
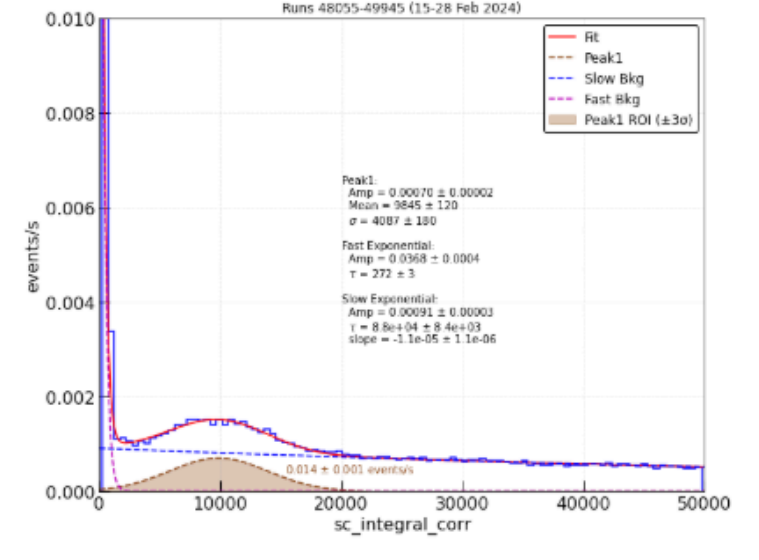
**RUN4: Jan 2024**  
filtered



**RUN4: Feb 2024**  
filtered



**RUN4: Feb 2024**  
filtered



A quantitative study on the correlation between the two regions is **under study**;  
Measurement campaign with a new radon-meter will start in 10 days;



# CYGNO04



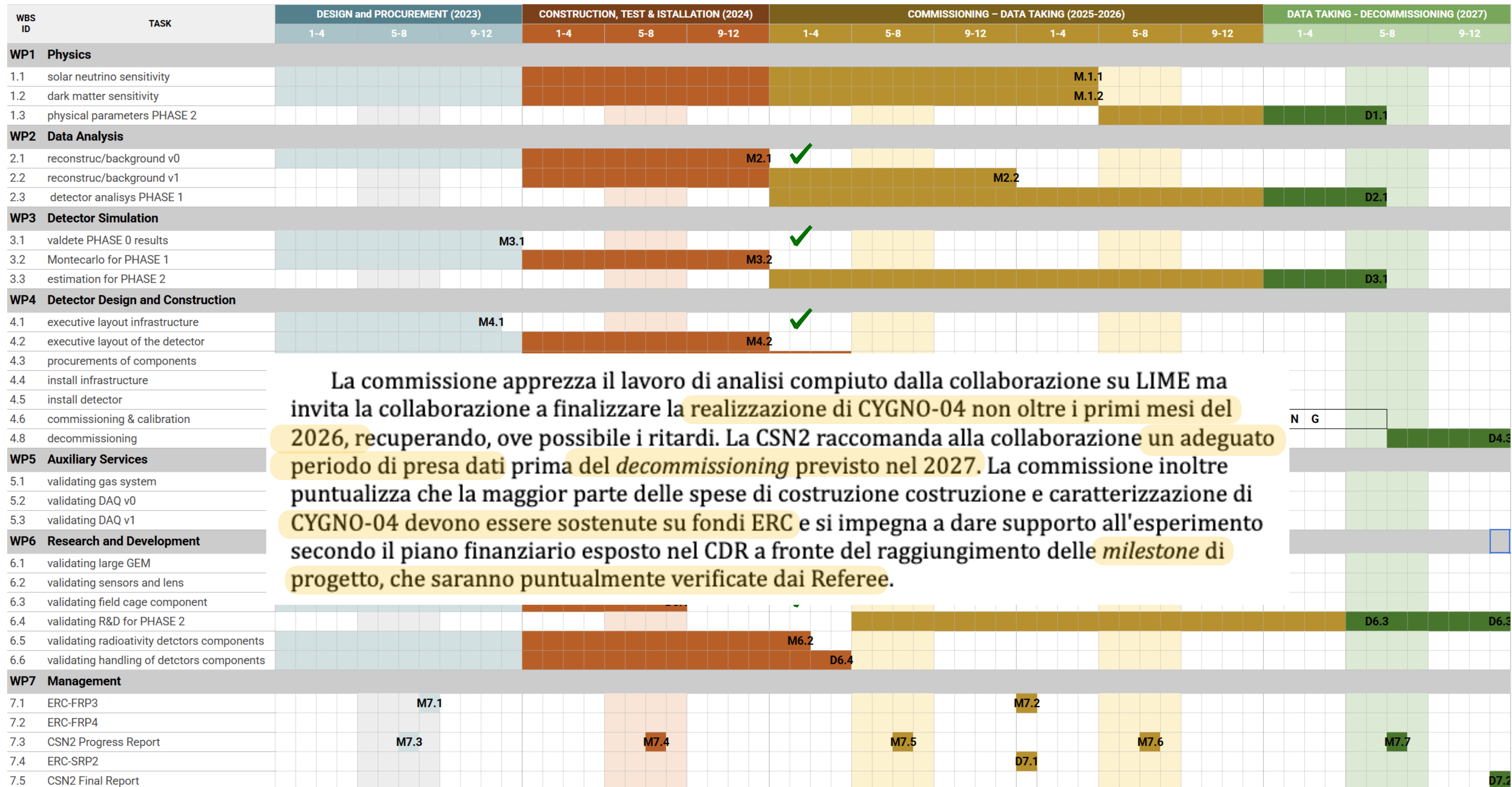
# CYGNO04 GANTT Progress Report 2024



WBS ID	TASK	DESIGN and PROCUREMENT (2023)			CONSTRUCTION, TEST & INSTALLATION (2024)			COMMISSIONING – DATA TAKING (2025-2026)						DATA TAKING - DECOMMISSIONING (2027)			
		1-4	5-8	9-12	1-4	5-8	9-12	1-4	5-8	9-12	1-4	5-8	9-12	1-4	5-8	9-12	
<b>WP1 Physics</b>																	
1.1	solar neutrino sensitivity														M.1.1		
1.2	dark matter sensitivity														M.1.2		
1.3	physical parameters PHASE 2																D1.1
<b>WP2 Data Analysis</b>																	
2.1	reconstruc/background v0						M2.1	✓									
2.2	reconstruc/background v1								M2.2								
2.3	detector analisys PHASE 1																D2.1
<b>WP3 Detector Simulation</b>																	
3.1	valdete PHASE 0 results			M3.1				✓									
3.2	Montecarlo for PHASE 1						M3.2										
3.3	estimation for PHASE 2																D3.1
<b>WP4 Detector Design and Construction</b>																	
4.1	executive layout infrastructure			M4.1				✓									
4.2	executive layout of the detector						M4.2										
4.3	procurements of components								M4.3								
4.4	install infrastructure						D4.1										
4.5	install detector									D4.2							
4.6	commissioning & calibration										M4.4	-> D A T A T A K I N G					
4.8	decommissioning																D4.3
<b>WP5 Auxiliary Services</b>																	
5.1	validating gas system	D5.1						✓✓									
5.2	validating DAQ v0			M5.1				✓									
5.3	validating DAQ v1						D5.2	✓									
<b>WP6 Research and Development</b>																	
6.1	validating large GEM			M6.1				✓✓									
6.2	validating sensors and lens						D6.2	✓									
6.3	validating field cage component						D6.1	✓									
6.4	validating R&D for PHASE 2															D6.3	D6.3
6.5	validating radioactivity detctors components								M6.2								
6.6	validating handling of detctors components									D6.4							
<b>WP7 Management</b>																	
7.1	ERC-FRP3			M7.1								M7.2					
7.2	ERC-FRP4																
7.3	CSN2 Progress Report			M7.3			M7.4			M7.5			M7.6			M7.7	
7.4	ERC-SRP2											D7.1					
7.5	CSN2 Final Report																D7.2



# CYGNO04 GANTT Progress Report 2024



La commissione apprezza il lavoro di analisi compiuto dalla collaborazione su LIME ma invita la collaborazione a finalizzare la realizzazione di CYGNO-04 non oltre i primi mesi del 2026, recuperando, ove possibile i ritardi. La CSN2 raccomanda alla collaborazione un adeguato periodo di presa dati prima del decommissioning previsto nel 2027. La commissione inoltre puntualizza che la maggior parte delle spese di costruzione costruzione e caratterizzazione di CYGNO-04 devono essere sostenute su fondi ERC e si impegna a dare supporto all'esperimento secondo il piano finanziario esposto nel CDR a fronte del raggiungimento delle milestone di progetto, che saranno puntualmente verificate dai Referee.

