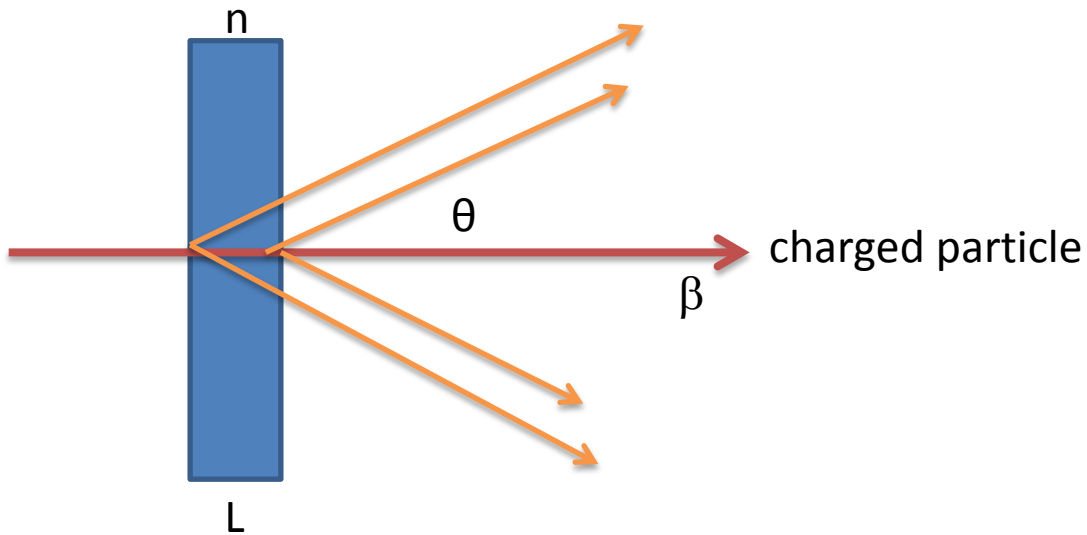


A RICH-like cosmic rays setup

F.Bucci, A.Cassese, R.Ciaranfi,
M.Lenti, S.Lami, F.Maletta, K.Xhani

Cherenkov effect in a nutshell



n : index of refraction
 θ : Cherenkov angle
 β : particle velocity ($c=1$)
 L : radiator length

$$\cos \theta = 1/n\beta$$
$$\cos \theta(\max) = 1/n$$
$$\beta(\text{threshold}) = 1/n$$

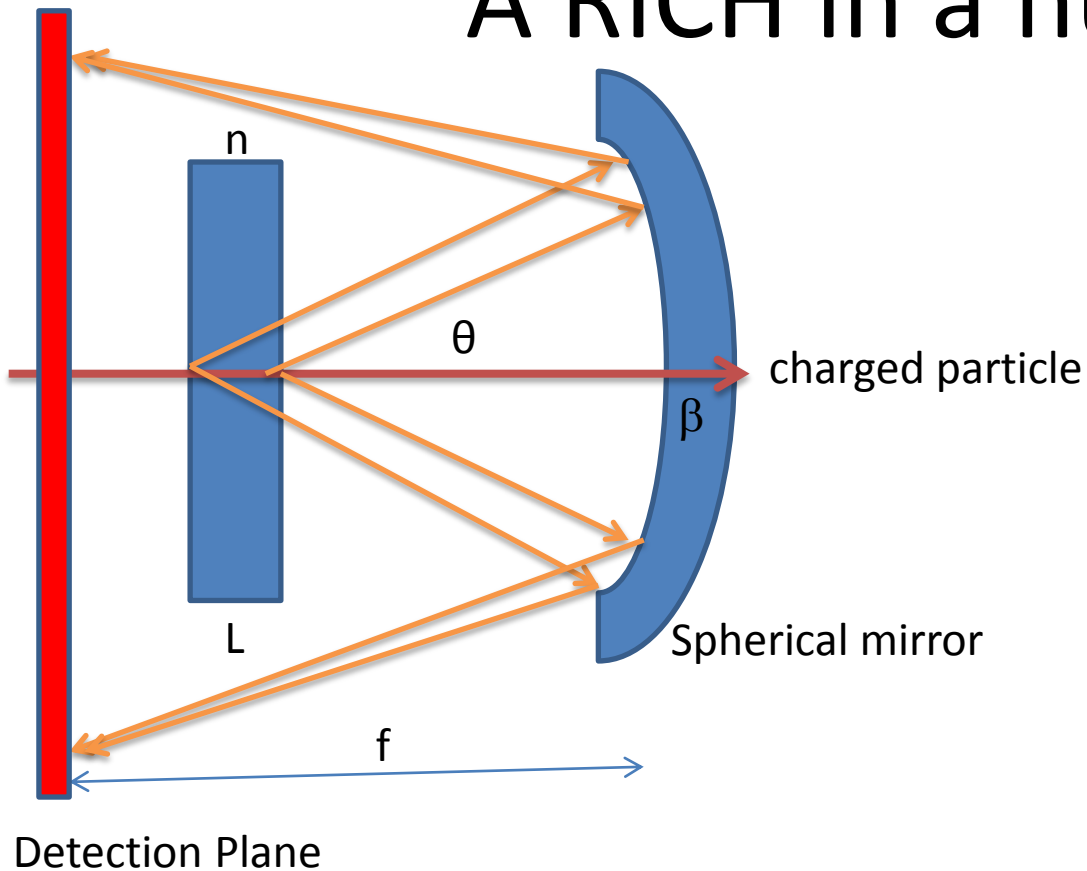
Frank-Tamm eq.:

$$\frac{d^2 N_\gamma}{dE dx} = 370 \sin^2 \theta \text{ eV}^{-1} \text{ cm}^{-1};$$

$$N_{p.e.} = N_0 L \sin^2 \theta;$$

$$N_0 \approx 100 \text{ cm}^{-1}$$

A RICH in a nutshell



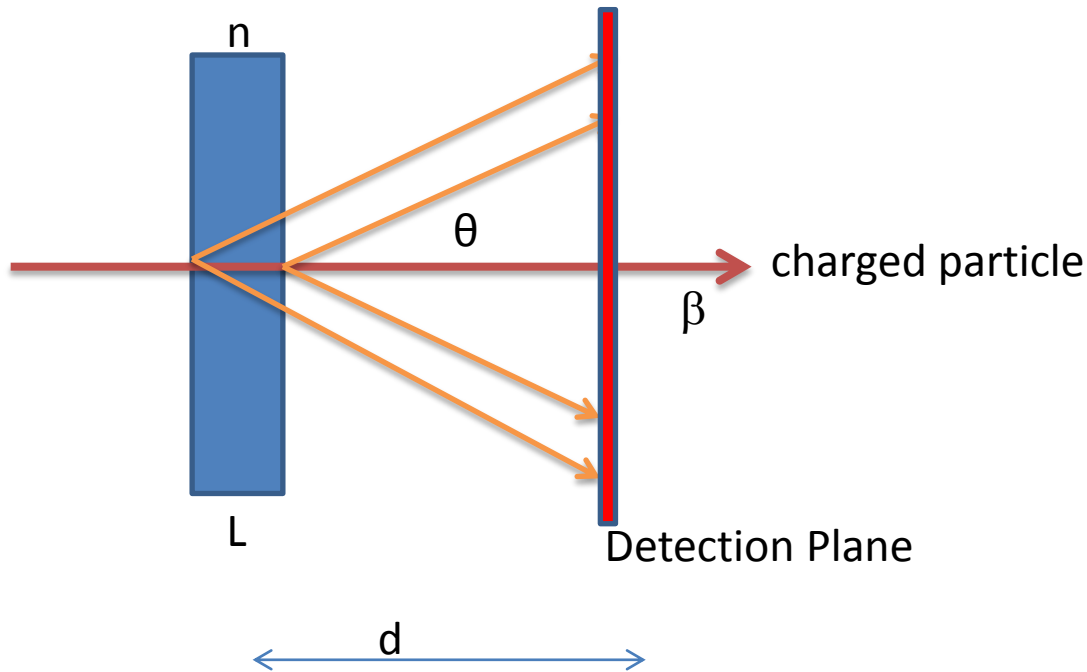
The Cherenkov cone is mapped (pure optical geometry) into a ring in the mirror focal plane

The sensors (PMT, etc) are placed in the mirror focal Plane

$$r_{ring} = f \tan \theta$$

Usually $L \approx f$

A proximity focusing cherenkov in a nutshell



If $L \ll d$ an approximate Ring is formed on the Detection plane

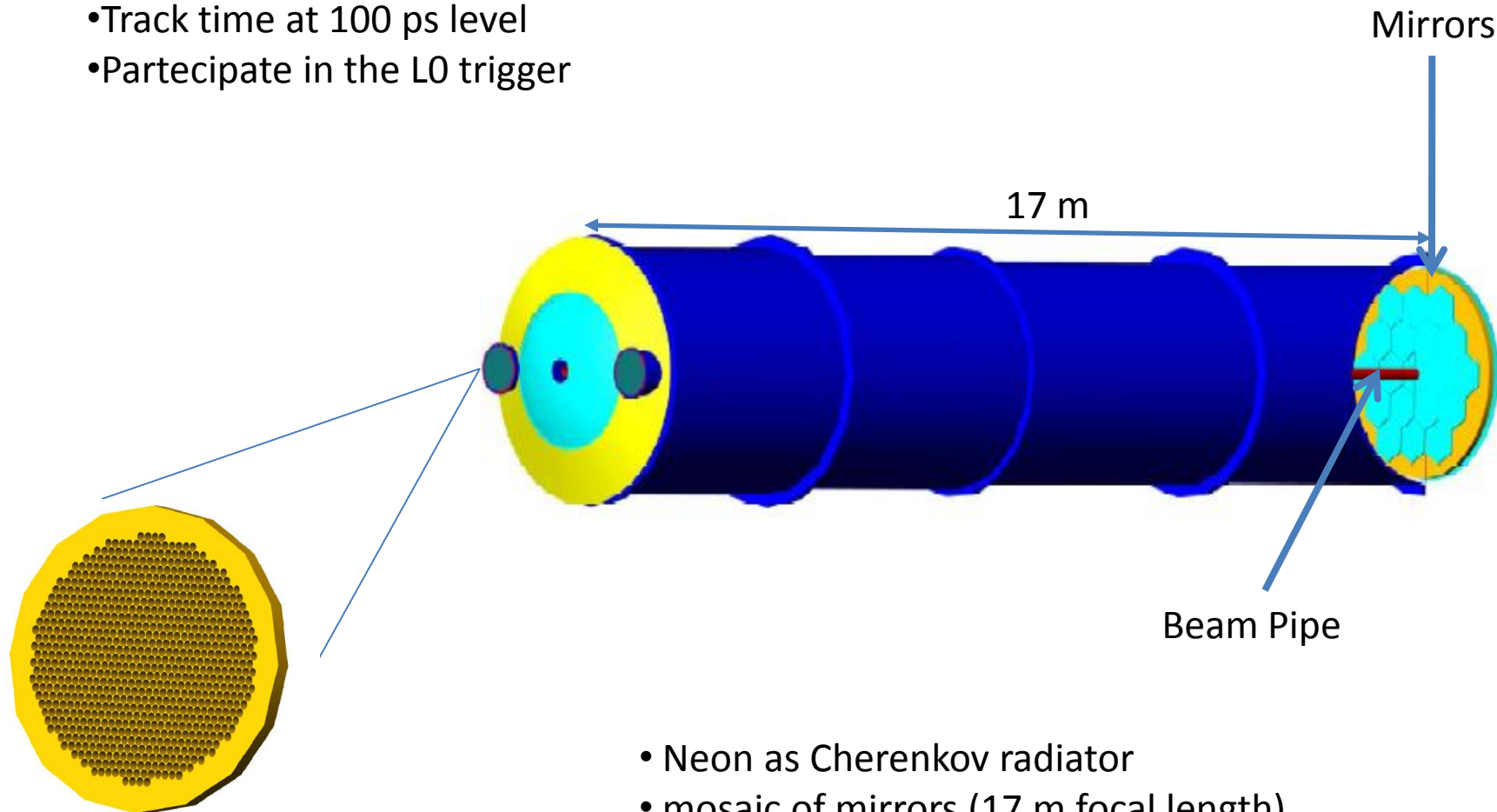
$$r_{ring} \approx d \tan \theta$$

The thickness of the ring depends on L

THE NA62 RICH

INFN Perugia
INFN Firenze
CERN

- Separate pions from muons <1% level between 15 and 35 GeV/c
- Track time at 100 ps level
- Participate in the L0 trigger



