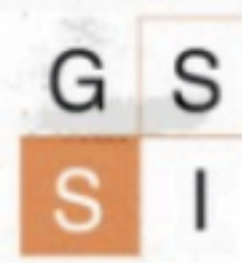




CYGNO status and plans Feb 2025



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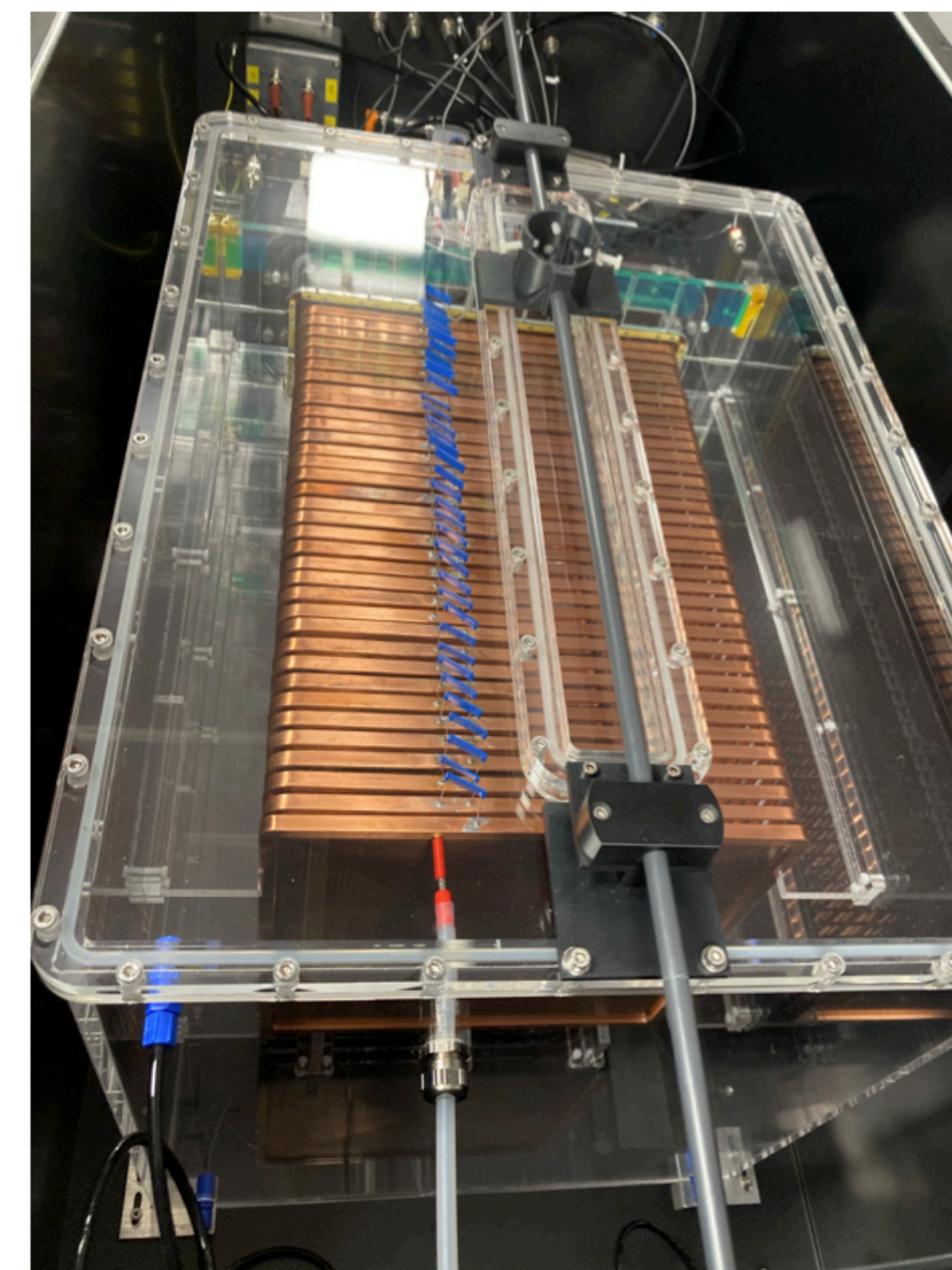
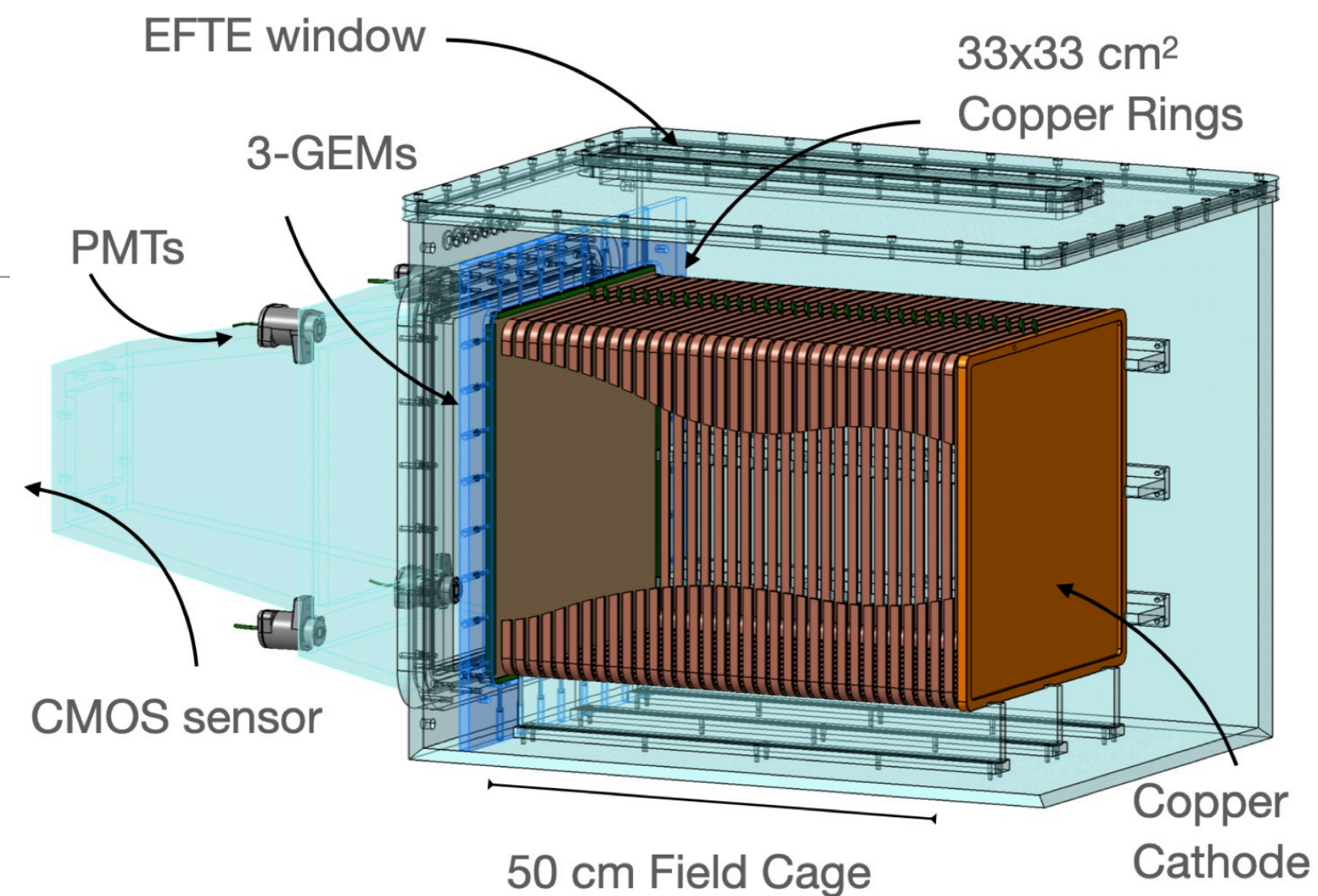
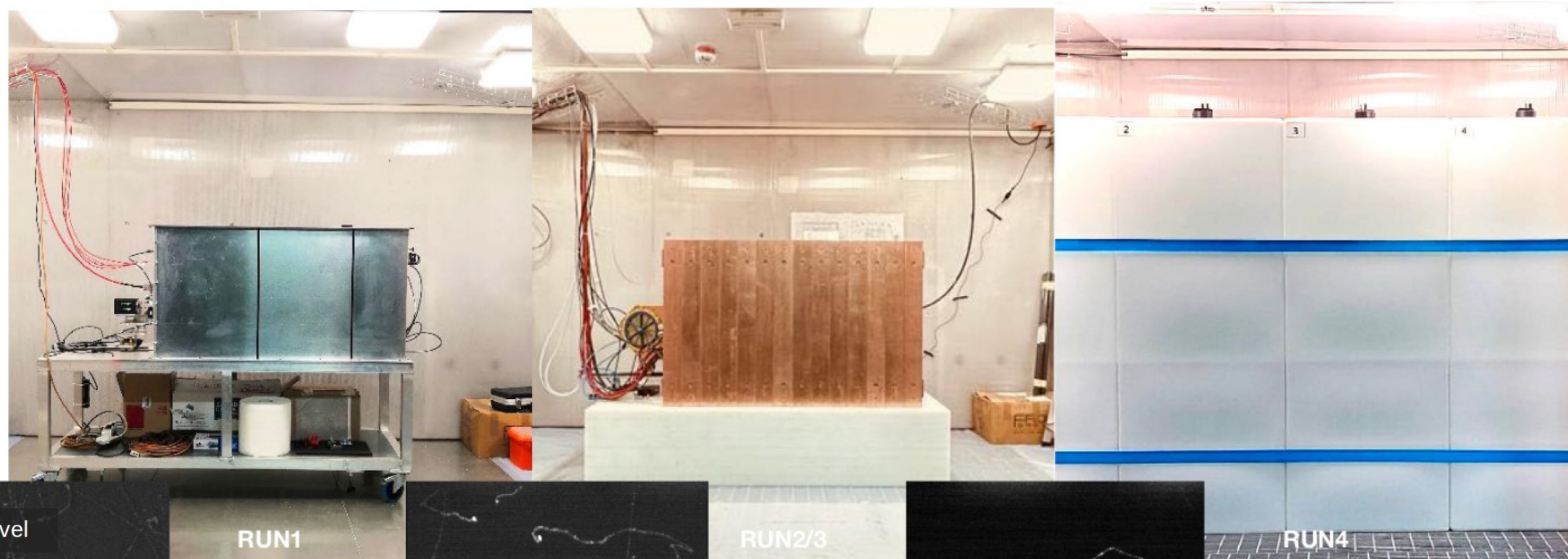
Collaboration Meeting 2024