N G



In Nov 2024 **Davide Fiorina** (GSSI) was appointed as Technical Coordinator for CYGNO, 6 months later Giovanni Mazzitelli resignation;

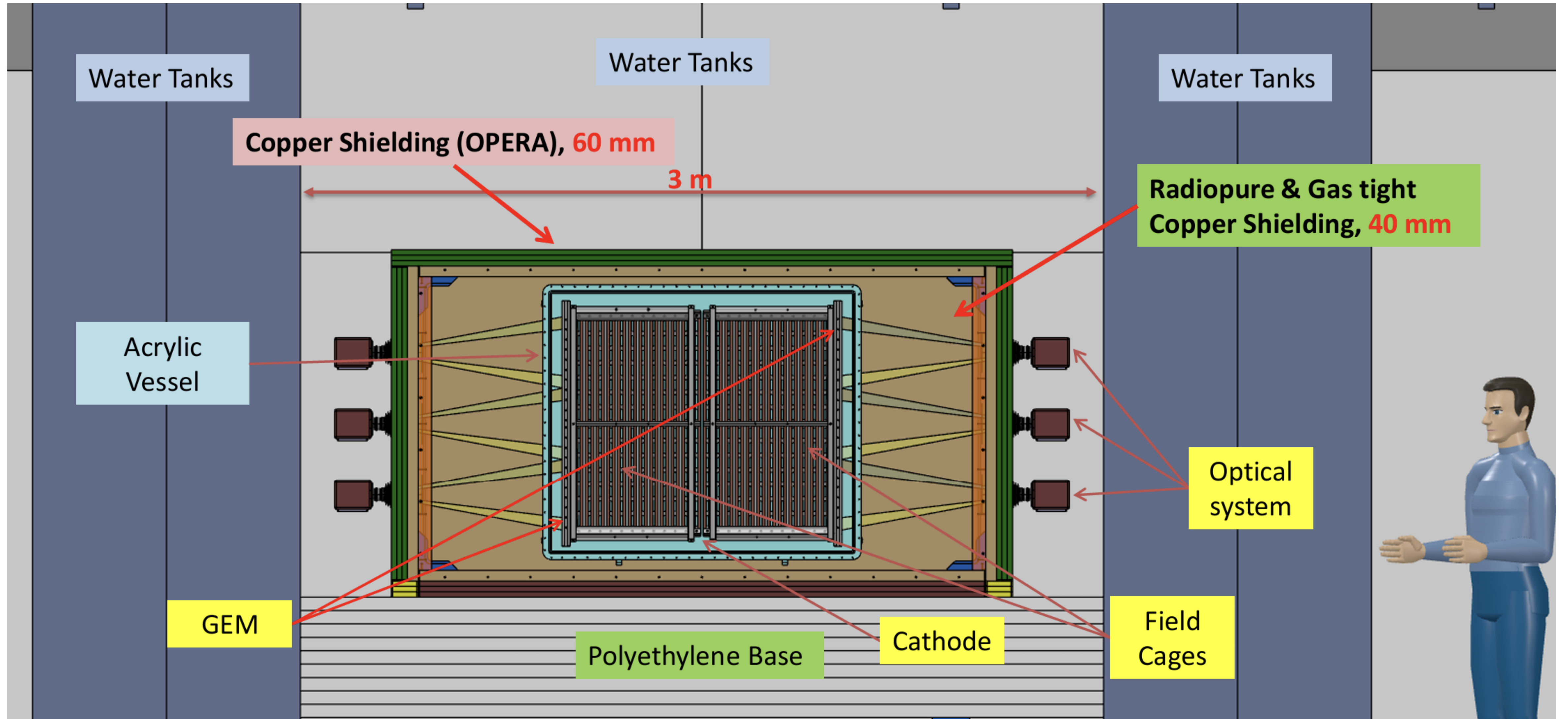
With the new year and a stabilised situation, it became clear that the **delay** in producing the **executive drawings for CYGNO04** was unrecoverable by relying solely on the forces available to LNF's Design Service

We therefore asked the **LNGS Design Service** for support: merging and sharing their resources with LNF's one can be the **opportunity to finalise the work** done so far in a reasonable time interval;

In order to reach this aim, the proposal issued by the LNGS+LNF Services is the creation of a **task force** including:

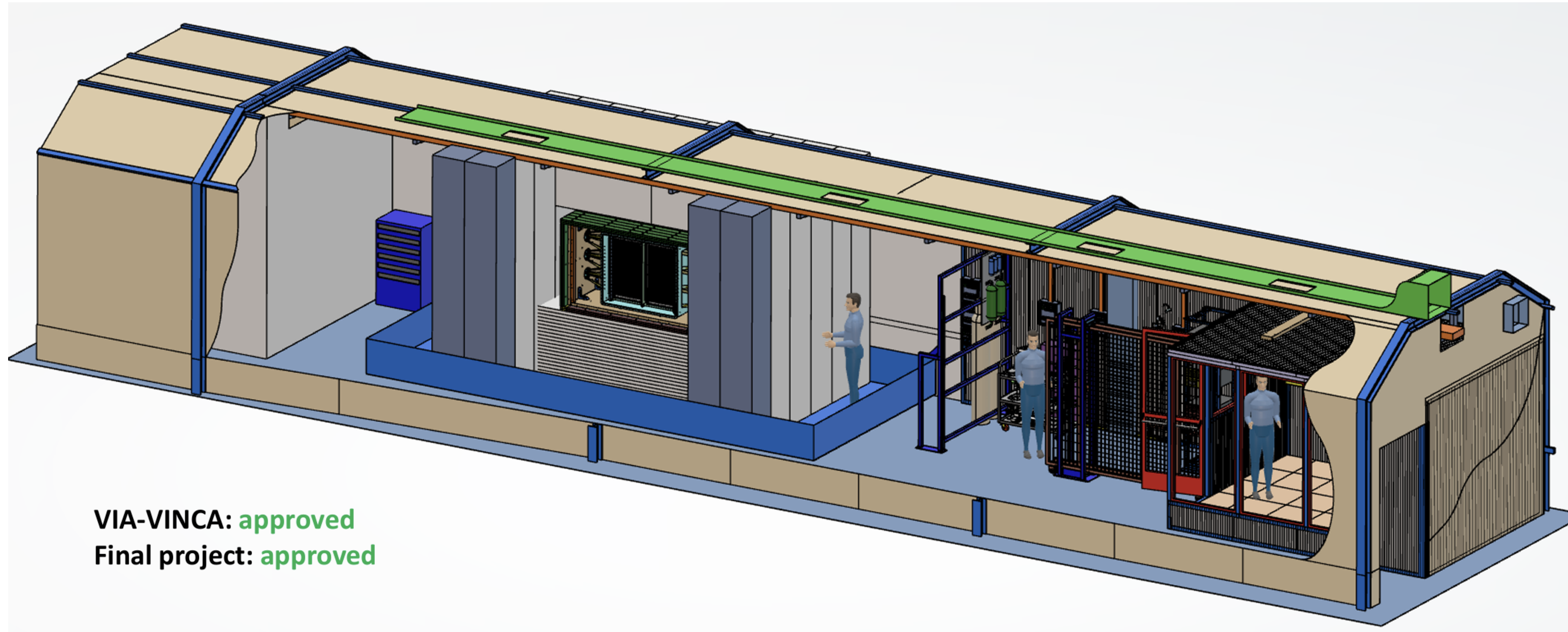
- **a designer** with the supervision of the **heads of the Services** from **both** the labs;
- having as **direct link to CYGNO** Collaboration the **Technical Coordinator** supported by Davide Pinci as CYGNO spokesperson;
- We are now on the correct path to deliver the detector in the **first trimester of 2026**







# Hall-F Infrastructure

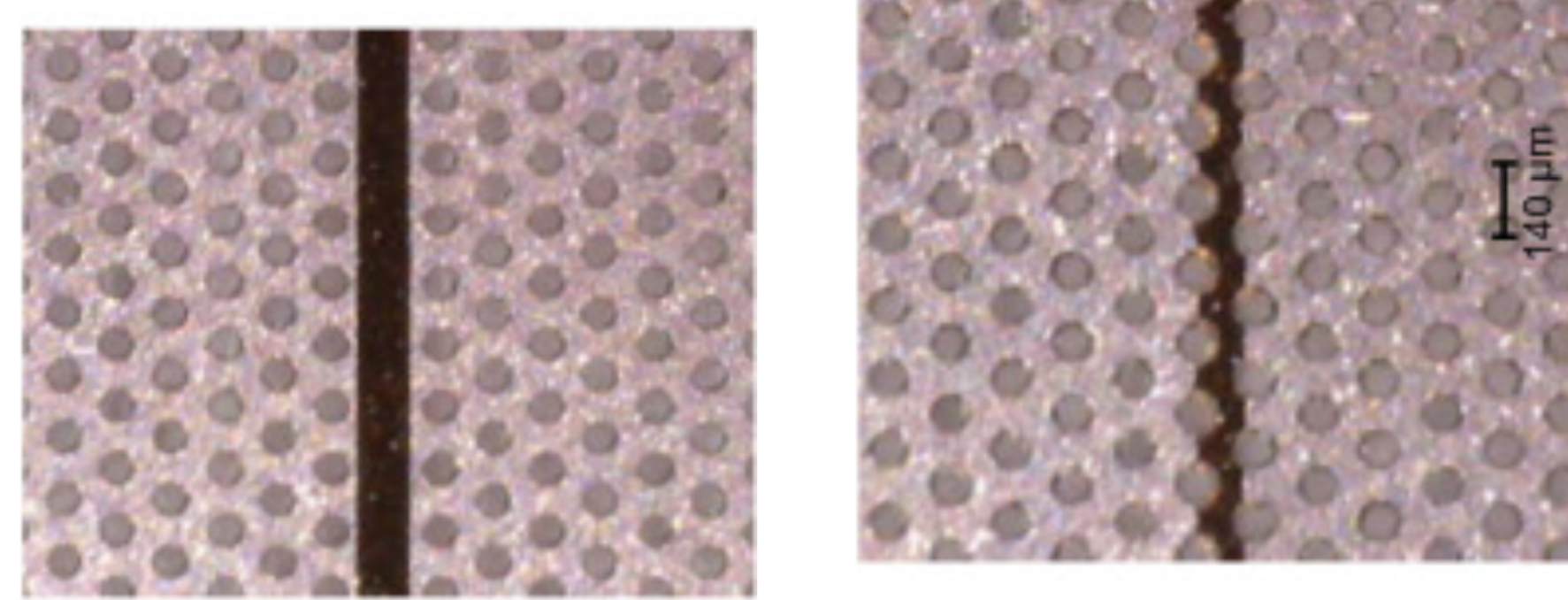
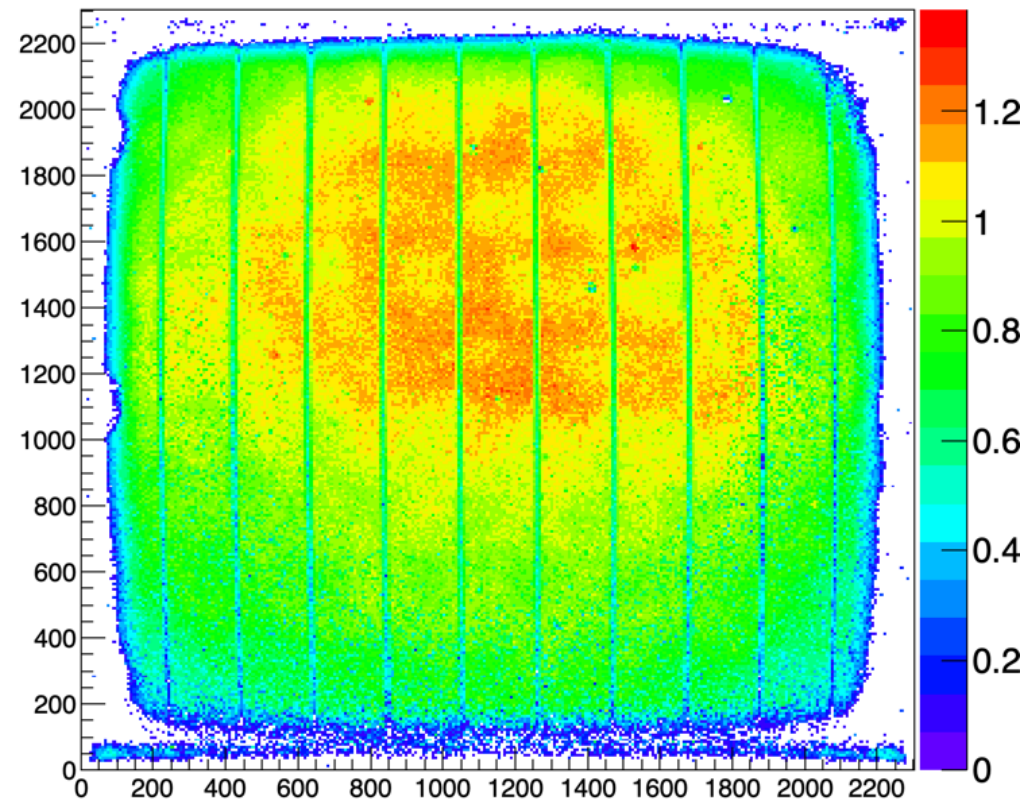