PM lodging disk

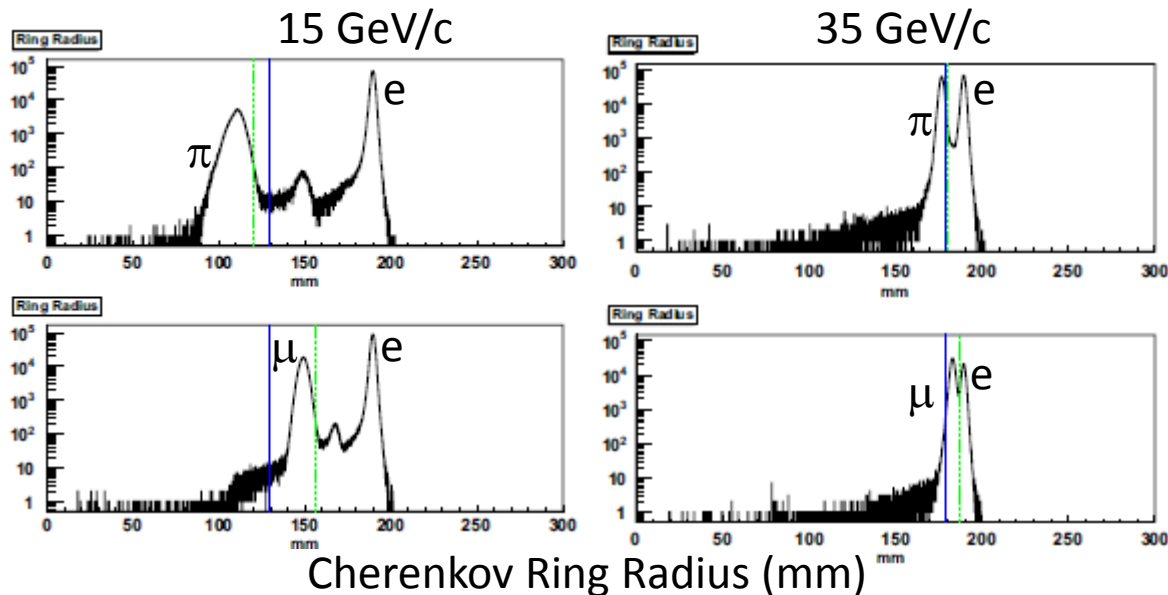
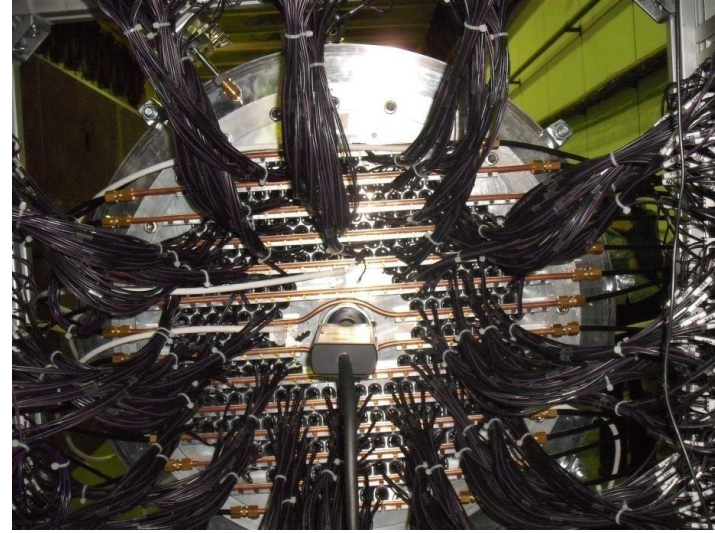
- Neon as Cherenkov radiator
- mosaic of mirrors (17 m focal length)
- two spots with PM (1000 PM per spot)
- beam pipe passing through

The NA62 RICH prototype

- A RICH prototype was built in 2007,
- Vessel: 17 m long, 0.6 m wide, filled with Neon
- Equipped with one mirror
- In a test beam in 2007 at CERN (TCC8) it was equipped with 96 PMTs, demonstrating the required time resolution
- In a test beam in 2009 at CERN (TCC8) it was equipped with 414 PMTs, demonstrating the required $\pi-\mu$ separation between 15 and 35 GeV/c
- 2 NIM papers published

NA62 RICH prototype publications

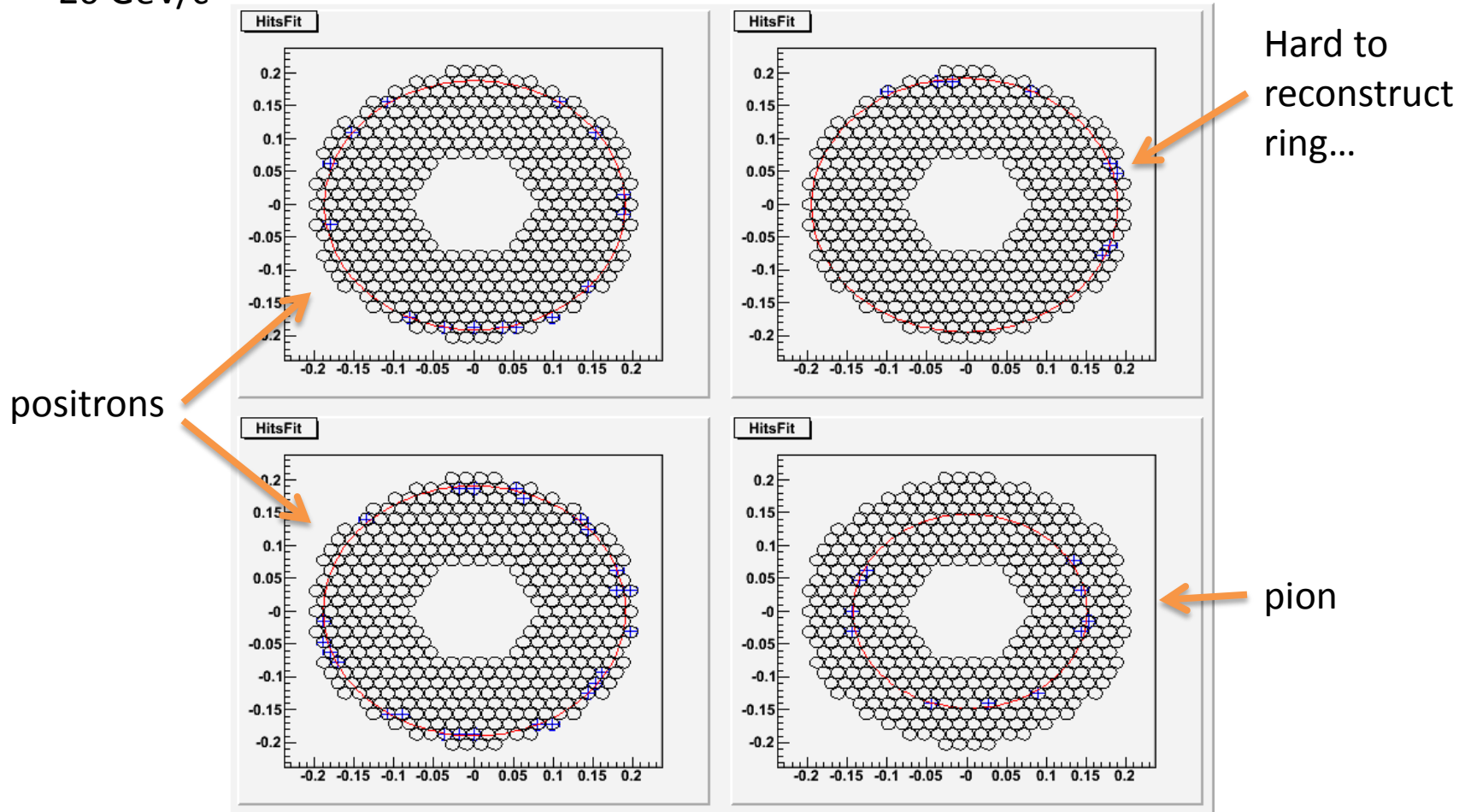
- “Construction and test of a RICH prototype”, G.Anzivino et al, NIMA, v.538, p.314-318, 2008.
- “Pion-Muon separation with a RICH prototype for the NA62 experiment”, B. Angelucci et al, NIMA, v.621, p-205-211, 2010



Cherenkov Ring Radius (mm)