Three days of discussion about status and plans of:

- LIME Data **acquisition** and **analysis**;
- CYGNO04 **drawing** and **simulation**;
- Component **validation**;
- Project management and **financial status**;

LIME: RUNs 1-5



- testing and optimisation of **ancillary systems** (HV, Gas, Slow Control, DAQ, Reco, Monitoring);
- development and optimisations of data **analysis algorithms** and performance study;
- development and validation of **detector** and **background simulation**;

LIME: RUNs 1-5



	Time slot	Number of pictures	Event rate	Number of events
RUN 1: No-shielding	3 Nov 2023 - 15 Dec 2023	$4 \cdot 10^5$	35 Hz	$4 \cdot 10^6$
RUN 2: 4 cm Cu shielding	15 Feb 2023 - 15 March 2023	$4.5 \cdot 10^5$	3.5 Hz	$5 \cdot 10^5$
RUN 3: 10 cm Cu shielding	5 May 2023 - 16 Nov 2023	$1.6 \cdot 10^6$	1.5 Hz	$7.3 \cdot 10^5$
RUN 4: 10 cm Cu + 40 cm water shielding	30 Nov 2023 - 31 March 2024	$2 \cdot 10^6$	1.0 Hz	$6 \cdot 10^5$
RUN 5: 10 cm Cu shielding (neutron flux measurements)	17 May 2024 - 1 Dec 2024	$12 \cdot 10^6$	1.5 Hz	$5.4 \cdot 10^6$

Special data takings

AmBe for Nuclear Recoils	2-4 Aug 2023	$2 \cdot 10^5$	0.04 Hz of NR	$2.5 \cdot 10^3$ NR
^{241}Am for Electron Recoils	7-16 Nov 2023	$7 \cdot 10^5$	50 Hz	10^6
AmBe for Nuclear Recoils	5-15 Dec 2024	$6 \cdot 10^5$	0.04 Hz of NR	$7.0 \cdot 10^3$ NR

LIME PLANS



With the **end of RUN4** in March 2024, LIME data taking for **PHASE_0** is considered closed;

The whole schedule successfully concluded with **3 months delay** (including the delayed start due to civil works) in over **27 months of operations**;

The water shielding was then removed and in **May 2024** and **RUN5** has been taken for 6 months to measure the **flux of underground neutrons**, in the framework of **PRIN “Zero Radioactivity on future experiments”**;

Good exercise to **tune and test simulation and analysis algorithms** for **NR**, propaedeutical to DM search;

In **December 2024** a 10 days long **AmBe campaign** was performed to have a good calibration on signal-like events;

From **January 2025** we are taking some **“technical runs”** to study the effects of **gas filters**, a **new camera**, and **calibration** with ^{85}Kr