VIA-VINCA: approved  
Final project: approved

- Civil works **started** this week and **delivery** expected **end of February**;
- In accordance with the CYGN004 timeline;

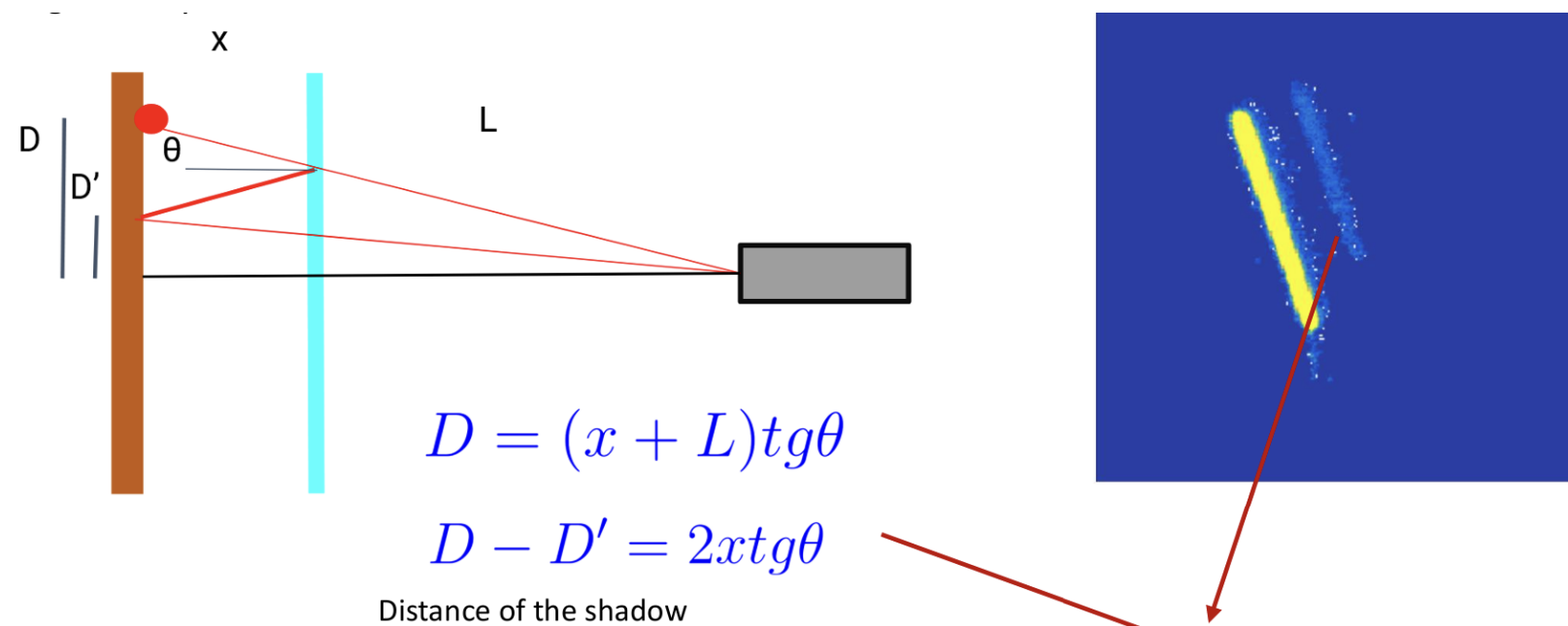


# Component Validation - GEM

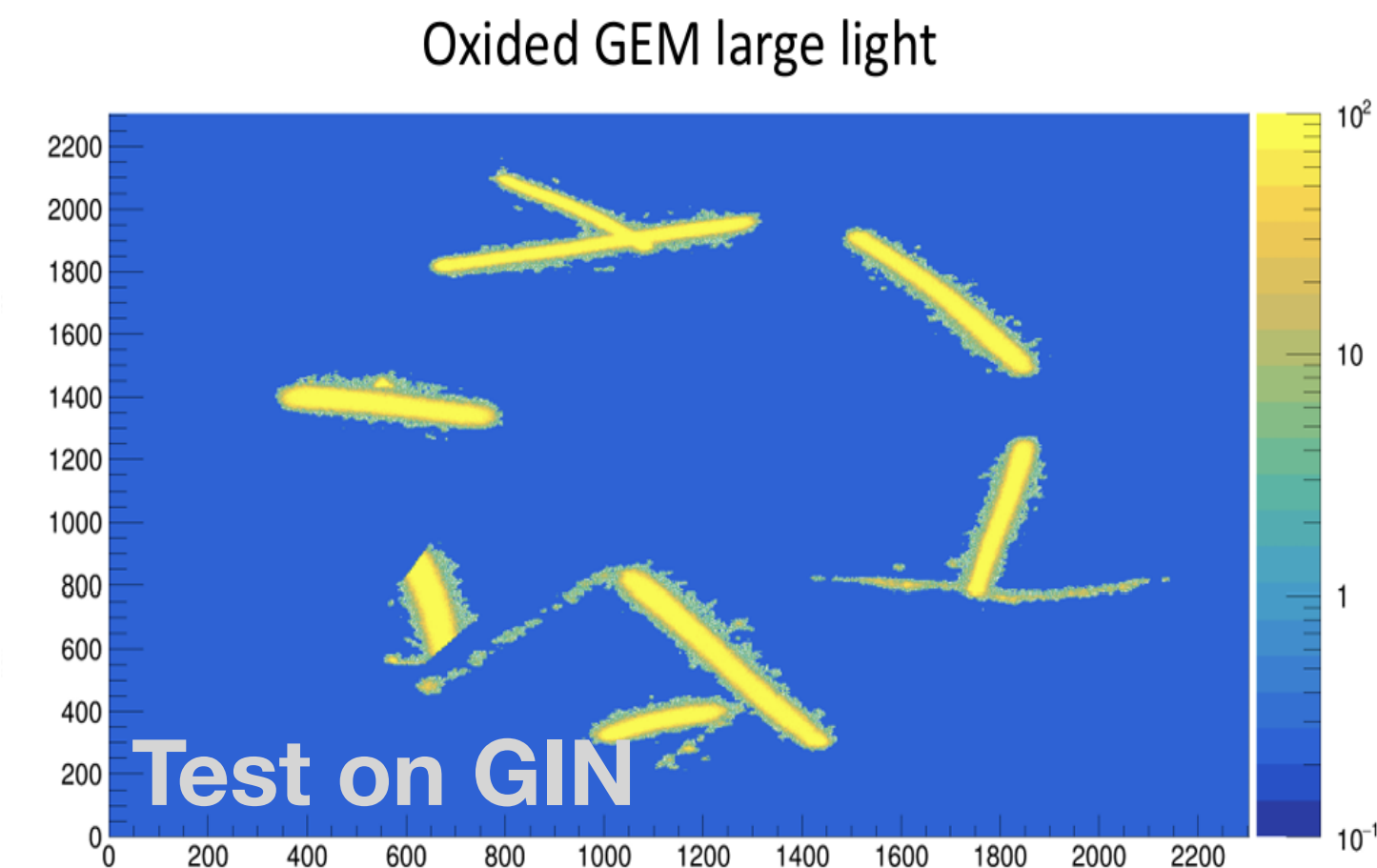
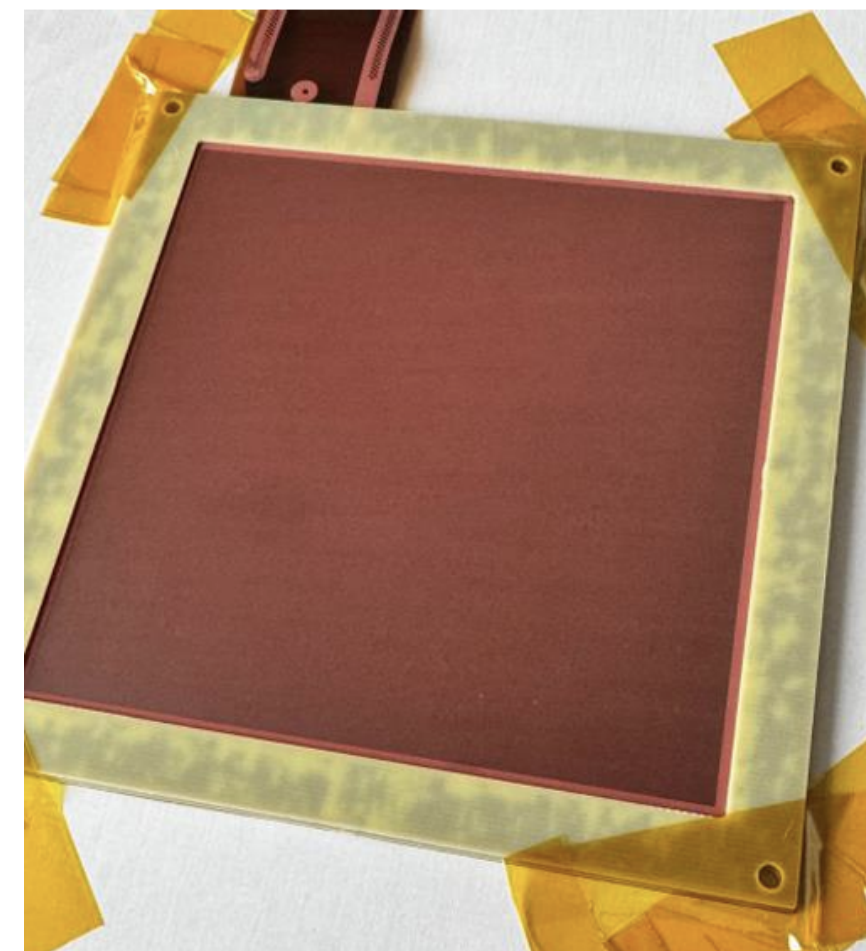