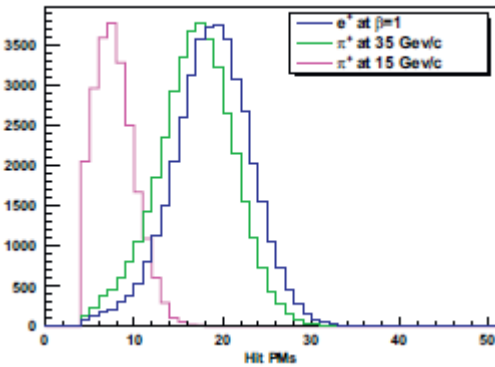
NA62 RICH prototype: rings

20 GeV/c

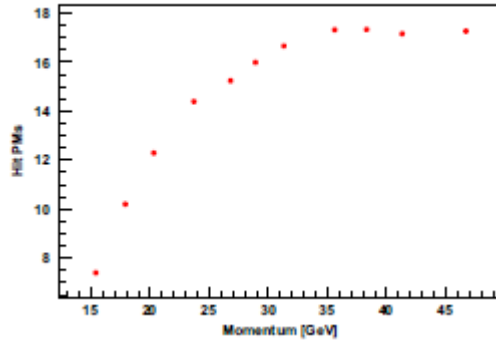


NA62 RICH prototype results

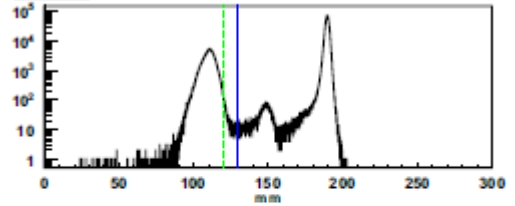
NHits Vs Ring Radius



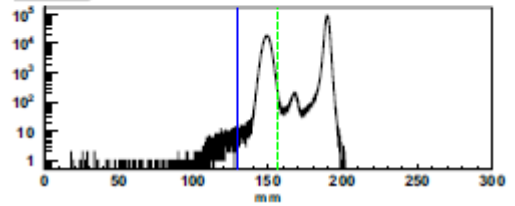
Hit PMs Vs Momentum



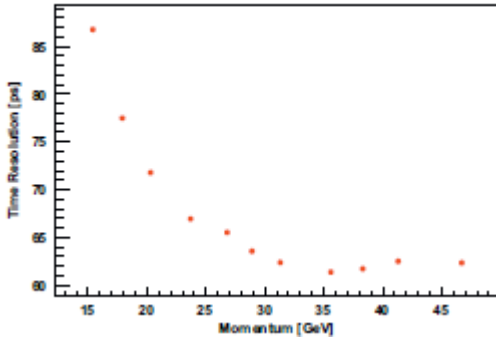
Ring Radius



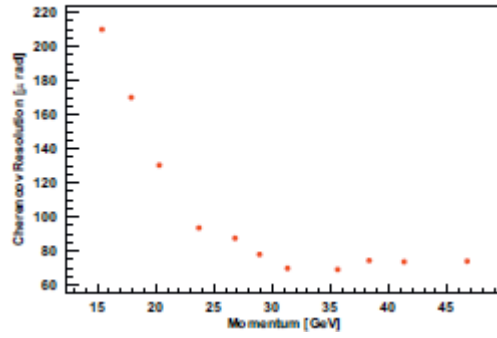
Ring Radius



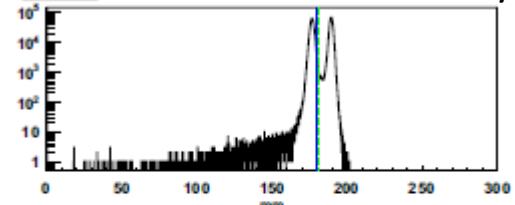
Time Resolution Vs Momentum



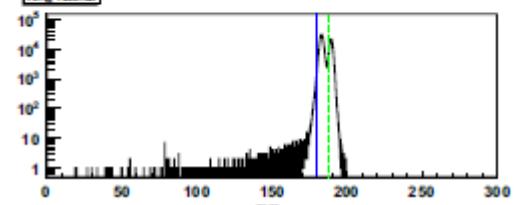
Cherenkov Resolution Vs Momentum



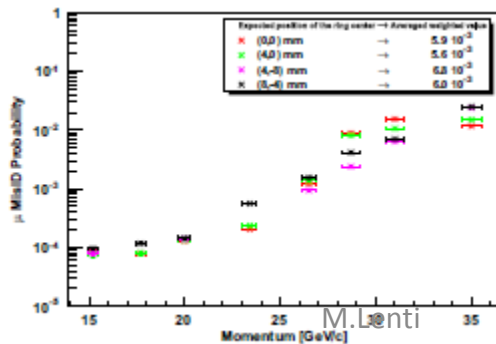
Ring Radius



Ring Radius



μ suppression: 0.7%
 ($15 < p < 35$ GeV/c)

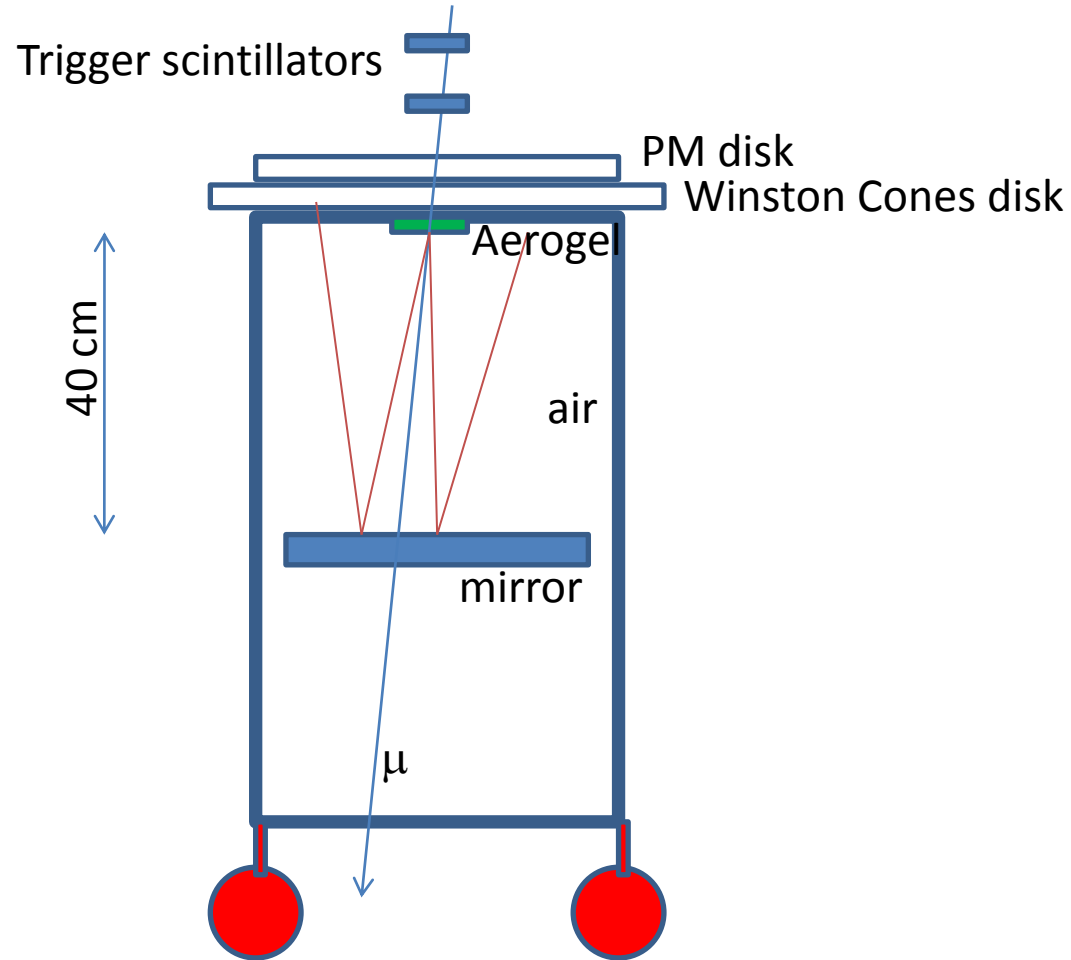


Cosmic rays setup

- Try to recycle part of the RICH prototype for a didactic cosmic rays setup
- Neon as radiator not practical for cosmics
- Use aerogel bricks as radiator
- Not a RICH but a proximity focusing Cherenkov detector
- Re-use the spherical mirror in “flat mode”

Cosmics rays setup: “COSMORICH”

“vessel” light tight,
not gas tight



PM disk

Electronics

Preparation



Winston Cones

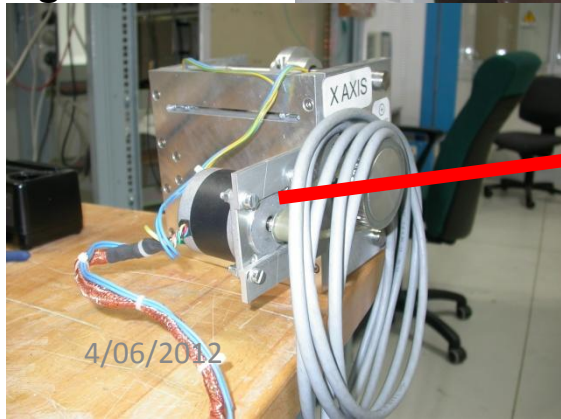


Cosmic chariot



Mirror cradle

Align motor



4/06/2012



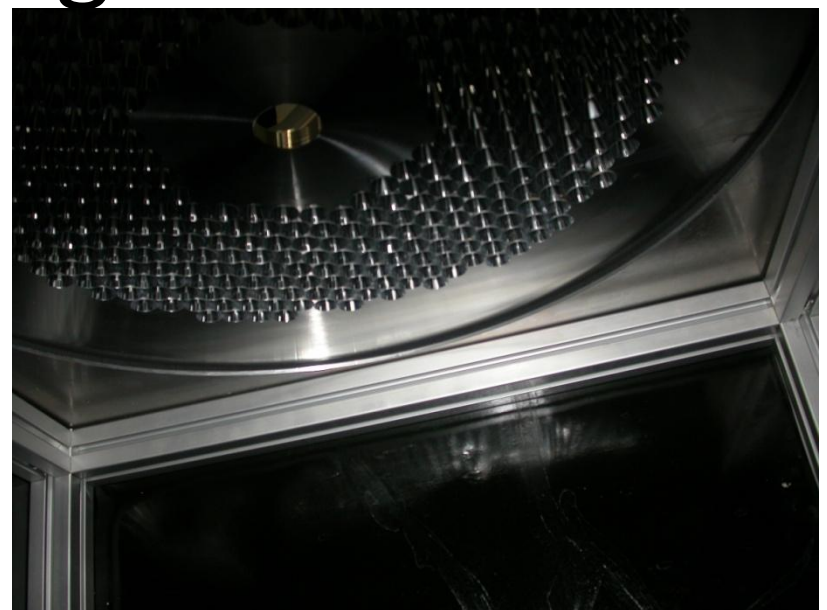
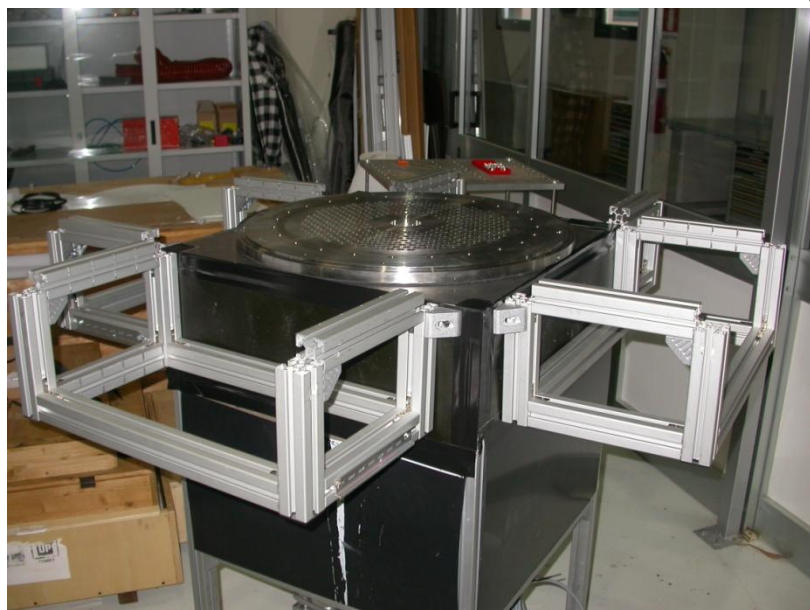
Closing flange

M.Lenti

12



Progressing...