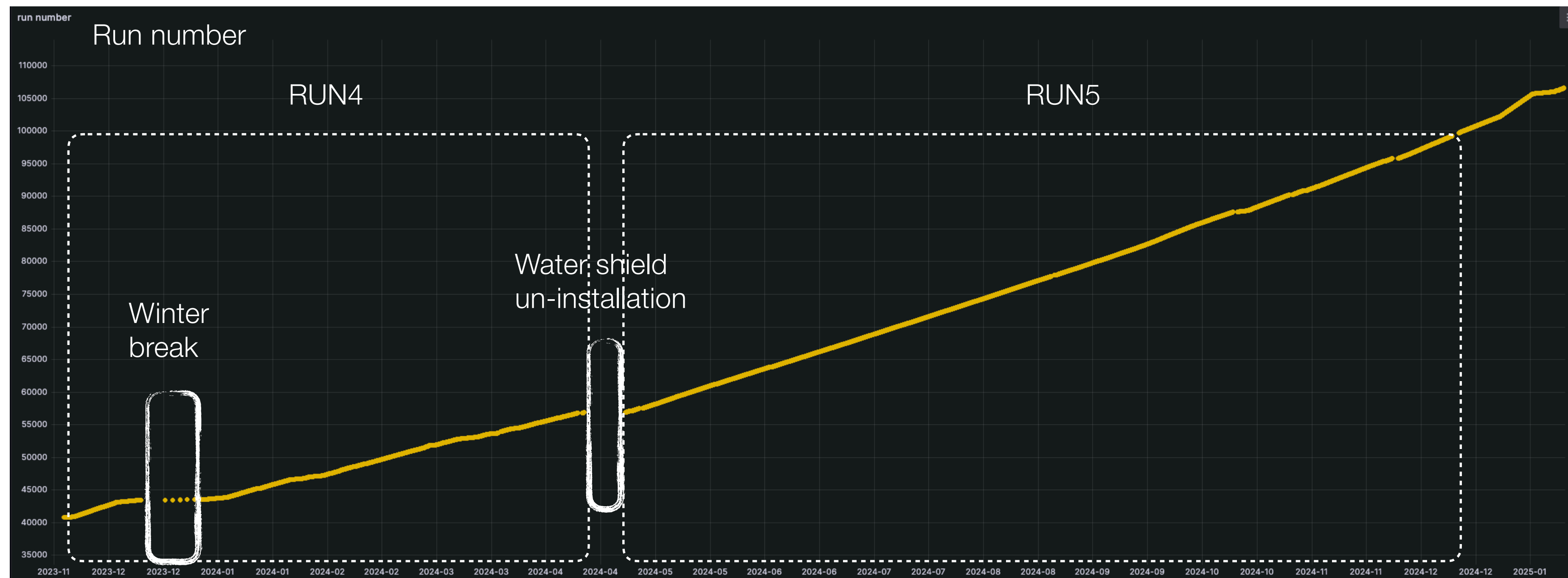
LIME DATA TAKING SUMMARY

Ancillary systems: DAQ and Slow control



Data taking started in October 2023 never stopped;

More than **7×10^4 runs** have been taken for a total of **28×10^6 pictures**;