Random **segmentation** should reduce the dead area while keeping the operational stability (tested by CMS-GEM)

**Oxidised GEM avoiding reflection** of large light emission on the PMMA window as observed in LIME



Theory and measurement agree

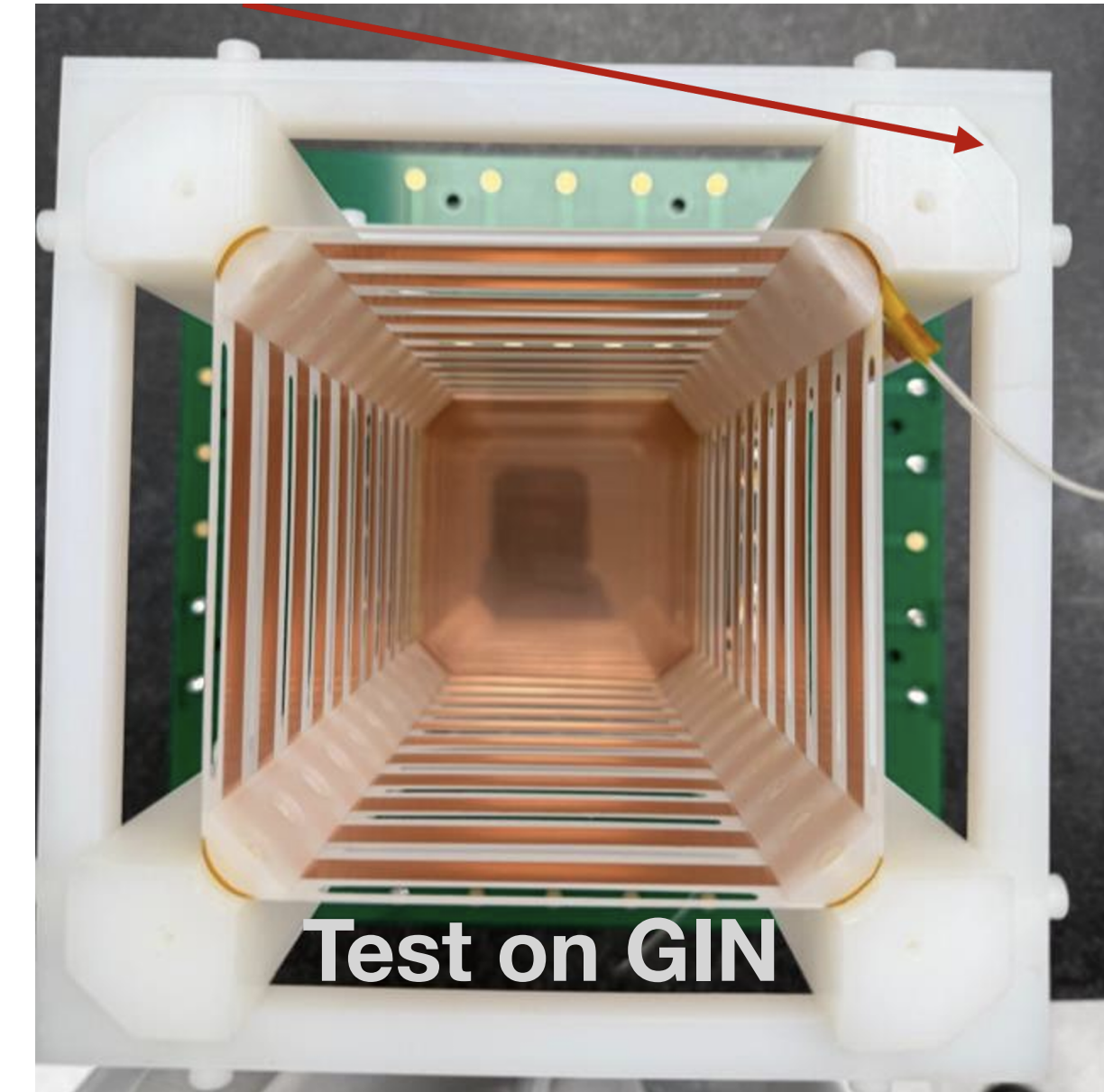
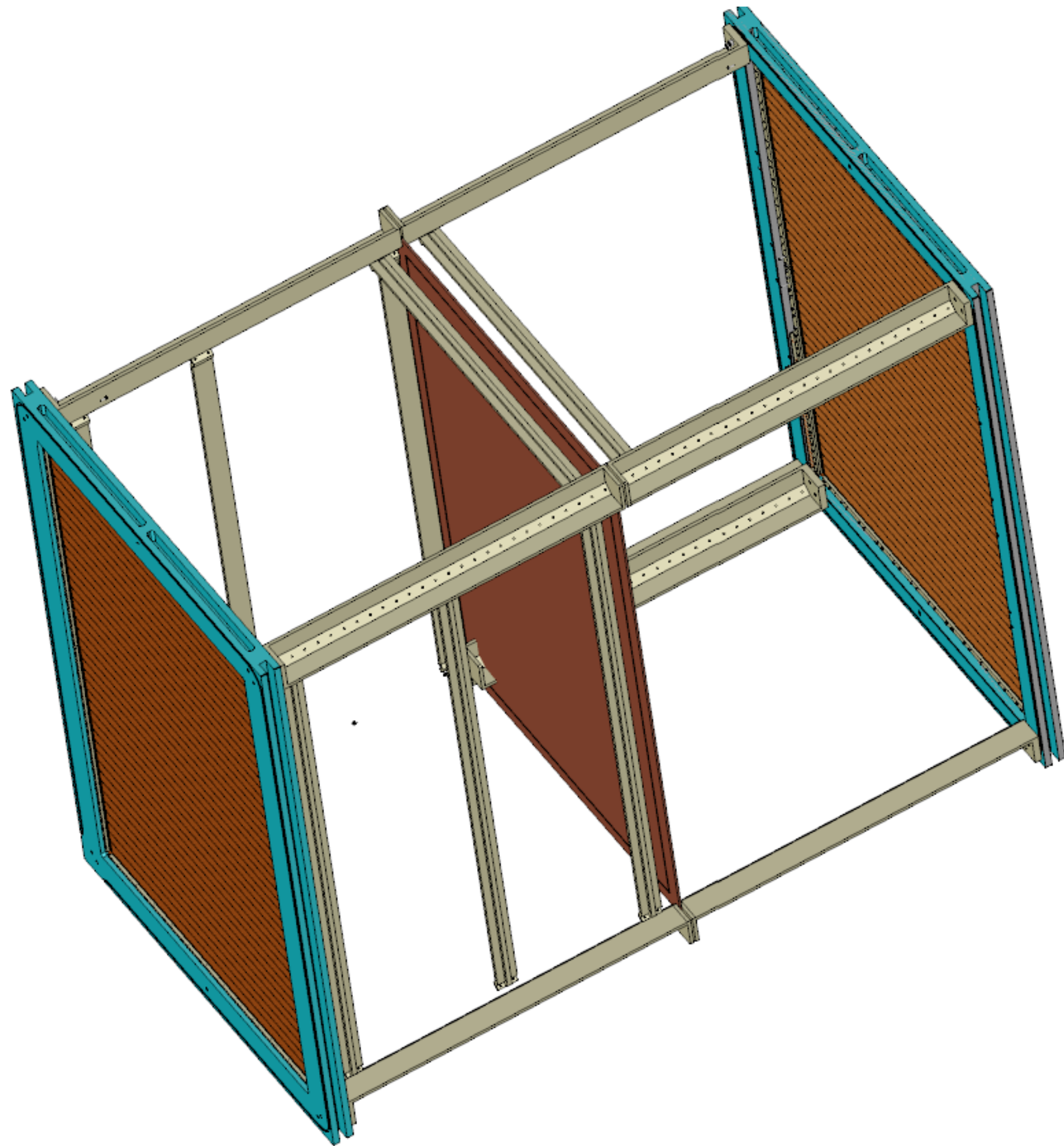


Done!