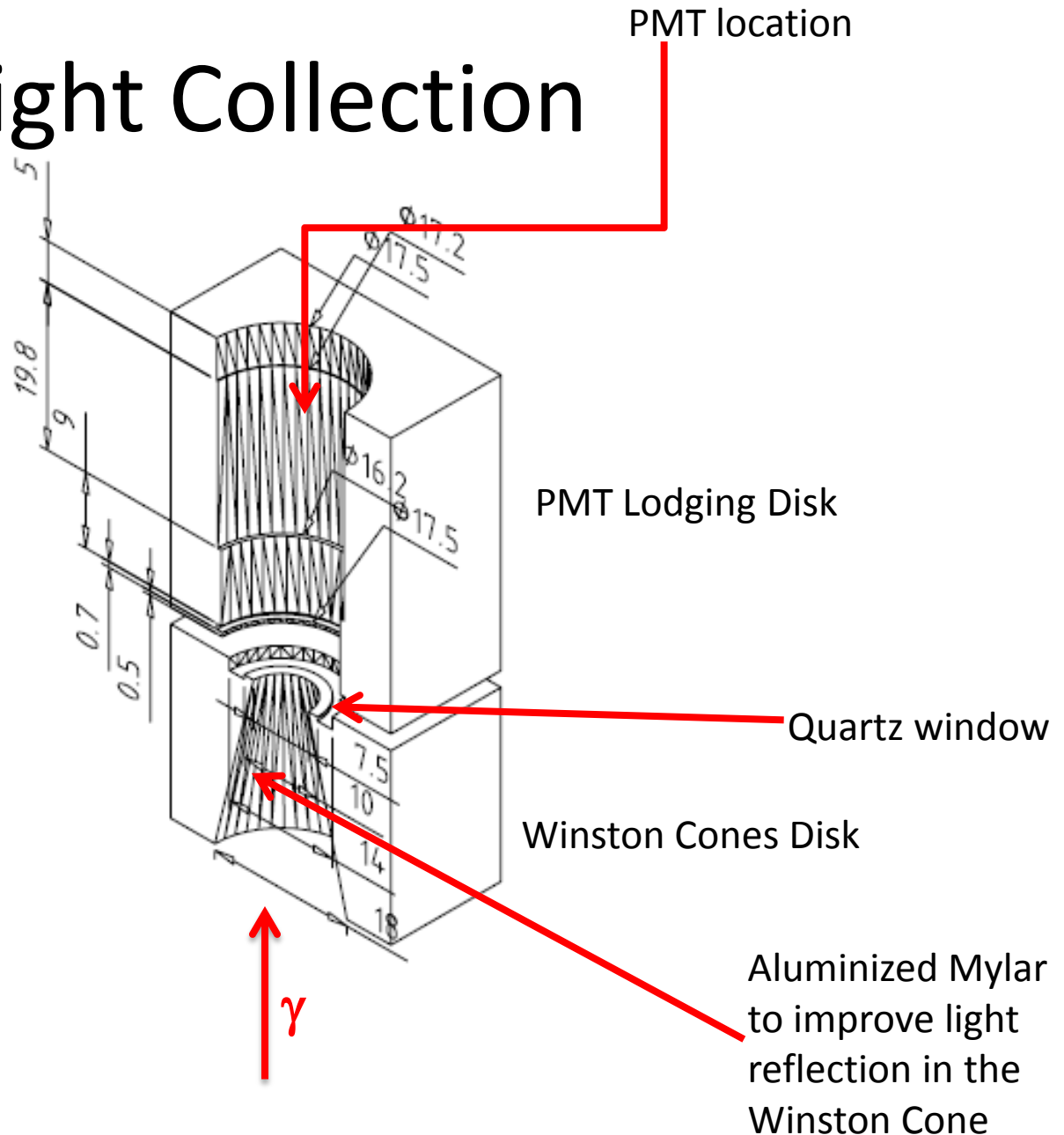


View from inside

Light Collection

PMT: 16 mm \varnothing
Active region: 8 mm \varnothing
Packing: 18 mm step

Need a light collection system from 18 mm \varnothing to 8 mm \varnothing :
Winston Cone



Quartz windows



- Praezisions Glas & Optik GmbH
- Windows made of UV-grade Quartz Glass:
 - double side polished
 - diameter. 12.7mm (+/-0,1)
 - thickness 1mm (+/-0,1)
 - ground edge, unbeveled
- Necessary to separate air from Neon with good UV transparency
- Not necessary for the cosmic ray setup.....

NA62 Mylar reflective coating with Al/MgF2



Fixation in the evaporator:

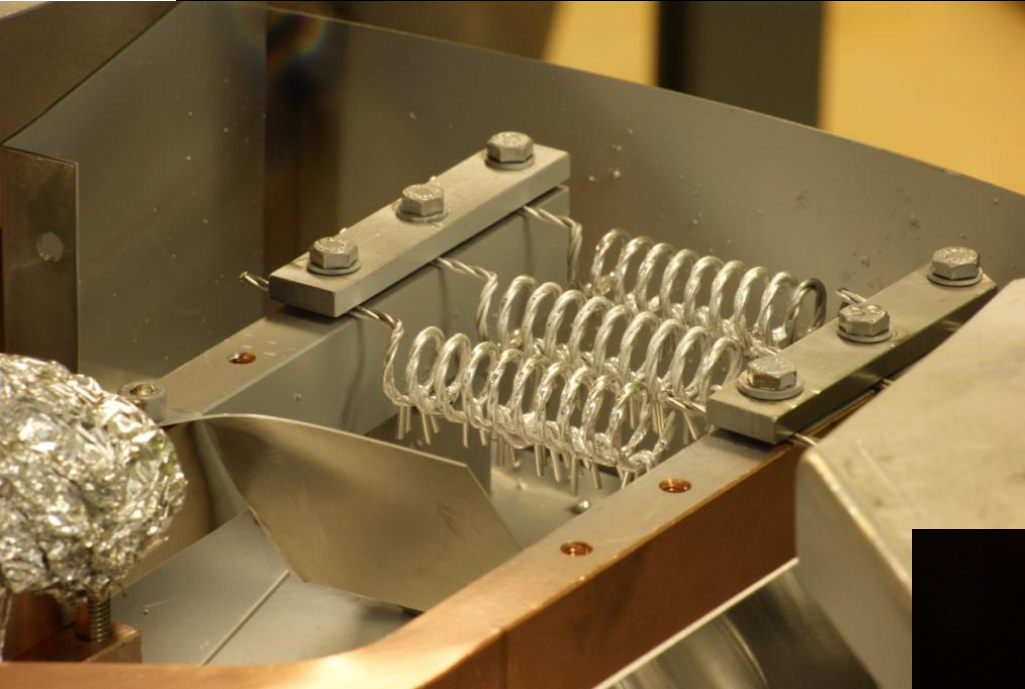
The clamped Mylar foil is fixed to the rotating plate in the upper part of the evaporation plant

Quality Control samples:

For each process (foil) we add a glass sample to assure the quality control of the process. An extra sample is cut out of the Mylar foil to document the result on the foil.



NA62 Mylar reflective coating with Al/MgF2



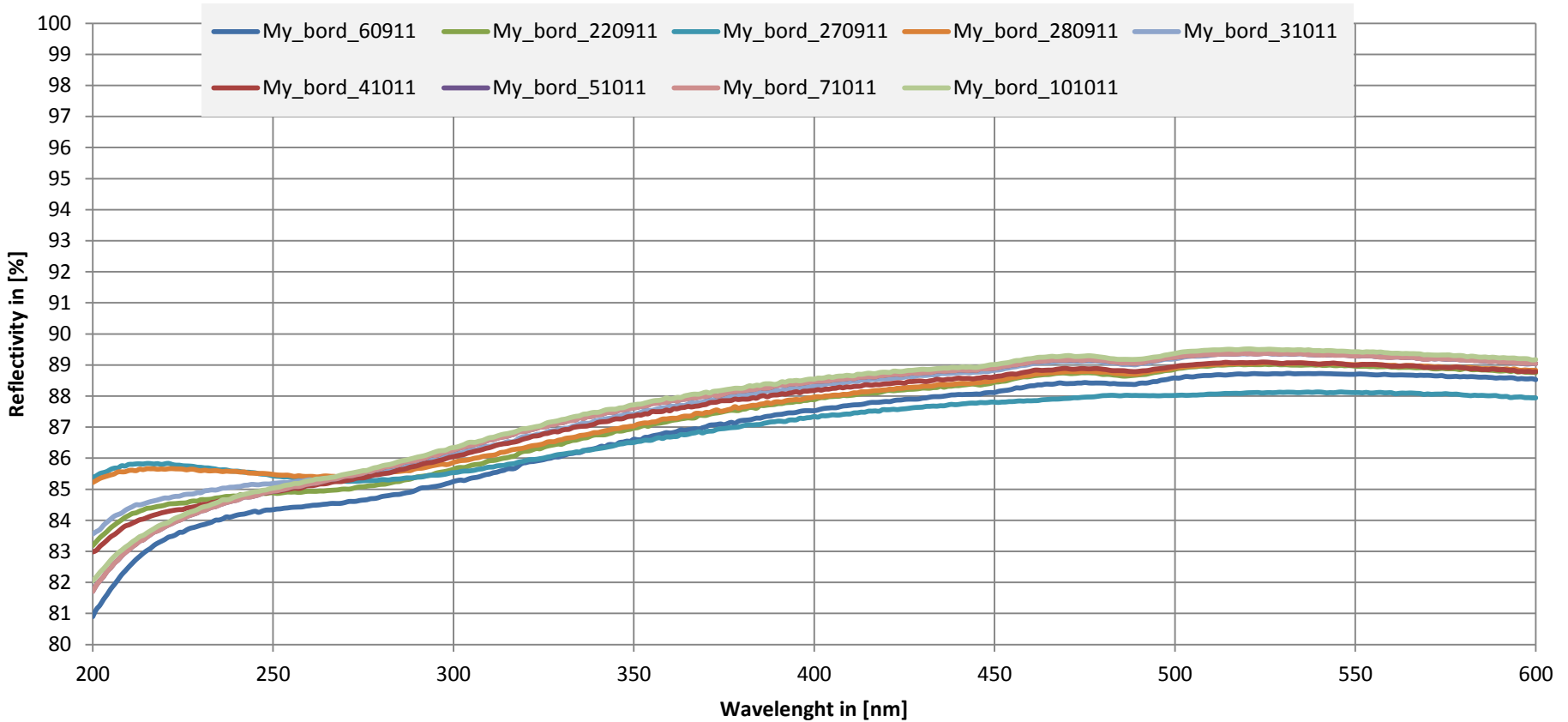
Aluminium evaporation:
Tungsten filaments are charged and pre-melted with pure Aluminium.

Via resistive heating (joule effect) the Aluminium is evaporated to the rotating substrate (Mylar foil). The vacuum during the process is in the order of 10^{-6} mbar.

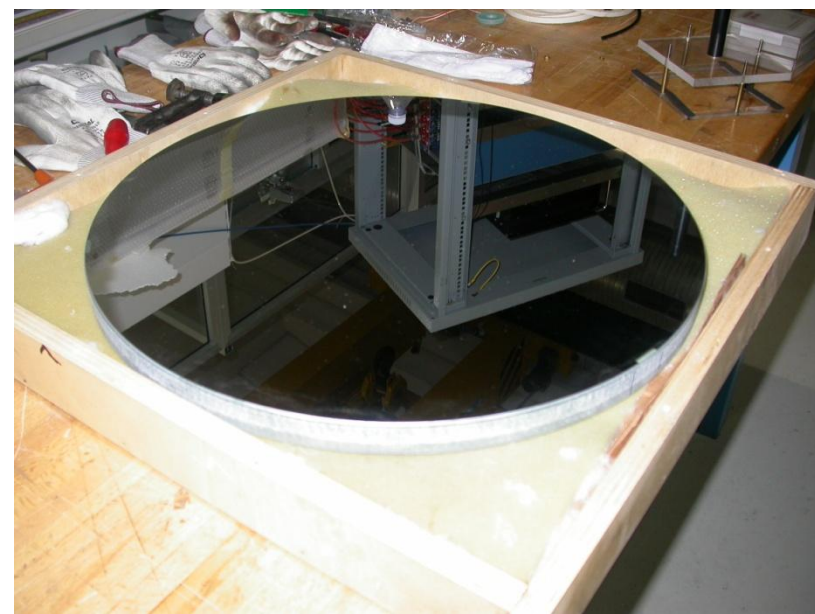


NA62 Mylar reflective coating with Al/MgF2

NA_62: Comparaison Mylar reference preproduction

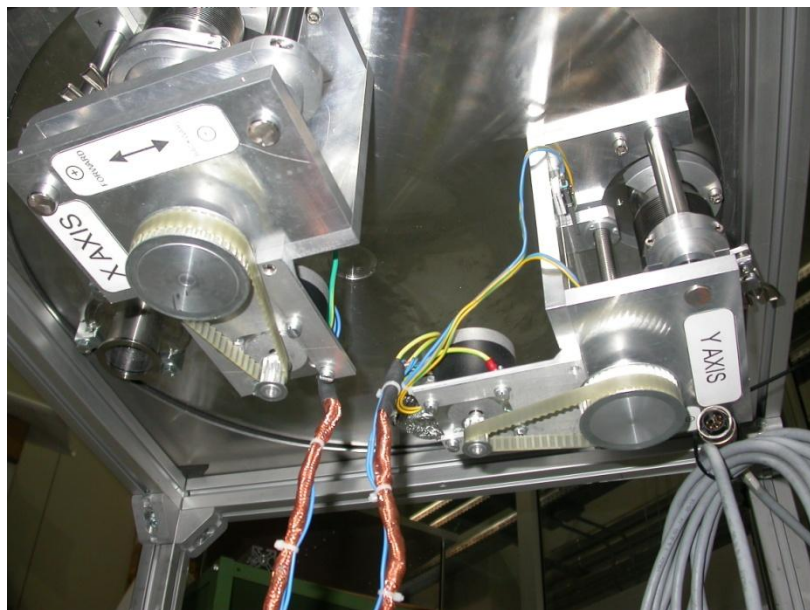


Reflectivity scan of some coated Mylar foils (scan from 200nm-600nm)



Back to preparation....