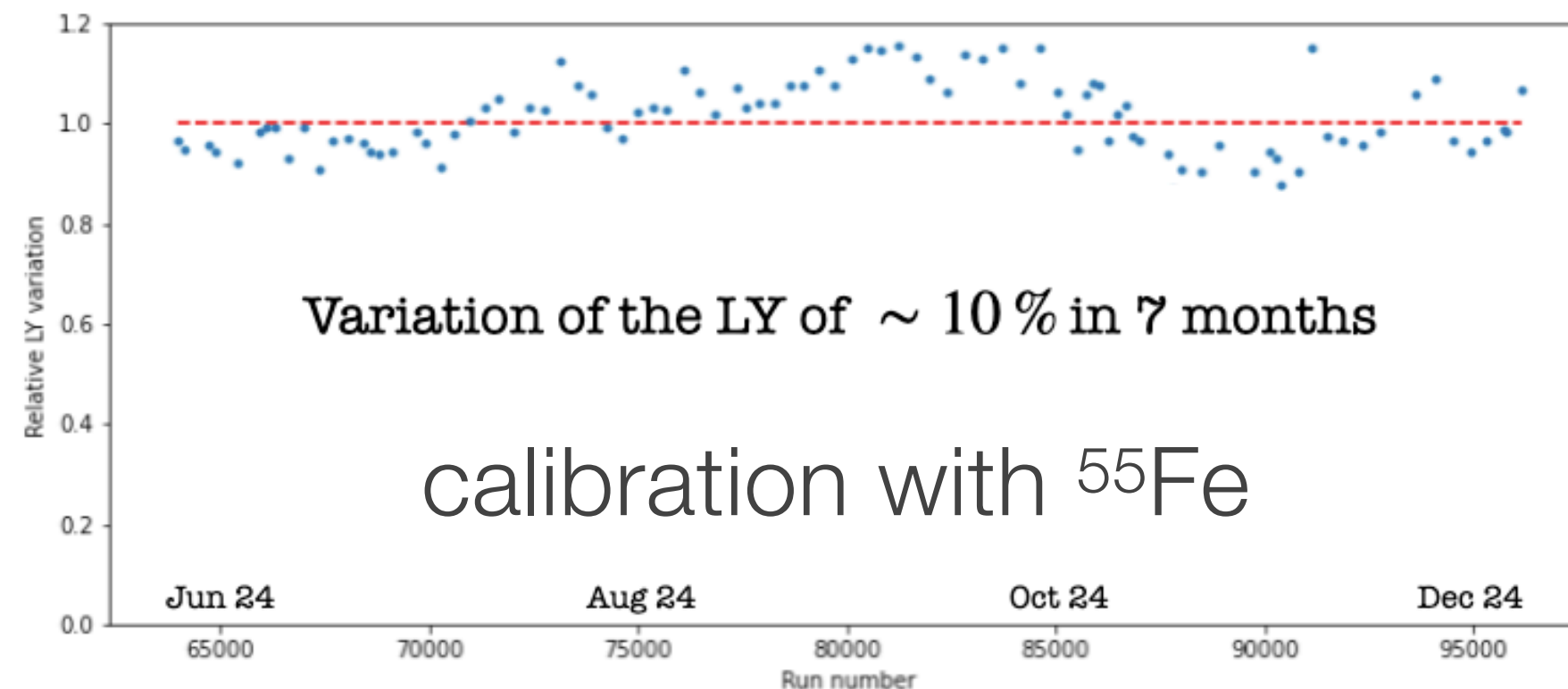
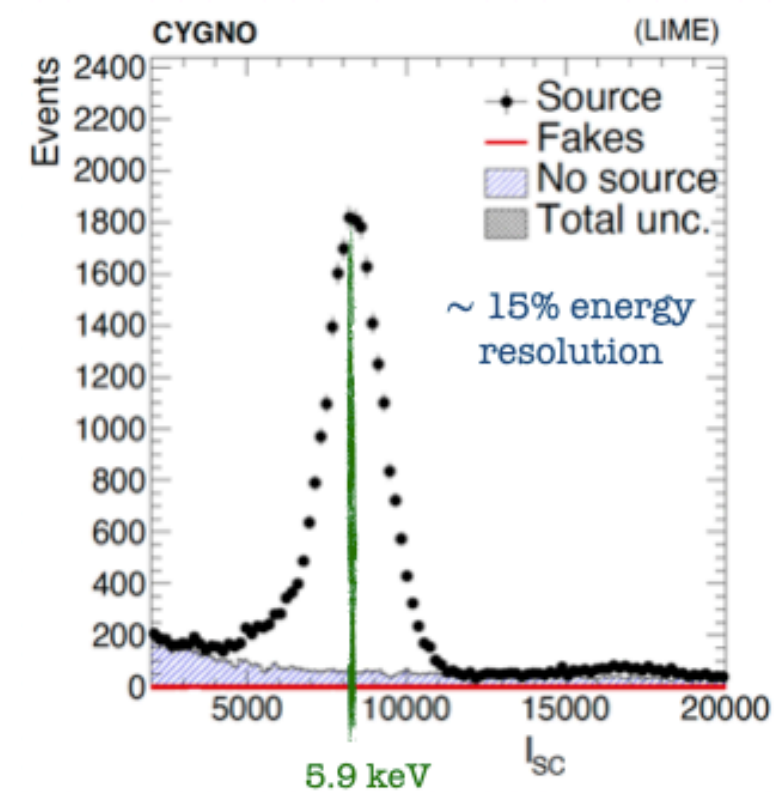
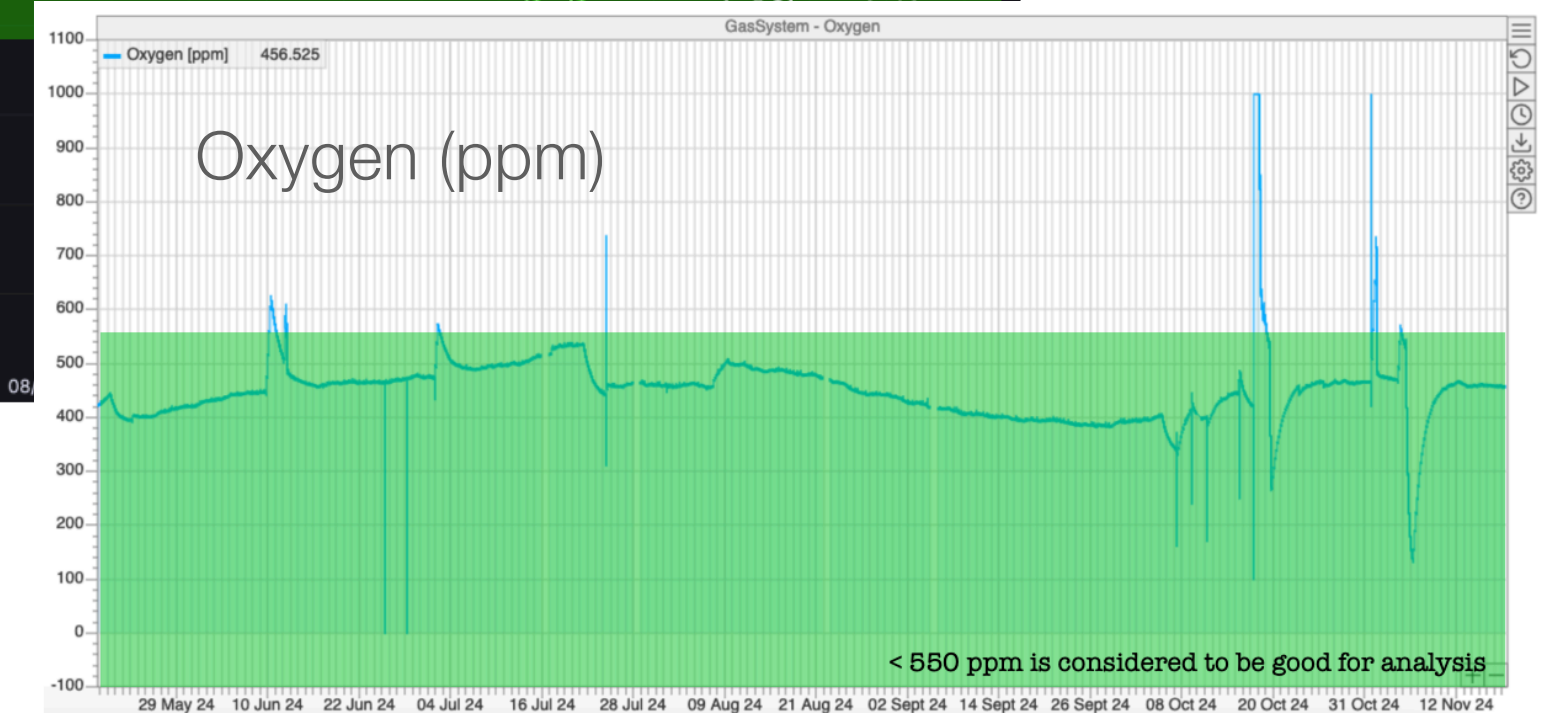
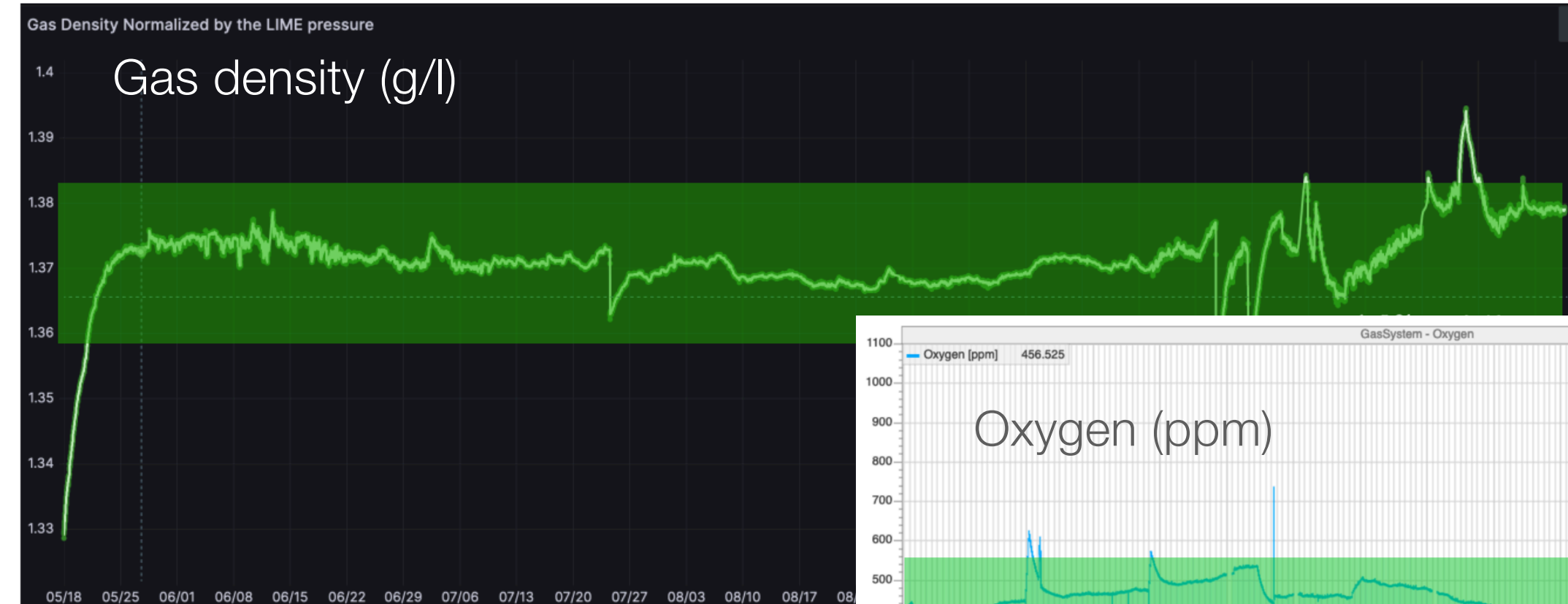
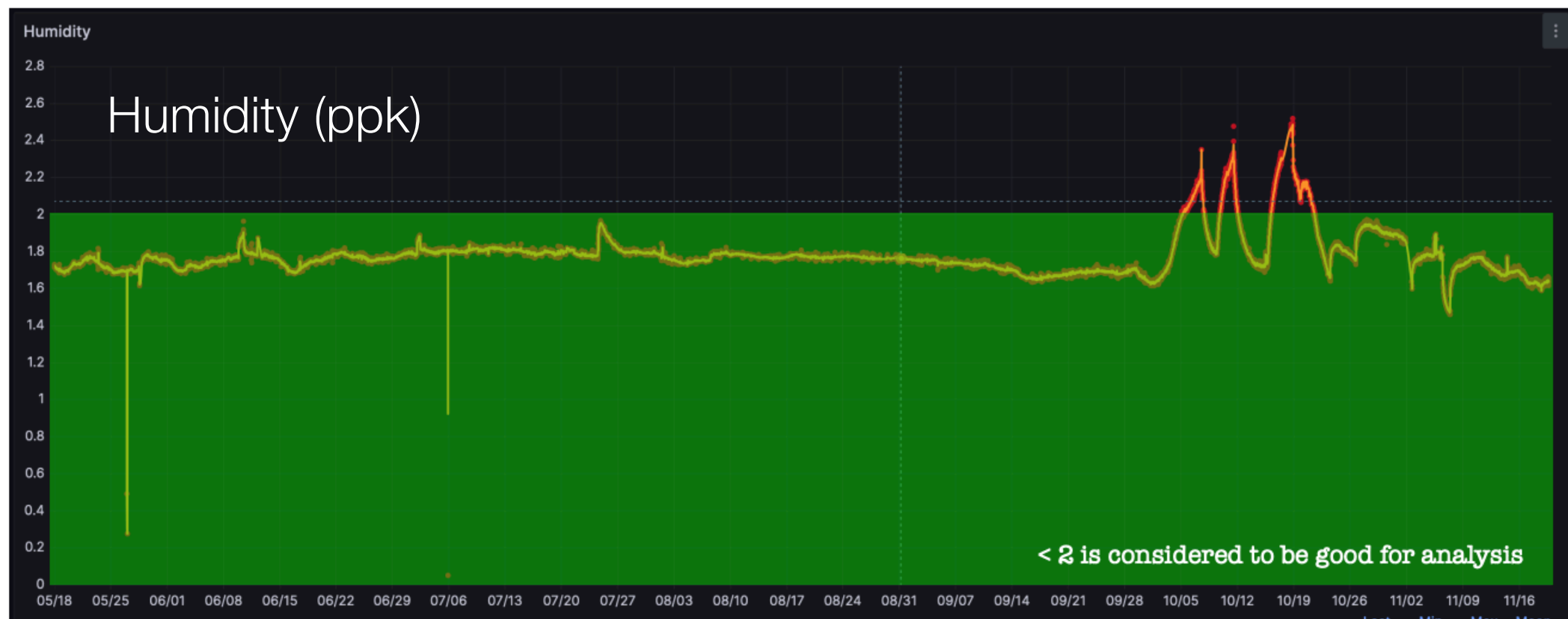
All of them have been **promptly transferred** on the INFN-Cloud and **reconstructed**;