**Studies** and developments **funded** by CSN2 in past years



# Component Validation - Field Cage



## Field Cage Structure in Nylon66

- Field rings made by copper strips on Kapton (PET as backup)
- Resistors soldered on one side

The cathode will be a Kapton foil copper clad on both sides

- Backup solution, full copper plate

Done!

**Studies** and developments **funded** by CSN2 in past years

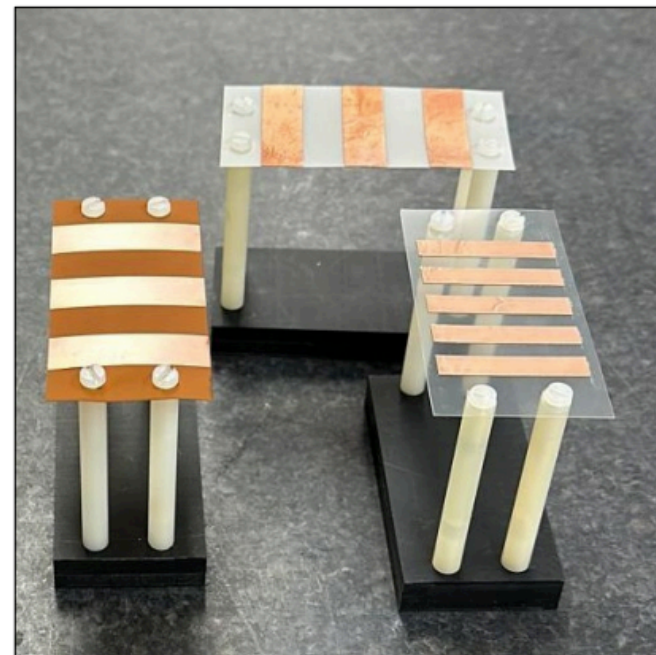


# Component Validation - Field Cage



## Field Cage Foils

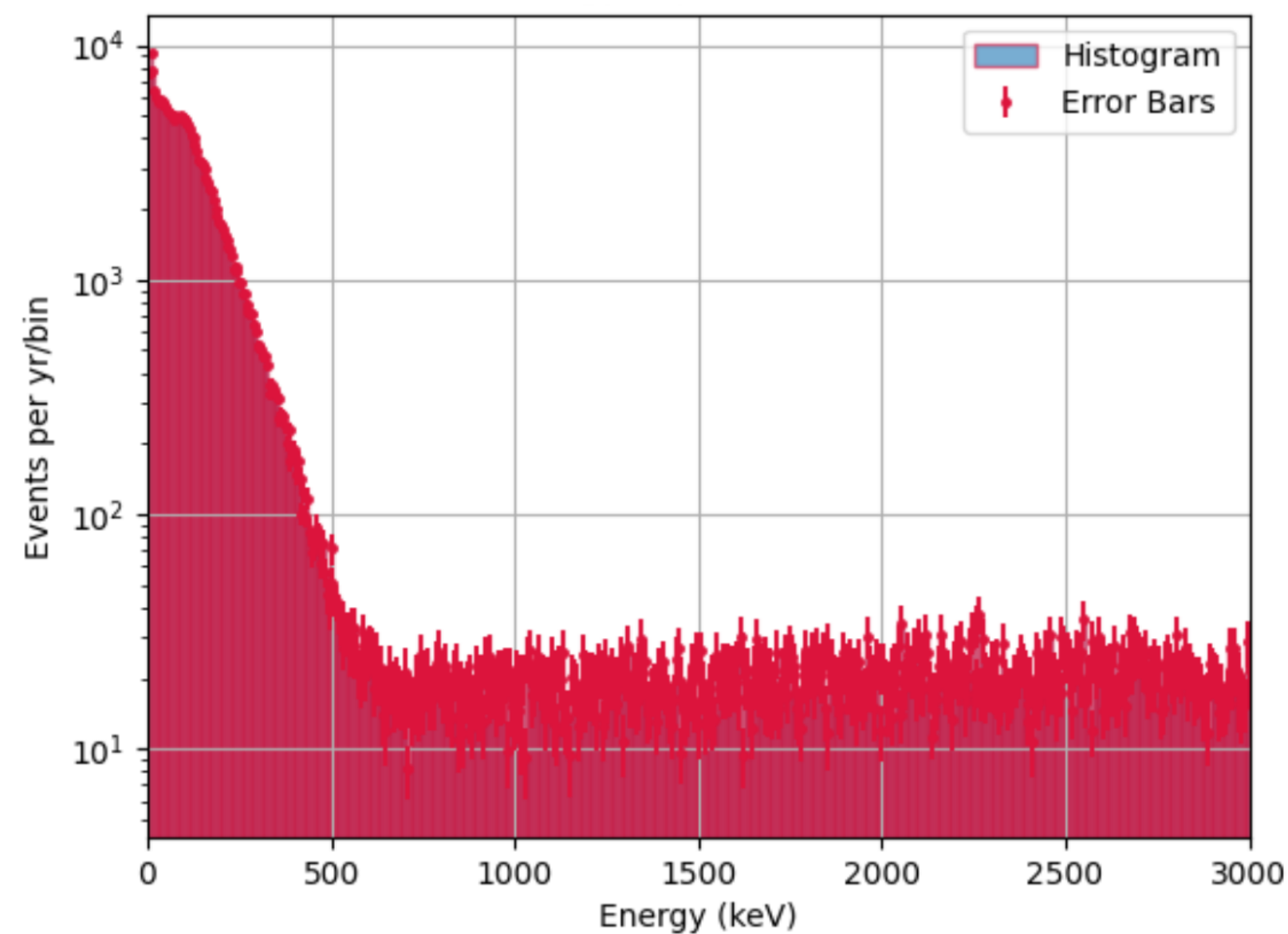
- Three different field cage foils:
  - > **PET+Cu with glue**: a polyethylene foil with Cu strips glued by spreading the adhesive over the entire foil
  - > **PET+Cu with no glue**: a polyethylene foil with Cu strips glued by applying the adhesive only between the strips and the foil
  - > **Kapton+Cu**: a kapton foil with Cu strips
- The foils were measured at LNGS



Donatella Tozzi - Cygno CM 2024

2

Sample Name	in measure	measured	analysed
FC-kapton+Cu	no	yes	yes
FC-PET+Cu	no	yes	yes
FC-PET+Cu no glue	no	yes	yes
nylon6	no	yes	on-going
GEM foil	no	yes	on-going
nylon screws	no	yes	on-going
steel screws	no	yes	on-going
SMD resistors	no	yes	on-going
CMOS camera	no	yes	yes
PMMA	yes	no	no



In the low energy range (1-20 keV):

- $(1.56 \pm 0.01) 10^5$  evt/year;
- $(1.20 \pm 0.08) 10^3$  NR evt/year;

Done!



# Component handling procedures



## Field Cage fo

- Ultrasonic bath with 200g/l
- Washing under running demineralized acid.
- Clean bath vessel and remove
- Ultrasonic bath with demineralized
- Degas mode operation for 5
- Drying with nitrogen under
- Put on granite table and set
- Operate the vacuum bag fo

## GEM Cleaning

- Ultrasonic bath with demineralised
- Washing under running demineralized contaminant.
- Drying
- Further drying with clean chamber cloth

## GEM Testing

### 1. Visual Inspection

- Mechanical damages (tears, wrinkles)
- Dust or irregularities on the surface
- High resistance test with Megger insu

### 2. Leakage Current test af

- High resistance test with Megger insu
- Place the GEM foil in a nitrogen flux
- Supply to high-voltage up to 600V
- Measure the current with a picoamm
- Acceptable current is usually in the n

## Copper Cleaning

- Ultrasonic bath with 1% acid soap (Elma clean 60, or other soaps acidic.....) for at least 20 minutes.
- Washing with 10% HNO<sub>3</sub> and 1% H<sub>2</sub>O<sub>2</sub> solution for 1 hour (takes off about 7/10 um). To shorten the time we can increase H<sub>2</sub>O<sub>2</sub>
- The use nitric acid, because citric acid needs the addition of a lot of H<sub>2</sub>O<sub>2</sub> to make it work;
- Washing with 1% citric acid solution and 0.5% H<sub>2</sub>O<sub>2</sub> for 15 seconds;
- Bath with demineralized water for 15 seconds.
- Washing under running (demineralized) water to remove the remaining traces of acid.
- Drying with clean room cloth
- Further drying with clean chamber cloth and nitrogen flow under laminar flow hood

### Sample list

- Cu piece CSN CARL SCHREIBER

Sample "Cu L"	Weight [g]	Sample treatment	Cu dissolved [g]	Note
Starting	14.10			
After Etching n 1	11.05	6 mL H <sub>2</sub> O + 10 mL HNO <sub>3</sub>	3.05	Waste
After Etching n 2	8.13	6 mL H <sub>2</sub> O + 10 mL HNO <sub>3</sub>	2.92	Measured
After Etching n 3	5.20	6 mL H <sub>2</sub> O + 10 mL HNO <sub>3</sub>	2.93	Measured
After Etching n 4	2.13	6 mL H <sub>2</sub> O + 10 mL HNO <sub>3</sub>	3.07	Spiked 100 ppt

Tab.1 Sample etching with HNO<sub>3</sub>

### Rinse and description

Cu sample was rinsed with 5% of acid soap "Decon" in ultra-sonic bath, nitric acid and citric acid, this procedure was performed by Roberto Cerroni together with samples cleaning for gamma-ray spectroscopy.