The mirror



4/06/2012

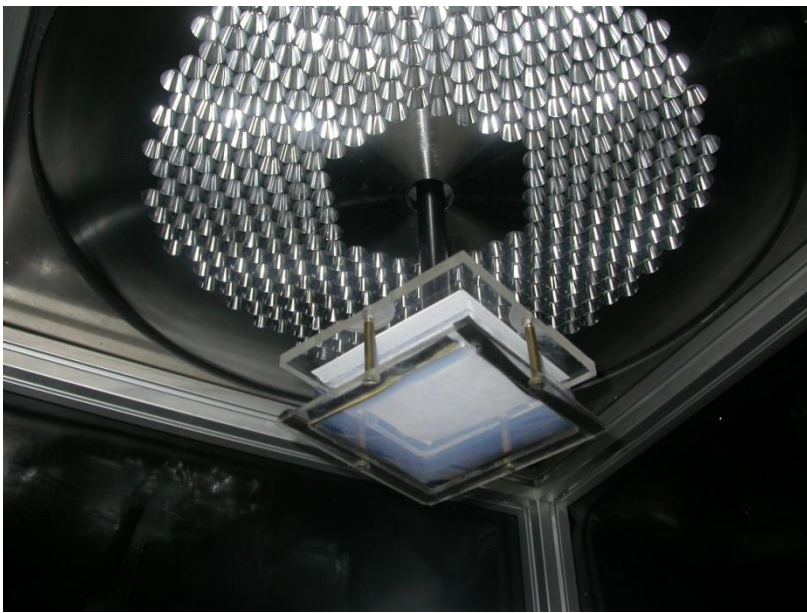
M.Lenti

Aerogel bricks

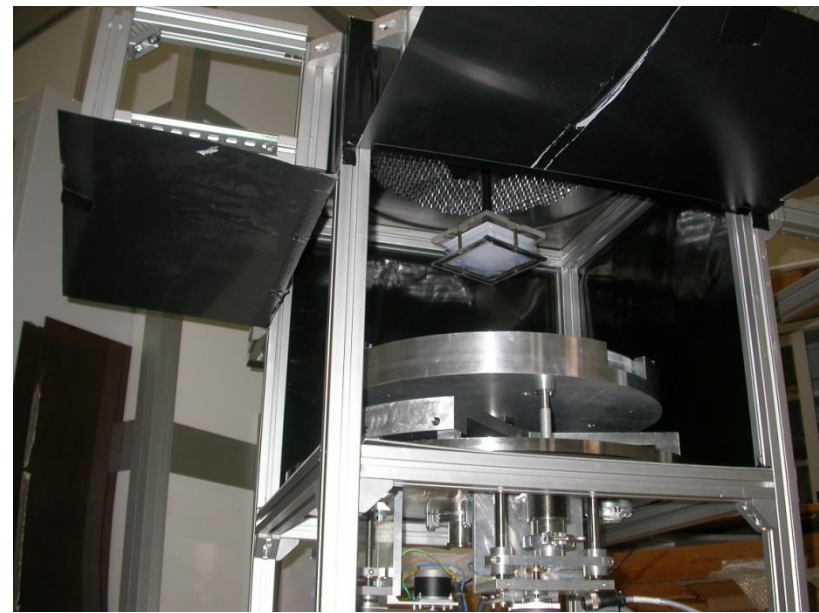
19

Aerogel

- Matsushita SP-30 (courtesy of **Clara Matteuzzi** from LHCb), hydrophobic
- 3 bricks of $12 \times 12 \times 1 \text{ cm}^3$ each
- $n=1.03$ ($\theta_{\text{max}} = 250 \text{ mrad}$, $p_{\mu}(\text{thresh})=440 \text{ MeV}/c$)
- Dist aerogel-mirror $\approx 40 \text{ cm}$
- $R_{\text{ring}} (\beta=1) \approx 2 \times 40 \times 0.250 \approx 20 \text{ cm}$ (as in test beam)



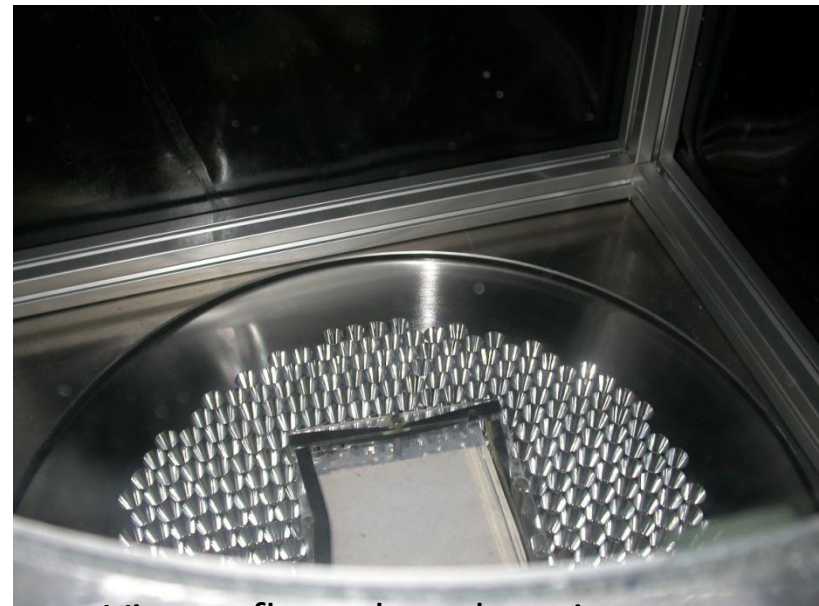
Inserting aerogel...



Again...



Final position



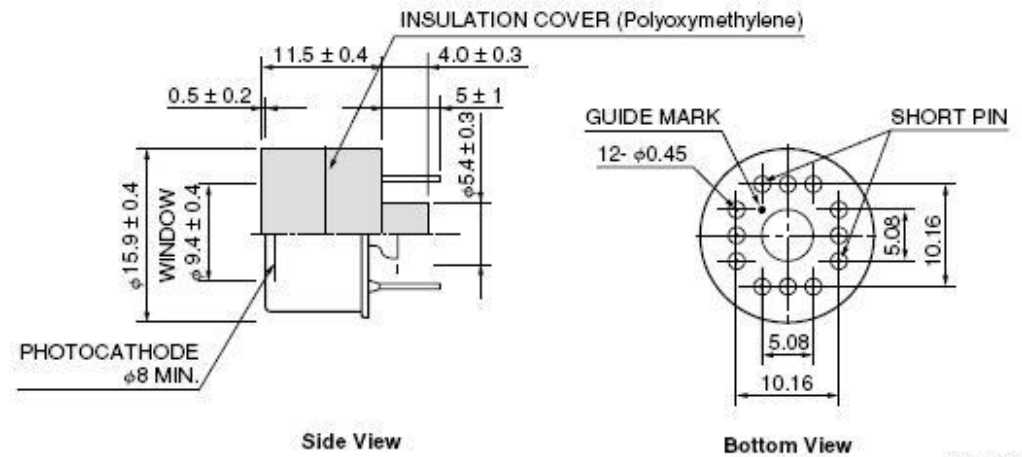
View reflected on the mirror

RICH PMT: Hamamatsu R7400U-3



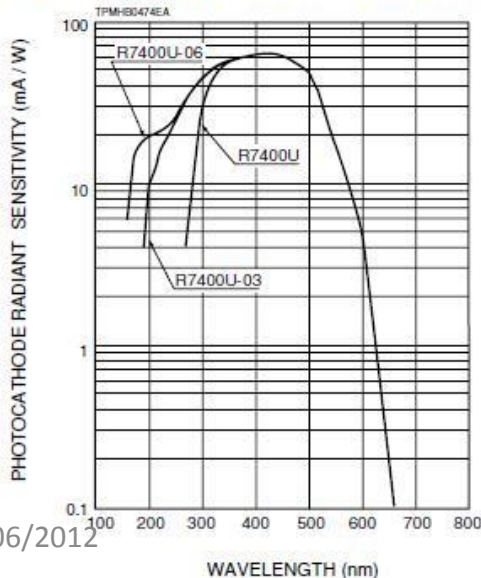
Figure 11: Dimensional Outline and Basing Diagram (Unit: mm)

①R7400U, -01, -02, -03, -04, -20, R7400P



TPMHA0411EC

Figure 2: Typical Spectral Response (Bialkali)

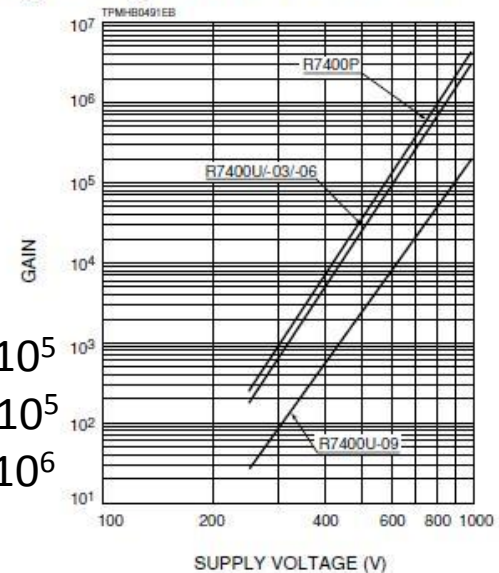


*Fast, low noise, single anode
PMT, 8 dynodes,
 $G \sim 1.5 \cdot 10^6$ @ -900 V*

**~2000 PMTs tested in
Perugia/Firenze**

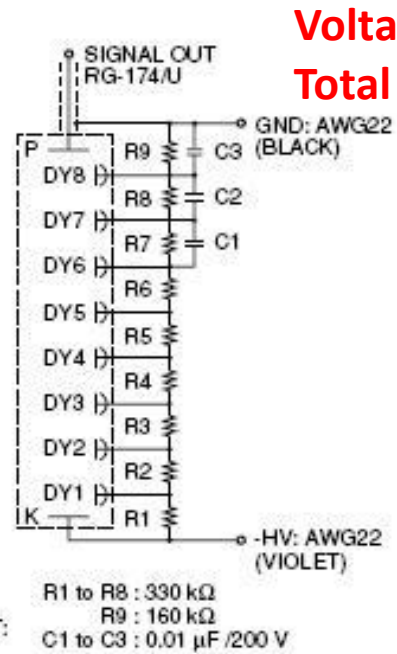
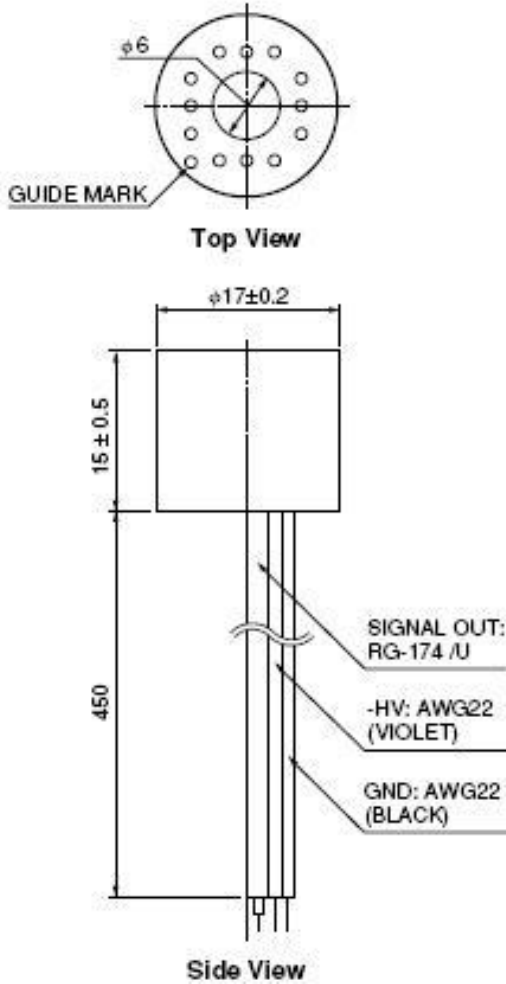
normal: gain @ 800 V > 1×10^5
good: gain @ 800 V > 5×10^5
best: gain @ 800 V > 1×10^6