Reco-files are made available to the CYGNO users **few minutes after** the run was taken;

Ancillary systems: Gas and HV



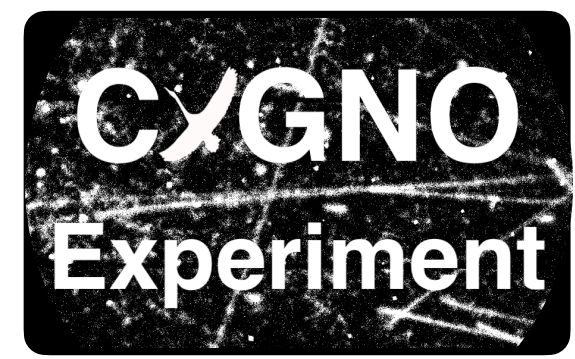
Less than one **spark**-like every 3 days was recorded in 2024 (12 times lesser than previous best limit ...)
Several **grafana** pages allows to monitor the **detector operation**, its main parameters and **data quality**;
Gas **humidity**, **oxygen** and other **contaminants** were **under control** for the whole **2024**



Detector calibration with ^{55}Fe is now a **completely automated** procedure and showed a **stability with 7% RMS** over 6 months

CYGNO04

TDR CYGNO financial plan and CSN2 proposal



The TDR submitted in **2022** included:

- infrastructure installation by Oct 2024;
- detector installation by Dec 2024;
- start of data taking Sept 2025;

Since the spring **2024** an important delay started to accumulate leading to 5-6 months delay

- infrastructure installation by the end of Feb 2025;
- detector installation by summer 2025;
- data taking to start at the beginning of 2026;

Financial profile for ERC and INFN from the **2022 TDR**

Year	INITIUM/ERC	CYGNO/INFN
2019	20	54
2020	201	44
2021	71	96
2022	40	96
2023	374	120
2024	302	125
2025	60	135
2026	0	95
2027	0	50
Tot 23-27	736	525
Tot	1068	815