- CYGNO: Procedures for material handling and cleaning

- 1

- CYGNO: Procedures for material handling and cleaning

- 2

- CYGNO: Procedures for material handling and cleaning

- 3

We are collecting all procedures materials and parts cleaning, handling, testing and quality assessment in a shared document



# Component Validation - Copper Shield

## Copper analysis with ICPMS

- They used the described procedure to clean it and measured the U and Th content

	Etching 2	Etching 3
	[pg * g <sup>-1</sup> ]	[pg * g <sup>-1</sup> ]
Th	9 ± 3	7 ± 2
U	5 ± 2	2 ± 1

- These were the Matthias results

radionuclide concentrations:

Th-232:		
Ra-228:	< 0.38 mBq/kg	<=> < 9.3 E-11 g/g
Th-228:	< 0.20 mBq/kg	<=> < 4.9 E-11 g/g
U-238:		
Ra-226:	< 0.44 mBq/kg	<=> < 3.5 E-11 g/g
Th-234:	< 17 mBq/kg	<=> < 9.3 E-10 g/g
Pa-234m:	< 11 mBq/kg	<=> < 6.5 E-10 g/g
U-235:	< 0.37 mBq/kg	<=> < 6.5 E-10 g/g
K-40:	< 3.2 mBq/kg	<=> < 1.0 E-7 g/g
Cs-137:	< 0.14 mBq/kg	
Co-60:	< 0.12 mBq/kg	
Co-58:	(0.8 +- 0.1) mBq/kg	
Mn-54:	(0.12 +- 0.05) mBq/kg	

<sup>58</sup>Co has an half life of 70 days

<sup>54</sup>Mn has an half life of 1 year

	October
Layer_0 (1-20 keV) [evts/yr]	14655 ± 347
Layer_1 (1-20 keV) [evts/yr]	4518 ± 176
Layer_2 (1-20 keV) [evts/yr]	10588 ± 2671
<b>TOTAL</b>	<b>29761 ± 2700</b>

Significant **improvement** thanks to ICP-MS measurement of Schrieber's copper

**Huge statistics** required to see first effects of Bi210 in OPERA's copper  
→ **more storage space needed**

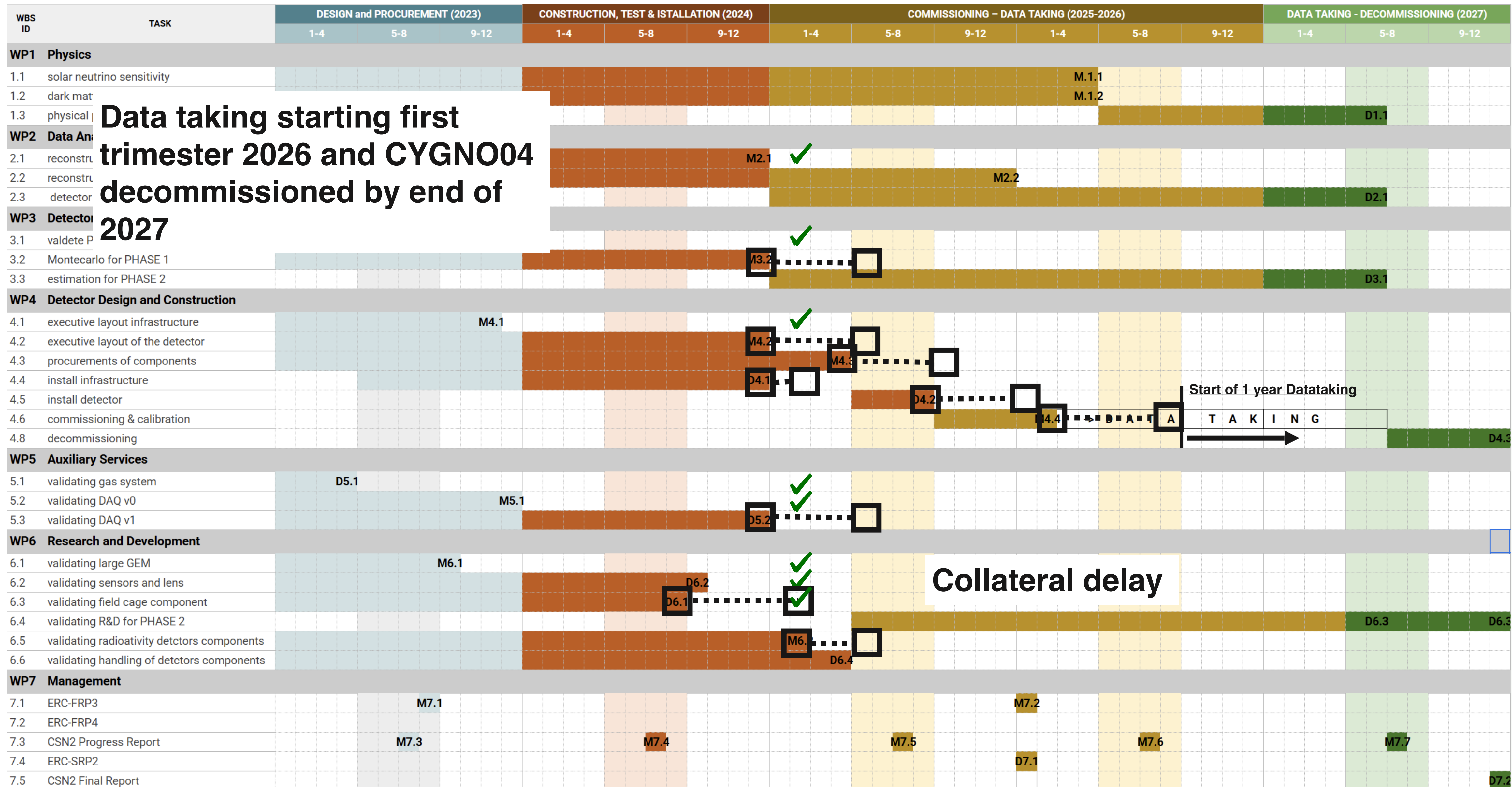
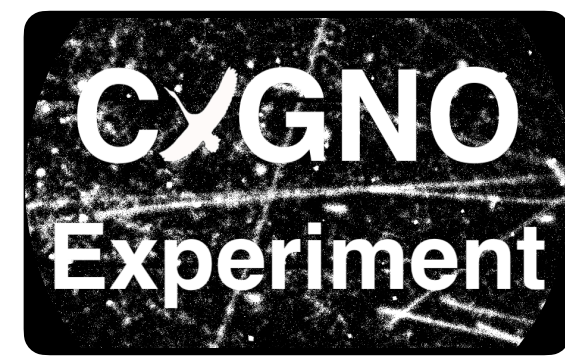
In light of the updated results, the most viable option is still **4 cm clean copper + 6 cm OPERA**

Upper limits on U and Th obtained with Germanium detector, 10 times larger than actual values