Figure 4: Typical Gain Characteristics

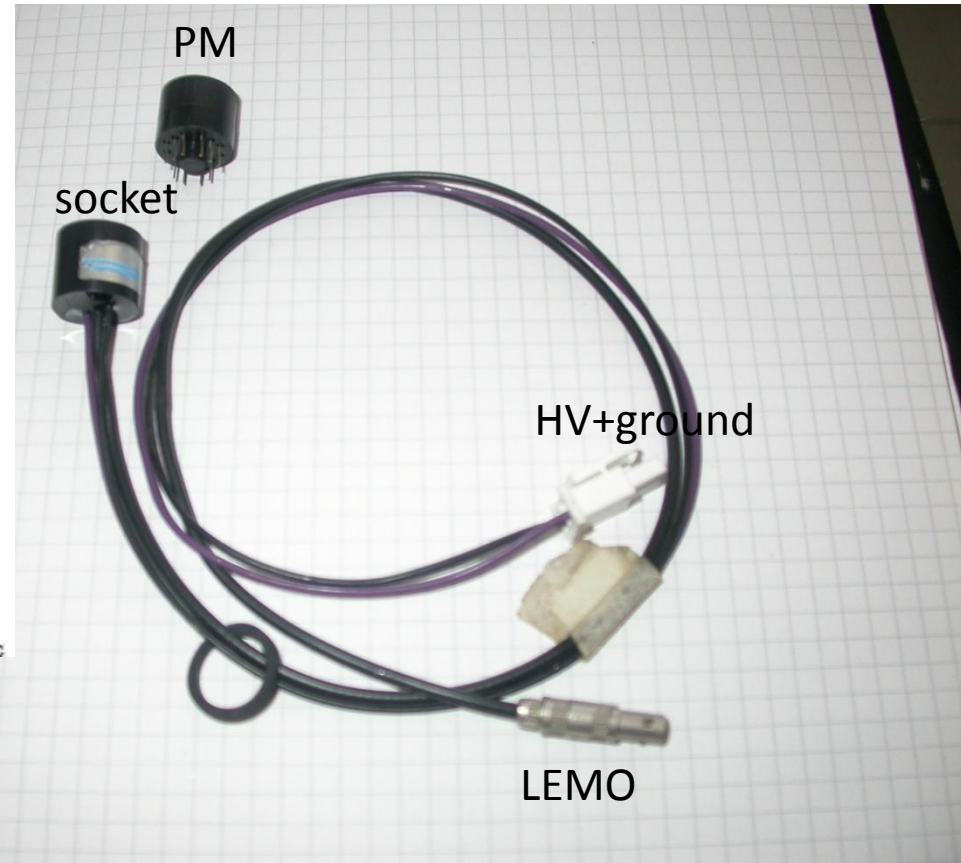


Hamamatsu E5780 Socket Assembly

Cable Output Type E5780



Voltage Divider Circuit
Total Resistance 2.8 MΩ



Systematic tests w. different resistors



2m long cables.
Molex connectors.

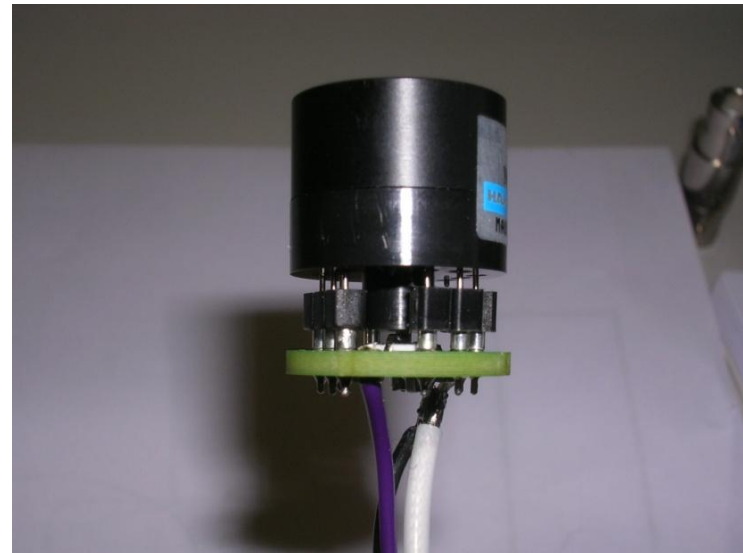
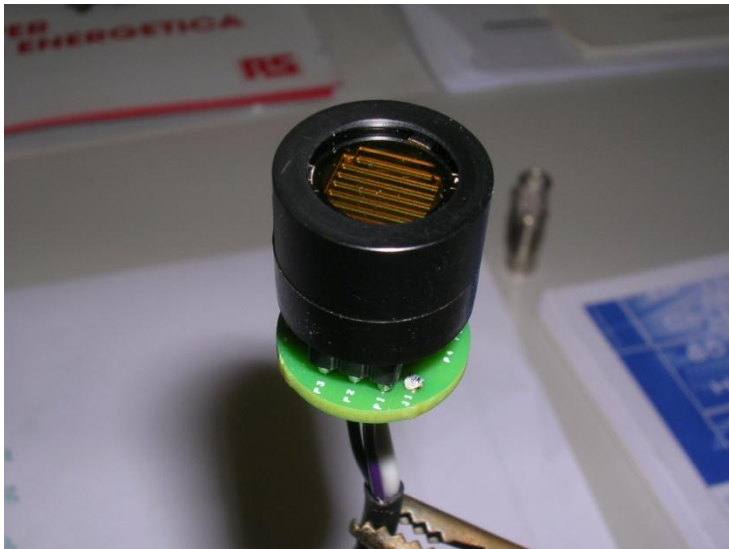
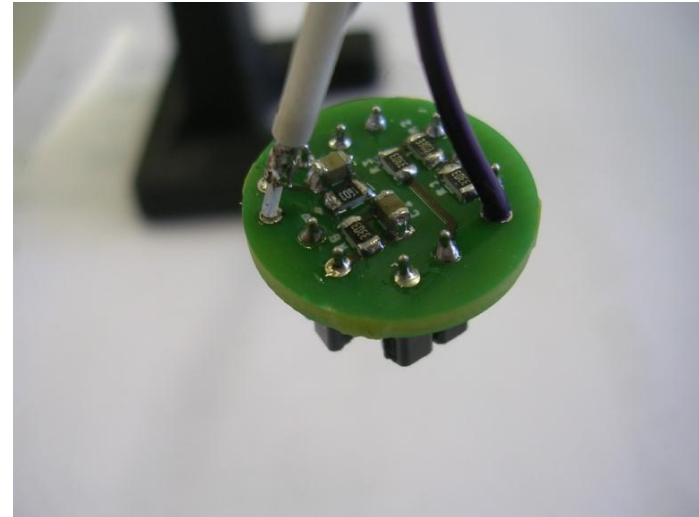
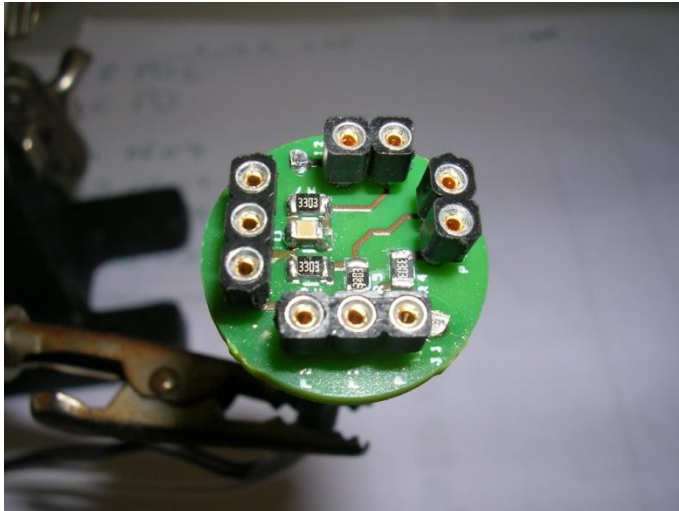
Comparison Hamamatsu - INFN
dividers with 3x, 5x, 6x, 8x and
10xR.

Tests with laser illuminating 2
PMTs at the same time, one
kept fixed at 900V as a
reference, to monitor laser
fluctuations.



4/06/2012

INFN – Firenze Socket Prototype



HV power supply

- mainframes CAEN SY1527 (16 slots each, 8U high)
- boards CAEN A1733N (12 ch, 1slot, 6U high)
- boards CAEN A1535SN (24 ch, 2slots, 6U high)

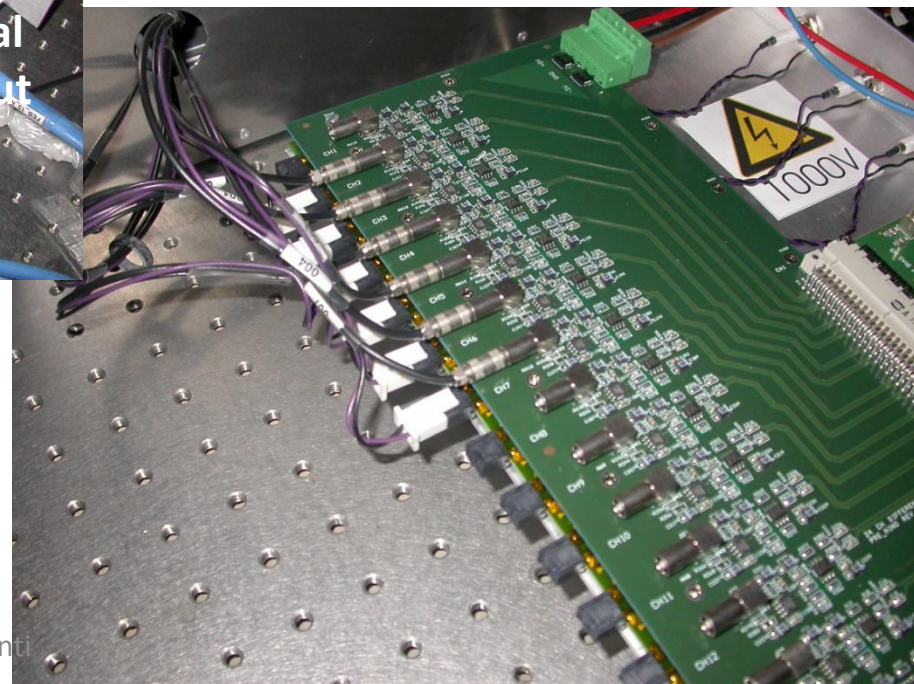
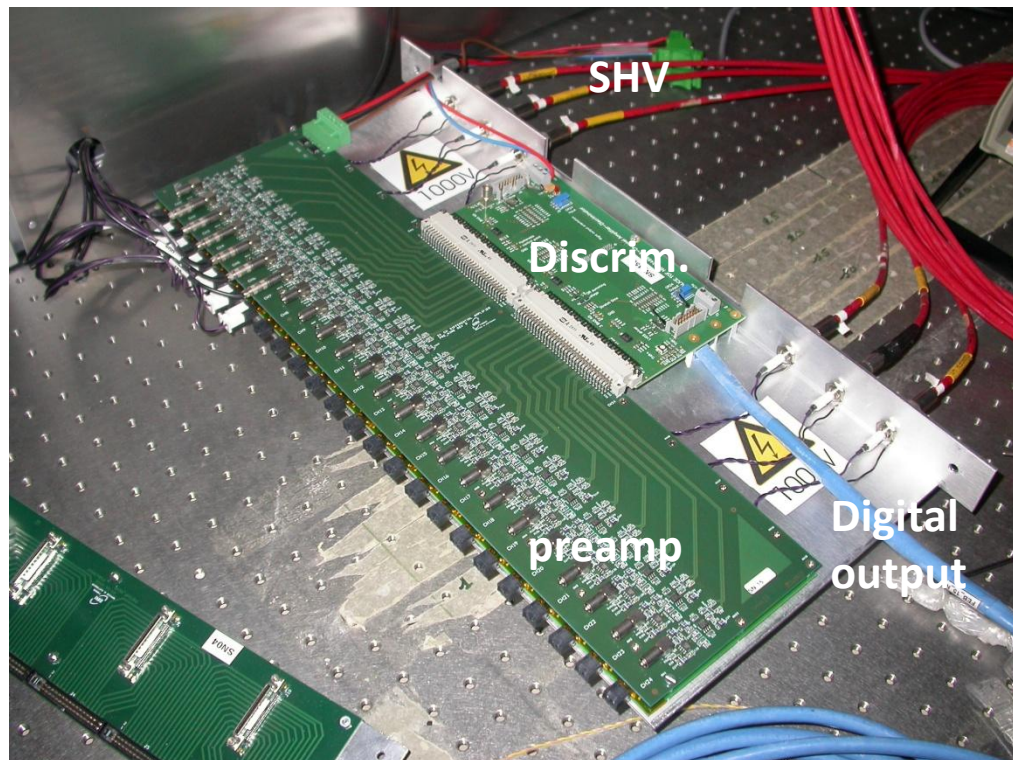
In the lab:

1 SY1527

14 A1733N (9 used)



RICH prototype FE board

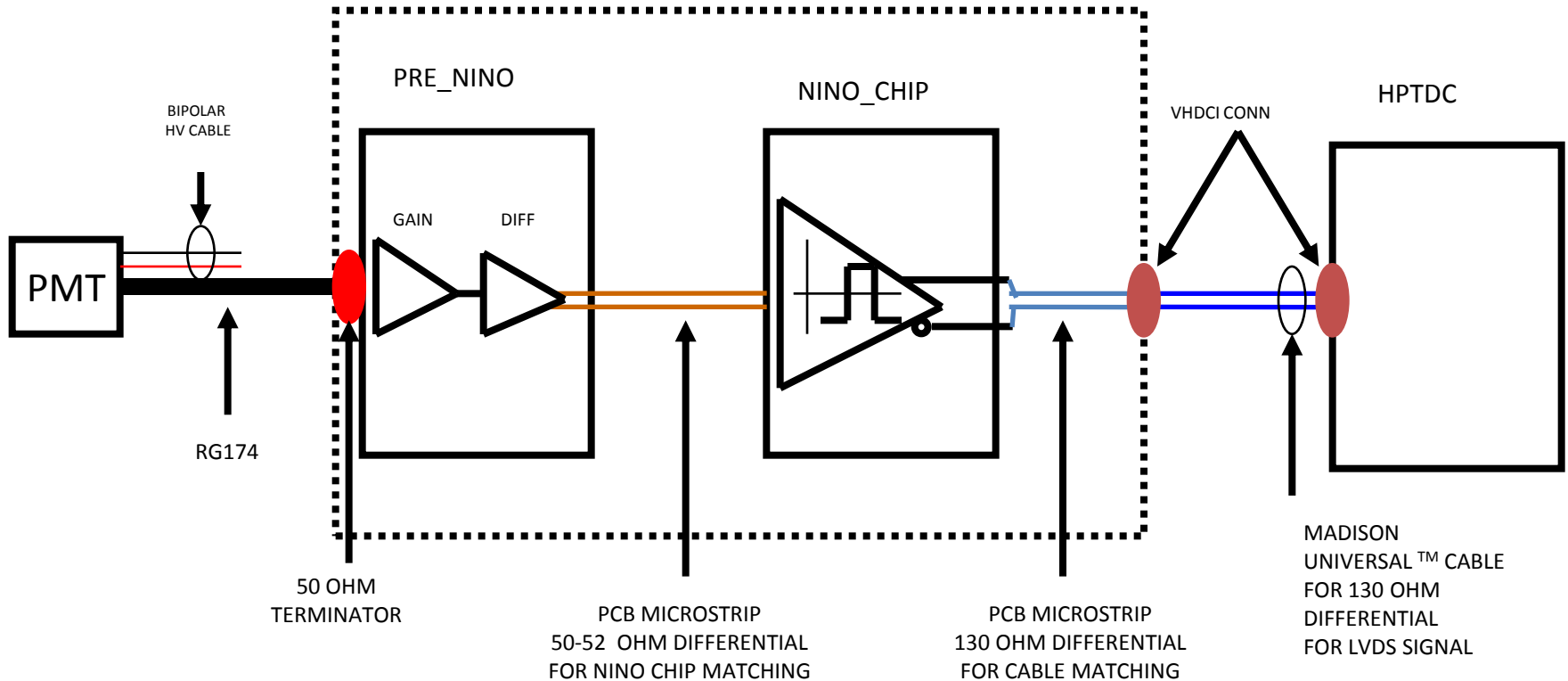


Prototype version: 24 channels
FINAL version: 32 channels

NA-62

SIGNAL CHAIN OVERVIEW

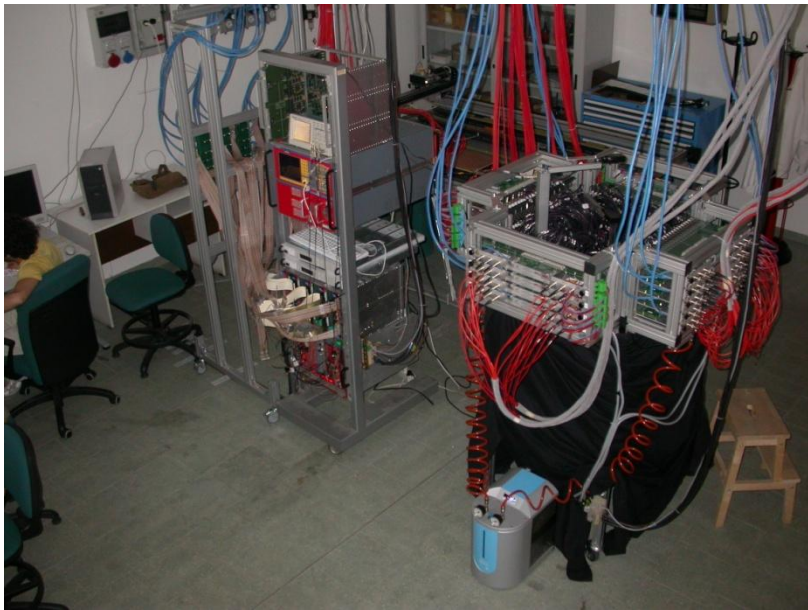
VME RICH FE 32 CH BOARD



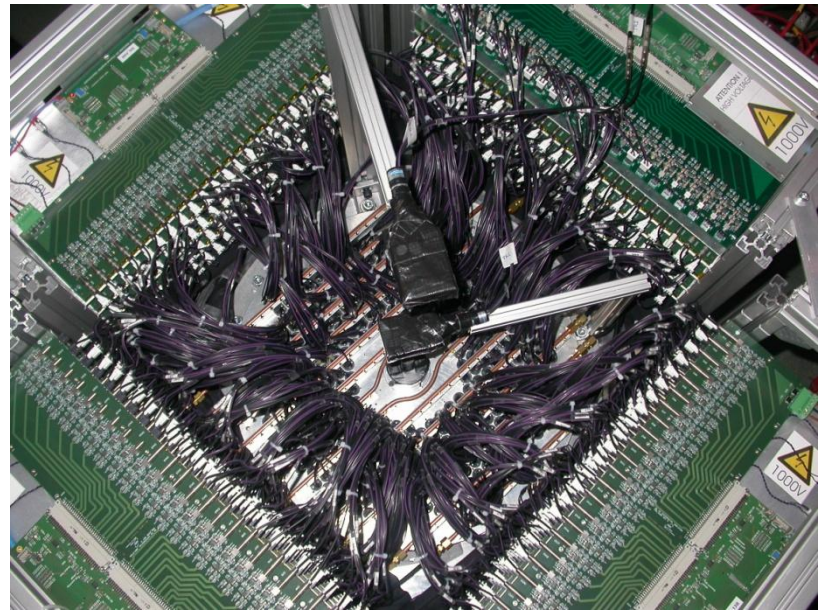
DAQ

- Crate VME
- 1 bridge V2718: controller and optical link to PC
- 3 TDC V1190A (128 ch), LSB:100 ps, 40 MHz clock
- 1 TDC V1290 (32 ch), used with LSB:100 ps, 40 MHz clock
- 2 PC for control (start daq), data transfer, analysis, etc