La commissione apprezza il lavoro di analisi compiuto dalla collaborazione su LIME ma invita la collaborazione a finalizzare la **realizzazione di CYGNO-04 non oltre i primi mesi del 2026, recuperando, ove possibile i ritardi.** La CSN2 raccomanda alla collaborazione **un adeguato periodo di presa dati prima del decommissioning previsto nel 2027.** La commissione inoltre puntualizza che la maggior parte delle spese di costruzione costruzione e caratterizzazione di **CYGNO-04 devono essere sostenute su fondi ERC** e si impegna a dare supporto all'esperimento secondo il piano finanziario esposto nel CDR a fronte del raggiungimento delle **milestone di progetto, che saranno puntualmente verificate dai Referee.**

Between **December and January**, it became clear that the **delay** in producing the **executive drawings for CYGNO04** was unrecoverable by relying solely on the forces available to LNF's Design Service

We therefore asked the **LNGS Design Service** for support: merging and sharing their resources with LNF's one can be the **opportunity to finalise the work** done so far in a reasonable time interval;

In order to reach this aim, the proposal issued by the LNGS+LNF Design Services is the creation of a **task force** including:

- **a designer from both the Labs** (Alessandro Lalli - LNGS and Antonio Croce - LNF)
- with the **supervision of the two heads of the Services** (Donato Orlandi - LNGS and Cesidio Capoccia - LNF, head of the service)
- having as **direct link to CYGNO** Collaboration the **Technical Coordinator (Davide Fiorina** since November after Giovanni Mazzitelli resigned in April) supported by Davide Pinci as CYGNO spokesperson;

CYGNO current financial plan



Financial plan for INFN from the 2024 Progress Report

The new situation had very negligible effects on the financial plans:

- A total of **490 k€** are expected to being funded by INFN for the quinquennium 2023-2027 (**515 k€** foreseen in the **TDR**)
- In particular **305 k€** foreseen by INFN in 25-27;
- A total of **1099 k€** are expected to being funded by ERC for the quinquennium 2023-2027 (**1068 k€** foreseen in the **TDR**)

INFN - CSN2	2025	2026	2027
Gas Bottles	16	20	10
Gas Recovery	15	20	10
Consumables	25	10	20
Optics and DAQ	50	0	0
Tot w/o Travels (k€)	106	50	40
Travels - Shift	14	20	10
Travels - Installation	25	10	30
Tot Travels (k€)	39	30	40
Tot (k€)	145	80	80

Year	INITIUM/ERC	CYGNO/INFN
2019	20	54
2020	201	44
2021	71	96
2022	40	96
2023	164	93.5
2024	528	91.5
2025	75	145
2026	0	80
2027	0	80
Tot 23-27	767	490
Tot	1099	780

CYGNO04 Economic sustainability



Currently, from the ERC fundings, to cover the core costs of CYGNO04, there are a total of **304 k€** available at **INFN** and **GSSI**

Part of the equipment is being used and validated on LIME:

- high voltage system;
- gas system;
- DAQ and trigger;

According to the latest quotations received, the **other costs** expected for the **construction** of CYGNO04 is of **304 k€**, therefore an expense that can **be fully covered with ERC funds**.

Deadline for finalizing purchases: **August 2025**

The **total value** of **CYGNO04** is of about **920 k€**.

		Unit cost	Needed Quantity	Total Cost	We Have	To Buy	Still To Spend	Still To Spend
GAS VOLUME	Cone and GEM holders	5	1	5	0	1	5	35
	Vessel	30	1	30	0	1	30	
READOUT	4-GEM Set	10	2	20	2	0	0	14
	Orca Quest	26	6	156	6	0	0	
	Camera mechanics	1	6	6	0	6	6	
	Lenses	2	6	12	2	4	8	
	PMT+Socket	1	16	16	16	0	0	
CATHODE	Frame+foil+feedthrough	17	1	17	0	1	17	17
FIELD CAGE	Field Cage	13	1	13	0	1	13	13
CALIBRATION SYSTEM	Krypton-Rubidium	6	1	6	0	1	6	8
	55-Fe system	2	1	2	0	1	2	
SHIELDING	4 ton Copper Elsasser	156	1	156	1	0	28	193
	copper precision machining	10	1	10	0	1	10	
	Opera copper refurbishing	30	1	30	0	1	30	
	High Voltage feed through	10	1	10	0	1	10	
	Optical windows	10	1	10	0	1	10	
	water tanks	65	1	65	0	1	65	
	frame	20	1	20	0	1	20	
polietilene	20	1	20	0	1	20		
GAS SYSTEM	gas system	85	1	85	1	0	0	0
	filters	1	8	8	8	0	0	
HV SYSTEM	GEM-HV	5	1	5	1	0	0	0
	PMT-HV	3	2	6	2	0	0	
	Cathode HV	5	1	5	1	0	0	
DAQ & SLOW CONTROLS	modules	45.5	1	45.5	1	0	0	0
COMPRESSED AIR SYSTEM	filters and pipes	5	1	5	0	1	5	5
ELECTRIC SERVICES	power distribution system	10	1	10	1	0	0	0
	cables & connectors	5	1	5	1	0	0	
	UPS	5	1	5	1	0	0	
NETWORK DISTRIBUTION	network distribution system	13	1	13	1	0	0	0
COOLING and CONDITIONING	Conditioning	20	1	20	1	0	0	4
	Cameras chiller	5	1	5	1	0	0	
	PMTs flow system	2	1	2	0	1	2	
	Cables & connectors	2	1	2	0	1	2	
SAFETY (PRA-VIA)	pra+vinca	6	1	6	1	0	0	0
	fire detection	10	1	10	1	0	0	
	gas monitor	15	1	15	1	0	0	
	safety design	10	1	10	1	0	0	
CIVIL WORK	total	40	1	40	1	0	0	0
DESIGN AND DOCUMENTATION	software	5	1	5	0	1	5	15
	final audit	10	1	10	0	1	10	
				Gran Total 921.5			304	304

Component Validation - Copper Shield



Copper analysis with ICPMS

- They used the described procedure to clean it and measured the U and Th content

	Etching 2	Etching 3
	[pg * g ⁻¹]	[pg * g ⁻¹]
Th	9 ± 3	7 ± 2
U	5 ± 2	2 ± 1

- These were the Matthias results

radionuclide concentrations:

Th-232:		
Ra-228:	< 0.38 mBq/kg	<=> < 9.3 E-11 g/g
Th-228:	< 0.20 mBq/kg	<=> < 4.9 E-11 g/g
U-238:		
Ra-226:	< 0.44 mBq/kg	<=> < 3.5 E-11 g/g
Th-234:	< 17 mBq/kg	<=> < 9.3 E-10 g/g
Pa-234m:	< 11 mBq/kg	<=> < 6.5 E-10 g/g
U-235:	< 0.37 mBq/kg	<=> < 6.5 E-10 g/g
K-40:	< 3.2 mBq/kg	<=> < 1.0 E-7 g/g
Cs-137:	< 0.14 mBq/kg	
Co-60:	< 0.12 mBq/kg	
Co-58:	(0.8 +- 0.1) mBq/kg	
Mn-54:	(0.12 +- 0.05) mBq/kg	

⁵⁸Co has an half life of 70 days

⁵⁴Mn has an half life of 1 year

	October
Layer_0 (1-20 keV) [evts/yr]	14655 ± 347
Layer_1 (1-20 keV) [evts/yr]	4518 ± 176
Layer_2 (1-20 keV) [evts/yr]	10588 ± 2671
TOTAL	29761 ± 2700

Significant **improvement** thanks to ICP-MS measurement of Schrieber's copper

Huge statistics required to see first effects of Bi210 in OPERA's copper
→ **more storage space needed**

In light of the updated results, the most viable option is still **4 cm clean copper + 6 cm OPERA**

Upper limits on U and Th obtained with Germanium detector, 10 times larger than actual values

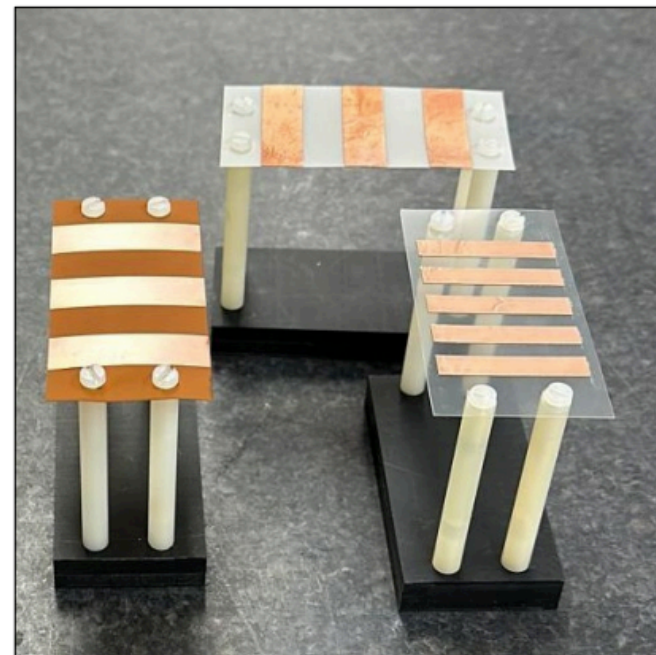
Field Cage and **GEM** foils are now on queue for the ICPMS measurements

Component Validation - Field Cage



Field Cage Foils

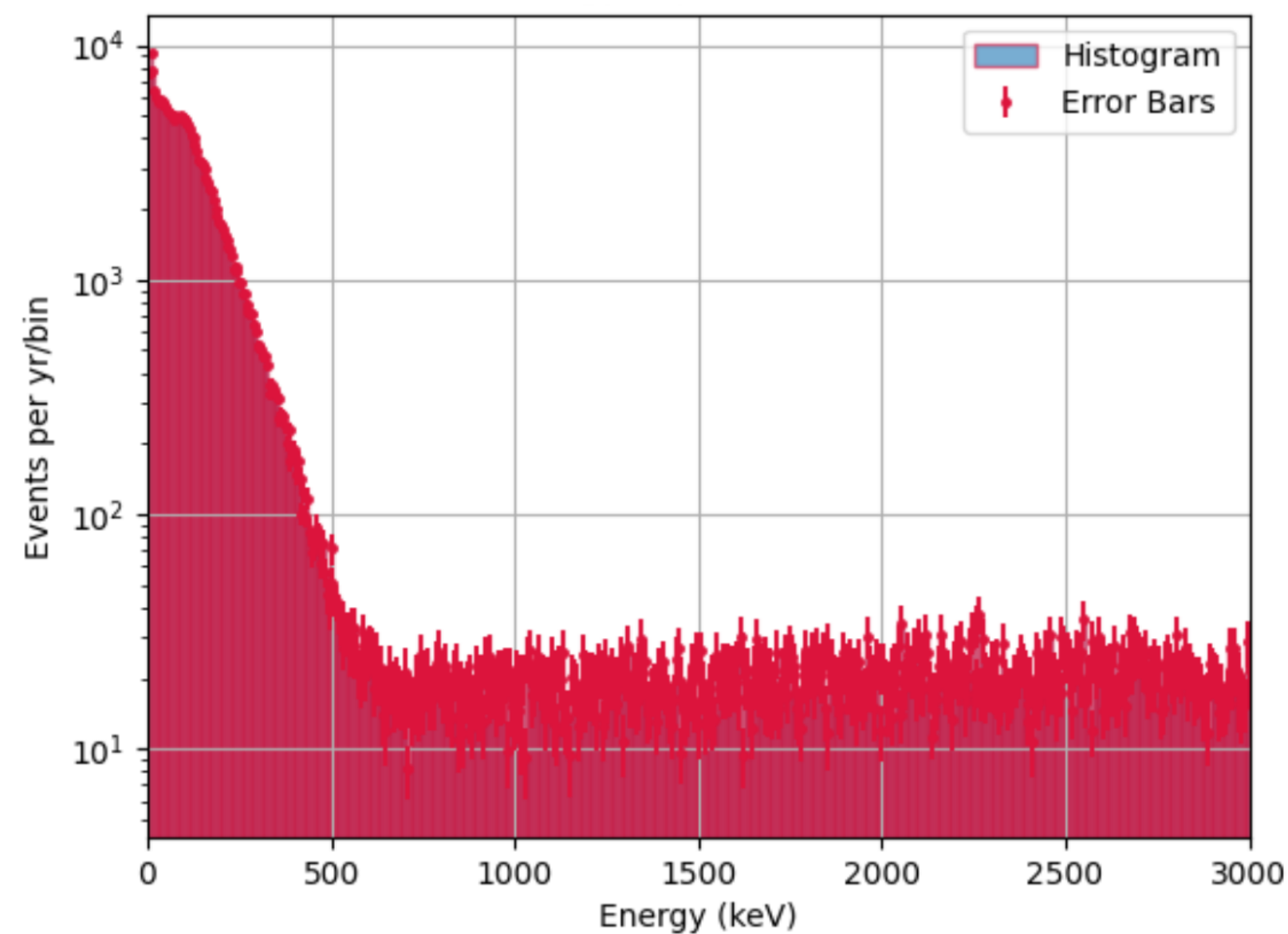
- Three different field cage foils:
 - **PET+Cu with glue**: a polyethylene foil with Cu strips glued by spreading the adhesive over the entire foil
 - **PET+Cu with no glue**: a polyethylene foil with Cu strips glued by applying the adhesive only between the strips and the foil
 - **Kapton+Cu**: a kapton foil with Cu strips
- The foils were measured at LNGS