**Field Cage** and **GEM** foils are now on queue for the ICPMS measurements



# CYGNO04 GANTT Up to date



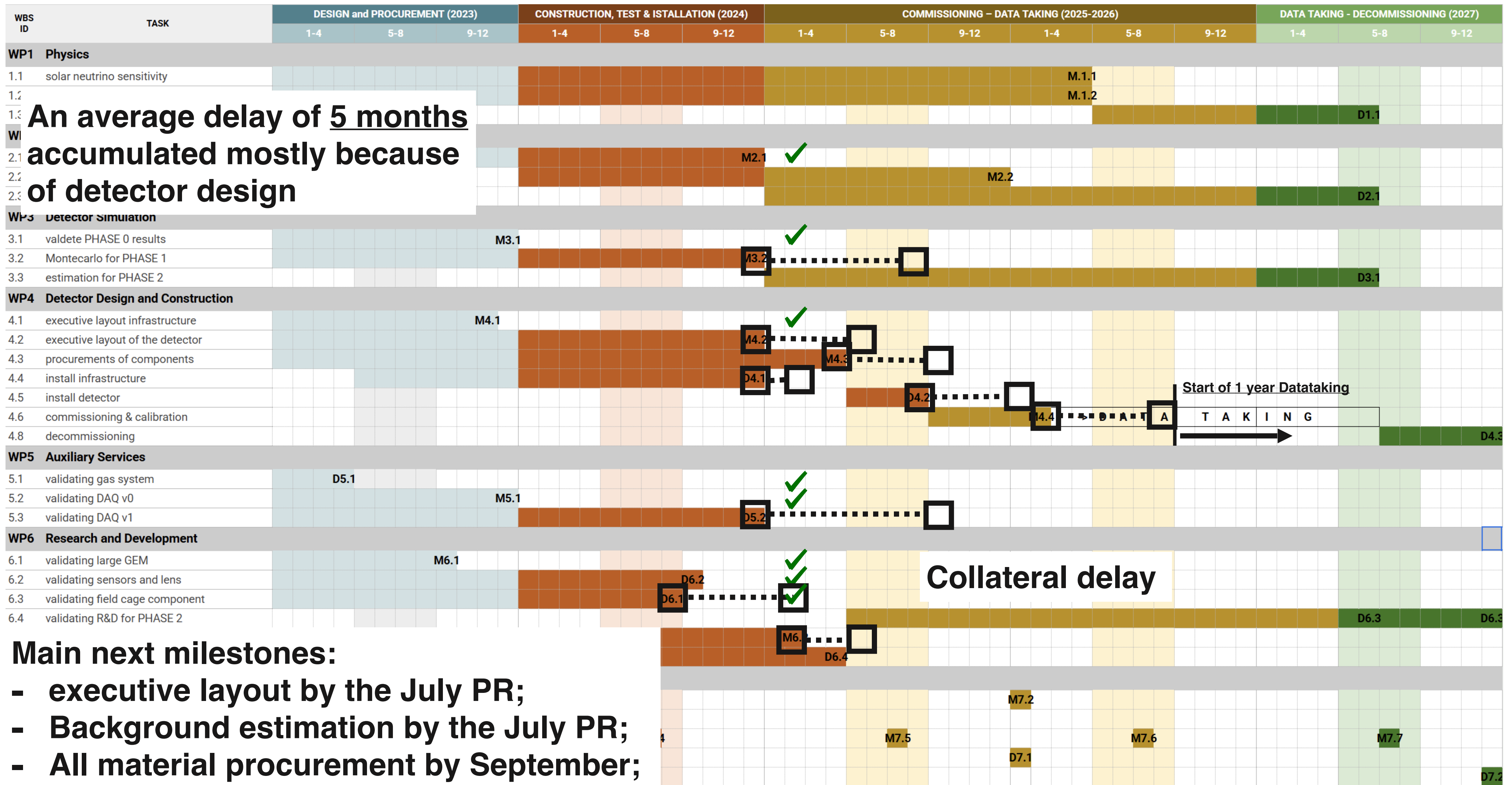
**Data taking starting first trimester 2026 and CYGNO04 decommissioned by end of 2027**

Start of 1 year Datataking

Collateral delay



# CYGNO04 GANTT Up to date



**An average delay of 5 months accumulated mostly because of detector design**

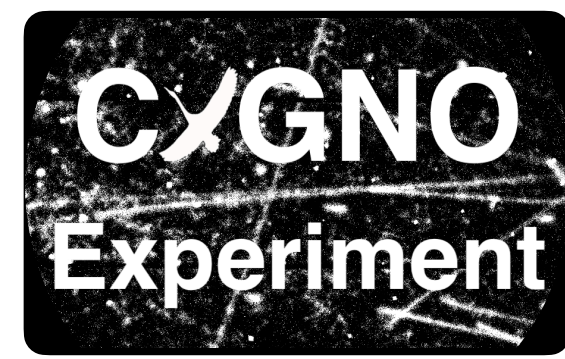
**Collateral delay**

## Main next milestones:

- executive layout by the July PR;
- Background estimation by the July PR;
- All material procurement by September;



# TDR CYGNO financial plan and CSN2 proposal



Financial profile for ERC and INFN  
from the **2022 TDR**

Year	INITIUM/ERC	CYGNO/INFN
2019	20	54
2020	201	44
2021	71	96
2022	40	96
2023	374	120
2024	302	125
2025	60	135
2026	0	95
2027	0	50
Tot 23-27	736	525
Tot	<b>1068</b>	<b>815</b>

The new situation had very negligible effects on the financial plans:

A total of **490 k€** are expected to being funded by INFN for the quinquennium 2023-2027 (**515 k€** foreseen in the **TDR**)

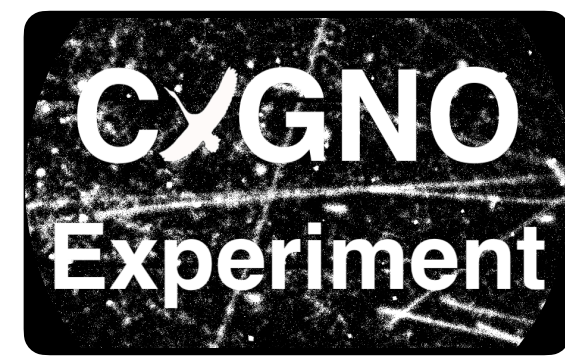
Financial profile for ERC and INFN  
from the up to date

Year	INITIUM/ERC	CYGNO/INFN
2019	20	54
2020	201	44
2021	71	96
2022	40	96
2023	164	93.5
2024	359	91.5
2025	276	145
2026	0	80
2027	0	80
Tot 23-27	799	490
Tot	<b>1131</b>	<b>780</b>

- In particular **305 k€** foreseen by INFN in 25-27;
- A total of **1131 k€** are expected to being funded by ERC for the quinquennium 2023-2027 (**1068 k€** foreseen in the **TDR**)



# CYGNO04 Economic sustainability



Currently, from the ERC fundings, to cover the core costs of CYGNO04, there are a total of **300 k€** available at **INFN** and **GSSI**

Part of the equipment is being used and validated on LIME:

- high voltage system;
- gas system;
- DAQ and trigger;

According to the latest quotations received, the **other costs** expected for the **construction** of CYGNO04 is of **276 k€**, therefore an expense that can **be fully covered with ERC funds**.

**Deadline for finalizing purchases: August 2025**

The **total value** of **CYGNO04** is of about **940 k€**.

GAS VOLUME	Cone and GEM holders	5	1	5	0	1	5	35		
	Vessel	30	1	30	0	1	30			
READOUT	4-GEM Set	10	2	20	2	0	0	14		
	Orca Quest	26	6	156	6	0	0			
	Camera mechanics	1	6	6	0	6	6			
	Lenses	2	6	12	2	4	8			
	PMT+Socket	1	16	16	16	0	0			
CATHODE	Frame+foil+feedthrough	17	1	17	0	1	17	17		
FIELD CAGE	Field Cage	13	1	13	0	1	13	13		
CALIBRATION SYSTEM	Krypton-Rubidium	6	1	6	0	1	6	8		
	55-Fe system	2	1	2	0	1	2			
SHIELDING	4 ton Copper Elsasser	176	1	176	1	0	0	165		
	copper precision machining	10	1	10	0	1	10			
	Opera copper refurbishing	30	1	30	0	1	30			
	High Voltage feed through	10	1	10	0	1	10			
	Optical windows	10	1	10	0	1	10			
	water tanks	65	1	65	0	1	65			
	frame	20	1	20	0	1	20			
polietilene	20	1	20	0	1	20				
GAS SYSTEM	gas system	85	1	85	1	0	0	0		
	filters	1	8	8	8	0	0			
HV SYSTEM	GEM-HV	5	1	5	1	0	0	0		
	PMT-HV	3	2	6	2	0	0			
	Cathode HV	5	1	5	1	0	0			
DAQ & SLOW CONTROLS	modules	45.5	1	45.5	1	0	0	0		
COMPRESSED AIR SYSTEM	filters and pipes	5	1	5	0	1	5	5		
ELECTRIC SERVICES	power distribution system	10	1	10	1	0	0	0		
	cables & connectors	5	1	5	1	0	0			
	UPS	5	1	5	1	0	0			
NETWORK DISTRIBUTION	network distribution system	13	1	13	1	0	0	0		
COOLING and CONDITIONING	Conditioning	20	1	20	1	0	0	4		
	Cameras chiller	5	1	5	1	0	0			
	PMTs flow system	2	1	2	0	1	2			
	Cables & connectors	2	1	2	0	1	2			
SAFETY (PRA-VIA)	pra+vinca	6	1	6	1	0	0	0		
	fire detection	10	1	10	1	0	0			
	gas monitor	15	1	15	1	0	0			
	safety design	10	1	10	1	0	0			
CIVIL WORK	total	40	1	40	1	0	0	0		
DESIGN AND DOCUMENTATION	software	5	1	5	0	1	5	15		
	final audit	10	1	10	0	1	10			
<b>Gran Total</b>								<b>941.5</b>	<b>276</b>	<b>276</b>



# Papers in last year



## 1 Secondary scintillation yield from GEM electron avalanches in He-CF<sub>4</sub> and He-CF<sub>4</sub>-CH<sub>4</sub> mixtures for CYGNO – Directional Dark Matter Search with an optical TPC

F.D. Amaro,<sup>a</sup> E. Baracchini,<sup>d,e</sup> S. Bianco,<sup>d</sup> C. Capoccia,<sup>d</sup> M. Caponero,<sup>d,e</sup> D.S. Carrone,<sup>d</sup> G. Cavoto,<sup>g,h</sup> A. Cortese,<sup>b,c</sup> L.A. Costa,<sup>f</sup> E. Dané,<sup>f</sup> G. Dho,<sup>b,c</sup> F. Di Giambattista,<sup>d,e</sup> E. Di Marco,<sup>h</sup> G. D'Imperio,<sup>i</sup> P. Lima Júnior,<sup>e</sup> M. Maccarrone,<sup>d</sup> R.D.P. Mano,<sup>d</sup> G. Mazzitelli,<sup>d</sup> A.G. Mclennan,<sup>b,c</sup> G. Nóbrega,<sup>i</sup> I.F. Pains,<sup>i</sup> E. Paoletti,<sup>g</sup> S. Piantoni,<sup>g,h</sup> D. Piccolo,<sup>d</sup> Pierluigi Ronga,<sup>g</sup> R.J.C. Roque,<sup>a,2</sup> F. Rosatelli,<sup>d</sup> A. Russo,<sup>d</sup> S. Santilli,<sup>b,c</sup> N.J.C. Spooner,<sup>k</sup> R. Tesaro,<sup>d</sup> S. Tomassini,<sup>d</sup> and D. Tozzi<sup>7,8</sup>