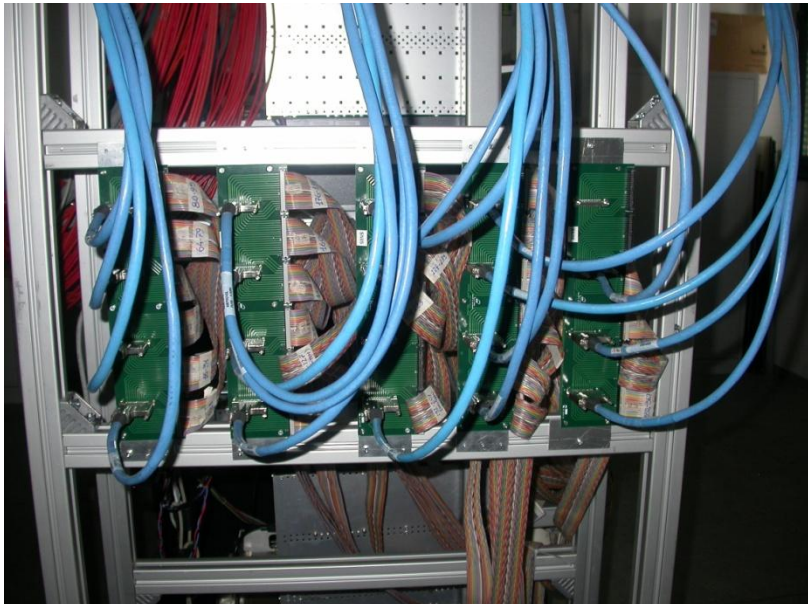


Bird eye view



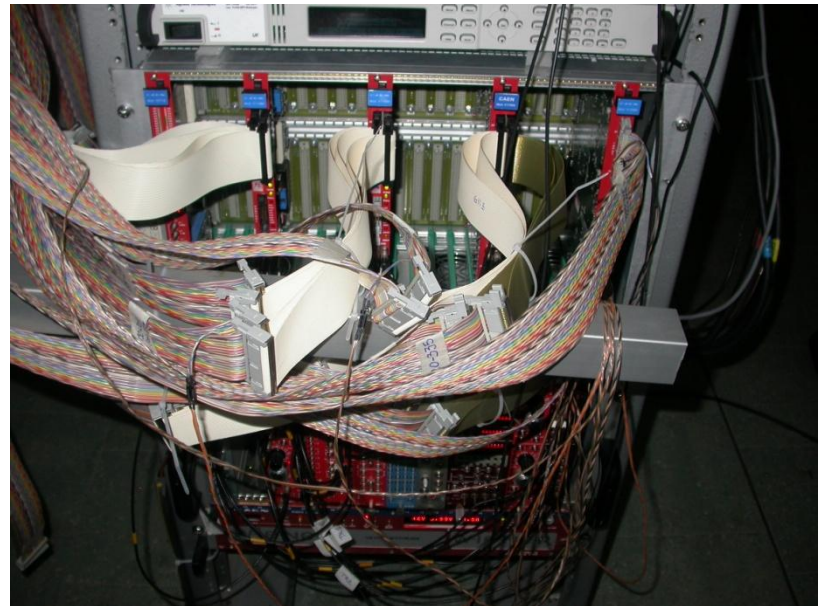
Trigger Scintillators, PMTs, FE

lab



4/06/2012

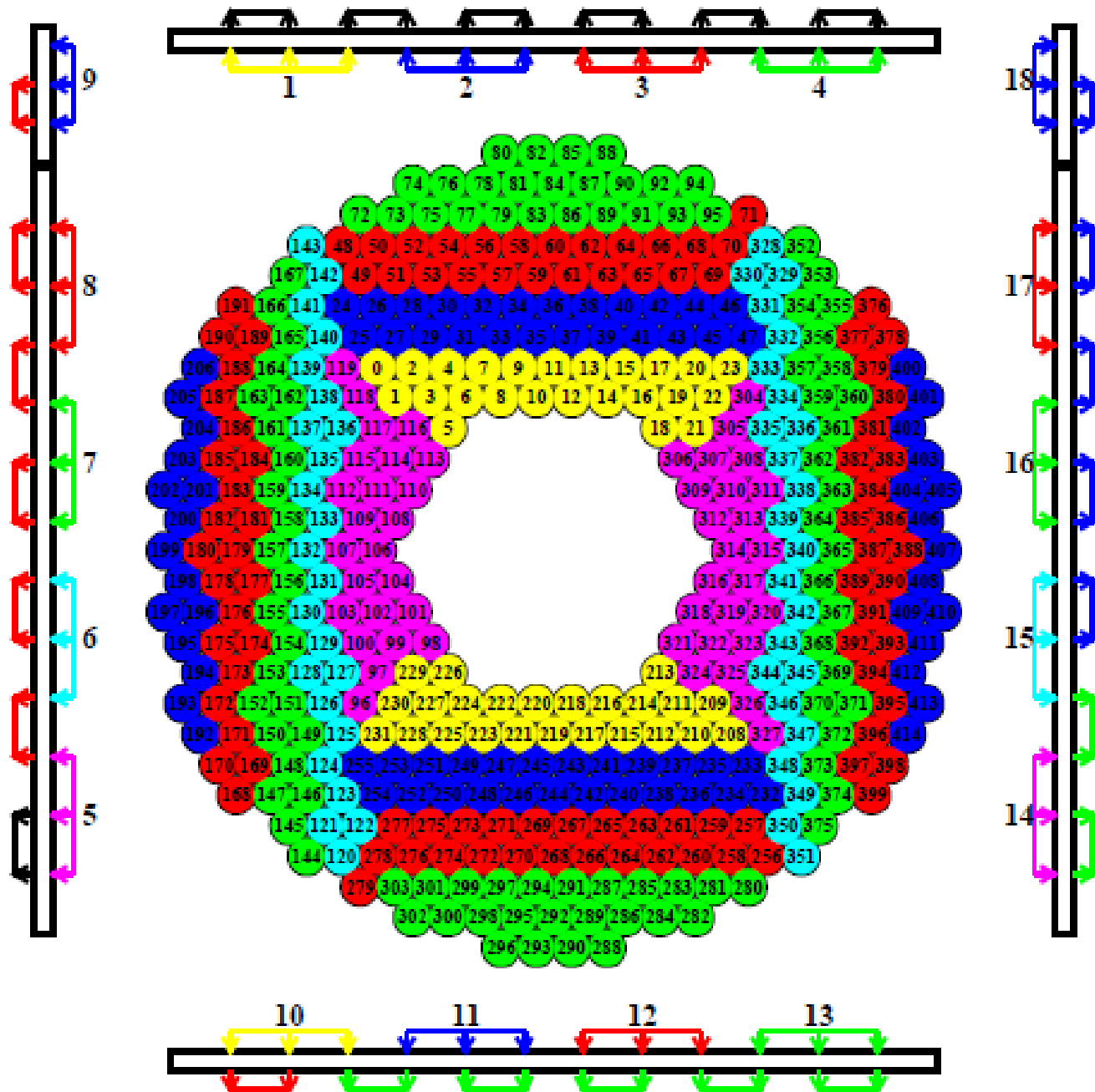
Patch panel



DAQ

M.Lenti

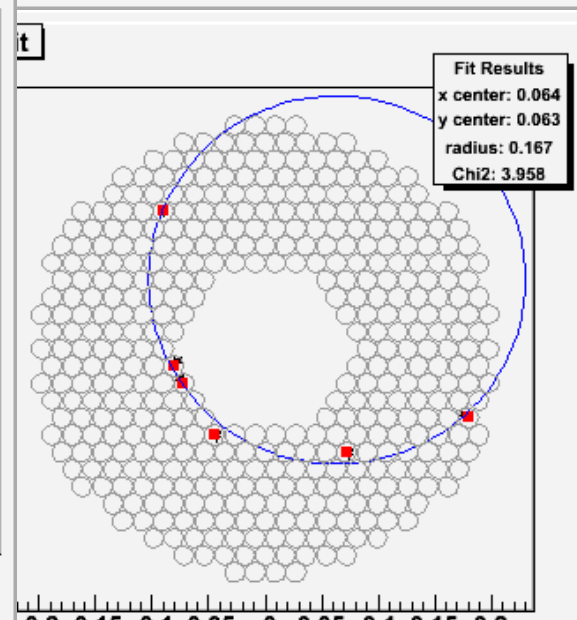
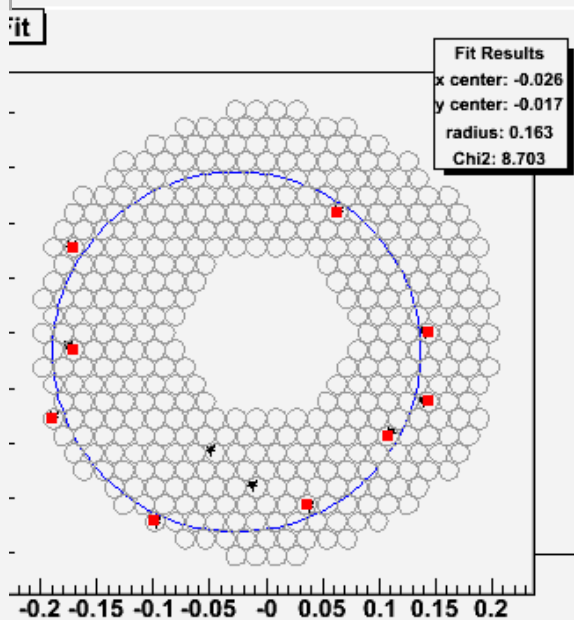
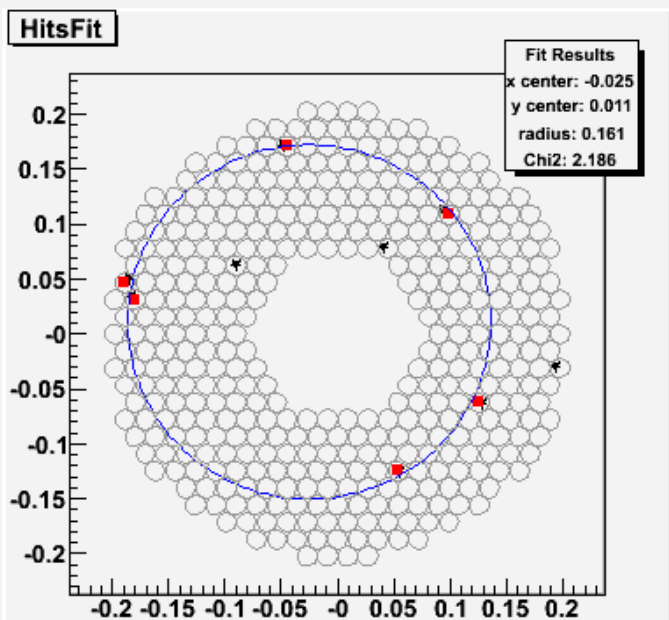
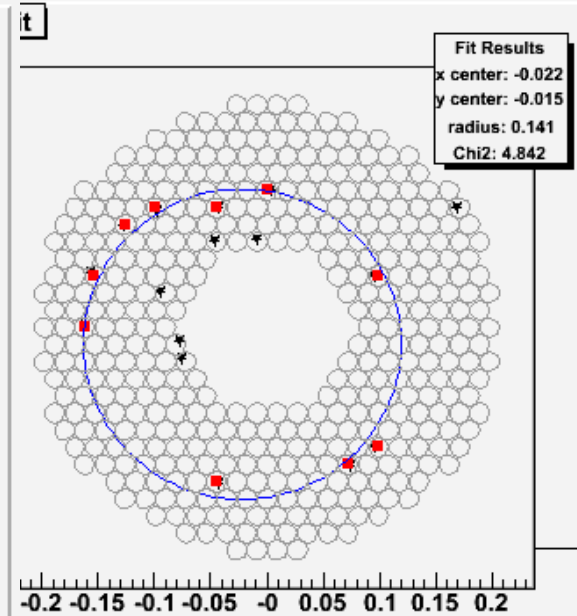
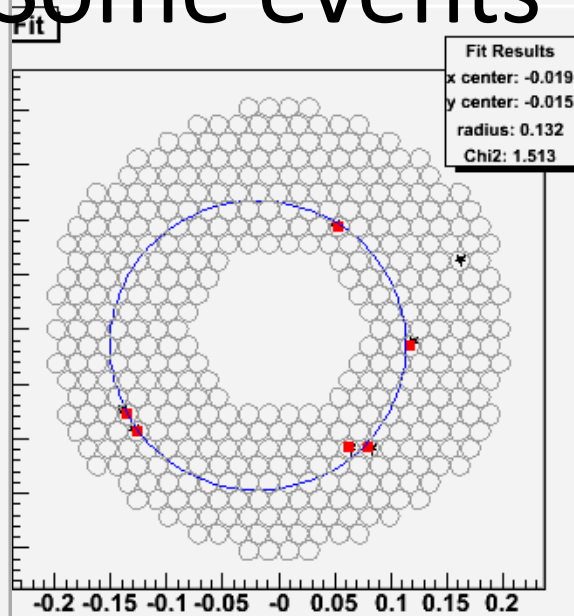
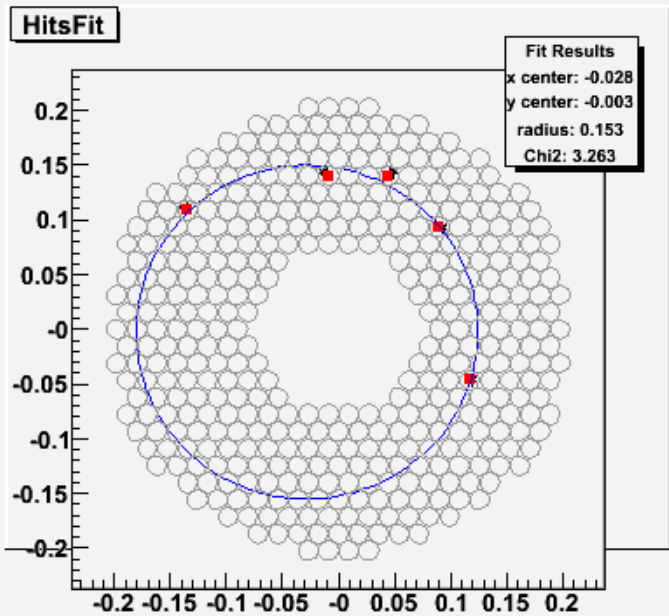
Channels and
FE-boards
map



In the lab....

- Trigger with a couple of scintillator
- TDC operated in Trigger Matching Mode
- TestBeam-like daq: enable for 7 secs (“spill”), disable for 3 secs (“interspill”)
- Each cycle is a “burst”
- Raw data produced....
- Make a true data acquisition
- Make offline data analysis
- Reconstruct cosmic rays induced Cherenkov ring
- Fit the ring (with a circle but it is an ellipse...)

Some events



Conclusions

- COSMORICH ready downstairs
- Enjoy the lab