Donatella Tozzi - Cygno CM 2024

2

Sample Name	in measure	measured	analysed
FC-kapton+Cu	no	yes	yes
FC-PET+Cu	no	yes	yes
FC-PET+Cu no glue	no	yes	yes
nylon6	no	yes	on-going
GEM foil	no	yes	on-going
nylon screws	no	yes	on-going
steel screws	no	yes	on-going
SMD resistors	no	yes	on-going
CMOS camera	no	yes	yes
PMMA	yes	no	no



In the low energy range (1-20 keV):

- $(1.56 \pm 0.01) 10^5$ evt/year;
- $(1.20 \pm 0.08) 10^3$ NR evt/year;

Papers in last year



1 Secondary scintillation yield from GEM electron avalanches in He-CF₄ and He-CF₄-CH₄ mixtures for CYGNO – Directional Dark Matter Search with an optical TPC

F.D. Amaro,^a E. Baracchini,^{d,e} S. Bianco,^d C. Capoccia,^d M. Caponero,^{d,e} D.S. Carrone,^d G. Cavoto,^{g,h} A. Cortese,^{b,c} L.A. Costa,^f E. Dané,^f G. Dho,^{b,c} F. Di Giambattista,^{d,e} E. Di Marco,^h G. D'Imperio,ⁱ P. Lima Júnior,^e G. Maccarrone,^d R.D.P. Mano,^d G. Mazzitelli,^d A.G. Mclennan,^{b,c} G. Nobrega,ⁱ I.F. Pains,ⁱ E. Paoletti,^g S. Piantoni,^{g,h} D. Piccolo,^d Pierluigi Ronga,^h R.J.C. Roque,^{a,2} F. Rosatelli,^d A. Russo,^d S. Scamporrino,^d N.J.C. Spooner,^k R. Tesaro,^d S. Tomassini,^d and S. Torelli,^{4,5}

^a LIBPhys, Department of Physics, University of Coimbra, 3004-516 Coimbra, Portugal
^b Gran Sasso Science Institute

He-CF₄-CH₄ ternary mixtures as target gas for the CYGNO directional dark matter experiment

R. J. da C. Roque, R. D. P. Mano, C. M. B. Monteiro (†) and CYGNO authorlist

Abstract

The CYGNO collaboration is developing a high-resolution optical Time Projection Chamber for directional dark matter neutrino spectroscopy, to be deployed at LNGS. He-40% CF₄ is being used as target gas, at atmospheric pressure, and a triple-GEM cascade is being used for the ionization signal, being the scintillation produced in the GEM readout by sCMOS cameras. High sensitivity to interactions in the few keV range for event tracking and particle identification are enabled with this solution. The addition of 3 to 10 percent of methane to He-40%CF₄ and demonstrate methane inclusion contributes to both the electrical stability of the TPC and to higher scintillation output from the GEM avalanches. In spite of the scintillation yield, the addition of methane increases the maximum voltage that can be applied before the onset of discharges, eventually resulting in higher scintillation yield. The visible component of the gas scintillation shows that isobutane photons emitted by He-CF₄. The maximum voltage that could be applied to be independent from the methane concentration in the scintillation gas. A hydrogen-based gas provides CYGNO with an even lighter target, reducing the dark matter detection threshold, in longer track lengths of light nuclear recoils and, thus, in a clearer direction discrimination.

Davide Pinci INFN - Roma

- 1 PREPARED FOR SUBMISSION TO JINST
- 2 Charge Amplification in Low Pressure CF₄/SF₆:He
- 3 Mixtures with a Multi-Mesh ThGEM for Directional Dark
- 4 Matter Searches



Baracchini,^{4,5} L. Benussi,⁶ ...
C. M. B. Monteiro,¹ R.A. Nobrega,ⁱ ...
F. Petrucci,^{2,3} S. Piantoni,^{g,h} ...
Renga,⁷ R. C. P. ...
Spooner.¹²

PREPARED FOR SUBMISSION TO JINST

Modeling the detector response of the CYGNO optical readout TPC

F.D. Amaro^a R. Antonietti^{b,c} E. Baracchini^{d,e} ...
C. Capoccia^f M. Caponero^{f,d} D.S. Carrone^d ...
F. Di Giambattista^{d,e} E. Di Marco^h ...
G.S.P. Lopes^m G. Maccarrone^d ...
G. Mazzitelli^d A.G. Mclennan^{b,c} ...
I.F. Painsⁱ E. Paoletti^g ...
D. Pierluigi^g ...
J.M.F. ...
N.J.C. Spooner^k R. Tesaro^d

^a Department of Physics, University of Coimbra, 3004-516 Coimbra, Portugal
^b Dipartimento di Matematica e Fisica, Università Roma TRE, 00146, Roma, Italy
^c Istituto Nazionale di Fisica Nucleare, Sezione di Roma Tre, 00146, Roma, Italy
^d Gran Sasso Science Institute, 67100, L'Aquila, Italy
^e Istituto Nazionale di Fisica Nucleare, Laboratori Nazionali del Gran Sasso, 67100, Assergi, Italy

Under internal review

Enhancing the light yield of He:CF₄ based gaseous detector

Fernando Domingues Amaro¹, Rita Antonietti^{2,3}, Elisabetta Baracchini^{4,5}, Luigi Benussi⁶, Stefano Bianco⁶, Roberto Campagnola⁶, Cesidio Capoccia⁶, Michele Caponero⁶, Danilo Santos Cardoso⁸, Luan Gomes Mattosinho⁸, Gianluca Cavoto^{10,11}, Igor Abritta Costa¹⁰, Emanuele Dané⁶, Giorgio Dho^{4,6,a}, ...
Rita Antonietti^{2,3}, Elisabetta Baracchini^{4,5}, Luigi Benussi⁶, Stefano Bianco⁶, Roberto Campagnola⁶, Cesidio Capoccia⁶, Michele Caponero⁶, Danilo Santos Cardoso⁸, Luan Gomes Mattosinho⁸, Gianluca Cavoto^{10,11}, Igor Abritta Costa¹⁰, Emanuele Dané⁶, Giorgio Dho^{4,6,a}, ...

An analytical model of the response of the optically readout GEM based TPC for the CYGNO experiment

Rita Antonietti^{2,3}, Elisabetta Baracchini^{4,5}, Luigi Benussi⁶, Stefano Bianco⁶, Francesco Borra⁶, Cesidio Capoccia⁶, Michele Caponero⁶, Danilo Santos Cardoso⁸, Gianluca Cavoto^{10,11}, Igor Abritta Costa¹⁰, Emanuele Dané⁶, Giorgio Dho^{4,6,a}, ...
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Under internal review