<sup>a</sup> LIBPhys, Department of Physics, University of Coimbra, 3004-516 Coimbra, Portugal  
<sup>b</sup> Gran Sasso Science Institute

## He-CF<sub>4</sub>-CH<sub>4</sub> ternary mixtures as target gas for the CYGNO directional dark matter experiment

R. J. da C. Roque, R. D. P. Mano, C. M. B. Monteiro (†) and CYGNO authorlist

**Abstract**  
The CYGNO collaboration is developing a high-resolution optical Time Projection Chamber for directional dark matter neutrino spectroscopy, to be deployed at LNGS. He-40% CF<sub>4</sub> is being used as target gas, at atmospheric pressure, and a triple-GEM cascade is being used for the ionization signal, being the scintillation produced in the GEM readout by sCMOS cameras. High sensitivity to interactions in the few keV range for event tracking and particle identification are enabled with this solution. The addition of 3 to 10 percent of methane to He-40%CF<sub>4</sub> and demonstrate methane inclusion contributes to both the electrical stability of the TPC and to higher scintillation output from the GEM avalanches. In spite of the scintillation yield, the addition of methane increases the maximum voltage that can be applied before the onset of discharges, eventually resulting in higher scintillation yield. The visible component of the gas scintillation shows that isobutane photons emitted by He-CF<sub>4</sub>. The maximum voltage that could be applied to be independent from the methane concentration in the scintillation gas. A hydrogen-based gas provides CYGNO with an even lighter target for dark matter detection threshold, in longer track lengths of light nuclear recoils and, thus, in a clearer direction discrimination.

Davide Pinci INFN - Roma

## 2 Charge Amplification in Low Pressure CF<sub>4</sub>/SF<sub>6</sub>:He Mixtures with a Multi-Mesh ThGEM for Directional Dark Matter Searches

F.D. Amaro,<sup>a</sup> R. Antonietti,<sup>b,c</sup> E. Baracchini,<sup>d,e</sup> F. Borra,<sup>g,h</sup> C. Capoccia,<sup>f</sup> M. Caponero,<sup>f,i</sup> D.S. Carrone,<sup>f</sup> L.A. Costa,<sup>b,c</sup> E. Dané,<sup>f</sup> G. Dho,<sup>d,e</sup> F. Di Giambattista,<sup>d,e</sup> E. Di Marco,<sup>h</sup> G. D'Imperio,<sup>i</sup> H.P. Lima Júnior,<sup>k</sup> E. Kemp,<sup>l,d</sup> G.S.P. Lopes,<sup>m</sup> G. Maccarrone,<sup>d</sup> M. Marcello Gregorio,<sup>n</sup> D.J.G. Marques,<sup>d</sup> G. Mazzitelli,<sup>d</sup> A.G. Mclennan,<sup>b,c</sup> S. Messina,<sup>g,h</sup> C.M.B. Monteiro,<sup>a</sup> R.A. Nobrega,<sup>m</sup> I.F. Pains,<sup>i</sup> E. Paoletti,<sup>g</sup> S. Pelosi,<sup>h</sup> F. Petrucci,<sup>b,c</sup> S. Piacentini,<sup>g,h</sup> D. Piccolo,<sup>f</sup> D. Pierluigi,<sup>g</sup> F. Ronga,<sup>g</sup> R.J.C. Roque,<sup>a</sup> F. Rosatelli,<sup>d</sup> A. Russo,<sup>d</sup> J.M.F. Santos,<sup>o</sup> N.J.C. Spooner,<sup>k</sup> R. Tesaro,<sup>d</sup> S. Tomassini,<sup>d</sup> and D. Tozzi<sup>7,8</sup>

## Modeling the detector response of the CYGNO optical readout TPC

F.D. Amaro,<sup>a</sup> R. Antonietti,<sup>b,c</sup> E. Baracchini,<sup>d,e</sup> F. Borra,<sup>g,h</sup> C. Capoccia,<sup>f</sup> M. Caponero,<sup>f,i</sup> D.S. Carrone,<sup>f</sup> L.A. Costa,<sup>b,c</sup> E. Dané,<sup>f</sup> G. Dho,<sup>d,e</sup> F. Di Giambattista,<sup>d,e</sup> E. Di Marco,<sup>h</sup> G. D'Imperio,<sup>i</sup> H.P. Lima Júnior,<sup>k</sup> E. Kemp,<sup>l,d</sup> G.S.P. Lopes,<sup>m</sup> G. Maccarrone,<sup>d</sup> M. Marcello Gregorio,<sup>n</sup> D.J.G. Marques,<sup>d</sup> G. Mazzitelli,<sup>d</sup> A.G. Mclennan,<sup>b,c</sup> S. Messina,<sup>g,h</sup> C.M.B. Monteiro,<sup>a</sup> R.A. Nobrega,<sup>m</sup> I.F. Pains,<sup>i</sup> E. Paoletti,<sup>g</sup> S. Pelosi,<sup>h</sup> F. Petrucci,<sup>b,c</sup> S. Piacentini,<sup>g,h</sup> D. Piccolo,<sup>f</sup> D. Pierluigi,<sup>g</sup> F. Ronga,<sup>g</sup> R.J.C. Roque,<sup>a</sup> F. Rosatelli,<sup>d</sup> A. Russo,<sup>d</sup> J.M.F. Santos,<sup>o</sup> N.J.C. Spooner,<sup>k</sup> R. Tesaro,<sup>d</sup> S. Tomassini,<sup>d</sup> and D. Tozzi<sup>7,8</sup>

<sup>a</sup>Department of Physics, University of Coimbra, 3004-516 Coimbra, Portugal  
<sup>b</sup>Departmento de Matemática e Física, Universidade Roma TRE, 00146, Roma, Italy  
<sup>c</sup>Istituto Nazionale di Fisica Nucleare, Sezione di Roma Tre, 00146, Roma, Italy  
<sup>d</sup>Gran Sasso Science Institute, 67100, L'Aquila, Italy  
<sup>e</sup>Istituto Nazionale di Fisica Nucleare, Laboratori Nazionali del Gran Sasso, 67100, Assergi, Italy

## Enhancing the light yield of He:CF<sub>4</sub> based gaseous detector

Fernando Domingues Amaro,<sup>1</sup> Rita Antonietti,<sup>2,3</sup> Elisabetta Baracchini,<sup>4,5</sup> Luigi Benussi,<sup>6</sup> Stefano Bianco,<sup>6</sup> Roberto Campagnola,<sup>6</sup> Cesidio Capoccia,<sup>6</sup> Michele Caponero,<sup>6</sup> Danilo Santos Cardoso,<sup>8</sup> Luan Gomes Mattosinho,<sup>9</sup> Igor Abritta Costa,<sup>10,11</sup> Emanuele Dané,<sup>6</sup> Giorgio Dho,<sup>4,6,a</sup> Flaminia Di Giambattista,<sup>4,5</sup> Joaquim Marques Ferreira dos Santos,<sup>1</sup> Herman Pessoa Lima,<sup>7</sup> R.D.P. Mano,<sup>1</sup> G. Mazzitelli,<sup>4,5</sup> I.F. Pains,<sup>1</sup> E. Paoletti,<sup>6</sup> Luciano Passarini,<sup>12</sup> Stefano Piacentini,<sup>4,5</sup> Davide Piccolo,<sup>13</sup> Atul Prajapati,<sup>4,5</sup> Francesco Rosatelli,<sup>14</sup> Alessandro Russo,<sup>1</sup> Joaquim Marques Ferreira dos Santos,<sup>1</sup> Giovanna Saviano,<sup>6,14</sup> Pedro Alberto Oliveira Costa Silva,<sup>1</sup> Neil John Curwen Spooner,<sup>13</sup> Roberto Tesaro,<sup>6</sup> Sandro Tomassini,<sup>6</sup> Samuele Torelli,<sup>4,5</sup> and Donatella Tozzi<sup>15</sup>

## An analytical model of the response of the optically readout GEM based TPC for the CYGNO experiment

Rita Antonietti<sup>1,2,3</sup>, Elisabetta Baracchini<sup>4,5</sup>, Luigi Benussi<sup>6</sup>, Stefano Bianco<sup>6</sup>, Francesco Borra<sup>7,8</sup>, Cesidio Capoccia<sup>6</sup>, Michele Caponero<sup>6</sup>, Danilo Santos Cardoso<sup>9</sup>, Gianluigi Dho<sup>4,5</sup>, Flaminia Di Giambattista<sup>4,5</sup>, Emanuele Di Marco<sup>10,11</sup>, Joaquim Marques Ferreira dos Santos<sup>1</sup>, Herman Pessoa Lima<sup>7</sup>, R.D.P. Mano<sup>1</sup>, G. Mazzitelli<sup>4,5</sup>, I.F. Pains<sup>1</sup>, E. Paoletti<sup>6</sup>, Luciano Passarini<sup>12</sup>, Stefano Piacentini<sup>4,5</sup>, Davide Piccolo<sup>13</sup>, Atul Prajapati<sup>4,5</sup>, Francesco Rosatelli<sup>14</sup>, Alessandro Russo<sup>1</sup>, Joaquim Marques Ferreira dos Santos<sup>1</sup>, Giovanna Saviano<sup>6,14</sup>, Pedro Alberto Oliveira Costa Silva<sup>1</sup>, Neil John Curwen Spooner<sup>13</sup>, Roberto Tesaro<sup>6</sup>, Sandro Tomassini<sup>6</sup>, Samuele Torelli<sup>4,5</sup>, and Donatella Tozzi<sup>15</sup>





**THANKS!**